# S3 multi-Higgs models: DM and leptogenesis

Myriam Mondragón UNAM

> Arturo Alvarez Estela Garcés Catalina Espinoza Humberto Reyes

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## Outline

- Our strategy
- S3 models with extended Higgs sector
- Quarks and leptons
- DM in S3-4H
- Leptogenesis in S3-3H
- Conclusions

# STANBARD MODEL $\begin{aligned} \mathcal{J} &= -\frac{1}{4} \int_{\infty}^{\infty} \int_{0}^{\infty} f^{**} \\ &+ i \mathcal{F} \mathcal{D} \mathcal{F} + h.c. \\ &+ \mathcal{F} \mathcal{H}_{ij} \mathcal{F}_{j} \mathcal{P} + h.c. \\ &+ |Q \mathcal{A}|^{2} - V(\mathcal{O}) \end{aligned}$ LAGRANGIAN

# WHAT PART OF

 $-\tfrac{1}{2}\partial_{\nu}g^{a}_{\mu}\partial_{\nu}g^{a}_{\mu} - g_{s}f^{abc}\partial_{\mu}g^{a}_{\nu}g^{b}_{\mu}g^{c}_{\nu} - \tfrac{1}{4}g^{2}_{s}f^{abc}f^{adc}g^{b}_{\mu}g^{c}_{\nu}g^{d}_{\mu}g^{e}_{\nu} + \tfrac{1}{2}ig^{2}_{s}(\bar{q}^{\sigma}_{i}\gamma^{\mu}q^{\sigma})g_{\mu}$  $\bar{G}^{a}\partial^{2}G^{a} + g_{a}f^{abc}\partial_{\mu}G^{a}G^{b}g^{c}_{\mu} - \partial_{\nu}W^{+}_{\mu}\partial_{\nu}W^{-}_{\mu} - M^{2}W^{+}_{\mu}W^{-}_{\mu} - \frac{1}{2}\partial_{\nu}Z^{0}_{\mu}\partial_{\nu}Z^{0}_{\mu} - \frac{1}{2c^{2}}M^{2}Z^{0}_{\mu}Z^{0}_{\mu}$  $\frac{1}{2}\partial_{\mu}A_{\nu}\partial_{\mu}A_{\nu} - \frac{1}{2}\partial_{\mu}H\partial_{\mu}H - \frac{1}{2}m_{h}^{2}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{-} - M^{2}\phi^{+} - M^{2}\phi^{$  $\frac{1}{2c_{L}}M\phi^{0}\phi^{0} - \beta_{h}[\frac{2M^{2}}{2} + \frac{2M}{9}H + \frac{1}{2}(H^{2} + \phi^{0}\phi^{0} + 2\phi^{+}\phi^{-})] + \frac{2M}{9^{2}}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}(W^{+}_{\mu}W^{-}_{\nu} - \phi^{-})] + \frac{2M}{9}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}(W^{+}_{\mu}W^{-}_{\mu}W^{-}_{\mu})] + \frac{2M}{9}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}(W^{+}_{\mu}W^{-}_{\mu}W^{-}_{\mu})] + \frac{2M}{9}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}(W^{+}_{\mu}W^{-}_{\mu}W^{-}_{\mu})] + \frac{2M}{9}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}(W^{+}_{\mu}W^{-}_{\mu}W^{-}_{\mu})] + \frac{2M}{9}\alpha_{h} - igc_{\omega}[\partial_{\nu}Z^{0}_{\mu}W^{ W_{\nu}^{+}W_{\mu}^{-}) - Z_{\nu}^{0}(W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+}) + Z_{\mu}^{0}(W_{\nu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}A_{\mu}(W_{\mu}^{-}W_{\nu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - igs_{\omega} \partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - igs_{\omega} \partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\nu}W_{\mu}^{-} - igs_{\omega} \partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\mu}W_{\mu}^{-} - igs_{\omega} \partial_{\mu}W_{\mu}^{-})] - igs_{\omega} \partial_{$  $W_{\nu}^{+}W_{\mu}^{-}) - A_{\nu}(W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+}) + A_{\mu}(W_{\nu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - \frac{1}{2}g^{2}W_{\mu}^{+}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-} + \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}^{-}W_{\nu}^{+}W_{\nu}^{-} + \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}^{-}W_{\nu}^{+}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-}W$  $\frac{1}{2}g^2W_{\mu}^+W_{\nu}^-W_{\mu}^+W_{\nu}^-+g^2c_{\omega}^2(Z_{\mu}^0W_{\mu}^+Z_{\mu}^0W_{\nu}^--Z_{\mu}^0Z_{\mu}^0W_{\nu}^+W_{\nu}^-)+g^2s_{\omega}^2(A_{\mu}W_{\mu}^+A_{\nu}W_{\nu}^- A_{\mu}A_{\mu}W_{\nu}^{+}W_{\nu}^{-}) + g^{2}s_{\omega}c_{\omega}A_{\mu}Z_{\nu}^{0}(W_{\mu}^{+}W_{\nu}^{-} - W_{\nu}^{-}W_{\mu}^{-}) - 2A_{\mu}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-}] - g\alpha[H^{3} + K_{\mu}^{-}W_{\nu}^{-}W_{\nu}^{-}] + g^{2}s_{\omega}c_{\omega}A_{\mu}Z_{\nu}^{0}(W_{\mu}^{+}W_{\nu}^{-} - W_{\nu}^{-}W_{\mu}^{-}) - 2A_{\mu}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-}] - g\alpha[H^{3} + K_{\mu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + K_{\mu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + K_{\mu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + K_{\mu}^{-}W_{\nu}^{-}W_$  $H\phi^{0}\phi^{0} + 2H\phi^{+}\phi^{-} - \frac{1}{2}g^{2}\alpha_{h}H^{4} + (\phi^{0})^{4} + 4(\phi^{+}\phi^{-})^{2} + 4(\phi^{0})^{2}\phi^{+}\phi^{-} + 4H^{2}\phi^{+}\phi^{-}$  $2(\phi^{0})^{2}H^{2}] - 9MW_{\mu}^{+}W_{\mu}^{-}H - \frac{1}{2}9\frac{4}{2}Z_{\mu}^{0}Z_{\mu}^{0}H - \frac{1}{2}ig[W_{\mu}^{+}(\phi^{0}\partial_{\mu}\phi^{-} - \phi^{-}\partial_{\mu}\phi^{0}) - W_{\mu}^{-}(\phi^{0}\partial_{\mu}\phi^{-})$  $\partial_{\mu}H) - W_{\mu}^{-}(H\partial_{\mu}\phi^{+} - \phi^{+}\partial_{\mu}H)] + \frac{1}{2}g_{e}^{-}(Z_{\mu}^{0}(H\partial_{\mu}\phi^{0} - \phi^{+}\partial_{\mu}H))$  $\phi^{+}\partial_{\mu}\phi^{0})] + \frac{1}{2}g[W^{+}_{\mu}(H\partial_{\mu}\phi^{-}-\phi^{-})]$  $\phi^{0}\partial_{\mu}H) ig \frac{f_{0}}{c_{\mu}}MZ^{0}_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+})+igs_{\mu}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+})-ig\frac{1-2c_{\mu}}{2c_{\mu}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}))$  $_{\mu}\phi^{-} - \phi^{-}\partial_{\mu}\phi^{+}) - \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu} H^{2} + (\phi^{0})^{2} + 2\phi^{+}\phi^{-}$  $(W_{+}^{+}\phi^{-}) - \frac{1}{2}g^{2} = Z_{\mu}^{0}\phi^{0}(W_{+}^{+}\phi^{-})$  $W^+_{\mu}\phi^- + W^-_{\mu}\phi^+) + \frac{1}{2}ig^2 s_{\omega}A_{\mu}H(W^+_{\mu}\phi^-)$  $-\bar{u}_{i}^{\lambda}(\gamma\partial + m_{\mu}^{\lambda})u_{j}^{\lambda} - d_{i}^{\lambda}(\gamma\partial + m_{d}^{\lambda}d_{j}^{\lambda} + igs_{w}A_{\mu}[-(e^{\lambda}\gamma^{\mu}e^{\lambda}) + \frac{2}{3}(\bar{u}_{i}^{\lambda}\gamma^{\mu}u_{j}^{\lambda})$  $\frac{1}{3}(d_j^{\lambda}\gamma^{\mu}d_j^{\lambda})] + \frac{4g}{4\varsigma_{\mu}}Z_{\mu}^{0}[(\bar{\nu}^{\lambda}\gamma^{\mu}(1+\gamma^5)\nu^{\lambda}) + (\bar{e}^{\lambda}\gamma^{\mu}(4s_{\mu}^2-1-\gamma^5)e^{\lambda}) - (\bar{u}_j^{\lambda}\gamma^{\mu})]$  $(1 - \gamma^5)\bar{u}_j^{\lambda}) + (d_j^{\lambda}\gamma^{\mu}(1 - \frac{8}{3}s_{\mu}^2 - \gamma^5)d_j^{\lambda})] + \frac{49}{2\sqrt{2}}W_{\mu}^+[(\nu^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda}) - (u_j^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda})] + (u_j^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda}) + (u_j^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda})$  $\gamma^{5} C_{\lambda\kappa} d_{3}^{\kappa} ] + \frac{i_{9}}{2\sqrt{2}} W_{\mu}^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) u_{3}^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} \frac{m_{e}}{M} [-\phi^{+} (\bar{\nu}^{\lambda} (1-\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} C_{\lambda\kappa} \gamma^{\mu} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} (1+\gamma^{5}) \nu^{\lambda})] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} (1+\gamma^{5}) \nu^{\lambda}] + \frac{i_{9}}{2\sqrt{2}} M^{-} [(\bar{e}^{\lambda} (1+\gamma^{5}) \nu^{\lambda}) + (d_{3}^{\kappa} (1+\gamma^{$  $\gamma^{5}(e^{\lambda}) + \phi^{-}(\bar{e}^{\lambda}(1+\gamma^{5})\nu^{\lambda})] - \frac{g}{2} \frac{m_{e}^{\lambda}}{M} \left[H(\bar{e}^{\lambda}e^{\lambda}) + i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{2}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\lambda})) + i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{2}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\lambda}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{2}(\bar{u}_{j}^{\lambda}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[$  $\gamma^{5})d_{j}^{\kappa}) + m_{u}^{\lambda}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1+\gamma^{5})d_{j}^{\kappa}] + \frac{i_{0}}{2M\sqrt{2}}\phi^{-}[m_{d}^{\lambda}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1+\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\star}(1-\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}$  $\gamma^{5}(u_{j}^{*}] = \frac{2}{2} \frac{m_{1}}{M} H(u_{j}^{\lambda}u_{j}^{\lambda}) - \frac{2}{2} \frac{m_{1}}{M} H(d_{j}^{\lambda}d_{j}^{\lambda}) + \frac{4}{2} \frac{m_{1}}{M} \phi^{0}(u_{j}^{\lambda}\gamma^{5}u_{j}^{\lambda}) + \frac{4}{2} \frac{m_{1}}{M} \phi^{0}(u_{$  $X^{+}(\partial^{2}-M^{2})X^{+}+X^{-}(\partial^{2}-M^{2})X^{-}+X^{0}(\partial^{2}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u}W^{+}(\partial_{u}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{u$  $\partial_{\mu}X^{+}X^{0}$  +  $igs_{\mu}W^{+}_{\mu}(\partial_{\mu}\bar{Y}X^{-} - \partial_{\mu}X^{+}Y) + igs_{\mu}W^{-}_{\mu}(\partial_{\mu}X X^{0} - \partial_{\mu}\bar{X}^{0}X^{+}) +$  $igs_{\omega}W^{-}_{\mu}(\partial_{\mu}X^{-}Y - \partial_{\mu}YX^{+}) + igc_{\omega}Z^{0}_{\mu}(\partial_{\mu}X^{+}X^{+} - \partial_{\mu}X^{-}X^{-}) + igs_{\omega}A_{\mu}(\partial_{\mu}X^{+}X^{+} - \partial_{\mu}X^{-}X^{-}) + igs_{\omega}A_{\mu}(\partial_{\mu}X^{+}X^{+}) + igc_{\omega}Z^{0}_{\mu}(\partial_{\mu}X^{+}X^{+}) + igc_{\omega}Z^{0}_{\mu}(\partial_{\mu}X^{+$  $\begin{array}{l} \partial_{\phi} \bar{X}^{-} X^{-} \rangle - \frac{1}{2} g M [\bar{X}^{+} X^{+} H + \bar{X}^{-} X^{-} H + \frac{1}{2^{2}} \bar{X}^{0} X^{0} H] + \frac{1 - 2c^{2}}{2c^{2}} i g M [\bar{X}^{+} X^{0} \phi^{+} - X^{0} X^{0} \phi^{-}] \\ X^{-} X^{0} \phi^{-}] + \frac{1}{2c^{2}} i g M [X^{0} X^{-} \phi^{+} - X^{0} X^{+} \phi^{-}] + i g M s_{*} [X^{0} X^{-} \phi^{+} - X^{0} X^{+} \phi^{-}] + \frac{1}{2} i g M \bar{X}^{+} X^{+} \phi^{0} - X^{-} X^{-} \phi^{0}] \end{array}$ 



