

# The Zee model: connecting neutrino masses to Higgs lepton flavor violation

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**CoEPP**  
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- 1 Introduction to neutrino mass models
- 2 The new observable: Higgs lepton flavor violation
- 3 Parameter scan of the Zee model
- 4 Summary and conclusions

# Introduction to neutrino mass models

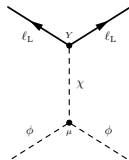
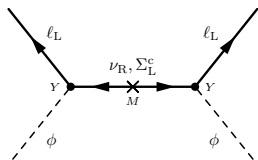
# The Weinberg operator: tree level completions– seesaws

- There is only one dimension 5 EFT operator, which violates  $\Delta L$  in 2.
- It generates Majorana neutrino masses ( $\alpha, \beta$  flavour indices):

$$\mathcal{O}_5 = \frac{1}{2} \frac{c_{\alpha\beta}}{\Lambda} (\overline{L}_\alpha \tilde{\Phi}) (\Phi^\dagger \tilde{L}_\beta) + \text{H.c.} \quad \longrightarrow \quad m_\nu = c \frac{v^2}{\Lambda}.$$

- Rewriting it, with  $\vec{\sigma} \equiv (\sigma_1, \sigma_2, \sigma_3)$ :

$$(\overline{L}_\alpha \tilde{\phi}) (\phi^\dagger \tilde{L}_\beta) = - (\overline{L}_\alpha \vec{\sigma} \tilde{\phi}) (\phi^\dagger \vec{\sigma} \tilde{L}_\beta) = 1/2 (\overline{L}_\alpha \vec{\sigma} \tilde{L}_\beta) (\phi^\dagger \vec{\sigma} \tilde{\phi}),$$

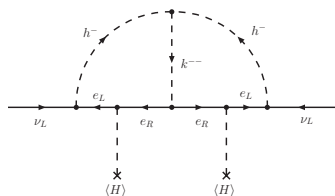


- Left) a  $Y = 0$  heavy fermion singlet (triplet), type I (III) seesaw.
- Right) a  $Y = 1$  heavy scalar triplet, type II seesaw.
- Drawbacks: typically difficult to test, problem of hierarchies.

- Prototype example: the Zee-Babu model, which adds a singly- and a doubly-charged scalar  $h^\pm, k^{\pm\pm}$ .  $\Delta L = 2$  in the  $\mu$  term:

$$\mathcal{L}_Y = \bar{L} Y e \phi + \bar{\tilde{L}} f L h^+ + \bar{e}^c g e k^{++} + (\mu h^2 k^{++} + \text{H.c.}).$$

- Neutrino masses are generated at two loops:



$$\mathcal{M}_\nu = \frac{v^2 \mu}{48\pi^2 M^2} \tilde{I} f Y g^\dagger Y^T f^T,$$

$$M \equiv \max(m_h, m_k).$$

## Clear predictions:

- $f$  is AS  $\rightarrow \det f = 0 \rightarrow \det \mathcal{M}_\nu = 0$ , so one  $\nu$  is massless.
- $k^{++}$  can be light enough to be searched for at the LHC.

# The new observable: Higgs lepton flavor violation

# HLFV as a test of new physics beyond the SM

- HLFV is a new observable to study, with upper limits:

Observable	ATLAS	CMS
$\text{Br}(h \rightarrow \tau\mu)$	1.43 %	1.20 %
$\text{Br}(h \rightarrow \tau e)$	1.04 %	0.69 %

- In the SM, Higgs couplings to the charged leptons are diagonal.
- HLFV implies NP beyond the SM, parameterized by EFT operators:

$$\mathcal{L}_{\text{leptons}} = \bar{L}i\not{D}L + \bar{e}_R i\not{D}e_R - (Y_e \bar{L}e_R \Phi + \sum_a \frac{C_a}{\Lambda^2} \mathcal{O}_a + \text{H.c.})$$

- HLFV occurs at  $D = 6$ :

$$\mathcal{O}_Y = \bar{L} C_Y e_R \Phi (\Phi^\dagger \Phi), \quad \mathcal{O}_{Di} = (\bar{e}_R \Phi^\dagger) C_{Di} i\not{D}(e_R \Phi).$$

- $\mathcal{O}_{Di}$  related by EOM to  $\mathcal{O}_Y$  + plus other non-HLFV operators.

