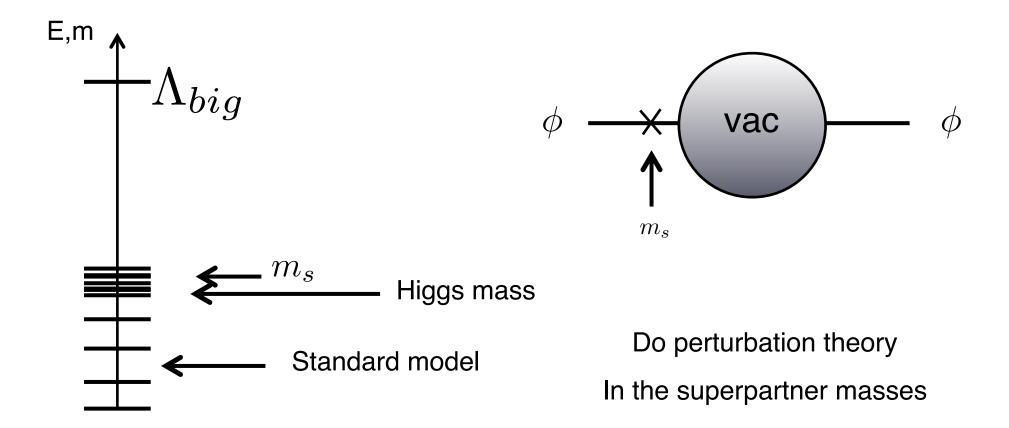
# Physics Beyond the Standard Model III: Composite Higgs and/or Extra Dimensions

David E Kaplan HCP Summer School 2007 CERN

#### Supersymmetry

Superpartner masses:  $m_s$ 

The symmetry is "softly broken", with  $m_s \ll \Lambda_{big}$ 



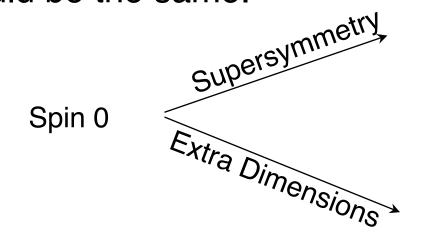
# Symmetries Between Particles with Different Spins

Masses and scattering cross sections

would be the same.

Spin 1/2

**Chiral Symmetry** 



Spin 1

Gauge Symmetry

A massless spin I field (e.g., the electromagnetic field) can be described by a 4-vector:

$$\mu = 0, 1, 2, 3$$
  $A_{\mu} \rightarrow \Phi, \vec{A}$ 

After imposing gauge invariance, and the equation of motion, there are two physical polarizations.

A massless spin I field in 4 space dimensions can be described by a 5-vector:

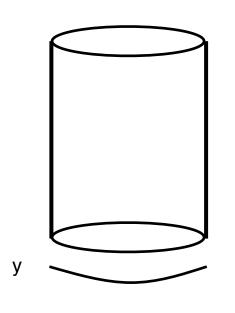
$$M=0,1,2,3,5$$
  $A_{M} \rightarrow \Phi, \vec{A}, A_{5}$ 

After imposing gauge invariance, and the equation of motion, there are *three* physical polarizations.

A massless spin I field in 4 space dimensions can be described by a 5-vector:

$$M=0,1,2,3,5$$
  $A_M \rightarrow \Phi, \vec{A}, A_5$ 

Now compactify.

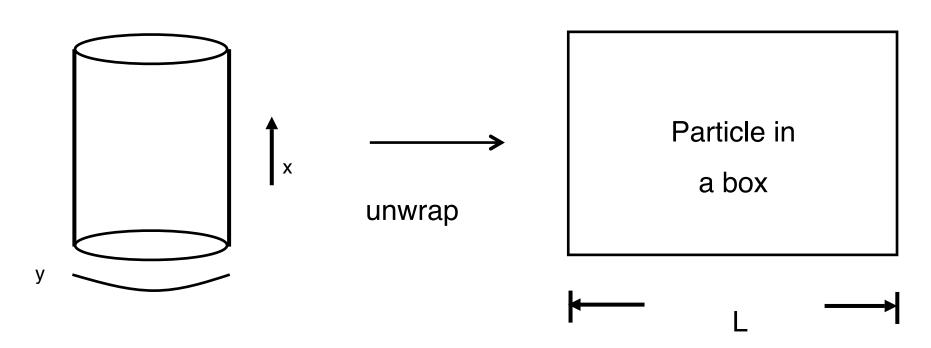


At distances much larger than the circumference, L, the new field  $A_5$  is not associated with a direction - it is a scalar (spinless) field.

A massless spin I field in 4 space dimensions can be described by a 5-vector:

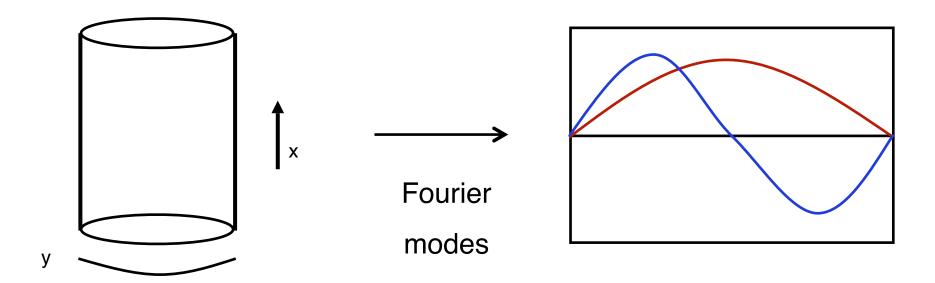
$$M=0,1,2,3,5$$
  $A_{M} \longrightarrow \Phi, \vec{A}, A_{5}$ 

Now compactify.



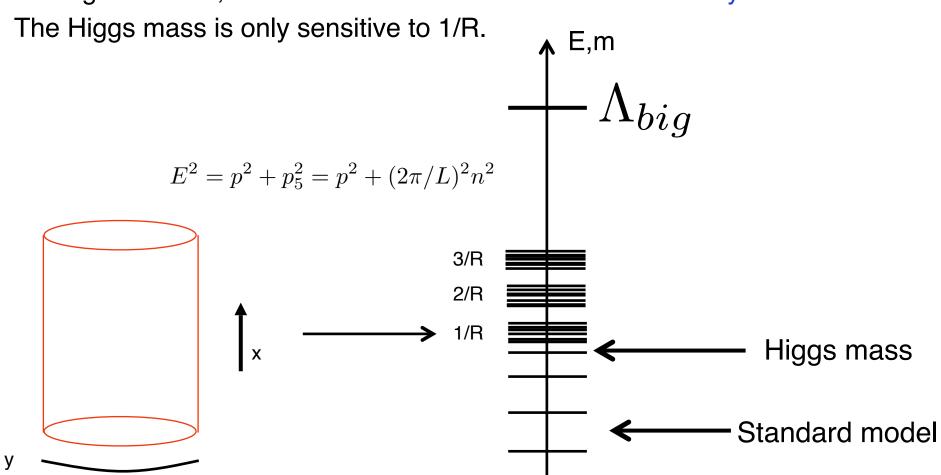
At short distances, 4-dimensional rotational invariance appears intact.

