

TESTING VECTOR AND SCALAR PORTAL COUPLINGS

Felix Yu
JGU Mainz

with Jia Liu, Xiao-Ping Wang, JHEP **1706**, 077 (2017) [1704.00730]

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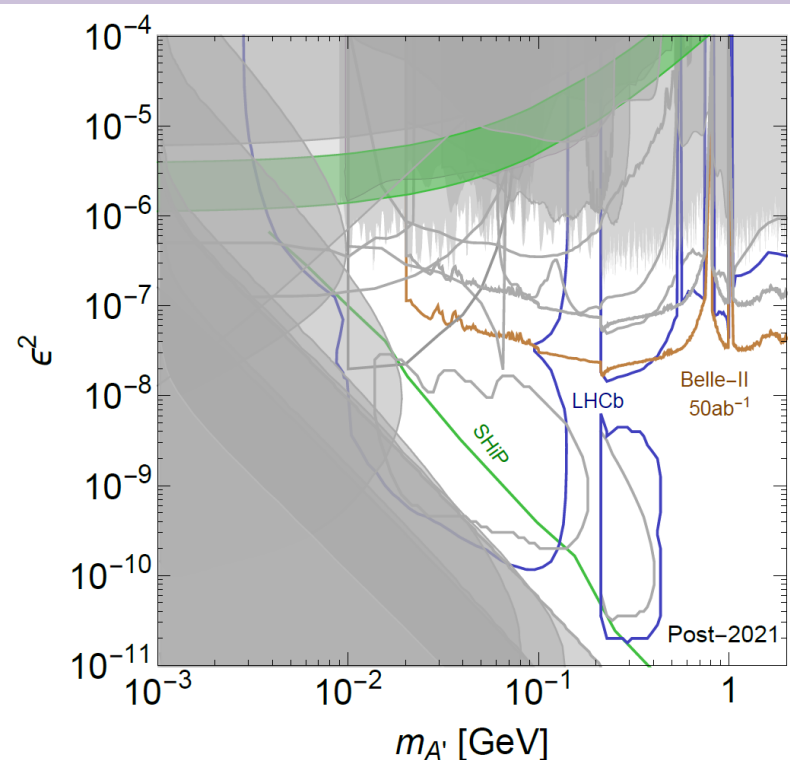
Direct production of New Particles at FCCee

- After Higgs discovery, particle physics has entered a distinct, exploratory phase
 - Outstanding problems, including dark matter and naturalness, are more acute
 - Mass scales of new physics unknown
- In conjunction, couplings of new physics particles are also unknown
 - TeV-scale strongly-coupled particles with prompt, cascade decays are strongly constrained
 - Weak-scale, weakly-coupled particles less constrained
 - Very weakly-coupled particles are very weakly constrained

Portal couplings

- Given direct probes at a given energy scale, sensitivity to UV scales follows NDA
 - Renormalizable, “portal” couplings are few (e.g. scalar Higgs portal, neutrino portal, vector portal, axion portal)

Alexander, et al. [1608.08632]



- Nevertheless, FCCee energies probe new portal couplings at mass scales untested by beam-dump experiments or LHC

FCCee can produce new particles directly here

Double Dark Portal model

Kinetic mixing of K with hypercharge gauge boson B

$$\mathcal{L} \supset -\frac{1}{4}B_{\mu\nu}B^{\mu\nu} - \frac{1}{4}W_{\mu\nu}^iW^{i\mu\nu} - \frac{1}{4}K_{\mu\nu}K^{\mu\nu} + \frac{\epsilon}{2\cos\theta_W}B_{\mu\nu}K^{\mu\nu} + |D_\mu H|^2 + |D_\mu \Phi|^2 + \mu_H^2|H|^2 - \lambda_H|H|^4 + \mu_D^2|\Phi|^2 - \lambda_D|\Phi|^4 - \lambda_{HP}|H|^2|\Phi|^2 + \bar{\chi}(i\not{D} - m_\chi)\chi$$

$U(1)_D$ charges $\Phi \sim +1$, $\chi \sim +1$

Scalar Higgs portal between dark Higgs Φ and SM H

- Two marginal operators: simultaneous vector portal and scalar portal couplings
 - Constraints driven by searches, not known from first principles (possible in UV completions)

Double Dark Portal model

- Fermion bilinears experience the new currents

$$\begin{aligned}\mathcal{L} &\supset gZ_{\mu, \text{SM}}J_Z^\mu + eA_{\mu, \text{SM}}J_{\text{em}}^\mu + g_D K_\mu J_D^\mu \\ &= \tilde{Z}_\mu \left(gJ_Z^\mu - g_D \frac{m_{Z, \text{SM}}^2 t_W}{m_{Z, \text{SM}}^2 - m_K^2} \epsilon J_D^\mu \right) \\ &\quad + \tilde{K}_\mu \left(g_D J_D^\mu + g \frac{m_K^2 t_W}{m_{Z, \text{SM}}^2 - m_K^2} \epsilon J_Z^\mu + e\epsilon J_{\text{em}}^\mu \right) \\ &\quad + \tilde{A}_\mu e J_{\text{em}}^\mu + \mathcal{O}(\epsilon^2)\end{aligned}$$

- U(1)_D- charged fermions pick up ϵ weak charge mediated by Z
- SM charged fermions pick up ϵ weak charge and ϵ electric charge mediated by dark photon
- Photon remains massless, long-range
 - (Singular behavior at $m_K = m_{Z, \text{SM}}$ is maximal mixing limit)

Double Dark Portal model

- Scalar boson mixing

- Higgs portal coupling leads to mass mixing between dark Higgs and SM Higgs

- Mixing angle

$$\tan 2\alpha = \frac{\lambda_{HP}v_Hv_D}{\lambda_Dv_D^2 - \lambda_Hv_H^2}$$

- Masses

$$m_{S, H_0}^2 = \lambda_Hv_H^2 + \lambda_Dv_D^2 \pm \sqrt{(\lambda_Hv_H^2 - \lambda_Dv_D^2)^2 + \lambda_{HP}v_H^2v_D^2}$$

- Dominant effect is $\cos \alpha$ -suppression of Higgs couplings to fermions, dark Higgs mass eigenstate S picks up $\sin \alpha$ -suppressed couplings to SM fermions

Double Dark Portal model

- Scalar-vector-vector interactions
 - Plays a key role in e^+e^- Higgs studies

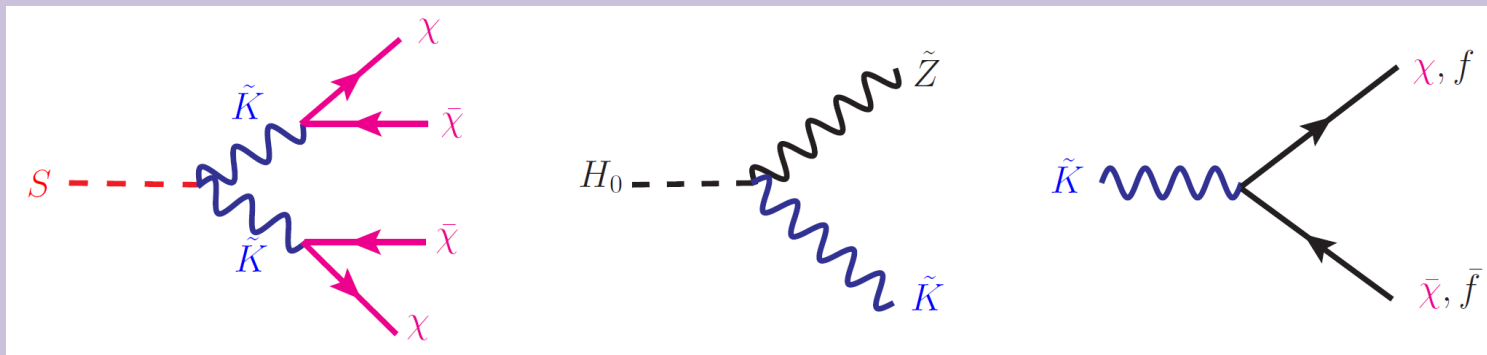
$$\begin{aligned}\mathcal{L} \supset & m_{Z,\text{SM}}^2 \left(\frac{\cos \alpha}{v_H} \right) \tilde{Z}_\mu \tilde{Z}^\mu H_0 \\ & + 2\epsilon t_W \frac{m_K^2 m_{Z,\text{SM}}^2}{(m_{Z,\text{SM}}^2 - m_K^2)} \left(\frac{\cos \alpha}{v_H} + \frac{\sin \alpha}{v_D} \right) \tilde{Z}_\mu \tilde{K}^\mu H_0 \\ & + m_K^2 \left(-\frac{\sin \alpha}{v_D} \right) \tilde{K}_\mu \tilde{K}^\mu H_0 \\ & + m_{Z,\text{SM}}^2 \left(\frac{\sin \alpha}{v_H} \right) \tilde{Z}_\mu \tilde{Z}^\mu S \\ & + 2\epsilon t_W \frac{m_K^2 m_{Z,\text{SM}}^2}{(m_{Z,\text{SM}}^2 - m_K^2)} \left(-\frac{\cos \alpha}{v_D} + \frac{\sin \alpha}{v_H} \right) \tilde{Z}_\mu \tilde{K}^\mu S \\ & + m_K^2 \left(\frac{\cos \alpha}{v_D} \right) \tilde{K}_\mu \tilde{K}^\mu S + \mathcal{O}(\epsilon^2)\end{aligned}$$

