Introduction to Neutrino Physics

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The 2nd Mini-School on "Experimental Tools in Particle Physics" at CFP at Zewail City of Science and Technology, Giza, Egypt

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1 Neutrinos Facts

2 Neutrino Masses and Mixing

3 Neutrino Oscillations

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1 Neutrinos Facts

- **2** Neutrino Masses and Mixing
- **3** Neutrino Oscillations
- 4 Questions, References & Thanks

Outline

1 Neutrinos Facts

2 Neutrino Masses and Mixing

3 Neutrino Oscillations

- Neutrinos are elementary particles.
- Neutrinos are electrically neutral.
- Neutrinos are fermions with spin- $\frac{1}{2}$.
- Neutrinos are leptons.
- Neutrinos interact only via weak interactions and gravity.
- Neutrinos exist in three flavors: electron neutrino ν_e , muon neutrino ν_{μ} and tau neutrino ν_{τ} .
- Neutrino flavors are created in weak interactions in association with the corresponding charged lepton.
- Neutrinos are tiny massive $\mathcal{O}(eV)$.
- Neutrinos are only left-handed.

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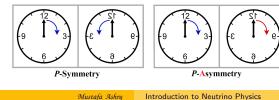
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Parity



Parity is the transformation under space reflection.

- Parity was assumed at the beginning to be a symmetry of nature.
- That's a 'mirrored' image of a natural system behaves in the same way does the 'mirror' image of that system.

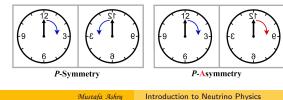


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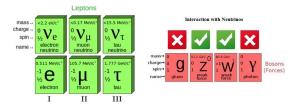


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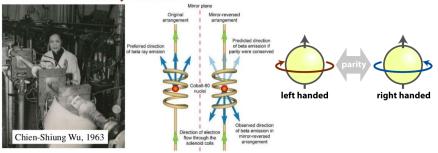


P-Symmetry

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Wu Experiment & Parity Violation

- Parity is a symmetry for both the strong and electromagnetic interactions.
- In 1963, Wu found that Parity is maximally violated in the weak interactions.



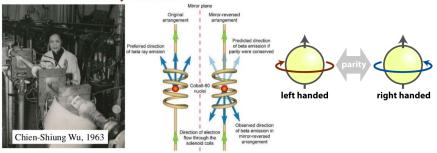
Parity Violation

Only left-handed fermions feel the weak interactions.

Mustafa Ashry Intro

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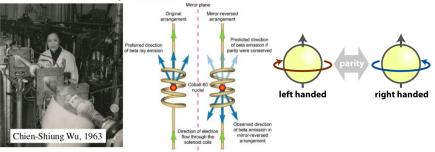
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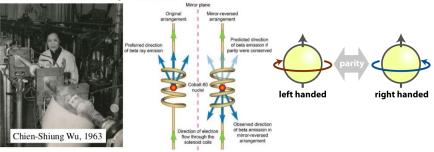
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Introduction to Neutrino Physics

-Neutrino Masses and Mixing

Outline

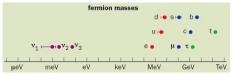
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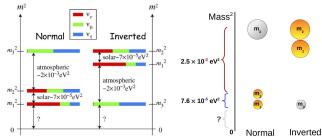
3 Neutrino Oscillations

Neutrino Masses Hierarchies

Neutrinos masses are of order of little eV's.



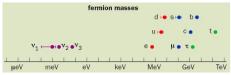
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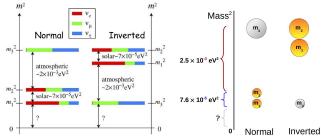
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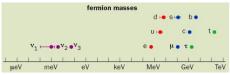
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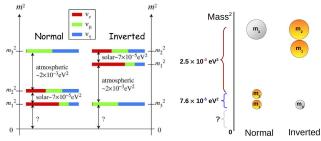
Introduction to Neutrino Physics

Neutrino Masses Hierarchies

Neutrinos masses are of order of little eV's.



■ Limits from solar and atmoshperic neutrino experiments propose the normal and inverted hierarchies for the neutrino masses.



Introduction to Neutrino Physics

Neutrino Masses->Neutrino Mixing

- One of the new phenomena that can occur if neutrinos have nonzero masses is neutrino mixing [1].
- This is the assumption that the neutrino states ν_e , ν_μ and ν_τ that couple to electrons, muons and tauons, respectively, do not have definite masses.
- They are linear combinations of three other states ν₁, ν₂ and ν₃ that do have definite masses m₁, m₂ and m₃.

$$\left(\begin{array}{c}\nu_e\\\nu_\mu\\\nu_\mu\\\nu_\tau\end{array}\right) = \left(\begin{array}{ccc}U_{e1}&U_{e2}&U_{e3}\\U_{\mu1}&U_{\mu2}&U_{\mu3}\\U_{\tau1}&U_{\tau2}&U_{\tau3}\end{array}\right) \left(\begin{array}{c}\nu_1\\\nu_2\\\nu_3\end{array}\right)$$

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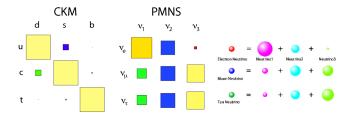
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CKM & PMNS Mixing matrices

The CKM quark mixing matrix is almost diagonal.

