

# INSTRUMENTATION & DETECTORS for HIGH ENERGY PHYSICS II

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# DETECTOR: INTRODUCTION QUIZZ

**What is a detector ?**

**What does a detector measure ?**

**(How is a detector designed ?)**

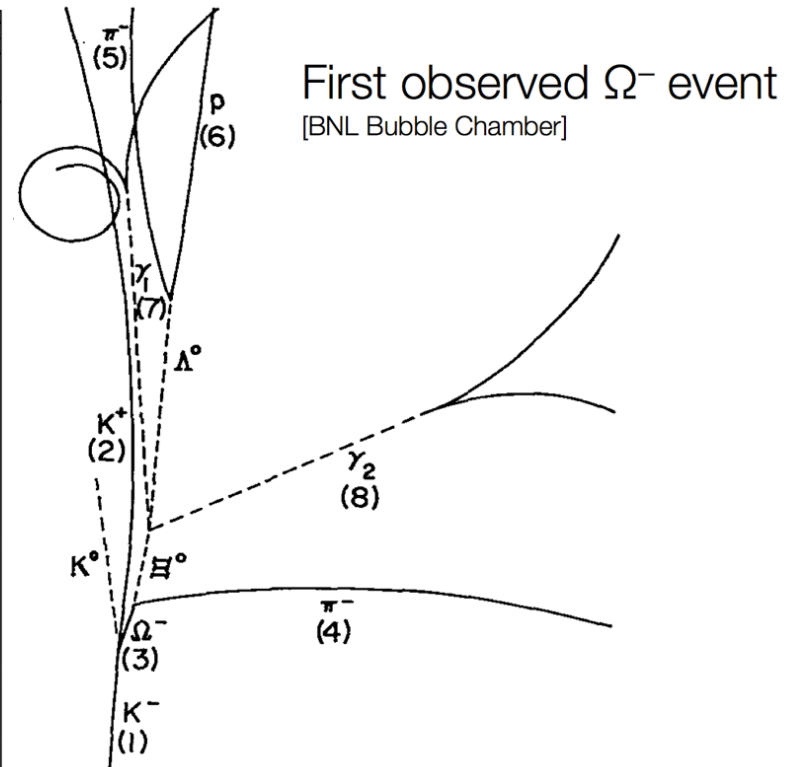
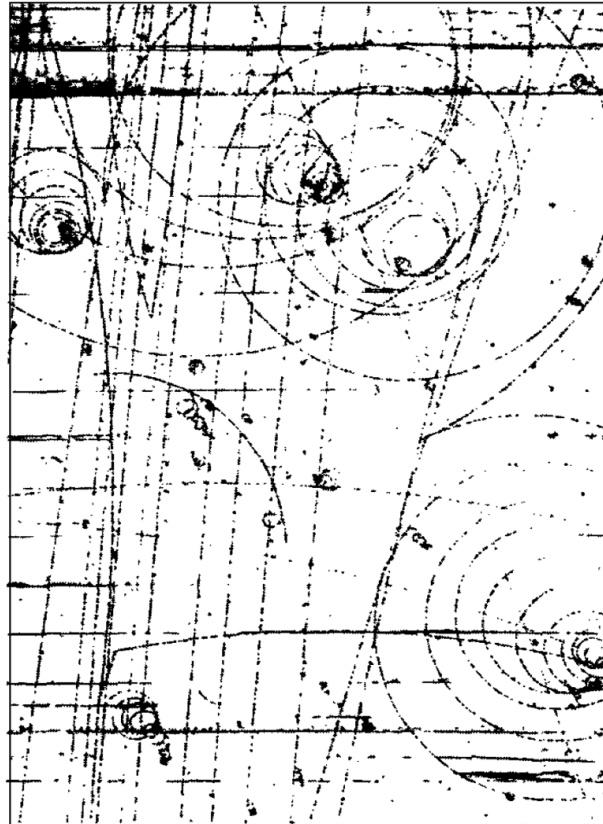
**Compare a digital camera with the ATLAS detector**

**Would you join an experiment where the calorimeter is in front of the tracking system ?**

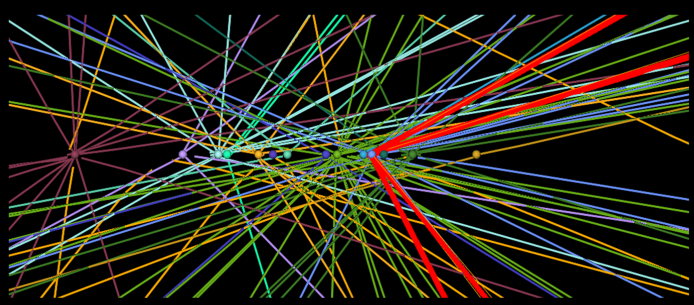
# WHAT IS A PARTICLE DETECTOR ?

An apparatus able to  
detect the passage of a particle  
and/or localise it  
and/or measure its momentum or energy  
and/or identify its nature  
and/or measure its time of arrival

.....

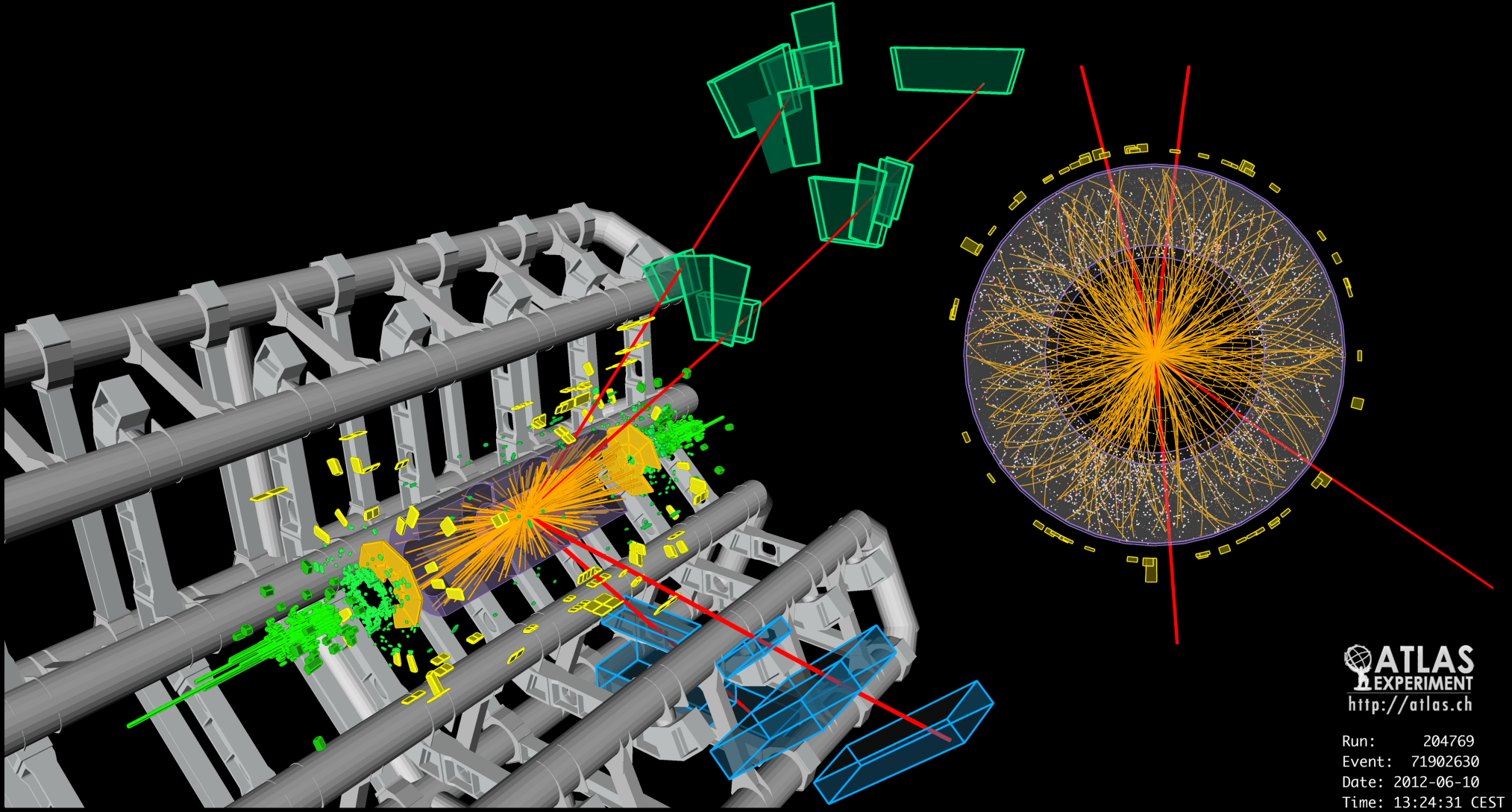


# ATLAS 4 $\mu$ event: LHC collision event




## Higgs Boson Discovery 2012

Higgs to 4 $\mu$  candidate event



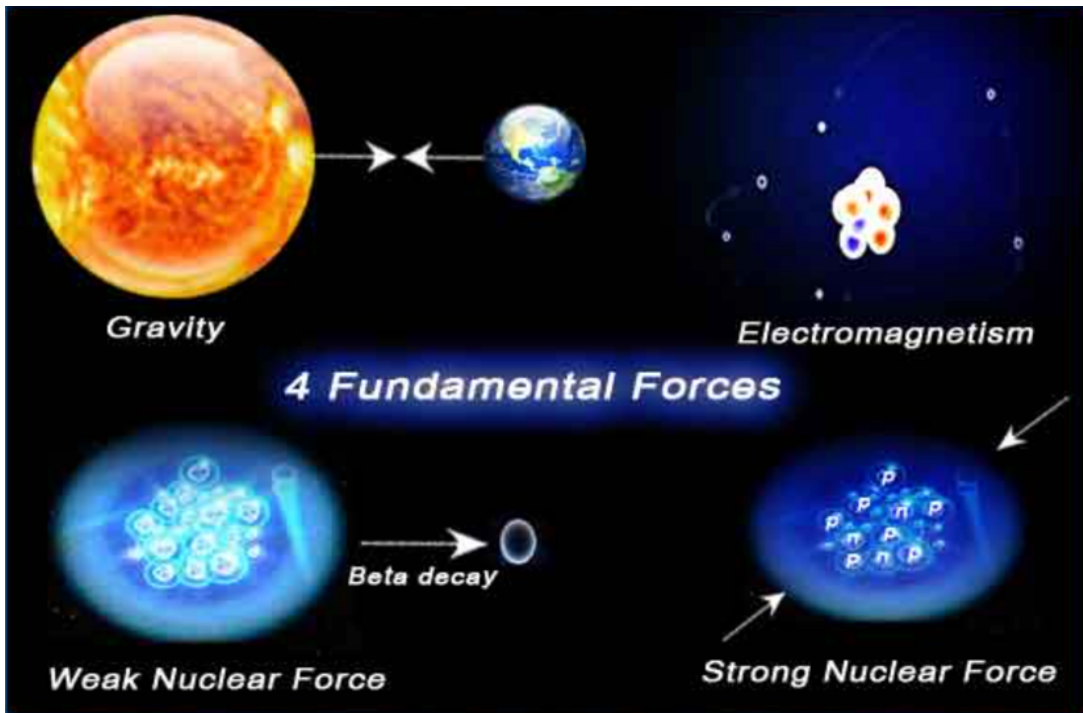
**ATLAS**  
EXPERIMENT  
<http://atlas.ch>

Run: 204769  
Event: 71902630  
Date: 2012-06-10  
Time: 13:24:31 CEST

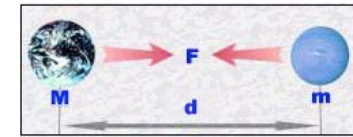


# TODAY INTERACTIONS

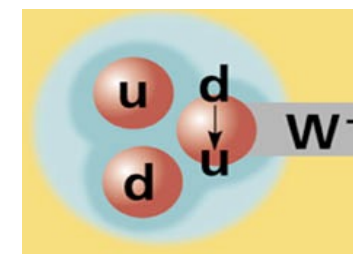
# INTERACTIONS



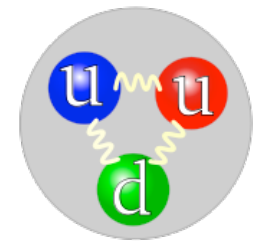
Gravity  
Graviton ?



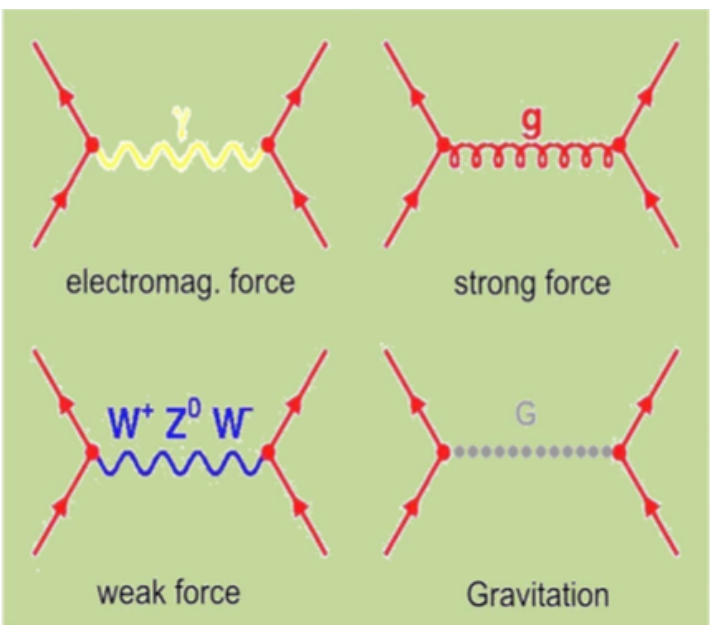
Electromagnetism  
Photon



Weak interaction  
W & Z bosons



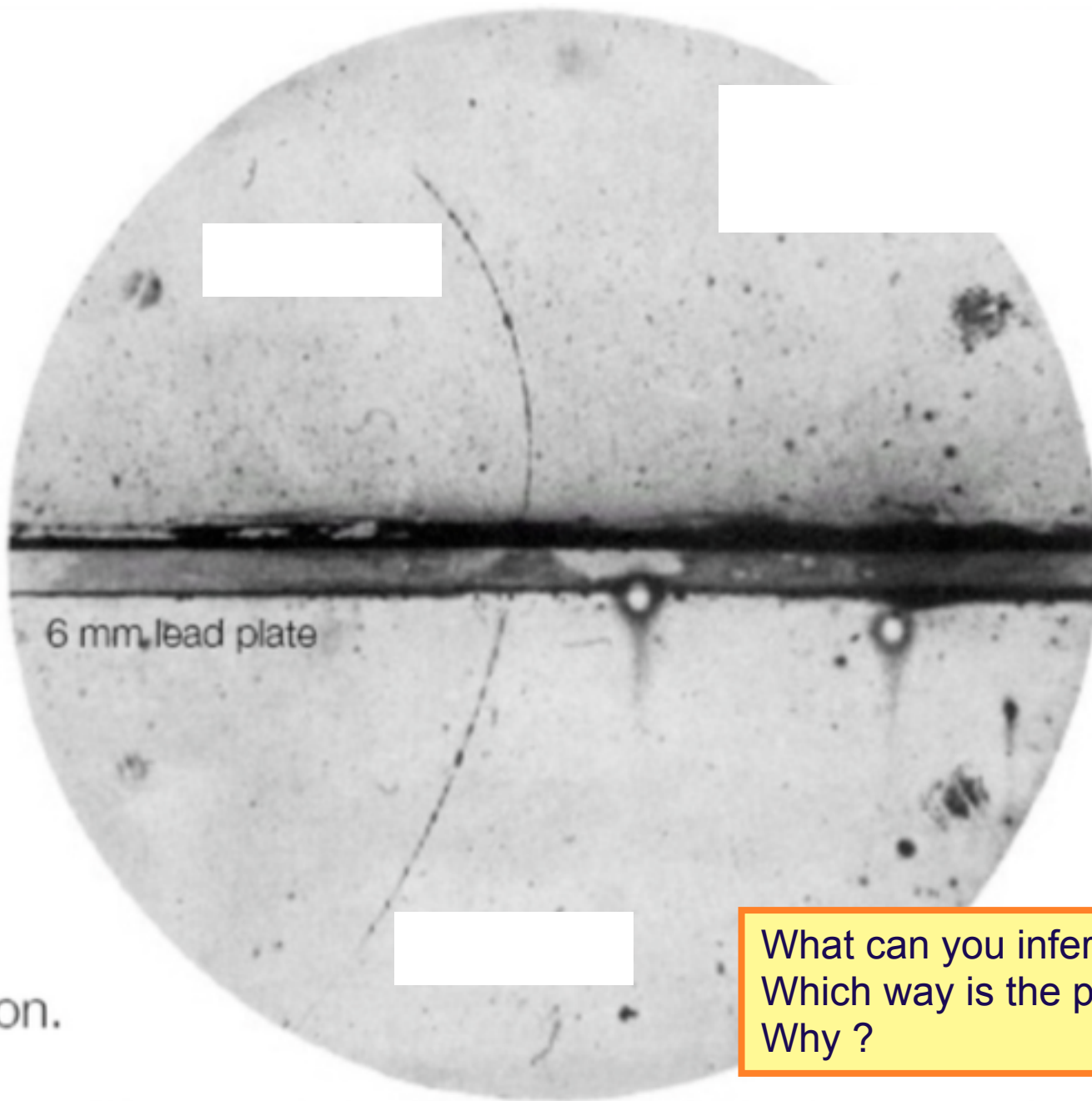
Strong interaction  
Gluons



In the Standard Model (SM) of particle physics, the electromagnetic and the weak forces are unified: electroweak interaction.



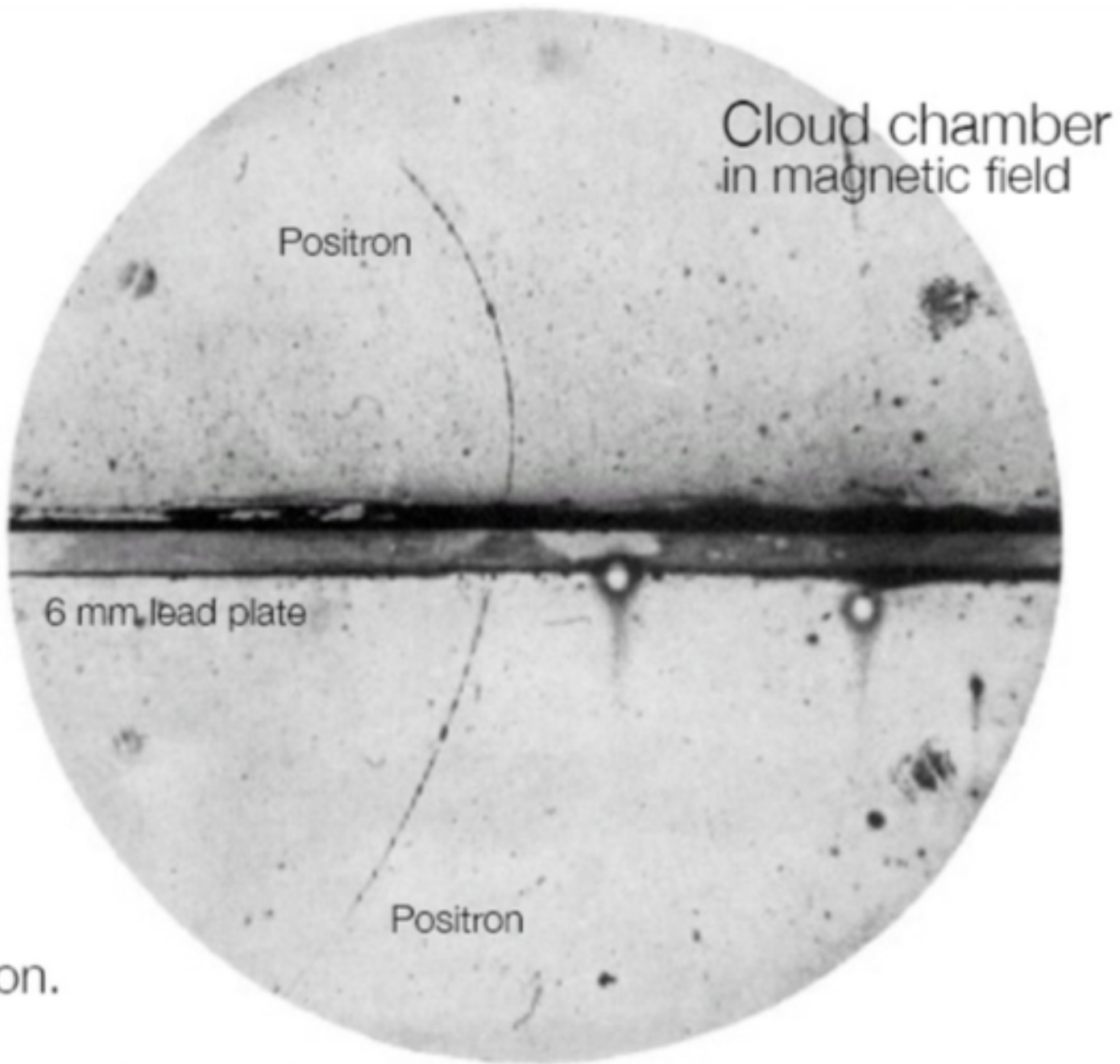
# HOW to DETECT and IDENTIFY a PARTICLE?



What can you infer from this picture about the setup ?  
Which way is the particle traversing the photograph ?  
Why ?

# POSITRON DISCOVERY in 1933

Positron discovery in 1933  
by Carl Andersen



ron.



# HOW ARE PARTICLES DETECTED ?

In order to detect a particle it must  
interact with the material of the detector  
transfer energy in some recognisable way and leave a *signal*.

Detection of particles happens via their energy loss in the material they traverse.

Charged particles

Photons

Hadrons

Neutrinos

Ionization, Bremsstrahlung, Cherenkov, ...