At long distances, the vector potential breaks up into a 3-vector and a scalar.



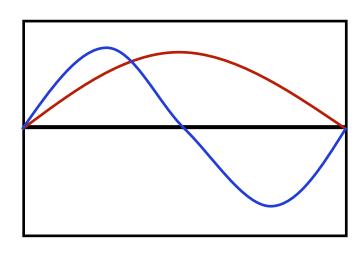
At short distances, 4-dimensional rotational invariance appears intact.

At long distances, 4-dimensional rotational invariance is softly broken.

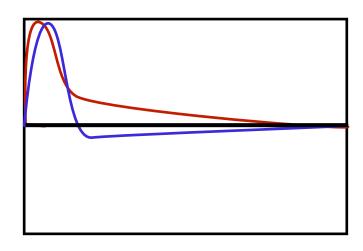


Some separation required between the Higgs mass and 1/R.

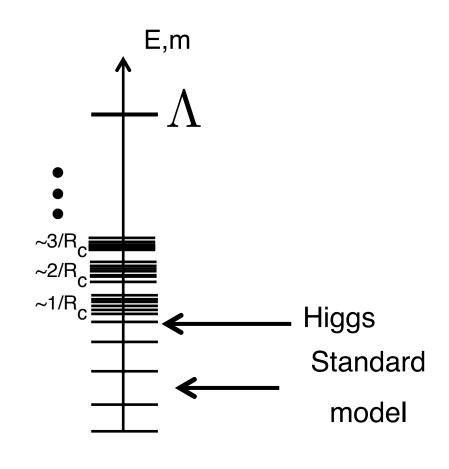


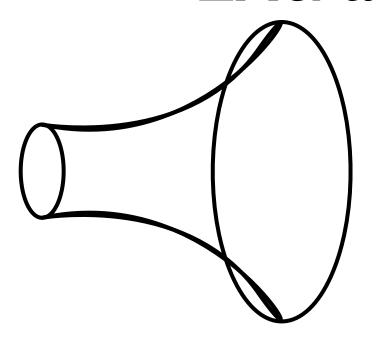




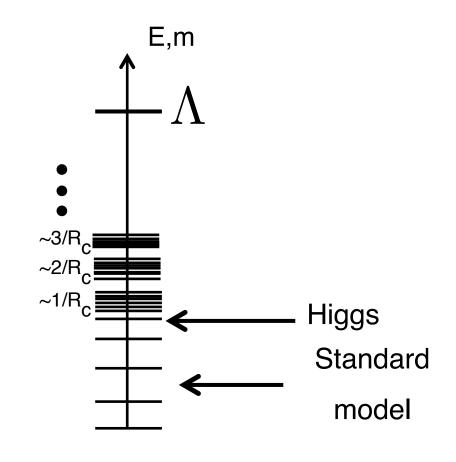








This 4D-5D "duality" also compares strong coupling with weak coupling.

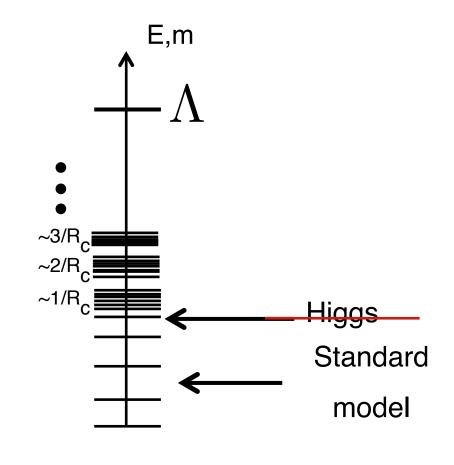




This 4D-5D "duality" also compares strong coupling with weak coupling.

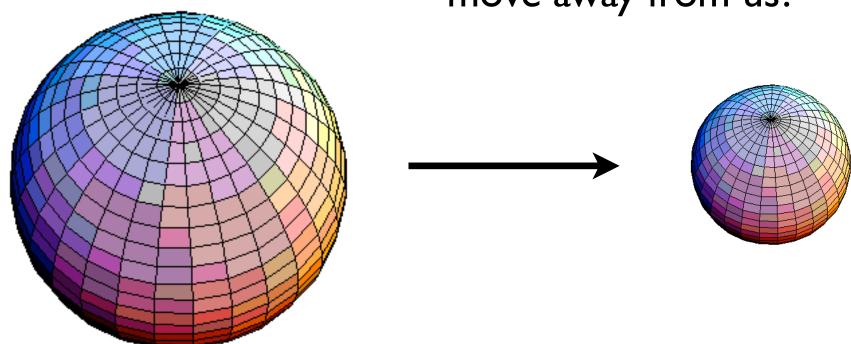
The theory should resemble

**Technicolor** 



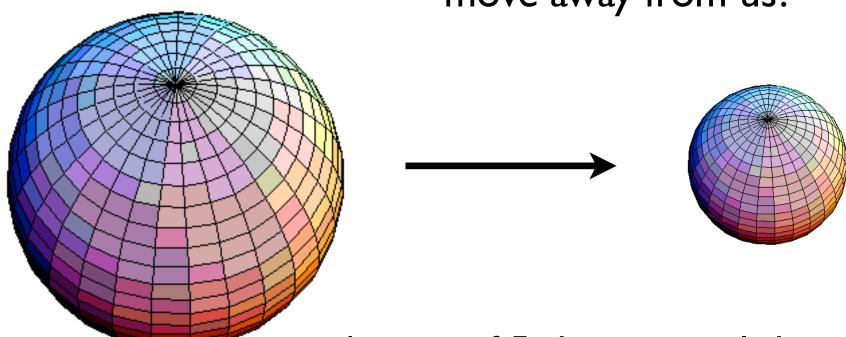
#### AdS/CFT

Did this object shrink or move away from us?