The PMNS mixing matrix of neutrinos is equibrated; i.e., all elements have approximately the same order.

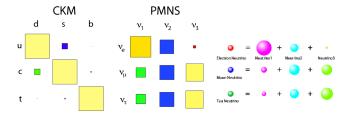
$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} = \begin{bmatrix} 0.82 \pm 0.01 & 0.54 \pm 0.02 & -0.15 \pm 0.03 \\ -0.35 \pm 0.06 & 0.70 \pm 0.06 & 0.62 \pm 0.06 \\ 0.44 \pm 0.06 & -0.45 \pm 0.06 & 0.77 \pm 0.06 \end{bmatrix}$$



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-Neutrino Oscillations

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Beats, Lagrangian

The classical Lagrangian for forced oscillations of a particle of mass m by a force F(t) in one dimension x is [2]

$$L = \frac{1}{2}m\dot{x}^{2} - \frac{1}{2}m\omega^{2}x^{2} + xF(t), \qquad (1)$$

- where ω is the frequency of the free oscillations.
- The *Euler-Lagrange* equation of motion of the system is

$$\ddot{x} + \omega^2 x = F(t)/m. \tag{2}$$

• The free oscillations $(F(t) \equiv 0)$ are

$$x(t) = A\cos(\omega t + \alpha), \qquad ($$

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where A is the amplitude and α is the initial phase.

• We consider an oscillatory force F(t) of the form

$$F(t) = f \cos(\gamma t + \beta). \tag{4}$$

Near the resonance (γ = ω), γ = ω + ε and ε is a small quantity, the motion may be regarded as small oscillations with variable amplitude

$$x(t) = C(t) \exp(i\omega t), \tag{5}$$

where

$$C^{2} = a^{2} + b^{2} + 2ab\cos\left(\epsilon t + \beta - \alpha\right), \tag{6}$$

and *a*, *b* are constants.

Clearly, the amplitude varies slowly with frequency ϵ between the limits

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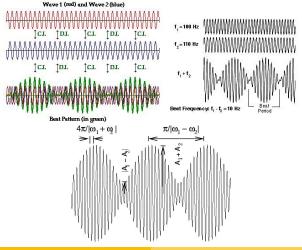
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Beats, Illustration

■ Click, enjoy ;) ● BEATS.



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Two flavor Mixing Approximation

• We consider the mixing between only two of the flavor states ν_{α} and ν_{β} :

$$\nu_{\alpha} = \nu_{i} \cos \theta_{ij} + \nu_{j} \sin \theta_{ij}, \qquad (8)$$

$$\nu_{\beta} = -\nu_{i} \sin \theta_{ij} + \nu_{j} \cos \theta_{ij}, \qquad (9)$$

where ν_i, ν_j are mass eigenstates involved.

- When a ν_{α} neutrino is produced with momentum \vec{p} at time t = 0, the ν_i and ν_j components will have slightly different energies E_i and E_j due to their slightly different masses.
- In quantum mechanics, their associated waves will therefore have slightly different frequencies, giving rise to the '*beats*' phenomenon.

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- In quantum mechanics, their associated waves will therefore have slightly different frequencies, giving rise to the 'beats' phenomenon.

- As a result, one finds that the original beam of ν_{α} particles develops a ν_{β} component whose intensity oscillates as it travels through space, while the intensity of the ν_{α} neutrino beam itself is correspondingly reduced.
- These are called 'neutrino oscillations' and their existence follows from simple quantum mechanics.
- We consider now that a ν_{α} produced with momentum \vec{p} at time t = 0, so that the initial state (8) is

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After time t this will become, using the equation of motion for the mass states ν_i, ν_j

$$a_i(t)|\nu_i,\vec{p}\rangle\cos\theta_{ij}+a_j(t)|\nu_j,\vec{p}\rangle\sin\theta_{ij},\qquad(11)$$

where

$$a_{i,j}(t) = \exp\left(-iE_{i,j}t\right) \tag{12}$$

are the usual oscillating time factors associated with any quantum mechanical stationary state.

Similarly the initial state (9)

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Propagation

■ Inverting eqs. (8,9) to obtain

$$\nu_{i} = \nu_{\alpha} \cos \theta_{ij} - \nu_{\beta} \sin \theta_{ij}, \qquad (14)$$
$$\nu_{j} = \nu_{\alpha} \sin \theta_{ij} + \nu_{\beta} \cos \theta_{ij}. \qquad (15)$$

■ Then substitute into eq. (11) to obtain

$$|\nu_{\alpha}(t),\vec{p}\rangle = A(t)|\nu_{\alpha}(0),\vec{p}\rangle + B(t)|\nu_{\beta}(0),\vec{p}\rangle, \tag{16}$$

where

$$A(t) = a_i(t)\cos^2\theta_{ij} + a_j(t)\sin^2\theta_{ij},$$

$$B(t) = \sin\theta_{ij}\cos\theta_{ij}[a_j(t) - a_i(t)].$$
(17)
(18)

