

XENON1T latest results

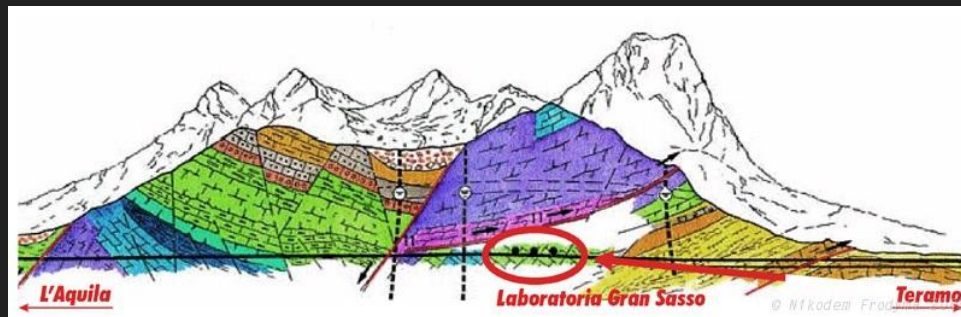
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XENON

“The XENON dark matter project aims at the detection of WIMP dark matter with dual-phase time projection chambers (TPCs), that are filled with a liquid xenon (LXe) target of increasing mass.”



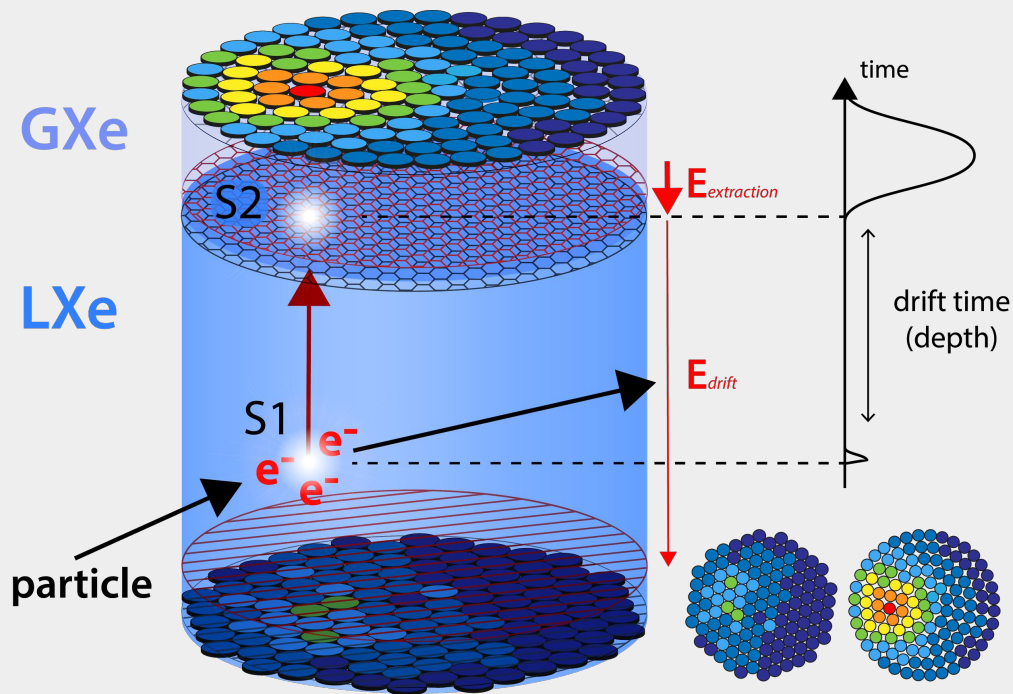
Detecting WIMPs

Dual phase TPCs allow detection of recoils of $O(10 \text{ keV})$ and provide:

- Discrimination between electronic and nuclear recoils.
- 3D position reconstruction for fiducialization.

