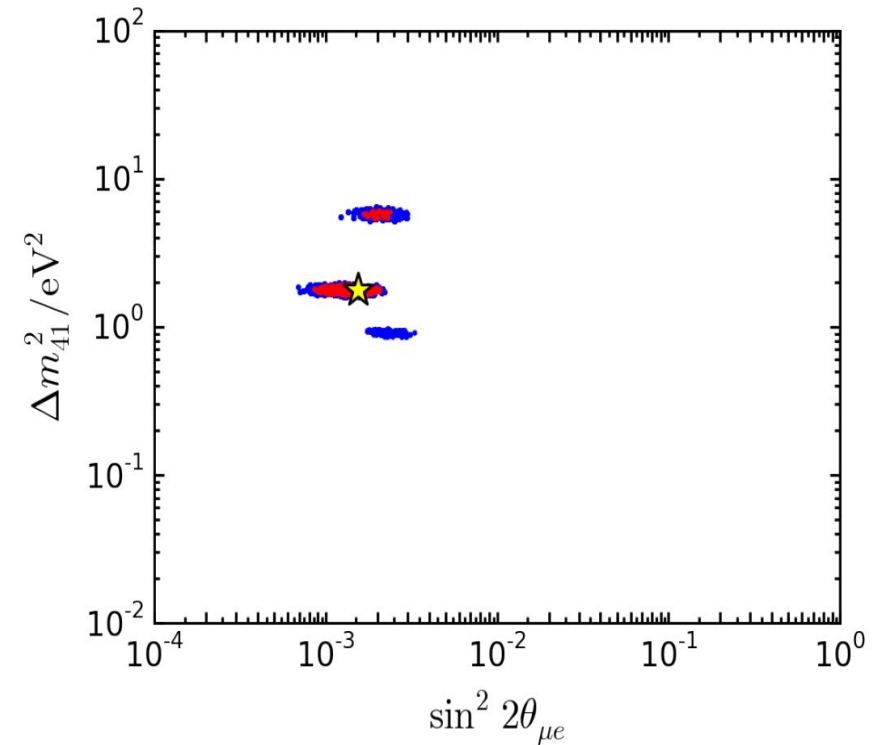


# Searching for Sterile Neutrinos at J-PARC with JSNS<sup>2</sup>

Johnathon Jordan - University of Michigan  
TeVPA 2017  
August 10, 2017

# Status of Sterile Neutrino Searches

- Several anomalous results in last few decades:
  - LSND ( $3.8\sigma$ )
  - MiniBooNE ( $3.8\sigma$ )
  - Gallium/SAGE ( $2.7\sigma$ )
  - Reactor ( $3.0\sigma$ )
- Results do not fit the standard 3 neutrino paradigm (solar + atmospheric)
- Also many null results (MINOS, MiniBooNE, and IceCube)
- Possible oscillation to sterile neutrinos
- JSNS<sup>2</sup> will be a direct test of past anomalous results



**Global fit to short baseline and IceCube data using a 3+1 model.**

arxiv:1607.00011

# JSNS<sup>2</sup> Experiment

- The J-PARC Sterile Neutrino Search at the J-PARC Spallation Neutron Source (JSNS<sup>2</sup>)

- Will look for sterile neutrinos using:

$$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$$

- Located at J-PARC Materials and Life Sciences Facility (MLF)
- 24m baseline, will look for oscillations with:

$$\Delta m^2 \approx 1 \text{ eV}^2$$

- Use distinct IBD signature to identify candidate events

$$\bar{\nu}_e + p \rightarrow e^+ + n$$

Technical Design Report (TDR):  
Searching for a Sterile Neutrino at J-PARC MLF  
(E56, JSNS<sup>2</sup>)

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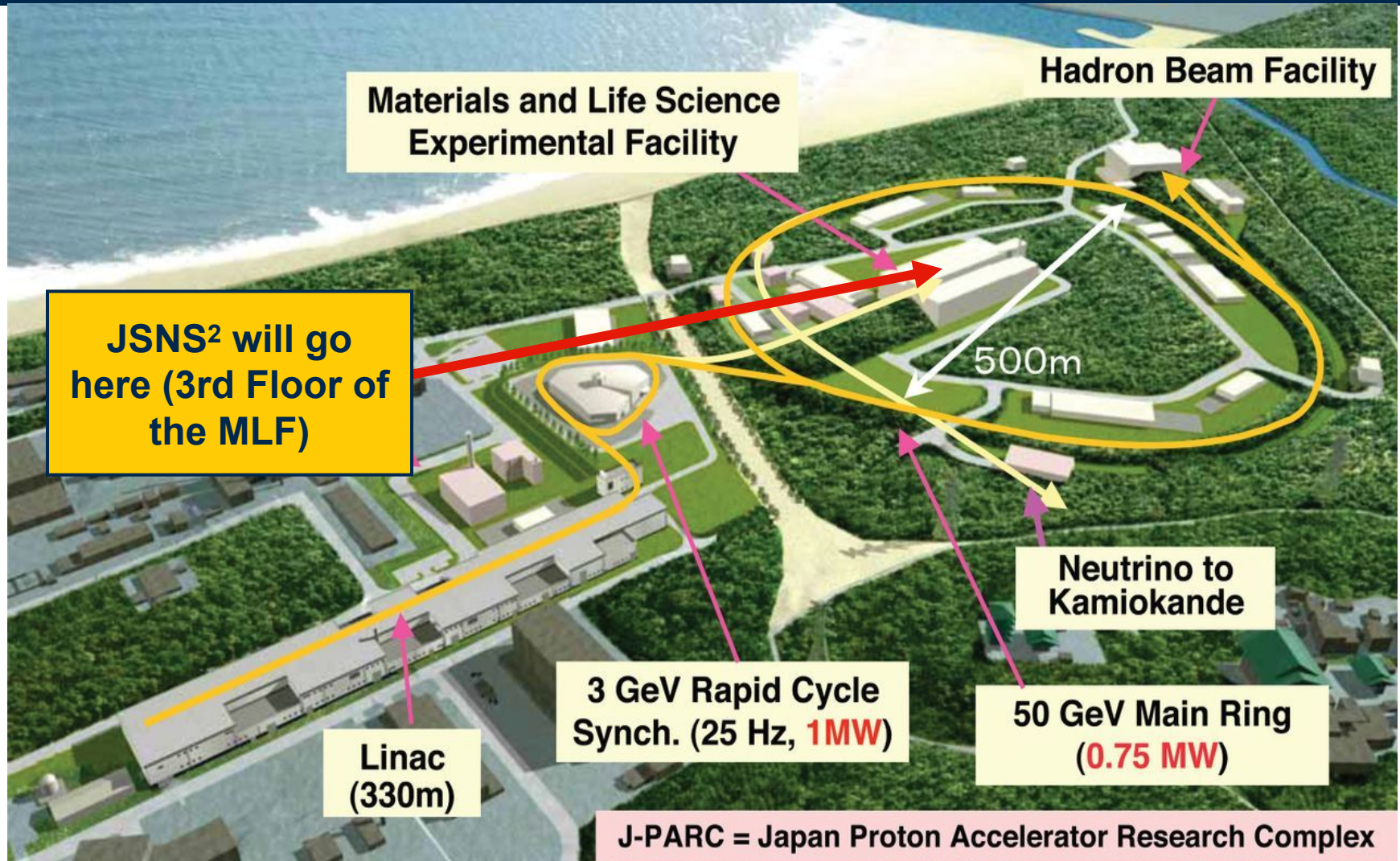
<sup>19</sup>University of Alabama, Tuscaloosa, AL, 35487, USA