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$$\nu_{j} = \nu_{\alpha} \sin \theta_{ij} + \nu_{\beta} \cos \theta_{ij}. \qquad (15)$$

■ Then substitute into eq. (11) to obtain

 $|
u_{\alpha}(t), \vec{p}\rangle = A(t)|
u_{\alpha}(0), \vec{p}\rangle + B(t)|
u_{\beta}(0), \vec{p}\rangle,$ (16)

where

$$A(t) = a_i(t)\cos^2\theta_{ij} + a_j(t)\sin^2\theta_{ij},$$

$$B(t) = \sin\theta_{ij}\cos\theta_{ij}[a_j(t) - a_i(t)].$$
(17)
(18)

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$$|
u_{lpha}(t),ec{
ho}
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u_{lpha}(0),ec{
ho}
angle+B(t)|
u_{eta}(0),ec{
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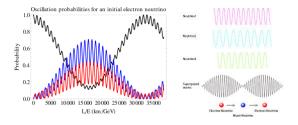
Probability

• The probability of finding a ν_{β} state is therefore

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = |\langle \nu_{\beta}(0), \vec{p} | \nu_{\alpha}(t), \vec{p} \rangle$$
$$= |B(t)|^{2} = \sin^{2}(2\theta_{ij}) \sin^{2}[\frac{1}{2}(E_{j} - E_{i})t].$$
(19)

The probability of finding a u_{α} state is therefore

 $P(\nu_{\alpha} \rightarrow \nu_{\alpha}) = |\langle \nu_{\alpha}(0), \vec{p} | \nu_{\alpha}(t), \vec{p} \rangle = |A(t)|^2 = 1 - |B(t)|^2.$ (20)



Mustafa Ashry

Introduction to Neutrino Physics

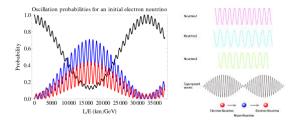
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$$egin{aligned} P(
u_lpha o
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u_eta(0), ec p |
u_lpha(t), ec p
angle \ &= |B(t)|^2 = \sin^2(2 heta_{ij})\sin^2[rac{1}{2}(E_j - E_i)t]. \end{aligned}$$

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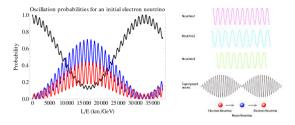


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- No oscillation if the mixing angle vanishes.
- For large energy difference $E_j E_i$, the oscillation may be within the uncertainty time (!).
- For small mixing angles and large energy difference, the oscillation is negligible.
- These formulas assume that the neutrinos are propagating in a vacuum. This is usually a very good approximation, because of the enormous mean free paths for neutrinos to interact with matter.
- It was shown that neutrino oscillations can be enhanced when neutrinos traverse very long distances in matter.
- This is the Mikheyev-Smirnov-Wolfenstein (MSW) effect, due to weak interactions with matter's electrons analogous to the electromagnetic process leading to the refractive index of light in a medium.
- The MSW effect was dramatically confirmed in the 'Sudbury Neutrino Observatory (SNO)', and resolved the solar neutrino problem.

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Detection

- The time t traveled by a neutrino is determined by the distance L of the neutrino detector from the source of the neutrinos.
- Neutrino momenta are always much greater than their possible masses and they travel, to a very good approximation, at the speed of light.
 In this approximation, t = L,

$$E_j - E_i = (m_j^2 + p^2)^{1/2} - (m_i^2 + p^2)^{1/2} \approx \frac{m_j^2 - m_i^2}{2p}.$$
 (21)

Accordingly, the probability (19) may be written

$$P(\nu_{\alpha} \to \nu_{\beta}) = \sin^2(2\theta_{ij})\sin^2[L/L_0], \qquad (22)$$

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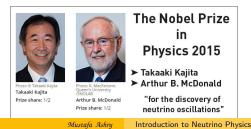
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Detection

Also

$$P(\nu_{\alpha} \to \nu_{\alpha}) = 1 - P(\nu_{\alpha} \to \nu_{\beta}).$$
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- The oscillation lengths are typically of order 100 km or more.
- Oscillations can be safely neglected under normal laboratory conditions.
- The Nobel Prize in Physics 2015 for the discovery of neutrino oscillations. Click NPNO, NPNO.pdf.

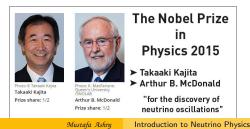


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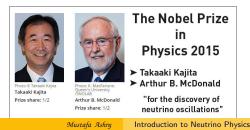
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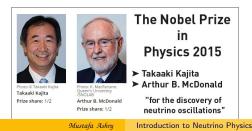
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Questions, References & Thanks

Outline

1 Neutrinos Facts

2 Neutrino Masses and Mixing

3 Neutrino Oscillations

4 Questions, References & Thanks

Questions, References & Thanks

Q?



Mustafa Ashry

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References

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Thanks

