

On Light Resonance Interpretations of the B Decay Anomalies

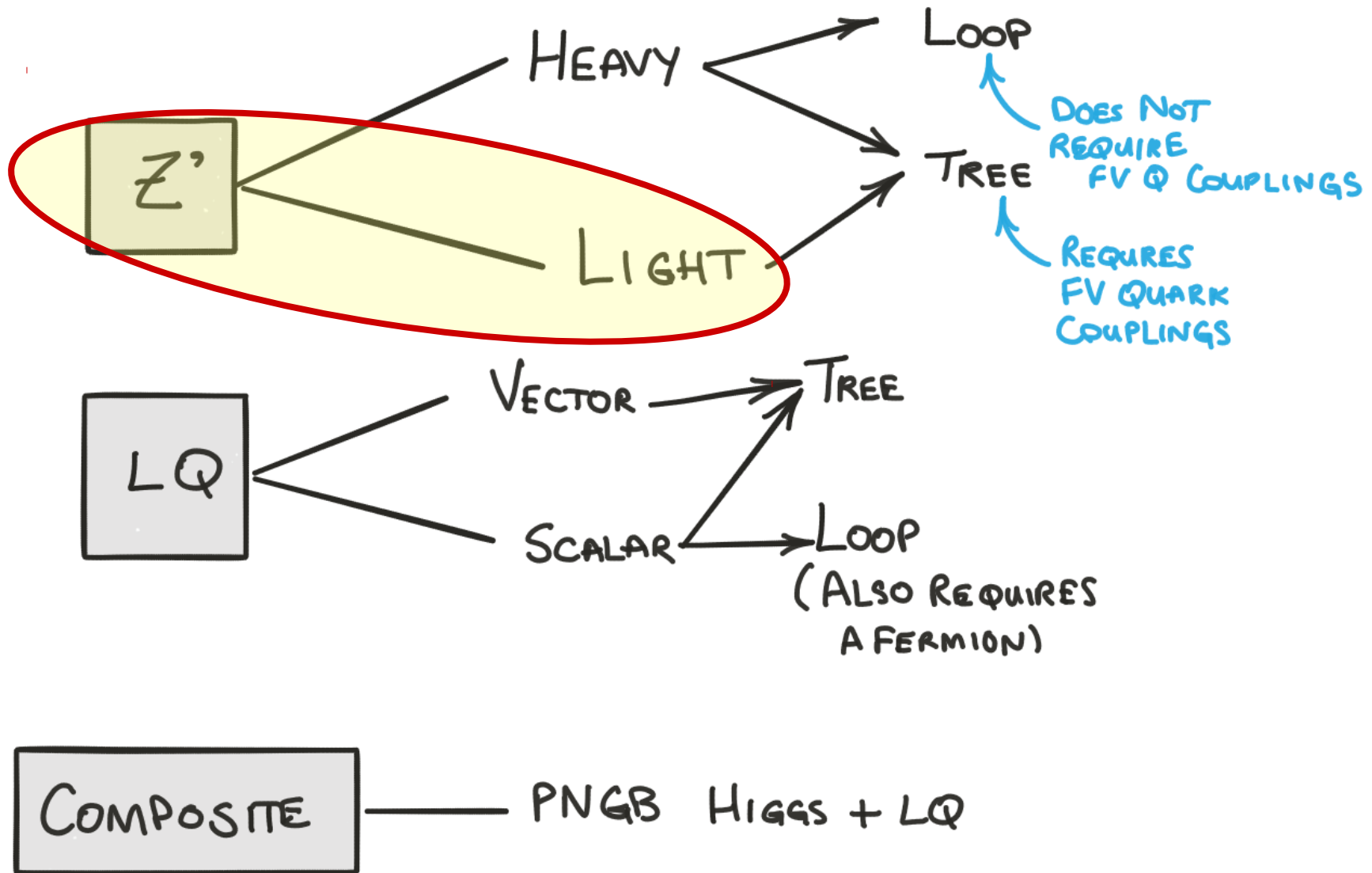
Fady Bishara



19 May 2017
CERN – Instant Workshop on B meson anomalies

Based on: 1705.03465 FB, U. Haisch, P. Monni

Focus of the talk



In particular, a new constraint from LHC

The light NP ansatz

The possibility that a light resonance could be responsible for the anomaly in P'_5 was mentioned by Amarjit Soni at 50th Rencontres de Moriond EW 2015, and subsequently re-emphasised to one of the authors by Brian Batell in a private conversation.

