



“Tell me that you have found no sign of  
New Physics again, I dare you.  
I double dare you. Tell me  
one more goddamn **time!**”

# Where to find SUSY?!

*Sven Heinemeyer, IFT/IFCA (CSIC, Madrid/Santander)*

Madrid, 06/2017

1. Introduction & Models
2. Results in GUT based models
3. Results in low-energy models
4. Conclusions



# 1. Introduction & Models

## Some “recent” measurements:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

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# 1. Introduction & Models

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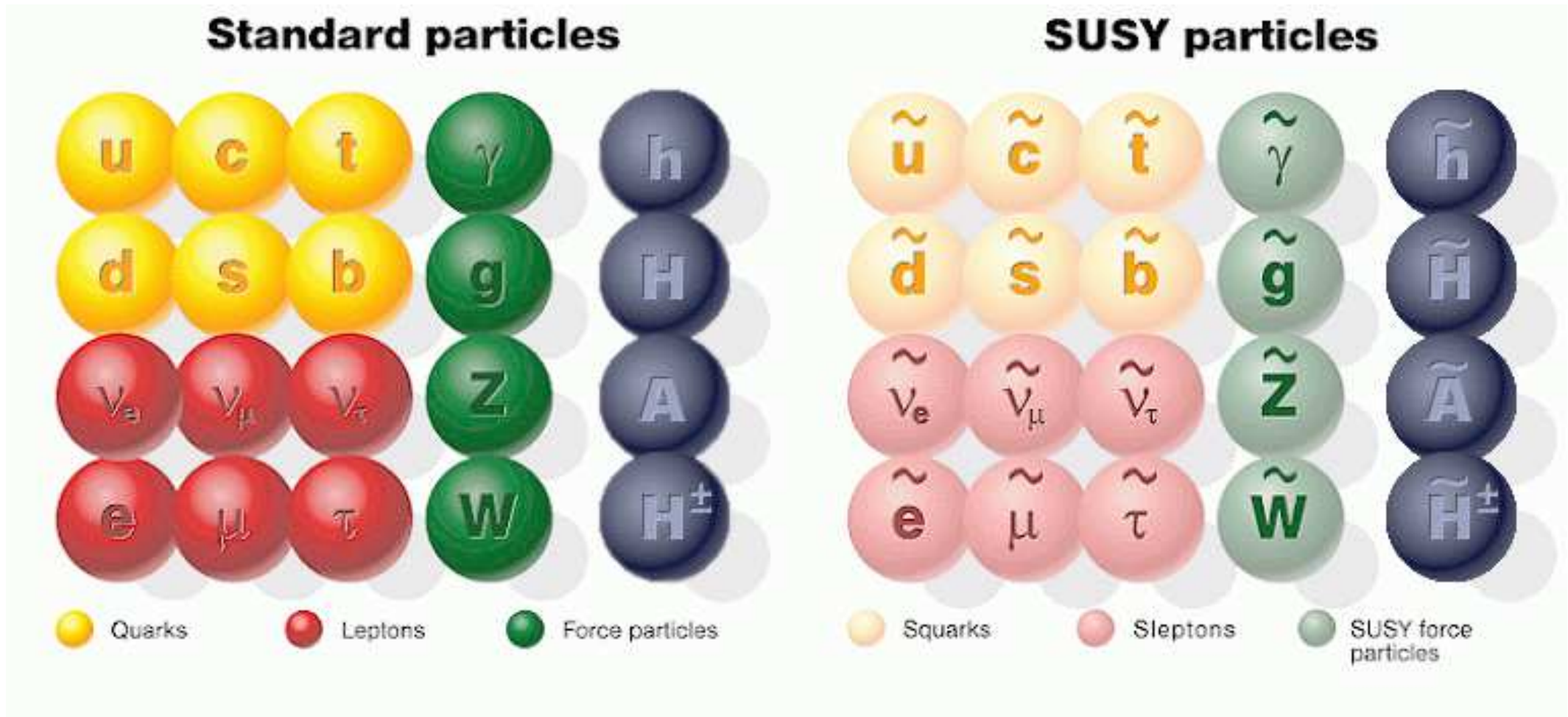
## Simple SUSY models predicted correctly:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

⇒ good motivation to look at SUSY!

# The Minimal Supersymmetric Standard Model (MSSM)

## Superpartners for Standard Model particles



Problem in the MSSM: more than 100 free parameters

Nobody(?) believes that a model describing nature has so many free parameters!

## A. Unconstrained models (MSSM):

agnostic about how SUSY breaking is achieved

no particular SUSY breaking mechanism assumed, parameterization of possible soft SUSY-breaking terms

most general case:

⇒ 105 new parameters: masses, mixing angles, phases

⇒ no model missed (within the MSSM)

⇒  $\mathcal{O}(100)$  parameters difficult to handle

## B. Constrained models:

CMSSM, NUHM1, NUHM2, SU(5), mAMSB, ...:

assumption on the scenario that achieves spontaneous SUSY breaking

⇒ prediction for soft SUSY-breaking terms  
in terms of small set of parameters

⇒ easy to handle

GUT based models: 1.) CMSSM (sometimes wrongly called mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

$m_0$  : universal scalar mass parameter

$m_{1/2}$  : universal gaugino mass parameter

$A_0$  : universal trilinear coupling

$\tan \beta$  : ratio of Higgs vacuum expectation values

$\text{sign}(\mu)$  : sign of supersymmetric Higgs parameter

} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

⇒ Lightest SUSY particle (LSP) is the lightest neutralino ⇒ DM!



GUT based models: 2.) NUHM1: (Non-universal Higgs mass model)

**Assumption:** no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively  $M_A$  as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu \text{ and } M_A$$

GUT based models: 3.) NUHM2: (Non-universal Higgs mass model 2)

**Assumption:** no unification of scalar Higgs parameter at the GUT scale

⇒ effectively  $M_A$  and  $\mu$  as free parameters at the EW scale

⇒ Scenario characterized by

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## GUT based models: 4.) SU(5) GUT:

### Assumption I:

no unification of scalar Higgs parameter at the GUT scale

( $\Rightarrow$  effectively  $M_A$  and  $\mu$  as free parameters at the EW scale)

### Assumption II:

$$(q_L, u_L^c, e_L^c)_i \in \mathbf{10}_i, (\ell_L, d_L^c)_i \in \bar{\mathbf{5}}_i$$

$\Rightarrow$  Scenario characterized by

$$m_5, m_{10}, m_{1/2}, A_0, \tan \beta, m_{H_u}, m_{H_d}$$

## GUT based models: 5.) mAMSB:

mAMSB scenario characterized by

$$m_{3/2}, m_0, \tan \beta, \text{sign}(\mu)$$

$m_{3/2} = \langle F \rangle / M_{\text{Planck}}$ : overall scale of SUSY particle masses

$m_0$ : phenomenological parameter: universal scalar mass term introduced in order to keep squares of slepton masses positive

typical feature: very small neutralino–chargino mass difference  
 $\Rightarrow \tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 + \pi^\pm$  with very soft pions

## Problem: We cannot be sure about the SUSY-breaking mechanism

- ⇒ it is possible that with the CMSSM, NUHM1, NUHM2, SU(5), mAMSB we missed the “correct” mechanism
- ⇒ hint: strong connection between colored and uncolored sector  
tension between low-energy EW effects and (colored) LHC searches

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- ⇒ it is possible that with the CMSSM, NUHM1, NUHM2, SU(5), mAMSB we missed the “correct” mechanism
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tension between low-energy EW effects and (colored) LHC searches

## Solution: investigate also the “general MSSM”

⇒ 10 parameters are manageable ⇒ pMSSM10

- squark mass parameters:  $m_{\tilde{q}_{1,2}} =: m_{\tilde{q}}, m_{\tilde{q}_3}$
- slepton mass parameter:  $m_{\tilde{l}}$
- gaugino masses:  $M_1, M_2, M_3$
- trilinear coupling:  $A$
- Higgs sector parameters:  $M_A, \tan \beta$
- Higgs mixing parameter:  $\mu$

## Our tool: **Mastercode**



⇒ collaborative effort of theorists and experimentalists

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*[Bagnaschi, Borsato, Buchmüller, Cavanaugh, Chobanova, Citron, Costa, De Roeck, Dolan, Ellis, Flücher, SH, Isidori, Liu, Lucio, Martinez Santos, Olive, Richards, Sakurai, Weiglein]*

### Über-code for the combination of different tools:

– Über-code original in Fortran, now re-written in C++

– tools are included as **subroutines**

– **compatibility** ensured by collaboration of authors of “MasterCode” and authors of “sub tools” **/SLHA(2)**

– sub-codes in Fortran or C++

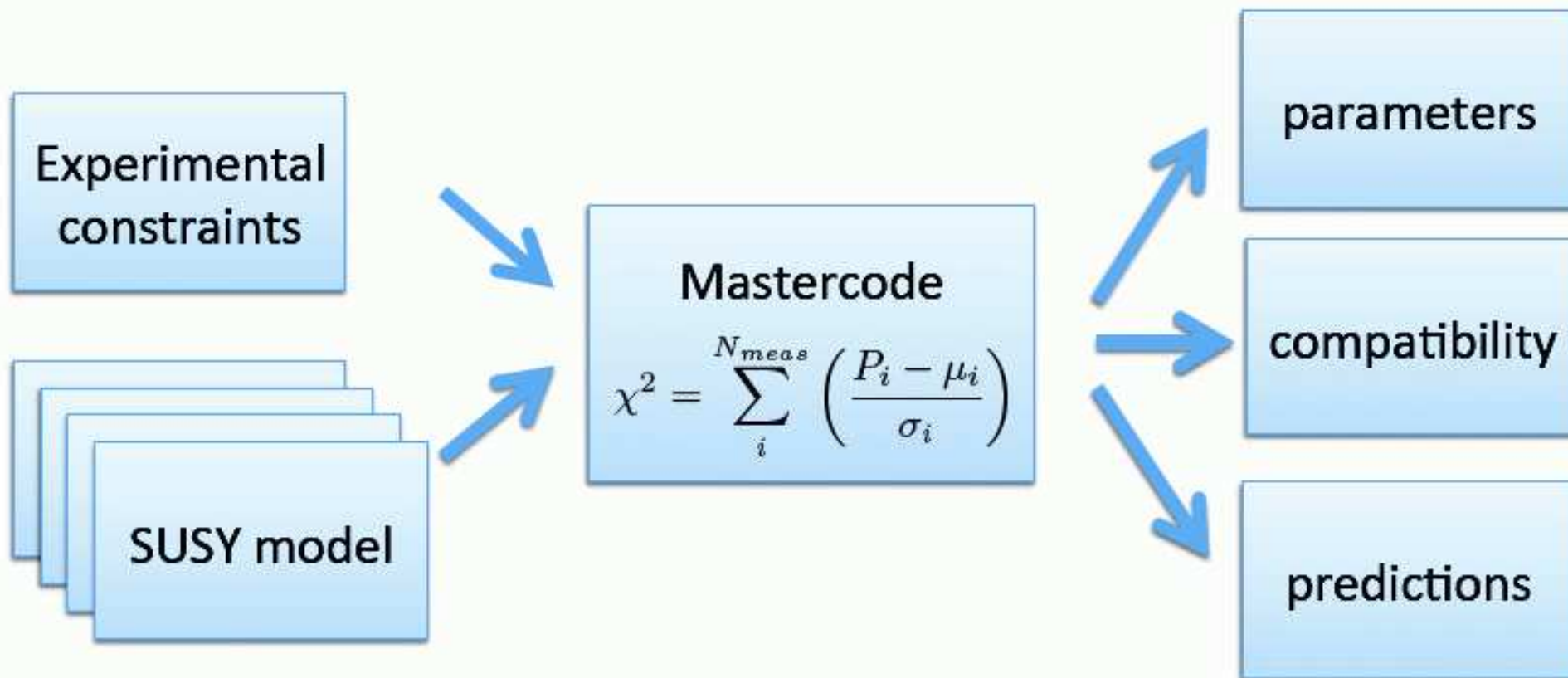
⇒ evaluate observables of one parameter point consistently with various tools

[cern.ch/mastercode](http://cern.ch/mastercode)

## The $\chi^2$ evaluation:



# Global fits of SUSY



## Data we have:

- Higgs boson mass (LHC)  $\Rightarrow$  FeynHiggs



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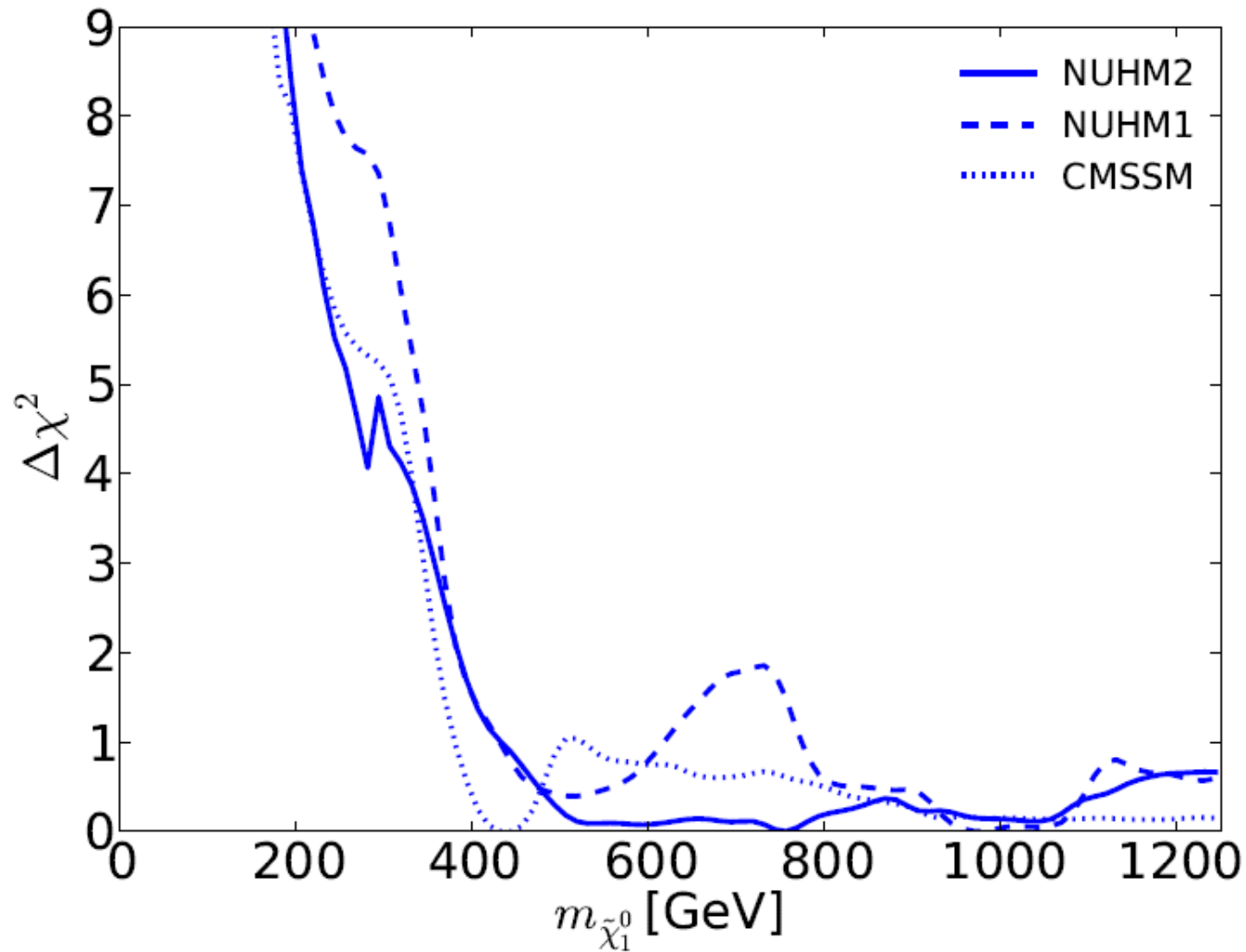
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- SUSY searches (LHC)  $\Rightarrow$  own re-cast
- electroweak precision data  $\Rightarrow$  FeynWZ, FeynHiggs
- flavor data  $\Rightarrow$  SuperIso, SuFla
- astrophysical data (DM properties)  $\Rightarrow$  MicrOMEGAs, SSARD

## 2. Results GUT based models

### Results in the CMSSM, NUHM1, NUHM2

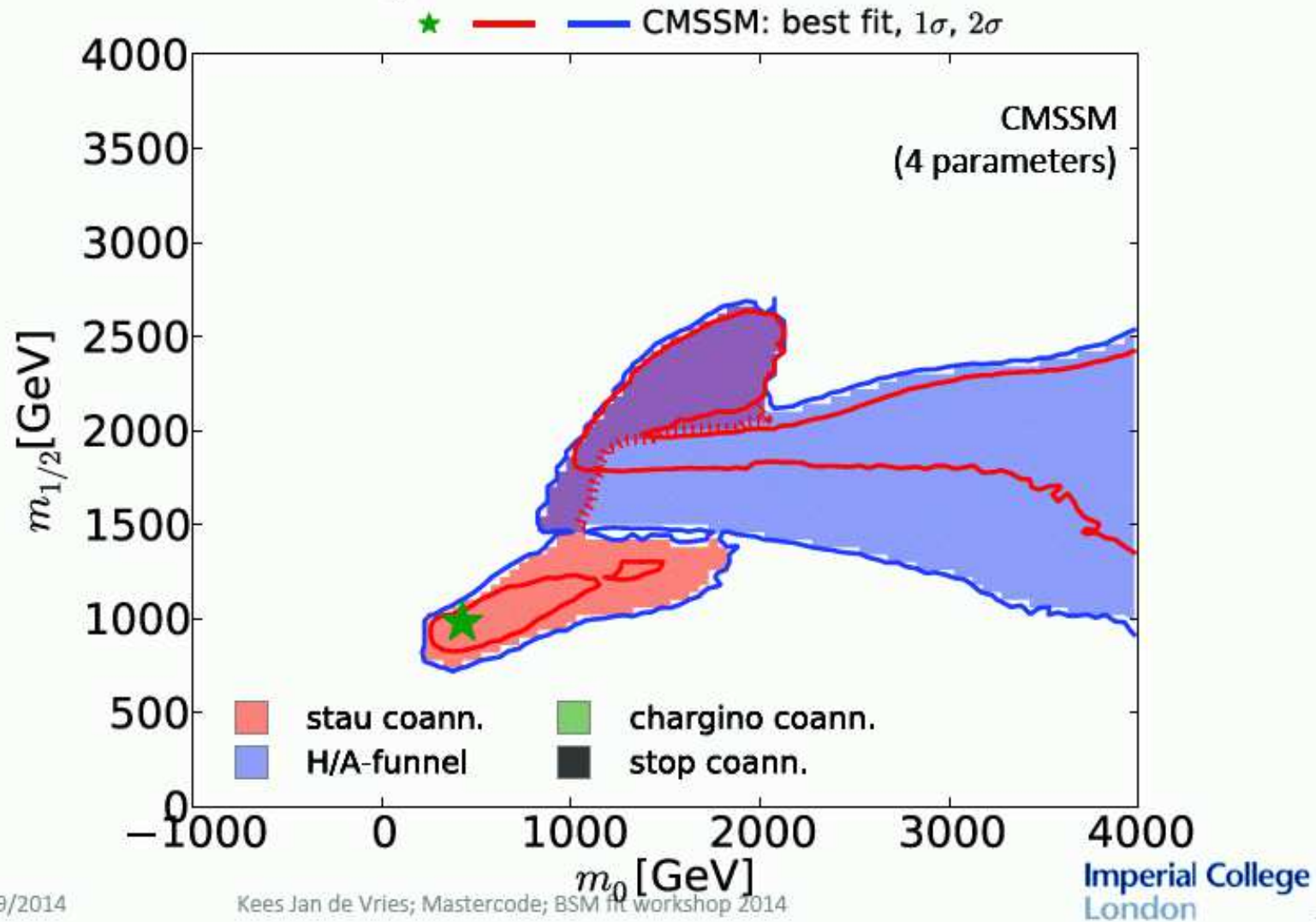


⇒ only very large values are favored

[2014]



