



Scintillators - Applications

Paul Lecoq CERN, Geneva

December 2017

ICFA School – Cuba – Nov 27th-Dec 8th, 2017



Outlook



• Lecture 1: Scintillator fundamentals

- Organic scintillators
- Inorganic scintillators
 - Scintillation mechanisms
 - Limits to the light yield and decay time
 - Energy resolution and non-proportionality

• Lecture 2: Scintillator applications

- Crystal growth techniques
- High energy physics and dark matter searches
- Space borne missions
- Medical applications
- Geophysical exploration
- Homeland security

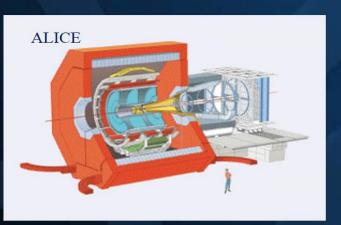
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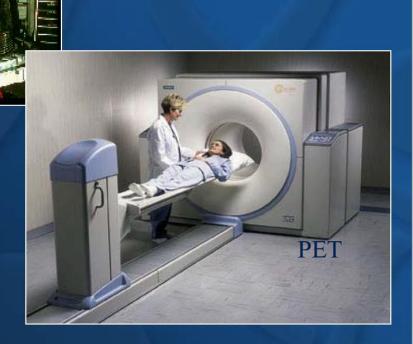


What drives sthe devlopment of new scintillators









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HEP and PET use the largest volume of scintillators



High energy physics (e.g. CMS) 80,000 crystals; 12,000 liters; highest production rate in 2005 4100 liters/yr (34 tons/yr)



Very low cost; high density; fast decay; radiation hard

Positron Emission Tomography in 2003, 450 sc/yr x 10 liters/sc = 4500 liters/yr (33 tons/yr)



High density and atomic number, fast decay, high light yield, low cost

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Various appelications of scintillators



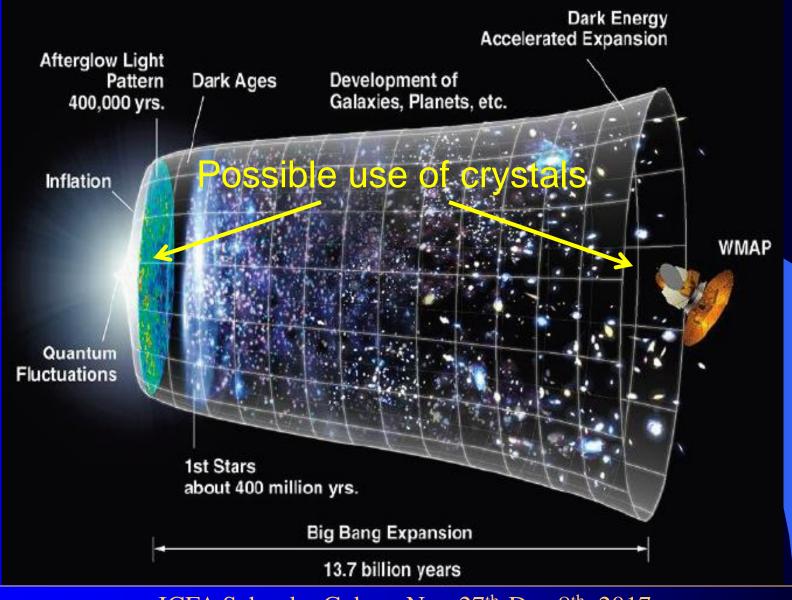
Detection and spectroscopy of energetic photons (and neutrons) in:

- High energy particle physics
- Nuclear physics
- Positron emission tomography
- X-ray computed tomography
- Security monitoring
- Treaty verification
- Geophysical exploration
- Non-destructive testing
- Radiation monitoring



Back to Creation





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Limits of scintillator properties



Scintillation eff. ¹ :	$\eta = \beta SQ$ where b is the number of e-h pairs S is the transfer efficiency to luminescent center Q is the quantum efficiency of the luminescent centers	
	$τ_r = 1.5 \times 10^{-5} \times \lambda^2 / ((f/9)(n^2 + 2)^2 \times n)$ where I is the emission wavelength f is the oscillator strength n is the refractive index	
Density:	limited by elemental composition and crystal structure, not many candidates > 10 g/cm ³	
Atomic number:	limited only by the periodic table	
Cost:	raw material abundance purification cost crystal growth or other manufacturing process	

¹ Lempicki, Nucl. Instr. Meth.
² Henderson and Imbush, Optical Spectroscopy of Inorganic Solids (1989)

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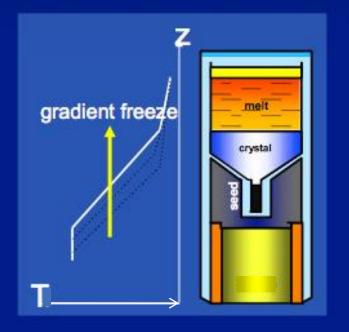




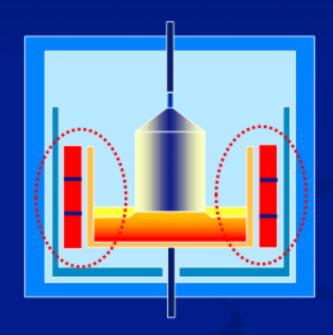
Most frequently used crystal growth technologies



Bridgman-techniques



Czochralski-techniques



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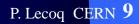
From principles to practice





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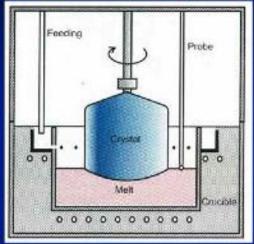
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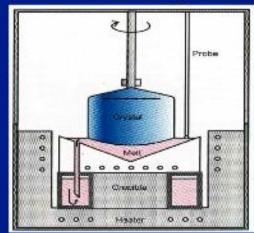




Powder feeding system



Melt feeding



Advantages:

- 1. Continuos growth process (large size of crystal)
- 2. Fixed "crystal-melt" interface
- 3. Feeding by raw material and activator
- 4. Good melt convection
- 5. Simple melt level control
- 6. Crystal and crucible rotation

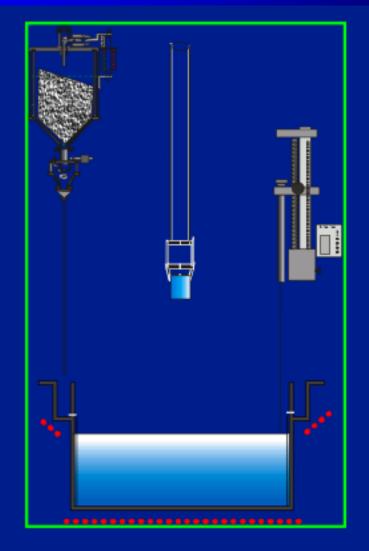
Disadvantages:

