



# LHC Luminosity Follow-up for 2016 - Summary

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Acknowledgements to G. Arduini, X. Buffat, K. Li, G. Trad, R. Tomas, ATLAS collaboration, CMS collaboration

Beam-beam and Luminosity studies meeting, 24 Feb. 2017

### Outline

Model description

.Emittance evolution from injection to stable beams

Peak luminosity along the year
 Calculated vs measured

Luminosity evolution

- -Luminosity evolution prediction
- -Emittance evolution during Stable Beams
- -Losses during the first hours at Stable Beams
- -Impact on integrated luminosity loss

.Summary

#### **Model description**

# Luminosity model description

# Self consistent bunch by bunch luminosity model Emittance evolution:

Intrabeam scattering (IBS), Synchrotron Radiation (SR) elastic scattering due to the proton-proton collisions

$$\left(\frac{d\epsilon_x}{dt}, \frac{d\epsilon_y}{dt}, \frac{d\sigma_s}{dt}\right)_{IBS+SR} = f(En, N_b(t_0), \epsilon_x(t_0), \epsilon_y(t_0), \sigma_s(t_0), dt)$$
$$\left(\frac{d\epsilon_{x,y}}{dt}\right)_{elastic} = N_{IP}\beta^*_{x,y}\mathcal{L}\sigma_{el} < \theta^2_{x,y} > /(n_bN_p).$$

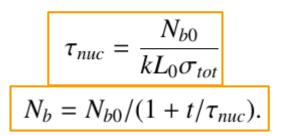
$$\frac{d\epsilon}{dt} = \left(\frac{d\epsilon}{dt}\right)_{IBS+SR} + \left(\frac{d\epsilon}{dt}\right)_{elastic}$$

.Or using the empirical evolution from the data

# Luminosity model description

**Bunch intensity evolution:** 

Luminosity burn-off due to the p-p collisions:



.Or using the empirical bunch intensity evolution from the data

#### **•Bunch length evolution:**

Intrabeam scattering and synchrotron radiation (see previous slide)Or using the empirical bunch length evolution from the data

- Iteration of the equations in a self-consistent way and in small timesteps (10-15min) such the bunch intensity can be considered constant
- Any of the modes can be called at each time step
  - Theoretical or empirical evolution of the bunch parameters and luminosity

# Luminosity model description

- The model can be applied bunch-by-bunch both for **colliding** and **non-colliding** bunches
- The emittance evolution function can be applied both at **injection** and **flat top energies**
- Can be applied under different assumptions:
  - Pure model:
    - Initial values of bunch intensities, emittances and bunch length taken from the data
    - Model iteration to compute intensity, emittance, bunch length and luminosity evolution
  - EmpiricalBlowUpBurnOff:
    - Emittance evolution taken from the data
    - Model iteration to compute bunch intensity, bunch length and luminosity evolution
  - IBSEmpiricalLosses:
    - Intensity evolution taken from the data
    - Model iteration to compute emittance, bunch length and luminosity evolution
  - EmpiricalBlowUpEmpiricalLosses:
    - Intensity and emittance evolution taken from the data
    - Model iteration to compute luminosity evolution

#### Data used as input

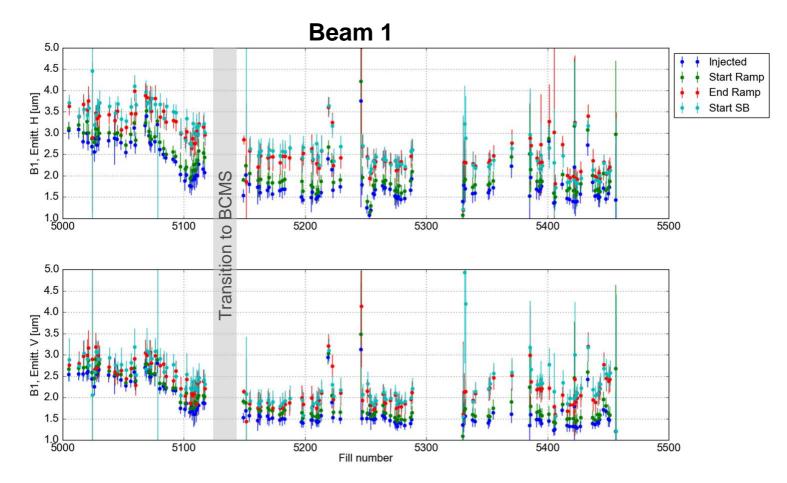
- Bunch by bunch intensity data from fBCT
- Bunch by bunch emittance data from BSRT
- Bunch by bunch bunch length data from BQM
- Bunch by bunch luminosities from ATLAS and CMS (Massi files are used)
- → A set of tools have been developed to ease the follow up of the beam quality and luminosity evolution in the LHC and the comparison against models.
  - Extended the python tools used for the scrubbing follow up and integrated the luminosity simulation in the same framework

Many thanks to Gianni for this!!!

→ The idea is to use the same tools for the weekly follow up of the fills of 2017

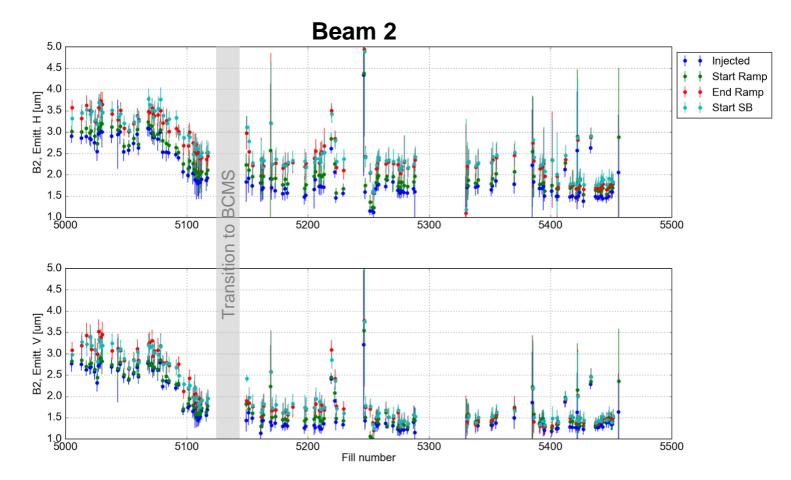
### Injection to Stable Beams Analysis

# **Emittance evolution from injection to stable beams**



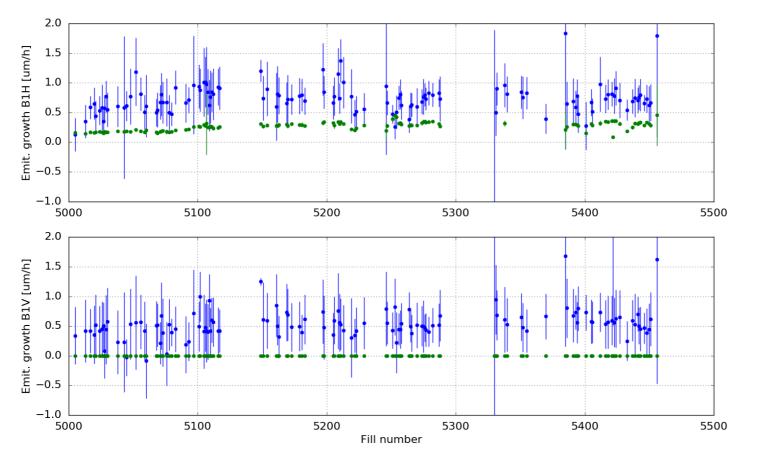
Larger part of the blow up induced during the Ramp
 Cannot be explained by IBS+SR

# **Emittance evolution from injection to stable beams**



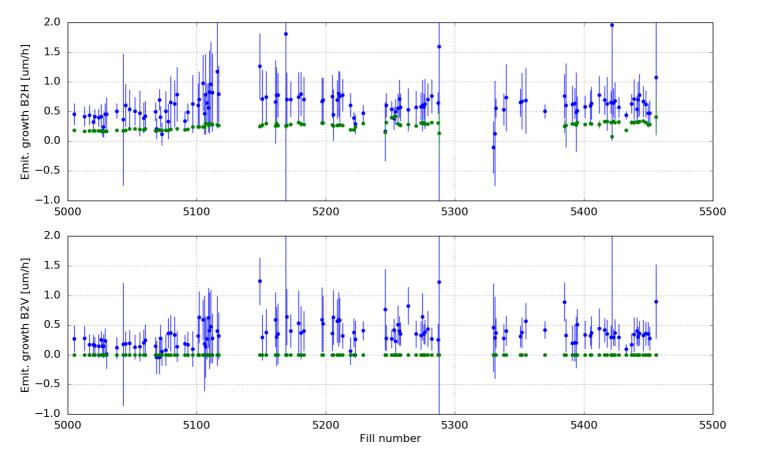
Larger part of the blow up induced during the Ramp
 Cannot be explained by IBS+SR

