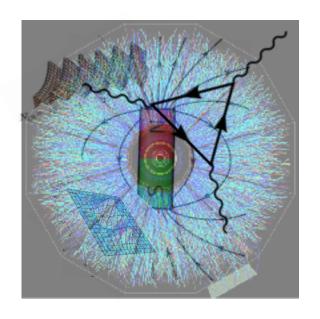
## Magnetic Field & Fluid Rotation in Heavy Ion Collisions







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Indiana University, Physics Dept. & CEEM

Central China Normal University

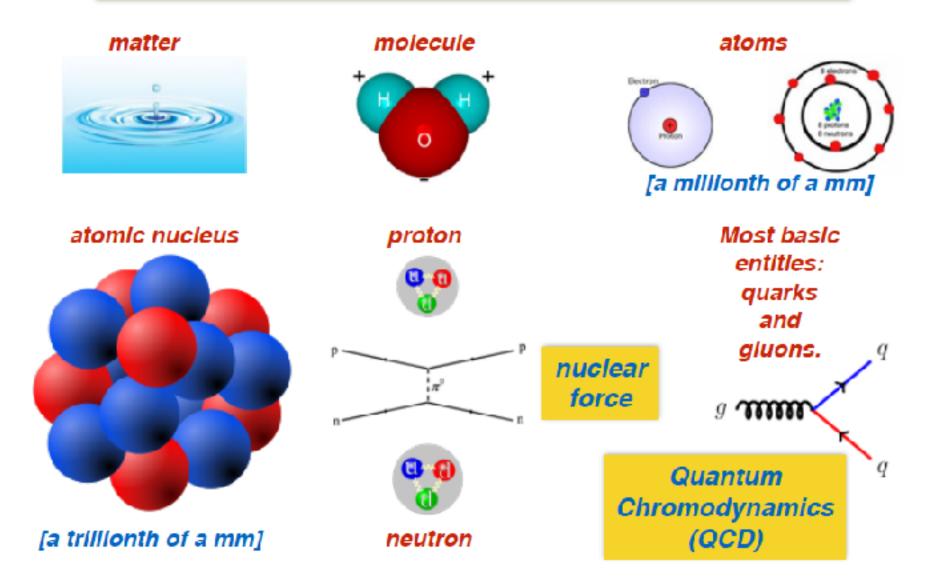
Research Supported by NSF & DOE



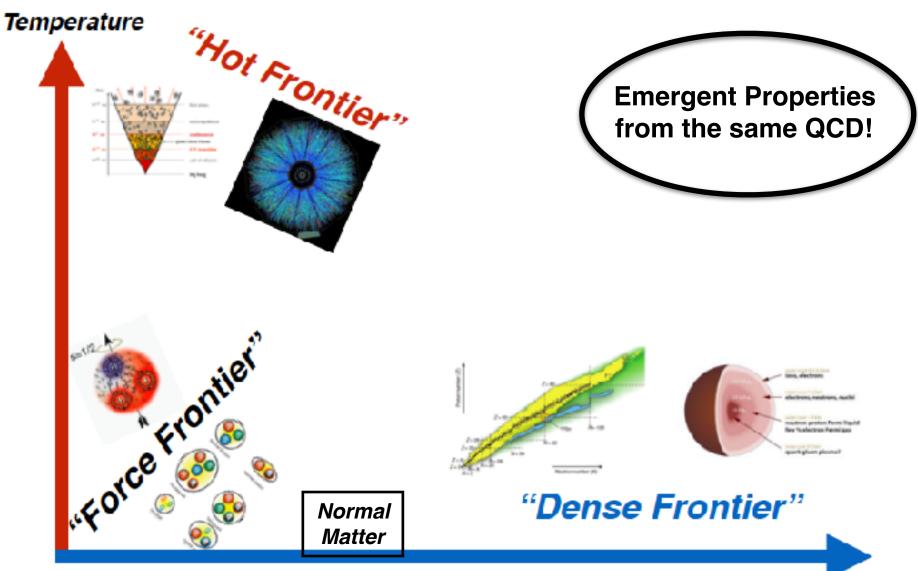
## The Bigger Picture

#### Nuclear Matter: At the Heart of All Matter

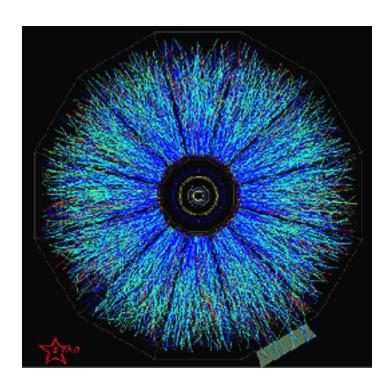
#### The physical world has a hierarchy of structures.



