

# *Long-lived-particles searches in low-scale seesaws*

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VNIVERSITAT  
DE VALÈNCIA



elusioes  
neutrinos, dark matter & dark energy physics

Talk based on A.C, P. Hernandez, J. Lopez-Pavon, J. Salvado (to appear..)

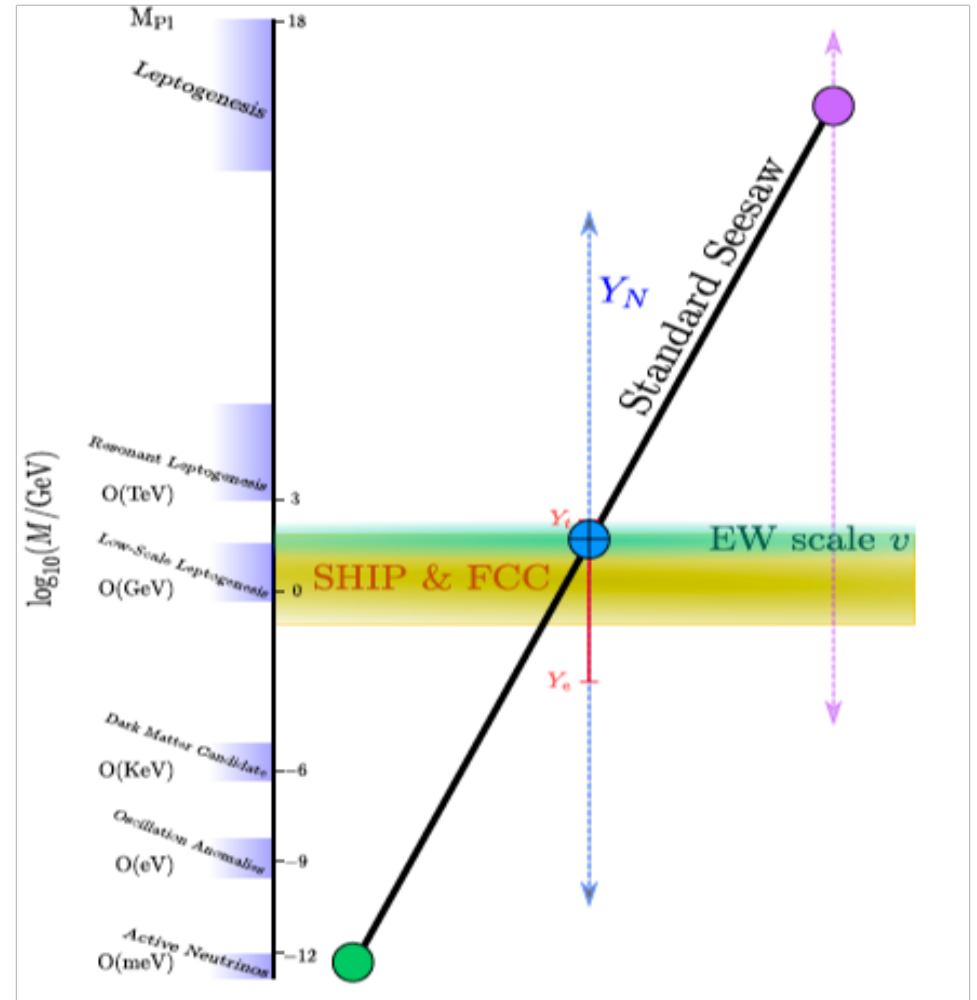
Cern 2017

# Motivation

$$\bullet \mathcal{L}_{seesaw} = \mathcal{L}_{SM} - \sum_{\alpha,i} \overline{L}_{\alpha} Y^{\alpha,i} \tilde{\phi} N_i - \sum_{i,j=1}^2 \frac{1}{2} \overline{N}_i^C M_N^{i,j} N_j + h.c$$

P. Minkowski(1977), M. Gell-Mann, P. Ramond and R. Slansky (1979),  
T. Yanagida (1979), R.N. Mohapatra and G. Senjanovic (1980)

- Observed neutrino masses
- Explain Matter-Antimatter asymmetry via neutrino oscillations if  $M_N \in [1, 10^2] GeV$
- Testable scenario in beam dump experiments and future colliders