<sup>20</sup>Colorado State University, Tuscaloosa, AL, 35487, USA

<sup>21</sup>Brookhaven National Laboratory, Upton, NY, 11973-5000, USA



# J-PARC RCS Beam



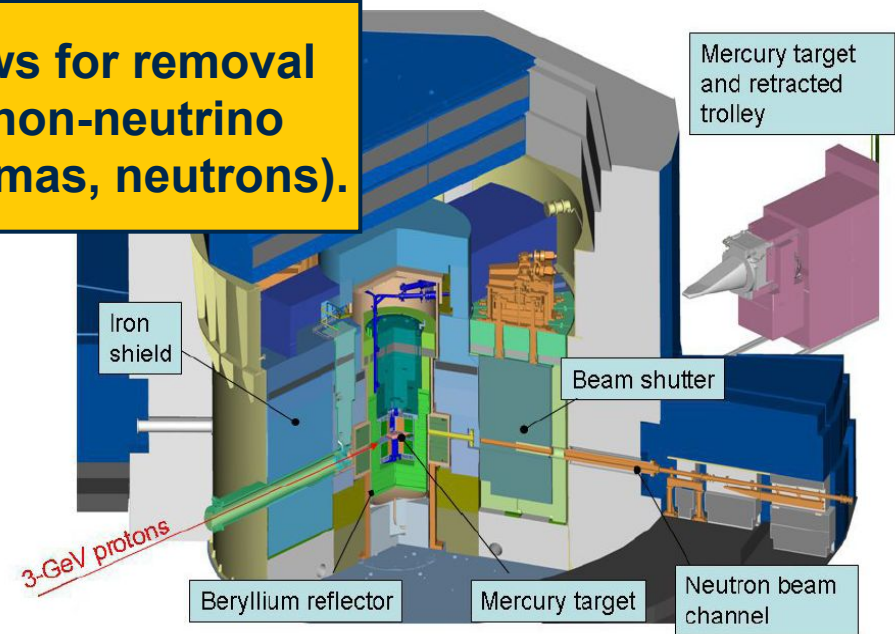
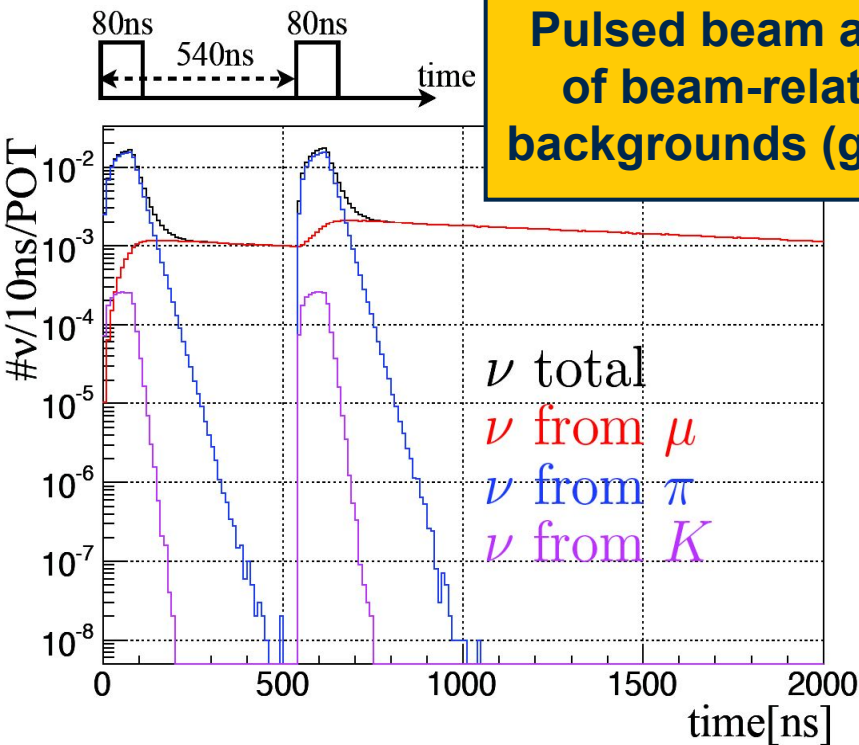