[Uli Haisch]

- Fuyuto, Hou, Kohda: [1512.09026]
- Sala, Straub: [1704.06188]
- Ghosh: [1704.06240]
- Alok, Bhattacharya, Datta, Kumar, Kumar, London: [1704.07397]
- Datta, Liao, Marfatia: [1702.01099]

A simplified model

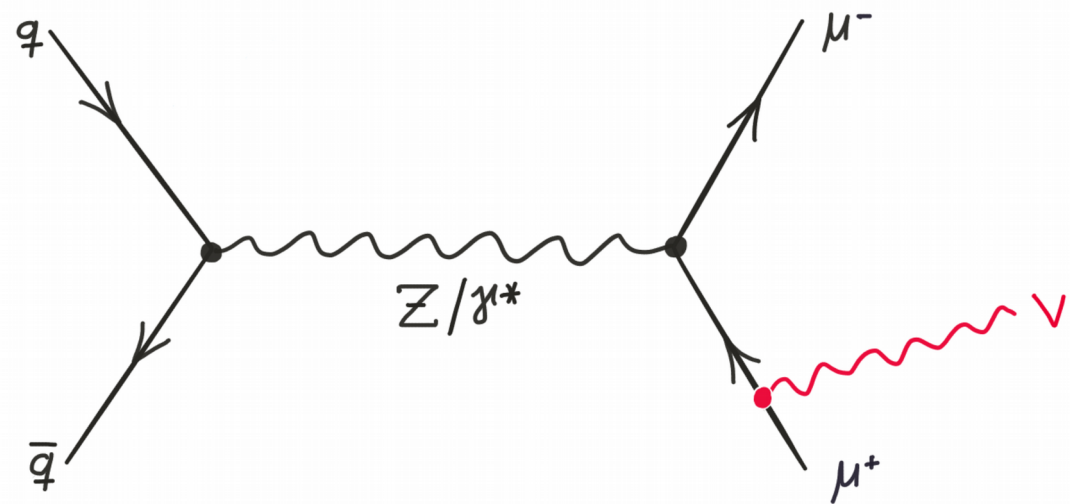
Sala, Straub: [1704.06188]

$$\mathcal{L} \supset (g_L^{sb} \bar{s}_L \not{V} b_L + \text{h.c.}) + \bar{\mu} (g_V^\mu - g_A^\mu \gamma_5) \not{V} \mu + g_V^\chi \bar{\chi} \not{V} \chi$$

$$g_\chi \gg g_V^\mu, g_A^\mu$$

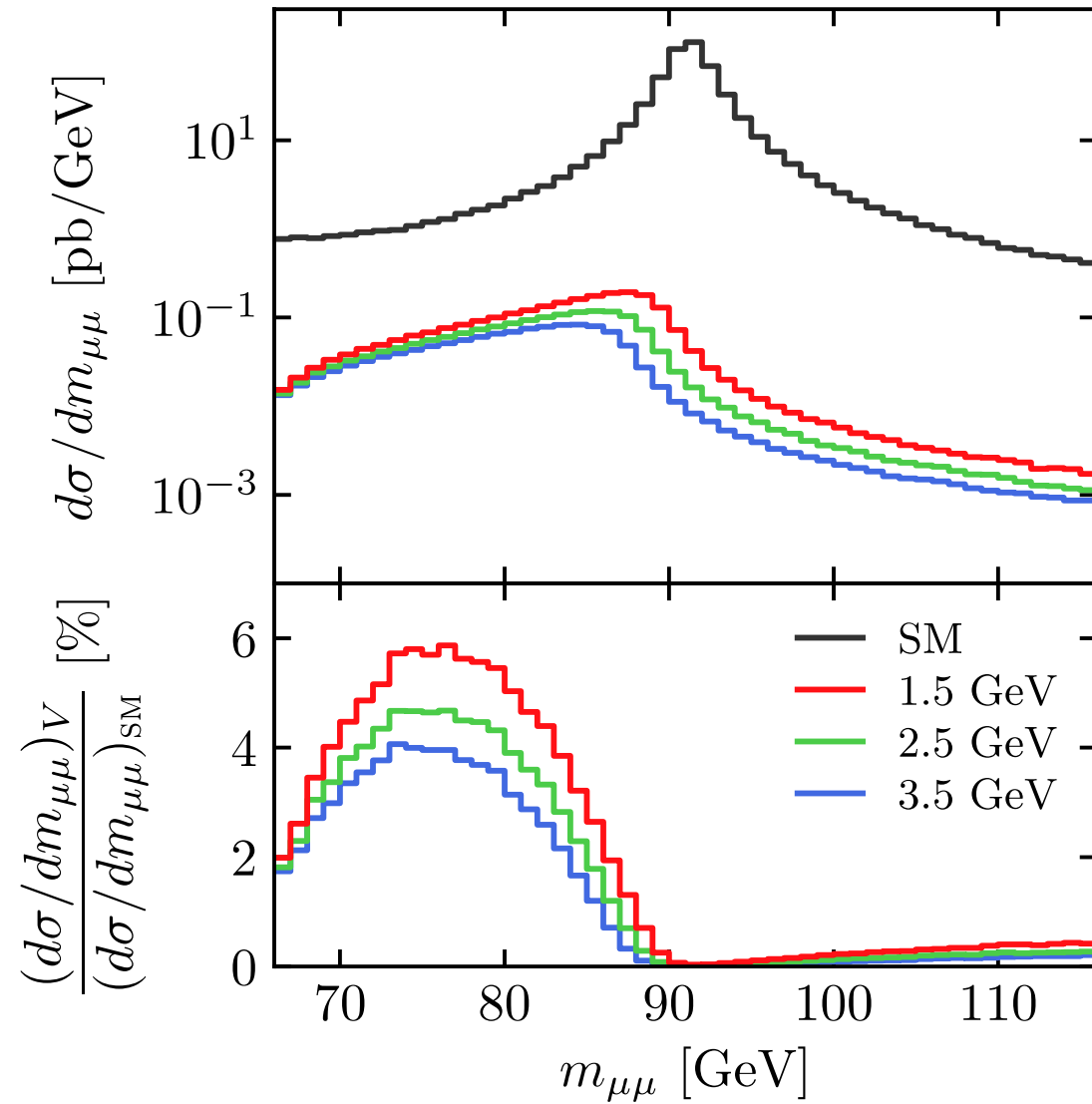
$$\frac{\Gamma_V}{m_V} \approx \frac{g_V^{\chi 2}}{12\pi}$$

