

ENERGY TRANSFER AND KALUZA-KLEIN MODE DECAY BETWEEN THROATS

Benedict v. Harling

Institute for Theoretical Physics
University of Heidelberg
Germany

SUSY 2007

Based on work
with A. Hebecker
and T. Noguchi.
arXiv:0705.3648

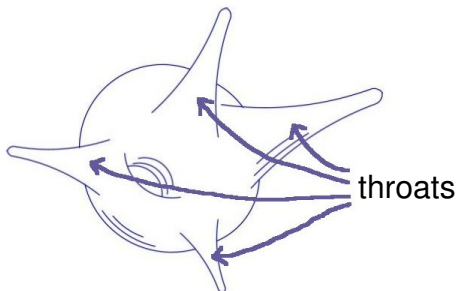
INTRODUCTION

- Flux compactifications of type IIB string theory have nice features, e.g. moduli stabilization.
- Backreaction of fluxes can create strongly warped regions
⇒ throats.
- Throats are common in the string theory landscape, see e.g. HEBECKER, MARCH-RUSSELL '06 on the distribution of Klebanov-Strassler throats.



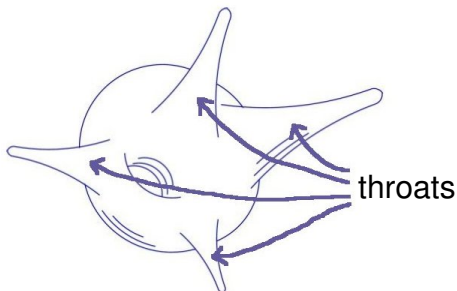
INTRODUCTION

- Flux compactifications of type IIB string theory have nice features, e.g. moduli stabilization.
- Backreaction of fluxes can create strongly warped regions
⇒ **throats**.
- **Throats are common in the string theory landscape**, see e.g. HEBECKER, MARCH-RUSSELL '06 on the distribution of Klebanov-Strassler throats.



INTRODUCTION

- Flux compactifications of type IIB string theory have nice features, e.g. moduli stabilization.
- Backreaction of fluxes can create strongly warped regions
⇒ **throats**.
- **Throats are common in the string theory landscape**, see e.g. HEBECKER, MARCH-RUSSELL '06 on the distribution of Klebanov-Strassler throats.



INTRODUCTION

- **Throats may be heated up in the early universe**, e.g. due to
 - ▶ brane-antibrane annihilation after inflation à la KKLMNT.
 - ▶ energy flow from the heated standard model sector.
- Interesting for cosmology: **Energy transfer rate** from a heated throat to other throats.



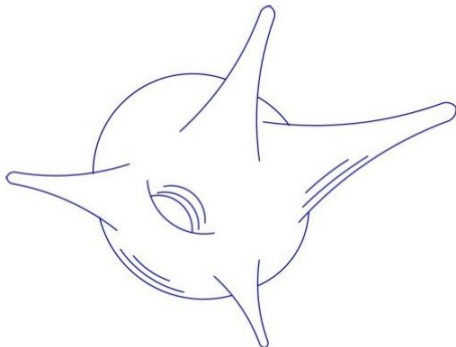
INTRODUCTION

- **Throats may be heated up in the early universe**, e.g. due to
 - ▶ brane-antibrane annihilation after inflation à la KKLMNT.
 - ▶ energy flow from the heated standard model sector.
- Interesting for cosmology: **Energy transfer rate** from a heated throat to other throats.



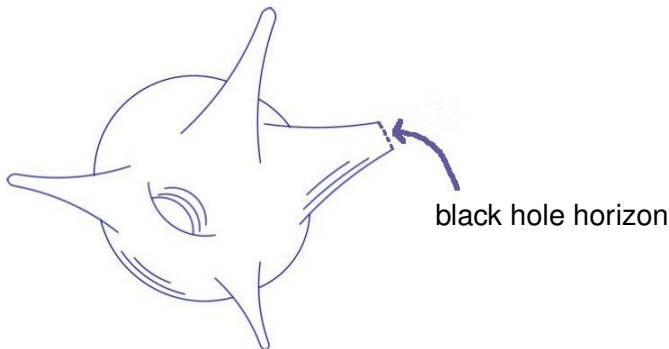
INTRODUCTION

- **Throats may be heated up in the early universe**, e.g. due to
 - ▶ brane-antibrane annihilation after inflation à la KKLMNT.
 - ▶ energy flow from the heated standard model sector.
- Interesting for cosmology: **Energy transfer rate** from a heated throat to other throats.