#### Emittance evolution during Flat Bottom – Beam 1



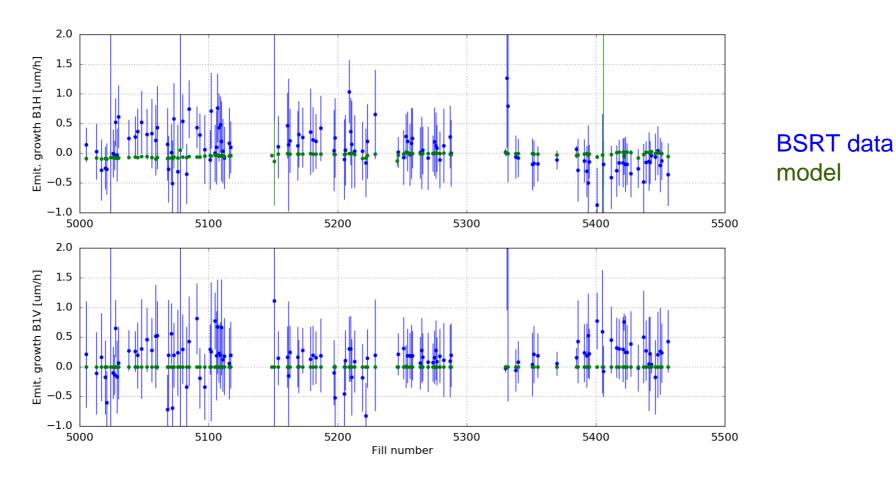
BSRT data model

#### Emittance evolution during Flat Bottom – Beam 2

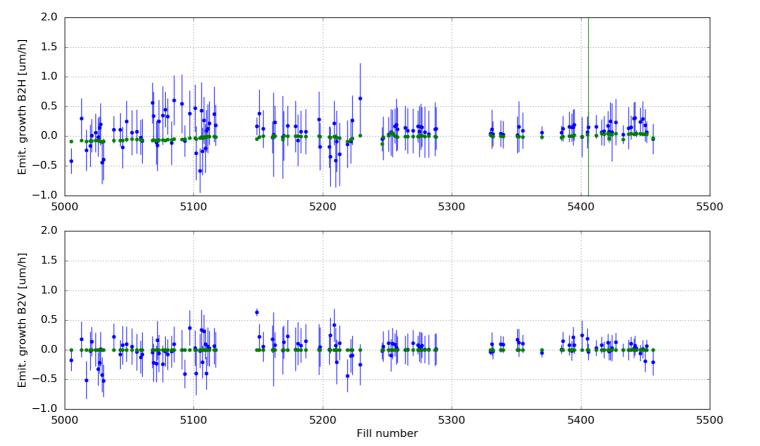


BSRT data model

#### Emittance evolution during Flat Top – Beam 1

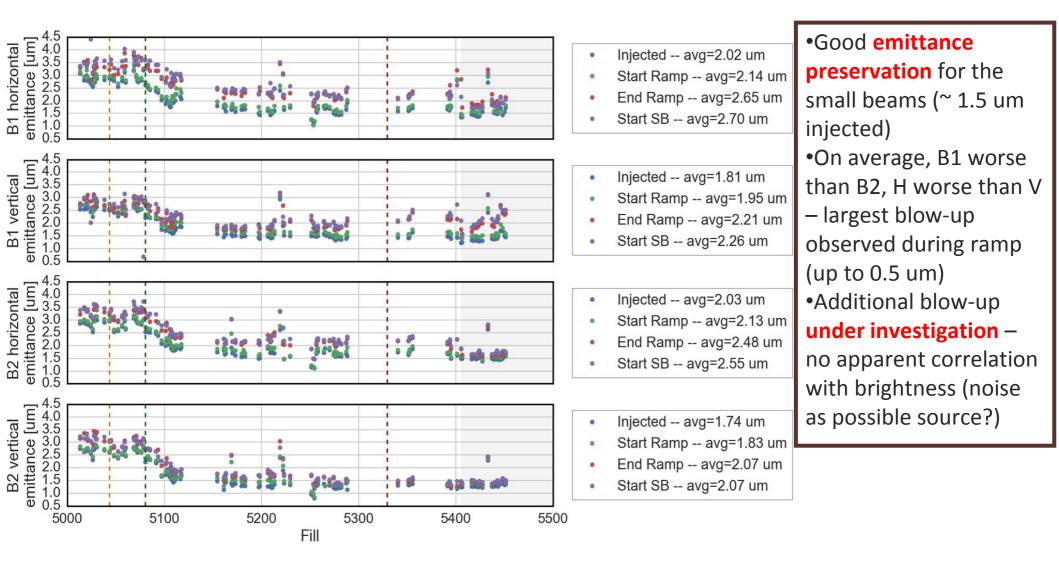


#### Emittance evolution during Flat Top – Beam 2

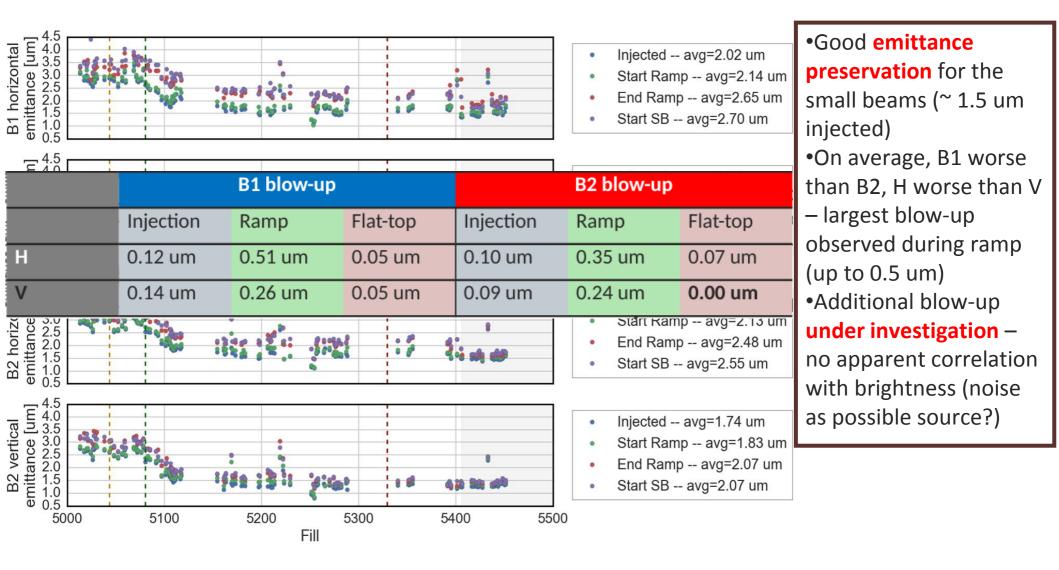


BSRT data model

# Emittance evolution from injection to stable beams – model corrected



# Emittance evolution from injection to stable beams – model corrected

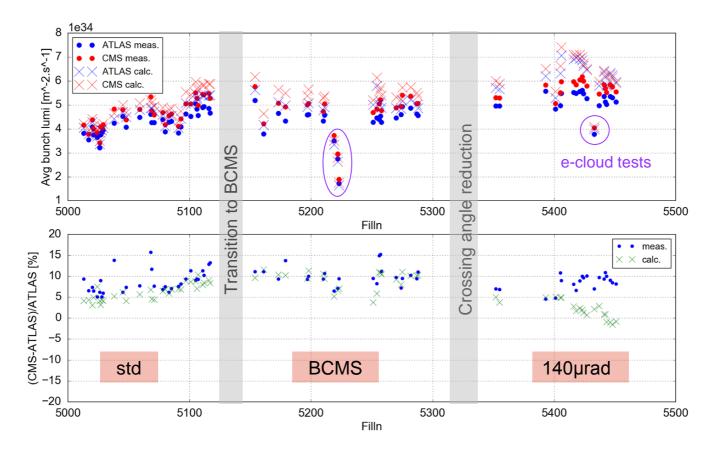


From K. Li, Chamonix 2017

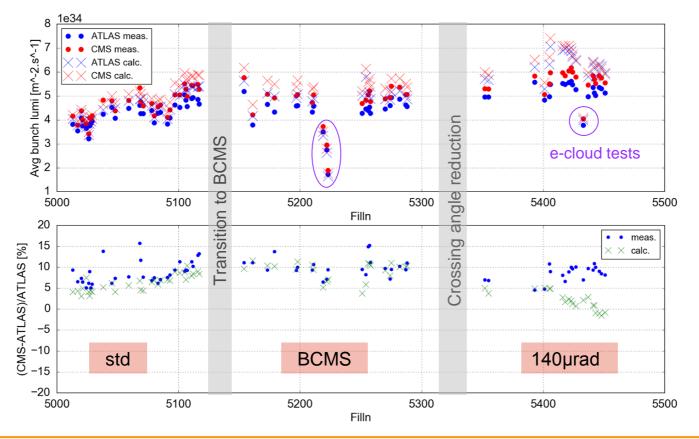
# **Emittance evolution from injection to stable beams**

- Remain to be done:
  - Add the model during the ramp (for the LHC parameters the effect is very small, of the order of 2-3%)
  - Estimate the peak luminosity loss per fill due to the extra emittance blow up during the cycle
    - It is almost ready.. needs to be finalized