- After SSB,  $\langle \Phi_0 \rangle = (h + v)/\sqrt{2}$ , diagonalize  $M_e$ :

$$(M_e)_{ii} \equiv \text{diag}(m_e, m_\mu, m_\tau) = \frac{1}{\sqrt{2}} V_L^\dagger \left( Y_e + C_Y \frac{v^2}{2\Lambda^2} \right) V_R v.$$

- Yukawas are no longer diagonal ( $V_L^\dagger C_Y V_R \approx C_Y$ ):

$$(y_e)_{ij} = \frac{m_i}{v} \delta_{ij} + (C_Y)_{ij} \frac{v^2}{\sqrt{2}\Lambda^2}, \quad \bar{C}_Y \equiv \sqrt{|(C_Y)_{\tau\mu}|^2 + |(C_Y)_{\mu\tau}|^2}.$$

- HLFV is given by:

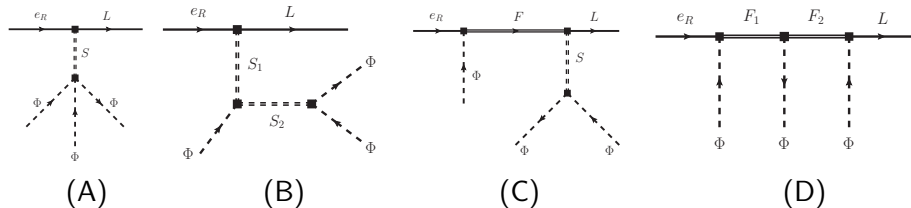
$$\text{BR}(h \rightarrow \tau\mu) = \frac{m_h}{8\pi\Gamma_h} \bar{y}^2, \quad \bar{y} \equiv \bar{C}_Y \frac{v^2}{\sqrt{2}\Lambda^2}.$$



# HLFV UV completions at tree level

[Rius, Santamaria and JHG, arXiv: 1605.06091]

Opening  $\mathcal{O}_Y = \bar{L} C_Y e_R \Phi (\Phi^\dagger \Phi)$  at tree level we obtain:

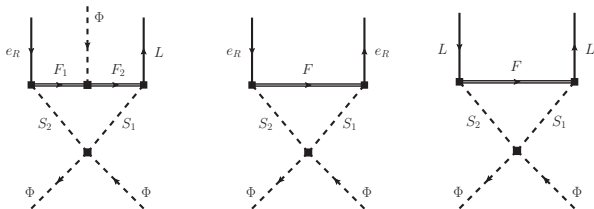


- Topology A (also B) is a two-Higgs doublet, with possible large HLFV.
- Topology C and D with VLL predict very small HLFV,  $< 10^{-6}$ .

We know that lepton flavor is violated in the neutrino sector, so:

**What are the HLFV rates in popular neutrino mass models?**

# Neutrino mass models typically generate HLFV at one loop



Top.	Part.	Representations	Neutrino mass models
LR	S, F	$(1, 0)_F, (3, 0)_F$	Dirac, SSI/III (ISS)
RR	S	$(1, 2)_S$	ZB (doubly-charged $k^{++}$ )
LL	S	$(1, 1)_S, (3, 1)_S$	ZB (singly-charged $h^+$ ), SSII
LL ( $Z_2$ )	$S \oplus F$	$(1, 1/2)_S \oplus (1, 0)_F, (3, 0)_F$	Scotogenic Model

→ In the Zee-Babu (with  $\lambda_{h\Phi}|h^+|^2\Phi^\dagger\Phi + \lambda_{k\Phi}|k|^2\Phi^\dagger\Phi + \text{h.c.}$ ):

$$A_{\text{ZB}}^{h \rightarrow \tau\mu} \sim \frac{m_\tau v}{(4\pi)^2} \left( \frac{\lambda_{h\Phi}}{m_{h^+}^2} (f_{e\mu}^* f_{e\tau}) + \frac{\lambda_{k\Phi}}{m_k^2} (g_{e\mu}^* g_{e\tau} + g_{\mu\mu}^* g_{\mu\tau} + g_{\mu\tau}^* g_{\tau\tau}) \right).$$