Photo/Compton effect, pair production

Nuclear interactions

Weak interactions

multiple  
interactions

single  
interactions...

multiple  
interactions

# THE 13 PARTICLES A DETECTOR MUST BE ABLE TO MEASURE AND IDENTIFY

$e^\pm$	$m_e = 0.511 \text{ MeV}$	} EM
$\mu^\pm$	$m_\mu = 105.7 \text{ MeV} \sim 200 m_e$	
$\gamma$	$m_\gamma = 0, Q = 0$	
$\pi^\pm$	$m_\pi = 139.6 \text{ MeV} \sim 270 m_e$	} EM, Strong $\sim 3.5 m_\pi$
$K^\pm$	$m_K = 493.7 \text{ MeV} \sim 1000 m_e$	
$p^\pm$	$m_p = 938.3 \text{ MeV} \sim 2000 m_e$	
$K^0$	$m_{K^0} = 497.7 \text{ MeV} \quad Q=0$	} Strong
$n$	$m_n = 939.6 \text{ MeV} \quad Q=0$	

The Difference in *Mass, Charge, Interaction*  
is the key to the Identification

# MEASURING PARTICLES

Particles are characterized by

Mass	[Unit: eV/c <sup>2</sup> or eV]
Momentum	[Unit: eV/c or eV]
Energy	[Unit: eV]
Charge	[Unit: e]

[+ Spin, Lifetime ...]

$$\begin{aligned} \text{eV} &= 1.6 \cdot 10^{-19} \text{ J} \\ c &= 299\,792\,458 \text{ m/s} \\ e &= 1.602176487(40) \cdot 10^{-19} \text{ C} \end{aligned}$$

Relativistic kinematics:

$$E^2 = \vec{p}^2 c^2 + m^2 c^4$$

$$\beta = \frac{v}{c} \quad \gamma = \frac{1}{\sqrt{1 - \beta^2}}$$

$$E = m\gamma c^2 = mc^2 + E_{\text{kin}}$$

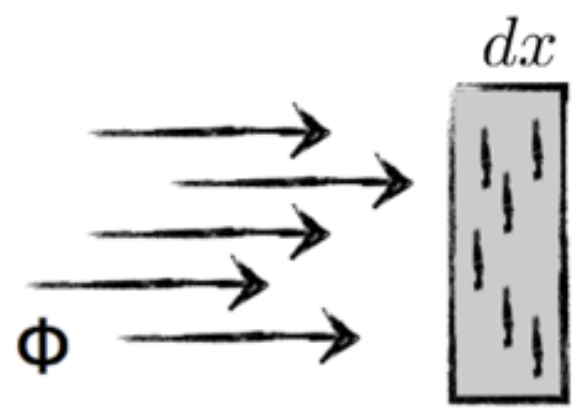
Particle Identification via  
measurement of

e.g.  $(E, \vec{p}, Q)$  or  $(\vec{p}, \beta, Q)$   
 $(\vec{p}, m, Q)$  ...

$$\vec{p} = m\gamma\vec{\beta}c \quad \vec{\beta} = \frac{\vec{p}c}{E}$$

# INTERACTION CROSS-SECTION

Flux  $\Phi = \frac{1}{S} \frac{dN_i}{dt}$   $[L^{-2} t^{-1}]$



Reactions per unit of time  $\frac{dN_{\text{reac}}}{dt} = \Phi \overbrace{\sigma N_{\text{target}}}^{\text{area obscured by target particle}} dx$   $[t^{-1}]$

$[L^{-2} t^{-1}]$   $[?]$   $[L^{-1}]$   $[L]$

Reaction rate per target particle  $W_{if} = \Phi \sigma$   $[t^{-1}]$

Cross section per target particle  $\sigma = \frac{W_{if}}{\Phi}$   $[L^2]$  = reaction rate per unit of flux

1b =  $10^{-28}$  m<sup>2</sup> (roughly the area of a nucleus with A = 100)

# CROSS-SECTION: ORDER OF MAGNITUDE

Standard cross section unit:

$$[\sigma] = \text{mb}$$

with  $1 \text{ mb} = 10^{-27} \text{ cm}^2$

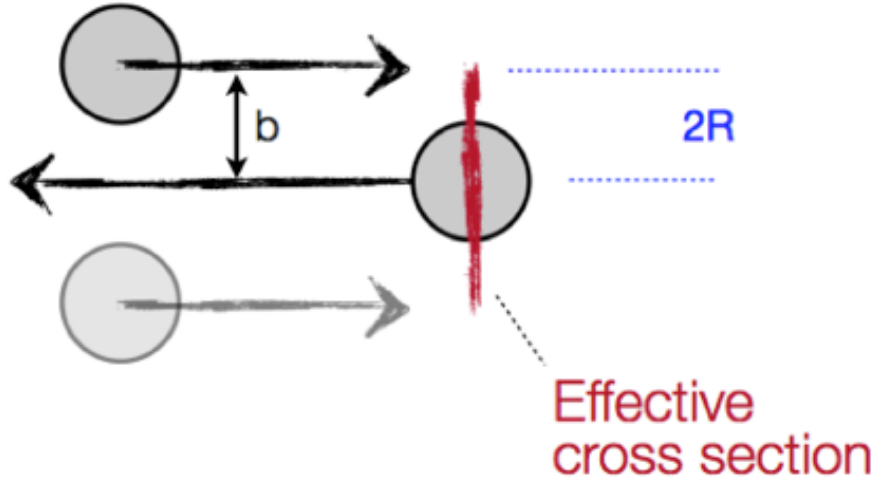
or in natural units:

$$[\sigma] = \text{GeV}^{-2}$$

with  $1 \text{ GeV}^{-2} = 0.389 \text{ mb}$   
 $1 \text{ mb} = 2.57 \text{ GeV}^{-2}$

Estimating the proton-proton cross section:

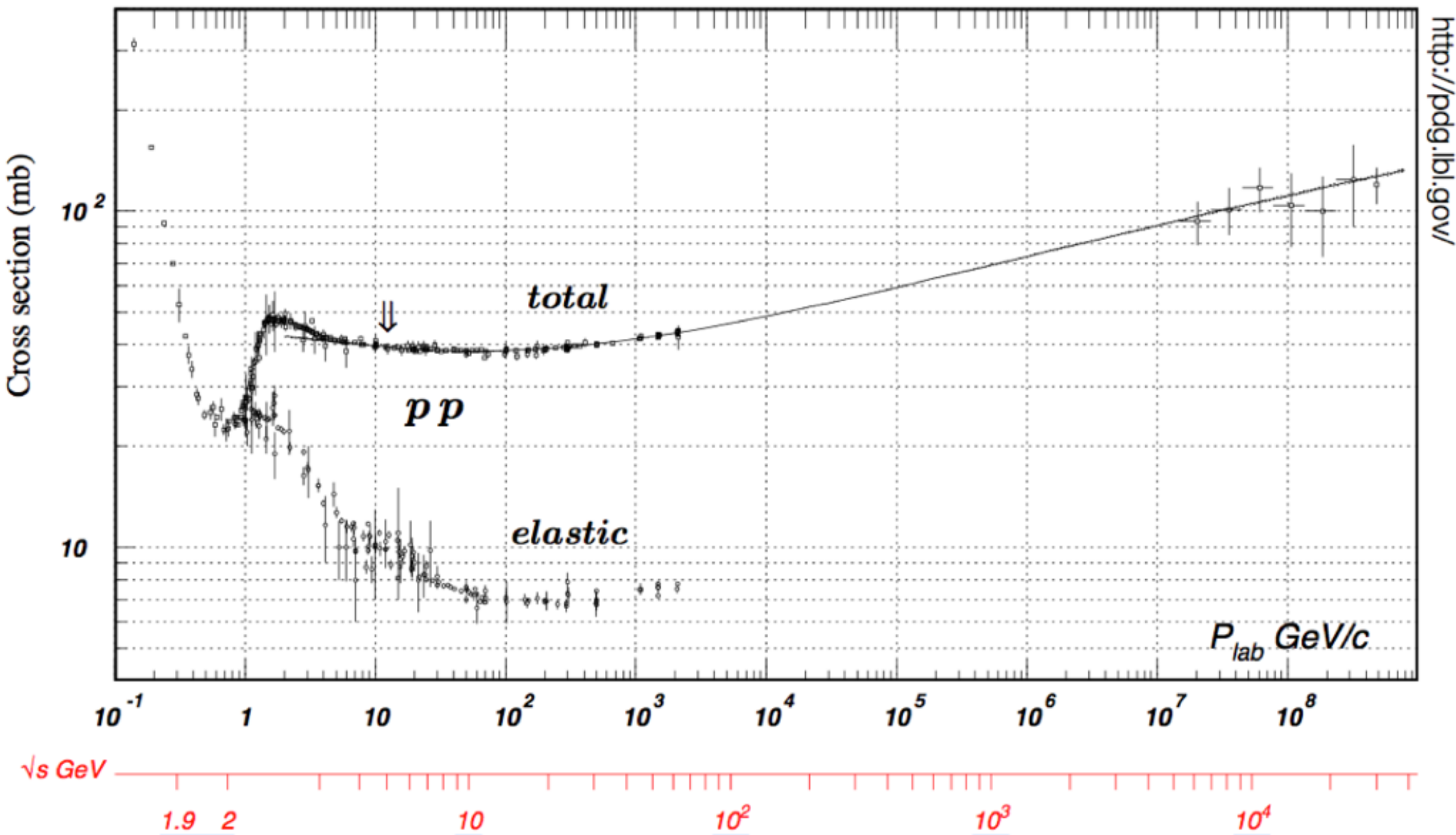
using:  $\hbar c = 0.1973 \text{ GeV fm}$   
 $(\hbar c)^2 = 0.389 \text{ GeV}^2 \text{ mb}$



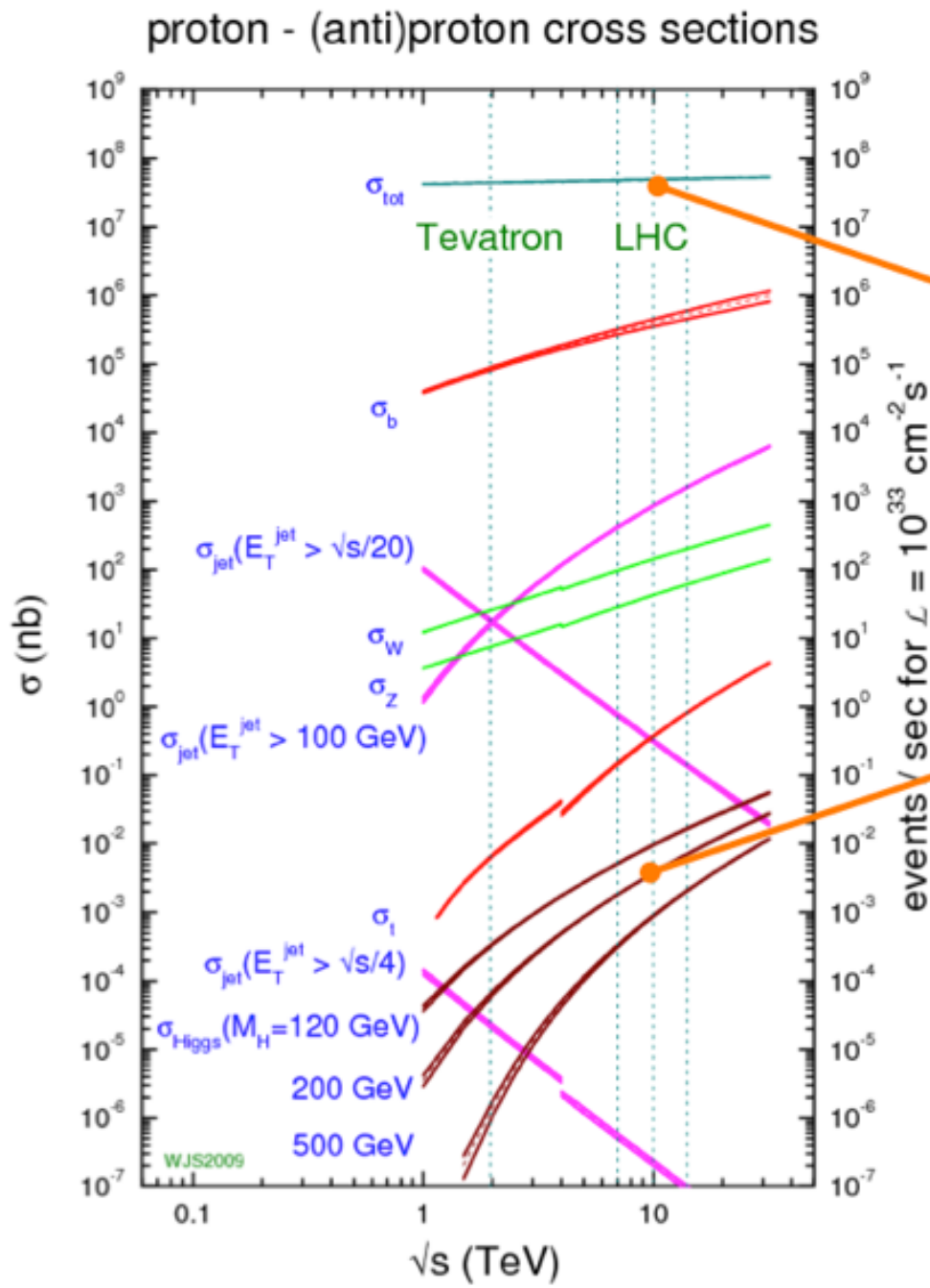
Proton radius:  $R = 0.8 \text{ fm}$   
 Strong interactions happens up to  $b = 2R$

$$\begin{aligned} \sigma &= \pi (2R)^2 = \pi \cdot 1.6^2 \text{ fm}^2 \\ &= \pi \cdot 1.6^2 \cdot 10^{-26} \text{ cm}^2 \\ &= \pi \cdot 1.6^2 \cdot 10 \text{ mb} \\ &= 80 \text{ mb} \end{aligned}$$

# PROTON-PROTON SCATTERING CROSS-SECTION



# CROSS-SECTIONS AT THE LHC



$10^8$  events/s

$\sim 10^{10}$

$10^{-2}$  events/s  $\sim$

10 events/min

$[m_H \sim 120 \text{ GeV}]$

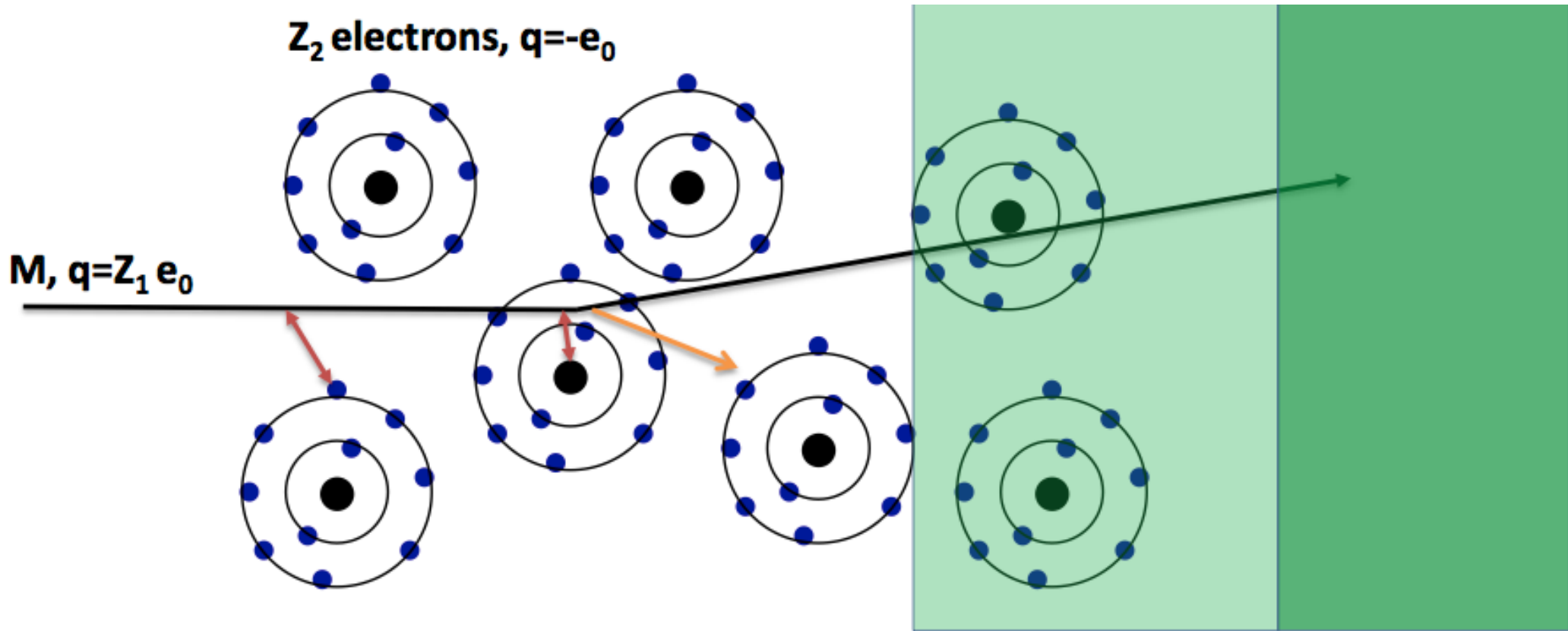
0.2%  $H \rightarrow \gamma\gamma$

1.5%  $H \rightarrow ZZ$

**TRIGGER !**

# ELECTROMAGNETIC INTERACTION

## PARTICLE - MATTER



### Interaction with the atomic electrons.

The incoming particle loses energy and the atoms are **excited or ionised**.

### Interaction with the atomic nucleus.

The incoming particle is deflected causing **multiple scattering** of the particle in the material.

During this scattering a **Bremsstrahlung photon** can be emitted

In case the particle's velocity is larger than the velocity of light in the medium, the resulting EM shockwave manifests itself as **Cherenkov radiation**. When the particle crosses the boundary between two media, there is a probability of 1% to produce an X-ray photon called **Transition radiation**.

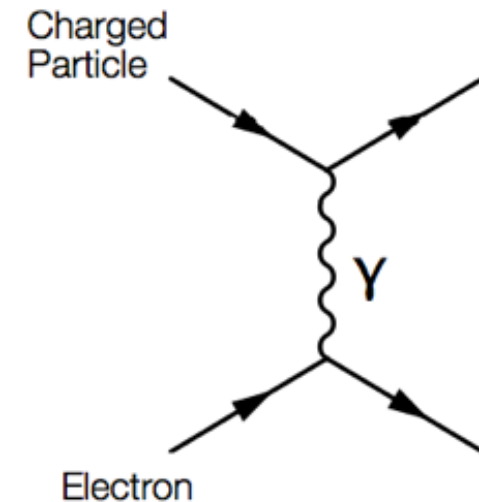


# ENERGY LOSS BY IONISATION: BETHE-BLOCH FORMULA

For now assume:  $Mc^2 \gg m_e c^2$

i.e. energy loss for heavy charged particles  
[dE/dx for electrons more difficult ...]

Interaction dominated  
by elastic collisions with electrons ...



Bethe-Bloch Formula

$$-\left\langle \frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

$$\propto 1/\beta^2 \cdot \ln(\text{const} \cdot \beta^2 \gamma^2)$$

# BETHE-BLOCH FORMULA

[see e.g. PDG 2010]

$$-\left\langle \frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

[· ρ]  
density

$K = 4\pi N_A r_e^2 m_e c^2 = 0.307 \text{ MeV g}^{-1} \text{ cm}^2$   
 $T_{\max} = 2m_e c^2 \beta^2 \gamma^2 / (1 + 2\gamma m_e/M + (m_e/M)^2)$   
 [Max. energy transfer in single collision]

$N_A = 6.022 \cdot 10^{23}$   
 [Avogadro's number]  
 $r_e = e^2 / 4\pi\epsilon_0 m_e c^2 = 2.8 \text{ fm}$   
 [Classical electron radius]

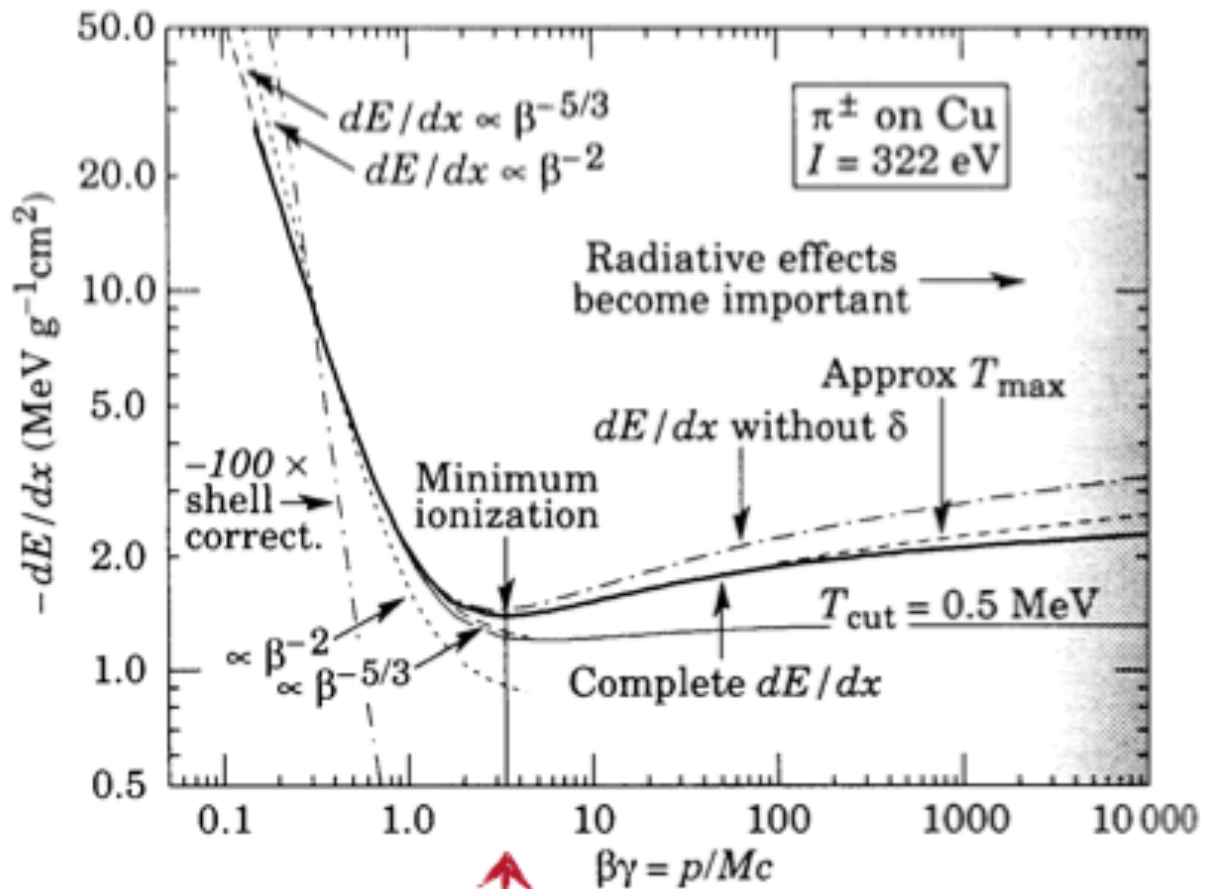
$z$  : Charge of incident particle  
 $M$  : Mass of incident particle  
 $Z$  : Charge number of medium  
 $A$  : Atomic mass of medium

$m_e = 511 \text{ keV}$   
 [Electron mass]  
 $\beta = v/c$   
 [Velocity]  
 $\gamma = (1 - \beta^2)^{-2}$   
 [Lorentz factor]

$I$  : Mean excitation energy of medium  
 $\delta$  : Density correction [transv. extension of electric field]

Validity:  
 $.05 < \beta\gamma < 500$   
 $M > m_\mu$

# ENERGY LOSS of PIONS in Cu



↑  
 $\beta\gamma = 3-4$

Minimum ionizing particles (MIP):  $\beta\gamma = 3-4$

$dE/dx$  falls  $\sim \beta^{-2}$ ; kinematic factor [precise dependence:  $\sim \beta^{-5/3}$ ]

$dE/dx$  rises  $\sim \ln(\beta\gamma)^2$ ; relativistic rise [rel. extension of transversal E-field]

Saturation at large  $(\beta\gamma)$  due to density effect (correction  $\delta$ ) [polarization of medium]