XENON1T:

- ~2000 kg of ultra pure liquid xenon target.
- 248 PMTs

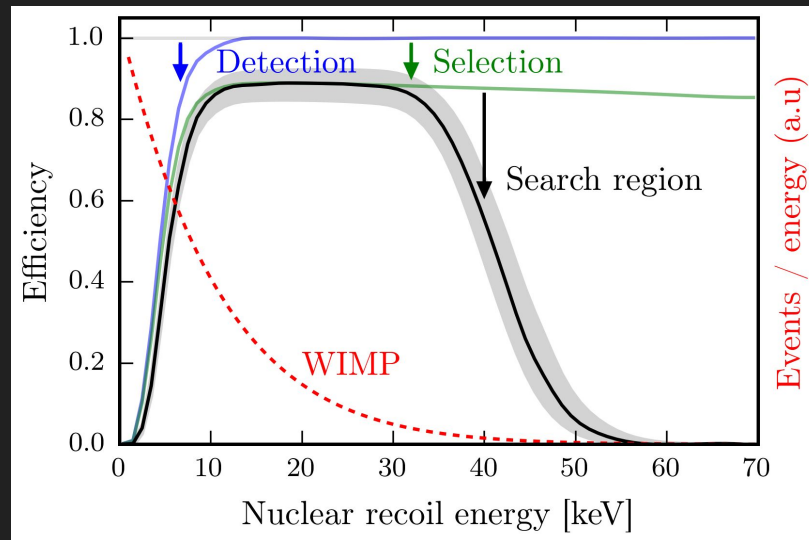


Data Selection

The first science run (SR0, from 22-11-2016 to 18-01-2017) uses 34.2 live days of blinded data.

Select single scatters in the LXe:

- 3-fold coincidence requirement
- Selection efficiencies estimated from control samples or simulation
- Search region defined within 3-70 PE in cS1



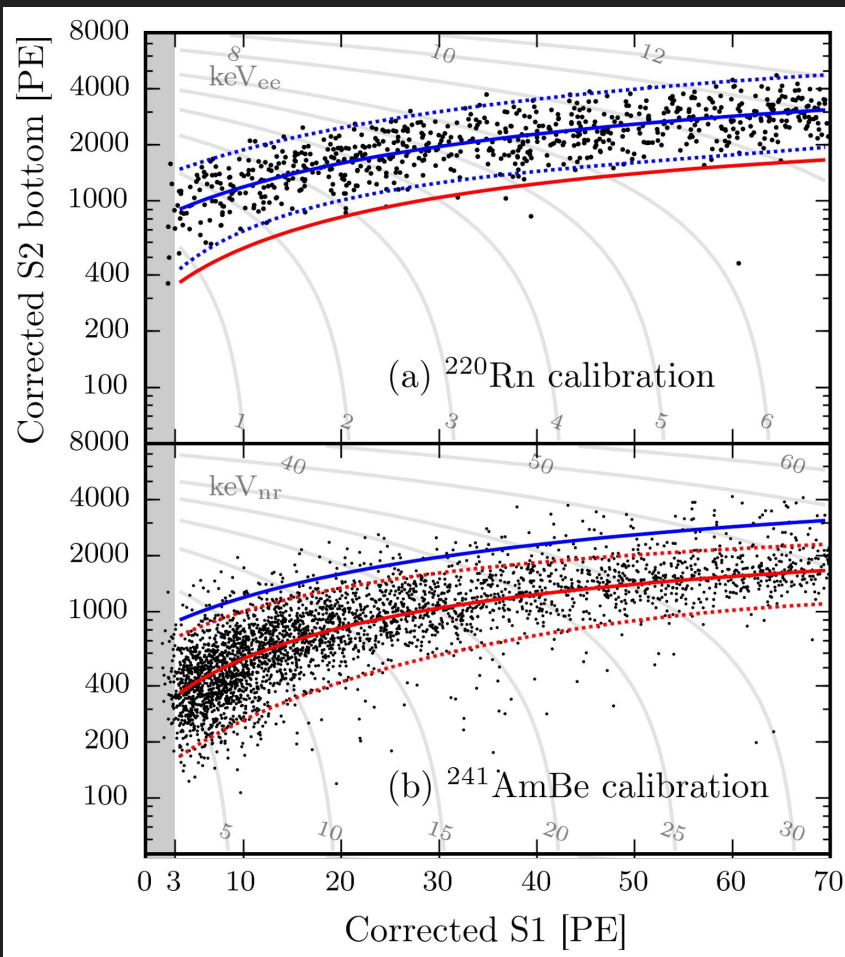
[arXiv:1705.06655](https://arxiv.org/abs/1705.06655)

Detector Calibration

Calibration data:

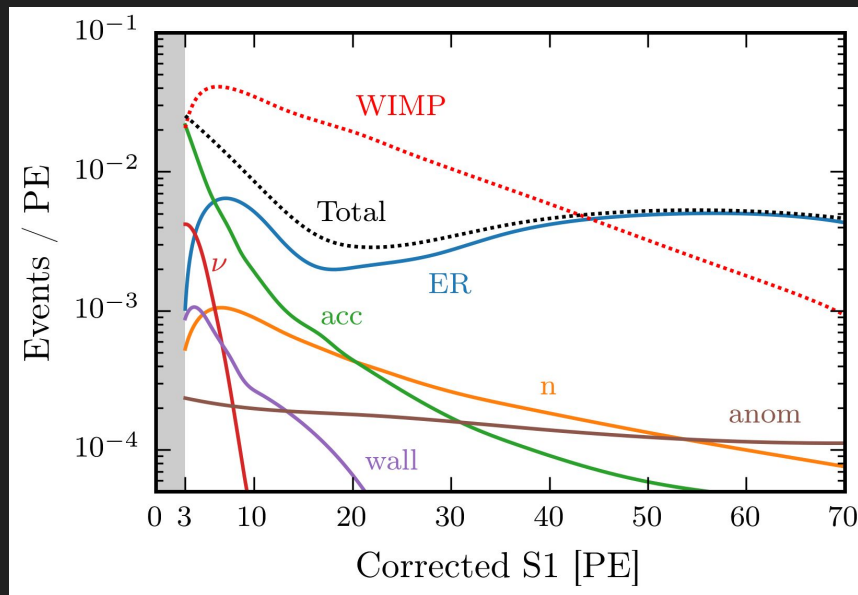
- Kr83m, 3.3 days (electronic recoils)
- Rn220, 3.0 days (electronic recoils)
- AmBe241, 16.3 days (nuclear recoils)
- LED data

The data is fitted with a physics motivated model incorporating many parameters. The two most important ones (efficiency and exciton-to-ion ratio) are propagated to the statistical framework.



Backgrounds

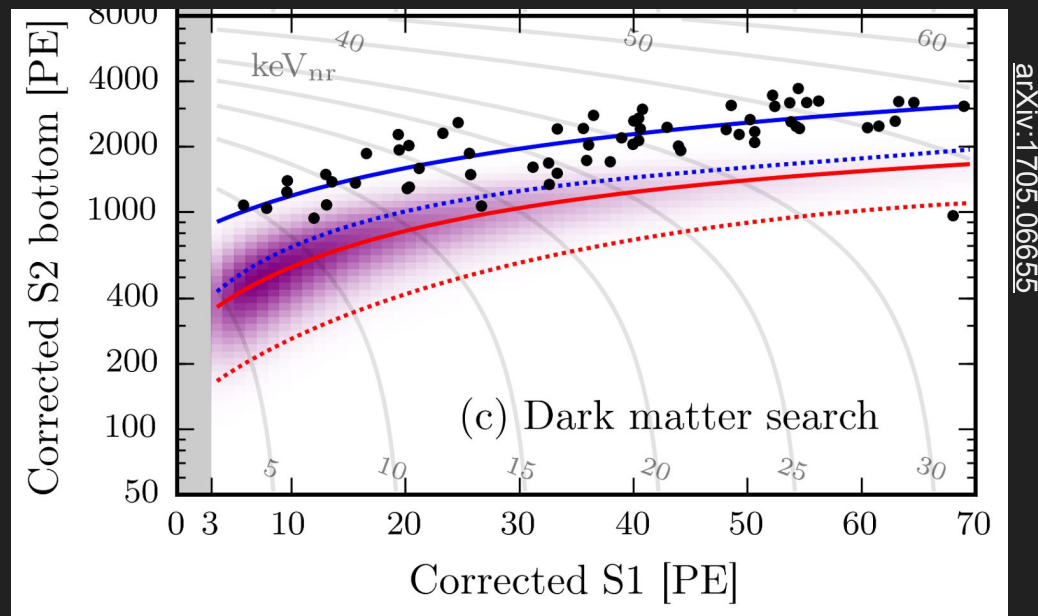
	Full
Electronic recoils (ER)	(62 ± 8)
Radiogenic neutrons (n)	0.05 ± 0.01
CNNS (ν)	0.02
Accidental coincidences (acc)	0.22 ± 0.01
Wall leakage ($wall$)	0.5 ± 0.3
Anomalous ($anom$)	$0.10^{+0.10}_{-0.07}$
Total background	63 ± 8
$50 \text{ GeV}/c^2, 10^{-46} \text{ cm}^2 \text{ WIMP}$	1.66 ± 0.01



Results

63 events pass selection in a 1T fiducial volume, compatible with background only hypothesis

Electronic recoil background rate of $(1.93 \pm 0.25) \times 10^{-4}$ events/(kg day keV_{ee}), the lowest in this type of detector.