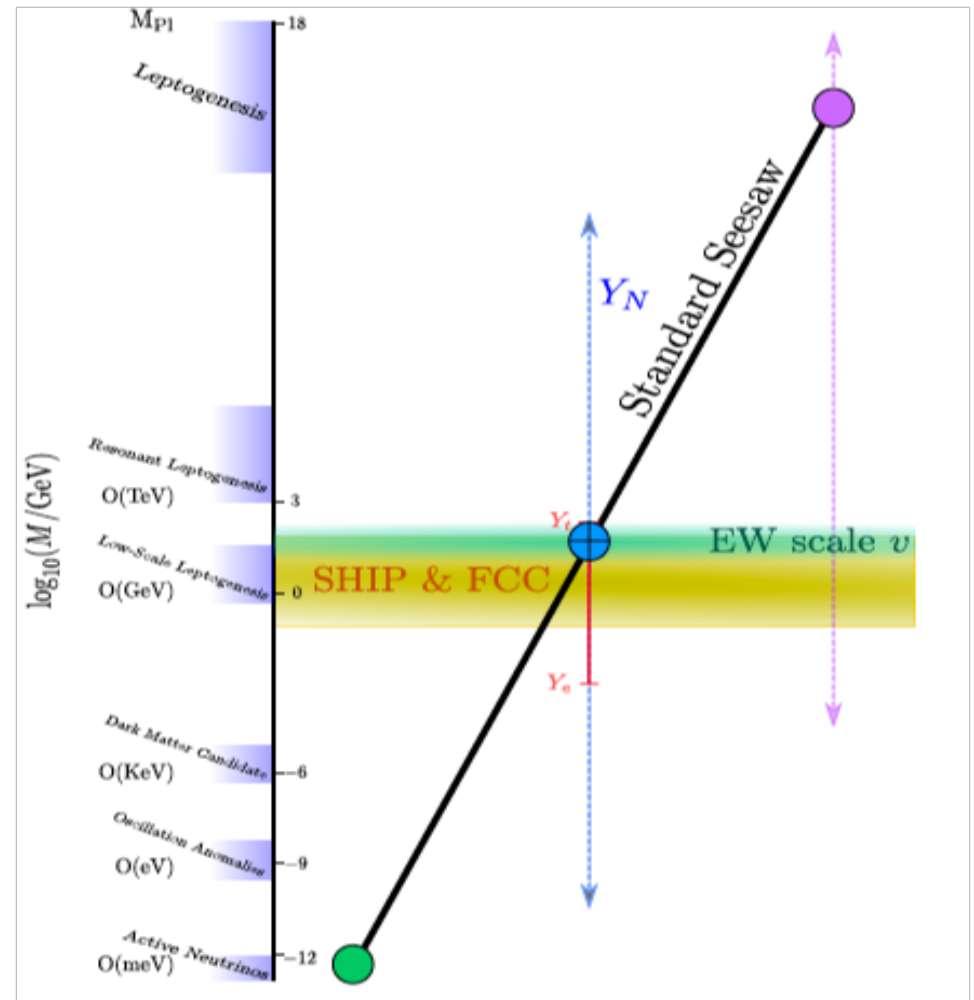


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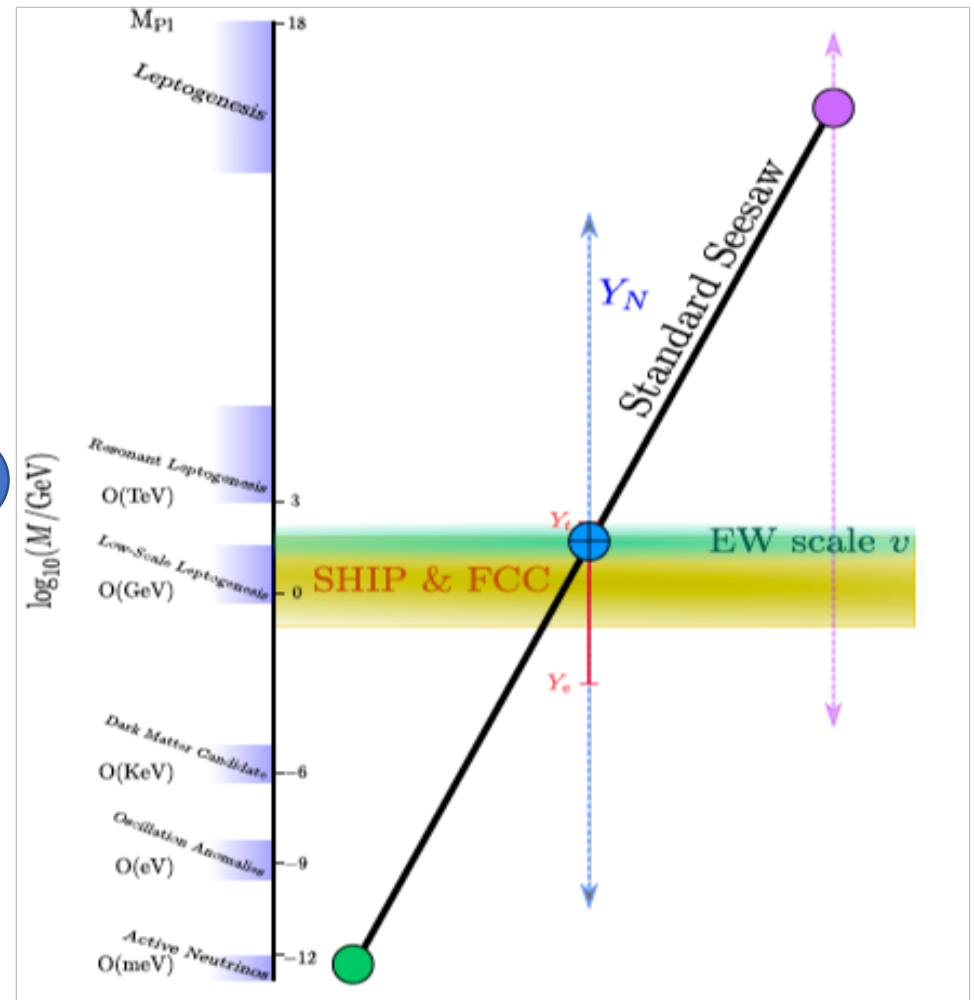
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E.K.Akhmedov, V.Rubakov, A.Y. Smirnov  
Asaka, Shaposhnikov