## A Map of Nuclear Matter



#### "Little Bang" in High Energy Nuclear Collision

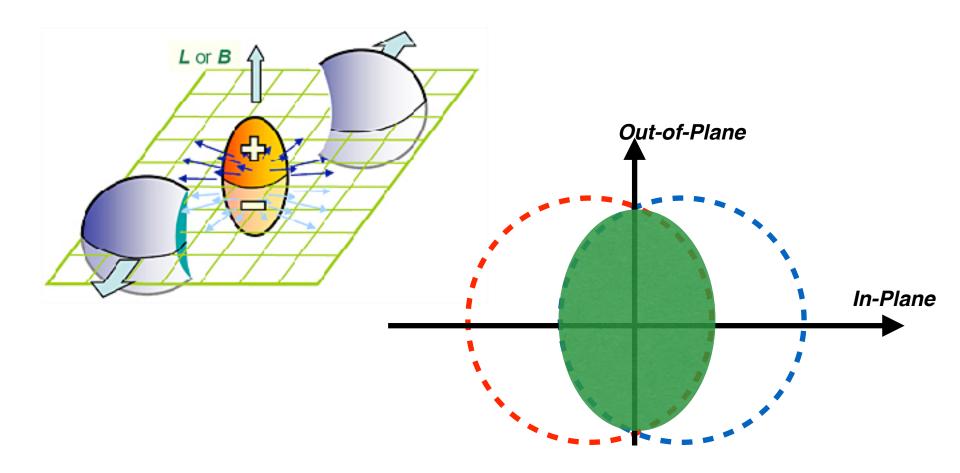




- \* Quark-gluon plasma (QGP) is created in such collisions.
- \* It is PRIMORDIALLY HOT ~ trillion degrees ~ early universe.



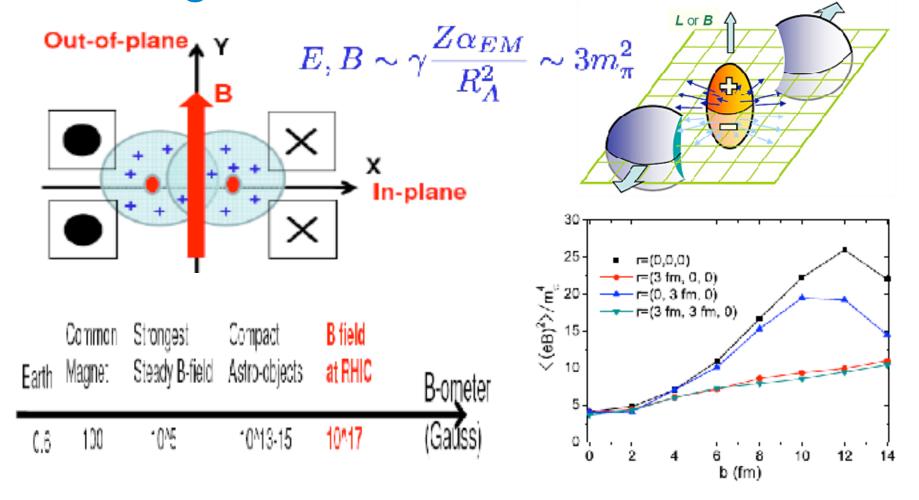
#### New Phase & New Extreme Conditions



The quark-gluon plasma is a type of CHIRAL MATTER, with (approximately) chiral quarks.

