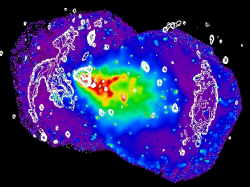
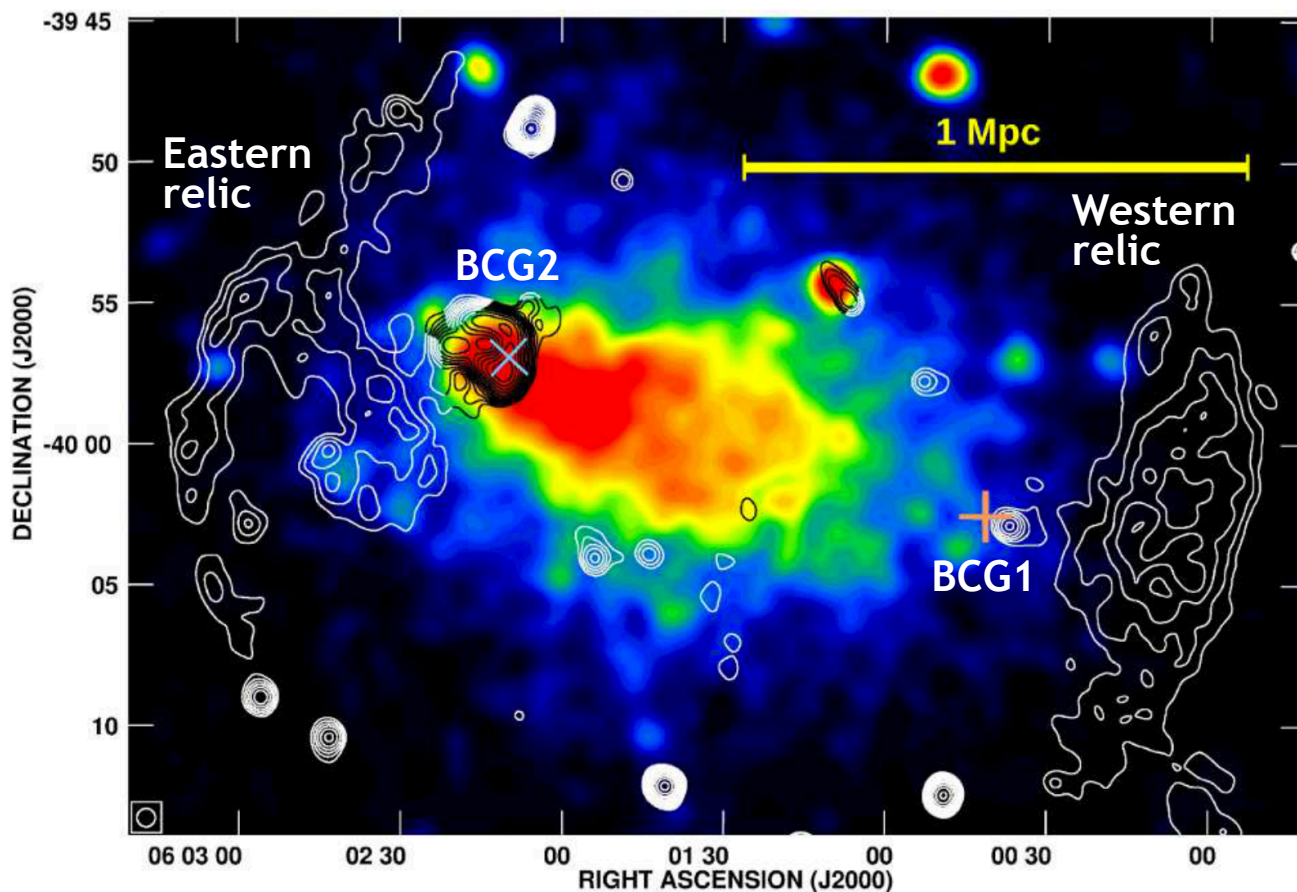


X-ray study of the merging cluster Abell 3376 with SUZAKU

Igone Urdampilleta



Abell 3376

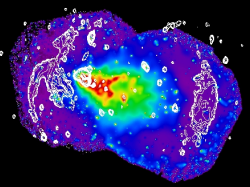


X-ray (ROSAT)
Radio (GMRT 325 MHz)

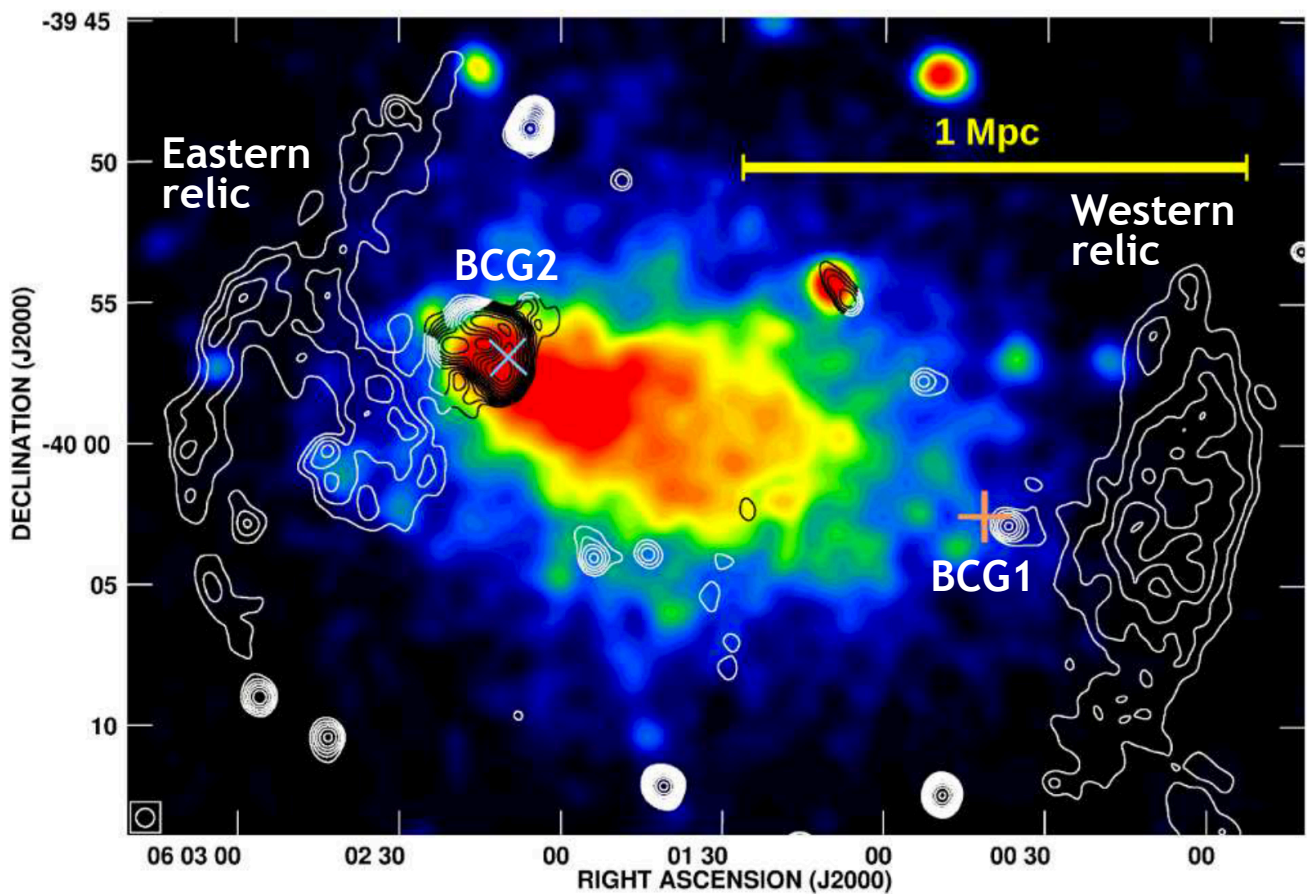
Kale et al. 2011

$z = 0.046$
 $M \sim 4-5 \cdot 10^{14} M_{\odot}$
 $\langle kT \rangle \sim 4.2 \text{ keV}$

Two giant arc-like ($\sim 2 \times 1.6 \text{ Mpc}$) radio relics
Discovered by Bagchi et al. 2006



Abell 3376



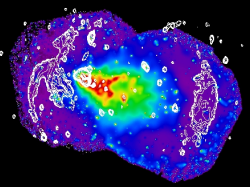
X-ray (ROSAT)
Radio (GMRT 325 MHz)