$$\text{Br}(V \rightarrow \mu\mu) \approx \frac{g_V^{\mu 2} + g_A^{\mu 2}}{g_V^{\chi 2}}$$

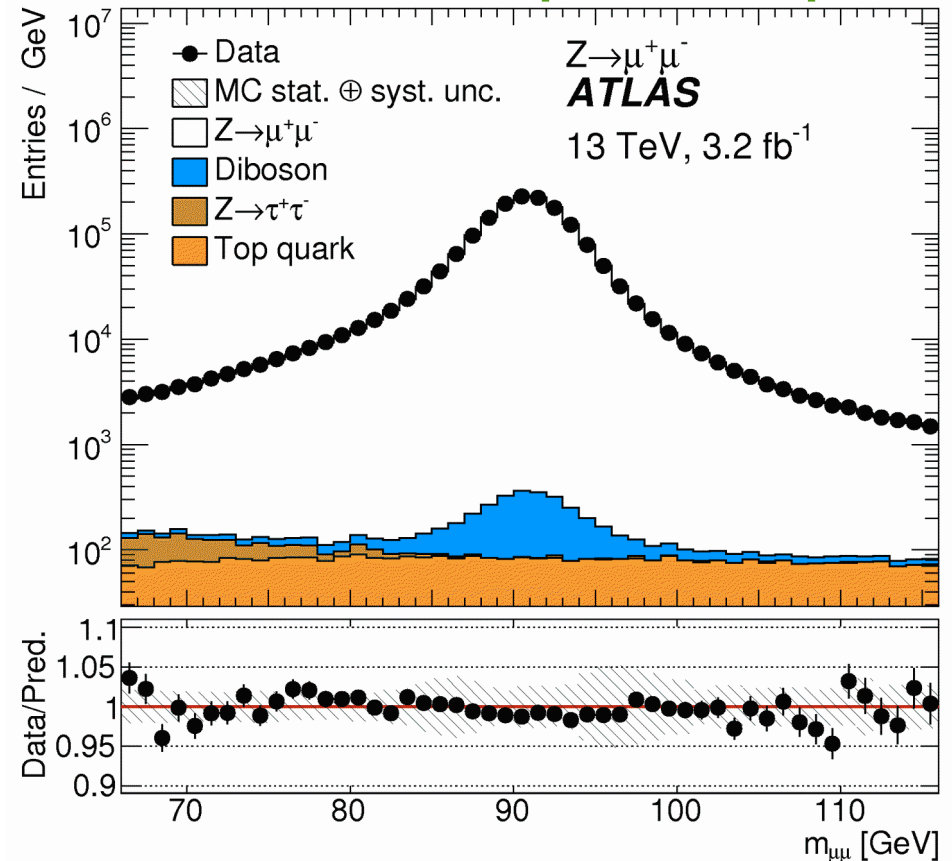


For $\Gamma_V/m_V \sim 20\%$, $g_V^\chi \sim 3$, and $\text{Br}(V \rightarrow \mu\mu) \sim 10^{-3}$

Z lineshape



ATLAS: [1612.03636]