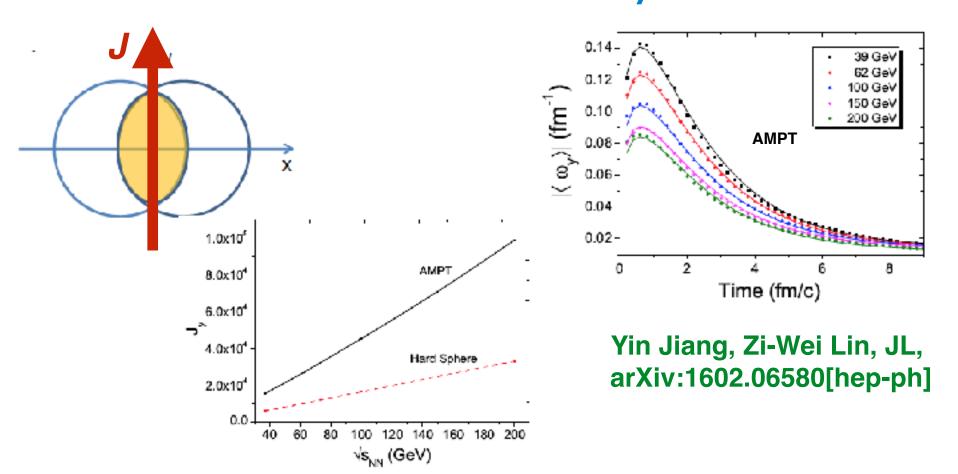
Heavy ion collision environment: extremely strong magnetic field and fluid rotation!

The Strongest EM Fields in Heavy Ion Collisions



- **Strongest** B field (and strong E field as well) naturally arises! [Kharzeev,McLerran,Warringa;Skokov,et al; Bzdak-Skokov; Deng-Huang; Bloczynski-Huang-Zhang-Liao; Skokov-McLerran;Tuchin; ...]
- "Out-of-plane" orientation (approximately)

#### The Most Vortical Fluid in Heavy Ion Collisions



The fluid created in heavy ion collisions carries extremely large angular momentum and local vorticity.

#### Properties of Matter under New Conditions

Very strong magnetic field —> Magnetic Polarization!

Very strong global rotation

-> Rotational Polarization!

Nontrivial interplay between external conditions & internal degrees of freedom (spin/momentum/angular momentum)





Thermodynamic properties of matter:

- \* Equation of State
- \* Phase Structure
- \* Macroscopic Polarization

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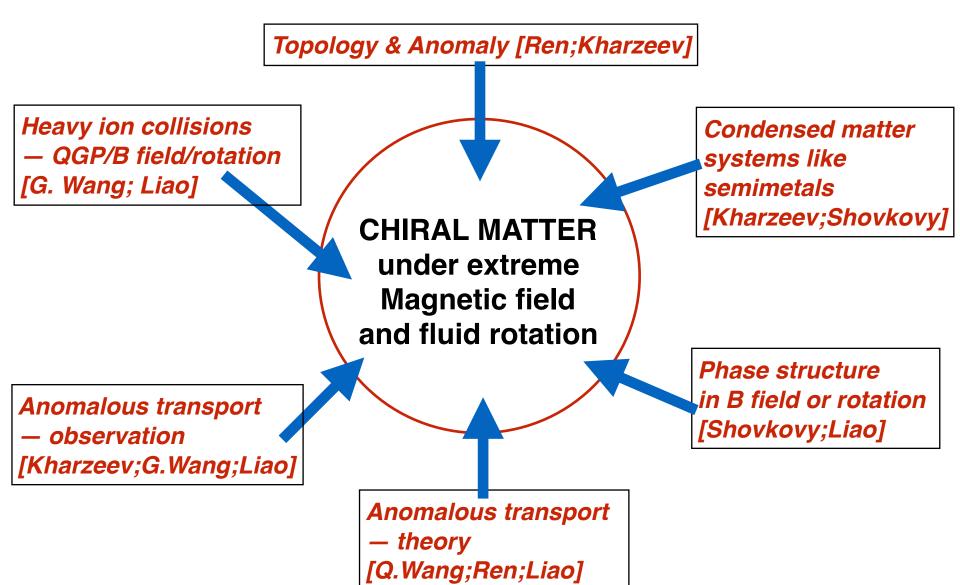
Transport properties of matter:

- \* Transport coefficients
- \* Anomalous transport

effects

\*

#### An Overall Look at the Present Summer School



#### Quantifying External Conditions in H.I.C.

In the first three lectures, I will give you a hands-on tutorial on how to compute such new external conditions as magnetic field and fluid rotation in heavy ion collisions.

