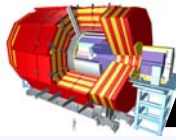


The Road to Discovery

Andy Parker
Cambridge University

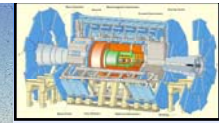
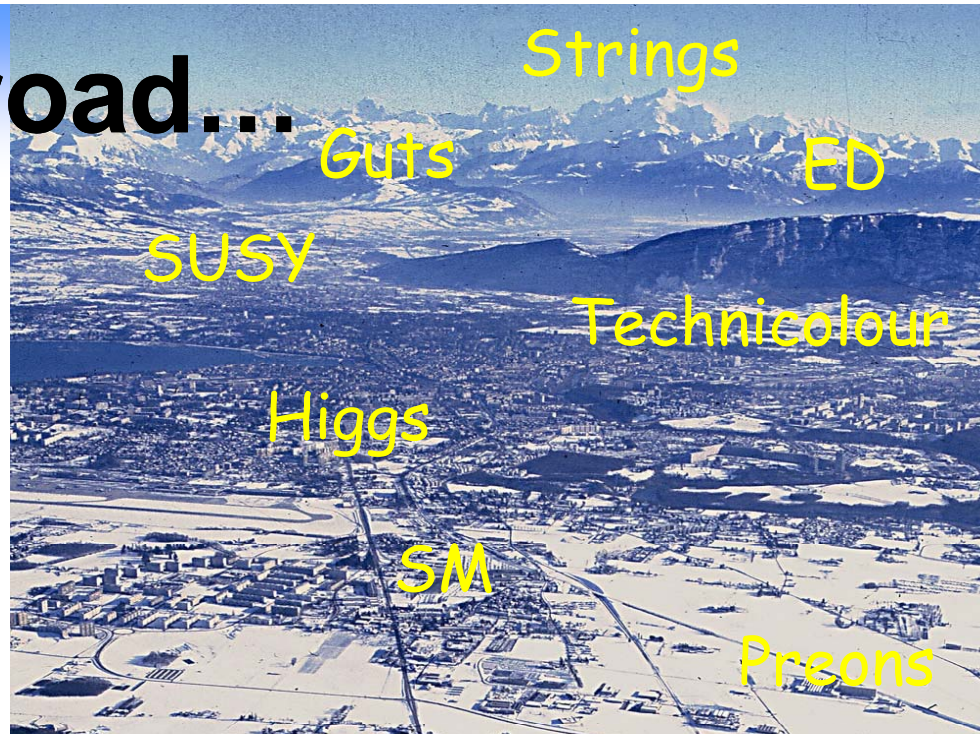


Along the road...

- Preparations - SM Physics, tools, problems
- The local road - hunting the Higgs
- The far horizon - SUSY
- The hidden road - Extra Dimensions

Thanks to Fabiola Gianotti and many authors of notes etc used in these lectures

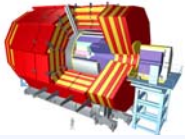
Andy Parker



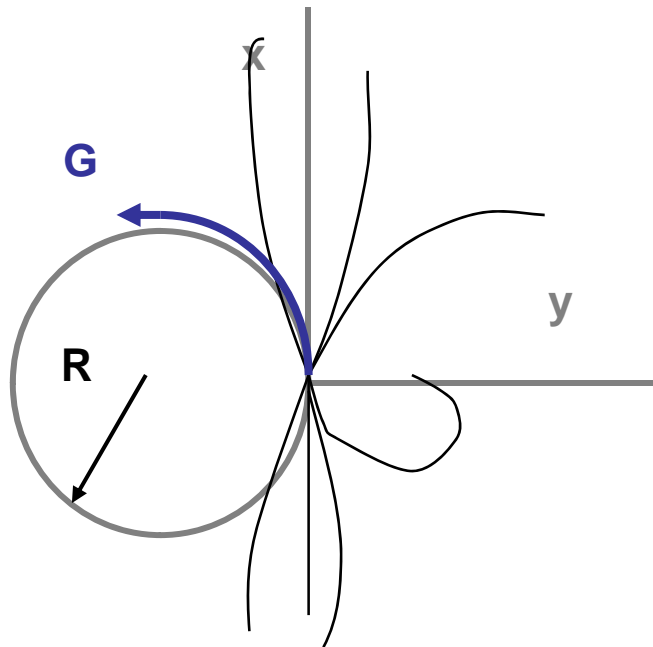
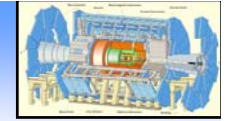
"The problem is not a want of a theory, but a want of evidence. If scientific advance really came from theorizing, natural scientists would have long ago wrapped up their affairs and gone on to more interesting matters"

Richard Lewontin NYRB

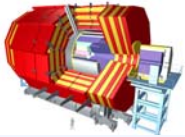
1995



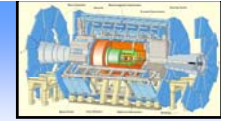
LED at LHC



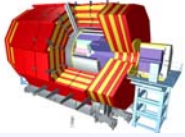
- ADD model (hep-ph/9803315)
- Each excited graviton state has normal gravitational couplings
- -> negligible effect
- LED: very large number of KK states in tower
- Sum over states is large.
- => Missing energy signature with massless gravitons escaping into the extra dimensions



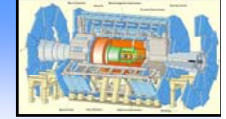
Signatures at LHC



- Good signatures are LBNL-45198
- Jet +missing energy channels: ATL-PHYS-2001-012
 - $gg \rightarrow gG$
 - $qg \rightarrow qG$
 - $q\bar{q} \rightarrow Gg$
- Photon channels
 - $q\bar{q} \rightarrow G\gamma$
 - $pp \rightarrow \gamma\gamma X$ Virtual graviton exchange
- Lepton channels
 - $pp \rightarrow ll X$ Virtual graviton exchange



Real graviton production



Cross section:

$$\frac{d^4 \sigma}{dm^2 dp_T^2 dy_{jet,\gamma} dy_G} = \frac{m_G^{n-2}}{2} \frac{S_{n-1}}{M_D^{n+2}} \frac{d\sigma_m}{dt} \sum_{i,j} \frac{f_i(x_1)}{x_1} \frac{f_j(x_2)}{x_2}$$

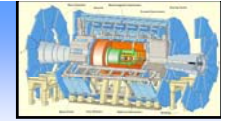
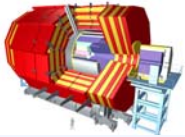
Note ED mass scale and n do not separate ->
difficult to extract n