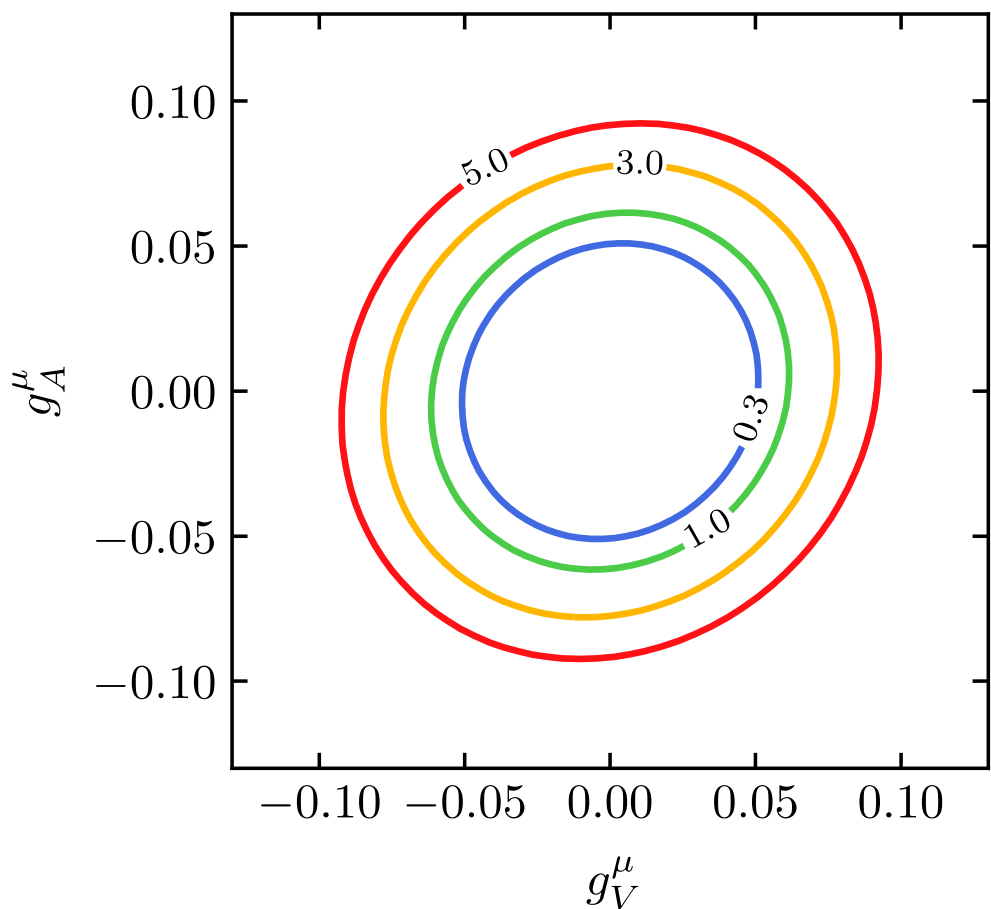
Fiducial phase space

$$p_T^\mu > 25 \text{ GeV}$$

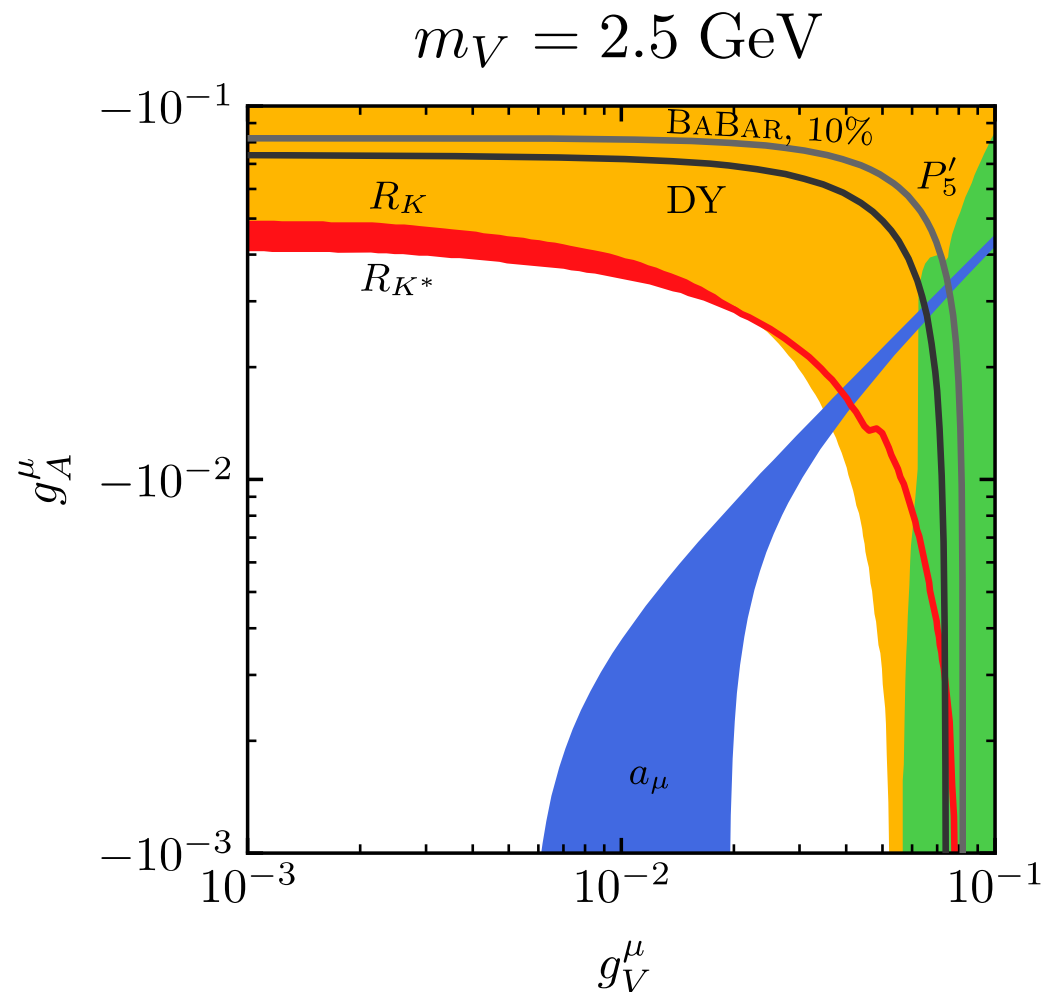
$$|\eta_\mu| < 2.5$$

