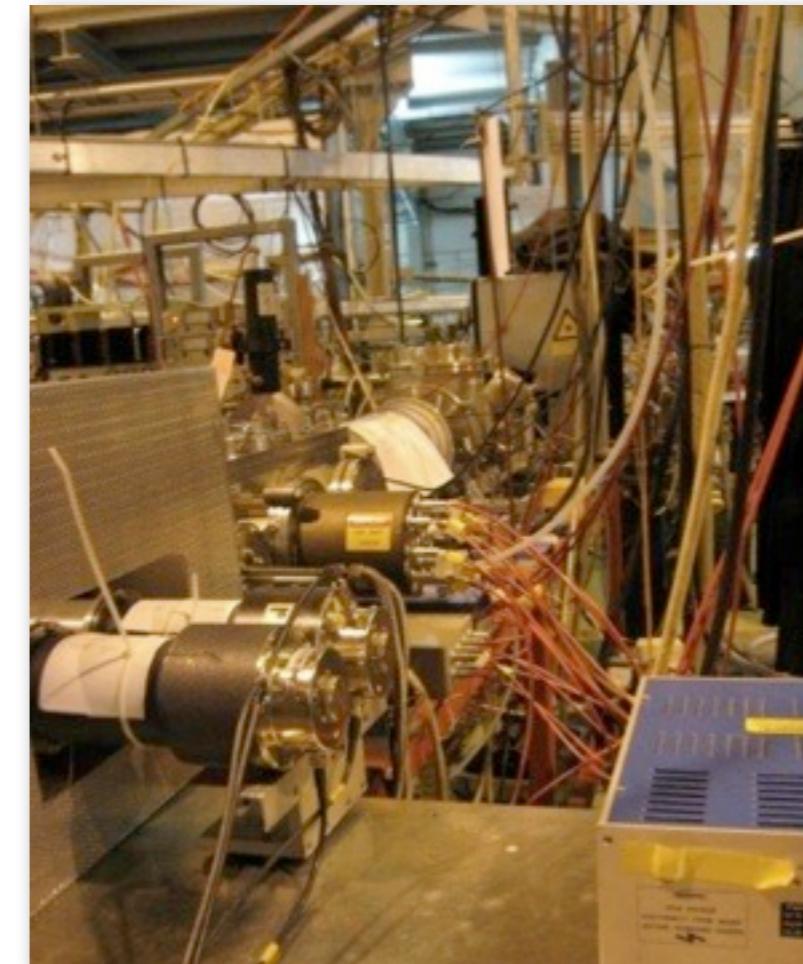
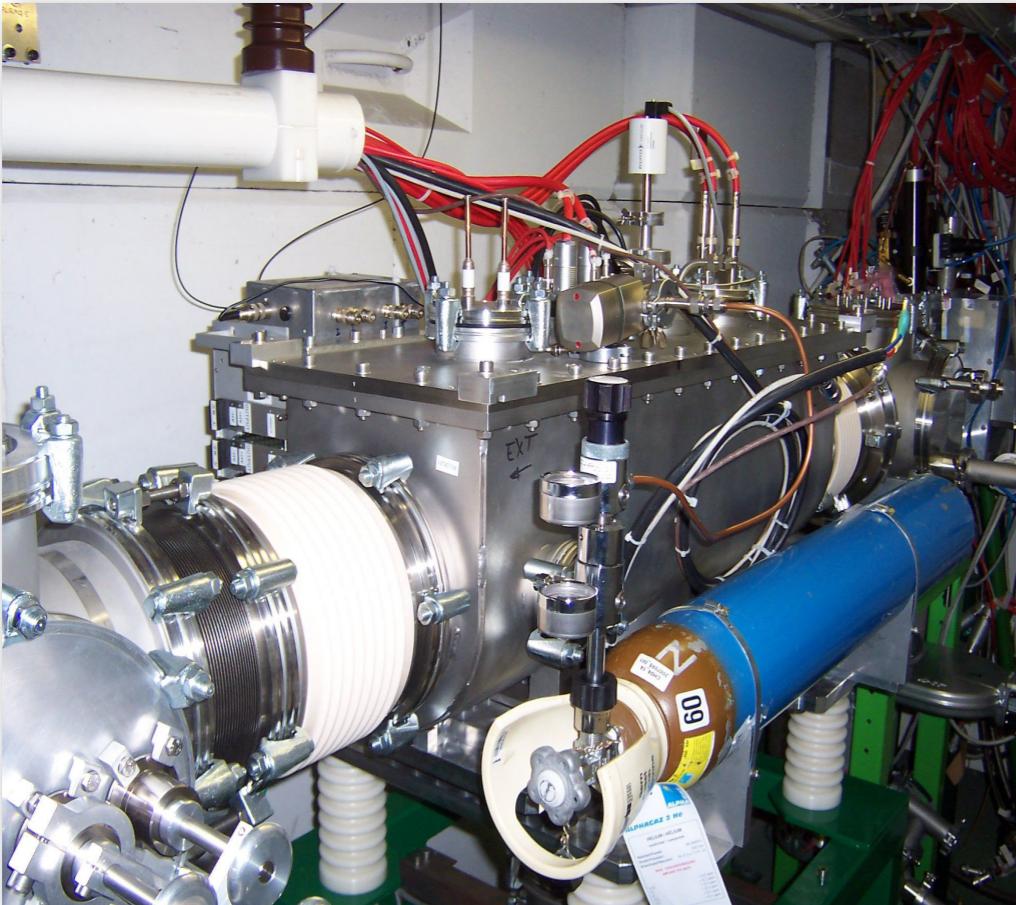
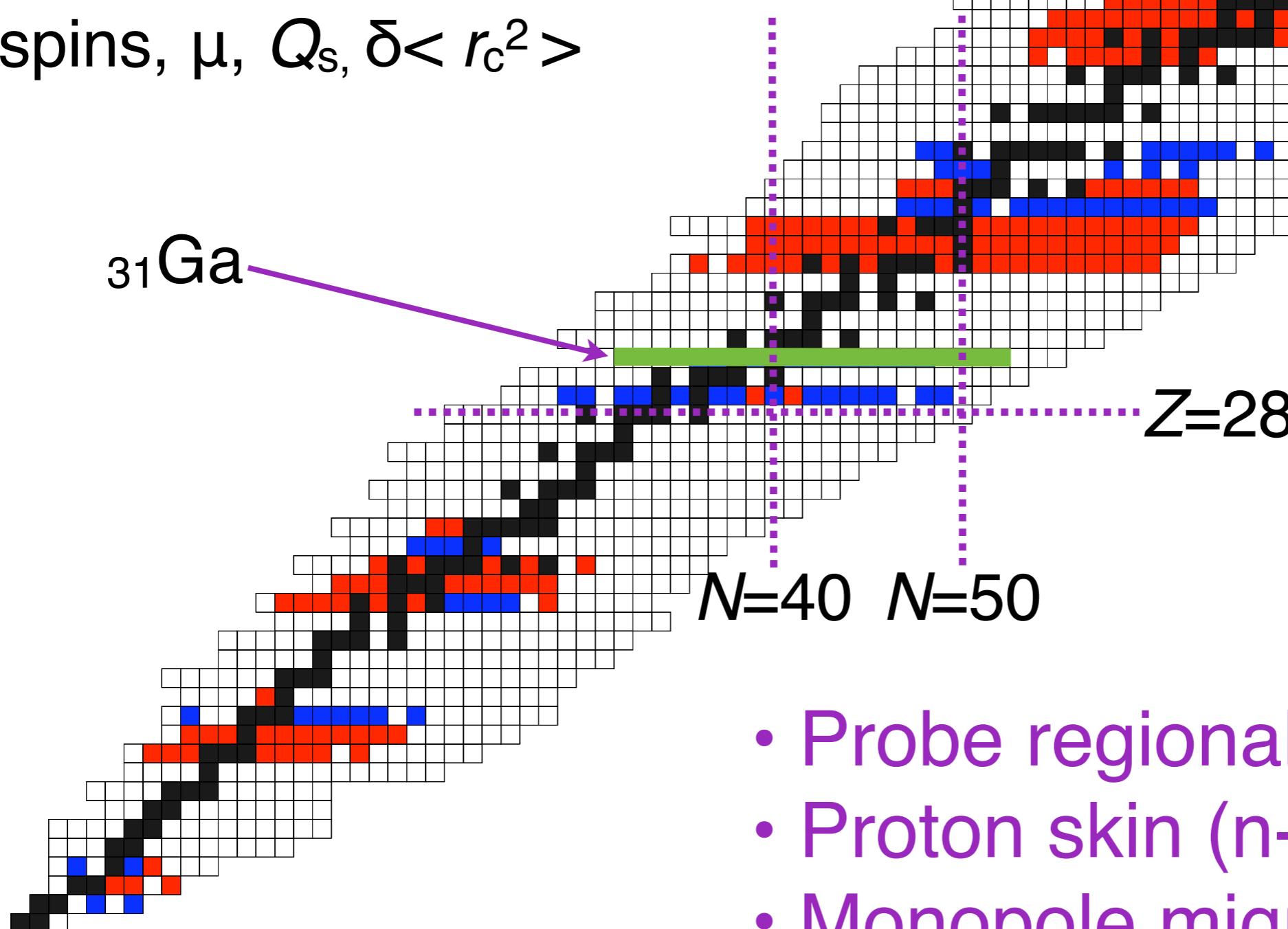