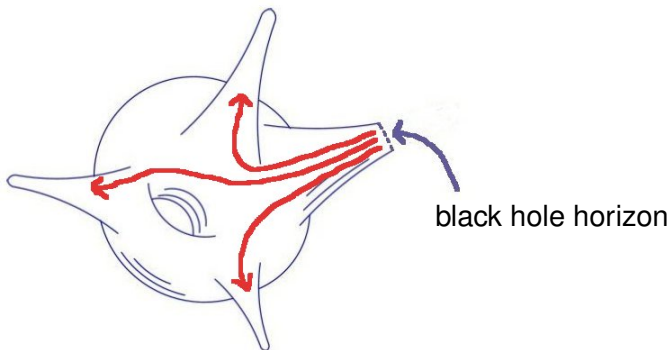
INTRODUCTION

- **Throats may be heated up in the early universe**, e.g. due to
 - ▶ brane-antibrane annihilation after inflation à la KKLMNT.
 - ▶ energy flow from the heated standard model sector.
- Interesting for cosmology: **Energy transfer rate** from a heated throat to other throats.

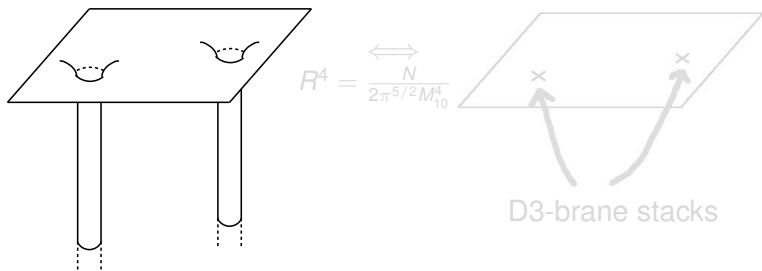


INTRODUCTION

- **Throats may be heated up in the early universe**, e.g. due to
 - ▶ brane-antibrane annihilation after inflation à la KKLMNT.
 - ▶ energy flow from the heated standard model sector.
- Interesting for cosmology: **Energy transfer rate** from a heated throat to other throats.

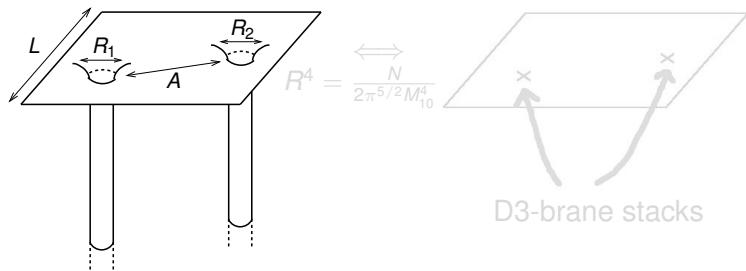


THE CALCULATION



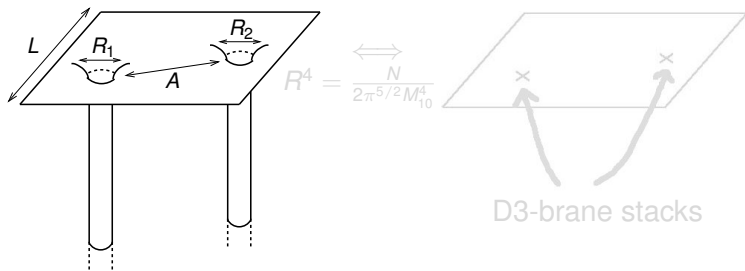
- Simple setup: Two AdS₅ × S⁵ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation ⇒ **multi-dimensional tunneling problem** ⇒ **difficult**.
- Way out: AdS₅ × S⁵ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes**.