- 1. Large melt volume
- 2. Large melt surface (activator evaporation)
- 3. Purification ?
- 4. Powder feeding

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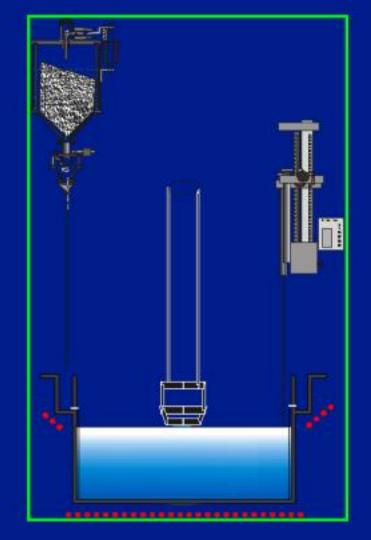










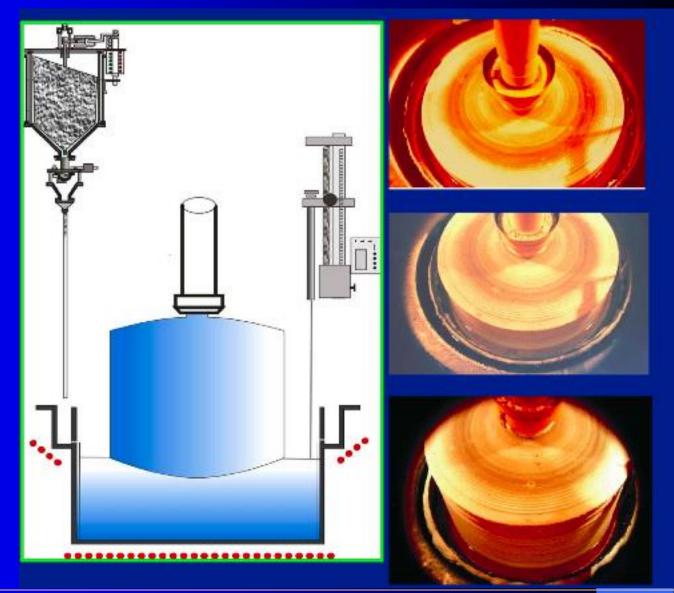


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Preparation of large size plates for SPECT cameras





Nal:TI single crystal \emptyset 520 mm, mass > 550 kg





• After growth the NaI:TI crystal is cut and pressed at high temperature to prepare large area plates for SPECT cameras

Institute of Single Crystals Kharkhov, Ukraine



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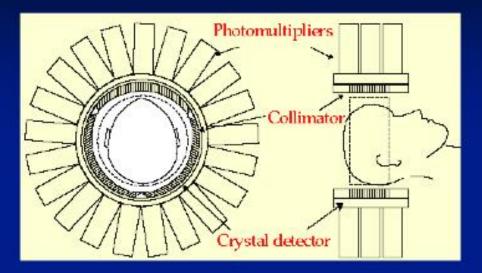
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Nal cylindrical detectors



Nal(TI) detector







Nal(TI) crystal

DSI, Boston USA

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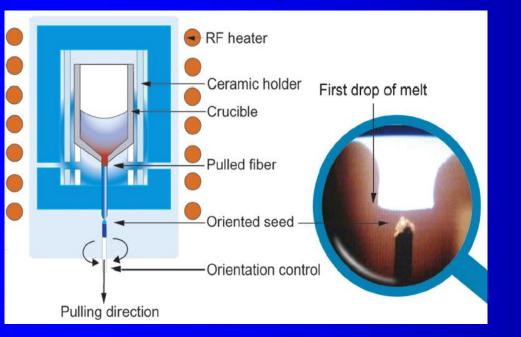
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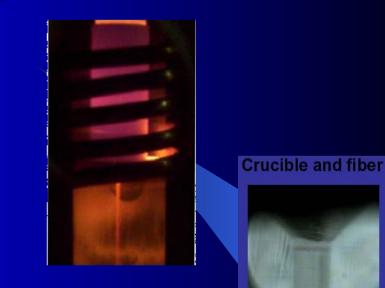
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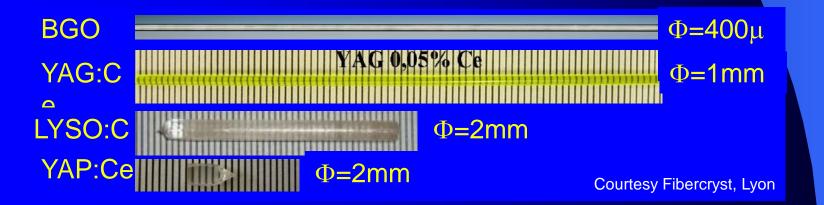


Micro-pulling-down crystal fiber growth









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Ceramics



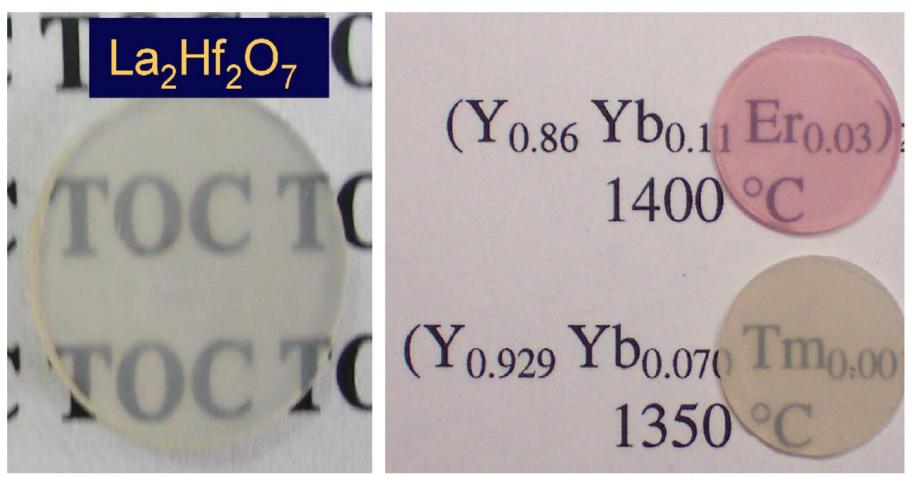


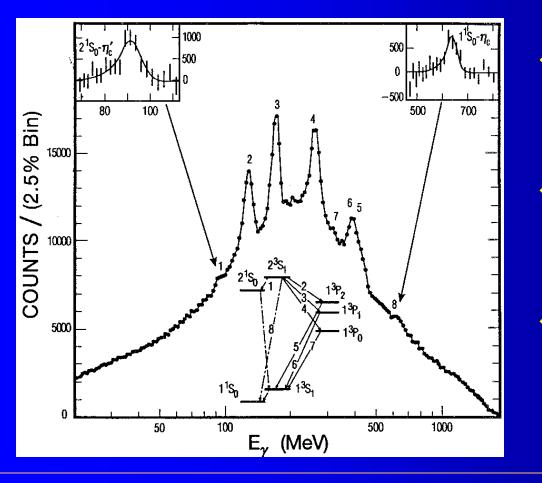
Fig. 17: Transparent ceramics of different heavy scintillators prepared with pre-reacted nanopowders