# Laser spectroscopy of gallium isotopes using ISCOOL



# Optical measurements in the region

→ spins,  $\mu$ ,  $Q_s$ ,  $\delta \langle r_c^2 \rangle$

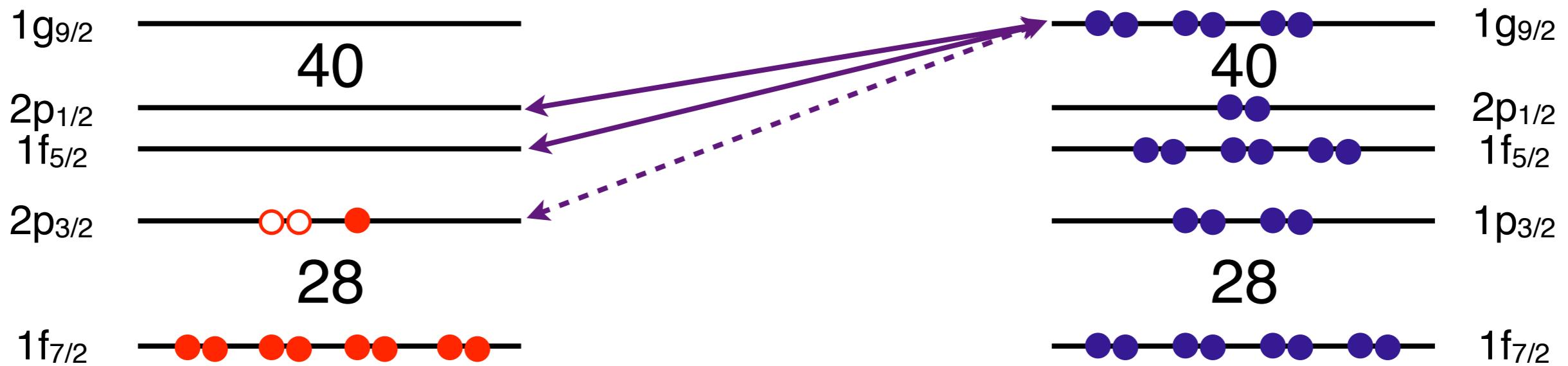


- Probe regional systematics
- Proton skin (n-def)
- Monopole migration (n-rich)

# Physics motivation (n-rich)

Otsuka (PRL 95 232502):-

Tensor force **attractive** between  $J=L+1/2$  and  $J=L-1/2$   
(otherwise **repulsive**)



Does  $5/2$  replace  $3/2$  as gs in Ga? When?

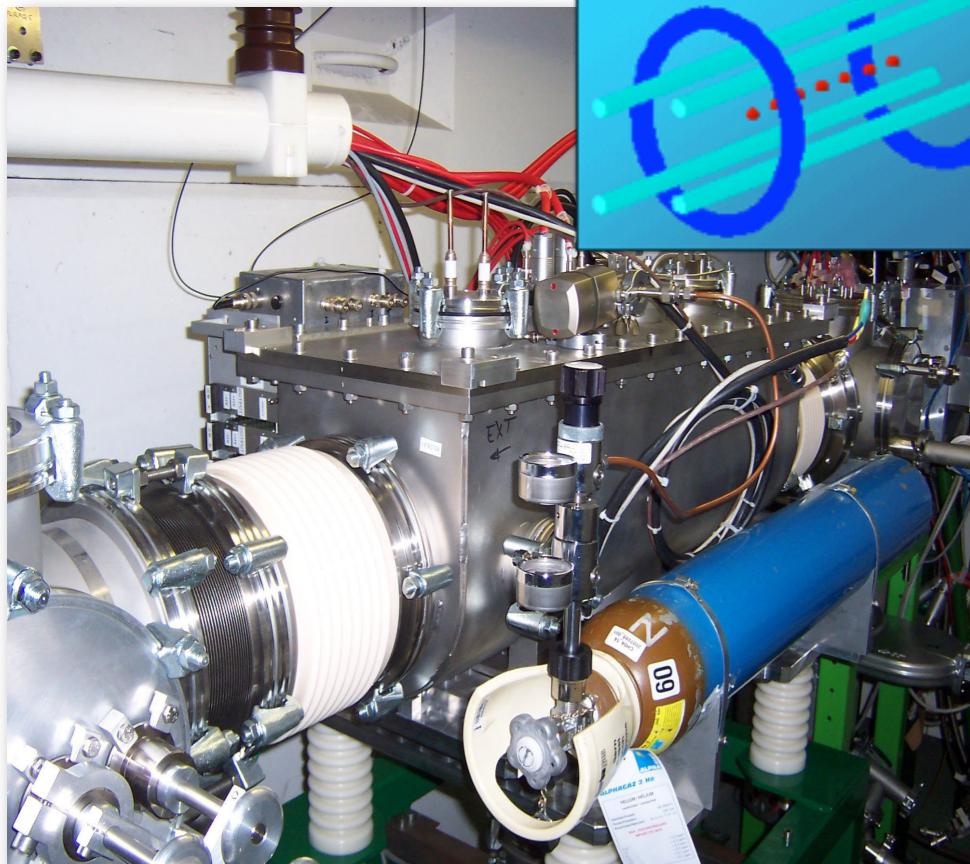
Use laser spectroscopy to measure the gs spins...

# Laser spectroscopy at ISOLDE

Target (Laser  
Ion Source)

Magnets

ISCOOL



(Gas filled RFQ)