# Mechanisms for relic dark matter density fulfillment in the CMSSM

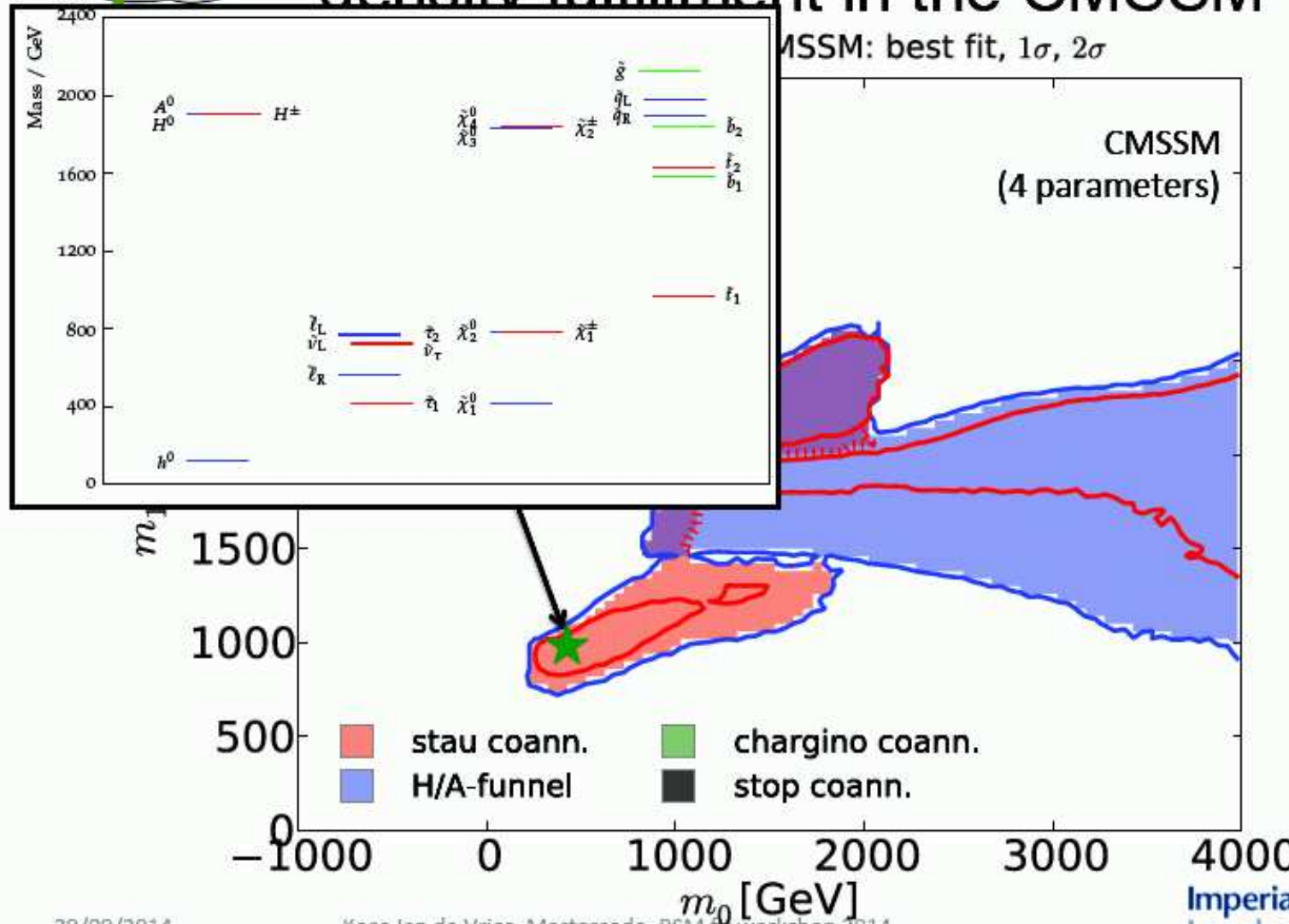


29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014



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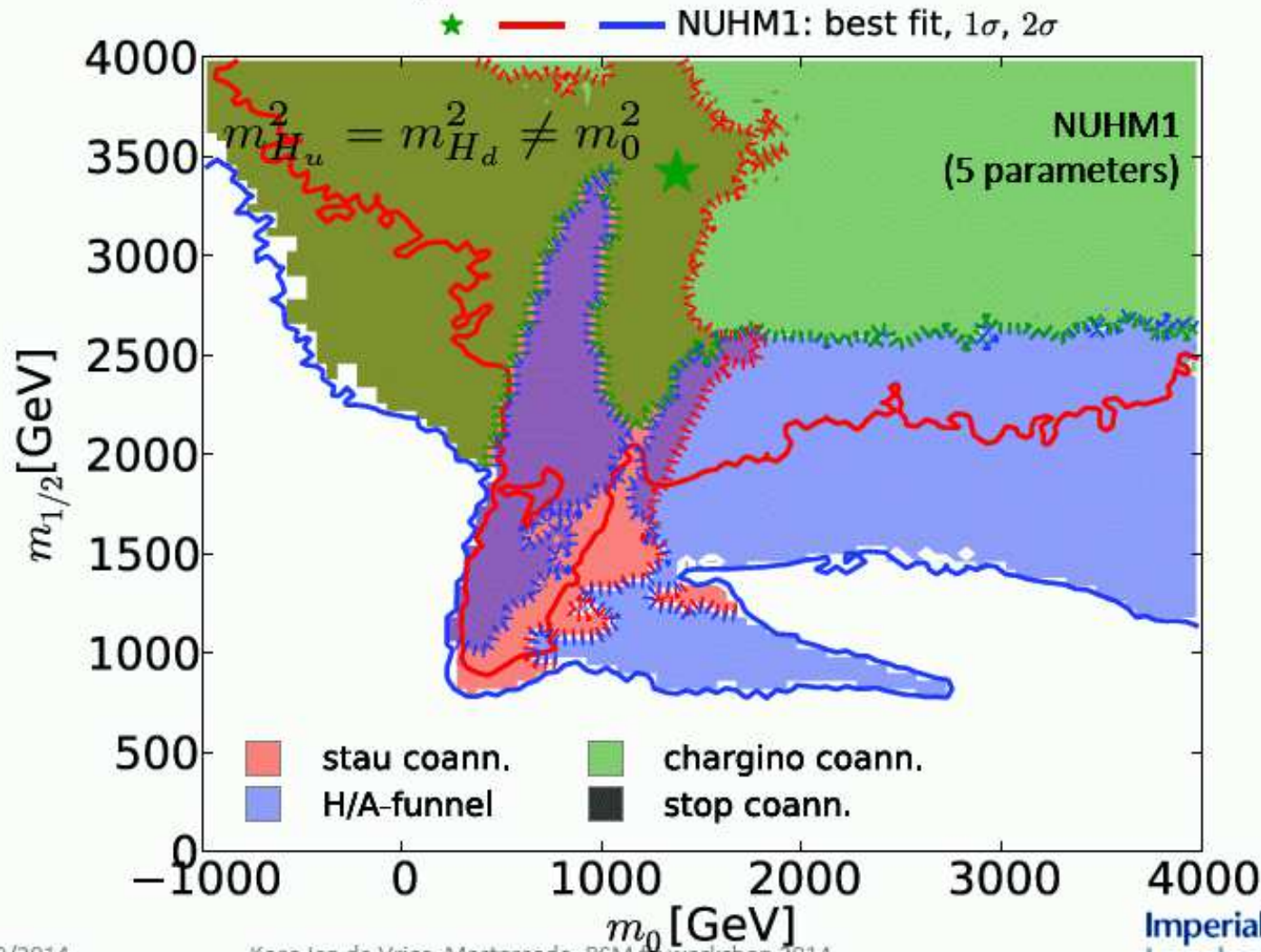
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# Mechanisms for relic dark matter density fulfillment in the NUHM1



29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014

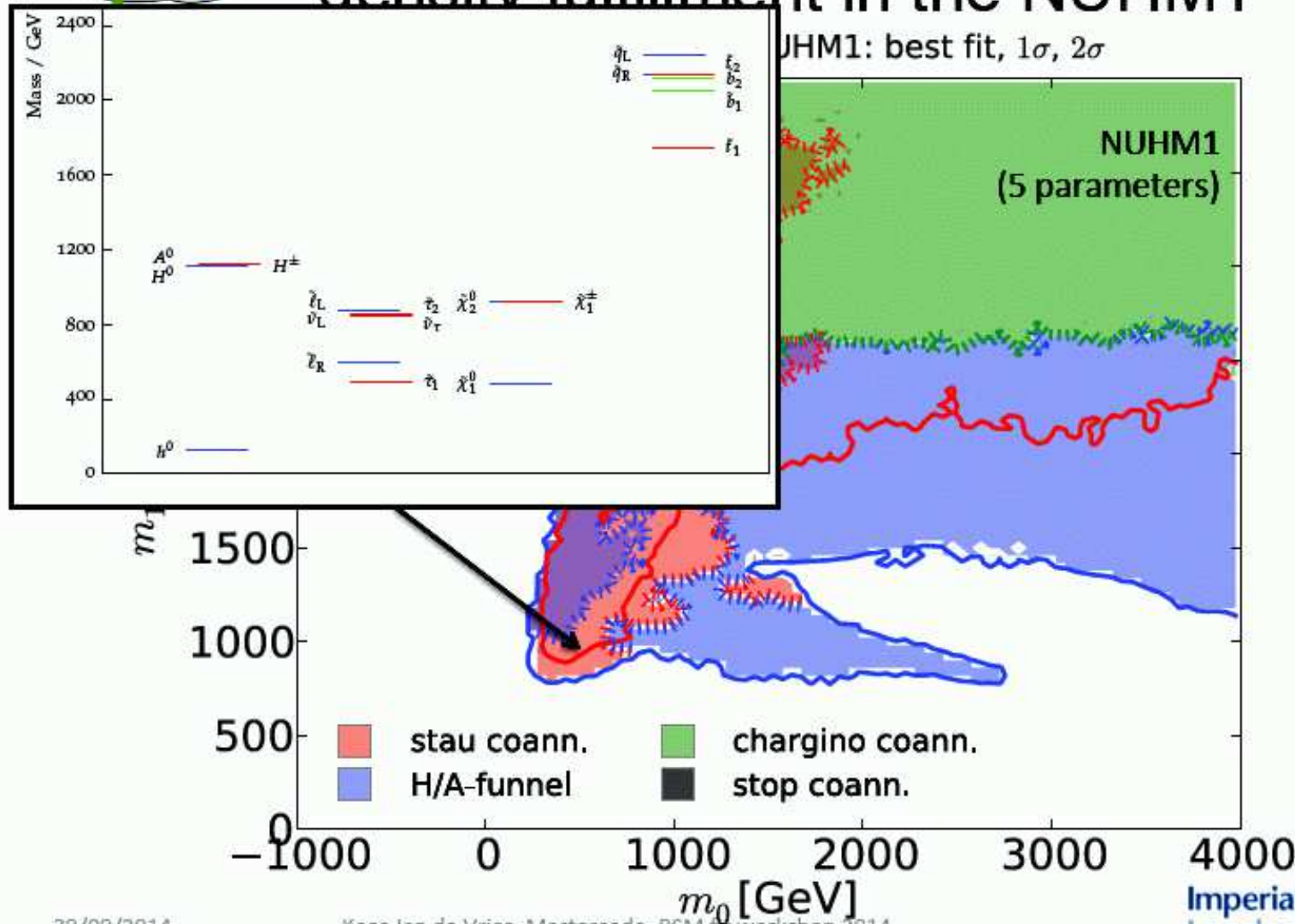
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# Mechanisms for relic dark matter density fulfillment in the NUHM1



29/09/2014

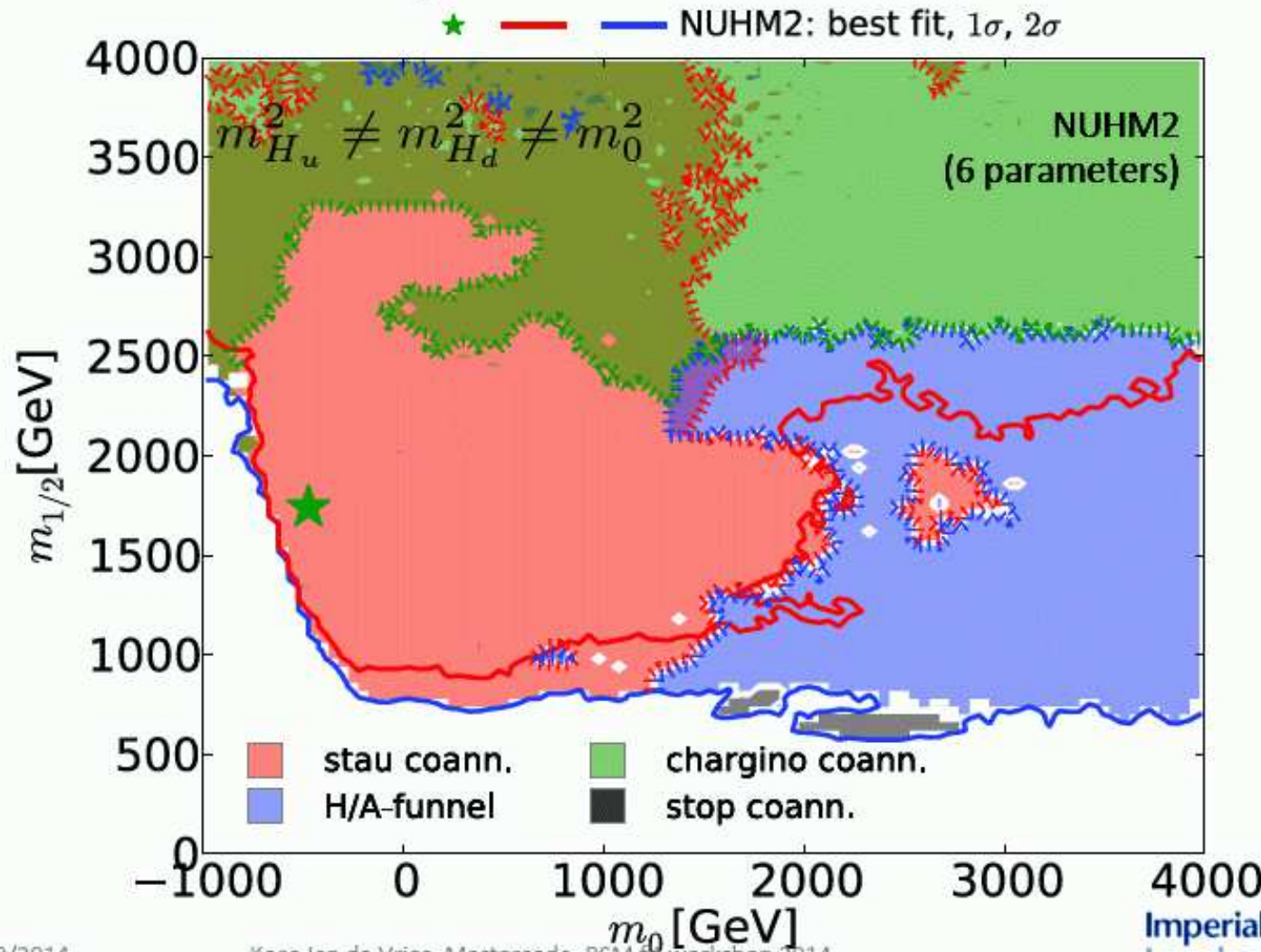
Kees Jan de Vries; Mastercode; BSM fit workshop 2014

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London

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# Mechanisms for relic dark matter density fulfillment in the NUHM2



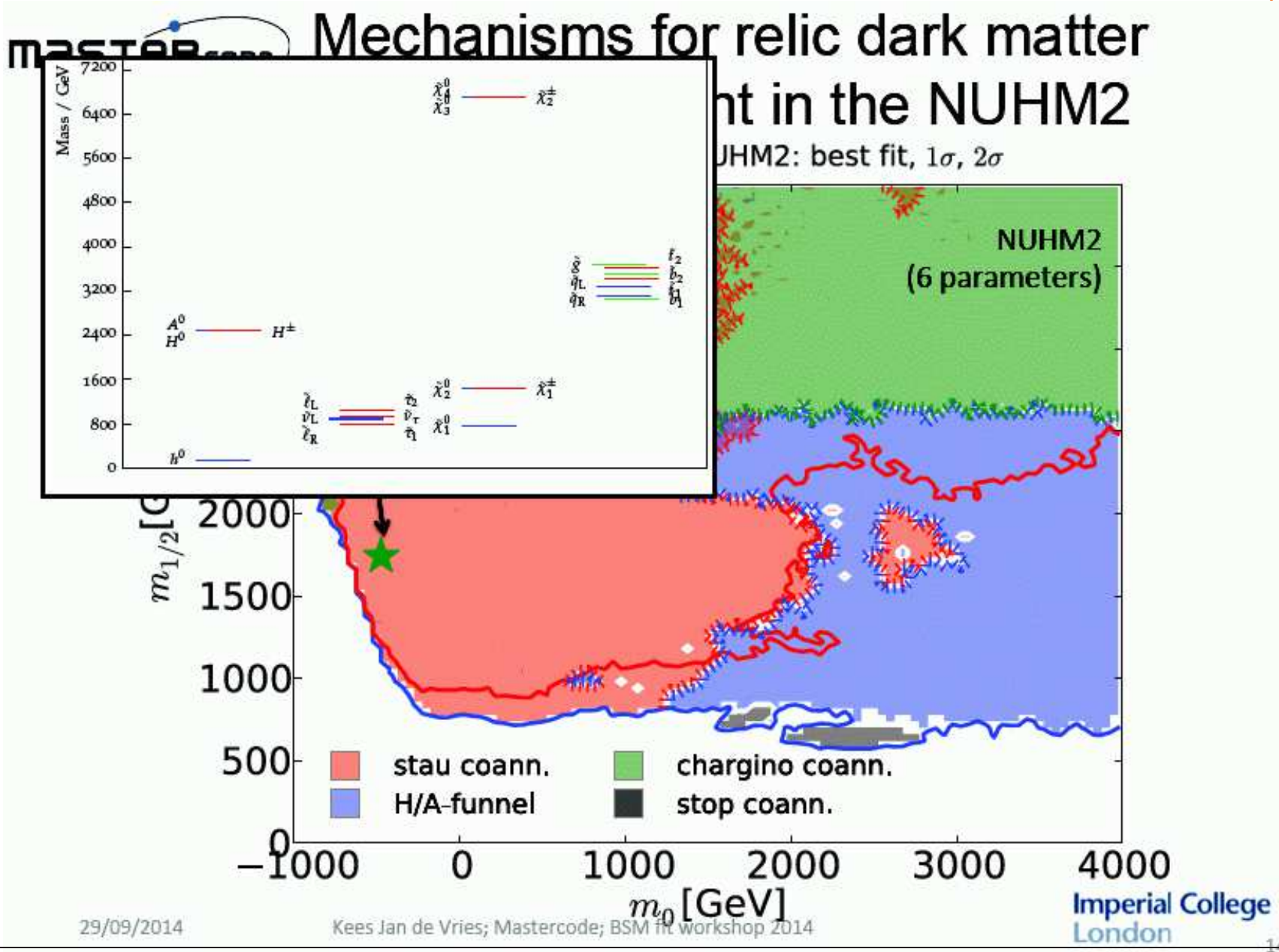
29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014

Imperial College London

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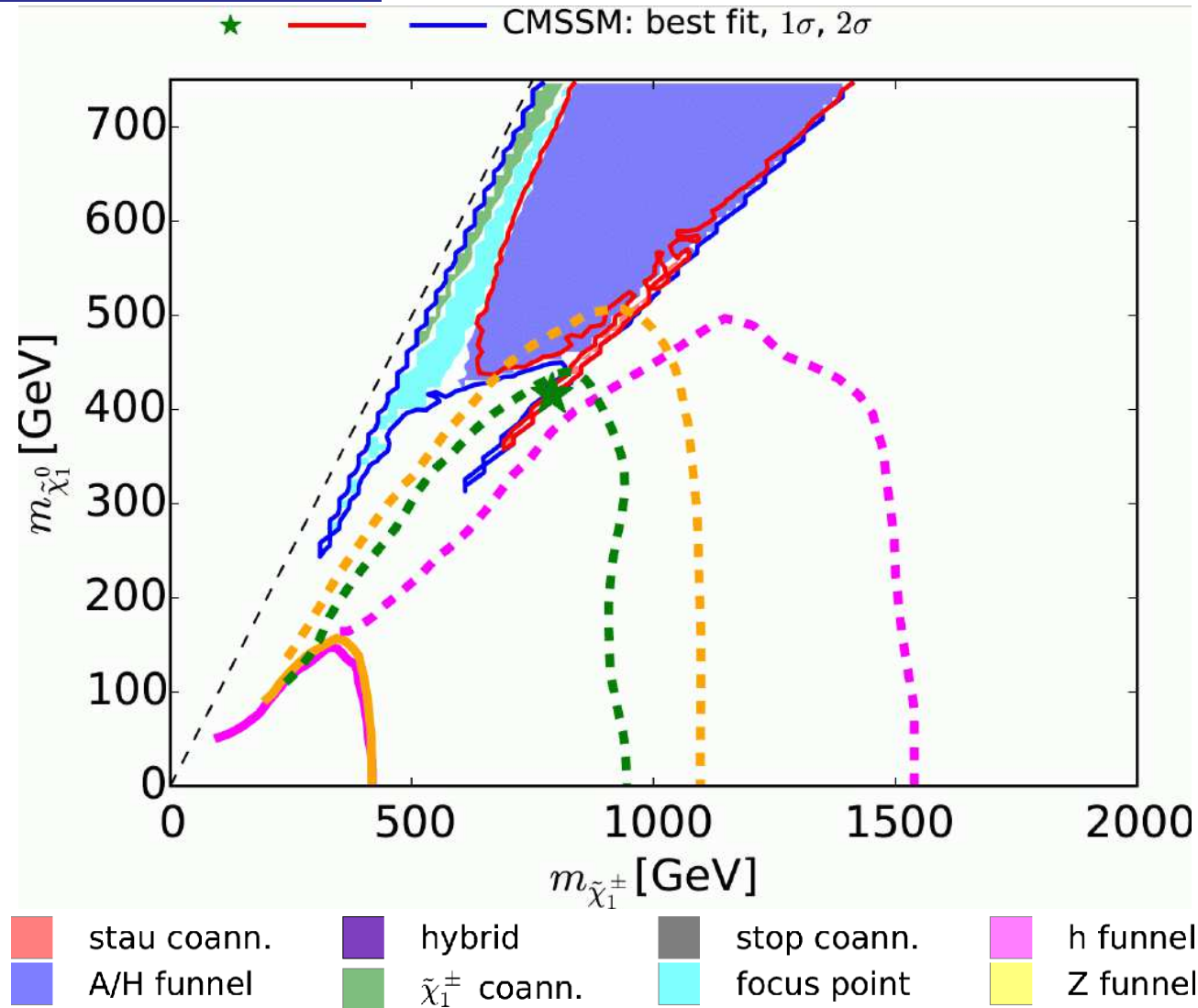
[2014]





