



# Diffuse radio emission in the Corona Borealis supercluster field

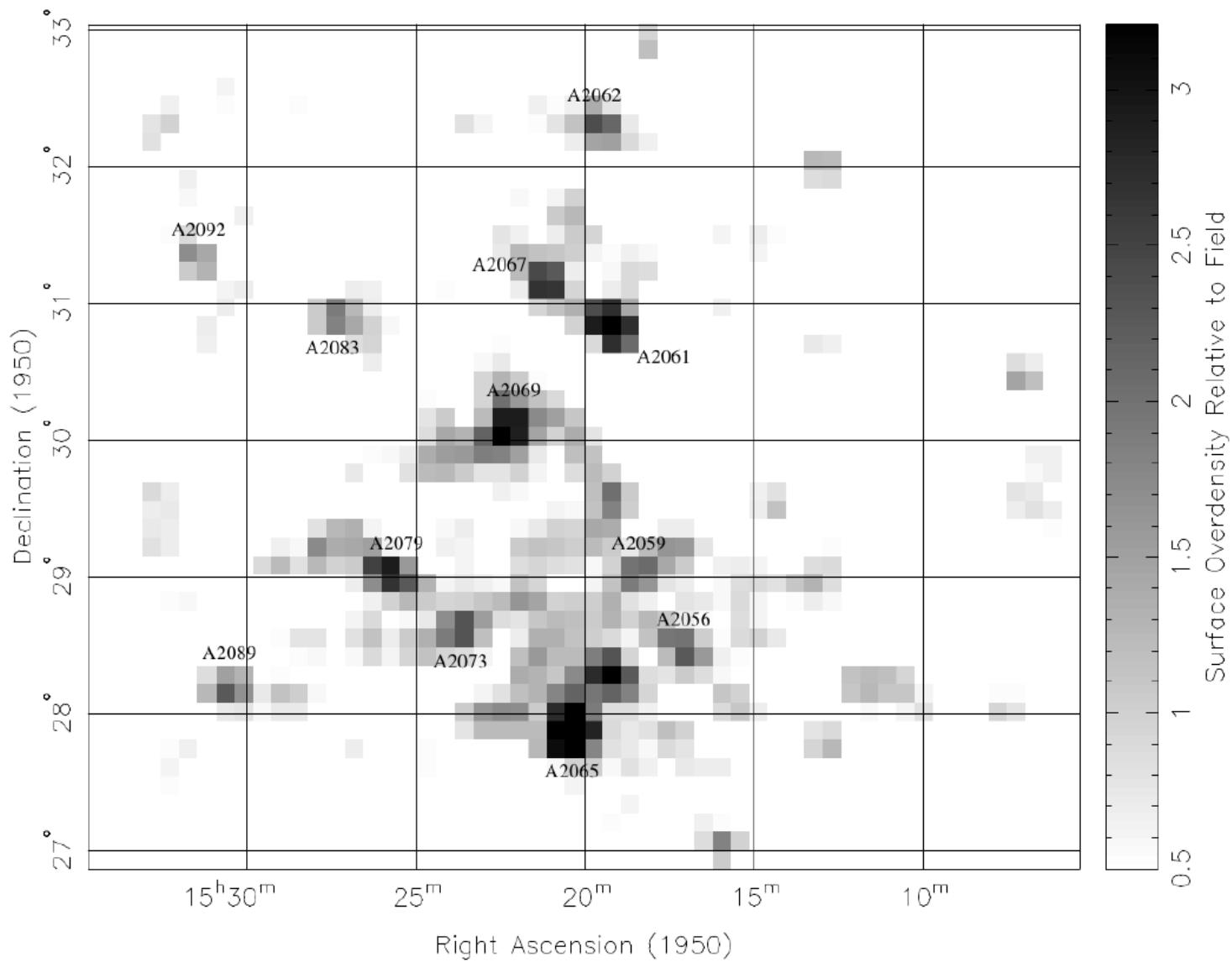


**Alexander Drabent**

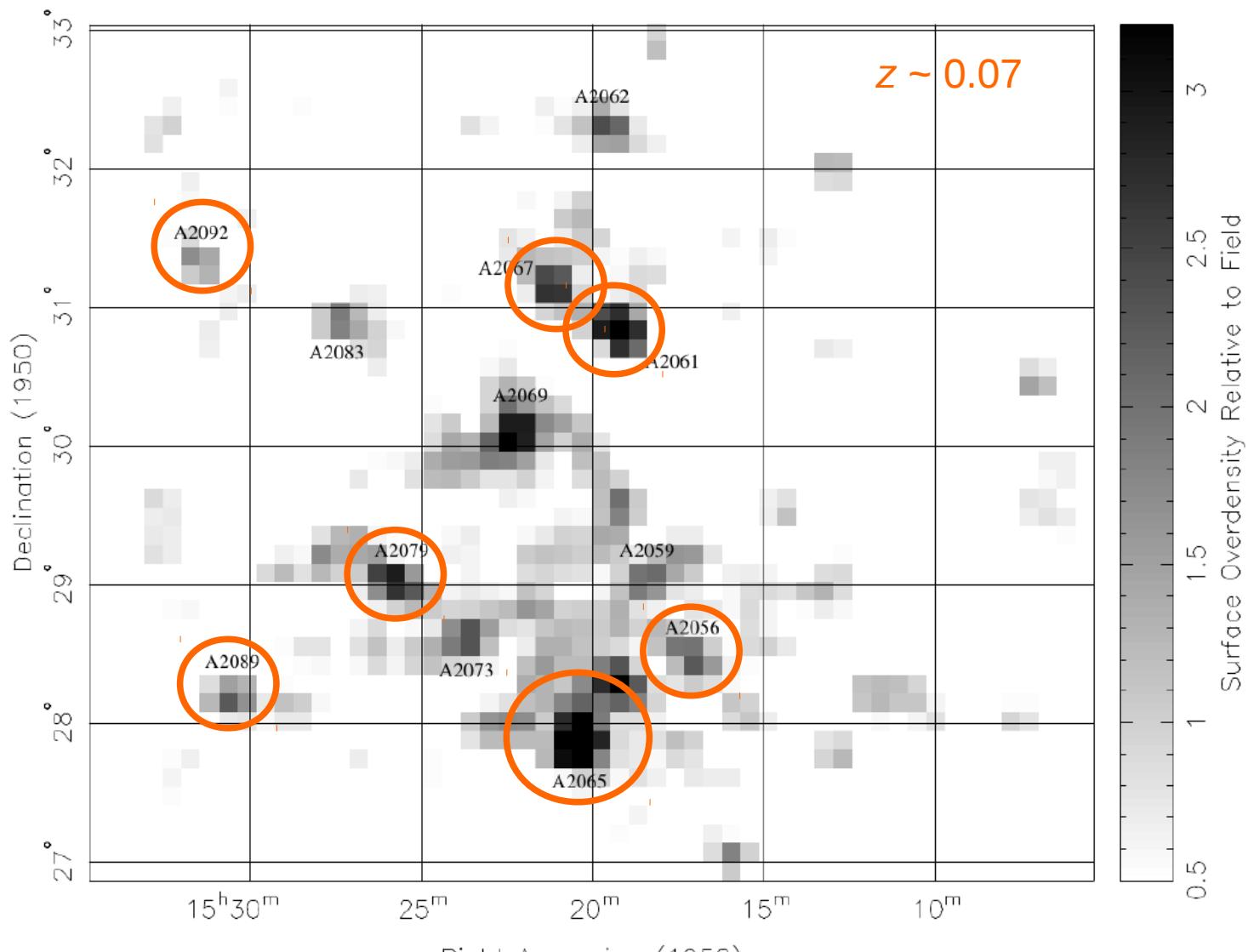
M. Hoeft, M. Brüggen, G. Brunetti, T. W. Shimwell  
and the LOFAR Surveys KSP Cluster working group