Kale et al. 2012

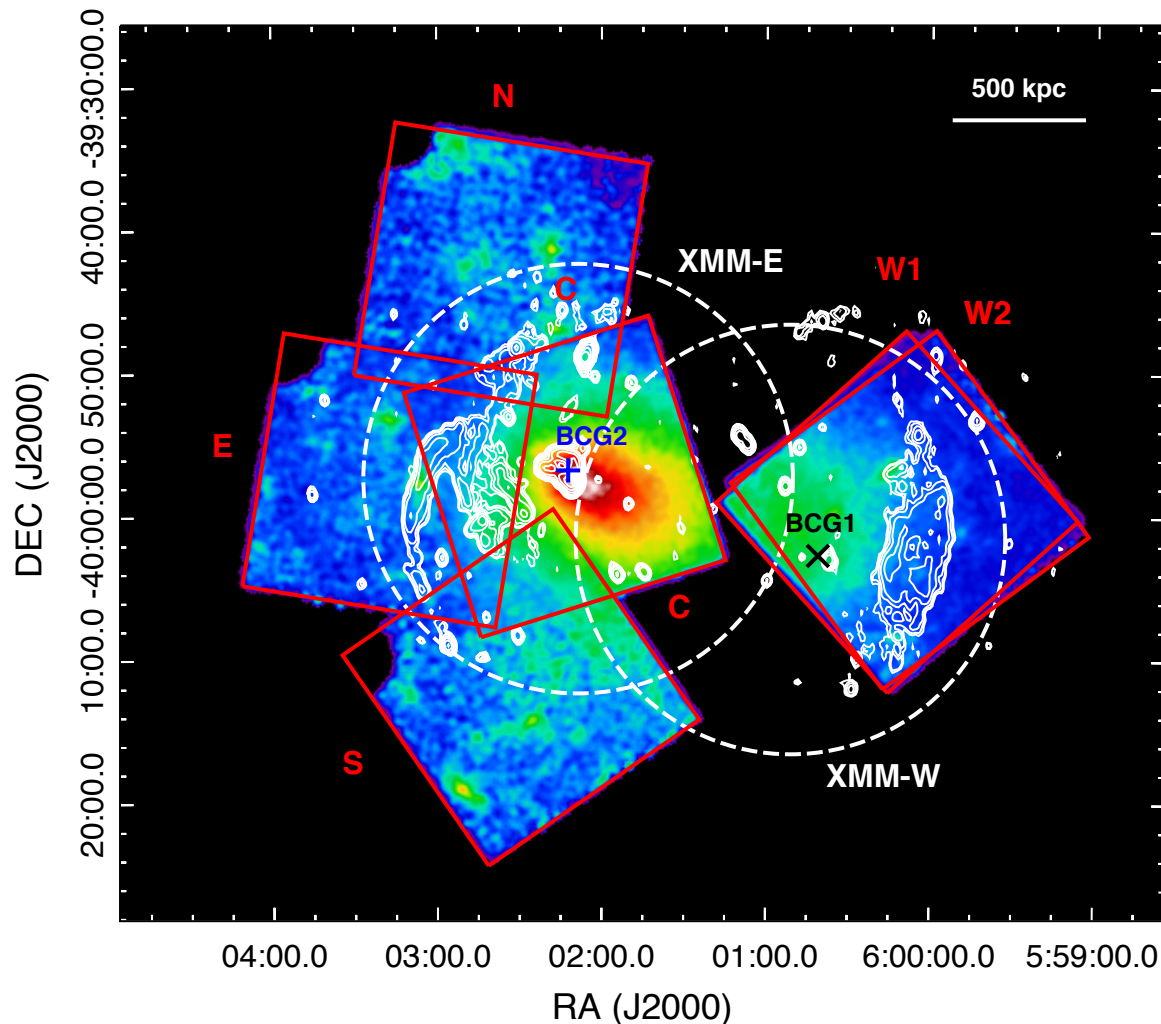
$z = 0.046$
 $M \sim 4-5 \cdot 10^{14} M_{\odot}$
 $\langle kT \rangle \sim 4.2 \text{ keV}$

Smaller and denser
A3376E crosses more
massive A3376W,
disrupting the core.

Two giant arc-like ($\sim 2 \times 1.6 \text{ Mpc}$) radio relics
Discovered by Bagchi et al. 2006



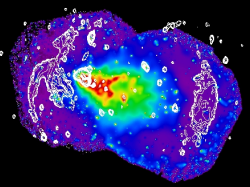
Abell 3376 observations



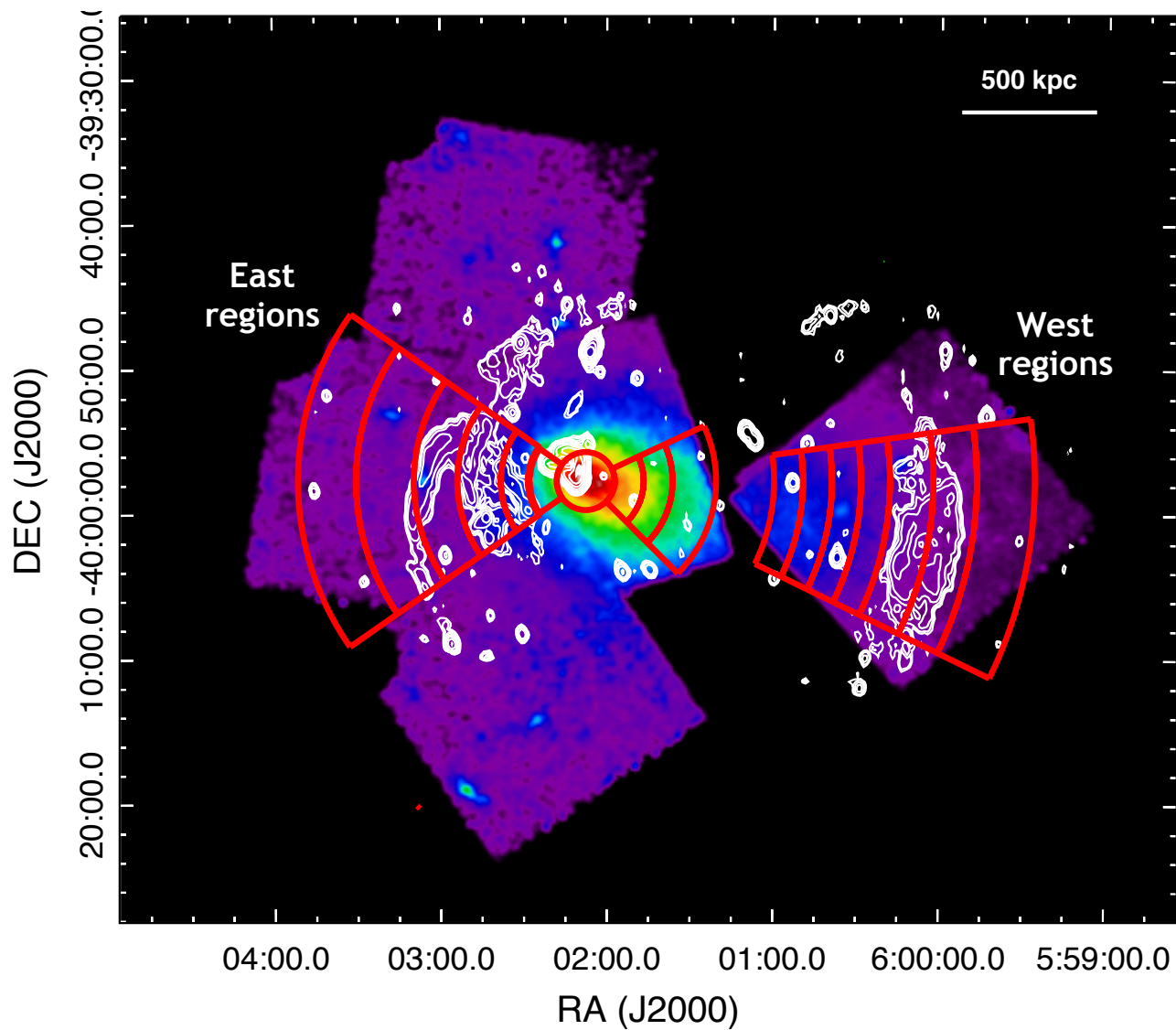
Obtain the spatial distribution of X-ray and radio component

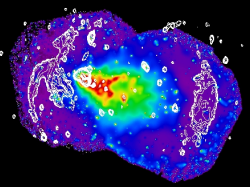
- Suzaku:
 - Western relic: C, W1 & W2
 - Eastern relic: C, N, E, S
- XMM-Newton:
 - Point sources
 - SB profiles

RADIO VLA (*Kale, R.*)

