

SUSY 07 – July 30, 2007

Christoph Luhn (University of Florida)

in collaboration with Herbi Dreiner, Hitoshi Murayama, and Marc Thormeier

**Proton hexality and neutrino masses
from an anomalous $U(1)$**

arXiv:0708.xxxx [hep-ph]

The supersymmetrized SM

renormalizable superpotential

$$W_{\text{ren.}} = \mu H_u H_d + h_{ij}^u H_u Q_i \bar{U}_j + h_{ij}^d H_d Q_i \bar{D}_j + h_{ij}^e H_d L_i \bar{E}_j \\ + \kappa_i H_u L_i + \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

non-renormalizable terms:

$$W_{\text{non-ren.}} = \frac{c_{ijkl}}{M_{\text{grav}}} Q_i Q_j Q_k L_l + \frac{c'_{ijkl}}{M_{\text{grav}}} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

Overview

extended gauge group: $\text{SM} \times U(1)_X$

anomalous $U(1)_X \longrightarrow$ broken spontaneously slightly below M_{grav}
(Dine-Seiberg-Wen-Witten mechanism)

proton-decay

$$U(1) \supset \mathbb{Z}_N$$

proton hexality

1.

family structure

Froggatt-Nielsen
mechanism

(also for neutrinos)

2.

μ -term

Giudice-Masiero
mechanism

anomalies

Green-Schwarz
mechanism

1. Proton hexality

Proton-decay and discrete symmetries

- $LQ\bar{D}$ and $\bar{U}\bar{D}\bar{D}$ \longrightarrow rapid p -decay
- introduction of matter parity (M_p)
- “dirty little secret”: $\frac{1}{M}QQQL \longrightarrow M \gtrsim 10^8 \cdot M_{\text{grav}}$
- use proton hexality (P_6) instead

	Q	\bar{U}	\bar{D}	L	\bar{E}	H_u	H_d
matter parity (M_p)	0	1	1	0	1	1	1
proton hexality (P_6)	0	1	5	4	1	5	1

$$M_p[QQQL] = 0 \quad M_p[\bar{U}\bar{U}\bar{D}\bar{E}] = 4 = 0 \pmod{2}$$

$$P_6[QQQL] = 4 \quad P_6[\bar{U}\bar{U}\bar{D}\bar{E}] = 8 = 2 \pmod{6}$$

Proton hexality

$$P_6 \cong M_p \times B_3$$

- $M_p \longrightarrow$ stable LSP
- $B_3 \longrightarrow$ stable proton

Proton hexality combines the attractive features of M_p and B_3

$$(M_p\text{-MSSM} \rightarrow P_6\text{-MSSM})$$

Origin of the discrete symmetry P_6

$U(1)_X$ breaking (flavon) field A : $\langle A \rangle \neq 0$, $X_A = -1$

$$U(1)_X \rightarrow P_6 \quad \boxed{X_\phi = \frac{q_\phi}{6} + s_\phi} \quad \begin{array}{l} q_\phi, s_\phi \in \mathbb{Z} \\ q_\phi = \text{discrete charge for } P_6 \end{array}$$

all P_6 **conserving** operators \longrightarrow **integer** overall X -charge

all P_6 **violating** operators \longrightarrow **fractional** overall X -charge

necessary and sufficient condition:

$$\boxed{X_{H_d} - X_{L_1} = \frac{1}{2} + \mathbb{Z} \quad \wedge \quad 3X_{Q_1} + X_{L_1} = \frac{1 \text{ or } 2}{3} + \mathbb{Z}}$$

tacit assumption: $\exists H_u H_d, H_u Q \bar{U}, H_d Q \bar{D}, H_d L \bar{E}$ effectively

2. Family structure

Family structure

- Froggatt-Nielsen: • hierarchy of masses and CKM structure due to family dependent X -charges
- $\mathcal{O}(1)$ coefficients and phases remain undetermined

q_ϕ family independent \rightsquigarrow discrete symmetry

↑

$$\boxed{X_\phi = \frac{q_\phi}{N} + s_\phi} \longrightarrow s_\phi \text{ family dependent } \rightsquigarrow \text{ family structure}$$

$$\left. \begin{array}{l} H_u Q_i \bar{U}_j \left(\frac{A}{M_{\text{grav}}} \right)^{X_{H_u} + X_{Q_i} + X_{\bar{U}_j}} \\ \text{with } \frac{\langle A \rangle}{M_{\text{grav}}} = \epsilon \approx \lambda_c \end{array} \right\} m_{ij}^{(u)} \sim \langle H_u \rangle \cdot \begin{pmatrix} \lambda_c^8 & \lambda_c^5 & \lambda_c^3 \\ \lambda_c^7 & \lambda_c^4 & \lambda_c^2 \\ \lambda_c^5 & \lambda_c^2 & 1 \end{pmatrix}_{ij}$$

λ_c = Wolfenstein parameter
Dine-Seiberg-Wen-Witten

$$m_u : m_c : m_t \sim \lambda_c^8 : \lambda_c^4 : 1$$

Origin of neutrino masses

without right-handed neutrinos

- $LH_u, LQ\bar{D}, LL\bar{E}$ - forbidden by P_6 ☹️
- $\frac{1}{M_{\text{grav}}} LH_u LH_u$ - too small ☹️

with right-handed neutrinos

- choose $q_{\bar{N}}$ so that \bar{N} couples to LH_u : $\longrightarrow LH_u\bar{N}$ and $\bar{N}\bar{N}$

- see-saw:
$$\begin{pmatrix} 0 & M_D \\ M_D^T & M_M \end{pmatrix} \longrightarrow M_\nu = M_D \cdot M_M^{-1} \cdot M_D^T$$

$$X_{L_i} + X_{H_u} + X_{\bar{N}_j} \geq 0 \quad M_D^{ij} \sim \langle H_u \rangle \cdot \epsilon^{X_{L_i} + X_{H_u} + X_{\bar{N}_j}}$$

$$X_{\bar{N}_i} + X_{\bar{N}_j} \geq 0 \quad M_M^{ij} \sim M_{\text{grav}} \cdot \epsilon^{X_{\bar{N}_i} + X_{\bar{N}_j}} \quad \longrightarrow \text{Case I}$$

$$X_{\bar{N}_i} + X_{\bar{N}_j} < 0 \quad M_M^{ij} \sim m_{\text{soft}} \cdot \epsilon^{-X_{\bar{N}_i} - X_{\bar{N}_j}} \quad \longrightarrow \text{Case II}$$

RESULTS

$U(1)_X$ charges of MSSM particles constrained by:

proton-hexality	}	48+504
family structure		X -charge assignments
Giudice-Masiero for μ -term		for Case I + II
Green-Schwarz anomaly cancellation		

in addition:

- no X -charged hidden sector fields \longrightarrow 48+24 models survive
- anomaly condition determine Kač-Moody level k_C of $SU(3)_C$, thus g_{string}

Enjoy your ...