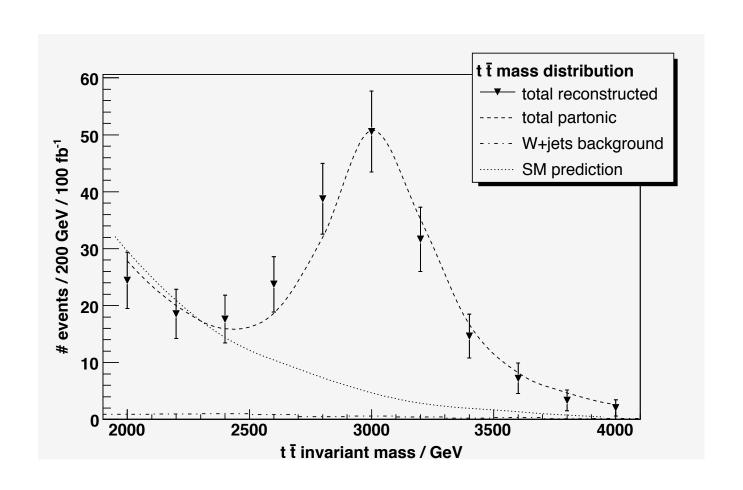
#### AdS/CFT

Did this object shrink or move away from us?



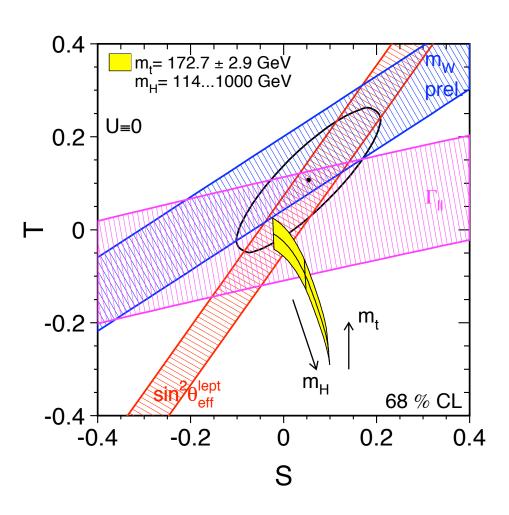
A type of 5-dimensional theory can be described as a 4-dimensional "scale-invariant" theory.

#### KK-gluon at the LHC

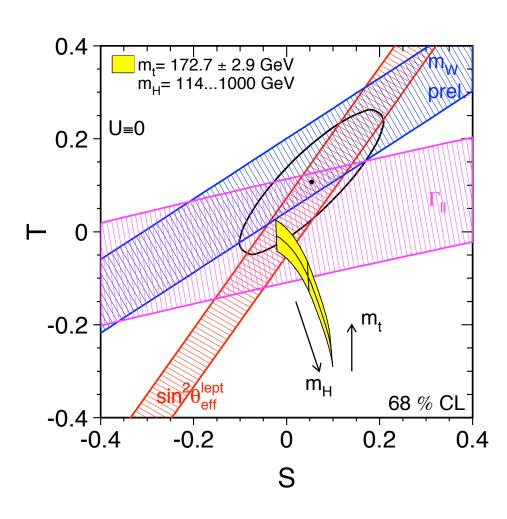


Agashe, Belyaev, Krupovnickas, Perez, Virzi

### What was wrong with Technicolor?



### What was wrong with Technicolor?





# How About Strong Coupling With a Higgs?

Pions are light, spinless composite particles.

How do they do that?

# How About Strong Coupling With a Higgs?

Pions are light, spinless composite particles.

How do they do that?

There is ample evidence that they are approximate Goldstone bosons from chiral-symmetry breaking in the theory of the strong nuclear force (QCD).

#### Goldstone Bosons

A "continuous" symmetry of the Lagrangian which is broken by

The ground state of the theory (the vacuum) is said to be

"spontaneously broken"

A spontaneously broken (internal) symmetry always produces a massless scalar particle.

If the symmetry is only approximate, the particle won't be massless, but can be very light.

#### Goldstone Bosons

A "continuous" symmetry of the Lagrangian which is broken by

The ground state of the theory (the vacuum) is said to be

"spontaneously broken"

A spontaneously broken (internal) symmetry always produces a massless scalar particle.

If the symmetry is only approximate, the particle won't be massless, but can be very light.

#### Do thing with the wire now.

#### Goldstone Bosons

Let's see the classical phenomenon using the wave description.

An infinite straight rope breaks translation invariance in directions perpendicular to the rope. The transverse waves are the Goldstone modes.

$$\mathcal{L} = \frac{1}{2}\sigma \left(\frac{\partial \phi}{\partial t}\right)^2 - \frac{1}{2}\tau \left(\frac{\partial \phi}{\partial x}\right)^2 \longrightarrow \frac{1}{c^2}\frac{\partial^2 \phi}{\partial t^2} = \frac{\partial^2 \phi}{\partial x^2}$$

Fourier transform:

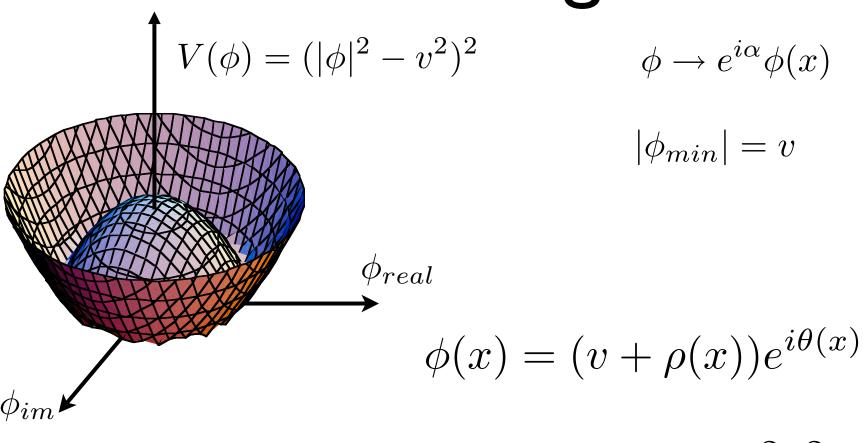
$$\phi(x,t) = \int \frac{dk \, d\omega}{4\pi^2} \tilde{\phi}(k,\omega) \, e^{i(kx-\omega t)} \longrightarrow \omega^2 = k^2 c^2$$

Can have waves with arbitrarily low frequency.

quantize: Particles with arbitrarily low energy —→massless particles

$$E^2 = p^2$$

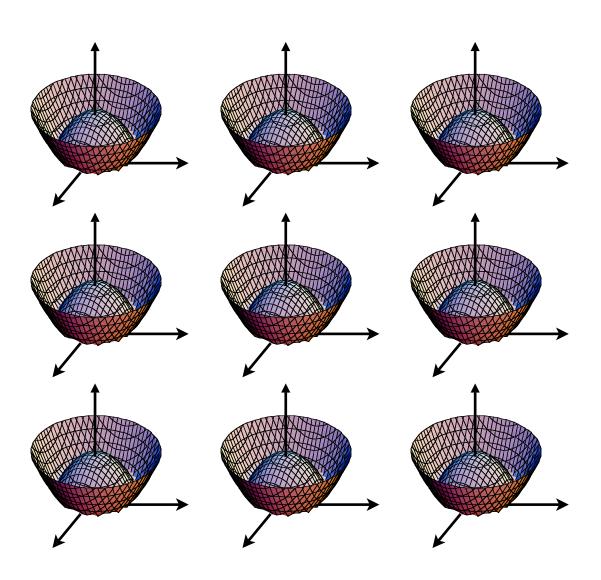
# Internal Symmetry Breaking



The potential is independent of theta.

