



# Non-minimal Dark Sector Phenomenology: Prelude

Phenomenology Symposium 2017

arXiv: 1612.06867 and arXiv: 1702.02944 with Jong-Chul Park and Seodong Shin



## DOOJIN KIM

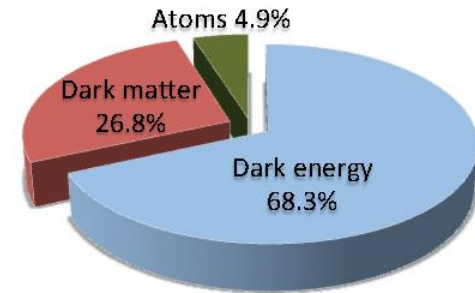
NON-MINIMAL DARK SECTOR

05.09.17

# Dark Matter

## ● Existence of dark matter

- ❑ Dark Matter (DM): **~25% of our universe**, existence supported by rotation curves, gravitational lensing, cosmic microwave background etc.



# Dark Matter

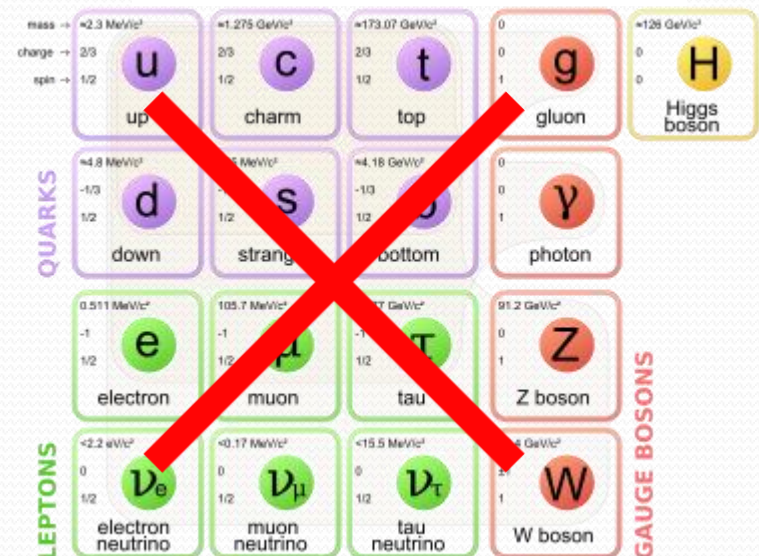
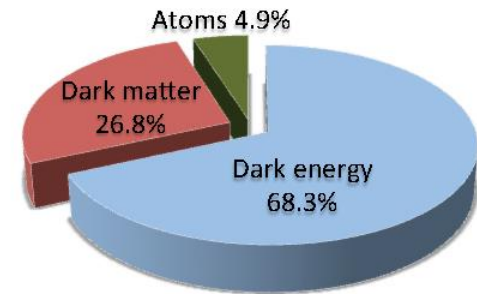
## Known properties

Dark Matter (DM): **~25% of our universe**, existence supported by rotation curves, gravitational lensing, cosmic microwave background etc.

Known DM properties (albeit few):

- ❖ gravitationally interacting
- ❖ not short-lived
- ❖ not hot
- ❖ not baryonic
- ❖ not electrically charged

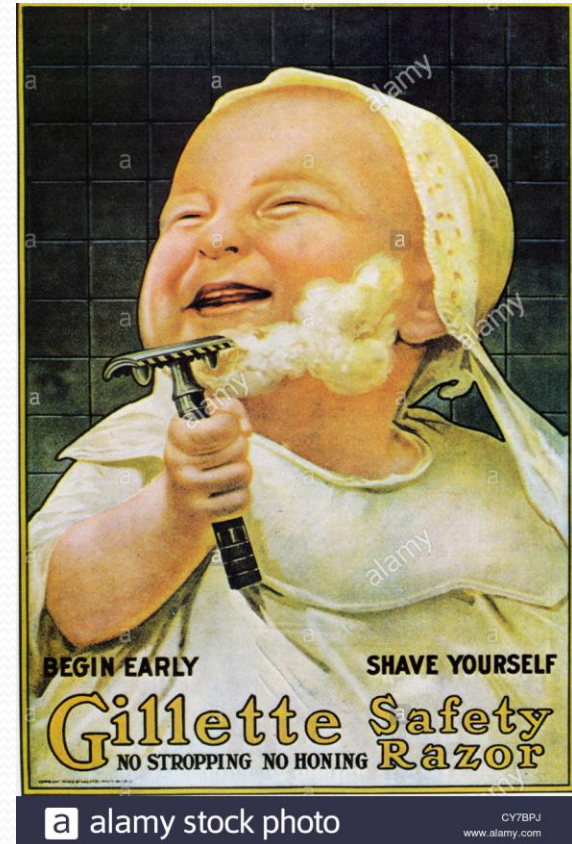
⇒ Need for **new physics/particles**



# “Minimal” Dark Sector

## ● Occam's razor(?)

- ❑ Many DM simplified models or new physics models including a DM candidate proposed:
  - ✓ Positing **single DM species**/focusing on **dominant DM** (with other dark sector particles put aside)
  - ✓ Good and economical approach toward the truth in the earlier stage!
  - ✓ Concentrating on **DM itself** and relevant **phenomenological implications**



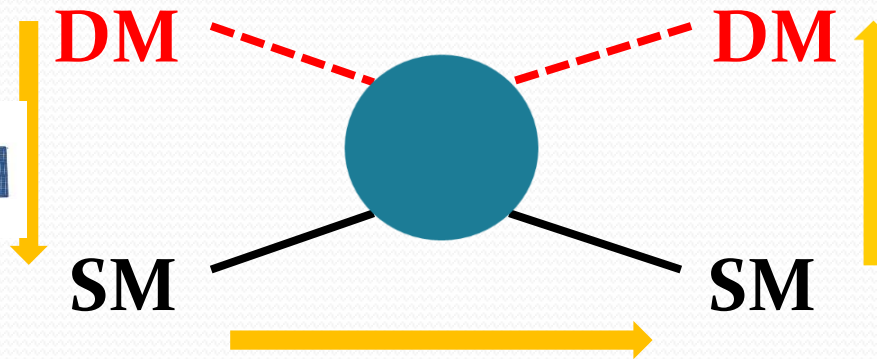
# “Minimal” Dark Sector

## ● “Minimal” phenomenological implications

### DM indirect search



- ✓ (Non-relativistic) DM annihilation/decay to  $\gamma, e^+, \bar{p}$ , etc.
- ✓  $\langle\sigma v\rangle \sim 10^{-26} \text{ cm}^3/\text{s}$