Phenomenology

- Three new states \tilde{K} , S , χ
- Many new interactions
 - Deviations in Z couplings
 - Deviations in Higgs couplings
 - Exotic Higgs decays (invisible, semi-visible, fully visible)
 - Interactions with dark matter mediated by dark photon
- Rich phenomenology for DM physics and colliders
 - Double Dark Portal model ties together two marginal couplings simultaneously
 - Attractive framework for marrying Higgs deviations and direct coupling to light, very-weakly coupled particles

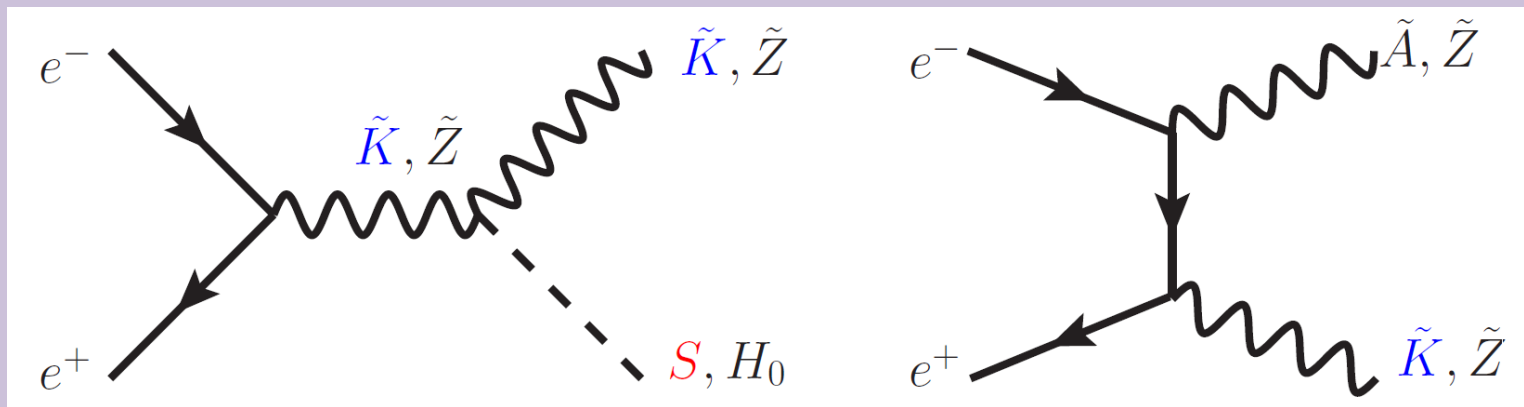
Collider phenomenology

- Modifications to Z couplings probed in precision electroweak observables
- Modifications to Higgs couplings tested by LHC and can be seen at a future Higgs factory
 - Also induce invisible and semi-visible exotic Higgs decays
- Will assume dark decays of S and K are on-shell
 - Ensured by kinematics and mild hierarchy for g_D and ε



Going beyond κ -framework, Higgs EFT

- New light states cause deviations in Higgs physics and can be directly produced



- Exploit radiative return process for hidden photon production
 - Recoil mass technique adapted to monophoton events and other SM candles as recoil taggers

Exploiting radiative return and recoil mass techniques at e^+e^- machines

- Radiative return – use ISR photon to make 2-2 production on-shell
 - At LHC, “radiative return” is better known as “mono-jet”
- Recoil mass method – use four-momentum conservation in 2-2 process
 - In case of invisible decay and radiative return, equivalent to searching for a monophoton peak
 - Design driver for e^+e^- electromagnetic calorimeter

$$E_{\text{vis}} = \frac{\sqrt{s}}{2} + \frac{m_{\text{vis}}^2 - m_X^2}{2\sqrt{s}}$$

$$m_{\text{recoil}} = m_X = \sqrt{s + m_{\text{vis}}^2 - 2E_{\text{vis}}\sqrt{s}}$$

Exotic invisible decay of Higgs

- Familiar case: Higgs recoiling against Z for invisible Higgs decays

– Invisible decay combines sensitivity to $\sin \alpha$ and ϵ , overall rate driven by g_D

$$\Gamma(H_0 \rightarrow \text{inv}) \approx \Gamma(H_0 \rightarrow SS) + \Gamma(H_0 \rightarrow \tilde{K}\tilde{K}) + 0.2 \times \Gamma(H_0 \rightarrow \tilde{K}\tilde{Z})$$

- Individual rates are

$$\Gamma(H_0 \rightarrow SS) = g_D^2 \sin^2 \alpha \frac{m_{H_0}}{32\pi} \sqrt{1 - \frac{4m_S^2}{m_{H_0}^2} \frac{(m_{H_0}^2 + 2m_S^2)^2}{m_{H_0}^2 m_K^2}},$$

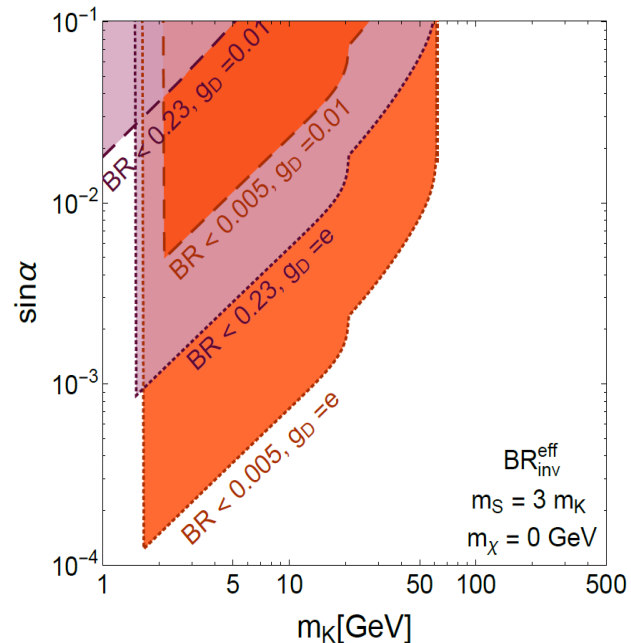
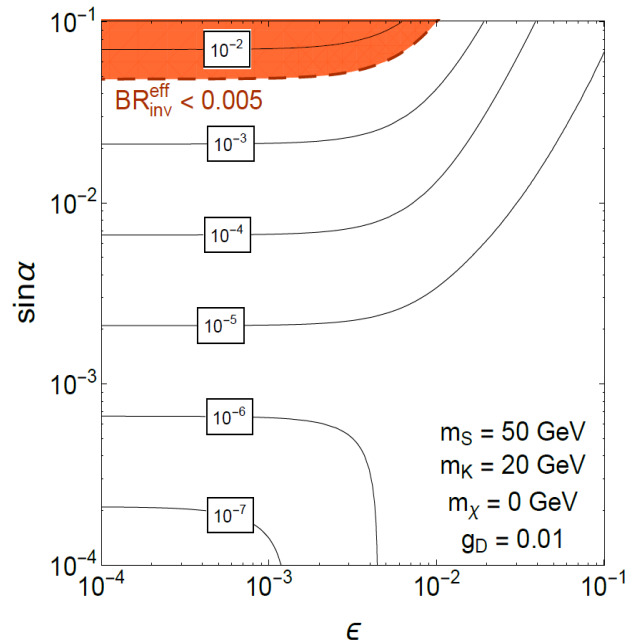
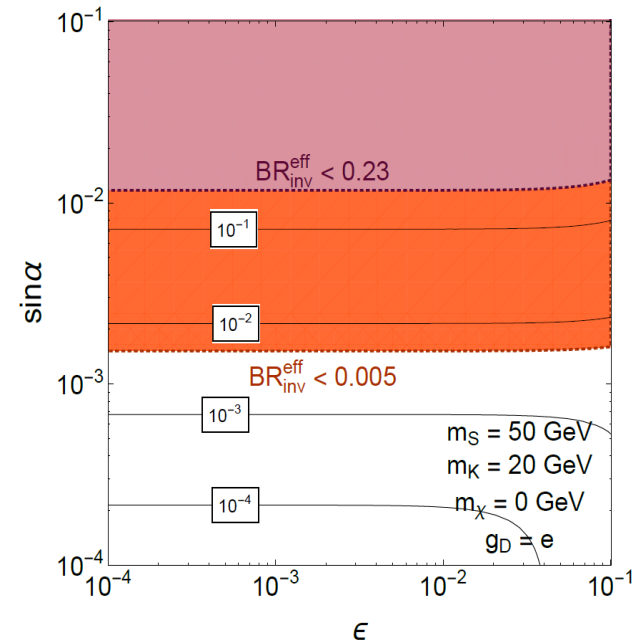
$$\Gamma(H_0 \rightarrow \tilde{K}\tilde{K}) = g_D^2 \sin^2 \alpha \frac{m_{H_0}}{32\pi} \sqrt{1 - \frac{4m_{\tilde{K}}^2}{m_{H_0}^2} \frac{m_{H_0}^4 - 4m_{H_0}^2 m_{\tilde{K}}^2 + 12m_{\tilde{K}}^4}{m_{H_0}^2 m_{\tilde{K}}^2} \frac{m_{\tilde{K}}^2}{m_{\tilde{K}}^2}},$$