Units:  $\text{MeV g}^{-1} \text{cm}^2$

MIP loses  $\sim 13 \text{ MeV/cm}$  [density of copper:  $8.94 \text{ g/cm}^3$ ]

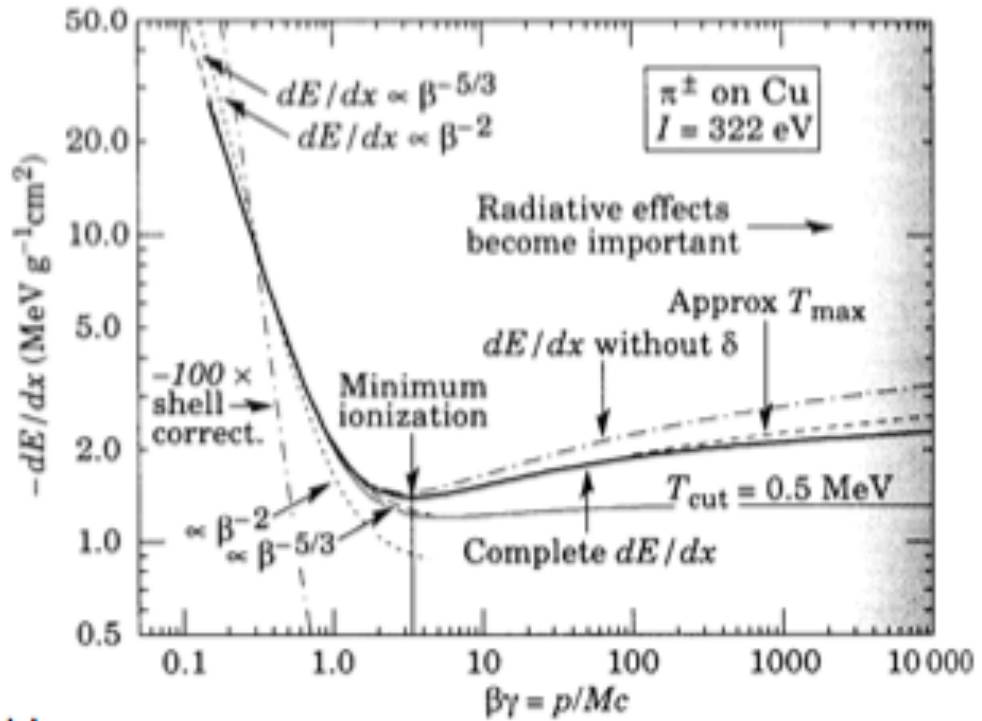
# UNDERSTANDING BETHE-BLOCH

## 1/β<sup>2</sup>-dependence:

Remember:

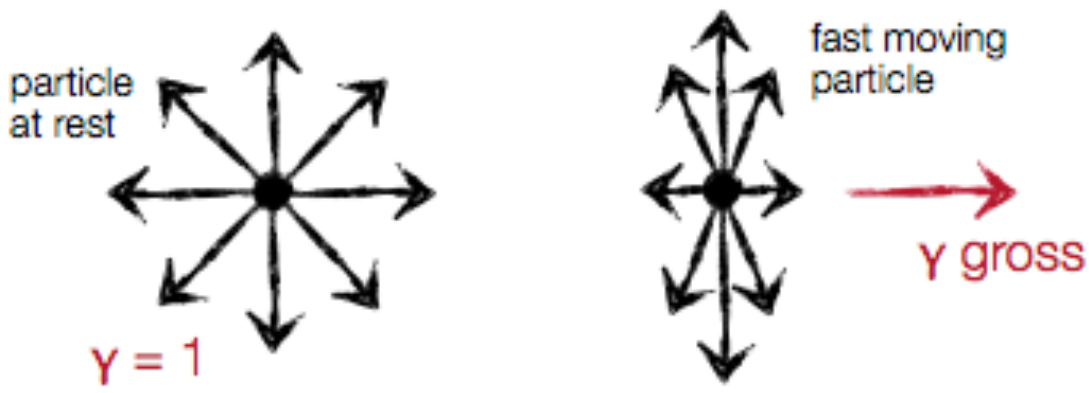
$$\Delta p_{\perp} = \int F_{\perp} dt = \int F_{\perp} \frac{dx}{v}$$

i.e. slower particles feel electric force of atomic electron for longer time ...



## Relativistic rise for βγ > 4:

High energy particle: transversal electric field increases due to Lorentz transform;  $E_y \rightarrow \gamma E_y$ . Thus interaction cross section increases ...



## Corrections:

- low energy : shell corrections
- high energy : density corrections

# UNDERSTANDING BETHE-BLOCH

## Density correction:

Polarization effect ...  
[density dependent]

→ Shielding of electrical field far from particle path; effectively cuts off the long range contribution ...

More relevant at high  $\gamma$  ...  
[Increased range of electric field; larger  $b_{max}$ ; ...]

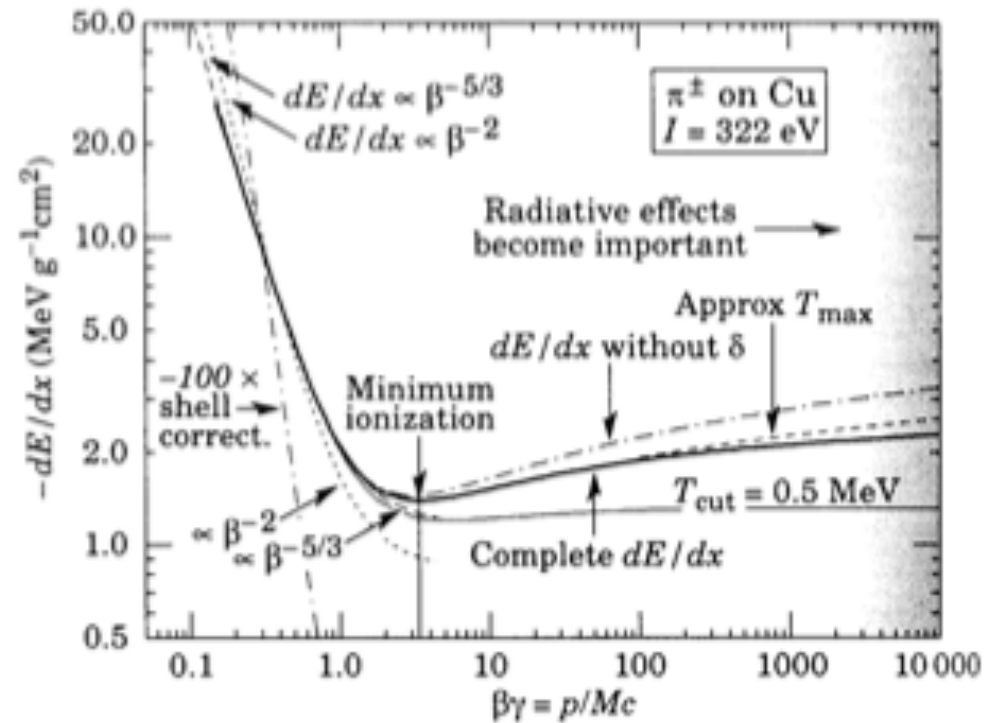
For high energies:

$$\delta/2 \rightarrow \ln(\hbar\omega/I) + \ln \beta\gamma - 1/2$$

## Shell correction:

Arises if particle velocity is close to orbital velocity of electrons, i.e.  $\beta c \sim v_e$ .

Assumption that electron is at rest breaks down ...  
Capture process is possible ...



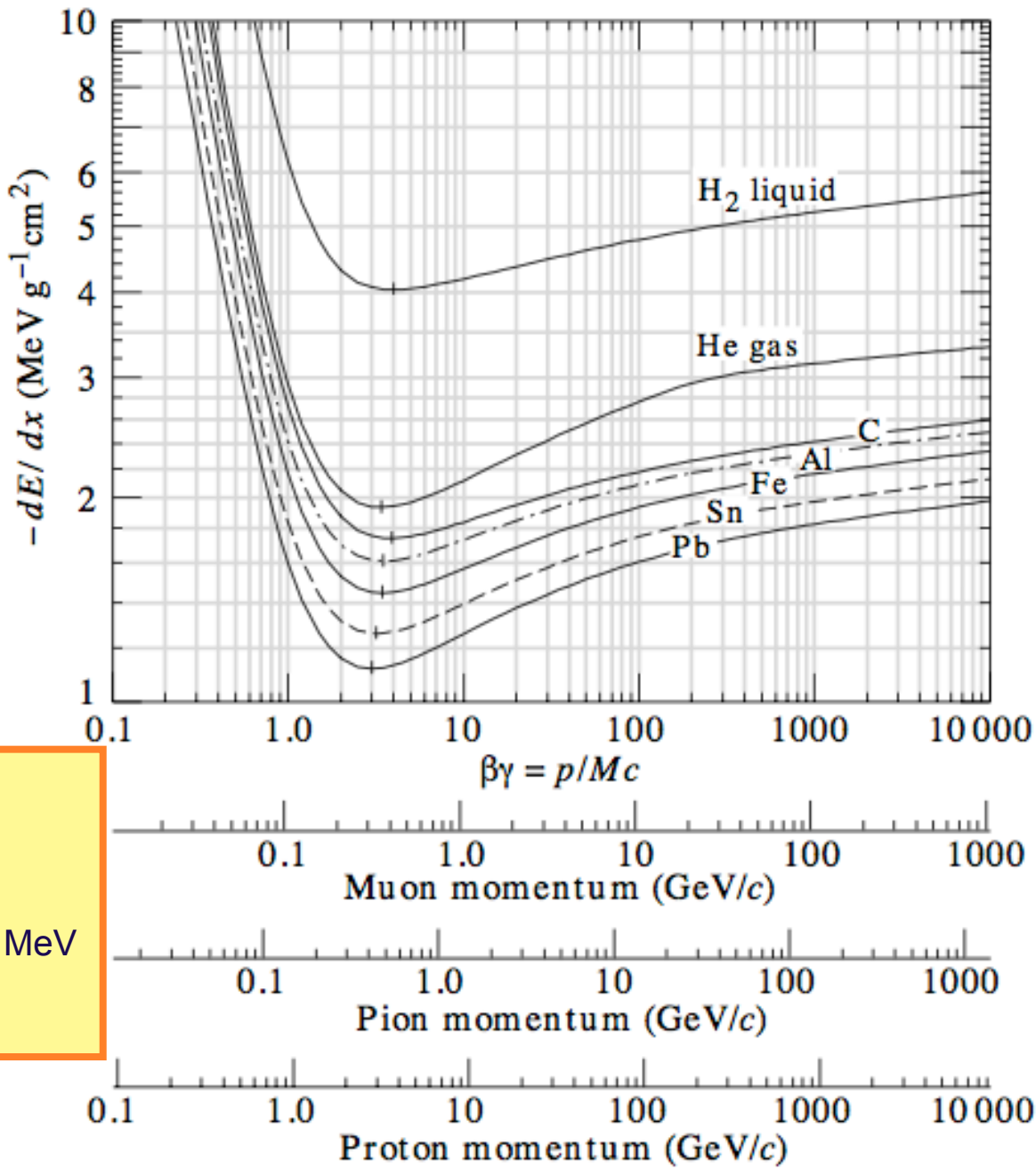
Density effect leads to saturation at high energy ...

Shell correction are in general small ...

# CHARGED PARTICLE ENERGY LOSS in MATERIALS

Dependance on target element  
 Mass A  
 Charge Z

Minimum Ionisation  
 $-dE/dx \sim 1-2 \text{ MeV g}^{-1}\text{cm}^2$   
 e.g. for Pb with  $\rho=11.35 \text{ g/cm}^3$ :  
 $-dE/dx \sim 13 \text{ MeV/cm}$



Can a 1 GeV muon traverse 1 meter of iron ?

$\rho_{\text{Fe}} = 7.87 \text{ g/cm}^3$   
 $dE/dx \sim 1.4 \text{ MeV cm}^2/\text{g}$  ( $p=1 \text{ GeV}$ )  
 $\Delta E = 7.87 \text{ g/cm}^3 \times 100\text{cm} \times 1.4 \text{ MeV cm}^2/\text{g} = 1102 \text{ MeV}$

For a 1 TeV muon ?  $\Delta E \sim 2 \text{ GeV}$

# MATERIAL PROPERTIES

## 6. ATOMIC AND NUCLEAR PROPERTIES OF MATERIALS

**Table 6.1.** Revised May 2002 by D.E. Groom (LBNL). Gases are evaluated at 20°C and 1 atm (in parentheses) or at STP [square brackets]. Densities and refractive indices without parentheses or brackets are for solids or liquids, or are for cryogenic liquids at the indicated boiling point (BP) at 1 atm. Refractive indices are evaluated at the sodium D line. Data for compounds and mixtures are from Refs. 1 and 2. Further materials and properties are given in Ref. 3 and at <http://pdg.lbl.gov/AtomicNuclearProperties>.

Material	$Z$	$A$	$\langle Z/A \rangle$	Nuclear collision length $\lambda_T$ {g/cm <sup>2</sup> }	Nuclear interaction length $\lambda_I$ {g/cm <sup>2</sup> }	Nuclear $dE/dx _{\min}$ <sup>b</sup> { $\frac{\text{MeV}}{\text{g/cm}^2}$ }	Radiation length $X_0$ {g/cm <sup>2</sup> }	Radiation length {cm}	Density {g/cm <sup>3</sup> } {(g/ℓ) for gas}	Liquid boiling point at 1 atm(K)	Refractive index $n$ (( $n - 1$ )×10 <sup>6</sup> for gas)
H <sub>2</sub> gas	1	1.00794	0.99212	43.3	50.8	(4.103)	61.28 <sup>d</sup>	(731000)	(0.0838)[0.0899]		[139.2]
H <sub>2</sub> liquid	1	1.00794	0.99212	43.3	50.8	4.034	61.28 <sup>d</sup>	866	0.0708	20.39	1.112
D <sub>2</sub>	1	2.0140	0.49652	45.7	54.7	(2.052)	122.4	724	0.169[0.179]	23.65	1.128 [138]
He	2	4.002602	0.49968	49.9	65.1	(1.937)	94.32	756	0.1249[0.1786]	4.224	1.024 [34.9]
Li	3	6.941	0.43221	54.6	73.4	1.639	82.76	155	0.534		—
Be	4	9.012182	0.44384	55.8	75.2	1.594	65.19	35.28	1.848		—
C	6	12.011	0.49954	60.2	86.3	1.745	42.70	18.8	2.265 <sup>e</sup>		—
N <sub>2</sub>	7	14.00674	0.49976	61.4	87.8	(1.825)	37.99	47.1	0.8073[1.250]	77.36	1.205 [298]
O <sub>2</sub>	8	15.9994	0.50002	63.2	91.0	(1.801)	34.24	30.0	1.141[1.428]	90.18	1.22 [296]
F <sub>2</sub>	9	18.9984032	0.47372	65.5	95.3	(1.675)	32.93	21.85	1.507[1.696]	85.24	[195]
Ne	10	20.1797	0.49555	66.1	96.6	(1.724)	28.94	24.0	1.204[0.9005]	27.09	1.092 [67.1]
Al	13	26.981539	0.48181	70.6	106.4	1.615	24.01	8.9	2.70		—
Si	14	28.0855	0.49848	70.6	106.0	1.664	21.82	9.36	2.33		3.95
Ar	18	39.948	0.45059	76.4	117.2	(1.519)	19.55	14.0	1.396[1.782]	87.28	1.233 [283]
Ti	22	47.867	0.45948	79.9	124.9	1.476	16.17	3.56	4.54		—
Fe	26	55.845	0.46556	82.8	131.9	1.451	13.84	1.76	7.87		—
Cu	29	63.546	0.45636	85.6	134.9	1.403	12.86	1.43	8.96		—
Ge	32	72.61	0.44071	88.3	140.5	1.371	12.25	2.30	5.323		—
Sn	50	118.710	0.42120	100.2	163	1.264	8.82	1.21	7.31		—
Xe	54	131.29	0.41130	102.8	169	(1.255)	8.48	2.87	2.953[5.858]	165.1	[701]
W	74	183.84	0.40250	110.3	185	1.145	6.76	0.35	19.3		—
Pt	78	195.08	0.39984	113.3	189.7	1.129	6.54	0.305	21.45		—
Pb	82	207.2	0.39575	116.2	194	1.123	6.37	0.56	11.35		—
U	92	238.0289	0.38651	117.0	199	1.082	6.00	≈0.32	≈18.95		—

Material	$Z$	$A$	$(Z/A)$	Nuclear $\alpha$ collision length $\lambda_T$ {g/cm <sup>2</sup> }	Nuclear $\alpha$ interaction length $\lambda_I$ {g/cm <sup>2</sup> }	$dE/dx _{\min}^b$ { $\frac{\text{MeV}}{\text{g/cm}^2}$ }	Radiation length $X_0$ {g/cm <sup>2</sup> }	{cm}	Density {g/cm <sup>3</sup> } {g/ℓ} for gas)	Liquid boiling point at 1 atm(K)	Refractive index $n$ {(n-1)×10 <sup>6</sup> } for gas)
Air, (20°C, 1 atm.), [STP]			0.49919	62.0	90.0	(1.815)	36.66	[30420]	(1.205)[1.2931]	78.8	(273) [293]
H <sub>2</sub> O			0.55509	60.1	83.6	1.991	36.08	36.1	1.00	373.15	1.33
CO <sub>2</sub> gas			0.49989	62.4	89.7	(1.819)	36.2	[18310]	[1.977]		[410]
CO <sub>2</sub> solid (dry ice)			0.49989	62.4	89.7	1.787	36.2	23.2	1.563	sublimes	—
Shielding concrete <sup>f</sup>			0.50274	67.4	99.9	1.711	26.7	10.7	2.5		—
SiO <sub>2</sub> (fused quartz)			0.49926	66.5	97.4	1.699	27.05	12.3	2.20 <sup>g</sup>		1.458
Dimethyl ether, (CH <sub>3</sub> ) <sub>2</sub> O			0.54778	59.4	82.9	—	38.89	—	—	248.7	—
Methane, CH <sub>4</sub>			0.62333	54.8	73.4	(2.417)	46.22	[64850]	0.4224[0.717]	111.7	[444]
Ethane, C <sub>2</sub> H <sub>6</sub>			0.59861	55.8	75.7	(2.304)	45.47	[34035]	0.509(1.356) <sup>h</sup>	184.5	(1.038) <sup>h</sup>
Propane, C <sub>3</sub> H <sub>8</sub>			0.58962	56.2	76.5	(2.262)	45.20	—	(1.879)	231.1	—
Isobutane, (CH <sub>3</sub> ) <sub>2</sub> CHCH <sub>3</sub>			0.58496	56.4	77.0	(2.239)	45.07	[16930]	[2.67]	261.42	[1900]
Octane, liquid, CH <sub>3</sub> (CH <sub>2</sub> ) <sub>6</sub> CH <sub>3</sub>			0.57778	56.7	77.7	2.123	44.86	63.8	0.703	398.8	1.397
Paraffin wax, CH <sub>3</sub> (CH <sub>2</sub> ) <sub>n≈23</sub> CH <sub>3</sub>			0.57275	56.9	78.2	2.087	44.71	48.1	0.93		—
Nylon, type 6 <sup>i</sup>			0.54790	58.5	81.5	1.974	41.84	36.7	1.14		—
Polycarbonate (Lexan) <sup>j</sup>			0.52697	59.5	83.9	1.886	41.46	34.6	1.20		—
Polyethylene terephthlate (Mylar) <sup>k</sup>			0.52037	60.2	85.7	1.848	39.95	28.7	1.39		—
Polyethylene <sup>l</sup>			0.57034	57.0	78.4	2.076	44.64	≈47.9	0.92–0.95		—
Polyimide film (Kapton) <sup>m</sup>			0.51264	60.3	85.8	1.820	40.56	28.6	1.42		—
Lucite, Plexiglas <sup>n</sup>			0.53937	59.3	83.0	1.929	40.49	≈34.4	1.16–1.20		≈1.49
Polystyrene, scintillator <sup>o</sup>			0.53768	58.5	81.9	1.936	43.72	42.4	1.032		1.581
Polytetrafluoroethylene (Teflon) <sup>p</sup>			0.47992	64.2	93.0	1.671	34.84	15.8	2.20		—
Polyvinyltolulene, scintillator <sup>q</sup>			0.54155	58.3	81.5	1.956	43.83	42.5	1.032		—
Aluminum oxide (Al <sub>2</sub> O <sub>3</sub> )			0.49038	67.0	98.9	1.647	27.94	7.04	3.97		1.761
Barium fluoride (BaF <sub>2</sub> )			0.42207	92.0	145	1.303	9.91	2.05	4.89		1.56
Bismuth germanate (BGO) <sup>r</sup>			0.42065	98.2	157	1.251	7.97	1.12	7.1		2.15
Cesium iodide (CsI)			0.41569	102	167	1.243	8.39	1.85	4.53		1.80
Lithium fluoride (LiF)			0.46262	62.2	88.2	1.614	39.25	14.91	2.632		1.392
Sodium fluoride (NaF)			0.47632	66.9	98.3	1.69	29.87	11.68	2.558		1.336
Sodium iodide (NaI)			0.42697	94.6	151	1.305	9.49	2.59	3.67		1.775
Silica Aerogel <sup>s</sup>			0.50093	66.3	96.9	1.740	27.25	136@ρ=0.2	0.04–0.6		1.0+0.21ρ
NEMA G10 plate <sup>t</sup>				62.6	90.2	1.87	33.0	19.4	1.7		—