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Charmonium spectroscopy – Crystal Ball - SLAC



♦ 50cm diameter spherical ball of NaI(Tl) crystals

672 crystals 42cm long, **PMT** readout

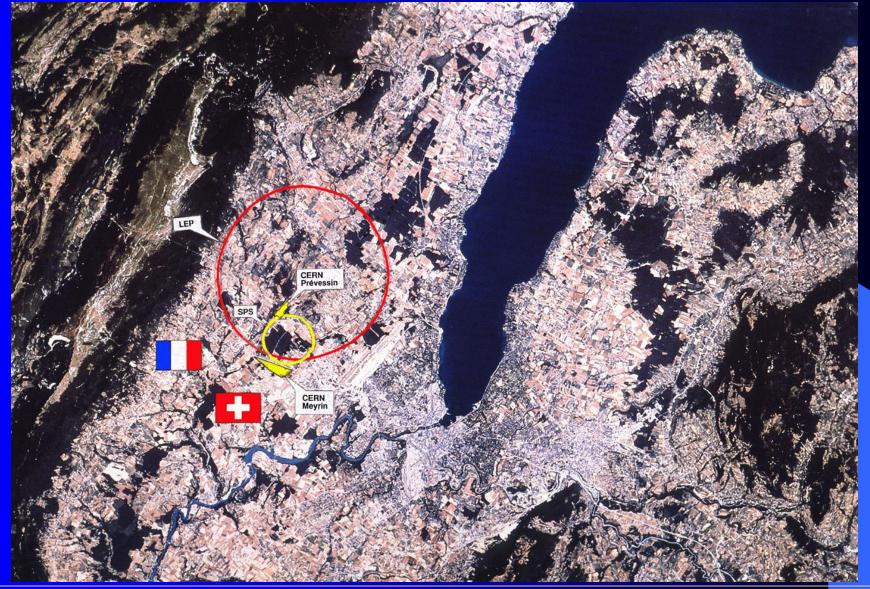
Very good resolution allowed precise spectroscopic study of charmonium states

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CERN in Geneva area





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The Large Hadron Collider LHC



Mt Blanc

Lake of Geneva

of the second second second second second

LHCb

CMS

Large Hadron Collider 27 km circumference



ATLAS

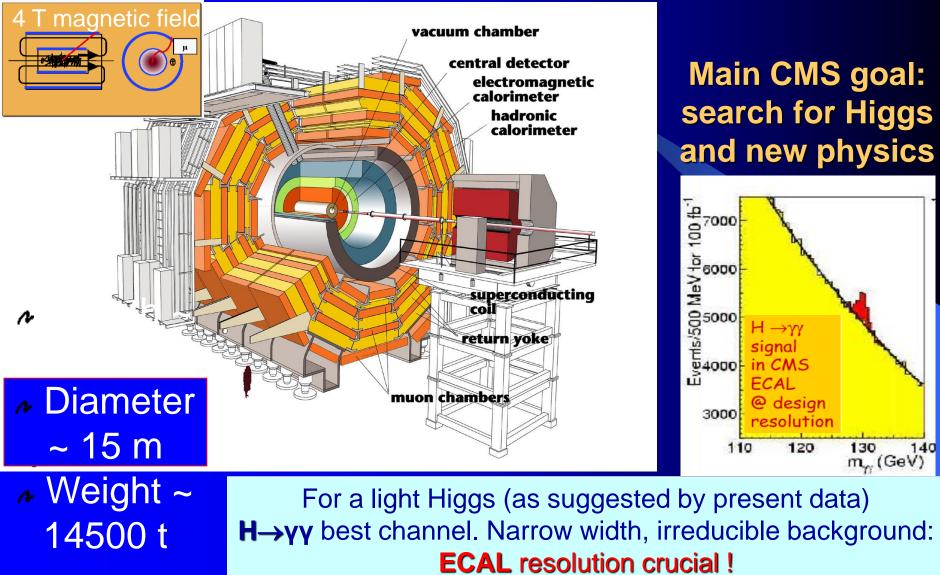
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Compact Muon Solenoid



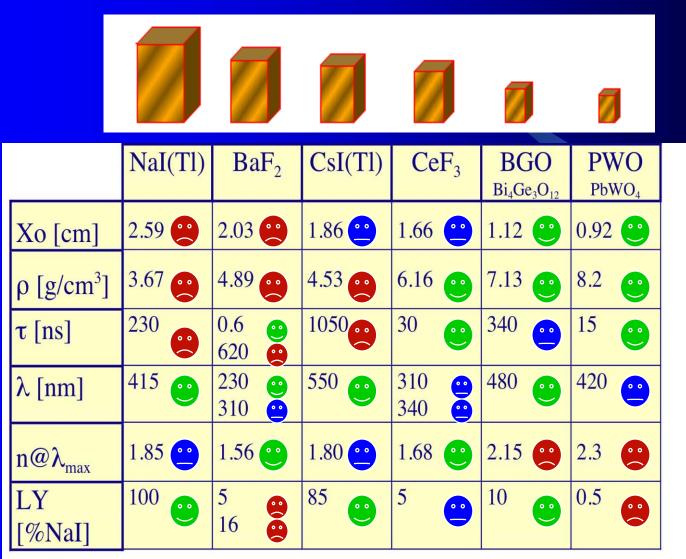


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Some popular crystals in HEP