# Detector Location

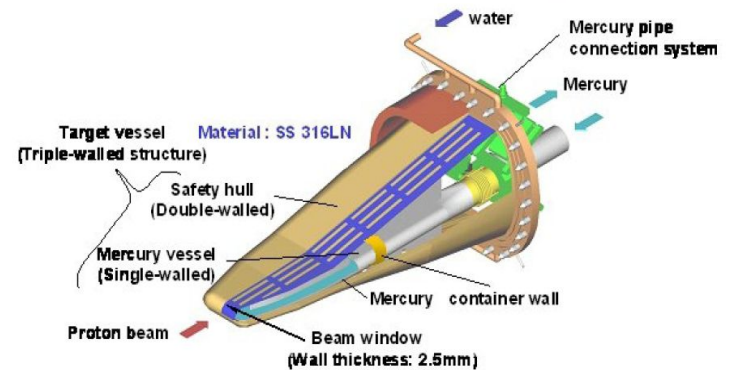
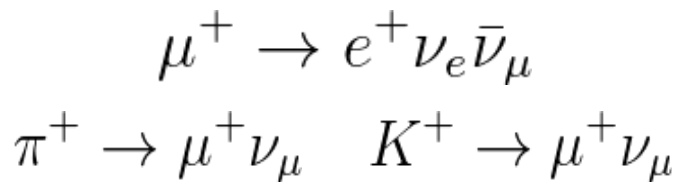
Detector Location  
on MLF 3rd Floor