Can use cutoff in M_D from parton distributions

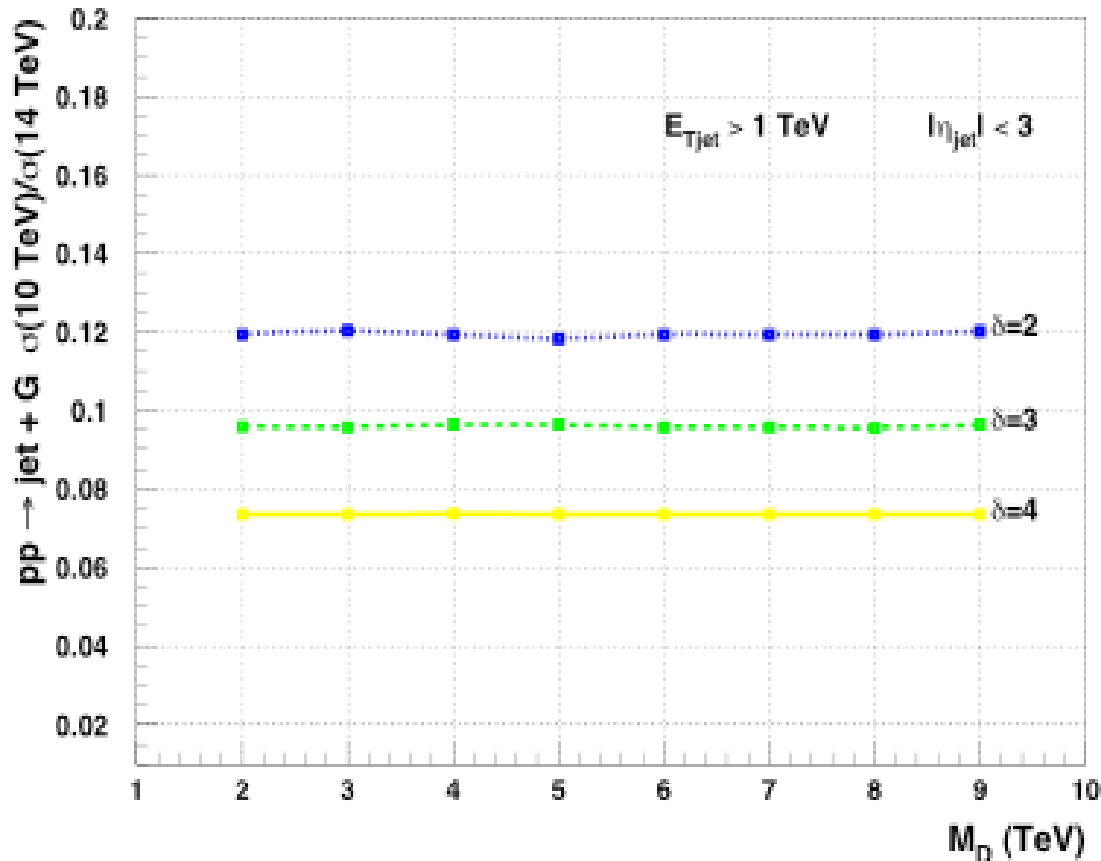
For $n > 6$, cross section unobservable at LHC

Quantum gravity theory unknown ->

Calculation only reliable at energies well away
from M_D



Energy variation of cross-section

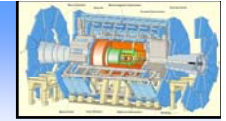
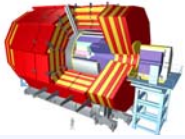


Cross section ratio
(10 TeV/14 TeV)

Need to measure to
5% to distinguish
 $n=2,3$

Need $O(10)$ more L at
10 TeV

Need luminosity to
<5%



Missing E_T analysis

$pp \rightarrow \text{jet} + E_T^{\text{Miss}}$

Jet energies $> 1 \text{ TeV}$

Dominant backgrounds:

$\text{Jet} + Z \rightarrow \nu \nu$

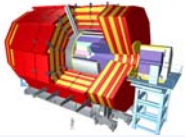
$\text{Jet} + W \rightarrow \tau \nu$

$\text{Jet} + W \rightarrow e \nu$

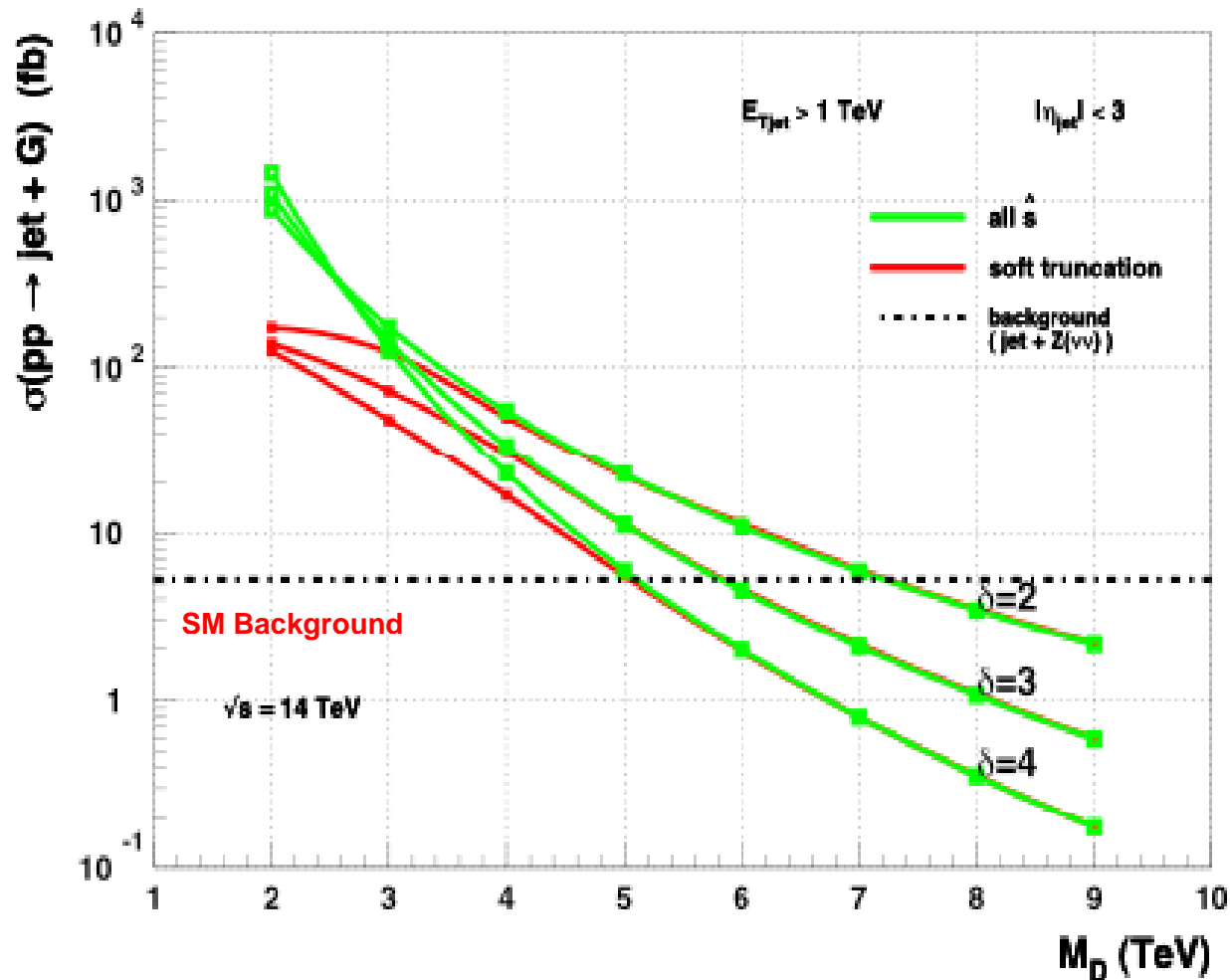
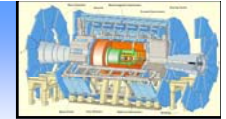
} Use lepton veto

Veto isolated leptons ($< 10 \text{ GeV}$ within $\Delta R = 0.2$)