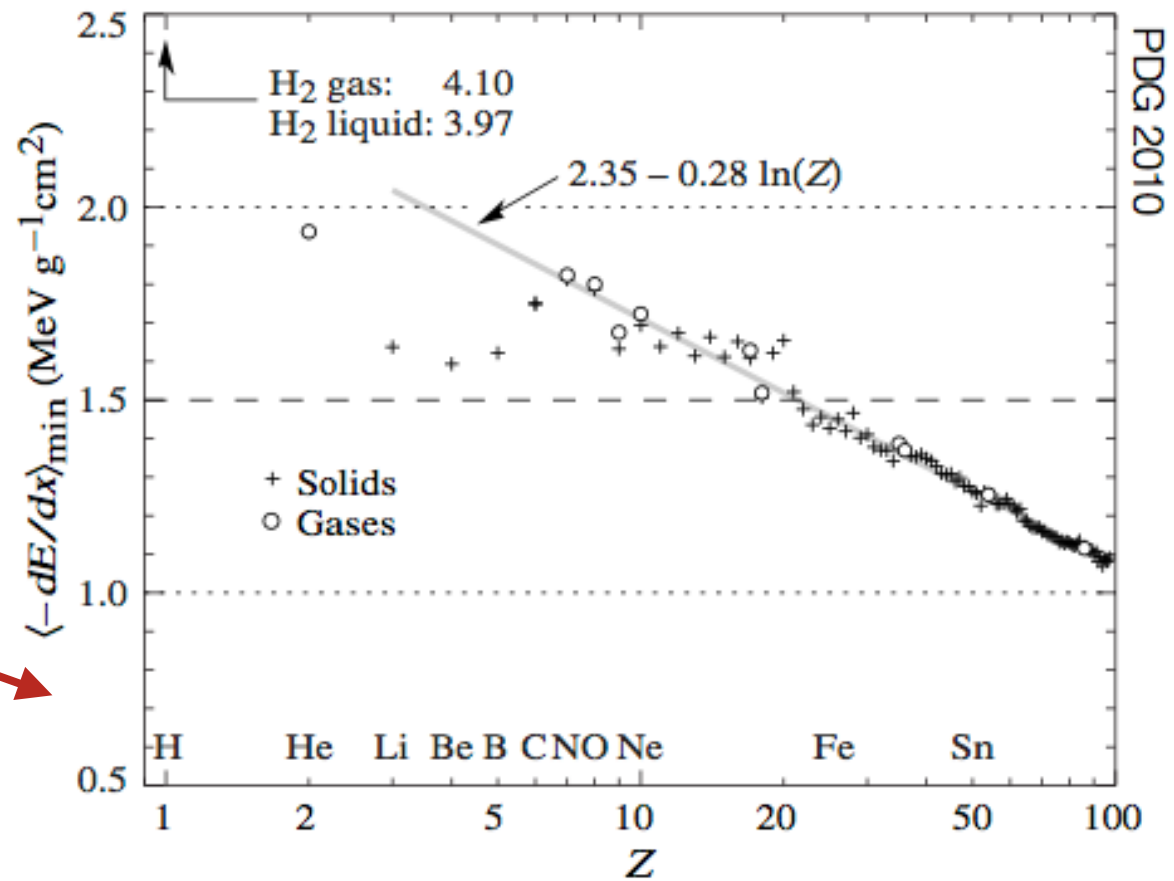
# STOPPING POWER AT MINIMUM IONISATION

$$-\left\langle \frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

[·ρ]  
density

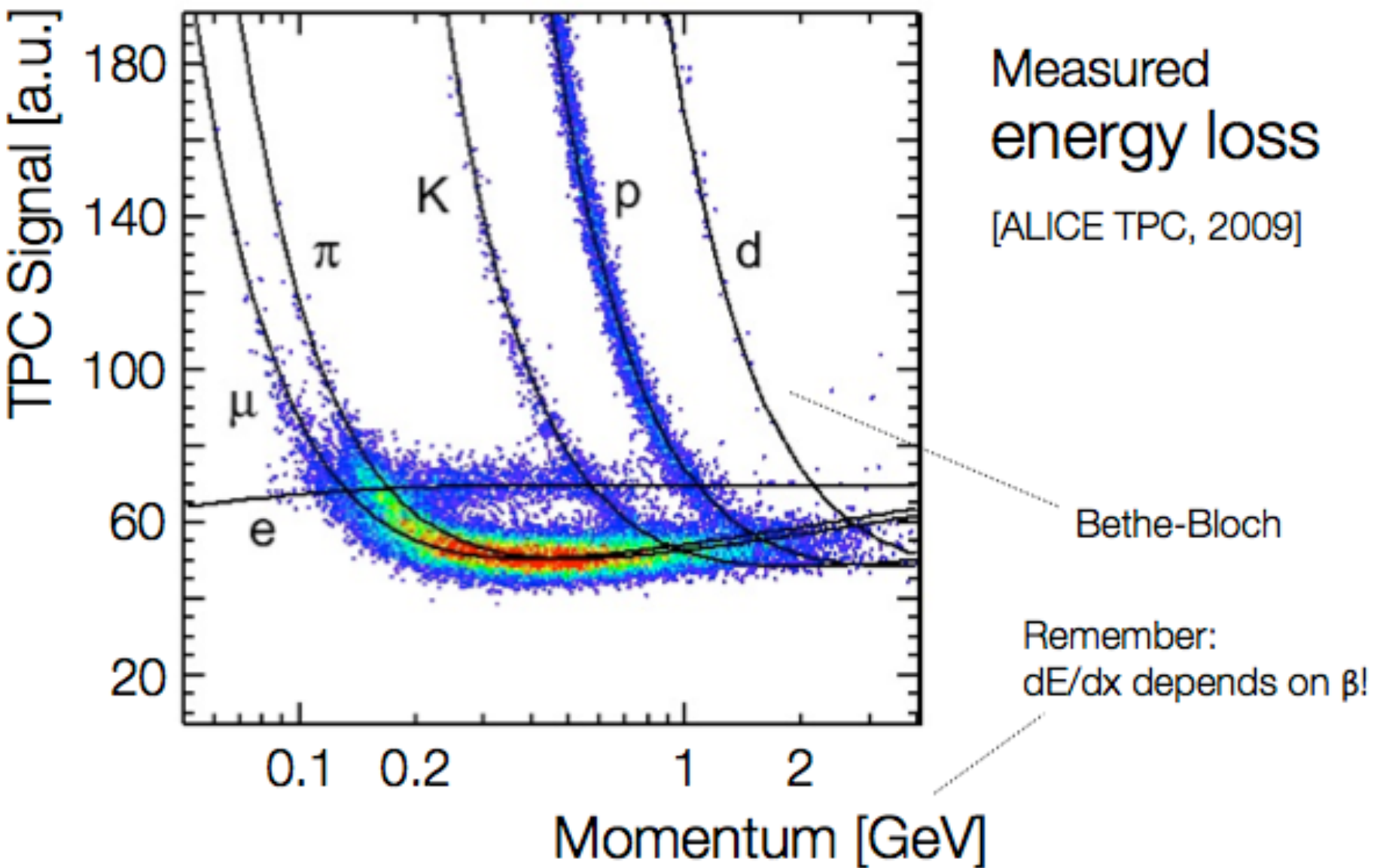
material  
dependance  
on Z/A ~ 1/2

Small  
dependance  
with Z



Stopping power at minimum ionization for the chemical elements. The straight line is fitted for Z > 6. A simple functional dependence on Z is not to be expected, since  $\langle -dE/dx \rangle$  also depends on other variables.

# dE/dX and PARTICLE IDENTIFICATION

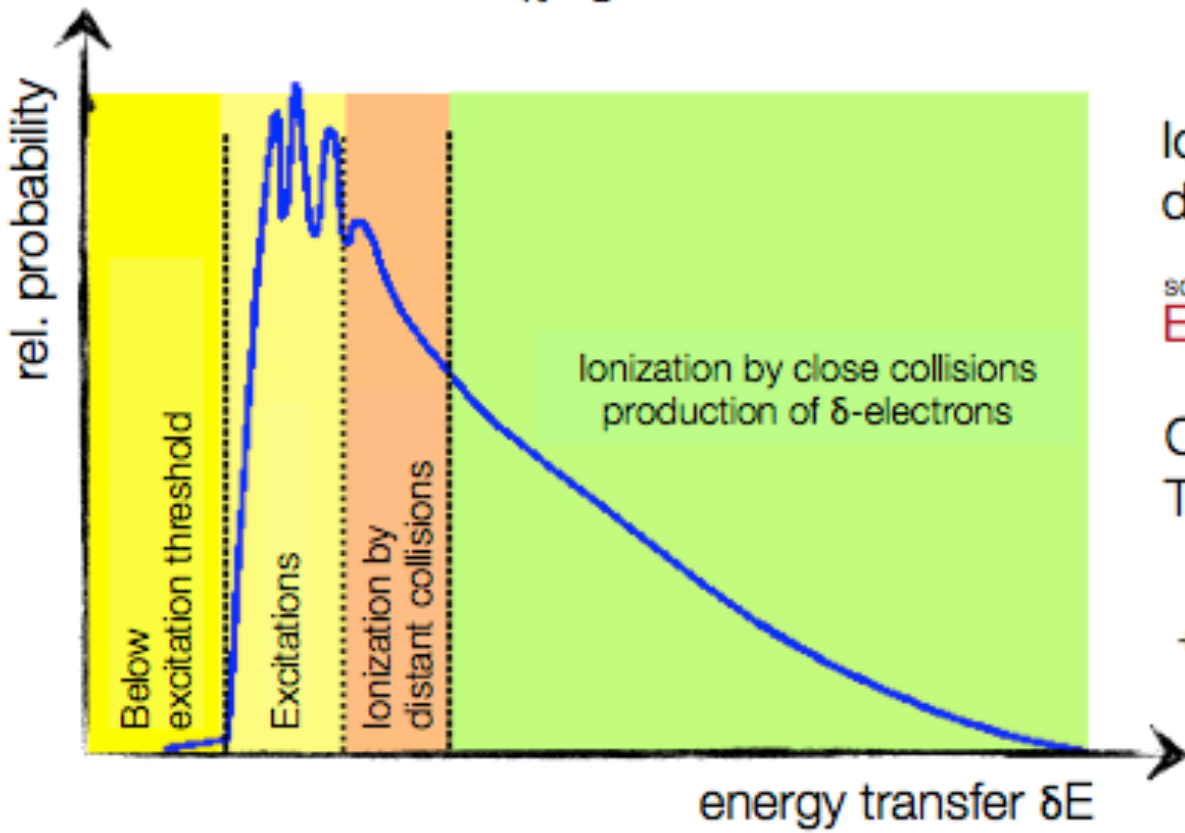


# dE/dx FLUCTUATIONS

Bethe-Bloch describes **mean** energy loss; measurement via energy loss  $\Delta E$  in a material of thickness  $\Delta x$  with

$$\Delta E = \sum_{n=1}^N \delta E_n$$

$N$  : number of collisions  
 $\delta E$  : energy loss in a single collision



Ionization loss  $\delta E$   
 distributed statistically ...

so-called  
**Energy loss 'straggling'**

Complicated problem ...  
 Thin absorbers: **Landau distribution**

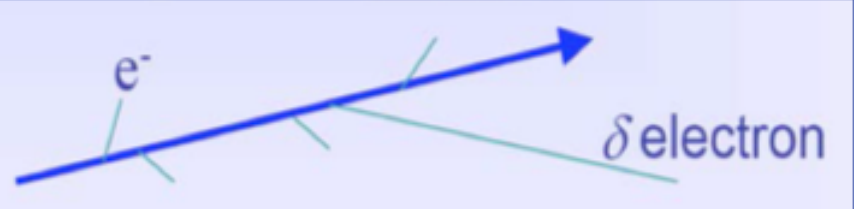
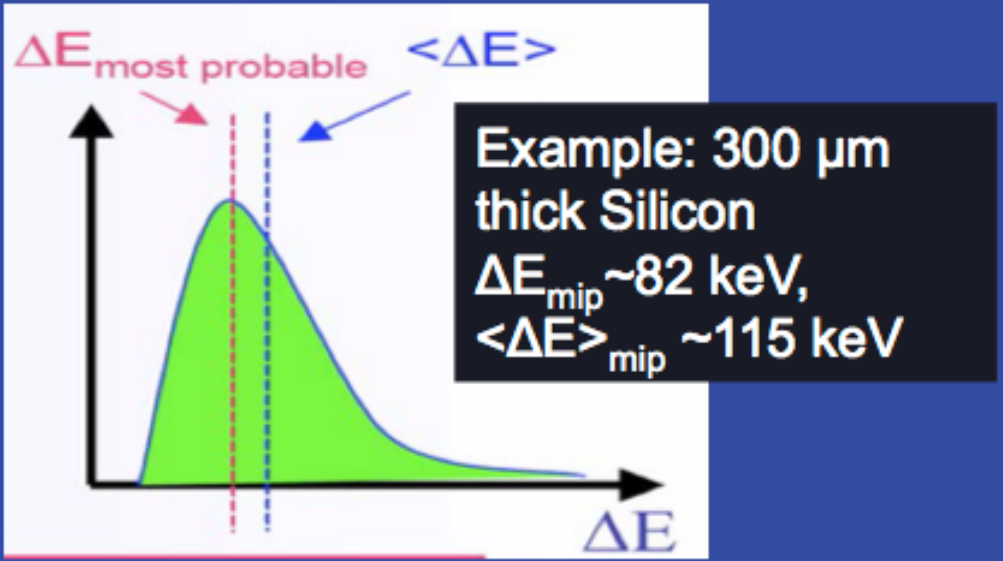
Standard Gauss with mean energy loss  $E_0$   
 + tail towards high energies due to  $\delta$ -electrons

see also Allison & Cobb  
 [Ann. Rev. Nucl. Part. Sci. 30 (1980) 253.]

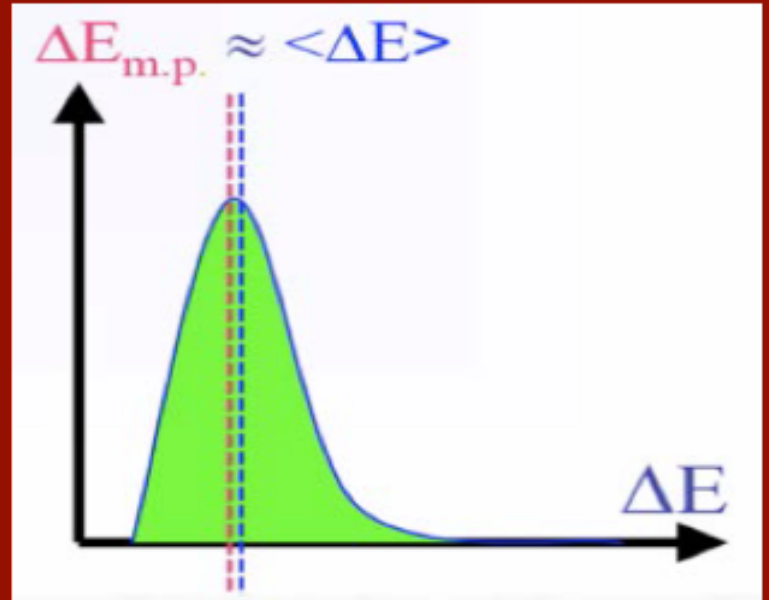
# dE/dx FLUCTUATIONS

In a detector, with limited granularity, one measures  $\Delta E/\Delta x$ , and not  $\langle dE/dx \rangle$   
i.e. the energy deposit in a thickness of material  
therefore multi-measurements are needed.

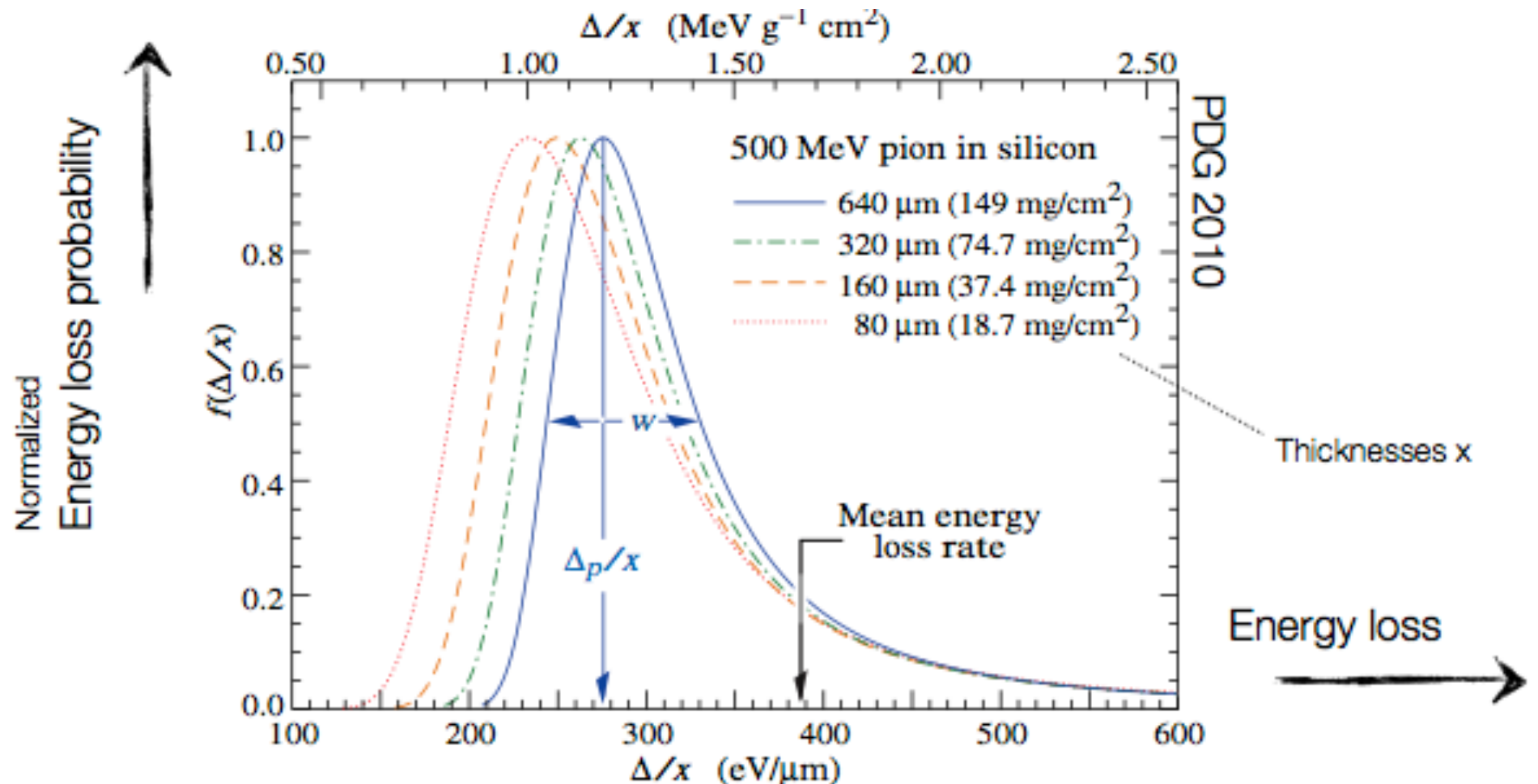
- Thin layers or low density materials: dE/dx has large fluctuations towards high losses (Landau)



- Thick layers and high density materials: the dE/dx is a more Gaussian-like (many collisions)



# dE/dx FLUCTUATIONS - LANDAU DISTRIBUTION



Approximation:

$$f(\Delta/x) = \frac{1}{\sqrt{2\pi}} \exp \left[ -\frac{1}{2} \left( \frac{\Delta/x - a(\Delta/x)_{mip}}{\xi} \right)^2 + e^{-\left( \frac{\Delta/x - a(\Delta/x)_{mip}}{\xi} \right)} \right]$$

for full form see e.g. Leo

$\xi$ : material constant

# ENERGY LOSS of ELECTRONS

Bethe-Bloch formula needs modification

Incident and target electron have same mass  $m_e$   
Scattering of identical, indistinguishable particles

$$-\left\langle \frac{dE}{dx} \right\rangle_{\text{el.}} = K \frac{Z}{A} \frac{1}{\beta^2} \left[ \ln \frac{m_e \beta^2 c^2 \gamma^2 T}{2I^2} + F(\gamma) \right]$$

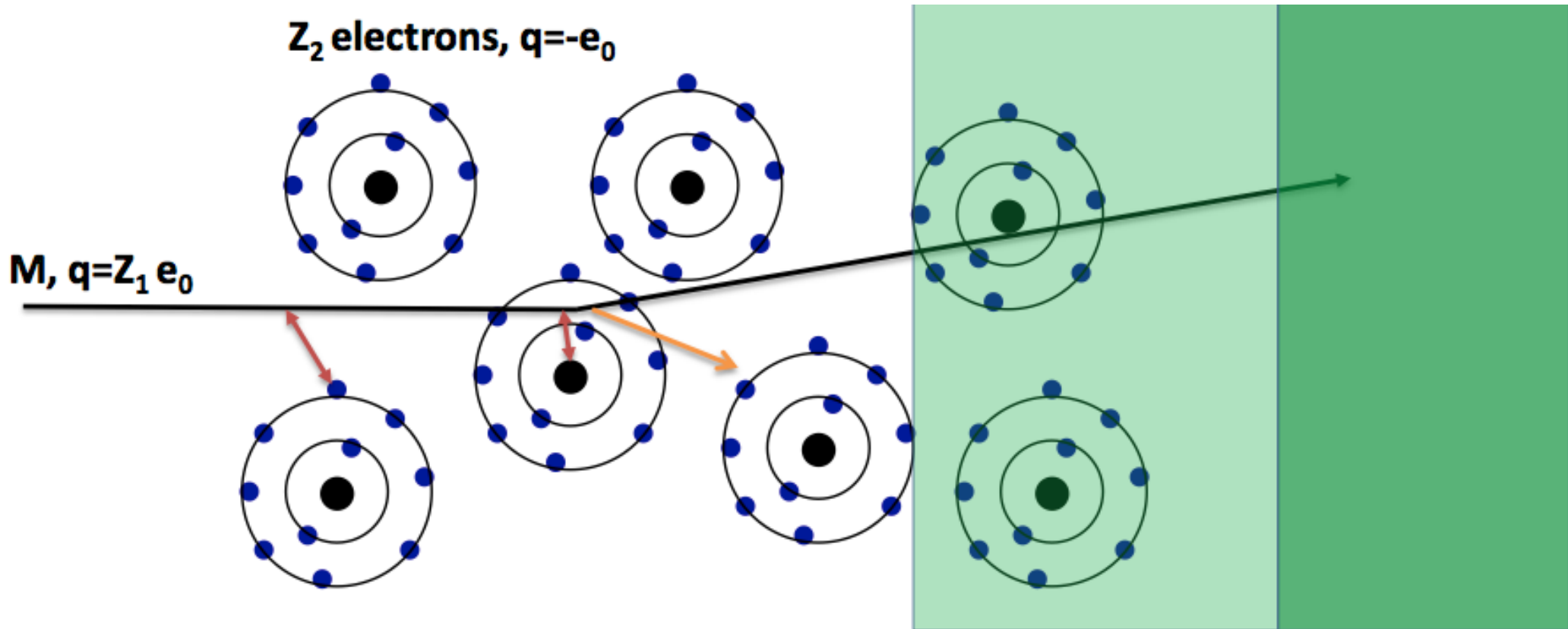
[T: kinetic energy of electron]

$$W_{\text{max}} = \frac{1}{2}T$$

Remark: different energy loss for electrons and positrons at low energy as positrons are not identical with electrons; different treatment ...

# ELECTROMAGNETIC INTERACTION

## PARTICLE - MATTER



**Interaction with the atomic electrons.**  
The incoming particle loses energy and the atoms are **excited** or **ionised**.

**Interaction with the atomic nucleus.**

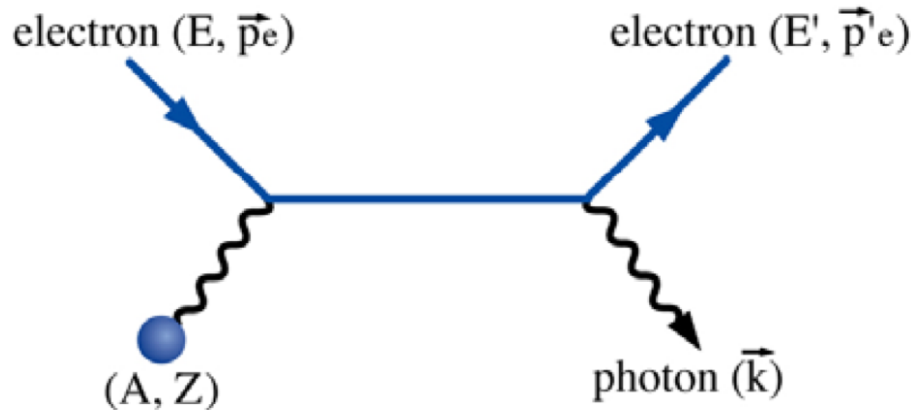
The incoming particle is deflected causing **multiple scattering** of the particle in the material.