1.4 GeV Protons

Proton Beam

GPS Target

HRS Target

HRS Separator

GPS Separator

GLM

GHM

Control Room

LA1

REX-ISOLDE

ASPIC

COLLAPS

COMPLIS

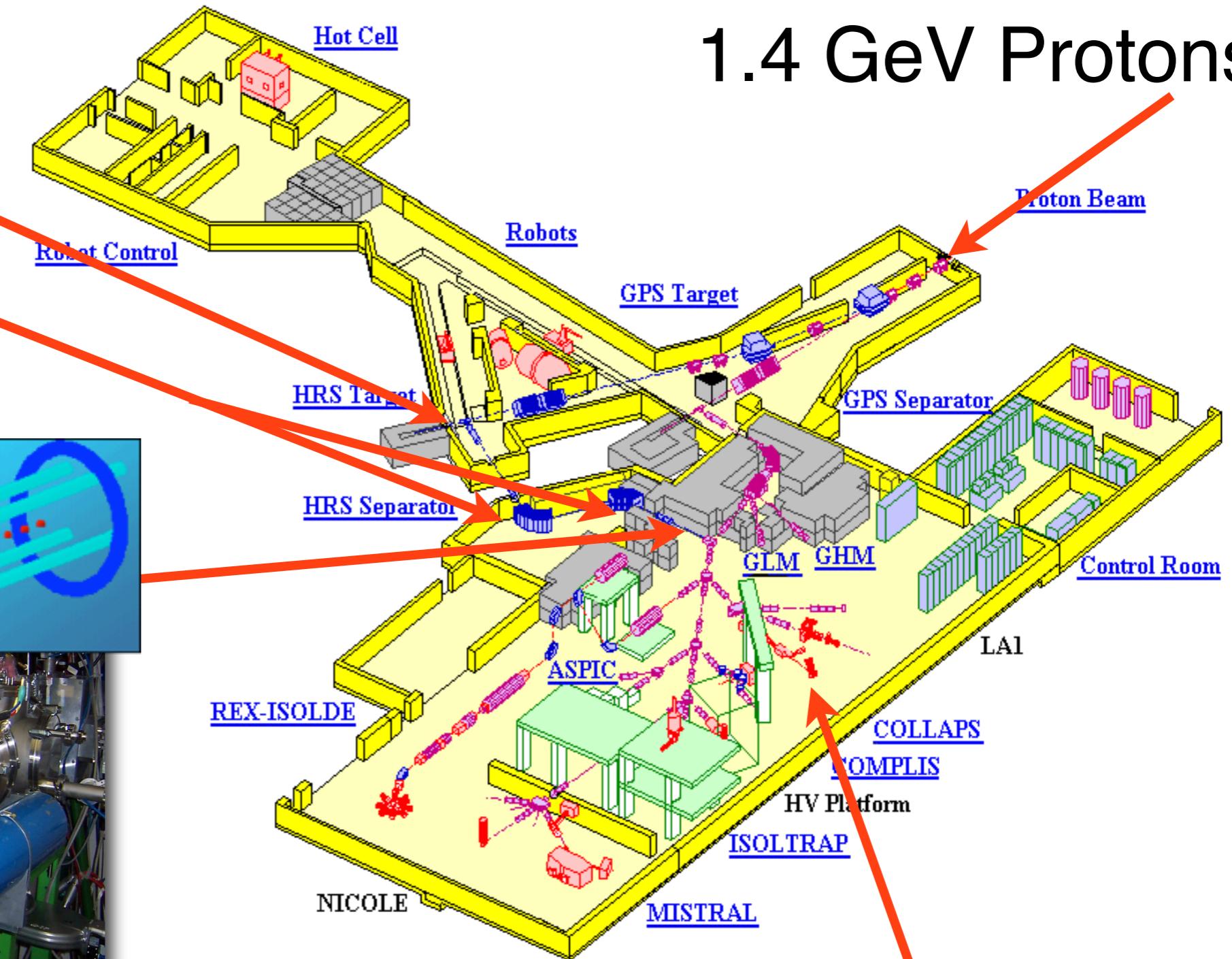
HV Platform

ISOLTRAP

NICOLE

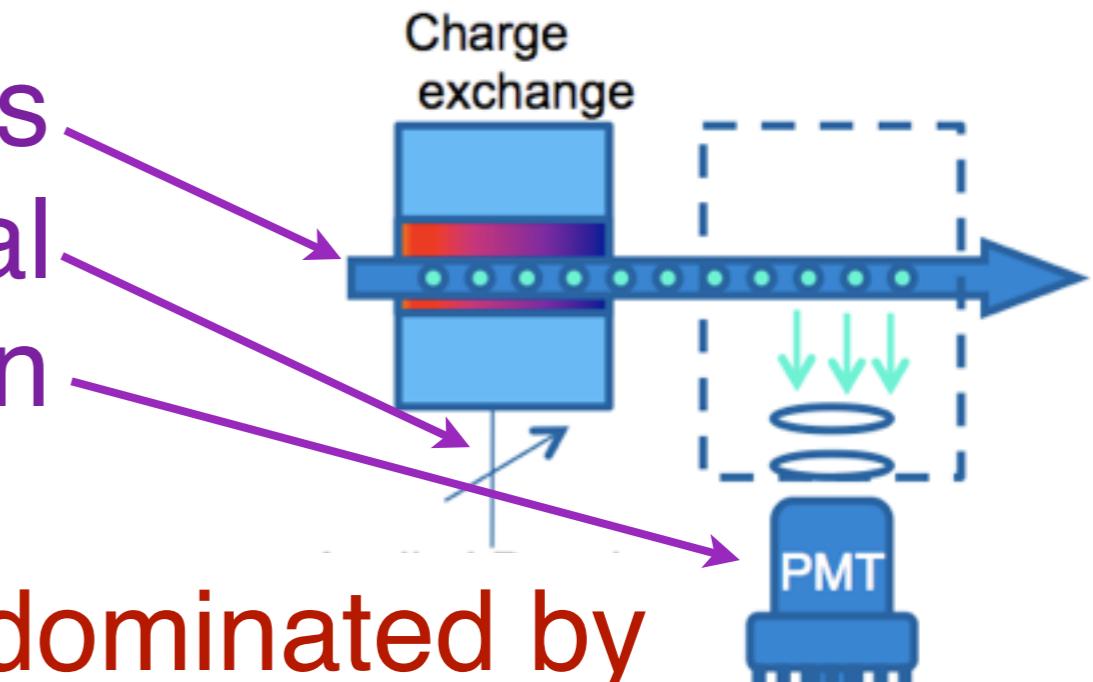
MISTRAL

Spectroscopy

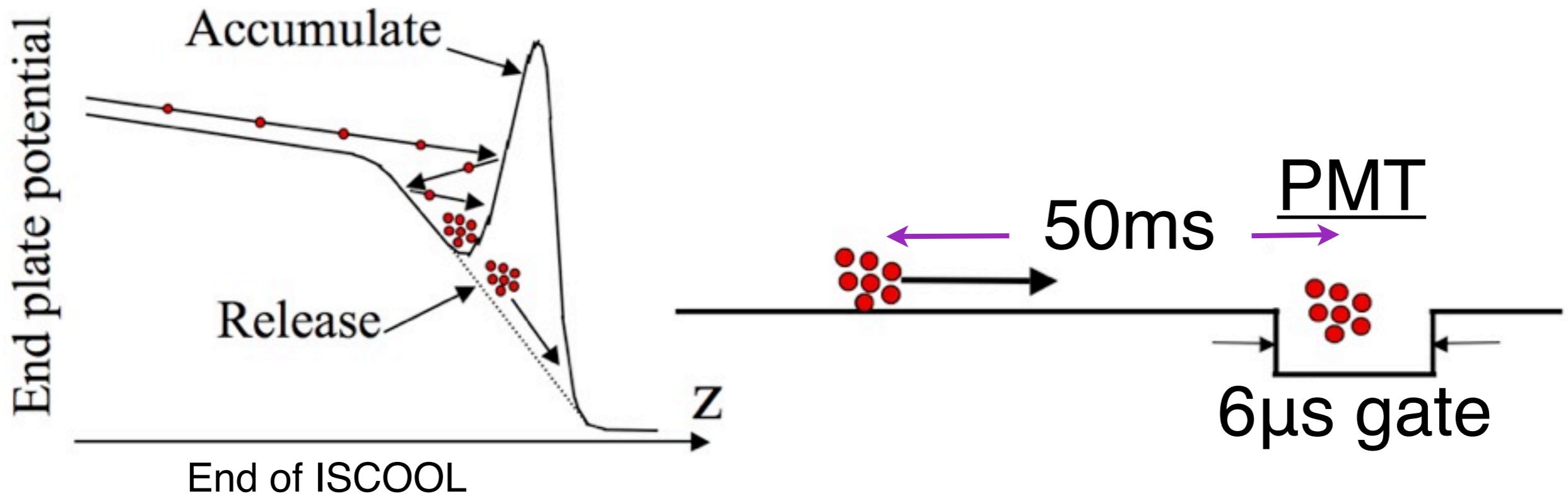


# Laser spectroscopy with ISCOOL

Collinear laser-atom beams  
Doppler tuning potential  
Photon detection

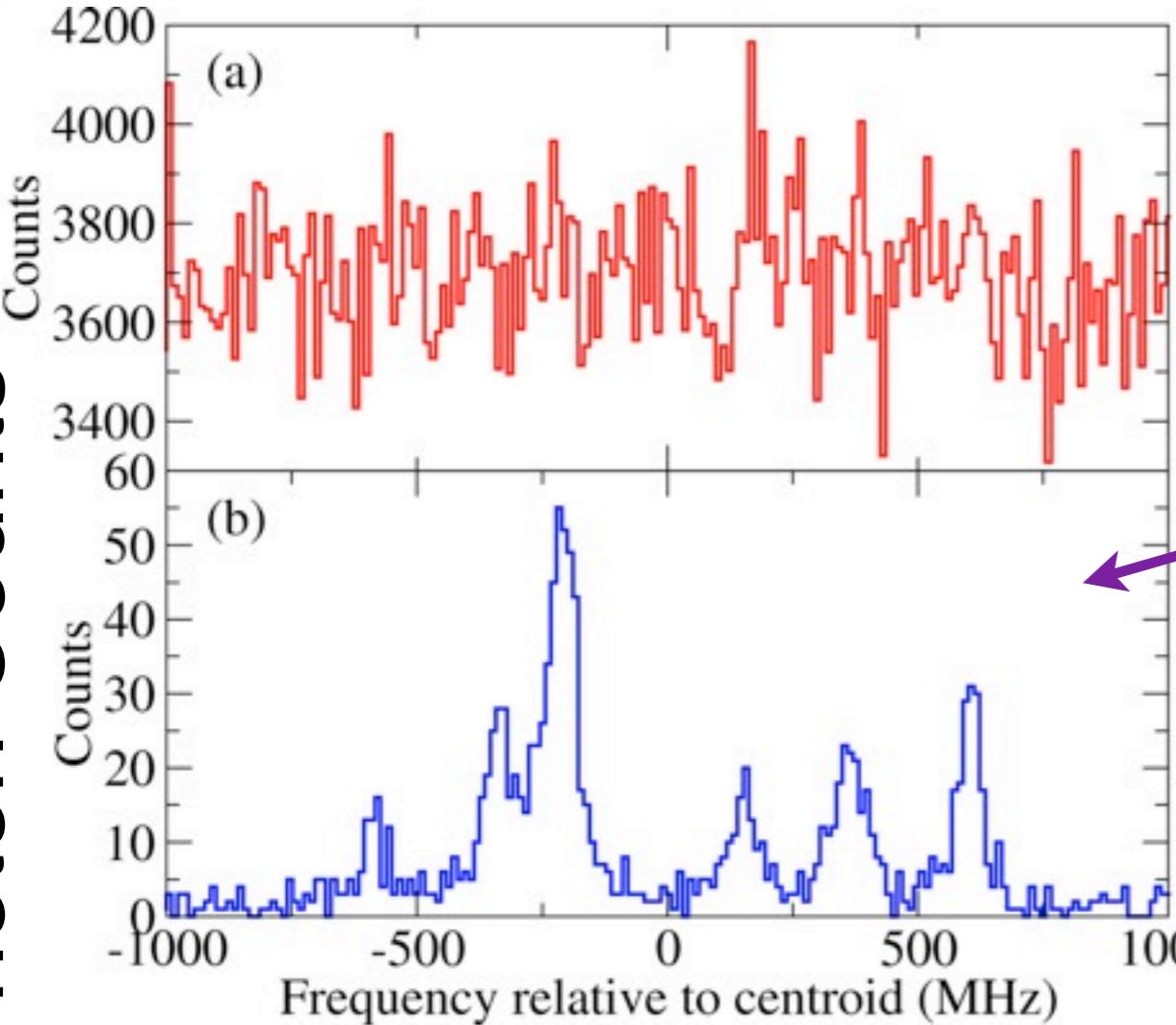


Photon background dominated by  
continuous laser scatter

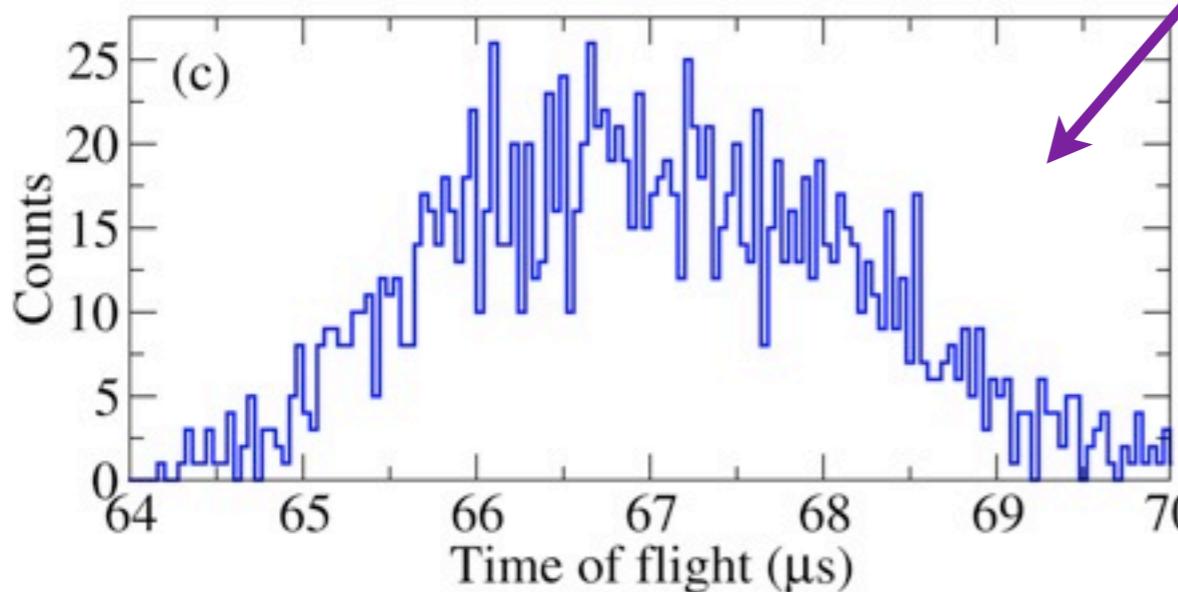


# Example spectrum - $^{76}\text{Ga}$

Photon Counts



Ungated



Gated (64 $\mu\text{s}$  - 70 $\mu\text{s}$ )