#### **Stable Beams Analysis**

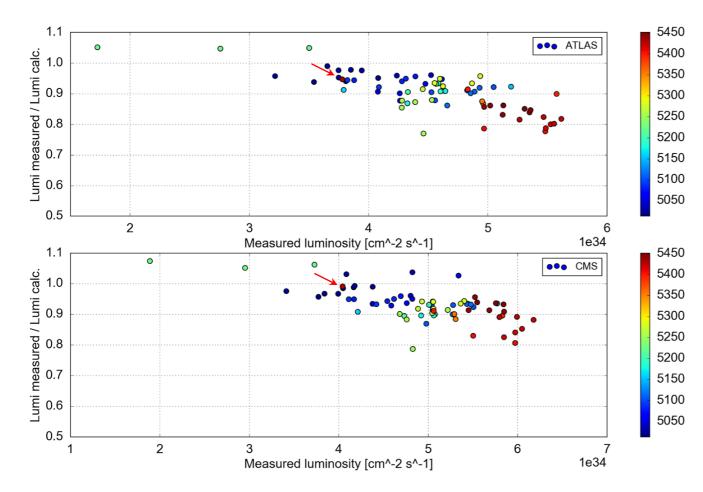


- Comparison of average peak luminosity as measured by the experiments (dots) and calculated by beam parameters (crosses)
- Fairly good agreement for large part of the run
  - Some discrepancy for the last fills (to be understood)



#### To be done:

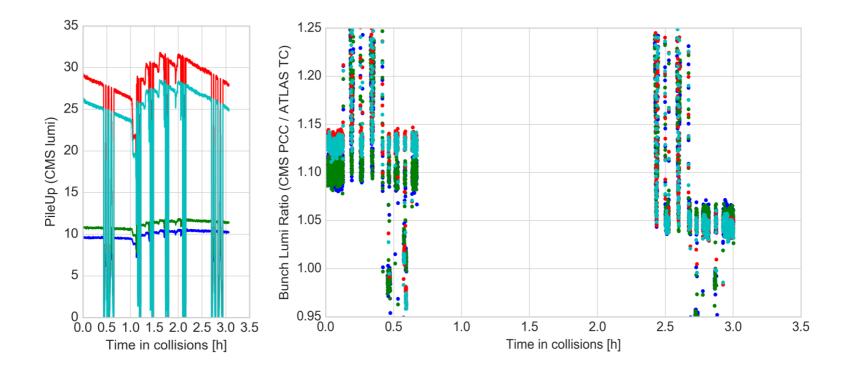
- Repeat the same plot with different set of BSRT calibration as requested by Jamie and Cristoph
  - Rescale the middle part to the last calibration factors
- Follow-up with the experiments to understand their limitations



- Correlation of the difference between the measured and calculated luminosity with the measured peak luminosity observed
  - Valid also in the same BSRT calibration factor periods

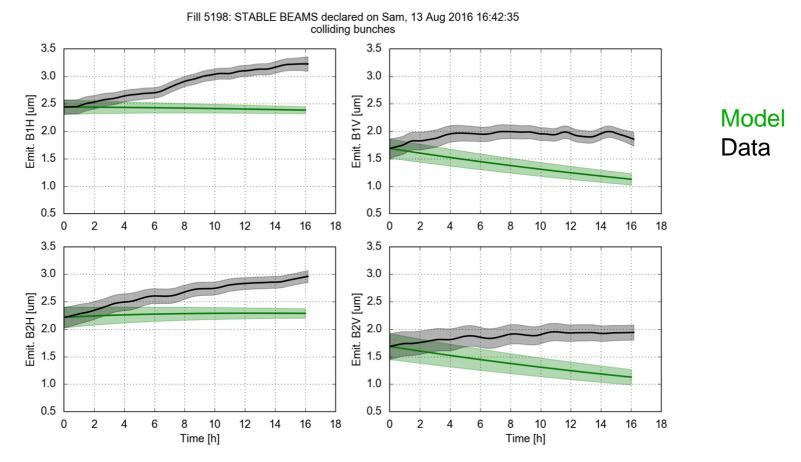
• To be done:

- verify with the other set of BSRT calibration factors
- Check within the same calibration periods



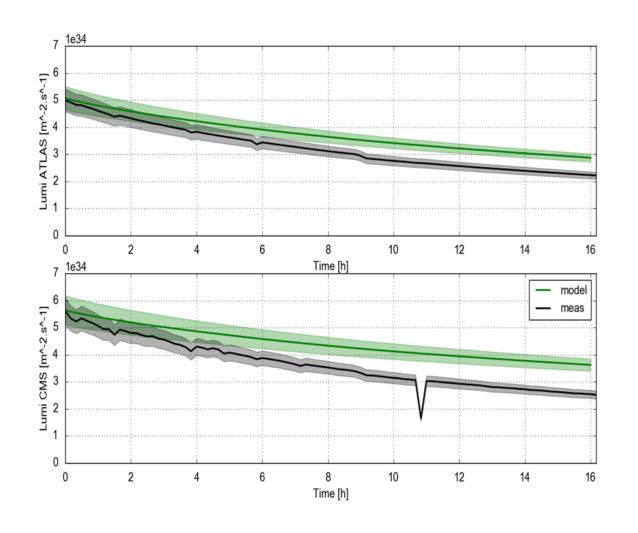
- A very interesting experiment followed up by Michi and W. Kozanecki
  - Crossing angle scan test
  - 4 bunches with different brightness were brought to collision
  - ~5-8% geometric effect (larger for lower emittance bunches)
  - ~5% imbalance still observed at zero crossing angle

### **Bunch parameters evolution**



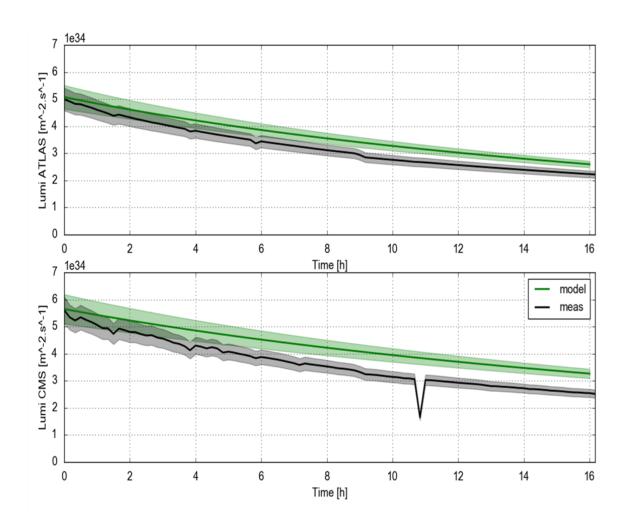
- Using as input only the initial values for the transverse emittance, bunch length and bunch intensity
- Larger emittance blow up observed than expected
- The prediction is not good also for the bunch length and bunch intensity
- Missing sources of extra emittance blow up and losses to be understood

# Luminosity evolution prediction



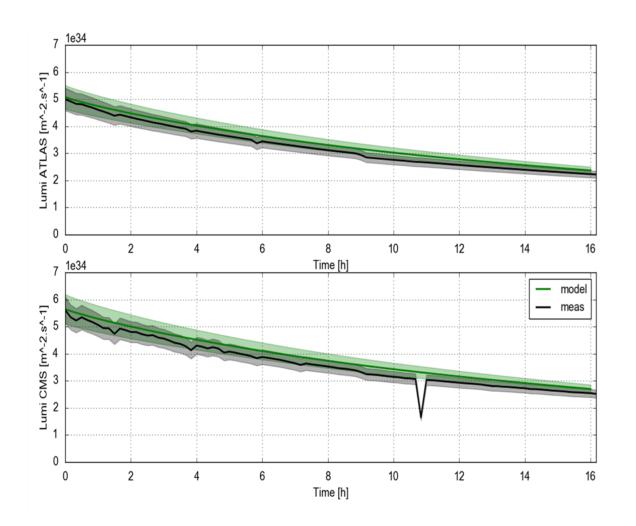
- For each Fill the model is applied under different assumptions and the luminosity evolution is calculated bunch-by-bunch
- Comparison of the average evolution assuming the ideal case: IBS+SR+Burnoff
- Only initial bunch parameters are taken from the data and then the model is iterated to predict the evolution of emittances, bunch length, bunch intensity and luminosity

# Luminosity evolution prediction



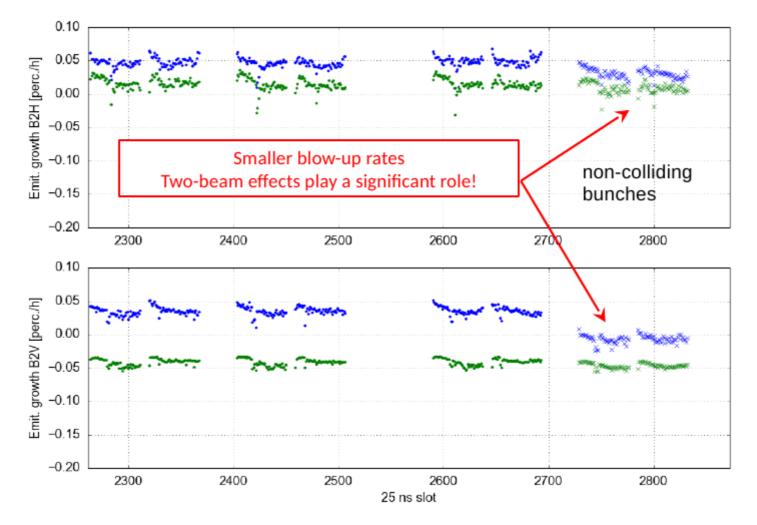
- For each Fill the model is applied under different assumptions and the luminosity evolution is calculated bunch-by-bunch
- Comparison of the averaged evolution using the empirical bunch intensity evolution: IBSEmpiricalLosses
- Bunch intensity evolution from the data. The model computes the emittance, bunch length and luminosity evolution