During this scattering a **Bremsstrahlung photon** can be emitted

In case the particle's velocity is larger than the velocity of light in the medium, the resulting EM shockwave manifests itself as **Cherenkov radiation**. When the particle crosses the boundary between two media, there is a probability of 1% to produce an Xray photon called **Transition radiation**.

# BREMSSTRAHLUNG

Real photon emission in the electromagnetic field of the atomic nucleus



Electric field of the nucleus + of the electrons  
 $Z(Z+1)$

At large radius, electrons screen the nucleus  
 $\ln(183Z^{-1/3})$

$$d\sigma/dk = 4 \alpha Z(Z+1)r_e^2 \ln(183Z^{-1/3})(4/3-4/3y+y^2)/k \quad [\text{D.F.}]$$

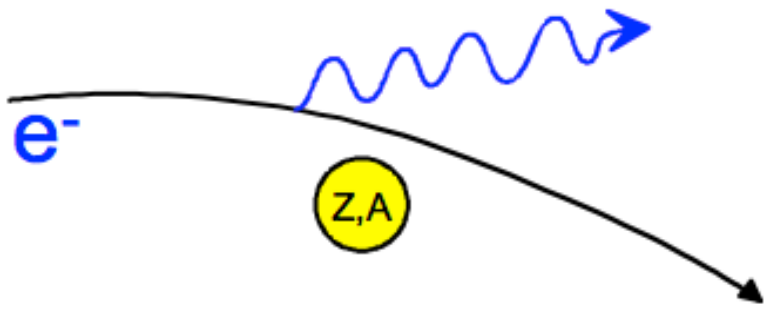
where  $y=k/E$  and  $r_e = \frac{1}{4\pi\epsilon_0} \cdot \frac{e^2}{m_e c^2} = 2.818 \cdot 10^{-15} \text{ m}$  classical radius of the electron.

For a given E, the average energy lost by radiation,  $dE$ , is obtained by integrating over  $y$ .



# BREMSSTRAHLUNG & RADIATION LENGTH

Bremsstrahlung arises if particles are accelerated in Coulomb field of nucleus



$$\frac{dE}{dx} = 4\alpha N_A \frac{z^2 Z^2}{A} \left( \frac{1}{4\pi\epsilon_0} \frac{e^2}{mc^2} \right)^2 E \ln \frac{183}{Z^{\frac{1}{3}}} \propto \frac{E}{m^2}$$

i.e. energy loss proportional to  $1/m^2$  → **main relevance for electrons ...**  
 ... or ultra-relativistic muons

Consider electrons:

$$\frac{dE}{dx} = 4\alpha N_A \frac{Z^2}{A} r_e^2 \cdot E \ln \frac{183}{Z^{\frac{1}{3}}}$$

$$\frac{dE}{dx} = \frac{E}{X_0} \quad \text{with} \quad X_0 = \frac{A}{4\alpha N_A Z^2 r_e^2 \ln \frac{183}{Z^{\frac{1}{3}}}}$$

[Radiation length in g/cm<sup>2</sup>]

$$\Rightarrow E = E_0 e^{-x/X_0}$$

After passage of one  $X_0$  electron has lost all but  $(1/e)^{\text{th}}$  of its energy  
 [i.e. 63%]

# RADIATION LENGTH

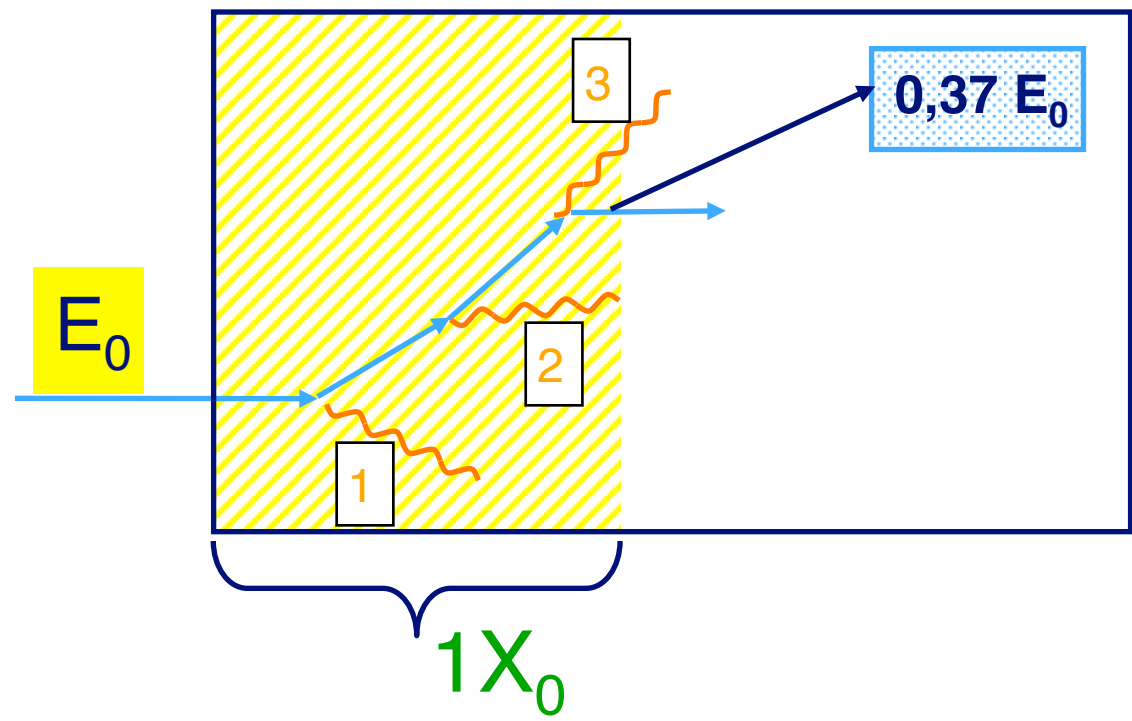
The radiation length is a “universal” distance, very useful to describe electromagnetic showers (electrons & photons)

$X_0$  is the distance after which the incident electron has radiated  $(1-1/e)$  63% of its incident energy

$$dE/dx = E/X_0$$

$$dE/E = dx/X_0$$

$$E = E_0 e^{-x/X_0}$$



	Air	Eau	Al	LAr	Fe	Pb	PbWO <sub>4</sub>	LAr/Pb
Z	-	-	13	18	26	82	-	-
$X_0$ (cm)	30420	36	8,9	14	1,76	0.56	0.89	1.9

# CRITICAL ENERGY

Critical energy:

$$\left(\frac{dE}{dx}\right)_{\text{Tot}} = \left(\frac{dE}{dx}\right)_{\text{Ion}} + \left(\frac{dE}{dx}\right)_{\text{Brems}}$$

$$\left.\frac{dE}{dx}(E_c)\right|_{\text{Brems}} = \left.\frac{dE}{dx}(E_c)\right|_{\text{Ion}}$$

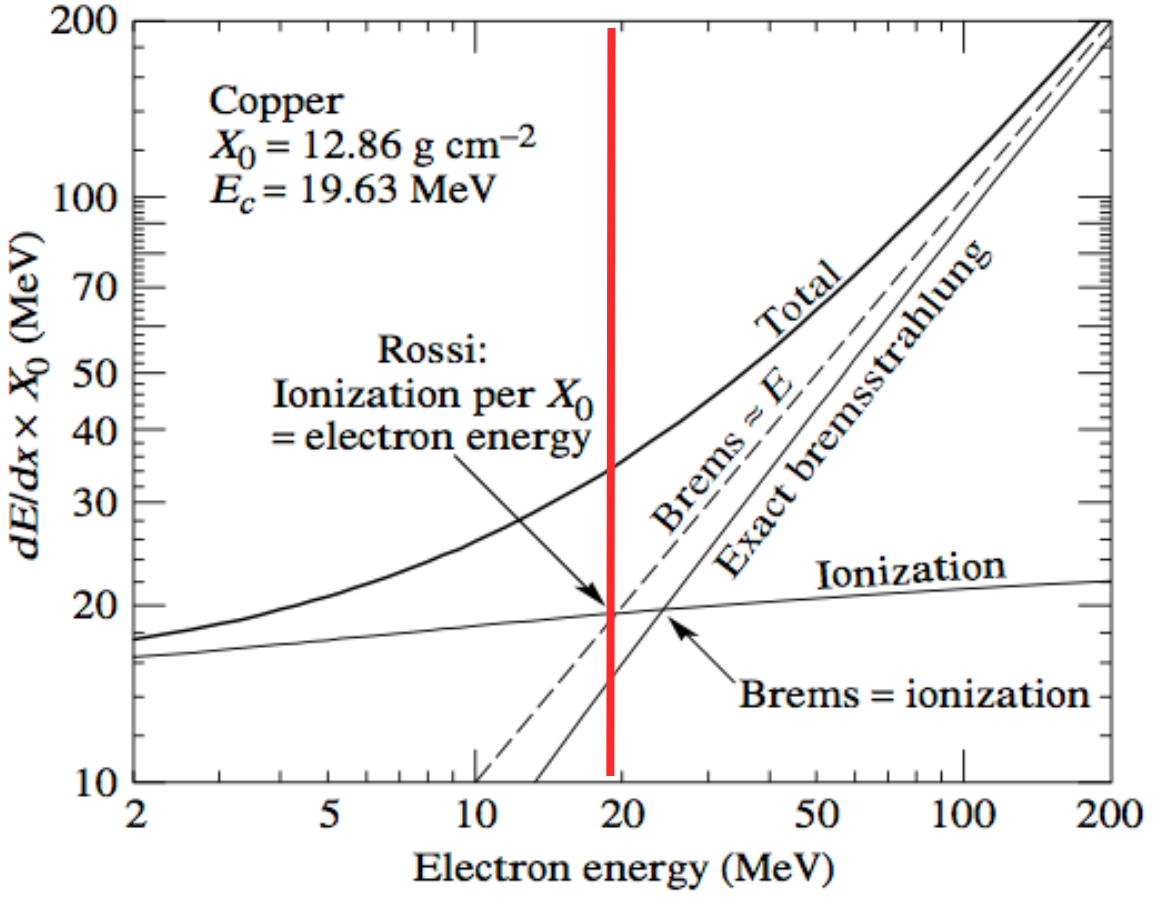
Approximation:

$$E_c^{\text{Gas}} = \frac{710 \text{ MeV}}{Z + 0.92}$$

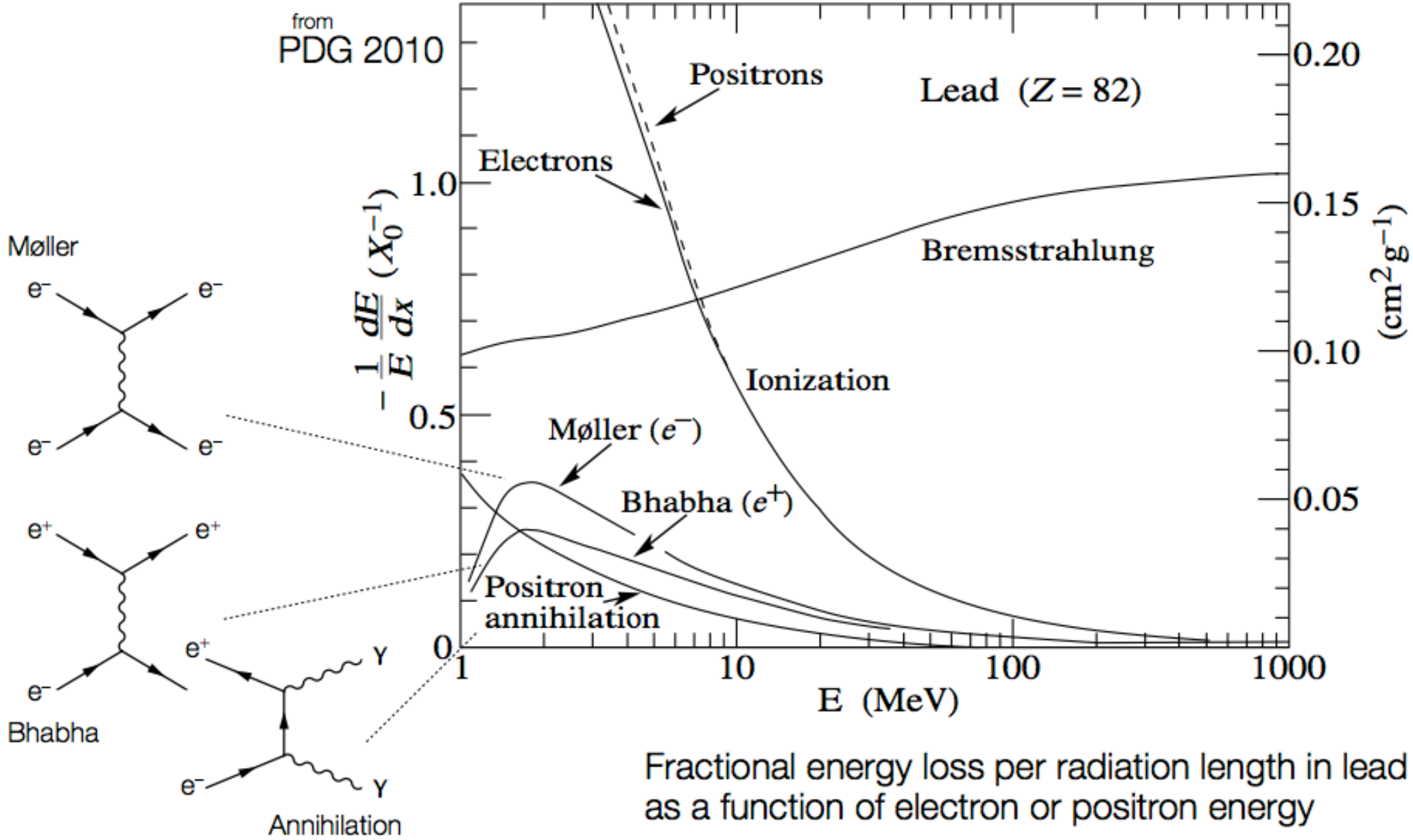
$$E_c^{\text{Sol/Liq}} = \frac{610 \text{ MeV}}{Z + 1.24}$$

Example Copper:

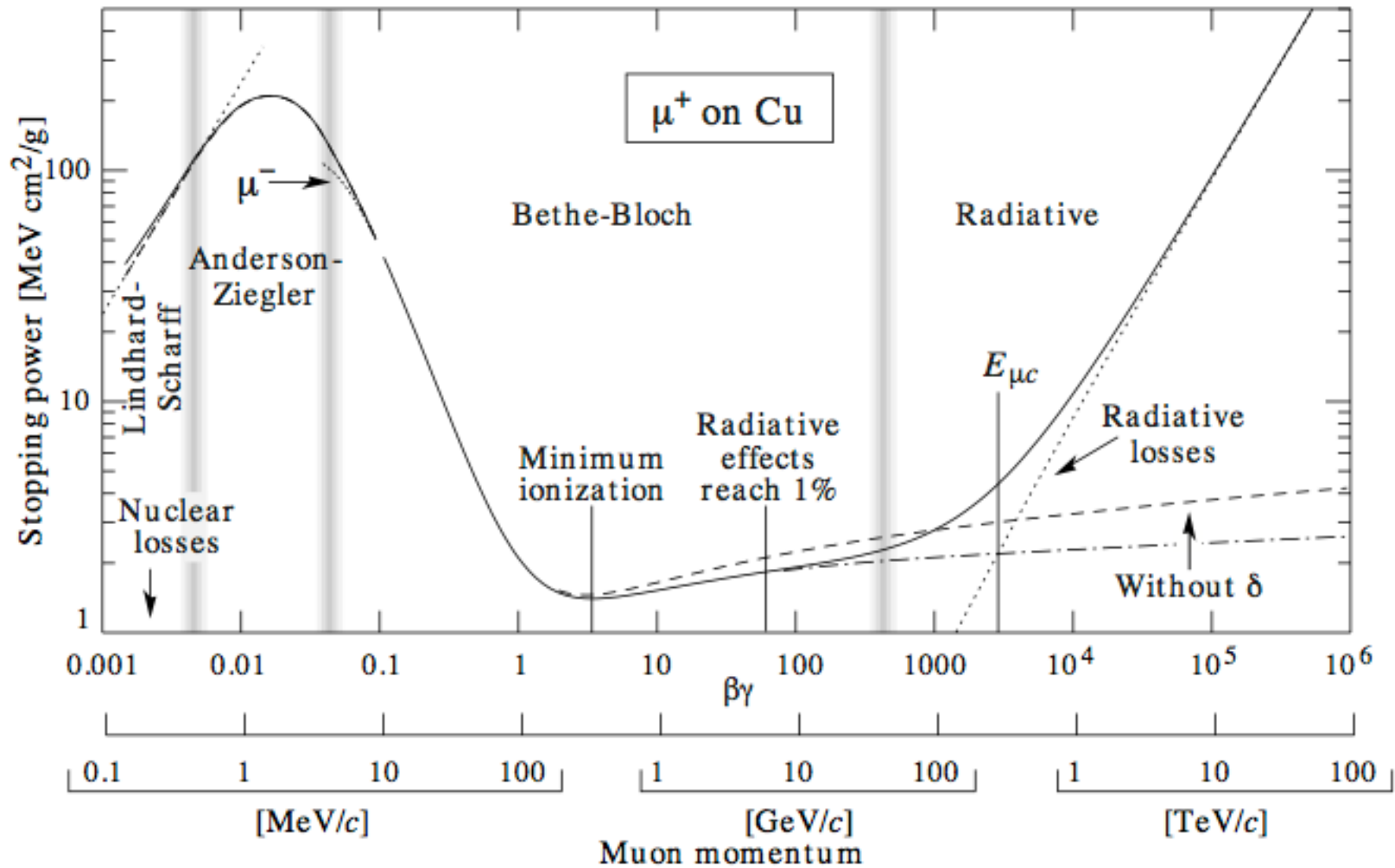
$$E_c \approx 610/30 \text{ MeV} \approx 20 \text{ MeV}$$



# TOTAL ENERGY LOSS FOR ELECTRONS



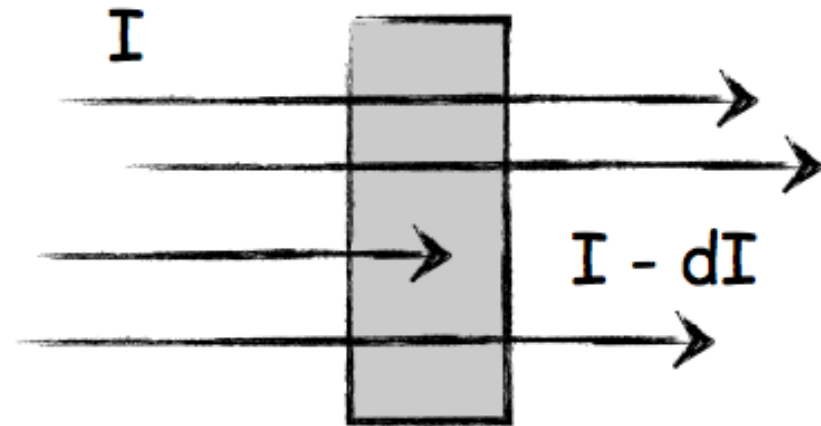
# $\mu^+$ in COPPER



# INTERACTION OF PHOTONS WITH MATTER

Characteristic for interactions of photons with matter:

A single interaction removes photon from beam !



Possible Interactions

Photoelectric Effect  
Compton Scattering  
Pair Production

Rayleigh Scattering ( $\gamma A \rightarrow \gamma A$ ; A = atom; coherent)  
Thomson Scattering ( $\gamma e \rightarrow \gamma e$ ; elastic scattering)  
Photo Nuclear Absorption ( $\gamma K \rightarrow pK/nK$ )  
Nuclear Resonance Scattering ( $\gamma K \rightarrow K^* \rightarrow \gamma K$ )  
Delbruck Scattering ( $\gamma K \rightarrow \gamma K$ )  
Hadron Pair production ( $\gamma K \rightarrow h^+h^-K$ )

$$dI = -\mu I dx$$

[  $\mu$ : absorption coefficient ]

depends on  
E, Z,  $\rho$

→ Beer-Lambert law:

$$I(x) = I_0 e^{-\mu x}$$

with  $\lambda = 1/\mu = 1/n\sigma$   
[ mean free path ]

# PHOTO-ELECTRIC EFFECT

Energy of outgoing electron:

$$E_e = h\nu - I_b$$

Photon energy

Binding energy  
[strongly Z dependent]

Typical energy dependence:

$$\sigma_{ph} = 2\pi r_e^2 \alpha^4 Z^5 (mc^2) / E_\gamma$$

[for  $E_\gamma \gg mc^2$ ]

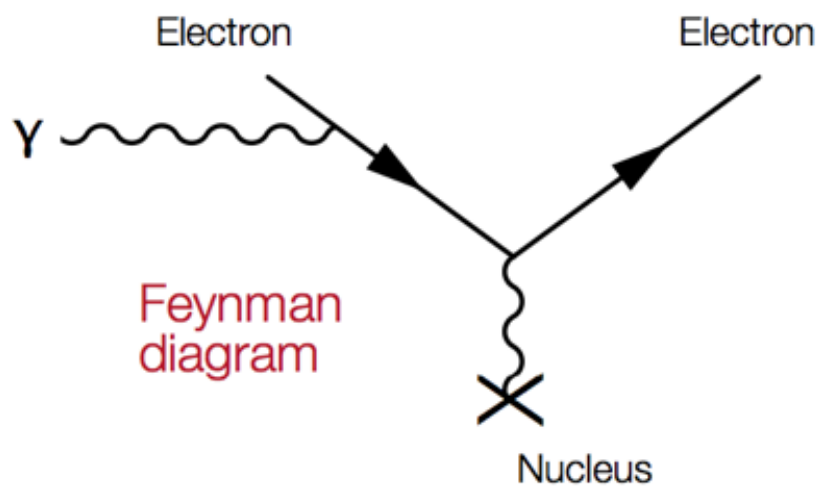
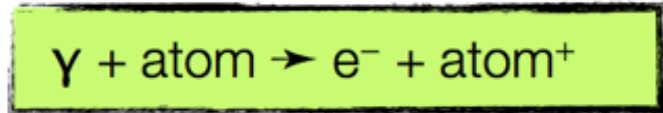
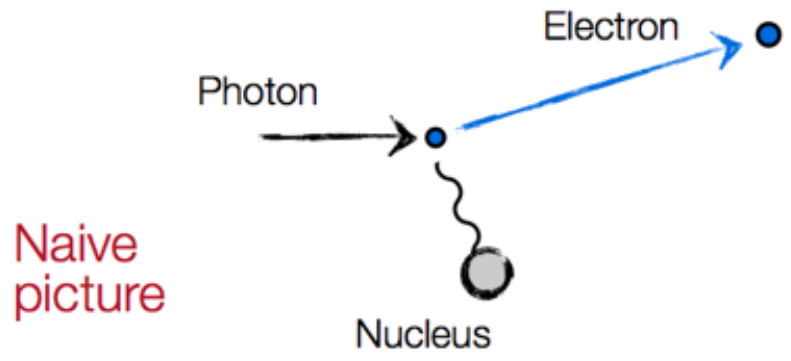
$$\sigma_{ph} = \alpha\pi a_B Z^5 (I_0 / E_\gamma)^{7/2}$$