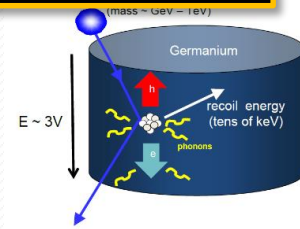


### DM production



- ✓ Active DM production at colliders
- ✓ Mono-X searches
- ✓ Expected rate inferred from/related to  $\langle\sigma v\rangle \sim 10^{-26} \text{ cm}^3/\text{s}$

### DM direct search

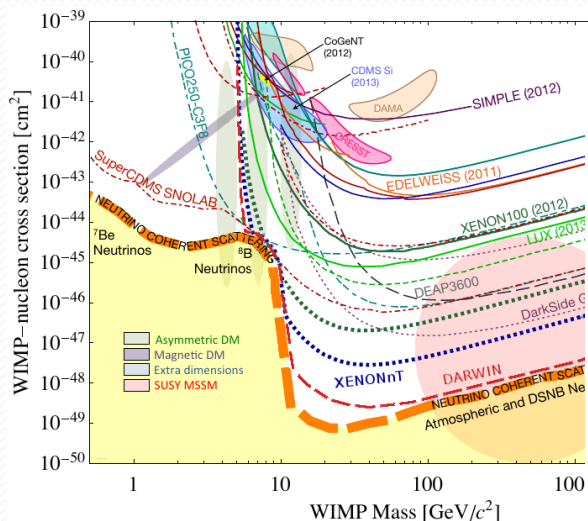


- ✓ (Non-relativistic) DM scattering off target nuclei
- ✓  $E_{\text{recoil}} \sim 1 - 100 \text{ keV}$

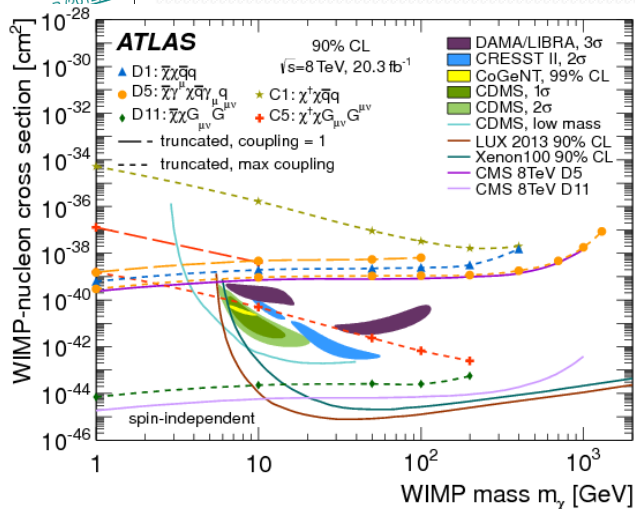
# “Non-minimal” Dark Sector

## Why flavorful?

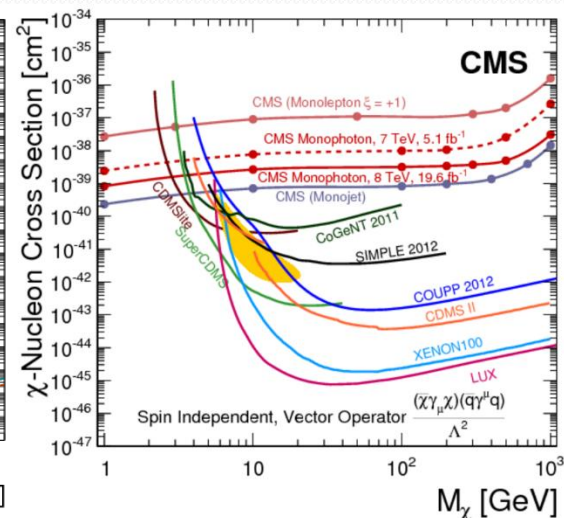
- ❑ No “unambiguous” observation of DM signatures via non-gravitational interactions (many searches/interpretations designed under minimal dark-sector scenarios)



[P. Cushman, C. Calbiati and D. N. McKinsey, (2013); L. Baudis (2014)]