# HLFV rates at one loop

- We can estimate HLFV in previous neutrino mass models:

$$\text{BR}(h \rightarrow \mu\tau) \sim 0.06 \frac{\lambda_{ih}^2}{(4\pi)^4} \left(\frac{v}{\text{TeV}}\right)^4 \left(\frac{Y}{M_i/\text{TeV}}\right)^4.$$

- $\tau \rightarrow \mu\gamma$  typically gives the constraint:

$$\left(\frac{Y}{M_i/\text{TeV}}\right)^4 \lesssim \mathcal{O}(0.01 - 1) \quad \longrightarrow \quad \text{BR}(h \rightarrow \mu\tau) \lesssim 10^{-8}.$$

Is  $\text{BR}(h \rightarrow \mu\tau) \sim 0.01$  possible, overcoming the loop  $\sim 1/(4\pi)^4$ ?

- Evade CLFV? NO, some of the new F and S in the loop are charged. One expects CLFV at the same level as HLFV [Dorsner].
- Large Yukawas with special textures:  $\text{BR} \lesssim 10^{-5}$  [ISS, Arganda].
- But: large  $Y, \lambda$  lead to instabilities/non-perturbative and  $h \rightarrow \gamma\gamma$ .

→ **Does any neutrino mass model generate HLFV at tree level?**

# Parameter scan of the Zee model

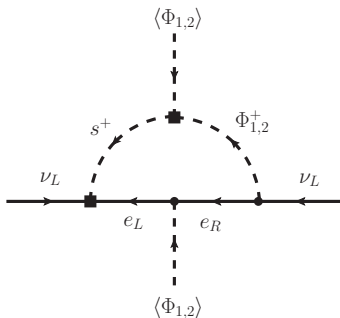
arXiv: 1701.05345

# The Zee Model [Zee, Cheng, Babu, Wolfenstein, Petcov, Kanemura, Aristizabal, He..]

- The Zee model (most general version),  $\Delta L = 2$  in the  $\mu$  term:

$$\mathcal{L}_Y = -\bar{L}(Y_1^\dagger \Phi + Y_2^\dagger \Phi_2)e_R - \bar{L}f Lh^+ + \mu \epsilon_{\alpha\beta} \Phi_1^\alpha \Phi_2^\beta h^- + \text{H.c.}$$

- Neutrino masses are generated via the one-loop diagram:



$$\mathcal{M}_\nu = \frac{s_{2\varphi} t_\beta}{8\sqrt{2}\pi^2 v} \left( f m_f^2 + m_f^2 f^T - \frac{v}{\sqrt{2}s_\beta} (f m_f Y_2 + Y_2^T m_f f^T) \right) \ln \frac{m_{h_2^+}^2}{m_{h_1^+}^2}$$

- The mass matrix (neglecting terms  $\propto m_e$ ) reads:

$$\mathcal{M}_\nu \propto \begin{pmatrix} -2f^{e\tau}Y_2^{\tau e} & -f^{e\tau}Y_2^{\tau\mu} - f^{\mu\tau}Y_2^{\tau e} & \frac{\sqrt{2}s_\beta}{v}m_\tau f^{e\tau} - f^{e\tau}Y_2^{\tau\tau} \\ -f^{e\tau}Y_2^{\tau\mu} - f^{\mu\tau}Y_2^{\tau e} & -2f^{\mu\tau}Y_2^{\tau\mu} & \frac{\sqrt{2}s_\beta}{v}m_\tau f^{\mu\tau} - f^{\mu\tau}Y_2^{\tau\tau} \\ \frac{\sqrt{2}s_\beta}{v}m_\tau f^{e\tau} - f^{e\tau}Y_2^{\tau\tau} & \frac{\sqrt{2}s_\beta}{v}m_\tau f^{\mu\tau} - f^{\mu\tau}Y_2^{\tau\tau} & 2\frac{m_\mu}{m_\tau}f^{\mu\tau}Y_2^{\mu\tau} \end{pmatrix}.$$