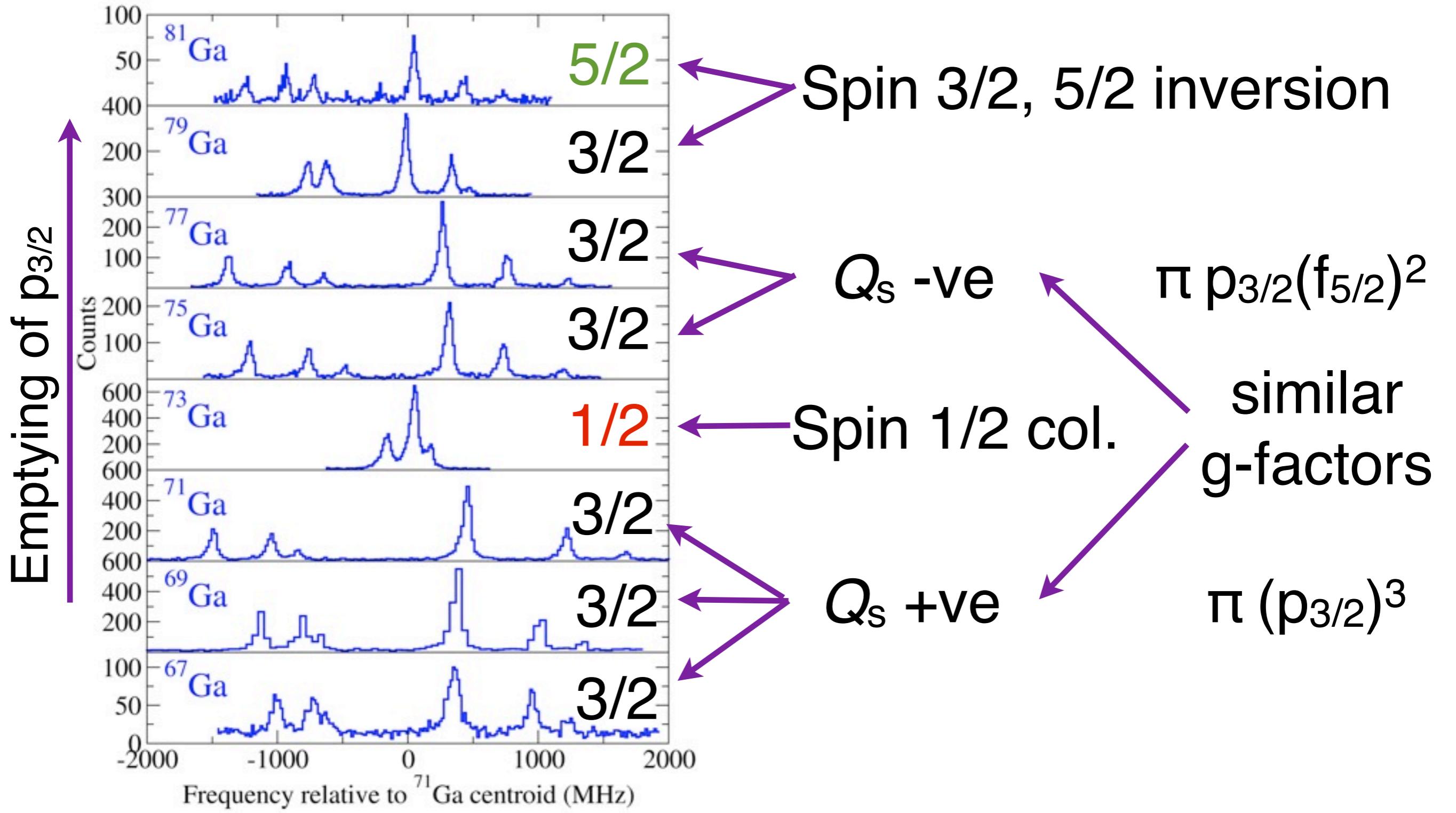
Time of flight  
(50ms accumulation)

Background suppression

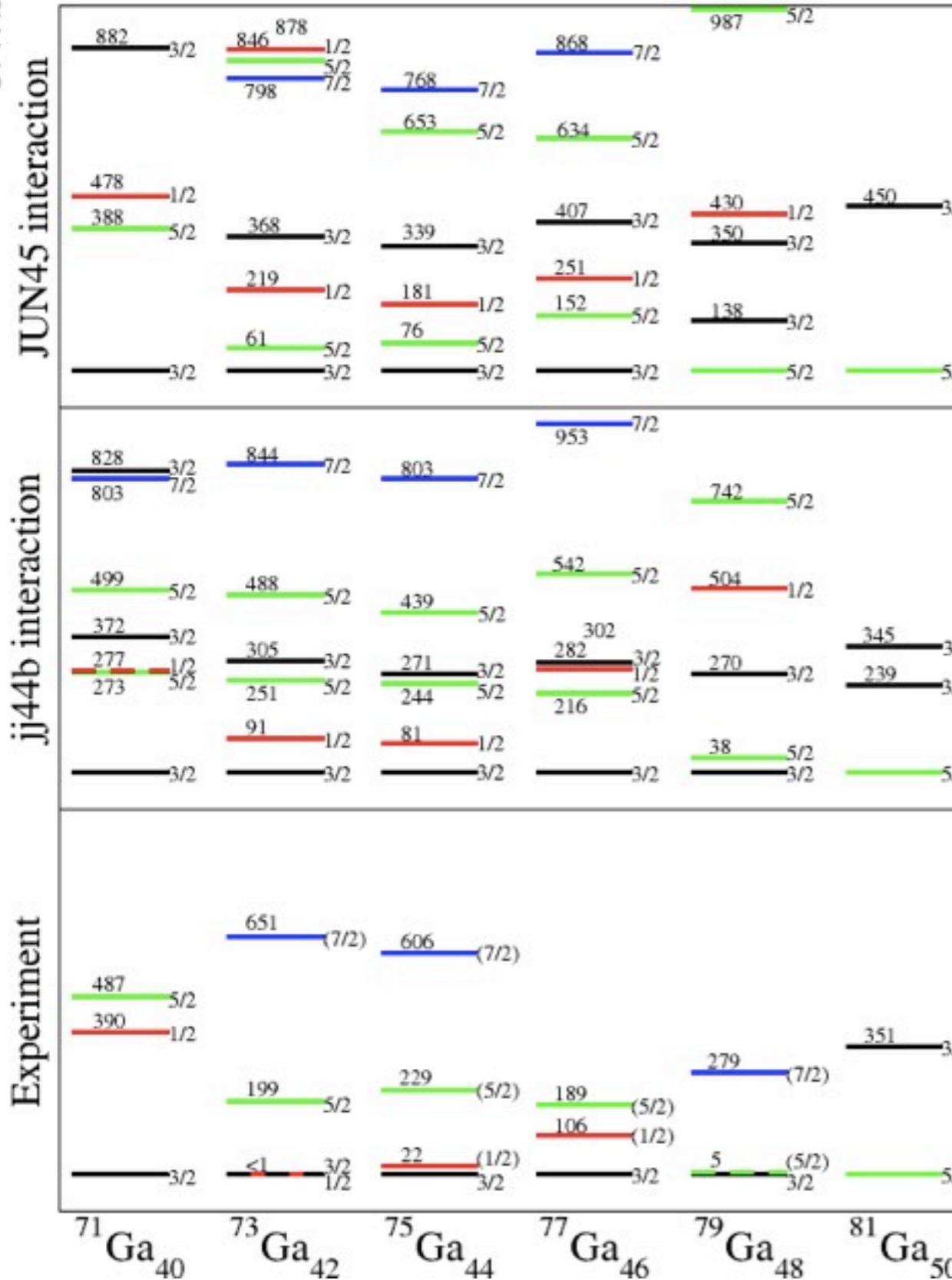
$$50\text{ms} / 6\mu\text{s} = \sim 10^4$$

# 417nm Ga I spectra

$826 \text{ cm}^{-1} 3d^{10}4s^24p \ ^2P_{3/2} \rightarrow 24789 \text{ cm}^{-1} 4s^25s \ ^2S_{1/2}$