# Computing Magnetic Fields in Heavy Ion Collisions

A (very incomplete!) list of recent references on computation of magnetic fields

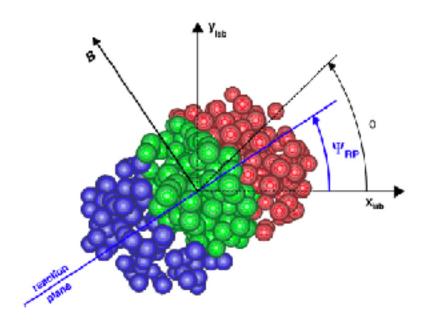
- \* Deng, Huang, arXiv:1201.5108 ; \* Bzak, Skokov, arXiv:1111.1949
- \* Blocynski, Huang, Zhang, Liao, 1209.6594
- \* McLerran, Skokov, arXiv:1305.0774; \* Tuichin, arXiv:1305.5806
- \* Li, Sheng, Q. Wang, 1602.02223
- \* Inghirami, Zanna, Beraudo, Moghaddam, Becattinni, Bleicher, arXiv:1609.03042

\*

### Magnetic Field in Heavy Ion Collisions

**Blackboard** 

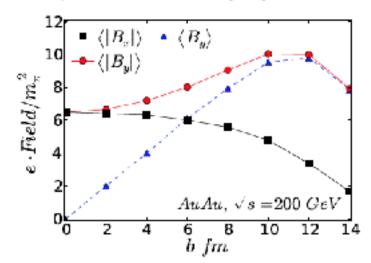
#### More Realistic Computations

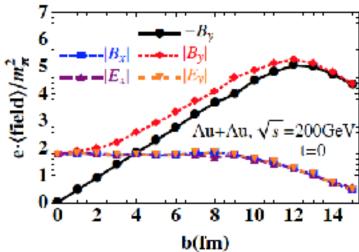


- \* Glauber model (with discrete nucleons)
- \* Event-by-event with Monte-Carlo
- \* Finite size of proton
- \* Azimuthal angule fluctuations
- \* Time evolution further in medium??????