# LHC prospects for CMSSM:

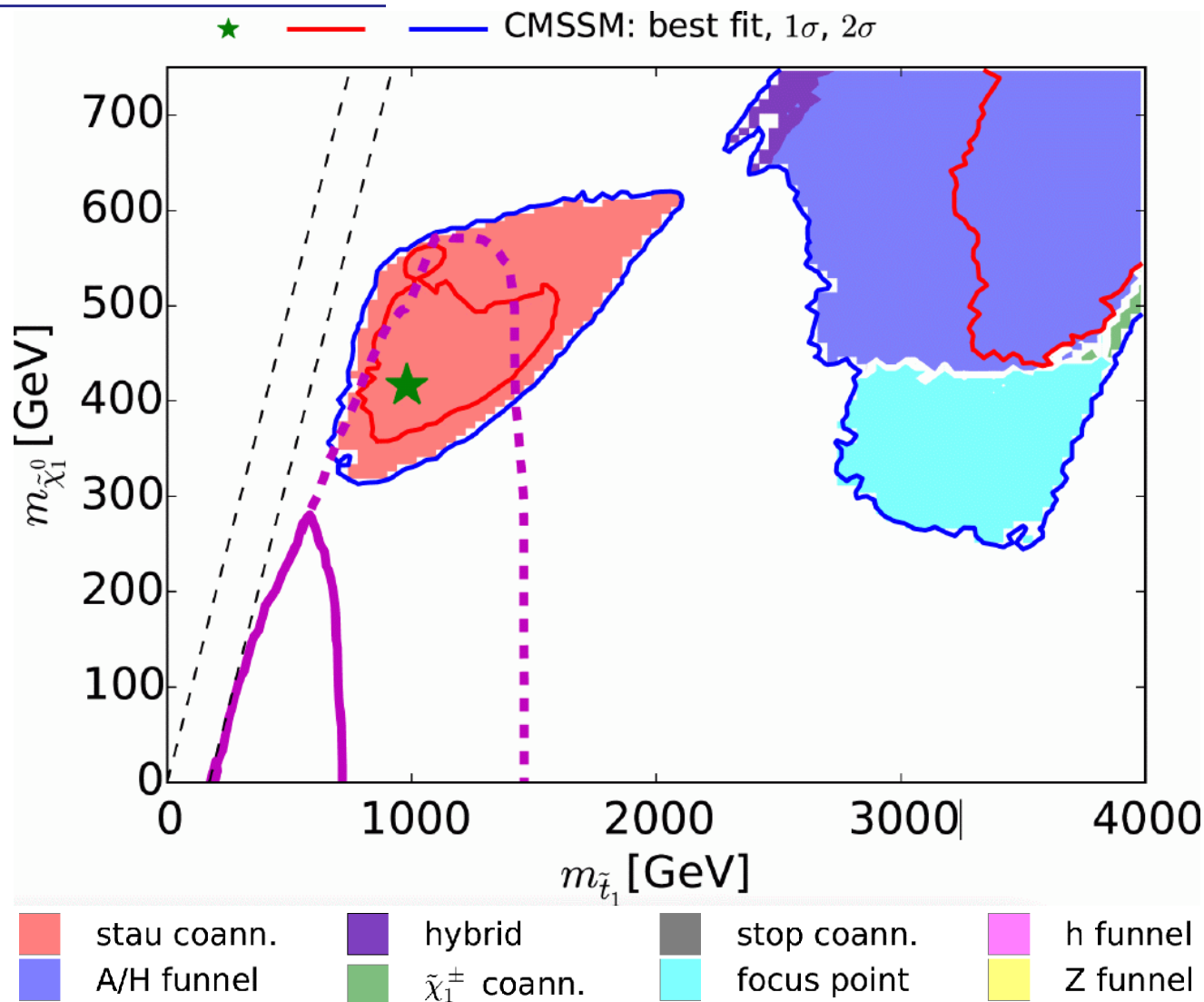
[2015]



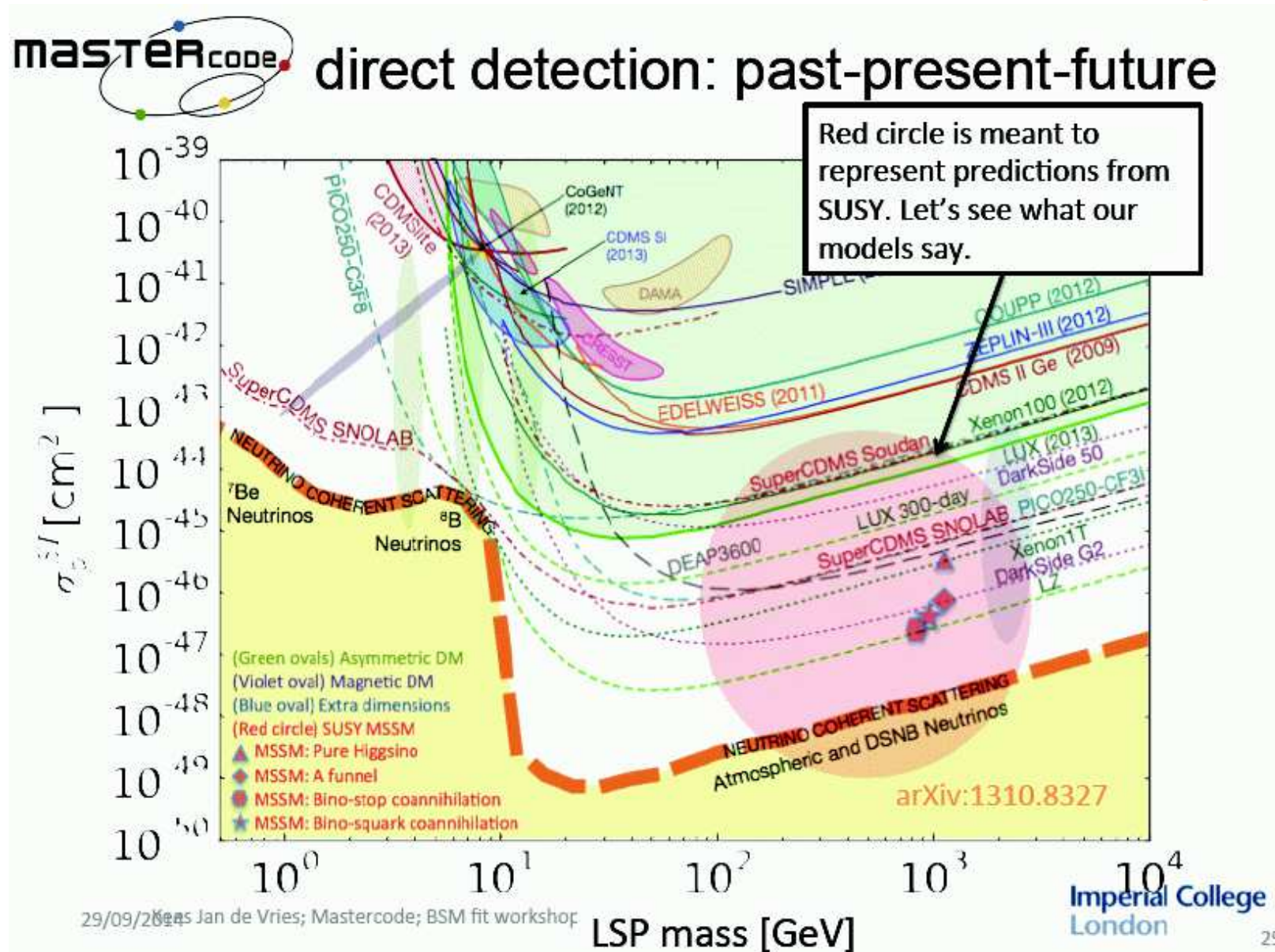
solid: current LHC limits, dashed: HL-LHC prospects  
 $\Rightarrow$  best-fit regions can be covered! (in EW searches)

# LHC prospects for CMSSM:

[2015]

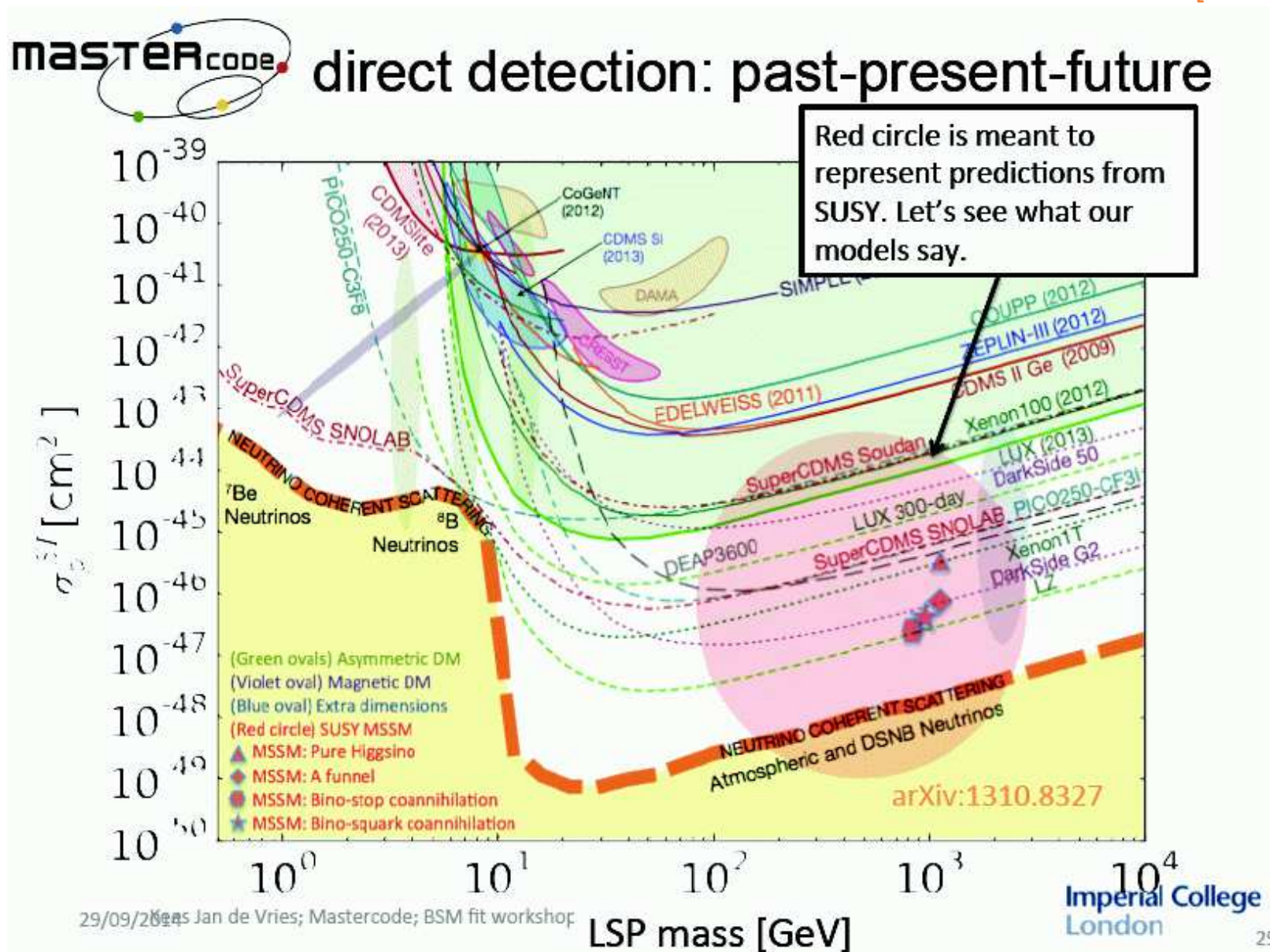


solid: current LHC limits, dashed: HL-LHC prospects  
 ⇒ best-fit regions can partially be covered! (in colored searches)



29/09/2014 Jan de Vries; Mastercode; BSM fit workshop



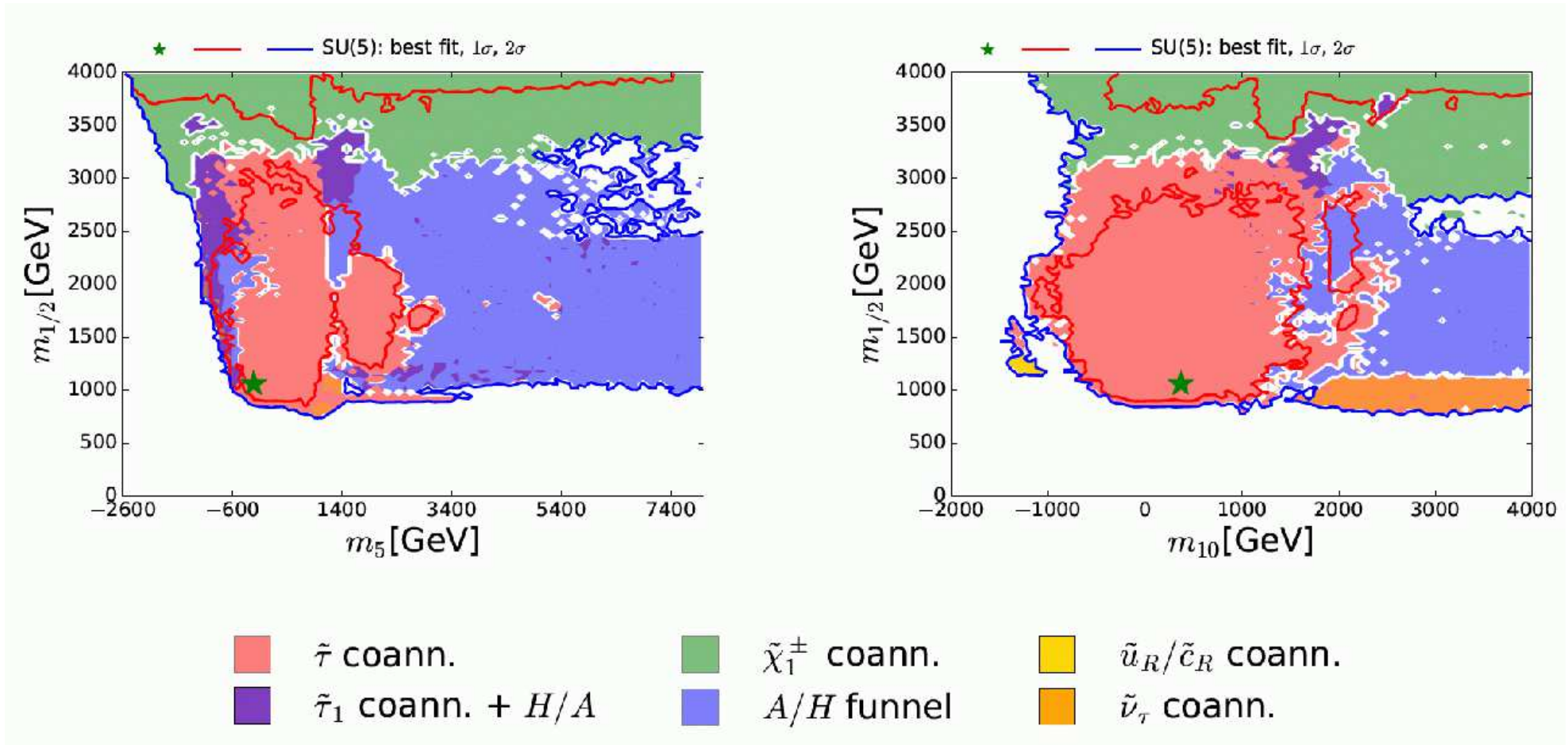


# Results in the SU(5)



Dark Matter annihilation mechanism:

[2016]

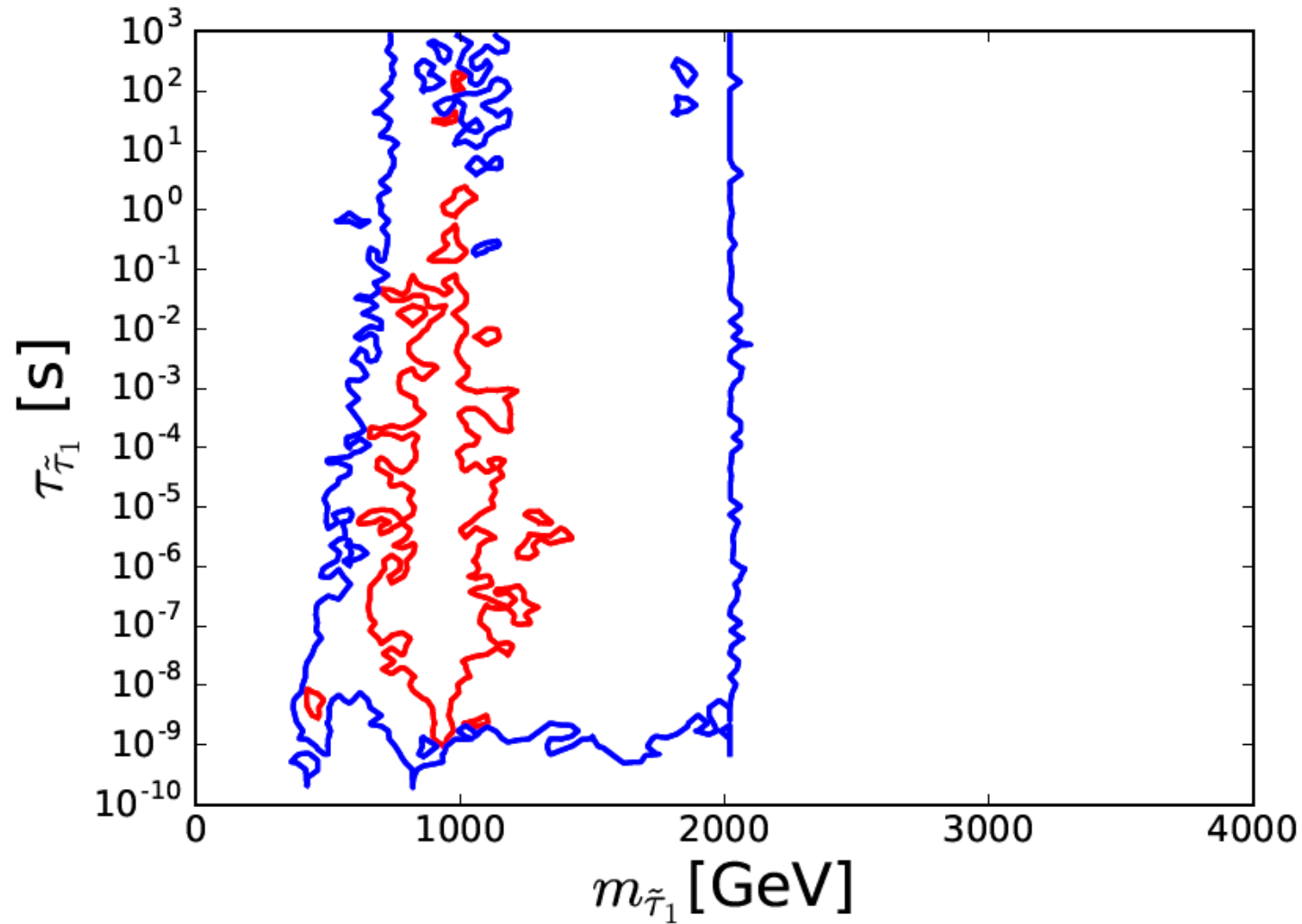


$\Rightarrow \tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$  co-ann. possible  $\Rightarrow$  but  $\tilde{\tau}_1$  co-ann. dominant!