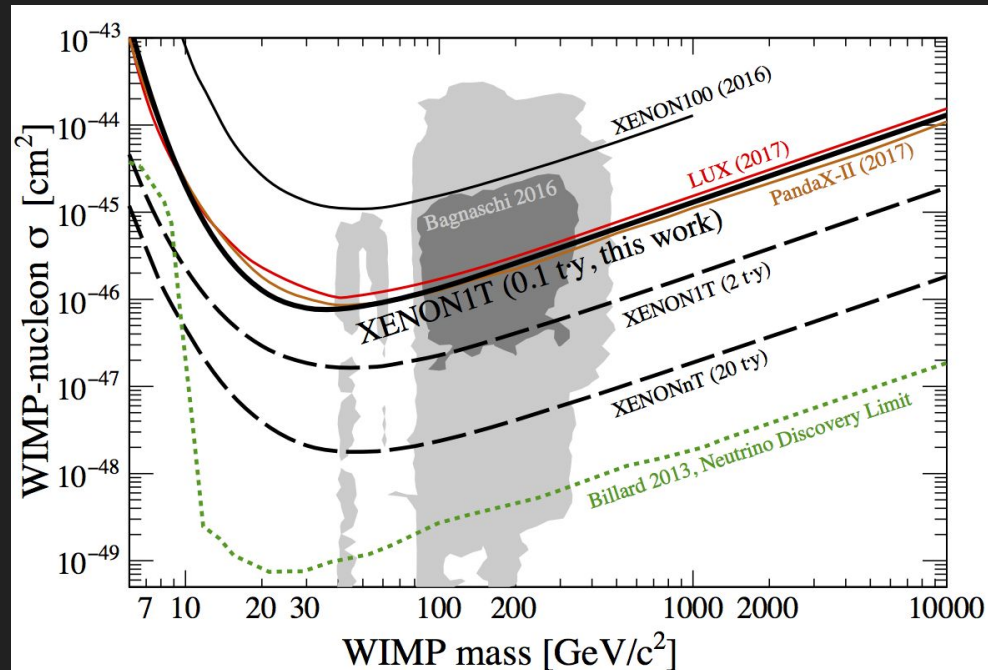


The Exclusion Limit

Using an extended unbinned profile likelihood.

90% limit on the spin independent WIMP-nucleon cross section.

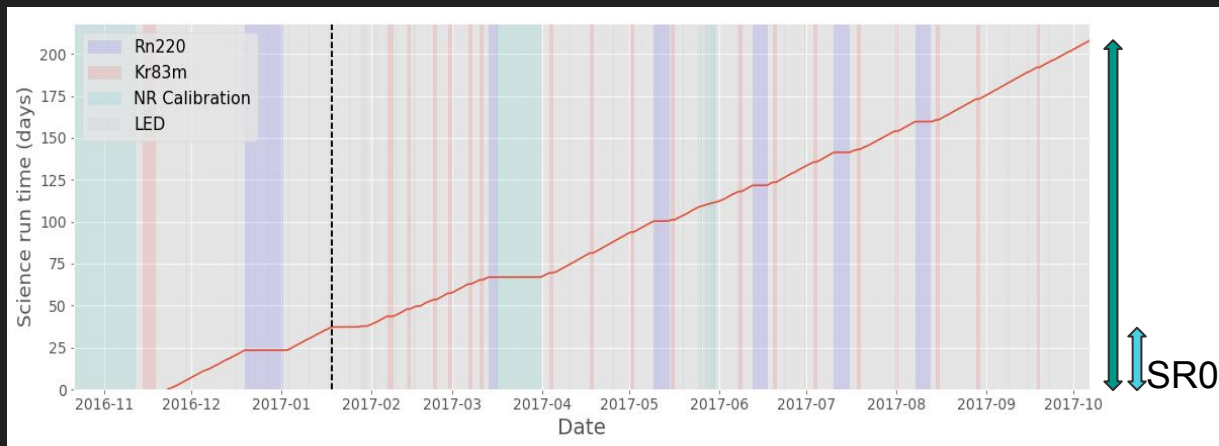
Strongest exclusion for 35 GeV/c^2 WIMPs at $7.7 \times 10^{-47} \text{ cm}^2$