**26<sup>th</sup> October, 2017**

# The Corona Borealis Supercluster field

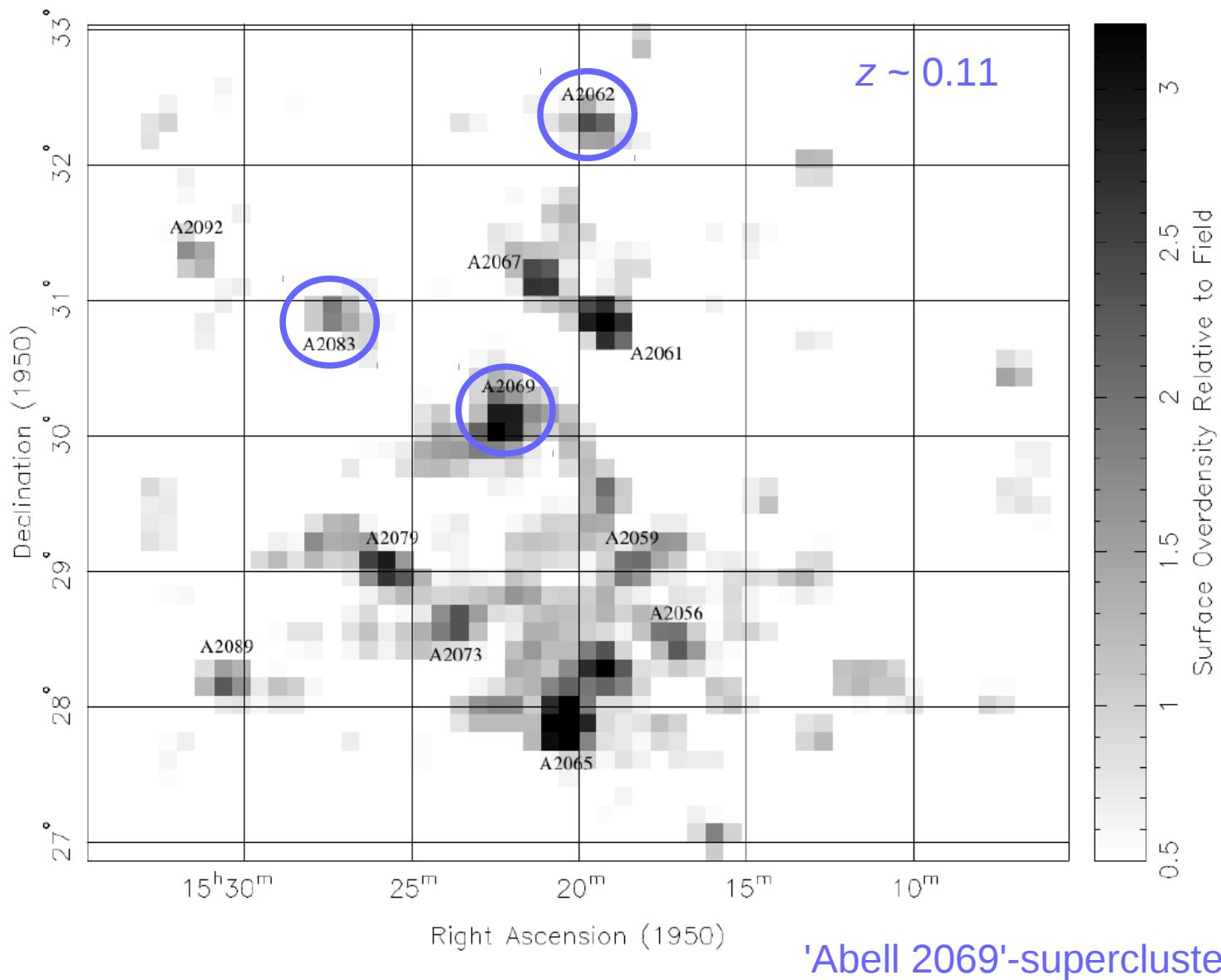


# The Corona Borealis Supercluster field



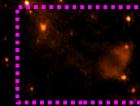
Corona Borealis supercluster

# The Corona Borealis Supercluster field

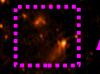


# Corona Borealis supercluster field

5°



A2061



A2069



A2065

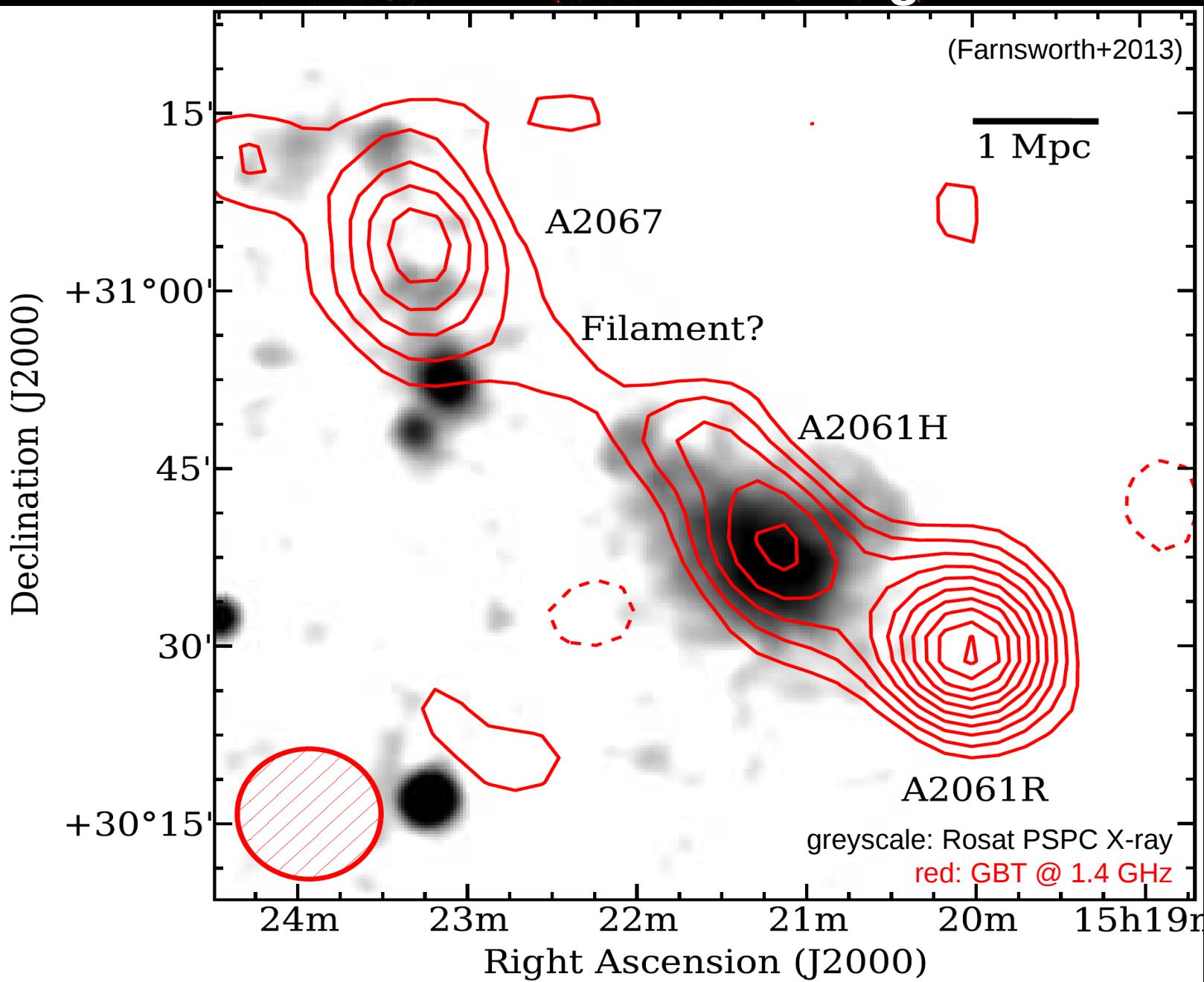
LOFAR HBA @ 153 MHz

beam:  $28'' \times 24''$

r.m.s. noise: 450  $\mu\text{Jy}/\text{beam}$

# Abell 2061-Abell 2067 bridge?

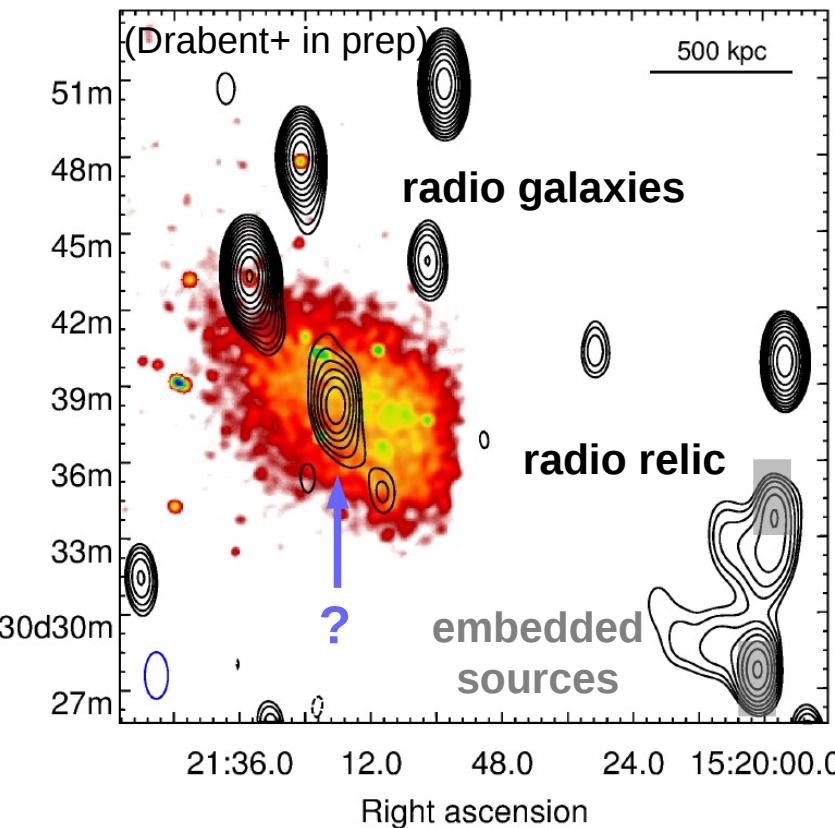
(Farnsworth+2013)



# Abell 2061

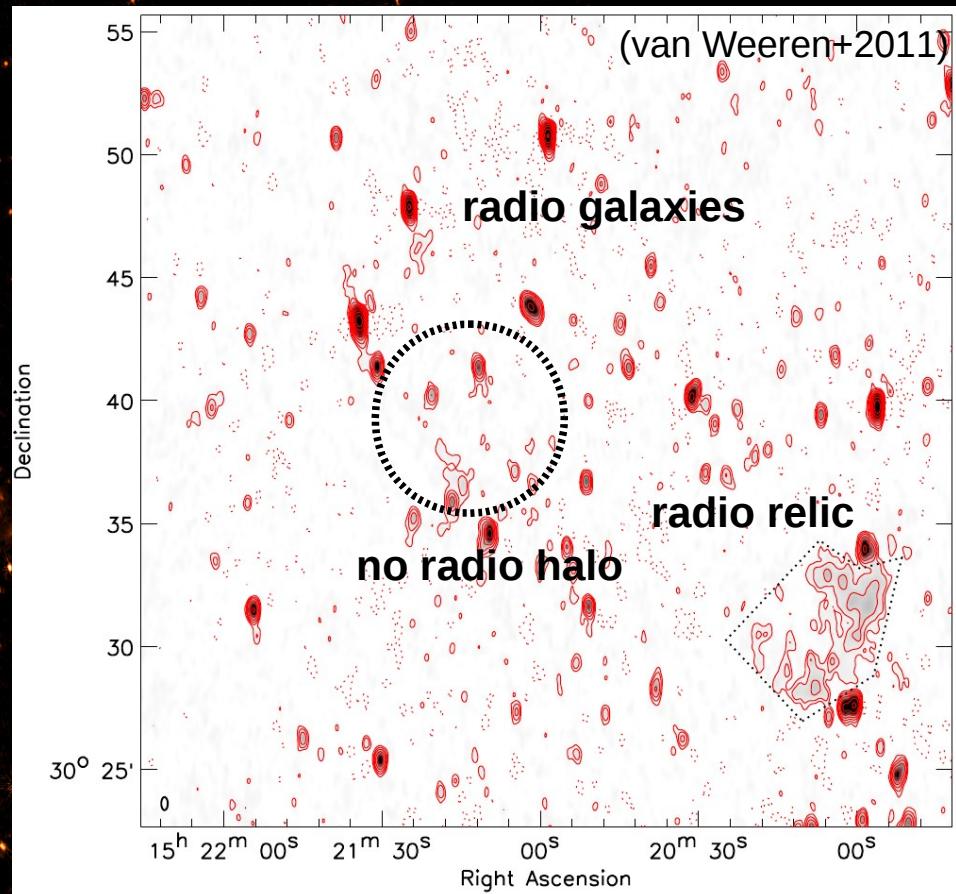
## steep spectrum radio source at cluster center