$$66 < m_{\mu\mu} < 116 \text{ GeV}$$

Derived constraint



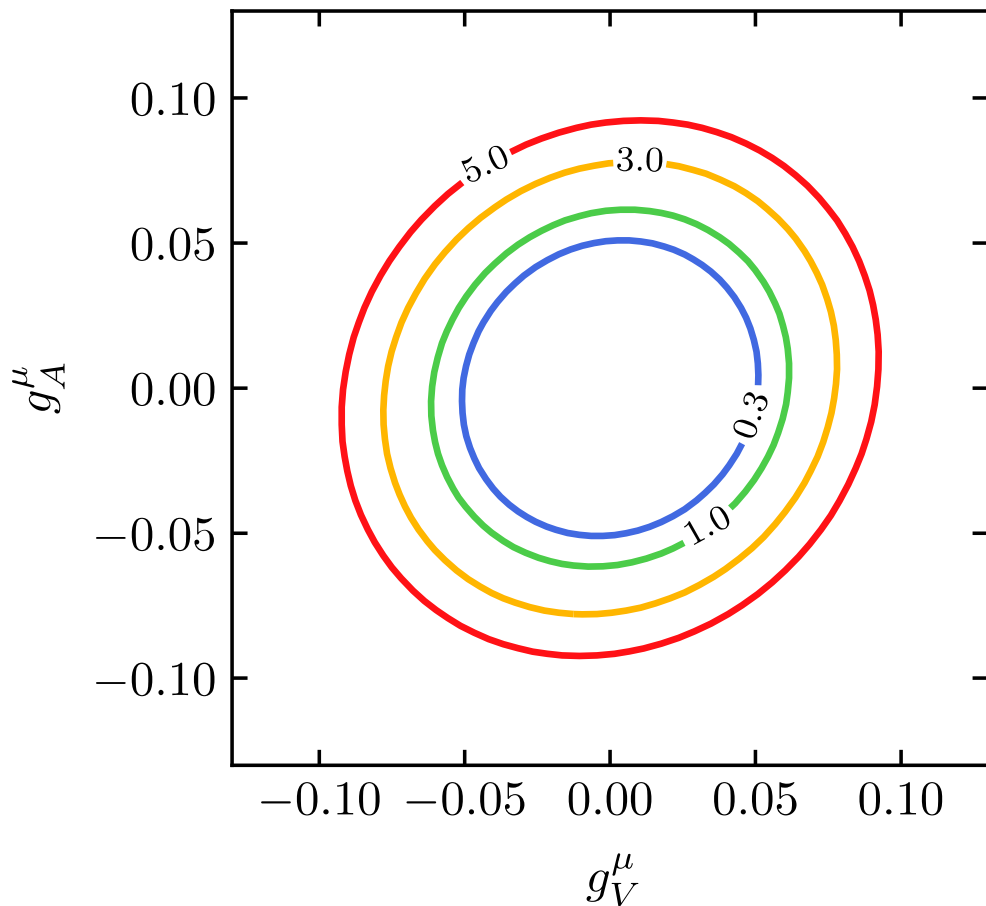
95% CL contours for various masses in GeV