- HLFV is given by:

$$\text{BR}(h \rightarrow \mu\tau) = \frac{m_h}{8\pi\Gamma_h} \left( \frac{c_{\beta-\alpha}}{\sqrt{2}c_\beta} \right)^2 (|Y_2^{\tau\mu}|^2 + |Y_2^{\mu\tau}|^2).$$

Reproducing correctly neutrino mixings implies  $Y_2^{e\tau}, Y_2^{\mu\tau} \neq 0$ , so:

- We expect a lower bound on HLFV and CLFV.
- Need full parameter scan to test if they can be observable.

# Full parameter scan of Zee model [with Multinest, Superplot]

[T. Ohlsson, S. Riad, J. Wiren and JHG, arXiv: 1701.05345]

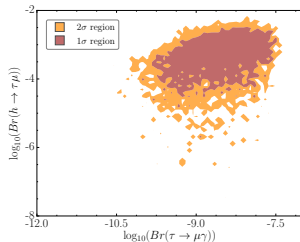
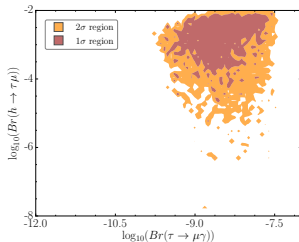
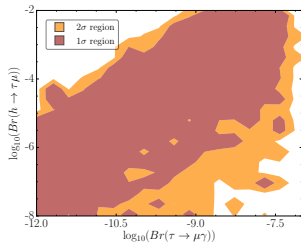
We show the one and two sigma profile likelihoods for:

- no neutrino masses ( $\mu = 0$ ),
- neutrino masses ( $\mu \neq 0$ ) in:
  - 1 Normal Ordering (NO)
  - 2 Inverted Ordering (IO)

$\mu = 0$

NO

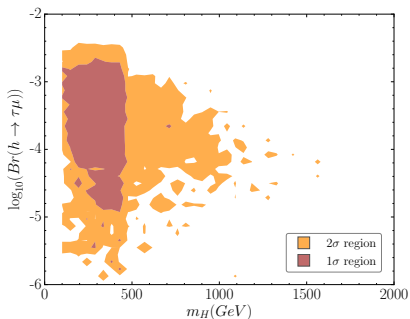
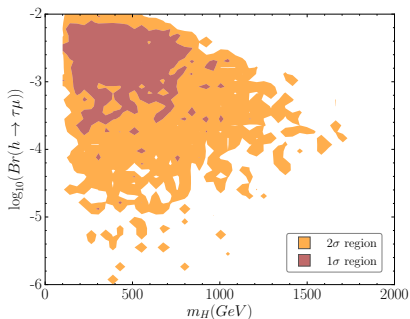
IO



# Decoupling of $h \rightarrow \tau\mu$ with $m_H$ (NO left, IO right)

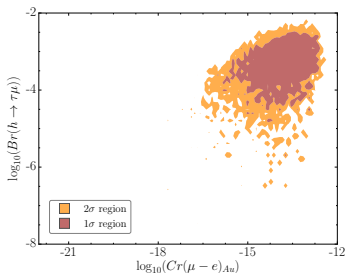
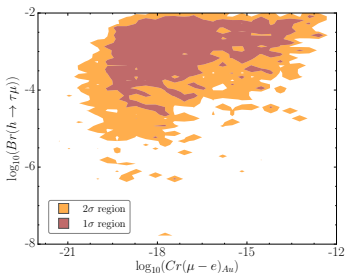
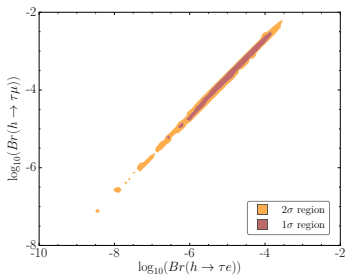
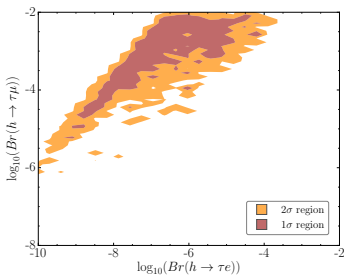
Having a sufficiently SM-like Higgs boson as observed demands  $s(\beta - \alpha) \approx 1$  (alignment limit). Expanding around  $\beta - \alpha \approx \pi/2$ :