Declination



contours: WSRT @ 346 MHz

Radio relic:  $(90 \pm 9)$  mJy

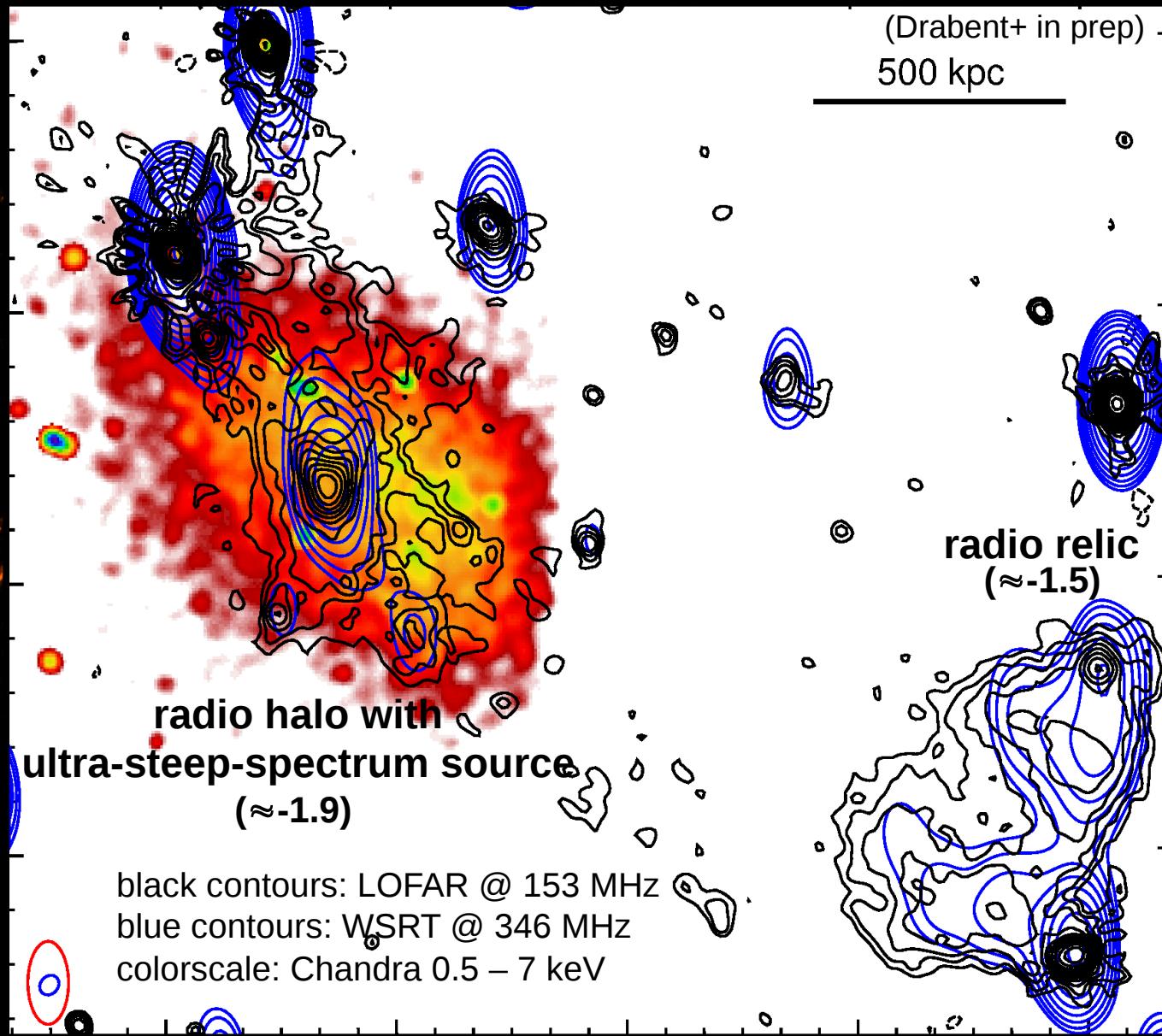


contours: WSRT @ 1.4 GHz  
 $(27 \pm 1)$  mJy

spectral index of radio relic:  $-0.9 \pm 0.1$

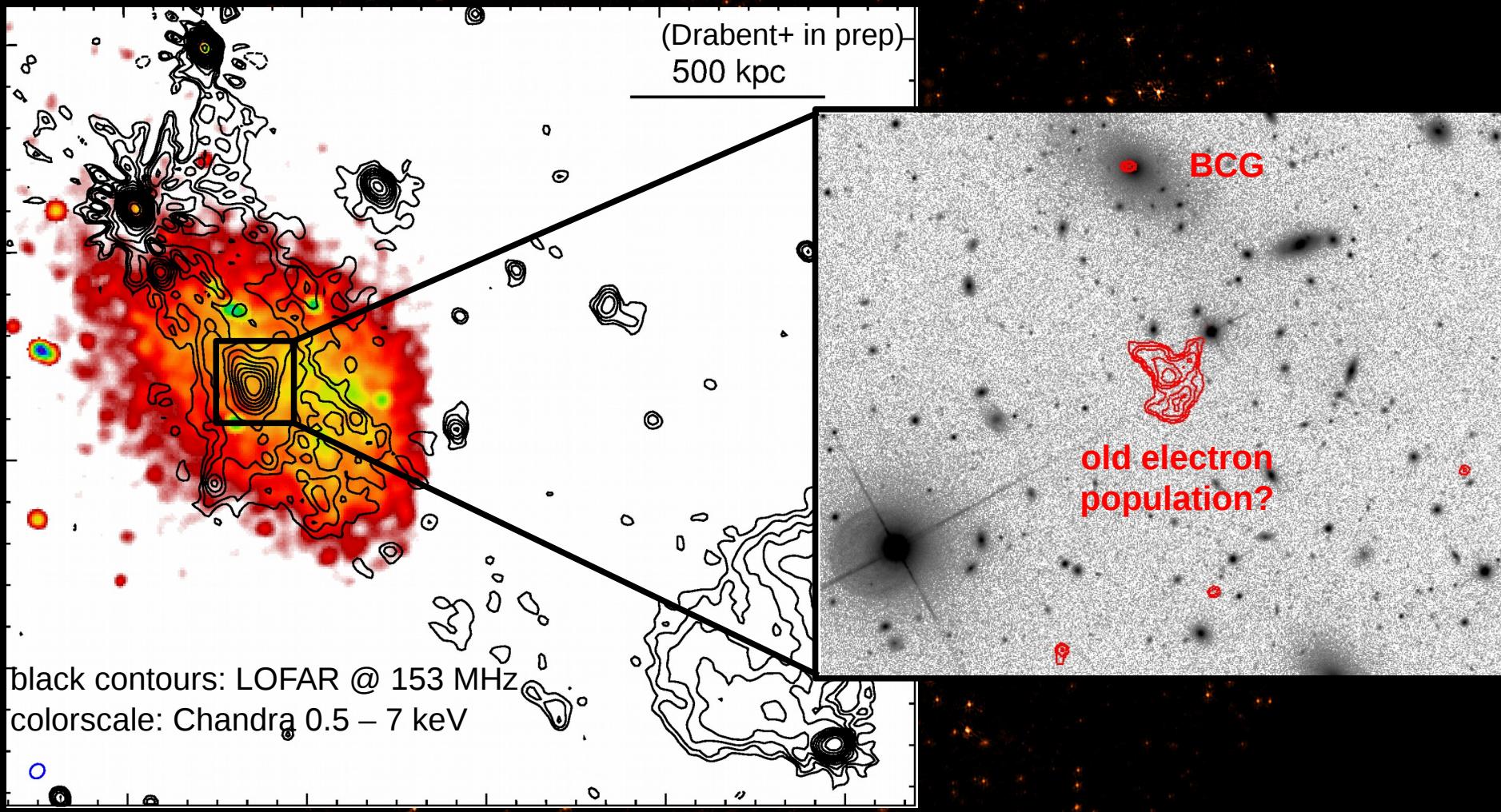