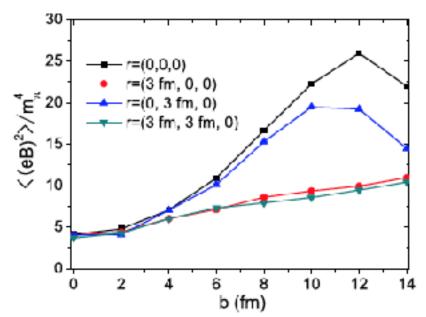
#### Event-By-Event Magnetic Fields

## Bzdak & Skokov, arXiv:1111.1949

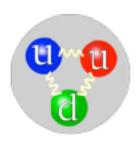




Deng & Huang, arXiv:1201.5108

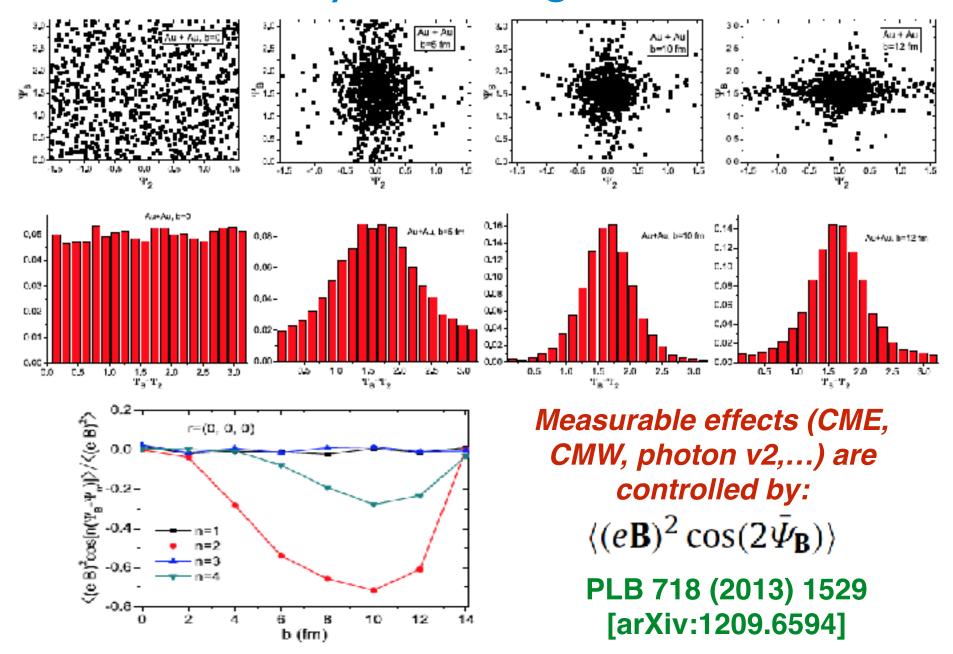


PLB 718 (2013) 1529 [arXiv:1209.6594]

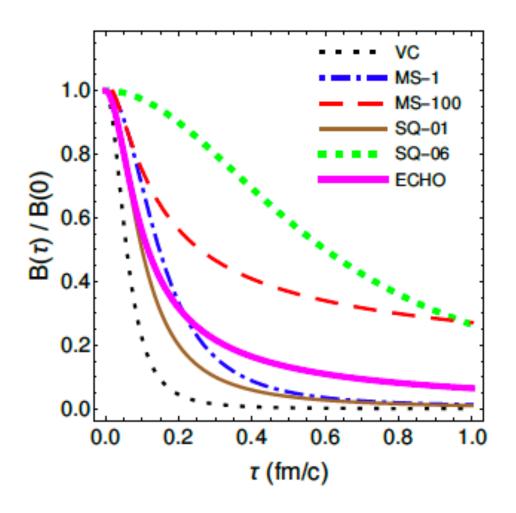


Proton is a finite size object!

#### Event-By-Event Magnetic Fields



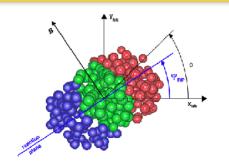
#### Time Evolution of Magnetic Field



The single biggest uncertainty in current understanding of magnetic fields in heavy ion collisions!

#### Small v.s. Large and Low v.s. High Energy

#### Why we do NOT expect CME in pA collisions?

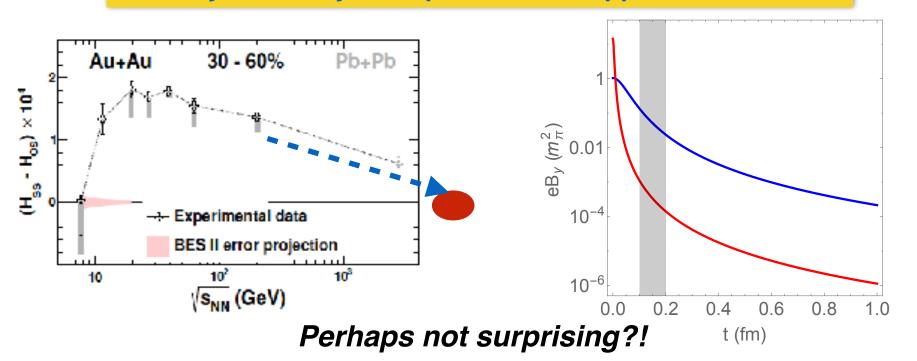


Angle de-correlation between B and event plane!!

$$\cos[2(\Psi_B - \Psi_2)]$$

CMS reported nice measurements at 5TeV.

