
Baryogenesis

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What is the problem?

Open questions

- Why there is basically only matter in the universe?

$$n(\bar{B}) \ll n(B)$$

- Quantitatively, can we explain

$$\eta \equiv \frac{n(B) - n(\bar{B})}{N(\gamma)} \approx 6 \times 10^{-10}$$

B stands for Baryons. γ stands for CMB photons

Remarks

$$n(\bar{B}) \ll n(B) \quad \& \quad \eta \equiv \frac{n(B) - n(\bar{B})}{N(\gamma)} \approx 6 \times 10^{-10}$$

- Rather amazing that we can even ask this question
- The SM predicts an hierarchy, but it falls short in getting the numerical value of η
- Many explanations beyond the SM. Nothing confirmed
- Not easy to probe any of these ideas as there is only one number to explain
- The solution must involve interplay of particle physics and cosmology
- Maybe neutrinos are the key to the solution

Some vocabulary

- Baryogenesis: The creation of net baryon number in the universe
- Leptogenesis: A nickname for “baryogenesis via leptogenesis” where lepton number is generated and this what drives baryogenesis

What about the lepton number of the universe?

Outline

- Very basic cosmology
- The Sakharov's conditions
- CPV
- GUT Baryogenesis
- SM Baryogenesis
- Leptogenesis

Very basic cosmology

The very basic of early cosmology

- Recall statistical mechanics
- The early universe was very hot
- While the universe expands it cools down
- While cooling down, when $T < M$ the equilibrium density of particles reduce exponentially

$$n \propto \exp(-T/M)$$

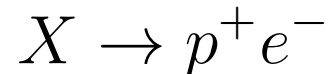
- The actual density does not follow the equilibrium one
- How close it does? It depends on the mass and strength of interactions
- The weaker the interaction, it deviates more from its equilibrium density

The Sakharov's conditions

Sakharov's conditions

Sakharov's conditions for dynamically generated baryon asymmetry

- Baryon number violating process



- C and CP violation

$$\Gamma(X \rightarrow p^+ e^-) \neq \Gamma(\bar{X} \rightarrow p^- e^+)$$

- Deviation from thermal equilibrium

$$\Gamma(X \rightarrow p^+ e^-) \neq \Gamma(p^+ e^- \rightarrow X)$$

Sakharov's conditions: Remarks

- B violation requires for microscopic baryogenesis
 - C and CP violation and out of equilibrium require for macroscopic baryogenesis
 - Out of equilibrium condition: Make sure nothing can bring it back to equilibrium
-

Eugene Wigner on Baryon conservation : “I can feel it in my bones”

- It is a proof that Baryon number is a very good symmetry today
 - In a way, it is also a proof that Baryon number must be broken
-

Total asymmetry

Consider decays of particle X that generate baryons

$$\eta = N_I \epsilon \eta_a$$

- N_I is the initial density of X at $T \gg m_X$
- $\epsilon < 1$ is the CP violation asymmetry. This factor is from particle physics
- $\eta_a < 1$ is an efficiency factor due to “washout” effects. This factor is from cosmology

CPV

What is CP

- A symmetry between a particle and its anti-particle
- CP is violated if we have

$$\Gamma(A \rightarrow B) \neq \Gamma(\bar{A} \rightarrow \bar{B})$$

- It is not easy to detect CPV
 - Always need interference of two (or more) diagrams
 - These two amplitudes must differ in both their weak and strong phases

All these phases

- Weak phase (CP-odd phase) changes sign under CP. It is a phase in \mathcal{L}

$$CP(Ae^{i\phi}) = Ae^{-i\phi}$$

- Strong phase (CP-even phase) is invariant under CP

$$CP(Ae^{i\delta}) = Ae^{i\delta}$$

- Due to time evolution $\psi(t) = e^{-iHt}\psi(0)$ and thus needs intermediate real states
- Check for yourself that to get CPV we need both weak and strong phases

GUT Baryogenesis

GUTs

An example of a way to get Baryogenesis

- I will not get into the full story of GUTs
- For our “story” we need to know that the $SU(3) \times SU(2) \times U(1)$ group of the SM is part of a bigger group, say $SU(5)$
- Baryons and leptons sit in the same representation. In the $SU(5)$ case, 5 and 10
- Breaking of the GUT group results in heavy particles that “break” baryon and lepton numbers

