

Cosmology with voids

A look at the underdense side of the Large Scale Structure

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Outline

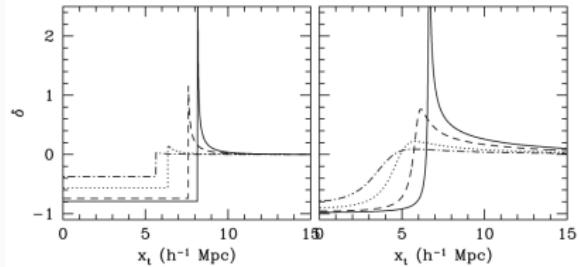
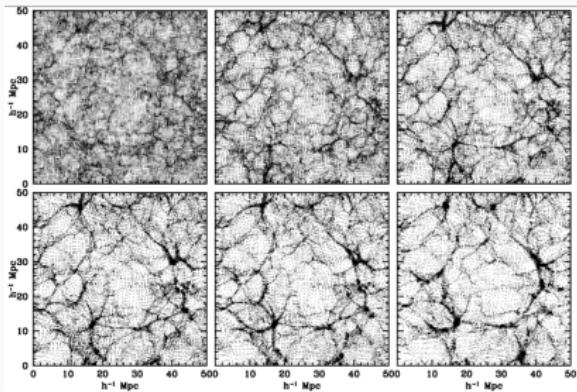
1. Cosmic voids in the Large Scale Structure
2. Towards a reliable cosmic void statistics
3. Conclusions & future developments

Cosmic voids in the Large Scale Structure

Cosmic Voids

Theoretical definition:

- \approx spherical top-hat shape (Icke, 1984)
- faster than Hubble flow $v_{exp} \propto \bar{\rho}(< r_{shell})$
- $v_{inner} > v_{outer} \Rightarrow$ shell crossing: $\delta_v = -2.71$
- formed void \Rightarrow Hubble expansion



- simple structure/shape
- spherical expansion
- $\delta_{min} = -1 \Rightarrow$ mildly non-linear

Potential cosmological probes:

AP-test, ISW, DE equation of state, theories of gravity, coupled DE, nature of DM, redshift space distortions...

Cosmic void size function

Linear theory

$$\frac{dn_L}{d \ln r_L} = \frac{\rho}{M} f_{\ln \sigma}(\sigma, \delta_c, \delta_v) \frac{d \ln \sigma^{-1}}{d \ln r_L} = \frac{f_{\ln \sigma}(\sigma, \delta_c, \delta_v)}{V(r_L)} \frac{d \ln \sigma^{-1}}{d \ln r_L}$$

$$\delta_v = -2.71 \rightarrow \delta_v^{NL} = -0.8 \rightarrow r/r_L = 1.71$$

SvdW (Sheth & van de Weijgaert, 2004)

$$\frac{dn}{d \ln r} = \left. \frac{dn_L}{d \ln r_L} \right|_{r_L=r/1.7}$$

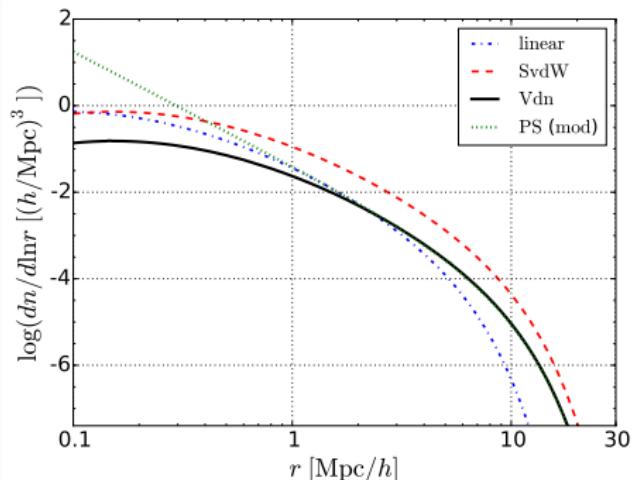
Vdn (Jennings et al., 2013)

$$V(r)dn = V(r_L)dn_L \Big|_{r_L(r)}$$

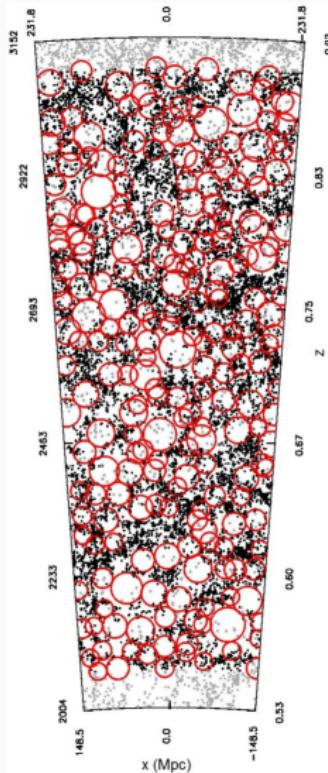
$$\frac{dn}{d \ln r} = \frac{V(r_L)}{V(r)} \left. \frac{dn_L}{d \ln r_L} \right|_{r_L(r)}$$

