

# LLP's @ LHC: Taking Stock

M.J. Ramsey-Musolf

*U Mass Amherst*



<http://www.physics.umass.edu/acfi/>

LLP Workshop,  
CERN April 2017

# *LLP's For Newcomers*



# *LLP's For Newcomers*



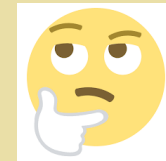
*LLP Dog Race*

# *LLP's: Roadmap for Theory & Expt*

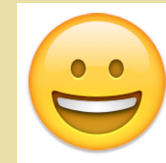


# *Experimental LLP Search: Motivation*

- *Theorists think it's interesting*



- *It's something we can do*



- *It addresses fundamental Q's*



# ***LLP's @ LHC: Motivation***

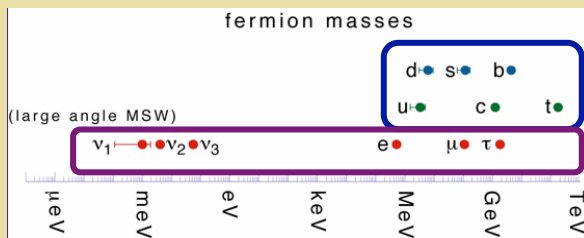
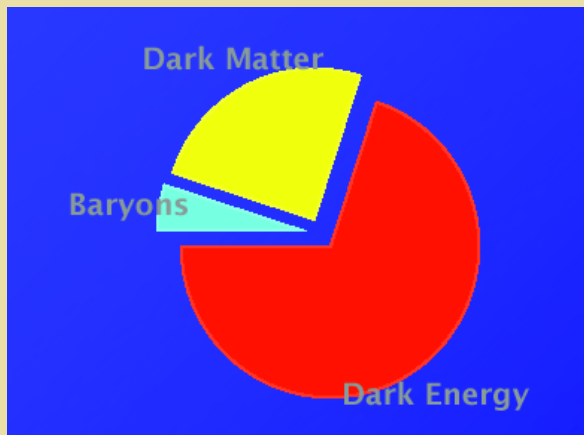
- *Discovery of LLP's may provide clues to key open questions in fundamental physics*
- *Consideration of physical scales  $\rightarrow$  LLP decay lengths  $\sim$  ATLAS, CMS & LHCb detectors*
- *Energy frontier capabilities are unique and complementary to those at Intensity & Cosmic frontiers*

# *LLP's @ LHC: Motivation*

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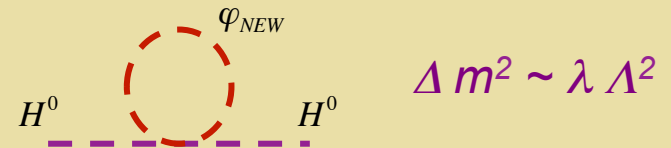
# Fundamental Questions

**MUST** answer



Origin of  $m_\nu$

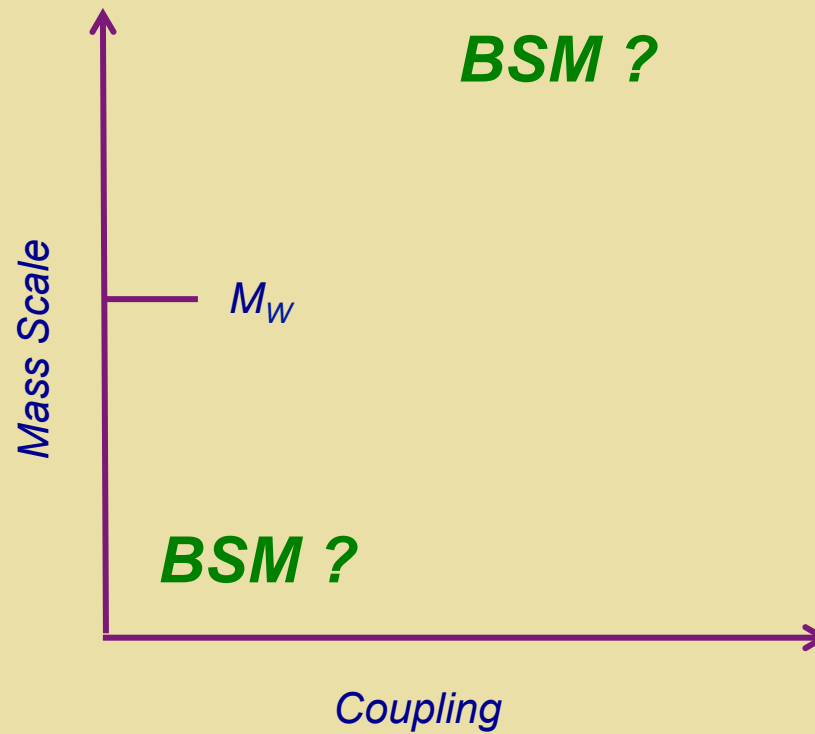
**SHOULD** answer



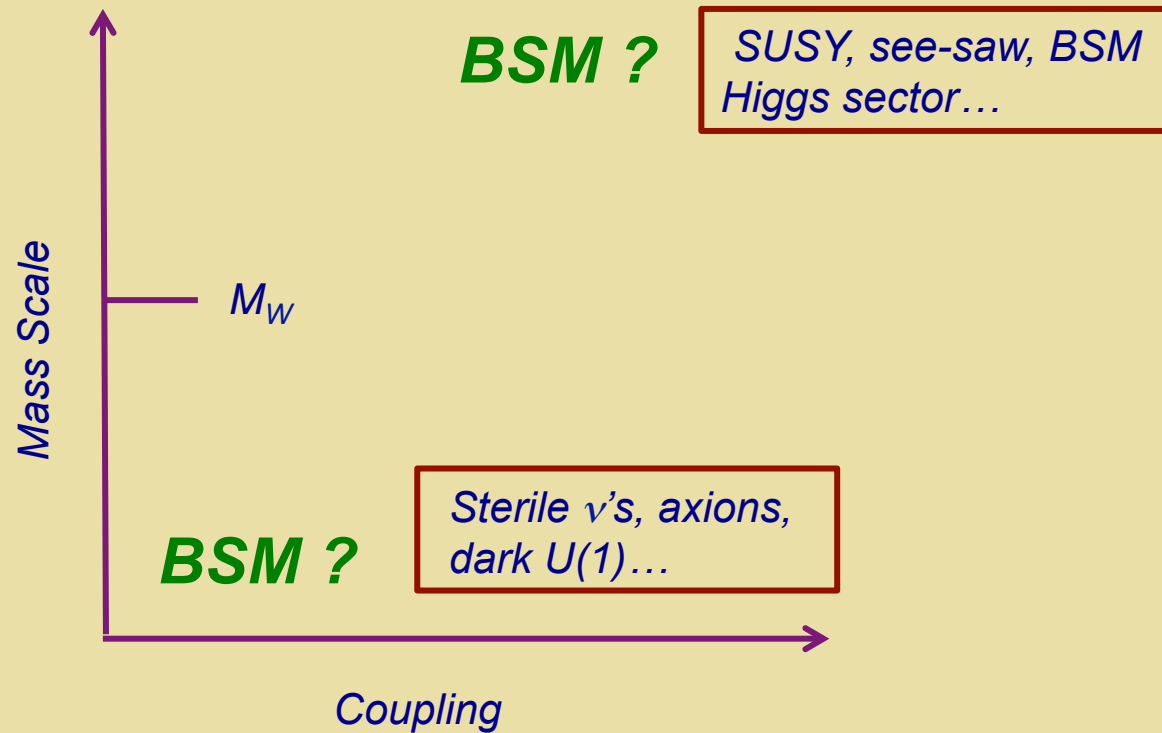
$\Lambda$  Cosmological



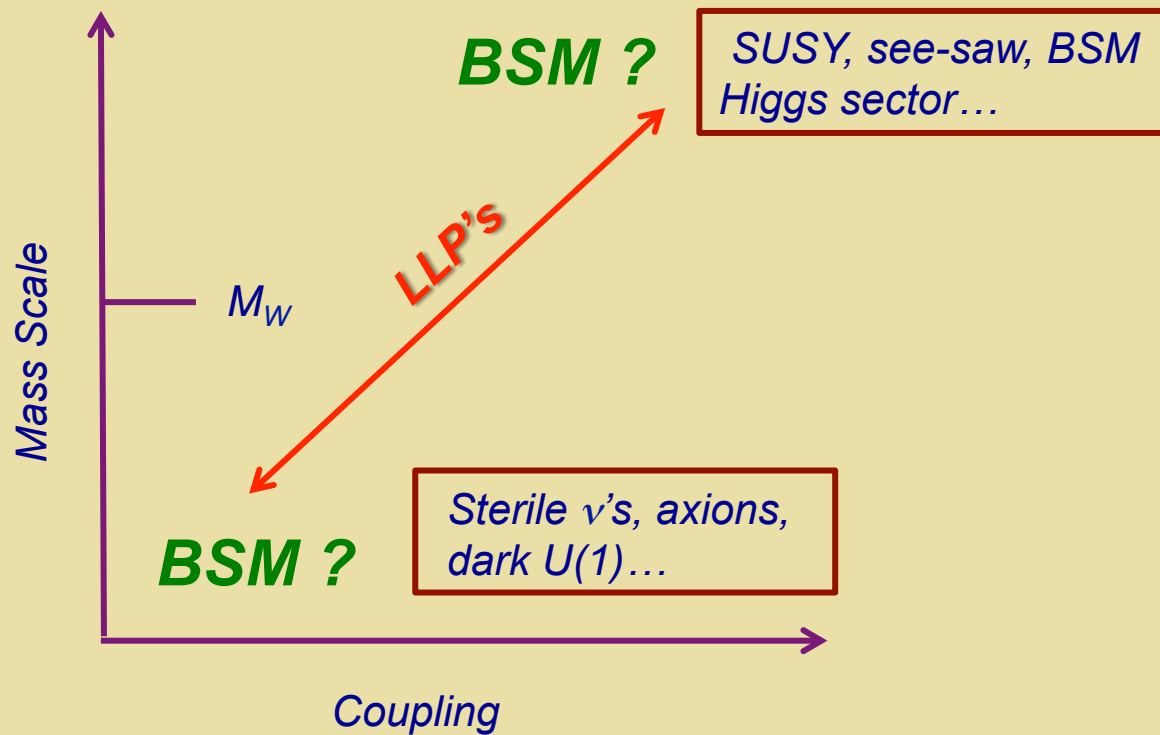
# ***BSM Physics: Where Does it Live ?***



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# BSM Physics: Where Does it Live ?