$$\text{Br}(h \rightarrow \tau\mu) \simeq \frac{m_h}{16\pi\Gamma_h} \frac{\lambda_6^2 v^4}{c_\beta^2 m_H^4} (|Y_2^{\tau\mu}|^2 + |Y_2^{\mu\tau}|^2).$$

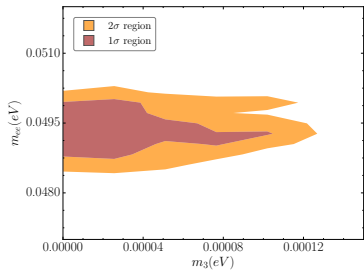
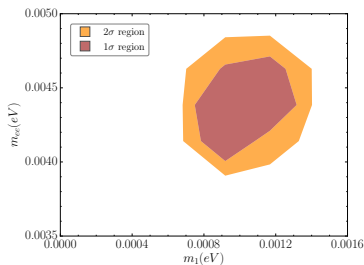
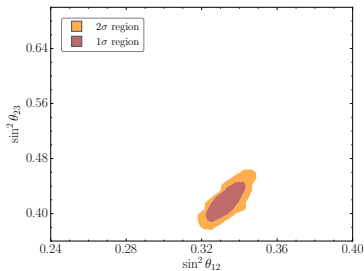
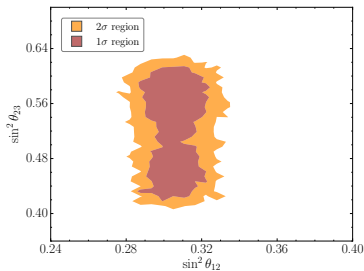




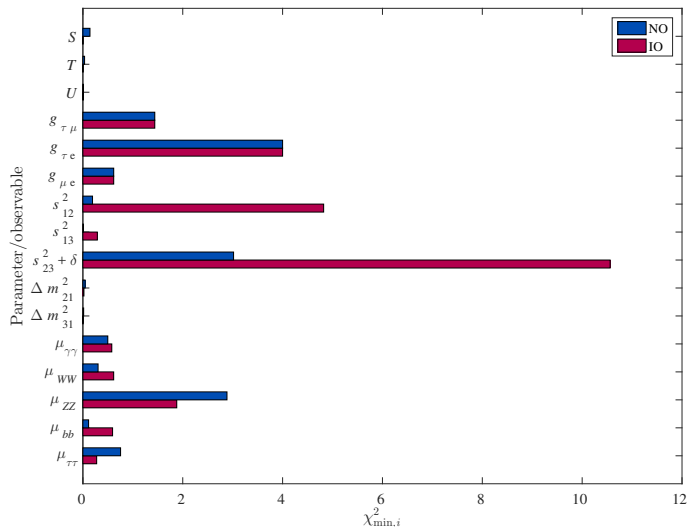
# Other HLFV and CLFV processes (NO left, IO right)



# Neutrino mixings, lightest mass, $m_{ee}$ (NO left, IO right)



# Individual contributions to the $\chi^2$



# Naturality limits from the Higgs mass. 95 % C.L. results

- $m_h$  gets a correction  $\delta m_h \propto \mu$ . Demanding  $\delta m_h/m_h \lesssim \kappa$ :

$$\mu \lesssim \kappa \frac{4\pi m_h}{s_{\beta-\alpha}} \simeq 1.5 \left( \frac{\kappa}{s_{\beta-\alpha}} \right) \text{ TeV}.$$

- $\kappa = 1$  (10) corresponds to *no* (10 %) fine-tuning.