Instrumental background to E_T^{Miss} is small



High P_T jet cross section

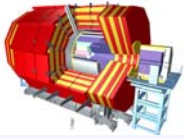


$E_T^{\text{Jet}} > 1 \text{ TeV}$

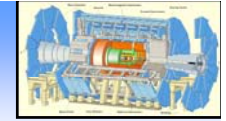
$|\eta_{\text{Jet}}| < 3$

100fb⁻¹ of data
SM Background
~500 events

No prediction
for $n > 4$



Missing E_T signal

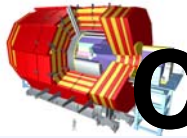


QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

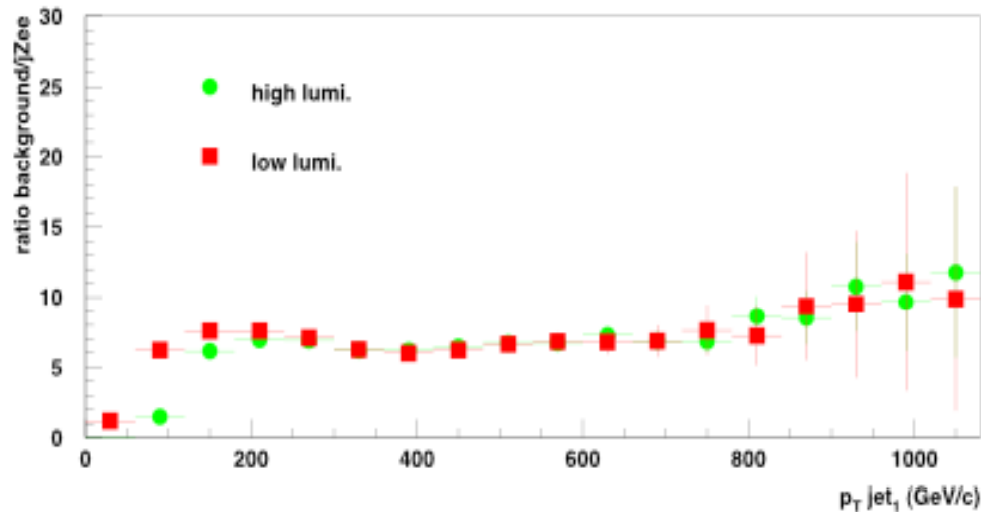
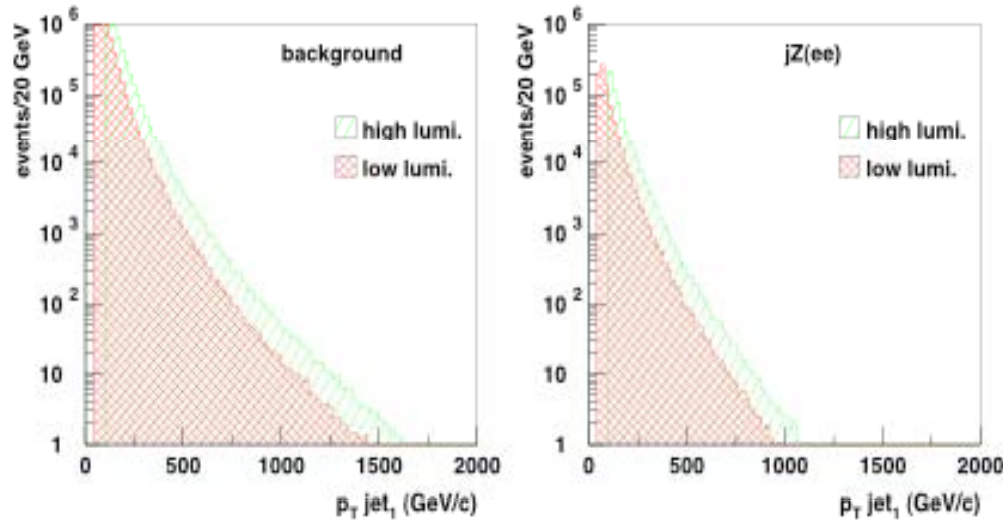
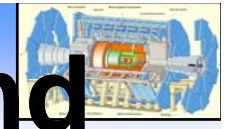
Signal:
Excess of events
at high E_T

Dominant
background

$Z \rightarrow \nu\nu$



Calibration of Z-> $\nu\nu$ background

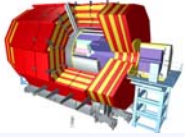


Use Z-> ee

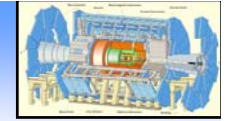
Two isolated electrons,
 $P_T > 15$, M_{ee} within 10
GeV of M_Z

Account for acceptance
differences e, μ , ν

BR's differ by factor 3,
so calibration sample
has less statistics

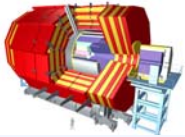


ATLAS Discovery potential

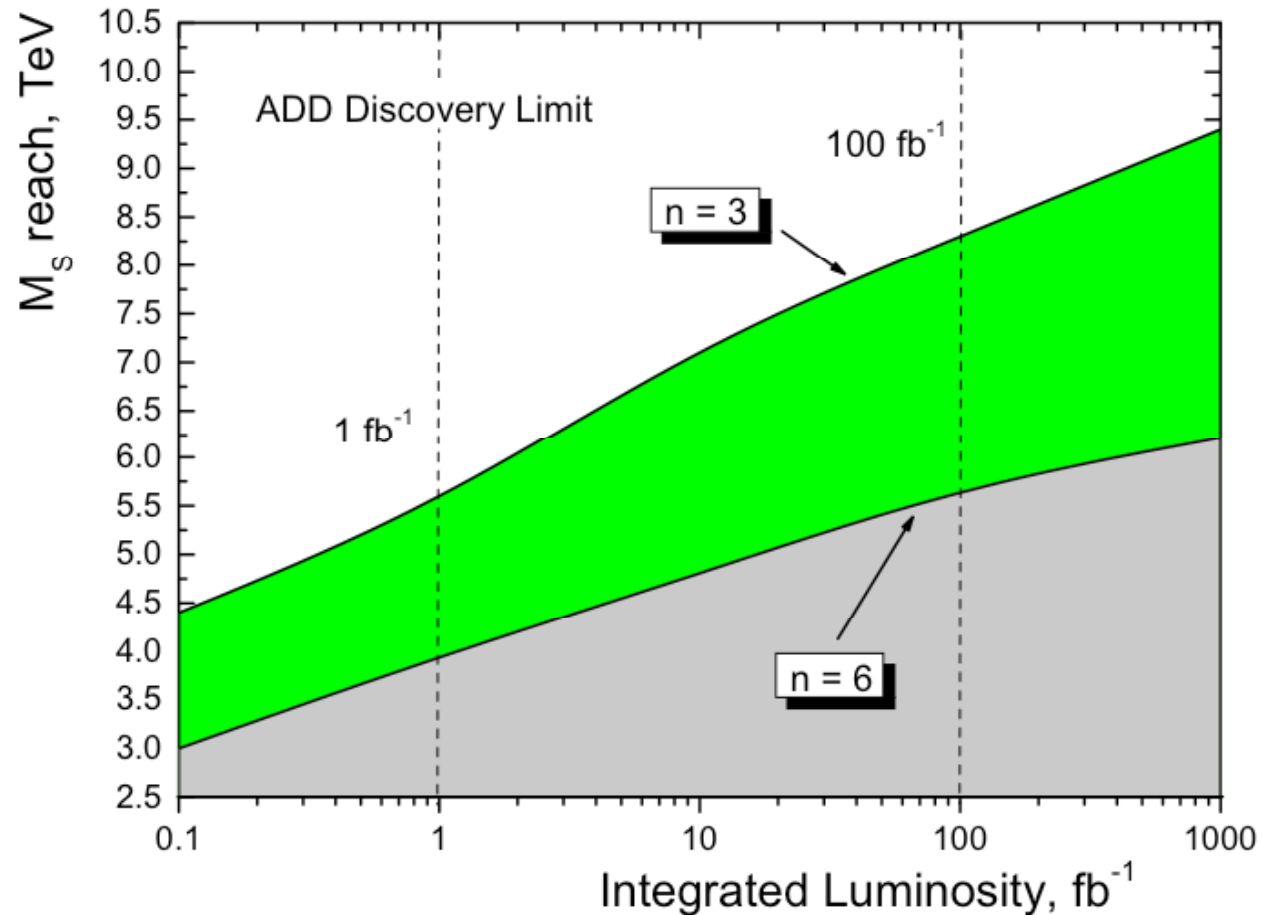
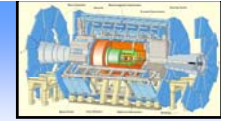


5σ discovery limits, ADD model, $E_T > 1$ TeV, 100fb^{-1}

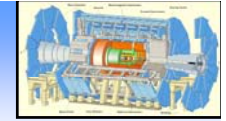
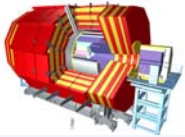
n	M_D^{min}	M_D^{Max} (TeV)	R
2	~4	7.5	10 μm
3	~4.5	5.9	300 μm
4	~5	5.3	1 μm