$$\Gamma(H_0 \rightarrow \tilde{K}\tilde{Z}) = \frac{\epsilon^2 t_W^2 \left(\frac{\cos \alpha}{v_H} + \frac{\sin \alpha}{v_D} \right)^2}{16\pi m_{H_0}^3 \left(m_K^2 - m_{Z, \text{SM}}^2 \right)^2} \frac{m_K^4 m_{Z, \text{SM}}^4}{m_{\tilde{K}}^2 m_{\tilde{Z}}^2} \sqrt{m_{H_0}^4 + \left(m_{\tilde{K}}^2 - m_{\tilde{Z}}^2 \right)^2 - 2m_{H_0}^2 \left(m_{\tilde{K}}^2 + m_{\tilde{Z}}^2 \right)} \\ \times \left((m_{H_0}^2 - m_{\tilde{K}}^2 - m_{\tilde{Z}}^2)^2 + 8m_{\tilde{K}}^2 m_{\tilde{Z}}^2 \right)$$

Exotic invisible decay of Higgs

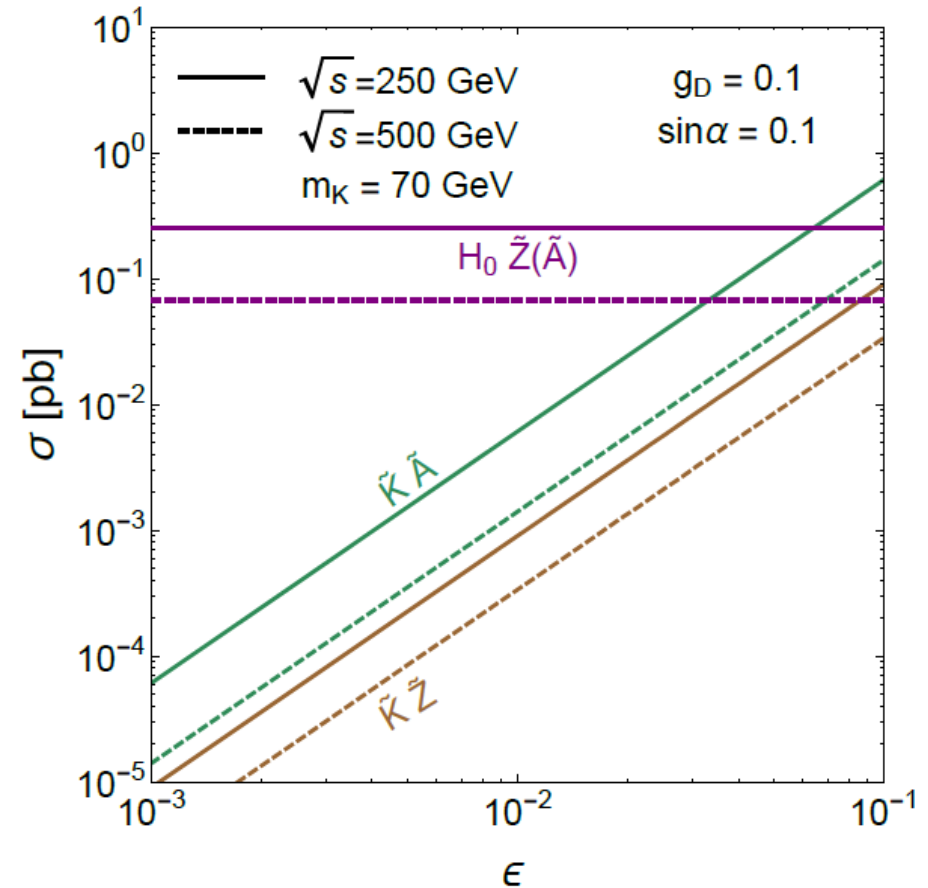
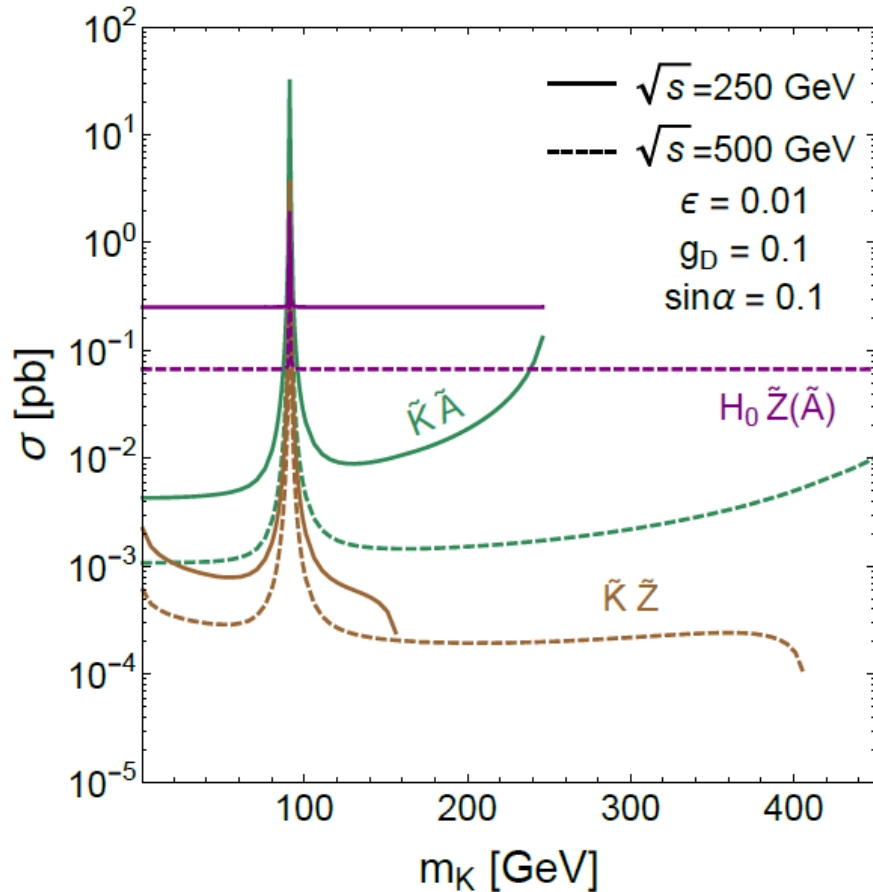
- Familiar case: Higgs recoiling against Z for invisible Higgs decays
 - Invisible decay combines sensitivity to $\sin \alpha$ and ϵ , overall rate driven by g_D

$$\Gamma(H_0 \rightarrow \text{inv}) \approx \Gamma(H_0 \rightarrow SS) + \Gamma(H_0 \rightarrow \tilde{K}\tilde{K}) + 0.2 \times \Gamma(H_0 \rightarrow \tilde{K}\tilde{Z})$$



Direct production of new light states

- Possible new physics within kinematic reach
 - Signatures too difficult at LHC, exploit e^+e^- capabilities