# NLSP stau searches at the LHC:

[2016]



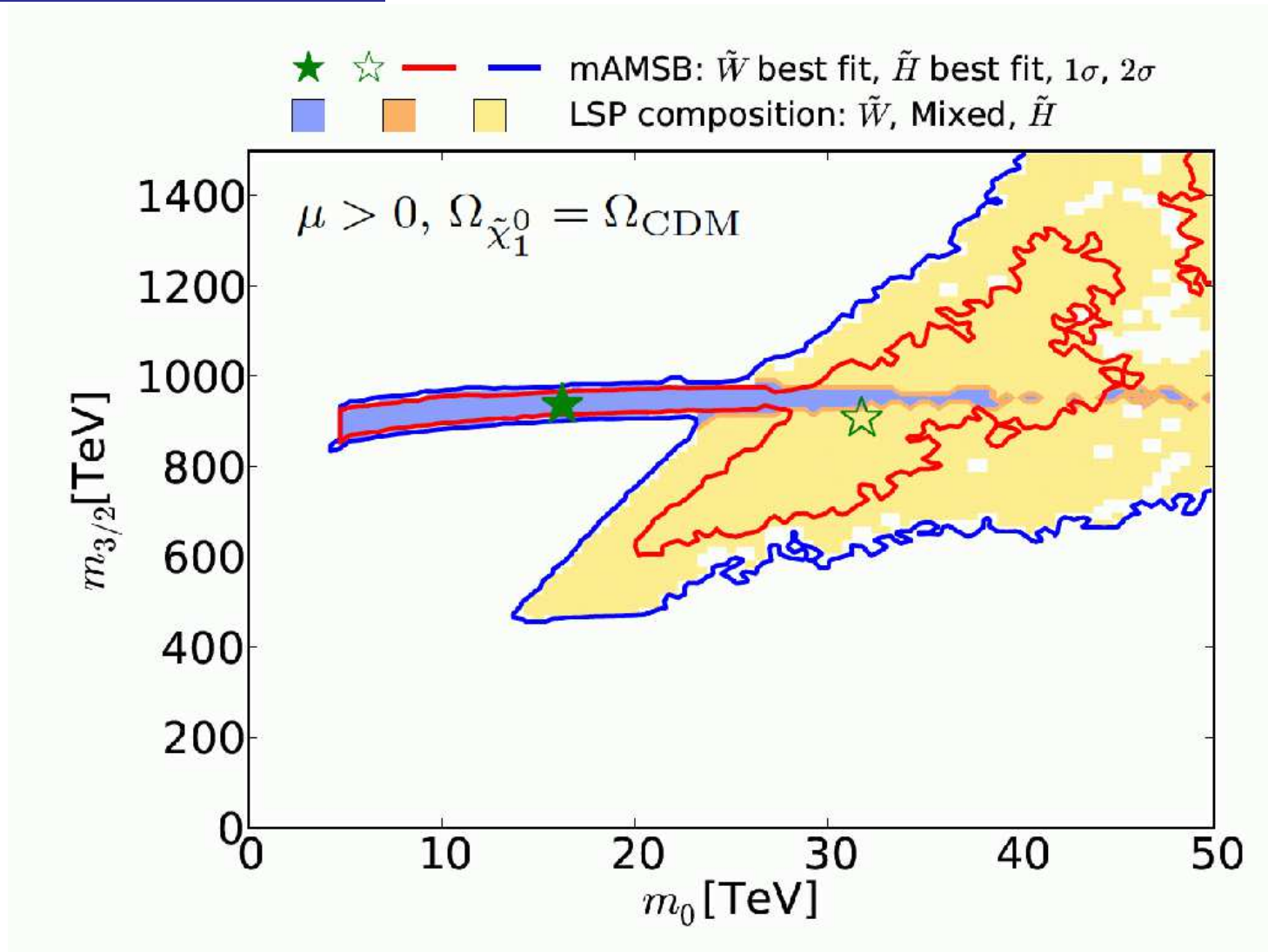
⇒ search for long-lived staus!

# Results in the mAMSB



## Dark Matter composition:

[2016]



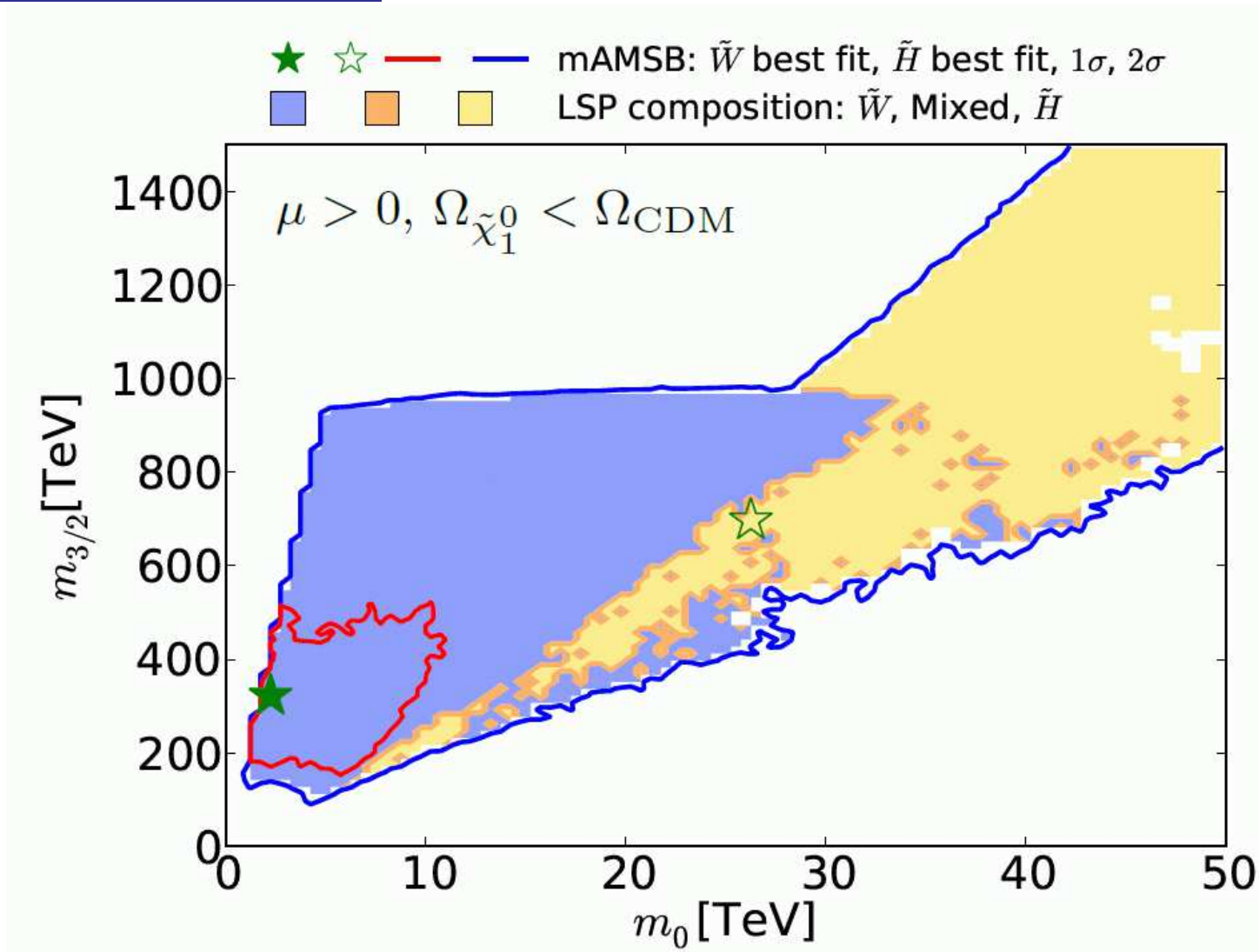
$\Rightarrow m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1$  TeV (wino),  $\sim 1.1 \pm 0.02$  TeV (higgsino)

# Results in the mAMSB



Dark Matter composition:

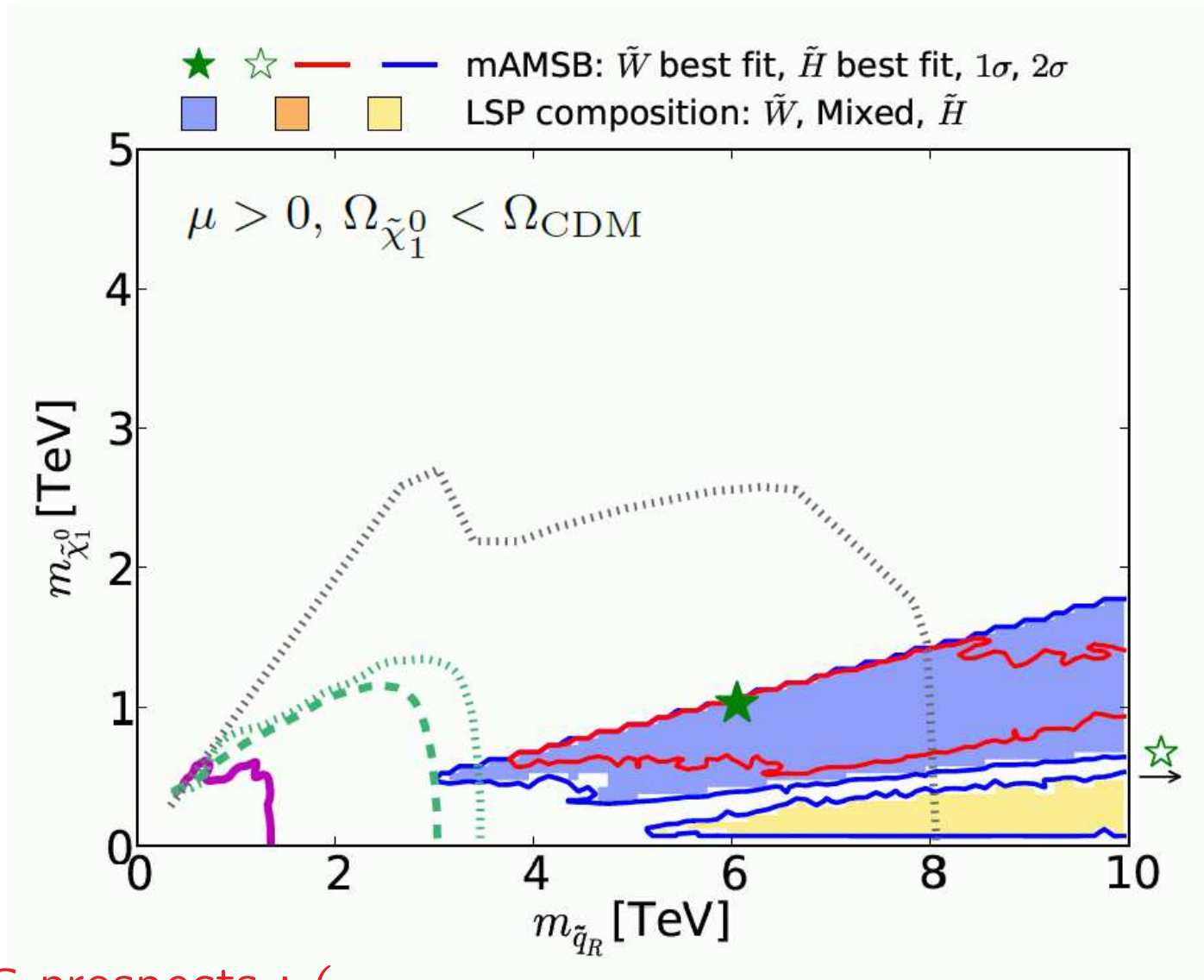
[2016]



⇒ very relaxed limits ⇒ lower masses

# Squark mass vs. DM mass:

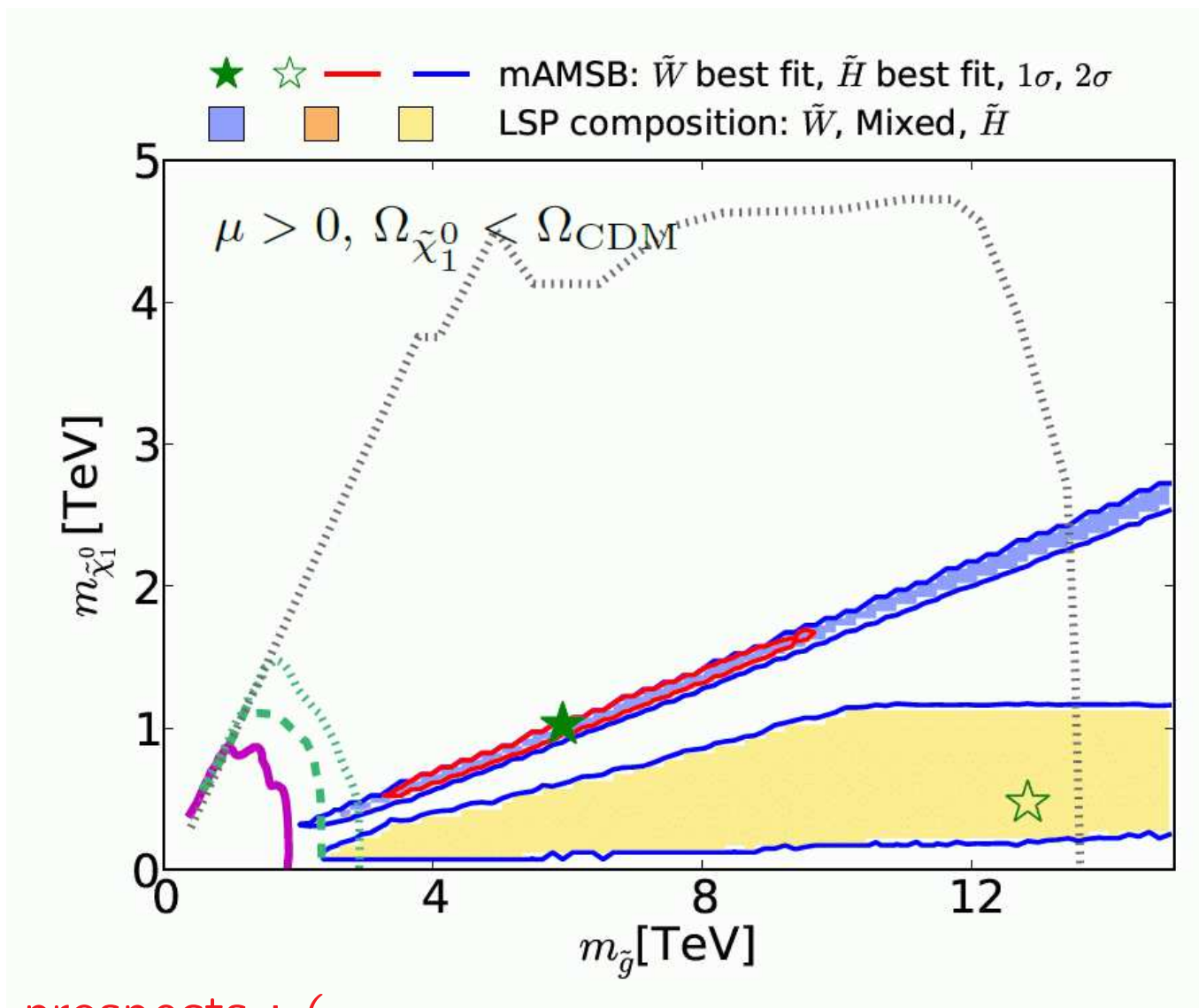
[2016]



⇒ bad LHC prospects :-(  
 ⇒ better FCC-hh prospects :-)

# Glino mass vs. DM mass:

[2016]



⇒ bad LHC prospects :-)

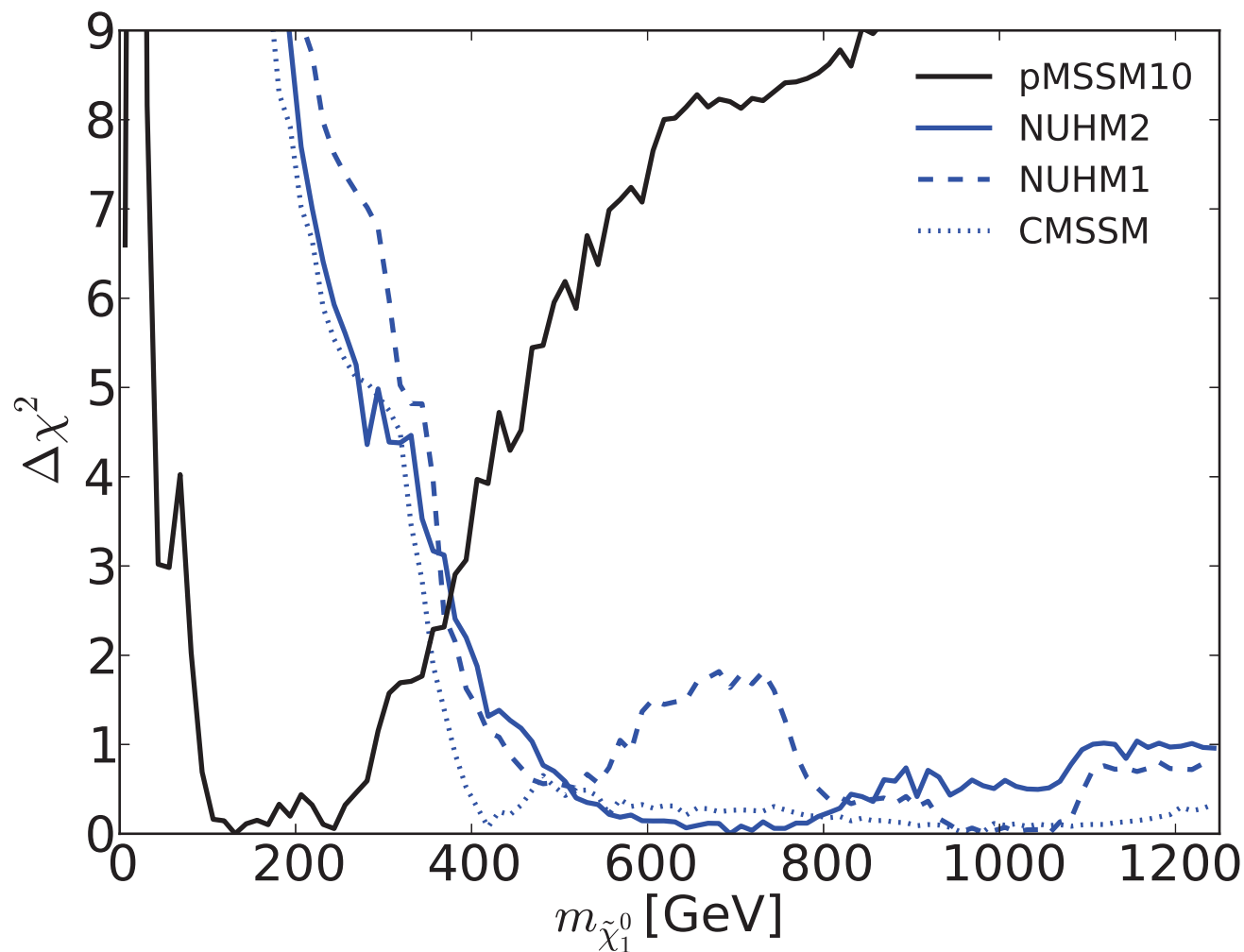
⇒ better FCC-hh prospects :-)

### 3. Results in low-energy models $\Rightarrow$ pMSSM10



DM mass: pMSSM10 vs. GUT based models prediction:

[2015]