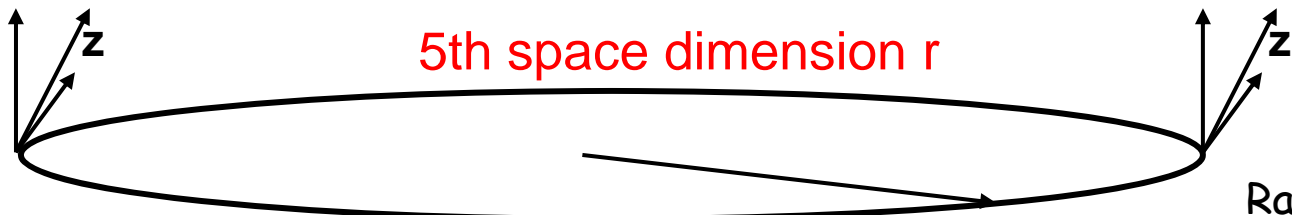
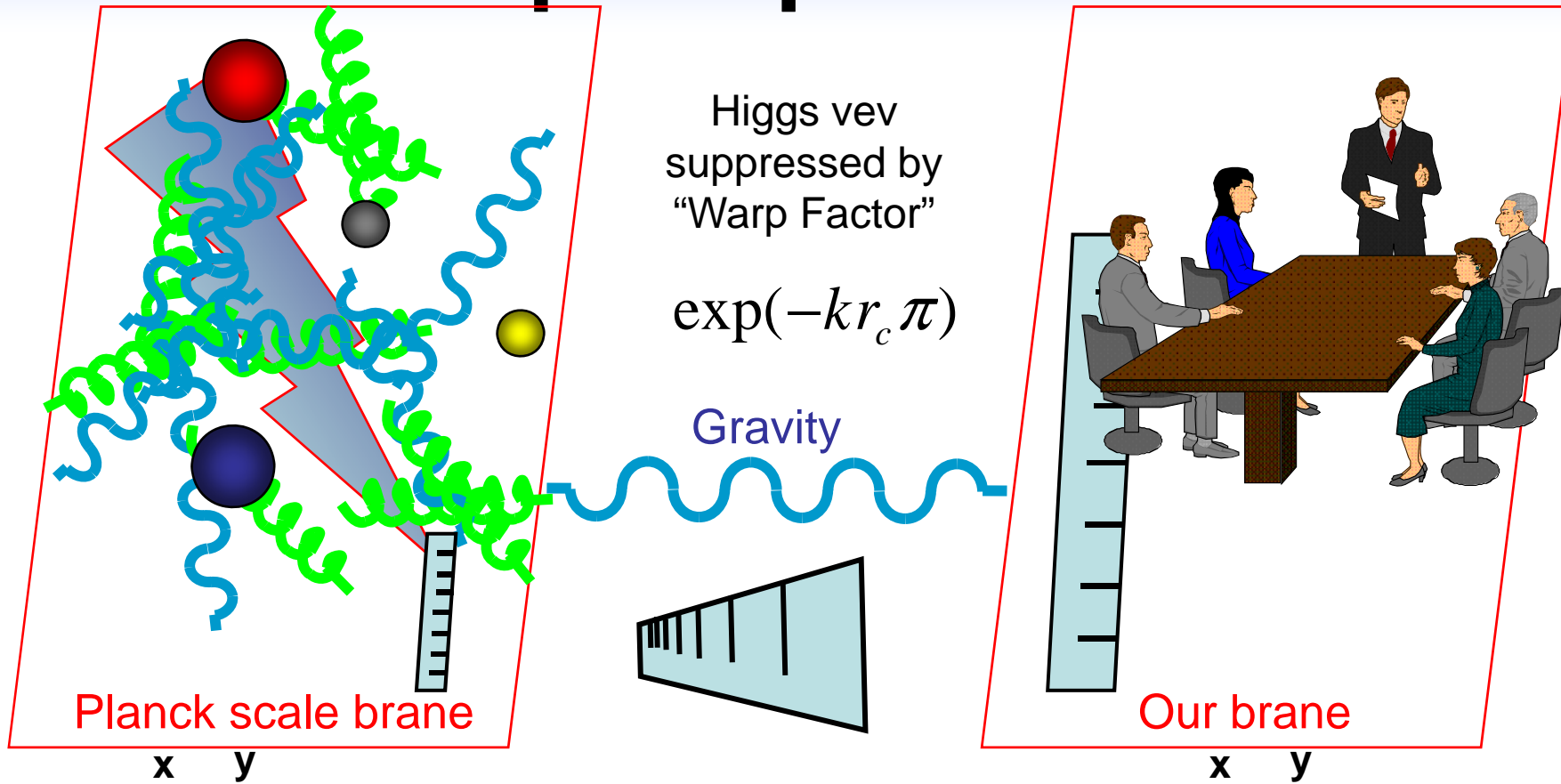
CMS discovery reach, ADD



Discovery reach in range 4.5-8.5 TeV for $n=3$.

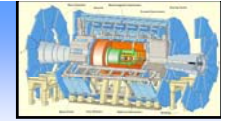
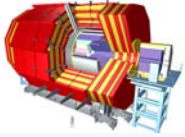


Warped space-times



$$r_c \approx 10^{-32} m$$

Randall and Sundrum
PRL83,3370(1999)



R-S models

Consider Randall and Sundrum type models

Gravity propagates in a 5-D non-factorizable geometry

Hierarchy between M_{Planck} and M_{Weak} generated by "warp factor"

Need : no fine tuning

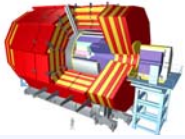
Gravitons $kr_c \approx 10$ have KK excitations with scale

Spectrum of graviton excitations π seen as resonances which decay to fermion pairs (eg e^+e^-)

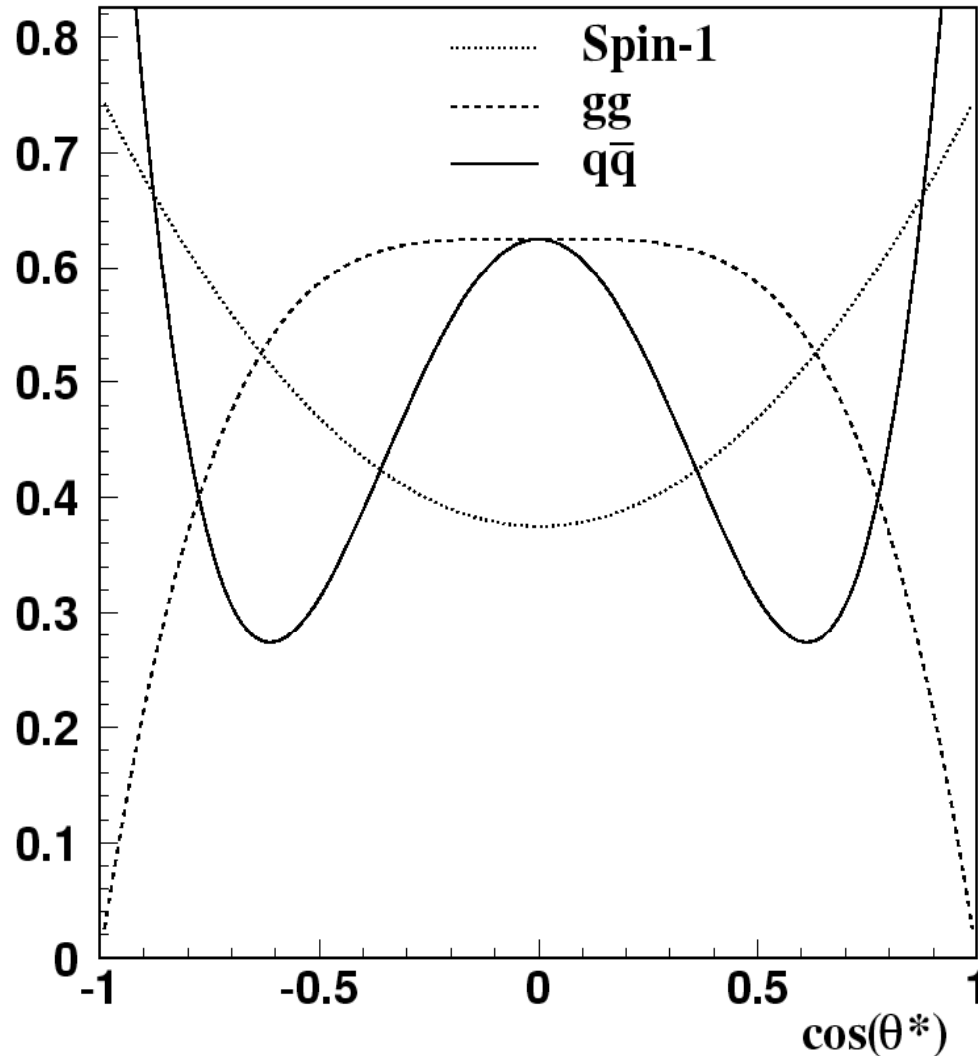
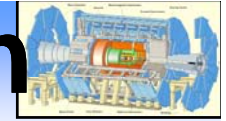
First excitation is at

