

Calorimeters and Pandora Performance

Matthias Weber (CERN)

CLIC calorimeters and PandoraPFA



Fine grain calorimeters at CLIC with high segmentation to achieve best possible performance of particle flow identification:

- PandoraPFA algorithms matching information of all detector subsystems to identify and reconstruct each particle correctly by its type: charged hadrons (assigned type: $\pi\pm$), muons, electrons, photons, neutral hadrons (assigned type: neutrons)
- important task: keep misidentification of particles as low as possible as well as wrong assignment/splitting of calorimeter clusters to avoid double counting
 →this confusion term becomes important contribution in jet energy resolution for high energies, low confusion term essential to achieve large discrimination power between boosted jets from W and Z
- tauFinder run after particle identification by PandoraPFA
- Ongoing work to integrate forward calorimeters (BeamCal and LumiCal) into reconstruction sequence



New CLIC default Pandora Settings

Previous CLIC PandoraSettings: CLIC workshop in March



Set a cut on MaxHCalHitHadronicEnergy of 1 GeV

 \rightarrow scales all Hadron Hcal hit energies beyond 1 GeV of hadrons down to 1 GeV

 \rightarrow cut likely too harsh for energy stages at 1.5 and 3 TeV, reconstructed energy largely asymmetric with long tail to lower energies, mean shifted to 5-7 % lower energies

Muon identification cuts not adopted to new detector model with reduced number of muon chambers (six instead of 9)

 \rightarrow Large muon identification inefficiency in barrel



New CLIC PandoraSettings



- Remove cut on MaxHCalHitHadronicEnergy (MHHHE) of 1 GeV
- Use SoftwareCompensation, which reweights hits with very large energy densities. These are more likely to come from electromagnetic component in hadron showers which have higher response in HCAL
- \rightarrow Proposed default settings by PandoraPFA experts from Cambridge
- Use SoftwareCompensation weights from ILD at the moment, weights not applies for hadrons beyond 100 GeV

 \rightarrow work on CLIC specific weights has been started

• New muon identification cuts adopted for CLIC

Use new cluster noise cleaning

 \rightarrow Rescale energy of HCAL hits with large contribution to cluster energy by using information from adjacent layers

Total energy reconstruction



Compare the reconstructed energy using the different settings



This behavior is gone with the software compensation (MHHHE removed), mean of reconstructed energy distribution within 1 % of true energy for all energy points considered, symmetric distribution

6

Jet Energy Resolution



Determined using RMS_{90} the total relative energy resolution

 $\Sigma E_{PFO}/\Sigma E_{MC}$, where all stable visible MC particles are used in the denominator, ignoring the energies of neutrinos

→ Multiply by a factor of $\sqrt{2}$ to get the jet energy resolution values

 \rightarrow Study as function for vs cos θ of leading quark for several energies



For very low jet energies (50 GeV) resolution around 5 %, for medium jet energies (100-250 GeV) jet energy values resolution around 3.5-4.0 %. Considerable improvement for high jet energies from previously 5 and 6 % reduced to 4-4.5 %

8



First look: Overlay in Z→uds dataset

Overlay: previous performance of PandoraPFA old detector model CLIC_ILD



Considerable increase of missing transverse p_x with overlay, tight selected PFOs are used for the overlay plot

Overlay: missing transverse momentum



Overlay increases missing $p_x RMS_{90}$ from 4.54 to 6.42 GeV for 380 GeV \rightarrow previously increased to 9 GeV (no overlay around 4.5 as well)

For 1500 GeV dataset RMS₉₀ increased from 19.7 to 21.1 GeV \rightarrow previously increased to 24 GeV (no overlay around 18.5)

→caveat: use old PandoraPFA settings, expect more improvement for higher energy CLICdp Collaboration Meeting 2017, Software & Validation Session, August 30
11
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Hadronic tau identification: TauFinder performance

TauFinder



0.82

0.8

0.78

0.76

0.74

0.72

0.7

0.68

0.66

0.64

Taufinder integrated into MarlinReco/Analysis, check efficiency for several samples: $e^+e^- \rightarrow \tau^+\tau^-$ at 200, 400, 1000, 2000 GeV, cut on theta between 80-100 degrees

- Require angular matching of reconstructed tau jet and MC truth tau jet of 10 • degree (very relaxed 30 degree angular cut does hardly do anything)
- Require charge matching (hardly any inefficiency <1%) ٠



Nice efficiencies for low energy samples (100,200 GeV taus)

3

Tau identification efficiency: issue at high energies



Efficiency for several tau samples: 200, 400, 1000, 2000 GeV Require angular matching of reconstructed tau jet and MC truth tau jet: 10 degree (very relaxed 30 degree angular cut does hardly do anything) Require charge matching (hardly any inefficiency <1 %)