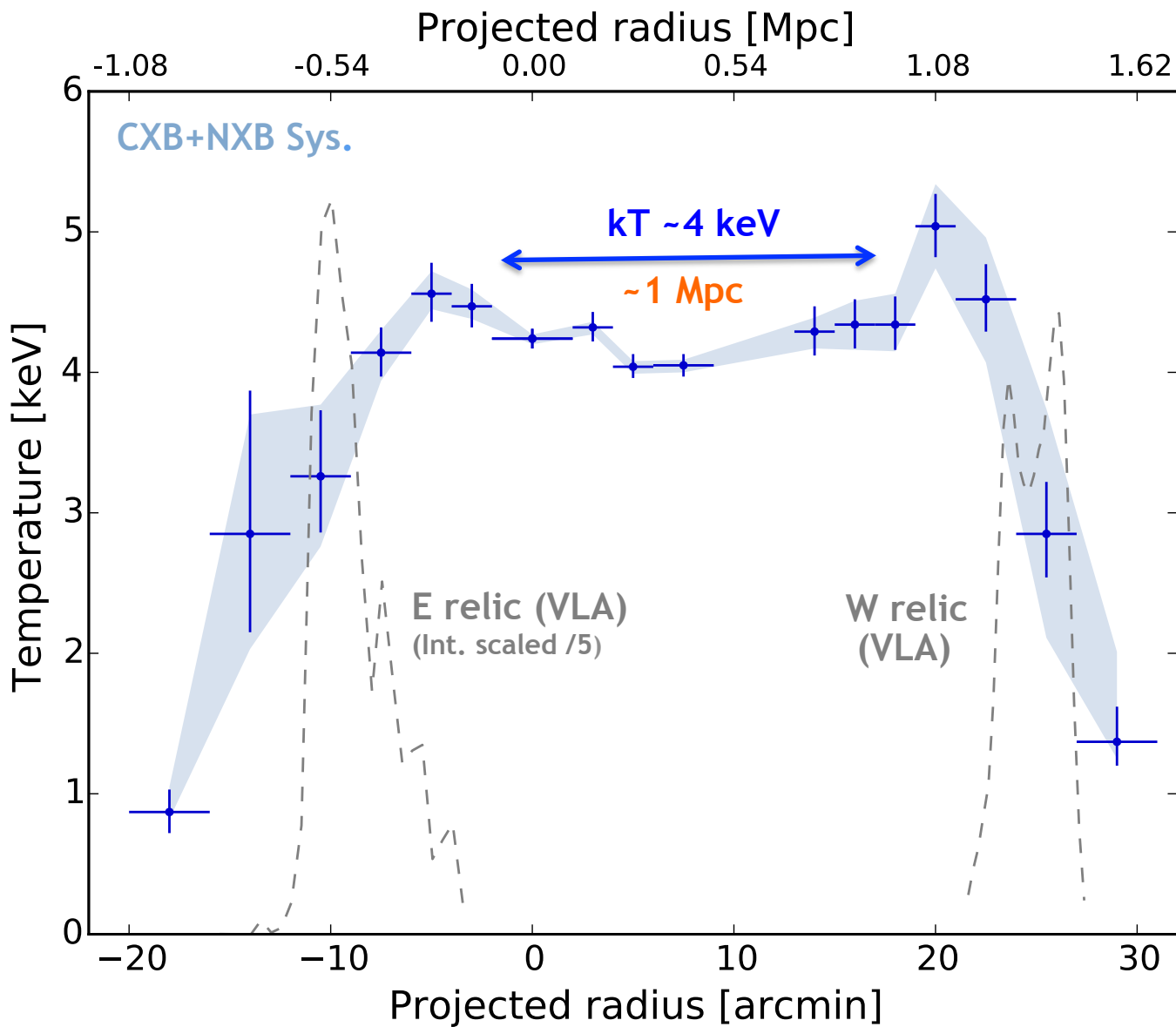


E & W Regions



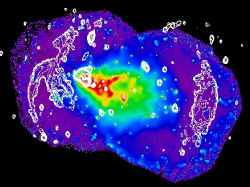


E & W T radial profile

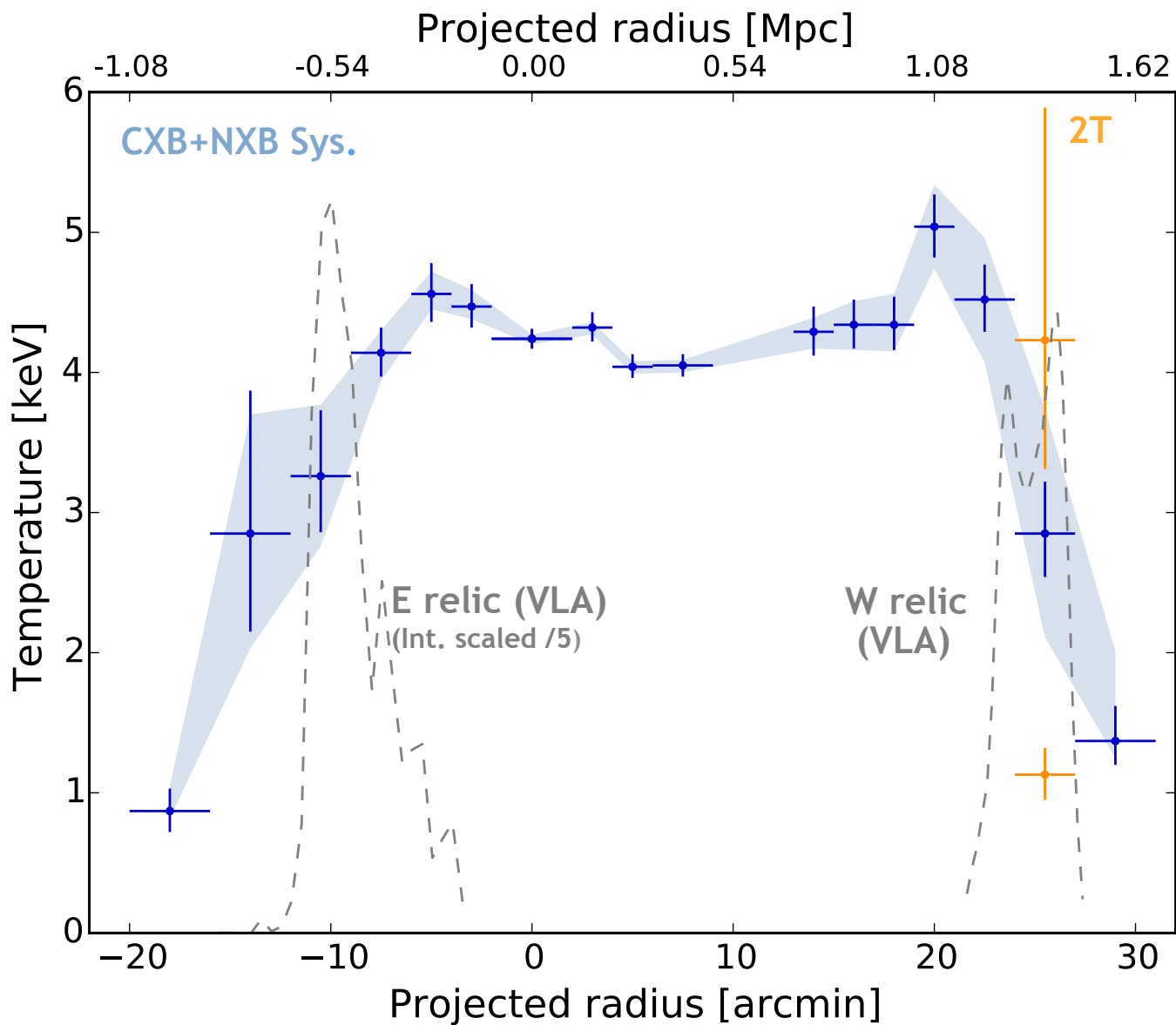


- $kT \sim 4 \text{ keV}$ in the center along $\sim 1 \text{ Mpc}$

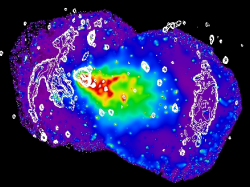
- T decreases along radio relics at West and East



E & W T radial profile

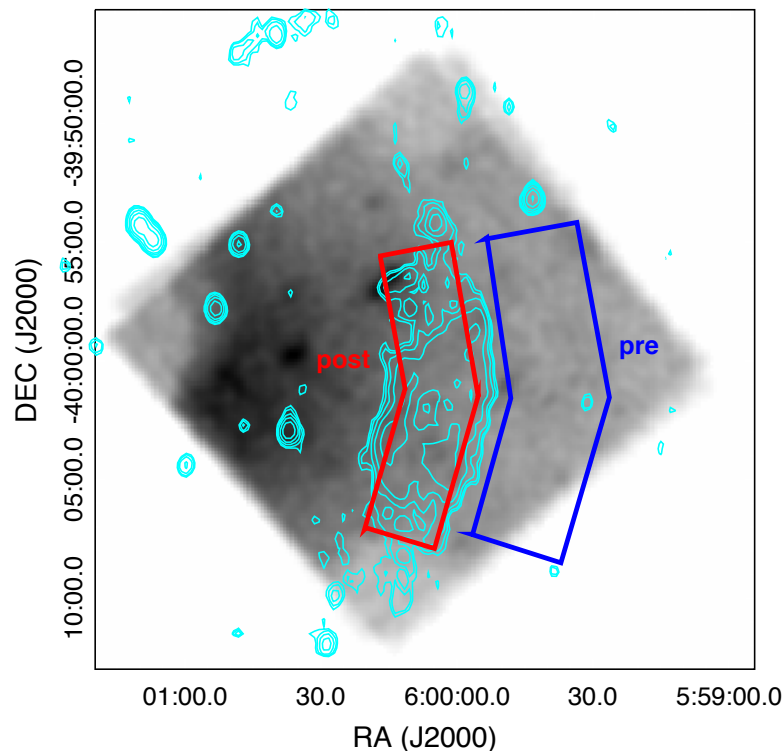
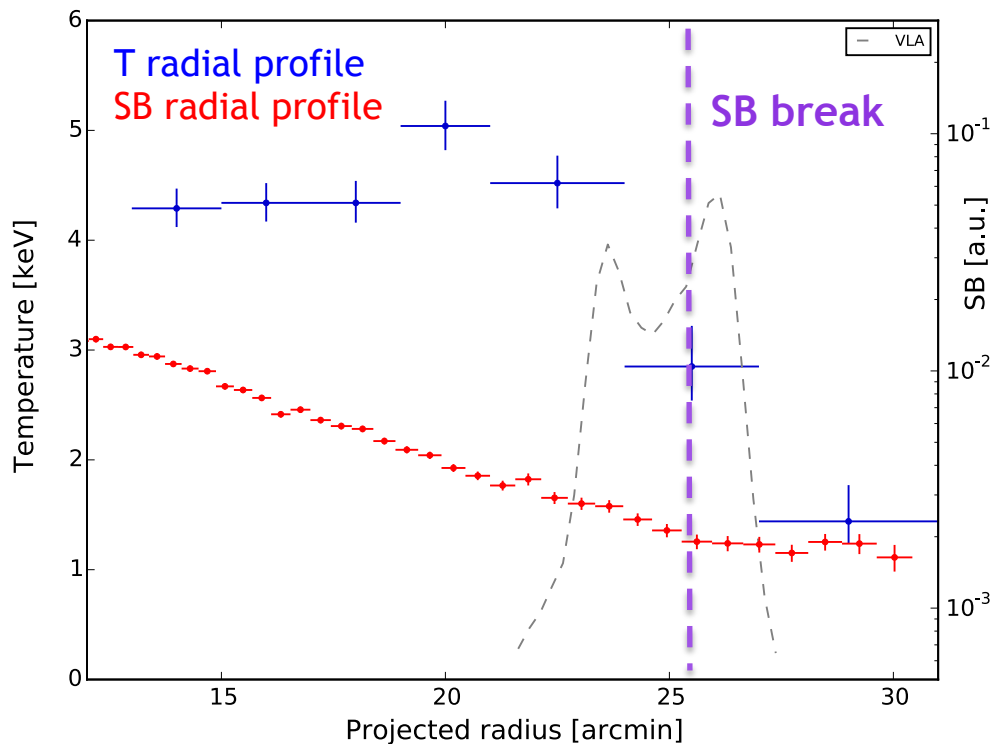