DO YOU NOT

**UNDERSTAND?** 

# How to go BSM?

- Many ways to go BSM
- Usually: add symmetries, add particles, add interactions
- All of the above
- Messy...
- I will concentrate on masses and mixings



### Some aspects of the flavour problem

• Quark and charged lepton masses very different, very hierarchical

$$m_u : m_c : m_t \sim 10^{-6} : 10^{-3} : 1$$
  
 $m_d : m_s : m_b \sim 10^{-4} : 10^{-2} : 1$   
 $m_e : m_\mu : m_\tau \sim 10^{-5} : 10^{-2} : 1$ 

- Neutrino masses unknown, only difference of squared masses.
- Type of hierarchy (normal or inverted) also unknown

• Quark mixing angles

 $\theta_{12} \approx 13.0^{o}$  $\theta_{23} \approx 2.4^{o}$  $\theta_{13} \approx 0.2^{o}$ 

• Neutrino mixing angles

 $\Theta_{12} \approx 34.0^{\circ}$  $\Theta_{23} \approx 45^{\circ}$  $\Theta_{13} \approx 8.5^{\circ}$ 

- Small mixing in quarks, large mixing in neutrinos.
   Very different
- Is there an underlying symmetry?

How do we choose a flavour symmetry?

- Several ways:
- Look for inspiration in a high energy extension of SM, i.e. strings or GUTs (see S. King talk)
- Look at low energy phenomenology
- At some point they should intersect...
- In here:
  - Find the smallest flavour symmetry suggested by data
  - Explore how generally it can be applied (universally)
  - Follow it to the end
  - Compare it with the data (Neruda's quote...)



### Logarithmic plot of masses



Plot of mass ratios

Suggests a  $2 \oplus I$  structure

### Flavour/family symmetries more symmetries

- Add the flavour symmetry of your choice
- Continuous → upon breaking might generate massless Goldstone bosons U(1), SU(3)...



- Discrete
   Z<sub>N</sub>, A4, S3, S4, Q6,
   Δ27...
  - → might also generate accidental continuous symmetries
- Explain masses and mixings of quarks and leptons
- Usually also add more particles → Higgs

### Multi-Higgs models more particles

- 2HDM without SUSY
  - widely studied
  - different versions depending on how they couple to the other SM particles
- 3 or more HDM also possible
- Extra singlet Higgs models NMSSM
- Complex MSSM
   CMSSM

• FCNC's

- More sources of CP violation
- Candidates for dark matter, with some discrete symmetry
- Might give explanation for the mass hierarchies and mixing of quarks and leptons
- More scalar particles give ways to test them

# N-Higgs doublet models — NHDM

Add more complex electroweak doublets
 All with same hyper charge Y=1

$$V(\phi) = Y_{ij}\phi_i^{\dagger}\phi_j + Z_{ijkl}(\phi_i^{\dagger}\phi_j)(\phi_k^{\dagger}\phi_l).$$

- $N^2 + N^2(N^2 + 1)/2$  real parameters: 12 for 2HDM, 54 for 3HDM...
- Potential must be bounded by below, no charge or colour breaking minima
- Must respect unitarity bounds
- Can have CP breaking minima  $\rightarrow$  baryogenesis (or disaster)

# 3HDM

- Even more Higgses  $\implies$  54 real parameters in potential
- Complemented with additional symmetry(ies)
- Studies started in the 70's, hope to find global symmetry that explains the mass and mixing patterns
- The first symmetries to be added were the permutational groups S3 and S4
- Another approach was to use  $\Delta(54)$  or  $\Delta(27)$  hierarchy in masses come from vevs. Did not generate a hierarchical vev structure v3>>v2>>v1
- Different modern versions of these models exist

e.g.,I.de Medeiros, D. Emmanuel-Costa (2011) for  $\Delta(54)$  ,

# 3HDM with S3

- Low-energy model
- Extend the concept of flavour to the Higgs sector by adding two more eW doublets