#### Why CME may be expected to disappear at 5TeV?



# Computing Fluid Rotation in Heavy Ion Collisions

A (very incomplete!) list of recent references on computation of fluid rotation

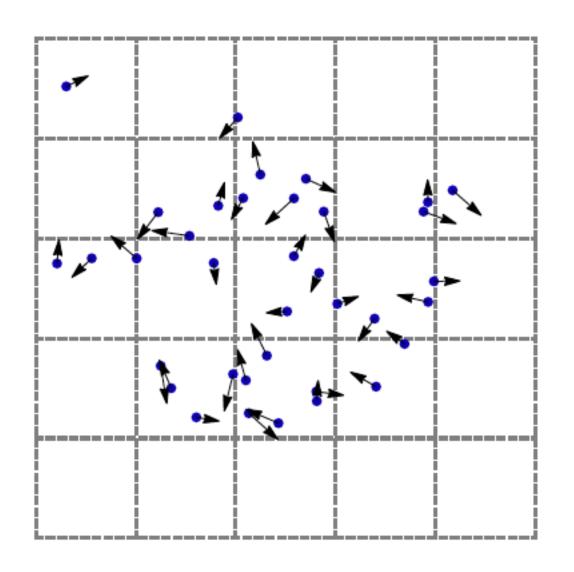
- \* Jiang, Lin, JL, arXiv:1602.06580
- \* Deng, Huang, arXiv:1603.06117
- \* Pang, Petersen, Q. Wang, X.-N. Wang, arXiv:1605.04024
- \* Becattinni, et al, arXiv:1501.04468
- \* Csernai, Magas, arXiv:1302.5310

\*

## Fluid Rotation in Heavy Ion Collisions

**Blackboard** 

#### Fluid Rotation in Heavy Ion Collisions



$$\vec{J} = \int d^3r \vec{r} \times \vec{p}(\vec{r}).$$

$$\vec{v}(\vec{r}) = \vec{p}(\vec{r})/\epsilon(\vec{r})$$

$$\vec{\omega} = \nabla \times \vec{v}$$

### Angular Momentum Carried by Fireball

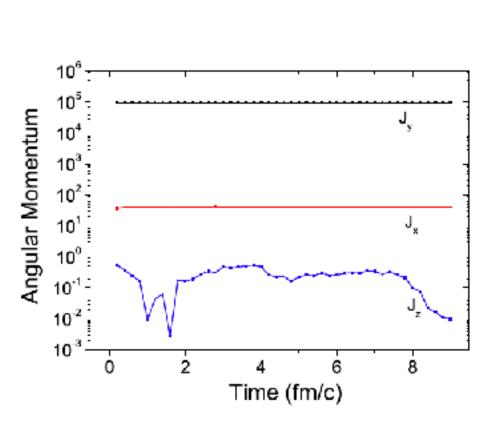
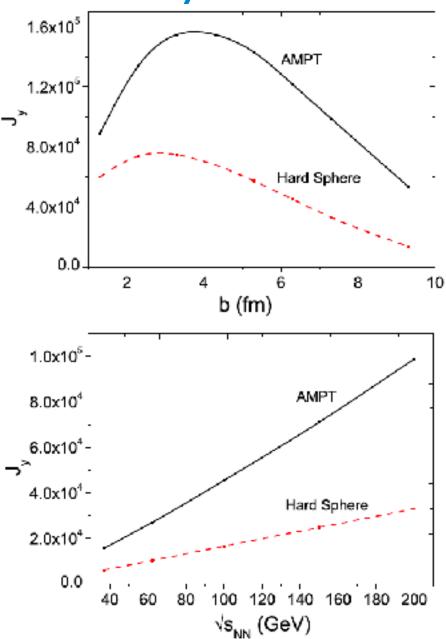


FIG. 2. Angular momentum from the AMPT model at b = 7 fm and  $\sqrt{s_{NN}} = 200$  GeV.



### Fluid Rotation in Heavy Ion Collisions

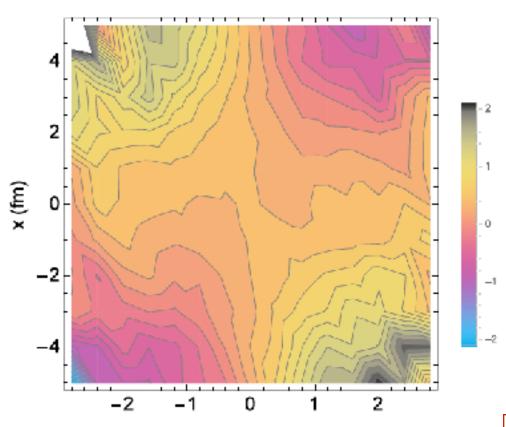


FIG. 10. Radial velocity profile at  $\eta = 1$  and t = 1 fm/c, with b = 7 fm and  $\sqrt{s_{VN}} = 200$  GeV.