# ***Solutions w/ LLP's: A Sampler***

*Apologies for omissions !*

## *Solutions w/ LLP's: A Sampler*

<i>LLP Scenario</i>	<i><math>m_H</math></i>	<i>BAU</i>	<i>DM</i>	<i><math>m_\nu</math></i>
<i><math>\nu</math>SM</i>				
<i>WIMPY baryogenesis</i>				
<i>Dark QCD</i>				
<i>Stealth SUSY</i>				
<i>Neutral Naturalness</i>				
<i>Dark U(1)</i>				

## Solutions w/ LLP's: A Sampler

<i>LLP Scenario</i>	$m_H$	<i>BAU</i>	<i>DM</i>	$m_\nu$
<i><math>\nu</math>SM</i>	✗	✓	✓	✓
<i>WIMPY baryogenesis</i>	✗	✓	?	✗
<i>Dark QCD</i>	✗	✓	✓	✗
<i>Stealth SUSY</i>	✓	✓	✓	✗
<i>Neutral Naturalness</i>	✓	✗	✗	✗
<i>Dark U(1)</i>	✗	✗	✓	✗

## Solutions w/ LLP's: A Sampler

LLP Scenario	$m_H$	BAU	DM	$m_\nu$
$\nu$ SM	✗	✓	✓	✓
WIMPY baryogenesis	✗	✓	?	✗
Dark QCD	✗	✓	✓	✗
Stealth SUSY	✓	✓	✓	✗
Neutral Naturalness	✓	✗	✗	✗
Dark U(1)	✗	✗	✓	✗

Hidden Valleys

Strassler, Zurek '06...

# Solutions w/ LLP's: $\nu$ SM

**$\nu$ MSM** P. Mermod

Spin-1/2 fermions

Quarks	Left	u	Right	Left	c	Right	Left	t	Right
	Left	d	Right	Left	s	Right	Left	b	Right
	Left	$\nu_1 N_1$	Right	Left	$\nu_2 N_2$	Right	Left	$\nu_3 N_3$	Right
Leptons	Left	e	Right	Left	$\mu$	Right	Left	$\tau$	Right

Spin-1 bosons

g
$\gamma$
$Z^0$
$W^\pm$

Force carriers

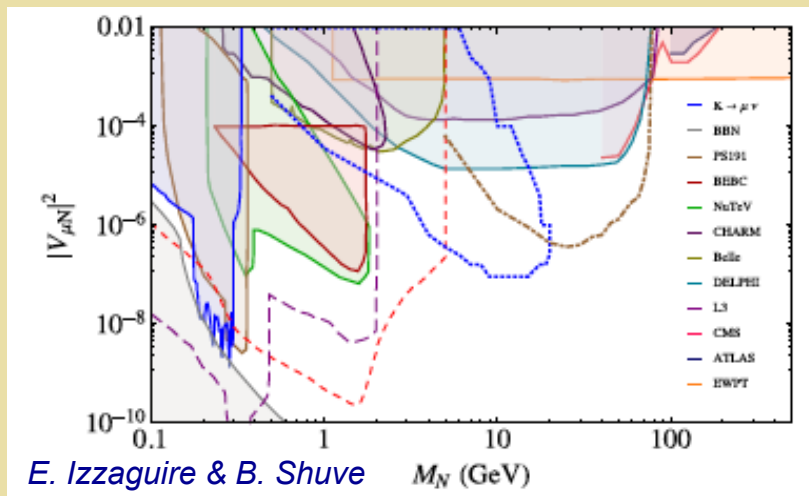
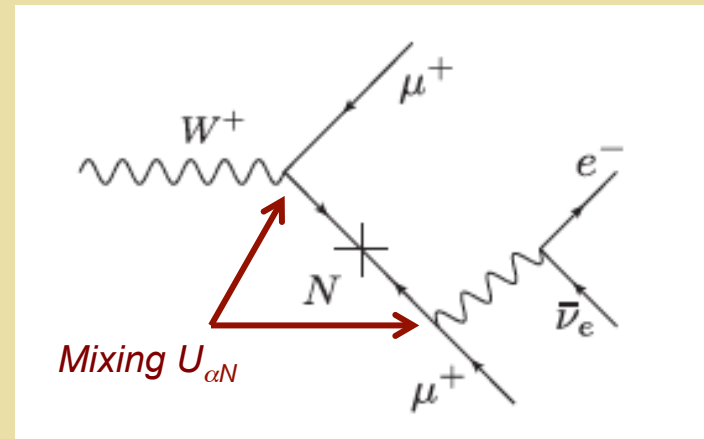
H
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Spin-0 Higgs boson

$N_1$  mass  $\sim$  keV  
→ dark matter

$N_{2,3}$  mass  $\sim$  GeV  
→ seesaw  
→ leptogenesis

Ann. Rev. Nucl. Part. Sci. 59, 191 (2009)





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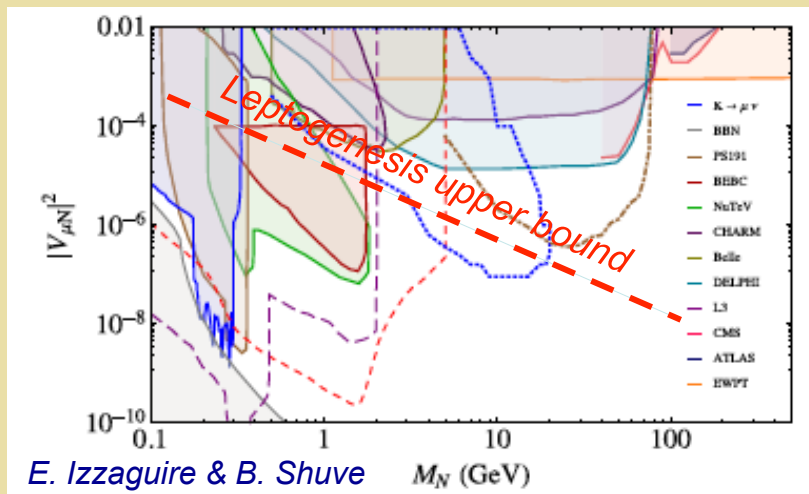
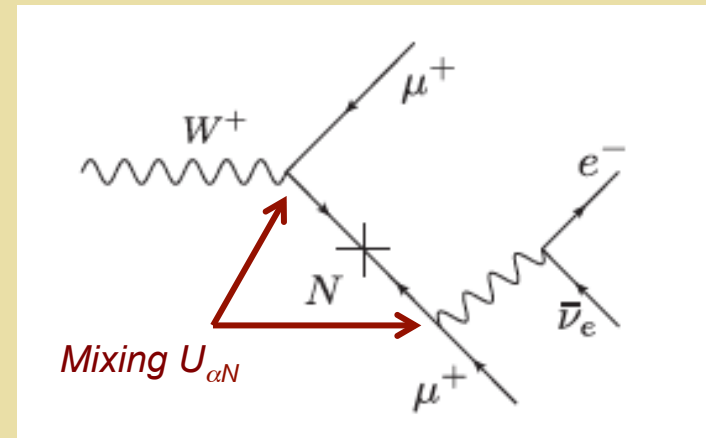
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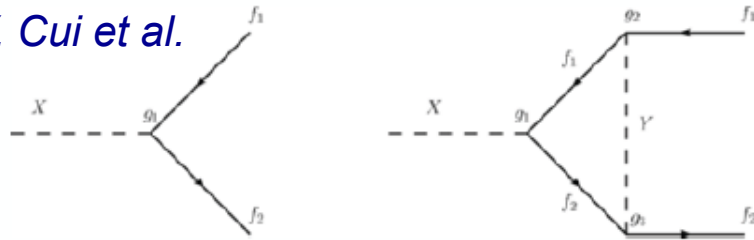


## BAU from Leptogenesis

- Drewes et al '16
- Lower bound  $< 10^{-10}$

# Solutions w/ LLP's: Wimpy Baryogenesis

Y. Cui et al.



$$\Gamma(\chi \rightarrow f) \neq \Gamma(\chi \rightarrow \bar{f})$$

$$n_f - n_{\bar{f}} \neq 0$$

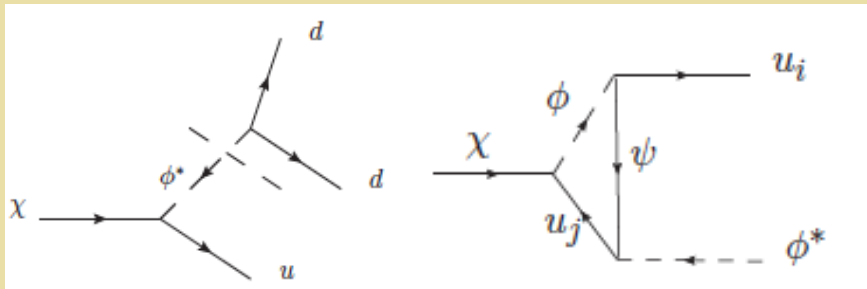
Baryon number violating:

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Lepton number violating:

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$$\chi \rightarrow L_i L_j \bar{E}_k$$