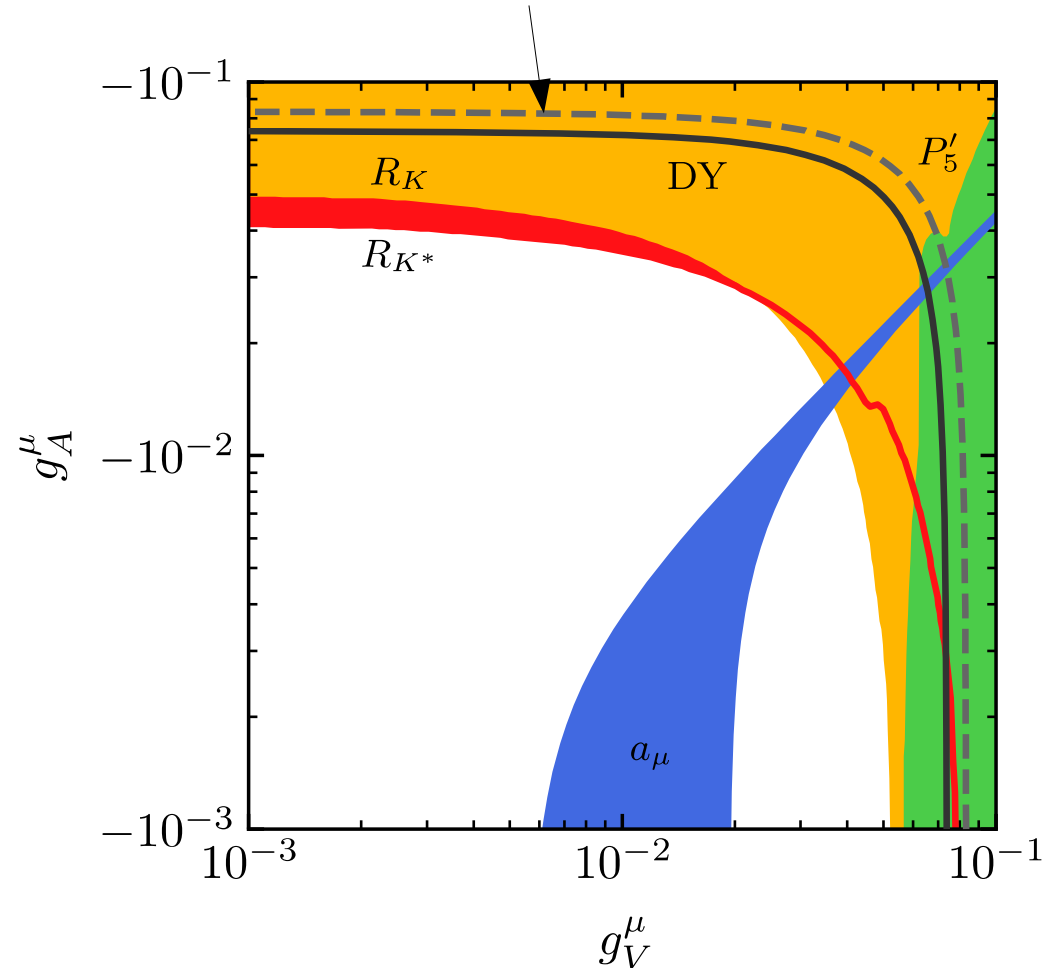
B anomaly preferred regions taken from: [Sala, Straub, \[1704.06188\]](#) $g_{bs} = -1.5 \times 10^{-8}$

Derived constraint

Adding uncorrelated 2.1% luminosity uncertainty



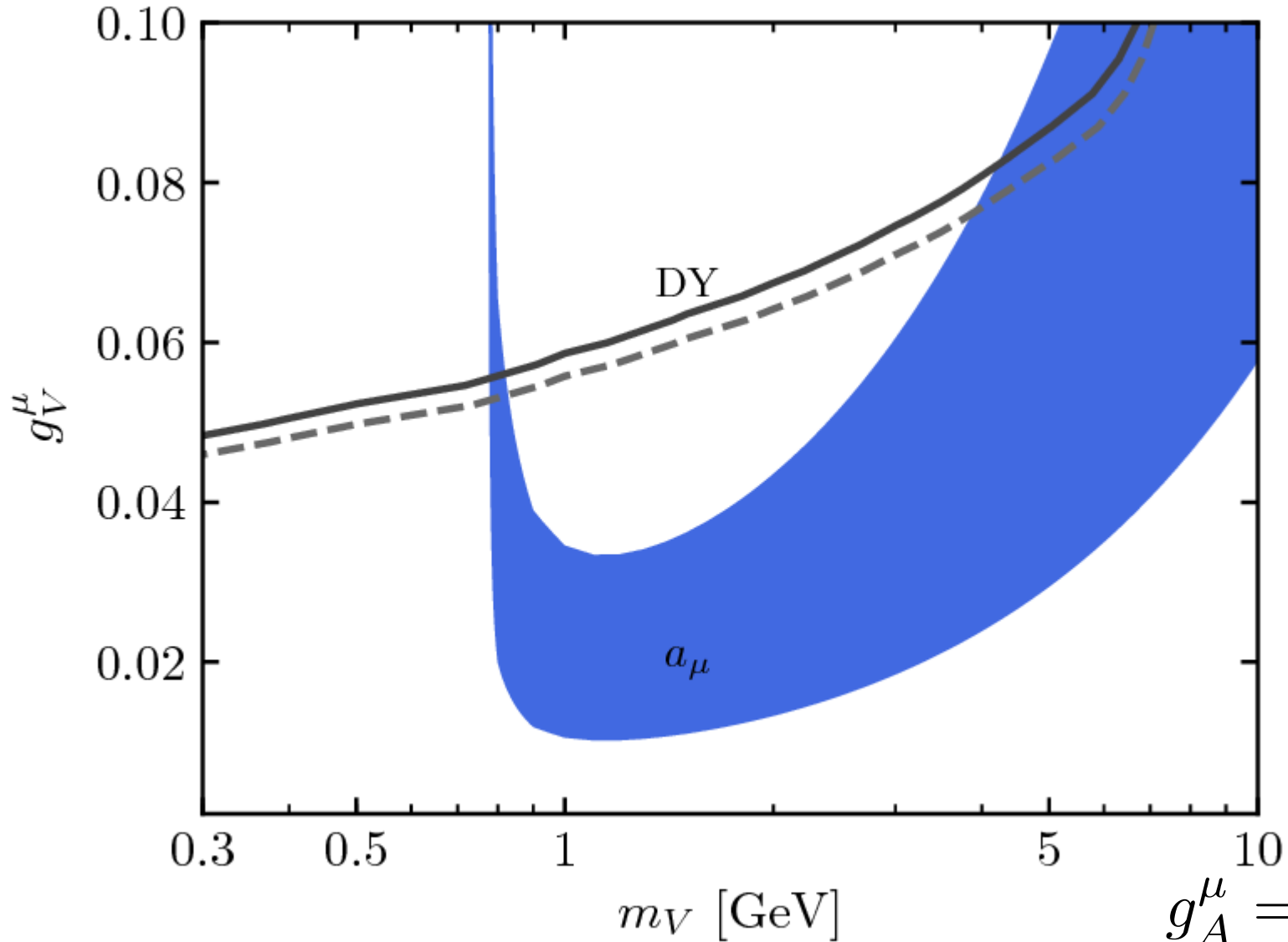
95% CL contours for various masses in GeV