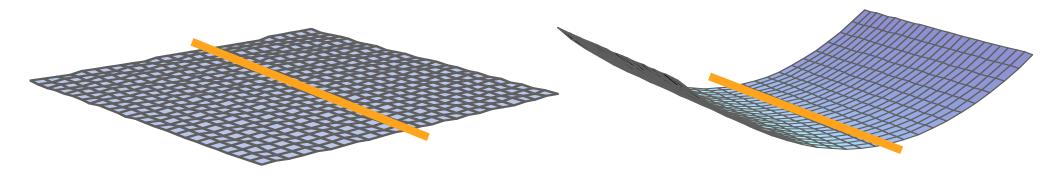
$$V = (2v\rho + \rho^2)^2$$

# Internal Symmetry Breaking



The potentials live at every point in space and waves of fluctuations between vacua move through space

#### Pseudo-Goldstone Bosons

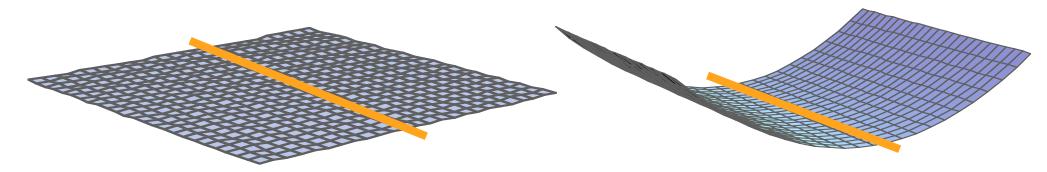


$$\mathcal{L} = \frac{1}{2}\sigma \left(\frac{\partial y}{\partial t}\right)^2 - \frac{1}{2}\tau \left(\frac{\partial y}{\partial x}\right)^2 - \frac{1}{2}\eta^2 y^2$$

#### Equation of motion:

$$\frac{1}{c^2} \frac{\partial^2 y}{\partial t^2} = \frac{\partial^2 y}{\partial x^2} - \mu^2 c^2 y^2$$

#### Pseudo-Goldstone Bosons

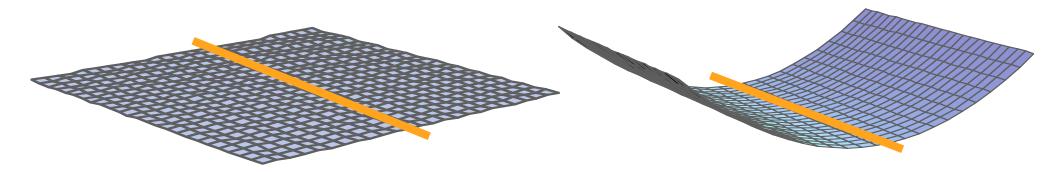


$$\mathcal{L} = \frac{1}{2}\sigma \left(\frac{\partial y}{\partial t}\right)^2 - \frac{1}{2}\tau \left(\frac{\partial y}{\partial x}\right)^2 - \frac{1}{2}\eta^2 y^2$$

A mass gap appears:

$$\omega^2 = k^2 + \mu^2$$

#### Pseudo-Goldstone Bosons



$$\mathcal{L} = \frac{1}{2}\sigma \left(\frac{\partial y}{\partial t}\right)^2 - \frac{1}{2}\tau \left(\frac{\partial y}{\partial x}\right)^2 - \frac{1}{2}\eta^2 y^2$$

A mass gap appears:

$$\omega^2 = k^2 + \mu^2$$

This is what we think a pion is.

# Technicolor vs. Composite Higgs

 $Q_i$ ,  $Q_i^c$  = techniquarks

$$\langle Q_i Q_j^c \rangle \sim \Lambda_{TC}^3 \delta_{ij}$$

$$Q = \begin{pmatrix} U \\ D \end{pmatrix}, \ Q_i^c = U^c, D^c$$

$$SU(2)_L \times SU(2)_R$$

$$\Pi = \left( \begin{array}{cc} \pi^0 & \pi^+ \\ \pi^- & -\pi^0 \end{array} \right)$$

without the Higgs, QCD breaks Electroweak symmetry, and the pions are eaten

$$M_W = g f_{\pi}$$

# Technicolor vs. Composite Higgs

 $Q_i$ ,  $Q_i^c$  = techniquarks

$$\langle Q_i Q_j^c \rangle \sim \Lambda_{TC}^3 \delta_{ij}$$
  $Q = \begin{pmatrix} U \\ D \\ S \end{pmatrix}, Q_i^c = \begin{pmatrix} U^c & D^c & S^c \end{pmatrix}$ 

$$\begin{pmatrix} \pi^0 - \frac{\eta^0}{\sqrt{c}} & \pi^+ & K^+ \end{pmatrix}$$

$$SU(3)_L \times SU(3)_R \qquad \Pi = \begin{pmatrix} \pi^0 - \frac{\eta}{\sqrt{3}} & \pi^+ & K^+ \\ \pi^- & -\pi^0 - \frac{\eta}{\sqrt{3}} & K^0 \\ K^- & \bar{K}^0 & \frac{2\eta^0}{\sqrt{3}} \end{pmatrix}$$

Georgi, Kaplan

### Technicolor vs. Composite Higgs

 $Q_i$ ,  $Q_i^c$  = techniquarks

$$\langle Q_i Q_j^c \rangle \sim \Lambda_{TC}^3 \delta_{ij}$$
  $Q = \begin{pmatrix} U \\ D \\ S \end{pmatrix}, Q_i^c = (U^c D^c S^c)$ 

$$SU(3)_L \times SU(3)_R$$

$$SU(3)_L \times SU(3)_R \qquad \Pi = \begin{pmatrix} \pi^0 - \frac{\eta^0}{\sqrt{3}} & \pi^+ & K^+ \\ \pi^- & -\pi^0 - \frac{\eta}{\sqrt{3}} & K^0 \end{pmatrix}$$

$$\equiv \left(\begin{array}{c} H^+ \\ H^0 \end{array}\right)$$

Then include weak couplings to give it a vev

The naturally light pion has significant couplings (e.g., electro-magnetic).

The naturally light pion has significant couplings (e.g., electro-magnetic).