2T structure
found at radio relic



Western shock

SB profile break found with XMM observations at radio relic



Rankine-Hugoniot

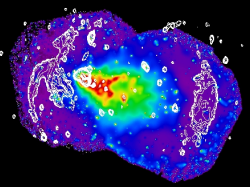
$$T_{\text{post}} = 4.2 \pm 0.3$$

$$T_{\text{pre}} = 1.3 \pm 0.3$$



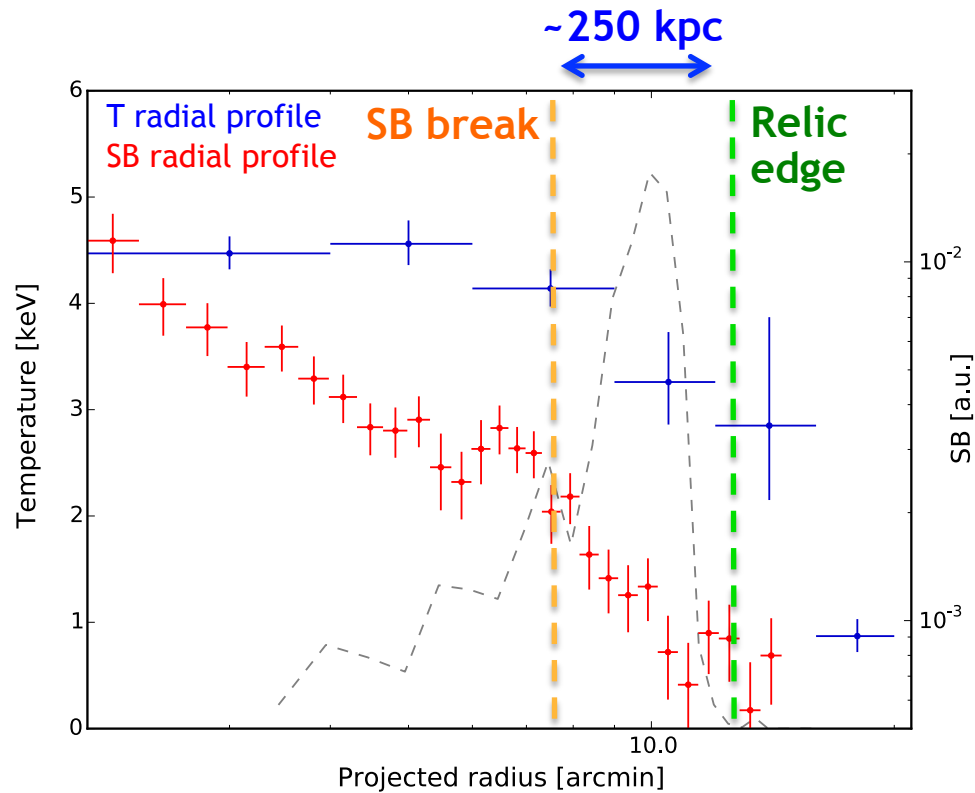
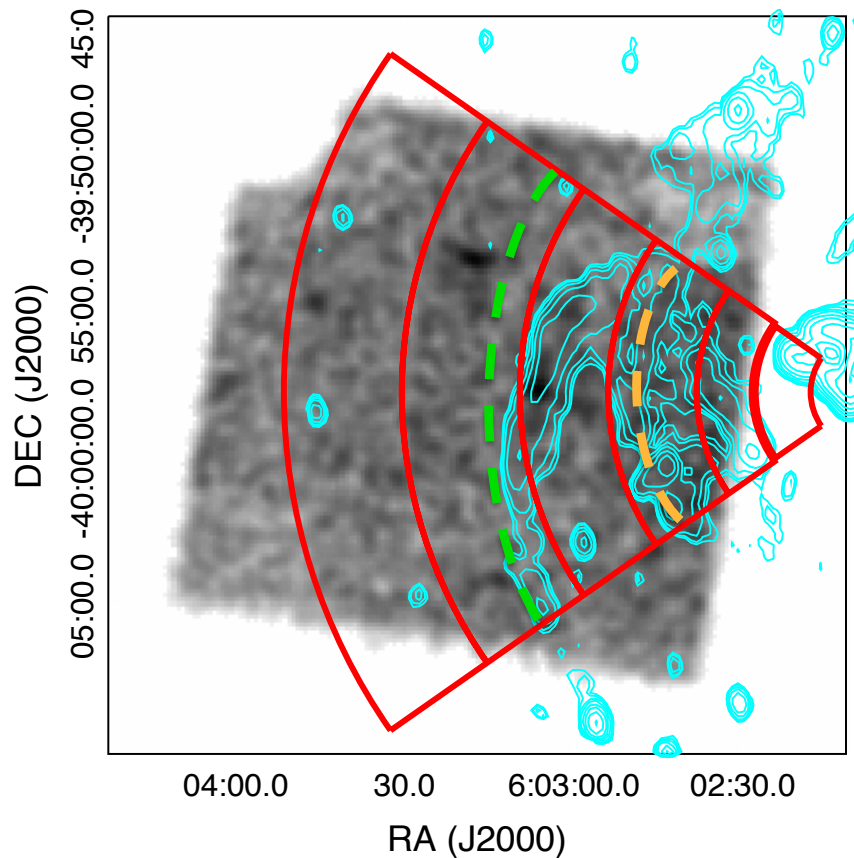
$$M_w \sim 2.8 \pm 0.4$$

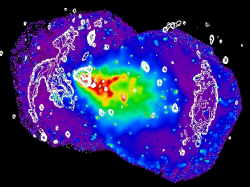
$$v_{\text{sw}} \sim 1630 \pm 250 \text{ km/s}$$



Eastern shock

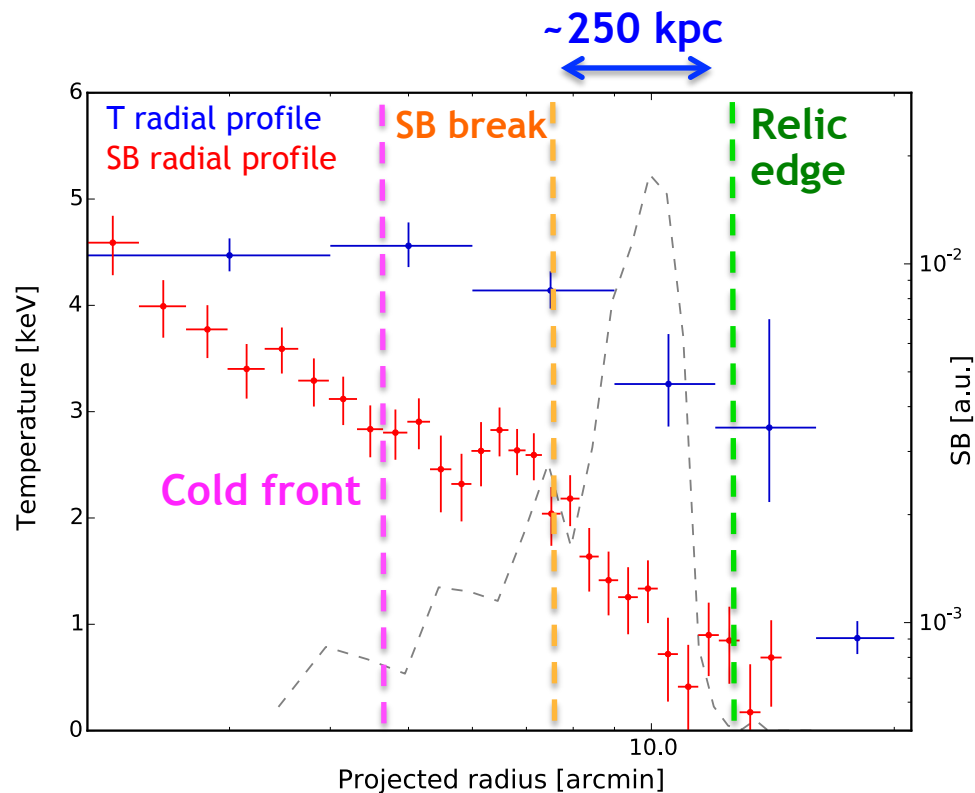
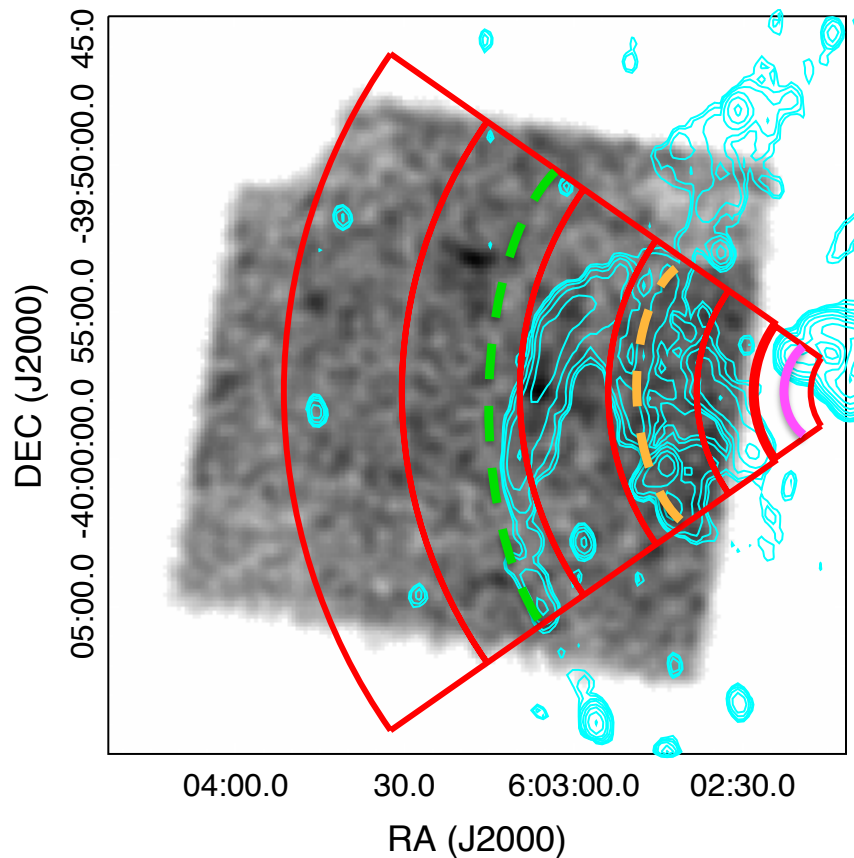
SB profile break found with XMM displaced from relic outer edge