Prospects for dark photon

- Many possible visible and invisible final states

$$e^+e^- \rightarrow \tilde{Z}H_0 \text{ Study } \tilde{Z} \rightarrow \ell\ell \text{ and semi-visible } H_0 \rightarrow (\ell\ell)_Z\chi\chi$$

$$e^+e^- \rightarrow \tilde{Z}\tilde{K} \text{ Study } \tilde{Z} \rightarrow \ell\ell \text{ and } \tilde{K} \rightarrow \bar{\chi}\chi \text{ or } \ell\ell$$

$$e^+e^- \rightarrow \gamma\tilde{K} \text{ Study } \tilde{K} \text{ inclusive decays, and exclusive } \tilde{K} \rightarrow \bar{\chi}\chi \text{ or } \ell\ell$$

$$e^+e^- \rightarrow \tilde{Z}S \text{ Study } \tilde{Z} \rightarrow \ell\ell \text{ and } S \rightarrow 4\chi$$

- Event simulation using MG5+Pythia+Delphes

- Use parametrized preliminary CEPC detector card

- SM backgrounds and cuts driven by e^+e^- environment

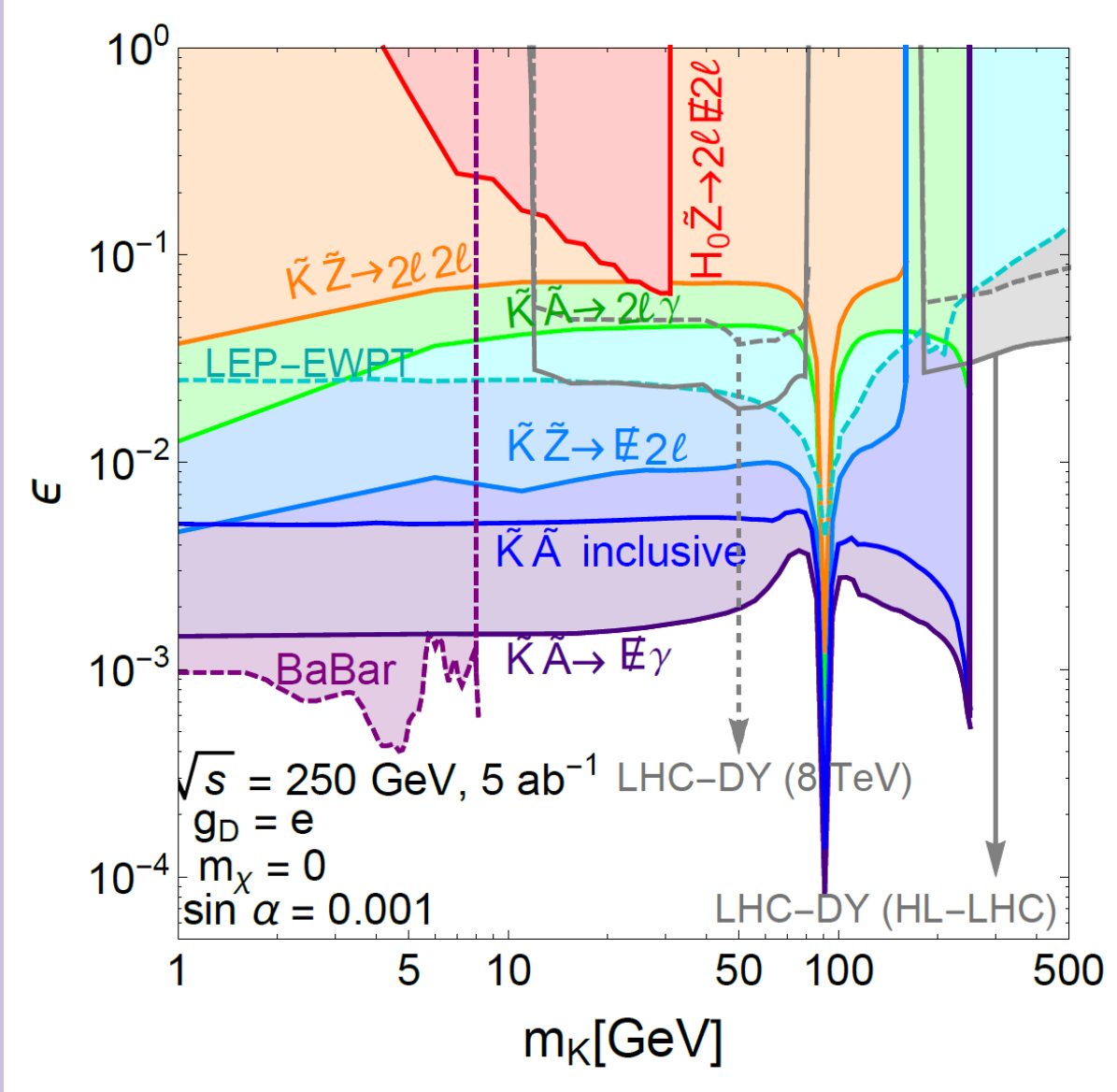
- Rates for visible states are lower by $(\epsilon/g_D)^2$, best sensitivity from requiring missing energy threshold