# Luminosity evolution prediction



- For each Fill the model is applied under different assumptions and the luminosity evolution is calculated bunch-by-bunch
- Comparison of the average evolution using the empirical bunch intensity and empirical emittance evolution:
  - EmpiricalBlowUpEm piricalLosses
- Bunch intensity and emittance evolution from the data. The model computes the bunch length and luminosity evolution

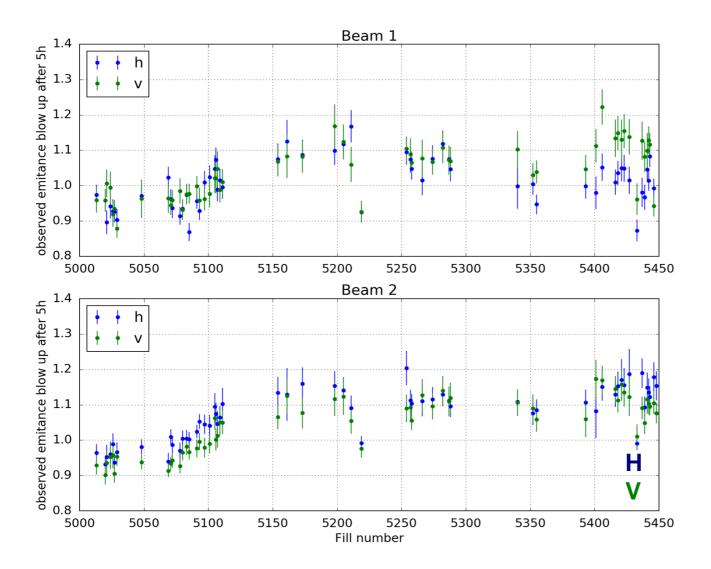
#### **Extra emittance blow-up**



•Fill 5205 went in collision with one non-colliding BCMS train in B2 Ideal to make comparisons

•Non-colliding bunches blow-up less  $\rightarrow$  2-beam effects play a significant role

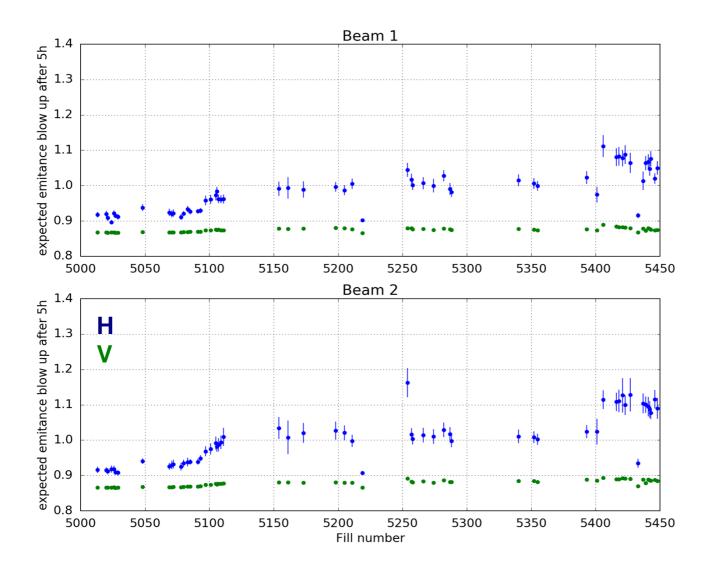
### **Observed emittance blow-up**



 Emittance growth within ±0.1 um/h

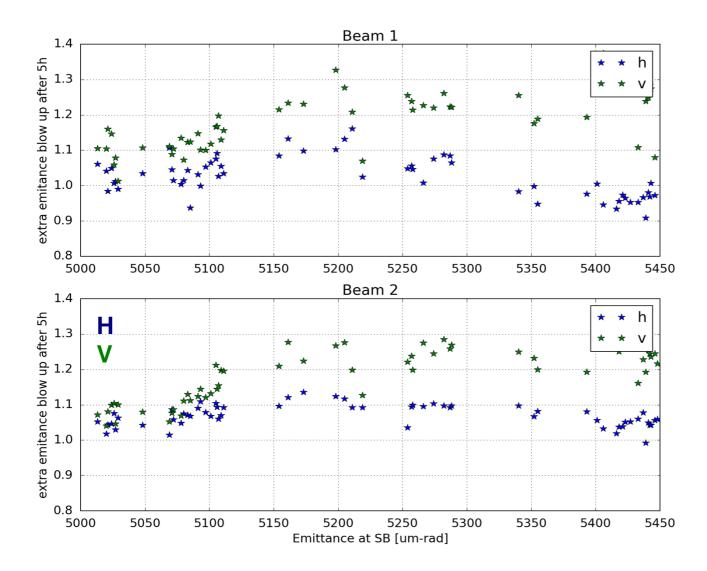
\*Observed emittance blow up = emittance after 5h in SB / emittance at the beginning of SB

#### **Expected emittance blow-up**



\*Expected emittance blow up = pure model prediction after 5h in SB / emittance at the beginning of SB

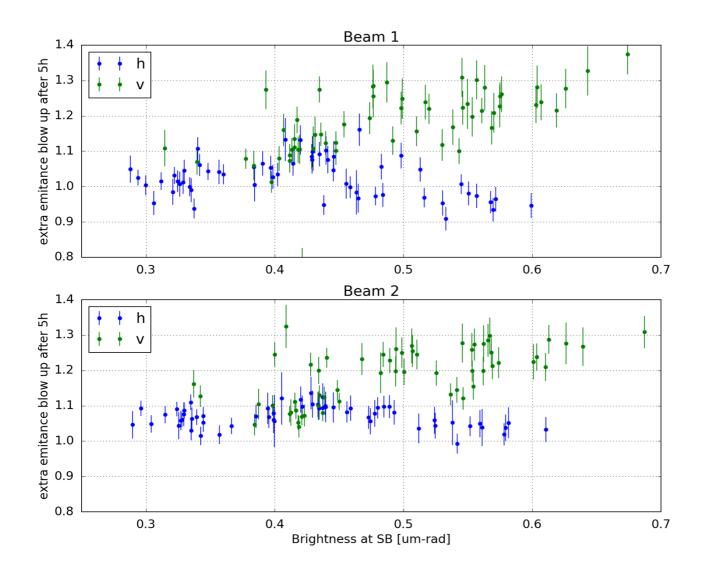
### **Extra emittance blow-up**



 More blow up observed in the vertical plane

\*Extra emittance blow up = observed emittance blow up / expected emittance blow up

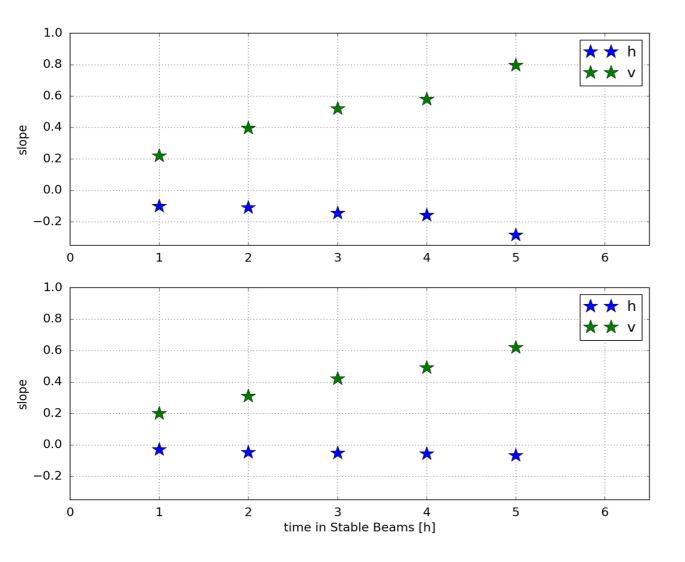
## **Brightness correlations**



- Very weak
  brightness
  dependance
  - Mainly in the vertical plane

\*Extra emittance blow up = observed emittance blow up / expected emittance blow up

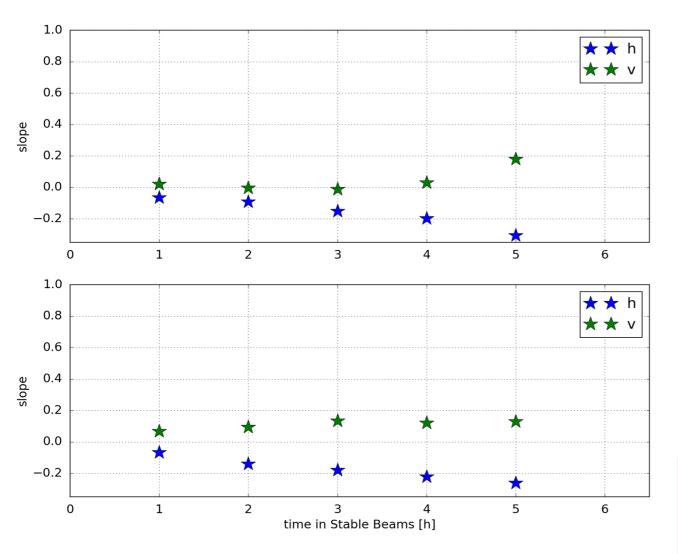
# Brightness correlations – colliding bunches



- For every fill a linear fit was applied to the extra emittance blow up versus initial brightness, after 1-5h in Stable Beams
- Only colliding bunches are used
- The slope of the linear fit versus the time in stable beams is shown in the plots
- Brightness depentence of the extra emittance blow up in the vertical plane