# Abell 2061

## radio halo found – filaments of radio relic visible

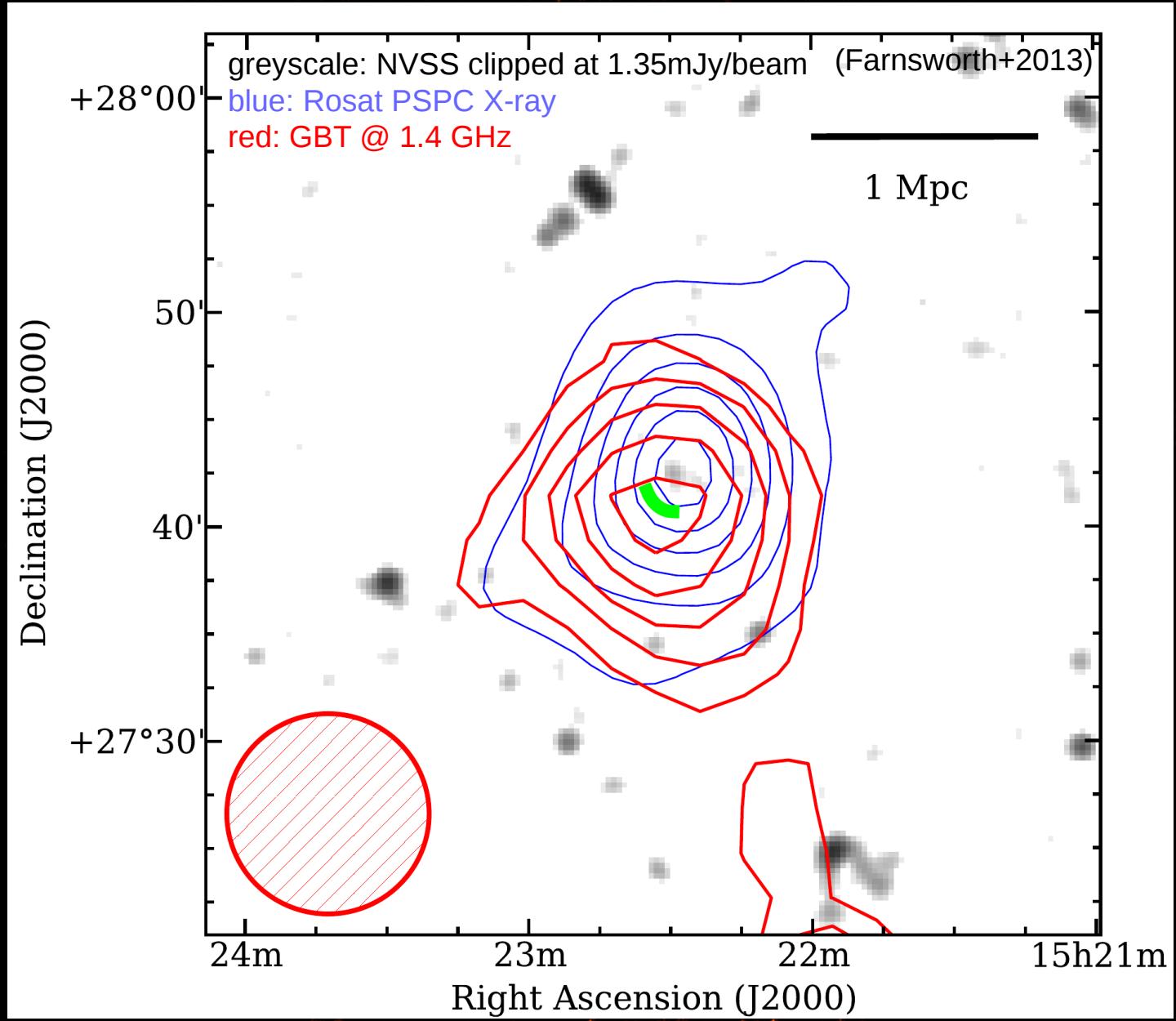


# Abell 2061

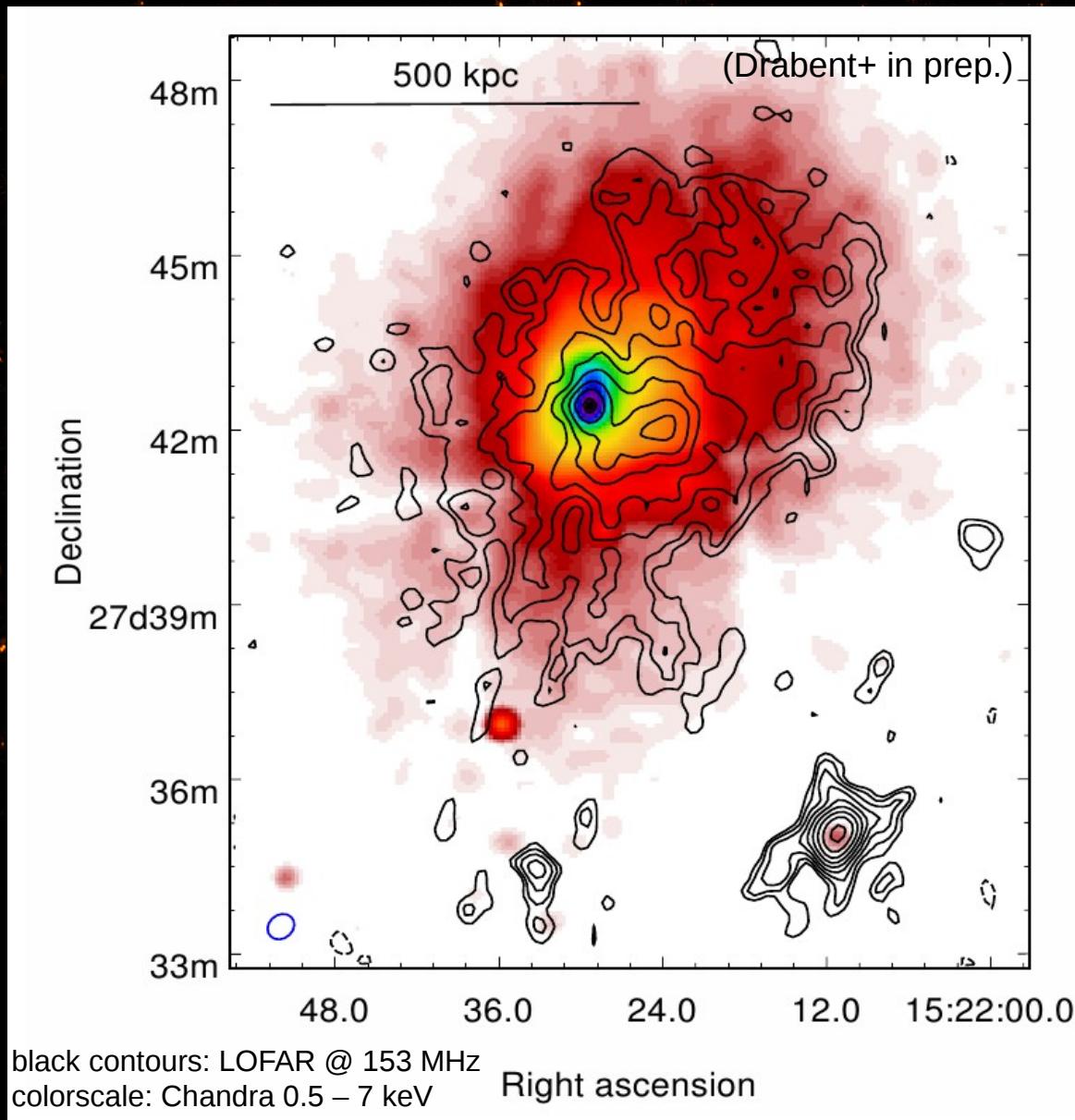
## radio halo + embedded ultra-steep spectrum source