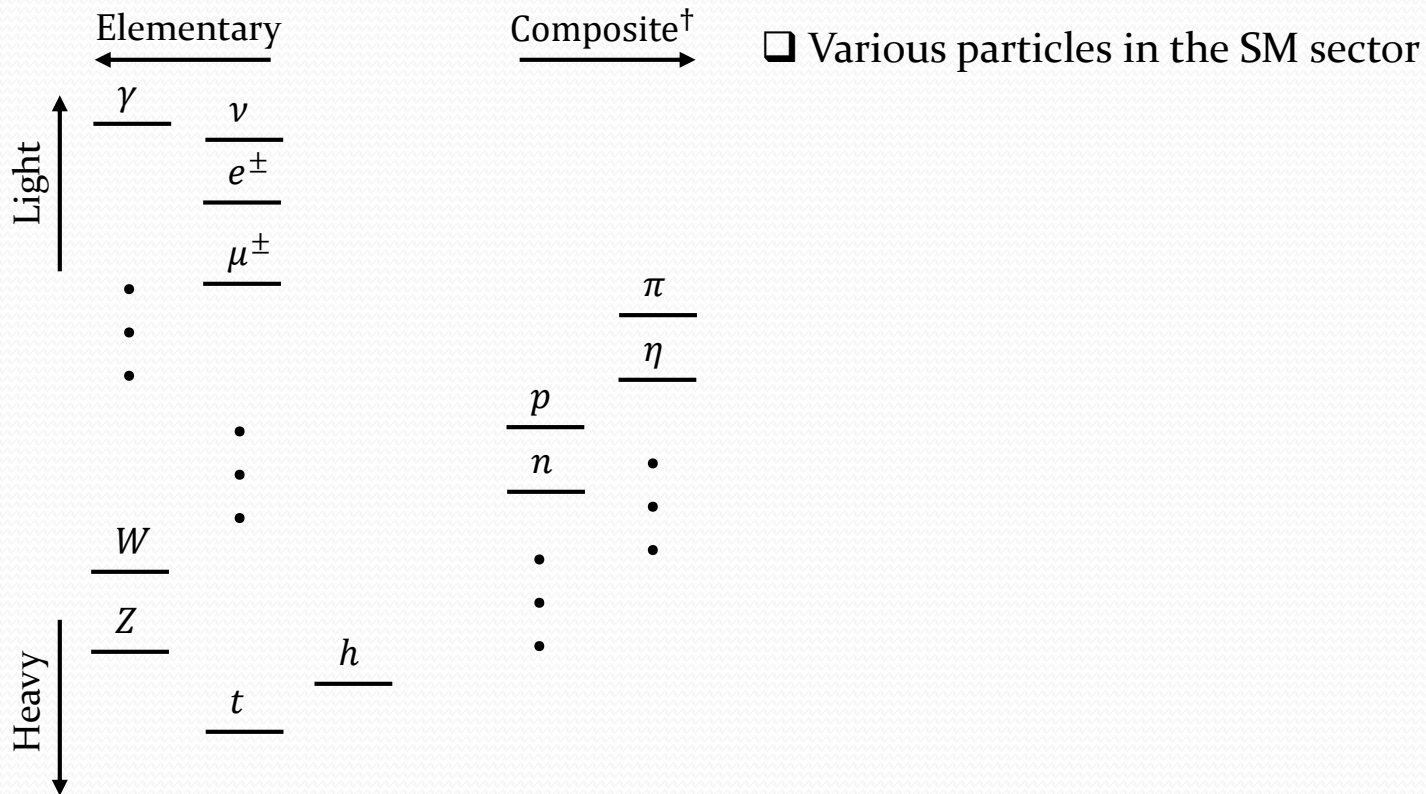
[ATLAS mono-jet search (2015)]



[CMS mono-photon search (2014)]

# “Non-minimal” Dark Sector

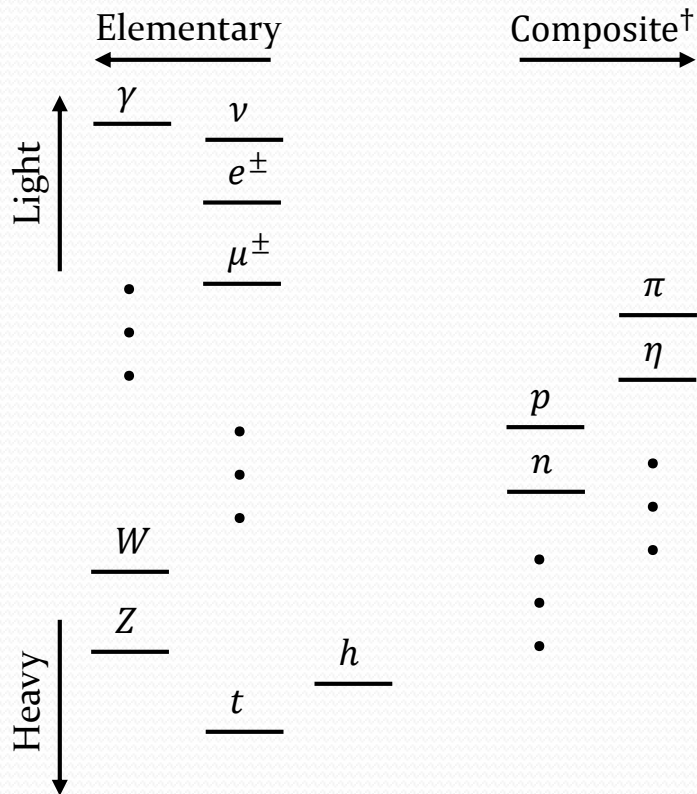
## ● Why flavorful?



†: here meaning the particles made of elementary ones

# “Non-minimal” Dark Sector

## ● Why flavorful?



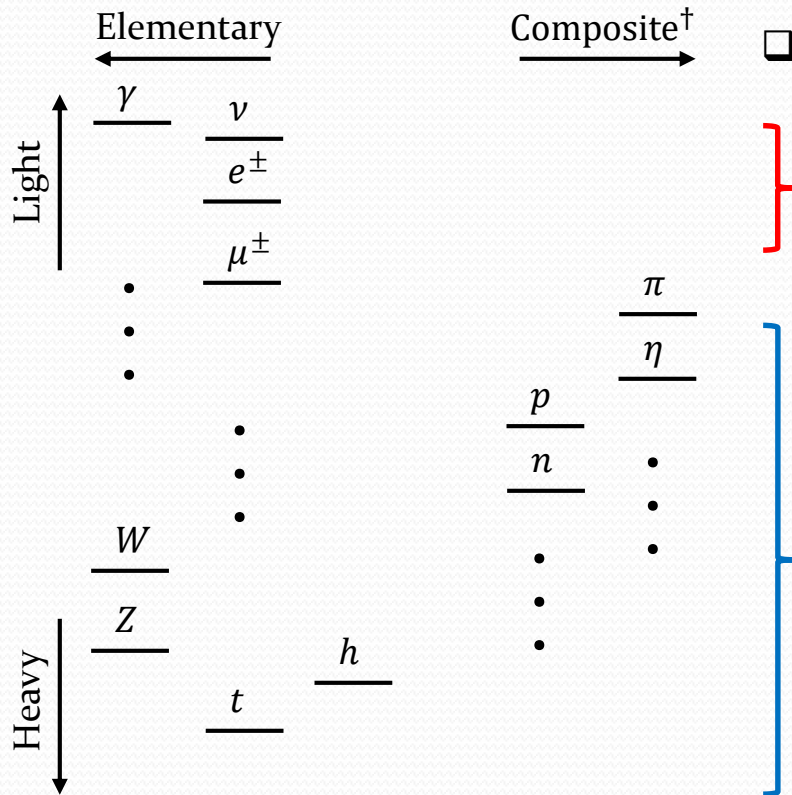
□ Various particles in the SM sector

✓ **Multiple stable particles** → **interesting physics** from other stable members which are **not difficult to detect** albeit not dominant