pMSSM10 theory shaded

Future Analyses

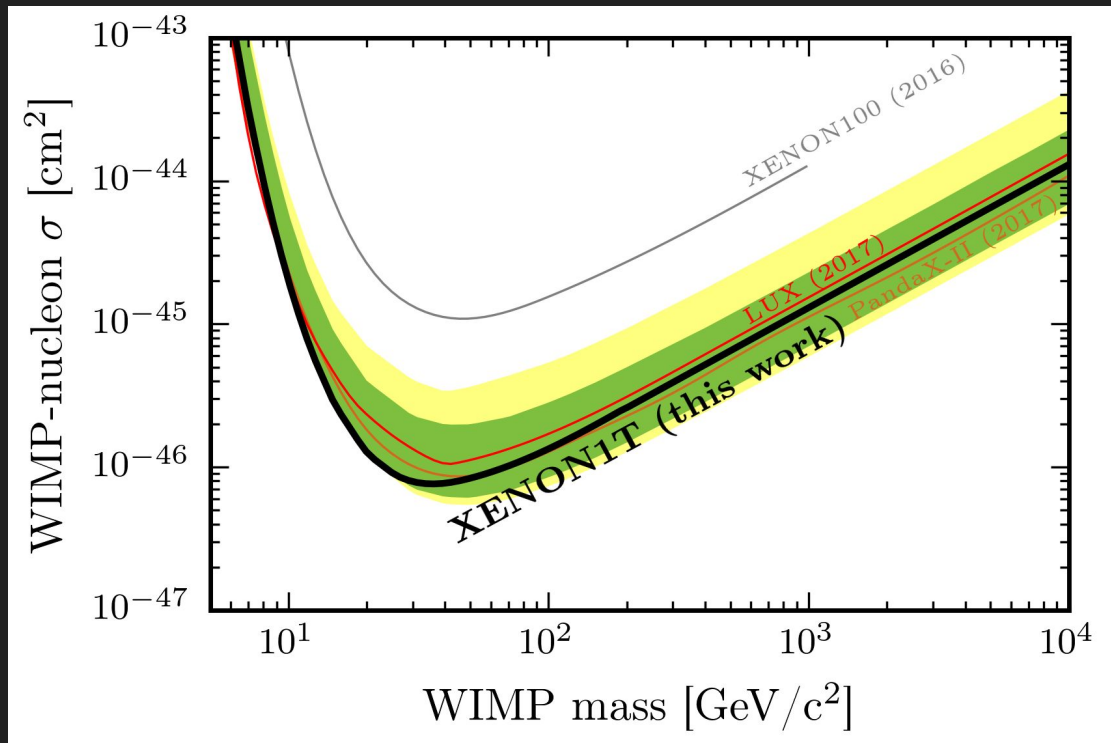
- XENON1T is still taking data, the next analysis is underway and will have a much larger exposure.
- XENONnT (volume ~8000 kg) is being built to replace XENON1T.



Conclusions

- XENON1T is the world's most sensitive dark matter detector of its kind.
- It has excluded WIMPs down to $7.7 \times 10^{-47} \text{ cm}^2$, (90% UL at 35 GeV/c²)
- XENON1T is still taking data and the next analysis is ongoing with a much larger exposure.
- XENONnT is being built to upgrade XENON1T and take over the search for WIMP dark matter.

Bonus Material



Why use liquid xenon?

- Intrinsic scintillator (178 nm, VUV)
 - Transparent
- Heavy (~ 125 GeV)
 - Elastic collisions with WIMPs of mass $O(100)$ GeV
- Dense (2.8 kg/L at -90 degC)
 - Provides self-shielding
- Stable, no long lived isotopes
- Electronic/Nuclear Recoil discrimination
- Allows for both spin-dependent and spin independent measurements.