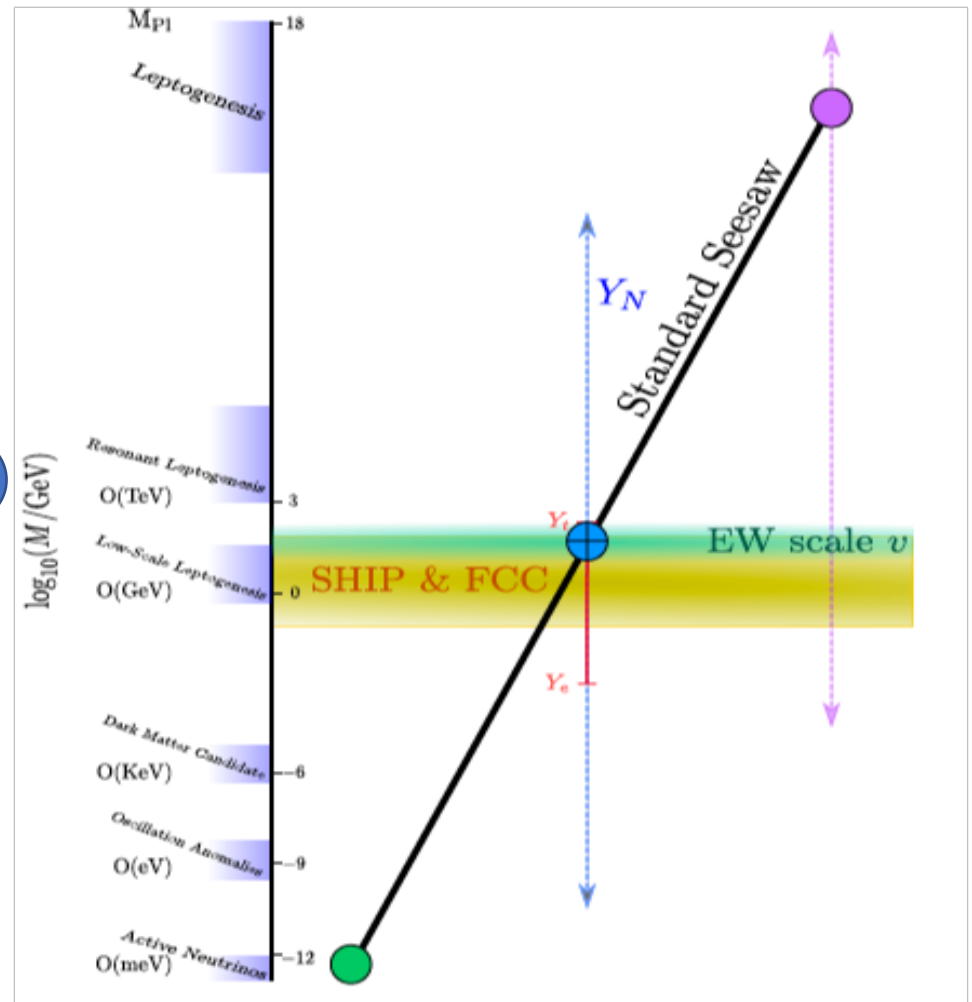
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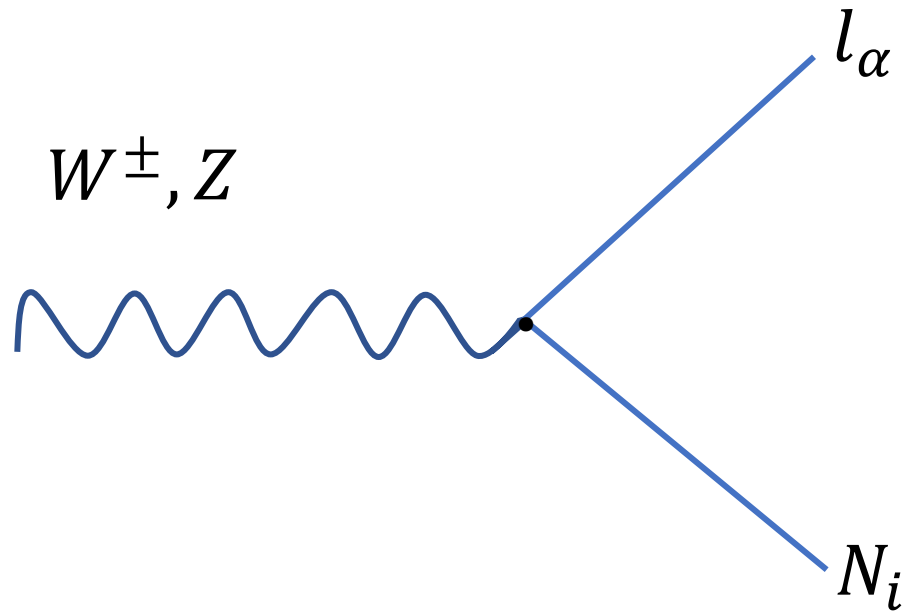
- Observed neutrino masses 😊
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- light neutrinos masses