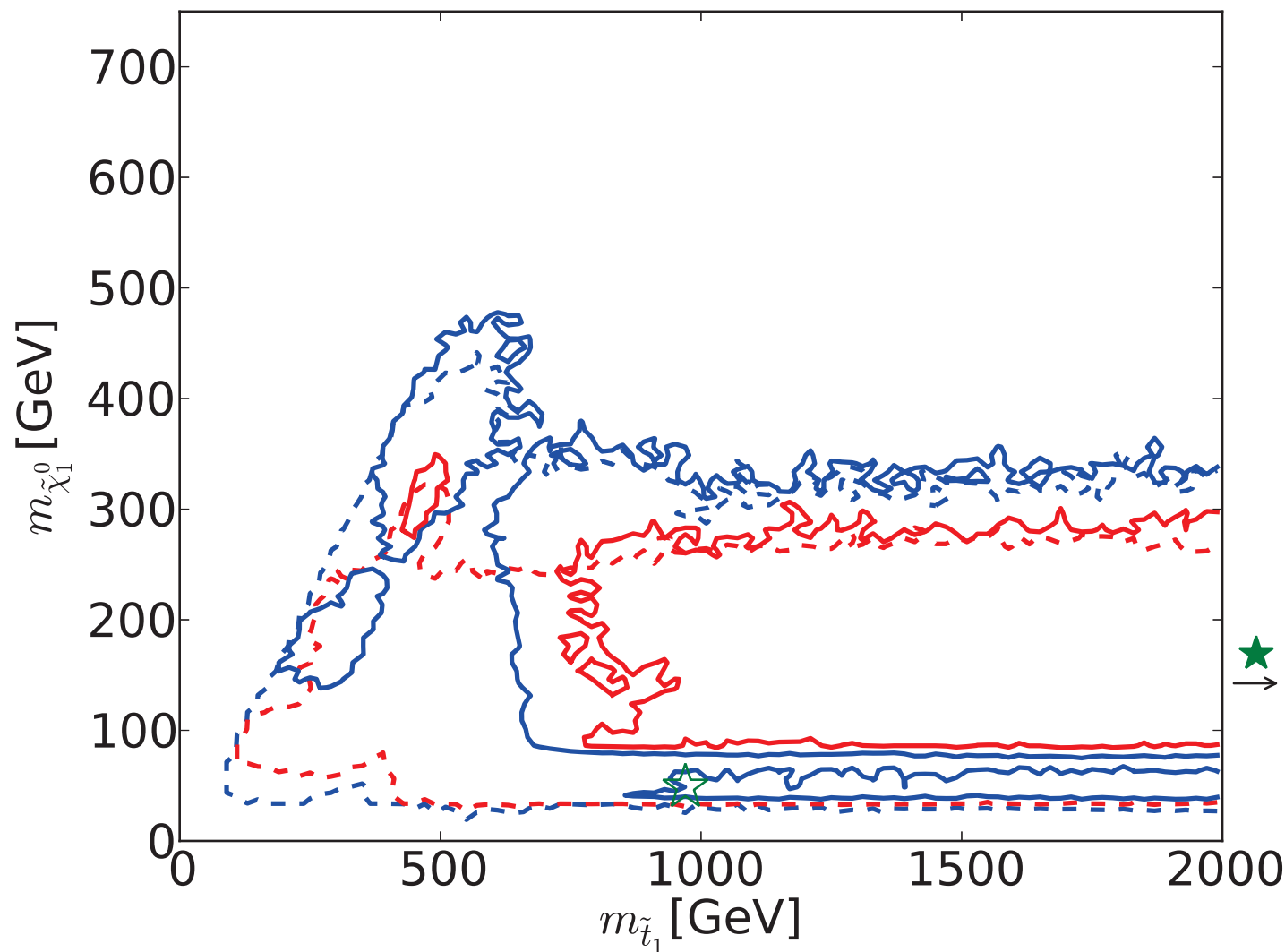


$\Rightarrow$  pMSSM10 predicts much lower DM mass than GUT-based models

pMSSM10 prediction: DM mass vs. light stop mass:

[2015]

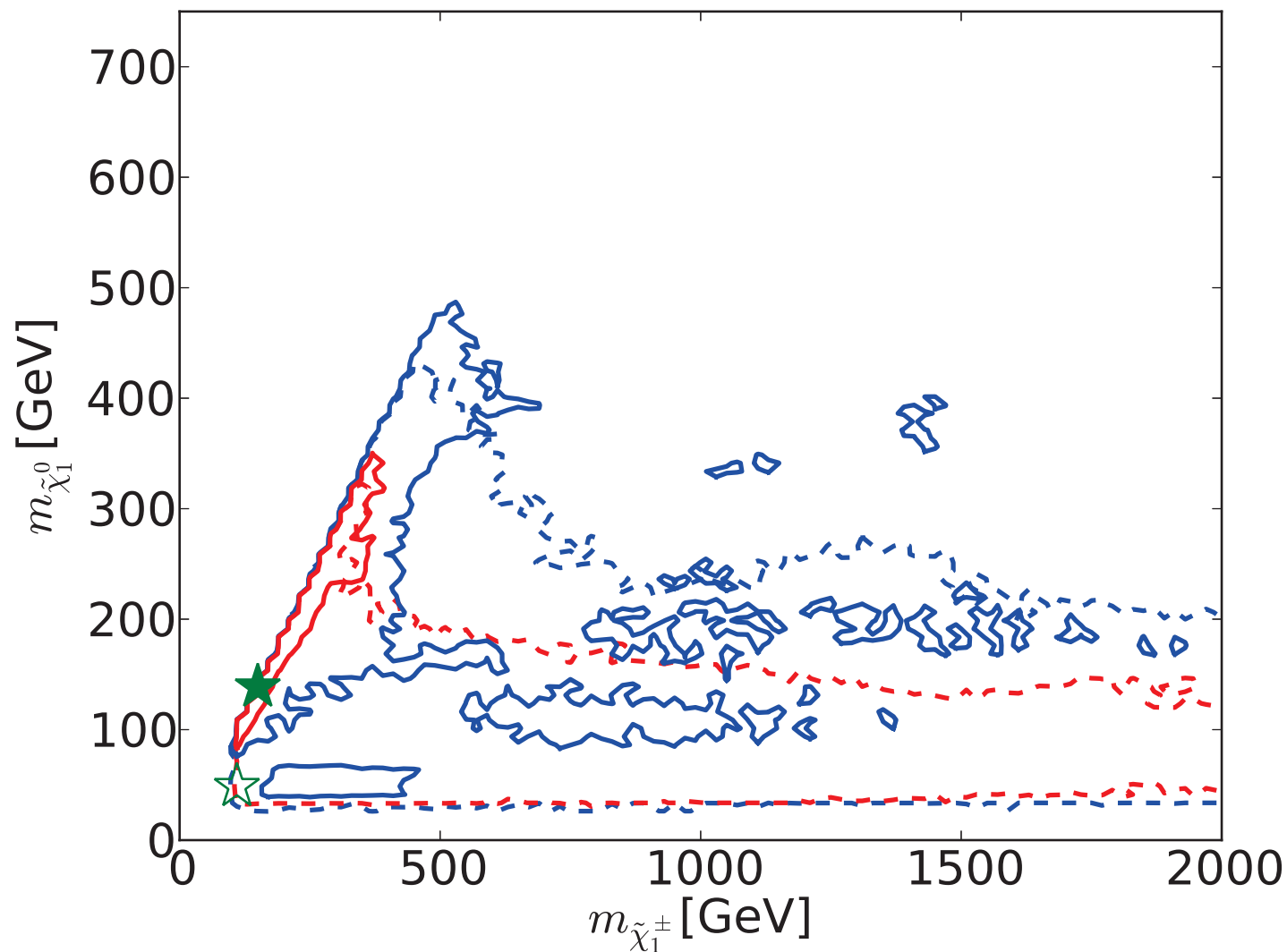
- ★ ——— pMSSM10 w LHC8: best fit,  $1\sigma$ ,  $2\sigma$
- ☆ - - - pMSSM10 w/o LHC8: best fit,  $1\sigma$ ,  $2\sigma$



# pMSSM10 prediction: DM mass vs. light stop mass:

[2015]

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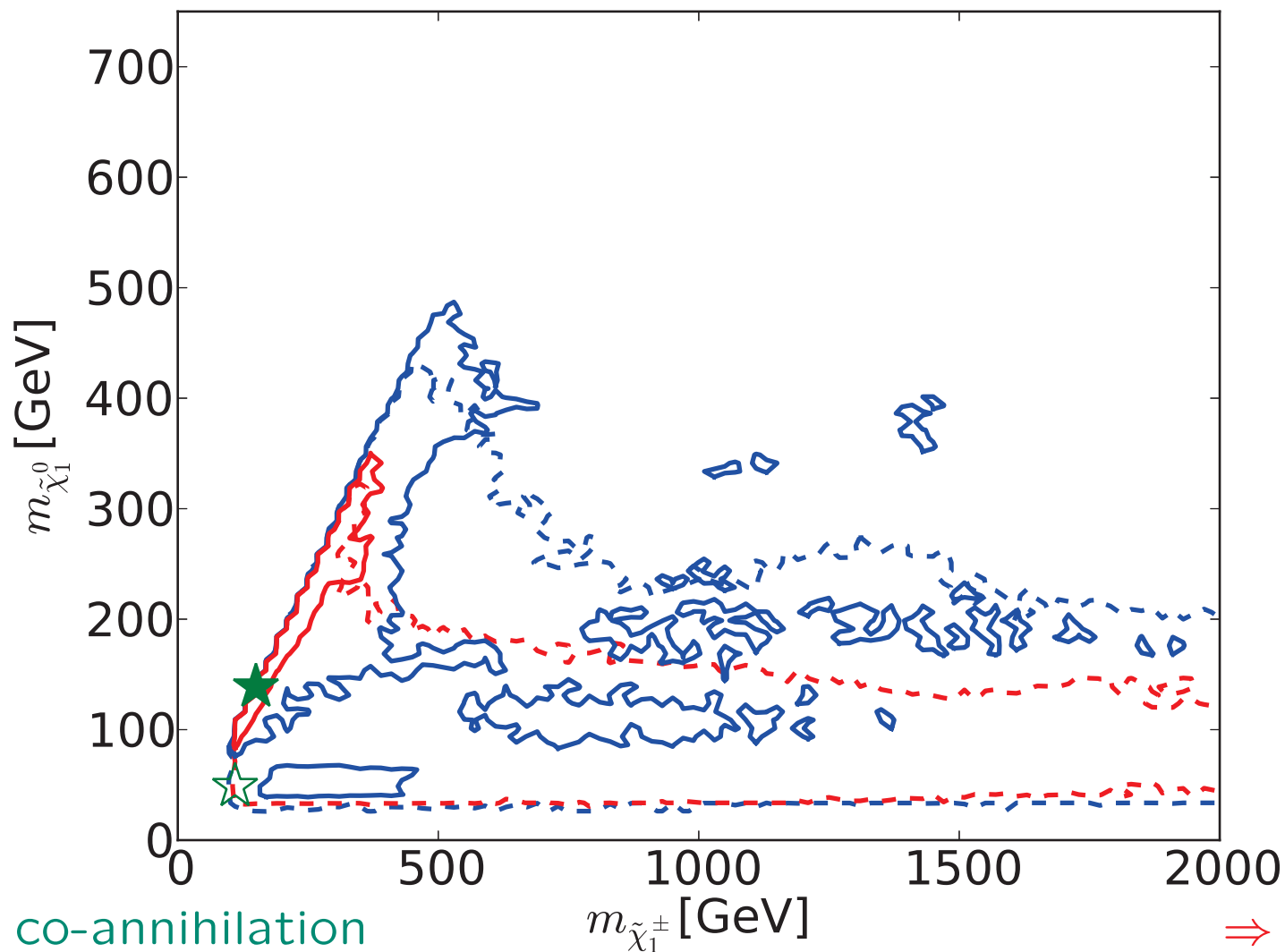




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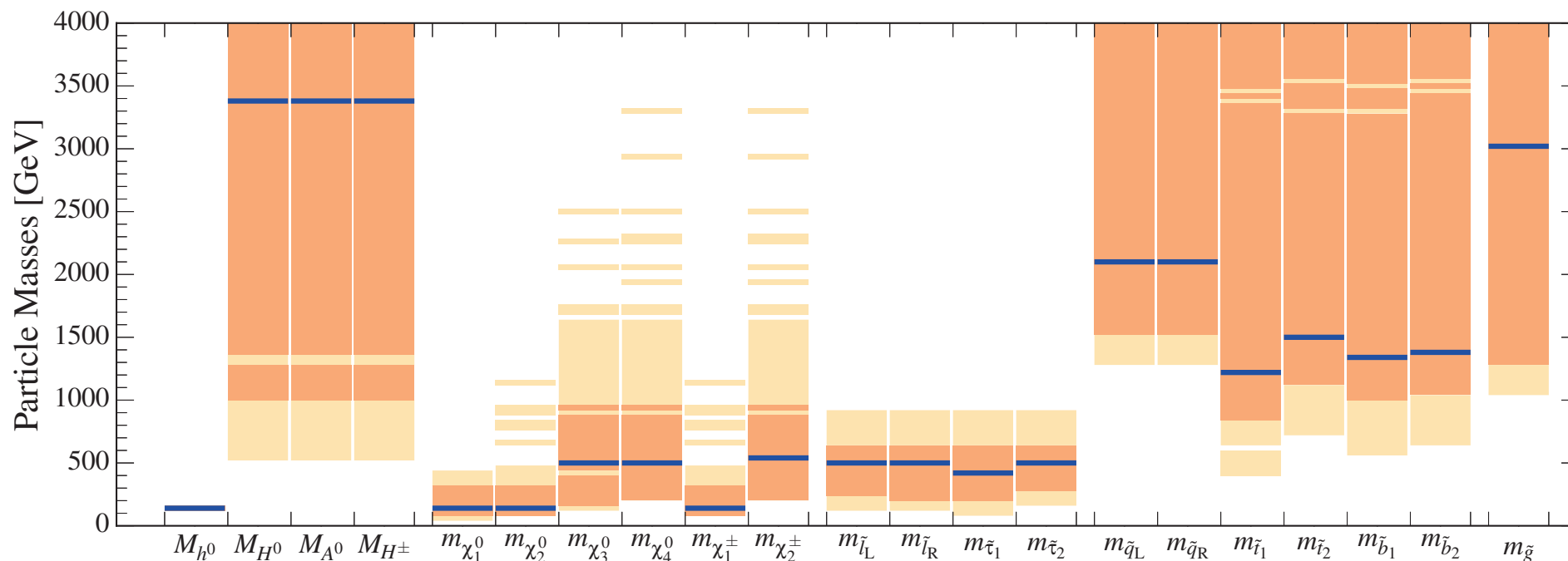
[2015]

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⇒ chargino co-annihilation

⇒  $M_1 \approx M_2$



⇒ high colored masses

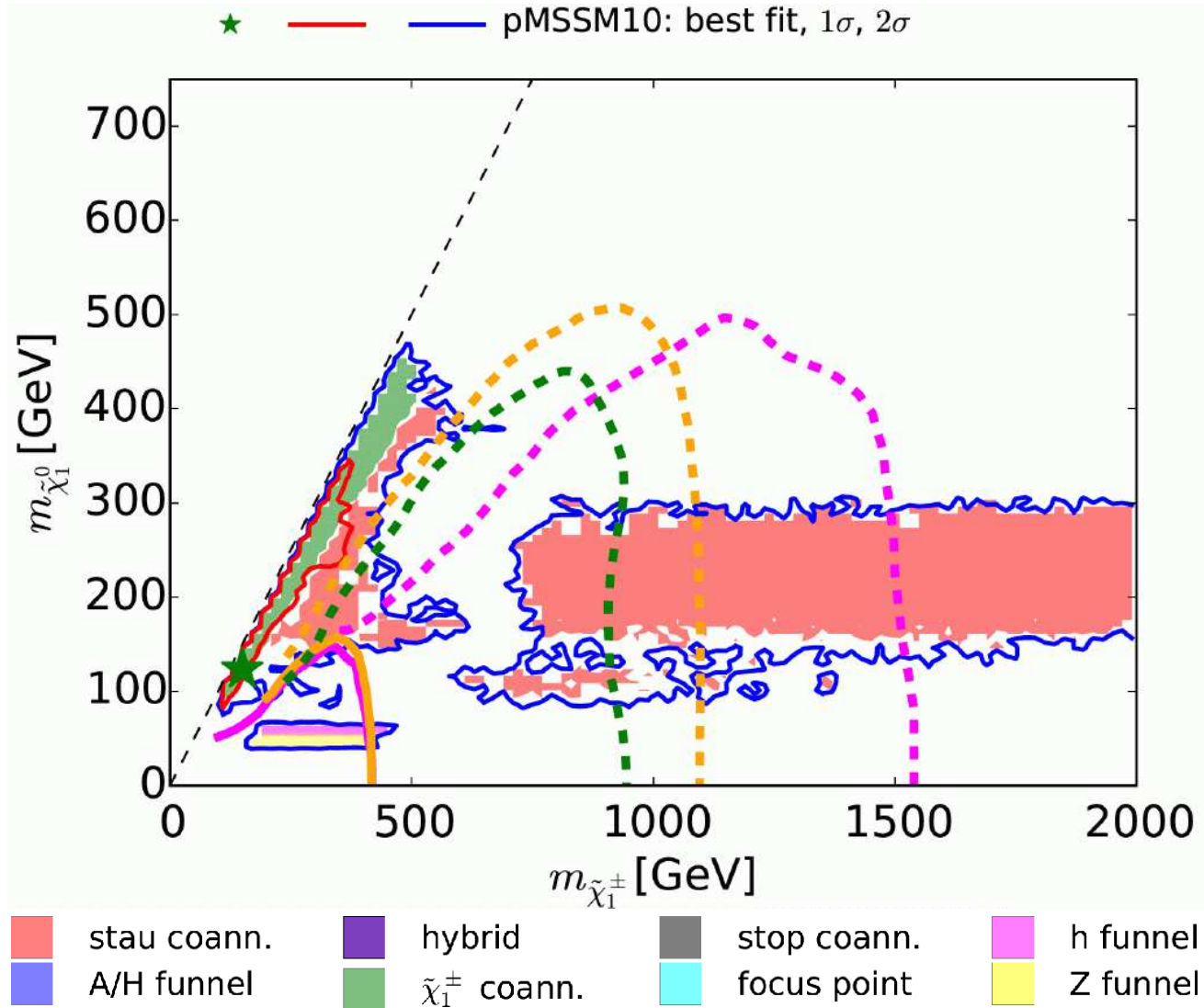
⇒ relatively low electroweak masses

partially with not too large ranges

⇒ clear prediction for  $m_{\tilde{\chi}_1^0}$

# LHC prospects for pMSSM10:

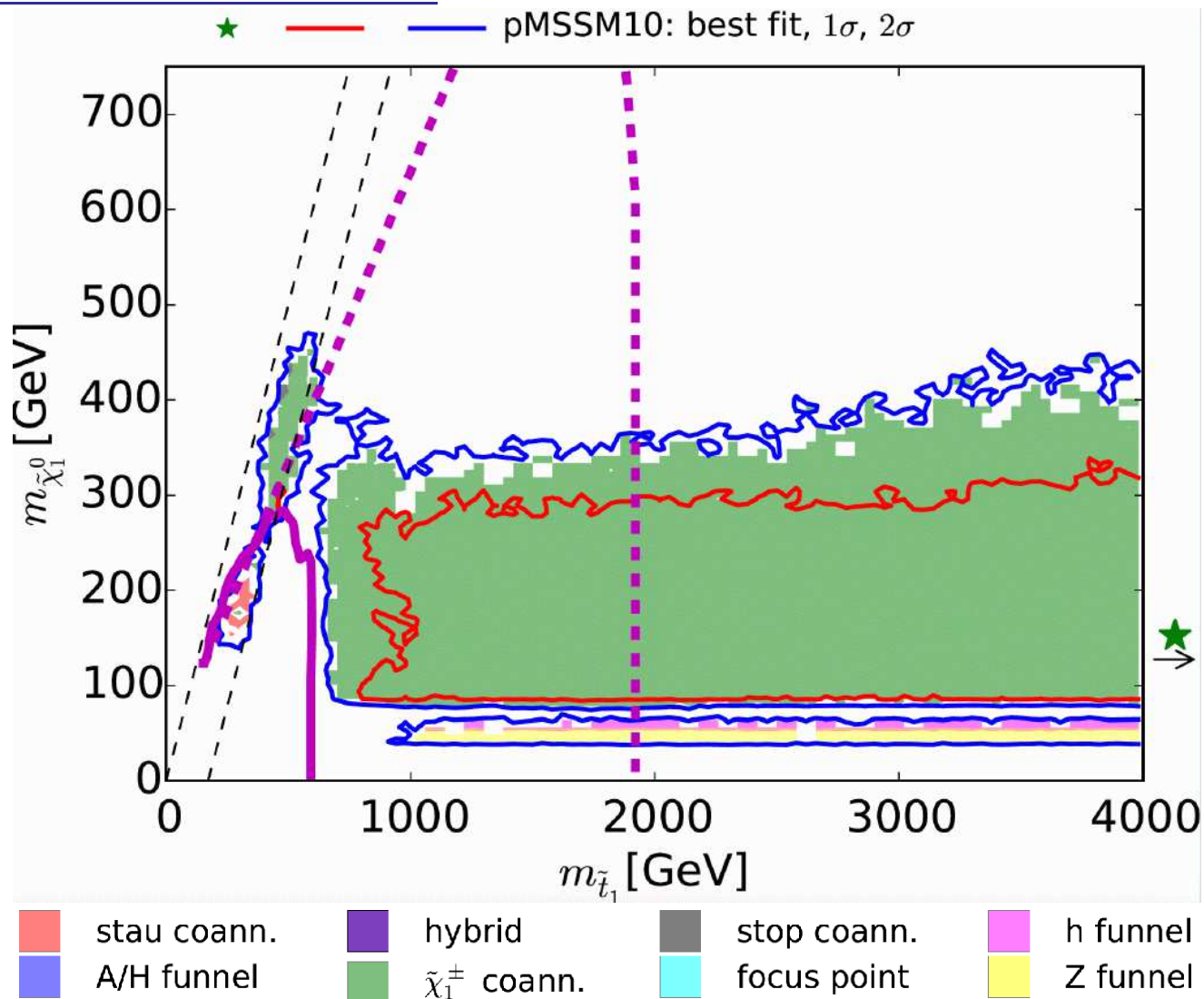
[2015]



solid: current LHC limits, dashed: HL-LHC prospects  
 $\Rightarrow$  best-fit regions not covered! (in EW searches)

# LHC prospects for pMSSM10:

[2015]



solid: current LHC limits, dashed: HL-LHC prospects  
 $\Rightarrow$  best-fit regions can partially be covered! (in colored searches)

## 4. Conclusinos

- **SUSY** is (still) the best-motivated BSM scenario
  - constrained models: **CMSSM, NUHM1, NUHM2, SU(5), mAMSB**
  - general models: **pMSSM10, ...**
- Our tool: **MasterCode**: combination of **LHC searches, Higgs measurements, EWPO, BPO, CDM**  $\Rightarrow \chi^2$  evaluation
- CMSSM, NUHM1, NUHM2:
  - some accessible spectra
  - many inaccessible spectra
  - CMSSM best-fit points  $\rightarrow$  HL-LHC
- SU(5):
  - some accessible spectra, many inaccessible spectra
  - long-lived staus ?!
- mAMSB:
  - $\Omega_{\tilde{\chi}_1^0} = \Omega_{\text{CDM}} \Rightarrow$  mostly inaccessible spectra
  - $\Omega_{\tilde{\chi}_1^0} < \Omega_{\text{CDM}} \Rightarrow$  lower masses allowed
  - still bad HL-LHC prospects, but good FCC-hh prospects
- pMSSM10:
  - $m_{\tilde{\chi}_1^0} \lesssim 400$  GeV; important: chargino co-annihilation
  - colored searches cover parts of 68% CL region
  - EW searches miss 68% CL region (compr. spectrum!)