- Add symmetry: permutation symmetry of three objects, symmetry operations (reflections and rotations) that leave an equilateral triangle invariant
- 3HDM with symmetry S3:
   8 couplings in the Higgs potential

### A sample of S3 models

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G. Bhattacharyya et al, Phys. Rev. D83, 011701 (2011)

D. Meloni, JHEP 1205 (2012) 124

S. Dev et al, Phys.Lett. B708 (2012) 284-289

S. Zhou, Phys.Lett. B704 (2011) 291-295

E. Barradas et al, 2014

P. Das et al, 2014

Just a sample, there are many more... I apologise for those not included

- Smallest non-Abelian discrete group
- Has irreducible representations, 2,  $I_s$  and  $I_A$
- We add three right-handed neutrinos to implement the see-saw mechanism
- We apply the symmetry "universally" to quarks, leptons and Higgs-es
  - First two families in the doublet
  - Third family in symmetric singlet
- Three sectors related, we treat them simultaneously

### Predictions, advantages?

- Possible to reparametrize mixing matrices in terms of mass ratios, successfully
- Reproduces well CKM → one less parameter as SM
- PMNS → fix one mixing angle, predictions for the other two within experimental range
- Predicts reactor mixing angle
   θI3 ≠ 0
- No extra flavons
- FCNCs suppressed by symmetry

- Higgs potential has 8 couplings
- Underlying symmetry in quark, leptons and Higgs → residual symmetry of a more fundamental one?
- Lots of Higgses:
   3 neutral, 4 charged,
   2 pseudoscalars
   Natural decoupling limit
- Further predictions will come from Higgs sector: decays, branching ratios
- A. Mondragón, M. M., F. González, E. Peinado, U. Saldaña, O. Félix, E. Rodríguez, A. Pérez, H. Reyes...; Teshima et al

### Quarks

3HDM: $G_{SM} \otimes S_3$									
	$\psi^f_L$	$\psi^f_R$	Mass matrix			Possible mass textures			
A	<b>2</b> , 1 <sub><b>S</b></sub>	<b>2</b> , 1 <sub><b>S</b></sub>	$\begin{pmatrix} \mu_1^f + \mu_2^f \\ \mu_4^f \\ \mu_8^f \end{pmatrix}$	$\begin{array}{c} \mu_4^f \\ \mu_1^f - \mu_2^f \\ \mu_9^f \end{array}$	$ \begin{array}{c} \mu_6^f \\ \mu_7^f \\ \mu_3^f \end{array} \right) \\$	$\begin{pmatrix} 0\\ \mu_2^f sc\left(3-t^2\right)\\ 0 \end{pmatrix}$	$ \begin{array}{c} \mu_{2}^{f}sc\left(3-t^{2}\right)\\ -2\mu_{2}^{f}c^{2}\left(1-3t^{2}\right)\\ \mu_{7}^{f*}/c \end{array} $	$\begin{pmatrix} 0 \\ \mu_7^f/c \\ \mu_3^f - \mu_1^f - \mu_2^f c^2 (1 - 3t^2) \end{pmatrix}$	
$A^{'}$							$\begin{pmatrix} 0 & \frac{2}{\sqrt{3}}\mu_2^f \\ \frac{2}{\sqrt{3}}\mu_2^f & 0 \\ 0 & \frac{2}{\sqrt{3}}\mu_9^f \end{pmatrix}$	$ \begin{pmatrix} 0 \\ \frac{2}{\sqrt{3}}\mu_7^f \\ \mu_3^f - \mu_1^f \end{pmatrix} $	NNI
В	$2, 1_{\mathbf{A}}$	<b>2</b> , 1 <sub><b>A</b></sub>	$\begin{pmatrix} \mu_1^f + \mu_2^f \\ \mu_4^f \\ - \mu_9^f \end{pmatrix}$	$\substack{\mu_1^f \\ \mu_1^f - \mu_2^f \\ \mu_8^f}$	$ \begin{pmatrix} \mu_7^f \\ -\mu_6^f \\ \mu_3^f \end{pmatrix} $	$\begin{pmatrix} 0\\ -\mu_4^f c^2 \left(1-3t^2\right)\\ 0 \end{pmatrix}$	$-\mu_4^f c^2 (1 - 3t^2) 2\mu_4^f sc (3 - t^2) -\mu_6^{f*}/c$	$\begin{pmatrix} 0 \\ -\mu_6^f/c \\ \mu_3^f - \mu_1^f +  \mu_4^f sc(3-t^2) \end{pmatrix}$	
$B^{'}$							$egin{pmatrix} 0 & -2\mu_4^f \ -2\mu_4^f & 0 \ 0 & 2\mu_8^f \ \end{pmatrix}$	$\begin{pmatrix} 0 \\ -2\mu_{6}^{f} \\ \mu_{3}^{f} - \mu_{1}^{f} \end{pmatrix}$	NNI

Table 2: Mass matrices in  $S_3$  family models with three Higgs  $SU(2)_L$  doublets:  $H_1$  and  $H_2$ , which occupy the  $S_3$  irreducible representation **2**, and  $H_S$ , which transforms as  $1_{\mathbf{S}}$  for the cases when both the left- and right-handed fermion fields are in the same assignment. The mass matrices shown here follow a normal ordering of their mass eigenvalues  $(m_1^f, m_2^f, m_3^f)$ . We have denoted  $s = \sin \theta$ ,  $c = \cos \theta$  and  $t = \tan \theta$ . The third column of this table corresponds to the general case, while the fourth column to a case where we have rotated the matrix to a basis where the elements (1, 1), (1, 3) and (3, 1) vanish. The primed cases, A' or B', are particular cases of the unprimed ones, A or B, with  $\theta = \pi/6$  or  $\theta = \pi/3$ , respectively.