[for  $I_0 \ll E_\gamma \ll mc^2$ ]

Example values:

$a_B = 0.53 \cdot 10^{-10}$  m;  $I_0 = 13.6$  eV;  $\alpha = 1/137$ ;  $1 \text{ b} = 10^{-24}$  m  
use  $E_\gamma = 100$  keV

→  $\sigma_{ph}(\text{Fe}) = 29$  barn  
 $\sigma_{ph}(\text{Pb}) = 5000$  barn



# PAIR PRODUCTION

Cross Section:  
[for  $E_\gamma \gg m_e c^2$ ]

$$\sigma_{\text{pair}} \approx \underbrace{\frac{7}{9} \left( 4 \alpha r_e^2 Z^2 \ln \frac{183}{Z^{1/3}} \right)}_{A/N_A X_0}$$

[ $X_0$ : radiation length]  
[in cm or g/cm<sup>2</sup>]

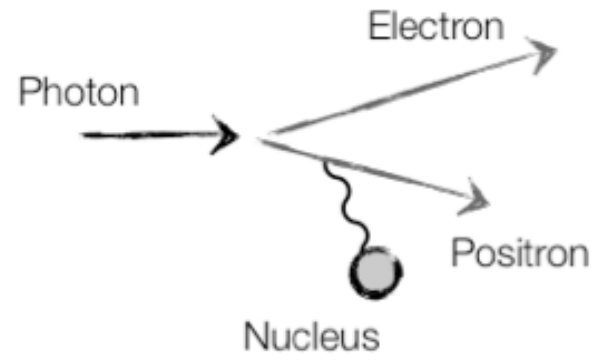
Absorption coefficient:

$$\mu = n\sigma \quad [\text{with } n: \text{particle density}]$$

$$\begin{aligned} \mu &= \rho \cdot N_A/A \sigma_{\text{pair}} \\ &= 7/9 \frac{1}{X_0} \end{aligned}$$

[where now  $X_0$  is in cm]

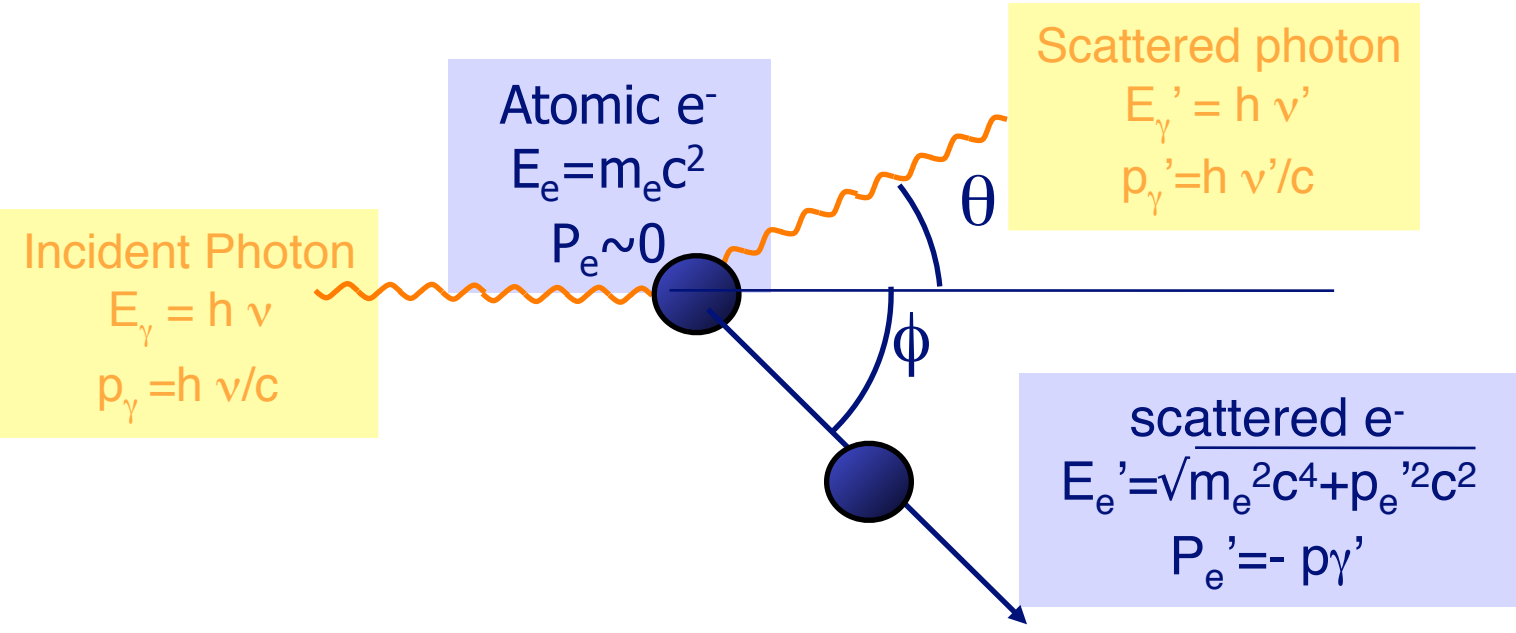
$$I(x) = I_0 e^{-\mu x}$$



	$\rho$ [g/cm <sup>3</sup> ]	$X_0$ [cm]
H <sub>2</sub> [fl.]	0.071	865
C	2.27	18.8
Fe	7.87	1.76
Pb	11.35	0.56
Air	$1.2 \cdot 10^{-3}$	$30 \cdot 10^3$



# COMPTON SCATTERING

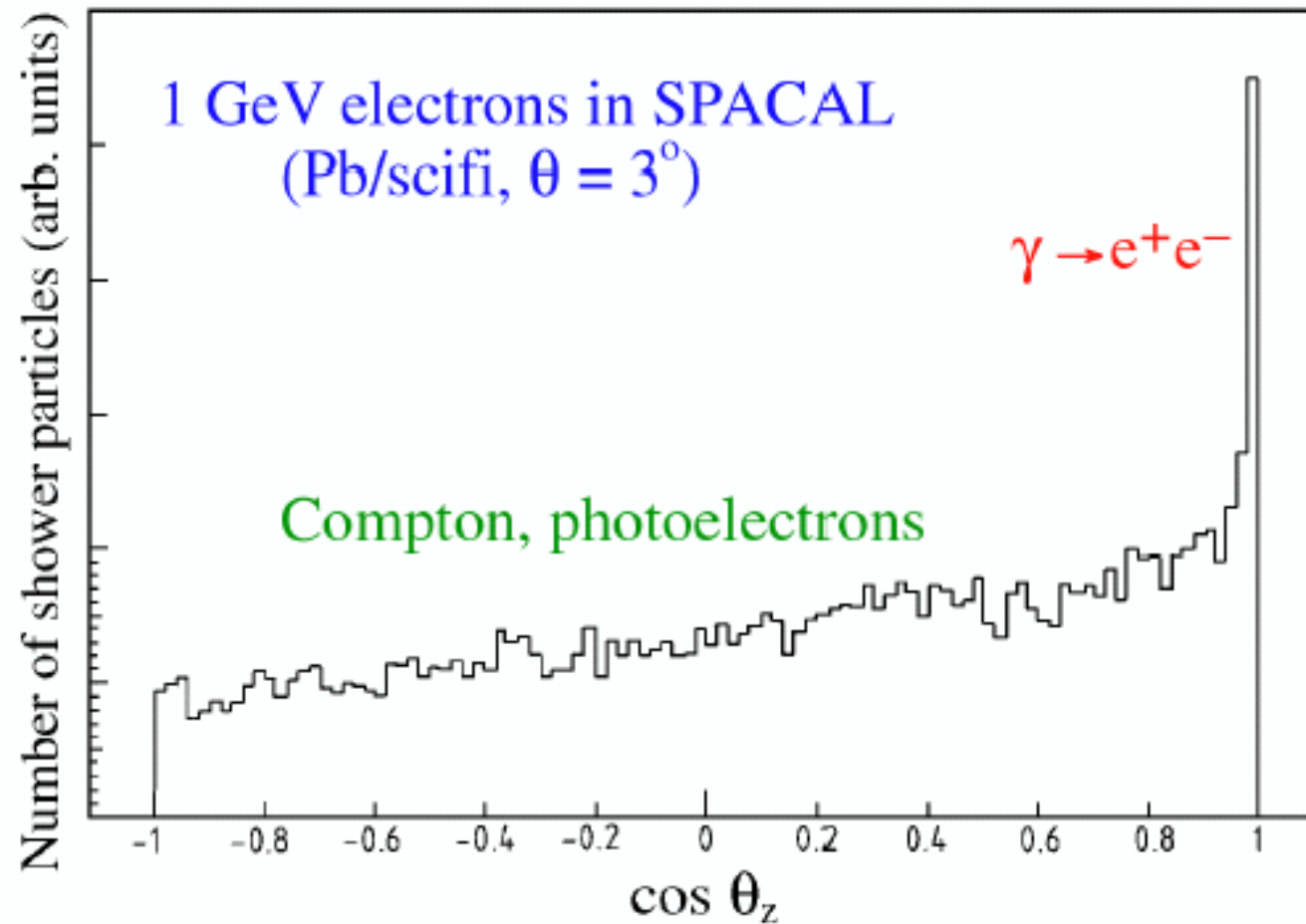


QED cross-section for  $\gamma$ -e scattering

$$\sigma_{\text{compton}} \sim Z \cdot \ln(E_\gamma) / E_\gamma$$

Process dominant at  $E_\gamma \approx 100 \text{ keV} - 5 \text{ GeV}$

# ANGULAR DISTRIBUTION



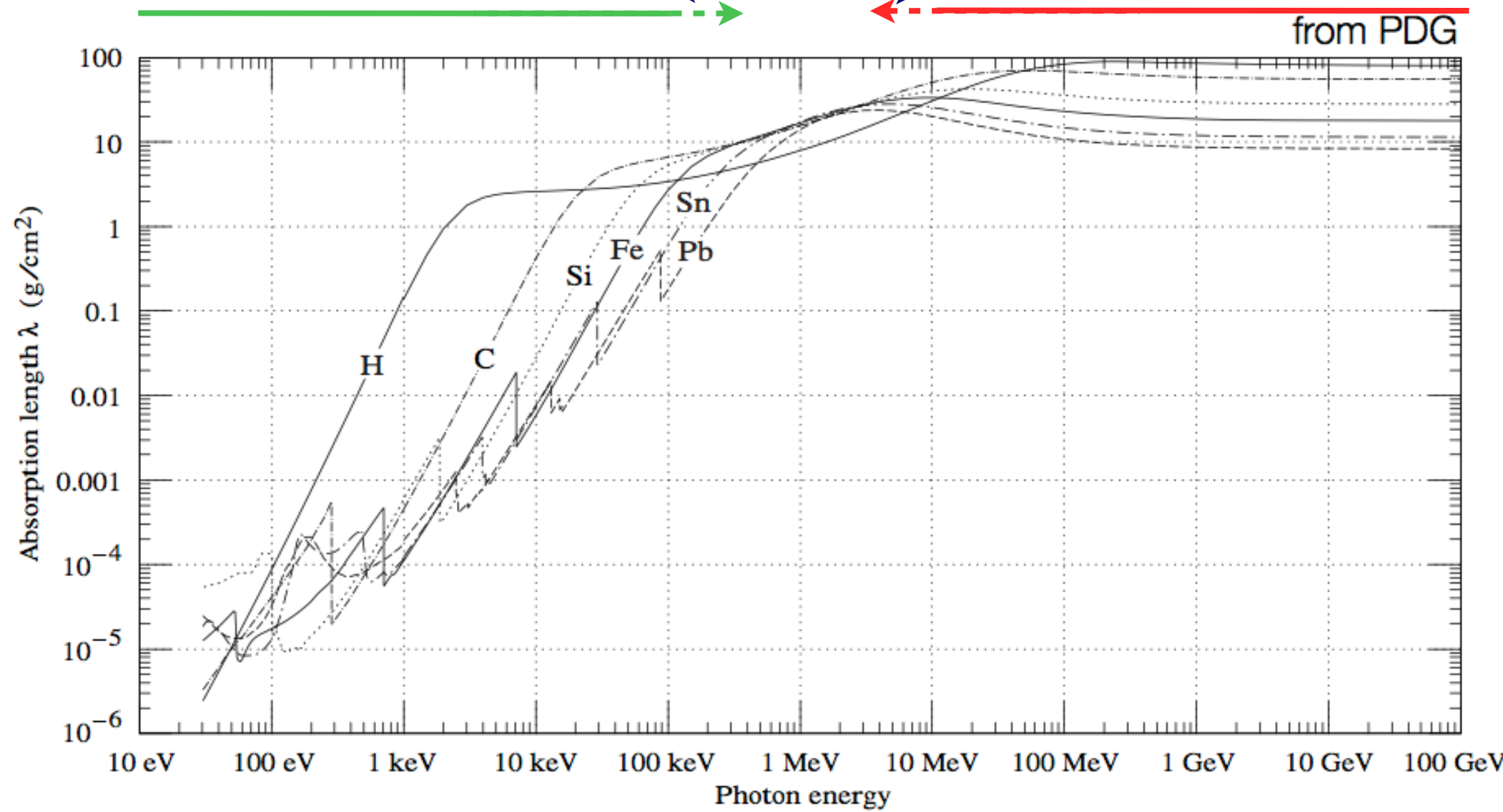
# INTERACTION OF PHOTONS WITH MATTER

Mass absorption coefficient  $\lambda = 1/(\mu/\rho)$  [g.cm<sup>2</sup>] with  $\mu = N_A \cdot \sigma/A$

$$\sigma_{Ph} \propto \frac{Z^5}{E^{3.5}}$$

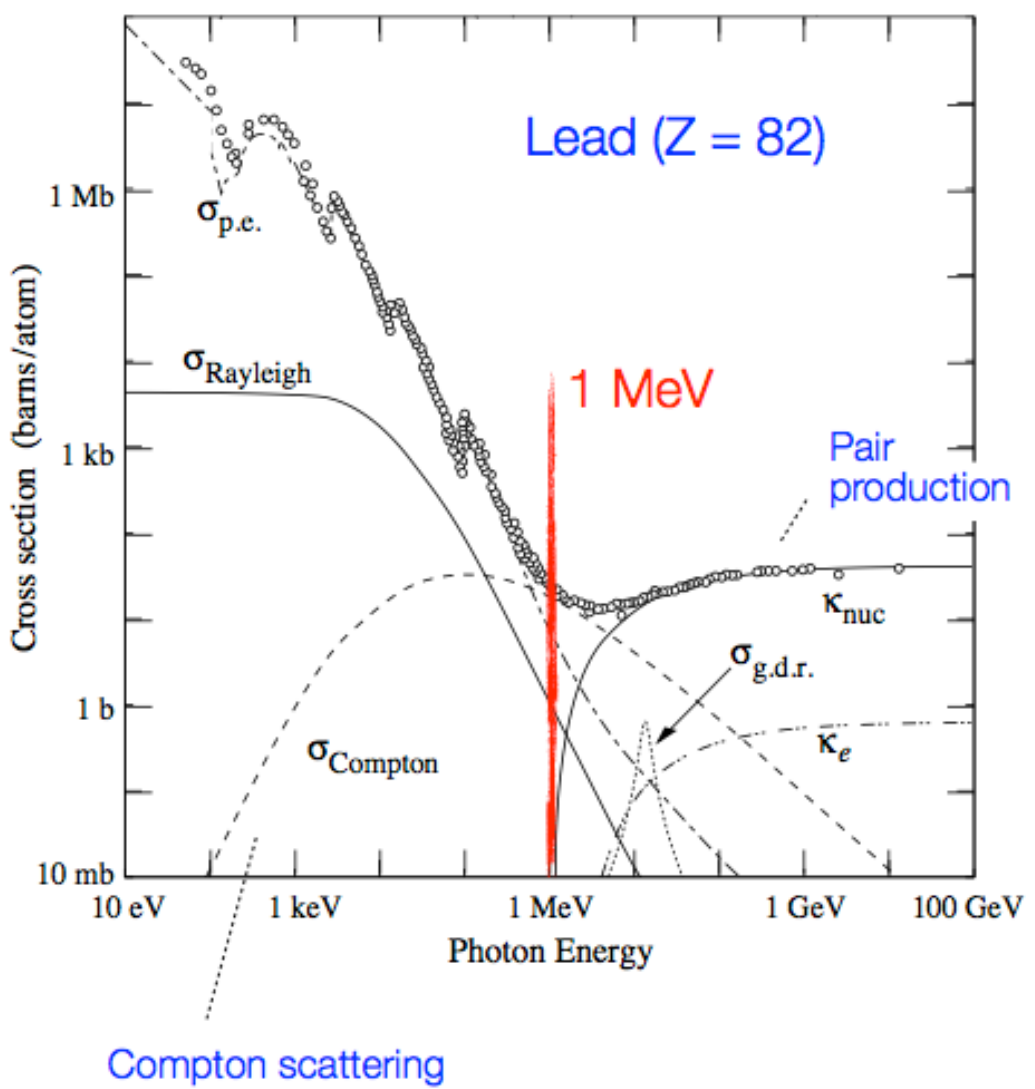
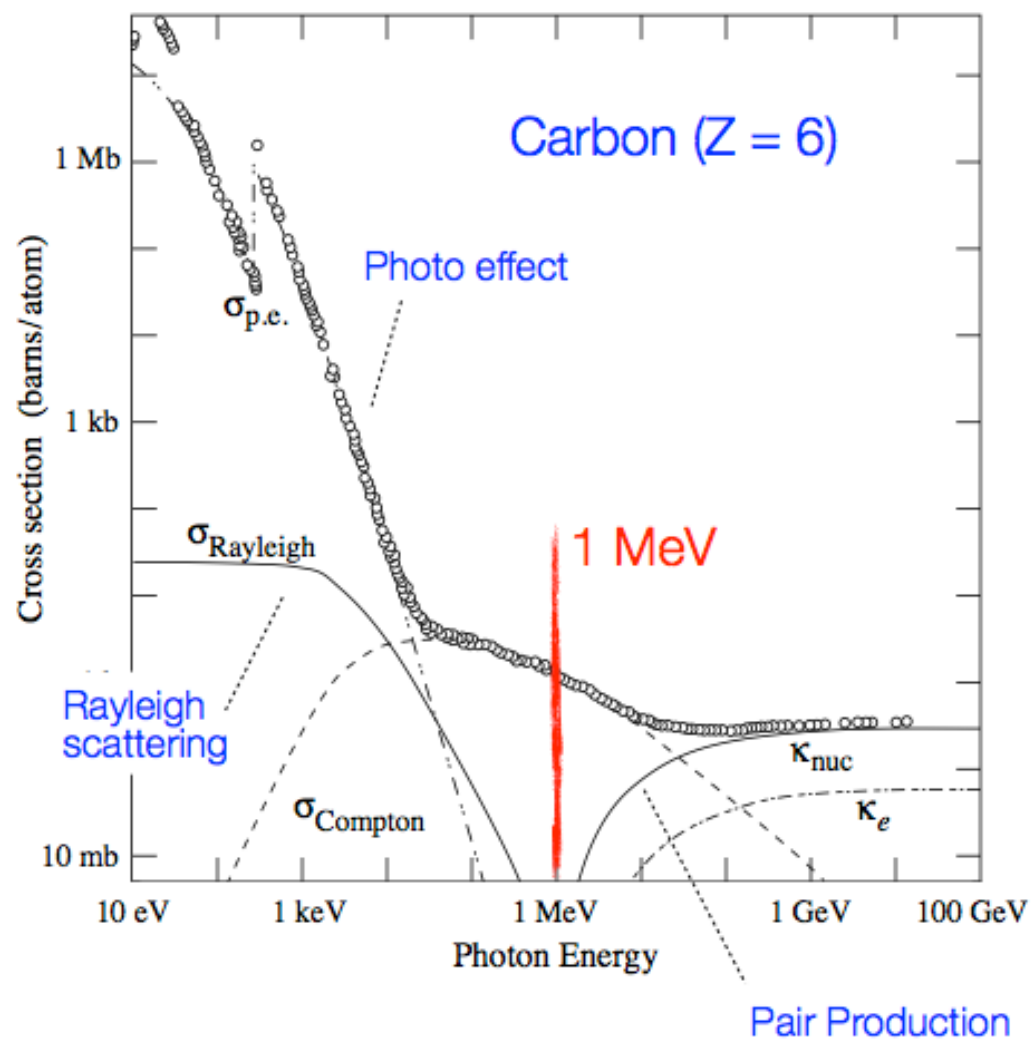
$$\sigma_{Compton} \propto \frac{\ln E}{E} \cdot Z$$

$$\sigma_{Pair} \propto Z^2$$

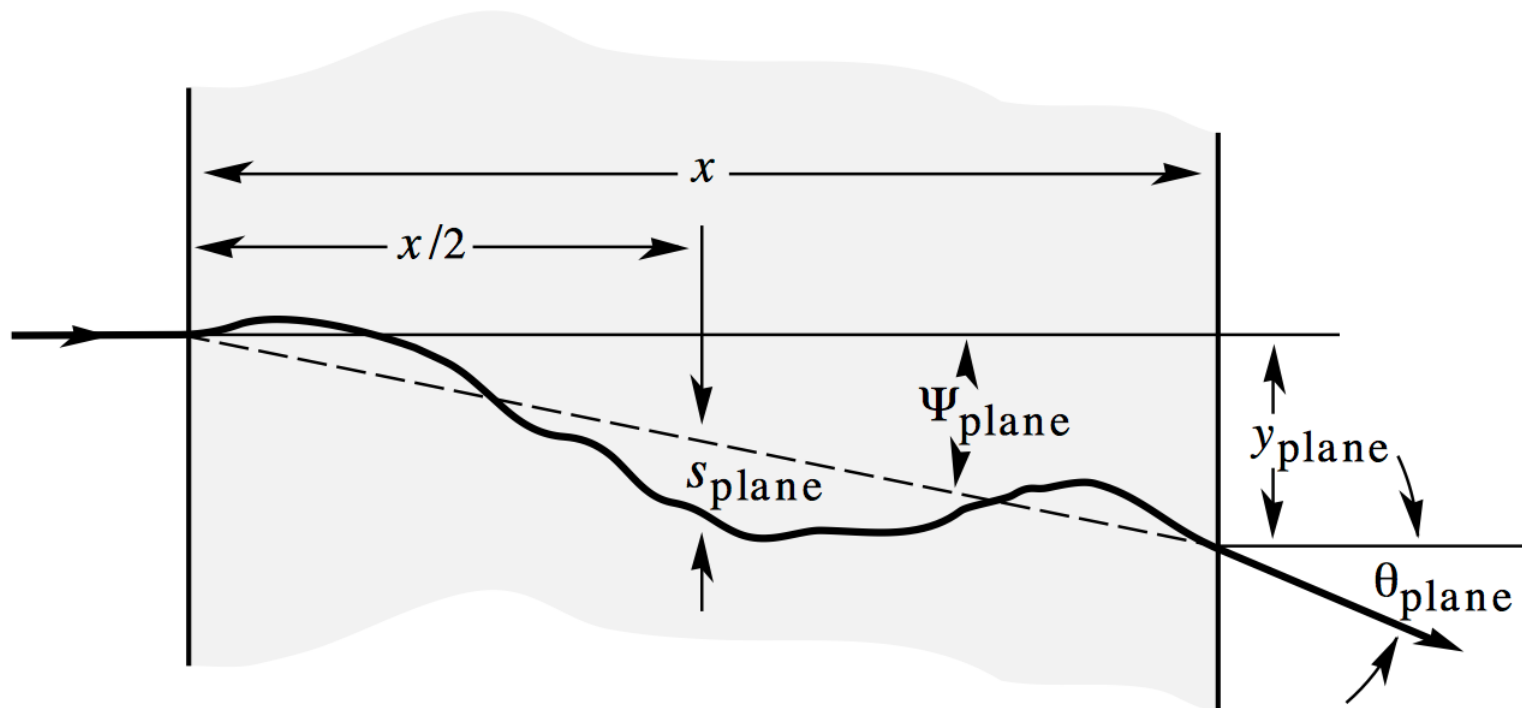


# INTERACTION OF PHOTONS WITH MATTER

## Photon Total Cross Sections



# MULTIPLE SCATTERING



Scattering of charged particles off the atoms in the medium causes a change of direction