$$m_\nu \approx -\frac{v^2}{2} Y \frac{1}{M} Y^T$$

- two heavy states with mass M



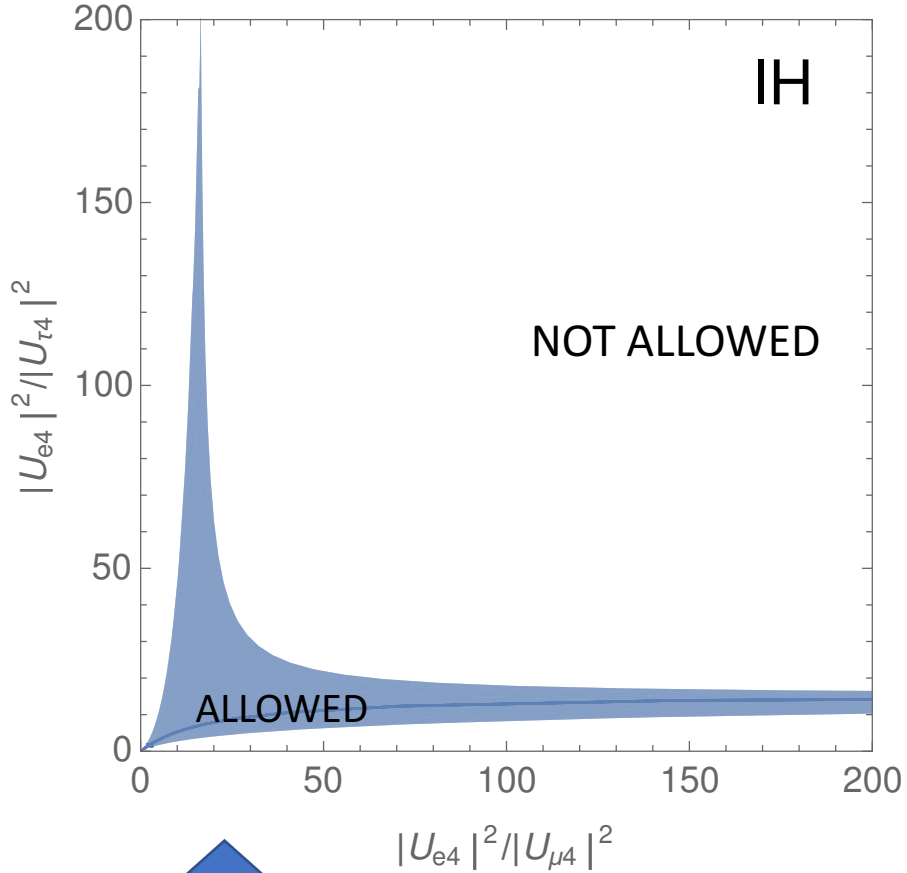
*mixing between leptons and sterile neutrinos*

$$U_{\alpha i} = \frac{v}{\sqrt{2}} Y M^{-1} \approx \sqrt{\frac{m_\nu}{M}}$$

Strongly correlated to the light neutrino masses!!

# High Predictivity

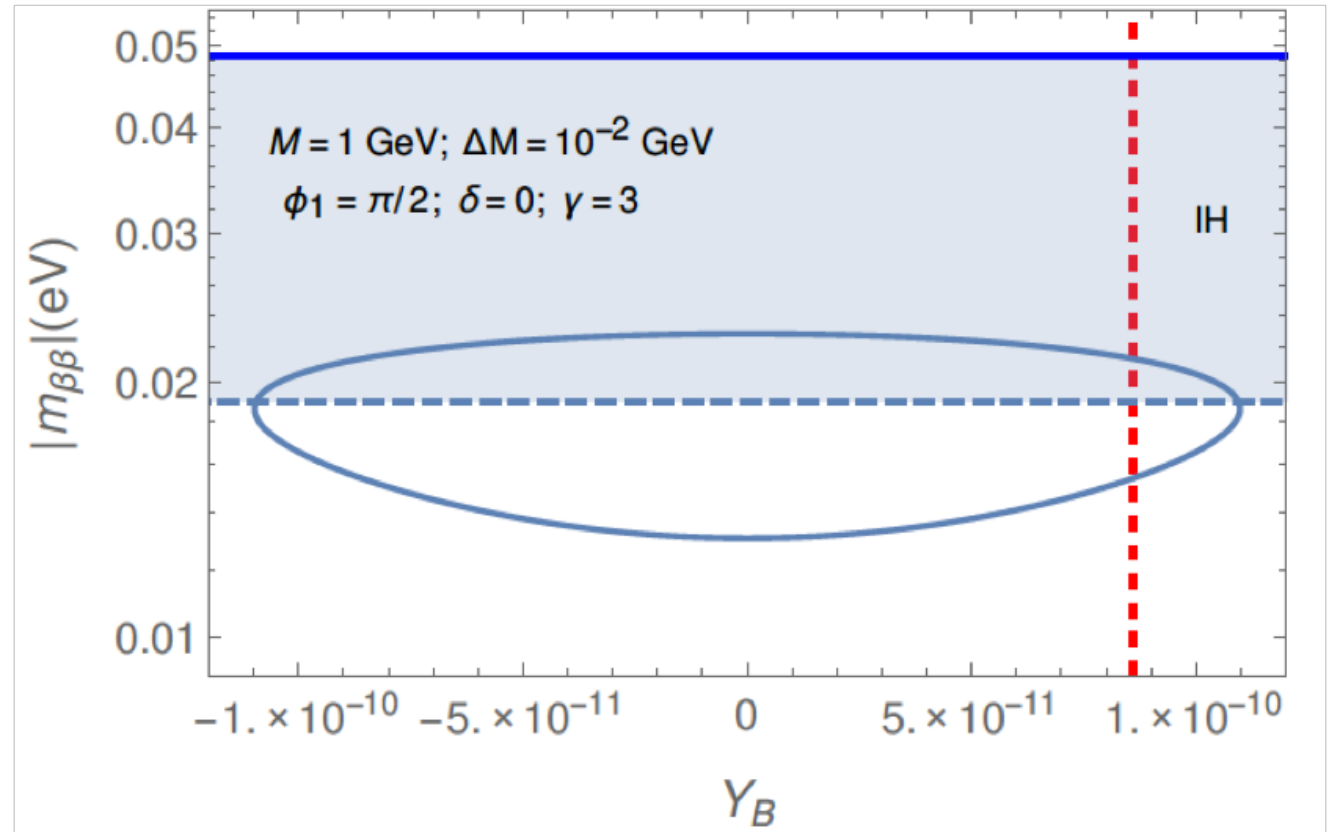
Flavour ratios of mixings



Strong constraints from **light neutrinos masses and mixings**

CP Phases  $(\delta, \phi)$

Non trivial correlation between baryon asymmetry and neutrinosless double beta decay amplitude



P. Hernandez, M. Kekic, J. Lopez-Pavon, J. Racker and J. Salvado, JHEP 08 (2016) 157

A.C, P. Hernandez, J.Lopez-Pavon, M. Kekic, J.Salvado, 1611.05000

These predictions rely to a large extent on its  
minimality



To what extent can they be modified in the the presence of additional new physics?

