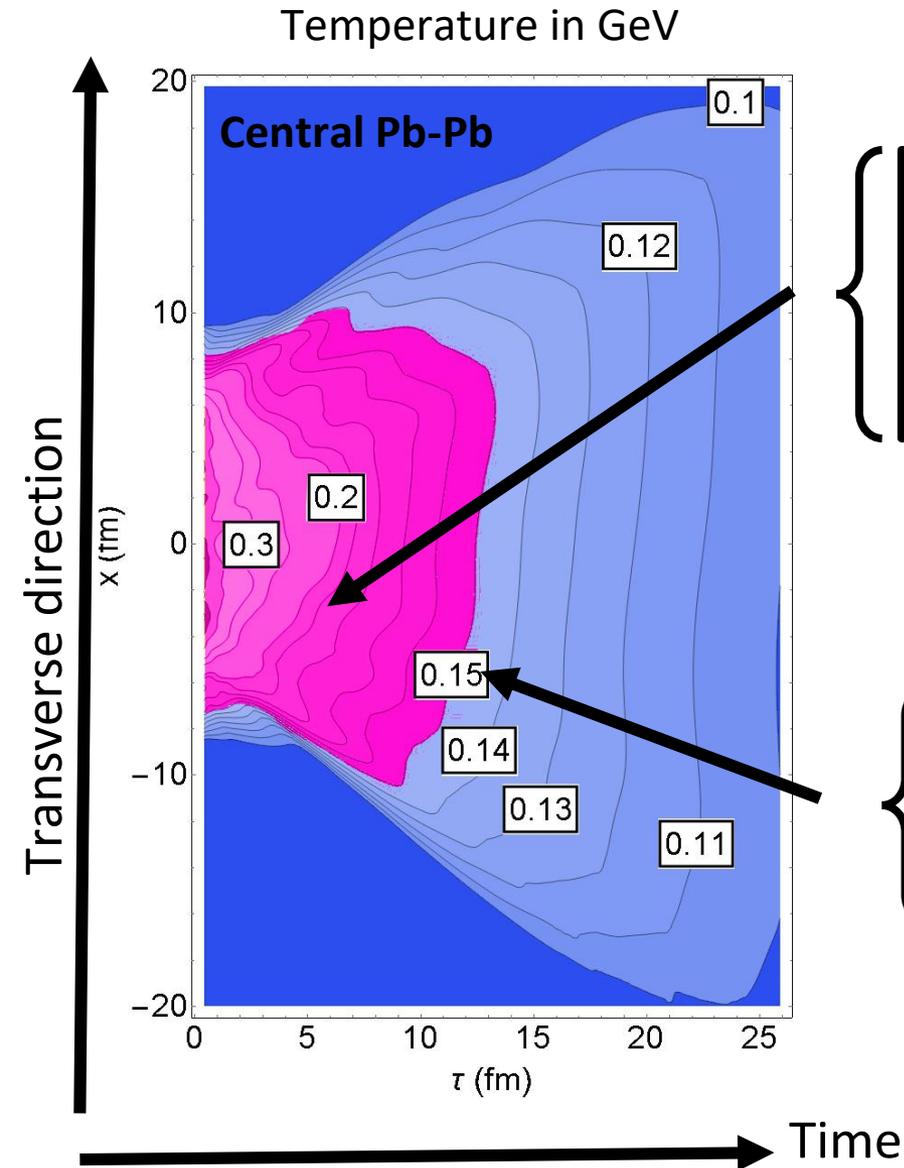# Probing heavy ion collisions with photons at low and high $\sqrt{s_{NN}}$

Jean-François Paquet

Duke
UNIVERSITY

in collaboration with
Charles Gale, Sangyong Jeon,
Scott McDonald & Chun Shen

Photons as probes of the plasma

Temperature in GeV

Electromagnetic probes
(along with jet energy loss, heavy quarks, ...):
More direct probes of the hot early plasma

As opposed to **soft hadron** observables:
Indirect probes of the hot plasma

Sources of photons in heavy ion collisions

Dominant sources of photons (excluding hadronic decays):
prompt & thermal photons

Prompt photons

- Produced in initial hard nucleon collisions

$$\frac{\mathrm{d}N_\gamma}{\mathrm{d}^3 p} = \frac{N_{binary}}{\sigma_{pp}^{inel}} \underbrace{f_{a/A} \otimes f_{b/B} \otimes \mathrm{d}\hat{\sigma}_{ab \to \gamma/c + d} [\otimes D_{\gamma/c}]}_{\substack{\text{NLO pQCD [INCNLO] } (p_T \text{ dependence}) \\ \text{[Could also use fit to p+p data]}}}$$

Thermal photons

- Radiated by the hot plasma

$$\frac{\mathrm{d}N_\gamma}{\mathrm{d}^3 p} = \int \mathrm{d}^4 X \frac{\mathrm{d}\Gamma_\gamma}{\mathrm{d}^3 p} (\mathrm{p}, \underbrace{T(X), u^\mu(X), \dots}_{\substack{\text{Spacetime profile of} \\ \text{plasma from} \\ \text{hydrodynamic simulation}}})$$

Photon emission rate
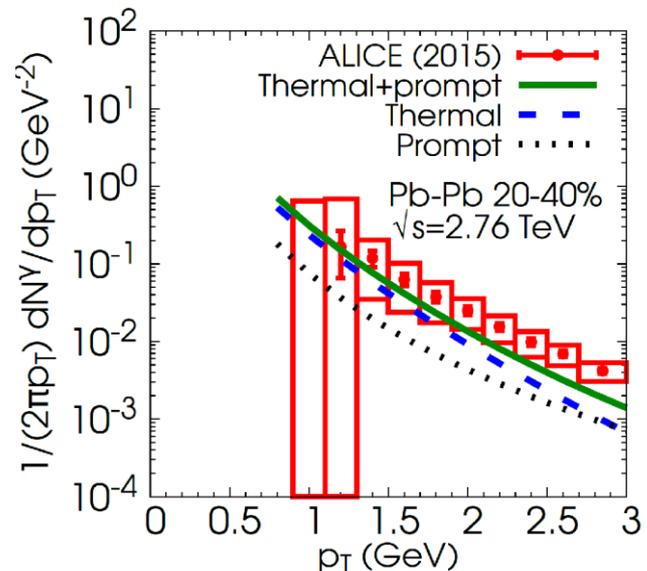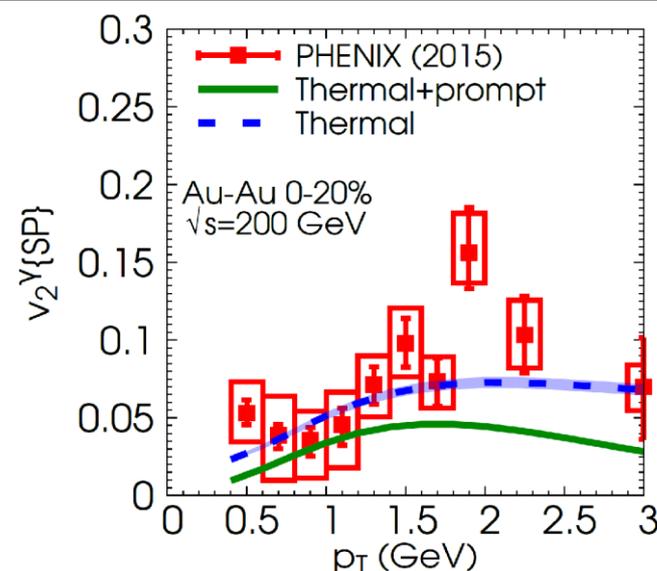(per volume)
for hot QCD plasma

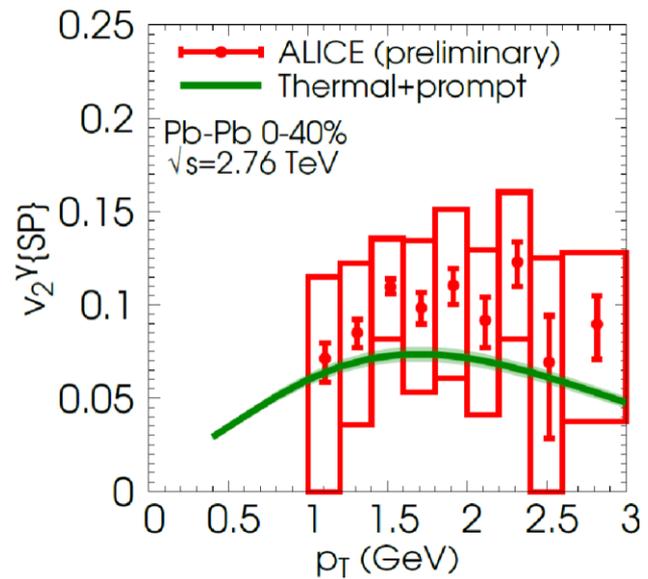Temperature [in GeV] and flow velocity profile

Differential spectra & v_2 at LHC & top RHIC $\sqrt{s_{NN}}$
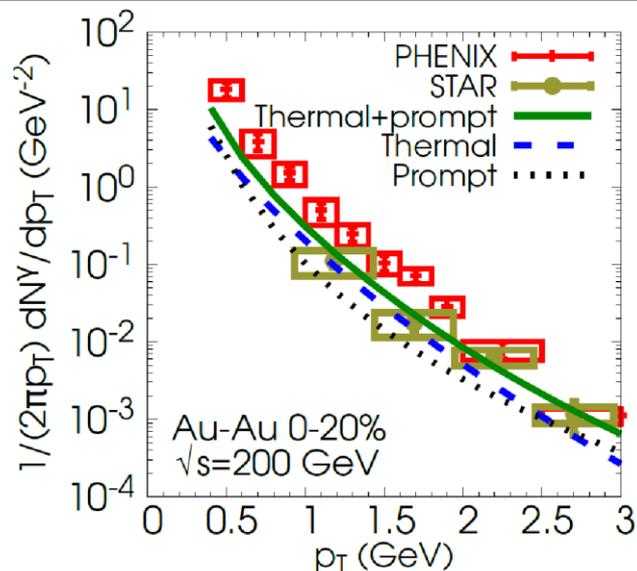
RHIC
$\sqrt{s_{NN}}=$
200 GeV

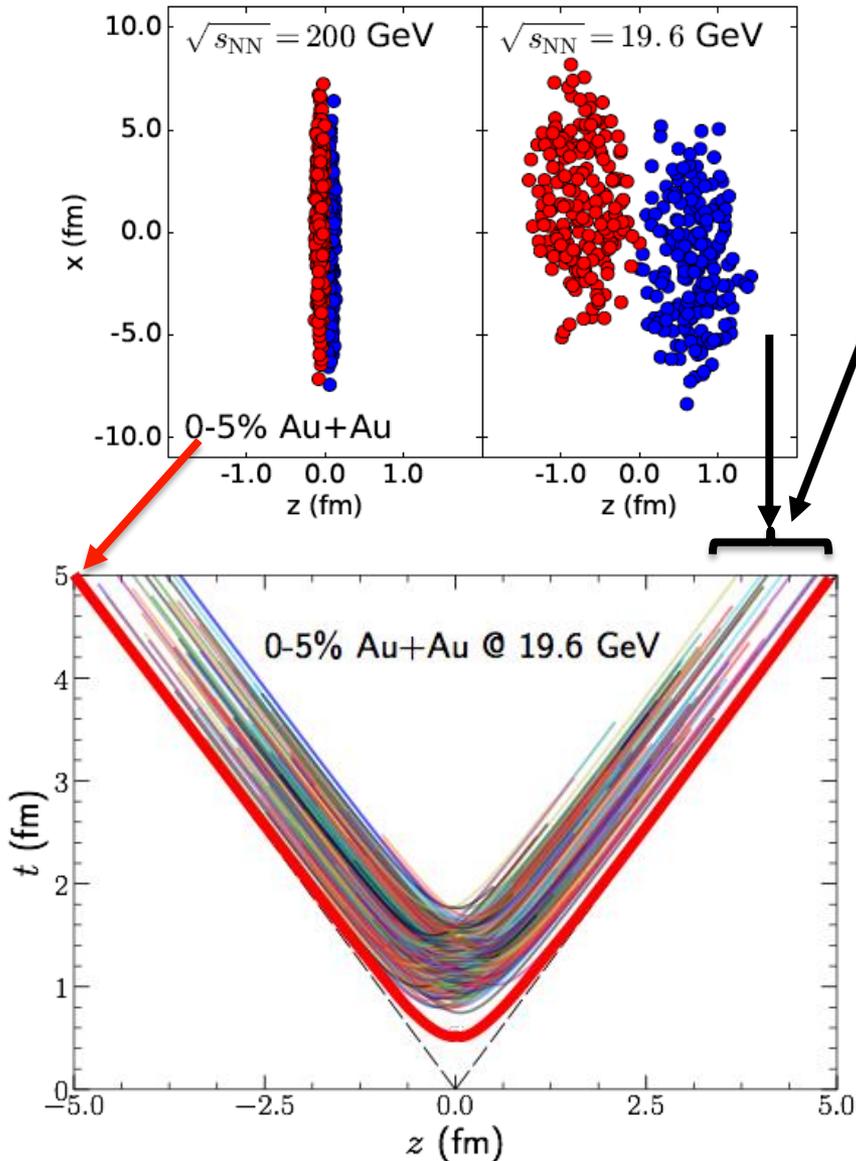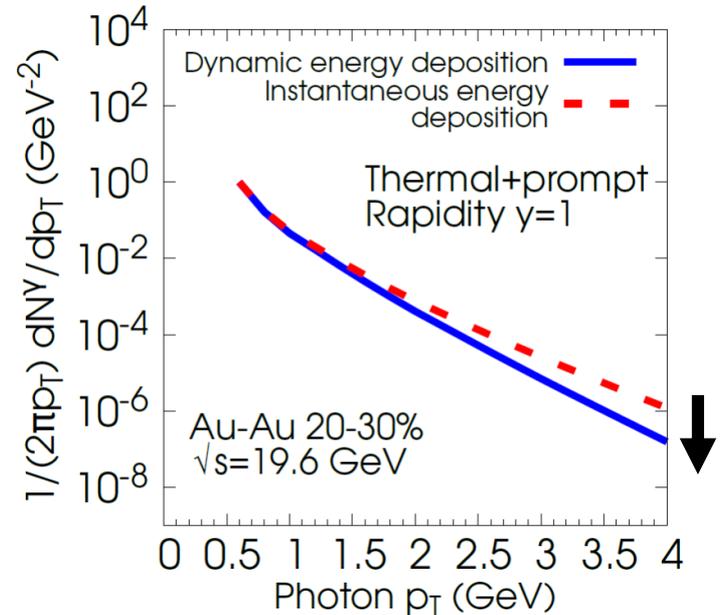LHC
$\sqrt{s_{NN}}=$
2760 GeV

Ref.: Gale, Jeon, McDonald, Paquet, Shen, in preparation.

Photons in the RHIC beam energy scan

Ref.: Schenke and Shen, PRC 97 (2018)

Energy **deposited over extended time** period **at lower** $\sqrt{s_{NN}}$
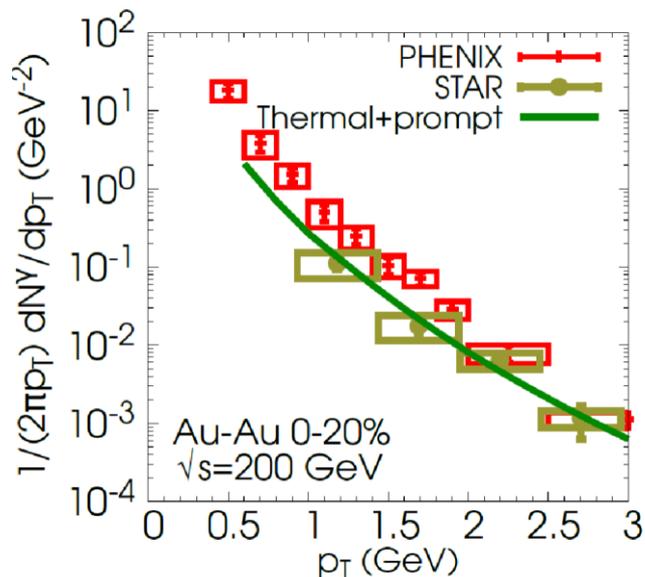
RHIC $\sqrt{s_{NN}}$=19.6 GeV

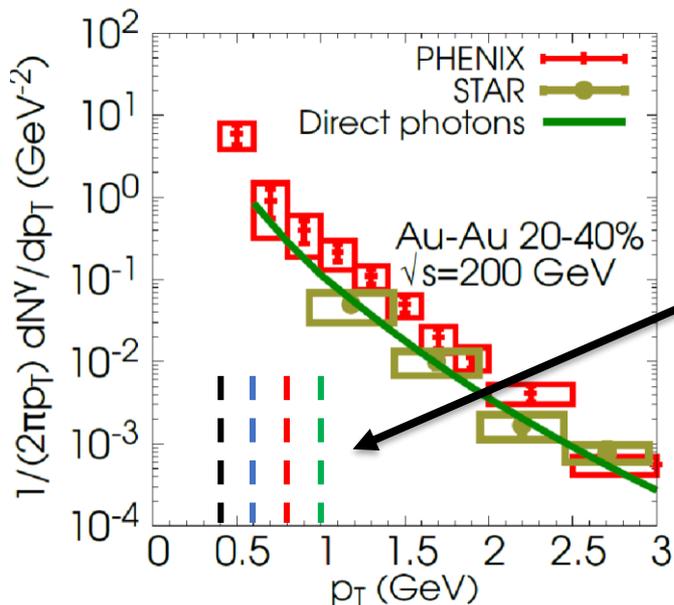Ref.: Gale, Jeon, McDonald, Paquet, Shen, in preparation.

Photons sensitive to initial energy deposition

Looking at the same data in a different light

Centrality&$\sqrt{s_{NN}}$ dependence?
Use p_T-integrated observables
(e.g. **photon multiplicity**)
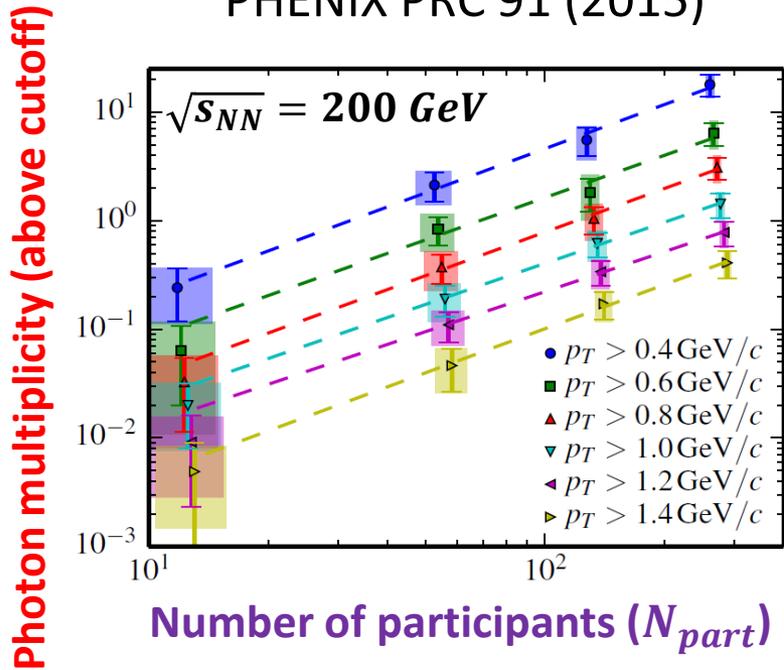
Can use different lower p_T
cut-offs

Jean-François Paquet (Duke)

Centrality&$\sqrt{s_{NN}}$ dependence of photon production

Photon multiplicity (above p_T cutoff)
vs proxy for centrality&$\sqrt{s_{NN}}$ (e.g. **hadron multiplicity**, N_{part})
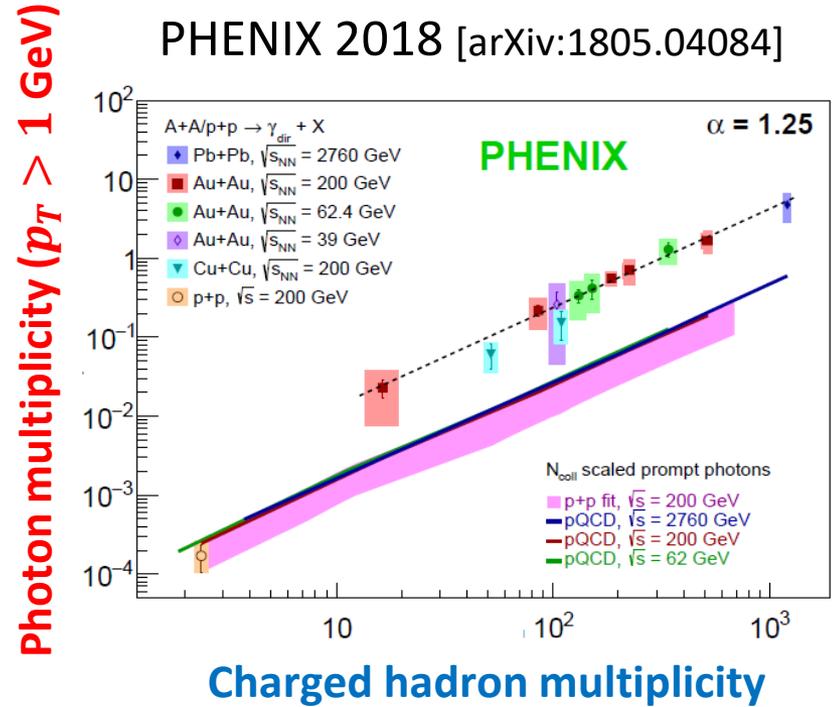[fixed & very low cut-off for hadron multiplicity]