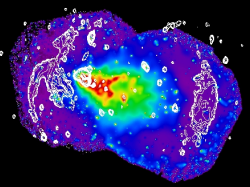




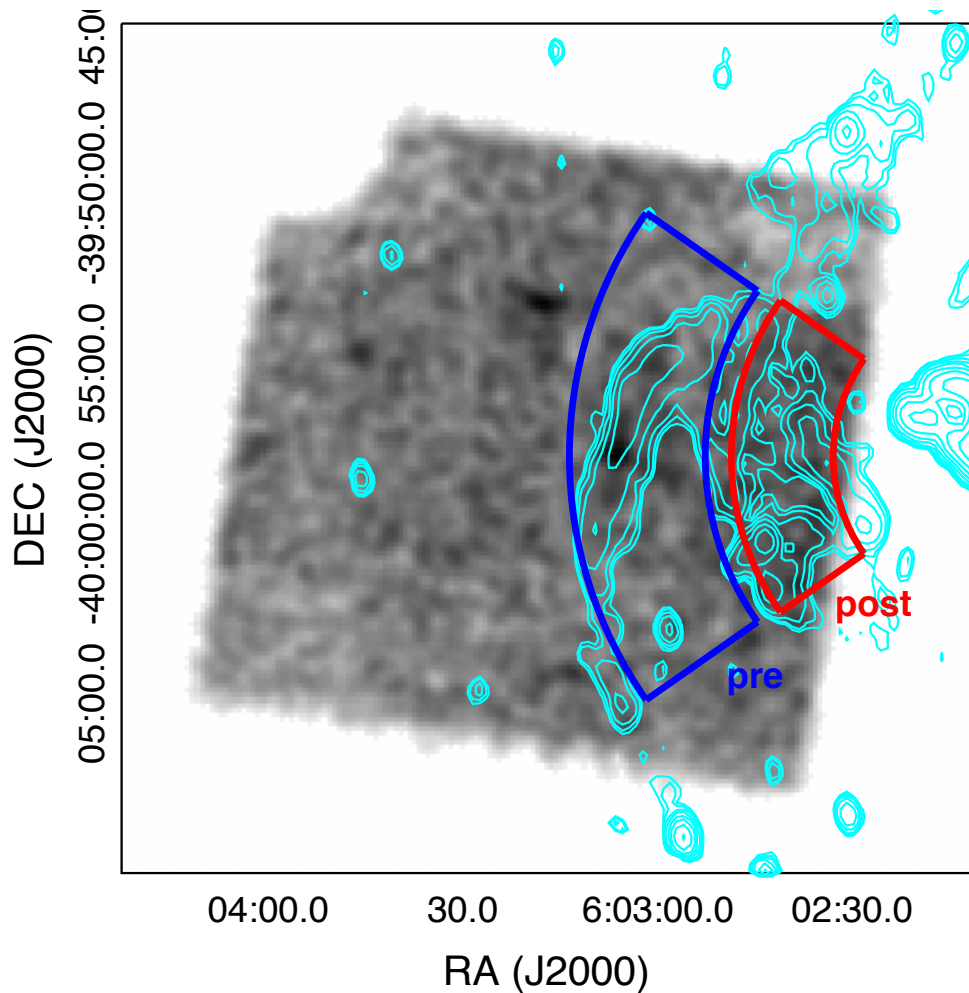
Eastern shock

SB profile break found with XMM displaced from relic outer edge





Eastern shock



$$T_{\text{post}} = 4.7 \pm 0.4$$

$$T_{\text{pre}} = 3.3 \pm 0.4$$

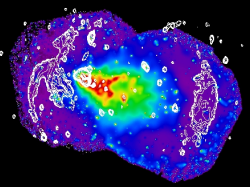


$$M_E \sim 1.5 \pm 0.1$$

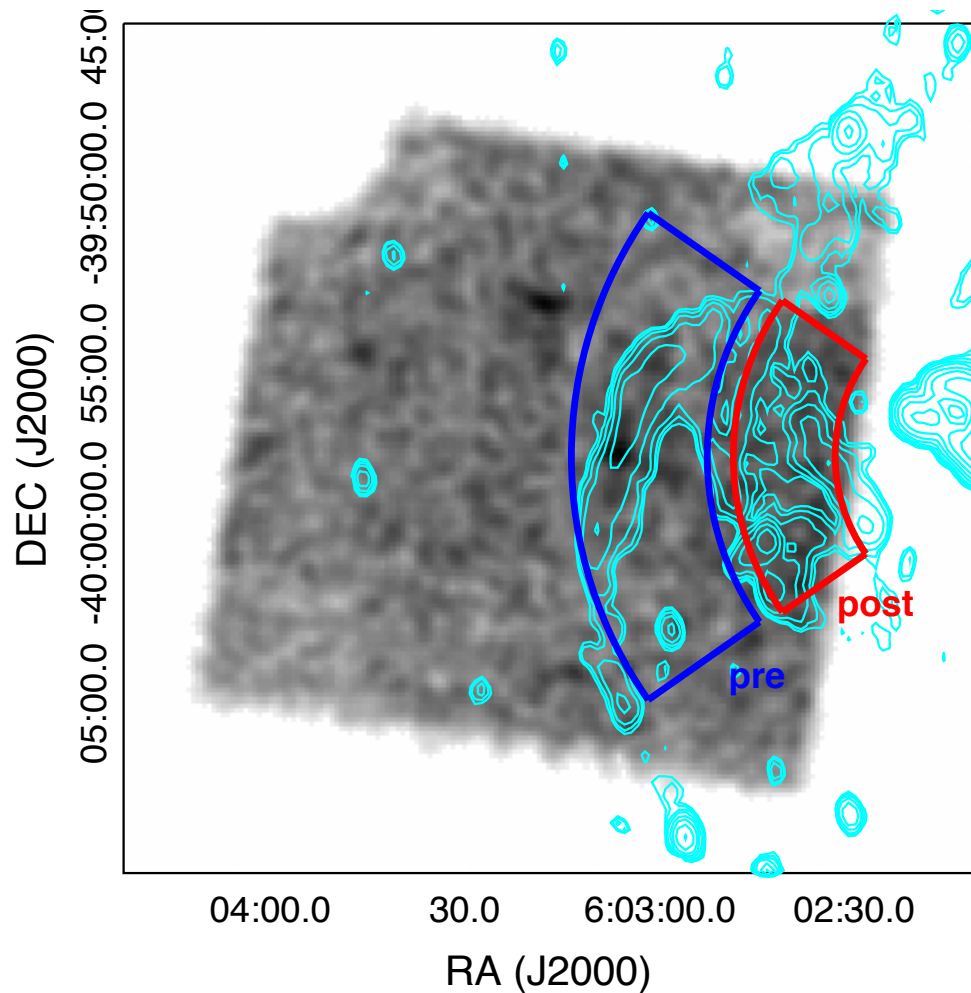
$$v_{\text{sE}} \sim 1450 \pm 150 \text{ km/s}$$

d shocks $\sim 1.9 \text{ Mpc}$:

$t_{\text{core passage}} \sim 0.6 \text{ Gyr}$



Eastern shock



$$T_{\text{post}} = 4.7 \pm 0.4$$

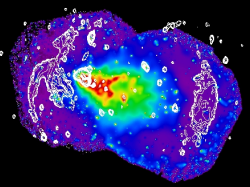
$$T_{\text{pre}} = 3.3 \pm 0.4$$



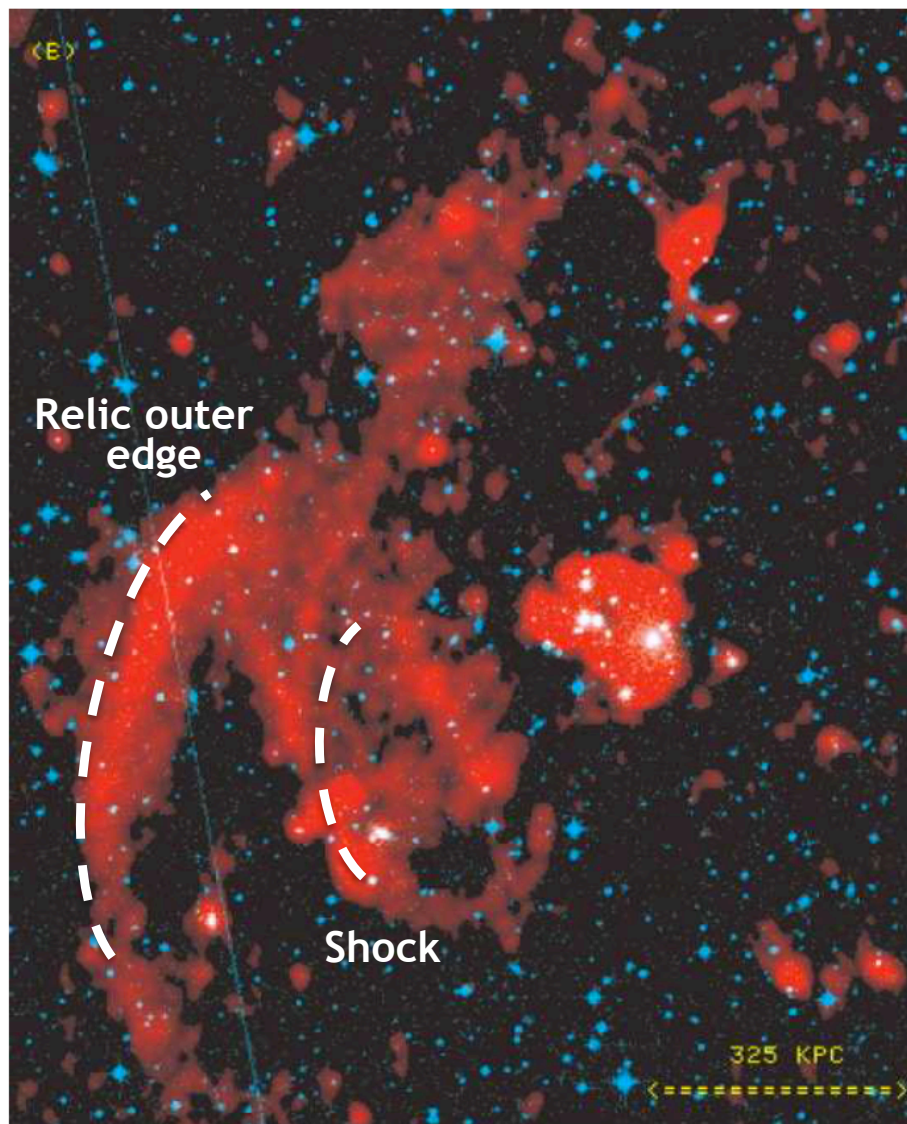
$$M_E \sim 1.5 \pm 0.1$$

$$v_{sE} \sim 1450 \pm 150 \text{ km/s}$$

*Still evolving
merging galaxy cluster*



Eastern shock/relic

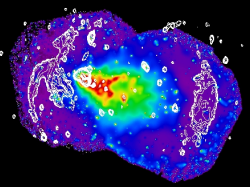


- ✓ Complex radio relic structure:
 - Northern E-relic
 - Elongated E-relic
 - Notch
- ✓ No X-ray SB break at relic outer edge limited by low SB emission

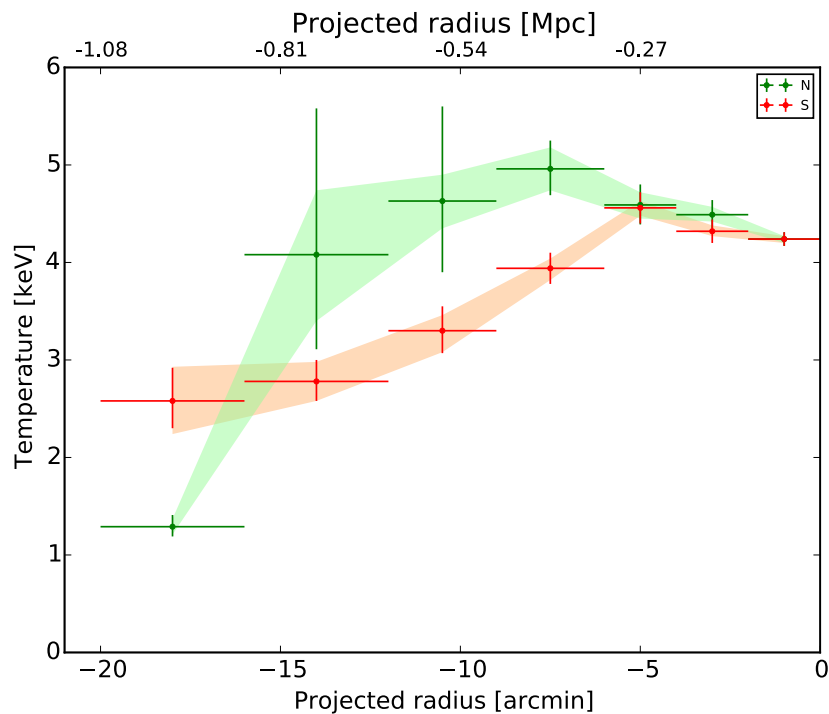
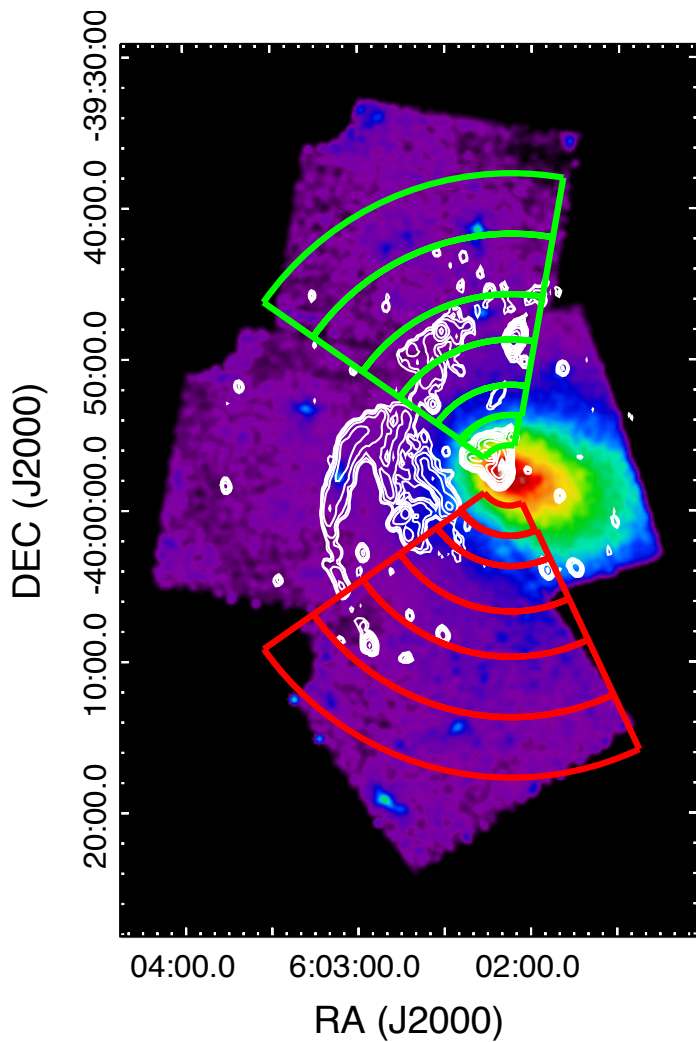
High-sensitivity X-ray images needed with Athena

Radio (VLA 1.4 GHz)

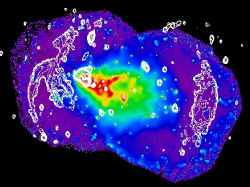
(Bagchi et al. 2006)