*Like leptogenesis*

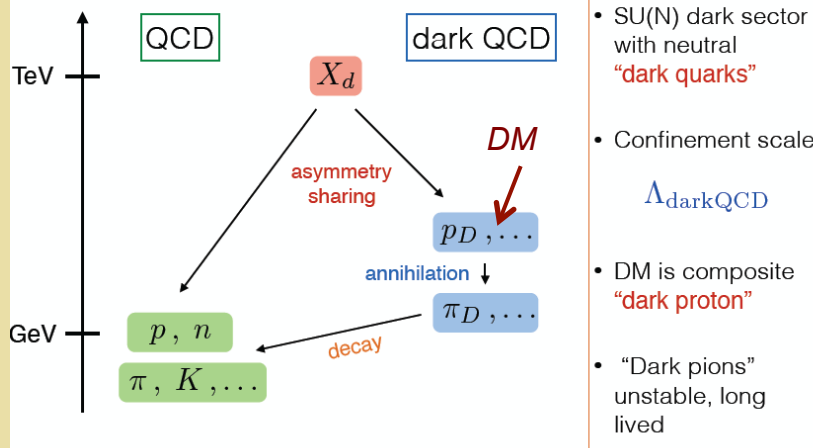
# *Solutions w/ LLP's: Hidden Valleys*



# Solutions w/ LLP's: Dark QCD

P. Schwaller

## Dark QCD



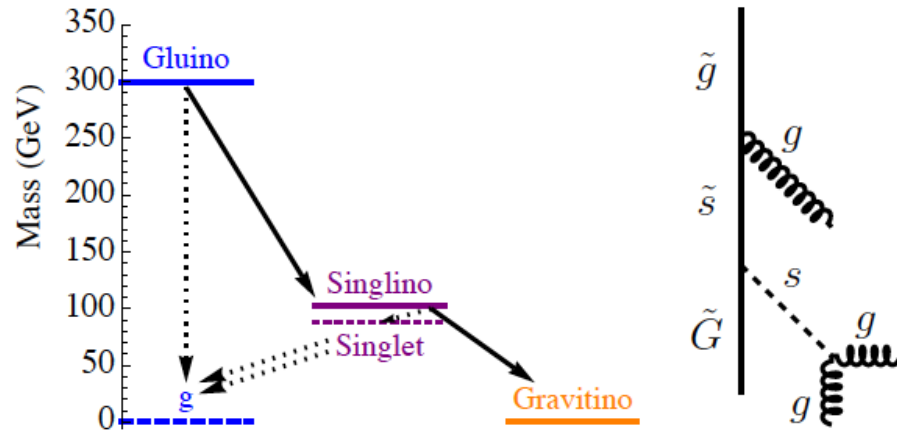
Asymmetric DM:  $m_{DM} n_{DM} \sim 5 m_N n_B \rightarrow$

- For  $n_{DM} \sim n_B \rightarrow m_{DM} \sim \text{few} \times m_B$
- $\Lambda_{\text{dark QCD}} \sim \text{few} \times \Lambda_{\text{QCD}}$



# Solutions w/ LLP's: Stealth SUSY

Fan, Reece, Ruderman



# Solutions w/ LLP's: Neutral Naturalness

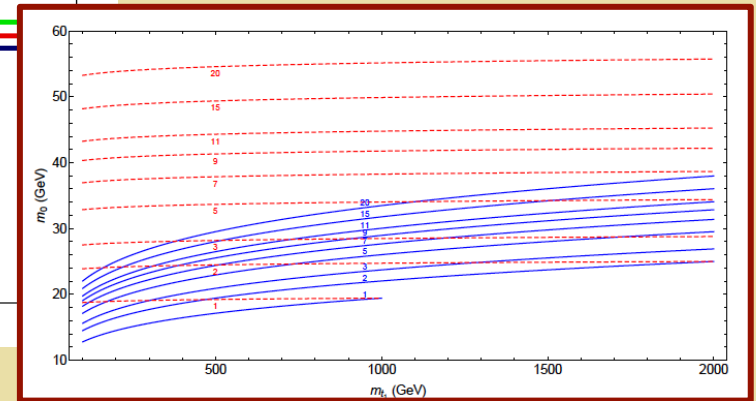
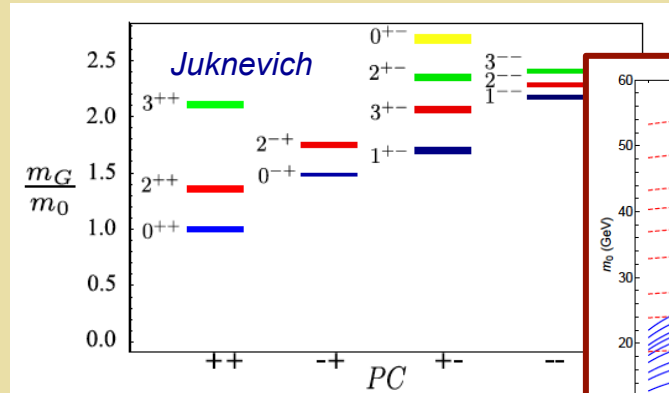
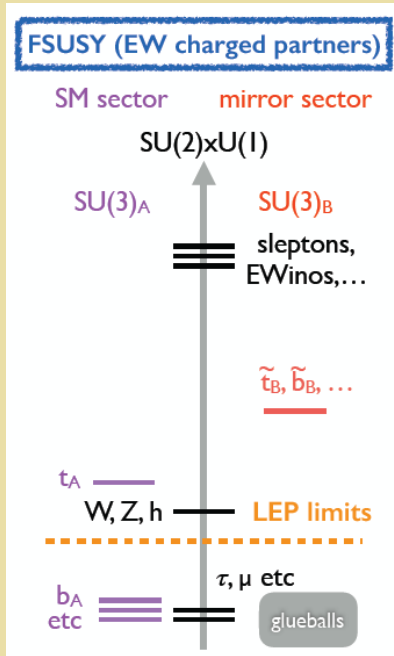
D. Curtin, C. Verhaaren

		<i>scalar</i>	<i>fermion</i>
<i>strong direct production</i> {	<i>QCD</i>	SUSY	Composite Higgs/ RS
<i>DY direct production</i> {	<i>EW</i>	folded SUSY	Quirky Little Higgs
<i>Higgs portal direct production</i> {	<i>singlet</i>	?	Twin Higgs

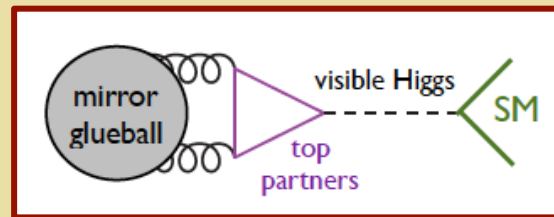
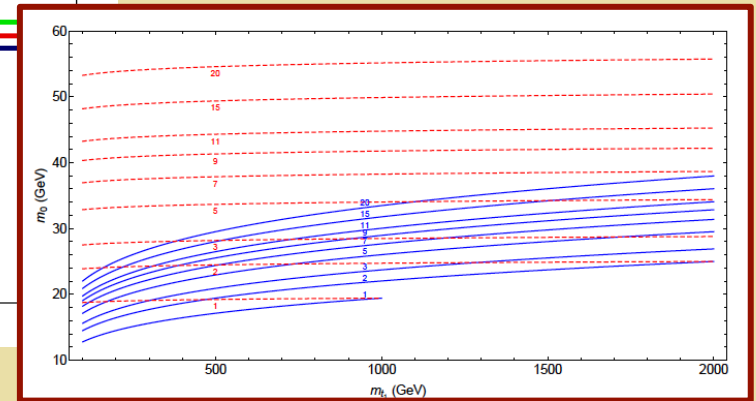
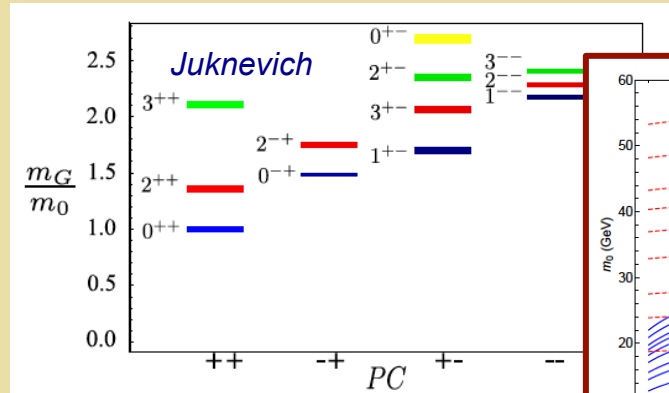
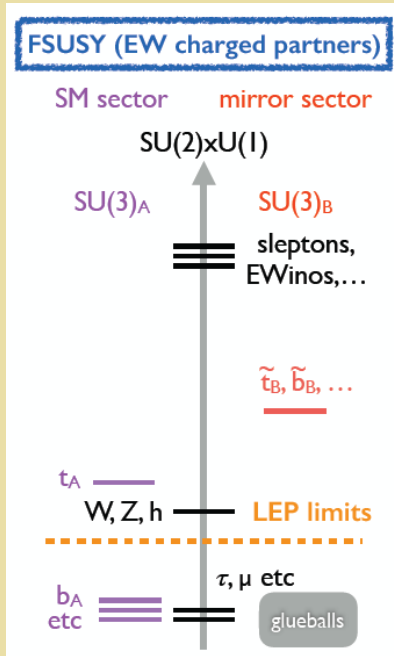
Mirror Glueballs      Higgs coupling shifts  
Higgs portal observables      ~ tuning

**Top partners**

# Solutions w/ LLP's: Neutral Naturalness



# Solutions w/ LLP's: Neutral Naturalness





## ***Solutions w/ LLP's: Dark U(1)***

$$SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_D$$

# Dark Z: Mechanism

$$\mathcal{L} \subset -\frac{1}{4} \hat{B}_{\mu\nu} \hat{B}^{\mu\nu} - \frac{1}{4} \hat{Z}_{D\mu\nu} \hat{Z}_D^{\mu\nu} + \frac{1}{2} \frac{\epsilon}{\cos\theta} \hat{Z}_{D\mu\nu} \hat{B}^{\mu\nu} + \frac{1}{2} m_{D,0}^2 \hat{Z}_D^\mu \hat{Z}_{D\mu}$$

$$V_0(H, S) = -\mu^2 |H|^2 + \lambda |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

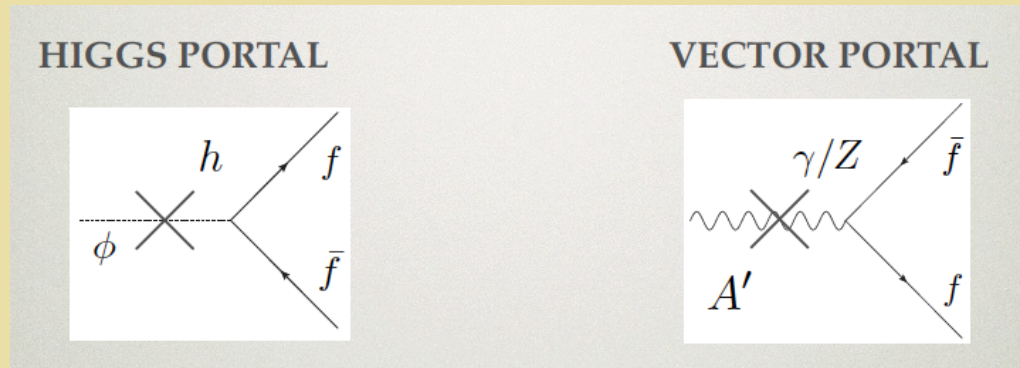
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*Abelian Kinetic Mixing*      *Mass Mixing*

$$V_0(H, S) = -\mu^2 |H|^2 + \lambda |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \boxed{\kappa |S|^2 |H|^2}$$