PHENIX PRC 91 (2015)

PHENIX 2018 [arXiv:1805.04084]

PHENIX 2015:
$$\log\left(N^{\gamma}_{p_T > p_T^{cut}}\right) \sim slope(p_T^{cut}) \log(N_{part})$$
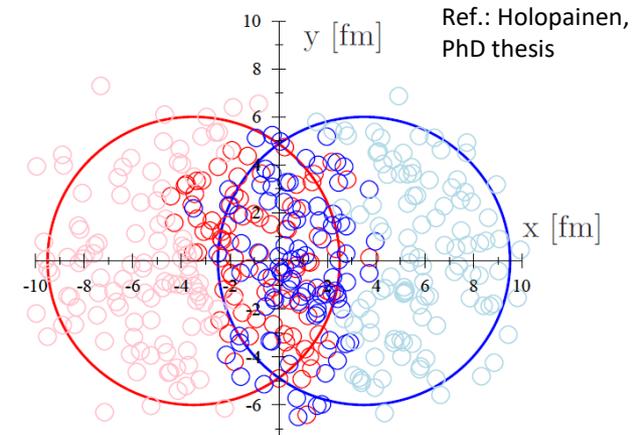
PHENIX 2018:
$$\log\left(N^{\gamma}_{p_T > 1\ GeV}\right) \sim 1.25 \log(N^{h^{\pm}})$$

Photons across centralities and $\sqrt{s_{NN}}$

Prompt photons

$$\frac{\mathrm{d}N_\gamma}{\mathrm{d}^3 k} = \frac{N_{binary}}{\sigma_{pp}^{inel}} \, f_{a/A} \otimes f_{b/B} \otimes \mathrm{d}\hat{\sigma}_{ab \to \gamma/c+d} [\otimes D_{\gamma/c}]$$

- Scale with the number of binary collisions
- Depends on $\sqrt{s_{NN}}$ through N_{binary}, σ_{pp}^{inel} and $f \otimes f \otimes \mathrm{d}\hat{\sigma} \otimes D$
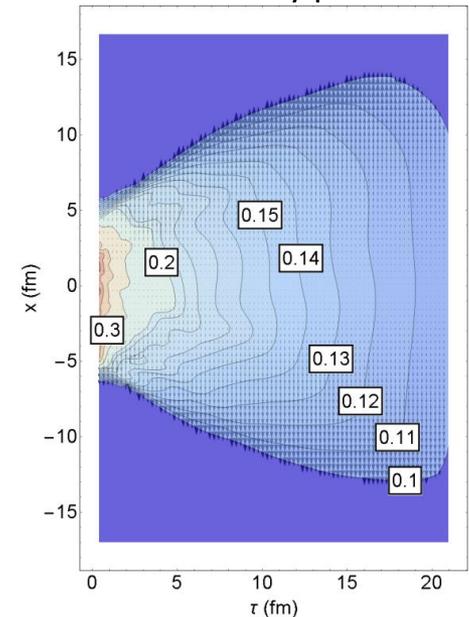
Ref.: Holopainen, PhD thesis

Thermal photons

$$\frac{\mathrm{d}N_\gamma}{\mathrm{d}^3 p} = \int \mathrm{d}^4 X \, \frac{\mathrm{d}\Gamma_\gamma}{\mathrm{d}^3 p} \left(k, T(X), u^\mu(X), \ldots \right)$$

Photon emission rate (per volume) for hot QCD plasma
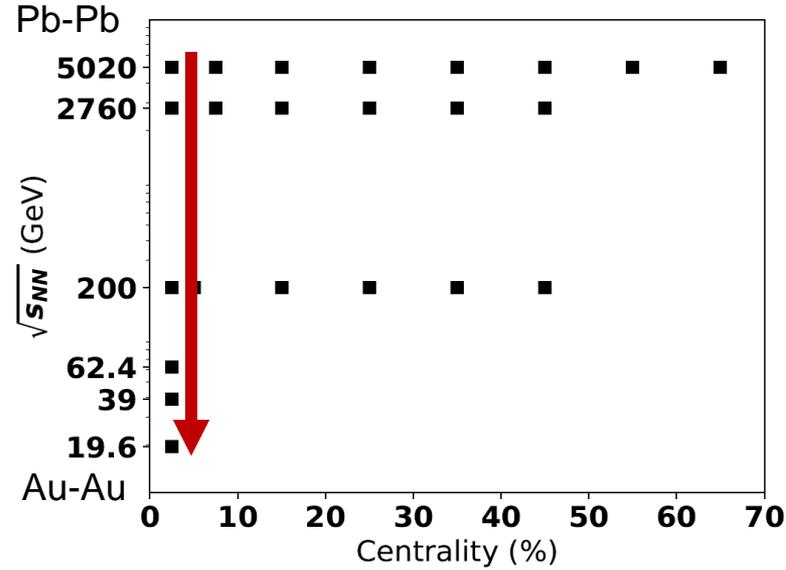
Spacetime profile of plasma from hydrodynamic simulation

- Different $\sqrt{s_{NN}}$ & centrality **=** different temperature and flow velocity profiles

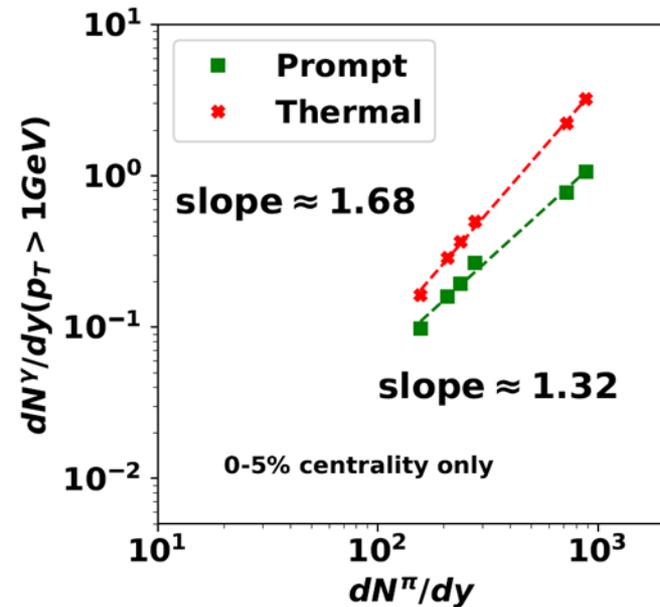Temperature [in GeV] and flow velocity profile

Ref.: QM2017 proceedings [1704.07842]

8

$\sqrt{s_{NN}}$ dependence

$$\log\left(N^{\gamma}_{p_T>p_T^{min}}\right) \sim slope(p_T^{min})\log(N^{\pi})$$

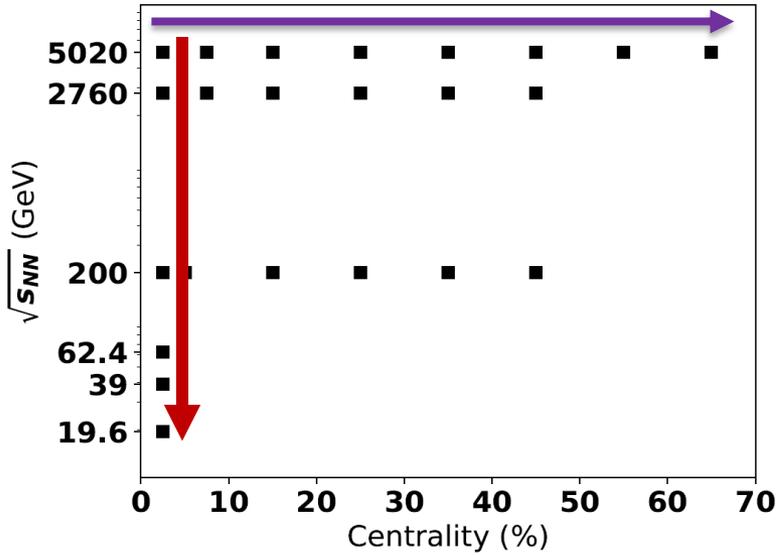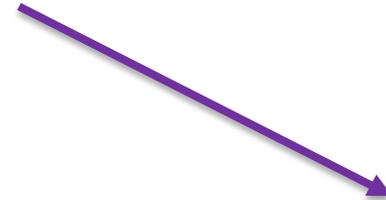**Varying $\sqrt{s_{NN}}$ at fixed centrality**

$\sqrt{s_{NN}}$ **dependence**
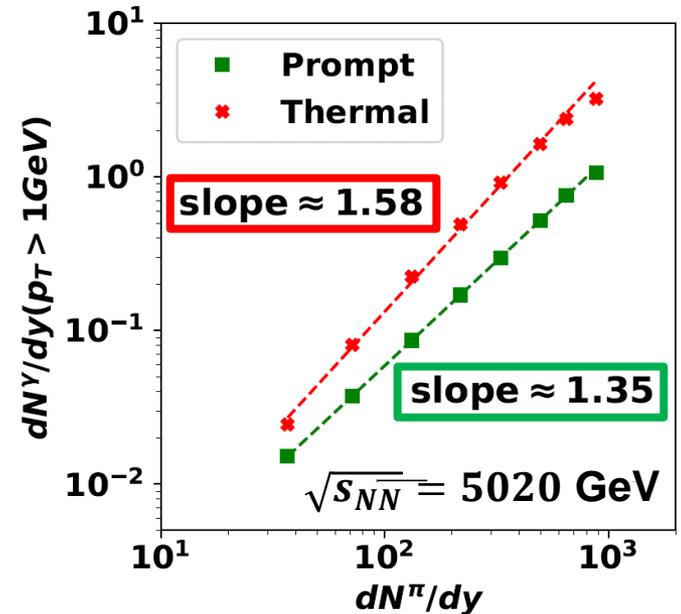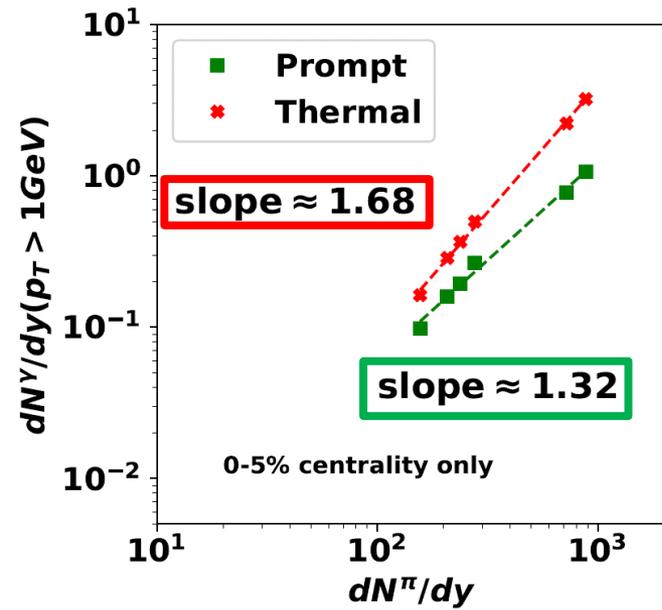
$$\log\left(N^{\gamma}_{p_T>p_T^{min}}\right) \sim slope(p_T^{min})\log(N^{\pi})$$

Varying centrality at fixed $\sqrt{s_{NN}}$

Varying $\sqrt{s_{NN}}$ at fixed centrality

slope ≈ 1.68

slope ≈ 1.32

0-5% centrality only

Prompt
Thermal

slope ≈ 1.58

slope ≈ 1.35

$\sqrt{s_{NN}} = 5020$ GeV

Prompt
Thermal

Jean-François Paquet (Duke)

$\sqrt{s_{NN}}$ dependence
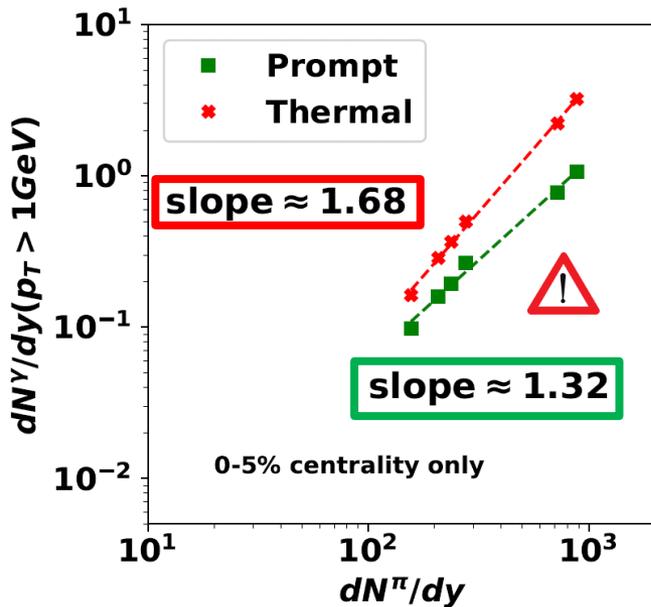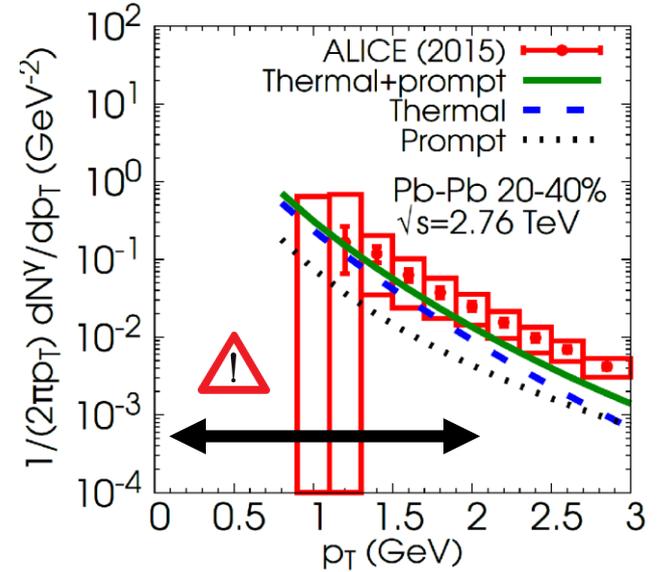
Preliminary results

Prompt photons

pQCD: not accurate at low p_T

$$\frac{dN_\gamma}{d^3p} = \frac{N_{binary}}{\sigma_{pp}^{inel}} \, f_{a/A} \otimes f_{b/B} \otimes d\hat{\sigma}_{ab \to \gamma/c+d} [\otimes D_{\gamma/c}]$$
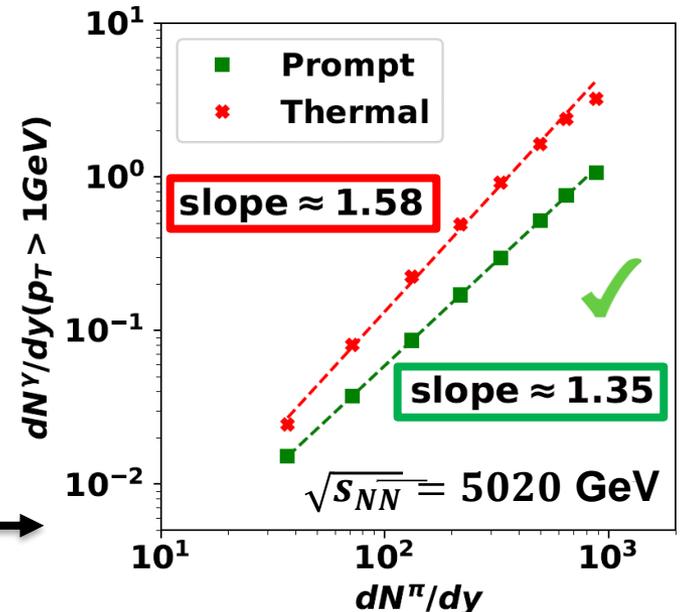
Fixed centrality:

$$\log\left(N^\gamma_{p_T > p_T^{min}}\right) = \log(N_{binary}) + \log(pQCD)$$

$$\approx 1.35 \log(N^\pi) + \log(pQCD)$$

ALICE (2015)
Thermal+prompt
Thermal
Prompt

Pb-Pb 20-40%
$\sqrt{s}$=2.76 TeV

Careful with low p_T prompt photon calculations (large uncertainty)

Centrality dependence under better control

Prompt
Thermal

slope ≈ 1.68

slope ≈ 1.32

0-5% centrality only

Prompt
Thermal
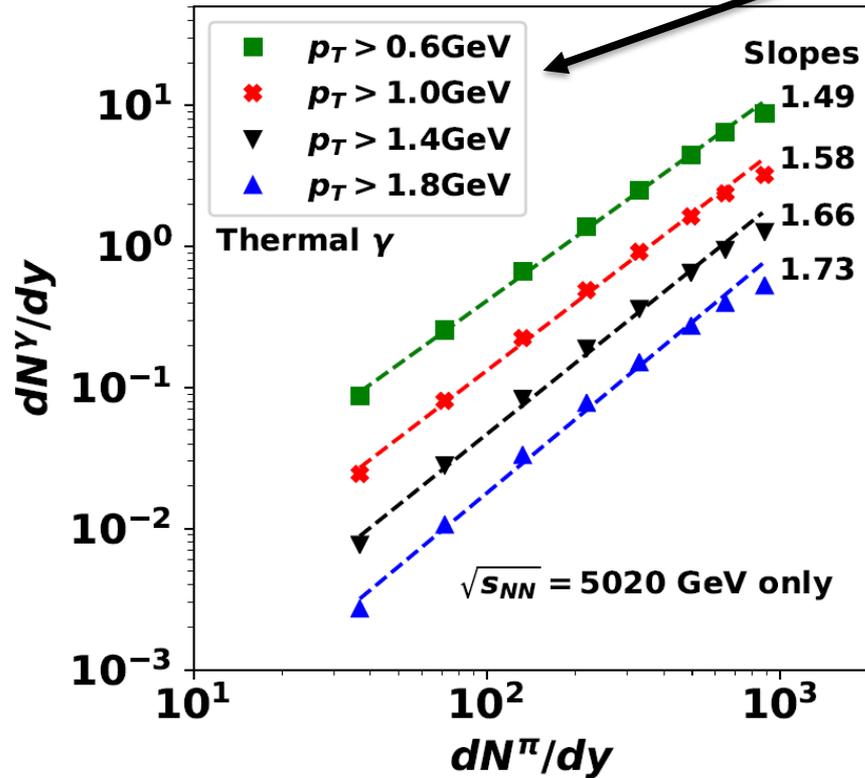
slope ≈ 1.58

slope ≈ 1.35

$\sqrt{s_{NN}} = 5020$ GeV
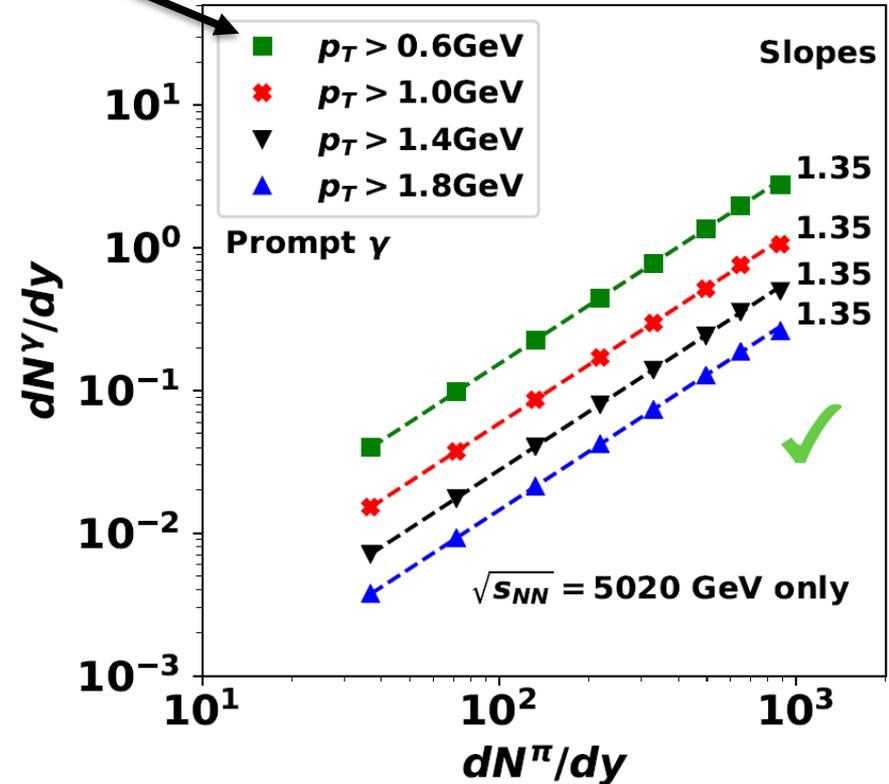
Photon vs pion multiplicity using **centrality**

$$p_T > p_T^{min}$$

Thermal photons
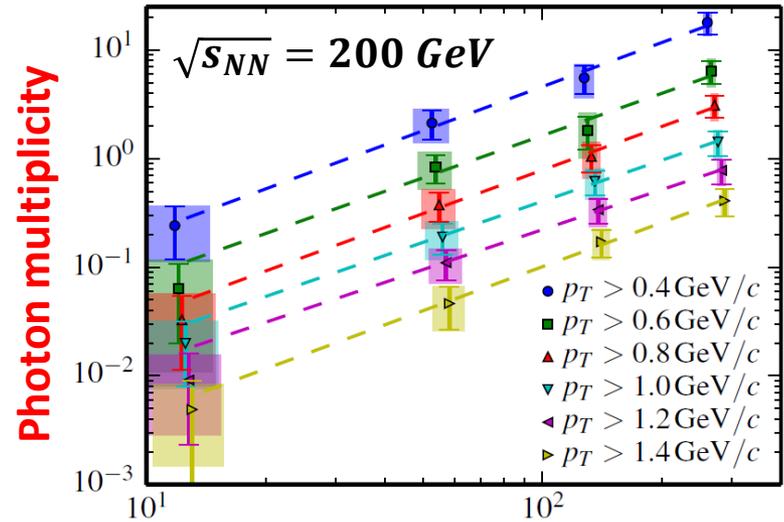
Prompt photons

$$\sqrt{s_{NN}} = 5020 \text{ GeV}$$

Simple test: p_T **cut-off dependence** for **thermal photons**, **NOT** for prompt