# Theory - energy levels



Shell model calculations,  
2 effective interactions:-

- JUN45
- jj44b

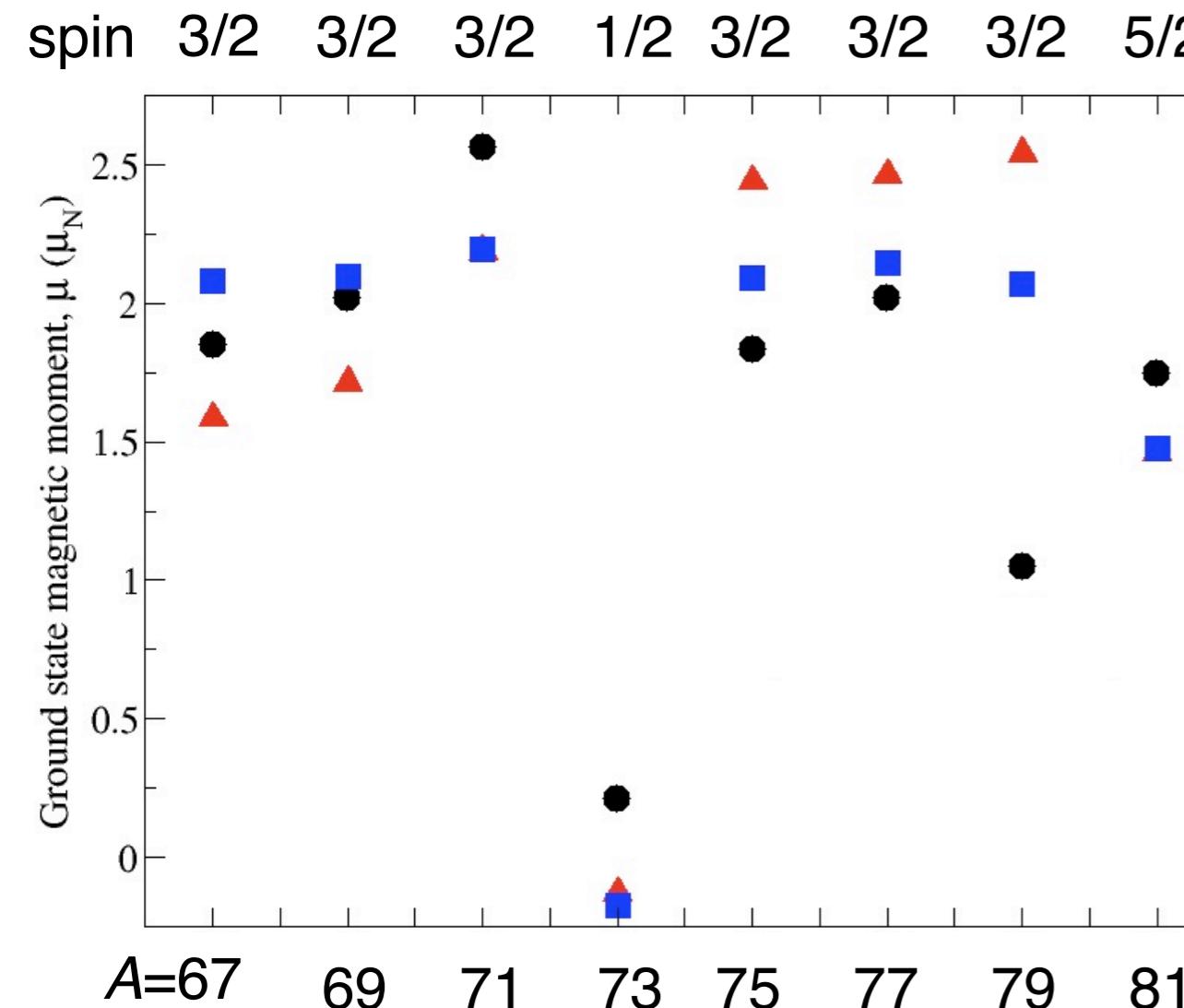
**1/2-, 3/2-, 5/2-, 7/2-**

jj44b - spin inversion

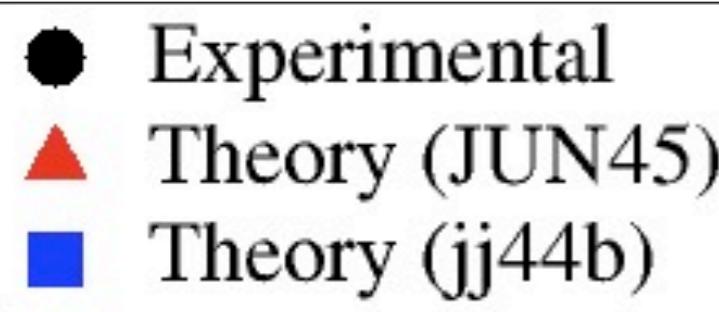
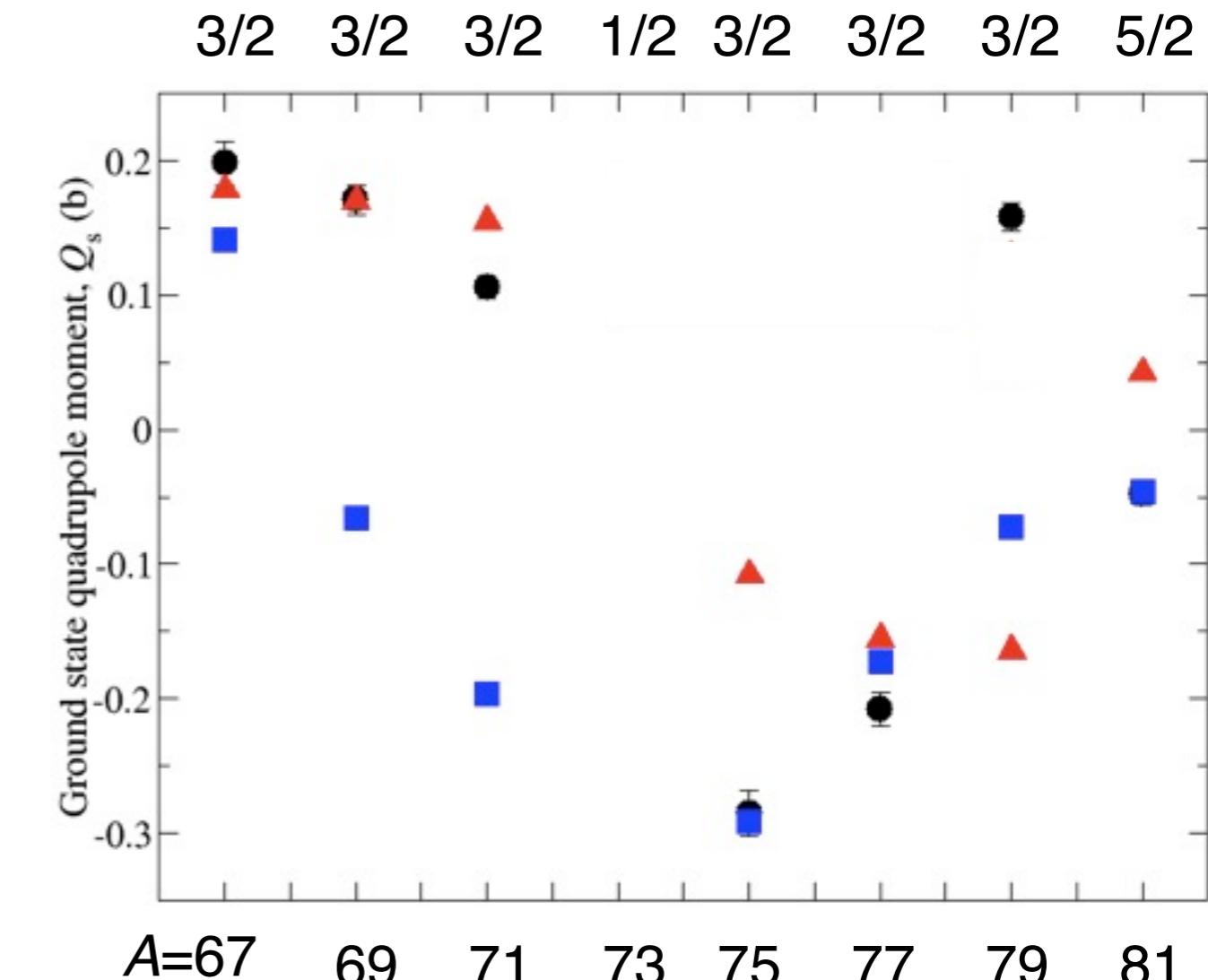
Fail to predict  $^{73}\text{Ga}$  ( $I=1/2$ )