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  - general models: **pMSSM10, ...**
- Our tool: **MasterCode**: combination of LHC searches, Higgs measurements, EWPO, BPO, CDM  $\Rightarrow \chi^2$  evaluation
- CMSSM, NUHM1, NUHM2:
  - $m_{\tilde{\chi}_1^0} \gtrsim 400$  GeV neutrino floor
  - best-fit regions (mostly) **above/at/below**
- SU(5):
  - stau co-ann., but also  $\tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$  co-ann. possible
  - $m_{\tilde{\chi}_1^0}$  as in **CMSSM, NUHM1, NUHM2**
  - best-fit region **at neutrino floor**
- mAMSB:
  - $m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1$  TeV (wino),  $\sim 1.1 \pm 0.02$  TeV (higgsino)
  - DD: wino **at neutrino floor**, higgsino **tested by next round**
- pMSSM10:
  - $m_{\tilde{\chi}_1^0} \lesssim 400$  GeV; important: chargino co-annihilation
  - at the 68% CL accessible at the **next generation of DD**
  - at the 95% CL even **below neutrino floor**

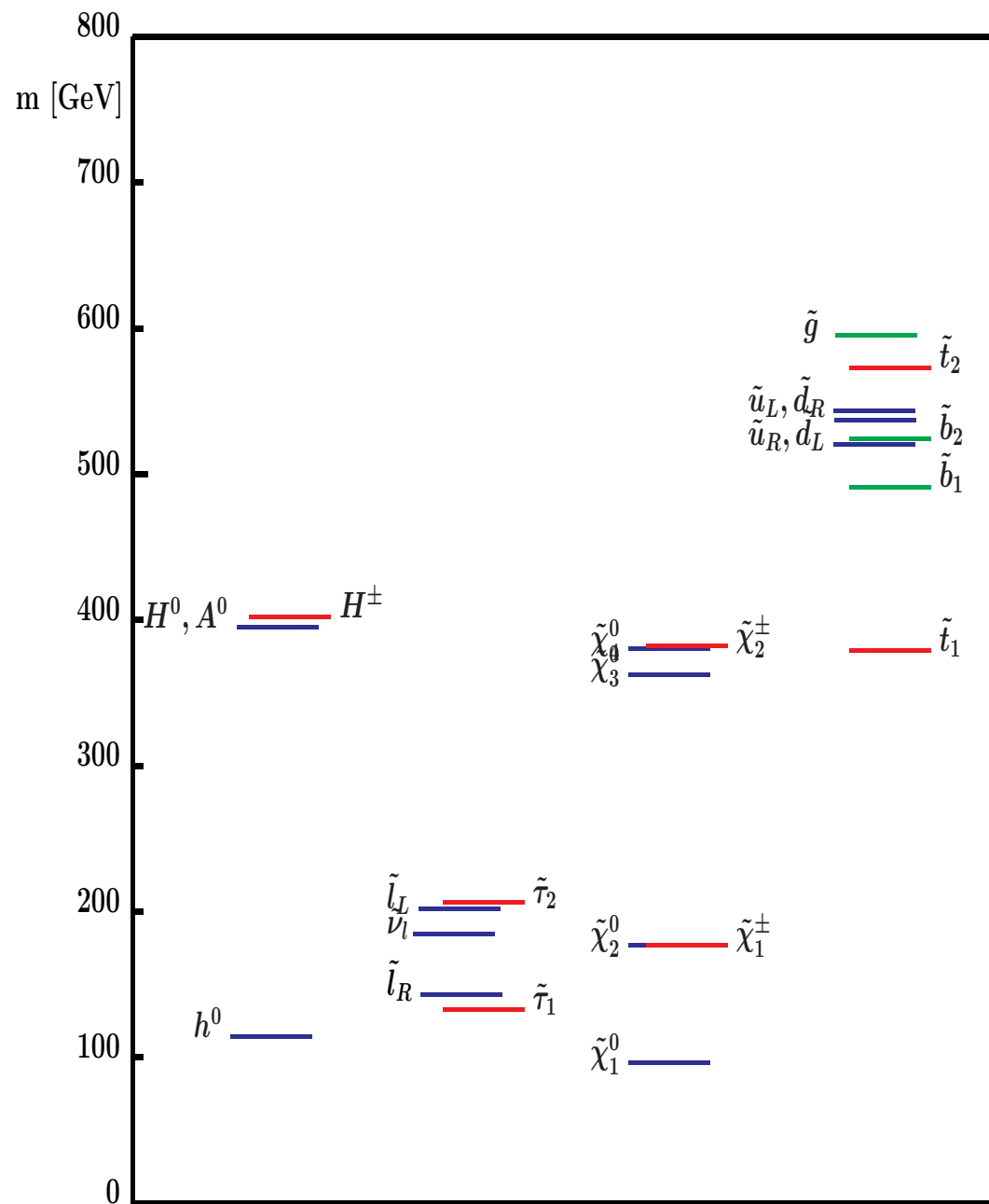


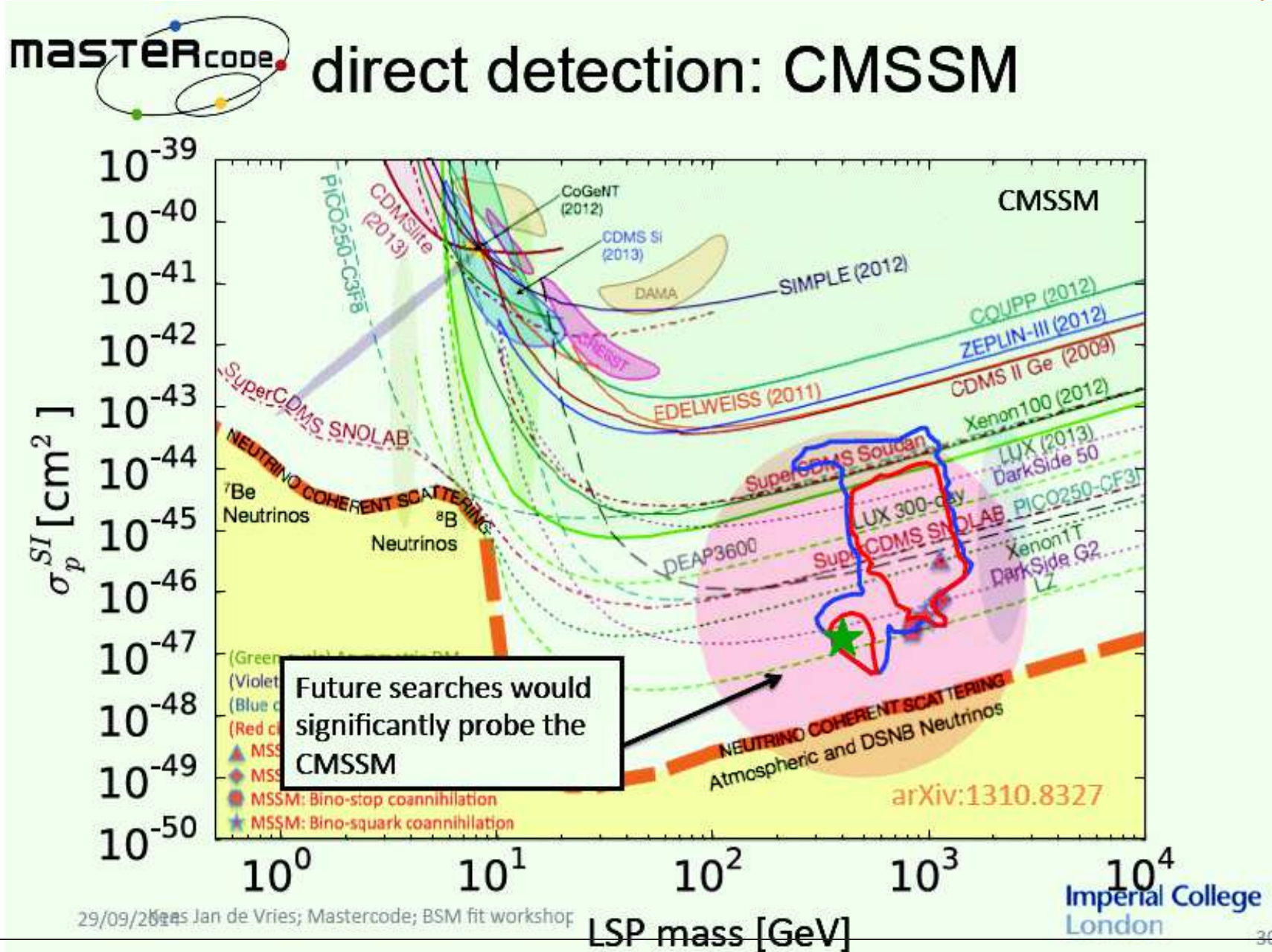
Further Questions?



“Typical” CMSSM scenario  
 (SPS 1a benchmark scenario):

Strong connection between  
 all the sectors





29/09/2015 Jan de Vries; Mastercode; BSM fit workshop

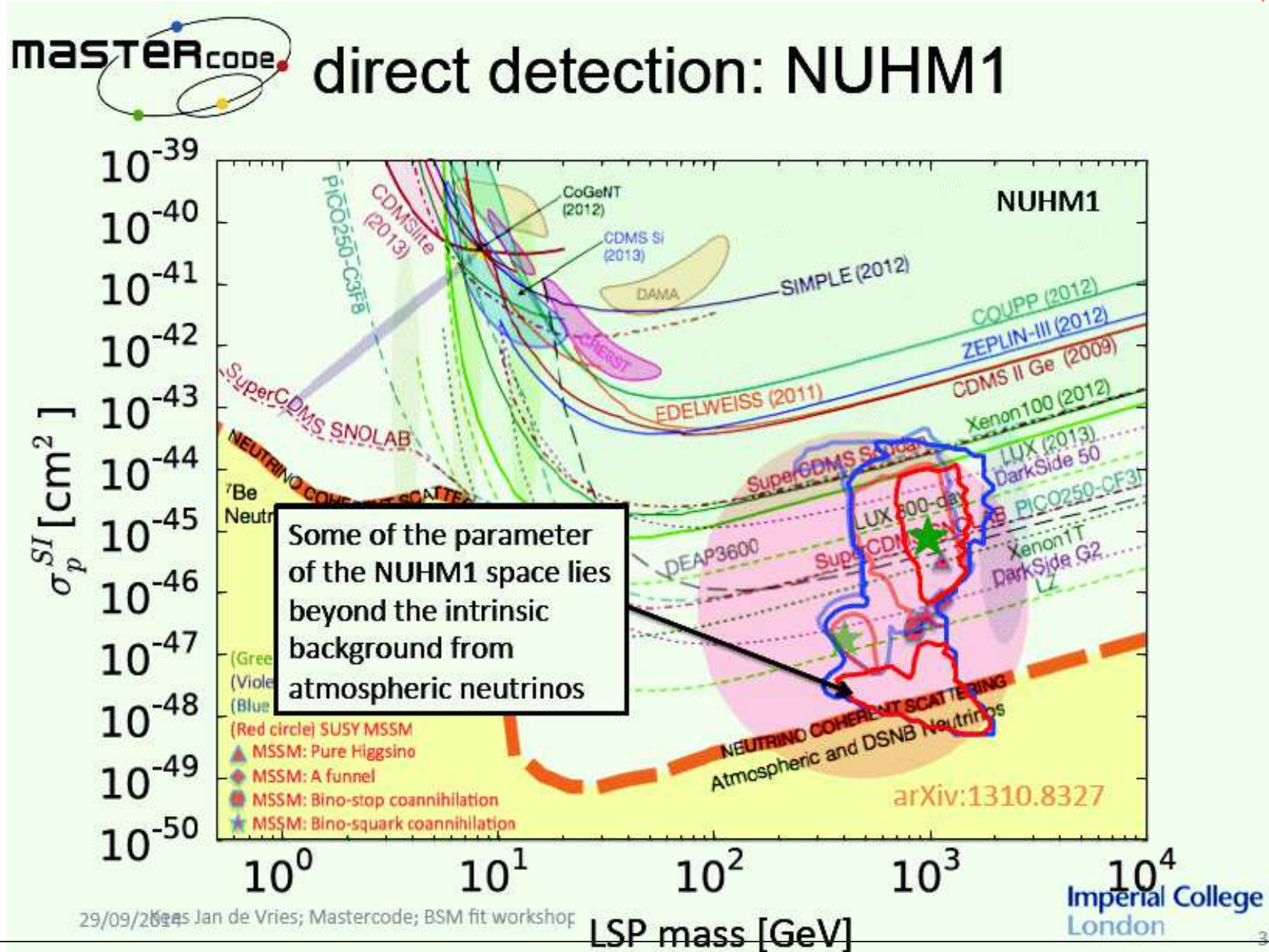
LSP mass [GeV]

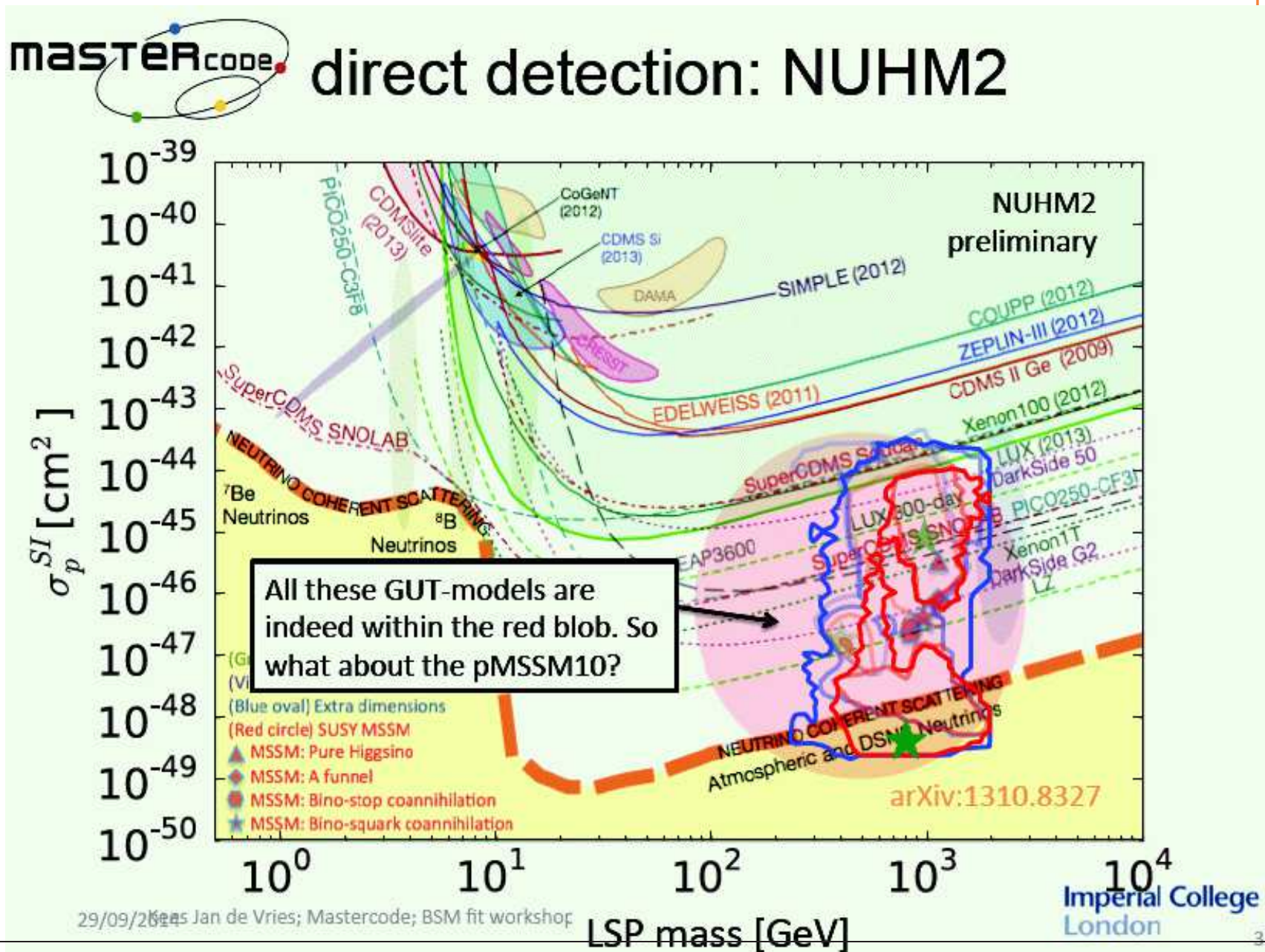
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[?2014]