Photon vs N_{part} using centrality

What does the data say?

$$\sqrt{s_{NN}} = 200 \text{ GeV}; \, p_T > p_T^{min}$$

Ref.: PHENIX PRC 91 (2015)

p_T^{min} (GeV/c)	Slopes vs N_{part}
0.4	$1.36 \pm 0.08 \pm 0.08$
0.6	$1.41 \pm 0.14 \pm 0.12$
0.8	$1.42 \pm 0.07 \pm 0.11$
1.0	$1.35 \pm 0.06 \pm 0.07$
1.2	$1.36 \pm 0.09 \pm 0.07$
1.4	$1.40 \pm 0.06 \pm 0.10$

Note: STAR also has data [PRB 770 (2017) 451] but uncertainties too large

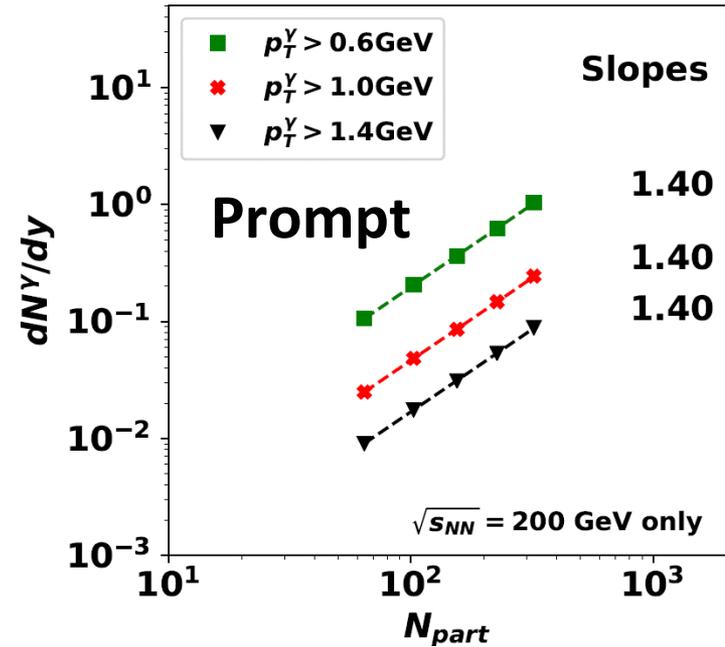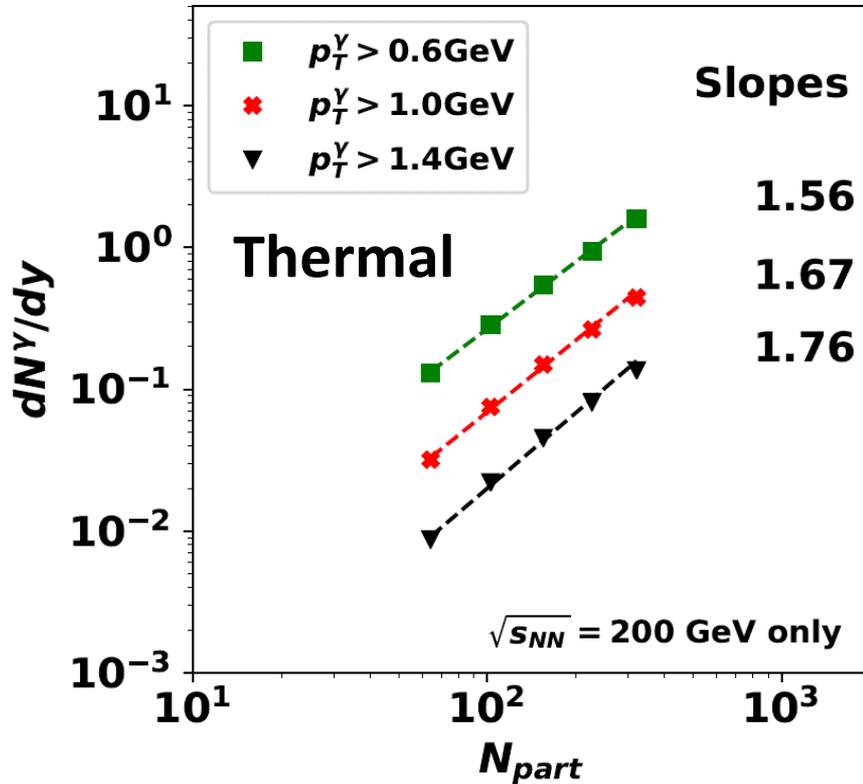# Photon vs pion multiplicity using centrality

What does the data say?
Few centralities and large uncertainties but little sign of p_T dependence

$$\sqrt{s_{NN}} = 200 \text{ GeV}; \; p_T > p_T^{min}$$

Thermal

$p_T^\gamma > 0.6\text{GeV}$
$p_T^\gamma > 1.0\text{GeV}$
$p_T^\gamma > 1.4\text{GeV}$

Slopes

1.56

1.67

1.76

$\sqrt{s_{NN}} = 200$ GeV only

dN^γ/dy vs N_{part}

Prompt

$p_T^\gamma > 0.6\text{GeV}$
$p_T^\gamma > 1.0\text{GeV}$
$p_T^\gamma > 1.4\text{GeV}$

Slopes

1.40

1.40

1.40

$\sqrt{s_{NN}} = 200$ GeV only

Ref.: PHENIX PRC 91 (2015)

p_T^{min} (GeV/c)	Slopes vs N_{part}
0.4	$1.36 \pm 0.08 \pm 0.08$
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1.4	$1.40 \pm 0.06 \pm 0.10$

Note: STAR also has data [PRB 770 (2017) 451] but uncertainties too large

Summary & outlook

- Realistic calculation of photon production across wide range of centrality/$\sqrt{s_{NN}}$

- Photon multiplicity vs centrality/$\sqrt{s_{NN}}$ can provide additional insights
  - Prediction for **centrality dependence** cleaner and more **robust**
  - Prediction for $\sqrt{s_{NN}}$ **difficult** because of prompt photons
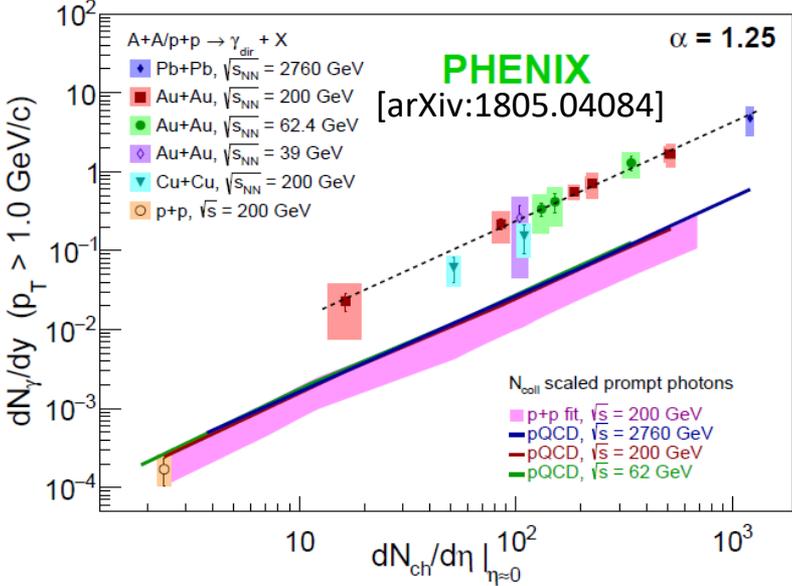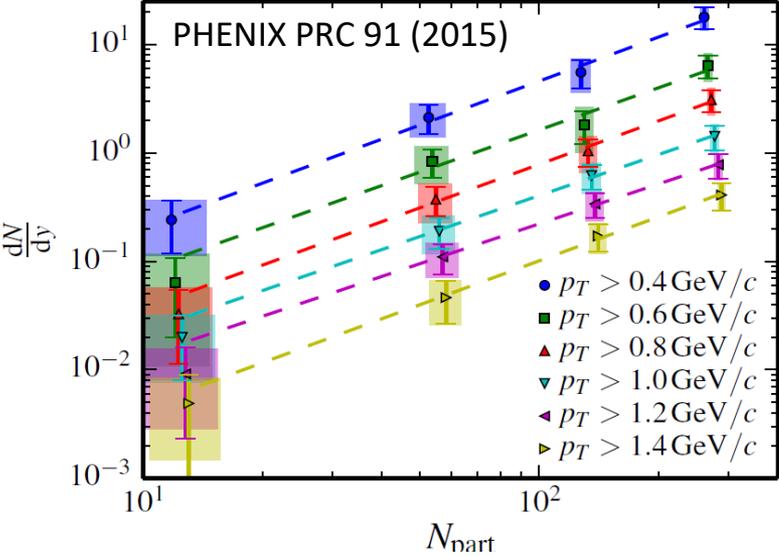
- **Jet-medium photons**: include medium effects to improve prompt photons
- Calculate centrality dependence of **other production mechanisms**

Ref.: Charles Gale, Sangyong Jeon, Scott McDonald (McGill), Jean-François Paquet (Duke) & Chun Shen (Wayne State/RIKEN BNL), in preparation

Questions?

Backup slides

QM2017

Photon vs pion multiplicity using $\sqrt{s_{NN}}$

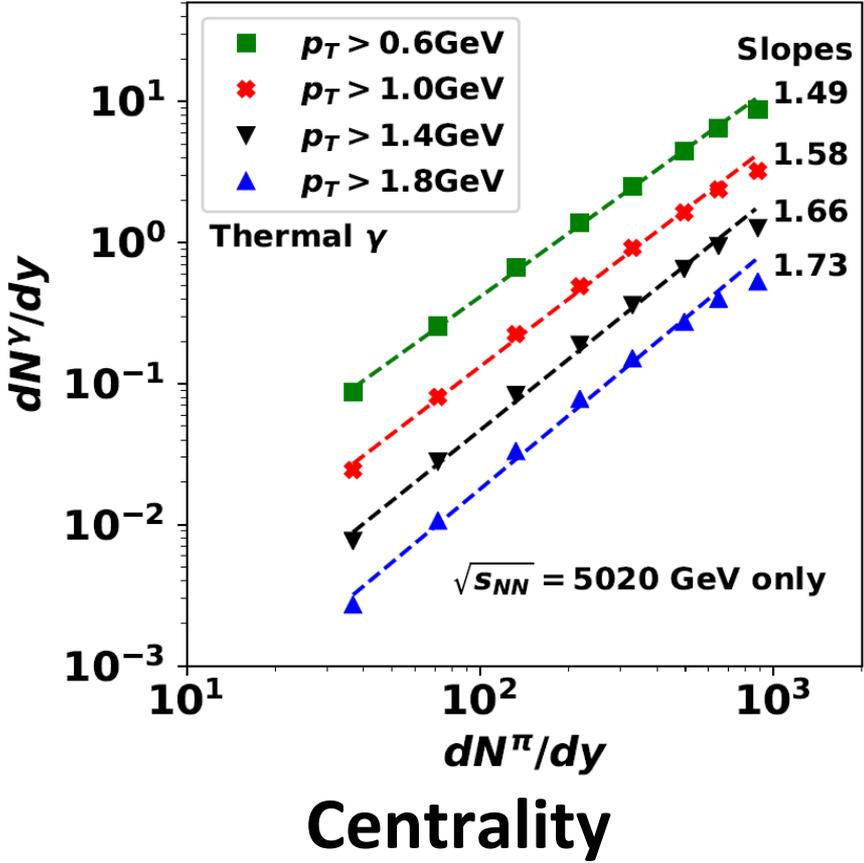$$0 - 5\% \text{ centrality; } p_T > p_T^{\min}$$

Thermal photons

Prompt photons

Difficult to draw conclusion from slope

Thermal photons

Sqrt(s)

Centrality

Prompt photons

Sqrt(s)

Centrality

$\sqrt{s_{NN}}$ dependence
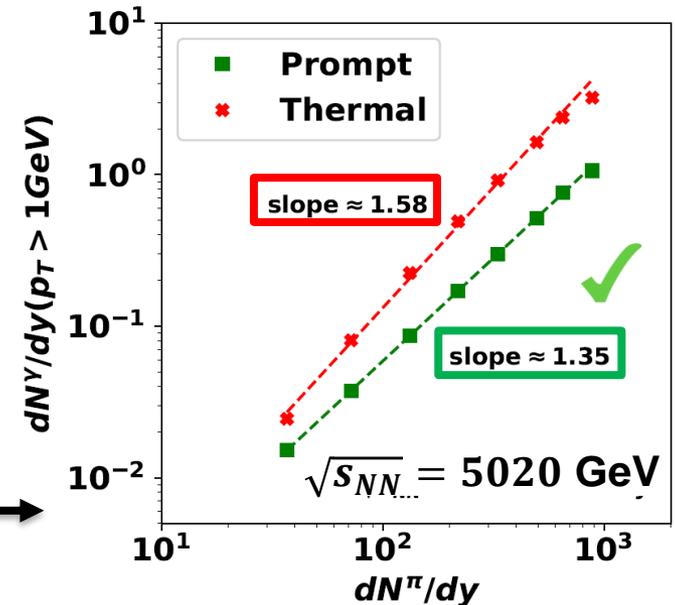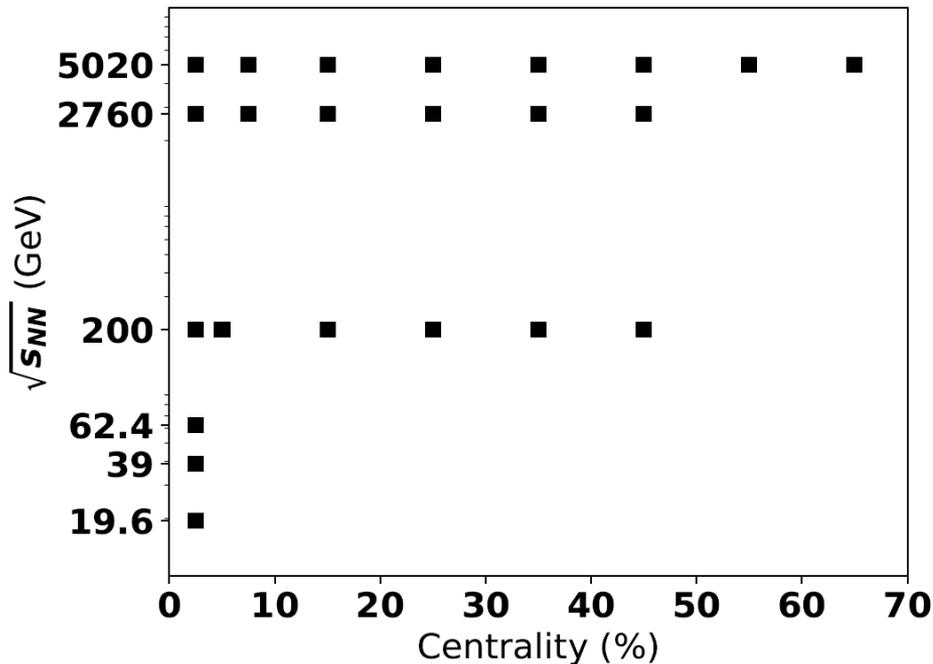
Preliminary results

Careful with low p_T prompt photon calculations (large uncertainty)

Centrality dependence under better control

Centrality & $\sqrt{s_{NN}}$ dependence of photon production

Photon multiplicity (above p_T cutoff)
vs **hadron multiplicity** (proxy for centrality & $\sqrt{s_{NN}}$)

Ref.: Gale, Jeon, McDonald, Paquet, Shen, in preparation.

RHIC $\sqrt{s_{NN}}$=19.6, 39, 62.4 & 200 GeV;
LHC $\sqrt{s_{NN}}$=2760 & 5020 GeV
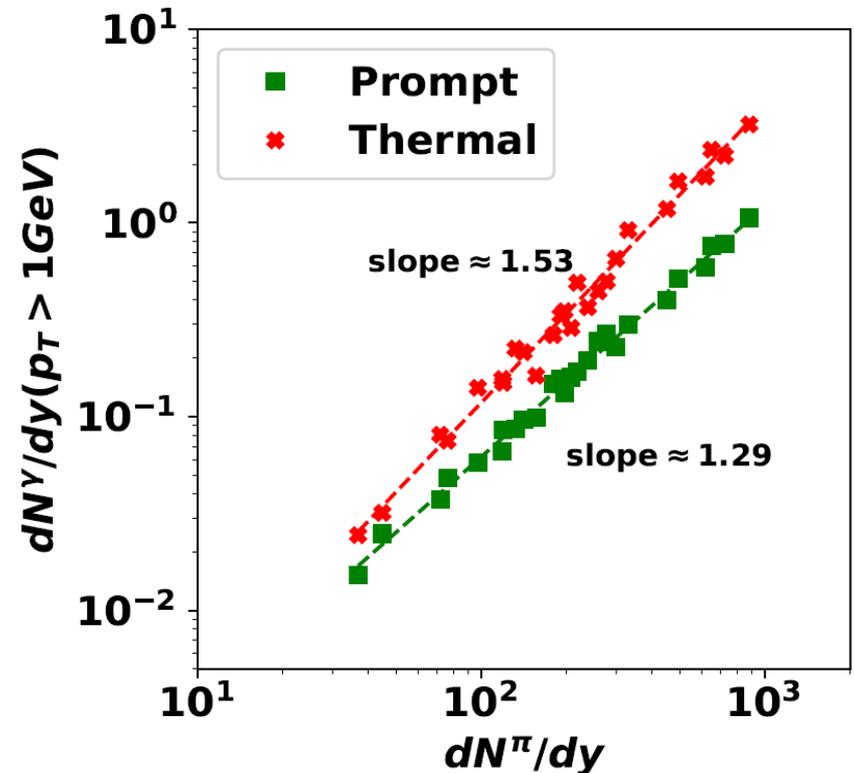
$$\log\left(N^{\gamma}_{p_T>1\ GeV}\right) \sim \boldsymbol{slope}\ \log(N^{\pi})$$
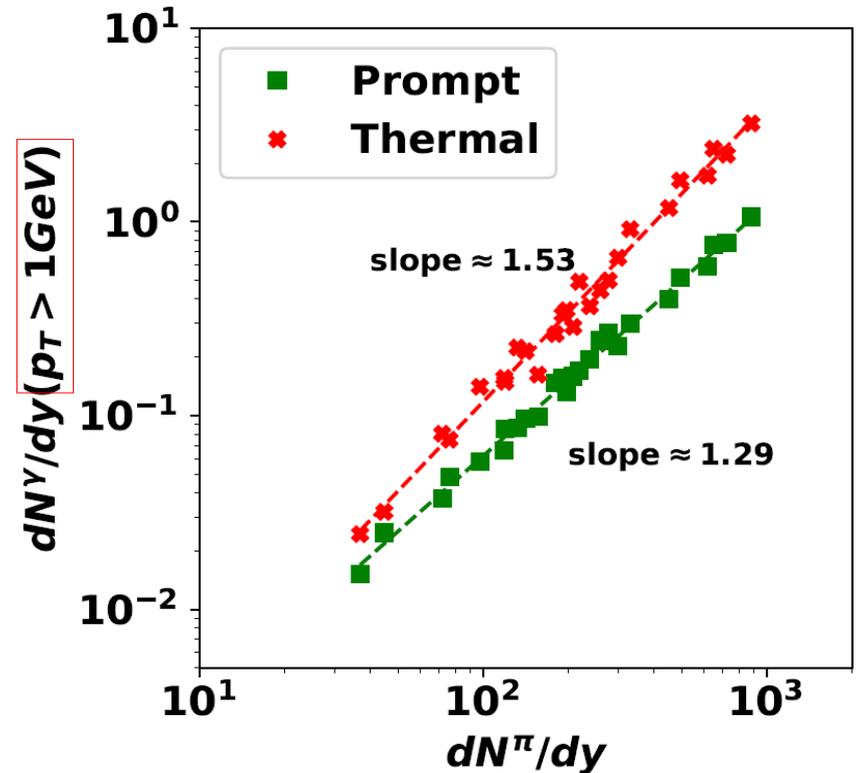
Centrality&$\sqrt{s_{NN}}$ dependence of photon production

Photon multiplicity (above p_T cutoff**)**
vs **hadron multiplicity** (proxy for centrality&$\sqrt{s_{NN}}$)
[fixed & very low cut-off for hadron multiplicity]

PHENIX 2018 [arXiv:1805.04084]
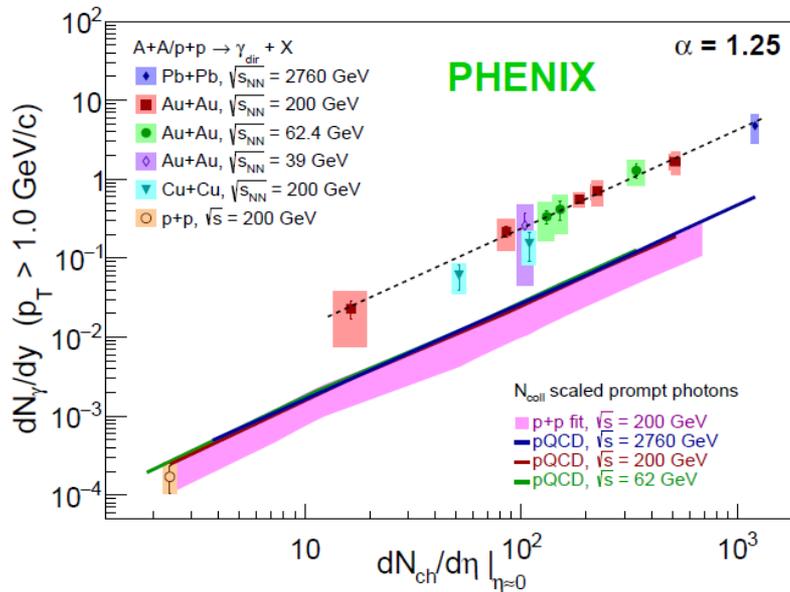
Ref.: Gale, Jeon, McDonald, Paquet, Shen, in preparation.