# Model independent approach: EFT

d=5 operators

- $O_W = \sum_{\alpha,\beta} \frac{(\alpha_W)_{\alpha,\beta}}{\Lambda} \overline{L}_\alpha \tilde{\phi} \phi^\dagger L_\beta^C$
- $O_{N\phi} = \sum_{i,j} \frac{(\alpha_{N\phi})_{i,j}}{\Lambda} \overline{N}_i N_j^C \phi^\dagger \phi$
- $O_{NB} = \sum_{i \neq j} \frac{(\alpha_{NB})_{i,j}}{\Lambda} \overline{N}_i \sigma_{\mu\nu} N_j^C B^{\mu\nu}$

S.Weinberg, Phys.Rev.Lett, 43(1979) 1566-1570

M.Graesser, Phys.Rev.D76 (2007) 075006

F. Del Aguila, S.Bar-Shalom, A.soni and J.Wudka, Phys. Lett. B670 (2009)

# Model independent approach: EFT

$$\bullet O_W = \sum_{\alpha,\beta} \frac{(\alpha_W)_{\alpha,\beta}}{\Lambda} \overline{L}_\alpha \tilde{\phi} \phi^\dagger L_\beta^C \quad \longrightarrow$$

Additional contribution  
to the light neutrino  
masses.

It can **modify the  
predictions**

$$\bullet O_{N\phi} = \sum_{i,j} \frac{(\alpha_{N\phi})_{i,j}}{\Lambda} \overline{N}_i N_j^C \phi^\dagger \phi$$

$$\bullet O_{NB} = \sum_{i \neq j} \frac{(\alpha_{NB})_{i,j}}{\Lambda} \overline{N}_i \sigma_{\mu\nu} N_j^C B^{\mu\nu}$$

# Model independent approach: EFT

$$\bullet O_W = \sum_{\alpha,\beta} \frac{(\alpha_W)_{\alpha,\beta}}{\Lambda} \overline{L}_\alpha \tilde{\phi} \phi^\dagger L_\beta^C$$

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New coupling  
to the Higgs

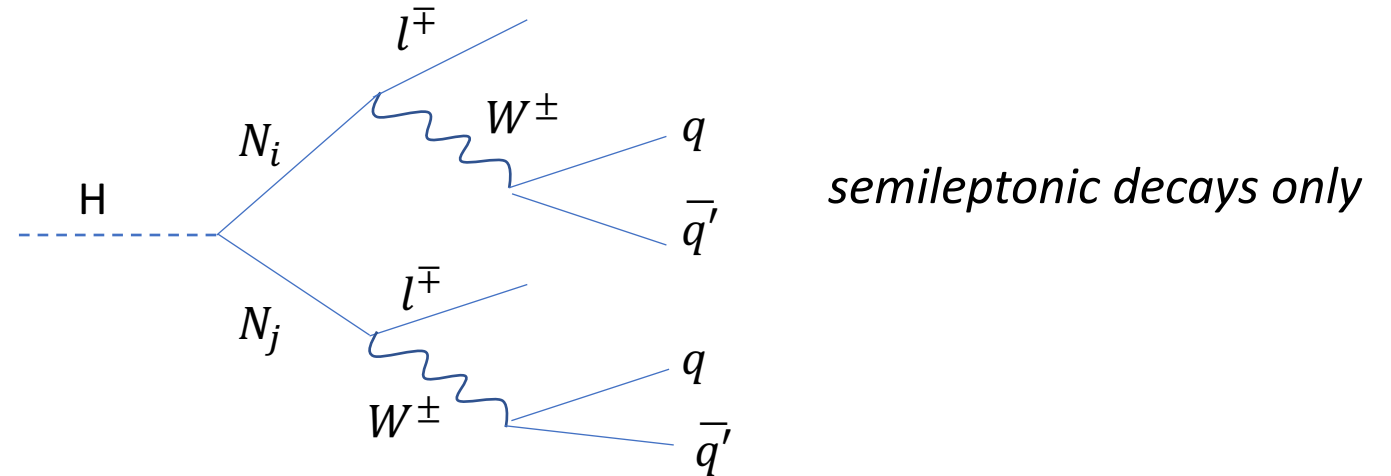
$$\bullet O_{NB} = \sum_{i \neq j} \frac{(\alpha_{NB})_{i,j}}{\Lambda} \overline{N}_i \sigma_{\mu\nu} N_j^C B^{\mu\nu}$$

# Higgs coupling

$$L \supset -\frac{v}{\sqrt{2\Lambda}} H \bar{N}^c \alpha_{N\phi} N + h.c. \quad \xrightarrow{M_i \leq \frac{M_H}{2}} \quad \text{Spectacular signal at LHC, pair of displaced vertices (DVs)}$$

*MadGraph5*, parton-level  
Monte Carlo analysis

$$E_{C.M} = 13 \text{ TeV}, \mathcal{L} = 300 \text{ fb}^{-1}$$



1. Search of displaced tracks in the inner tracker where at least one displaced lepton,  $e$  or  $\mu$ , is reconstructed from each vertex
2. Search for displaced tracks in the muon chambers and outside the inner tracker, where at least one  $\mu$  is reconstructed from each vertex

## Kinematical cuts:

- $p_T(l) > 26\text{GeV}$
- $|\eta| < 2$
- $\Delta R > 0.2$
- $\cos \theta_{\mu\mu} > -0.75$