# Theory - nuclear moments

## Magnetic dipole



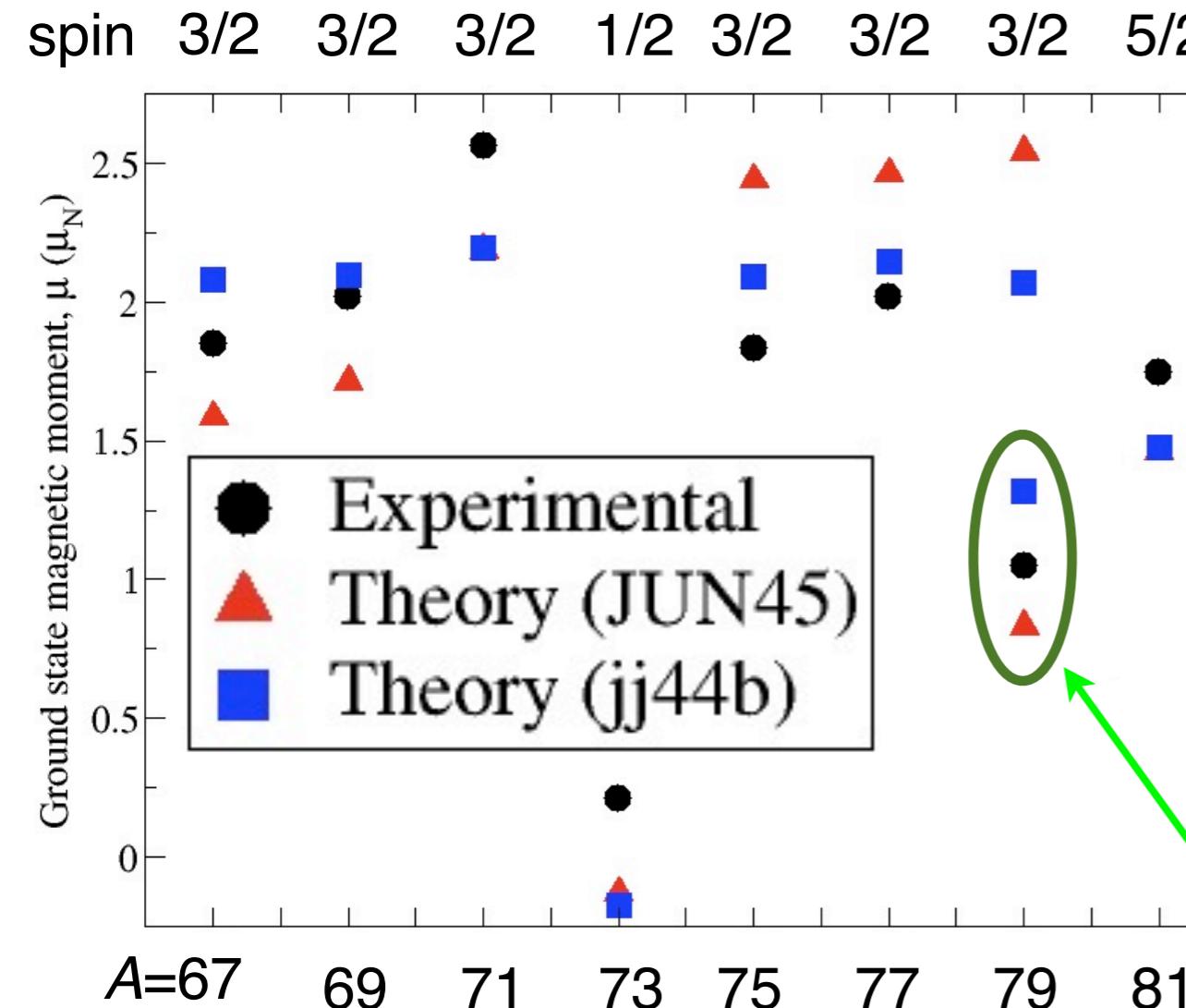
## Electric quadrupole



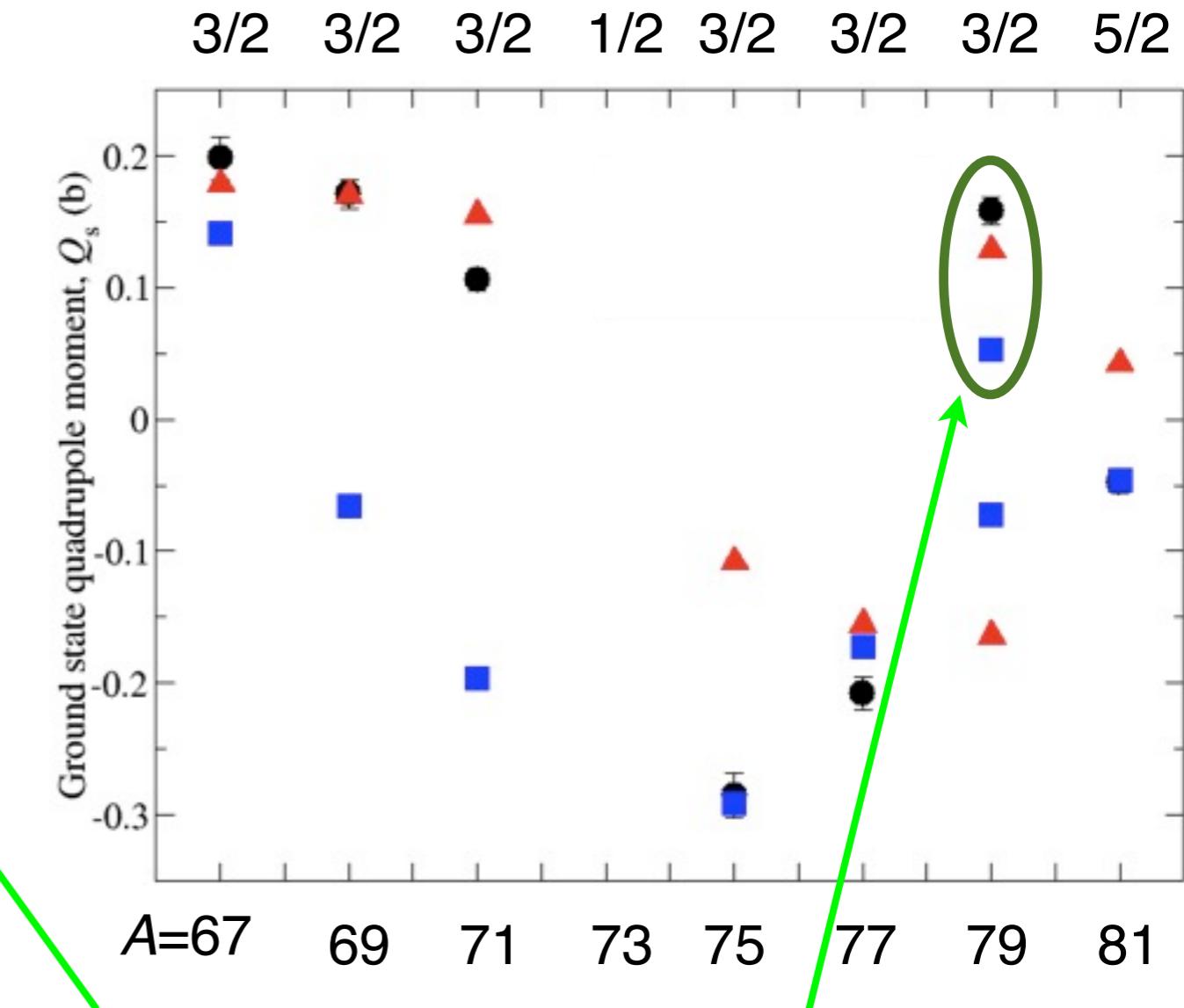
Poor match for  $^{79}\text{Ga}$

# Theory - nuclear moments

## Magnetic dipole



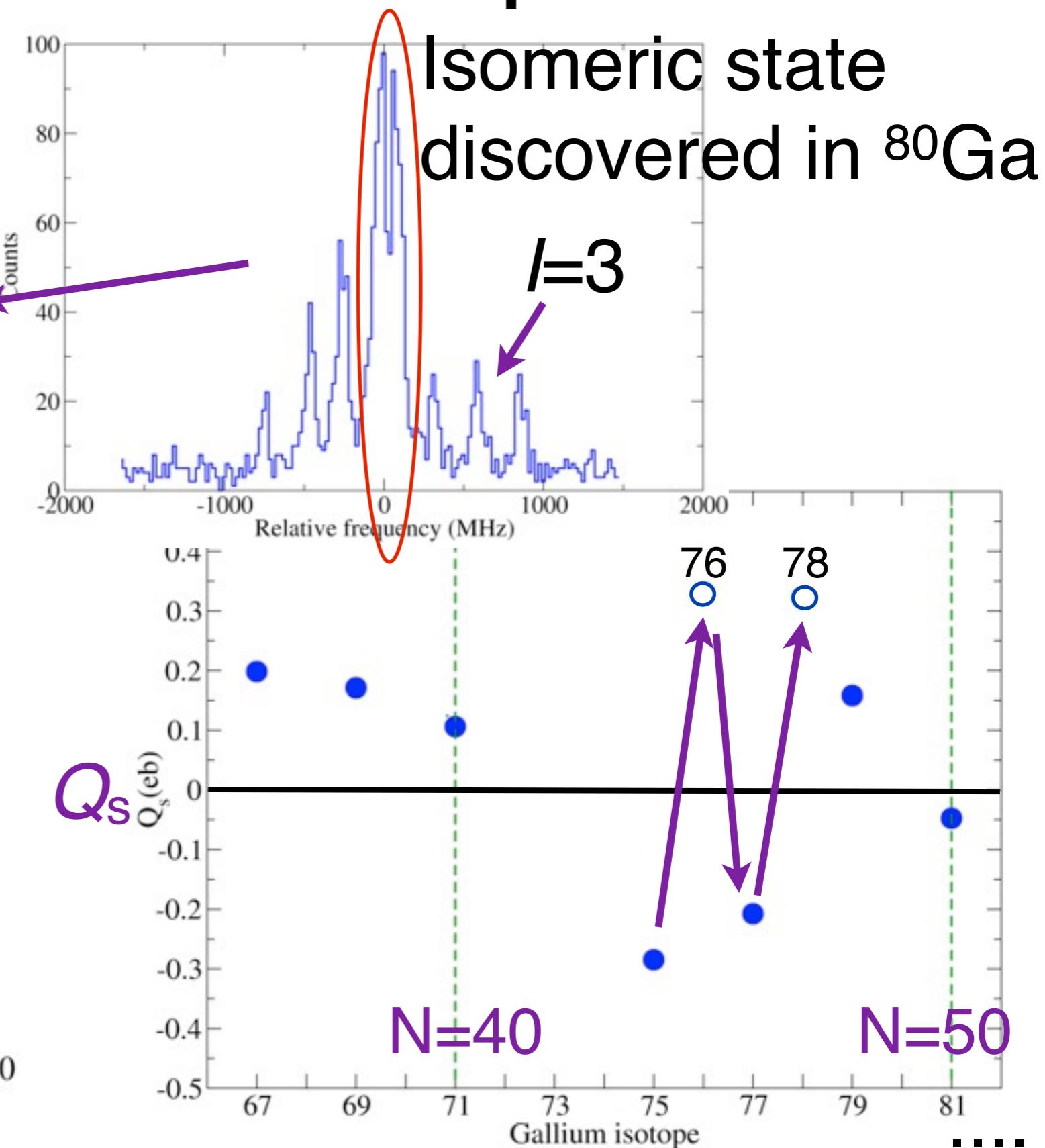