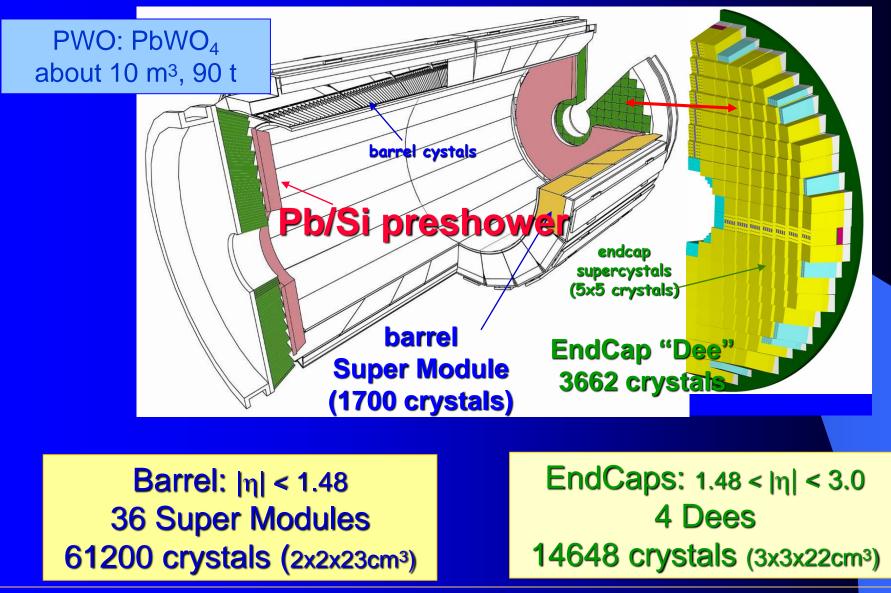
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ECAL layout





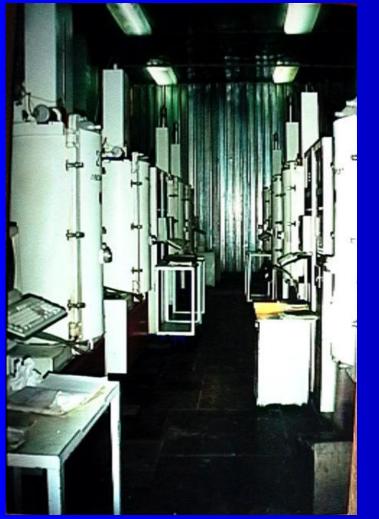
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Production facilities in Bogoroditsk





159 Lazurit 3M growth ovens



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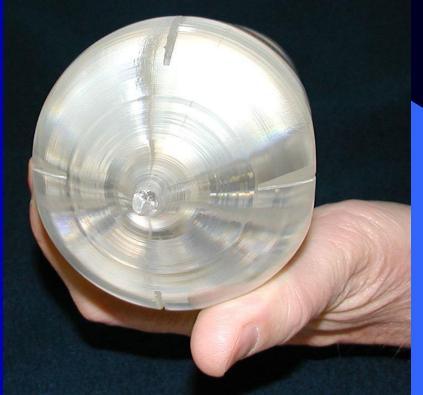
PWO growth of large ingots



• Replace 120mm Pt crucible by new 170mmcomposite Pt crucibles



- Upgrade the Rf power cycle for heating, smelting, pulling and cooling
- Optimization of procedure for crystal ingot extraction from crystallization unit



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Square end cut 85mm

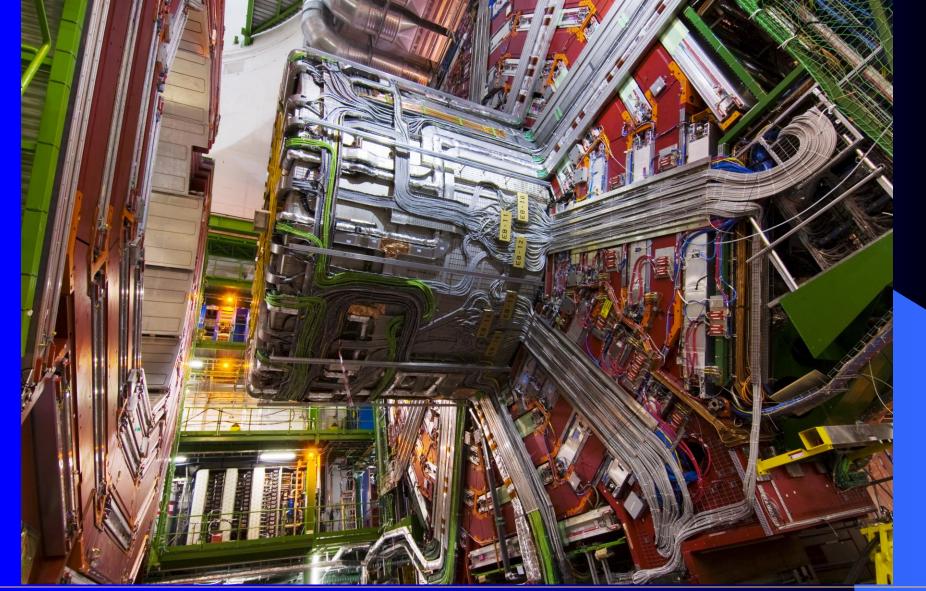






CMS Barrel





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CMS Endcaps





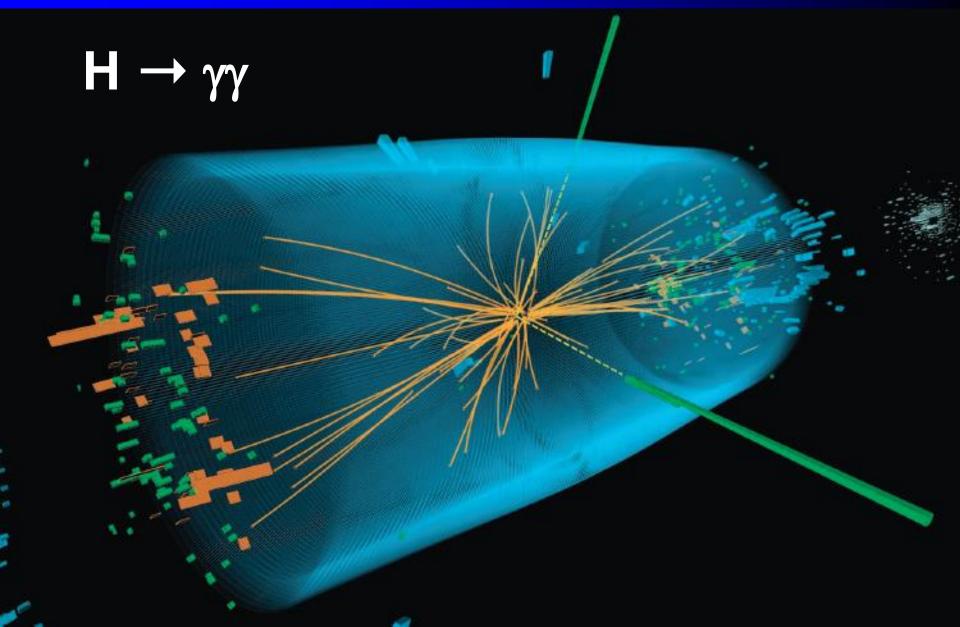
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And finally!!