where
$$m_1 = kx_1 \exp(-kr_c \pi) = 3.83 \frac{k}{M_{Pl}} \Lambda_\pi$$

$$0.01 \leq \frac{k}{M_{Pl}} \leq 1$$

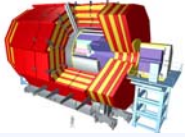


Angular distributions of e^+e^- in graviton frame

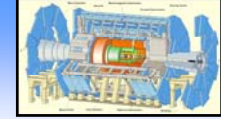


Angular distributions are very different depending on the spin of the resonance and the production mechanism.

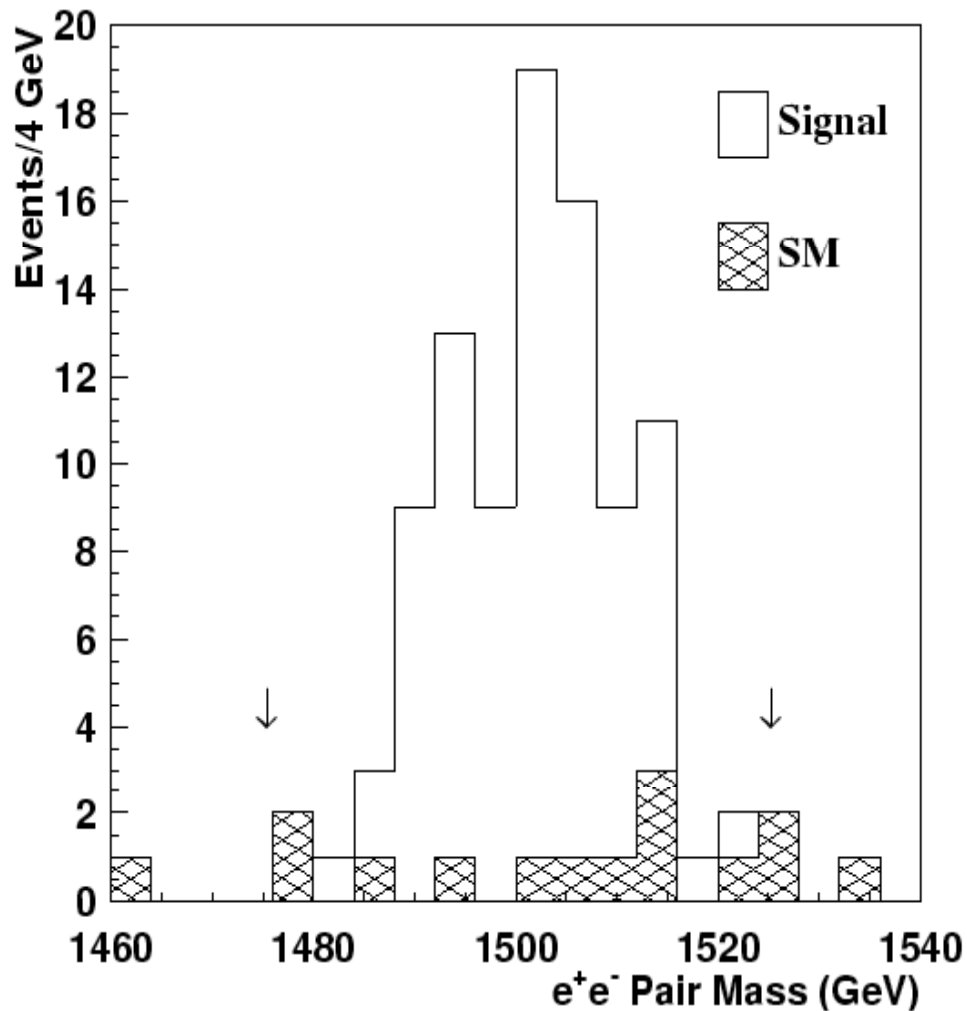
=>get information on the spin and couplings of the resonance



Graviton Resonance

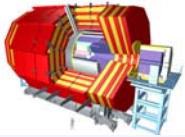


$$G \rightarrow e^+e^-$$

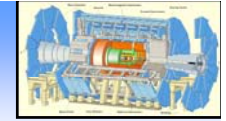


Graviton resonance is very prominent above small SM background, for 100fb^{-1} of integrated luminosity

Plot shows signal for a 1.5 TeV resonance, in RS model. The Drell-Yan background can be measured and subtracted from the sidebands.



CMS e^+e^- resonance search



e^+e^- channel favoured for G resonance as well as Z' and Z^* searches - reach depends on coupling