X and Y

- Heavy spin one doublet

$$q(X) = 4/3 \quad q(Y) = 1/3$$

- The couplings give rise to

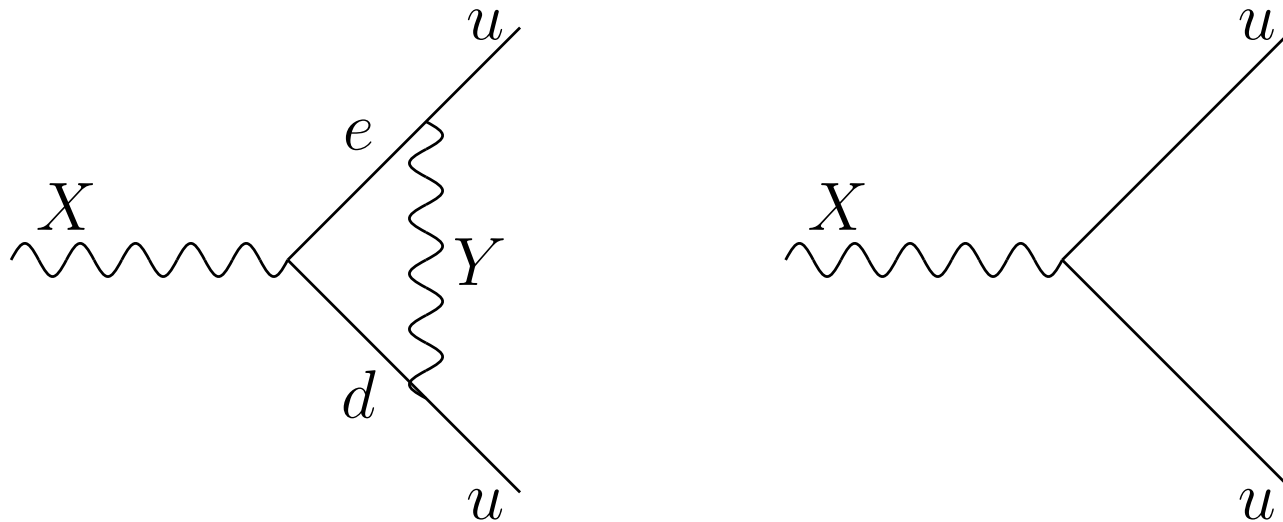
$$X \rightarrow u + u \quad X \rightarrow e^+ + \bar{d}$$

and

$$Y \rightarrow e^+ + \bar{u} \quad Y \rightarrow d + u \quad Y \rightarrow \bar{d} + \bar{\nu}_e$$