THE CALCULATION



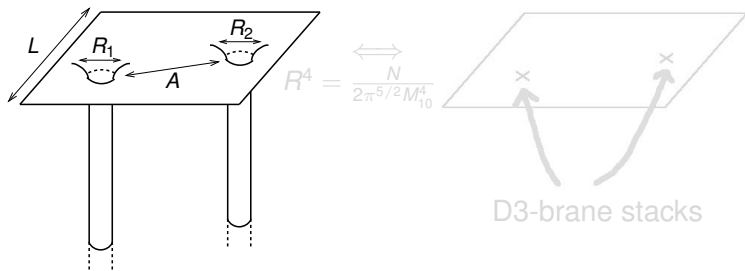
- Simple setup: Two $\text{AdS}_5 \times S^5$ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation \Rightarrow **multi-dimensional tunneling problem** \Rightarrow **difficult**.
- Way out: $\text{AdS}_5 \times S^5$ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes**.

THE CALCULATION



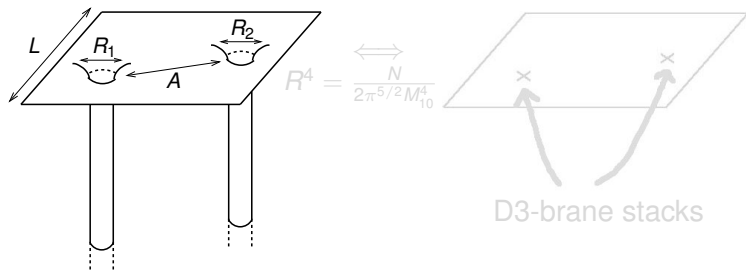
- Simple setup: Two $\text{AdS}_5 \times S^5$ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation \Rightarrow multi-dimensional tunneling problem \Rightarrow difficult.
- Way out: $\text{AdS}_5 \times S^5$ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes.**

THE CALCULATION



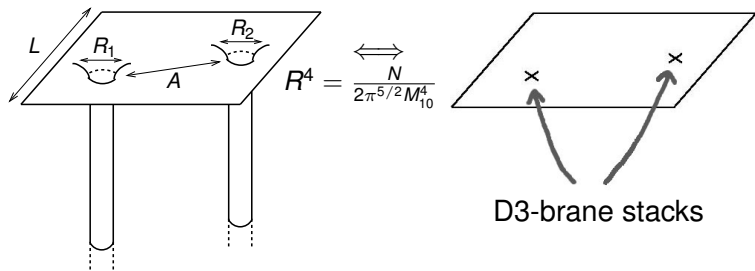
- Simple setup: Two $\text{AdS}_5 \times S^5$ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation \Rightarrow **multi-dimensional tunneling problem** \Rightarrow **difficult**.
- Way out: $\text{AdS}_5 \times S^5$ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes**.

THE CALCULATION

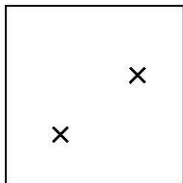


- Simple setup: Two $\text{AdS}_5 \times S^5$ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation \Rightarrow **multi-dimensional tunneling problem** \Rightarrow **difficult**.
- Way out: $\text{AdS}_5 \times S^5$ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes**.

THE CALCULATION



- Simple setup: Two $\text{AdS}_5 \times S^5$ throats embedded in a 6d torus.
- Energy transfer rate determined by **transition probability of Hawking radiation** between the two throats.
- Restrict to the dilaton. Its equation of motion can be written in form of a Schrödinger equation \Rightarrow **multi-dimensional tunneling problem** \Rightarrow **difficult**.
- Way out: $\text{AdS}_5 \times S^5$ throats are near-horizon geometry of black 3-branes, **equivalent to stack of D3-branes**.



- World-volume theory on stack of N D3-branes is a $\mathcal{N} = 4$ $U(N)$ super Yang-Mills.
- The two brane stacks are coupled by the supergravity fields in the embedding space.
- Gauge theories are strongly coupled, but relevant process can be described by a weak-coupling calculation (KLEBANOV '97).

$$\sim \frac{1}{M_{10}^4} \phi \text{tr} F_{\alpha\beta}^2$$

- World-volume theory on stack of N D3-branes is a $\mathcal{N} = 4$ $U(N)$ super Yang-Mills.
- **The two brane stacks are coupled by the supergravity fields in the embedding space.**
- Gauge theories are strongly coupled, but relevant process can be described by a weak-coupling calculation (KLEBANOV '97).

$$\sim \frac{1}{M_{10}^4} \phi \text{tr} F_{\alpha\beta}^2$$

- World-volume theory on stack of N D3-branes is a $\mathcal{N} = 4$ $U(N)$ super Yang-Mills.
- The two brane stacks are coupled by the supergravity fields in the embedding space.
- Gauge theories are strongly coupled, but relevant process can be described by a weak-coupling calculation (KLEBANOV '97).