Just what we need for the Higgs

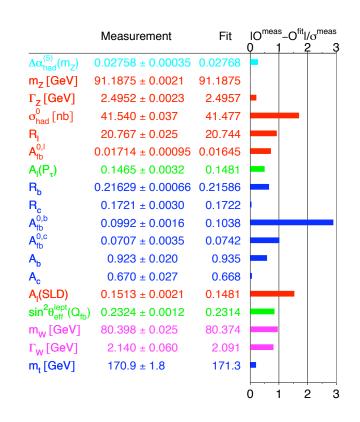
The naturally light pion has significant couplings (e.g., electro-magnetic).

Just what we need for the Higgs

heuristically:

$$\left(\frac{m_{\pi}}{m_{\rho}}\right)^2 \sim \left(\frac{m_{\pi}}{\Lambda}\right)^2 \sim \left(\frac{140 \,\mathrm{MeV}}{770 \,\mathrm{MeV}}\right)^2 \sim 4\%$$

but per mil required:

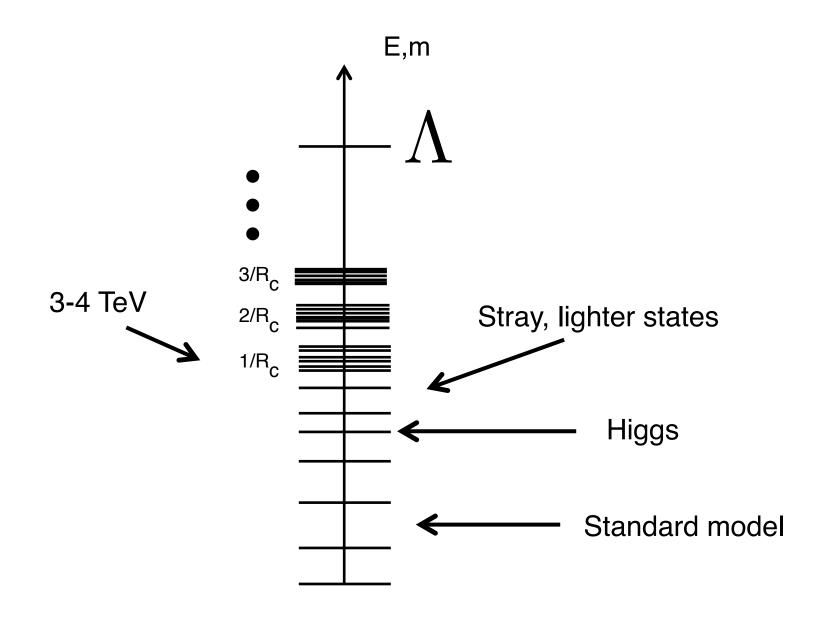


## Answer - Some Tuning

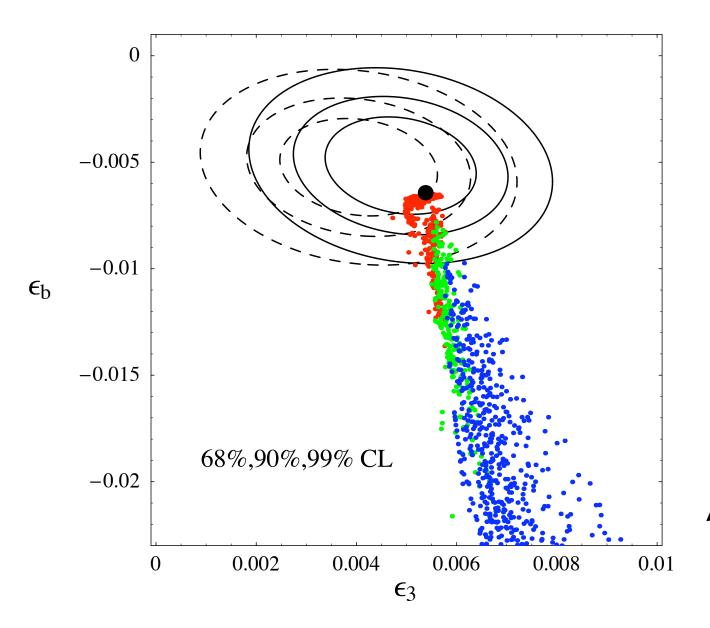
Worst problem is a contribution to  $\,Z 
ightarrow b \overline{b} \,$ 

In current models, must push the "Techni-"rho mass to 3-4 TeV, requiring a fine-tuning in the Higgs mass of  $\sim$ 5%.

## Answer - Some Tuning



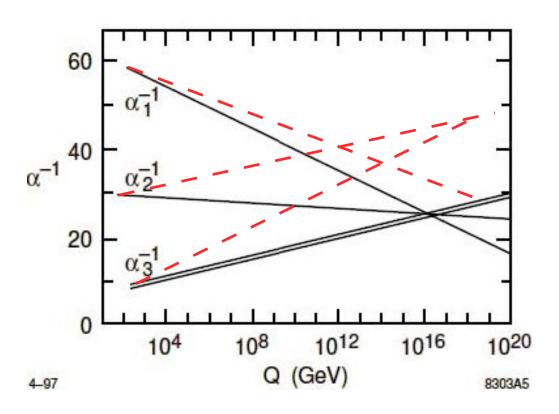
## RS w/ composite Higgs



< 4% <10%, >4% > 10%

Agashe, Contino

$$\alpha(M_Z) = \alpha_{unified} + SM + superpartners$$

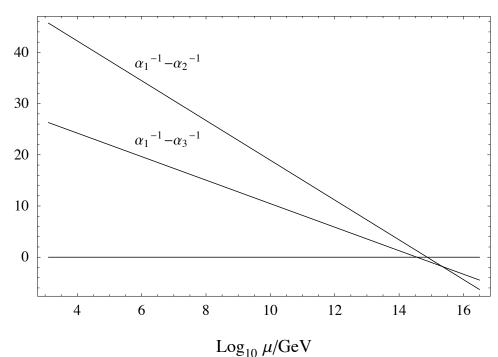


$$\alpha(M_Z) = \alpha_{unified} + \text{SM} + \text{superpartners}$$

$$\alpha(M_Z) = \alpha_{unified} + \text{SM} - \{H, t_R, t_R^c\}$$

$$\alpha(M_Z) = \alpha_{unified} + SM + superpartners$$

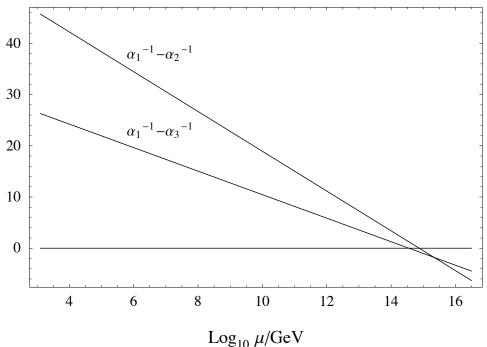
$$\alpha(M_Z) = \alpha_{unified} + SM - \{H, t_R, t_R^c\}$$