A different detector concept for precise jet calorimetry

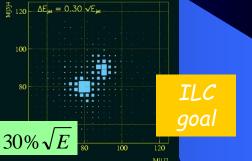
 A conventional approach is limited to about 60% - 70%/ √E

New approaches have been proposed

- Particule Flow for a 3D imaging calorimeter
 - Each particle in a jet measured individually
 - Requires very high granularity
 - Limitations
 - Complex Integration due to huge number of channels
 - Identification of individual particles in a jet challenging at very high energy
- Dual readout method
 - Measure event by event the electromagnetic fraction of the hadronic shower by separating Cerenkov and scintillation light



 $\frac{120}{100} \Delta E_{\mu} = 0.60 \forall E_{\mu}$



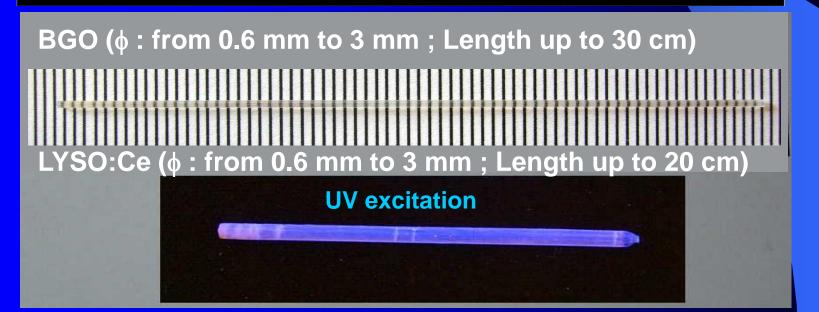




Some crystal fibers







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Dual readout based on Metamaterials



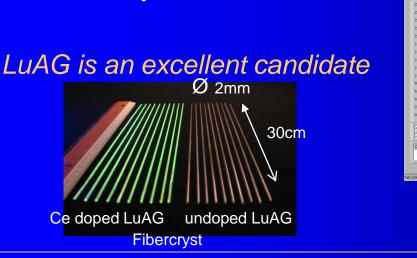
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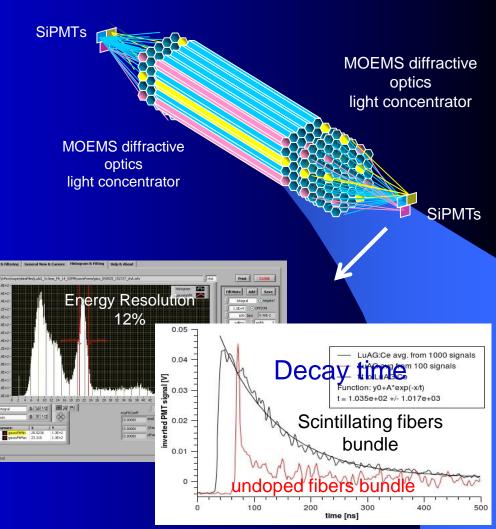
• Select a non-intrinsic scintillating material (unlike BGO or PWO) with high bandgap for low UV absorption

• The undoped host will behave as an efficient Cerenkov: heavy material, high refraction index n, high UV transmission

• Cerium or Praesodinum doped host will act as an efficient and fast scintillator

 \approx 60ns decay for Ce \approx 20ns decay for Pr





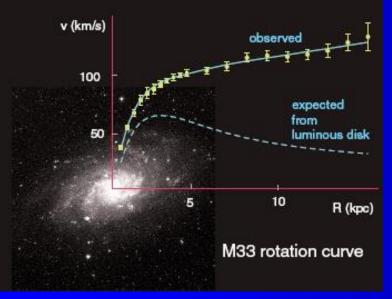
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WIMP-Dark matter searches



Evidence of Dark Matter Rotation curve of spiral galaxies



L.Bergstrom Rep.Progr.Phys.63 (2000) 793

V=(GM/r)^{1/2} $\Omega_{\rm M}$ ~0.3 $\Omega_{\rm X}$ ~0.7 Direct detection - elastic scattering off nuclei

Features:

 Low energy → Low energy threshold (~10 keV)

2. Expected event rate < 0.1/kg day

Detector mass 1kg-100 kg

 High radiopurity of detector and construction materials

Underground location

Background rejection

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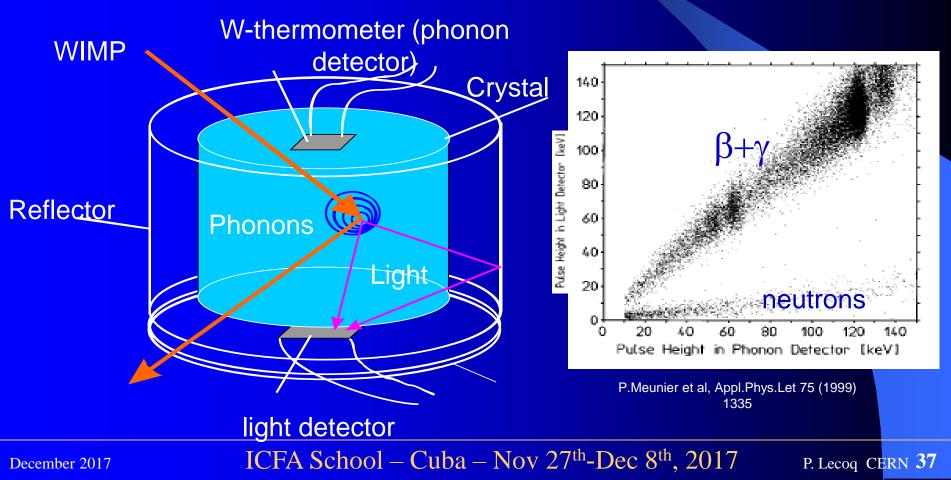


Why cryogenic scintillator? CREST detector



Sensitivity + Selectivity

high-energy resolution of cryogenic phonon detector + discrimination of events with low detection threshold ($\leq 10 \text{ keV}$)





Scintillator requirements



- high light yield at low temperatures
 - \checkmark low threshold,
 - \checkmark good energy resolution
- Radiopurity (ex Lu, Rb, K, U, Th)
- Suitable thermodynamics characteristics
- Possible candidates
 - \checkmark CaWO₄ satisfactory choice, currently in use, large ongoing effort to improve the material
 - \checkmark ZnWO₄ scintillator under development for cryogenic application
 - \checkmark CaMoO₄ and CdMoO₄ material under investigation

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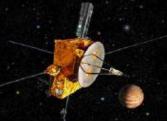
Scintillation Detectors on Space Missions













Past missions		GRS	NS	PS
-	Phobos	Csl	Stilben,plastic	yes
=	Lunar Prospector	BGO/BC454	BC454	
<u></u>	Near	Nal/BGO		
	Mars Observer	HPGe/BC454	BC454	
-	Mars Odyssey	HPGe	BC454,Stilben,plastic	Csl
~			19 1	