# “Non-minimal” Dark Sector

## ● Why flavorful?



□ Various particles in the SM sector

✓ **Multiple stable particles** → **interesting physics** from other stable members which are **not difficult to detect** albeit not dominant (proton is dominant in the visible sector)

✓ **Many heavier (unstable) states** → **interesting signatures/phenomenology** stemming from their decays (e.g., at lepton/hadron colliders)

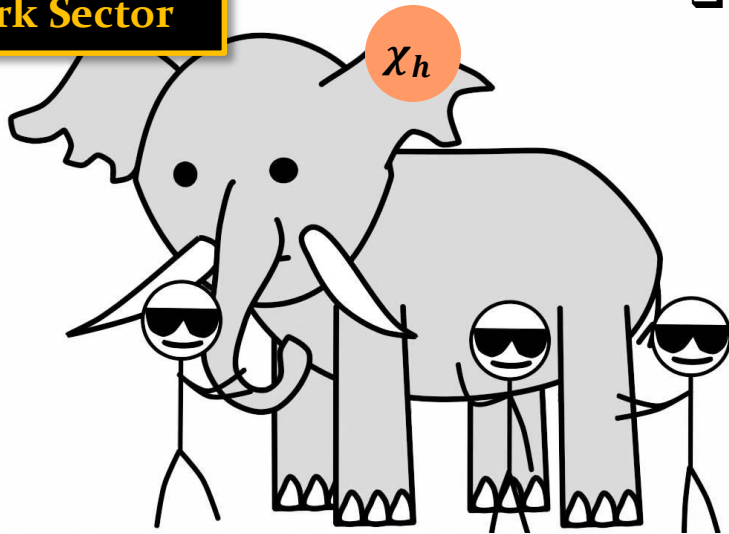
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# “Non-minimal” Dark-sector Scenarios

- In what sense?

**Dark Sector**

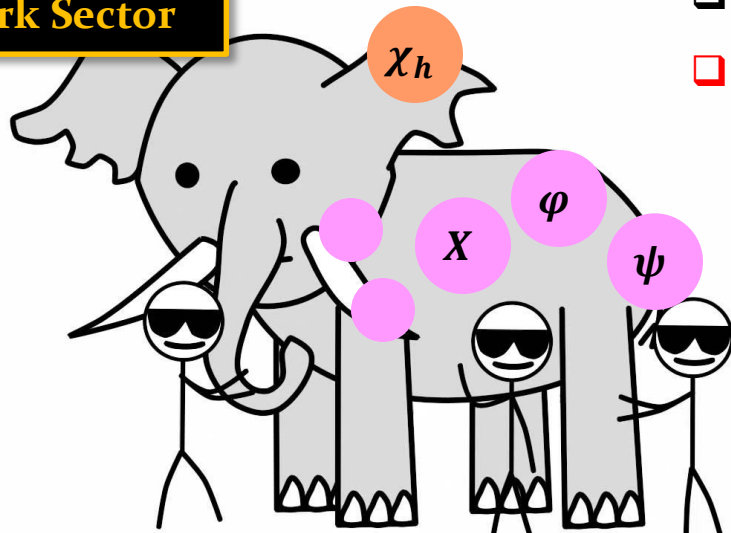
- $\chi_h$ : dominant relic (as in the minimal setup)



# “Non-minimal” Dark-sector Scenarios

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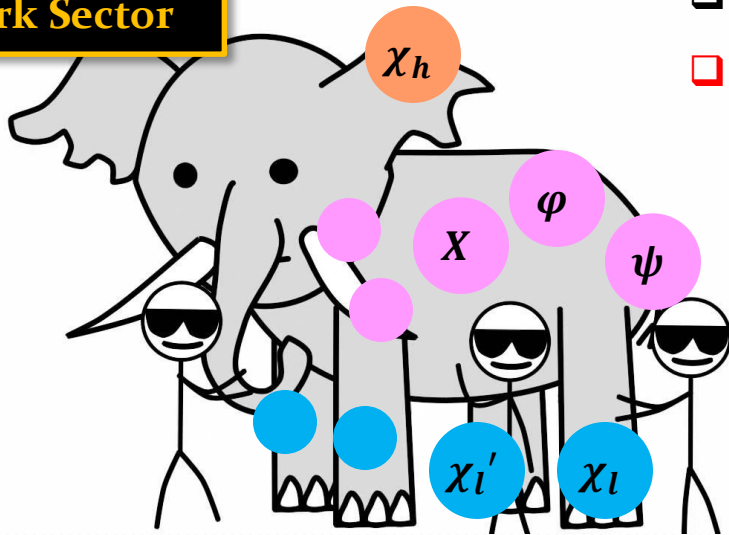


- $\chi_h$ : dominant relic (as in the minimal setup)
- **More members** in the dark sector
  - ✓ Unstable members, say  $\psi$ ,  $\varphi$ ,  $X$ , ... (e.g., cosmic ray excess interpretations [DK and J.-C. Park (2015)])

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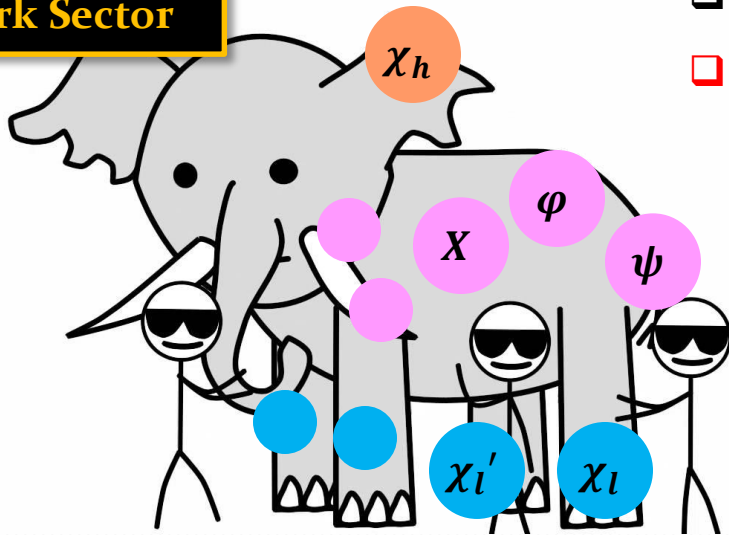