- LEP direct constraints ($\epsilon < 0.03$) not competitive

Collider study cuts

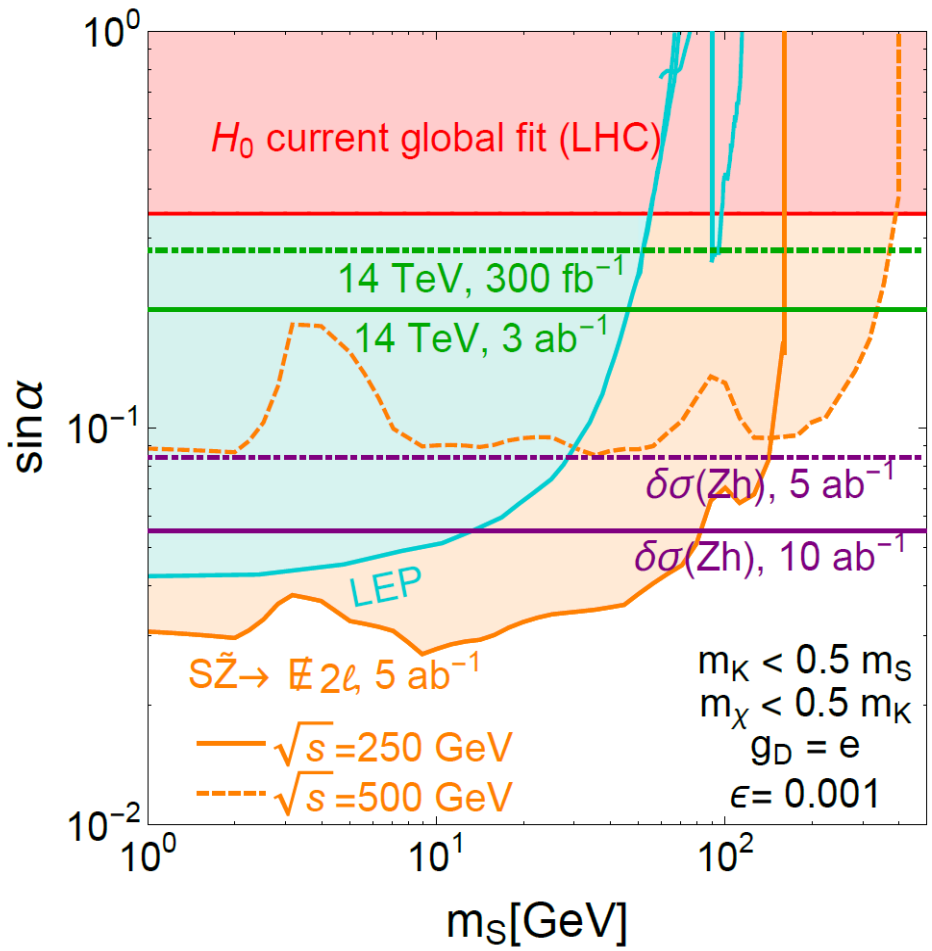
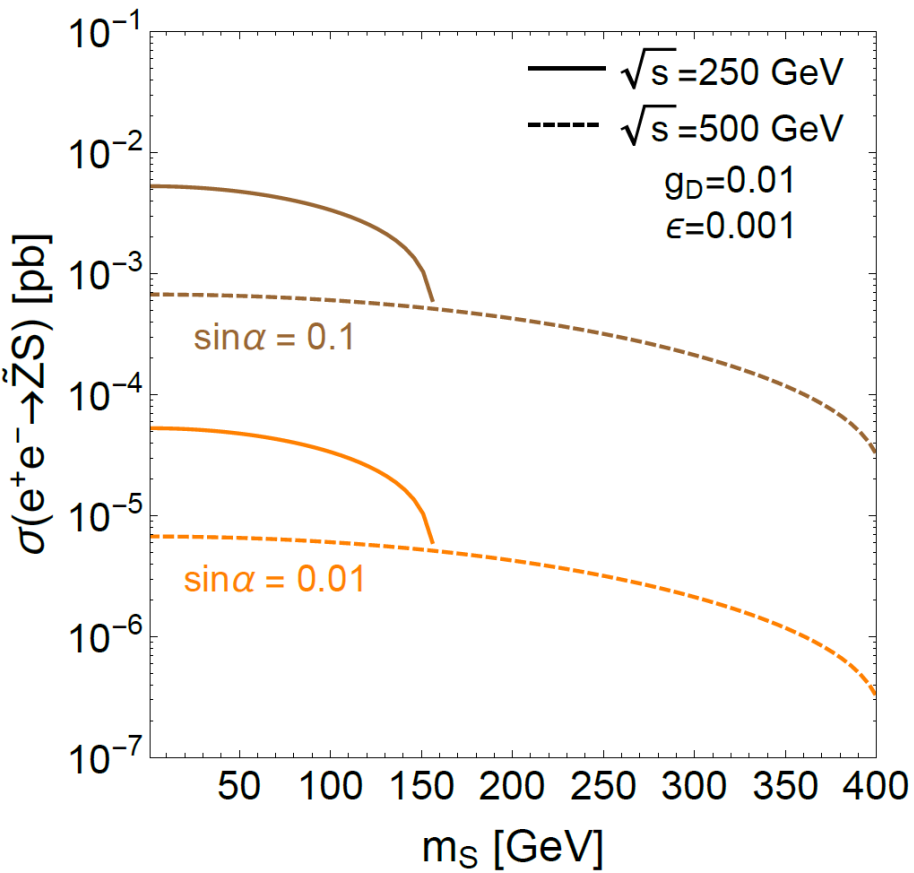
Parameter	Signal process		Background (pb)		Signal region	
ϵ	$\tilde{Z}\tilde{K}$	$\tilde{Z} \rightarrow \bar{\ell}\ell, \tilde{K} \rightarrow \bar{\chi}\chi$	$\bar{\ell}\bar{\ell}\bar{\nu}\nu$	0.929 (250 GeV)	$N_\ell \geq 2, m_{\ell\ell} - m_Z < 10 \text{ GeV},$ and $ m_{\text{recoil}} - m_{\tilde{K}} < 2.5 \text{ GeV}$	
				0.545 (500 GeV)		
		$\tilde{Z} \rightarrow \bar{\ell}\ell, \tilde{K} \rightarrow \bar{\ell}\ell$	$\bar{\ell}\bar{\ell}\bar{\ell}\bar{\ell}$	0.055 (250 GeV)		$N_\ell \geq 4, m_{\ell\ell} - m_Z < 10 \text{ GeV},$ and $ m_{\ell\ell} - m_{\tilde{K}} < 2.5 \text{ GeV}$
				0.023 (500 GeV)		
	$\tilde{A}\tilde{K}$	\tilde{K} inclusive decay	$\gamma\bar{f}f$	23.14 (250 GeV)	$N_\gamma \geq 1, \text{ and}$ $ E_\gamma - (\frac{\sqrt{s}}{2} - \frac{m_{\tilde{K}}^2}{2\sqrt{s}}) < 2.5 \text{ GeV}$	
				8.88 (250 GeV)		
		$\tilde{K} \rightarrow \bar{\ell}\ell$	$\gamma\bar{\ell}\bar{\ell}$	12.67 (250 GeV)	$N_\gamma \geq 1, N_\ell \geq 2,$ $ E_\gamma - (\frac{\sqrt{s}}{2} - \frac{m_{\tilde{K}}^2}{2\sqrt{s}}) < 2.5 \text{ GeV},$ and $ m_{\ell\ell} - m_{\tilde{K}} < 5 \text{ GeV}$	
				4.38 (500 GeV)		
		$\tilde{K} \rightarrow \bar{\chi}\chi$	$\gamma\bar{\nu}\nu$	3.45 (250 GeV)	$N_\gamma \geq 1,$ $ E_\gamma - (\frac{\sqrt{s}}{2} - \frac{m_{\tilde{K}}^2}{2\sqrt{s}}) < 2.5 \text{ GeV},$ and $\cancel{E} > 50 \text{ GeV}$	
				2.92 (500 GeV)		
		$\tilde{Z}H_0$	$H_0 \rightarrow \tilde{K}\tilde{Z}$ with $\tilde{K} \rightarrow \bar{\chi}\chi, \tilde{Z} \rightarrow \bar{\ell}\ell$	$\bar{\ell}\bar{\ell}\bar{\ell}\bar{\ell}\bar{\nu}\nu$	1.8×10^{-5} (250 GeV)	$N_\ell \geq 4, m_{\ell\ell} - m_Z < 10 \text{ GeV},$ and $ m_{\text{recoil}} - m_{\tilde{K}} < 2.5 \text{ GeV}$
					3.5×10^{-4} (500 GeV)	
$\sin \alpha$	$\tilde{Z}S$	$\tilde{Z} \rightarrow \bar{\ell}\ell$ $S \rightarrow \tilde{K}\tilde{K} \rightarrow 4\chi$	$\bar{\ell}\bar{\ell}\bar{\nu}\nu$	0.87 (250 GeV)	$N_\ell \geq 2, m_{\ell\ell} - m_Z < 10 \text{ GeV},$ and $ m_{\text{recoil}} - m_S < 2.5 \text{ GeV}$	
				0.87 (250 GeV)		

Dark photon sensitivity



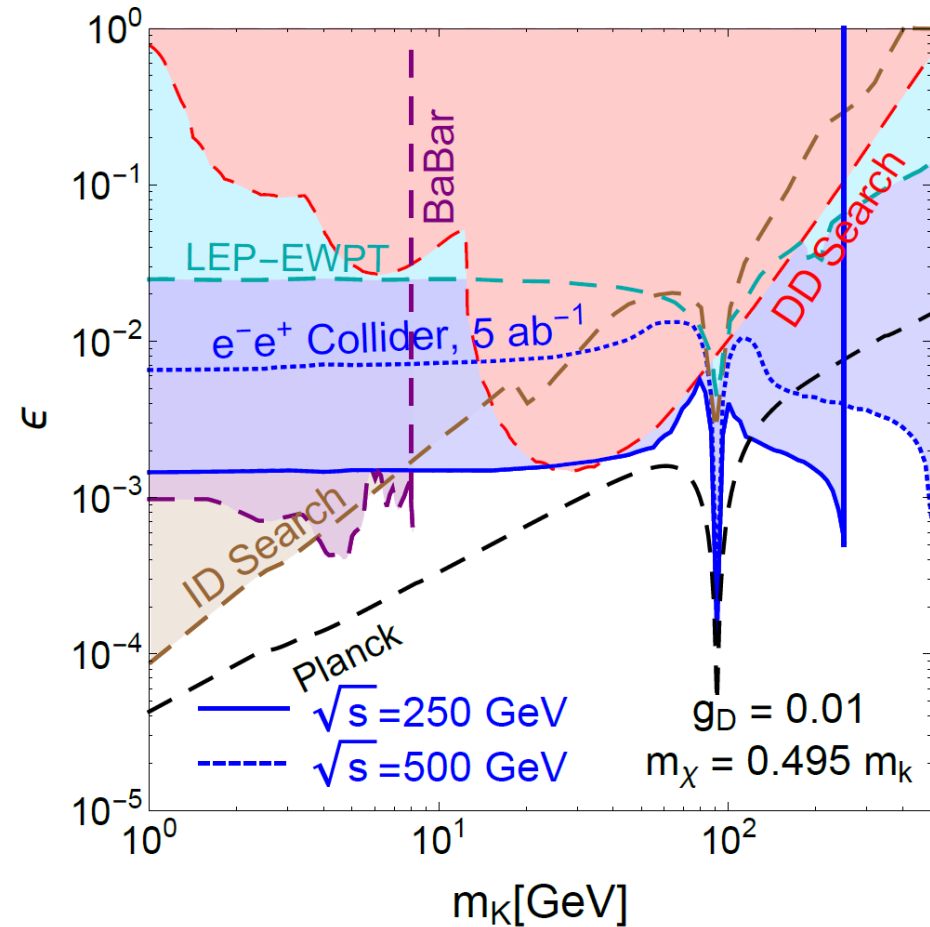
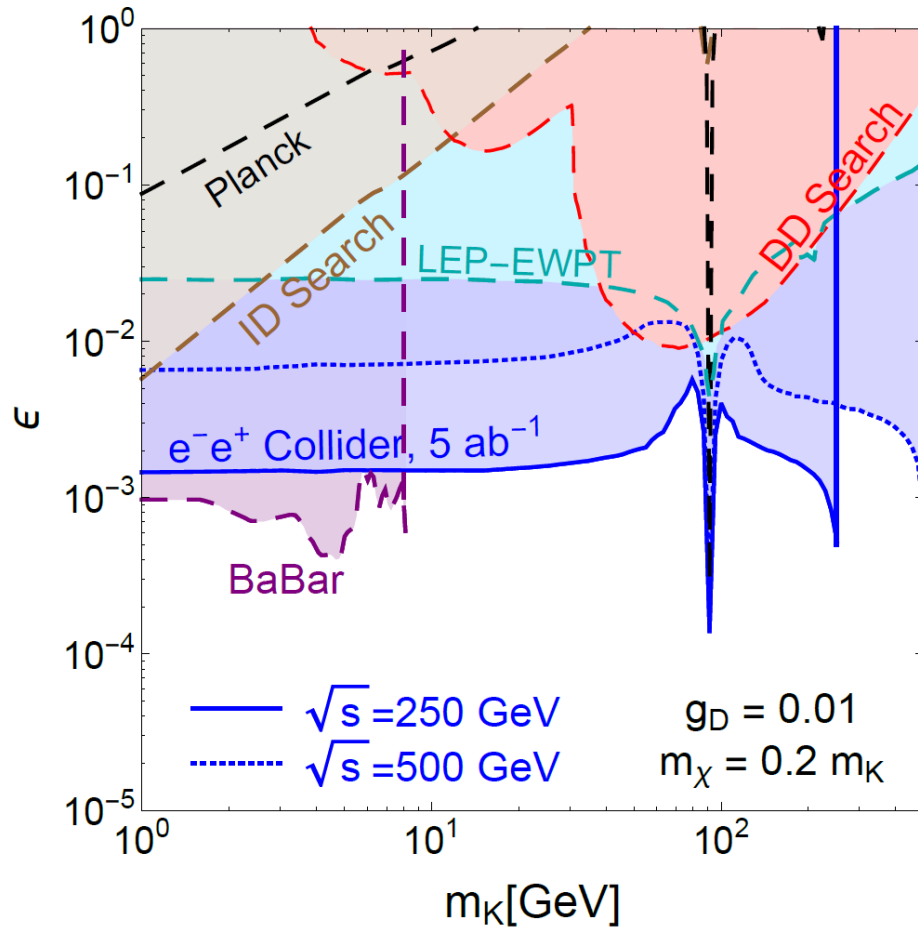
Prospects for dark scalar