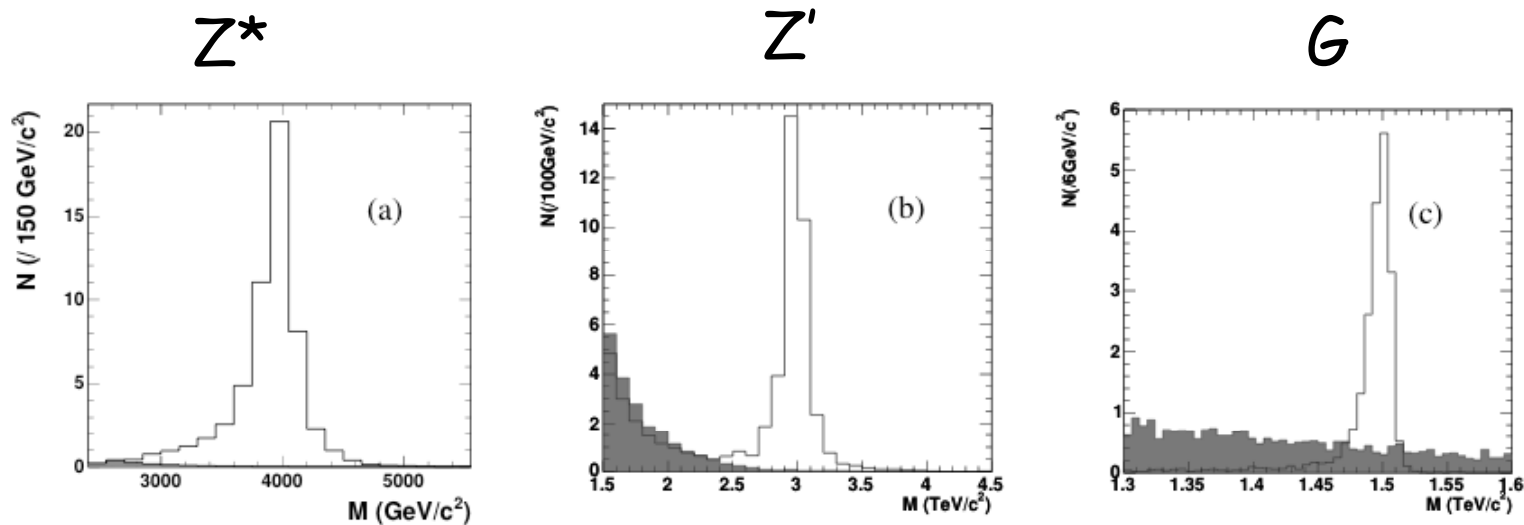
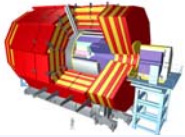
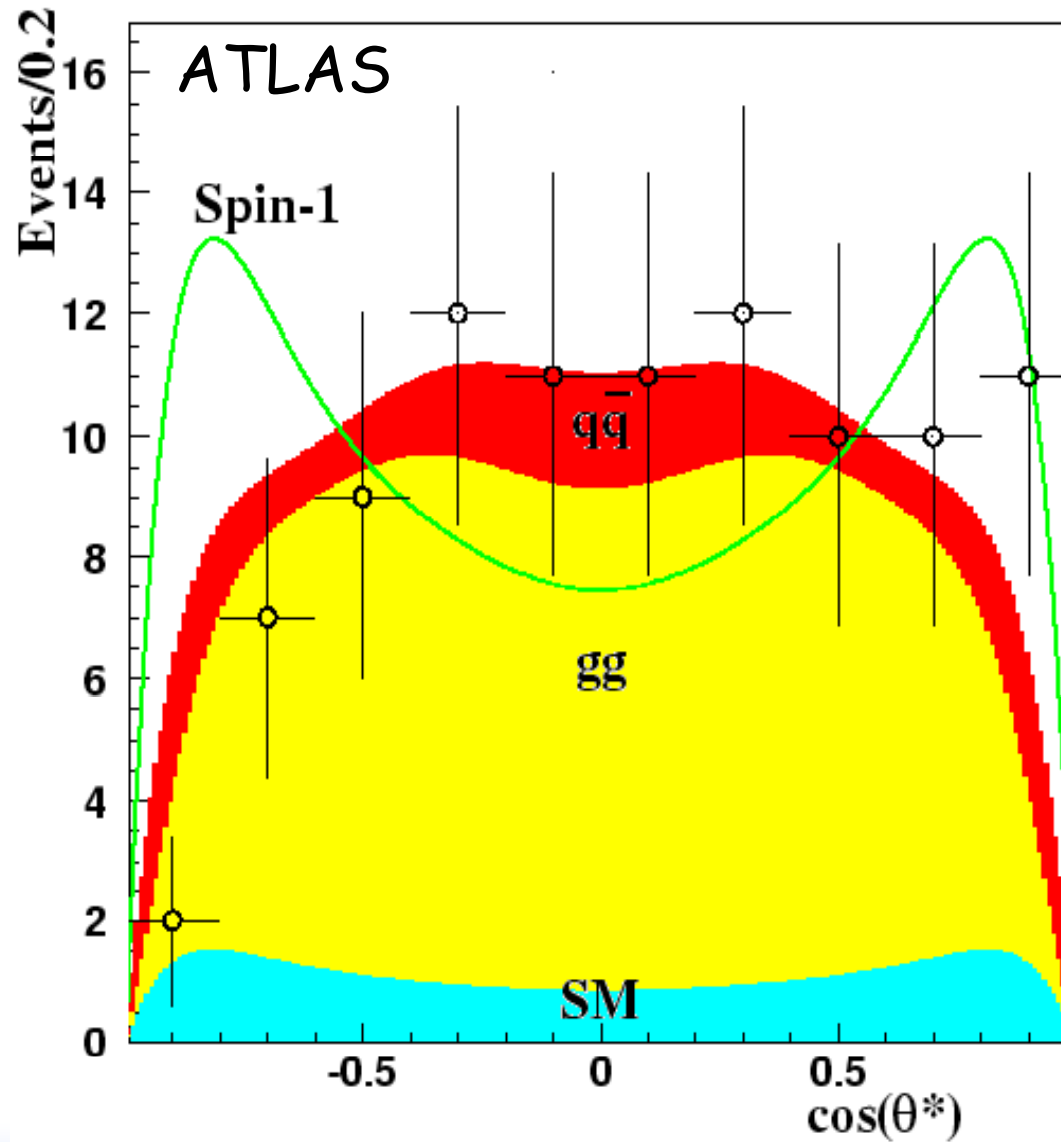
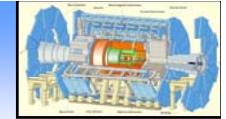


Figure 14.3: Resonance signal (white histograms) and Drell-Yan background (shaded histograms) for KK Z boson production with $M = 4.0 \text{ TeV}/c^2$ (a), SSM Z' boson production with $M = 3.0 \text{ TeV}/c^2$ (b), and graviton production with $M = 1.5 \text{ TeV}/c^2$, coupling parameter $c = 0.01$ (c), for an integrated luminosity of 30 fb^{-1} .



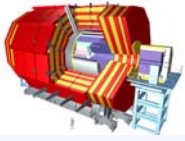
Angular distribution



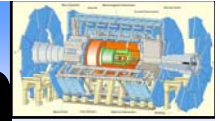
$$G \rightarrow e^+e^-$$

1.5 TeV resonance
Production dominantly
from gluon fusion
Statistics for 100fb^{-1}

Acceptance removes
events at high $\cos \theta^*$



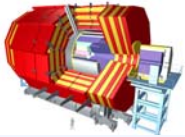
Exploring the extra dimension



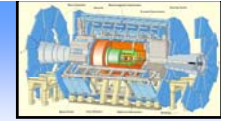
- Check that the coupling of the resonance is universal: measure rate in as many channels as possible:
 $\mu\mu, \gamma\gamma, jj, bb, tt, WW, ZZ$
- Use information from angular distribution to separate gg and qq couplings
- Estimate model parameters k and r_c from resonance mass and $\sigma.B$
- In model with $M_G=1.5$ TeV, measure mass to ± 1 GeV and $\sigma.B$ to 14% from ee channel alone (dominated by statistics).
- Then measure

$$r_c = (8.2 \pm 0.6) \times 10^{-32} m$$

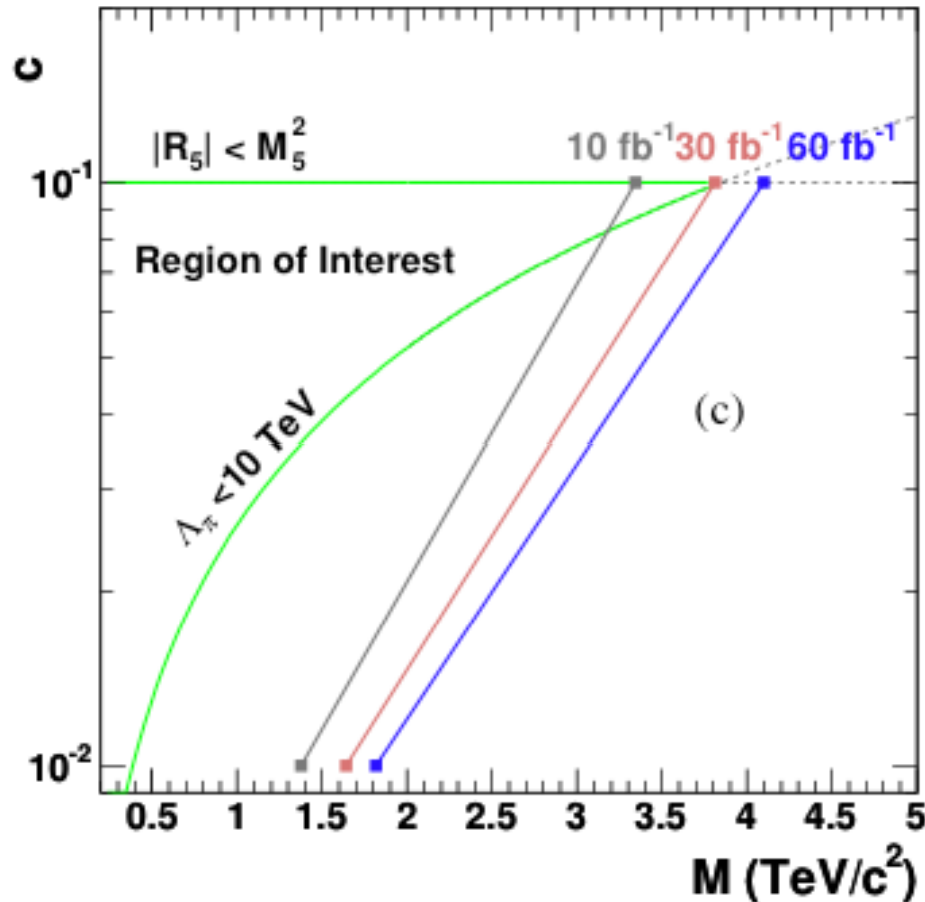
$$k = (2.43 \pm 0.17) \times 10^{16} \text{ GeV}$$