# Abell 2065

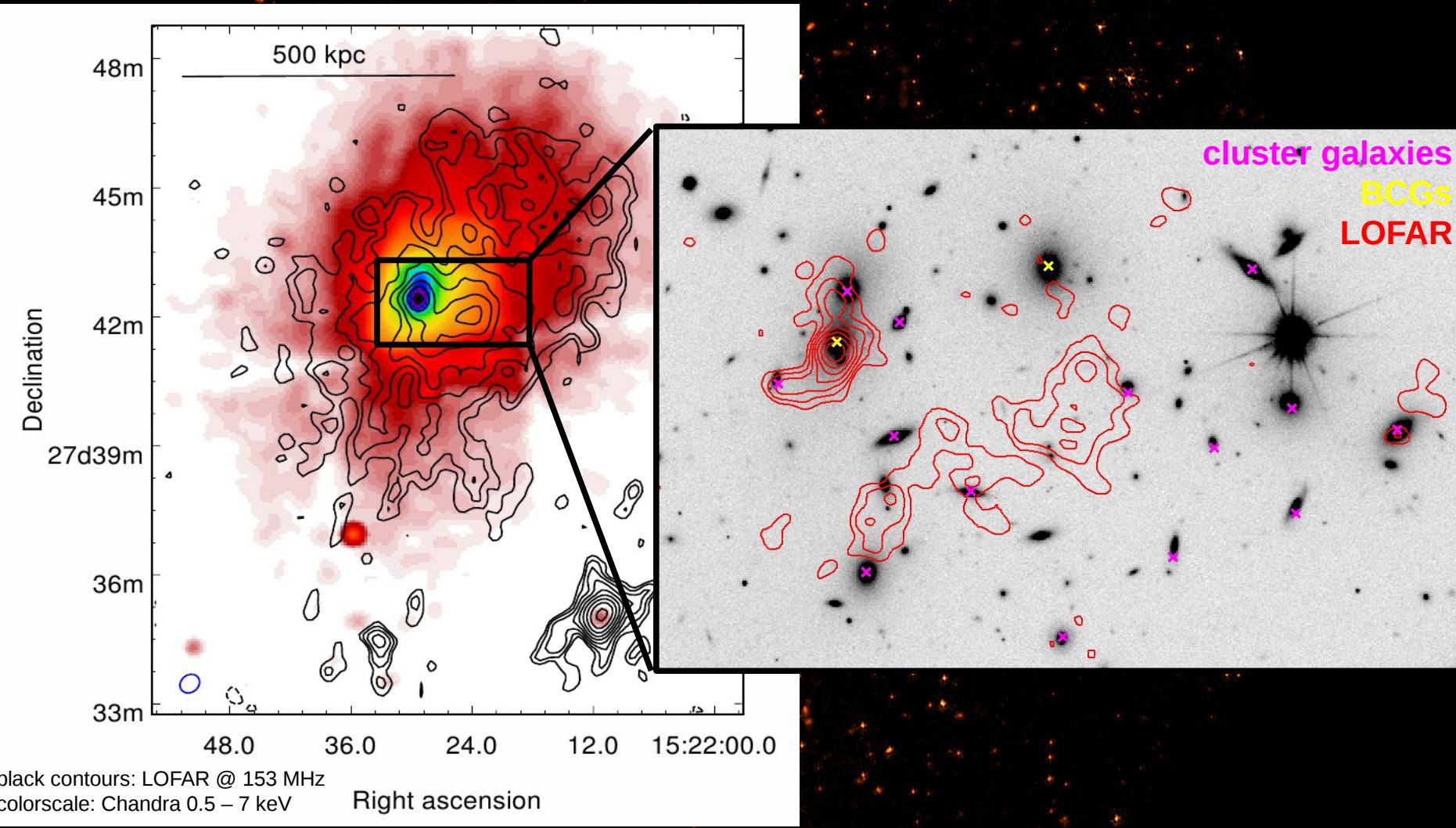


# Abell 2065 radio halo structure recovered

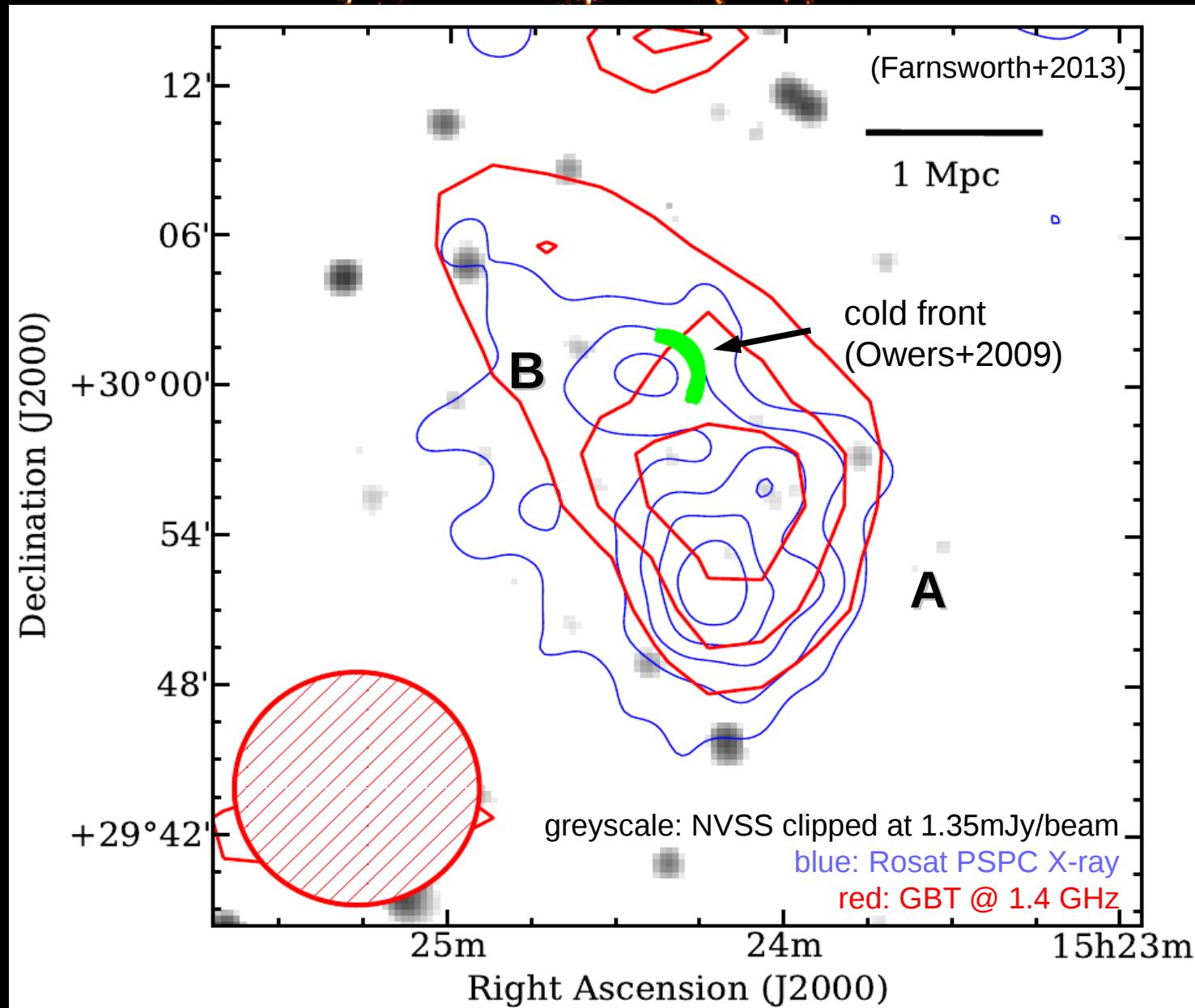


# Abell 2065

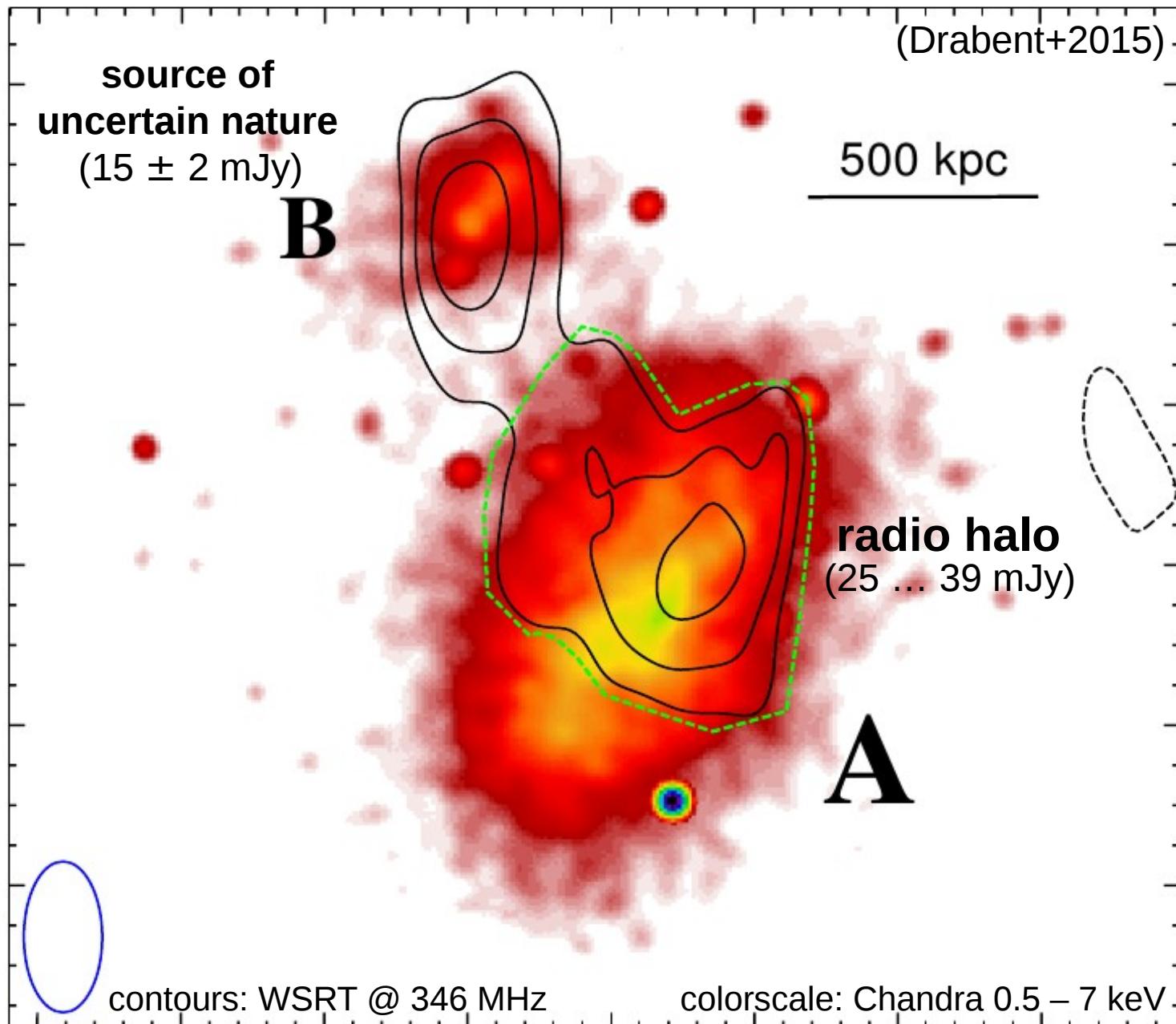