$$\vec{v}(\rho,\phi,\eta) = \hat{e}_{\rho}v_0(\rho,\eta)[1 + 2c_2(\rho,\eta)\cos 2\phi],$$

FIG. 5. 
$$\omega_y$$
 (in the unit of fm<sup>-1</sup>) profile at  $y = 0$  and  $t = 1$  fm/c, with  $b = 7$  fm and  $\sqrt{s_{NN}} = 200$  GeV.

$$\omega_y = \frac{\partial v_\rho}{\partial z} \cos \phi$$

$$= \frac{2}{t} (ch\eta)^2 \, \partial_\eta (v_0 + 2v_0 c_2 \cos 2\phi) \cos \phi$$

$$= \frac{2}{t} (ch\eta)^2 \left(\frac{x}{\rho}\right) \, \partial_\eta \left[v_0 + 2v_0 c_2 \left(2\frac{x^2}{\rho^2} - 1\right)\right].$$

#### Fluid Rotation in Heavy Ion Collisions

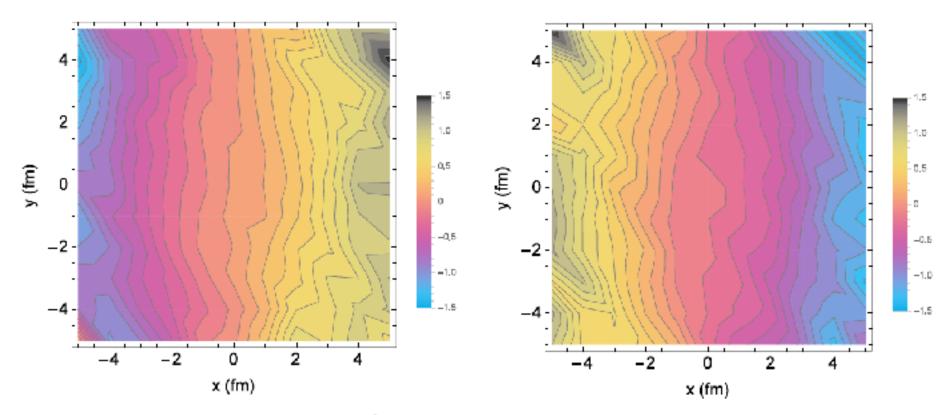
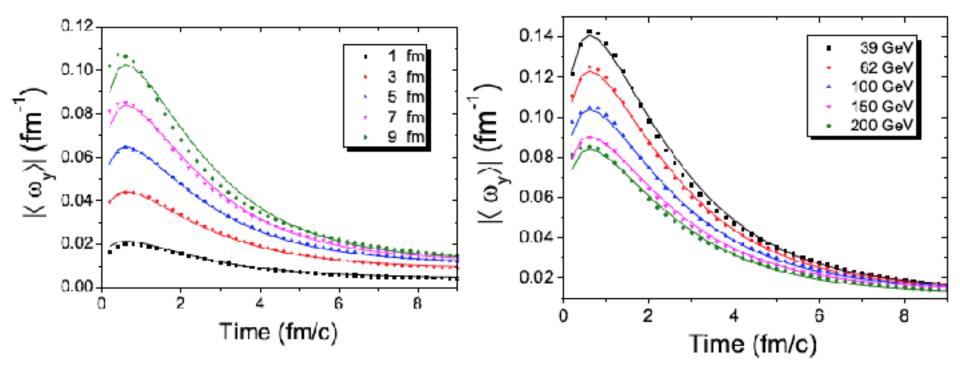


FIG. 7.  $\omega_{\tau}$  profile at  $\eta = -1$  and t = 1 fm/c, with b = 7 fm and  $\sqrt{s_{NN}} = 200$  GeV.

FIG. 6.  $\omega_y$  profile at  $\eta=1$  and t=1 fm/c, with b=7 fm and  $\sqrt{s_{NN}}=200$  GeV.