#### Mass matrices reproduce the NNI or the Fritzsch forms

F. González et al, Phys.Rev. D88 (2013) 096004

### Leptons - S3xZ2

- Charged leptons can be also parameterised successfully, no extra free parameters
- Neutrinos: Fixing one mixing angle we obtain the other two
- Neutrinos: S3 predicts ∂<sub>13</sub> ≠ 0
   M<sub>1</sub> and M<sub>2</sub> equal → ∂<sub>13</sub> too small
   M<sub>1</sub> and M<sub>2</sub> different → ∂<sub>13</sub> in experimental range

### Higgs-es this is where the model can be tested

### General potential

$$V = \mu_{1}^{2} \left( H_{1}^{\dagger} H_{1} + H_{2}^{\dagger} H_{2} \right) + \mu_{0}^{2} \left( H_{s}^{\dagger} H_{s} \right) + a \left( H_{s}^{\dagger} H_{s} \right)^{2} + b \left( H_{s}^{\dagger} H_{s} \right) \left( H_{1}^{\dagger} H_{1} + H_{2}^{\dagger} H_{2} \right) + c \left( H_{1}^{\dagger} H_{1} + H_{2}^{\dagger} H_{2} \right)^{2} + d \left( H_{1}^{\dagger} H_{2} - H_{2}^{\dagger} H_{1} \right)^{2} + e f_{ijk} \left( \left( H_{s}^{\dagger} H_{i} \right) \left( H_{j}^{\dagger} H_{k} \right) + h.c. \right) + f \left\{ \left( H_{s}^{\dagger} H_{1} \right) \left( H_{1}^{\dagger} H_{s} \right) + \left( H_{s}^{\dagger} H_{2} \right) \left( H_{2}^{\dagger} H_{s} \right) \right\} + g \left\{ \left( H_{1}^{\dagger} H_{1} - H_{2}^{\dagger} H_{2} \right)^{2} + \left( H_{1}^{\dagger} H_{2} + H_{2}^{\dagger} H_{1} \right)^{2} \right\} + h \left\{ \left( H_{s}^{\dagger} H_{1} \right) \left( H_{s}^{\dagger} H_{1} \right) + \left( H_{s}^{\dagger} H_{2} \right) \left( H_{s}^{\dagger} H_{2} \right) + \left( H_{1}^{\dagger} H_{s} \right) \left( H_{1}^{\dagger} H_{s} \right) + \left( H_{2}^{\dagger} H_{s} \right) \left( H_{2}^{\dagger} H_{s} \right) \right\}$$
(1)

Felix-Beltrán, Rodríguez-Jáuregui, M.M (2006), Das (2010), Félix-Beltrá

- The minimum of potential can be parameterised in spherical coordinates, two angles and v
- Minimisation fixes one angle to  $v1/v2 = \pi/6$ consistent with NNI form of mass matrices
- Has natural decoupling limit
- Predictions for decays and branching rations

work in progress

### 4HDM -S3 with DM

- We add another doublet, inert, to have a DM candidate. We assign it to the IA, and thus "saturate" the irreps
- A lot of Higgses (13), but the good features of 3H-S3 remain
- S3 symmetry constrains strongly the allowed couplings

### 4H-S3 potential

- Most general Lagrangian with 4 Higgs doublets and S3 symmetry, plus Z2 symmetry
- First two generations in a flavour doublet, third in a singlet, extra antisymmetric singlet is inert → DM candidates

$$\begin{split} -\mathcal{L}_{Y_{f}} &= Y_{1}^{f}(\bar{\psi}_{S,L}^{f}\psi_{S,R}^{f}H_{s}) + \frac{1}{\sqrt{2}}Y_{2}^{f}(\bar{\psi}_{1,L}^{f}\psi_{1,R}^{f} + \bar{\psi}_{2,L}^{f}\psi_{2,R}^{f})H_{s} \\ &\quad + \frac{1}{2}Y_{3}^{f}[(\bar{\psi}_{1,L}^{f}H_{2} + \bar{\psi}_{2,L}H_{1})\psi_{1,R}^{f} + (\bar{\psi}_{1,L}H_{1} - \bar{\psi}_{2,L}^{f}H_{2})\psi_{2,R}^{f} \\ &\quad + \frac{1}{\sqrt{2}}Y_{4}^{f}(\bar{\psi}_{1,L}^{f}\psi_{2,R}^{f} - \bar{\psi}_{2,L}^{f}\psi_{1,R}^{f})H_{a} \\ &\quad + \frac{1}{\sqrt{2}}Y_{5}^{f}(\bar{\psi}_{1,L}^{f}H_{1} + \bar{\psi}_{1,L}^{f}H_{1} + \bar{\psi}_{2,L}^{f}H_{2})\psi_{S,R}^{f} \\ &\quad + \frac{1}{\sqrt{2}}Y_{6}^{f}(\bar{\psi}_{S,L}^{f}(H_{1}\psi_{1,R}^{f} + H_{2}\psi_{2,R}^{f})] + \text{h.c.} \\ &\quad f = d, e. \end{split}$$

# Higgs potential 4H-S3

• We need to find the minima of the potential S3xZ2, which satisfy the stability and unitarity conditions