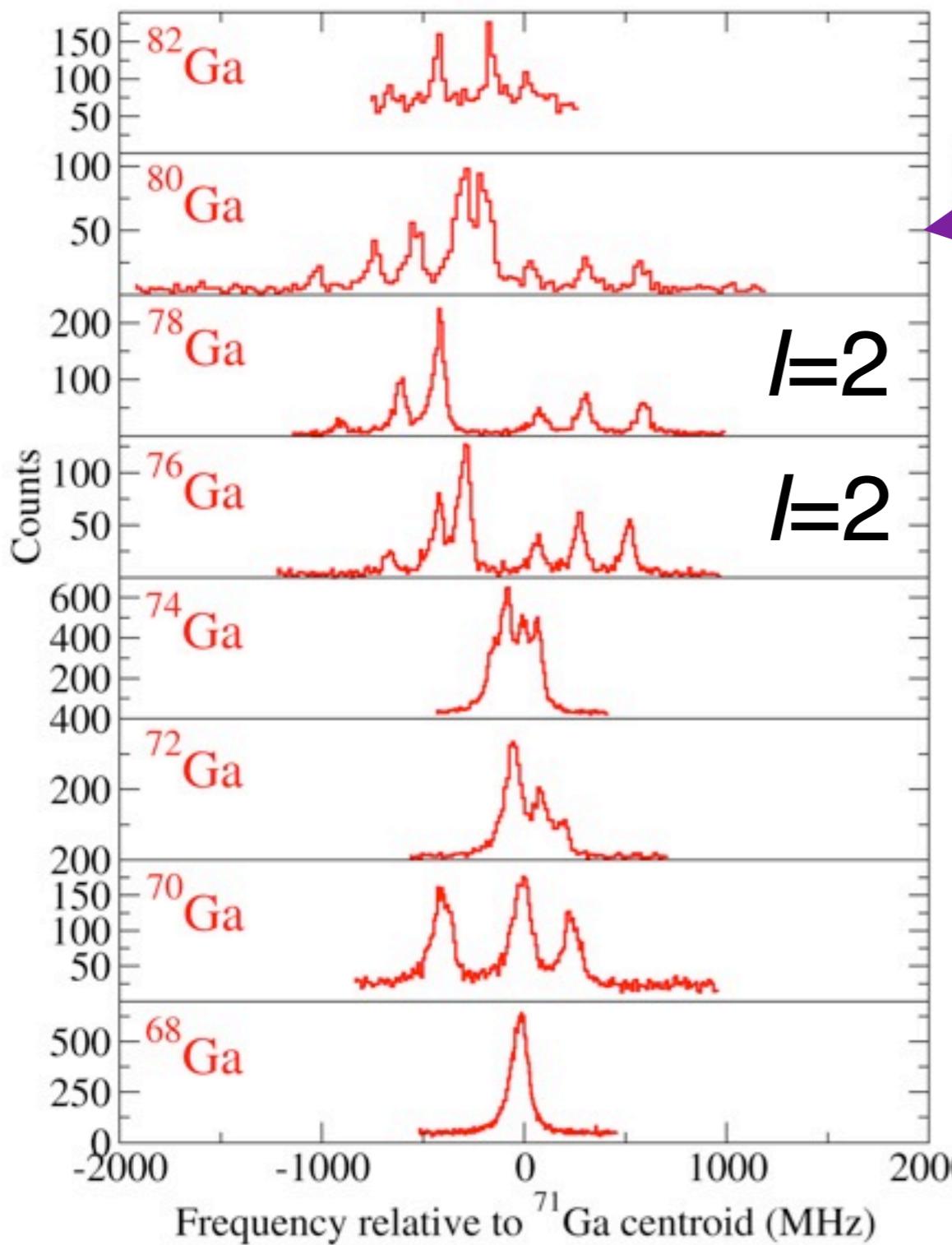
## Electric quadrupole



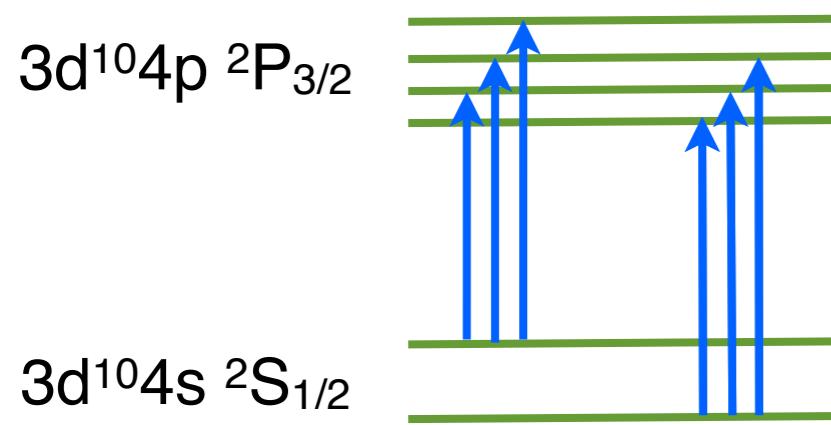
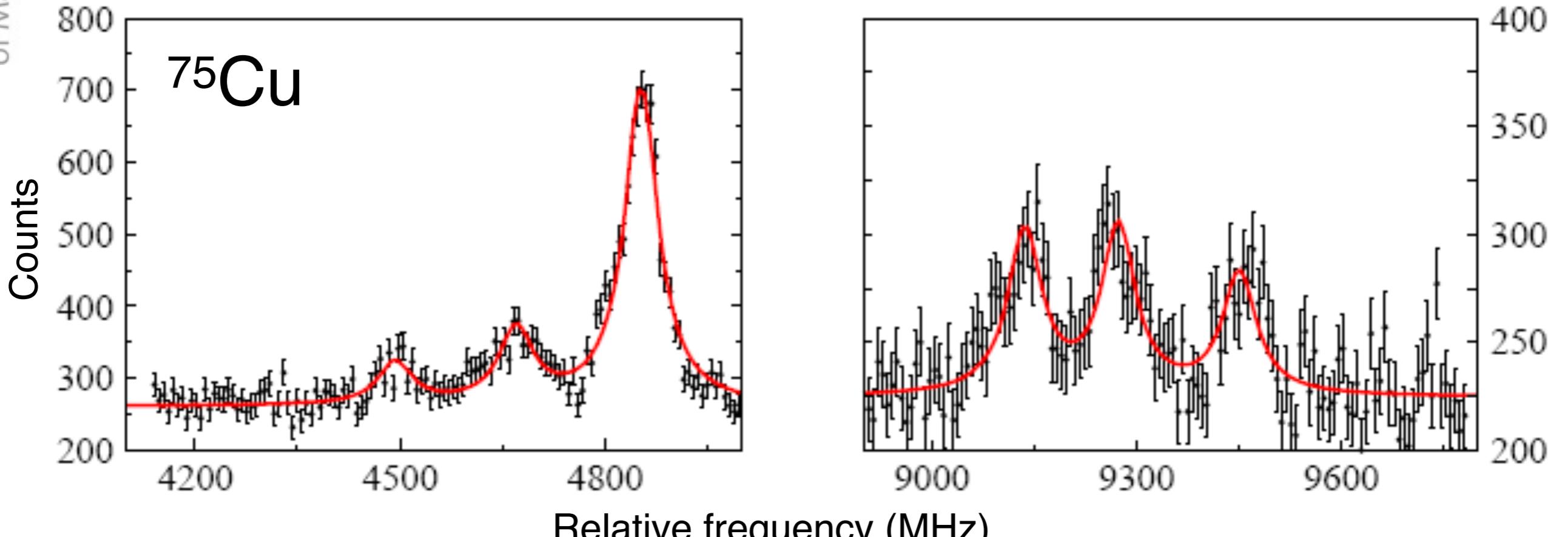
Predictions for second 3/2 states (300 keV)

It is these which match gs properties ( $f_{5/2}^3$  dominated)