## radio halo structure recovered



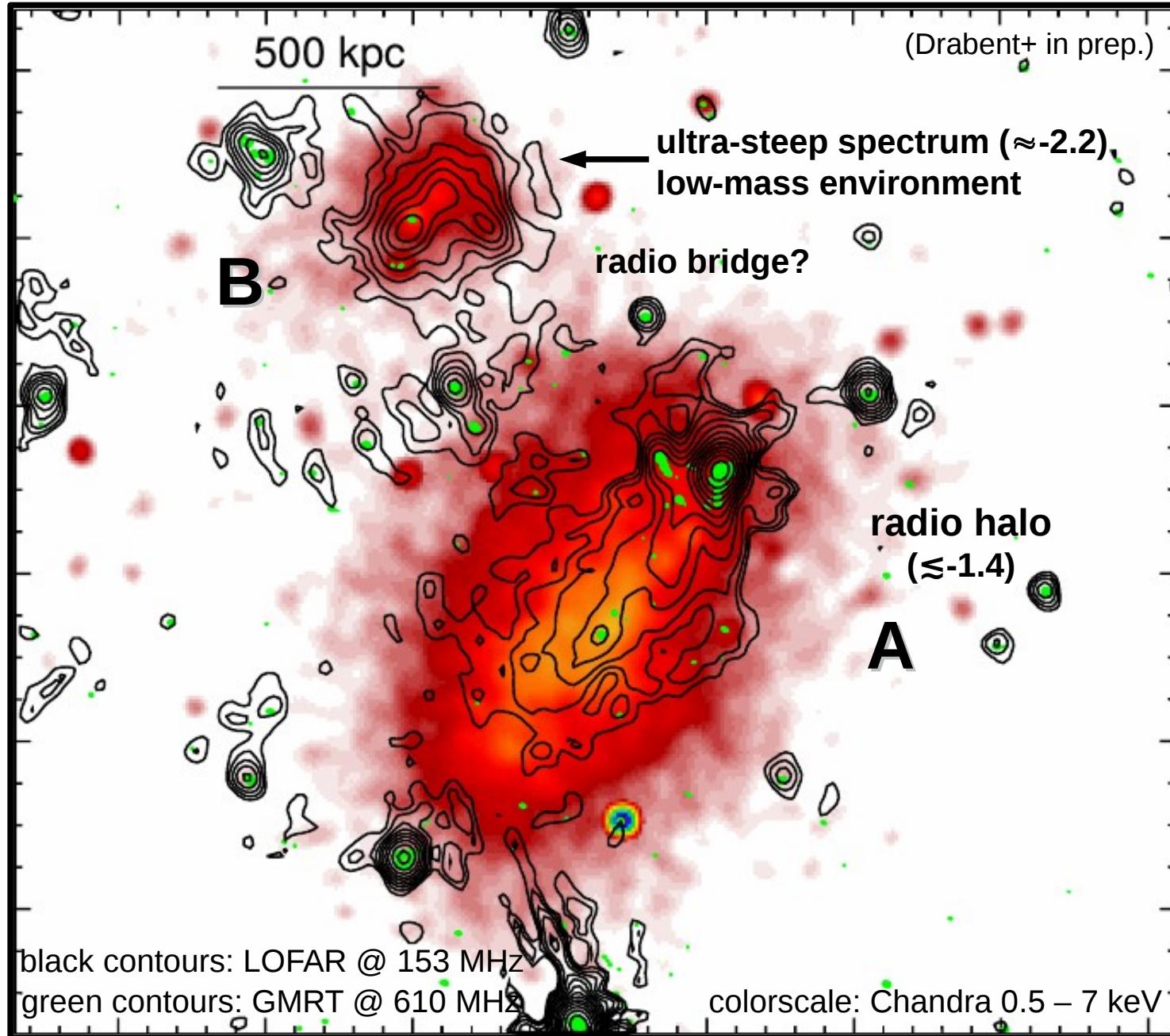
# Abell 2069



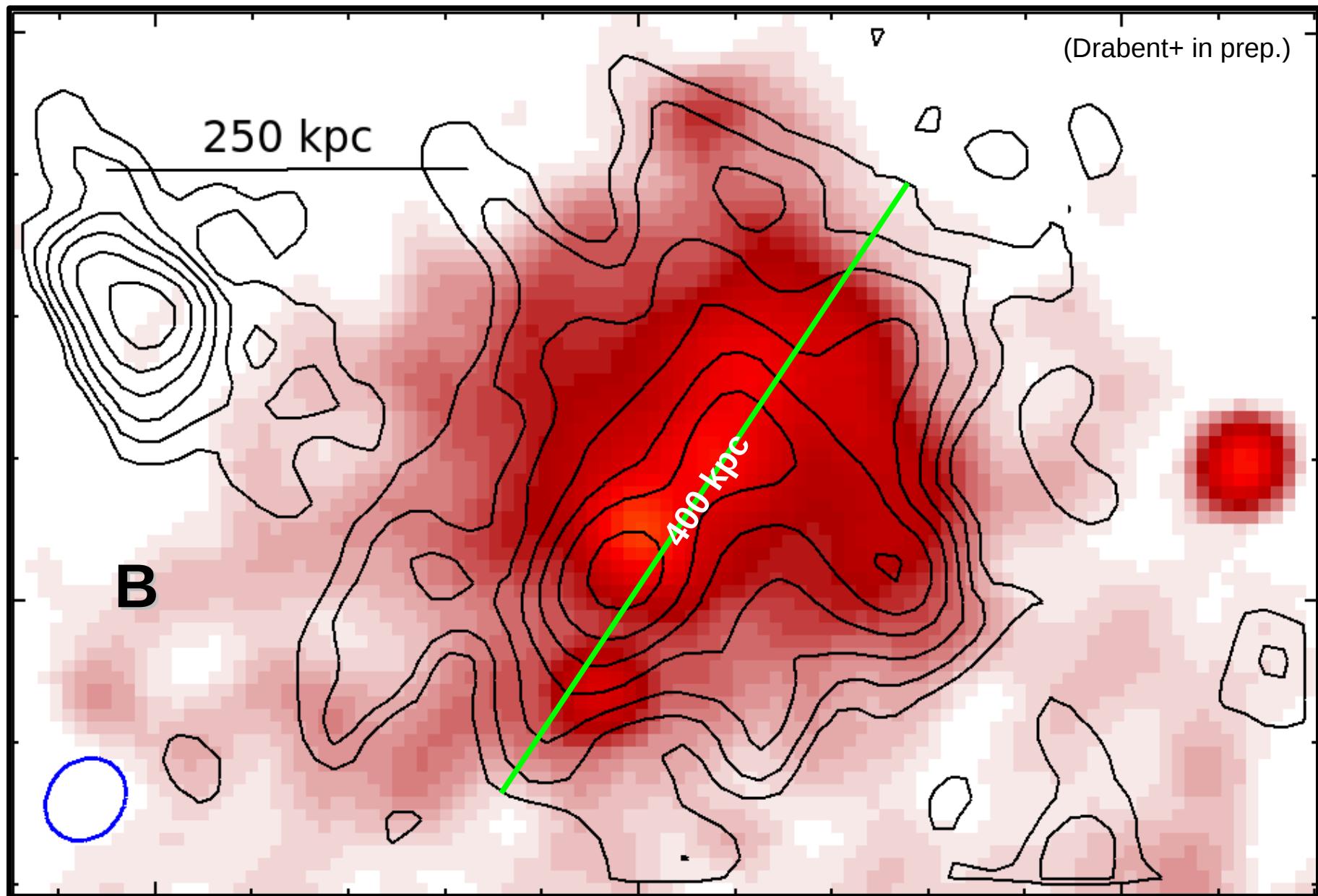
# Abell 2069: diffuse emission in both components