PHENIX:
Measured photon multiplicity in Pb+Pb, Au+Au and Cu+Cu has **slope**
$\sim\mathbf{1.25}$ vs **charged hadron multiplicity**

Photon vs pion multiplicity using centrality

What does the data say?
(few centralities and large uncertainties)

$$\sqrt{s_{NN}} = 200 \text{ GeV}; \ p_T > p_T^{min}$$

Thermal_Prompt γ

Slopes: 1.49, 1.56, 1.59, 1.60

$\sqrt{s_{NN}} = 200$ GeV only

Prompt

Slopes: 1.40, 1.40, 1.40, 1.40

$\sqrt{s_{NN}} = 200$ GeV only

Ref.: PHENIX PRC 91 (2015)

p_T^{min} (GeV/c)	Slopes vs N_{part}
0.4	$1.36 \pm 0.08 \pm 0.08$
0.6	$1.41 \pm 0.14 \pm 0.12$
0.8	$1.42 \pm 0.07 \pm 0.11$
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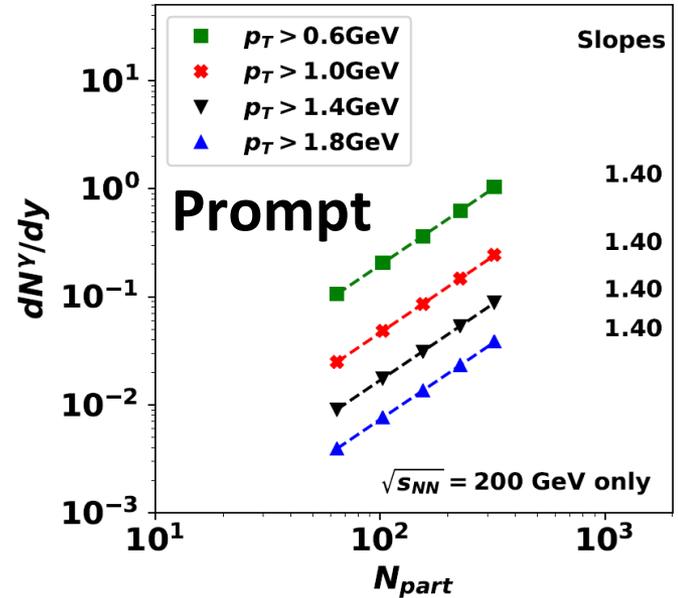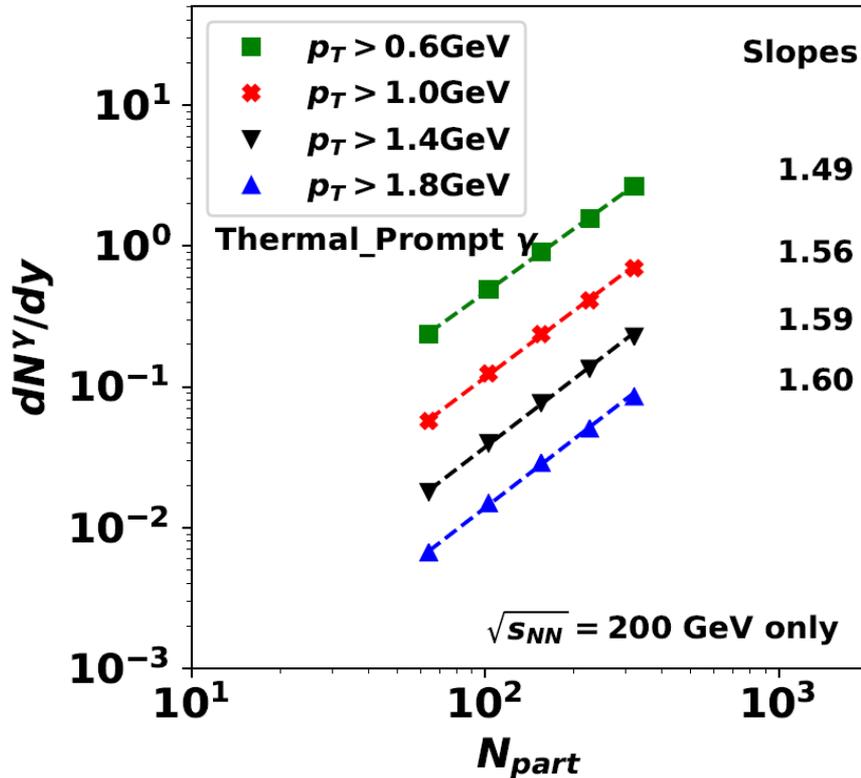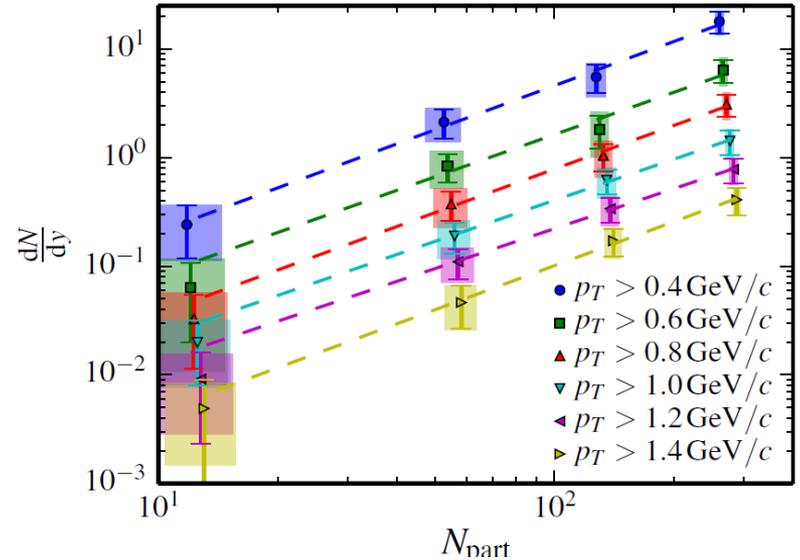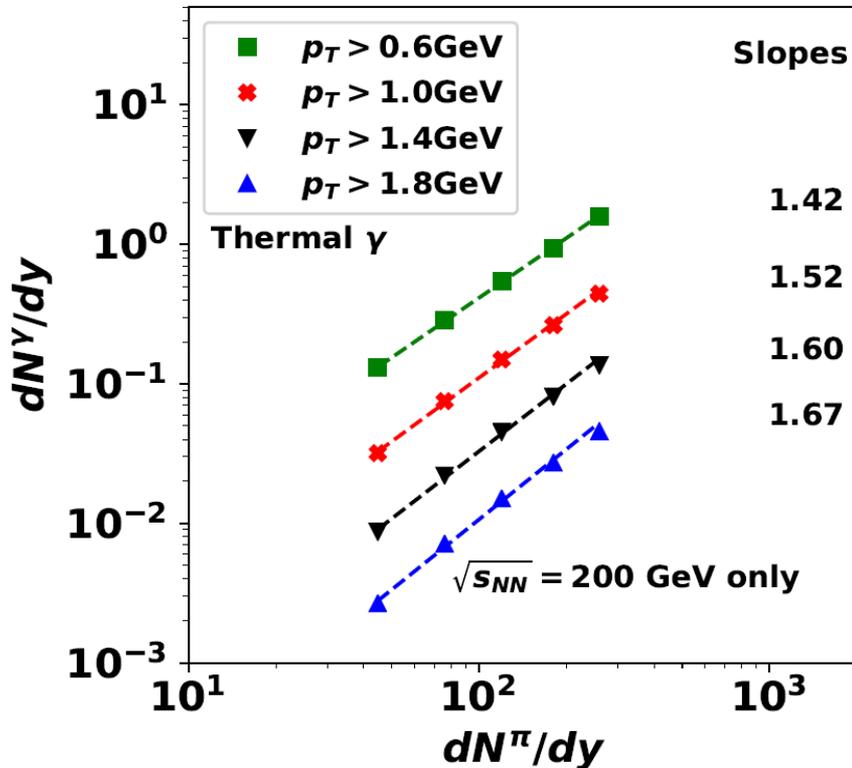Note: STAR also has data [PRB 770 (2017) 451]
but uncertainties too large

Photon vs pion multiplicity using centrality

What does the data say?

Few centralities and large uncertainties

$$\sqrt{s_{NN}} = 200 \text{ GeV}; \ p_T > p_T^{min}$$

Ref.: PHENIX PRC 91 (2015)

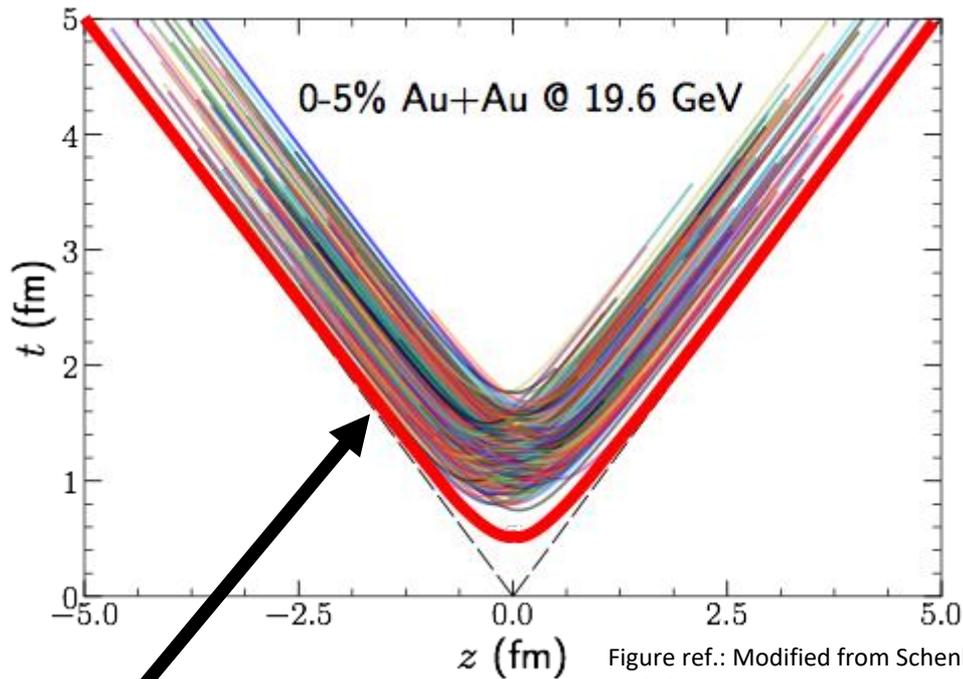| p_T^{min} (GeV/c)	Slopes vs N_{part}
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1.4	$1.40 \pm 0.06 \pm 0.10$

Note: STAR also has data [PRB 770 (2017) 451] but uncertainties too large

Differential spectra & v_2 in RHIC beam energy scan

Energy & baryon number **deposited over extended time** period **at lower** $\sqrt{s_{NN}}$

0-5% Au+Au @ 19.6 GeV

Figure ref.: Modified from Schenke and Shen

"Instantaneous" energy deposition (high collision energy limit)

Dynamic energy deposition
Instantaneous energy deposition

Thermal+prompt
Rapidity y=1

Au-Au 20-30%
$\sqrt{s}$=19.6 GeV

Dynamic energy deposition
Instantaneous energy deposition

Thermal+prompt
Rapidity y=1

Au-Au 20-30%
$\sqrt{s}$=19.6 GeV
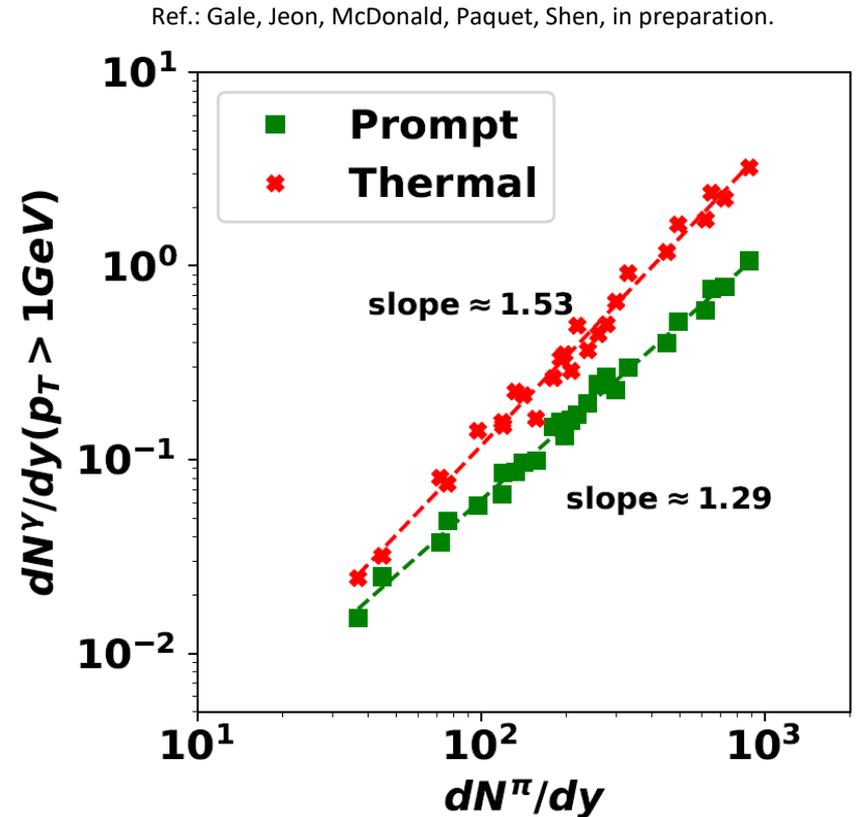
Jean-François Paquet (Duke)
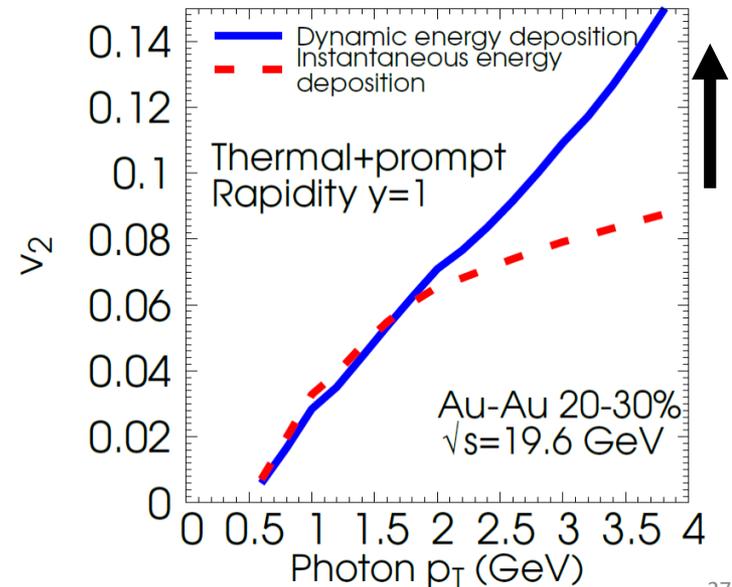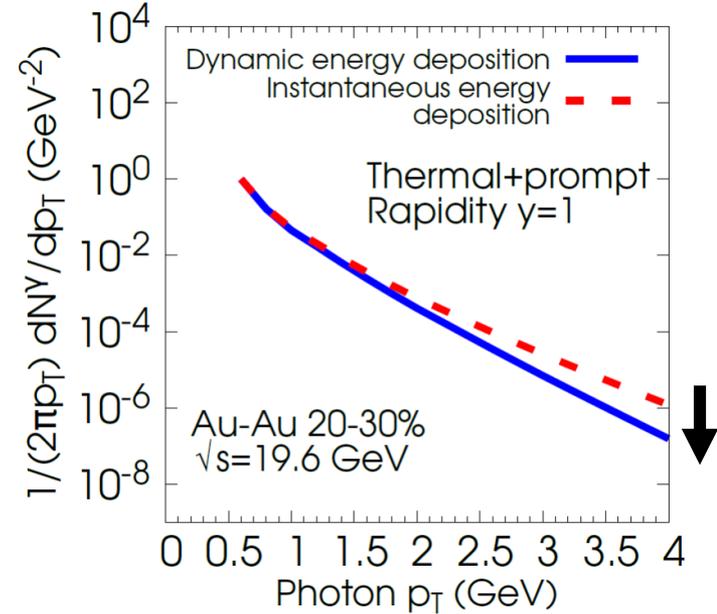
Centrality& $\sqrt{s_{NN}}$ dependence of photon production

Another approach: integrate photon spectra dN/dp_T **above** p_T **cutoff &
plot with respect to hadron multiplicity** (proxy for centrality& $\sqrt{s_{NN}}$)

Ref.: QM2015 proceedings [1704.07842]

Ref.: Gale, Jeon, McDonald, Paquet, Shen, in preparation.
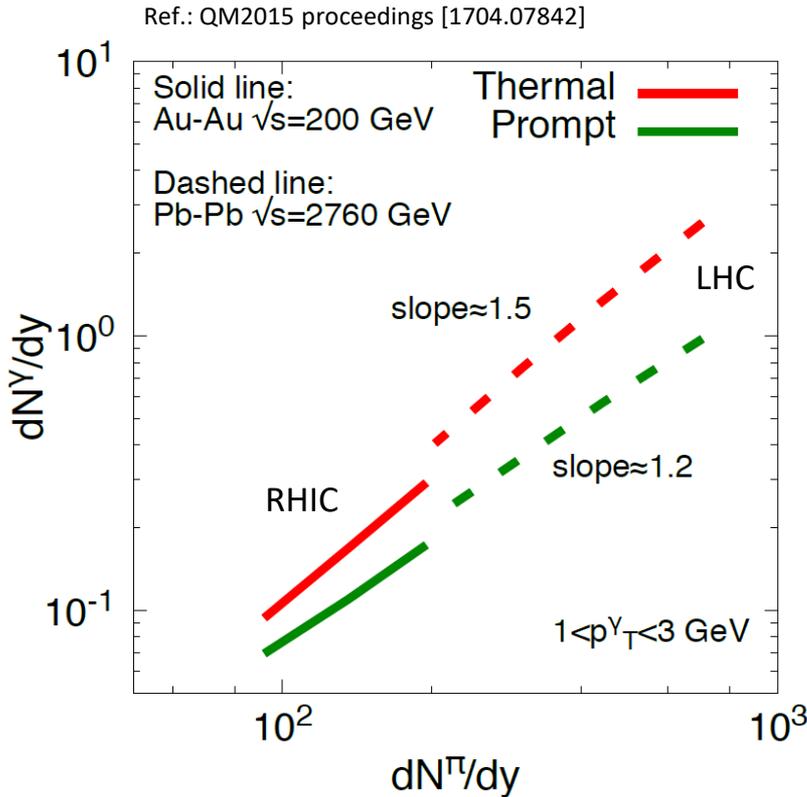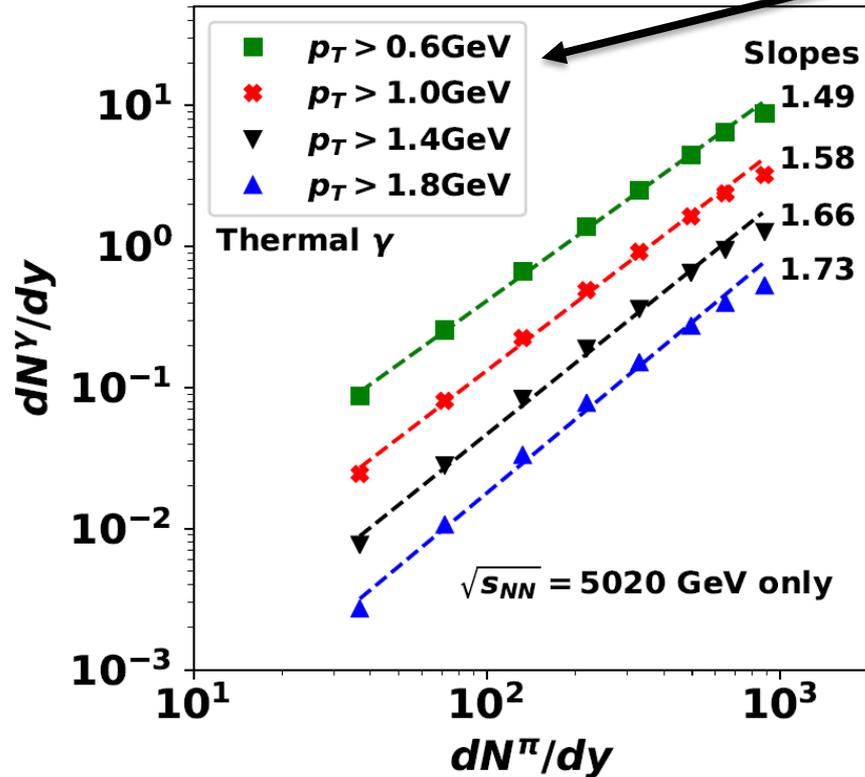