\*Extra emittance blow up = observed emittance blow up / expected emittance blow up

# **Brightness correlations – noncolliding bunches**



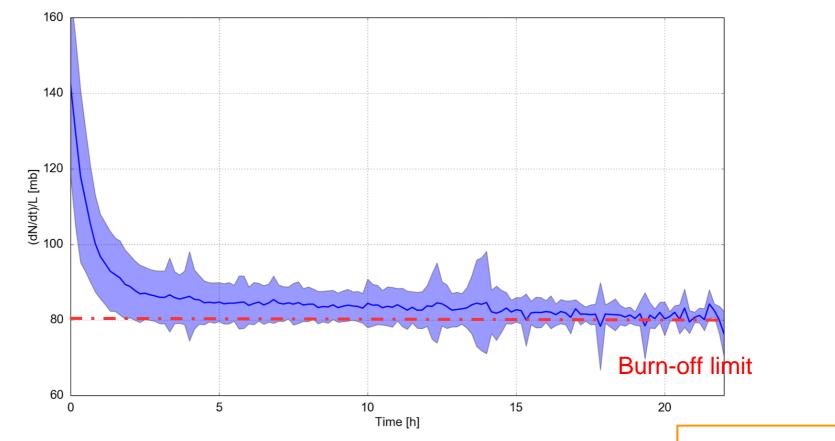
\*Extra emittance blow up = observed emittance blow up / expected emittance blow up

- For every fill a linear fit was applied to the extra emittance blow up versus initial brightness, after 1-5h in Stable Beams
- Only non-colliding bunches are used
- The slope of the linear fit versus the time in stable beams is shown in the plots
- No brightness dependence is observed in both planes

#### To be done:

Further investigation of the observations More detailed bbb analysis

#### **Beam Losses**

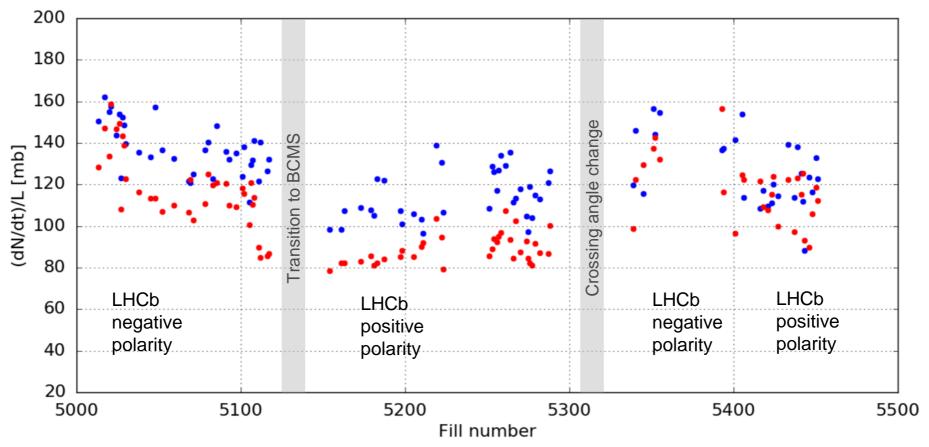


- Normalized loss rate for all fills
- Losses on-top of Burn-off were observed for many fills
- Mainly the first 3h and then become burn off dominated

To be done: Check the convergence limit bunch-by-bunch and fill by fill

#### **Beam Losses**

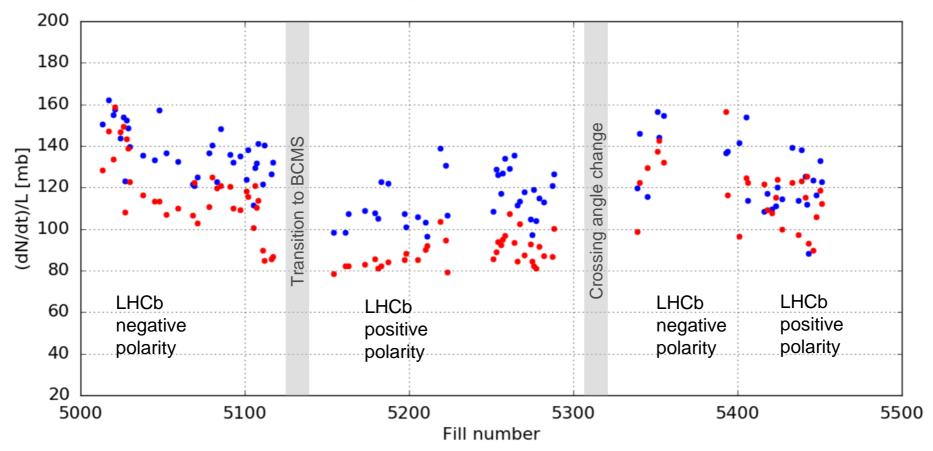
Averaged over the first 1.0h



- Evolution of the average normalized losses (after one hour in SB) along the run
- Beam 1 losses higher than Beam 2 losses
- Minimum losses after the transition to BCMS (Beam 2 losses become burn-off dominated)
- Increase of losses after the crossing angle change followed by an improvement trend
  - Clear impact of the LHCb polarity changes

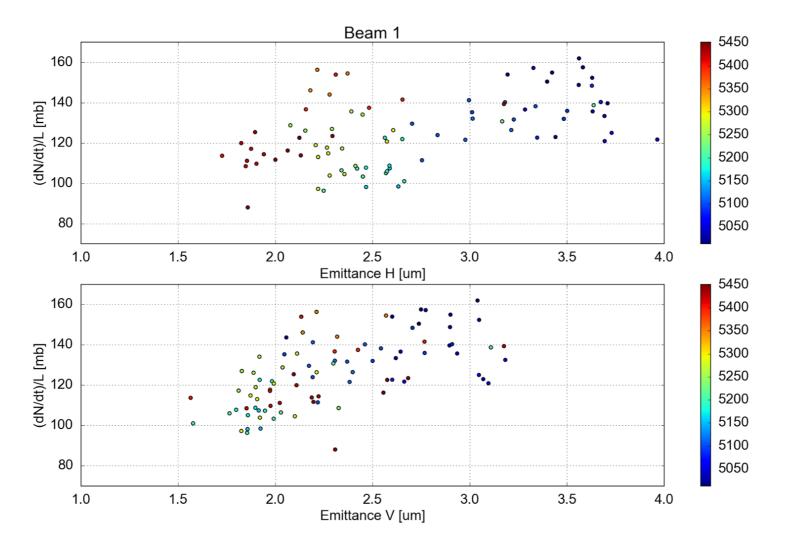
#### **Beam Losses**

Averaged over the first 1.0h



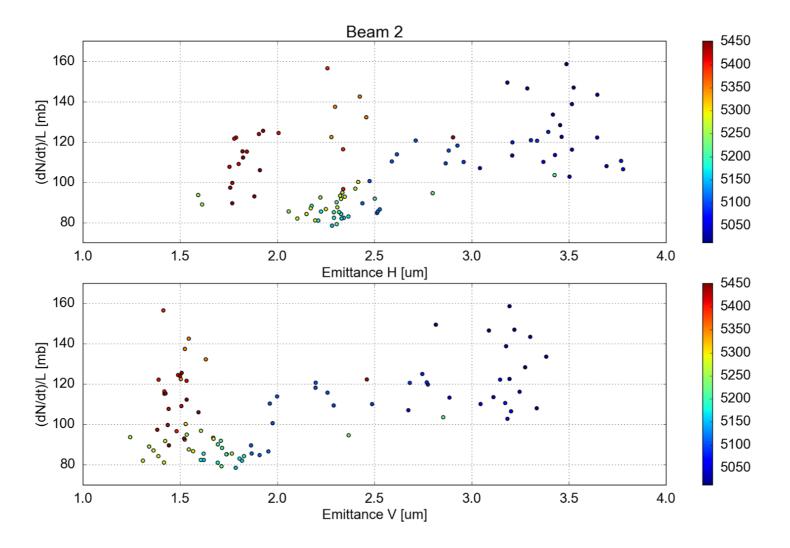
To be done: Estimate the integrated losses bunch-by-bunch and fill-by-fill at different time moments and the end of each fill. How this evolves in time and when does it reach the theoretical burn-off limit?