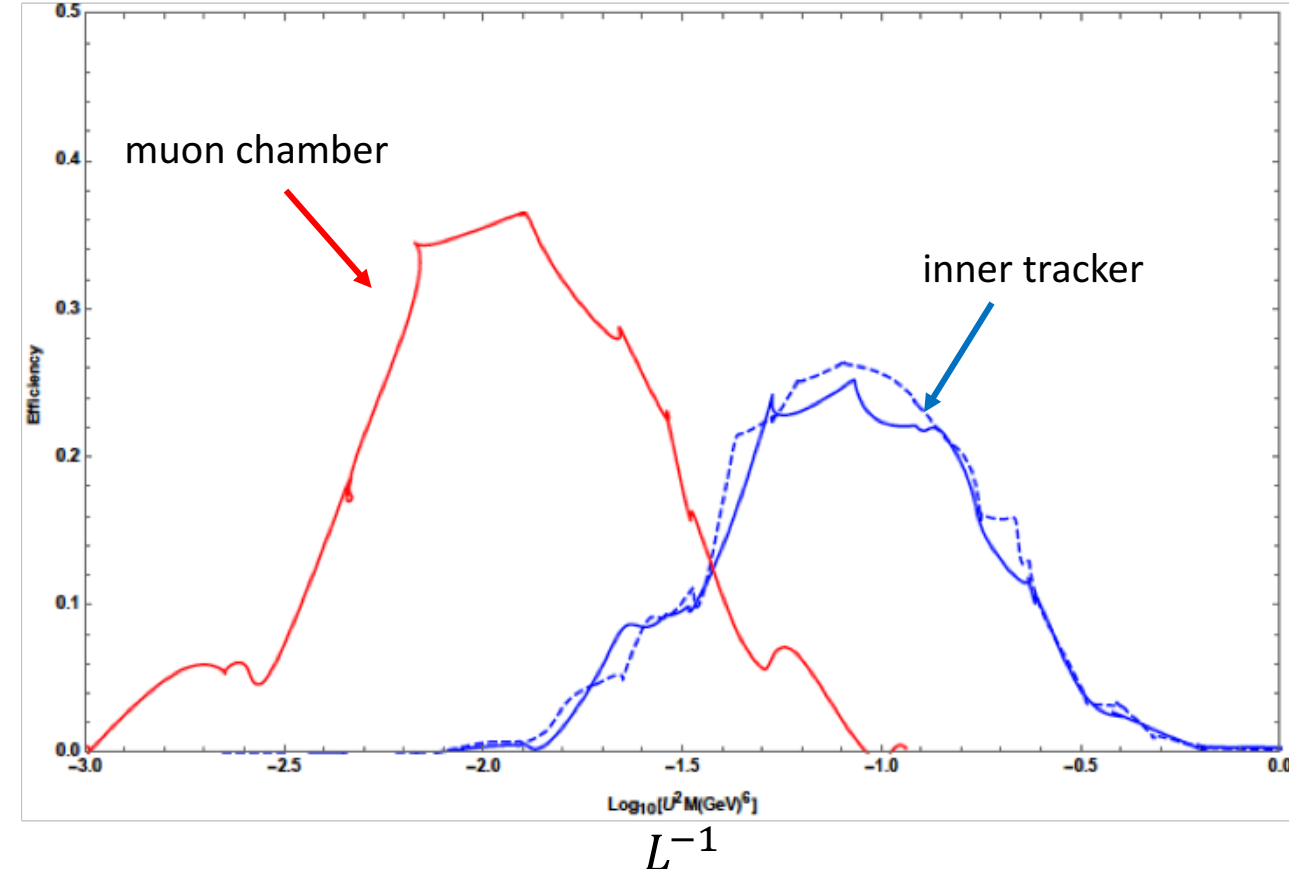
$\mu\mu$	10 GeV	20GeV	30GeV	40 GeV
$p_T$	7.0%	6.8%	6.0%	4.7 %
$\eta$	4.7%	4.9%	4%	3.2%
$\Delta R$	4.7%	4.9%	4%	3.2%
$\cos \theta_{\mu\mu}$	3.2%	3.6%	3.0%	2.7%

→ Sterile Neutrino Mass

## Cuts associated to displaced tracks:

1.  $10\text{cm} < |L_{xy}| < 50\text{cm}, |L_z| \leq 1.4\text{m}, \frac{d_0}{\sigma_d^t} > 12$  ( $\sigma_d^t \approx 20\mu\text{m}$ ) Inner Tracker (IT)
2.  $|L_{xy}| < 5\text{m}, |L_z| \leq 8\text{m}, \frac{d_0}{\sigma_d^\mu} > 4$  ( $\sigma_d^\mu \approx 2\text{cm}$ ) Muon Chambers (MC)