THE CALCULATION

- KK expansion of the dilaton \Rightarrow from 4d viewpoint there are two gauge theories coupled by tower of KK modes.
- Heated throat corresponds to heated gauge theory. Energy transfer due to processes of the type



- Thermally averaging the above process gives energy transfer rate (per 4d volume):

$$\dot{\rho} \sim \frac{R_1^8 R_2^8}{A^8} T^{13} + \frac{R_1^8 R_2^8}{L^{12}} T^9.$$

- KK expansion of the dilaton \Rightarrow from 4d viewpoint there are two gauge theories coupled by tower of KK modes.
- Heated throat corresponds to heated gauge theory. Energy transfer due to processes of the type

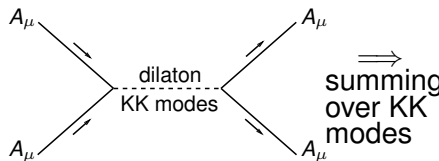
$$\mathcal{M} \sim \frac{s^2}{M_{10}^8 L^6} \sum_{\vec{n} \in \mathbb{Z}^6} \frac{e^{2\pi i \vec{n} \vec{A}/L}}{s - m_{\vec{n}}^2 + i\epsilon}$$

$$\sim \frac{s^2}{M_{10}^8 A^4} + \frac{s}{M_{10}^8 L^6}.$$

- Thermally averaging the above process gives energy transfer rate (per 4d volume):

$$\dot{\rho} \sim \frac{R_1^8 R_2^8}{A^8} T^{13} + \frac{R_1^8 R_2^8}{L^{12}} T^9.$$

- KK expansion of the dilaton \Rightarrow from 4d viewpoint there are two gauge theories coupled by tower of KK modes.
- Heated throat corresponds to heated gauge theory. Energy transfer due to processes of the type



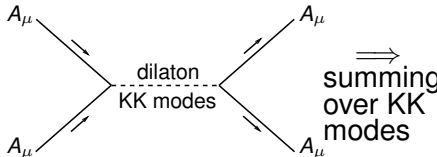
$$\mathcal{M} \sim \frac{s^2}{M_{10}^8 L^6} \sum_{\vec{n} \in \mathbb{Z}^6} \frac{e^{2\pi i \vec{n} \vec{A}/L}}{s - m_{\vec{n}}^2 + i\epsilon}$$

$$\sim \frac{s^2}{M_{10}^8 A^4} + \frac{s}{M_{10}^8 L^6}.$$

- Thermally averaging the above process gives energy transfer rate (per 4d volume):

$$\dot{\rho} \sim \frac{R_1^8 R_2^8}{A^8} T^{13} + \frac{R_1^8 R_2^8}{L^{12}} T^9.$$

- KK expansion of the dilaton \Rightarrow from 4d viewpoint there are two gauge theories coupled by tower of KK modes.
- Heated throat corresponds to heated gauge theory. Energy transfer due to processes of the type



$$\mathcal{M} \sim \frac{s^2}{M_{10}^8 L^6} \sum_{\vec{n} \in \mathbb{Z}^6} \frac{e^{2\pi i \vec{n} \vec{A}/L}}{s - m_{\vec{n}}^2 + i\epsilon}$$

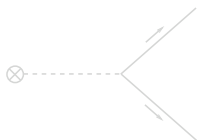
$$\sim \frac{s^2}{M_{10}^8 A^4} + \frac{s}{M_{10}^8 L^6}.$$

- Thermally averaging the above process gives energy transfer rate (per 4d volume):

$$\dot{\rho} \sim \frac{R_1^8 R_2^8}{A^8} T^{13} + \frac{R_1^8 R_2^8}{L^{12}} T^9.$$

KK MODE DECAY RATE

- When throat temperature drops below IR scale \Rightarrow black hole horizon is replaced by IR cutoff region \Rightarrow from 4d viewpoint there is a non-relativistic gas of **throat-localized KK modes**.
- KK modes localized in one throat can decay to other throats \Rightarrow **decay rate**.
- **Throat-localized KK mode is dual to glueball state on equivalent stack of D-branes**. Decay rate given by the process:



- **Glueball-dilaton vertex** can be determined as follows: Calculate decay rate for a simpler setup in the gravity picture. Then match the vertex such that the decay rate is reproduced in the gauge theory picture.