RHIC $\sqrt{s_{NN}}$ =200 GeV
LHC $\sqrt{s_{NN}}$ =2760 GeV

RHIC $\sqrt{s_{NN}}$ =19.6, 39, 62.4 & 200 GeV
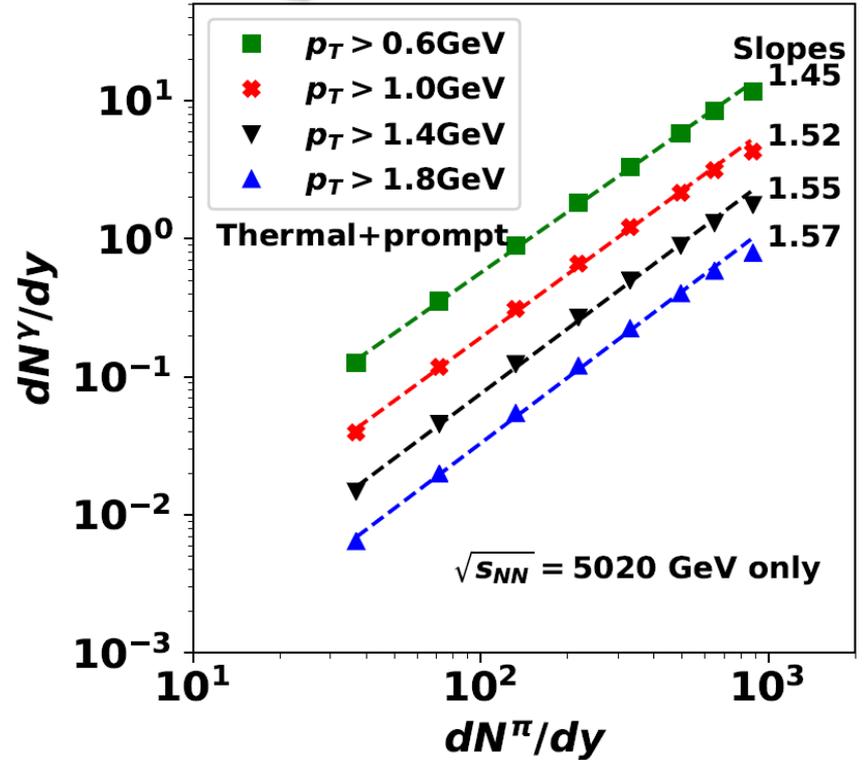LHC $\sqrt{s_{NN}}$ =2760 & 5020 GeV

Photon vs pion multiplicity using centrality

$$\sqrt{s_{NN}} = 5020 \text{ GeV}; \ p_T > p_T^{min}$$

Thermal photons

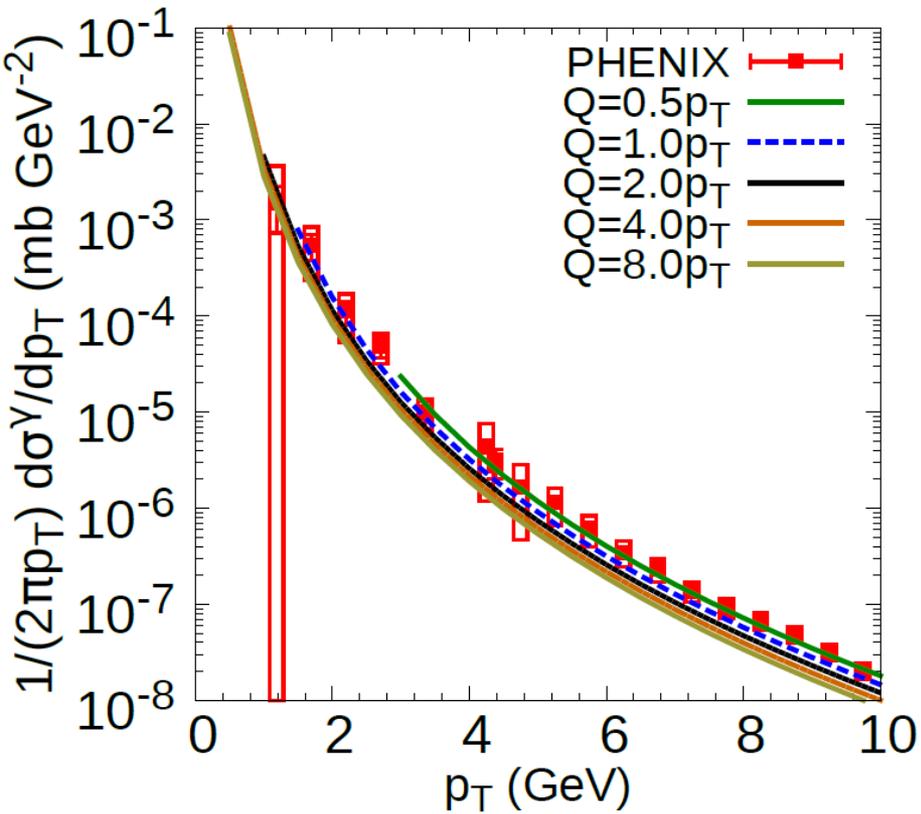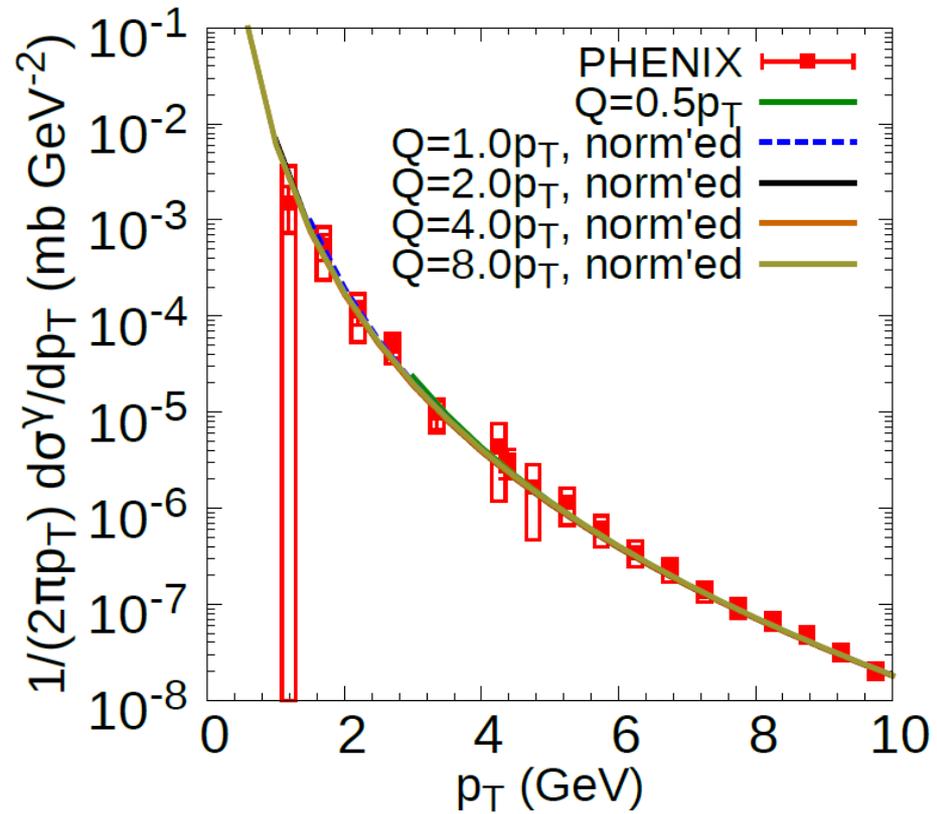**Thermal+prompt photons**

Prompt photon slope: 1.35

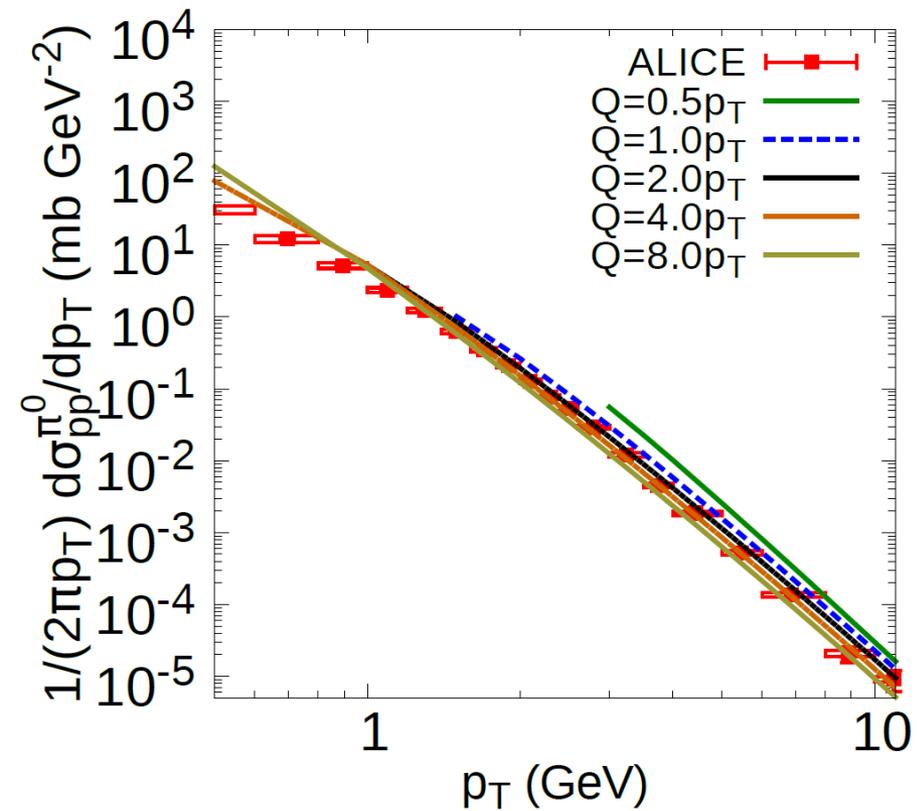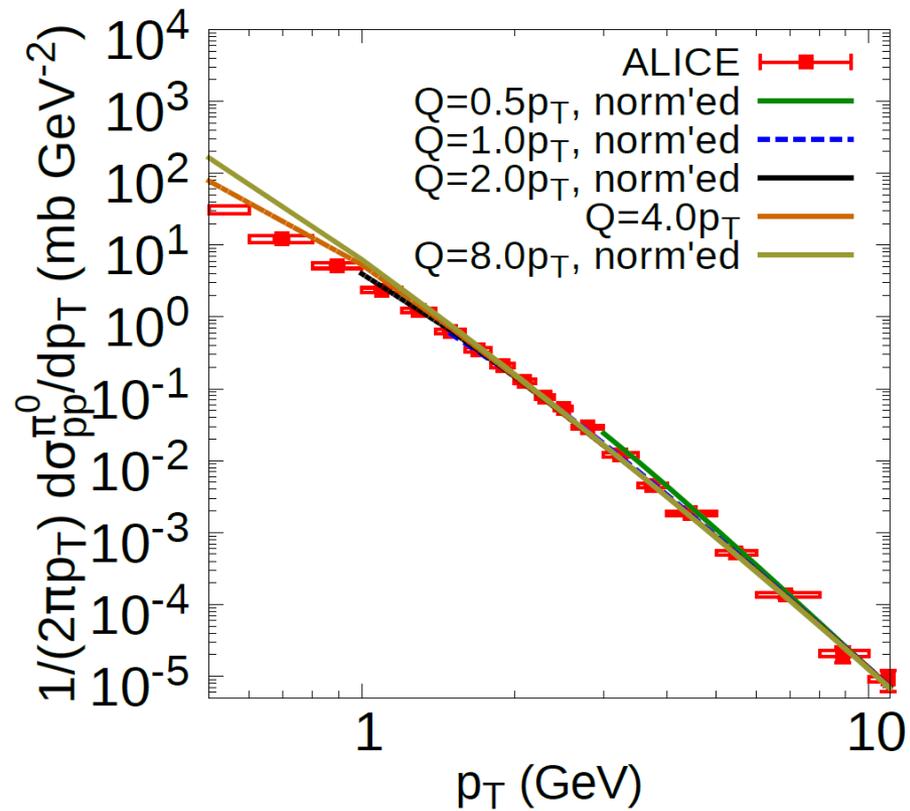Simple test: p_T **cut-off dependence** for **thermal photons**, **NOT** for prompt

Prompt photons

Figure A.1: Direct photon spectrum measured in $\sqrt{s_{NN}} = 200$ GeV proton-proton collisions at RHIC compared with (a) perturbative QCD calculations made with different scales Q (b) normalised perturbative QCD calculations (Equation A.1)
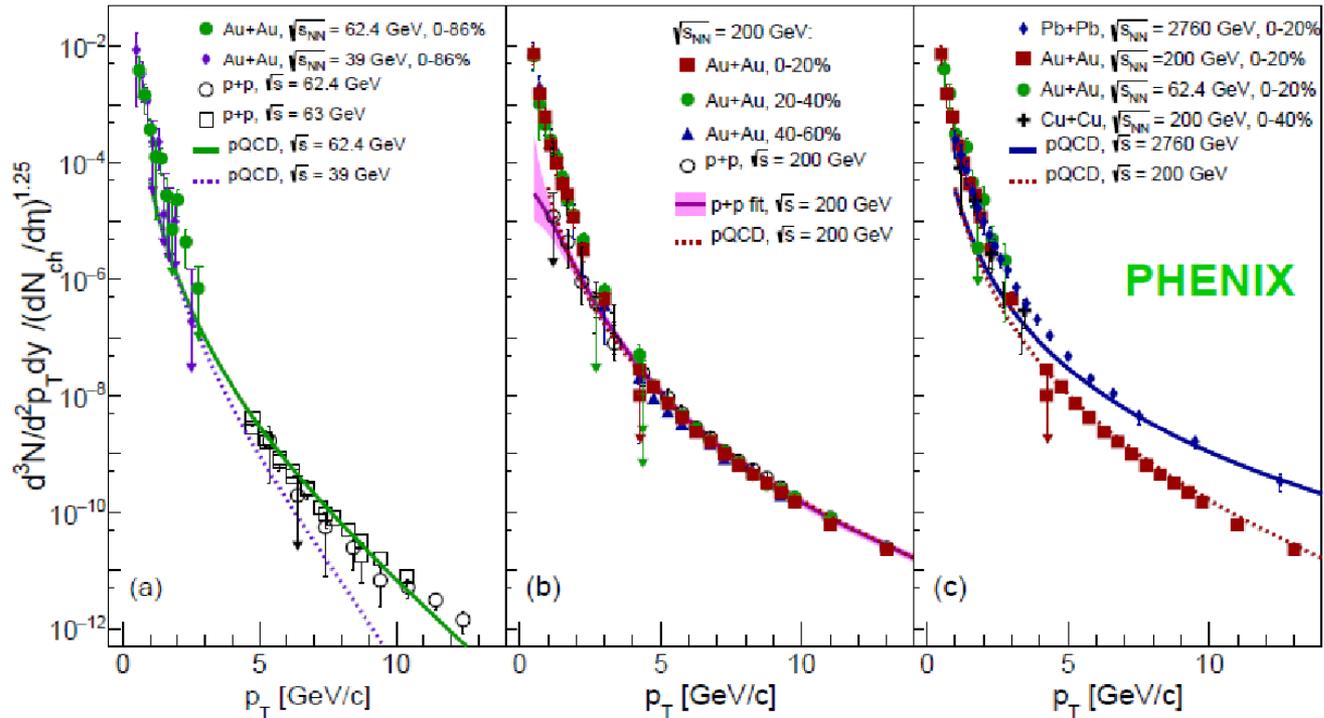
Figure A.2: Same as Figure A.1 for pions measured in $\sqrt{s_{NN}} = 2.76$ TeV proton-proton collisions at the LHC

PHENIX collaboration

Beam-energy and centrality dependence of direct-photon emission from ultra-relativistic heavy-ion collisions
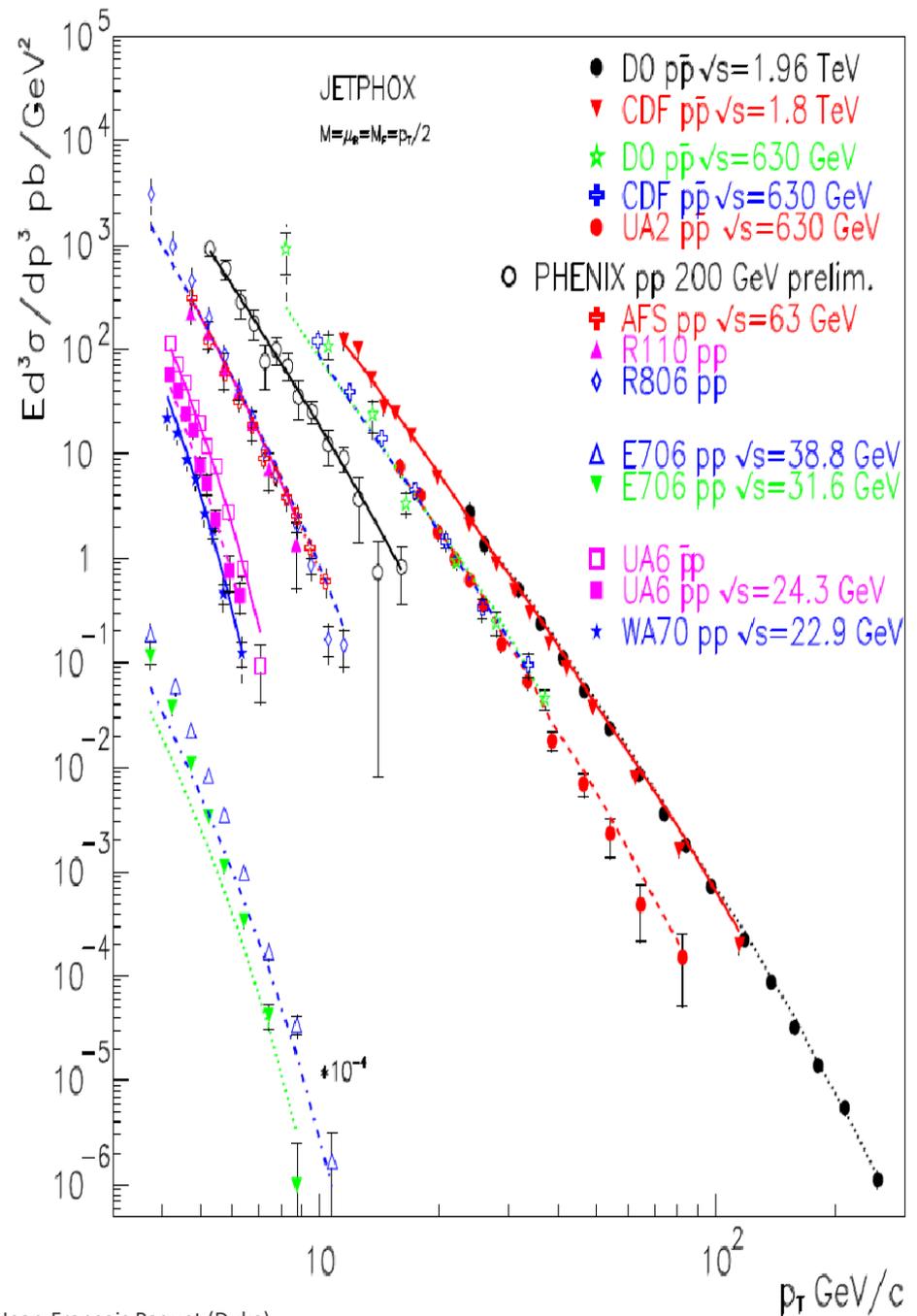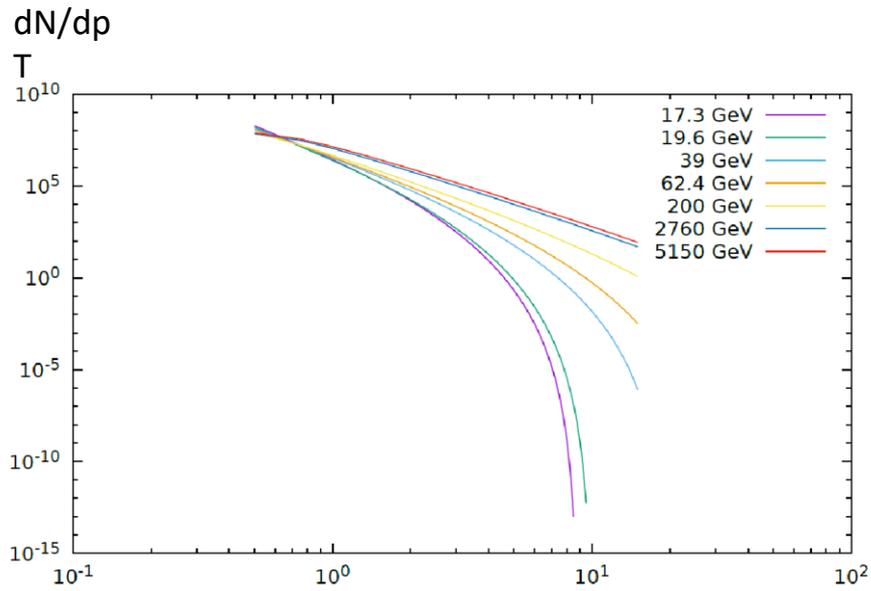
Aurenche et al: prompt photons in p+p