# J-PARC MLF/Beam

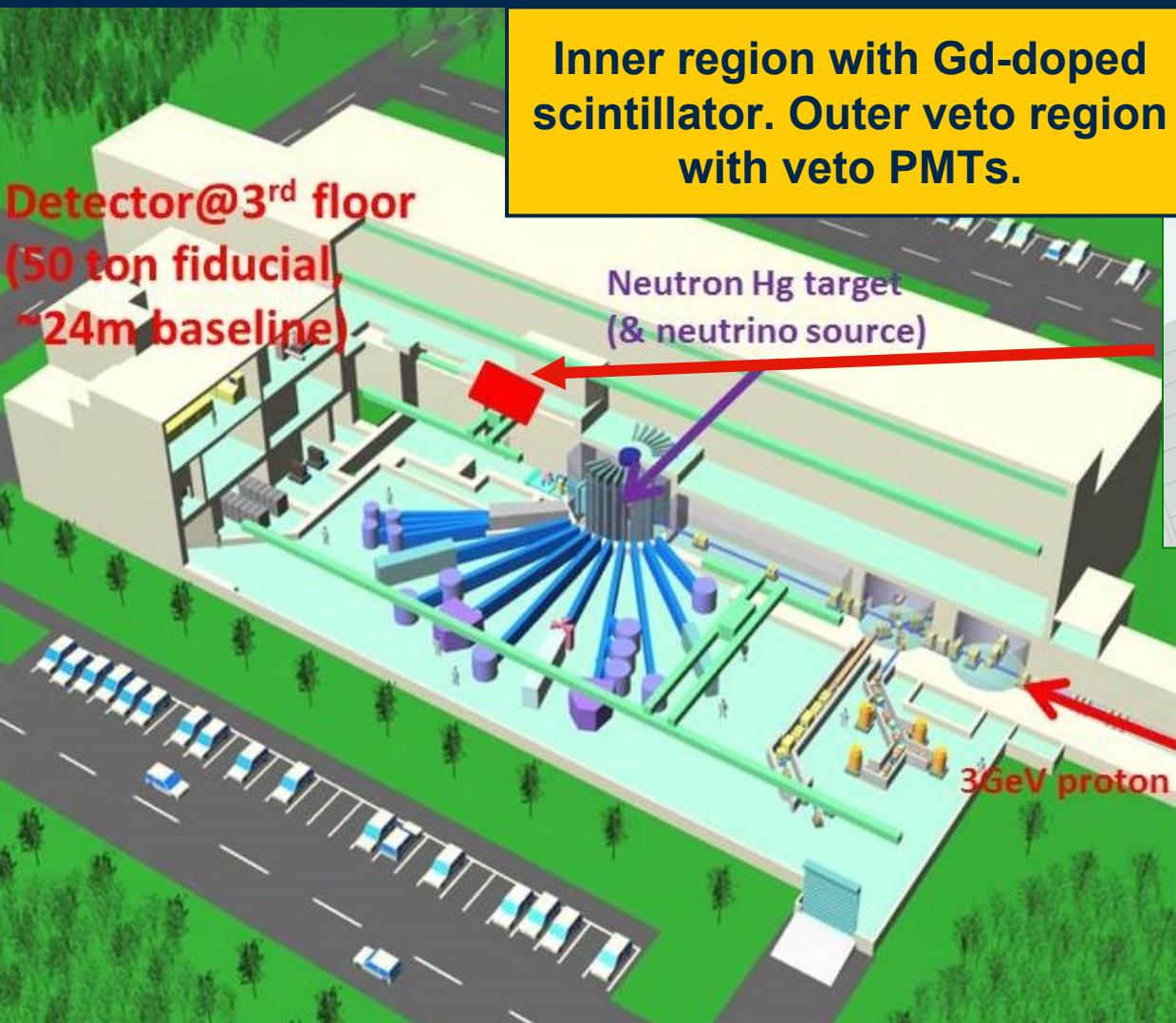


## DAR Production Mechanisms:

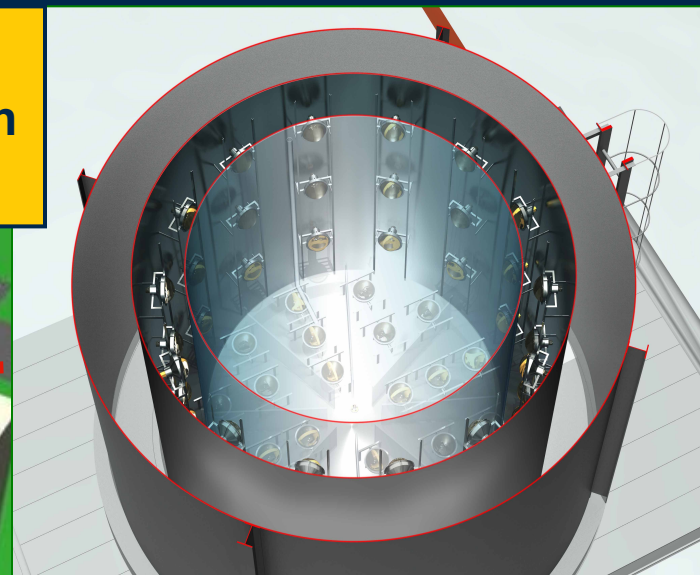




# JSNS<sup>2</sup> Detector



Inner region with Gd-doped scintillator. Outer veto region with veto PMTs.



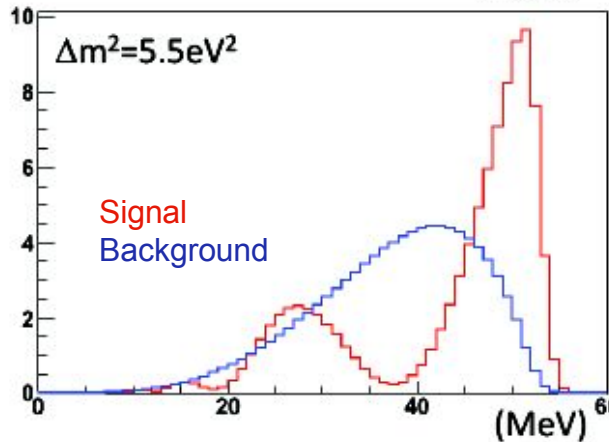
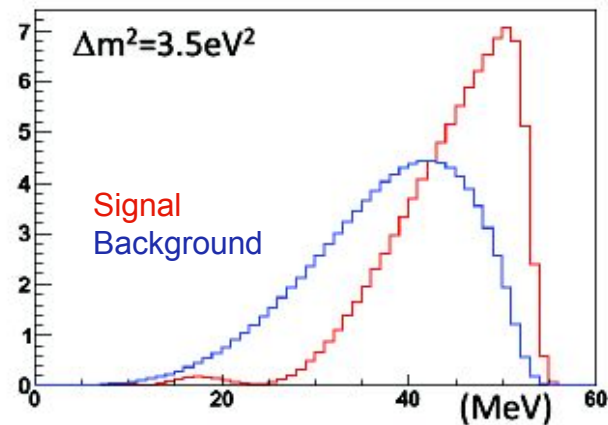
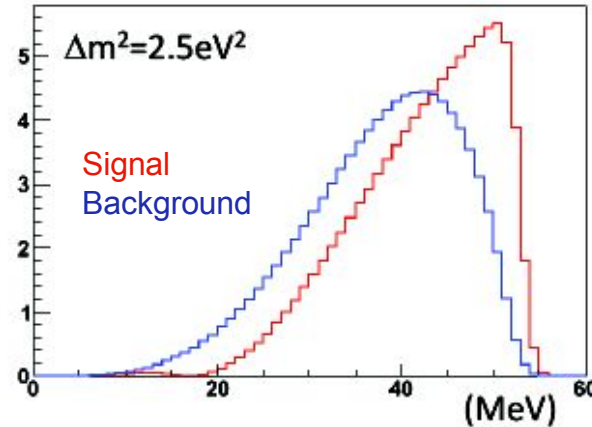
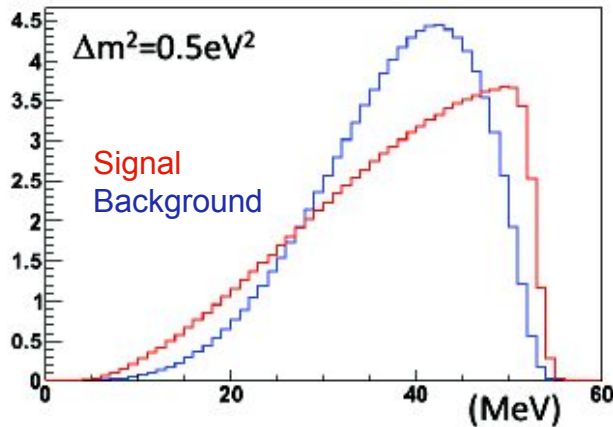
193 PMTs per inner detector

LAB-based liquid scintillator

Phase 0: 17t fiducial mass

Phase 1: 34t fiducial mass

# Intrinsic Background



Dominant irreducible intrinsic background:

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

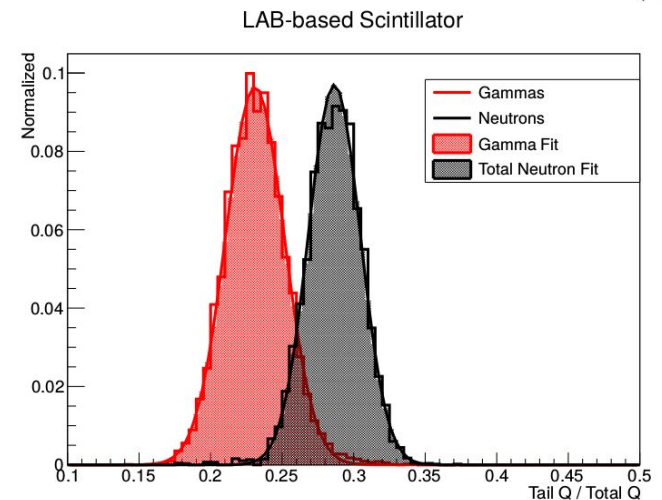
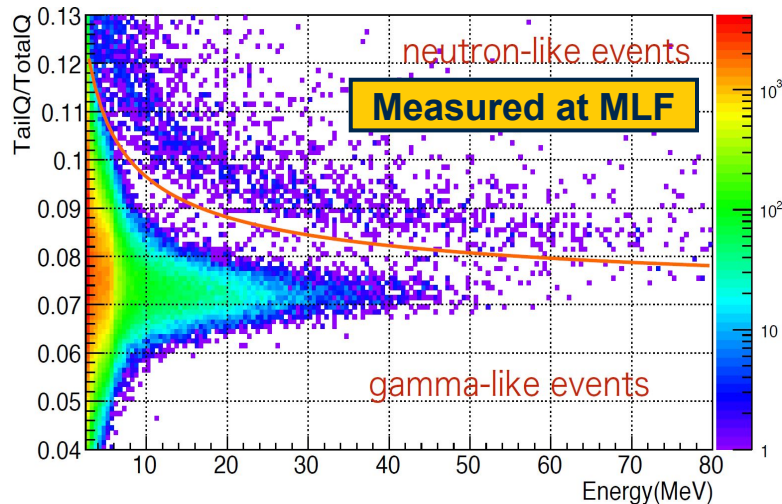
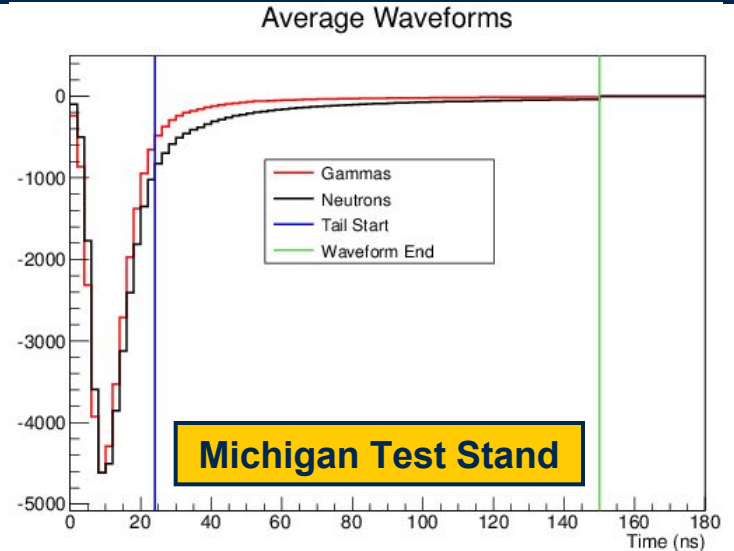
Background shape is different than the signal shape for all oscillation scenarios.

M. Harada et al, arXiv:1310.1437 [physics.ins-det]

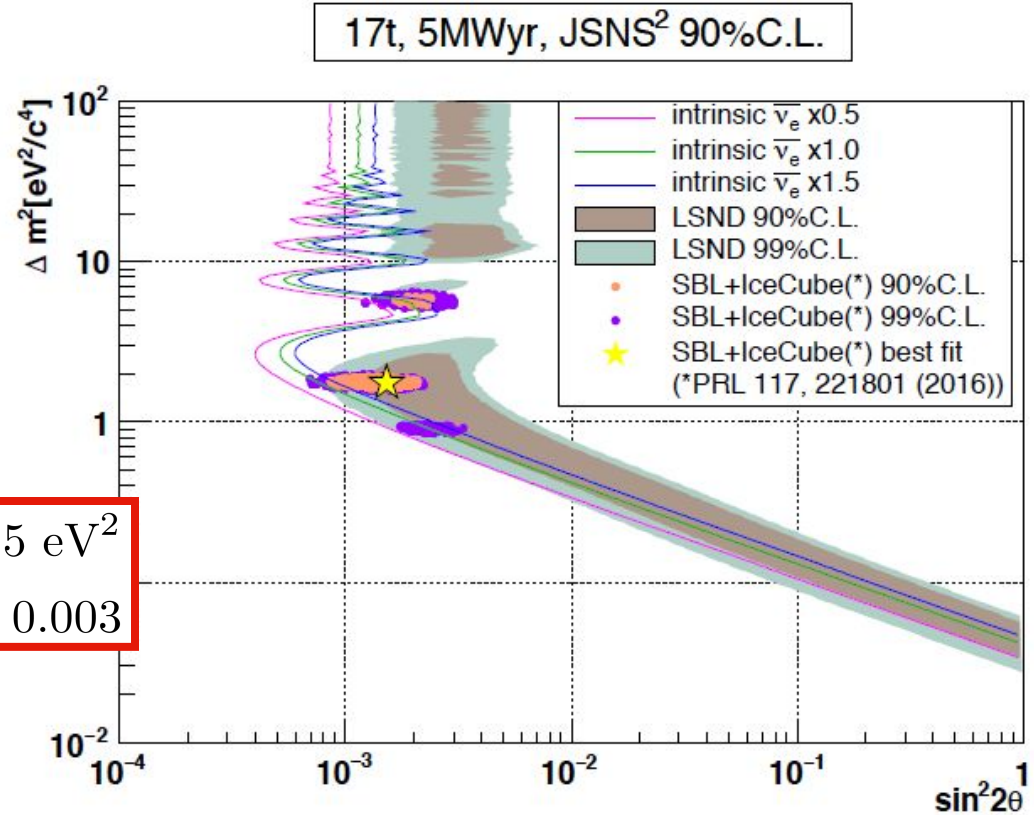
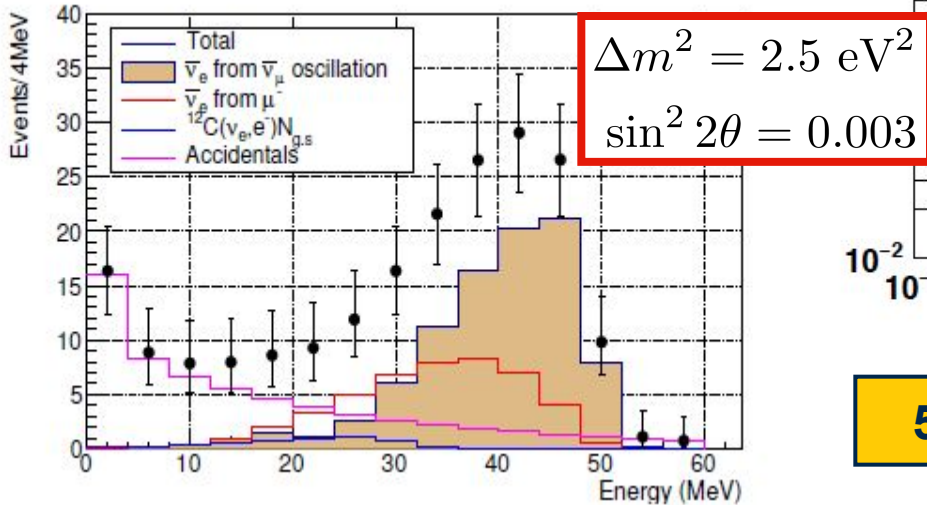
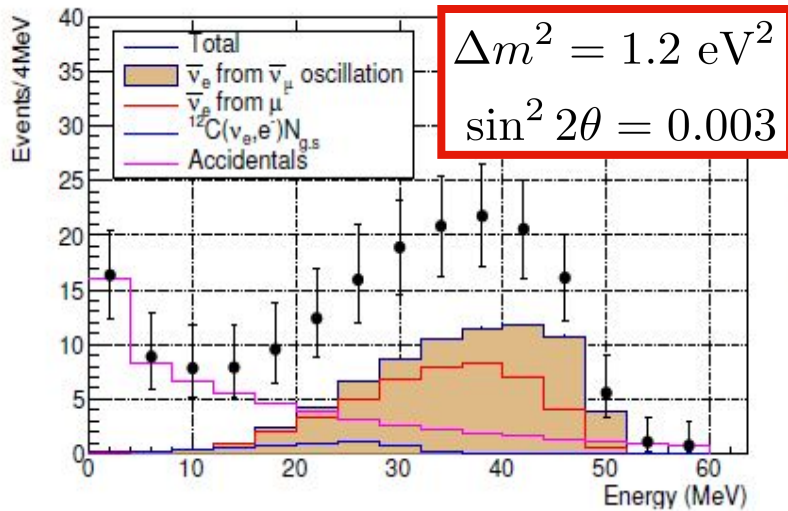