## **Beam Losses**



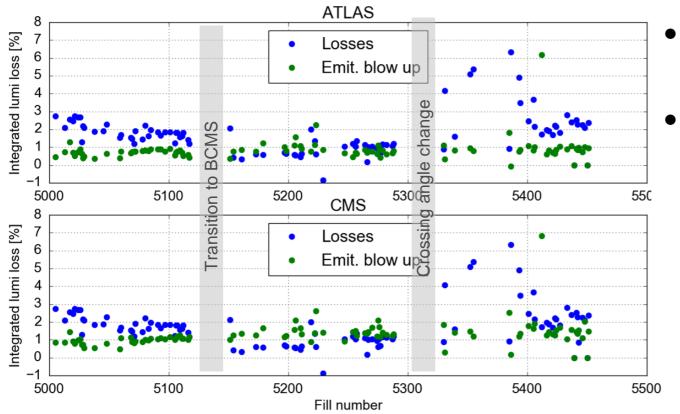
.Losses correlation with the emittance at the beginning of Stable Beams

## **Beam Losses**



.Losses correlation with the emittance at the beginning of Stable Beams

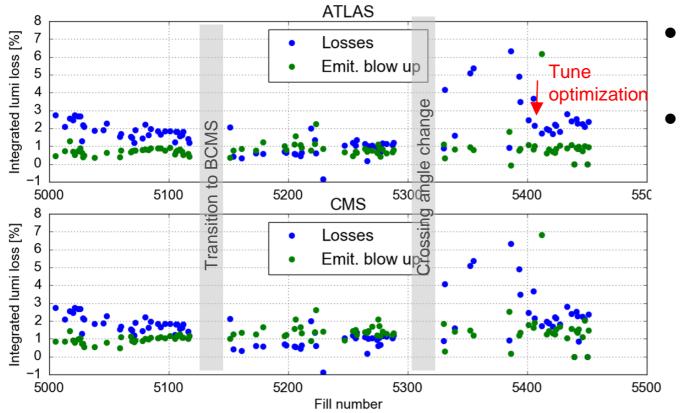
# Integrated luminosity loss due to different degradation mechanisms



- The integrated luminosity over the first 3h is calculated for each model assumption
   Integrated luminosity loss
  - due to:
    - extra losses
    - extra emittance blow up

- Contribution of the extra emittance blow-up is constant over the year
- Contribution of extra losses is sensitive to changes in the machine

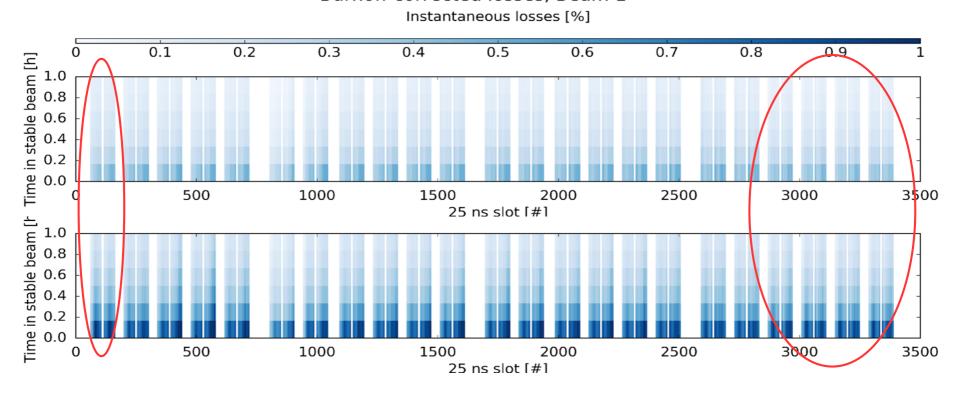
# Integrated luminosity loss due to different degradation mechanisms



- The integrated luminosity over the first 3h is calculated for each model assumption
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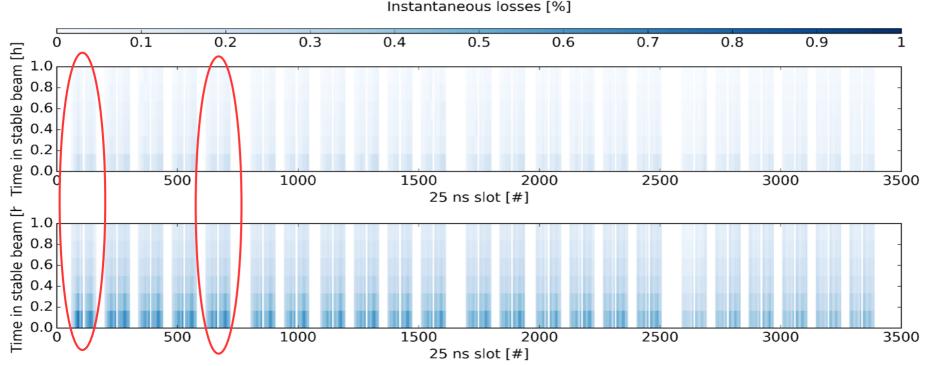
- Contribution of the extra emittance blow-up is constant over the year
- Contribution of extra losses is sensitive to changes in the machine

## Instantaneous beam losses before and after the crossing angle change



- Burn-off corrected losses averaged over many fills for Beam 1
  - Top: crossing angle of 185 murad
  - Bottom: crossing angle of 140 murad
- More losses observed at the end of the trains for the large crossing angle  $\rightarrow$  e-cloud traces
- More losses at the middle of many trains (with full LR encounters) are observed after the crossing angle change
  - The effect is more pronounced during the first 30 min.

## Instantaneous beam losses before and after the crossing angle change



- Burn-off corrected losses averaged over many fills for Beam 2
  - Top: crossing angle of 185 murad
  - Bottom: crossing angle of 140 murad
- More losses observed at the end of the trains for the large crossing angle  $\rightarrow$  e-cloud traces
- More losses at the middle of many trains (with full LR encounters) are observed after the crossing angle change
   To be done: Extend the analy
  - The effect is more pronounced during the first 30 min.
  - Less pronounced than for Beam 1

To be done: Extend the analysis trying to verify correlations with long ranges

## **Other interesting studies**

#### • To be done..

- Application of the analysis to the 2015 data and compare observations
- Can we put all Run II together?
- Same with 2012 data
- To be followed up
  - Luminosity predictions for HL-LHC, HE LHC, etc..
    - Nikos already took over this part

## Summary

- Luminosity follow up tools are set up profiting from the consolidated experience with the scrubbing follow up (thanks to Gianni!)
  - Cleaning up is needed!
- There are three main parts of the analysis
  - Emittance evolution from injection to stable beams
    - Is the observed extra emittance blow-up due to noise effects?
  - Emittance evolution during stable beams
    - Interesting observations
    - The effect is different between colliding and non-colliding bunches
    - Is the non-brightness dependent part coming from noise effects?
  - Extra losses during collisions
    - After how long in SB the theoretical limit is reached for each fill?
    - Long range correlations? Can we verify the correlation of losses with long ranges?

## Summary

- It has been a very interesting path, quite busy most of the times :)
- Lots of work has been already done and many open fronts to be continued
- Ilias Efthymiopoulos is taking over the activity with the usual suspects around him and hopefully new people to join
  I am still around if support is needed
- Some documentation has been put together and I will try to clean it up soon
- ★ I would like to take this opportunity to give my special thanks to Yannis for his support and to Gianni for all the help and input!!

## **Relevant presentations and proceedings**

https://indico.cern.ch/event/368172/contributions/1784182/attachments/732013/100431 1/LHC\_Luminosity\_modeling2.pdf

IPAC 2015: <u>http://accelconf.web.cern.ch/AccelConf/IPAC2015/papers/tupty020.pdf</u>

Evian 2015: https://indico.cern.ch/event/434129/contributions/1917206/attachments/1205402/18302 41/LumiModel-Evian2015.pdf

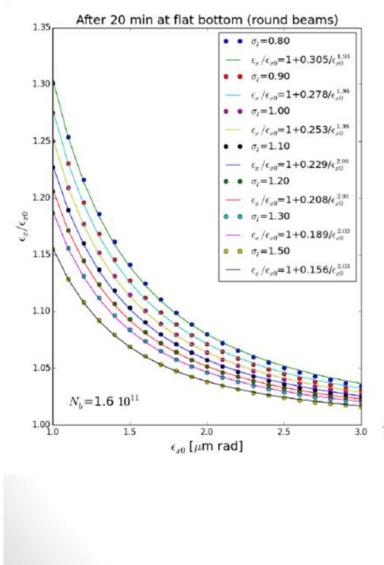
IPAC 2016: <a href="https://cds.cern.ch/record/2207357/files/tupmw002.pdf">https://cds.cern.ch/record/2207357/files/tupmw002.pdf</a>

Evian 2016: <u>https://indico.cern.ch/event/578001/contributions/2366376/attachments/1388316/21137</u> <u>83/F.Antoniou\_Evian2016.pdf</u> Evian 2016 proceedings to be published soon

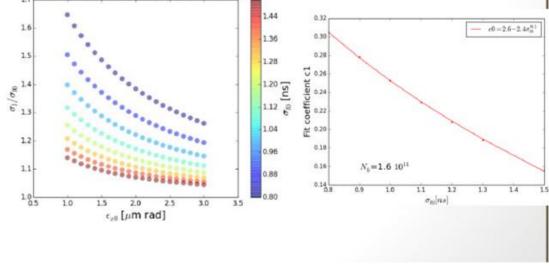
#### Thank you for your attention!