Quantity	NO		IO	
	$\kappa = 1$	$\kappa = 10$	$\kappa = 1$	$\kappa = 10$
$\chi^2_{\min}$	10.7	11.0	21.7	24.0
$\text{Br}_{h \rightarrow \tau \mu}$	$[1 \cdot 10^{-6}, 1 \cdot 10^{-2}]$	$[1 \cdot 10^{-6}, 1 \cdot 10^{-2}]$	$[2 \cdot 10^{-7}, 4 \cdot 10^{-3}]$	$[1 \cdot 10^{-7}, 5 \cdot 10^{-3}]$
$\text{Br}_{h \rightarrow \tau e}$	$[1 \cdot 10^{-10}, 2 \cdot 10^{-4}]$	$[1 \cdot 10^{-10}, 2 \cdot 10^{-4}]$	$[6 \cdot 10^{-9}, 3 \cdot 10^{-4}]$	$[3 \cdot 10^{-9}, 3 \cdot 10^{-4}]$
$\text{Br}_{\tau \rightarrow \mu \gamma}$	$[8 \cdot 10^{-10}, 3 \cdot 10^{-8}]$	$[1 \cdot 10^{-10}, 3 \cdot 10^{-8}]$	$[3 \cdot 10^{-11}, 3 \cdot 10^{-8}]$	$[3 \cdot 10^{-11}, 4 \cdot 10^{-8}]$
$\text{Br}_{\mu \rightarrow e \gamma}$	$[10^{-21}, 6 \cdot 10^{-13}]$	$[3 \cdot 10^{-22}, 6 \cdot 10^{-13}]$	$[1 \cdot 10^{-31}, 1 \cdot 10^{-12}]$	$[1 \cdot 10^{-34}, 1 \cdot 10^{-12}]$
$\text{Cr}_{\mu \rightarrow e}$	$[10^{-21}, 4 \cdot 10^{-13}]$	$[1 \cdot 10^{-21}, 4 \cdot 10^{-13}]$	$[3 \cdot 10^{-17}, 3 \cdot 10^{-13}]$	$[3 \cdot 10^{-17}, 3 \cdot 10^{-13}]$
$m_{H,A}$	$< 1.7$ TeV	$< 2.5$ TeV	$< 1.1$ TeV	$< 1.5$ TeV
$m_{h^\pm}$	$< 1.7$ TeV	$< 2.5$ TeV	$< 1.1$ TeV	$< 1.5$ TeV
$s_{\beta-\alpha}$	[0.98, 1.0]	[0.98, 1.0]	[0.97, 1.0]	[0.97, 1.0]

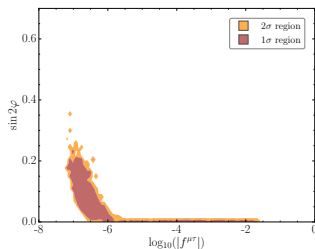
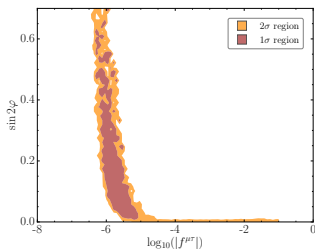
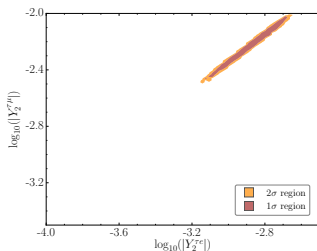
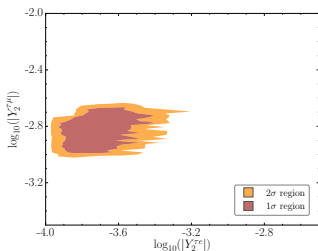
## Summary and conclusions

# Summary and conclusions

- 1 Radiative models are a simple explanation for the lightness of neutrino masses, they do not have large hierarchies and are testable.
- 2 HLFV implies BSM physics, maybe related to neutrino masses.
- 3 However, most neutrino models generate one-loop suppressed HLFV.
- 4 On the other hand, the Zee model generates HLFV at tree level:
  - Large  $h \rightarrow \tau\mu$  is possible.
  - NO gives a good fit, IO is disfavoured.
  - If  $\theta_{23}$  happens to be in the second octant, then IO will be excluded.
  - Lightest neutrino compatible with massless only in IO.
  - Scalar masses have to be below  $\sim 2$  TeV.
  - Future  $\tau \rightarrow \mu\gamma$  ( $\mu e$  conversion) completely test NO (IO).