# Backgrounds

- Other major background is from fast (10-100 MeV) neutrons:
  - Cosmic neutrons
  - Beam neutrons (accidentals)
- Reject cosmic neutrons with pulse shape discrimination (PSD)
- Beam and cosmic based background have been measured *in situ*



# Sensitivity

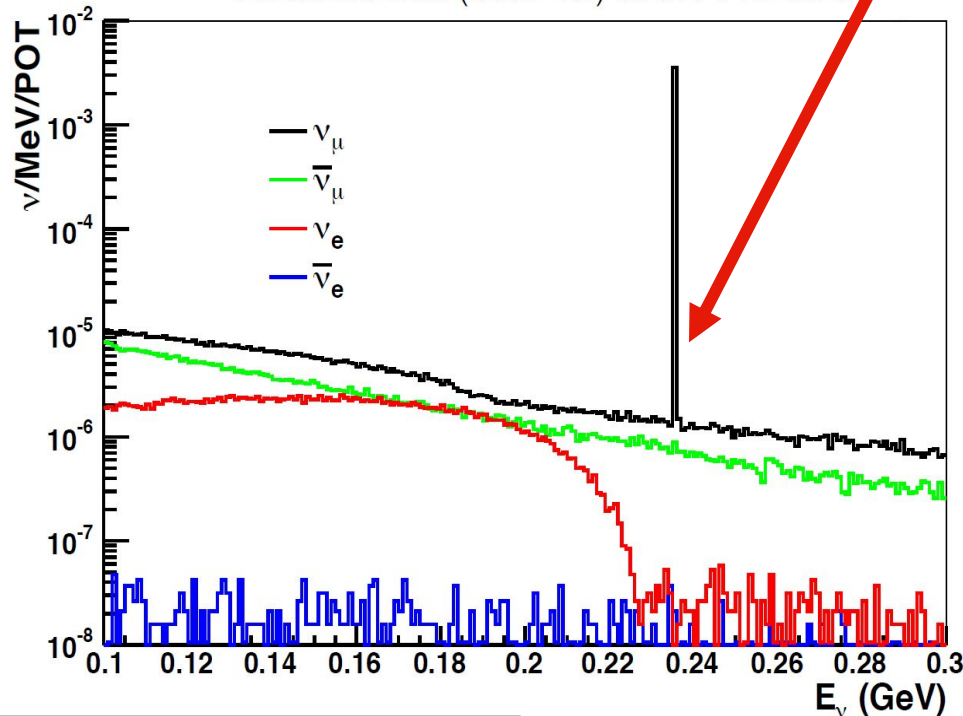


**5 years, 1 MW beam power, 17t fiducial.**

# KDAR Neutrinos

236 MeV  $\nu_\mu$  from  $K^+ \rightarrow \mu^+ \nu_\mu$  (BR = 63.5%) decay at rest

Neutrino flux (over  $4\pi$ ) at JPARC-MLF



- Can be used as a standard candle for reconstruction and cross sections near 236 MeV
- Probe the nucleus with a known energy, weak-interaction-only particle
- Measure energy transfer to test underlying nuclear models