$$\begin{split} V_4 &= \mu_0^2 H_s^{\dagger} H_s + \mu_1^2 (H_1^{\dagger} H_1 + H_2^{\dagger} H_2) + \mu_2^2 H_a^{\dagger} H_a \\ &+ \lambda_1 (H_1^{\dagger} H_1 + H_2^{\dagger} H_2)^2 + \lambda_2 (H_1^{\dagger} H_2 - H_2^{\dagger} H_1)^2 \\ &+ \lambda_3 [(H_1^{\dagger} H_1 - H_2^{\dagger} H_2)^2 + (H_1^{\dagger} H_2 + H_2^{\dagger} H_1)^2] \\ &+ \lambda_4 [(H_s^{\dagger} H_1) (H_1^{\dagger} H_2 + H_2^{\dagger} H_1) + (H_s^{\dagger} H_2) (H_1^{\dagger} H_1 - H_2^{\dagger} H_2) + \text{h.c.}] \\ &+ \lambda_5 (H_s^{\dagger} H_s) (H_1^{\dagger} H_1 + H_2^{\dagger} H_2) \\ &+ \lambda_6 [(H_s^{\dagger} H_1) (H_1^{\dagger} H_s) + (H_s^{\dagger} H_2) (H_2^{\dagger} H_s)] \\ &+ \lambda_7 [(H_s^{\dagger} H_1) (H_s^{\dagger} H_1) + (H_s^{\dagger} H_2) (H_s^{\dagger} H_2) + \text{h.c.}] \\ &+ \lambda_8 (H_s^{\dagger} H_s)^2 \\ &+ \lambda_9 [(H_a^{\dagger} H_2) (H_1^{\dagger} H_2 + H_2^{\dagger} H_1) - (H_a^{\dagger} H_1) (H_1^{\dagger} H_1 - H_2^{\dagger} H_2) + h.c.] \\ &+ \lambda_{10} (H_a^{\dagger} H_a) (H_1^{\dagger} H_1 + H_2^{\dagger} H_2) \\ &+ \lambda_{11} [(H_a^{\dagger} H_1) (H_1^{\dagger} H_a) + (H_a^{\dagger} H_2) (H_2^{\dagger} H_a)] \\ &+ \lambda_{12} [(H_a^{\dagger} H_1) (H_a^{\dagger} H_1) + (H_a^{\dagger} H_2) (H_a^{\dagger} H_2) + \text{h.c.}] \\ &+ \lambda_{13} (H_a^{\dagger} H_a)^2 + \lambda_{14} (H_s^{\dagger} H_a H_a^{\dagger} H_s) + \lambda_{15} [(H_1^{\dagger} H_s) (H_2^{\dagger} H_a) + h.c.]. \end{split}$$

4H-S3 potential

- We have to analyse the Higgs potential to find the proper minima:
  - Stability conditions potential bounded from below
  - Scalar S-matrix is unitary
  - Quartic Higgs couplings are perturbative
- These conditions restrict the allowed parameter space, i.e. couplings
- Check masses of DM candidates
- Calculate relic density

#### Stability conditions

 $\lambda_8 > 0$ 

 $\lambda_1 + \lambda_3 > 0$ 

 $\lambda_1 - \lambda_2 > 0$ 

 $\lambda_{13} > 0$ 

 $\lambda_5 > -2\sqrt{(\lambda_1 + \lambda_3)\lambda_8}$ 

 $\lambda_{10} > -2\sqrt{(\lambda_1 + \lambda_3)\lambda_{13}}$ 

 $\lambda_{14} > -2\sqrt{\lambda_8\lambda_{13}}.$ 

 $\lambda_5 + \lambda_6 - 2|\lambda_7| > \sqrt{(\lambda_1 + \lambda_3)\lambda_8}$ 

 $\lambda_1 + \lambda_3 + |2\lambda_4| + \lambda_5 + 2\lambda_7 + \lambda_8 > 0$ 

 $\lambda_{10} + \lambda_{11} - 2|\lambda_{12}| > \sqrt{(\lambda_1 + \lambda_3)\lambda_{13}}$ 

#### Unitarity conditions

$$a_{1}^{\pm} = (\lambda_{1} - \lambda_{2} + \frac{\lambda_{5} + \lambda_{6}}{2}) \\ \pm \sqrt{(\lambda_{1} - \lambda_{2} + \frac{\lambda_{5} + \lambda_{6}}{2})^{2} - 4[(\lambda_{1} - \lambda_{2})(\frac{\lambda_{5} + \lambda_{6}}{2}) - \lambda_{4}^{2}]} \\ a_{2}^{\pm} = (\lambda_{1} + \lambda_{2} + 2\lambda_{3} + \lambda_{8}) \\ \pm \sqrt{(\lambda_{1} + \lambda_{2} + 2\lambda_{3} + \lambda_{8})^{2} - 4[\lambda_{8}(\lambda_{1} + \lambda_{2} + 2\lambda_{3}) - 2\lambda_{7}^{2}]} \\ a_{3}^{\pm} = (\lambda_{1} - \lambda_{2} + 2\lambda_{3} + \lambda_{8}) \\ \pm \sqrt{(\lambda_{1} - \lambda_{2} + 2\lambda_{3} + \lambda_{8})^{2} - 4[\lambda_{8}(\lambda_{1} + \lambda_{2} + 2\lambda_{3}) - \frac{\lambda_{6}^{2}}{2}]} \\ a_{4}^{\pm} = (\lambda_{1} + \lambda_{2} + \frac{\lambda_{5}}{2} + \lambda_{7}) \\ \pm \sqrt{(\lambda_{1} + \lambda_{2} + \frac{\lambda_{5}}{2} + \lambda_{7})^{2} - 4[(\lambda_{1} - \lambda_{2})(\frac{\lambda_{5}}{2} + \lambda_{7}) - \lambda_{4}^{2}]}$$
(6.132)

$$a_{5}^{\pm} = (5\lambda_{1} - \lambda_{2} + 2\lambda_{3} + 3\lambda_{8})$$
  

$$\pm \sqrt{(5\lambda_{1} - \lambda_{2} + 2\lambda_{3} + 3\lambda_{8})^{2} - 4[3\lambda_{8}(5\lambda_{1} - \lambda_{2} + 2\lambda_{3}) - \frac{1}{2}(2\lambda_{5} + \lambda_{6})^{2}]}$$
  

$$a_{6}^{\pm} = (\lambda_{1} + \lambda_{2} + 4\lambda_{3} + \frac{\lambda_{5}}{2} + \lambda_{6} + 3\lambda_{7}) \pm ((\lambda_{1} + \lambda_{2} + 4\lambda_{3} + \frac{\lambda_{5}}{2} + \lambda_{6} + 3\lambda_{7})^{2} - 4[(\lambda_{1} + \lambda_{2} + 4\lambda_{3})(\frac{\lambda_{5}}{2} + \lambda_{6} + 3\lambda_{7}) - 9\lambda_{4}^{2}])^{1/2}$$