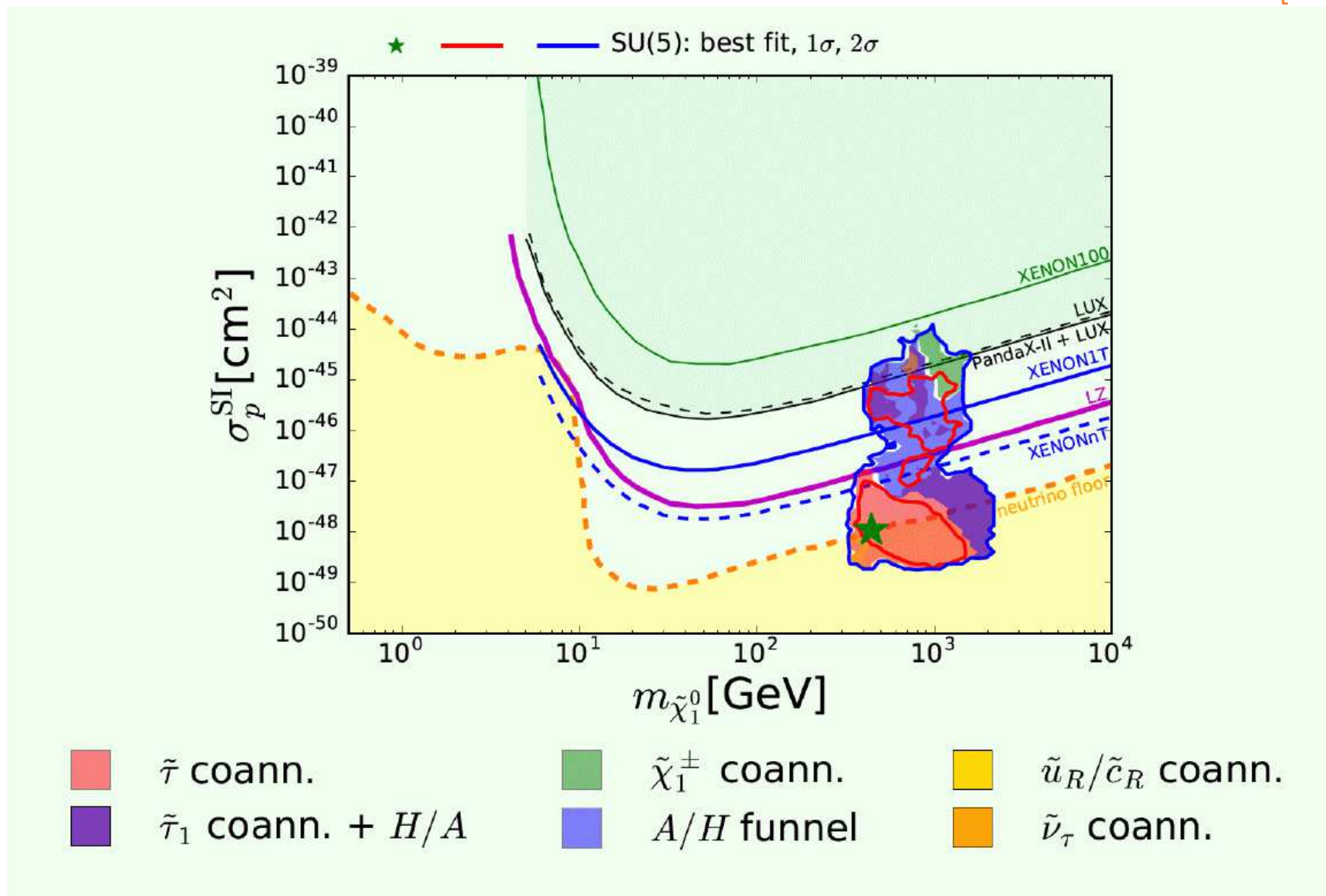






# Dark Matter Direct Detection prospects in SU(5):

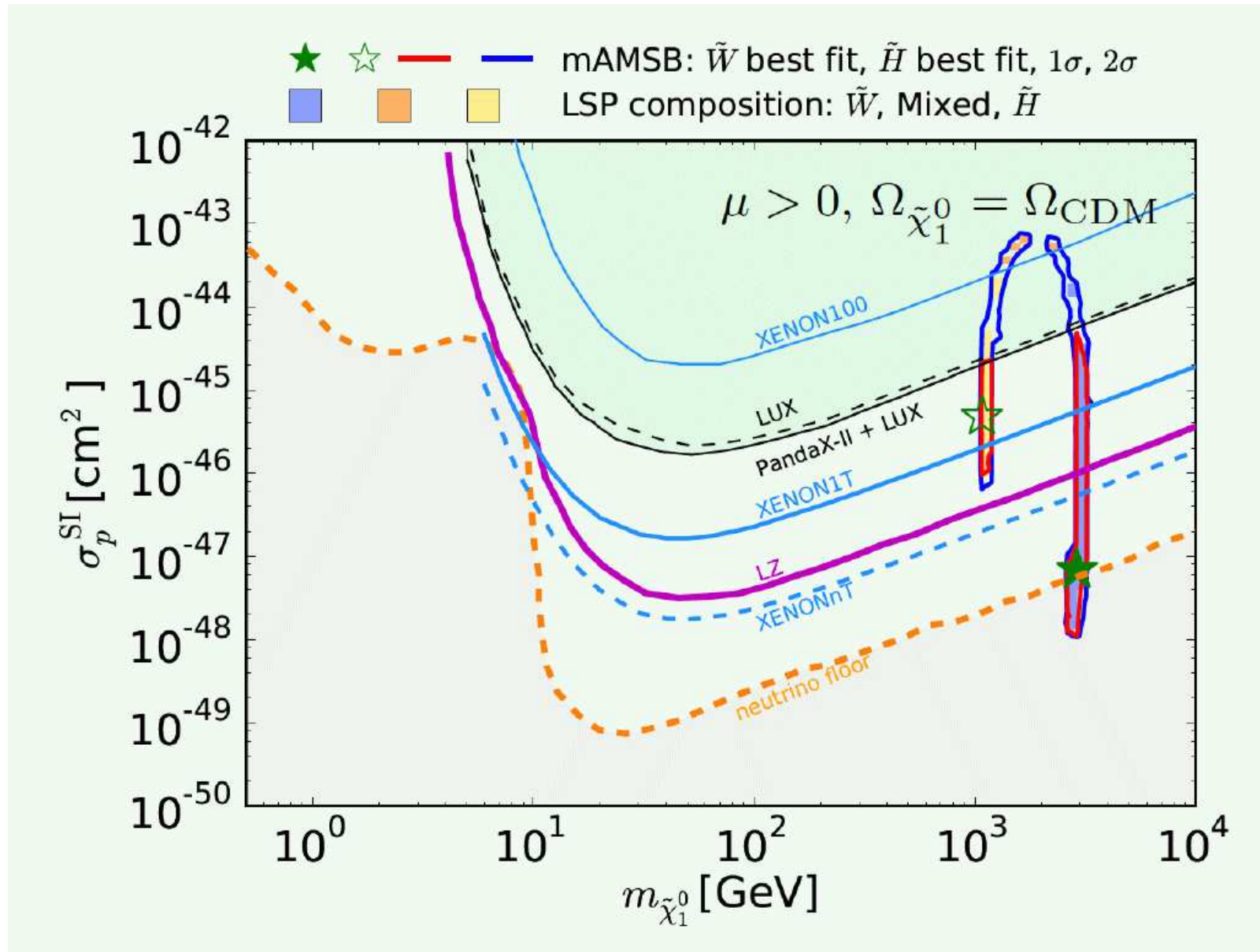
[2016]





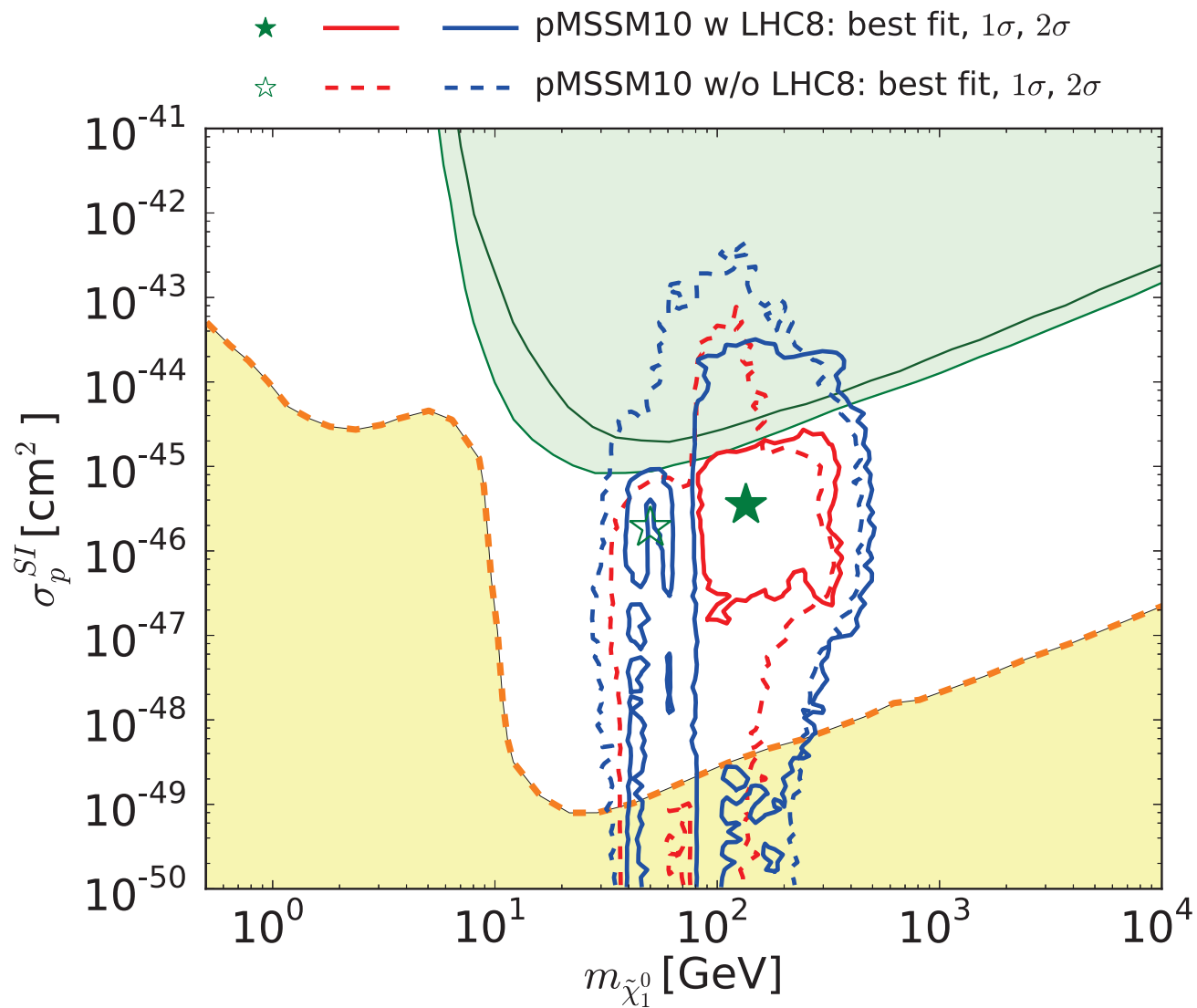
# Dark Matter Direct Detection prospects in mAMSB:

[2016]



pMSSM10 prediction:  $m_{\tilde{\chi}_1^0}$  vs.  $\sigma_p^{SI}$ :

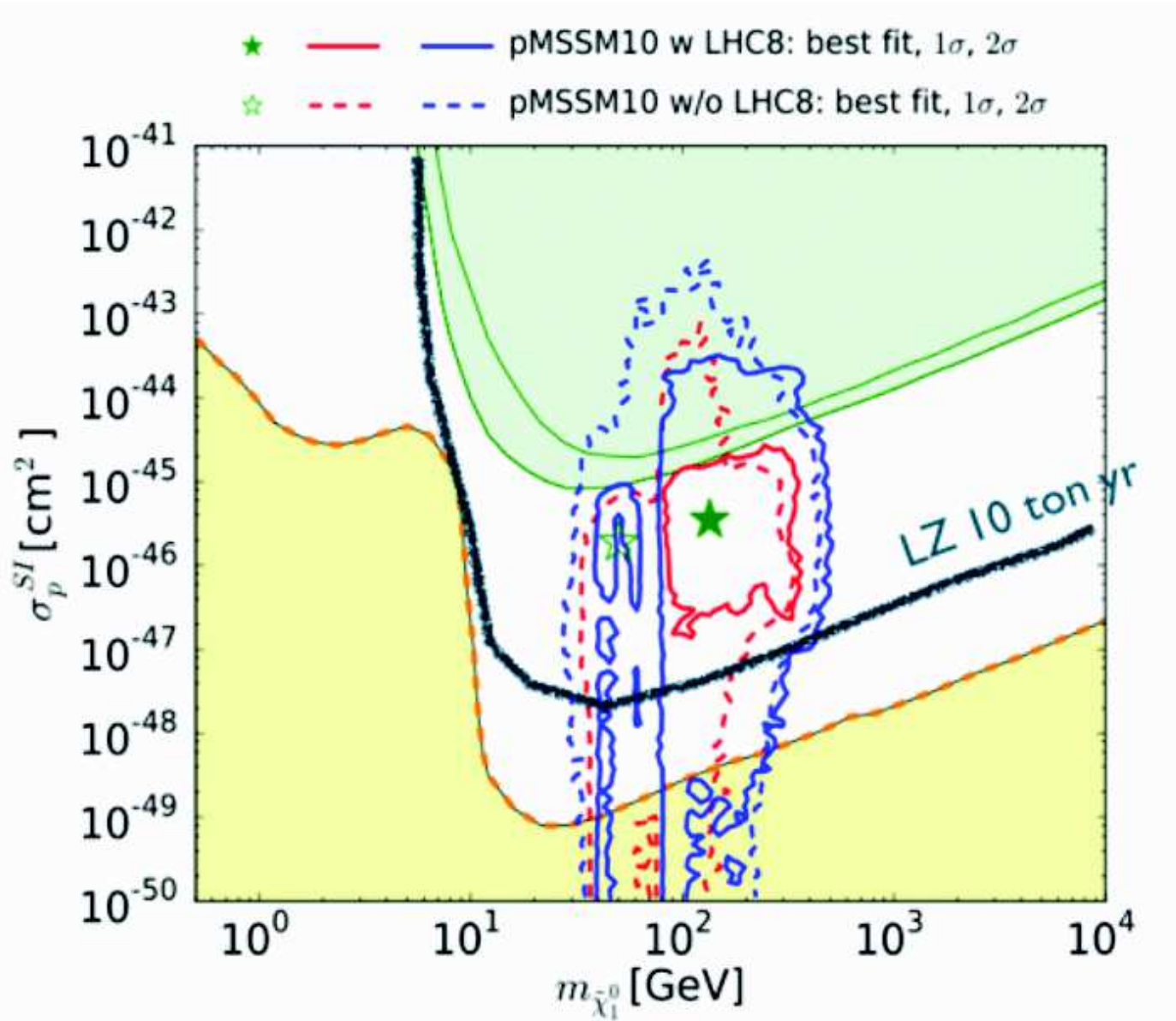
[2015]



⇒ LHC bounds try to “rescue” DD experiments!

pMSSM10 prediction:  $m_{\tilde{\chi}_1^0}$  vs.  $\sigma_p^{SI}$ : future expectations

[2015]



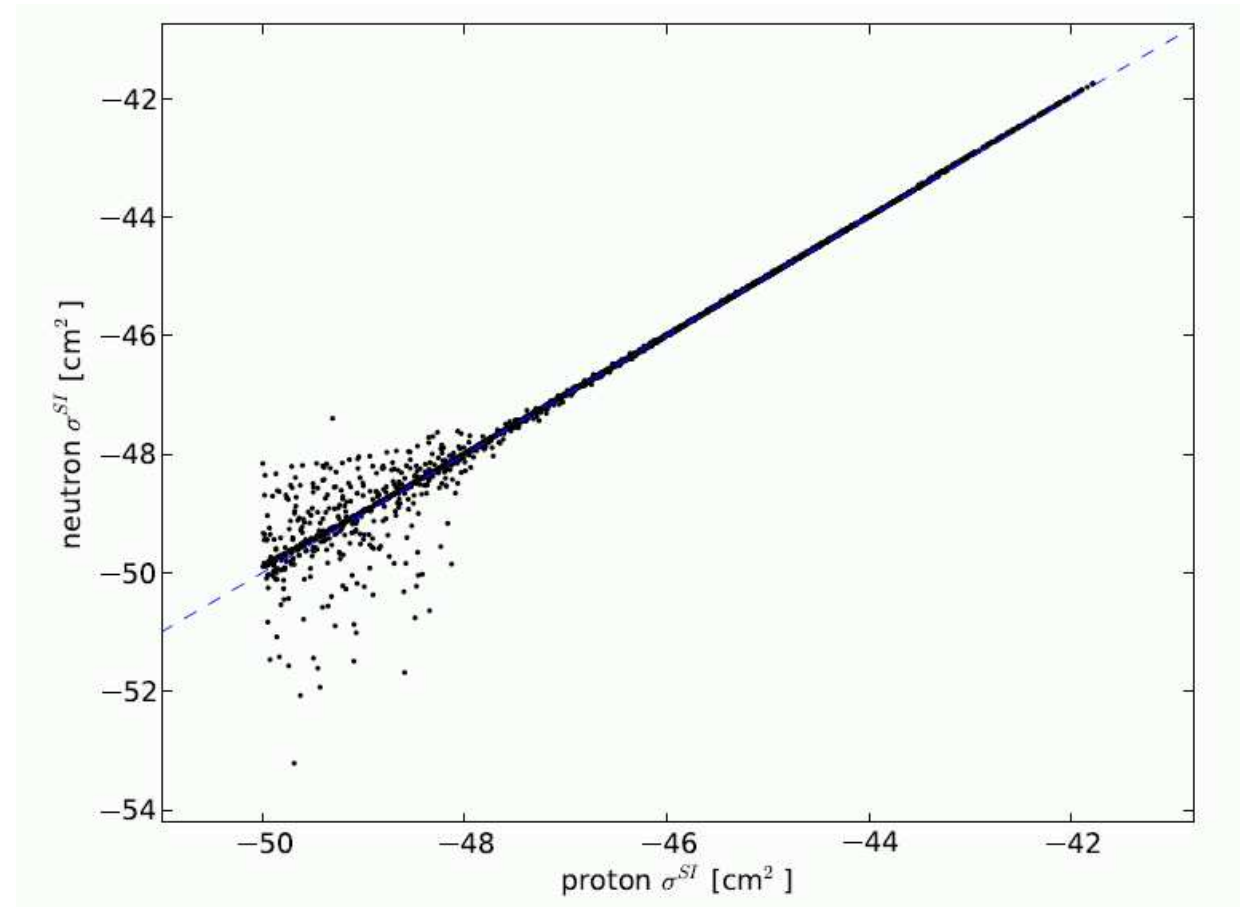
⇒ 68% CL areas covered by next round of DD experiments

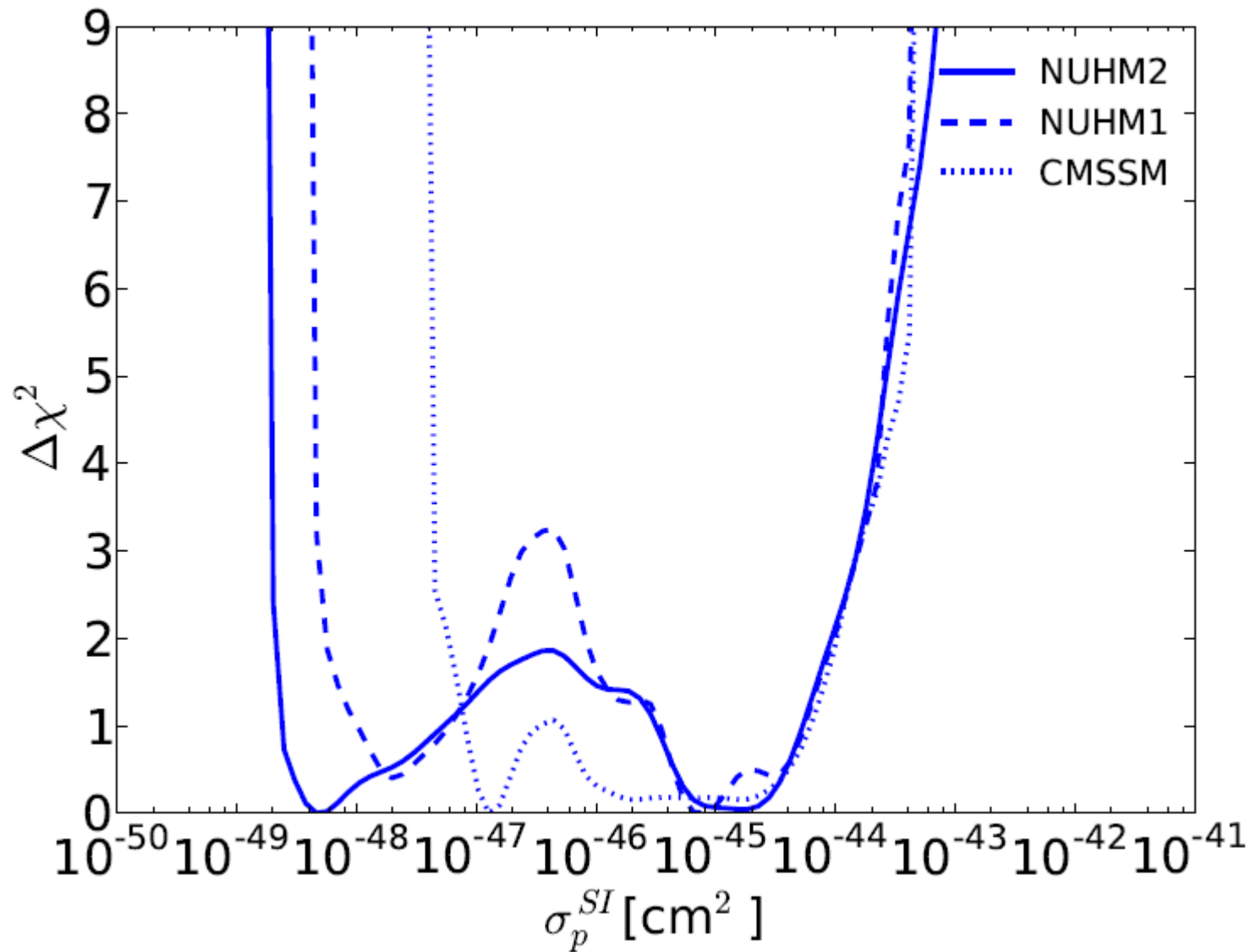
$\sigma_p^{SI}$  is evaluated for  
 $p$ -scattering

Can  $n$ -scattering come  
to rescue?

Some points with low  $\sigma_p^{SI}$   
have even lower  $\sigma_n^{SI}$

⇒ no “no-lose theorem”  
for DD experiments!