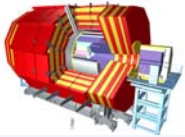
Graviton search reach



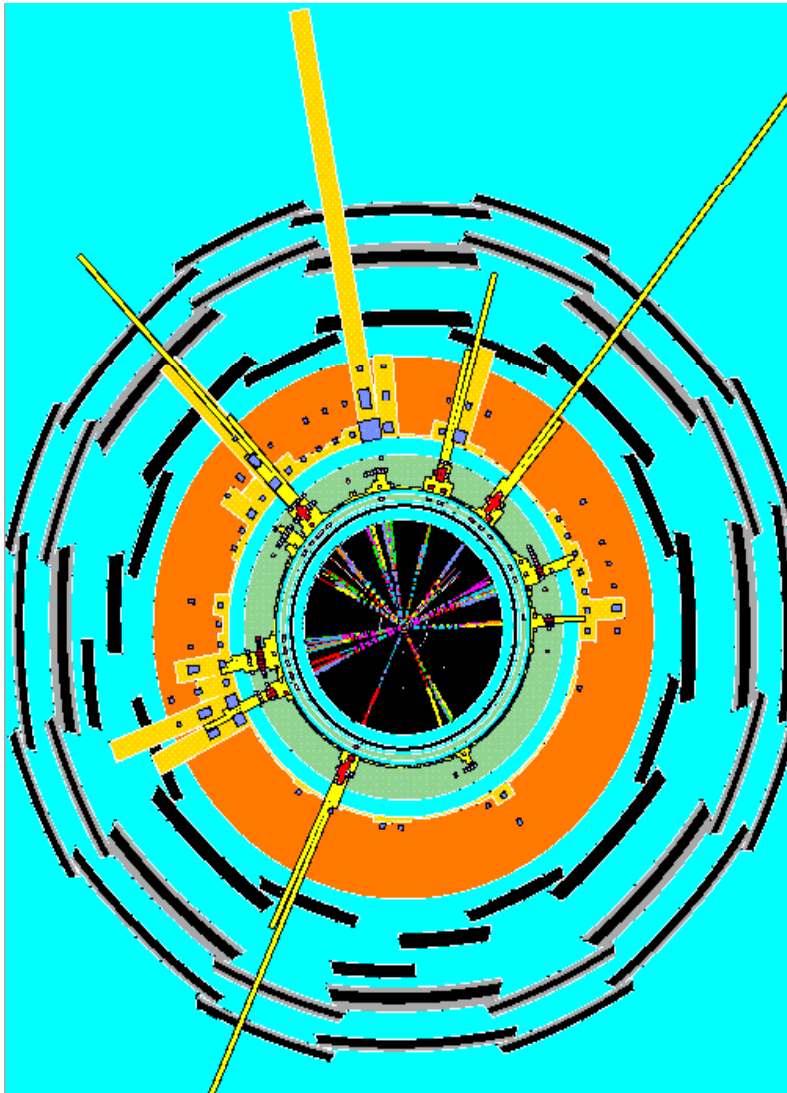
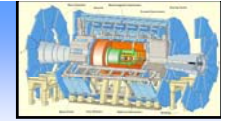
CMS



All of the theoretically interesting parameter space can be excluded with 30 fb⁻¹



Black Hole Production



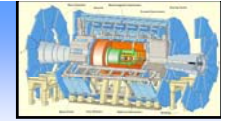
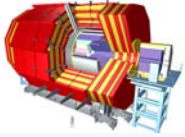
Low scale gravity in extra dimensions allows black hole production at colliders.

Decay by Hawking radiation (without eating the planet)

8 TeV mass black hole decaying to leptons and jets in ATLAS

8 partons produced with $p_T > 500 \text{ GeV}$

Richardson, Harris, Palmer:
JHEP 05(2005)053

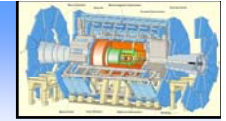
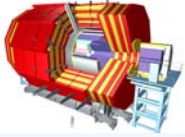


Black Hole production cross-section

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

$$\sigma_{BH} \sim \pi r_h^2$$

- Classical approximation to cross-section
- Controversial...see review by Gingrich [hep-ph/0609055](https://arxiv.org/abs/hep-ph/0609055)
- Very large rates for $n=2-6$ See [hep-ph/0111230](https://arxiv.org/abs/hep-ph/0111230)
- Almost independent of n



Black Hole Decay

Decay occurs by Hawking radiation, modified by "grey body" factors

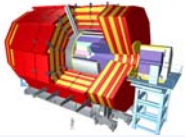
Hawking Temperature T_H

$$T_H = (n + 1) / 4\pi r_h$$

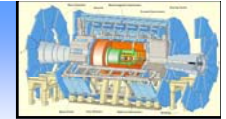
Black Hole radius r_h

$$r_h \sim \frac{\hbar}{M_{Pl} C} \left(\frac{m_{BH}}{M_{Pl}} \right)^{1/n+1}$$

Use observed final state energy spectrum to measure T_H and hence n ?



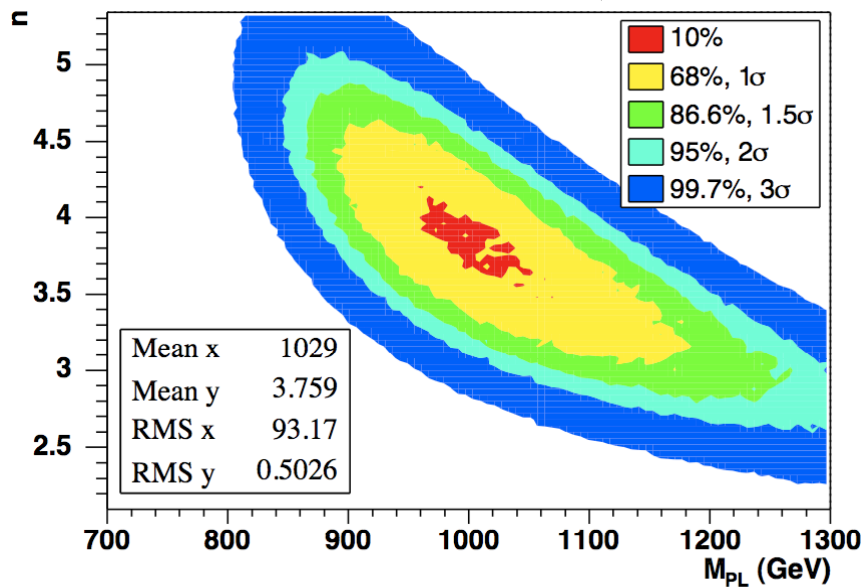
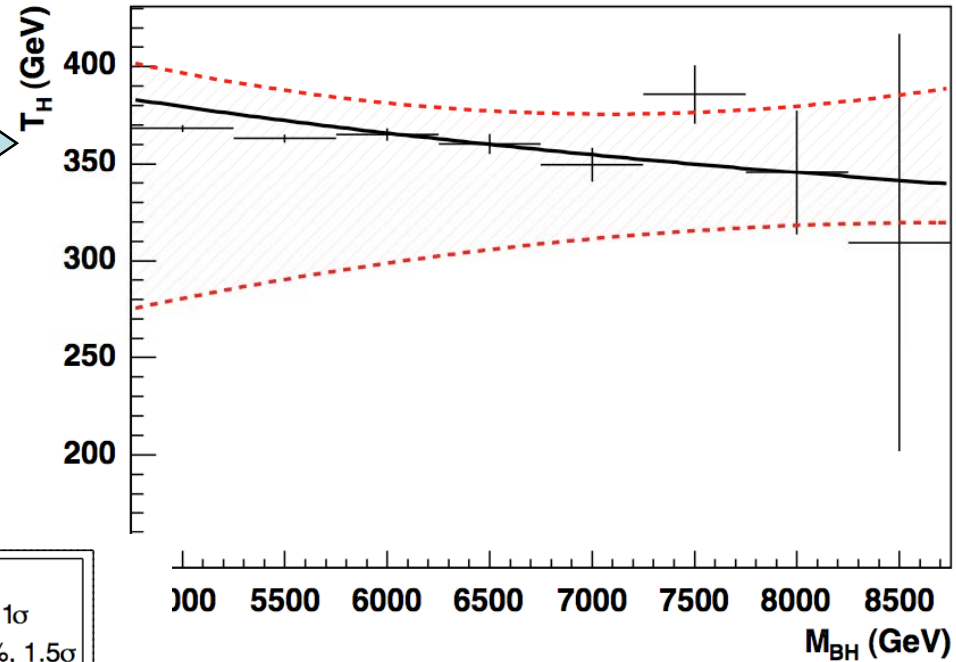
Black Hole Searches