## Expected Event Rate

Detector (source)	Target (mass)	Exposure	Distance from source	236 MeV $\nu_\mu$ CC events
JSNS <sup>2</sup> (JPARC-MLF)	Gd-LS (17 ton)	$1.125 \times 10^{23}$ POT (3 years)	24 m	30-60k



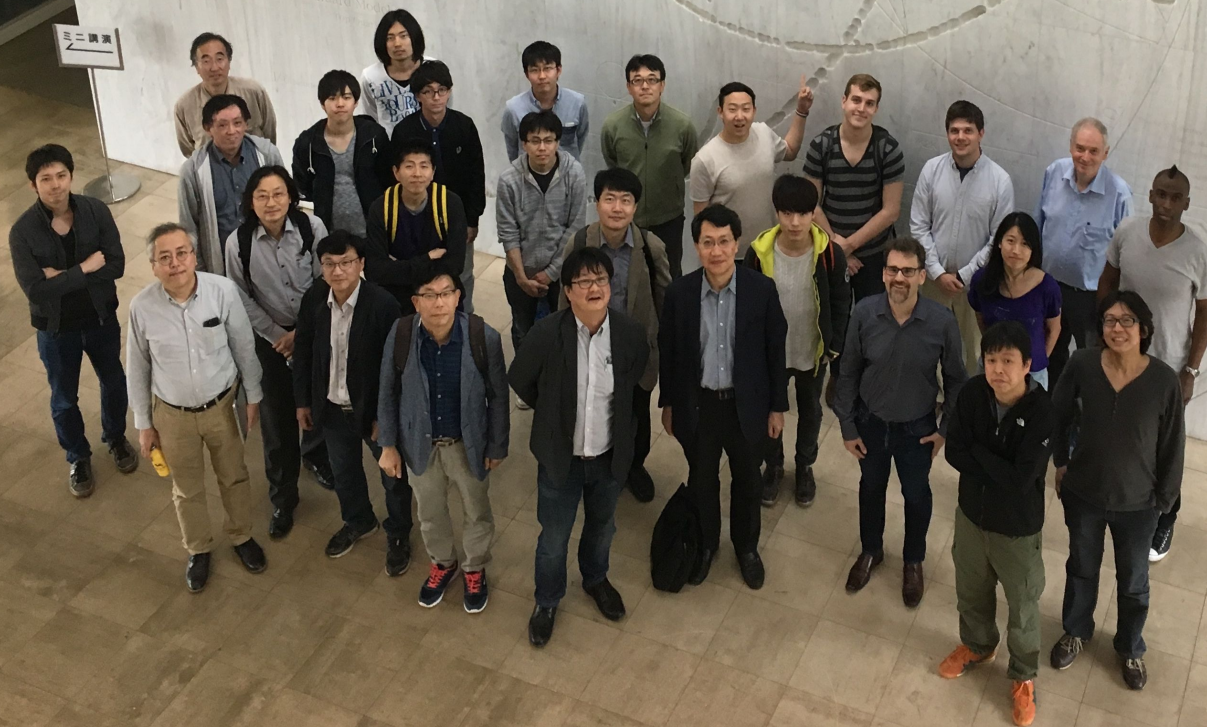
# Timeline and Outlook

- JSNS<sup>2</sup> was conceived in 2013.
- The experiment received Stage 1 approval in Feb. 2015.
- Funding for construction of the first 17-ton detector received in June 2016.
- TDR submitted in May for Stage 2 approval
- Construction has already started
- First data in late 2018