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  - ✓ More dark matter species, say  $\chi_l, \chi_l'$  ... (e.g., dynamical dark matter models [K. Dienes and B. Thomas, (2011)])

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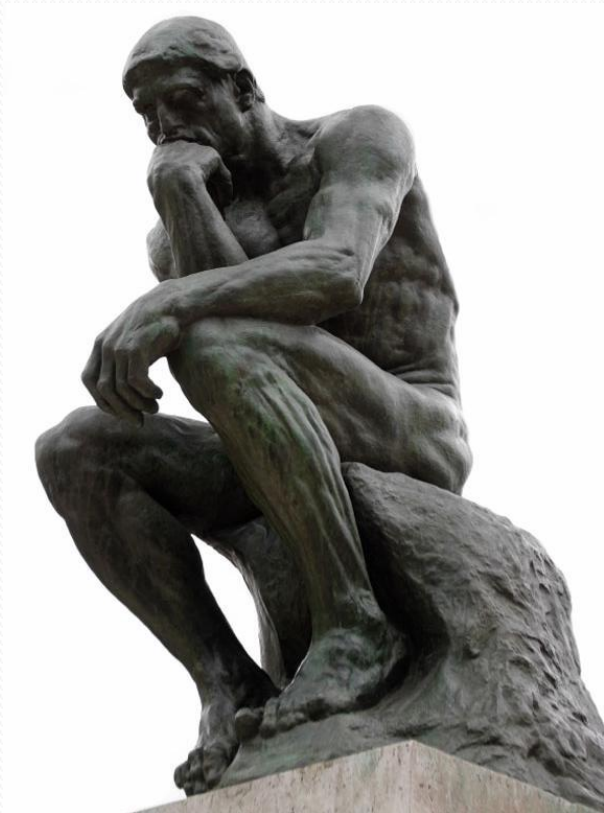
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## □ Rising interest

- ❖ Boosted dark matter scenarios [K. Agashe et al., (2014); K. Kong, G. Mohlabeng, J.-C. Park (2014)]
- ❖ Assisted freeze-out mechanism [G. Belanger and J.-C. Park (2011)]

# “Non-conventional” Implications?

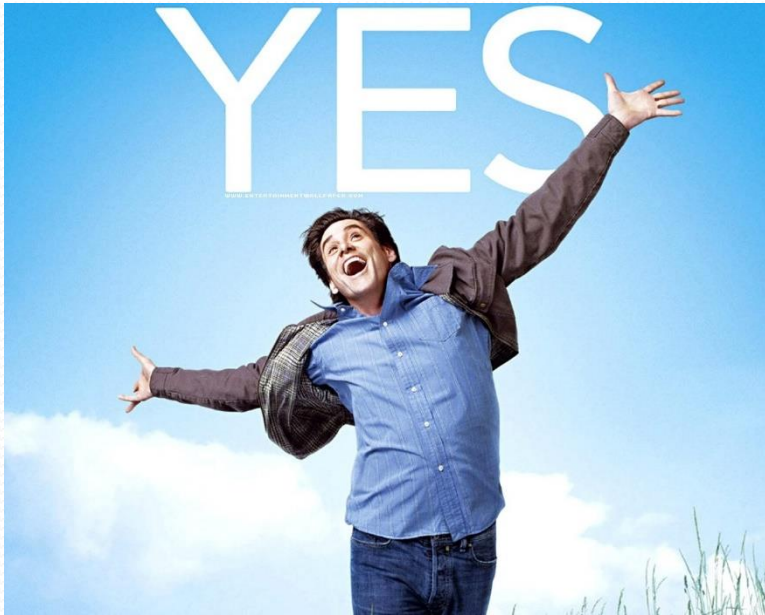
- Big question



- Existence of more members in the dark sector  
→ are there any **non-trivial/non-conventional implications not available in the minimal setup?**

# “Non-conventional” Implications!

## ● Big question



- ❑ Existence of more members in the dark sector  
→ are there any **non-trivial/non-conventional implications not available in the minimal setup?**
  - ✓ **New dark matter search strategies: dark matter “colliders”** (see Seodong’s talk for details) [DK, J.-C. Park and S. Shin (2016)]
  - ✓ **New interpretations for existing/future data: dark matter “transporting” mechanism** (see Jong-Chul’s talk for details) [DK, J.-C. Park and S. Shin (2017)]