## **Extra slides**

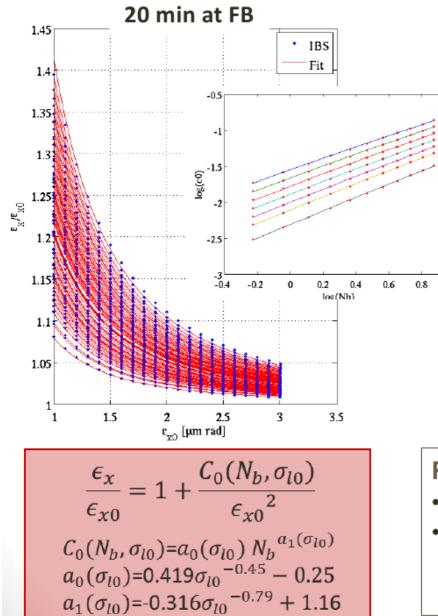
## The IBS + SR model

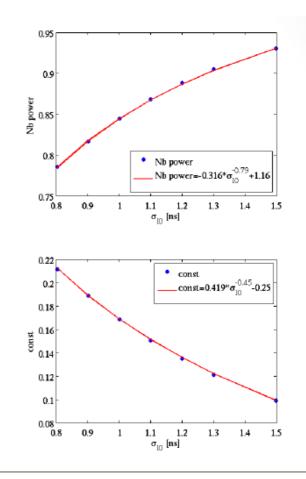


- Parameterization of the IBS effect at the horizontal and longitudinal plane with different initial horizontal emittance and bunch length (N<sub>b</sub>=1.6e11)
- Simple scaling laws can be found
- For ε<sub>t0</sub>=2µm and σ<sub>l</sub>=1.2ns (red line) the effect at flat bottom is of the order of 5%



## The IBS + SR model

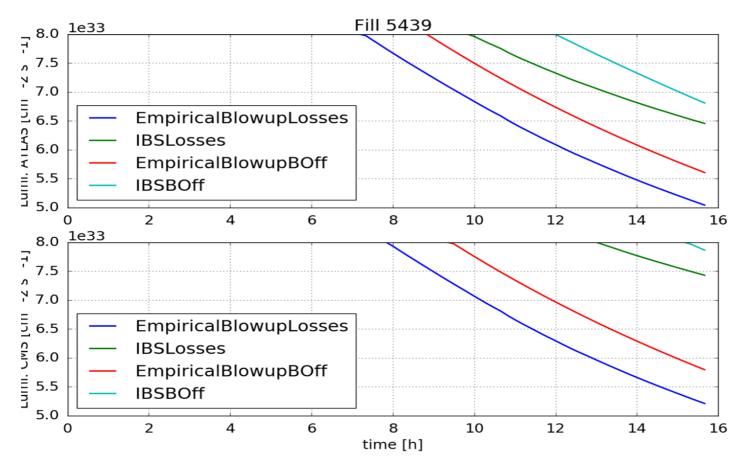




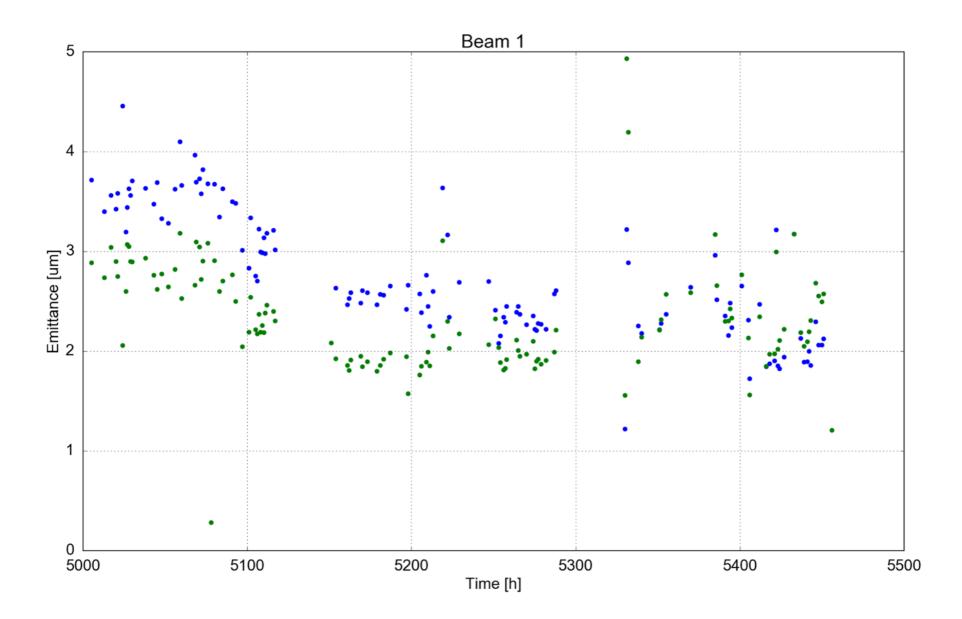
#### **Further comments**

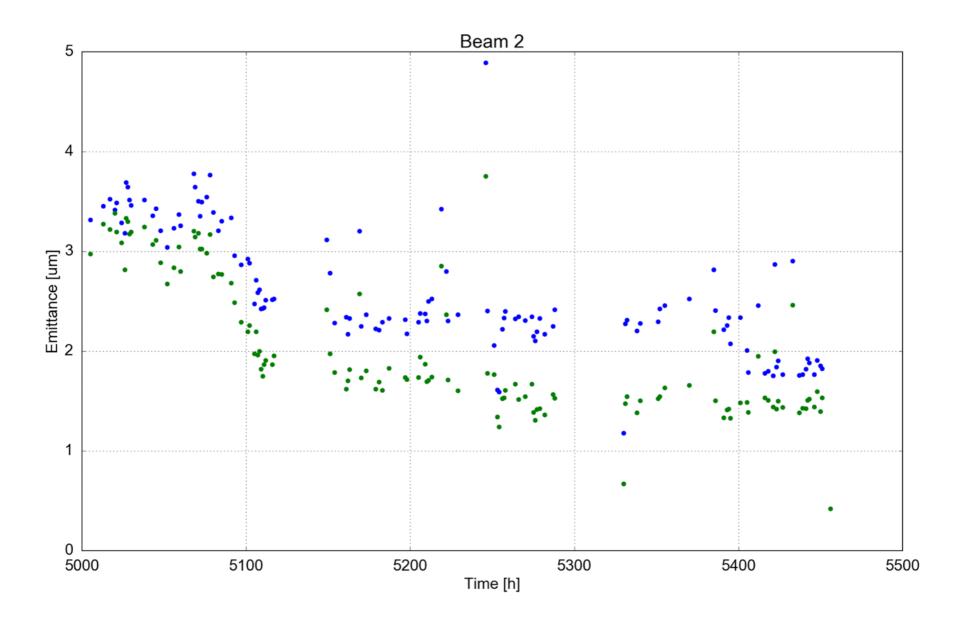
- No effect in the vertical plane
- The memory of the IBS effect in the longitudinal plane will be "erased" during the ramp