Agashe, Contino, Sundrum

$$\alpha(M_Z) = \alpha_{unified} + SM + superpartners$$

$$\alpha(M_Z) = \alpha_{unified} + SM - \{H, t_R, t_R^c\}$$



Agashe, Contino, Sundrum

Beautiful Flavor Structure Too

## RS summary

Strongly coupled theories are making a comeback as AdS/CFT suggests more theoretical control

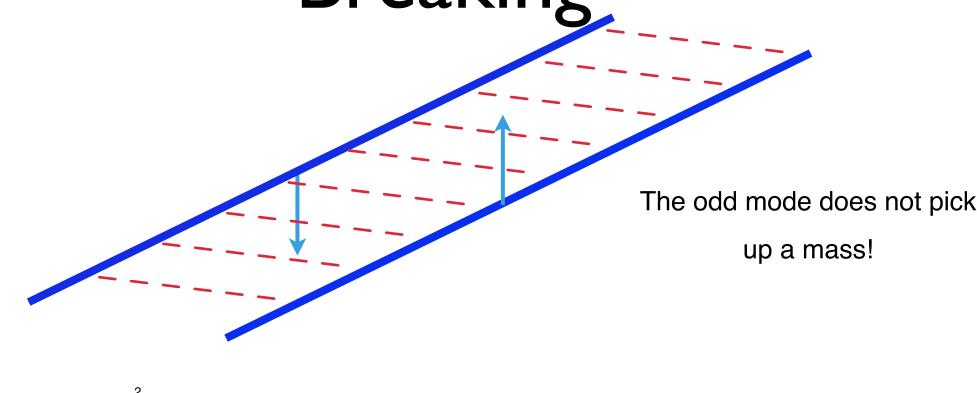
Best versions are when the Higgs is a pseudo-Goldstone boson similar to the pion (Composite Higgs Theories)

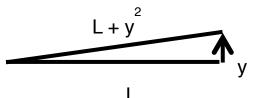
Significant effort made to simplify the phenomenology: Contino, Kramer, Son, Sundrum; hep-ph/0612180

## Little Higgs

The goal of the Little Higgs models is to remove that last bit of tuning in composite Higgs theories - at the price of complicating the models.

# Collective Symmetry Breaking

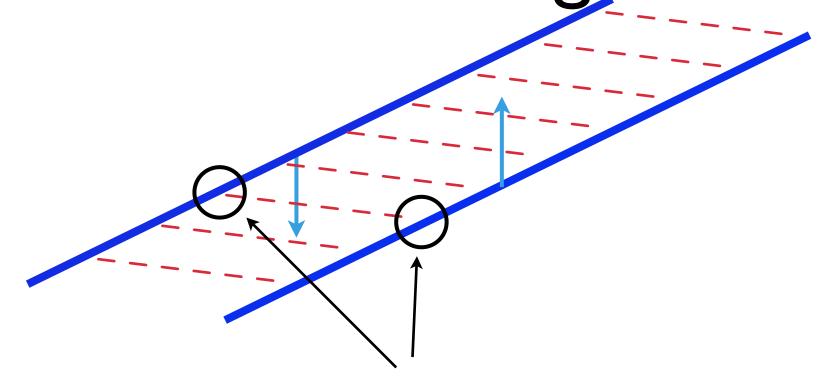




Force goes like displacement

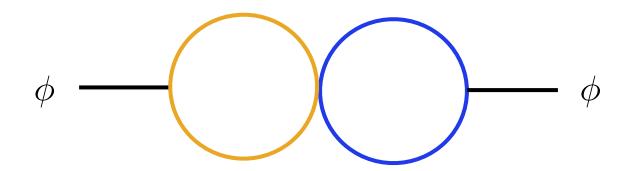
Higher order effect - no mass gap! Quantum corrections induce

# Collective Symmetry Breaking



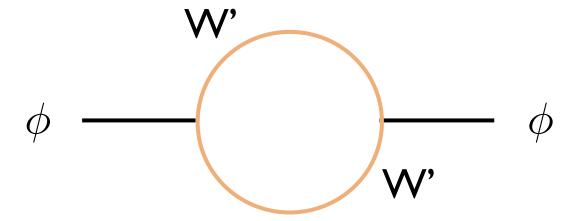
Two couplings required

## The "Little Higgs"



$$m_{\phi} \sim \epsilon^2 \Lambda$$

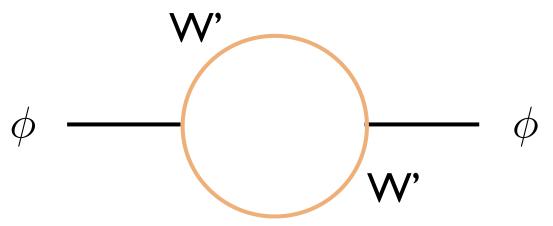
## The "Little Higgs"

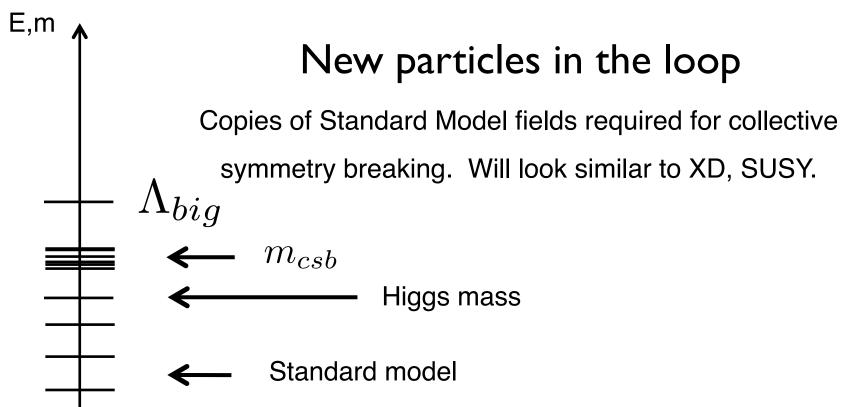


#### New particles in the loop

Copies of Standard Model fields required for collective symmetry breaking. Will look similar to XD, SUSY.