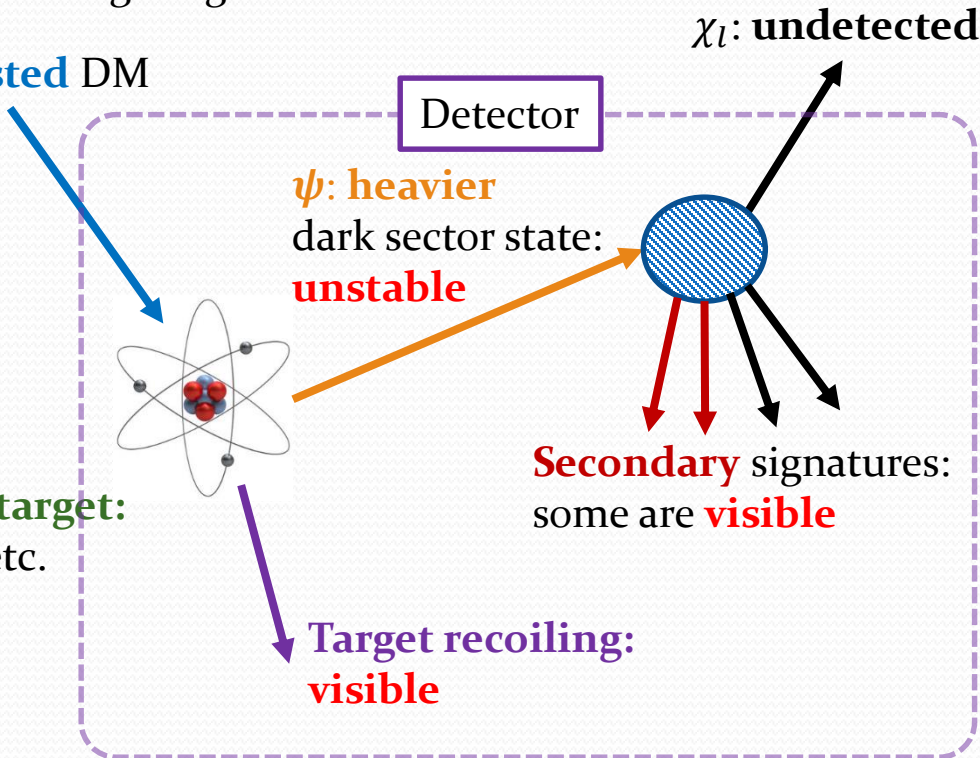
# Dark Matter “Colliders”

- **Basic idea** [DK, J.-C. Park and S. Shin (2016)]

□ We are imagining the situation in which

$\chi_l$ : **boosted** DM

**Fixed target:**  
 $e^-$ ,  $p$ , etc.





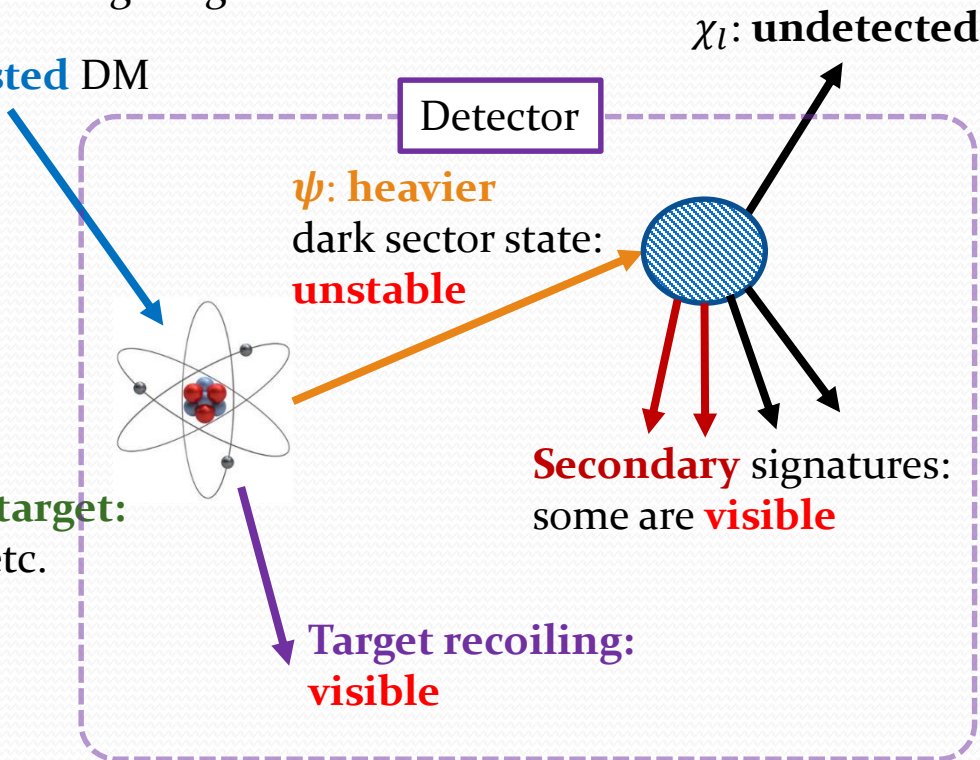
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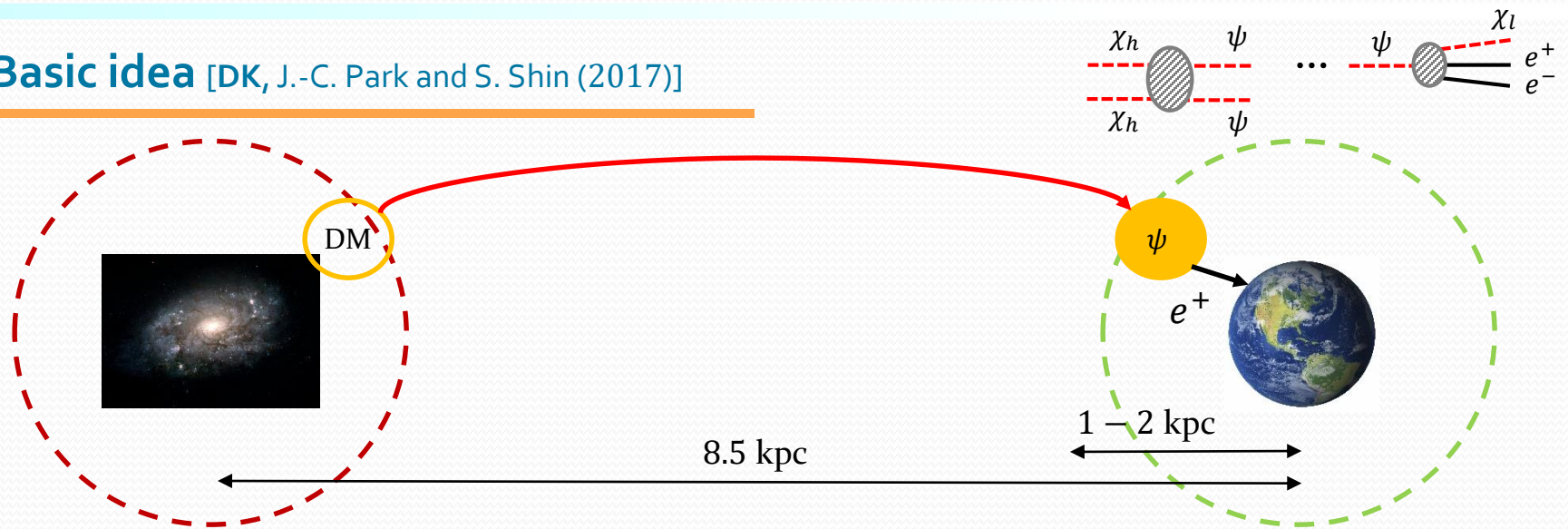
**Fixed target:**  
 $e^-$ ,  $p$ , etc.