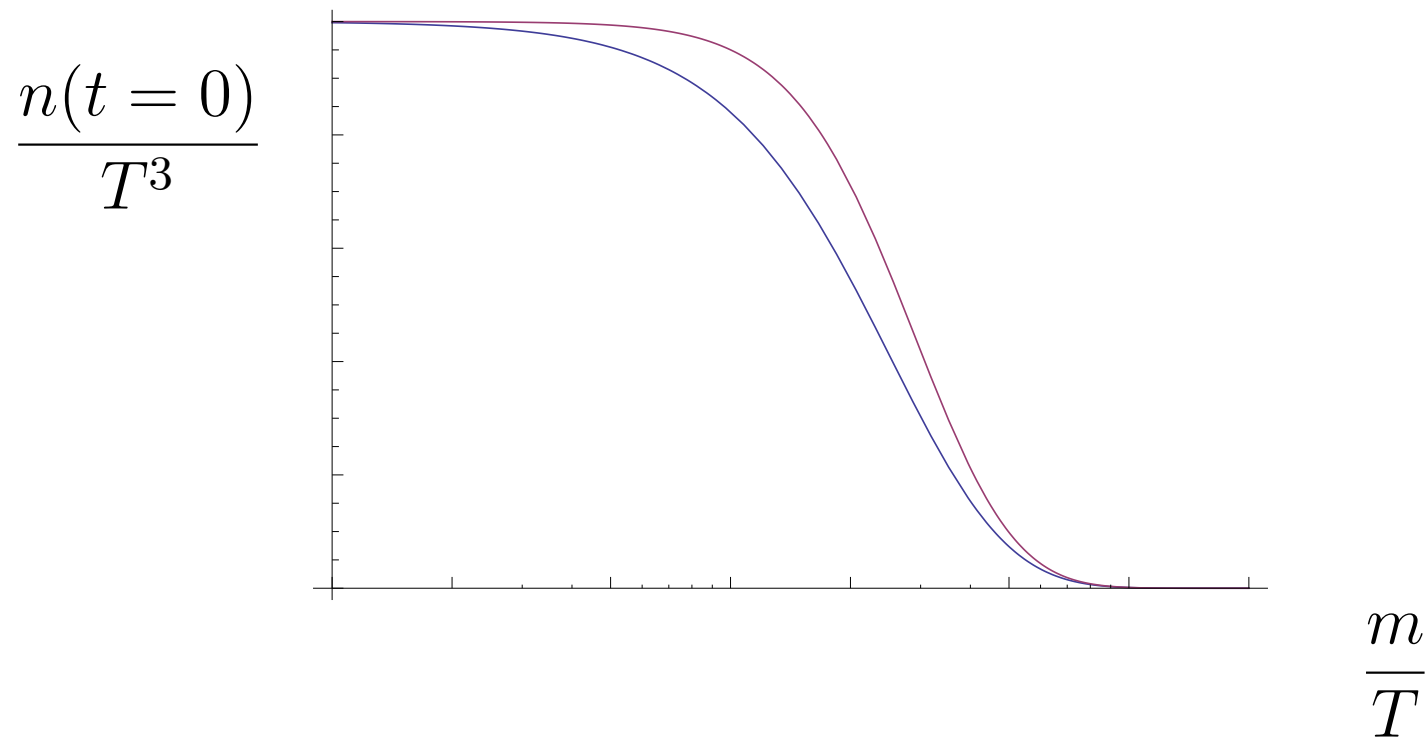
- Decays with different final state baryon number \Rightarrow
 X and Y violate baryon number
- B-L is conserved. X and Y have charge $2/3$ under it

GUT diagrams



- e and d can be on shell \Rightarrow strong phase
- Weak phase from the product of the four vertexes
- Similar diagrams for Y decay

Out of equilibrium



- Blue: equilibrium, red: actual
- Baryogenesis by decays out of equilibrium

SM Baryogenesis

SM baryogenesis

The three Sakharov's conditions are satisfied in the SM

- Baryon number violating process: sphalerons
- The weak interaction violates C and CP
- Out of equilibrium from the electroweak phase transition

The SM, however, is not enough

$$\eta_{\text{SM}} \sim 10^{-25} \ll 10^{-10}$$

Baryon number violation in the SM

- At the classical renormalizable level, baryon and lepton numbers are conserved in the SM
- Non perturbative operators, however, breaks it in a very interesting way
- The processes associate with this breaking are called sphalerons
- Sphalerons conserve $B-L$ and break $B+L$

The very basic of sphalerons

Just the very basic...

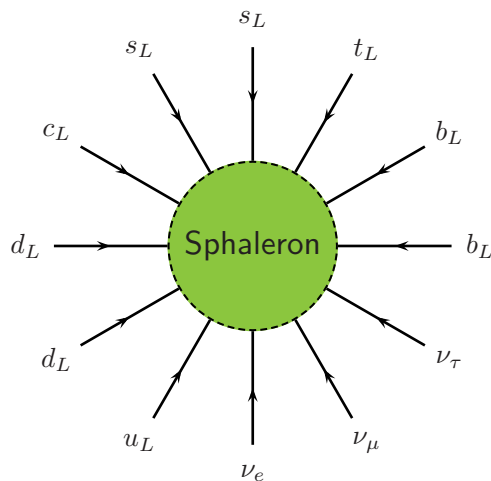
- Related to the chiral anomaly of a non abelian gauge group
- In the SM it is the SU(2)
- Non-perturbative effect. Similar to tunneling between different vacua of the theory
- At $T = 0$, because it is tunneling, it is exponential suppressed and negligible
- At $T \sim v$ no need to tunnel and sphalerons are important
- At $10^2 \lesssim T \lesssim 10^{12}$ GeV the sphalerons are in equilibrium

Sphalerons

Non-perturbative, “tunneling”, effect which involve 3 leptons and 9 quarks

- Can lead to

$$p^+ p^+ \rightarrow p^- e^+ e^+ e^+$$



- The rate depend exponentially on the temperature, T
- Too small to see it today
- Very important in the early universe
- Conserve B–L