B anomaly preferred regions taken from: [Sala, Straub, \[1704.06188\]](#) $g_{bs} = -1.5 \times 10^{-8}$

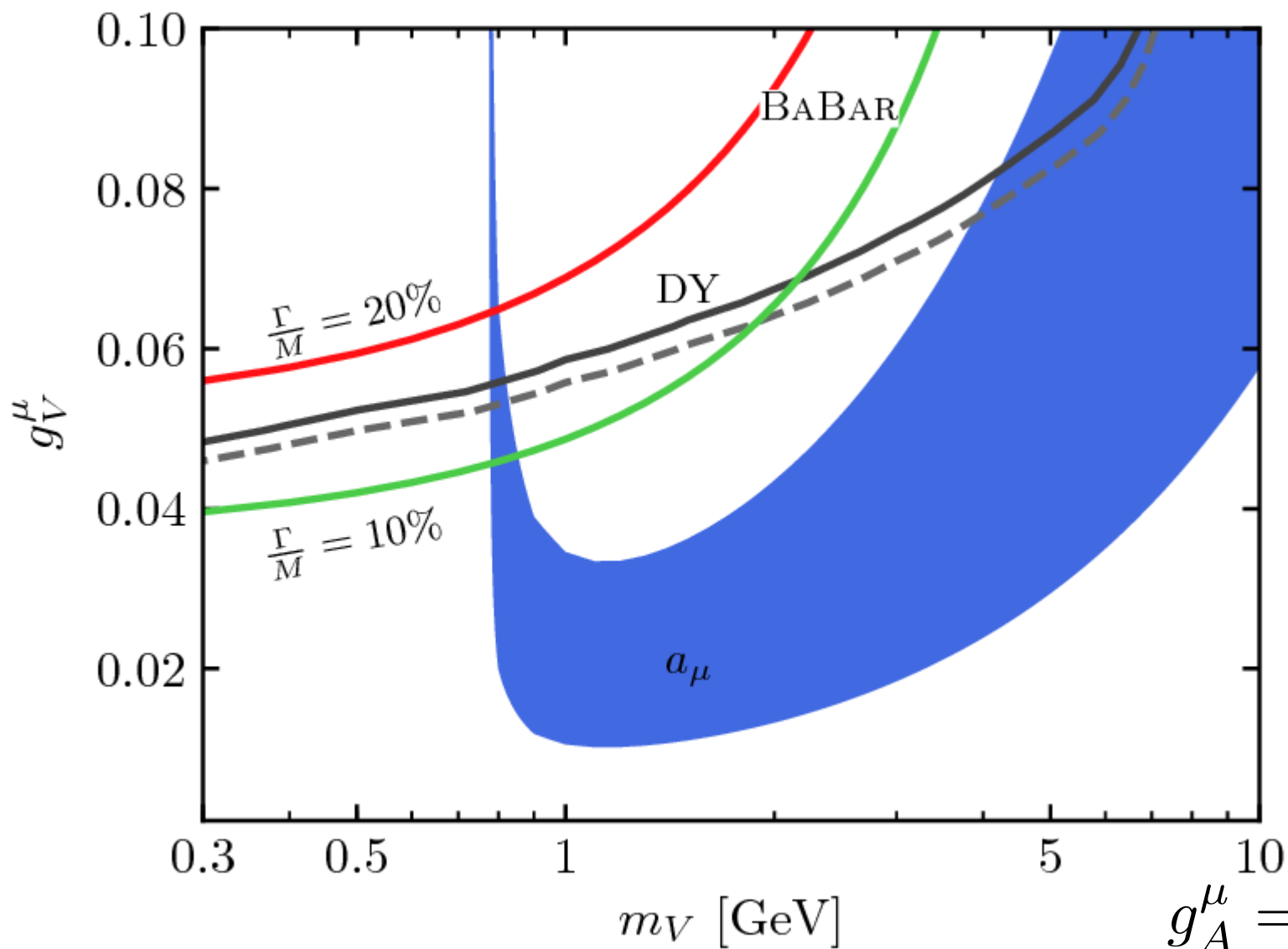
$(g - 2)\mu$

$$\Delta a_\mu \equiv a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = \frac{x}{12\pi^2} [g_V^2 f(x) + g_A^2 g(x)], \quad \text{with } x \equiv \frac{m_\mu^2}{m_V^2}$$