The statistical sum of many such small angle scattering results in a gaussian angular distribution with a width given by

$$\theta_0 = \frac{13.6 \text{ MeV}}{\beta c p} z \sqrt{x/X_0} \left[ 1 + 0.038 \ln(x/X_0) \right]$$

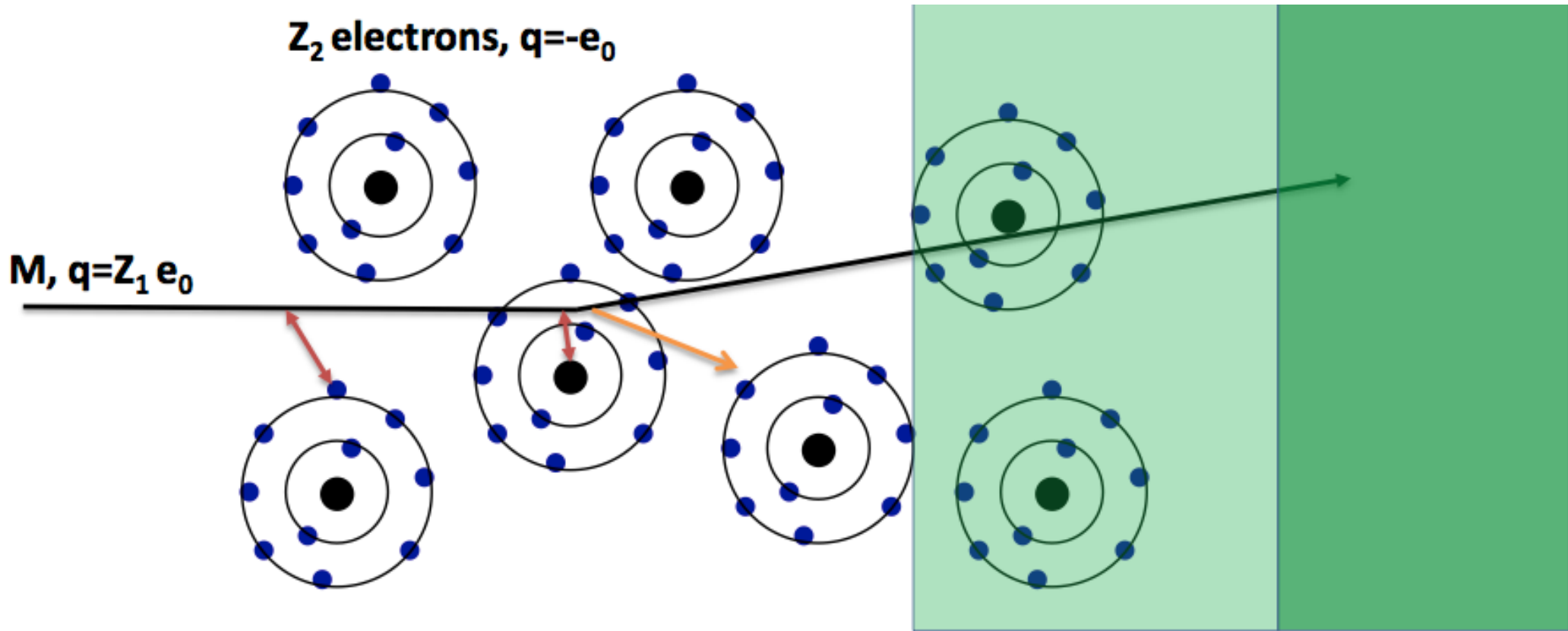
## Example

$p=1 \text{ GeV}$ ,  $x=300\mu\text{m}$ ,  $\text{Si } X_0=9.4 \text{ cm} \rightarrow \theta_0=0.8 \text{ mrad}$

For a distance of 10 cm this corresponds to  $80 \mu\text{m}$ , which is significantly larger than typical resolution of Si-strip detector.

# ELECTROMAGNETIC INTERACTION

## PARTICLE - MATTER



### Interaction with the atomic electrons.

The incoming particle loses energy and the atoms are **excited** or **ionised**.

### Interaction with the atomic nucleus.

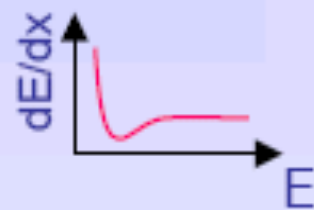
The incoming particle is deflected causing **multiple scattering** of the particle in the material.

During this scattering a **Bremsstrahlung photon** can be emitted

In case the particle's velocity is larger than the velocity of light in the medium, the resulting EM shockwave manifests itself as **Cherenkov radiation**. When the particle crosses the boundary between two media, there is a probability of 1% to produce an Xray photon called **Transition radiation**.

$e^+ / e^-$

■ Ionisation



■ Bremsstrahlung



$\gamma$

■ Photoelectric effect



■ Compton effect

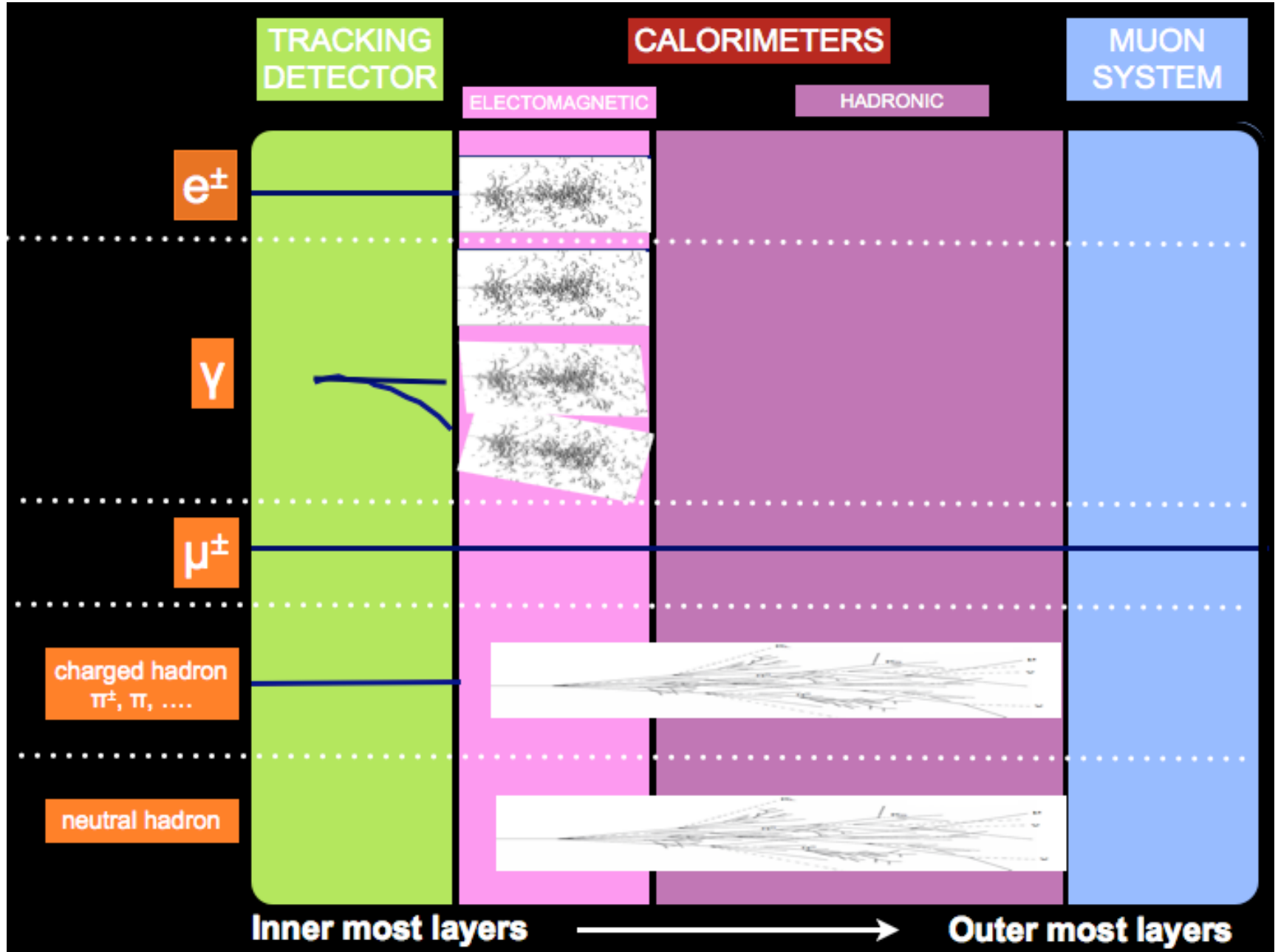


■ Pair production



# DETECTOR QUIZZ II : explain this schematic

INTERACTIONS → DETECTORS



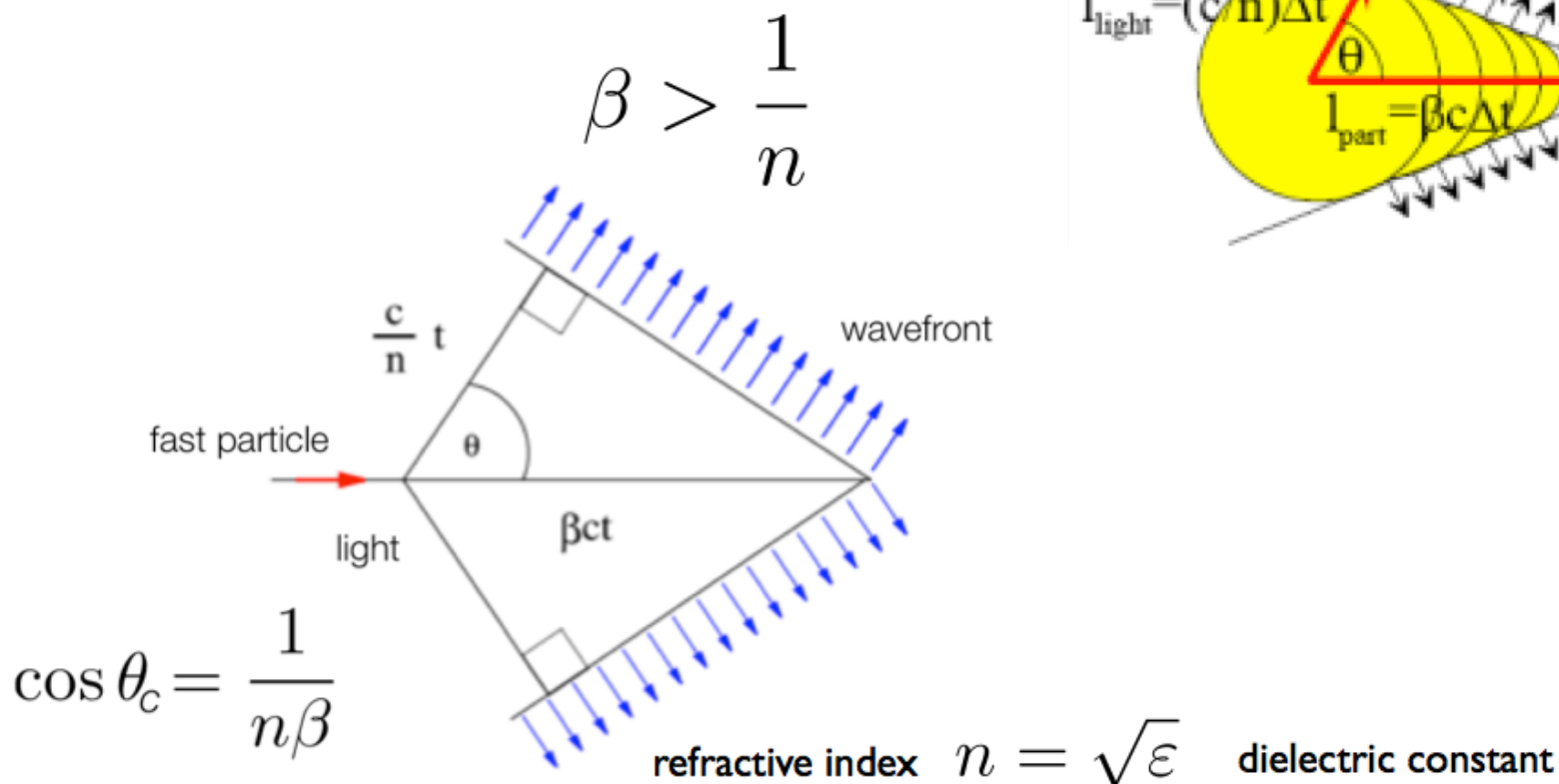




# CERENKOV RADIATION

Particles moving in a medium with **speed larger than speed of light in that medium** lose energy by emitting electromagnetic radiation

Charged particles polarise the medium generating an electrical dipole varying with time  
 Every point in the trajectory emits a spherical EM wave; waves constructively interfere



# CERENKOV RADIATION

## Parameters of Typical Radiator

Medium	n	$\beta_{thr}$	$\theta_{max} [\beta=1]$	$N_{ph} [eV^{-1} cm^{-1}]$
Air	1.000283	0.9997	1.36	0.208
Isobutan	1.00127	0.9987	2.89	0.941
Water	1.33	0.752	41.2	160.8
Quartz	1.46	0.685	46.7	196.4

Note: Energy loss by Cherenkov radiation very small w.r.t. ionization (< 1%).

### Example:

[Proton with  $E_{kin} = 1$  GeV passing through 1 cm water]

$$\beta = p/E \approx 0.875; \cos\theta_c = 1/n\beta = 0.859 \rightarrow \theta_c = 30.8^\circ$$

$$d^2N/(dEdx) = 370 \sin^2\theta_c eV^{-1} cm^{-1} \approx 100 eV^{-1} cm^{-1}$$

$$\rightarrow \Delta E_{loss} = \langle E \rangle \frac{d^2N}{dEdx} \Delta E \Delta x$$

$$= 2.5 eV \cdot 100 eV^{-1} cm^{-1} \cdot 5 eV \cdot 1 cm = 1.25 keV$$

Visible light only!

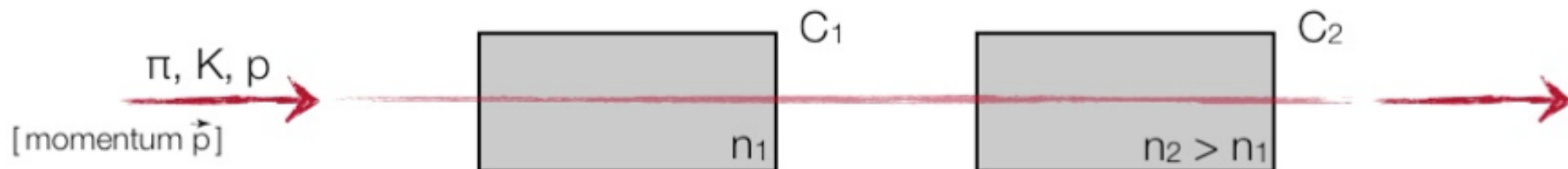
[ $E = 1 - 5 eV; \lambda = 300 - 600 nm$ ]

$$\Delta E_{loss} < 1.25 keV$$

# IDENTIFYING PARTICLES with CERENKOV RADIATION

Threshold detection:

Observation of Cherenkov radiation  $\rightarrow \beta > \beta_{\text{thr}}$



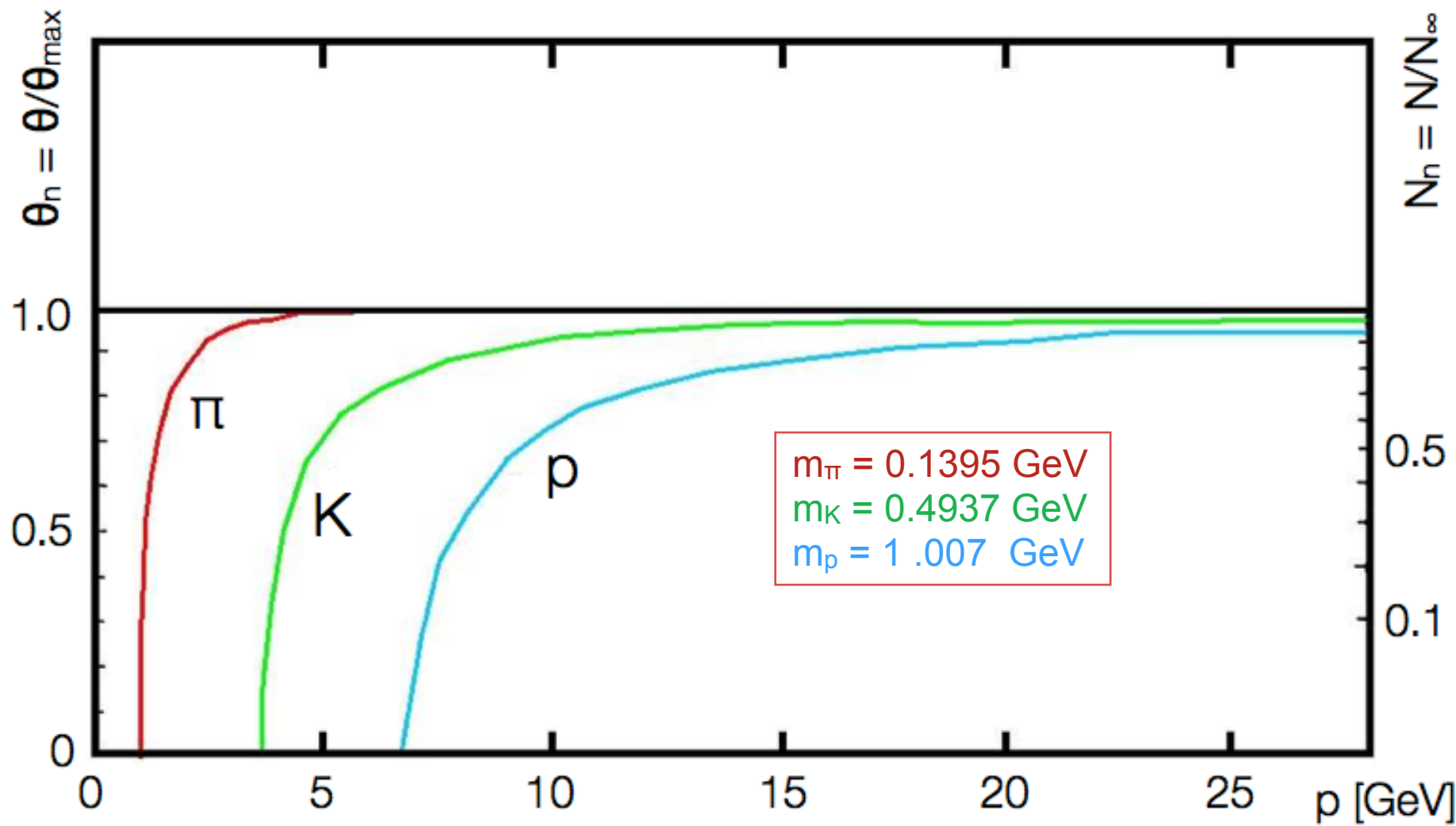
Choose  $n_1, n_2$  in such a way that for:

$$n_2 : \quad \beta_{\pi}, \beta_K > 1/n_2 \text{ and } \beta_p < 1/n_2$$

$$n_1 : \quad \beta_{\pi} > 1/n_1 \text{ and } \beta_K, \beta_p < 1/n_1$$

- |                                  |               |                   |
|----------------------------------|---------------|-------------------|
| Light in $C_1$ and $C_2$         | $\rightarrow$ | identified pion   |
| Light in $C_2$ and not in $C_1$  | $\rightarrow$ | identified kaon   |
| Light neither in $C_1$ and $C_2$ | $\rightarrow$ | identified proton |

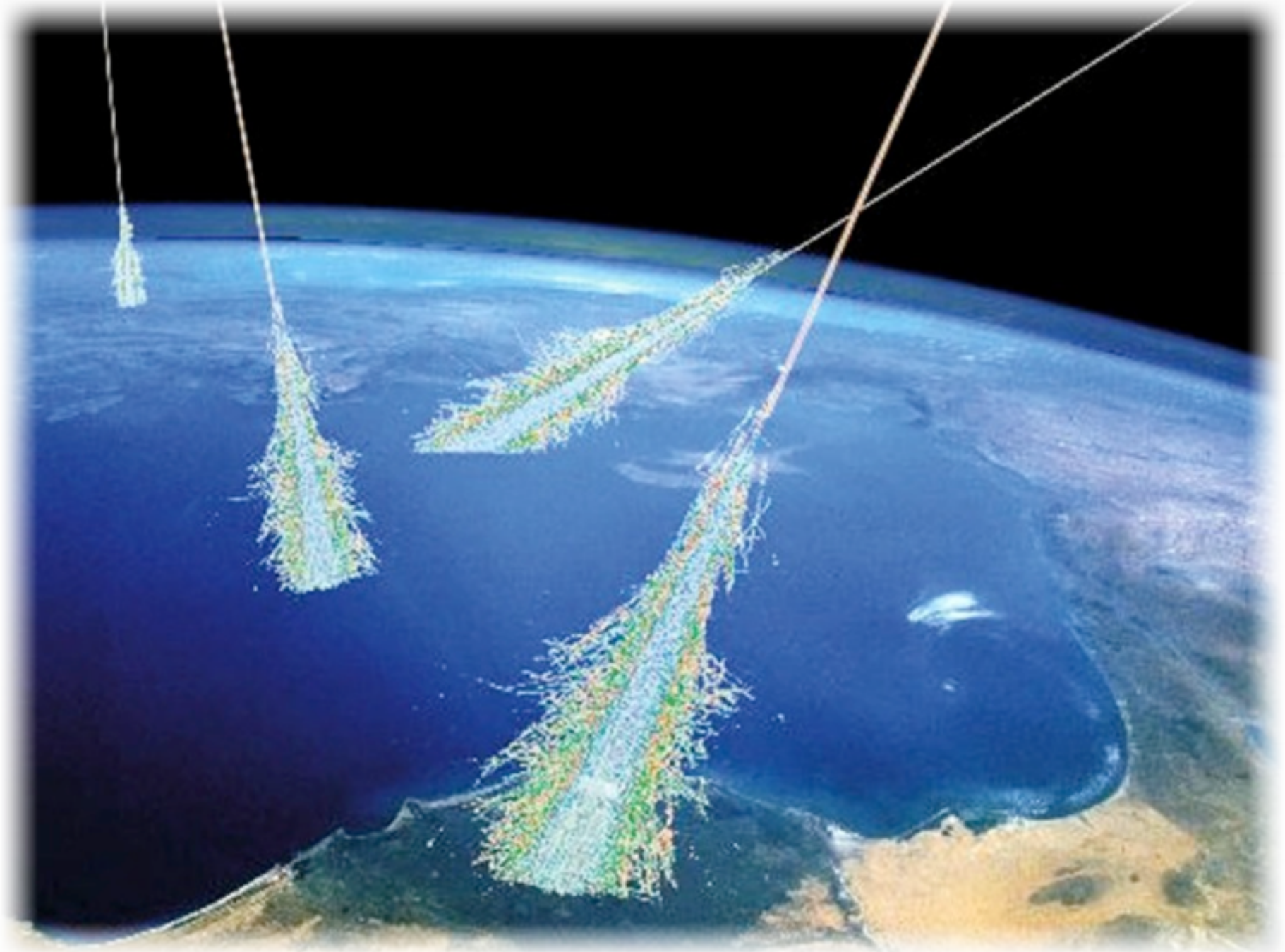
# CERENKOV RADIATION: MOMENTUM DEPENDENCE