- Similarly, direct dark Higgs production and precision Higgs measurements



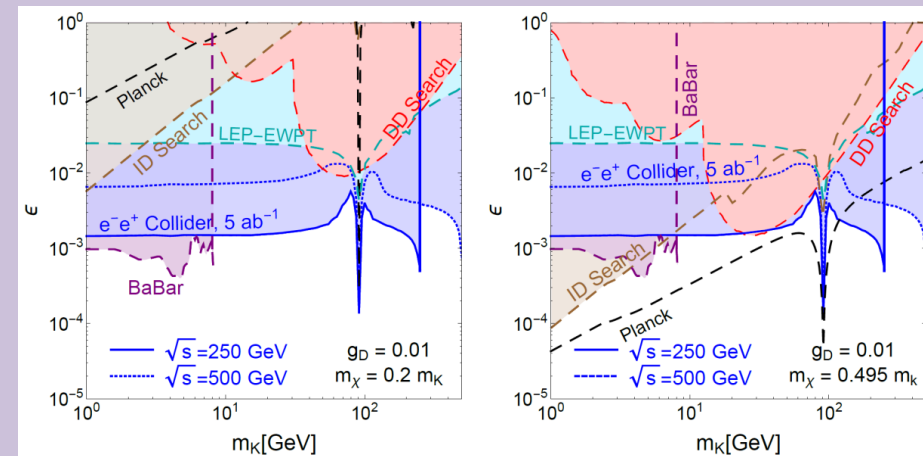
Comparing to complementary DM probes

- Dark matter discovery possible at e^+e^- machines



Conclusions

- Physics potential of e^+e^- machine goes well beyond precision Standard Model program
- Direct production of new, light, very weakly-coupled hidden particles possible
- Double Dark Portal model is a concrete framework for studying two marginal couplings in tandem



Double Dark Portal model

- Steps for solving the neutral vector Lagrangian (pedagogical)

- Diagonalize gauge boson mass matrix

- Usual $t_W = g'/g$ rotation corresponds to

$$\mathcal{L} \supset \frac{-1}{4} \begin{pmatrix} Z_{SM}^{\mu\nu} & A_{SM}^{\mu\nu} & K^{\mu\nu} \end{pmatrix} \begin{pmatrix} 1 & 0 & \epsilon t_W \\ 0 & 1 & -\epsilon \\ \epsilon t_W & -\epsilon & 1 \end{pmatrix} \begin{pmatrix} Z_{\mu\nu, SM} \\ A_{\mu\nu, SM} \\ K_{\mu\nu} \end{pmatrix}$$

$$+ \frac{1}{2} \begin{pmatrix} Z_{SM}^\mu & A_{SM}^\mu & K^\mu \end{pmatrix} \begin{pmatrix} m_{Z, SM}^2 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & m_K^2 \end{pmatrix} \begin{pmatrix} Z_{\mu, SM} \\ A_{\mu, SM} \\ K_\mu \end{pmatrix}$$

- Require $|\epsilon| < c_W$ for positive kinetic mixing determinant
- Field strengths are Abelian kinetic terms, non-Abelian interactions inherited from transformations

Double Dark Portal model

- Steps for solving the neutral vector Lagrangian (pedagogical)

– Remove kinetic mixing and canonically normalize

$$U_1 = \begin{pmatrix} 1 & 0 & 0 \\ -\epsilon^2 t_W & 1 & \epsilon \\ -\epsilon t_W & 0 & 1 \end{pmatrix} \quad U_2 = \begin{pmatrix} \sqrt{\frac{1-\epsilon^2}{1-\epsilon^2 c_W^{-2}}} & 0 & 0 \\ 0 & 1 & 0 \\ \frac{-\epsilon^3 t_W}{\sqrt{(1-\epsilon^2)(1-\epsilon^2 c_W^{-2})}} & 0 & \frac{1}{\sqrt{1-\epsilon^2}} \end{pmatrix}$$

$$\begin{aligned} \mathcal{L} \supset & \frac{-1}{4} \begin{pmatrix} Z_{SM}^{\mu\nu} & A_{SM}^{\mu\nu} & K^{\mu\nu} \end{pmatrix} (U_1^T)^{-1} (U_2^T)^{-1} \mathbb{I}_3 U_2^{-1} U_1^{-1} \begin{pmatrix} Z_{\mu\nu, SM} \\ A_{\mu\nu, SM} \\ K_{\mu\nu} \end{pmatrix} \\ & + \frac{1}{2} \begin{pmatrix} Z_{SM}^\mu & A_{SM}^\mu & K^\mu \end{pmatrix} (U_1^T)^{-1} (U_2^T)^{-1} \begin{pmatrix} \frac{m_{Z, SM}^2 (1-\epsilon^2)^2 + m_K^2 \epsilon^2 t_W^2}{(1-\epsilon^2)(1-\epsilon^2 c_W^{-2})} & 0 & \frac{-m_K^2 \epsilon t_W}{(1-\epsilon^2) \sqrt{1-\epsilon^2 c_W^{-2}}} \\ 0 & 0 & 0 \\ \frac{-m_K^2 \epsilon t_W}{(1-\epsilon^2) \sqrt{1-\epsilon^2 c_W^{-2}}} & 0 & \frac{m_K^2}{1-\epsilon^2} \end{pmatrix} \\ & \times U_2^{-1} U_1^{-1} \begin{pmatrix} Z_{\mu, SM} \\ A_{\mu, SM} \\ K_\mu \end{pmatrix} \end{aligned}$$