Leptogenesis

The SM+N

We know neutrinos have tiny masses

- What is the mechanism that give them mass?
- One option is the see-saw mechanism
- We add to the SM three SM singlets, N_i
- The N_i s acquire large Majorana masses
- The neutrinos masses are then

$$m_\nu \sim \frac{m_W^2}{M_N}$$

SM+N leptogenesis

Leptogenesis in decays of heavy singlet neutrino

- Similar to GUT, here N generate lepton asymmetry
- Sakharov's conditions for generating lepton asymmetry
 - Lepton number violating process: $N \rightarrow H^+ e^-$
 - C and CP violation

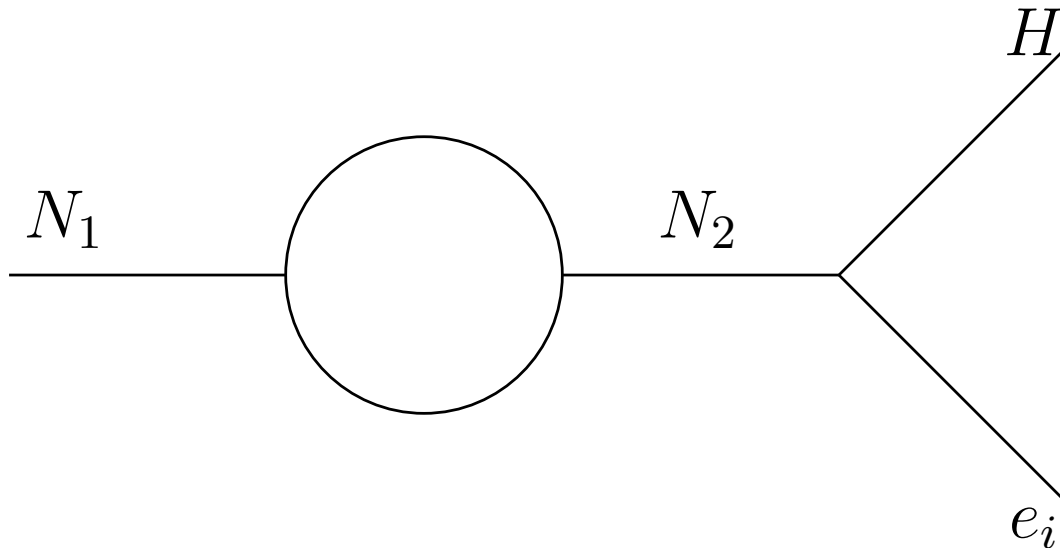
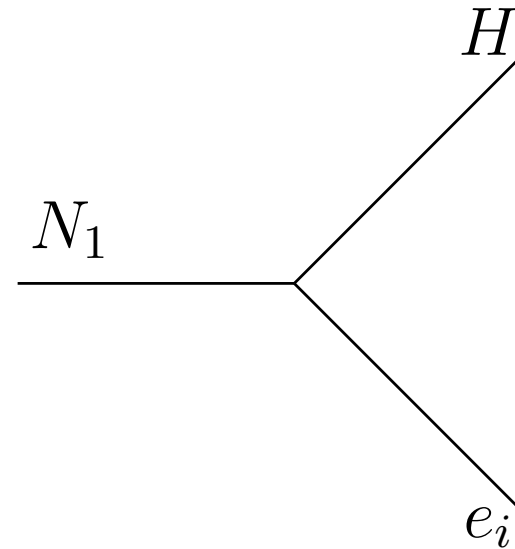
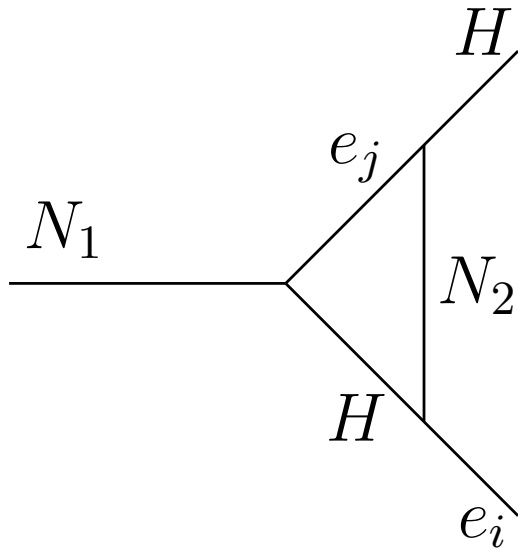
$$\Gamma(N \rightarrow H^+ e^-) \neq \Gamma(N \rightarrow H^- e^+)$$