# Odd-odd Ga isotopes



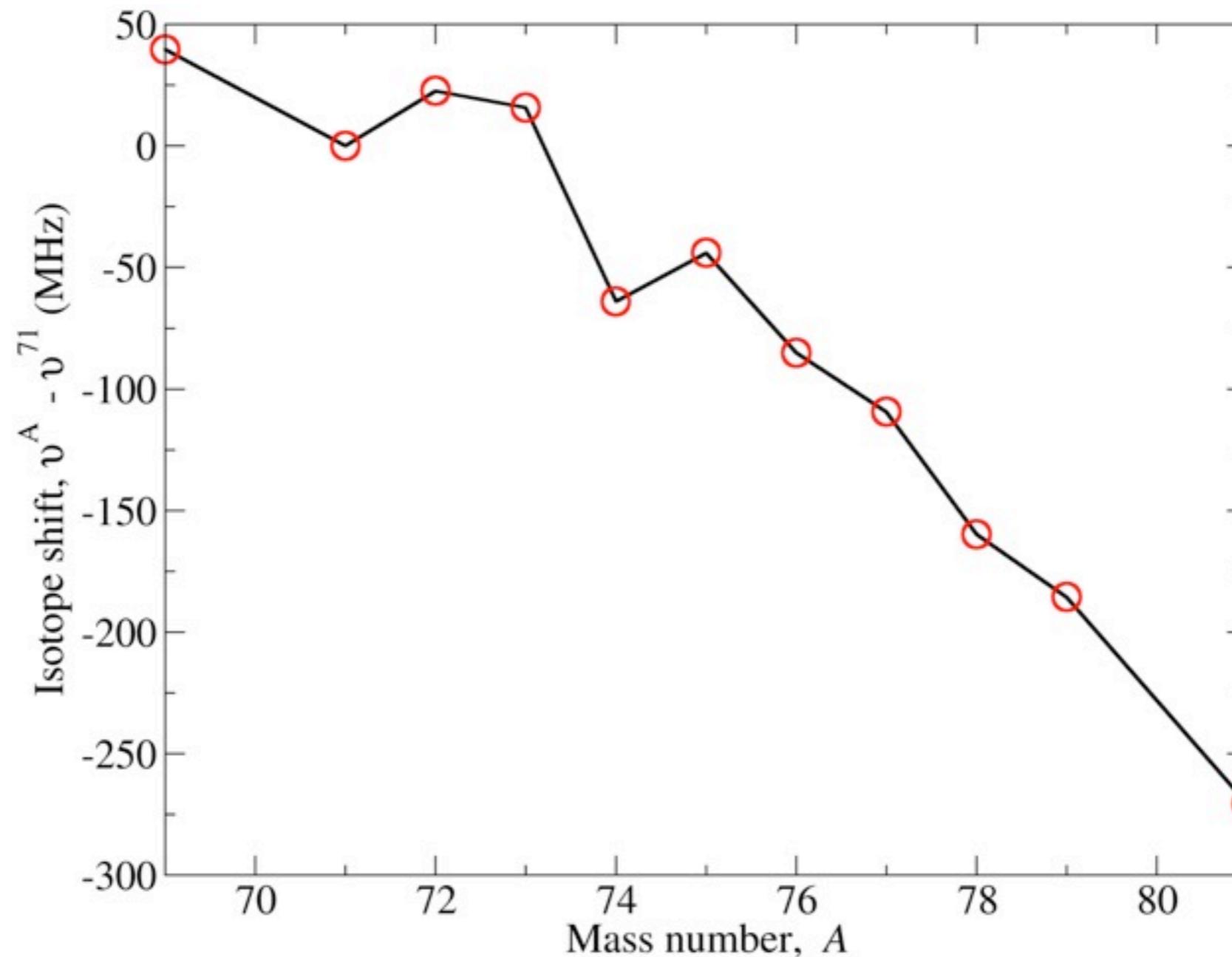
# Inversion in copper



(odd) 59-73Cu are  $I=3/2$   
but: 75Cu is  $I=5/2$

Inversion between  $N=44,46$  (Cu)  
 $N=48,50$  (Ga)

# Isotope shift data

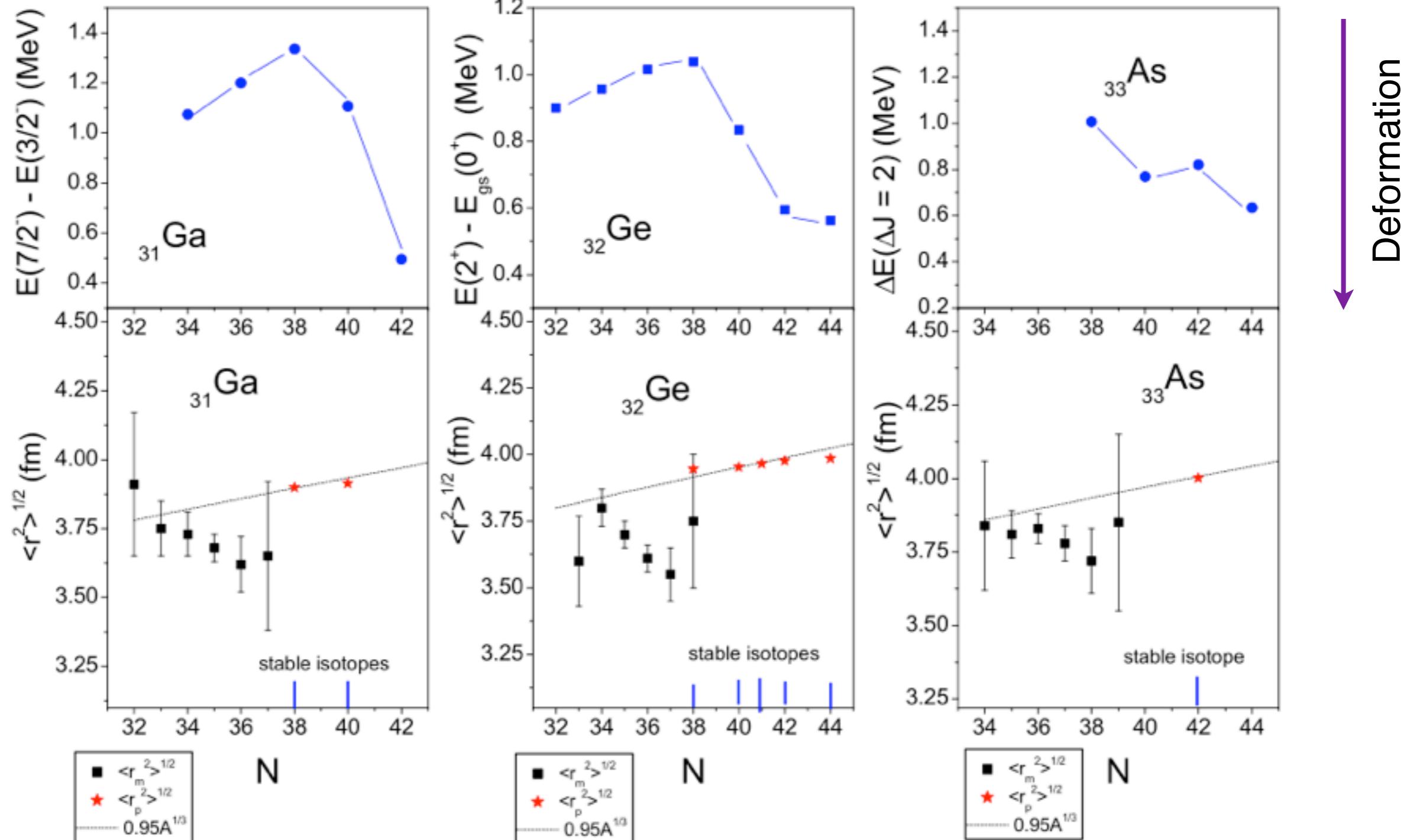


Need to calculate atomic factors...

$$\delta\nu^{A,A'} = M_i \frac{A' - A}{AA'} + F_i \boxed{\delta\langle r_c^2 \rangle^{A,A'}}$$

# Case for neutron-deficient study

A. Lépine-Szily *et al.*  
Eur. Phys. J. A. 25, s01, 227 (2005)



Matter radii *increase* with decreasing neutron number  
→ Suggest formation of a proton skin →  $\delta\langle r_c^2 \rangle$  → IS.

# In conclusion...

- Commissioned ISCOOL
  - Critically suppresses photon background
- Measured n-rich spins, moments, isotope shifts
  - Spin inversion due to monopole migration
  - Gradual emptying of proton  $p_{3/2} \rightarrow f_{5/2}$  domination
  - $^{73}\text{Ga}$  has anomalous spin of  $I=1/2$
- Will study of  $^{62-69}\text{Ga}$ 
  - Measure charge radii (proton skin?)

# Acknowledgements

B. Cheal,<sup>1,\*</sup> E. Mané,<sup>1</sup> J. Billowes,<sup>1</sup> M.L. Bissell,<sup>2</sup> K. Blaum,<sup>3</sup> B.A. Brown,<sup>4</sup> F.C. Charlwood,<sup>1</sup> K.T. Flanagan,<sup>5</sup> D.H. Forest,<sup>6</sup> C. Geppert,<sup>7,8</sup> M. Honma,<sup>9</sup> A. Jokinen,<sup>10,11</sup> M. Kowalska,<sup>12,3</sup> A. Krieger,<sup>8</sup> J. Krämer,<sup>8</sup> I.D. Moore,<sup>10,11</sup> R. Neugart,<sup>8</sup> G. Neyens,<sup>2</sup> W. Nörtershäuser,<sup>7,8</sup> M. Schug,<sup>3</sup> H.H. Stroke,<sup>13</sup> P. Vingerhoets,<sup>2</sup> D.T. Yordanov,<sup>3</sup> and M. Žáková<sup>8</sup>

<sup>1</sup>*School of Physics and Astronomy, The University of Manchester, Manchester, M13 9PL, UK*

<sup>2</sup>*Instituut voor Kern- en Stralingsfysica, Katholieke Universiteit Leuven, B-3001 Leuven, Belgium*

<sup>3</sup>*Max-Planck-Institut für Kernphysik, D-69117 Heidelberg, Germany*

<sup>4</sup>*National Superconducting Cyclotron Laboratory and Department of Physics and Astronomy,  
Michigan State University, East Lansing, Michigan 48824-1321, USA*

<sup>5</sup>*I.P.N. Orsay, F-91940 Orsay Cedex, France*

<sup>6</sup>*School of Physics and Astronomy, The University of Birmingham, Birmingham, B15 2TT, UK*

<sup>7</sup>*GSI Helmholtzzentrum für Schwerionenforschung GmbH, D-64291 Darmstadt, Germany*

<sup>8</sup>*Institut für Kernchemie, Universität Mainz, D-55128 Mainz, Germany*

<sup>9</sup>*Center for Mathematical Sciences, University of Aizu, Tsuruga,  
Illi-machi, Aizu-Wakamatsu, Fukushima 965-8580, Japan*

<sup>10</sup>*Department of Physics, University of Jyväskylä, FIN-40014 Jyväskylä, Finland*

<sup>11</sup>*Helsinki Institute of Physics, FIN-00014, University of Helsinki, Finland*

<sup>12</sup>*Physics Department, CERN, CH-1211 Geneva 23, Switzerland*

<sup>13</sup>*Department of Physics, New York University, New York, NY 10003, USA*