*Higgs Mixing*

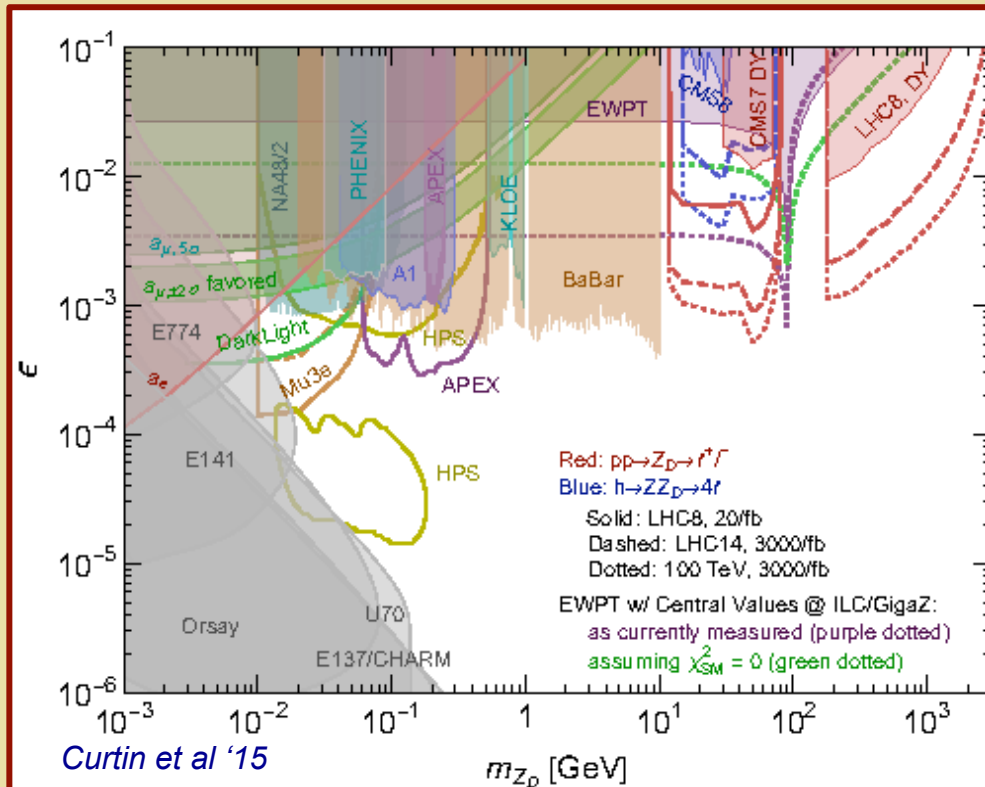


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Abelian Kinetic Mixing

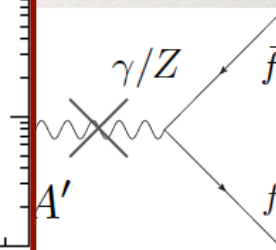
Mass Mixing



$$+ \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

Higgs Mixing

CTOR PORTAL



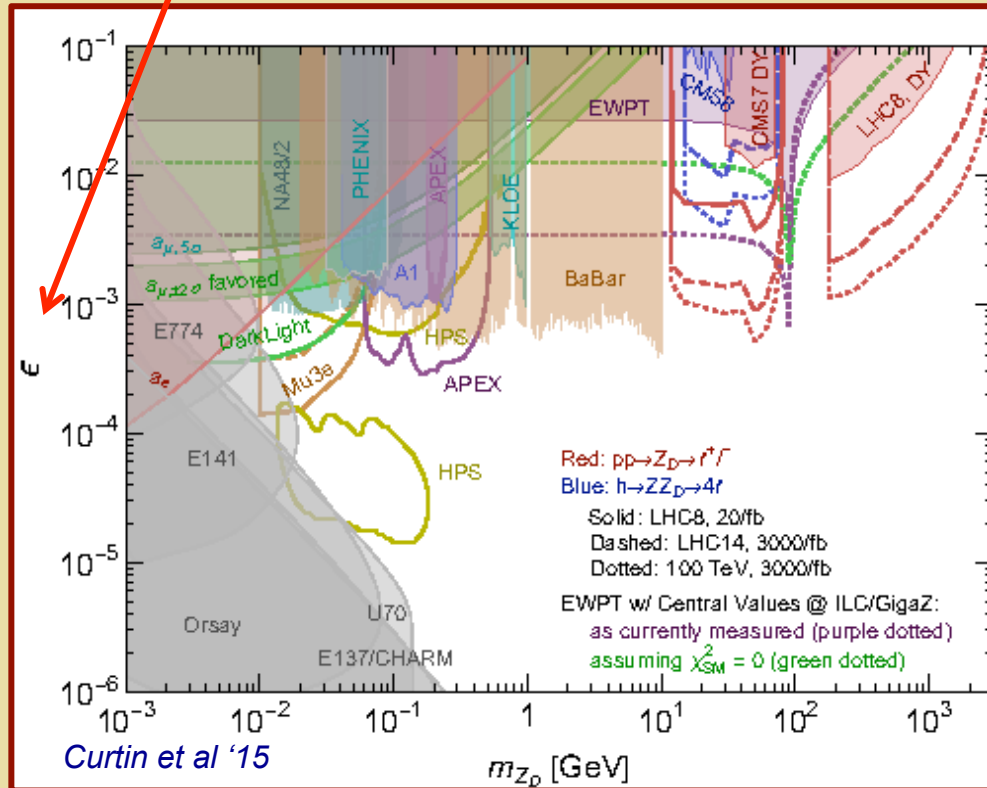
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Why so tiny?

Abelian Kinetic Mixing

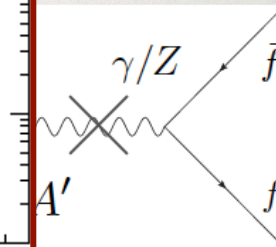
Mass Mixing



$$+ \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

Higgs Mixing

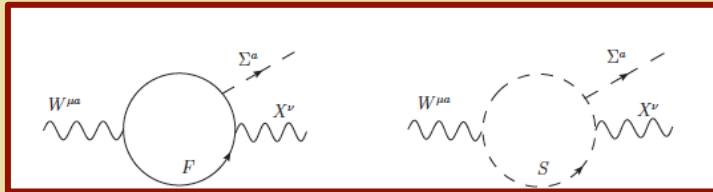
CTOR PORTAL



# Dark Z: Mechanism

*Non-Abelian Kinetic Mixing*

$$\mathcal{O}_{WX}^{(5)} = -\frac{\beta}{\Lambda} \text{Tr} (W_{\mu\nu} \Sigma) X^{\mu\nu}$$



*SU(2)<sub>L</sub> x U(1)<sub>D</sub> mediators*

$$\epsilon = \beta \sin \theta_W \left( \frac{v\Sigma}{\Lambda} \right)$$

*Small  $\epsilon$  from scale ratio;*

$$\beta \sim O(1)$$

*Arguelles, He, Ovaneysan, Peng, MRM '16  
See also Barello, Chang, Newby '15*

# *LLP's @ LHC: Motivation*

- *Discovery of LLP's may provide clues to key open questions in fundamental physics*
- *Consideration of physical scales  $\rightarrow$  LLP decay lengths  $\sim$  ATLAS, CMS & LHCb detectors*
- *Energy frontier capabilities are unique and complementary to those at Intensity & Cosmic frontiers*

# Solutions w/ LLP's: $\nu$ SM

**$\nu$ MSM** P. Mermod

Spin-1/2 fermions

Quarks	Left	u	Right	c	Right	t	Right
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Spin-1 bosons

g
$\gamma$
$Z^0$
$W^\pm$

Force carriers

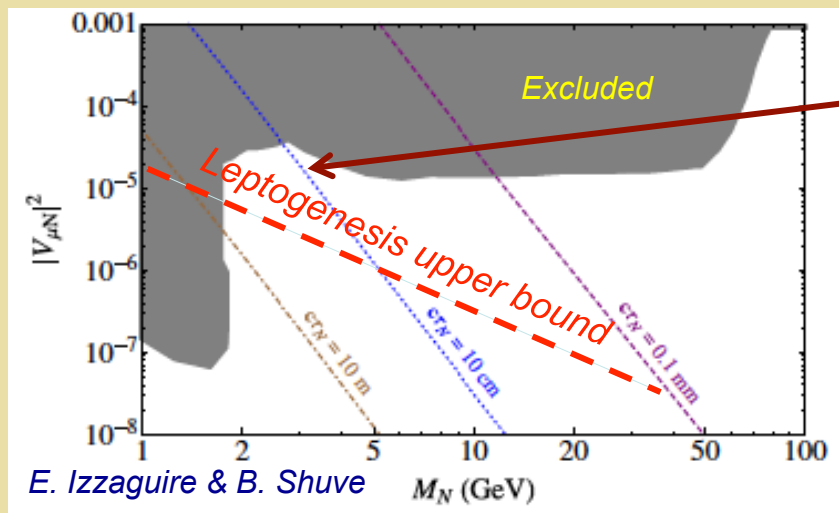
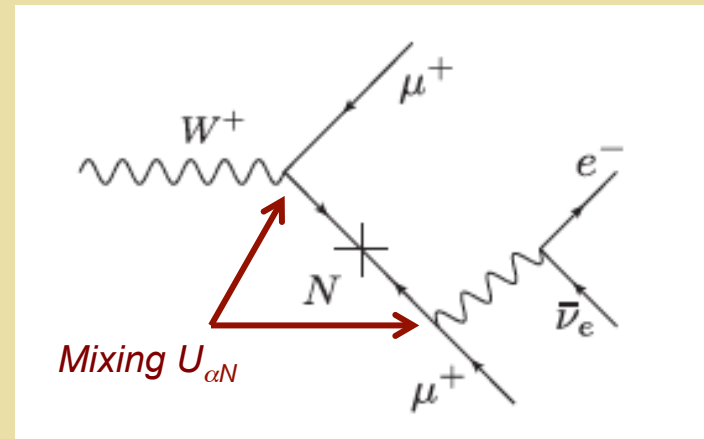
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Spin-0 Higgs boson

$N_1$  mass  $\sim$  keV  
 $\rightarrow$  dark matter

$N_{2,3}$  mass  $\sim$  GeV  
 $\rightarrow$  seesaw  
 $\rightarrow$  leptogenesis