## The "Little Higgs"

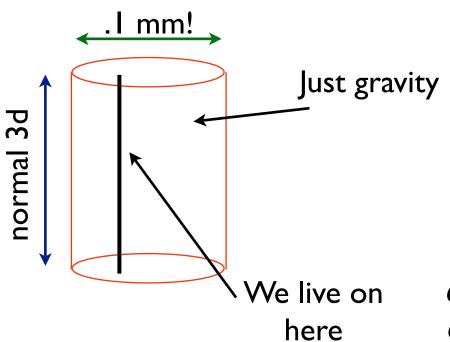




## Large Extra Dimensions

$$M_{pl}^2 = M_G^{D-2} V_{D-4}$$

Perhaps M<sub>G</sub> ~ I TeV!



D=6 severely constrained by short-dist. gravity tests and supernova 1987A

only D=6 has a nice stabilization mechanism

$$egin{array}{l} q ar q &
ightarrow \gamma + invisible \ gravitons \ gg &
ightarrow g + invisible \ gravitons \end{array}$$

# Should the world be "Natural"?

$$(m_h^{phys})^2 = M_{planck}^2 + M_{qtm\,corr}^2$$

Cancellation of one part in 10<sup>34</sup>

This doesn't happen in condensed matter system unless we force it to happen.

# The Cosmological Constant

Fine tuning of one part in  $10^{120}$ ! (one part in  $10^{60}$  with SUSY)

In 1987 Weinberg suggested that if our universe was one of a large selection, we would live in one where the cosmological constant was just small enough to allow for structure to form, and if we don't measure one, we can rule out the anthropic principle.

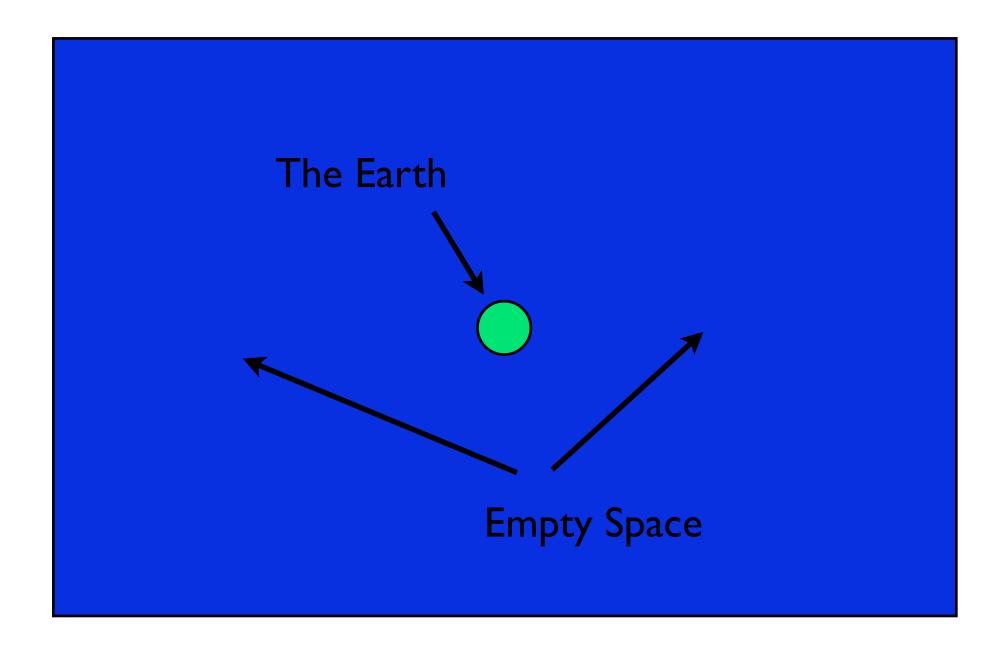
# The Cosmological Constant

Fine tuning of one part in  $10^{120}$ ! (one part in  $10^{60}$  with SUSY)

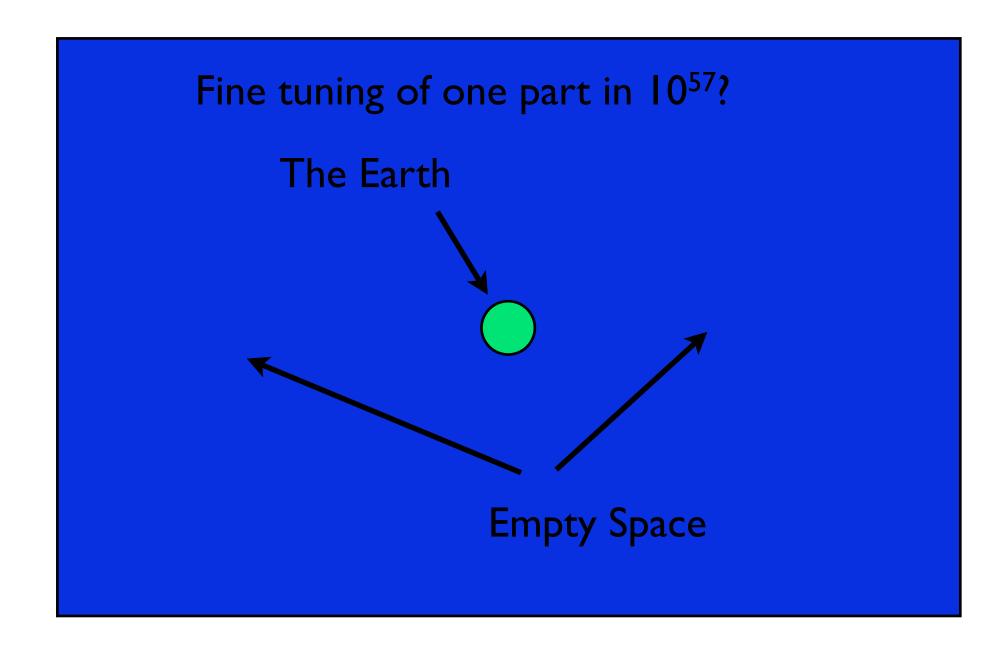
In 1987 Weinberg suggested that if our universe was one of a large selection, we would live in one where the cosmological constant was just small enough to allow for structure to form, and if we don't measure one, we can rule out the anthropic principle.

But then we did measure one...

### **Environmental Selection**



### **Environmental Selection**



## Split Supersymmetry

 $M_{
m Pl.}$ 

 $10^{16} \text{ TeV}$ 

Scalars (Squarks, sleptons, ...)