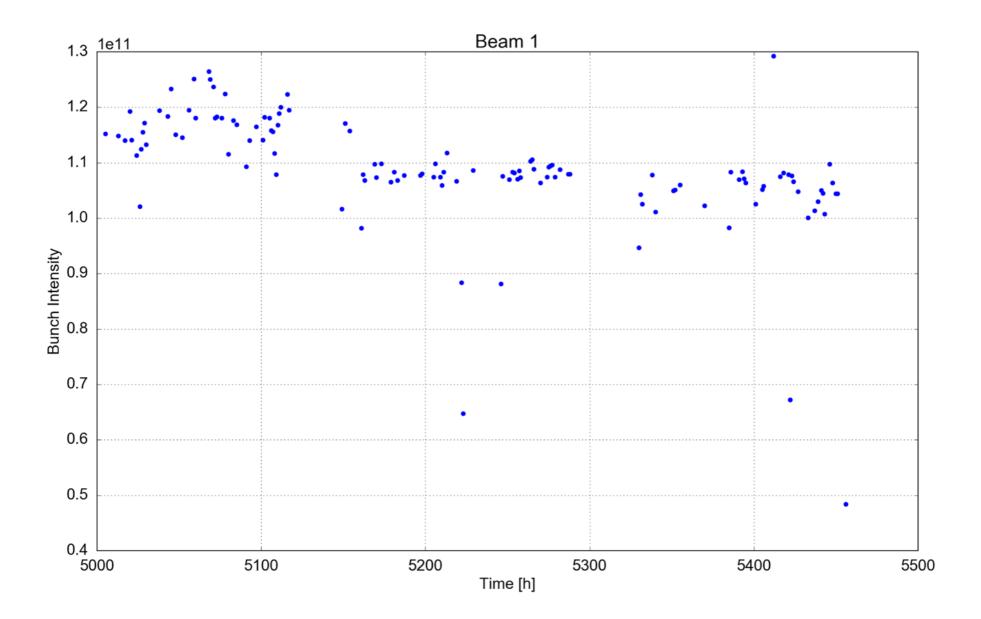
## **Leveling Fills**

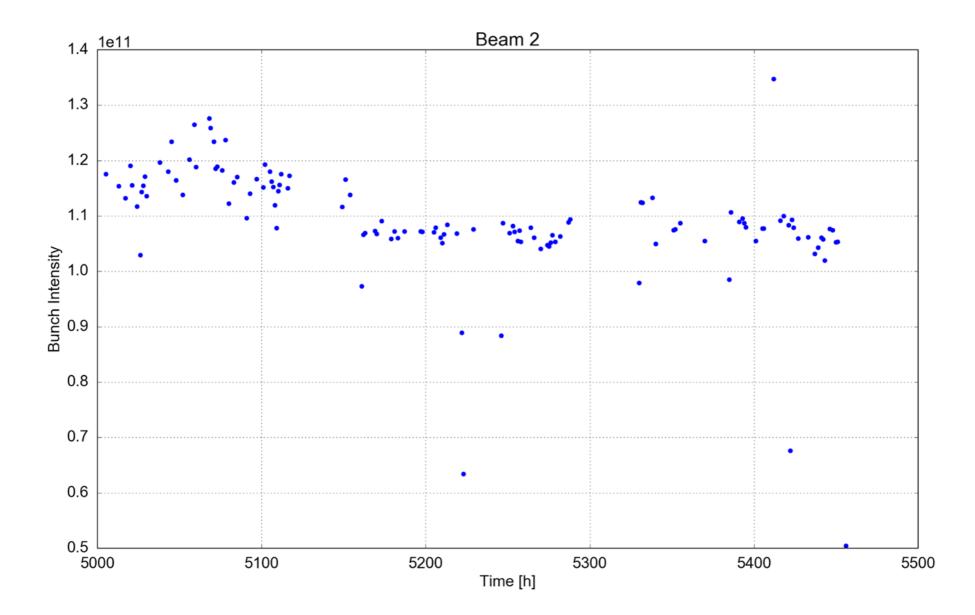


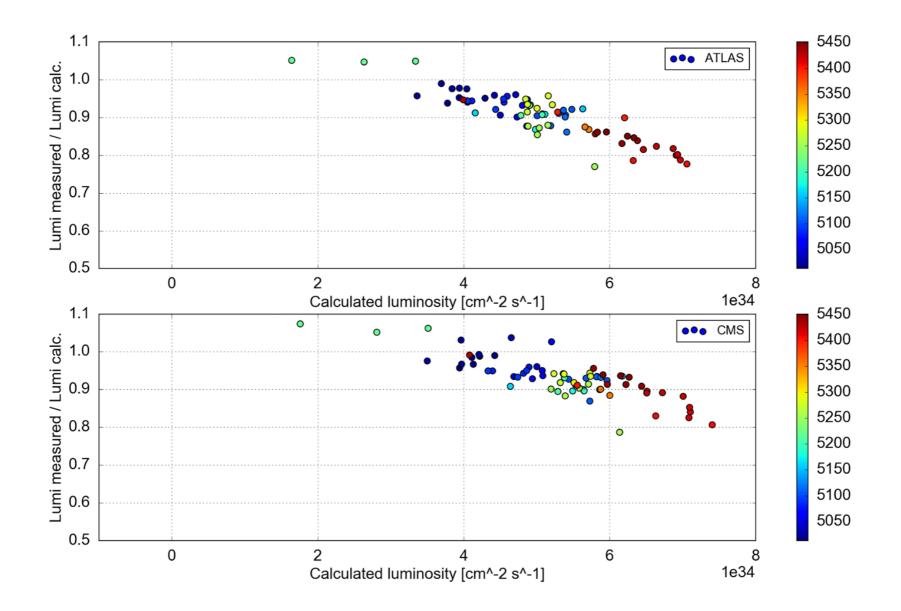
 Impact of the different degradation mechanisms on the leveling time based on the lumi model



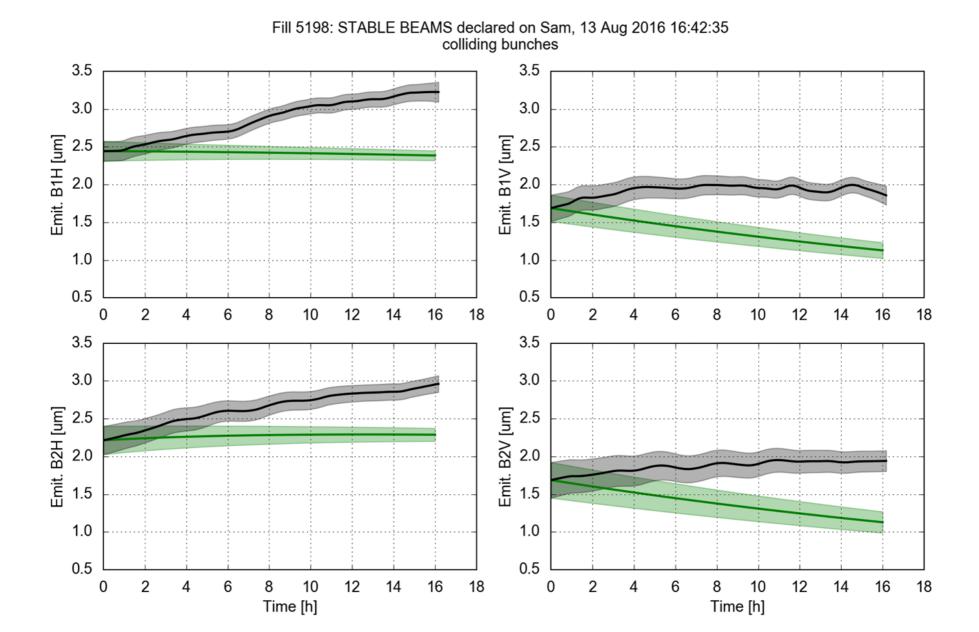




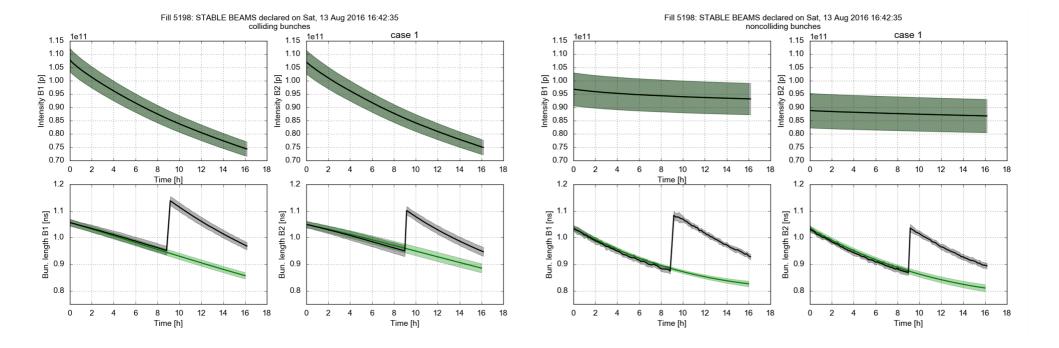




#### Luminosity modeling



# Luminosity model comparison with data: Bunch length

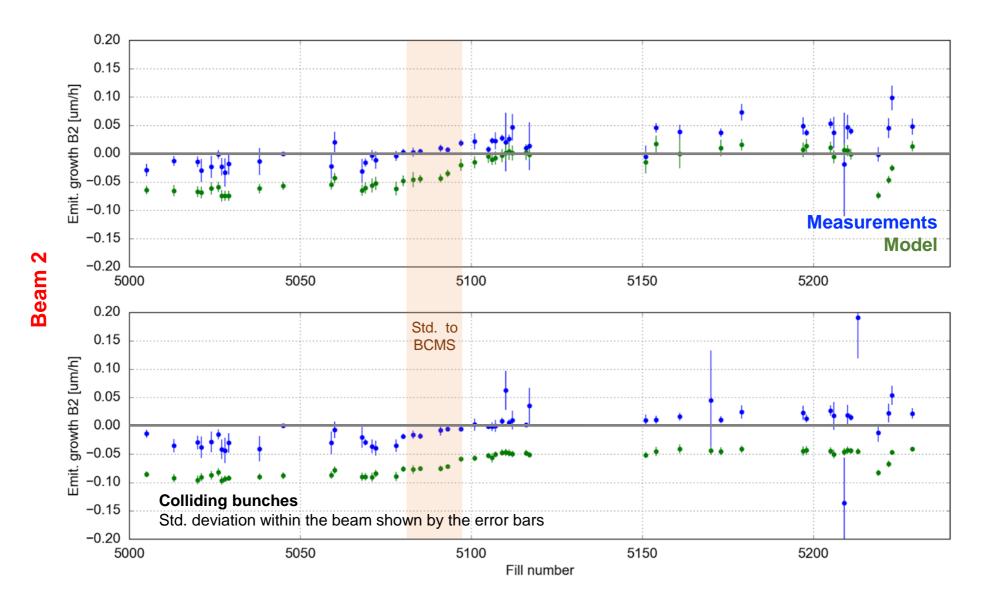


#### **Emittance evolution in Stable Beams**

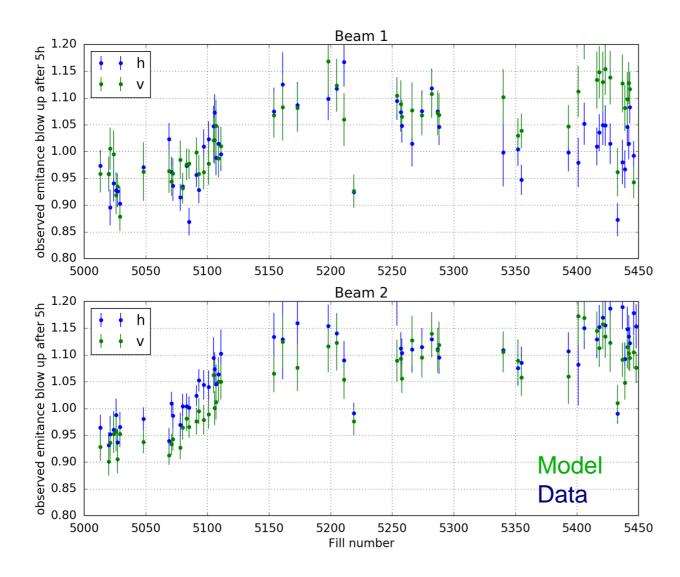
- CERN

•Emittance growth within ±0.1 um/h (~10 times less than injection), changing with the beam brightness •Both planes show an additional blowup of ~0.5 um/h with respect to the model

 $\rightarrow$  The difference between H and V is consistent with IBS



## **Observed emittance blow-up**



•Emittance growth within ±0.1 um/h

Both planes show an
 additional blowup of ~0.05
 um/h with respect to the
 model
 →The difference

between H and V is consistent with IBS