Modified halo-mass function

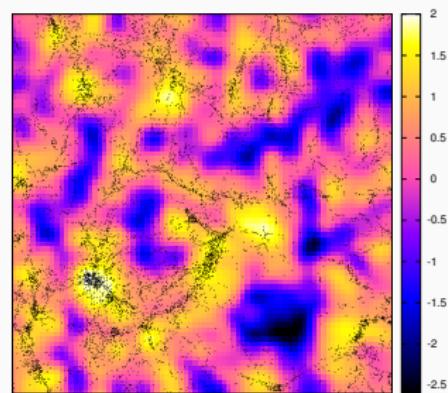
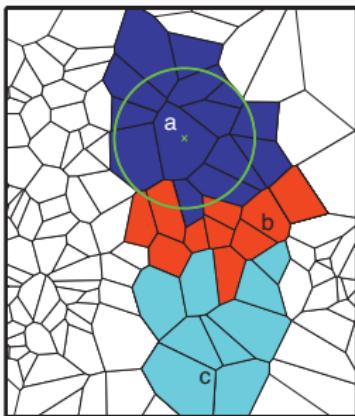
$$\frac{dn_L}{d \ln r_L} \equiv \frac{d \ln M}{d \ln r} \left(M \frac{dn}{dM} \right) = 3M \frac{dn}{dM}$$



Void finders



- density** regions with no tracers or $\rho < \rho_{\text{lim}}$
(Elyiv et al., 2013, Micheletti et al., 2014)
- geometry** Voronoi tessellation field estimator
(ZOBOV-Neyrinck, 2008; VIDE-Sutter et al., 2014)
- dynamics** recover underlying velocity field
(Lavaux & Wandelt, 2010; Elyiv et al., 2015)



Is this the end of the story?

Models + measured distribution → **NO MATCH**
GAP between theory & observations

- are the models reliable?
 - calibrate δ_v with simulations
- are the void finders doing their job?
 - search for objects coherent with models

Towards a reliable cosmic void statistics

Algorithm

input: **void-centre catalogue** (optional: scale radius, central density, density contrast)

tracer positions (DM-particles, halo/galaxy mock catalogues, galaxy surveys)

output: **new void catalogue** (spherical, shell-crossed, not overlapping)