Ann. Rev. Nucl. Part. Sci. 59, 191 (2009)



$$\Gamma(N \rightarrow \ell_a^- \ell_b^+ \nu_\beta) = \frac{G_F^2 M_N^5 |V_{\alpha N}|^2}{192\pi^3}$$

See also: Helo, Kovalenko & Hirsch



# Solutions w/ LLP's: Wimpy Baryogenesis

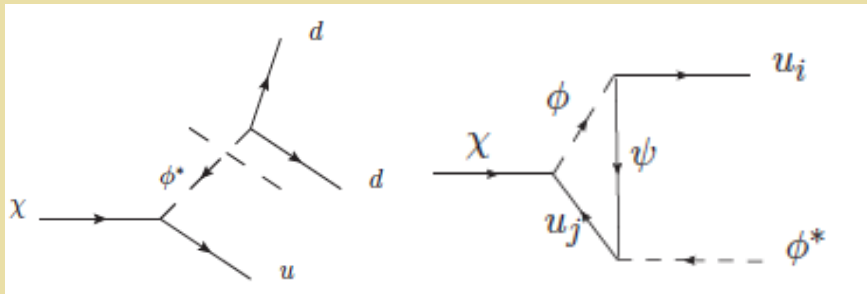
Baryon number violating:

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Lepton number violating:

$$\begin{aligned} \chi &\rightarrow L_i Q_j \bar{d}_k \\ \chi &\rightarrow L_i L_j \bar{E}_k \end{aligned}$$

3-body phase space



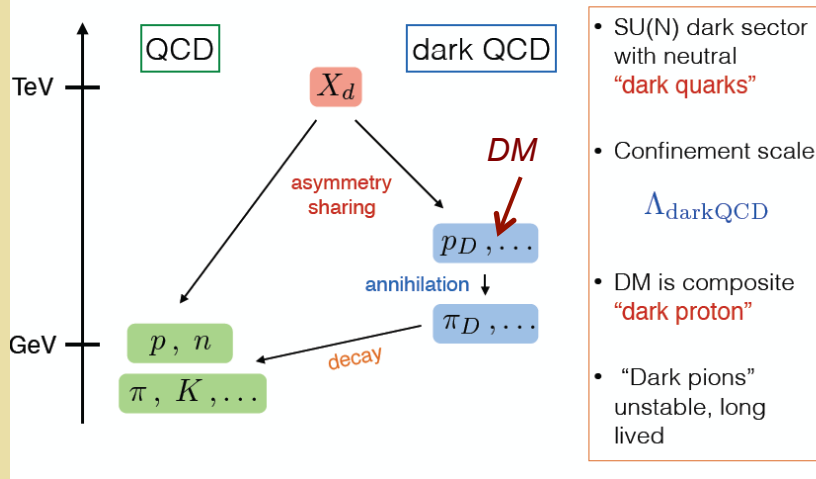
Sakharov: out-of-eq condition

$$\Gamma_\chi < H(T = M_\chi) \longleftrightarrow c\tau_\chi \gtrsim \text{mm}$$

# Solutions w/ LLP's: Dark QCD

P. Schwaller

## Dark QCD



Asymmetric DM:  $m_{DM} n_{DM} \sim 5 m_N n_B \rightarrow$

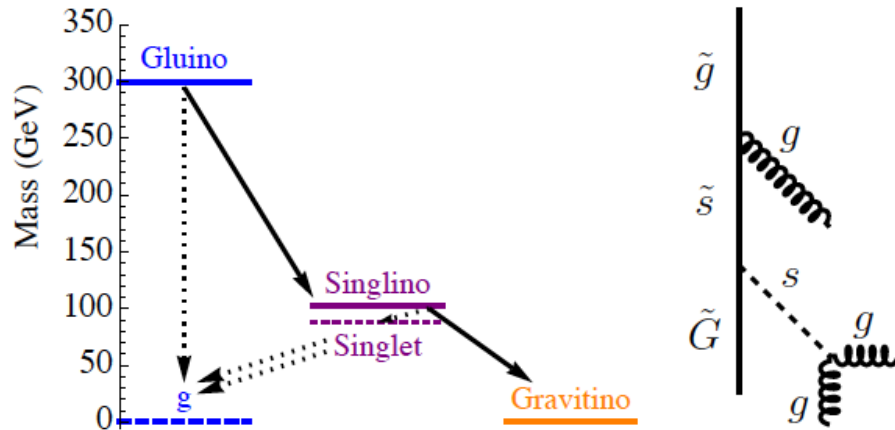
- For  $n_{DM} \sim n_B \rightarrow m_{DM} \sim \text{few} \times m_B$
- $\Lambda_{\text{dark QCD}} \sim \text{few} \times \Lambda_{\text{QCD}}$

$$c\tau(\pi_D \rightarrow \text{SM}) \sim \frac{M_X^4}{m_{\pi_D}^5} \sim \text{cm} \times \left(\frac{M_X}{\text{TeV}}\right)^4 \left(\frac{\text{GeV}}{m_{\pi_D}}\right)^5$$



# Solutions w/ LLP's: Stealth SUSY

Fan, Reece, Ruderman



$$\Gamma_{\tilde{X}} = \frac{m_{\tilde{X}}^5}{16\pi F^2} \left(1 - \frac{m_{\tilde{X}}^2}{m_{\tilde{X}}^2}\right)^4 \approx \frac{m_{\tilde{X}} (\delta m)^4}{\pi F^2}$$

$$m_{\tilde{X}} = 100 \text{ GeV}$$

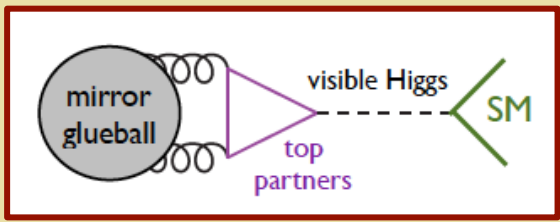
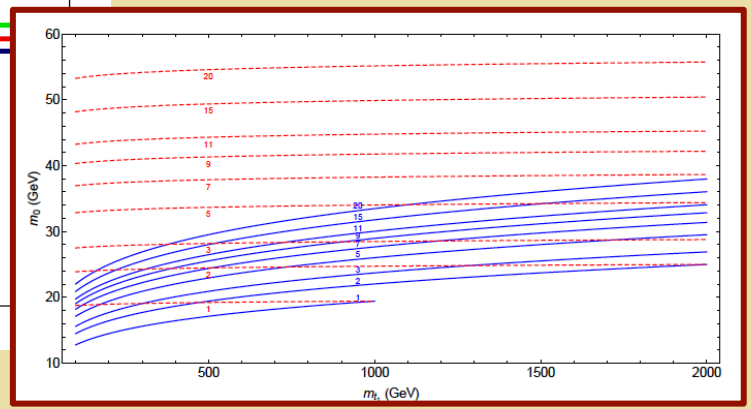
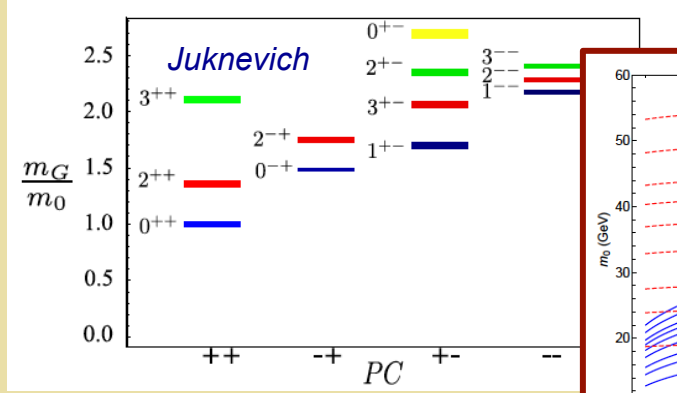
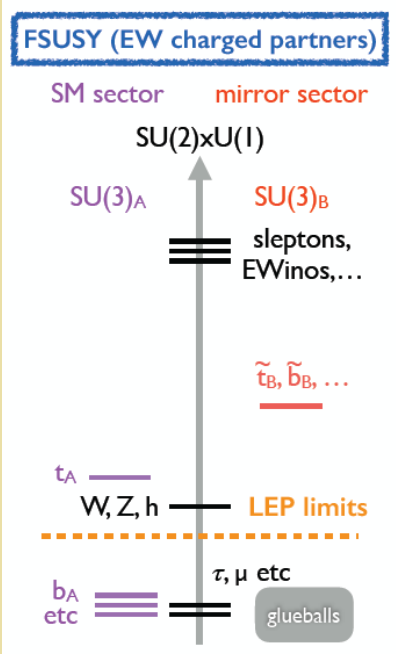
$$\delta m \sim 10 \text{ GeV}$$