# Abell 2069: diffuse emission is ultra-steep

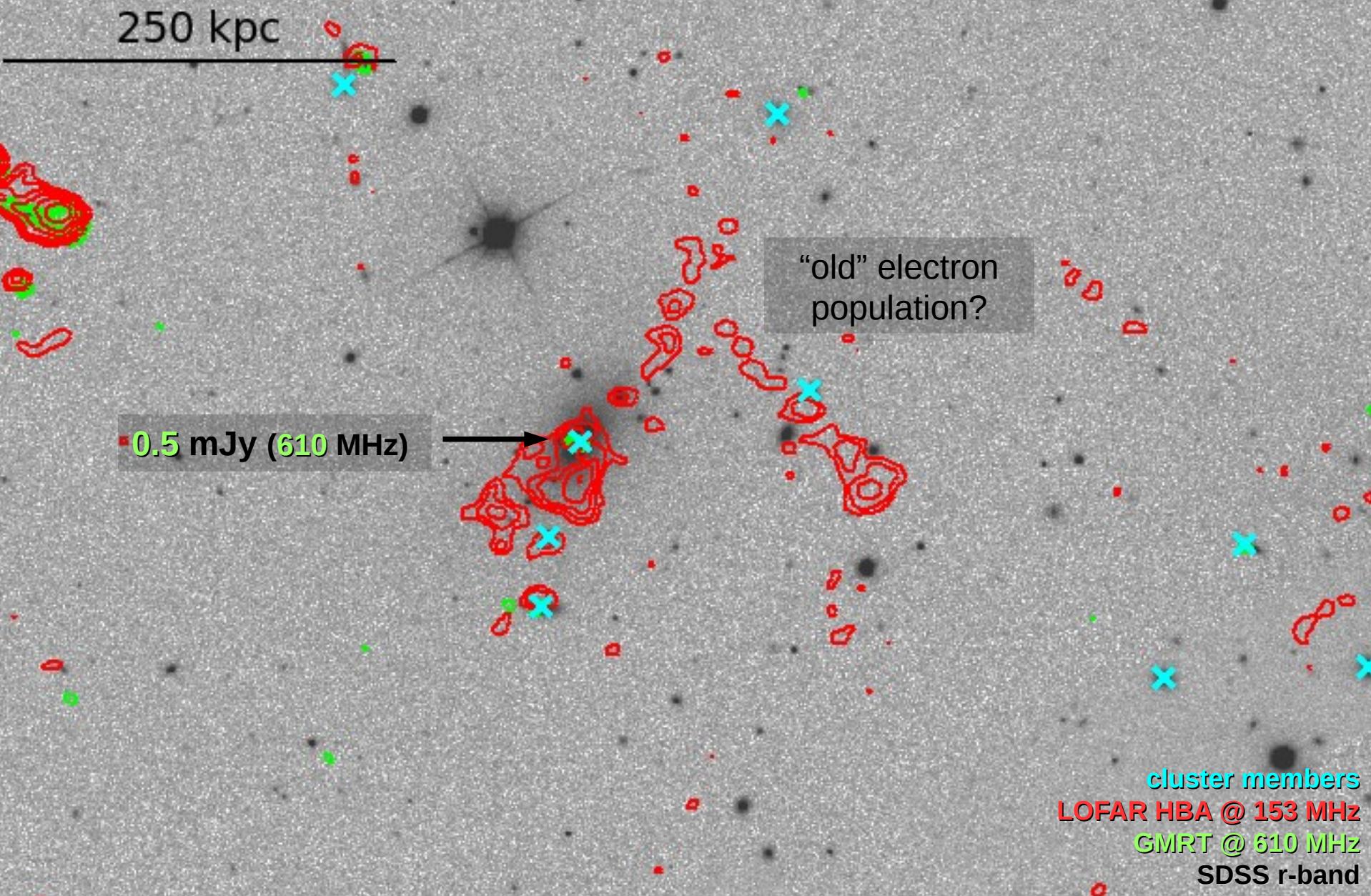


# Abell 2069B: Too small for a radio halo

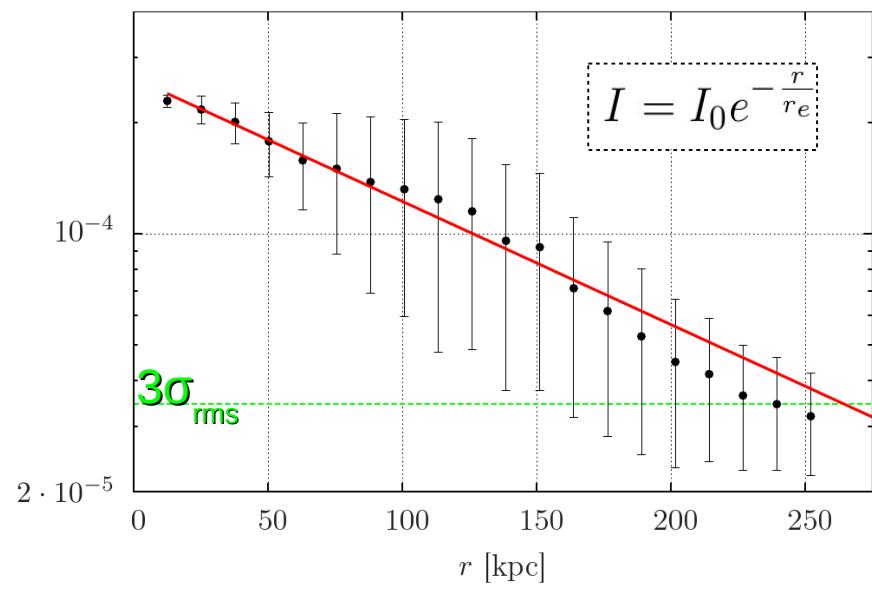
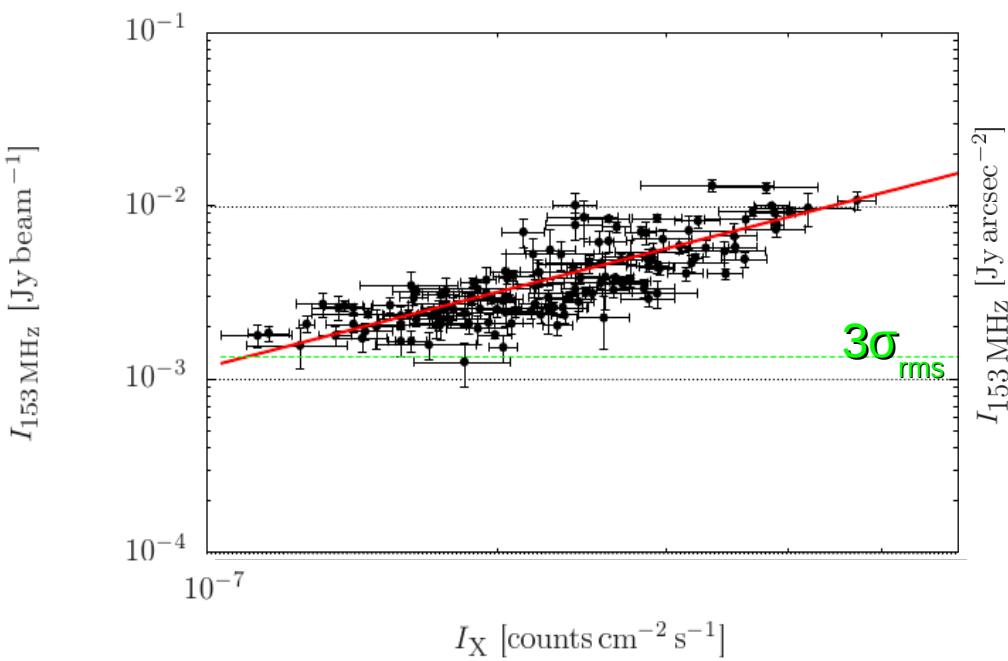
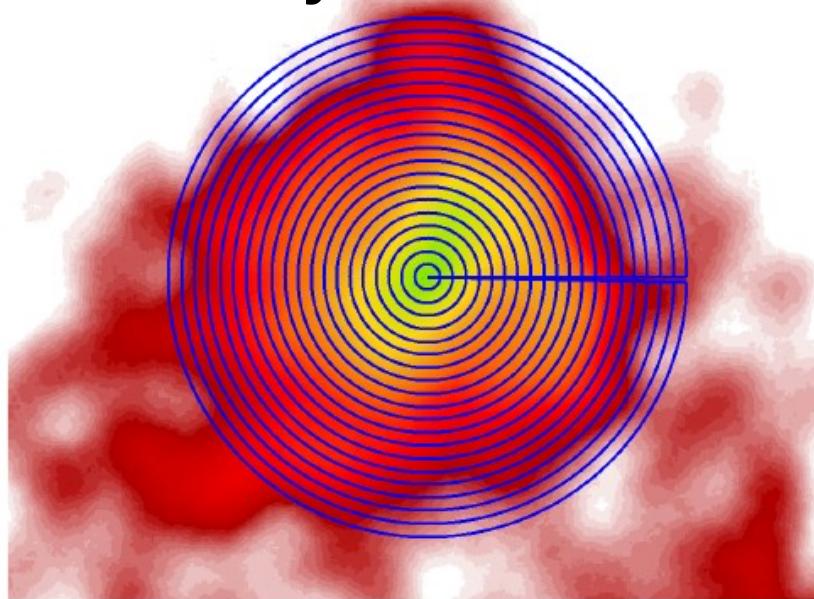
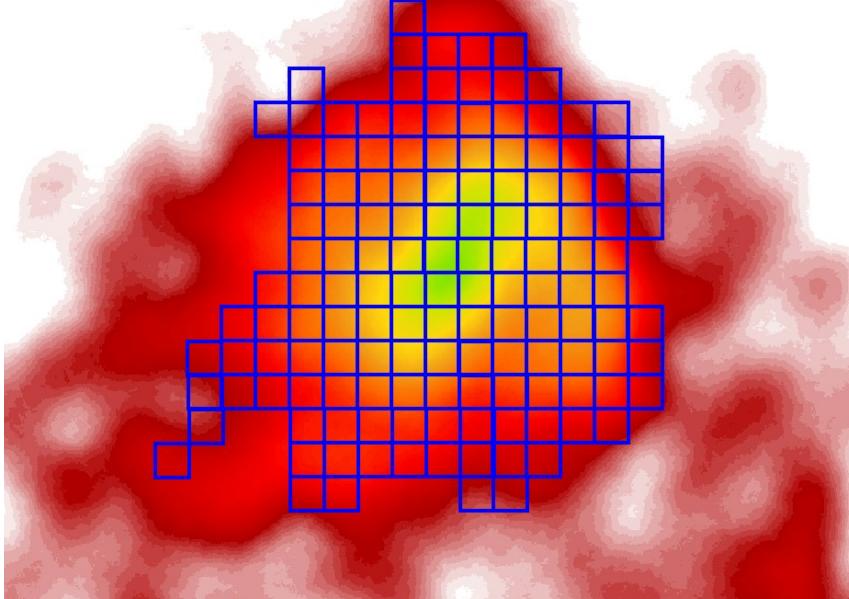


# Abell 2069B: no powerful AGN present

(Drabent+ in prep.)