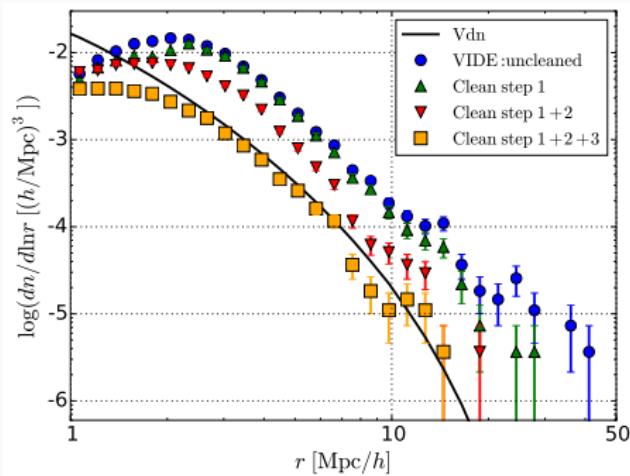
1. clean the catalogue

- $r \in [r_{\min}, r_{\max}]$
- $\delta_{\text{in}} < \delta_v^{\text{NL}}$

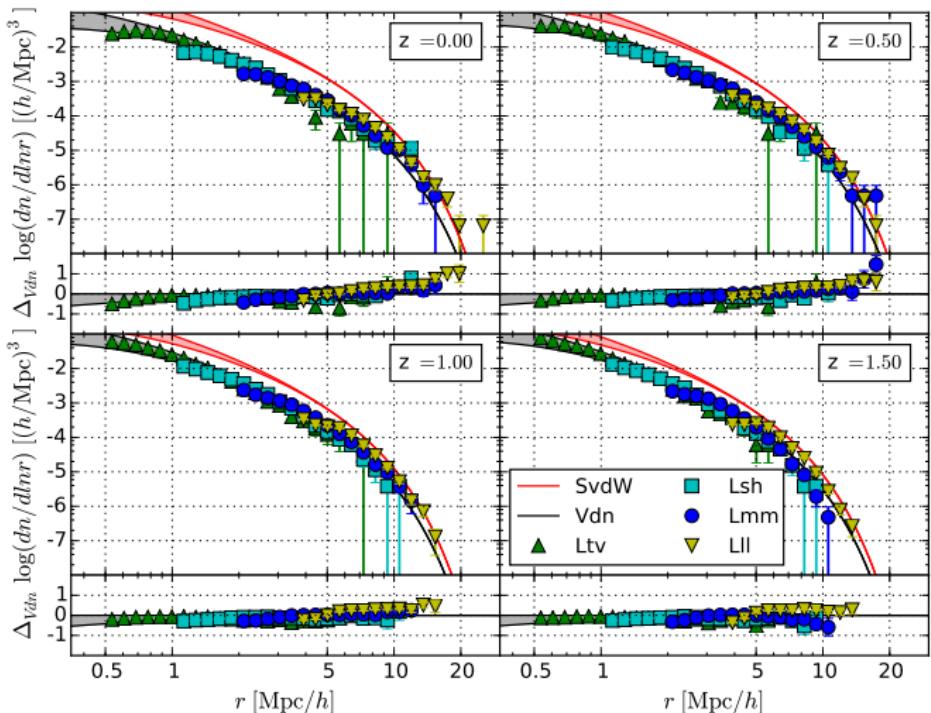
2. find shell-crossing radius

- get density profile
- find r ; $\delta(r) \equiv \delta_v^{\text{NL}}$

3. remove overlapping voids



Validation on unbiased tracers



(Ronconi et al., *in prep.*)

When it comes to bias...

Void-catalogue in a biased sample

- fix threshold:
 $\delta_{h,v}^{\text{NL}} \equiv \delta_v^{\text{NL}}$
- find radius:
 $r_{\text{eff},h} \neq r_{\text{eff,DM}}$

Void size-function

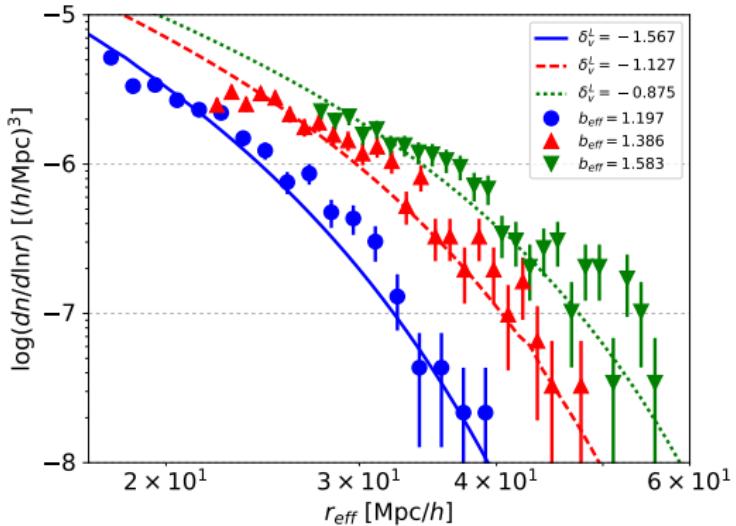
Underdensity threshold:

$$\delta_{v,\text{DM}}^{\text{NL}} = b_{\text{eff}} \delta_{v,\text{tr}}^{\text{NL}}$$

(Pollina et al., 2016)

Vdn model:

$$\frac{dn}{d \ln r} = \frac{f_{\ln \sigma}(\sigma, \delta_c, \delta_v)}{V(r)} \left. \frac{d \ln \sigma^{-1}}{d \ln r_L} \right|_{r_L(r)}$$



(Ronconi et al., *in prep.*)

Conclusions & future developments

Summary

- Developed an algorithm capable of
 - finding cosmic voids as coherent as possible with the theoretical definition (**spherical**, gone through **shell-crossing, not overlapping**)
 - allow comparison between different void-catalogues
- Validated the Vdn model without free parameters
- Extended the formalism to biased samples

Future developments:

- Parameters extraction from cosmic void number counts (Fisher information matrix)
- Void detection in HI intensity mapping (M.Viel, A.Lapi)
- Growing interest for the Euclid mission (In preparation: White paper on voids from the Euclid Void Group)