$$F \sim (100 \text{ TeV})^2$$

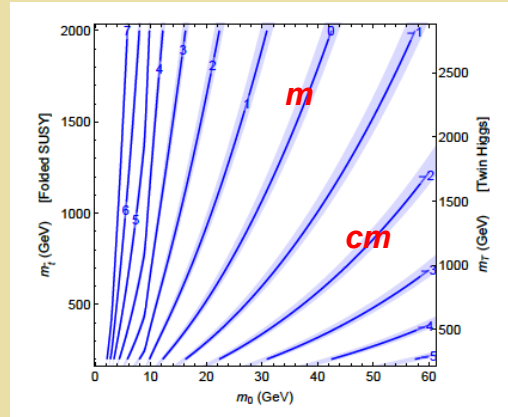
$$c \tau \sim 10 \text{ cm}$$



# Solutions w/ LLP's: Neutral Naturalness



$$\Gamma_{0^{++}} \sim m_0^7 / (M^4 m_h^2)$$



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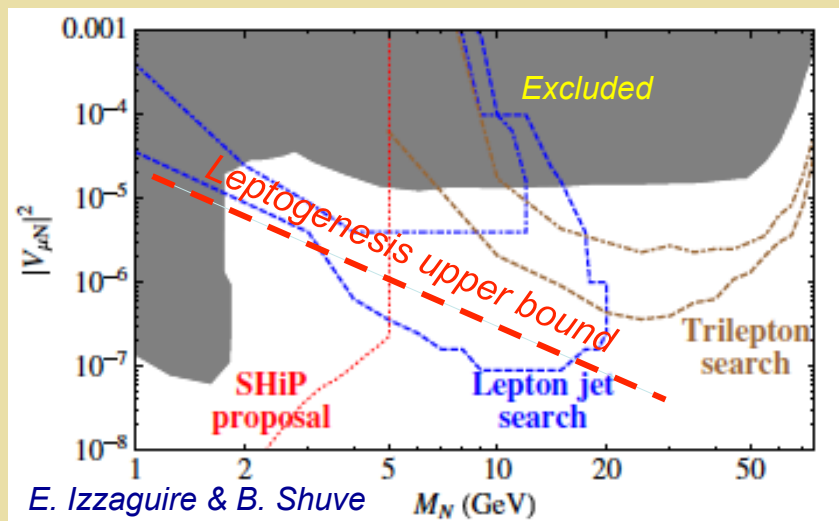
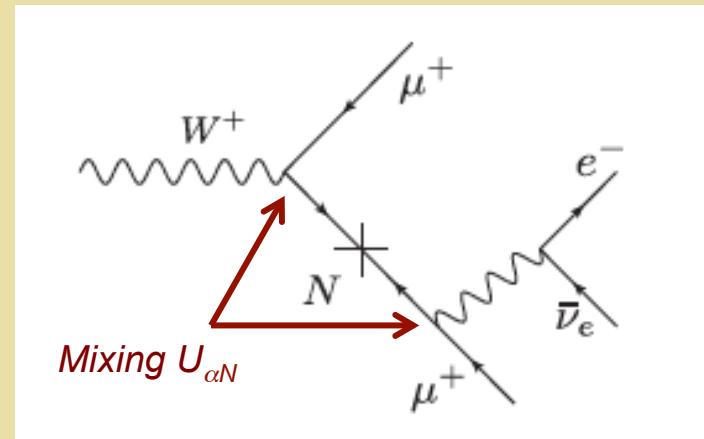
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$$\Gamma(N \rightarrow \ell_a^- \ell_b^+ \nu_\beta) = \frac{G_F^2 M_N^5 |V_{\alpha N}|^2}{192\pi^3}$$

- Displaced LJ +  $\mu$
- 3 resolved prompt leptons

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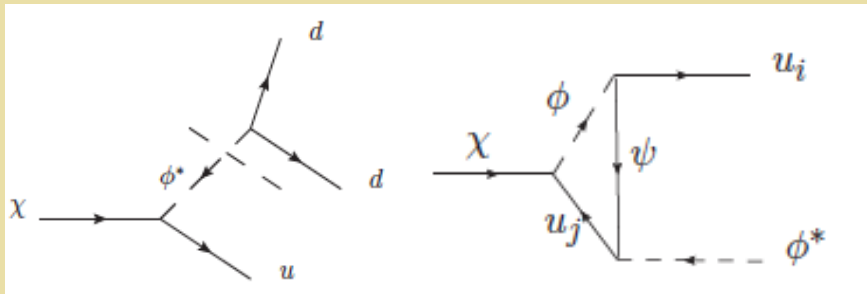
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3-body phase space



- *BNV: displaced jets*
- *LNV: displaced  $\mu$  + tracks*

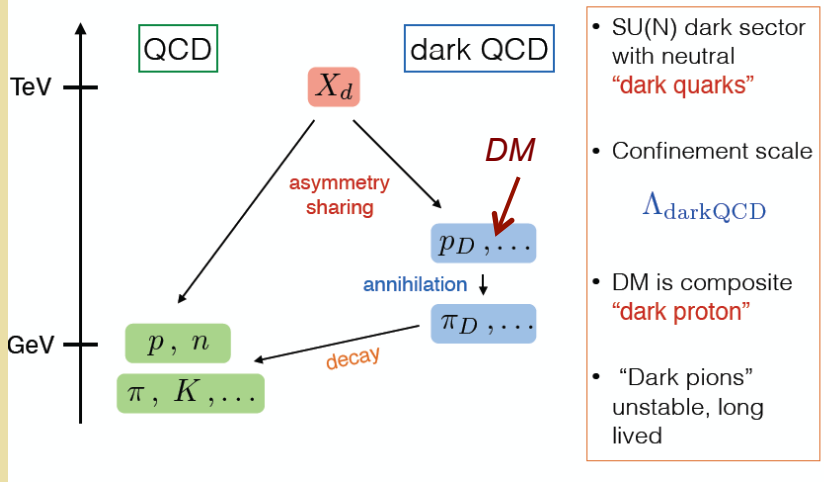
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P. Schwaller

## Dark QCD



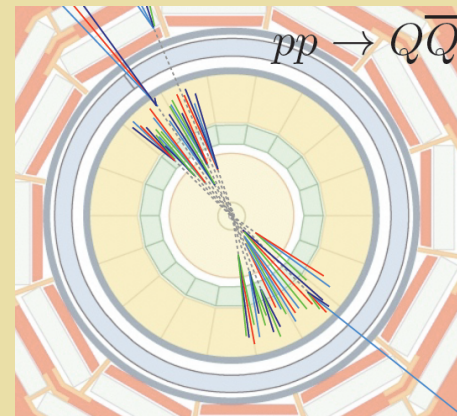
Asymmetric DM:  $m_{DM} n_{DM} \sim 5 m_N n_B \rightarrow$

- For  $n_{DM} \sim n_B \rightarrow m_{DM} \sim \text{few} \times m_B$
- $\Lambda_{\text{dark QCD}} \sim \text{few} \times \Lambda_{\text{QCD}}$

$$c\tau(\pi_D \rightarrow \text{SM}) \sim \frac{M_X^4}{m_{\pi_D}^5} \sim \text{cm} \times \left(\frac{M_X}{\text{TeV}}\right)^4 \left(\frac{\text{GeV}}{m_{\pi_D}}\right)^5$$



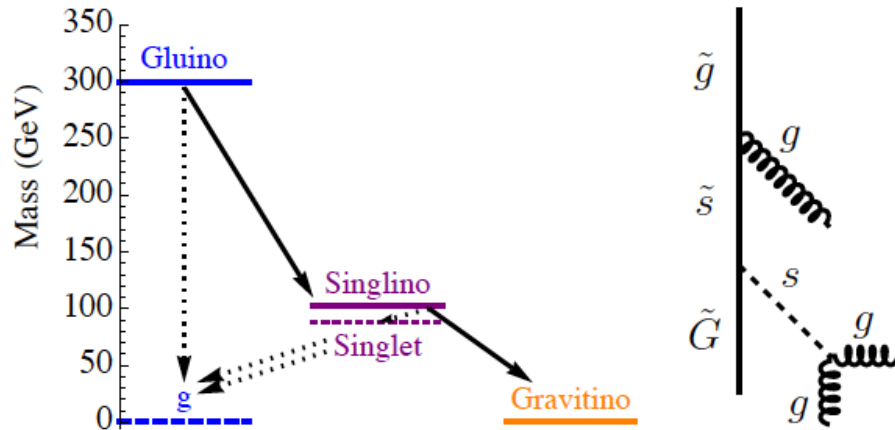
## Emerging Jets





# Solutions w/ LLP's: Stealth SUSY

Fan, Reece, Ruderman



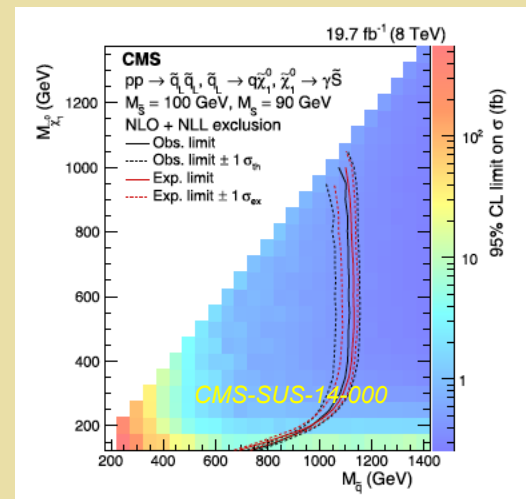
$$\Gamma_{\tilde{X}} = \frac{m_{\tilde{X}}^5}{16\pi F^2} \left(1 - \frac{m_{\tilde{X}}^2}{m_{\tilde{X}}^2}\right)^4 \approx \frac{m_{\tilde{X}} (\delta m)^4}{\pi F^2}$$

$$m_{\tilde{X}} = 100 \text{ GeV}$$

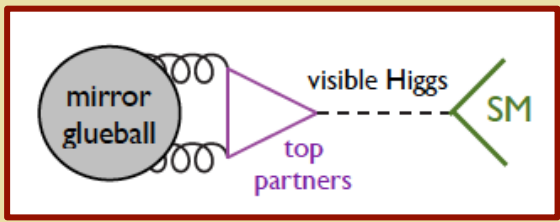
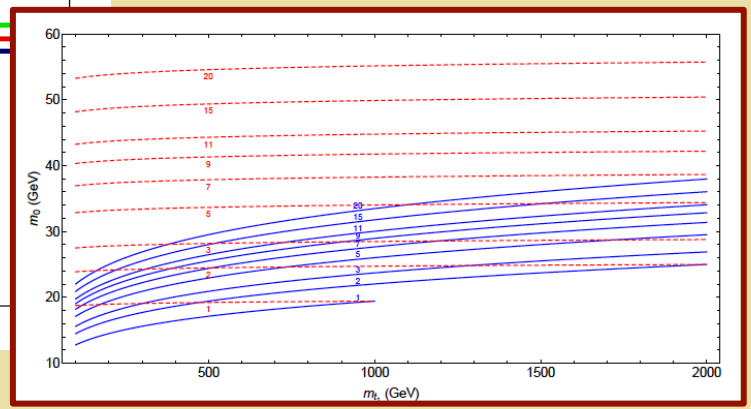
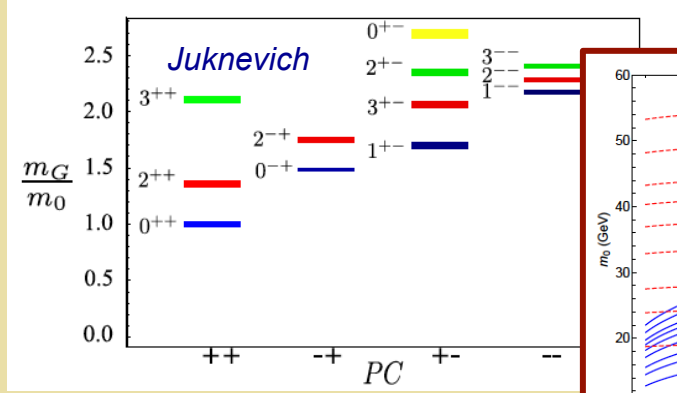
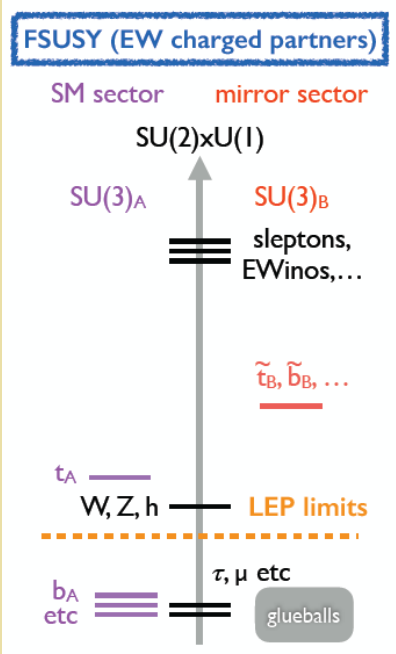
$$\delta m \sim 10 \text{ GeV} \quad c\tau \sim 10 \text{ cm}$$