 $M_{\text{\tiny susy}}$  ?  $\left\{ \begin{array}{c} 10^{15} \text{ TeV} \\ 10 \text{ TeV} \end{array} \right.$ 

Fermions (Higgsinos, gauginos) +SM Higgs

 $M_{
m weak}$ 

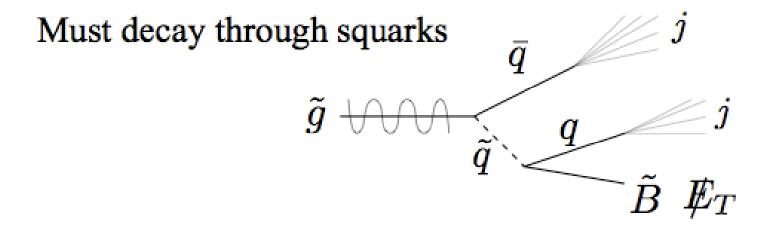
 $\sim 1 \text{ TeV}$ 

Preserves Unification and Dark Matter

 $M_{\scriptscriptstyle 
m CC}$ 

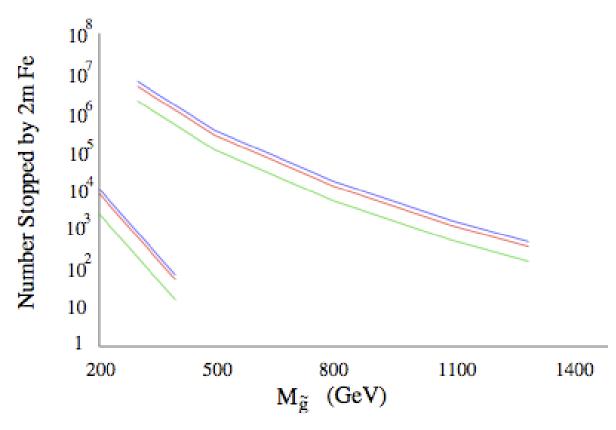
 $10^{-15} {
m TeV}$ 

## Gluinos are Long-Lived

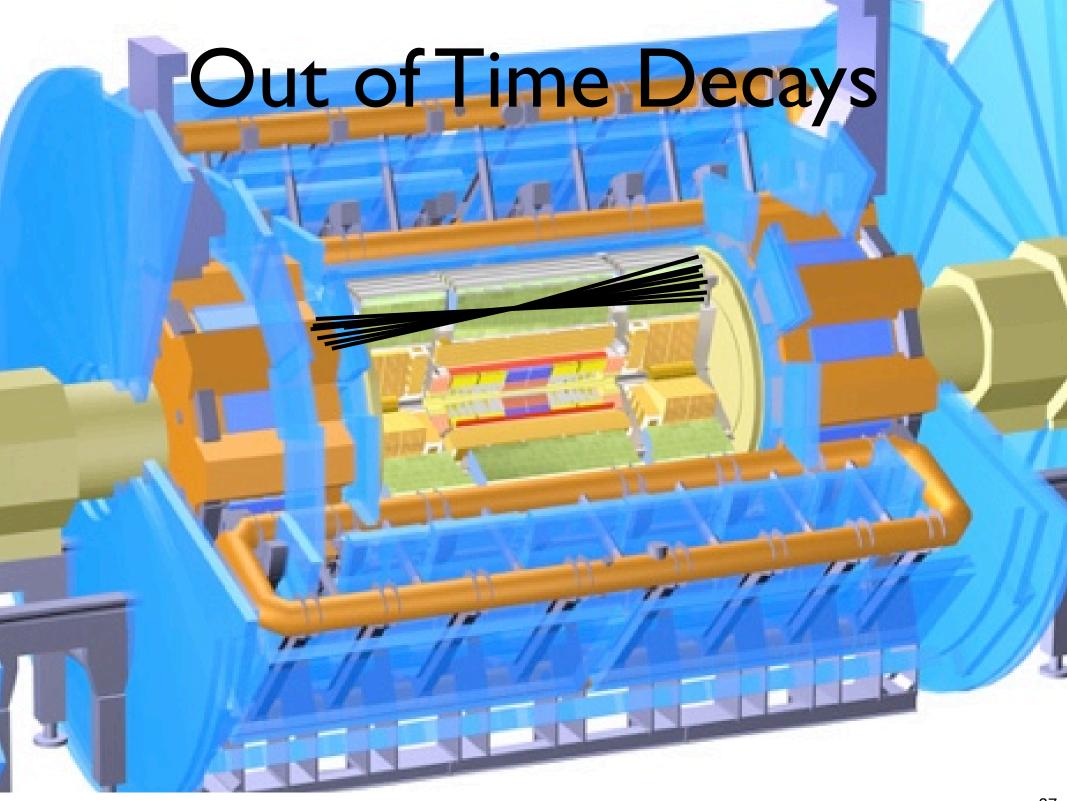


$$au_{ ilde{g}} \simeq 2 \; {
m sec.} \left(rac{350 \; {
m GeV}}{m_{ ilde{g}}}
ight)^5 \left(rac{M_{
m Susy}}{10^6 \; {
m TeV}}
ight)^4$$

## Stopped Gluinos



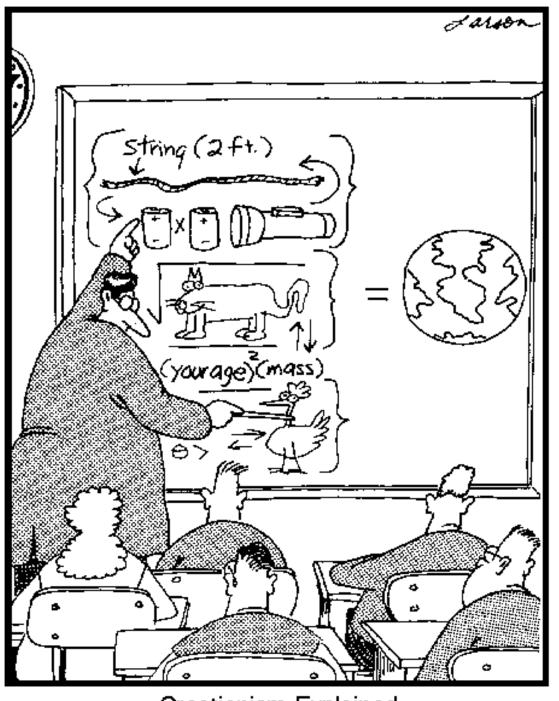
$2 \; { m fb^{-1}}$	$200~{ m GeV}$	$300~{ m GeV}$	$400~{ m GeV}$
CDF	$4.1 \times 10^{3}$	$3.1 \times 10^{2}$	$3.3 \times 10^{1}$
D0	$4.5 \times 10^{3}$	$3.3 \times 10^{2}$	$3.4 \times 10^{1}$
$100 \; {\rm fb^{-1}}$	300  GeV	$800~{ m GeV}$	$1300~{ m GeV}$
ATLAS	$5.8 \times 10^{6}$	$1.8 \times 10^{4}$	$6.2 \times 10^{2}$
CMS	$3.7 \times 10^{6}$	$1.2 \times 10^{4}$	$3.9 \times 10^{2}$



# Even if you HATE use of the anthropic principle...

It points us to a new signal to look for - very long lived stopped particles.

Time to figure out what is going on



Creationism Explained