Cherenkov angle  
 Number of photons

grows with  $\beta$  and reaches asymptotic value for  $\beta = 1$   
 $[\theta_{\max} = \arccos(1/n); N_{\infty} = x \cdot 370/\text{cm} (1 - 1/n^2)]$

# COSMIC RAYS

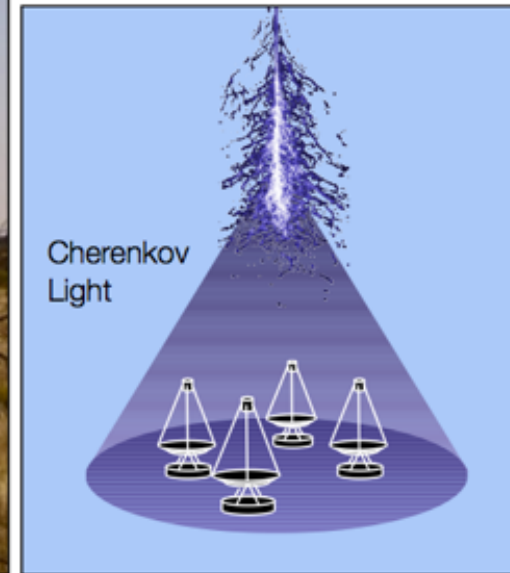


# HESS EXPERIMENT



Hess Telescopes  
Namibia

$\gamma$ -ray detection



# Transition radiation

Transition radiation occurs if a relativistic particle (**large  $\gamma$** ) passes the **boundaries between two media with different refraction indices.**

Intensity of radiation is **logarithmically proportional to  $\gamma$**

Angular distribution strongly forward peaked  
[Interference; coherence condition]

$$\theta \leq 1/\gamma$$

Coherent radiation is generated only over a very small formation length

$$D = \gamma c / \omega_p$$

Plasma frequency  
[from Drude model]

Volume element from which coherent radiation is emitted ...

$$V = \pi D \rho_{\max}^2$$

$\rho_{\max} = \gamma v / \omega$   
[transversal range ...  
... with large polarization]

Maximum energy of radiated photons limited by plasma frequency ...  
[results from requiring  $V \neq 0 \rightarrow \omega = \gamma \omega_p$ ]

$$E_{\max} = \gamma \hbar \omega_p$$

[X-Rays  $\rightarrow$  large  $\gamma$ !!]

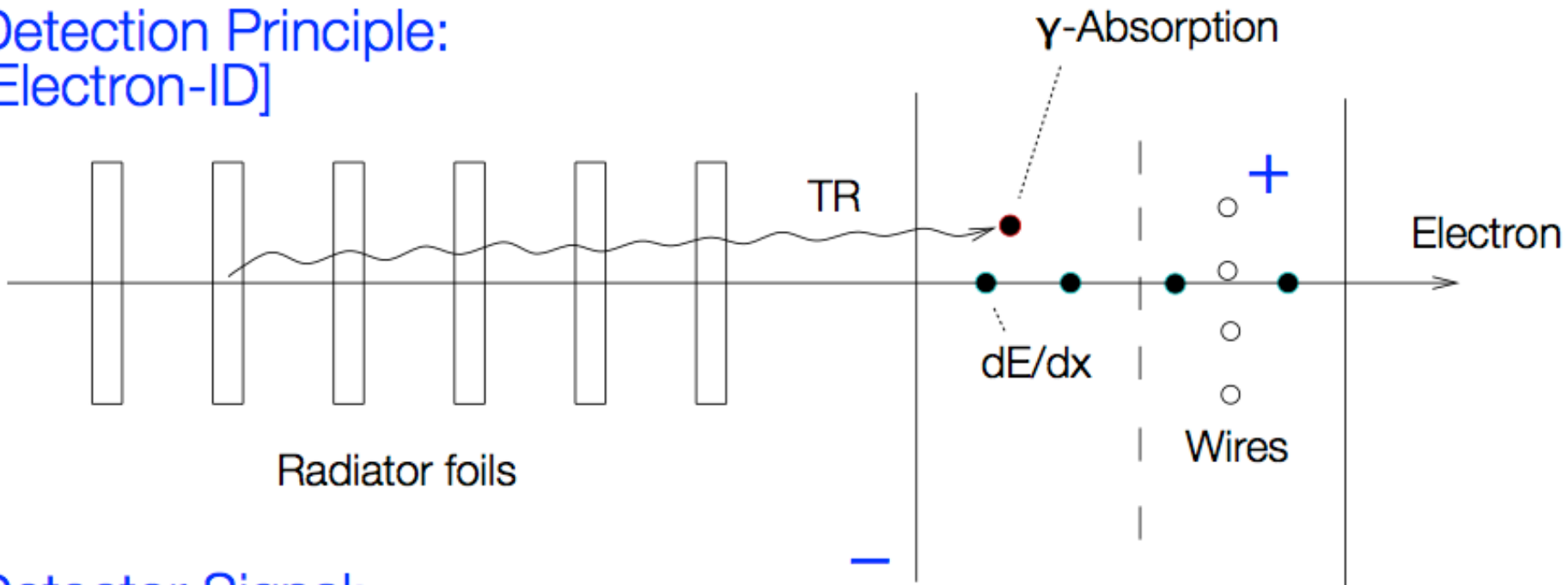
Typical values:      CH<sub>2</sub>:     $\hbar \omega_p = 20 \text{ eV}; \gamma = 10^3$  ]  
                             [ Air:     $\hbar \omega_p = 0.7 \text{ eV}$  ]

$D = 10 \mu\text{m}$   
[d > D: absorption dominates]

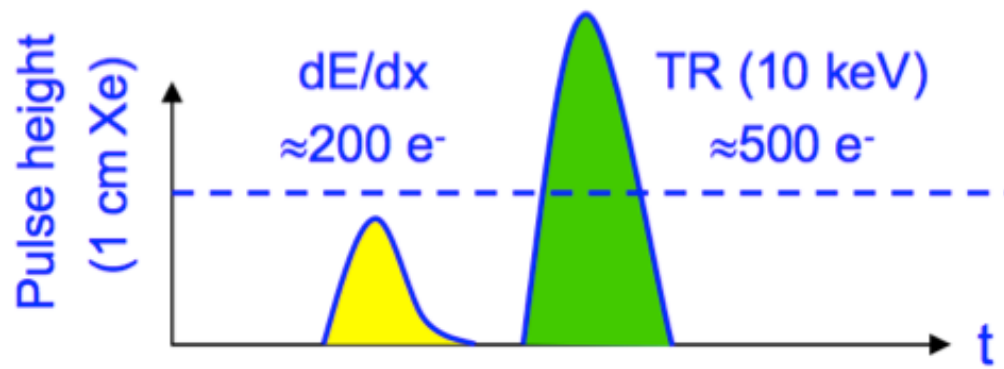


# IDENTIFYING PARTICLES WITH TRANSITION RADIATION

Detection Principle:  
[Electron-ID]



Detector Signal:



- Detector should be sensitive to  $3 \leq E_\gamma \leq 30$  keV.
  - ✓ Gaseous detectors
- In gas  $\sigma_{\text{photo effect}} \propto Z^5$
- Gases with high Z are required
  - ✓ e.g. Xenon (Z=54)

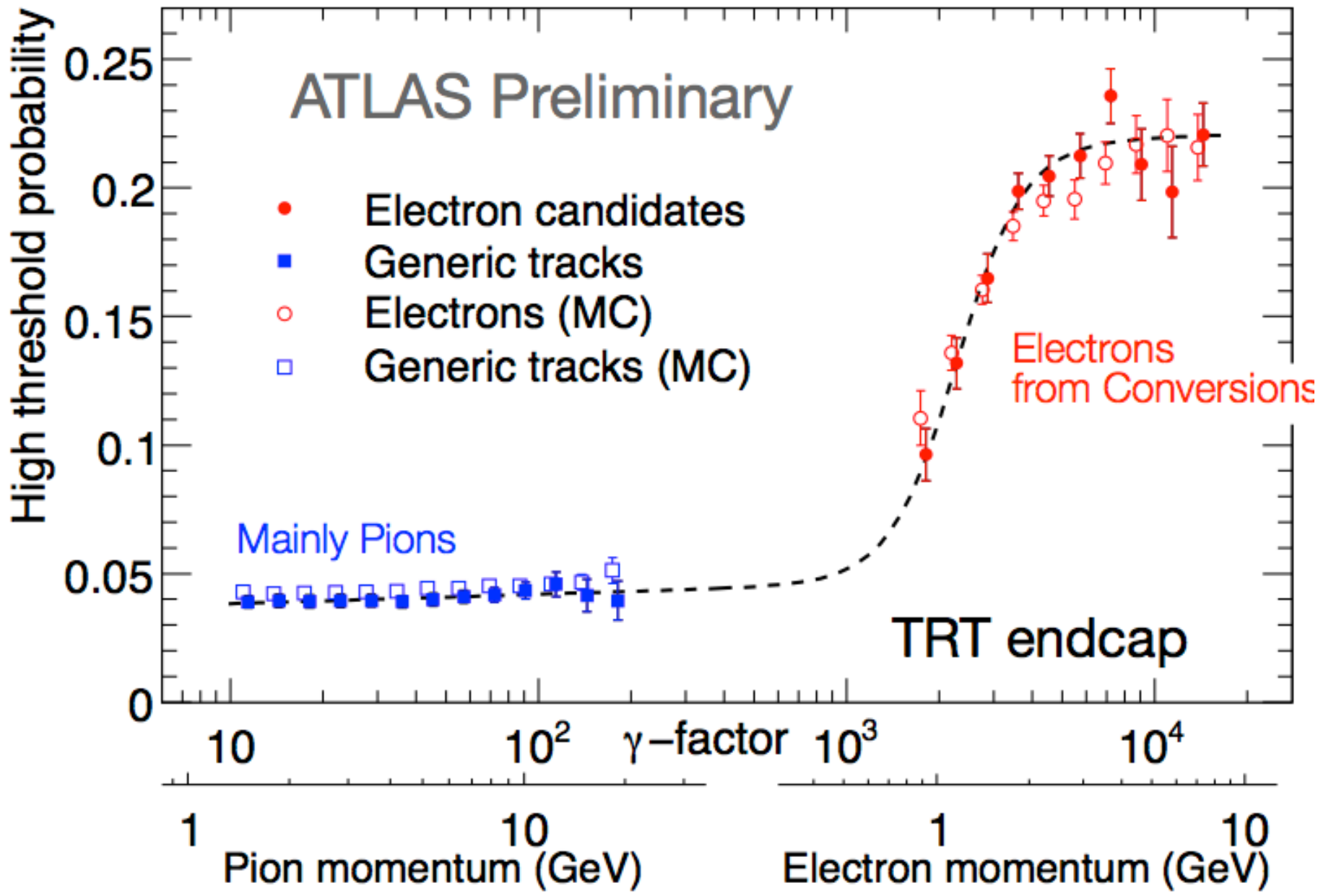
# ATLAS TRANSITION RADIATION TRACKER

Straw Tube Tracker  
with interspace filled with foam

→ Tracking & transition radiation



# IDENTIFYING PARTICLES WITH TRANSITION RADIATION



# CREDIT and BIBLIOGRAPHY

A lot of material in these lectures are from:

Daniel Fournier @ EDIT2011

Marco Delmastro @ ESIPAP 2014

Weiner Raigler @ AEPSHEP2013

Hans Christian Schultz-Coulon's lectures

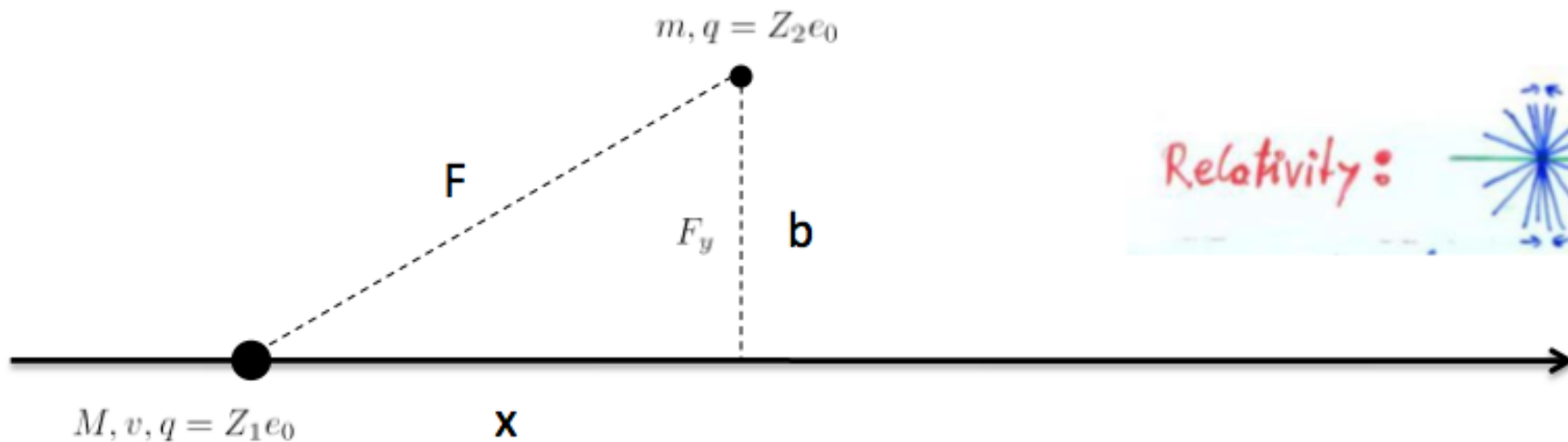
Carsten Niebuhr's lectures [1][2][3]

Georg Streinbrueck's lecture

Pippa Wells @ EDIT2011

Jérôme Baudot @ ESIPAP2014

# IONISATION & EXCITATION



While the charged particle is passing another charged particle the Coulomb force is acting, resulting in momentum transfer.

The relativistic form of the transverse electric field does not change the momentum transfer. The transverse field is stronger, but the time of action is shorter.

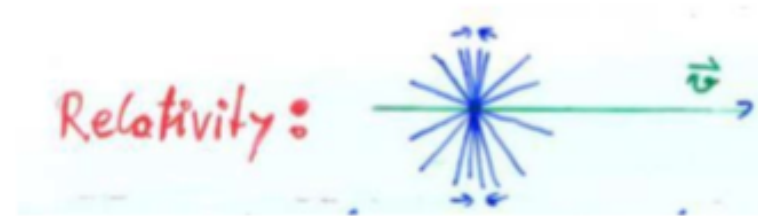
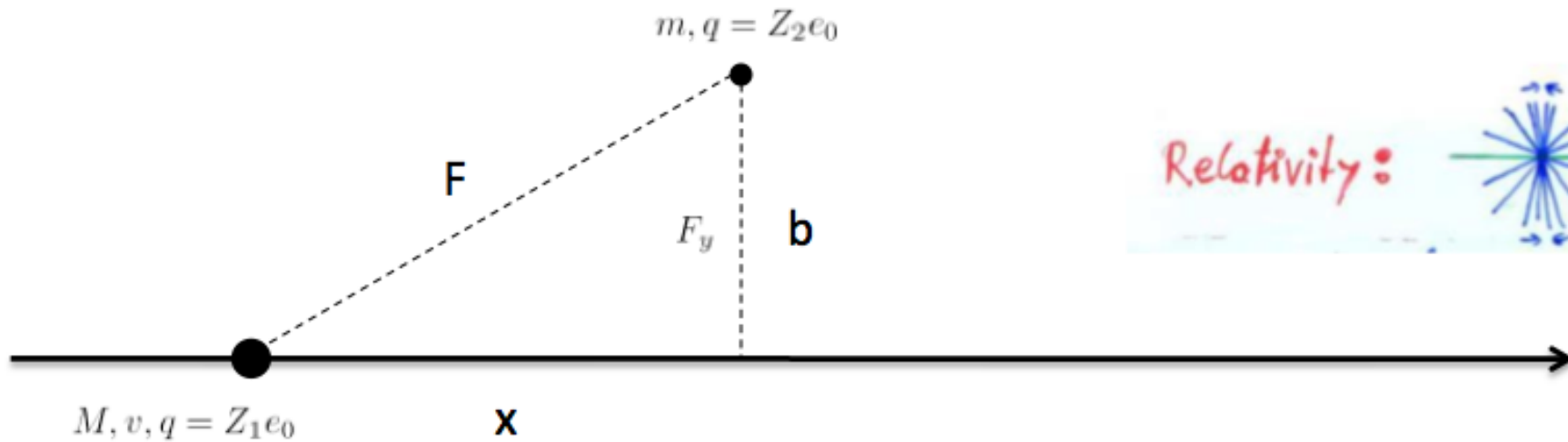
$$F_y = \frac{Z_1 Z_2 e_0^2}{4\pi\epsilon_0(b^2 + v^2 t^2)} \frac{b}{\sqrt{b^2 + v^2 t^2}}$$

$$\Delta p = \int_{-\infty}^{\infty} F_y(t) dt = \frac{2Z_1 Z_2 e_0^2}{4\pi\epsilon_0 v b}$$

$$F_y = \frac{\gamma Z_1 Z_2 e_0^2 b}{4\pi\epsilon_0(b^2 + \gamma^2 v^2 t^2)^{3/2}}$$

$$\Delta p = \int_{-\infty}^{\infty} F_y(t) dt = \frac{2Z_1 Z_2 e_0^2}{4\pi\epsilon_0 v b}$$

# IONISATION & EXCITATION



The transferred energy

The incoming particle transfers energy mainly/only to the atomic electrons.

$$\Delta E = \frac{(\Delta p)^2}{2m} = \frac{Z_2^2}{m} \frac{2Z_1^2 e_0^4}{(4\pi\epsilon_0)^2 v^2 b^2}$$

$$\Delta E(\text{electrons}) = Z_2 \frac{1}{m_e} \frac{2Z_1^2 e_0^4}{(4\pi\epsilon_0)^2 v^2 b^2}$$

$$\Delta E(\text{nucleus}) = \frac{Z_2^2}{2Z_2 m_p} \frac{2Z_1^2 e_0^4}{(4\pi\epsilon_0)^2 v^2 b^2}$$

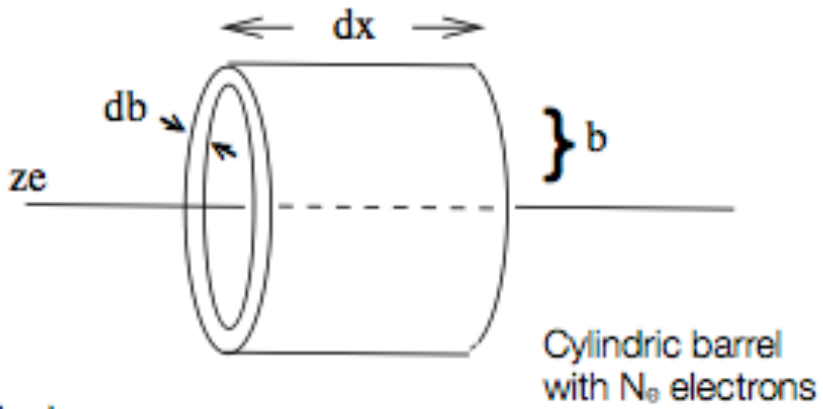
$$\frac{\Delta E(\text{electrons})}{\Delta E(\text{nucleus})} = \frac{2m_p}{m_e} \approx 4000$$

# BETHE-BLOCH FORMULA - CLASSICAL DERIVATION

Bohr 1913

Energy transfer onto **single** electron for **impact parameter b**:

$$\Delta E(b) = \frac{\Delta p^2}{2m_e}$$



Consider cylindric barrel  $\rightarrow N_e = n \cdot (2\pi b) \cdot db dx$

Energy loss per path length  $dx$  for distance between  $b$  and  $b+db$  in medium with electron density  $n$ :

Energy loss!

$$-dE(b) = \frac{\Delta p^2}{2m_e} \cdot 2\pi n b db dx = \frac{4z^2 e^4}{2b^2 v^2 m_e} \cdot 2\pi n b db dx = \frac{4\pi n z^2 e^4}{m_e v^2} \frac{db}{b} dx$$

Diverges for  $b \rightarrow 0$ ; integration only for relevant range  $[b_{min}, b_{max}]$ :

Bohr 1913

$$-\frac{dE}{dx} = \frac{4\pi n z^2 e^4}{m_e v^2} \cdot \int_{b_{min}}^{b_{max}} \frac{db}{b} = \frac{4\pi n z^2 e^4}{m_e v^2} \ln \frac{b_{max}}{b_{min}}$$

# BETHE-BLOCH FORMULA - CLASSICAL DERIVATION

Determination of relevant range  $[b_{\min}, b_{\max}]$ :

Bohr 1913

[Arguments:  $b_{\min} > \lambda_e$ , i.e. de Broglie wavelength;  $b_{\max} < \infty$  due to screening ...]

$$b_{\min} = \lambda_e = \frac{h}{p} = \frac{2\pi\hbar}{\gamma m_e v}$$

Use Heisenberg uncertainty principle or that electron is located within de Broglie wavelength ...

$$b_{\max} = \frac{\gamma v}{\langle \nu_e \rangle}; \quad \left[ \gamma = \frac{1}{\sqrt{1-\beta^2}} \right]$$

Interaction time ( $b/v$ ) must be much shorter than period of the electron ( $\gamma/v_e$ ) to guarantee relevant energy transfer ...

[adiabatic invariance]

$$-\frac{dE}{dx} = \frac{4\pi z^2 e^4}{m_e c^2 \beta^2} n \cdot \ln \frac{m_e c^2 \beta^2 \gamma^2}{2\pi\hbar \langle \nu_e \rangle}$$

Deviates by factor 2 from QM derivation

Electron density:  $n = N_A \cdot \rho \cdot Z/A$  !!  
 Effective ionization potential:  $I \sim h \langle \nu_e \rangle$