Current missions

÷	Ulysses	Csl/GRB		Plastic
-	Messenger	HPGe/BGO	GS20,BC454	

Missions in implementation

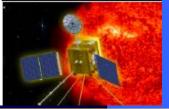
	Dawn	CZT/BGO	BGO,BC454, G20
\odot	Phobos Grunt	LaBr ₃	BC454,Stilben,plastic
Ξ	Solar Orbiter	LaBr ₃	plastic
-	BepiColombo	LaBr ₃	

GRS=gamma-ray spectrometer NS=neutron spectrometer PS=particle spectrometer





Csl



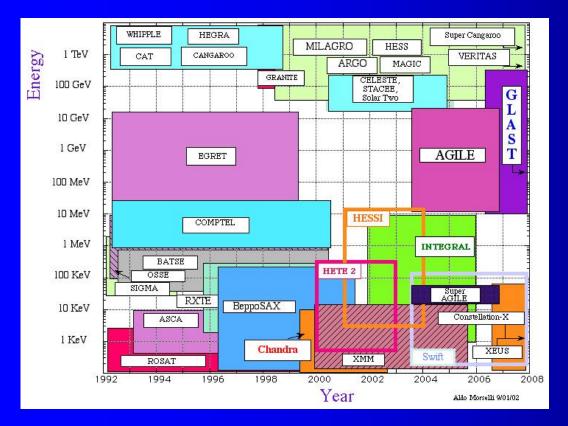
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Important Projects for Gamma Astrophysics





Parameter	EGRET	GLAST (Minimum Spec.)		
Energy Range	20 MeV - 30 GeV	20 MeV – 300 GeV		
Peak Effective Area ¹	1500 cm ²	>8000 cm ²		
Field of View	0.5 sr	>2 sr		
Angular. Resolution ²	5.8° (100 MeV)	<3.5° (100 MeV) <0.15° (>10 GeV)		
Energy Resolution ¹	10%	<10%		
Deadtime per Event	100 ms	<100 μs		
Source Location Determination ⁴	15'	<0.5'		
Point Source Sensitivity⁵	-I x 10 ⁻⁷ cm ⁻² s ⁻¹	<6 x 10 ⁻⁹ cm ⁻² s ⁻¹		

EGRET – Energetic Gamma Ray Experiment Telescope GLAST – Gamma Ray Large Area Space Telescope

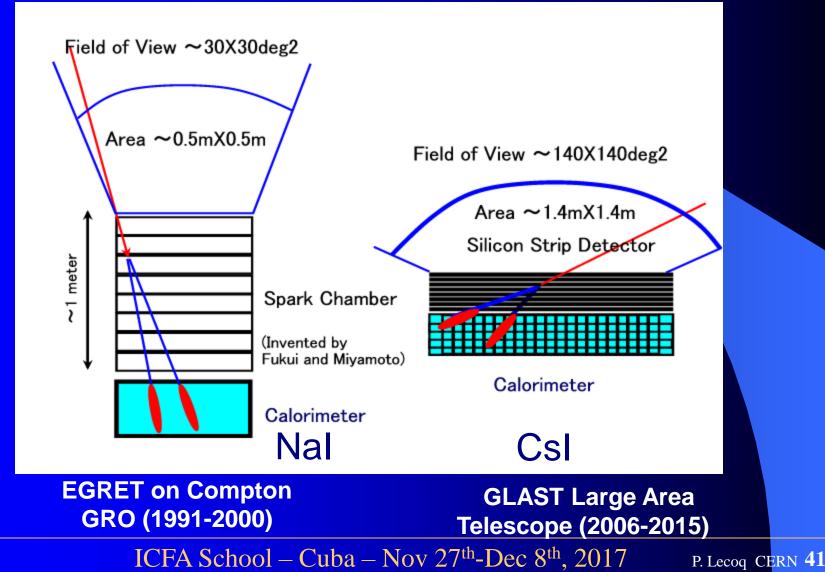
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EGRET versus GLAST





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The BepiColombo Mercury mission





Goal: The gamma and neutron mapping of Mercury surface **Science objectives:**

- ★ The mapping of water content in Mercury subsurface
- * The mapping of Mercury soil composition

Parameters:

PARAMETER	VALUE			
Mass	5.2 kg			
Power	5 W			
Volume				
Surface Resolution	400 km			
Minimal time resolution	2-4 sec			
Energy range, neutrons	Multi energy bands covering 10 ⁻³ eV – 15 κeV			
Energy range, gamma	300 keV – 10 MeV			
Energy resolution, gamma	3% at 660 keV			
Detectors	3He – proportional counters, stilben crystal, LaBr ₃ crystal			
Temperature range	(-20C, 40C)			
Position	ESA: BepiColombo			
Altitude	400 km – 1500 km			

Science Payload and Advanced Concepts

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History- LaBr₃ development program





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Scintillators in nuclear medicine

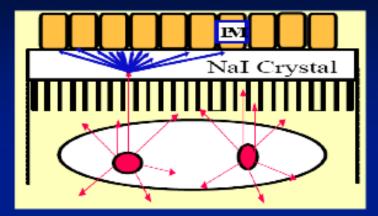


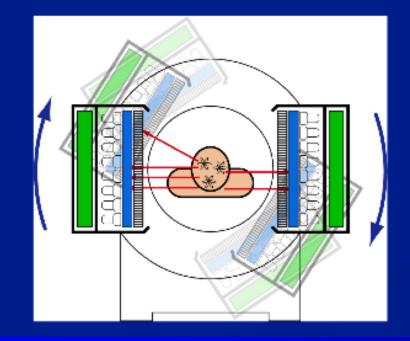
- High detection efficiency for different type of radiations, particularly γ-rays
- Good energy resolution
- Very good timing
- High counting rate capability
- Great variety in size and composition