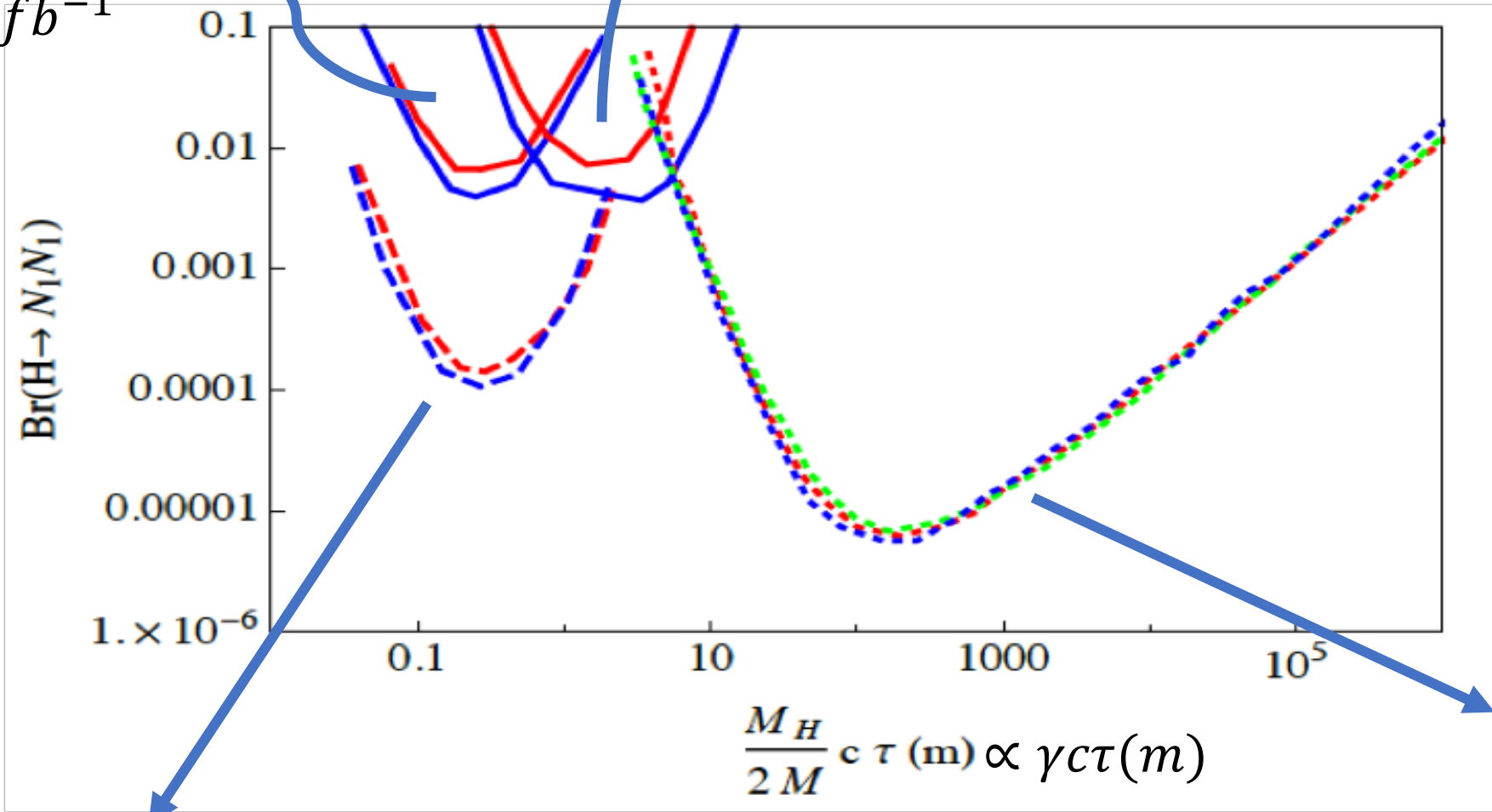
$$L^{-1} \propto U^2 M^6$$



NH, InnerTracker

$$\mathcal{L} = 300 fb^{-1}$$

NH, Muon Chamber



$M=20\text{GeV}$

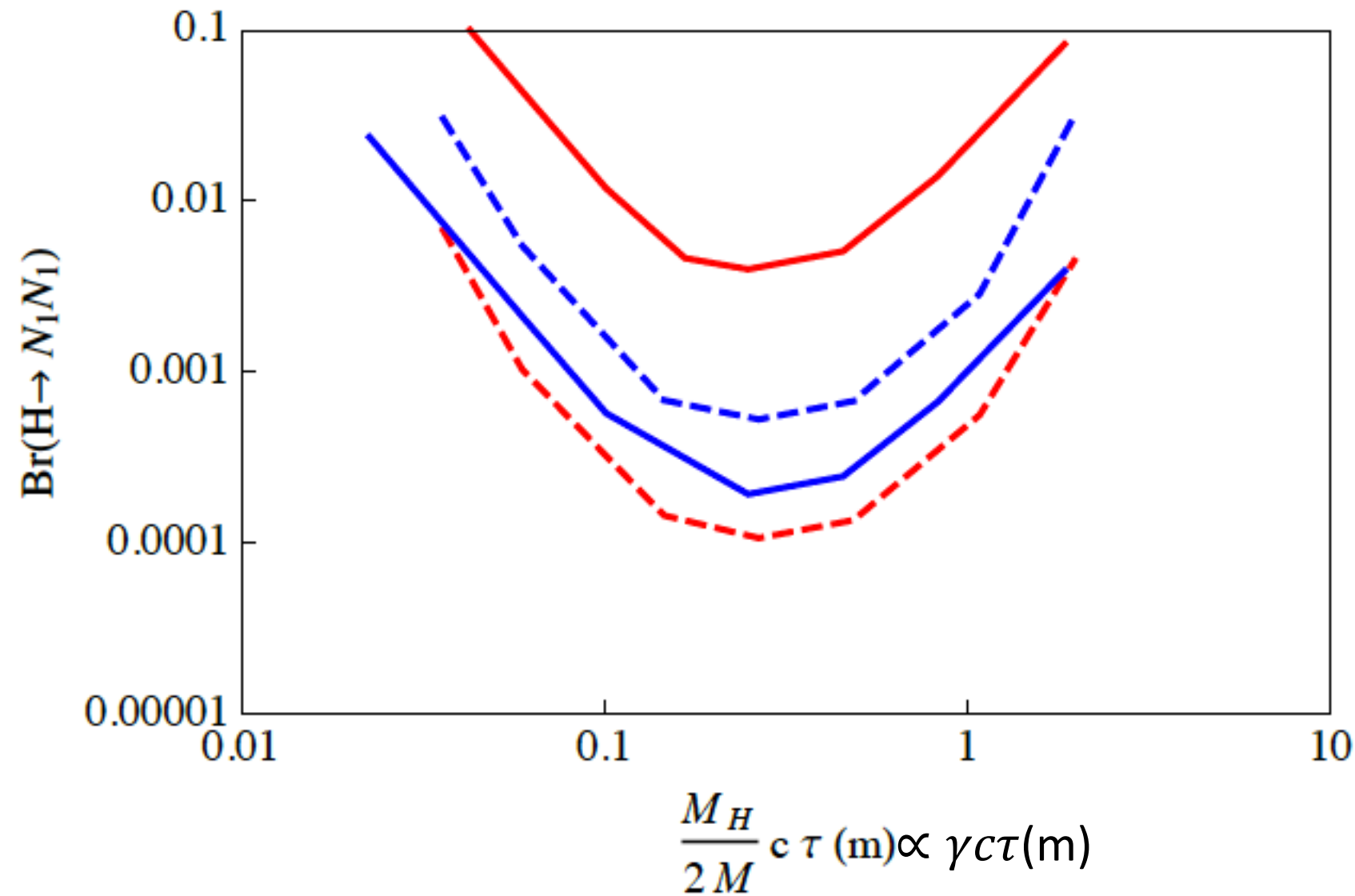
$M=35\text{GeV}$

$$(\delta, \phi) = \left(0, +\frac{\pi}{2}\right)$$

IH

Mathusla (J.P. Chou,  
D.Curtin and H.J.Lubatti,  
Phys.Lett.B767(2017) 29-  
36)

$$\mathcal{L} = 3000 fb^{-1}$$



$$M_N = 20 \text{ GeV}$$

*Solid*  $\rightarrow$  *NH*

*Dashed*  $\rightarrow$  *IH*

*Blue*  $\rightarrow (\delta, \phi) = \left(0, -\frac{\pi}{2}\right)$

*Red*  $\rightarrow (\delta, \phi) = \left(0, +\frac{\pi}{2}\right)$

$$\frac{\alpha_{N\phi}}{\Lambda} \leq 6 \times (10^{-3} - 10^{-2}) \text{ TeV}^{-1}$$



# Conclusion

- If the coefficients of the d=5 operators are all of the same order, the strongest bounds come from the bound on the lightest neutrino mass:

$$\left| \frac{\alpha_W v^2}{2\Lambda} \right| \leq O(1) m_{lightest} \leq 0.2 eV \rightarrow \frac{\alpha_W}{\Lambda} \leq 6 \times 10^{-9} TeV^{-1}$$

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- If we want the predictions on flavour mixings not to be lost, the bounds are **even stronger**