$$F \sim (100 \text{ TeV})^2$$

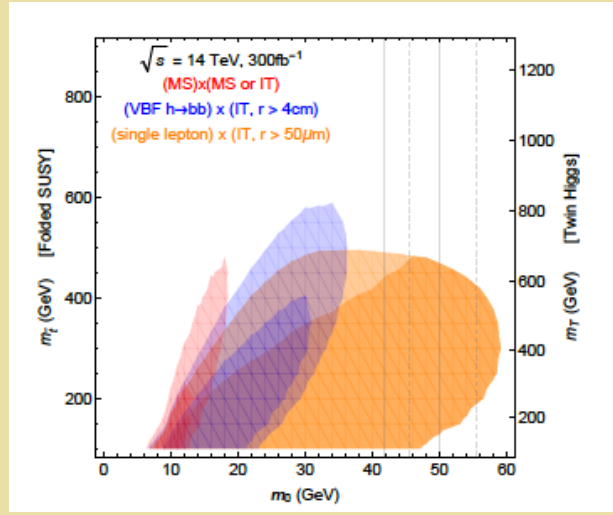
- Prompt  $V$  + displaced  $jj$  ("false resonances")
- DV's + high multiplicity  $b$ -jets
- ...



# Solutions w/ LLP's: Neutral Naturalness



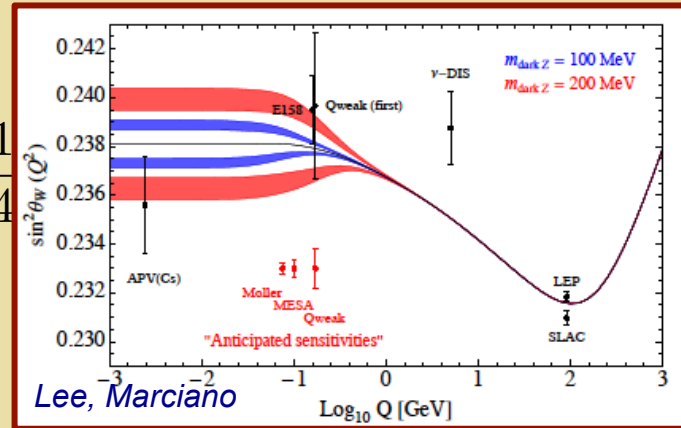
Exotic Higgs decays:  $h \rightarrow 0^{++} 0^{++}$   
w/ 2 DV's or 1 DV + ...



D. Curtin,  
C. Verhaaren

# Dark Z: Mechanism

$$\mathcal{L} \subset -\frac{1}{4} \hat{B}_{\mu\nu} \hat{B}^{\mu\nu} - \frac{1}{4}$$



PVES

$$\frac{1}{2} m_{D,0}^2 \hat{Z}_D^\mu \hat{Z}_{D\mu}$$

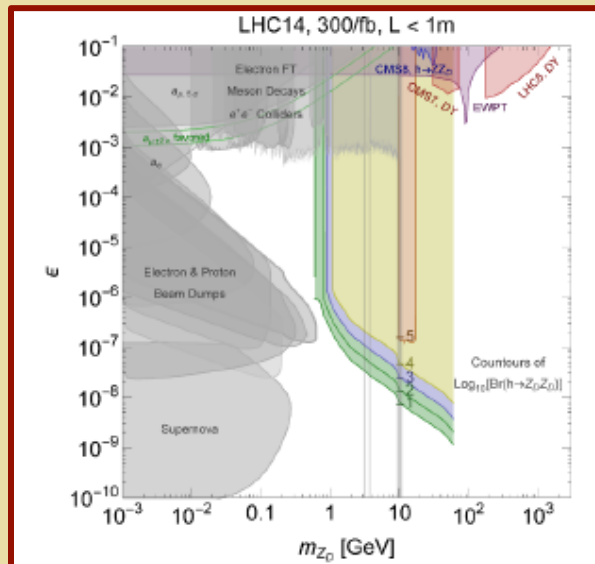
Mass Mixing

$$V_0(H, S) = -\mu^2 |H|^2 + \lambda |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

$$\kappa |S|^2 |H|^2$$

Higgs Mixing

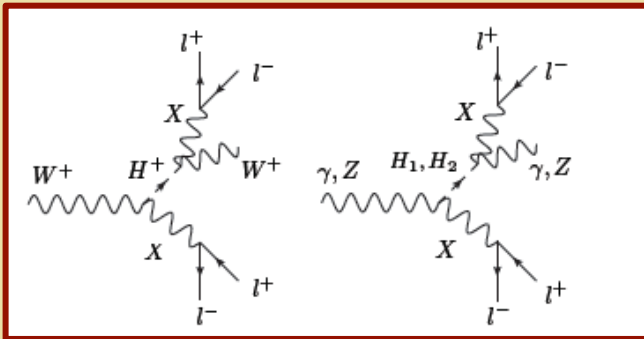
$h \rightarrow Z_D Z_D$



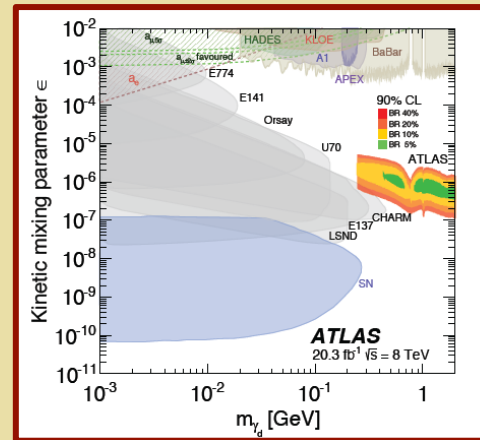
# Dark Z: Mechanism

Non-Abelian Kinetic Mixing

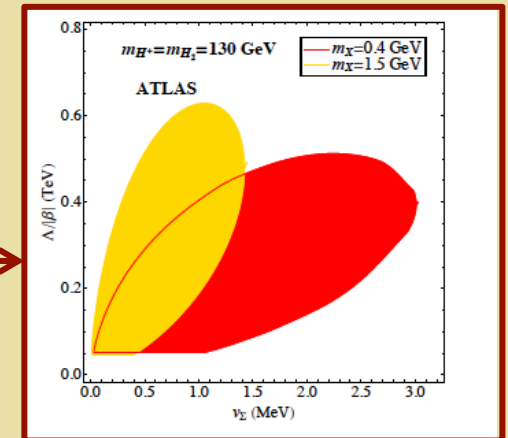
$$\mathcal{O}_{WX}^{(5)} = -\frac{\beta}{\Lambda} \text{Tr} (W_{\mu\nu} \Sigma) X^{\mu\nu}$$



Prompt V + 2 displaced LJ's



Recast ATLAS '14  
(no prompt V)



# *LLP's: Roadmap for Theory & Expt*



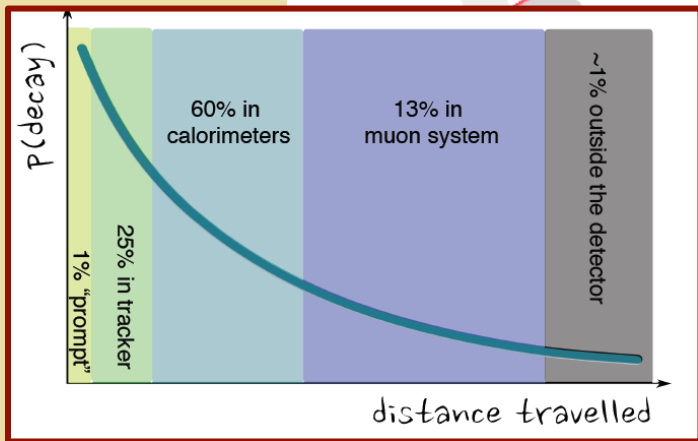
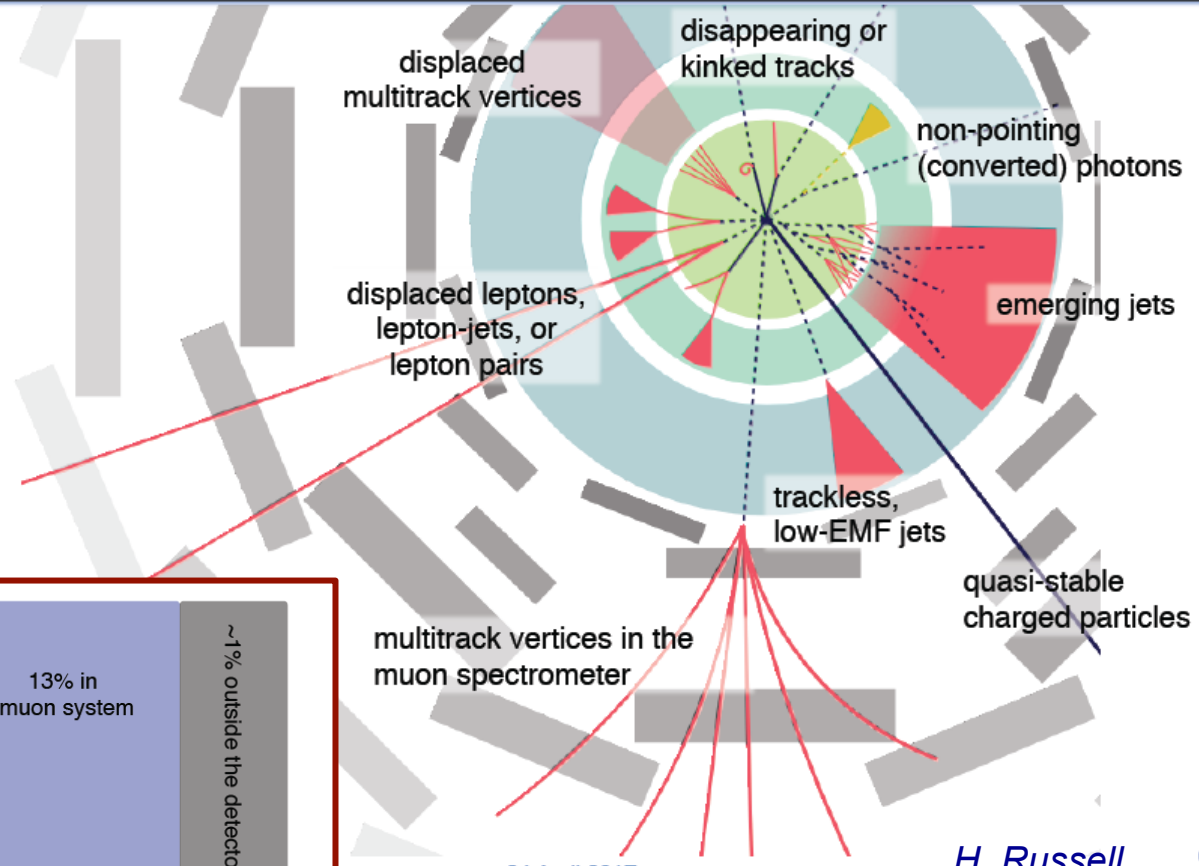
*Time to extend the coverage & reach !*