- Deviation from thermal equilibrium. N decays when it is no more in contact with the thermal bath

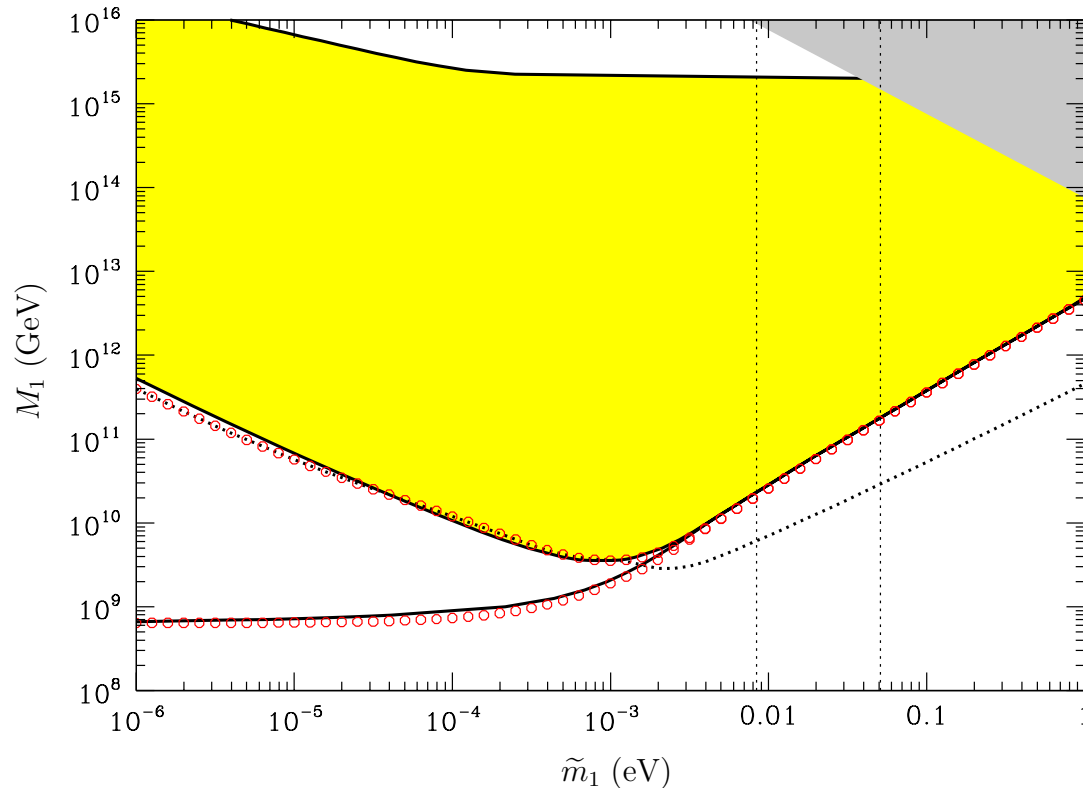
$$\Gamma(N \rightarrow H^+ e^-) \neq \Gamma(H^+ e^- \rightarrow N)$$

- Violates B–L. Sphalerons convert some L into B

Diagrams



Numerical prediction



- The right side of the plot is “theoretically” motivated
- The fact that both leptogenesis and oscillation data are consistent is not a trivial test

Tests of leptogenesis

- It is not easy to directly test leptogenesis as the heavy neutrino is too heavy, $m_N > 10^{11}$ GeV
- Leptogenesis predicts very small lepton asymmetry in the universe. Very hard to even detect the neutrino background
- Since leptogenesis requires CP violation, we like to find CP violation also in neutrino oscillation
- Majorana mass for the neutrinos can be probed with neutrinoless double beta decay
- The big “issue” with LG is that it is not easy to test
- Yet it passed one non trivial test!

Conclusions

Conclusions

- Baryogenesis is an open question in physics
- Many ideas, but no conclusion yet
- It will be so nice if we find more ways to find the answer