$$\left| \frac{\alpha_W v^2}{2\Lambda} \right| \leq \sqrt{\Delta m_{sol}^2}$$

In the presence, instead, of **large hierarchies**  $\alpha_W \ll \alpha_{N\phi} \sim \alpha_{NB}$   
(that could be protected by **global symmetries**:  $U(1)_L, MFV$ )

LHC  $\rightarrow$   $\left| \frac{\alpha_{N\phi} v^2}{\sqrt{2}\Lambda} \right| \leq 10^{-3} - 10^{-2} \rightarrow \frac{\alpha_{N\phi}}{\Lambda} \leq 6 \times (10^{-3} - 10^{-2}) TeV^{-1}$

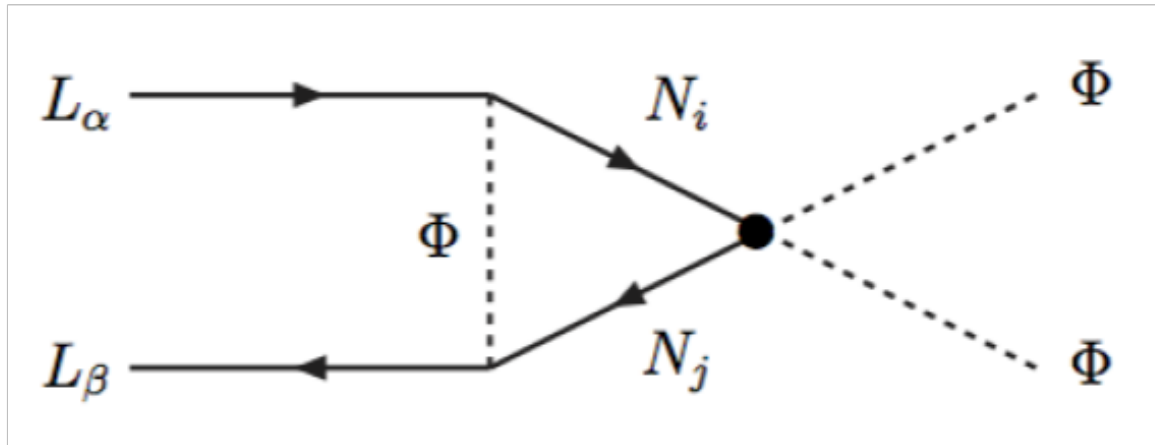
$$\frac{\alpha_{NB}}{\Lambda} \leq 10^{-2} - 10^{-1} TeV^{-1}$$

$\downarrow$

A.Aparici, K.Kim, A.Santamaria and J.Wudka, Phys. Rev. D80(2009) 013010

Thanks for your attention

# Radiative corrections



$$\delta \left( \frac{\alpha_W}{\Lambda} \right) \propto \frac{1}{4\pi^2} \frac{Y^2}{4} \frac{\alpha_{N\Phi}}{\Lambda} \log \frac{\mu^2}{M^2}$$



This contribution has to be equal or smaller than the tree-level contribution

$$\frac{\alpha_{N\phi}}{\Lambda} \lesssim \frac{2 \cdot 10^{13}}{\log \frac{\mu^2}{M^2}} \left( \frac{10^{-6}}{U^2} \right) \left( \frac{\text{GeV}}{M} \right)^2 \frac{\alpha_W}{\Lambda}$$