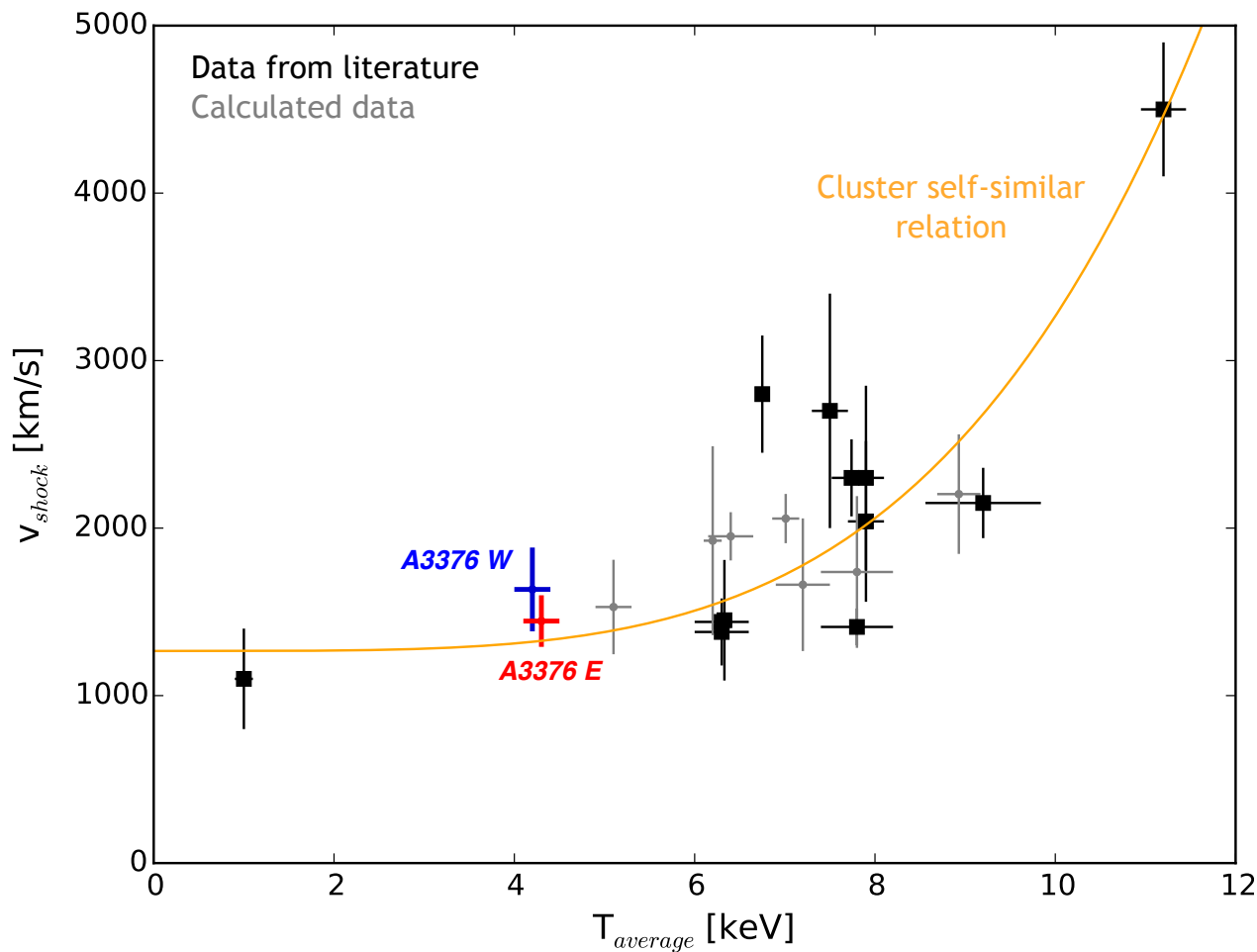
N & S T radial profile



- ✓ Radio features in the **North**,
no X-ray detection due to weakness
signal
- ✓ No radio signatures in **South**
- ✓ **South** consistent with relaxed cluster
Universal profile (Burns et al. 2010)

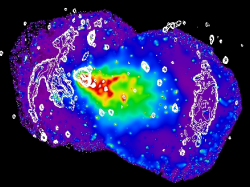


Vshock vs Taverage



$$v_{shock} = M \cdot c_s$$
$$c_s = \sqrt{\frac{\gamma k T_{pre}}{\mu m_p}}$$

v_{shock} appears correlated to $T_{average}$



Summary

A3376 is an **evolving** merging galaxy cluster with **double** giant shocks/radio relics with $M \sim 2-3$.

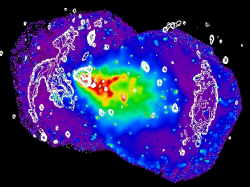
Azimuthal T radial distribution in X-ray:
W, E and N consistent with radio relics presence.

- **W**: shock and radio relic co-spatially located
- **E**: shock located at the “notch”

S agrees with relaxed clusters profile.