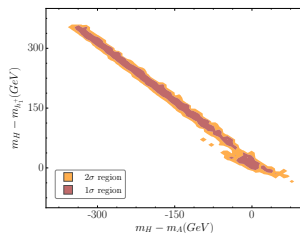
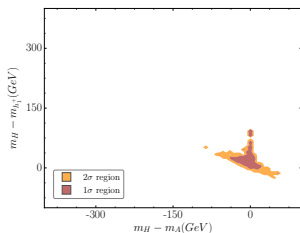
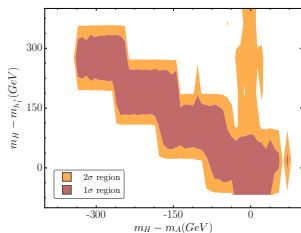
# Back-up slides

# Yukawa couplings (NO left, IO right)

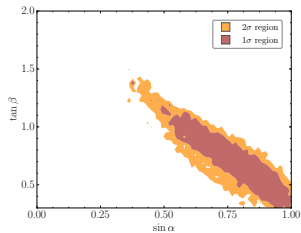
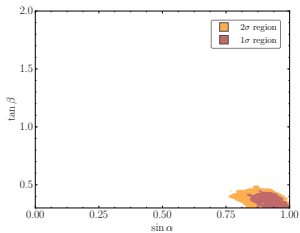
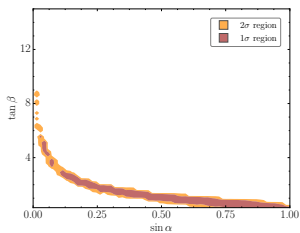




# Splittings of the scalar masses ( $\mu = 0$ , NO left, IO right)



# $\sin \alpha$ and $\tan \beta$ ( $\mu = 0$ , NO left, IO right)



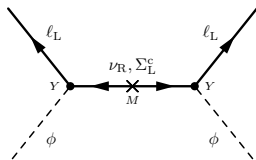
# Adding right-handed neutrinos: seesaw type I

$$\mathcal{L}_{\nu_R} = i \bar{\nu}_R \gamma^\mu \partial_\mu \nu_R - \left( \bar{L} \tilde{\phi} Y \nu_R + \frac{1}{2} \bar{\nu}_R^c m_R \nu_R + \text{H.c.} \right),$$

where  $m_R$  is a  $n \times n$  symmetric matrix. After SSB:

$$\mathcal{L}_{\nu \text{ mass}} = -\frac{1}{2} \begin{pmatrix} \bar{\nu}_L & \bar{\nu}_R^c \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D^T & m_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix} + \text{H.c.},$$

where  $m_D = Y \frac{v}{\sqrt{2}}$ .



If  $m_R \gg m_D$ , one gets  $n$   $m_R$  leptons (mainly singlets) and

$$m_\nu \simeq -m_D m_R^{-1} m_D^T.$$

# Adding a $Y = 1$ scalar triplet: seesaw type II

- The triplet is a  $2 \times 2$  matrix:

$$\chi = \begin{pmatrix} \chi^+/\sqrt{2} & \chi^{++} \\ \chi_0 & -\chi^+/\sqrt{2} \end{pmatrix}$$

- Yukawa coupling to the lepton doublets:

$$\mathcal{L}_\chi = - \left( (Y_\chi)_{\alpha\beta} \bar{L}_\alpha \chi L_\beta + \text{H.c.} \right) - V(\phi, \chi),$$

where  $Y_\chi$  is a symmetric matrix and  $\tilde{L} = i\tau_2 L^c$ . The scalar potential:

$$V(\phi, \chi) = m_\chi^2 \text{Tr}[\chi\chi^\dagger] + \left( \mu \tilde{\phi}^\dagger \chi^\dagger \phi + \text{H.c.} \right) + \dots$$