Double Dark Portal model

- Steps for solving the neutral vector Lagrangian (pedagogical)
 - Rediagonalize mass matrix via Jacobi rotation (exact)
 - To $O(\epsilon^3)$, masses and fields are

$$m_{\tilde{K}}^2 = m_K^2 + \frac{m_K^2 c_W^{-2} \epsilon^2 (m_{Z, \text{SM}}^2 c_W^2 - m_K^2)}{m_{Z, \text{SM}}^2 - m_K^2}, \quad m_{\tilde{Z}}^2 = m_{Z, \text{SM}}^2 + \frac{m_{Z, \text{SM}}^4 t_W^2 \epsilon^2}{m_{Z, \text{SM}}^2 - m_K^2}$$

$$\begin{pmatrix} \tilde{Z}_\mu \\ \tilde{A}_\mu \\ \tilde{K}_\mu \end{pmatrix} = \begin{pmatrix} Z_{\mu, \text{SM}} - \frac{t_W m_K^2}{m_{Z, \text{SM}}^2 - m_K^2} \epsilon K_\mu - \frac{m_{Z, \text{SM}}^4 t_W^2}{2(m_{Z, \text{SM}}^2 - m_K^2)^2} \epsilon^2 Z_{\mu, \text{SM}} \\ A_{\mu, \text{SM}} - \epsilon K_\mu \\ K_\mu + \frac{t_W m_{Z, \text{SM}}^2}{m_{Z, \text{SM}}^2 - m_K^2} \epsilon Z_{\mu, \text{SM}} - \left(\frac{1}{2} + \frac{m_K^4 t_W^2}{2(m_{Z, \text{SM}}^2 - m_K^2)^2} \right) \epsilon^2 K_\mu \end{pmatrix}$$

- Singular behavior at $m_K = m_{Z, \text{SM}}$ is maximal mixing limit
 - Effects from field redefinitions seen in dark, SM currents

Dark matter direct detection

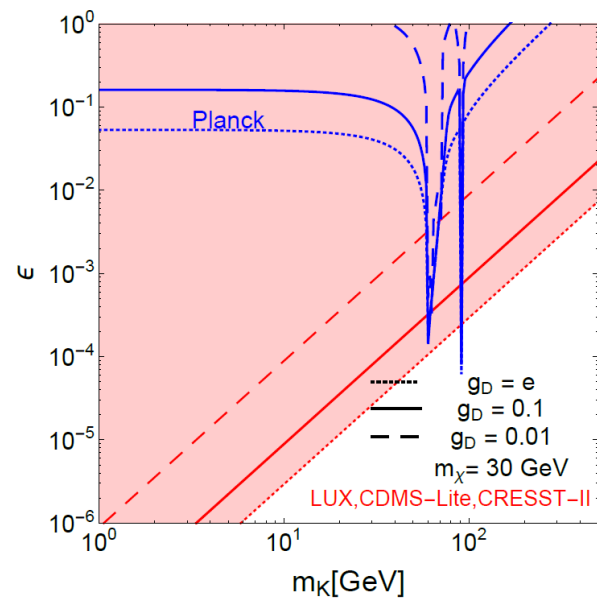
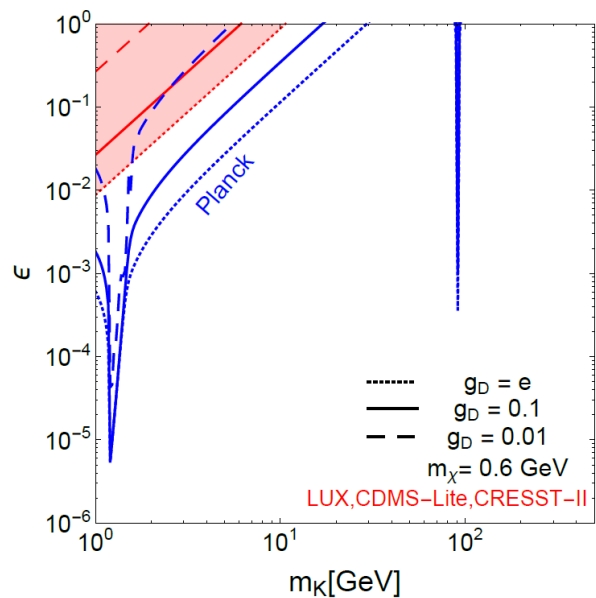
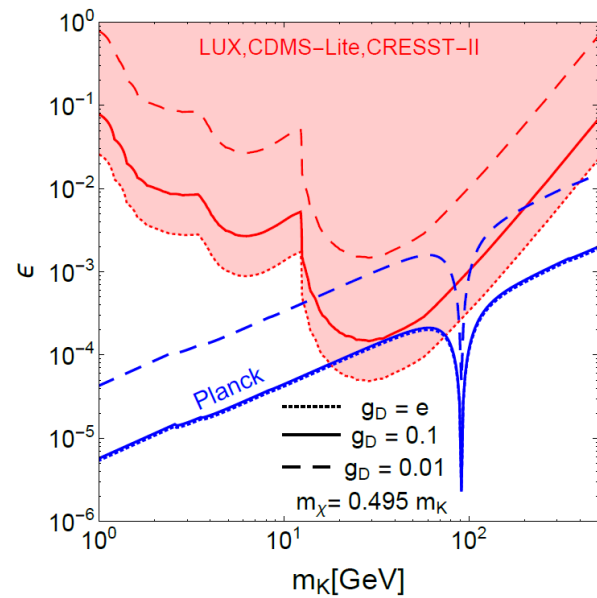
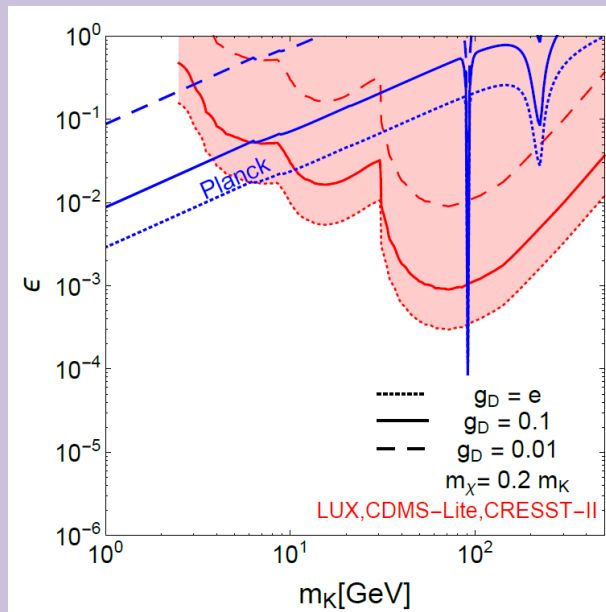
- Dark matter scattering off protons dominantly from dark photon exchange, suppressed by $(\epsilon e)^2$
 - Intrinsic cancellation between weak charged currents mediated by massive Z and K vectors (at this order in ϵ)
 - Dark matter does not interact with photon, hence only protons contribute to direct detection

$$\sigma_p \simeq \frac{\epsilon^2 g_D^2 e^2}{\pi} \frac{\mu_{\chi p}^2}{m_{\tilde{K}}^4} \approx 10^{-44} \text{ cm}^2 \left(\frac{g_D}{e}\right)^2 \left(\frac{\epsilon}{10^{-5}}\right)^2 \left(\frac{10 \text{ GeV}}{m_{\tilde{K}}}\right)^2$$