SPECT principle Spatial resolution





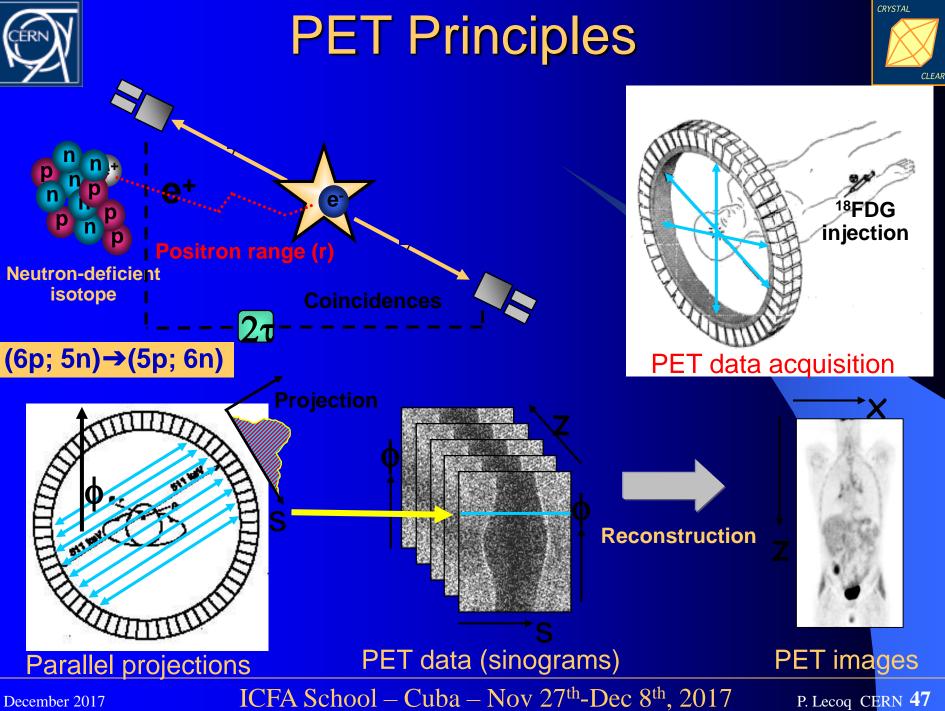


• Collimator – Ability of the collimator to localize the γ-ray source in the patient (~6-12 mm)

 Intrinsic – Ability of the Nal(*TI*) crystal and PMT to localize the γ-ray interactions in the crystal (~3-4 mm)

• Extrinsic – Overall system resolution combining collimator and intrinsic factors. Quadratic sum of FWHM of intrinsic and collimator resolution.

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Scintillators for PET

	1962	1977	1995	1999	2001	2003	<u>200</u> 7
	NaI	BGO	GSO:Ce	LSO:Ce	HHAP:Ce	LaBr ₃ :Ce	LuAG:Ce
Density (g/cm³)	3.67	7.13	6.71	7.40	8.34	5.29	6.73
Atomic number	51	75	59	66	65	47	63
Photofraction	0.17	0.35	0,25	0.32	0.30	0.13	0.30
Decay time (ns)	230	300	30-60	45	17	18	60
Light output (hv/MeV)	4300 0	8200	12500	27000	11400	70000	>25000
Peak emission (nm)	115	480	430	420	365	356	535
Refraction index	415 <u>1.85</u>	2.15	1.85	1.82	1.97	1.88	1.84

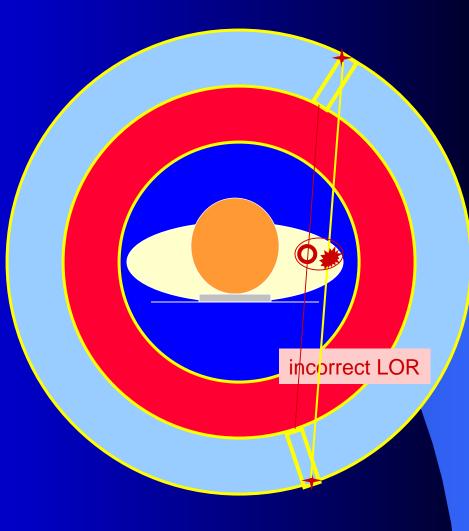


Parallax error in PET



Spatial Resolution -Depth-of-interaction DOI

Depth-of-interaction (DOI) information is needed to maintain good resolution at off-centre positions



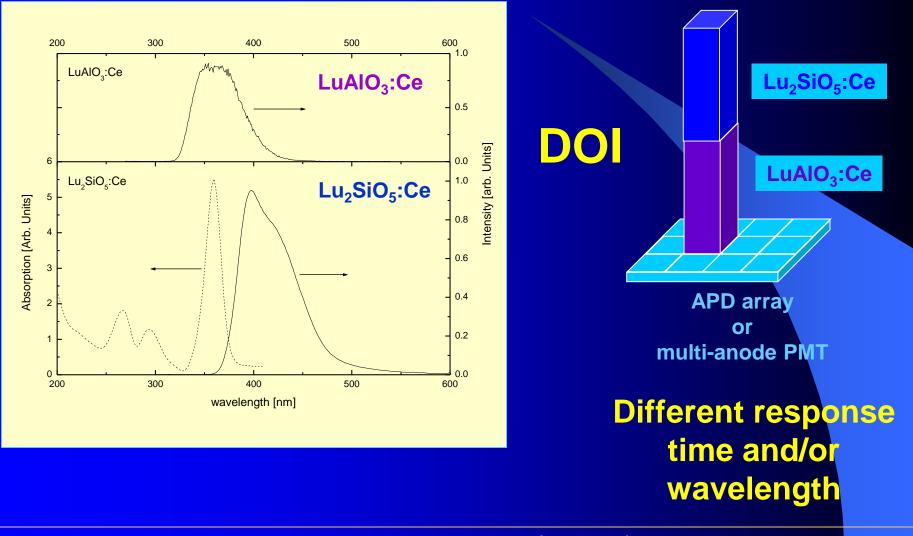
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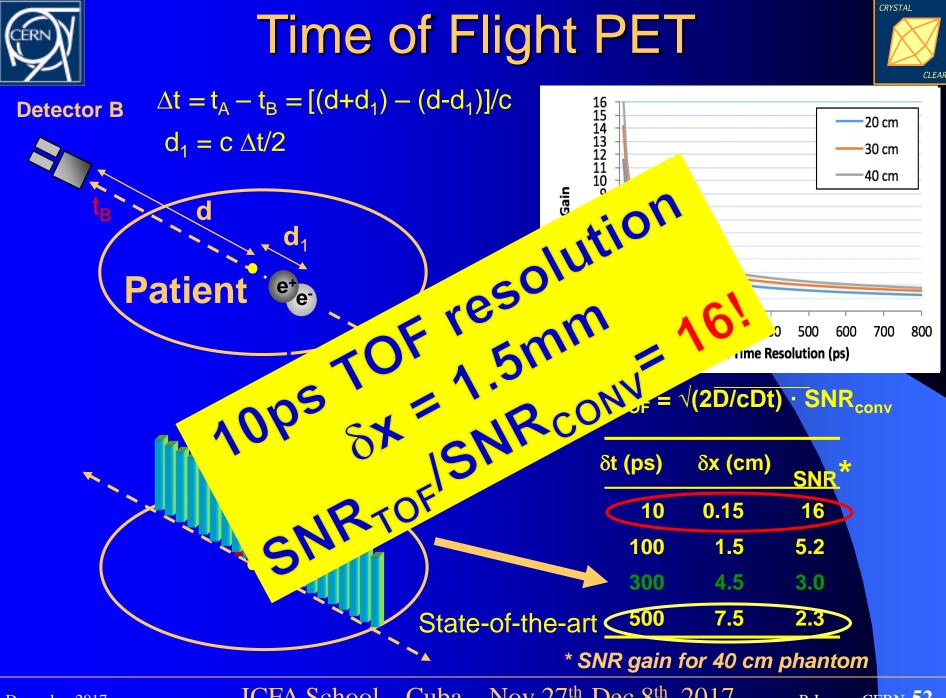