A NEW CRITICAL STUDY OF PHOTON
PRODUCTION IN HADRONIC COLLISIONS

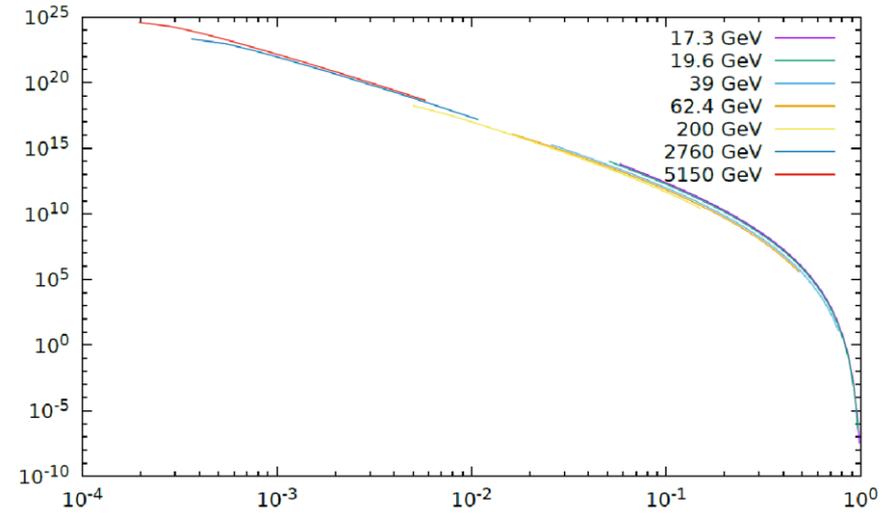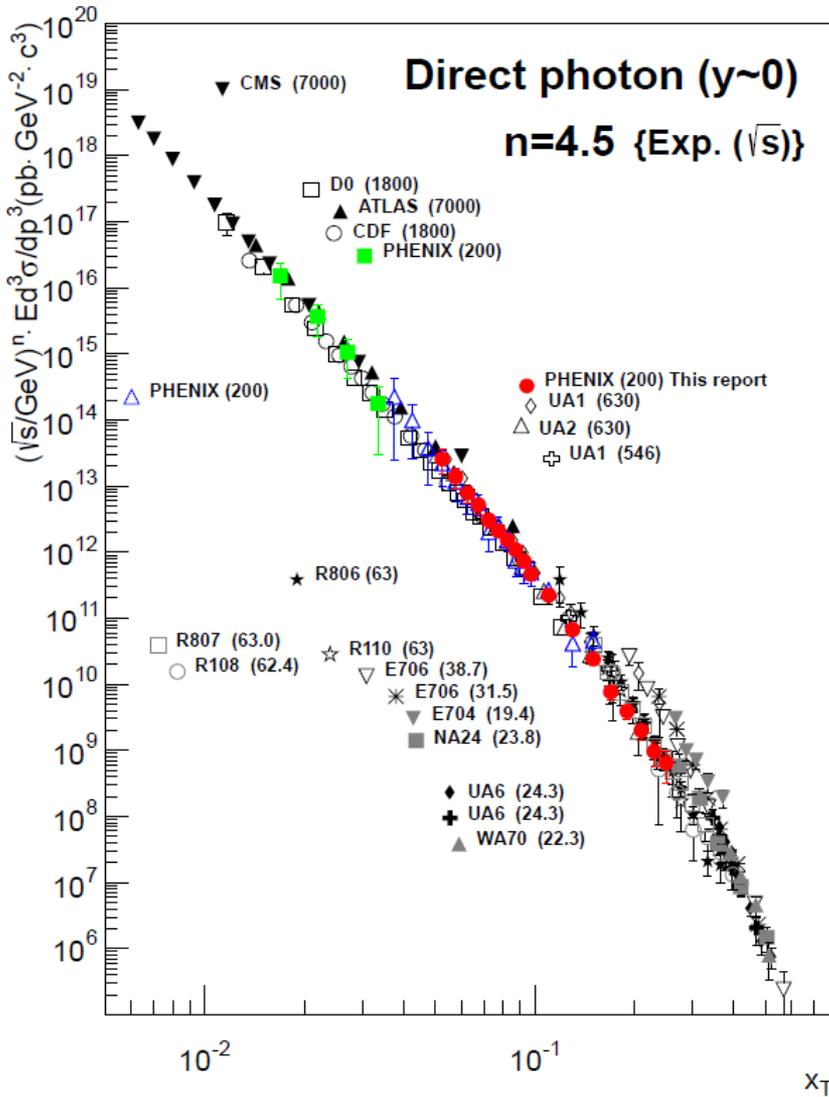P. Aurenche[1], M. Fontannaz[2], J.Ph. Guillet[1], E. Pilon[1], M. Werlen[1]

Prompts

dN/dp
T

Sqrt(s)^4.5*dN/dpT

pT

xT

Prompts

Sqrt(s)^4.5*dN/dpT

xT

PHENIX, PRD86 (2012) 072008   Jean-François Paquet (Duke)
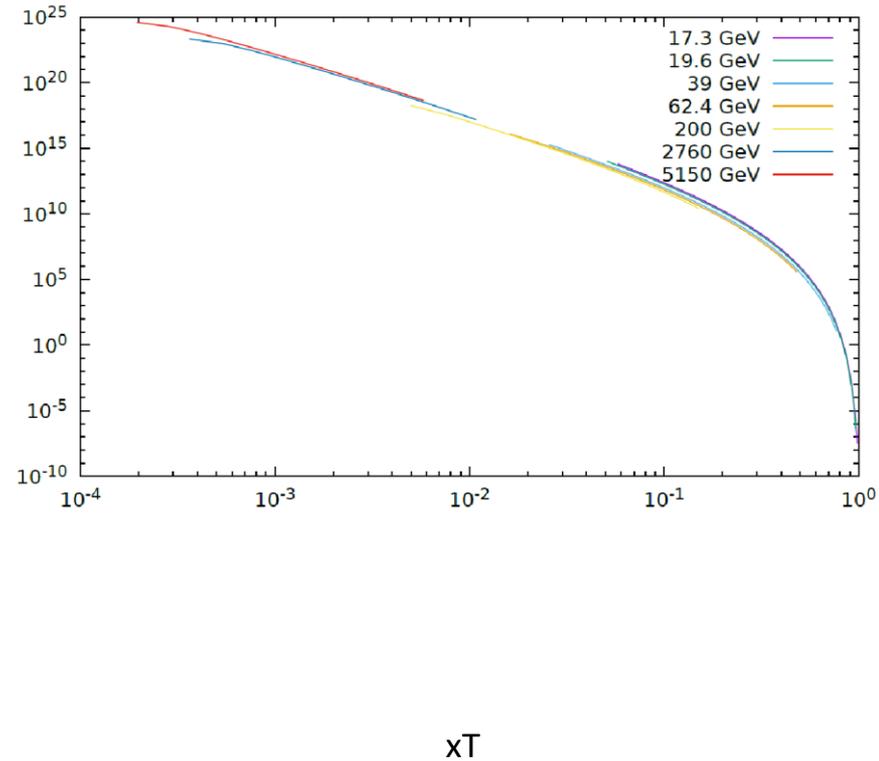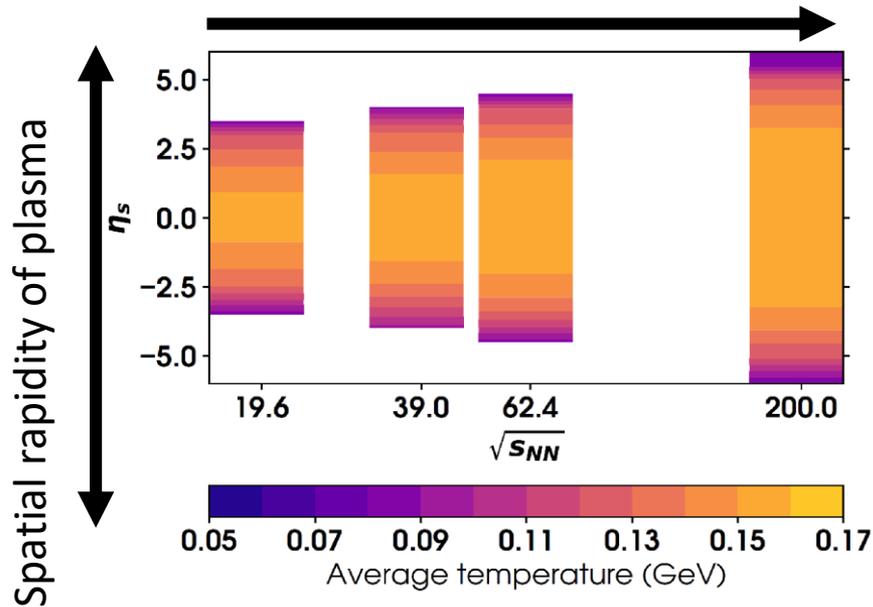
More about photons in BES

Medium properties for Au+Au at low $\sqrt{s_{NN}}$

Center-of-mass energy of collision

Center-of-mass energy of collision

Temperature averaged over the plasma's lifetime
(depends on freeze-out energy density)
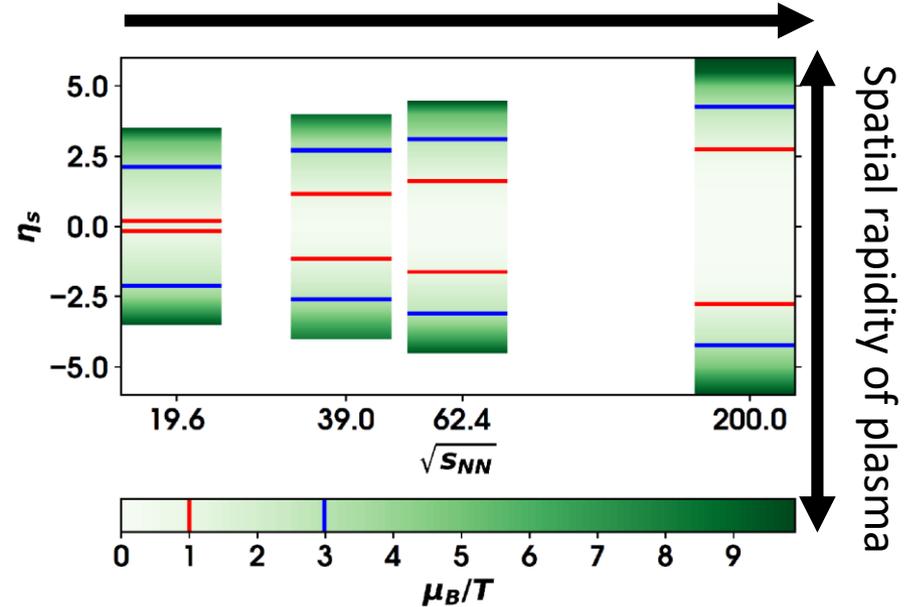
Baryon chemical potential to temperature ratio

(markers at μ_B/T=1 and μ_B/T= 3 to guide the eye)

Spacetime description at low $\sqrt{s_{NN}}$

Thermal photons:
radiated by the hot plasma

$$\frac{dN_\gamma}{d^3 k} = \int d^4 X \frac{d\Gamma_\gamma}{d^3 k}(k, T(X), u^\mu(X), \dots)$$

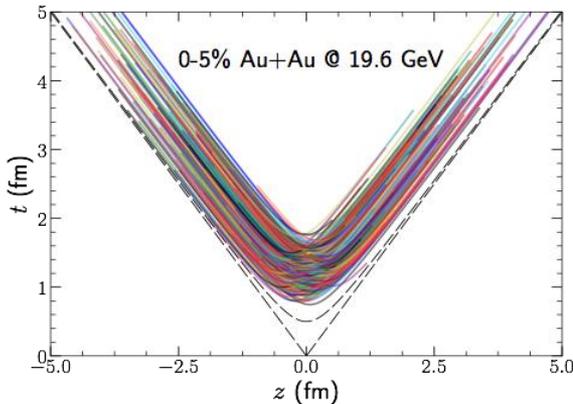**Spacetime profile of plasma from hydrodynamic simulation**

0-5% Au+Au @ 19.6 GeV

Spacetime evolution of plasma described with hydrodynamic model

See **Chun Shen's talk on Wednesday at 11h30**
(Collective dynamics III Session)

- Viscous 3+1D hydrodynamics with **conserved baryon current**   Ref.: Denicol, Gale, Jeon, Monnai, Schenke, Shen (2018)

- Finite μ_B equation of state

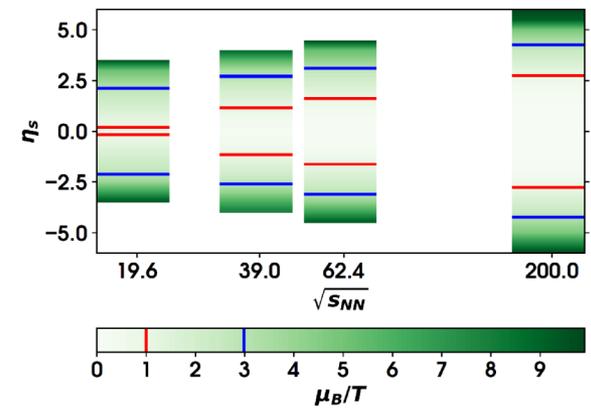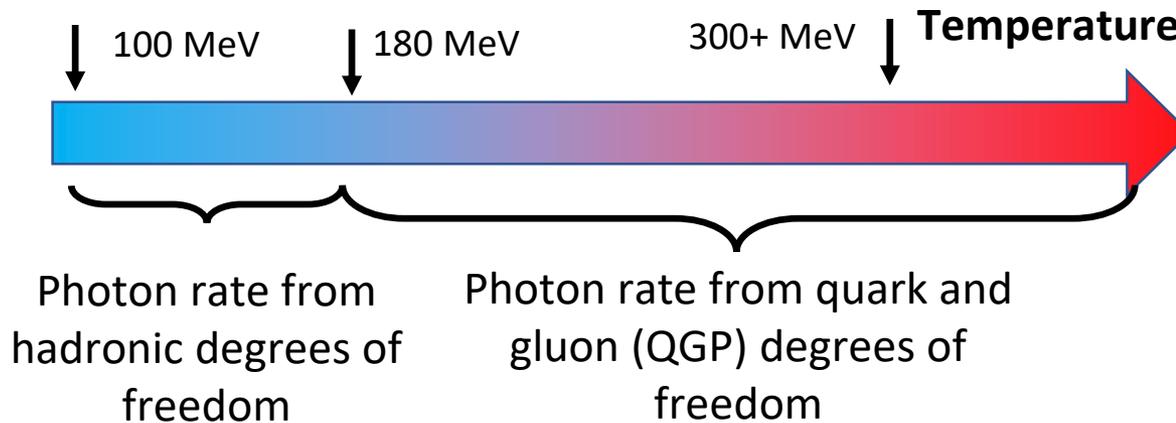- **Dynamical initialization** of hydrodynamics over extended period of time (at small $\sqrt{s_{NN}}$) using sources:
$$\partial_\mu T^{\mu\nu} = S^\nu$$
$$\partial_\mu J_B^\mu = \rho_B$$
Ref.: Shen & Schenke(2018)

- Freeze-out at **constant energy density** 0.16 GeV/fm^3

- Tuned to soft hadron data

Photon thermal rates

100 MeV   180 MeV   300+ MeV   **Temperature**

Photon rate from hadronic degrees of freedom

Photon rate from quark and gluon (QGP) degrees of freedom

Baryon chemical potential

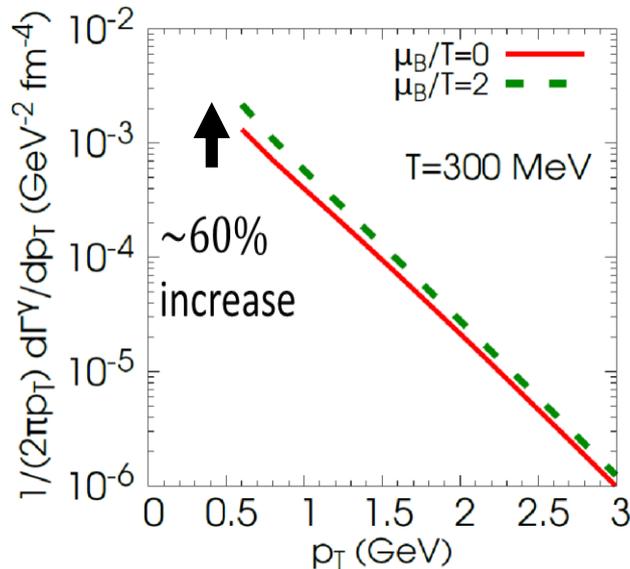**QGP rate:** Compton scattering, $q\bar{q}$ annihilation & bremsstrahlung (with LPM) at finite μ_B

Refs.: Traxler, Vija, Thoma (1995); Gervais, Jeon (2012); This work

Hadronic rates: meson interaction, baryon interaction (at finite μ_B)
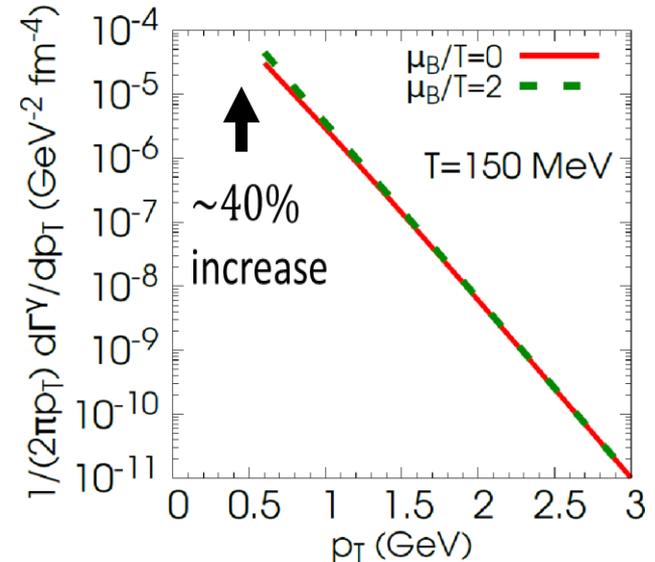
Refs.: Turbide, Rapp, Gale (2004); Heffernan, Hohler, Rapp (2014); Holt, Hohler, Rapp (2016)

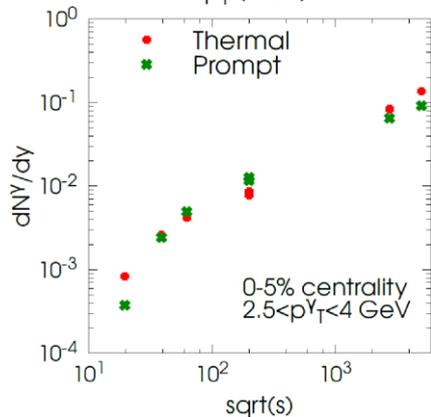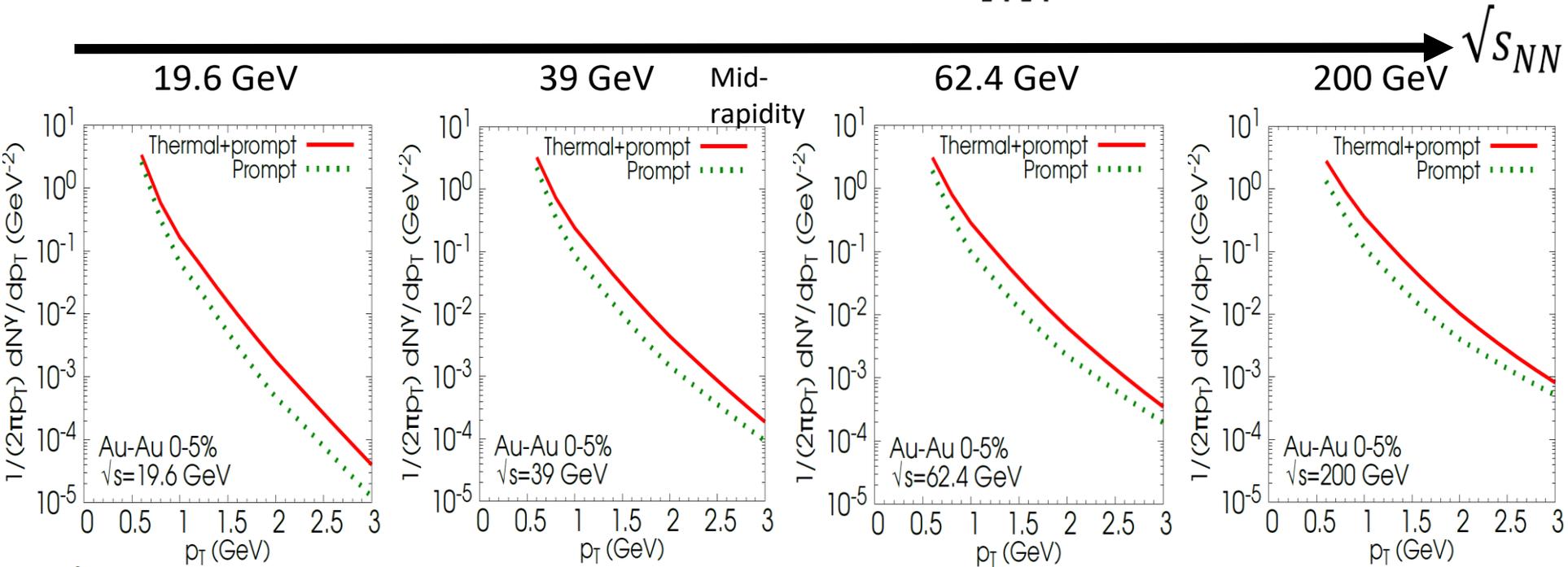# Effect of finite μ_B on thermal photon rate

QGP photon rate

Hadronic photon rate

⬆ Baryon chemical potential μ_B **increases the photon rate at** <u>**low momenta**</u>

$$e^{-\left(\frac{E - \mu_B}{T}\right)}$$

Photon production in low $\sqrt{s_{NN}}$ collisions

$\sqrt{s_{NN}}$ →

19.6 GeV	39 GeV	Mid-rapidity	62.4 GeV	200 GeV

High p_T thermal photon visible against prompt photons estimate as $\sqrt{s_{NN}}$ is decreased