- $\mu$  induces  $v_\phi$ , even if  $m_\chi > 0$ . From the  $\rho$  parameter,  $v_\chi \lesssim 5 \text{ GeV}$ .
- $Y_\chi$  and  $\mu$  violate LN. In the limit  $m_\chi \gg v_\phi$ , we get:

$$m_\nu = 2Y_\chi v_\chi = 2Y_\chi \frac{\mu v_\phi^2}{m_\chi^2}.$$

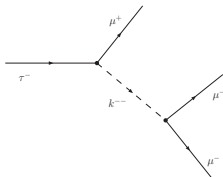
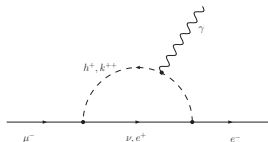
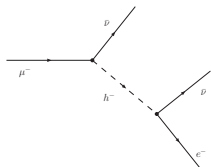
# Zee-Babu strongest constraints: cLFV and universality

- $|V_{ud}^{exp}|^2 + |V_{us}^{exp}|^2 + |V_{ub}^{exp}|^2 = 0.9999 \pm 0.0006$   
 $\approx 1 - \frac{\sqrt{2}}{G_F m_h^2} |f_{e\mu}|^2 \rightarrow |f_{e\mu}|^2 < 0.007 \left(\frac{m_h}{\text{TeV}}\right)^2$

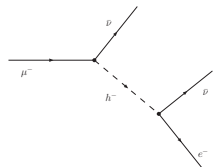
- $\tau/\mu$  universality:  $\frac{G_\tau^{exp}}{G_\mu^{exp}} = 0.9998 \pm 0.0013$   
 $||f_{e\tau}|^2 - |f_{e\mu}|^2| < 0.035 \left(\frac{m_h}{\text{TeV}}\right)^2$

- $\text{BR}(\tau \rightarrow \mu\gamma) < 4.4 \times 10^{-8}$   
 $\frac{|f_{e\tau}^* f_{\mu\tau}|^2}{(m_h/\text{TeV})^4} + \frac{16|g_{ee}^* g_{e\mu} + g_{e\mu}^* g_{\mu\mu} + g_{e\tau}^* g_{\mu\tau}|^2}{(m_k/\text{TeV})^4} < 0.7$

- $\text{BR}(\tau^- \rightarrow \mu^+ \mu^- \mu^-) < 2.1 \times 10^{-8}$   
 $|g_{\mu\tau} g_{\mu\mu}^*| < 0.008 \left(\frac{m_k}{\text{TeV}}\right)^2$

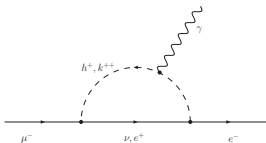


# Zee Babu universality and LFV constraints



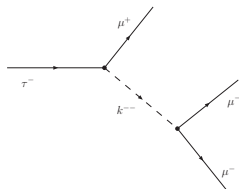
- $$|V_{ud}^{exp}|^2 + |V_{us}^{exp}|^2 + |V_{ub}^{exp}|^2 = 0.9999 \pm 0.0006$$

$$\approx 1 - \frac{\sqrt{2}}{G_F m_h^2} |f_{e\mu}|^2 \rightarrow |f_{e\mu}|^2 < 0.007 \left(\frac{m_h}{\text{TeV}}\right)^2$$



- $$\text{BR}(\mu \rightarrow e\gamma) < 5.7 \times 10^{-13}$$

$$\frac{|f_{e\tau}^* f_{\mu\tau}|^2}{(m_h/\text{TeV})^4} + \frac{16|g_{ee}^* g_{e\mu} + g_{e\mu}^* g_{\mu\mu} + g_{e\tau}^* g_{\mu\tau}|^2}{(m_k/\text{TeV})^4} < 1.6 \cdot 10^{-6}$$



- $$\text{BR}(\mu^- \rightarrow e^+ e^- e^-) < 1.0 \times 10^{-12}$$

$$\rightarrow |g_{e\mu} g_{ee}^*| < 2.3 \cdot 10^{-5} \left(\frac{m_k}{\text{TeV}}\right)^2$$