These are for 3H 4H ones similar

$$b_{1} = \lambda_{5} + 2\lambda_{6} - \lambda_{7}$$

$$b_{2} = \lambda_{5} - 2\lambda_{7}$$

$$b_{3} = 2(\lambda_{1} - 5\lambda_{1} - 2\lambda_{3})$$

$$b_{4} = 2(\lambda_{1} - \lambda_{1} - 2\lambda_{3})$$

$$b_{5} = 2(\lambda_{1} + \lambda_{1} - 2\lambda_{3})$$

$$b_{6} = \lambda_{5} - \lambda_{6}.$$
(6.133)

### Masses

- After electroweak symmetry breaking (Higgs mechanism) we are left with
   I3 massive particles!
- One has to be the SM Higgs boson two candidates
- Another three can be DM particles
- Check SM Higgs boson first

The masses are found by diagonalizing the matrix:

$$(\mathcal{M}_H^2)_{ij} = \frac{1}{2} \frac{\partial^2 V}{\partial H_i \partial H_j} \mid_{min}$$
.

Which is block diagonal, and all the submatrices have the following form:

$$m_{H^S}^2 = \begin{pmatrix} m_{h_s^n h_s^n} & m_{h_1^n h_s^n} & m_{h_2^n h_s^n} & 0 \\ m_{h_s^n h_1^n} & m_{h_1^n h_1^n} & m_{h_2^n h_1^n} & 0 \\ m_{h_s^n h_2^n} & m_{h_1^n h_2^n} & m_{h_2^n h_2^n} & 0 \\ 0 & 0 & 0 & 0 & m_{h_a^n h_a^n} \end{pmatrix}$$
The fields corresponding to Ha are decoupled.

Neutral scalar Higgses mass range Hs as the SM Higgs.



#### Charged Higgses mass range



Reparametrizing the vevs as

 $v_0 = v \cos \theta$  $v_1 = v \sin \theta \cos \phi$  $v_2 = v \sin \theta \sin \phi$ 

 $v_2 = \frac{1}{2}v\sin\theta$  and  $v_3 = v\cos\theta$ 

From the minimisation conditions  $tan^2\phi=\tfrac{1}{3}$ 





Neutral scalar Higgses mass range, scalar H2 as the SM Higgs.



Less points, no natural decoupling limit, but most studied solution so far





Charged scalar Higgses mass range



### Viable dark matter candidates

$$m_{h_a^n}^2 = \mu_2^2 + \lambda_{14}v_0^2 + 4(\lambda_{10} + \lambda_{11} - 2\lambda_{12})v_2^2$$
  
$$m_{h_a^n}^2 = \mu_2^2 + \lambda_{14}v_0^2 + 4(\lambda_{10} + \lambda_{11} + 2\lambda_{12})v_2^2.$$

### Theoretical restrictions

- The potential must have a finite vacuum
- The dispersion matrix must be unitary.
- Quartic couplings must be perturbative

Experimental restrictions

 The mass of the SM Higgs

 $m_{h_s^n} = 125.09 \pm .21 \text{GeV}.$ 

 The measured relic density

 $\Omega_{
m nbm} h^2 = 0.1186 \pm 0.0020$ 

# Allowed Feynman diagrams for the DM candidate



To find the relic density we need to calculate all the annihilation cross sections of the model

As a first test, we only used the circled Feynman Diagram.

This allows us to compare with the more studied 2HDM, and get a first estimate of relic density

# Relic Density scan, using MICROMEGAS.



The black line represents the measured relic density, approximately 0.118. All points in this graph are below this value.

- Simplified model to check viability
- Adding all contributions: still not saturated
- Very promising, since it is below the measured  $\boldsymbol{\Omega}$

G. Belanger, F. Boudjema, and A. Pukhov. micrOMEGAs : a code for the calculation of Dark Matter properties in generic models of particle interaction. In *The Dark Secrets of the Terascale*, pages 739–790, 2013.

# Preliminary

- Calculation done in approximation close to SM
- $\Omega$  still not saturated
- Allows for more than one DM candidate: other scalar, right-handed neutrino...
- Allows for Sommerfeld
   enhancement
- Need to consider coannihilation and decay of heavier scalars into DM candidate



### What else we need to worry about...

- Adding all contributions all Feynman diagrams works still well
- Check both candidates for DM
- Possibility: one of them decaying into the other
   → more contributions
- Vacuum might not be the deepest → check for metastable solutions, i.e. vacuum tunnelling (Vevacious)

 $\rightarrow$  will constrain more the allowed parameter space

### Leptogenesis in S3-3H

• The model lends itself naturally to leptogenesis and associated baryogenesis:

Heavy right handed neutrinos Majorana neutrinos Decay of right-handed neutrinos into the left-handed ones

$$\epsilon_1 = \frac{\sum_{\alpha} \Gamma(N_1 \to \ell_{\alpha} H) - \Gamma(N_1 \to \overline{\ell}_{\alpha} \overline{H})}{\sum_{\alpha} \Gamma(N_1 \to \ell_{\alpha} H) + \Gamma(N_1 \to \overline{\ell}_{\alpha} \overline{H})}.$$



$$\epsilon \simeq -\frac{3}{8\pi} \frac{1}{(h_{\nu} h_{\nu}^{\dagger})} \sum_{i=2,3} \operatorname{Im}\{(h_{\nu} h_{\nu}^{\dagger})_{1i}^{2}\} [f(\frac{M_{i}^{2}}{M_{1}^{2}}) + g(\frac{M_{i}^{2}}{M_{1}^{2}})].$$

$$f(x) = \sqrt{x} [1 - (1 + x) \ln(\frac{1 + x}{x})] \qquad \qquad q(x) = \frac{\sqrt{x}}{1 - x}$$