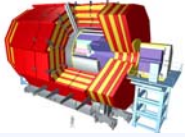


Can measure characteristic T
at average mass

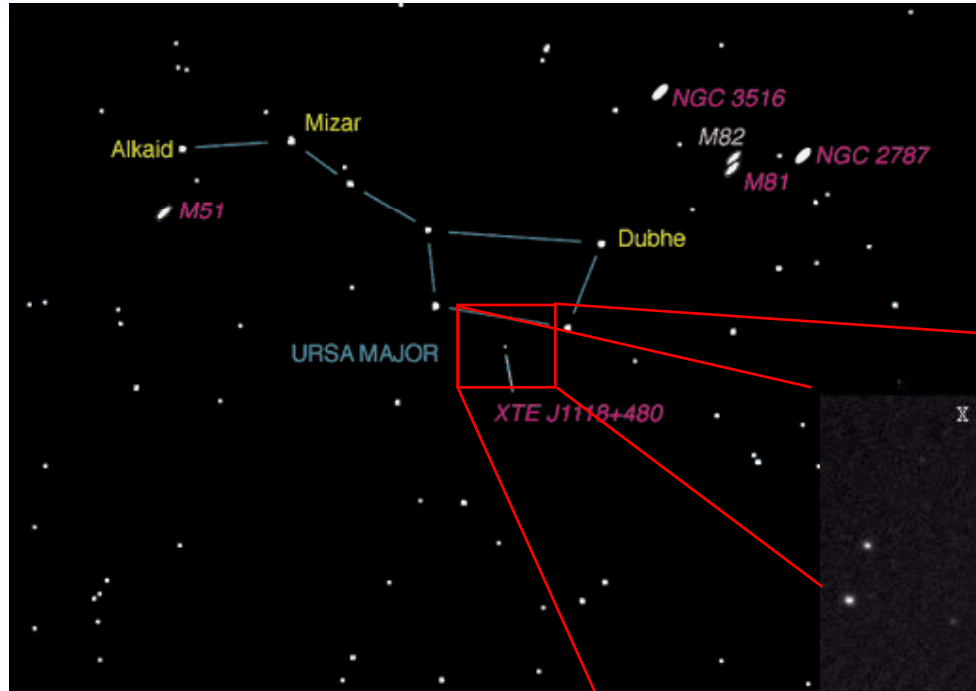
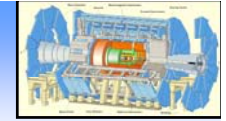
-> combine this with cross
section data to extract n .

Assume 20% error on σ





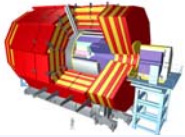
XTE J1118 +480



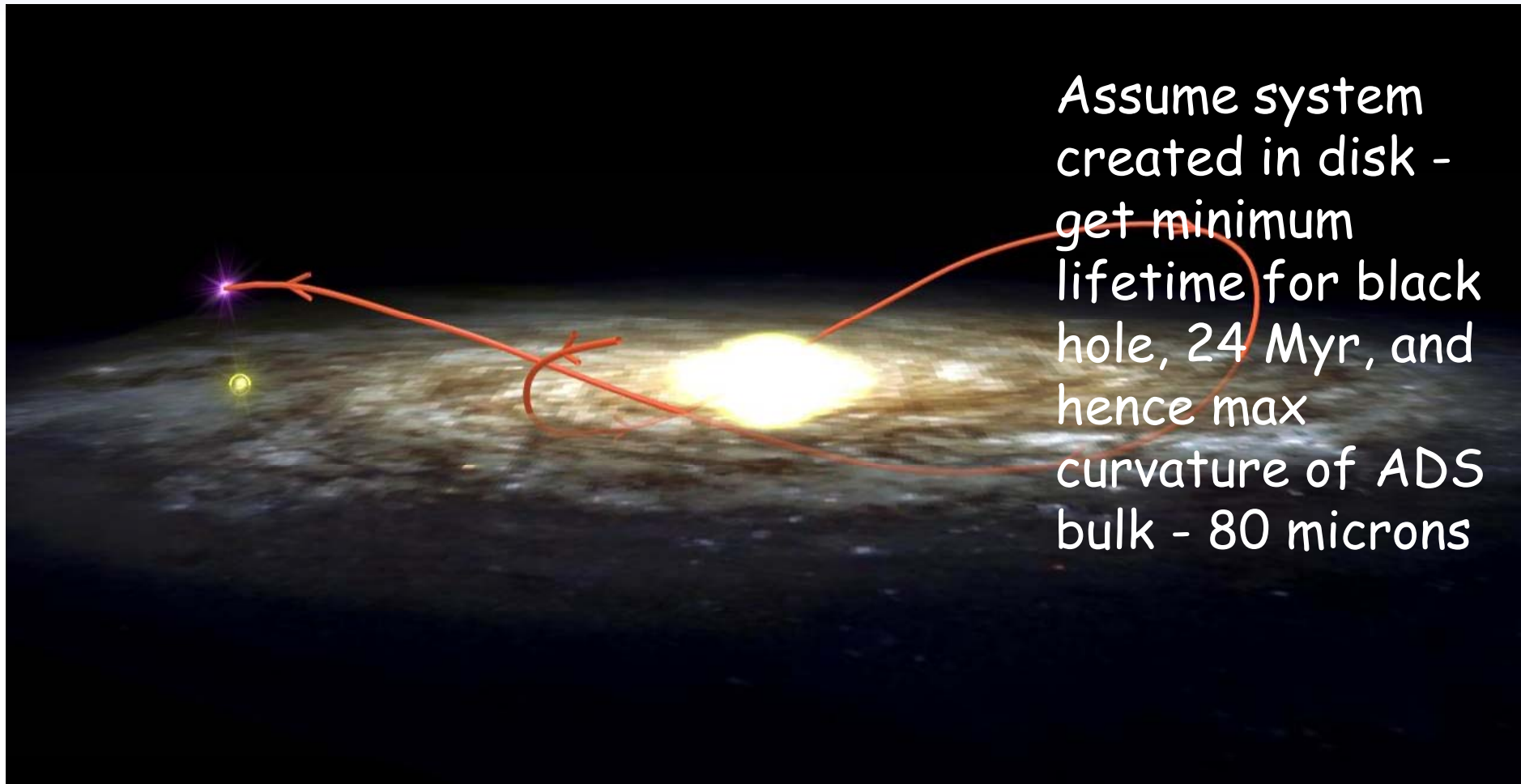
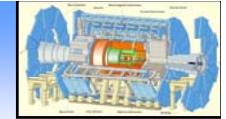
X-ray nova
observed in 2000
7 solar masses

Black hole with
companion star,
moving at 145 km/s
in galactic halo



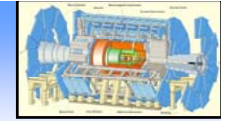
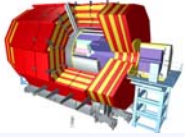


XTE J1118 +480



Assume system
created in disk -
get minimum
lifetime for black
hole, 24 Myr, and
hence max
curvature of ADS
bulk - 80 microns

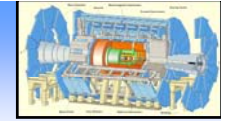
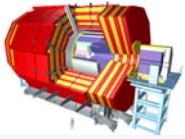
Inferred orbit passes through galactic plane several times.



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BACK-UP SLIDES