KK MODE DECAY RATE

- When throat temperature drops below IR scale \Rightarrow black hole horizon is replaced by IR cutoff region \Rightarrow from 4d viewpoint there is a non-relativistic gas of **throat-localized KK modes**.
- KK modes localized in one throat can decay to other throats \Rightarrow **decay rate**.
- **Throat-localized KK mode is dual to glueball state on equivalent stack of D-branes**. Decay rate given by the process:



- **Glueball-dilaton vertex** can be determined as follows: Calculate decay rate for a simpler setup in the gravity picture. Then match the vertex such that the decay rate is reproduced in the gauge theory picture.

KK MODE DECAY RATE

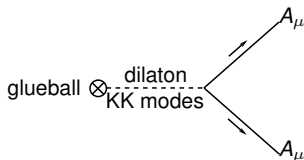
- When throat temperature drops below IR scale \Rightarrow black hole horizon is replaced by IR cutoff region \Rightarrow from 4d viewpoint there is a non-relativistic gas of **throat-localized KK modes**.
- KK modes localized in one throat can decay to other throats \Rightarrow **decay rate**.
- Throat-localized KK mode is dual to glueball state on equivalent stack of D-branes. Decay rate given by the process:



- **Glueball-dilaton vertex** can be determined as follows: Calculate decay rate for a simpler setup in the gravity picture. Then match the vertex such that the decay rate is reproduced in the gauge theory picture.

KK MODE DECAY RATE

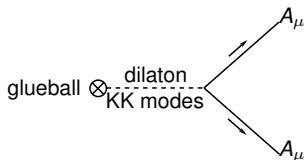
- When throat temperature drops below IR scale \Rightarrow black hole horizon is replaced by IR cutoff region \Rightarrow from 4d viewpoint there is a non-relativistic gas of **throat-localized KK modes**.
- KK modes localized in one throat can decay to other throats \Rightarrow **decay rate**.
- **Throat-localized KK mode is dual to glueball state on equivalent stack of D-branes**. Decay rate given by the process:



- **Glueball-dilaton vertex** can be determined as follows: Calculate decay rate for a simpler setup in the gravity picture. Then match the vertex such that the decay rate is reproduced in the gauge theory picture.

KK MODE DECAY RATE

- When throat temperature drops below IR scale \Rightarrow black hole horizon is replaced by IR cutoff region \Rightarrow from 4d viewpoint there is a non-relativistic gas of **throat-localized KK modes**.
- KK modes localized in one throat can decay to other throats \Rightarrow **decay rate**.
- **Throat-localized KK mode is dual to glueball state on equivalent stack of D-branes**. Decay rate given by the process:



- **Glueball-dilaton vertex** can be determined as follows: Calculate decay rate for a simpler setup in the gravity picture. Then match the vertex such that the decay rate is reproduced in the gauge theory picture.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- In which case is a 5d model a good approximation? If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, we get the same result:

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAH AND TYE '05, '06 only in special case.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- In which case is a 5d model a good approximation? If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, we get the same result:

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAHI AND TYE '05, '06 only in special case.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- **In which case is a 5d model a good approximation?** If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, **we get the same result:**

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAH AND TYE '05, '06 only in special case.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- **In which case is a 5d model a good approximation?** If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, **we get the same result:**

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAHI AND TYE '05, '06 only in special case.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- **In which case is a 5d model a good approximation?** If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, **we get the same result:**

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAHI AND TYE '05, '06 only in special case.