Efficiency decrease by over 50 %

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Tau inefficiency for one prong point towards issues in pion identification: Particle gun samples produced at fixed energy, homogeneous in $\cos\theta$ Pion identification efficiency drops to 80 % around 75-105 degrees. \rightarrow possibly also an effect of pion Bremsstrahlung \rightarrow Check with muons if tracking is the issue here



In non correctly identified pion events track exists, but for big majority tracks stop picking up hits well before Calorimeter face \rightarrow clusters not matched to track by PandoraPFA

In majority of events Pandora reconstructs two or more particles in the event





For Muon identification efficiency is over 98 % using newly adopted CLIC Muon identification criteria (besides exactly at 90 degrees) from 10 GeV up to 1000 GeV → For muons tracking works perfectly fine up to highest energies

 \rightarrow issue for pions not only related purely to tracking, physics effects do play a role



For muons at 1 TeV often an additional neutron is created As expected hardly changes anything on the total energy

Electron identification efficiency at large energies



Electrons are identified with efficiencies typically beyond 95 % For all events at least 1, typically more than two photons are created



Electron energy shows long Bremsstrahlung tail to lower energies, total energy of additional PFO photons overcompensates for this effect



Move to new default PandoraSetting in clicReconstruction

- Recover energy response deficit of 5 %, mainly a result of removed cut on HCAL hit energy for hadrons with impact on high energetic showers
- Software compensation enabled \rightarrow started work on CLIC specific weights

Changed muon identification parameters recover identification inefficiency in barrel

• Good identification efficiency over wide energy range

Tau identification efficiency for hadronic tau decays (both 1 and 3 prong decays) around 80 % for tau's up to 200 GeV, tau fake rate typically between 1.5 and 3 % in Z dijet events

• Issue at for very high energetic tau's, traced back to issues in reconstruction for high energetic pions

Check identification efficiency and energy reconstruction for high energetic pions and electrons

- For electrons impact of bremsstrahlung photons
- For pions split of clusters leads to creation of additional neutrons CLICdp Collaboration Meeting 2017, 22

Plans and outlook



- Work on CLIC specific software compensation weights → could further improve jet energy resolution
- Investigate issues of pion reconstruction at very large energies and physics cause of split into two calorimeter clusters
- Run overlay with new PandoraSettings and in study impact in depth
- Forward calorimeters now in default reconstruction sequence
 - Necessary background input file available in CLICPerformance/examples
- Further issues under investigation (not discussed in this talk)
 - Identification inefficiencies of high energetic photons in jet environment
 - Muon identification dip at exactly 90 degrees
 - Pion/Electron identification inefficiency at barrel-endcap transition (cause identified, possible solutions in works)



BACKUP

Remaining issues at high energies: contribution from photons



Check what might cause the higher jet energy resolution values for high energetic jets: photon energy distributions



For low energetic photons contribution well reconstructed, contribution from photons between 10 and 50 GeV vastly underestimated

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PhotonEnergy component in jets

PandoraPerfectPhoton (cheats photo pattern recognition) gives far better result of energy sum of high energetic photons



Only (cheated) perfect photon pattern recognition achieves to correctly identify the contribution of high energetic photons to the event energy

All other settings (CLIC default settings, default setting removing MHHHE cut, software compensation) shifted to lower values, but agreeing with each other

Photon Identification tuning in PandoraPFA



For each model the photon identification is retuned

- → Use PandoraSettingsPhotonTraining configuration file, result is a likelihood file for photons
- → Binned in terms of energy (0, 0.2, 0.5, 1, 1.5, 2.0,5,10,20 GeV)
- \rightarrow After photon likelihood file has been determined rerun calibration

 \rightarrow tested increase in photon identification bins to 11 to account more for possible differences at high photon energies (0, 0.2, 0.5, 1, 1.5, 2.0,5,10,20,35,50 GeV)

 \rightarrow Unfortunately no sizeable improvement observed

Pion Energy component in jets





Overlay: missing transverse momentum 1500 GeV





Overlay increases missing $p_x RMS_{90}$ from 4.54 to 6.42 GeV for 380 GeV \rightarrow previously increased to 9 GeV For 1500 GeV dataset RMS₉₀ increased from 19.7 to 21.1 GeV \rightarrow previously increased to 24 GeV

Hadron spectrum for CLIC (Zuds 500 vs Zuds 3000 GeV)



For 500 GeV dataset neutral hadron energies beyond 90 GeV are 1.9 %, for 3000 dataset 13.7 % \rightarrow if we want same coverage of neutral hadron energy spectrum need to calculate weights for samples up to 400 GeV (1.7 % beyond that point for 3000 GeV sample)

 \rightarrow For high end tail definitely need CLIC specific weights



Hit energy density for 1000 GeV neutral hadrons

Software compensation weights are determined by minimisation of a χ^2 like function using all energy points and densities up to 30 GeV/1000 cm³



Quite a few hits with high energy density beyond 100 GeV/dm³ (cutoff until when weights are applied)