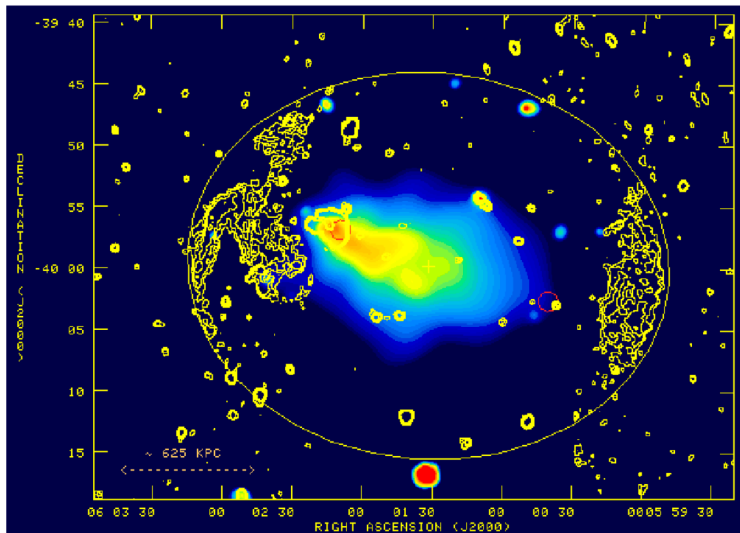


BACK UP

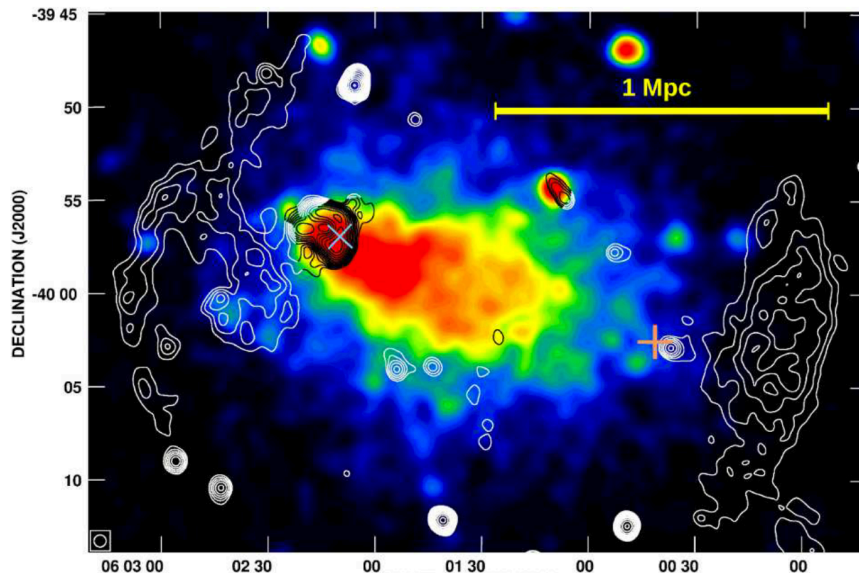


Radio emission

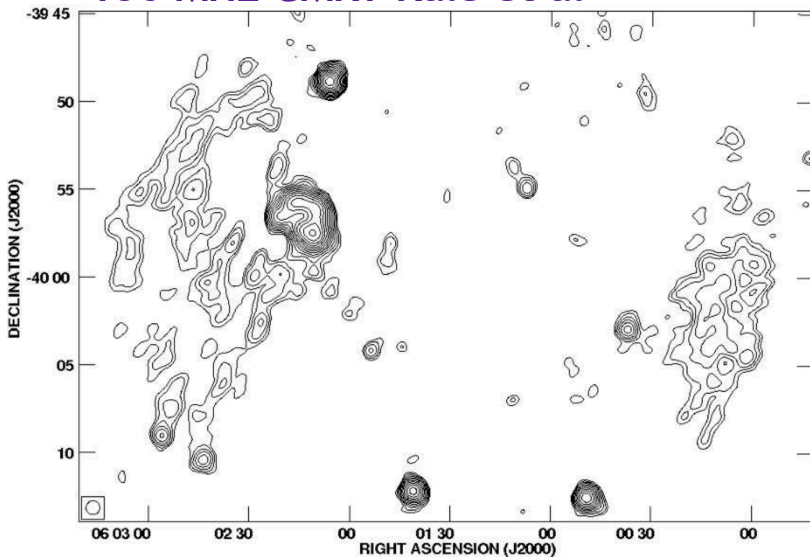
1.4 GHz VLA Bagchi et al



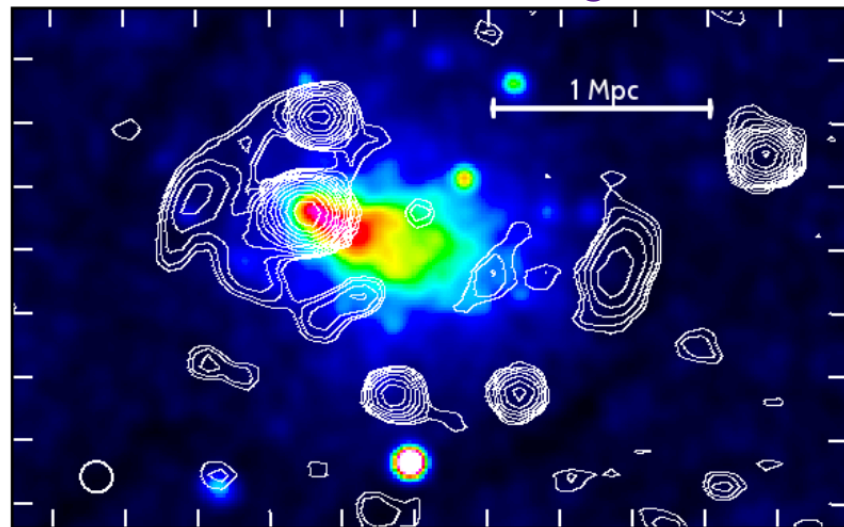
325 MHz GMRT Kale et al

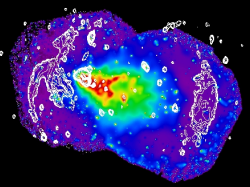


150 MHz GMRT Kale et al



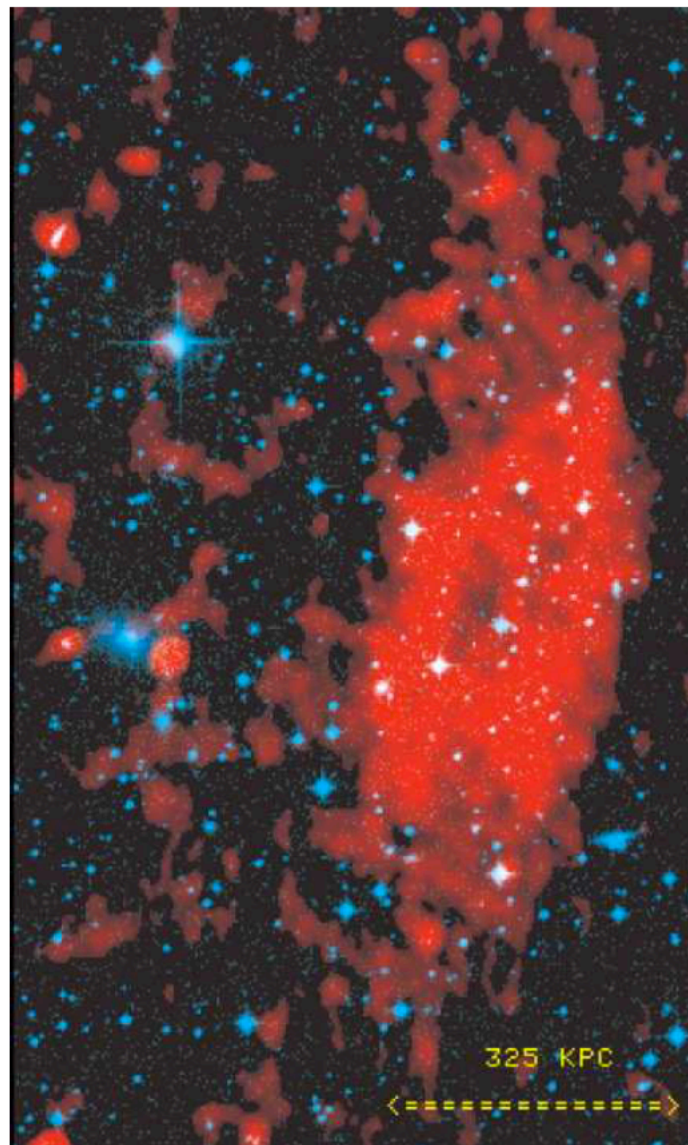
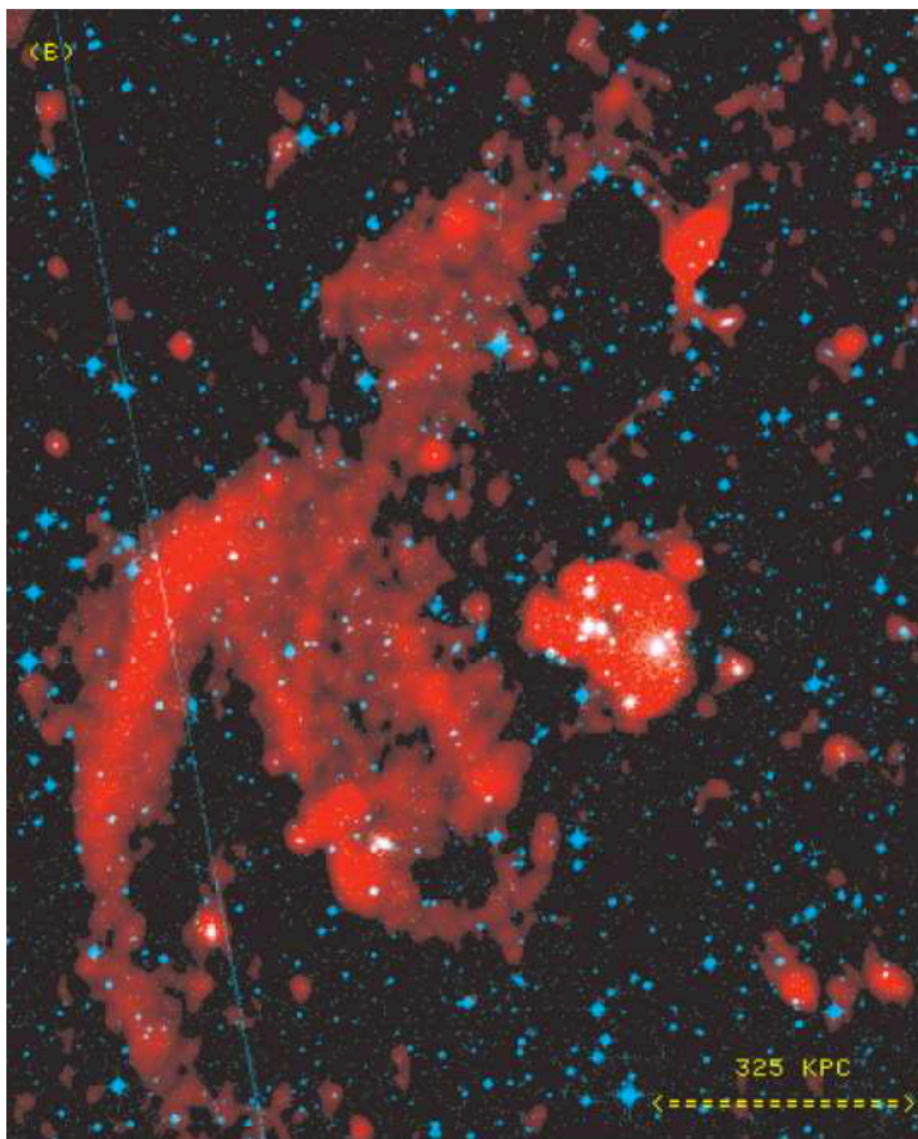
154-215 MHz MWA George et al

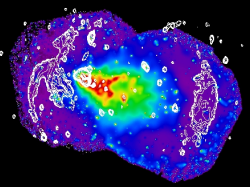




Radio map 1.4 GHz VLA

(Bagchi et al. 2006)





Polarization maps

(Kale et al. 2012)

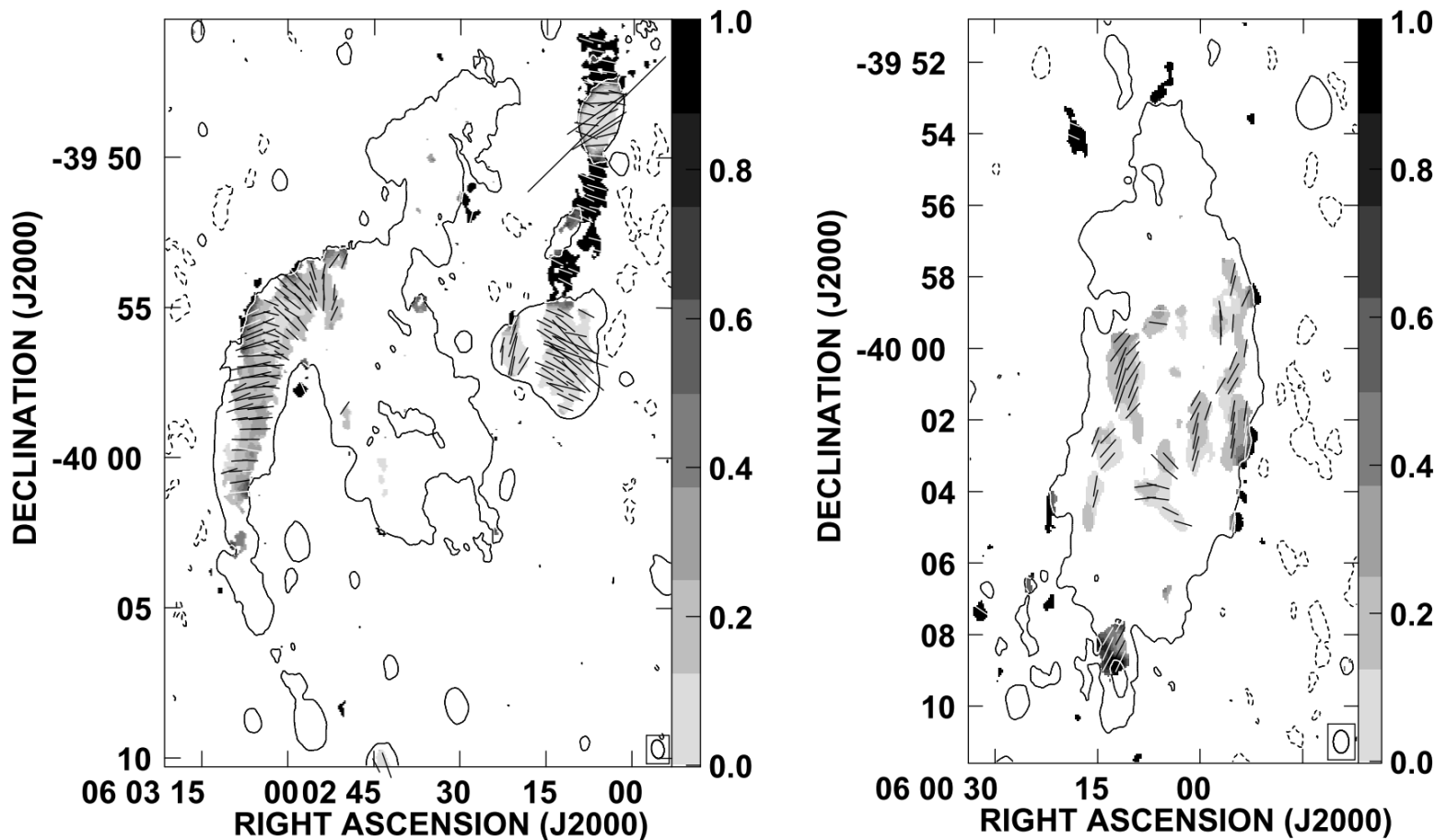
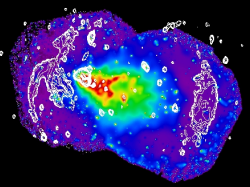


Figure 3. VLA 1400 MHz: Electric field vectors overlaid on fractional polarized intensity shown in grey-scale and Stokes I contours shown at $-0.16, 0.16 \text{ mJy beam}^{-1}$. The length of the vectors is proportional to the polarized intensity. The synthesized beam is $37'' \times 25''$ (P. A. 9.22°) in the left panel and $38'' \times 26''$ (P. A. 0.17°) in the right panel.



Radio spectral index map

EAST

WEST

(Kale et al. 2012)

325MHz GMRT
1.4 GHz VLA