$$(g - 2)\mu$$

$$\Delta a_\mu = (287 \pm 80) \times 10^{-11} \quad \text{PDG 2016}$$



$$g_A^\mu = 0.41 g_V^\mu$$

The BaBar bound

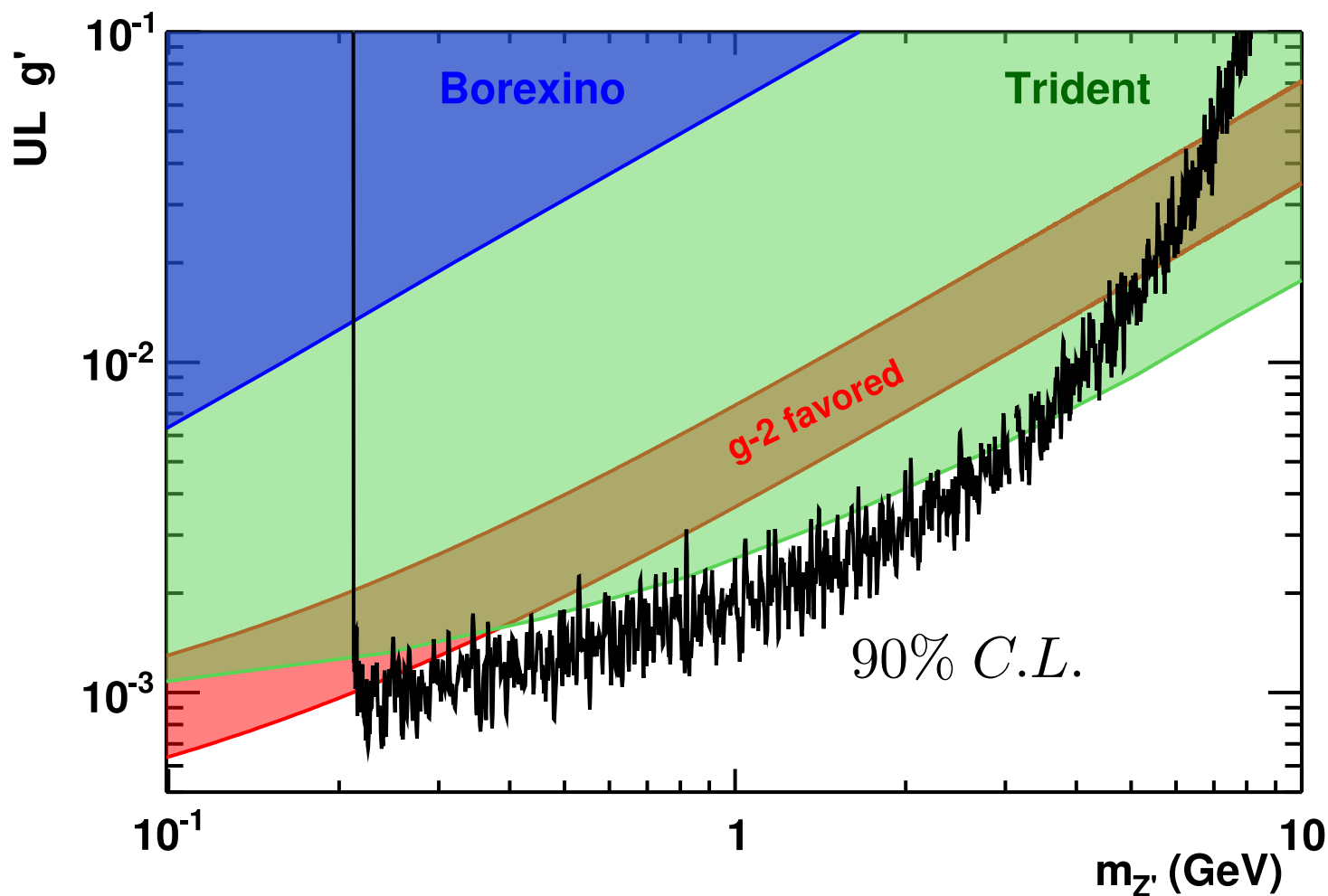
Brian Shuve [private comm.]

BaBar [1606.03501]

Benchmark model: gauged $L_\mu - L_\tau$

Look for $e^+e^- \rightarrow 4\mu$

For us, recast it by Br



Summary and outlook

- LHC measurements provide relevant constraints for light NP
- Many precision observables: $m_{\ell\ell}$, A_{FB} , ϕ^* , ...
- Use ratios (like normalized distributions) to reduce luminosity & PDF uncertainties
- Careful treatment of higher order effects needed for precise bound

Thank you!