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Scintillators for Geophysical Exploration





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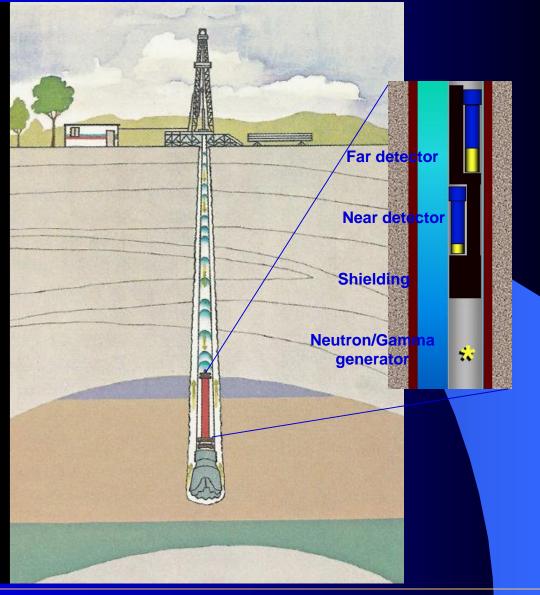
Oil Well Logging



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Measurement Issues

- Source and sensor both in borehole
- Usually want to measure Formation
- Need to make measurement with 1-3 seconds of data

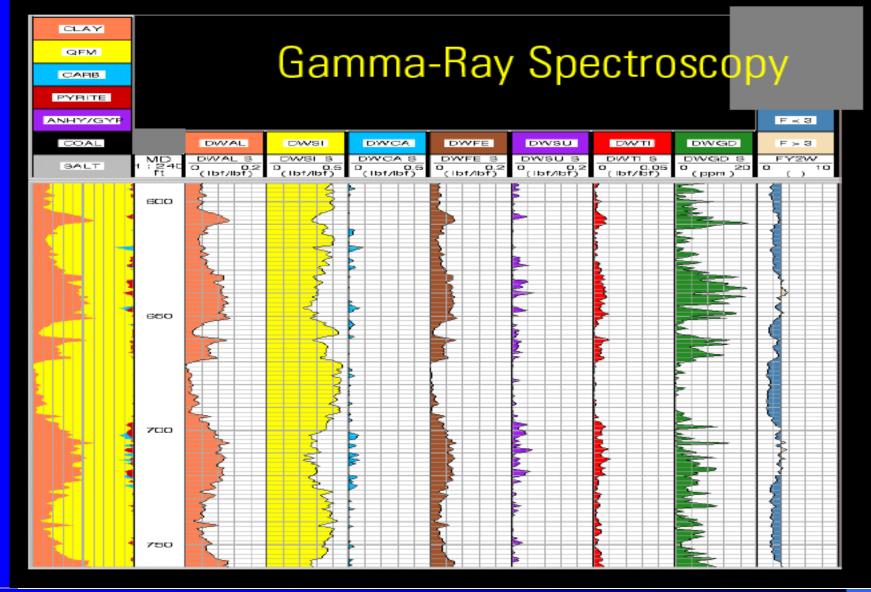


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Nuclear Tools





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Desirable Scintillator Properties



- High Density, High Z
- Good Resolution (all Energies)
- High Countrate Capability
- Reasonable Size
- Non-Hydroscopic
- Light Output (350 450 nm)
- Rugged
- Good High Temperature Performance
 - Resolution
 - Wavelength
 - Decay Time

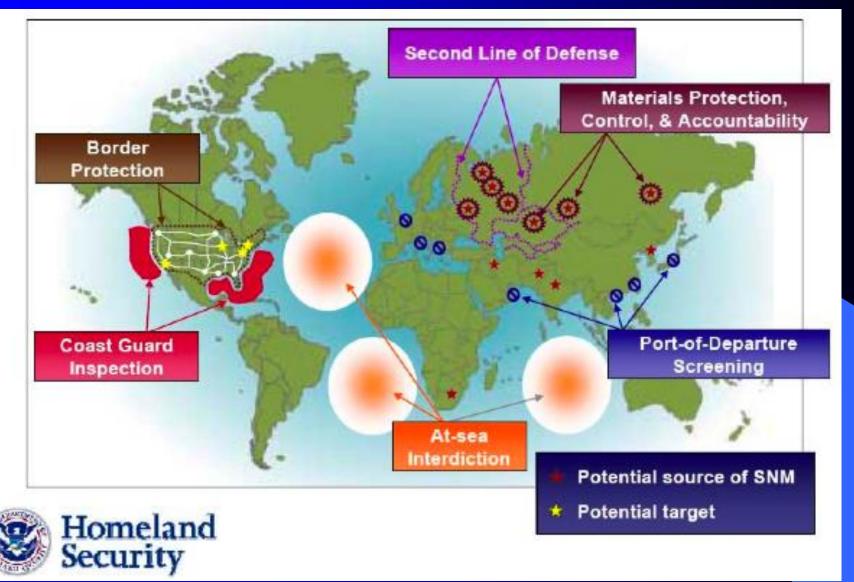
NaIBGOGSO

• LuAP ?



Homeland security





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Passive instrumentation













Hand-held, mobile, transportable, and fixed position Nal, LaBr₃, PVT

Spectrometers and counters

7 Managed by UT-Battelle for the Department of Energy



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Scientifiator Applications in Biomeland Seconds



Active Instrumentation











Mobile and fixed position; X ray, ⁶⁰Co, ¹³⁷Cs Nal, CdWO₄, BGO

Spectrometers, counters, imagers

8 Managed by UT-Battelle for the Department of Energy



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Scinificator Applications in Newsland Security



Border control





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Conclusion 1



- Scintillators are widely used in a large number of scientific and industrial domains
- The ideal scintillator does not exist and research should continue on new materials and new production technologies (micro-pulling down)
- Structuring at the nanoscale gives access to quantum confinment phenomena and increases the range of optical electric and magnetic properties that can be engineered
- Metamaterials based on quantum dots, photonic crystals and photonic crystal fibers can open the way to new detection paradigms with huge design flexibility



Conclusion 2



- A number of good scintillators have been discovered that have proved difficult to manufacture – e.g. LuPO4:Ce
- An alternative approach is to start with materials of known manufacturability, and improve the scintillation properties - e.g. PWO, YAG, GGG
- Total Set "One of the continuing scandals ciences is that it remains in ger structure of ever knowl