- ✓ Probing **heavier dark-sector states**
- ✓ Target recoil (like in typical DM direct detection exp.) + secondary visible signatures  $\Rightarrow$  **more** handles, (relatively) **background-free**
- ✓ **Complementary** to standard DM direct searches
- ✓ **Boosted DM sources** needed: BDM scenarios, fixed target experiments, etc.

# Dark Matter “Transporting” Mechanism

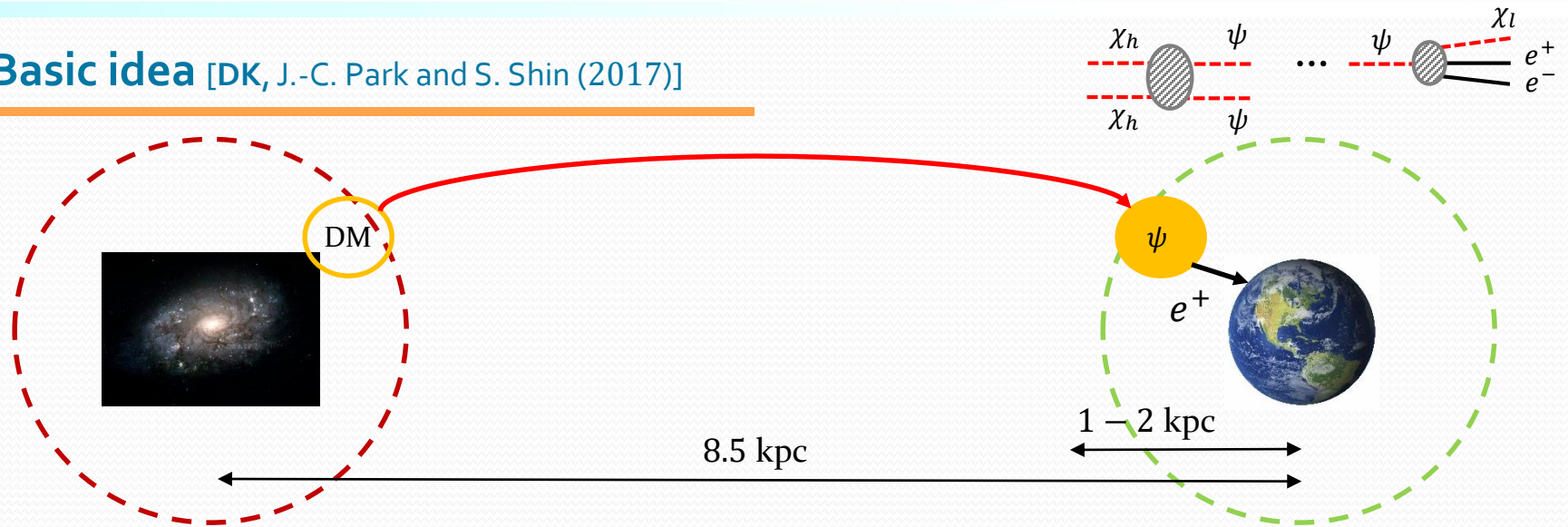
- **Basic idea** [DK, J.-C. Park and S. Shin (2017)]



- DM “transport” mechanism to explain cosmic positron excesses

# Dark Matter “Transporting” Mechanism

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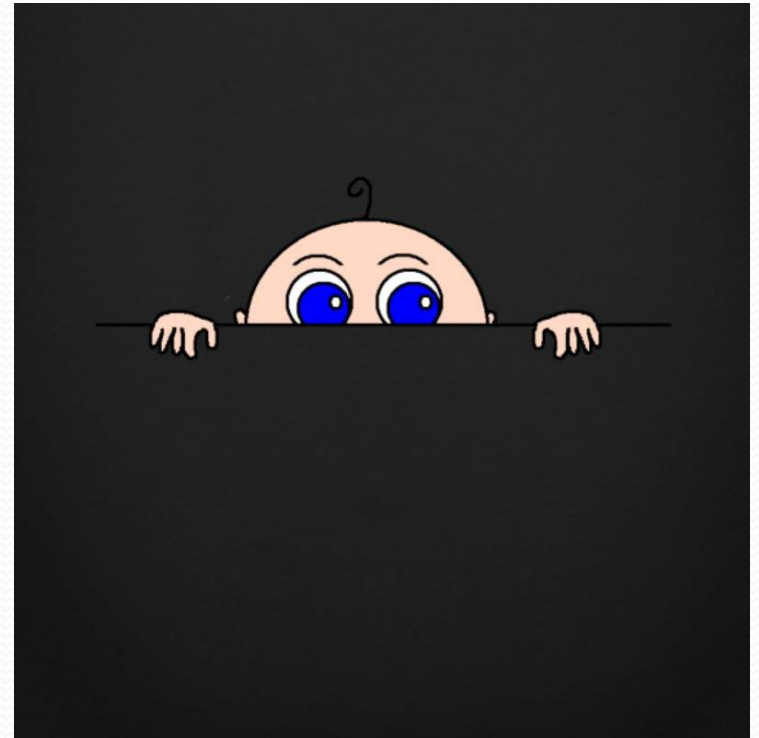


- ❑ DM “transport” mechanism to explain cosmic positron excesses
- ❑ Positron flux  $\Phi \sim \rho^2 \langle \sigma v \rangle$  : effectively enhancing **density** itself (a big DM clump at the GC)
  - ✓ “**Transporting**” (effectively) DM at the GC to the vicinity of the Earth via DM “proxy”  $\psi$
  - ✓ Consistent with various astrophysical and cosmological constraints (see Jong-Chul’s talk for more detailed discussion)