## Quantifying Rotation of QGP



#### Convenient parameterization:

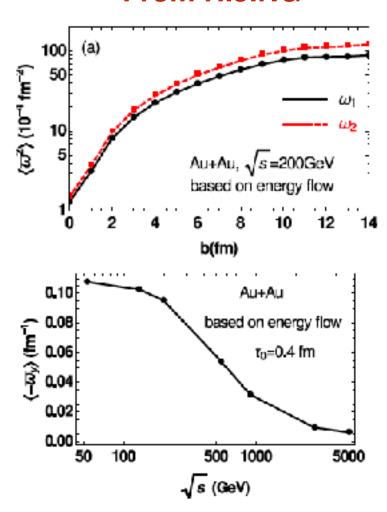
$$\langle \omega_y \rangle (t, b, \sqrt{s_{NN}}) = A(b, \sqrt{s_{NN}}) + B(b, \sqrt{s_{NN}}) (0.58t)^{0.35} e^{-0.58t}$$

$$\begin{split} A & \left[ e^{-0.016\,b\,\sqrt{s_{NN}}} + 1 \right] \times \tanh(0.28\,b) \\ & \times \left[ 0.001775\, \tanh(3 - 0.015\sqrt{s_{NN}}) + 0.0128 \right] \\ B & = \left| e^{-0.016\,b\,\sqrt{s_{NN}}} + 1 \right| \times \left[ 0.02388\,b + 0.01203 \right] \\ & \times \left[ 1.751 - \tanh(0.01\sqrt{s_{NN}}) \right] \;. \end{split}$$

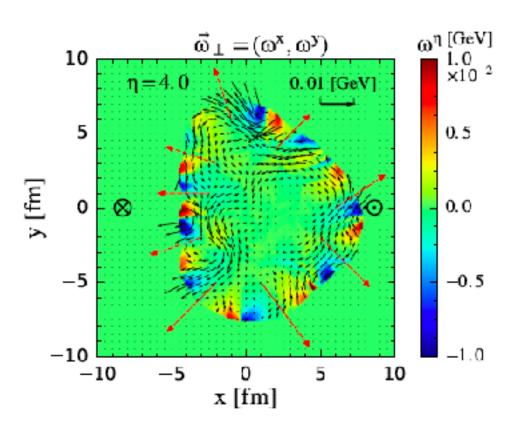
Yin Jiang, Zi-Wei Lin, JL, arXiv:1602.06580[hep-ph]

### Quantifying Rotation of QGP

#### From HIJING



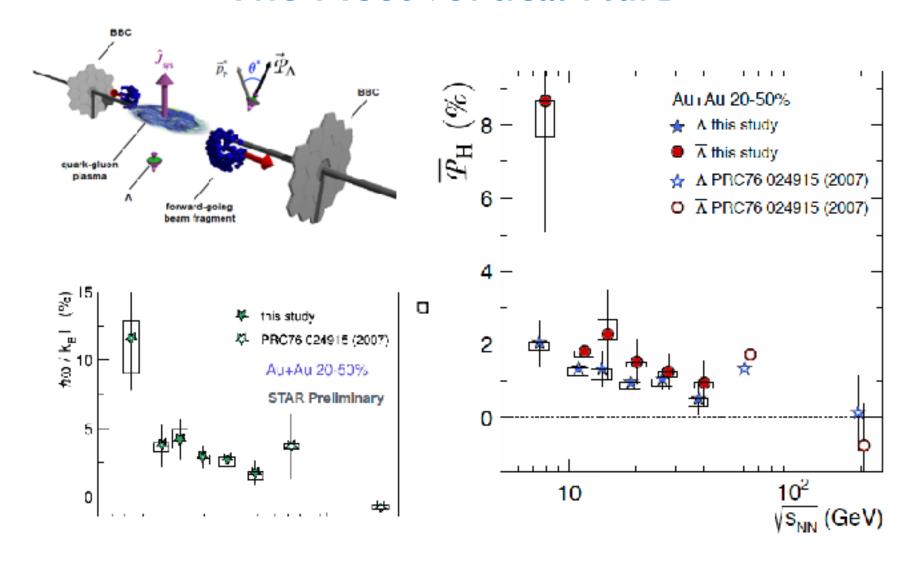
From 3D hydro



Deng, Huang, arXiv:1603.06117

Pang, Petersen, Wang, Wang arXiv:1605.04024

#### The Most Vortical Fluid



STAR Collaboration, arXiv:1701.06657

## Some Concluding Remarks

Q & A