DECAY RATE/COMPARISON WITH LITERATURE

- Same steps as before (ℓ angular quantum number, m_{IR} IR scale):

$$\Rightarrow \Gamma \sim \frac{R_1^{8+4\ell} R_2^8}{A^8} m_{\text{IR}} m^{8+4\ell} + \frac{R_1^{8+4\ell} R_2^8}{L^{12}} m_{\text{IR}} m^{4+4\ell}.$$

- From a 5d model with $R_1 = R_2$ and $\ell = 0$, DIMOPOULOS ET AL. '01 get

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- **In which case is a 5d model a good approximation?** If the compact embedding manifold is very small. Minimal size $L \sim R_1, R_2$, since otherwise throats could not be glued into compact manifold.
- Taking $R_1 = R_2 \sim L \sim A$ and $\ell = 0$, **we get the same result:**

$$\Gamma \sim (mR)^4 m_{\text{IR}}.$$

- We find agreement with a result by CHEN, FIROUZJAH AND TYE '05, '06 only in special case.

- We have determined **energy loss rate** of heated throat and **decay rate** of throat-localized KK modes.
- Genuine 10d calculation using the dual gauge theory picture. Simple setup, but **order of magnitude stays correct for more general geometries**.
- Results also applicable to small brane stacks \Rightarrow e.g. standard model branes in the bulk.
- Interesting for cosmology with multi-throat manifolds:
 - ▶ Throat-localized KK particles are **dark matter candidate**.
 - ▶ Can be very light \Rightarrow **constraints from cosmology**.
- We reproduce existing result by DIMOPOULOS ET AL. '01 in a limiting case. In general the **decay rate is much lower** \Rightarrow important for reheating after brane-antibrane inflation.

- We have determined **energy loss rate** of heated throat and **decay rate** of throat-localized KK modes.
- Genuine 10d calculation using the dual gauge theory picture. Simple setup, but **order of magnitude stays correct for more general geometries**.
- Results also applicable to small brane stacks \Rightarrow e.g. standard model branes in the bulk.
- Interesting for cosmology with multi-throat manifolds:
 - ▶ Throat-localized KK particles are **dark matter candidate**.
 - ▶ Can be very light \Rightarrow **constraints from cosmology**.
- We reproduce existing result by DIMOPOULOS ET AL. '01 in a limiting case. In general the **decay rate is much lower** \Rightarrow important for reheating after brane-antibrane inflation.

- We have determined **energy loss rate** of heated throat and **decay rate** of throat-localized KK modes.
- Genuine 10d calculation using the dual gauge theory picture. Simple setup, but **order of magnitude stays correct for more general geometries**.
- Results also applicable to small brane stacks \Rightarrow e.g. standard model branes in the bulk.
- Interesting for cosmology with multi-throat manifolds:
 - ▶ Throat-localized KK particles are **dark matter candidate**.
 - ▶ Can be very light \Rightarrow **constraints from cosmology**.
- We reproduce existing result by DIMOPOULOS ET AL. '01 in a limiting case. In general the **decay rate is much lower** \Rightarrow important for reheating after brane-antibrane inflation.

- We have determined **energy loss rate** of heated throat and **decay rate** of throat-localized KK modes.
- Genuine 10d calculation using the dual gauge theory picture. Simple setup, but **order of magnitude stays correct for more general geometries**.
- Results also applicable to small brane stacks \Rightarrow e.g. standard model branes in the bulk.
- Interesting for cosmology with multi-throat manifolds:
 - ▶ Throat-localized KK particles are **dark matter candidate**.
 - ▶ Can be very light \Rightarrow **constraints from cosmology**.
- We reproduce existing result by DIMOPOULOS ET AL. '01 in a limiting case. In general the **decay rate is much lower** \Rightarrow important for reheating after brane-antibrane inflation.

- We have determined **energy loss rate** of heated throat and **decay rate** of throat-localized KK modes.
- Genuine 10d calculation using the dual gauge theory picture. Simple setup, but **order of magnitude stays correct for more general geometries**.
- Results also applicable to small brane stacks \Rightarrow e.g. standard model branes in the bulk.
- Interesting for cosmology with multi-throat manifolds:
 - ▶ Throat-localized KK particles are **dark matter candidate**.
 - ▶ Can be very light \Rightarrow **constraints from cosmology**.
- We reproduce existing result by DIMOPOULOS ET AL. '01 in a limiting case. In general the **decay rate is much lower** \Rightarrow important for reheating after brane-antibrane inflation.