# Conclusions

## ● Summary

- ❑ The more, the messier? The more, the merrier! ⇒ Peeping into the dark sector through “flavorful” scenarios
- ❑ Rising interest in non-minimal dark sector physics
- ❑ Non-minimal/flavorful dark sector scenarios may provide spectacular phenomenology different from that in the minimal setup.
  - ✓ **New dark matter search strategies: dark matter “collider”** (see Seodong’s talk for detail)
  - ✓ **New mechanism for positron excesses: dark matter “transporting” mechanism** (see Jong-Chul’s talk for detail)



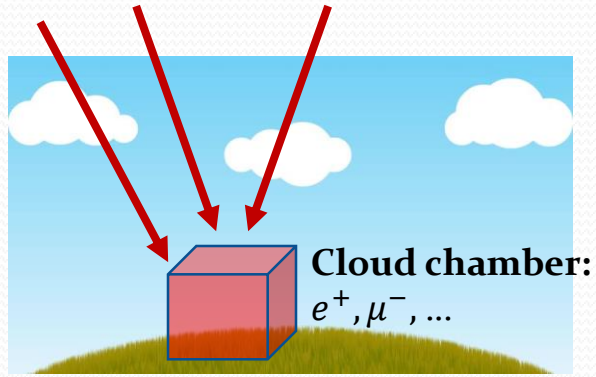


thank you!

**Don't leave... more exploration to dark world!**

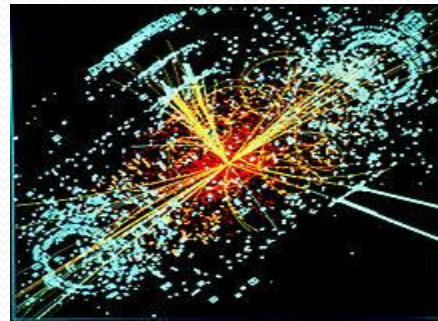
# Dark Matter “Colliders”

~5% visible sector



Passive searches

Collider: **controlled** environment

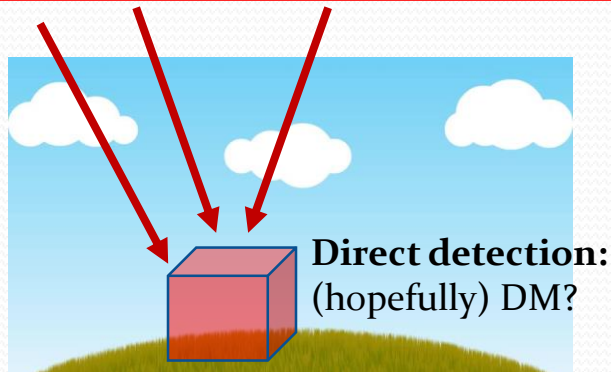


Conventional colliders

- Head-on collision of light SM-sector (stable) particles
- to produce heavier states
- and study resulting pheno.

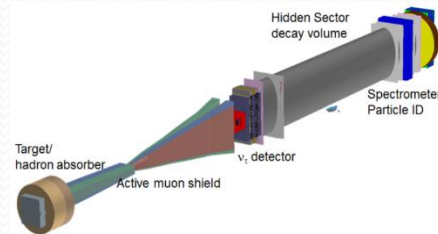
Active searches

~25% dark sector



Passive searches

DM “Collider” (e.g. fixed target exp.)  
: **controlled** environment



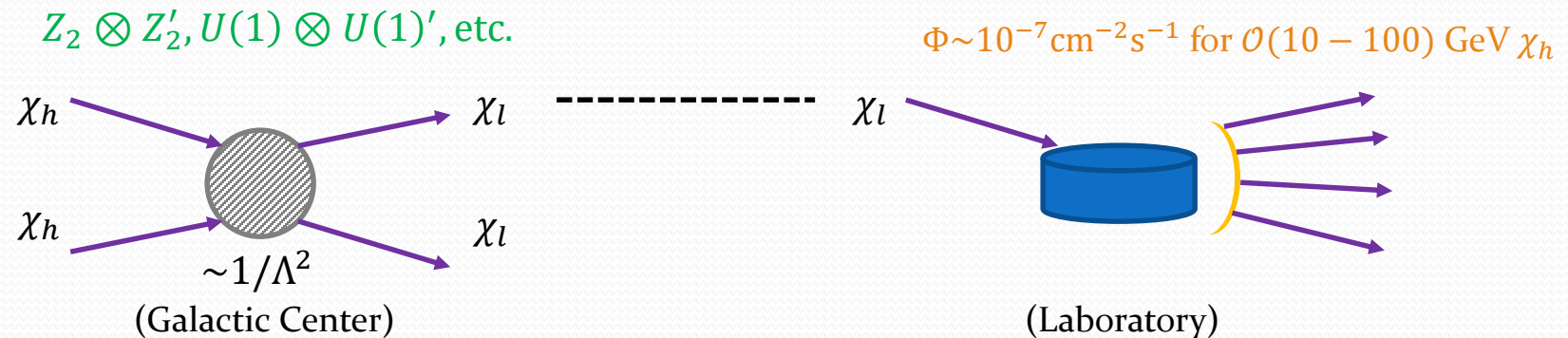
Dark matter colliders

- Collision of **light dark-sector (stable)** particles onto a target
- to produce **heavier dark-sector** states
- and study resulting pheno.

# A Boosted DM Source

## ● Boosted dark matter scenarios

- Boosted DM scenarios [K. Agashe et al., (2014); K. Kong, G. Mohlabeng, J.-C. Park (2014)]



- $\chi_h$ : heavier DM, dominant relic, non-relativistic, **not directly** communicating with SM
- $\chi_l$ : lighter DM, subdominant relic, **relativistic** at the current universe (non-relativistic at the early universe), **directly** communicating with SM
- Thermal relics determined by the Assisted Freeze-out mechanism [G. Belanger and J.-C. Park (2011)]