# LLP's: Challenges

- 1 or 2 displaced LJ's + prompt L or V ( $N_R$ , non-Abelian  $Z_D$ )
- Displaced jets (WIMPY baryogenesis, neutral naturalness)
- Displaced V + jets (Stealth SUSY)
- Displaced  $\mu$  + tracks (WIMPY baryogenesis)
- Emerging jets (Dark QCD)
- High multiplicity b-jets + displaced jets (Stealth SUSY, hidden valleys...)
- Disappearing or kinked charged tracks (SUSY, quirks, EW multiplet DM...)
- ...

# LLP's: Challenges

so where do we start?



24 April 2017

H. Russell

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# *LLP's: Challenges*

## **WG 1**

*Simp models & recast*

## **WG 2**

*Backgrounds*

*Triggering*

## **WG 3**

*Dark Shower*

## **WG 4**



# LLP's: Challenges

## WG 1

*Simp models & recast*

*( Common production ) x  
( Variety of decay objects )*

*L1 in good shape, but  
exploit associated objects*

*Triggering*

## WG 3

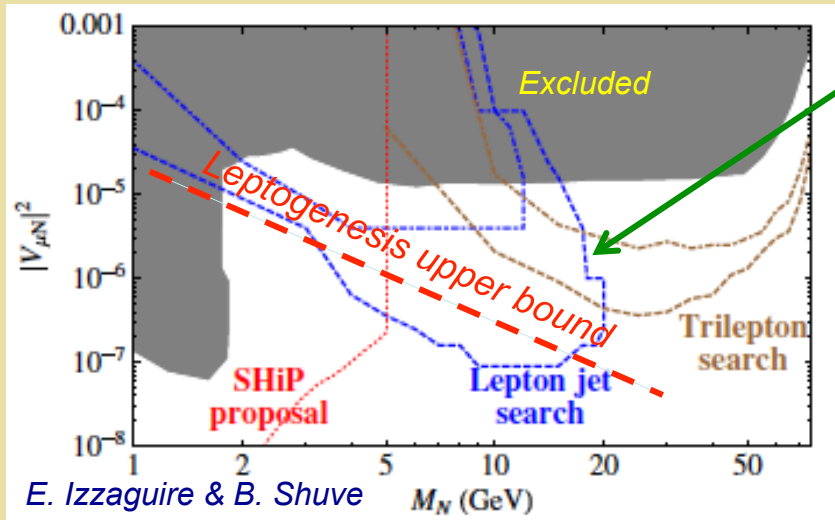
## WG 2

*Backgrounds*

*Dark Shower*

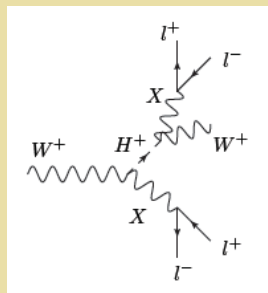
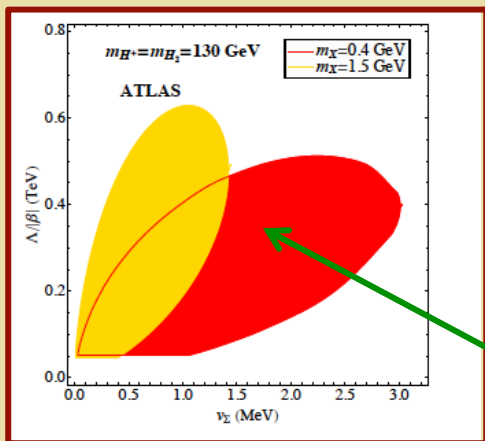
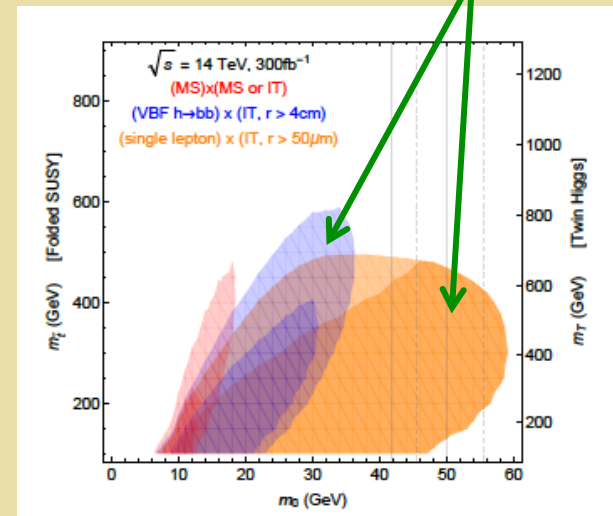
## WG 4

# Push the Reach w/ Prompt Objects



Prompt  $\mu$  + semilept  $N_R$  decay

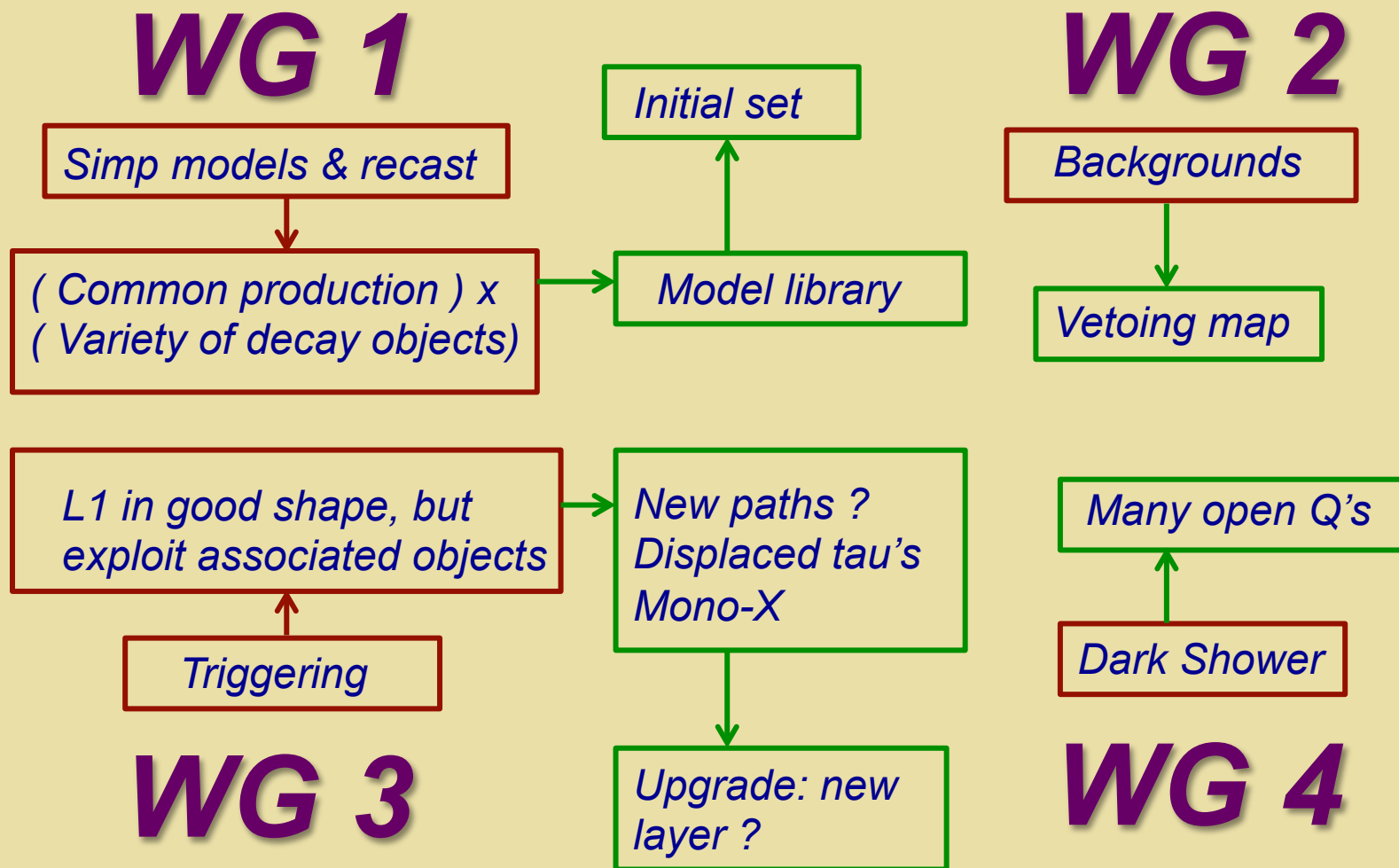
VBF jets or prompt  $V$



Prompt  $W$  + 2 DV lepton jets

Associated prompt objects !

# LLP's: Challenges



## *LLP's: Outlook*