Dark matter direct detection

- Exclusion limits are highly sensitive to the dark matter mass

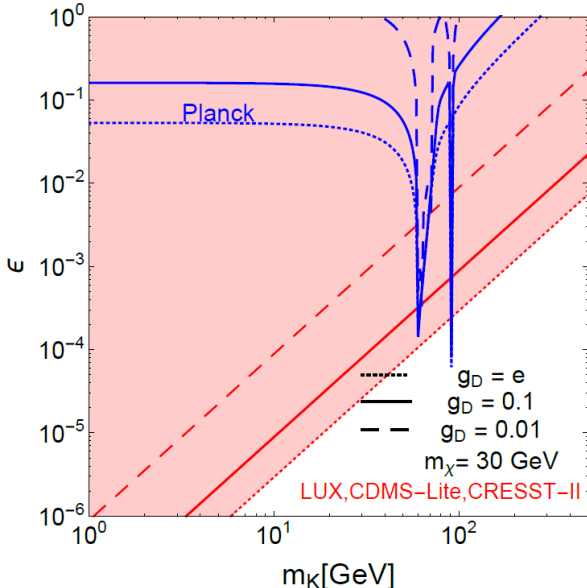
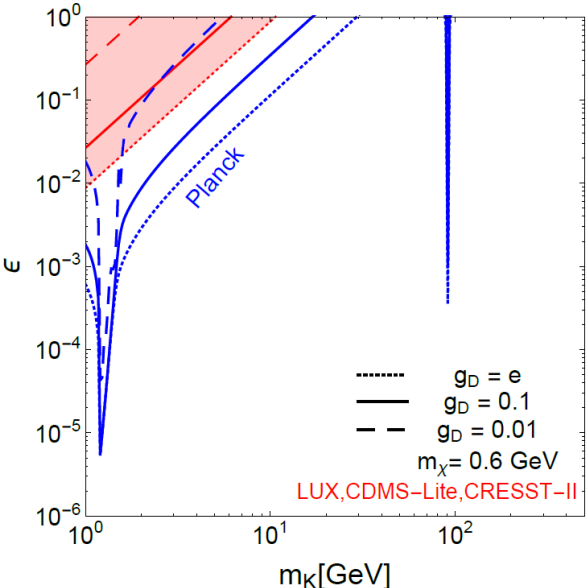
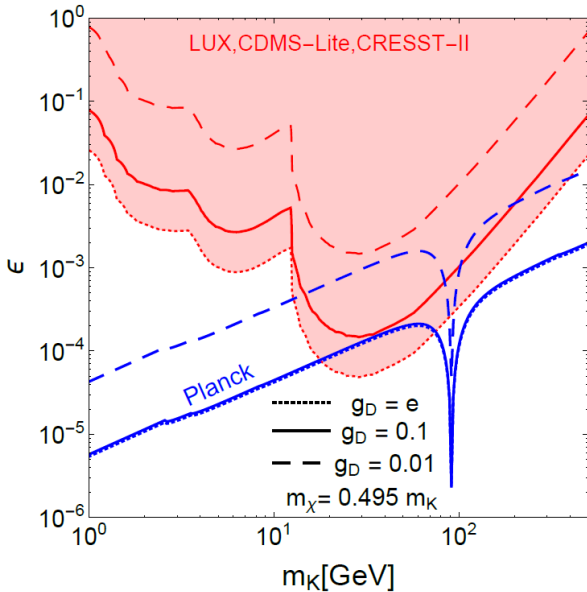
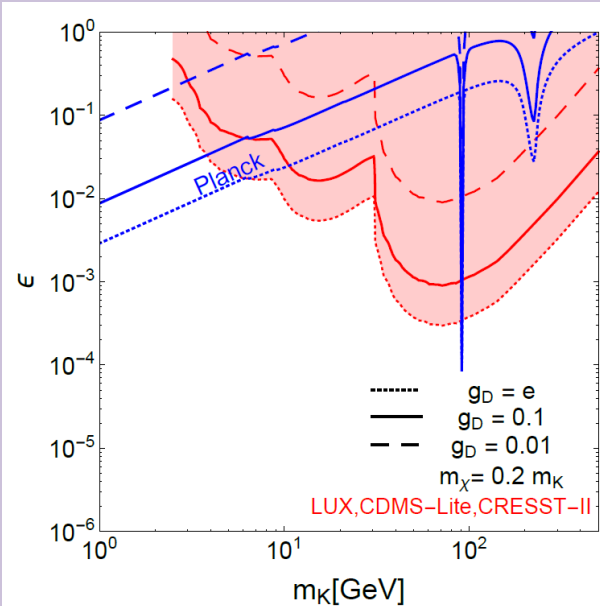
– Nuclear recoil energy threshold becomes too soft for light dark matter (about 5 GeV)



Dark matter direct detection

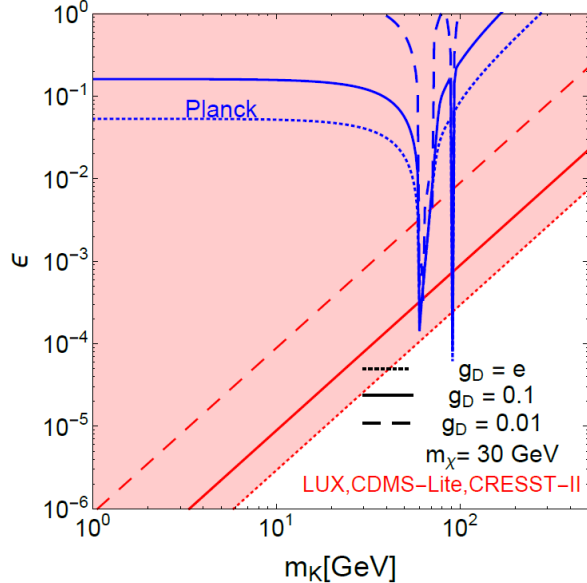
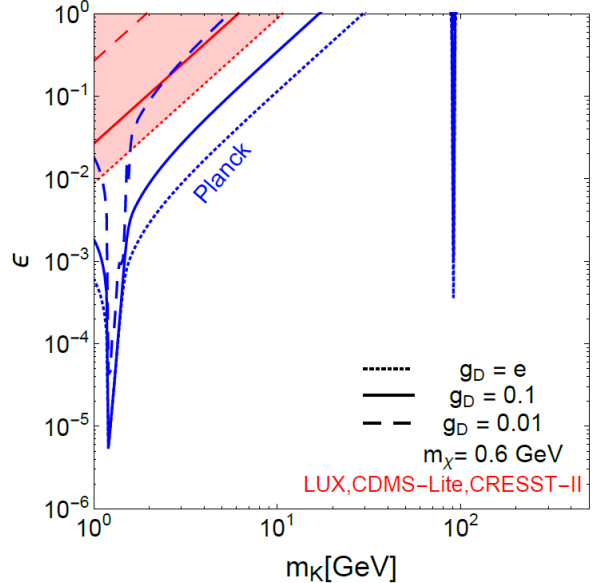
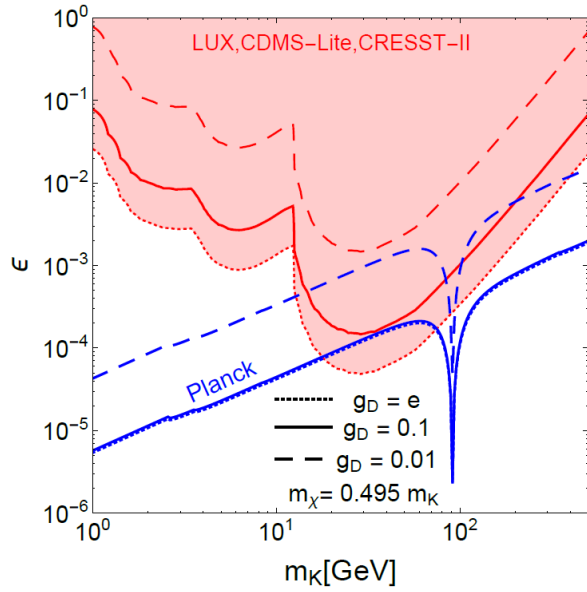
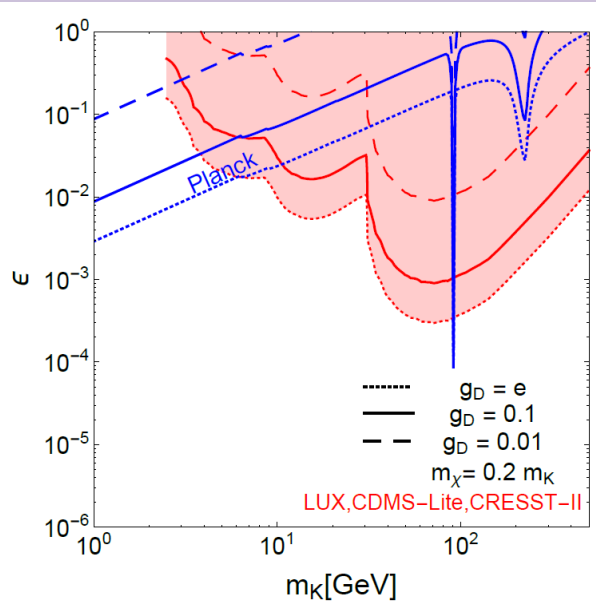
- Relic abundance (blue line) shows resonances at dark photon and Z masses

- DM is underabundant above blue line, overabundant below blue line



Dark matter direct detection

- Dark matter experiments fix the local relic abundance to 0.3 GeV/cm^3
 - On the other hand, the predicted dark matter relic abundance scales as ϵ^{-2} , while the scattering rate scales as ϵ^2
- Ratio of DD limits to relic abundance curve (for fixed m_K) gives the limit on local abundance



Dark matter indirect detection

- Present day annihilation constrained by observations of gamma ray spectra
- Early universe annihilation constrained by energy injection in CMB
- Strongest limits when DM mass is close to Z or dark photon resonance