# Asymmetry depends on the right- and left-handed neutrino masses, as well as the phases

Mi, right handed neutrinos hv mass matrices for Yuk couplings of Dirac neutrinos



Table 1: Z2 assignment in the leptonic sector.

$$M_e \simeq m_{\tau} \begin{pmatrix} \frac{1}{\sqrt{2}} \frac{\tilde{m_{\mu}}}{\sqrt{1+x^2}} & \frac{1}{\sqrt{2}} \frac{\tilde{m_{\mu}}}{\sqrt{1+x^2}} & \frac{1}{\sqrt{2}} \frac{\sqrt{1+x^2-\tilde{m_{\mu}}^2}}{\sqrt{1+x^2}} \\ \frac{1}{\sqrt{2}} \frac{\tilde{m_{\mu}}}{\sqrt{1+x^2}} & -\frac{1}{\sqrt{2}} \frac{\tilde{m_{\mu}}}{\sqrt{1+x^2}} & \frac{1}{\sqrt{2}} \frac{\sqrt{1+x^2-\tilde{m_{\mu}}^2}}{\sqrt{1+x^2-\tilde{m_{\mu}}^2}} \\ \frac{\tilde{m_e}(1+x^2)}{\sqrt{1+x^2-\tilde{m_{\mu}}^2}} e^{i\delta_e} & \frac{\tilde{m_e}(1+x^2)}{\sqrt{1+x^2-\tilde{m_{\mu}}^2}} e^{i\delta_e} & 0 \end{pmatrix}.$$

$$\tilde{m}_i = m_i / m_\tau, \ x^4 = (m_e / m_\tau)^4$$

 Mass matrices of charged leptons and neutrinos are parameterised in terms of physical masses

Similarly for the Majorana neutrino mass matrix —  $\mu$  function of the masses, Majorana phases  $\phi$  and Dirac phase  $\delta$ 

$$M_{\nu} = M_{\nu D} \tilde{M}^{-1} (M_{\nu D})^{T} = \begin{pmatrix} (\frac{1}{M_{1}} + \frac{1}{M_{2}})\mu_{2}^{2} & (\frac{1}{M_{1}} - \frac{1}{M_{2}})\mu_{2}^{2} & (\frac{1}{M_{1}} + \frac{1}{M_{2}})\mu_{2}\mu_{4} \\ (\frac{1}{M_{1}} - \frac{1}{M_{2}})\mu_{2}^{2} & (\frac{1}{M_{1}} + \frac{1}{M_{2}})\mu_{2}^{2} & (\frac{1}{M_{1}} - \frac{1}{M_{2}})\mu_{2}\mu_{4} \\ (\frac{1}{M_{1}} + \frac{1}{M_{2}})\mu_{2}\mu_{4} & (\frac{1}{M_{1}} - \frac{1}{M_{2}})\mu_{2}\mu_{4} & \frac{\mu_{4}^{2}}{M_{2}} + \frac{\mu_{3}^{2}}{M_{3}} \end{pmatrix}$$

$$(46)$$

#### Preliminary update of arXiv:1701.07929

#### Solving the Boltzmann equations

$$\epsilon = \frac{\operatorname{Im}[e^{2i\delta^*} M_2 m_3 \frac{\sqrt{M_2(m_2 - m_3)(m_3 - m_1)}}{\sqrt{m_3}}](f[\frac{M_3^2}{M_1^2}] + g[\frac{M_3^2}{M_1^2}])}{8\pi |M_2 m_3|} \cdot$$

$$Y_B = aY_{B-L} = \frac{a}{a-1}Y_L \quad a = \frac{(8N_f + 4N_H)}{(22N_f + 13N_H)},$$

$$\delta$$
 = Dirac phase, Y<sub>L,B</sub> = baryon, lepton asymmetry, N<sub>f,H</sub> = number of families and Higgs doublets

Asymmetry for the degenerate case  $M_1=M_2$ Only inverted hierarchy for light neutrinos possible

For the non-degenerate case it depends also on the phases  $\mu_{ia}$ ,  $\mu_i = |\mu_i| e^{i\mu_{ia}}$ 



Inverted hierarchy m1  $\approx$  .055 eV m2  $\approx$  .056 eV m3  $\approx$  .022 eV

Red - too much baryon asymmetry Darker shades of green - more asymmetry but not yet saturated

Solution to baryon asymmetry can have R-neutino masses as low as  $\sim 10^7$  GeV





Asymmetry as function of M<sub>1</sub> and M<sub>3</sub>.

### How can you go up?

Propose a model with a similar symmetry, but at high energies

SUSY GUT SU(5)×Q6

- Q6 is the double covering of S3, has similar structure
- Neutrino masses through extra singlets (right handed), U(1) or nonrenormalizable interactions

- Now we start from GUT scale and go down and we recover nice features of previous model (not identical)
- CKM and PMNS ok
- Neutrino mixing angles also within experimental bounds



## Conclusions

- Multi-Higgs models with S3 symmetry predictive, provide viable DM candidates
- Models have to be consistent from the SM point of view → extra Higgses sufficiently decoupled or inert
- Vacuum much more complicated than in SM, all checks necessary: stability, perturbative unitarity
- 4H-3S (based on 3H-3S) model promising:
  - explains well masses and mixings in quark and lepton sectors
  - can provide with two promising DM candidates
  - enough leptogenesis to explain baryogengesis (3H-3S) A.Alvarez & M.M
  - Possible to have more DM candidates in the R-neutrino sector, not yet explored