**TDR - arXiv:1705.08629**

# Questions

**JSNS<sup>2</sup> Collaboration at Collaboration Meeting in May 2017**



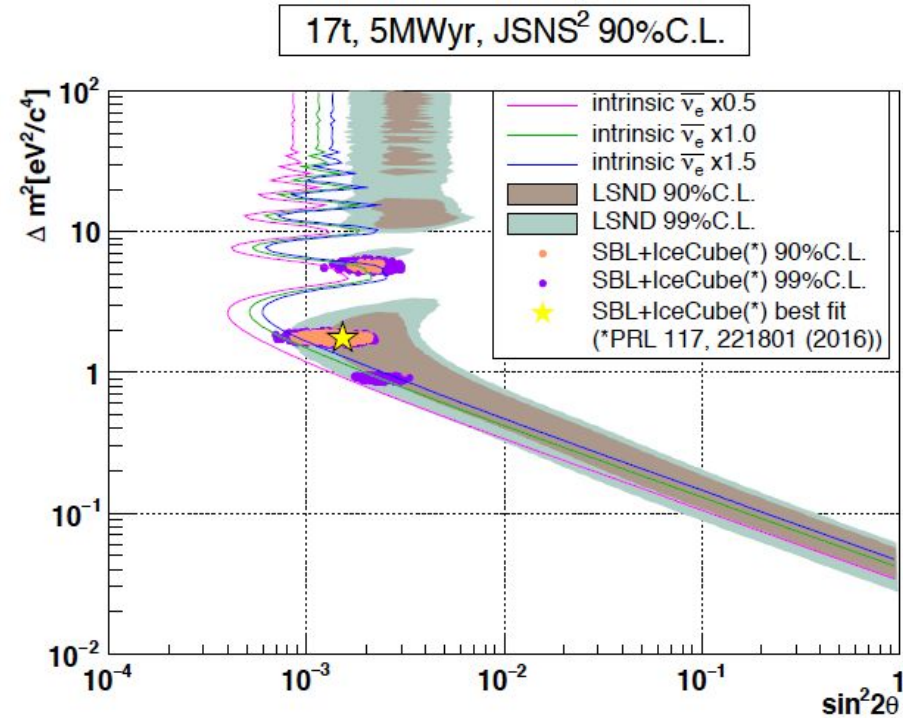
# Backup



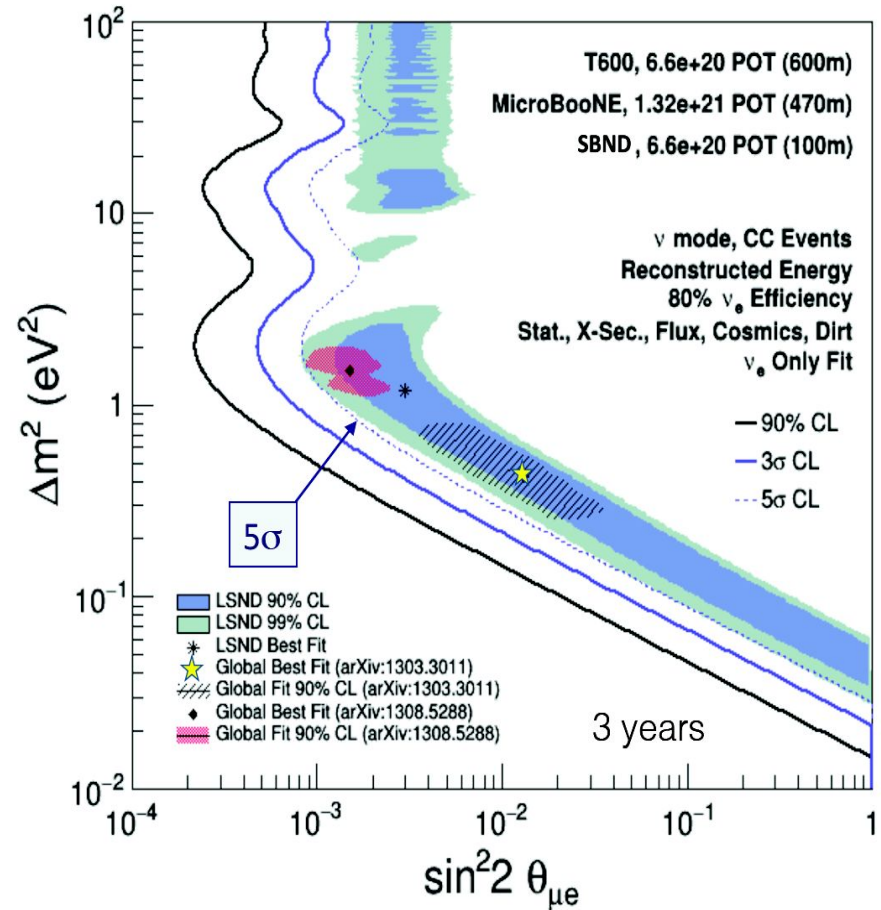
# LSND Comparison

	LSND	JSNS <sup>2</sup>	Advantage of JSNS <sup>2</sup>
<b>Detector Mass</b>	167 Tons	17 Tons	
<b>Baseline</b>	30 m	24 m	
<b>Beam Kinetic Energy</b>	0.8 GeV	3.0 GeV	Allows for KDAR measurements.
<b>Beam Power</b>	0.056 MW	1.0 MW	Much more intense beam
<b>Beam Pulse</b>	600 $\mu$ s, 120 Hz	80 ns (x2), 25 Hz	300 times less steady state background for IBD.
<b>Capture Nucleus</b>	H (2.2 MeV)	Gd ( $\sim$ 8 MeV)	Shorter capture time, higher signal to noise ratio.

# SBN Comparison



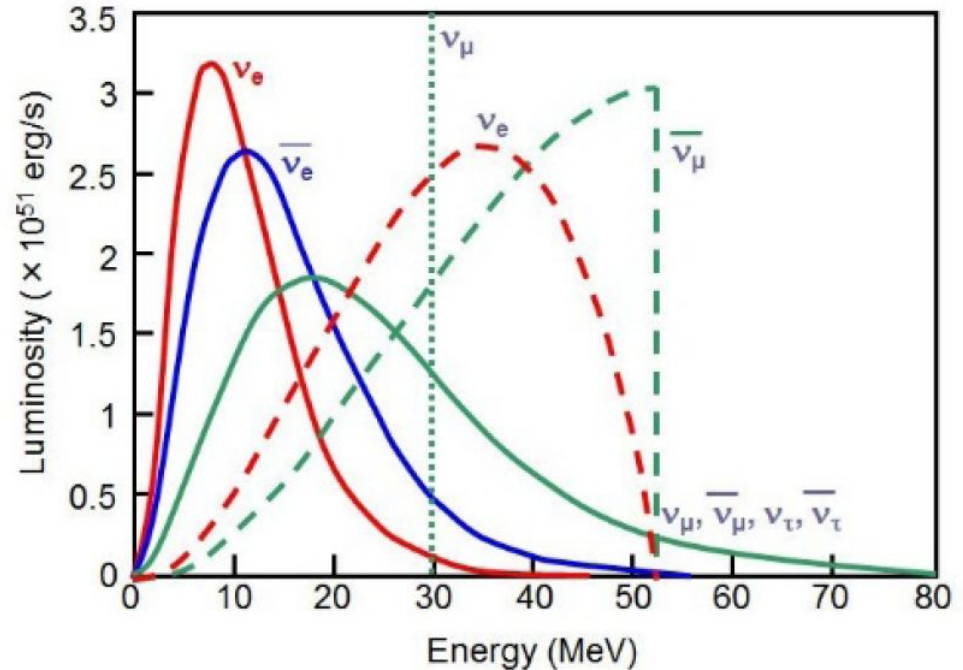
JSNS<sup>2</sup> Sensitivity Curves



SBN Sensitivity Curves

# Supernova Neutrinos

- Neutrino-driven wind is important in supernova simulations
- We need to know the neutrino nucleus cross sections better
- JSNS<sup>2</sup> can measure the neutrino nucleus rate for neutrinos on carbon 12
- With 3 years of data, we can measure the cross section to within 6%



**Decay at rest pions + muons (dashed)**  
**Supernova neutrinos (solid)**  
**Units of flux are arbitrary**