Prompt photon calculations difficult at low p_T and $\sqrt{s_{NN}}$: **p+p measurements valuable**

Photon v_2 in low $\sqrt{s_{NN}}$ collisions

$$\xrightarrow{\hspace{10cm}} \sqrt{s_{NN}}$$

| 19.6 GeV | 39 GeV | Mid-rapidity | 62.4 GeV | 200 GeV |
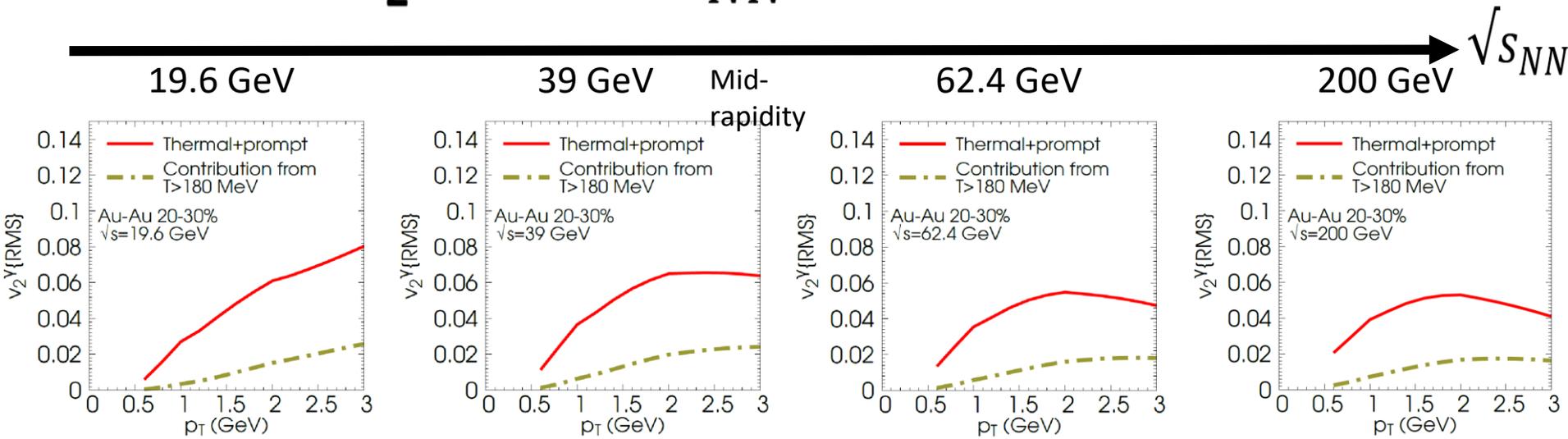
Significant v_2 across $\sqrt{s_{NN}}$

Larger photon v_2

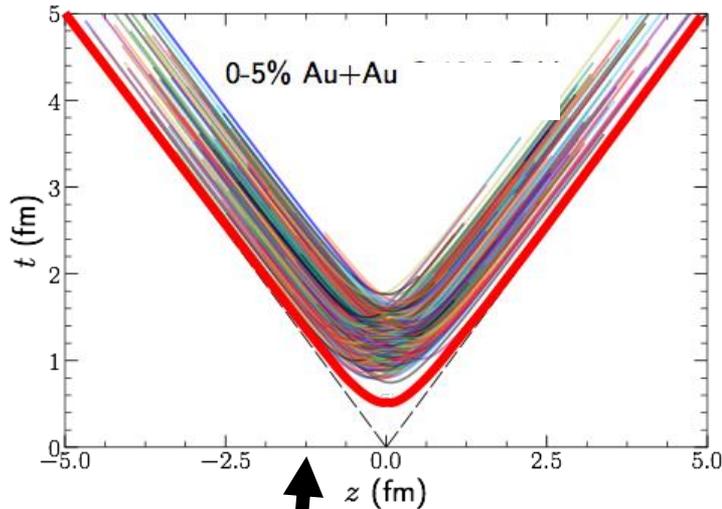**Higher fraction of v_2 from high T photons**

$$v_n^\phi = \frac{\left\langle \dfrac{dN^{th,T>180\,MeV}}{d\phi}\right\rangle_\phi v_n^{th,T>180\,MeV} + \left\langle \dfrac{dN^{th,T<180\,MeV}}{d\phi}\right\rangle_\phi v_n^{th,T<180\,MeV}}{\left\langle \dfrac{dN^{th,T>180\,MeV}}{d\phi}\right\rangle_\phi + \left\langle \dfrac{dN^{th,T<180\,MeV}}{d\phi}\right\rangle_\phi + \left\langle \dfrac{dN^{prompt}}{d\phi}\right\rangle_\phi}$$
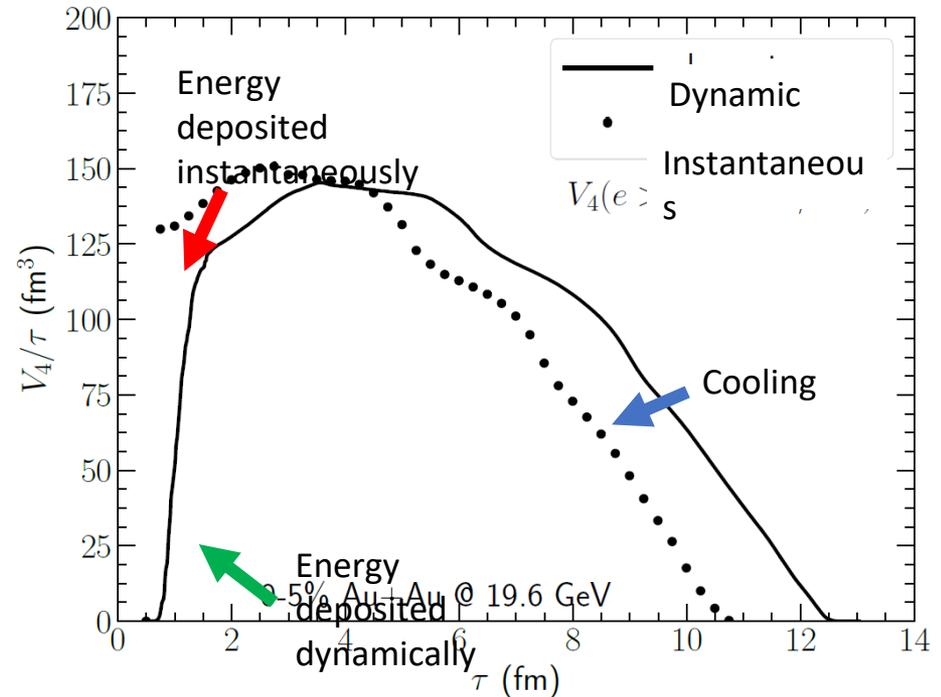
Probing the dynamical plasma

Energy & baryon number **deposited over extended time** period **at lower** $\sqrt{s_{NN}}$
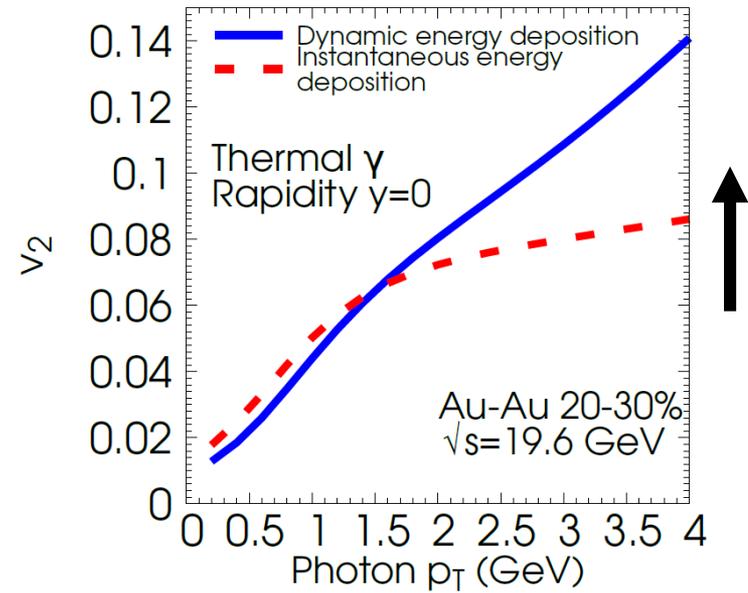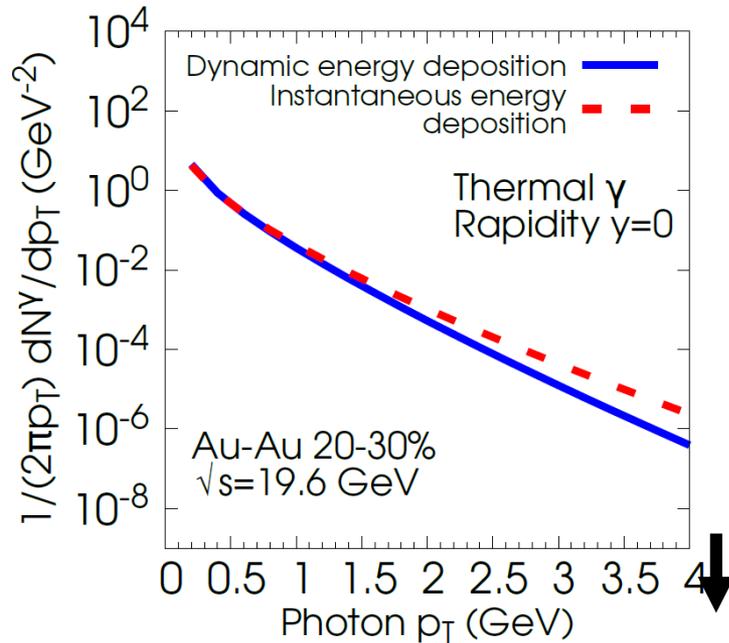

$\sqrt{s_{NN}}$ =19.6 GeV

"Instantaneous" energy deposition
(high collision energy limit)

Space-time 4-volume (above freeze-out) versus time

Energy deposited instantaneously

Dynamic

Instantaneous

$V_4(e$

Cooling

Energy deposited dynamically

0-5% Au+Au @ 19.6 GeV

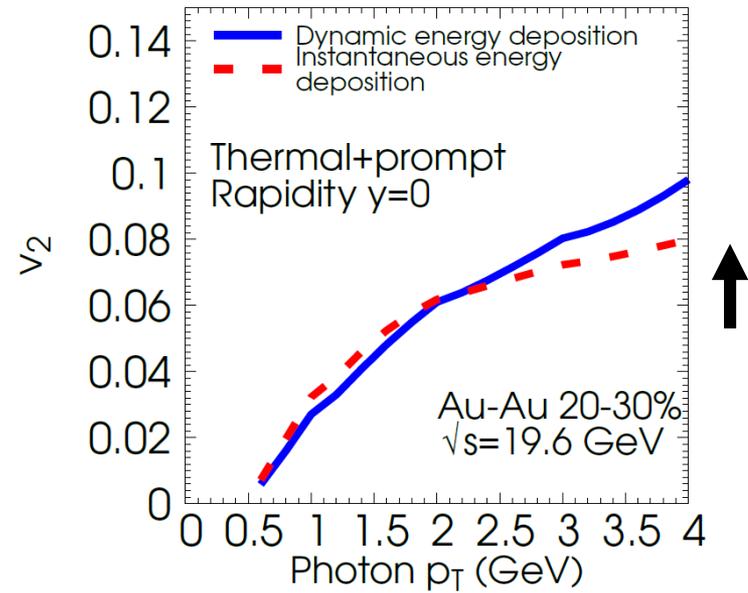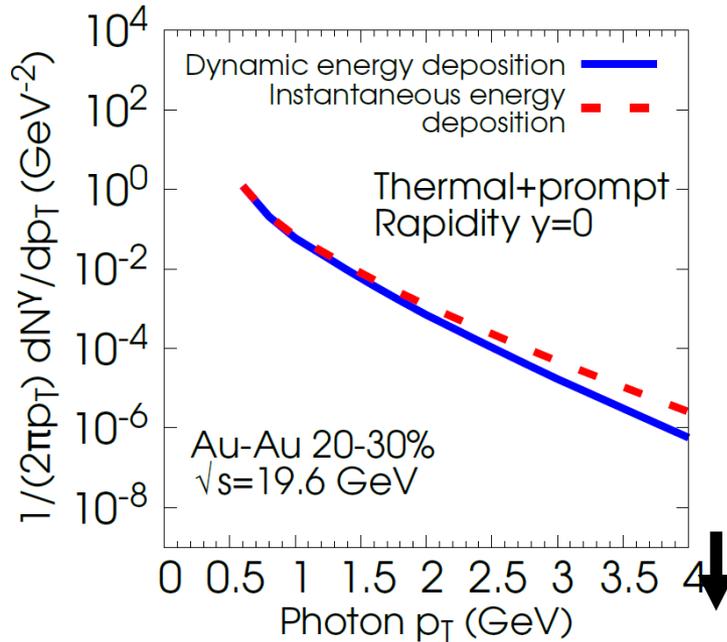# Early plasma dynamics at low $\sqrt{s_{NN}}$ (19.6 GeV)

Thermal photons

Effect of energy deposition (instantaneous vs dynamic) visible in photons

At low $\sqrt{s_{NN}}$ (19.6 GeV): thermal+prompt
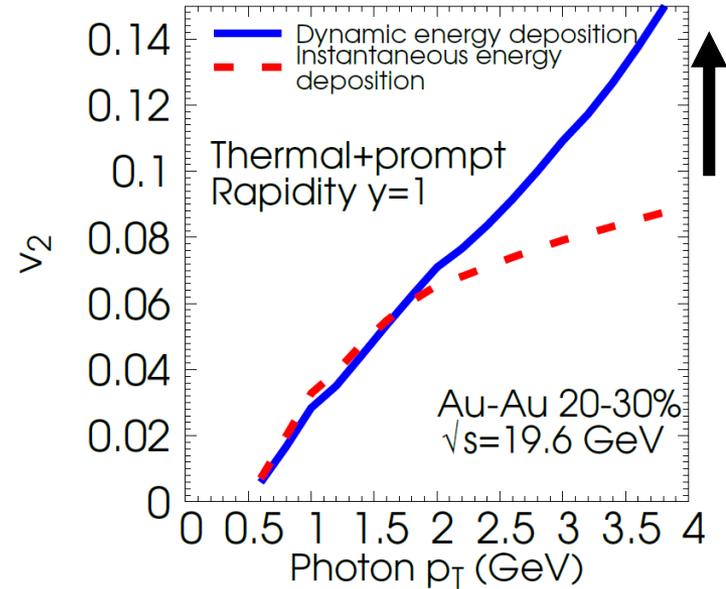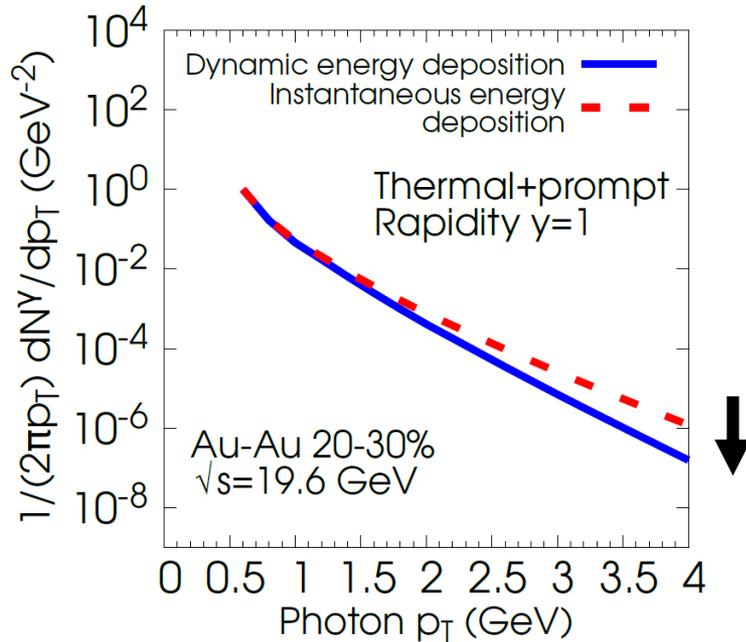
Thermal+prompt photons

Effect of energy deposition (instantaneous vs dynamic) visible in photons

At low $\sqrt{s_{NN}}$ (19.6 GeV): away from mid-rapidity
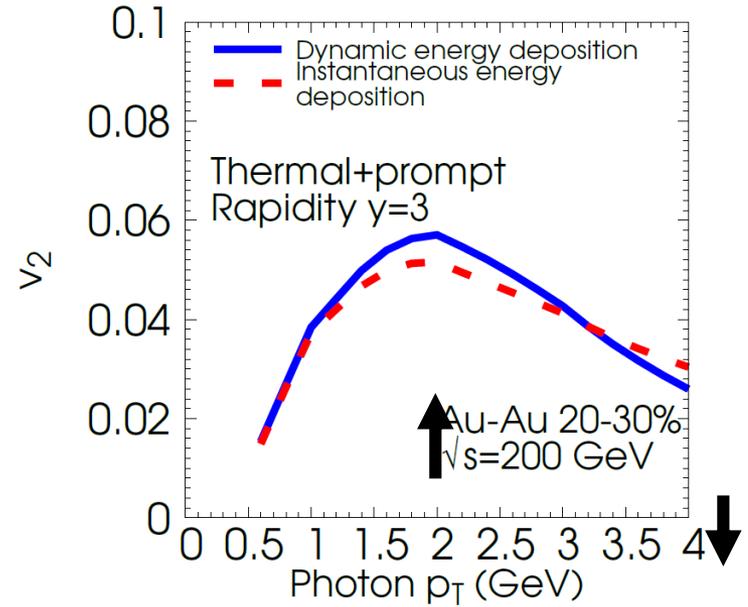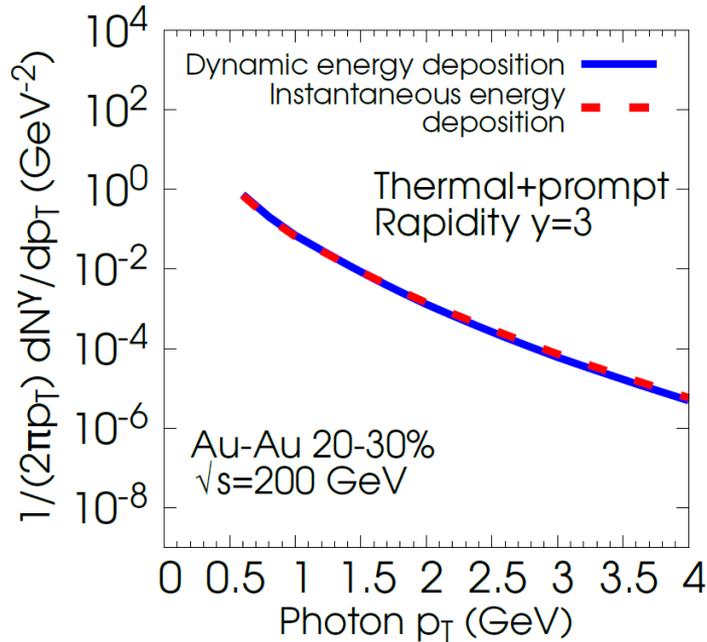
Thermal+prompt photons

Effect of energy deposition (instantaneous vs dynamic) visible in photons

At high $\sqrt{s_{NN}}$ (200 GeV) & high rapidity

Thermal+prompt photons

At high $\sqrt{s_{NN}}$, effect of dynamic energy deposition visible at **higher rapidity**

Summary

- **Significant** thermal **photon signal** at low $\sqrt{s_{NN}}$

- Small effect of μ_B on photon rate;
  dominant effect is distinct spacetime profile of medium

- **Early initial conditions** can be studied
  with photons using
  - Low $\sqrt{s_{NN}}$ collisions
  - Rapidity $y \gtrsim 2$ for higher $\sqrt{s_{NN}}$ collisions
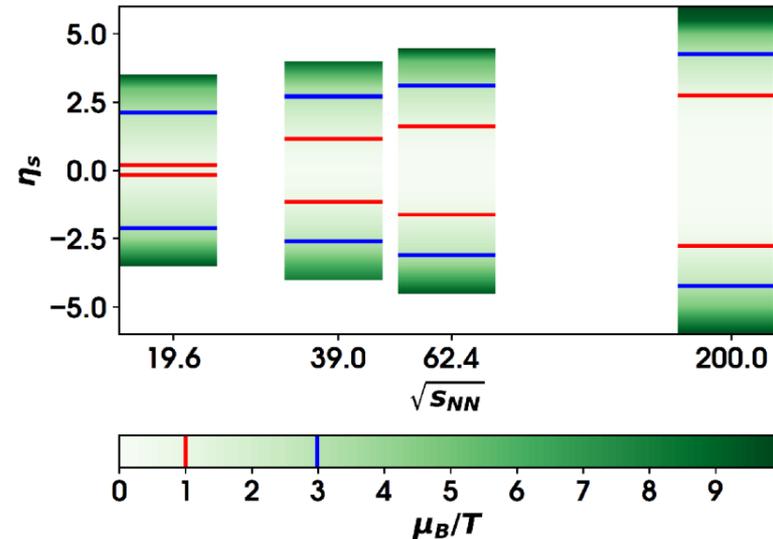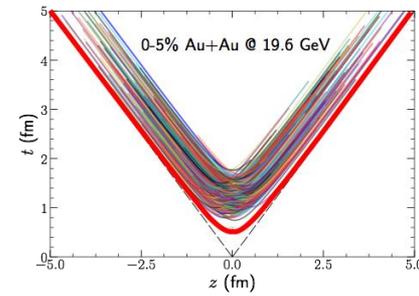
Photons are unique **direct** probes
of complex dynamics of
low $\sqrt{s_{NN}}$ collisions