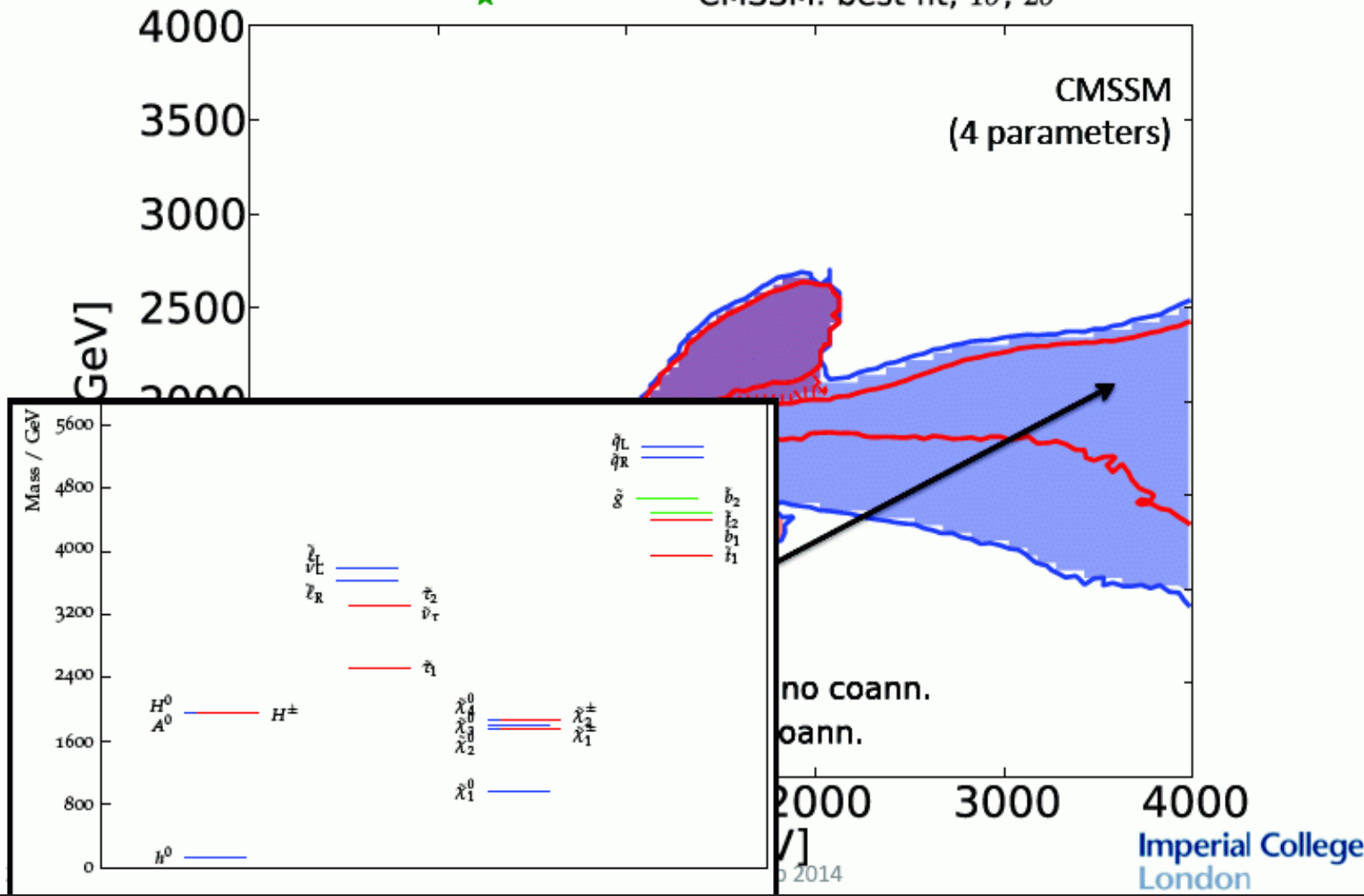


⇒ only very small values are favored



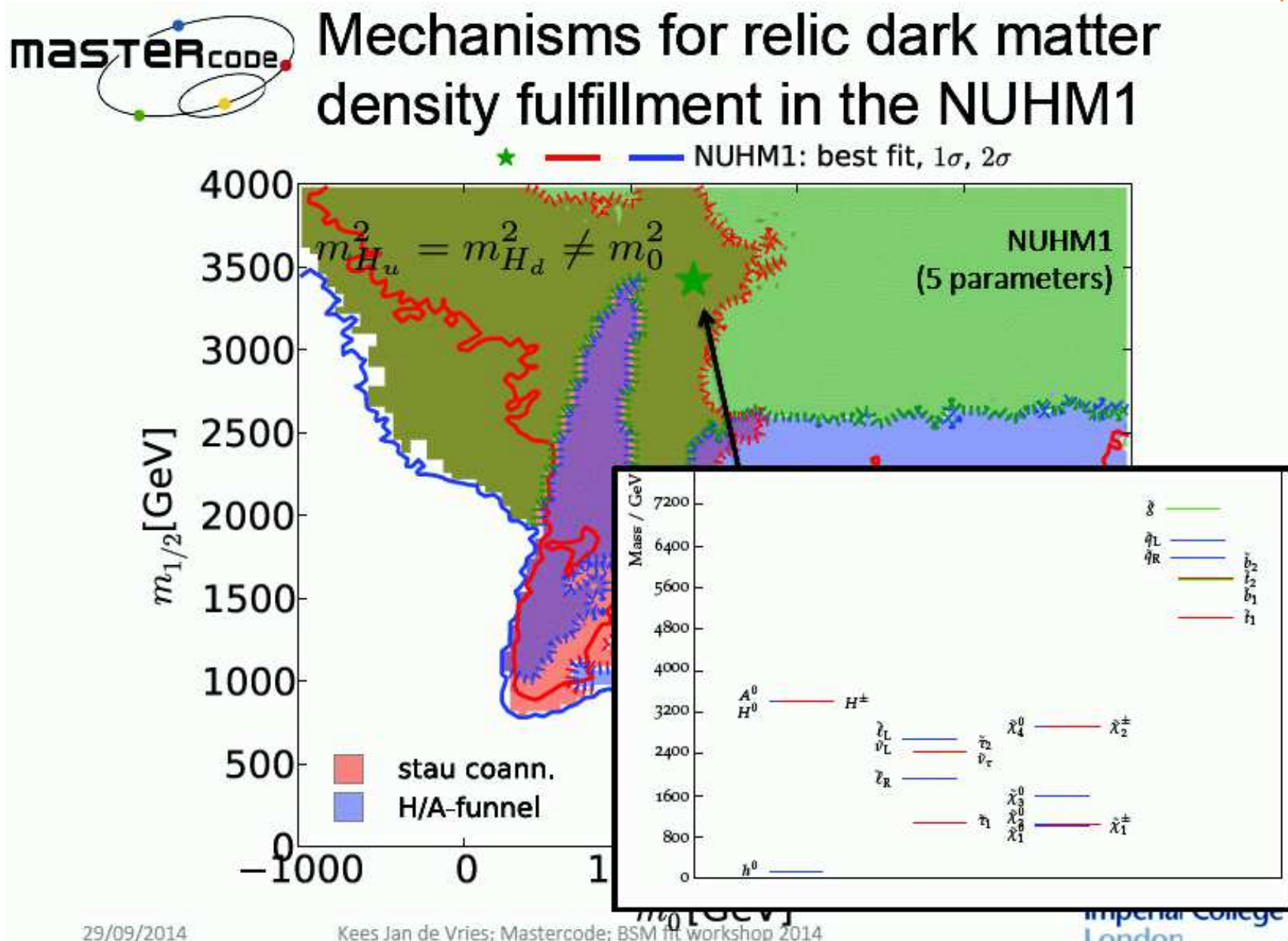
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★ — CMSSM: best fit, 1σ, 2σ

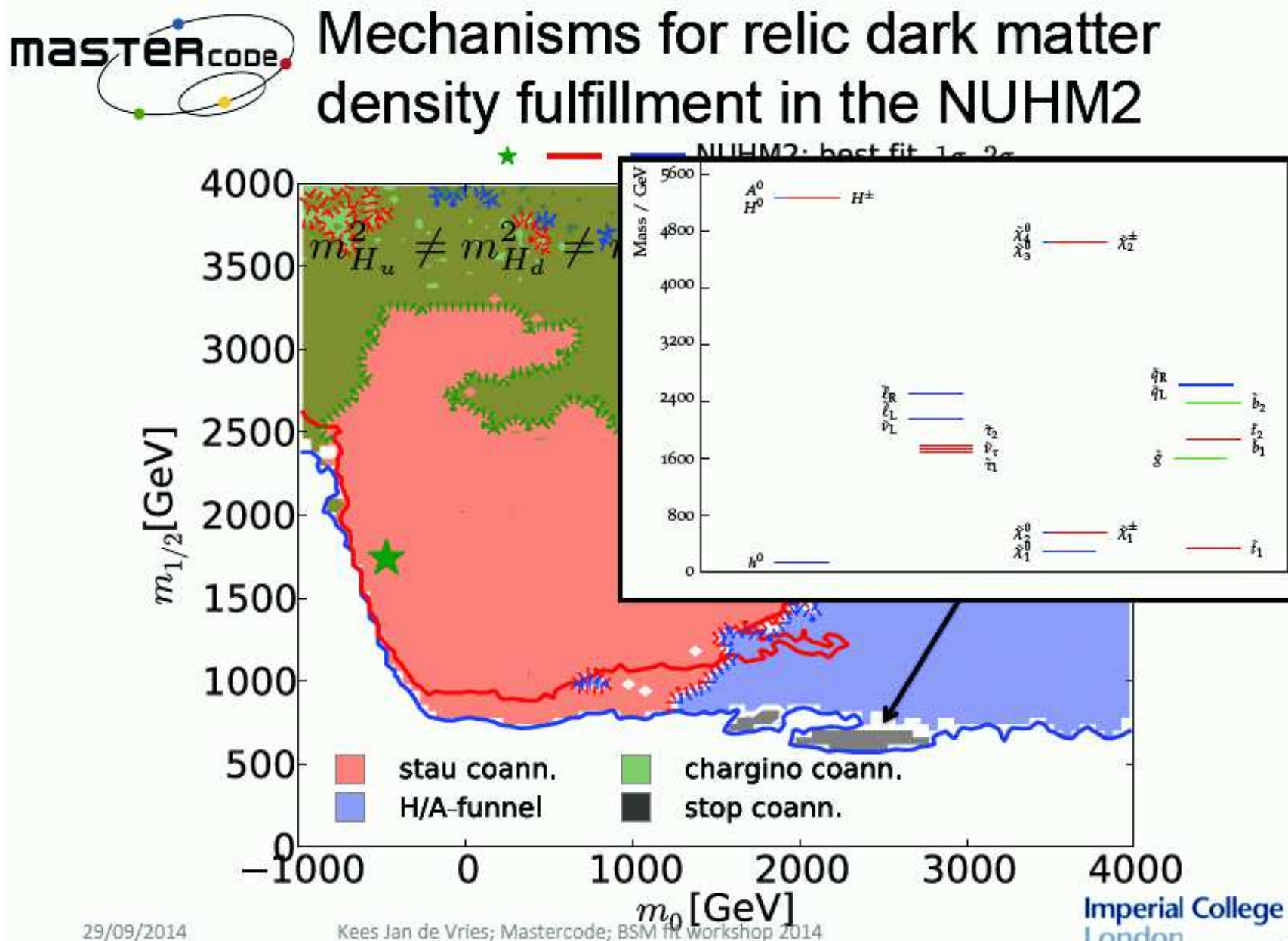




[2014]

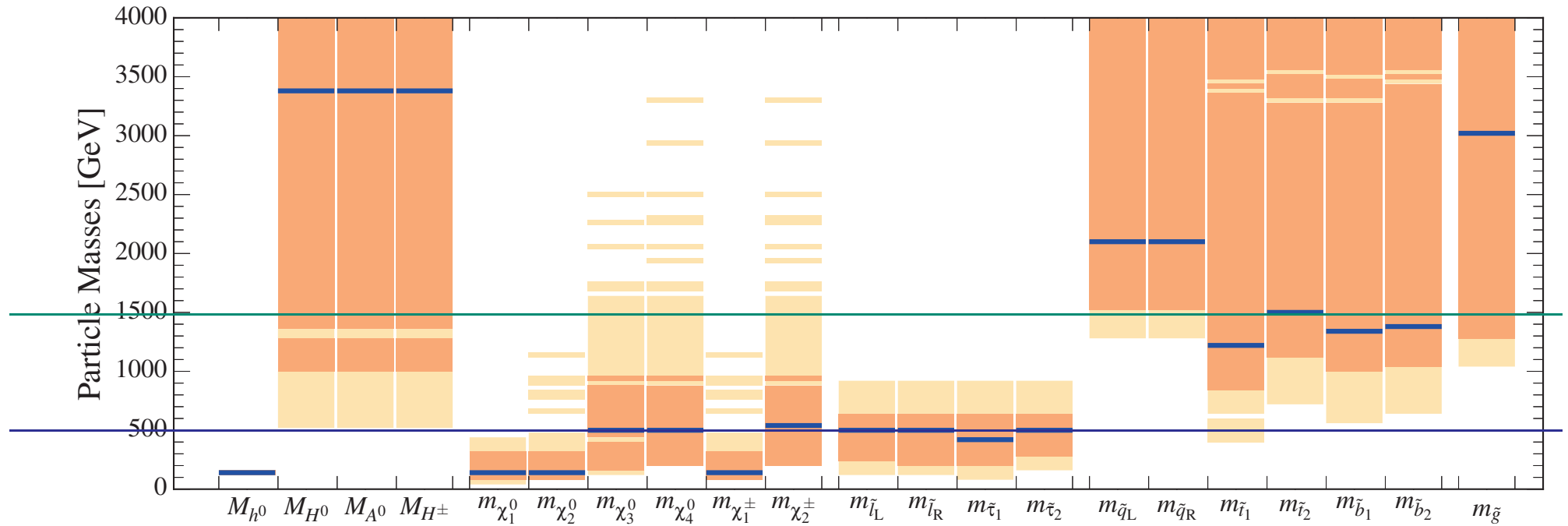






29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014

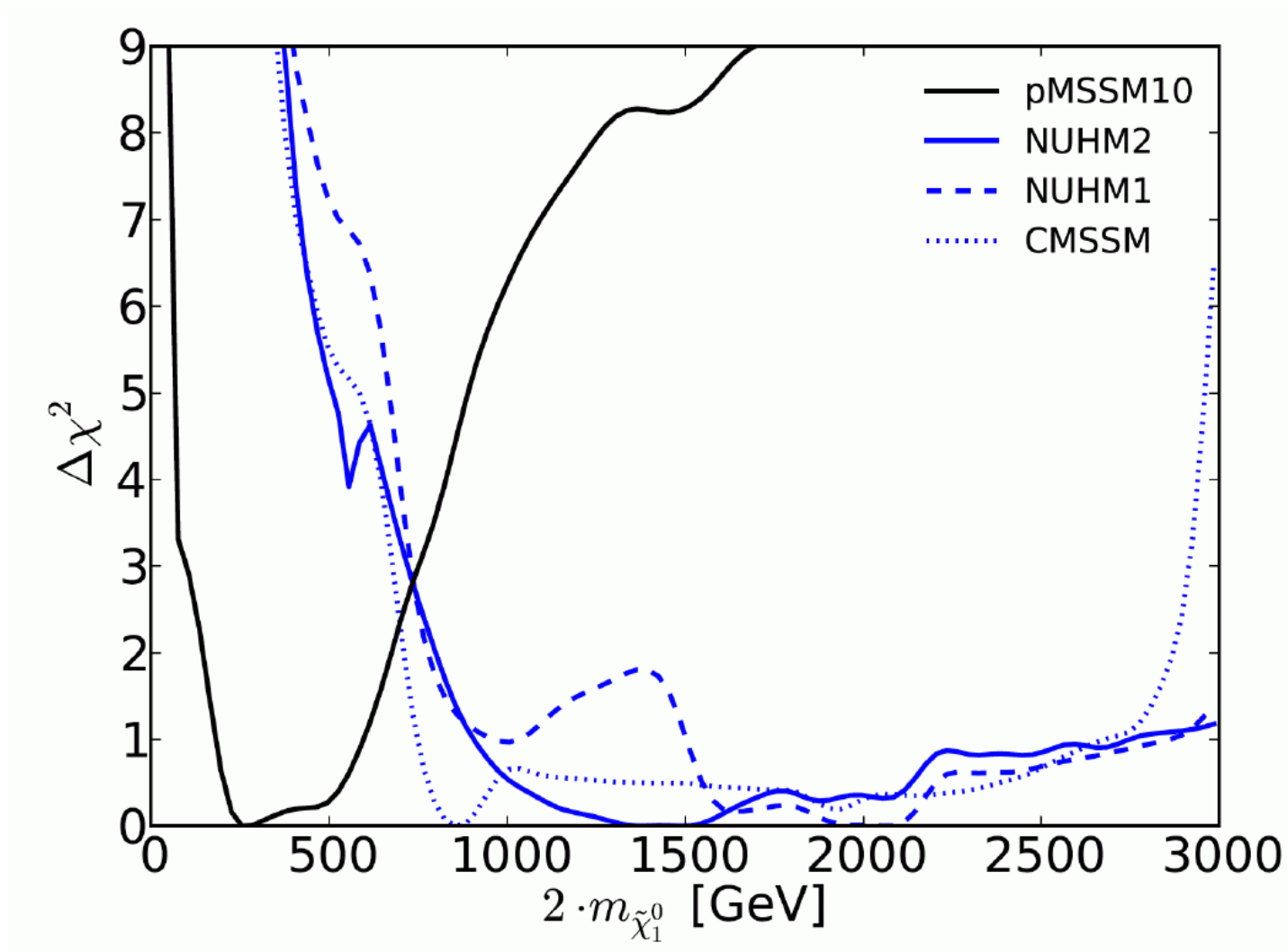


ILC:  $\sqrt{s} = 1000$  GeV  $\Rightarrow$  precision analysis of DM particle easy!

CLIC:  $\sqrt{s} = 3000$  GeV  $\Rightarrow$  precision analysis of DM particles easy!

# DM production cross sections: $e^+e^- \rightarrow \tilde{\chi}_1^0\tilde{\chi}_1^0(+\gamma)$

[2014]

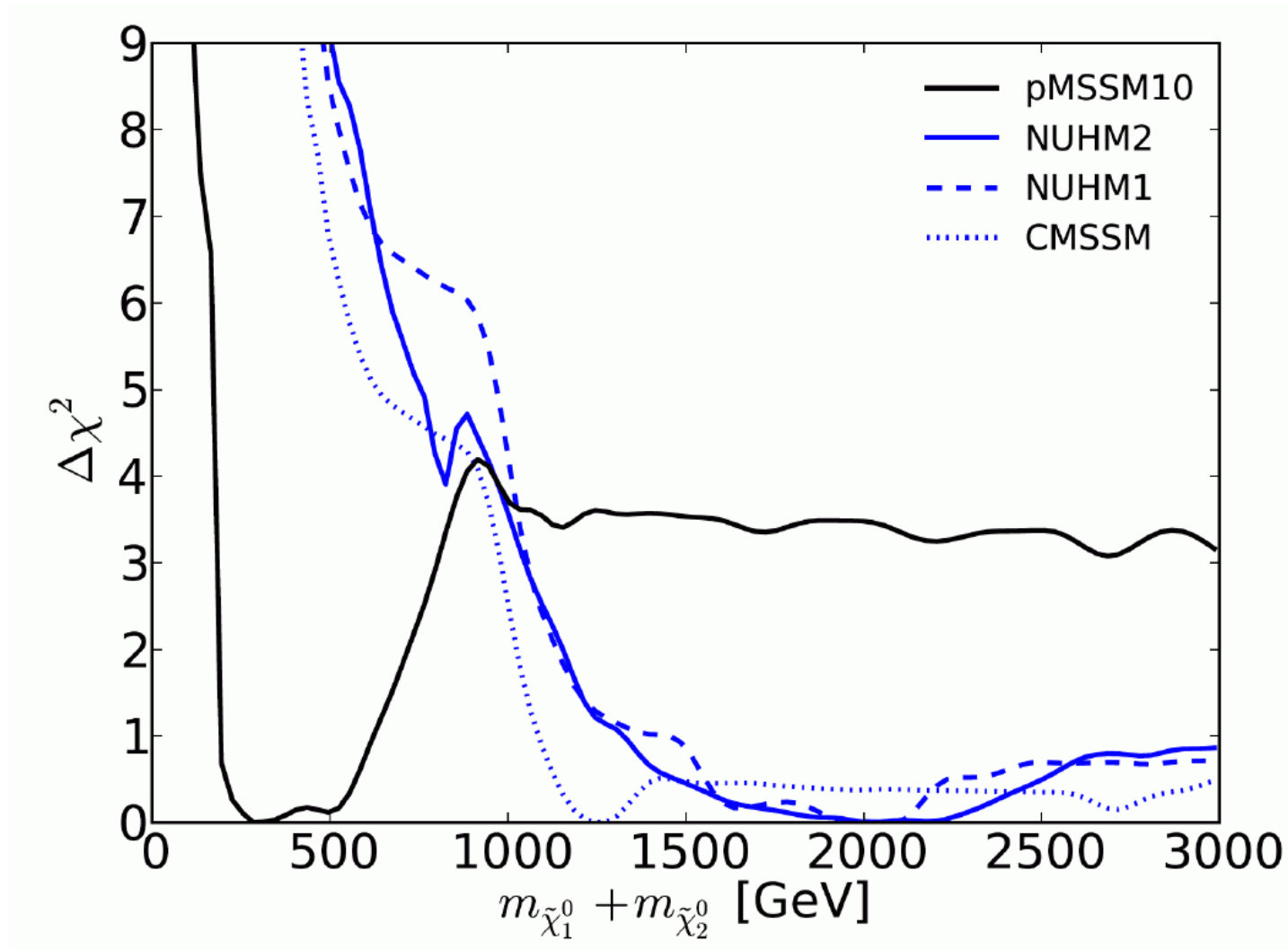


⇒ GUT based models: ILC :- ( , CLIC possible

⇒ pMSSM10: easy at the ILC

# DM production cross sections: $e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0$

[2014]



⇒ GUT based models: ILC :- ( , CLIC possible

⇒ pMSSM10: easy at the ILC - but no real upper limit