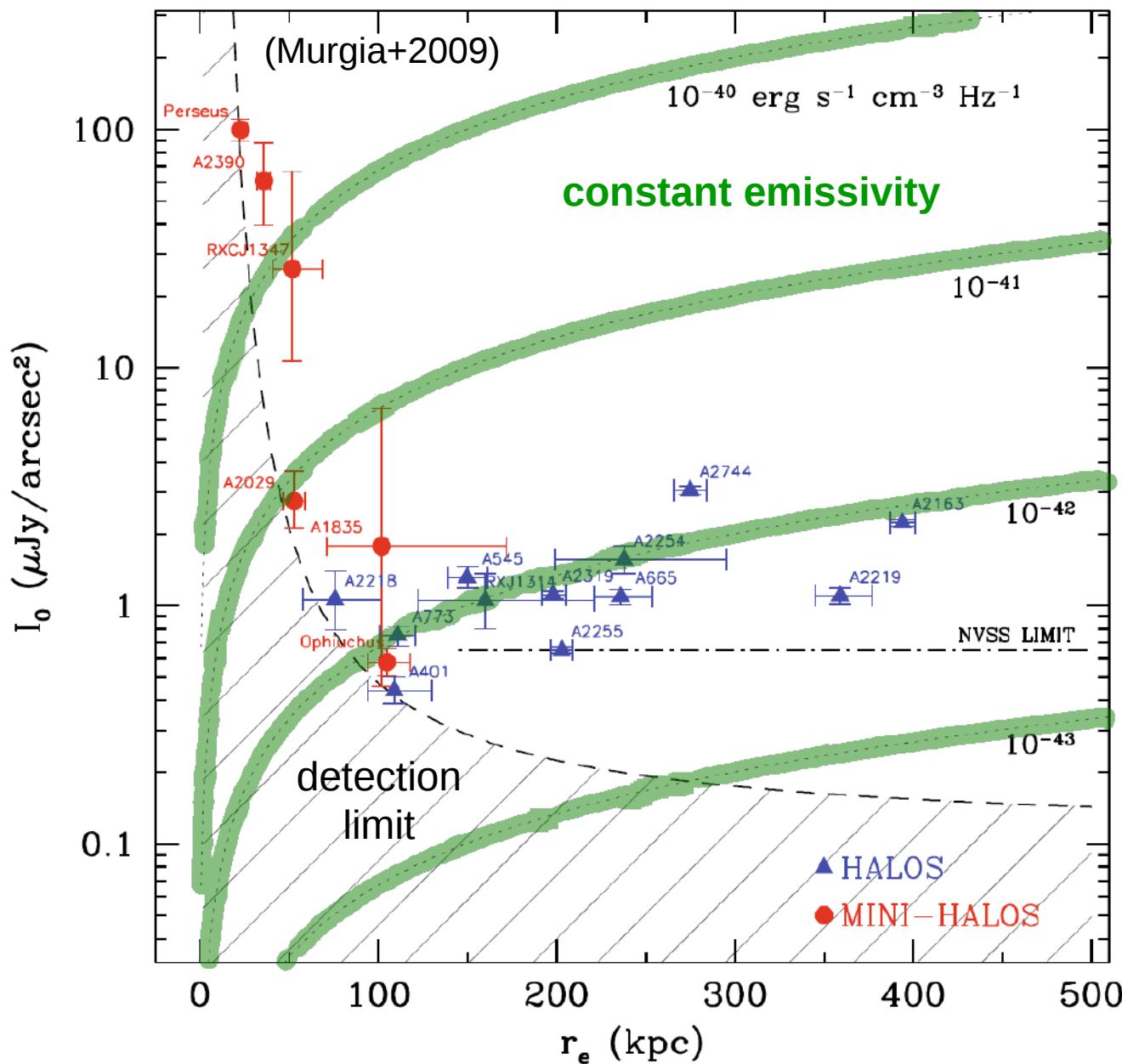


# Abell 2069B: indication for radio – X-ray correlation



→ derive radio emissivity

# Radio mini-halos are not small radio halos

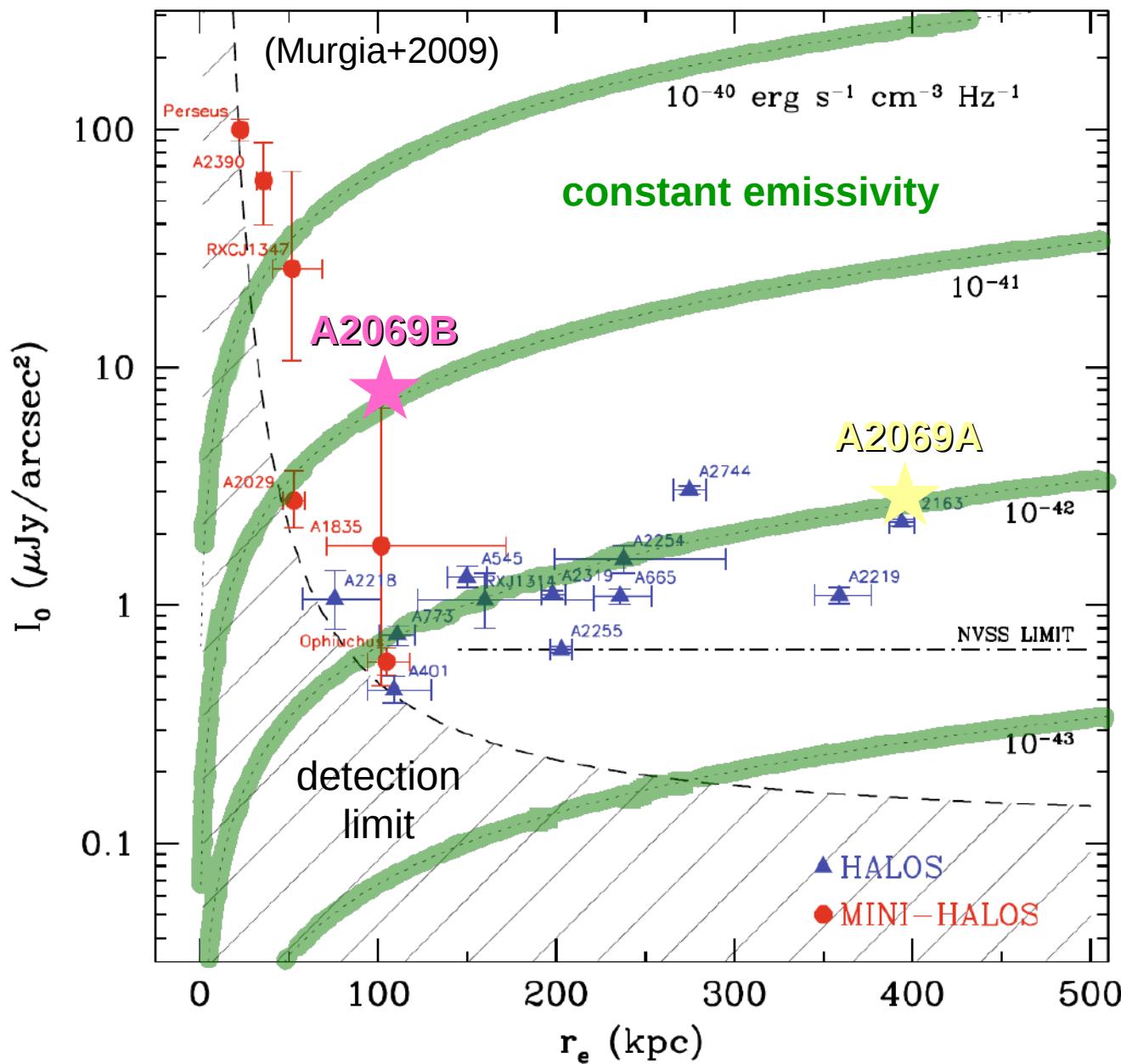


$$I = I_0 e^{-\frac{r}{r_e}}$$

volume-averaged  
radio emissivity

Radio mini-halos:  
wide span of  
emissivities

# Radio mini-halos are not small radio halos



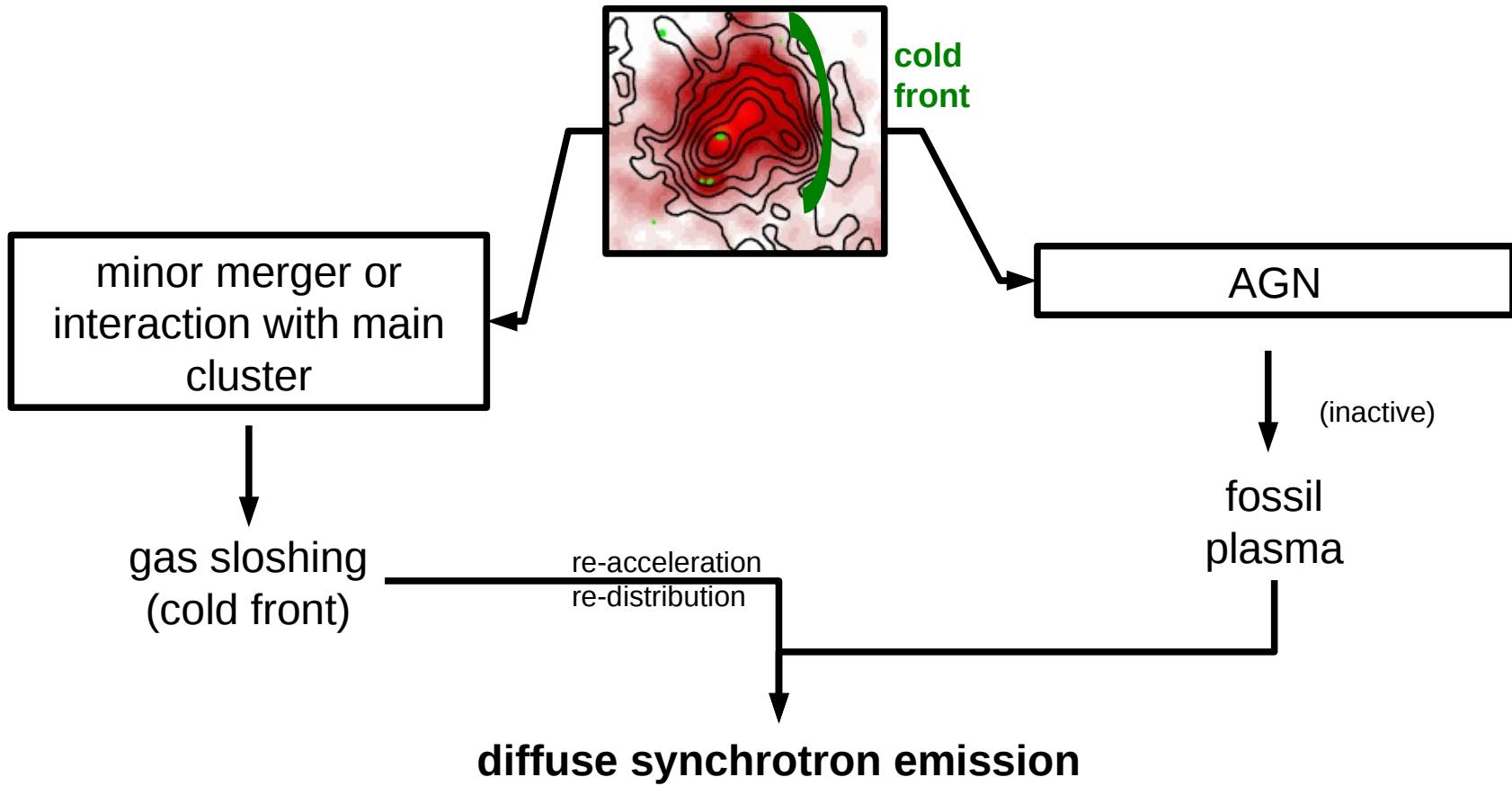
$$I = I_0 e^{-\frac{r}{r_e}}$$

volume-averaged  
radio emissivity

A2069B  
comparable to  
A2029

radio mini-halo?

# Abell 2069B: potential scenario for its origin



→ Abell 2069 is a two-halo system

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