Contributions to v_n

Simple derivation

$$v_n = \frac{\int d\phi \frac{dN}{d\phi} e^{in(\phi - \Psi_n)}}{\int d\phi \frac{dN}{d\phi}}$$

$$\frac{dN}{d\phi} = \frac{dN^{th,T>180\,MeV}}{d\phi} + \frac{dN^{th,T<180\,MeV}}{d\phi} + \frac{dN^{prompt}}{d\phi}$$

$$v_n = \frac{\int d\phi \left(\frac{dN^{th,T>180\,MeV}}{d\phi} + \frac{dN^{th,T<180\,MeV}}{d\phi} + \frac{dN^{prompt}}{d\phi} \right) e^{in(\phi - \Psi_n)}}{\int d\phi \left(\frac{dN^{th,T>180\,MeV}}{d\phi} + \frac{dN^{th,T<180\,MeV}}{d\phi} + \frac{dN^{prompt}}{d\phi} \right)}$$

$$v_n = \frac{\int d\phi \frac{dN^{th,T>180\,MeV}}{d\phi} e^{in(\phi - \Psi_n)} + \int d\phi \frac{dN^{th,T<180\,MeV}}{d\phi} e^{in(\phi - \Psi_n)}}{\int d\phi \left(\frac{dN^{th,T>180\,MeV}}{d\phi} + \frac{dN^{th,T<180\,MeV}}{d\phi} + \frac{dN^{prompt}}{d\phi} \right)}$$

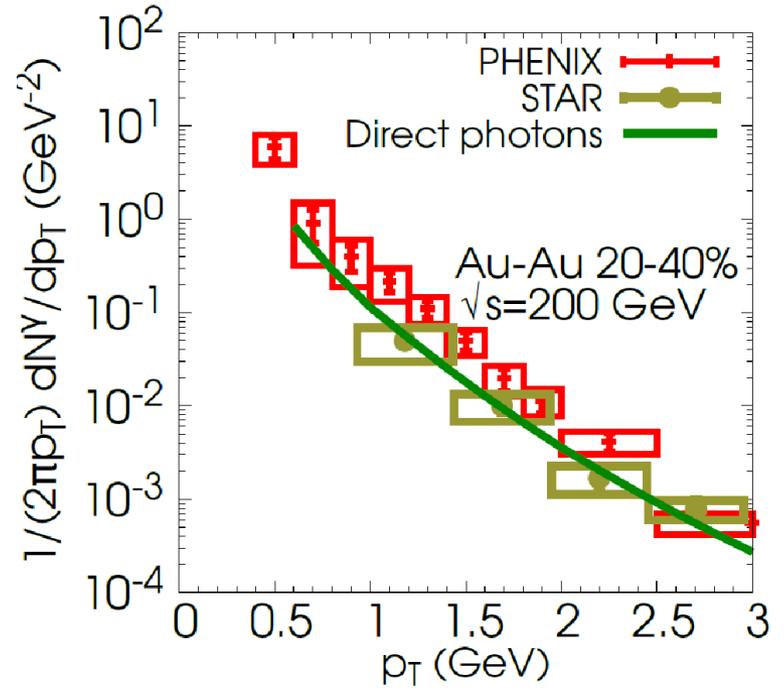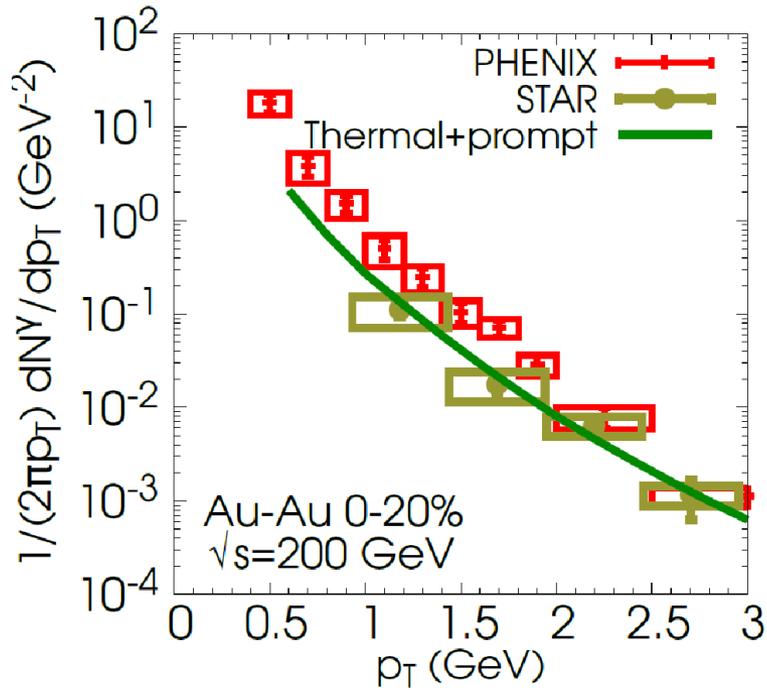$$v_n = \frac{\left\langle \frac{dN^{th,T>180\,MeV}}{d\phi} \right\rangle_\phi v_n^{th,T>180\,MeV} + \left\langle \frac{dN^{th,T<180\,MeV}}{d\phi} \right\rangle_\phi v_n^{th,T<180\,MeV}}{\left\langle \frac{dN^{th,T>180\,MeV}}{d\phi} \right\rangle_\phi + \left\langle \frac{dN^{th,T<180\,MeV}}{d\phi} \right\rangle_\phi + \left\langle \frac{dN^{prompt}}{d\phi} \right\rangle_\phi}$$
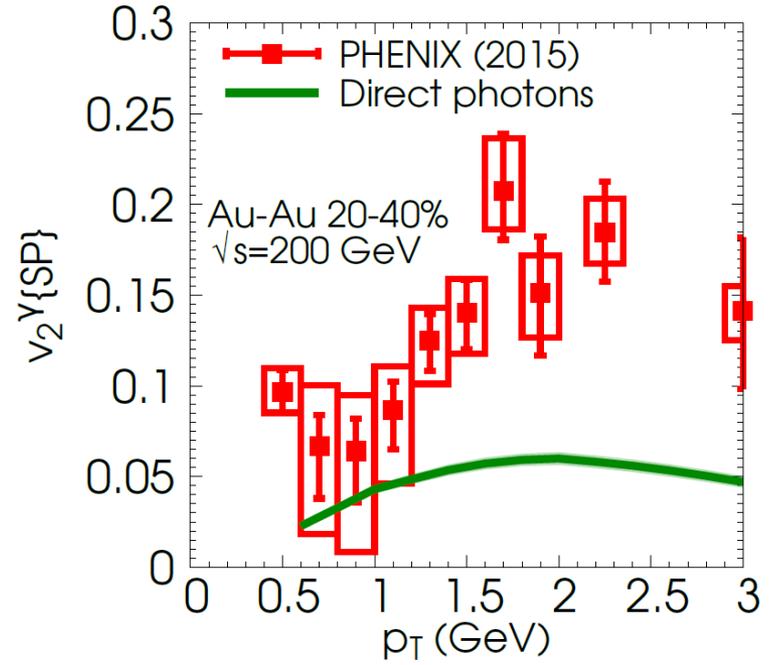
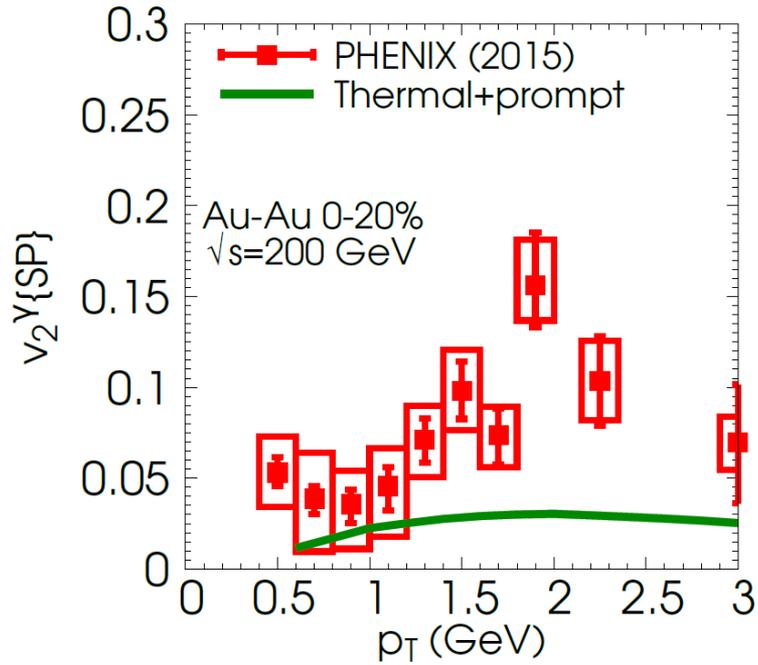$$\langle ... \rangle_\phi = \frac{1}{2\pi} \int d\phi \; ...$$

Comparison with available data

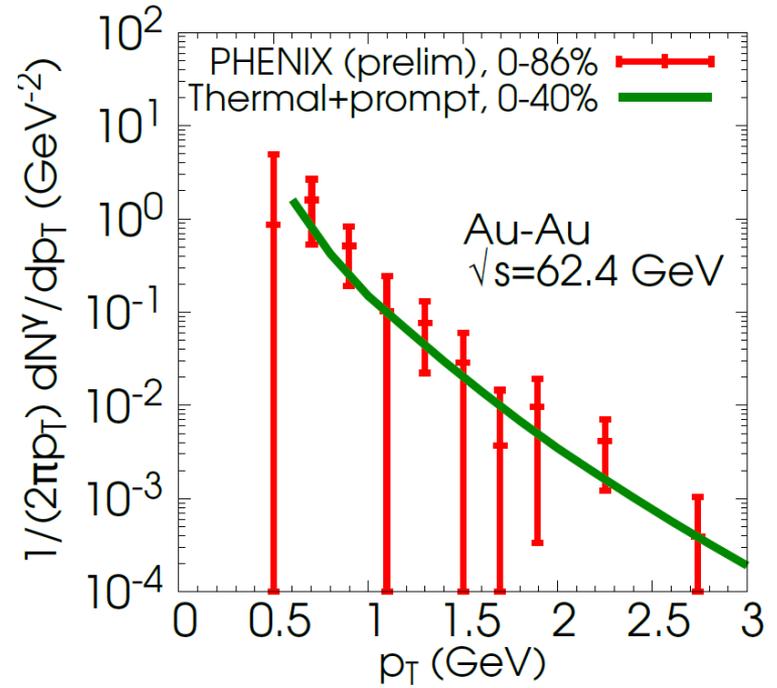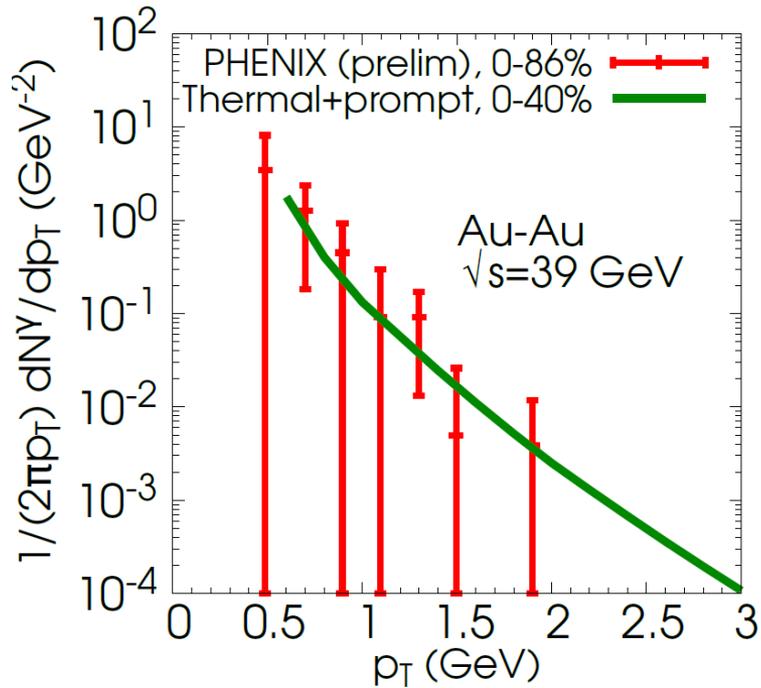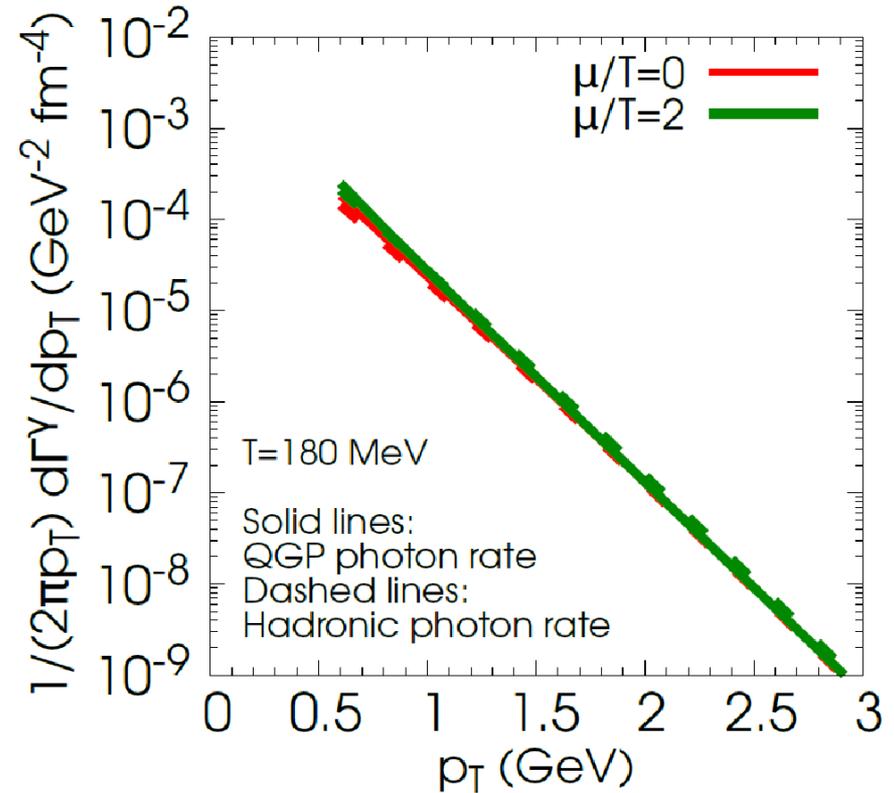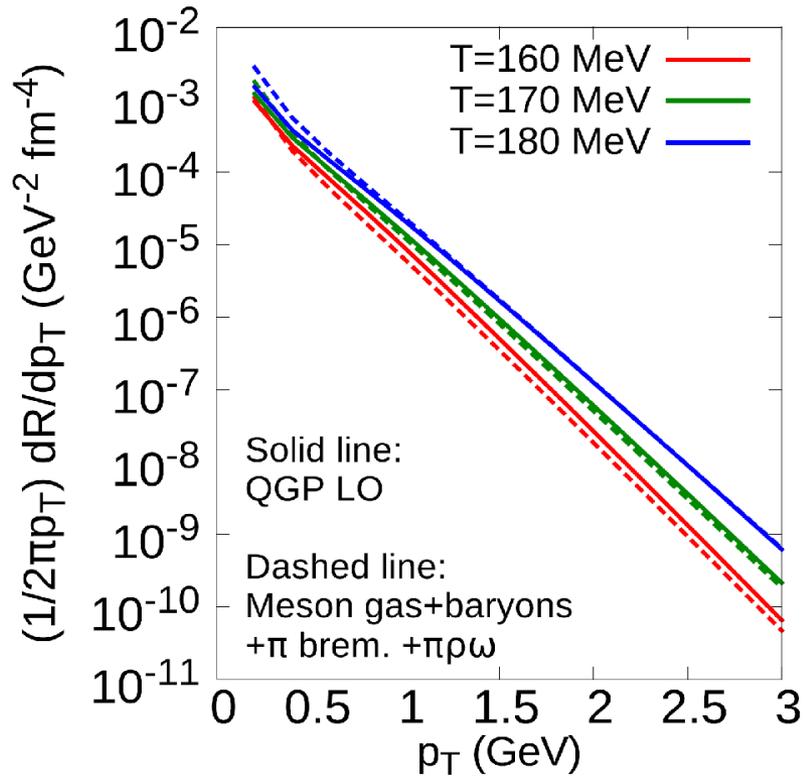# Spectra at 200 GeV

v₂ at 200 GeV

Spectra at 39 & 62.4 GeV:
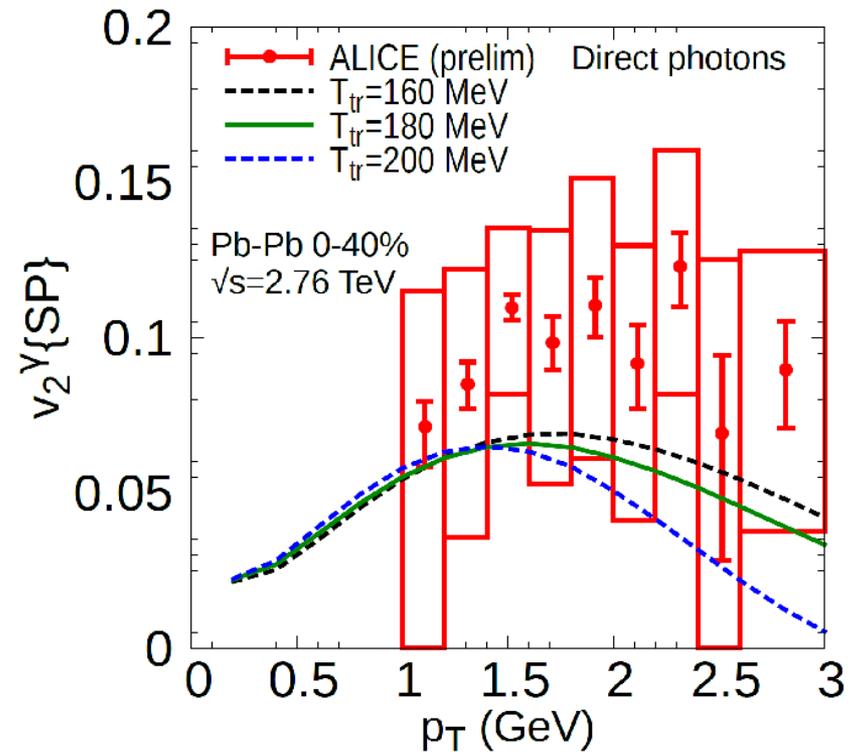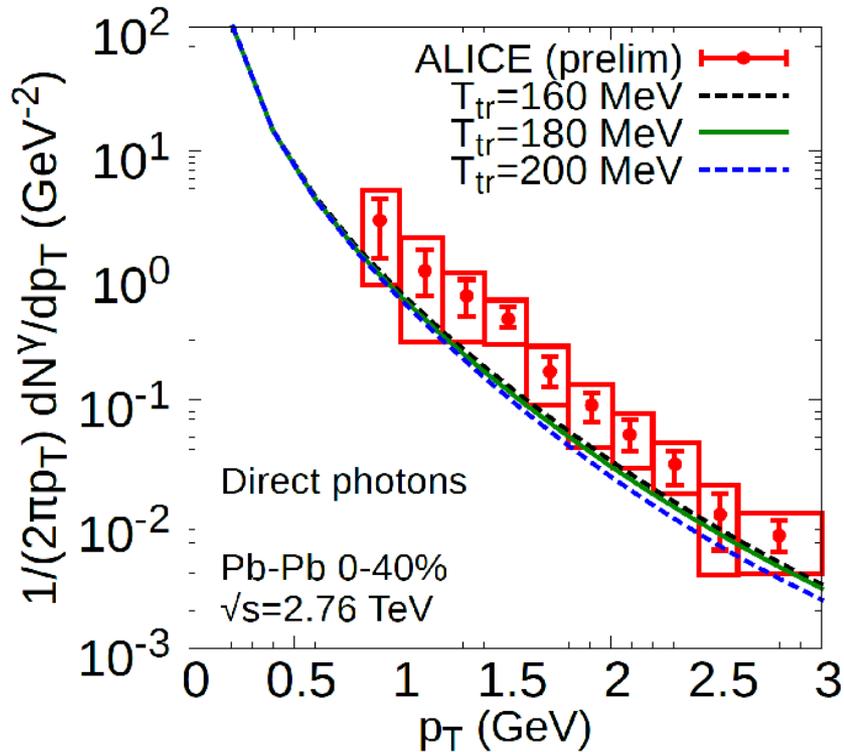# 0-40% centrality calculation vs 0-86% data

Transition temperature for photon rates

QGP vs hadronic photon rates around T_c

Photon production for different transition T

Thermal EM probes: thermal rate

$$\frac{d^4 N_{\gamma/l^+l^-}}{d^4 k} = \int d^4 X \boxed{\frac{d^4 \Gamma_{\gamma/l^+l^-}}{d^4 k}} (K^\mu, u^\mu(X), T(X), \pi^{\mu\nu}(X), \Pi(X))$$

Gas of hadrons: ~100 MeV

Deconfinement: ~160 MeV

Max T at RHIC: ~400 MeV

Max T at LHC: ~600 MeV:

Toward asymptotic QGP

Effective hadronic models

Texas A&M/McGill rates; Stony Brook rates; ...

AdS/CFT and other holography

Effective QCD models

Caron-Huot et al (AdS/CFT); Finazzo and Rougemont (bottom-up holography), BNL et al (semi-QGP), ...

Perturbative QCD

Arnold, Moore, Yaffe (AMY); Ghiglieri, Teaney; Laine, ...
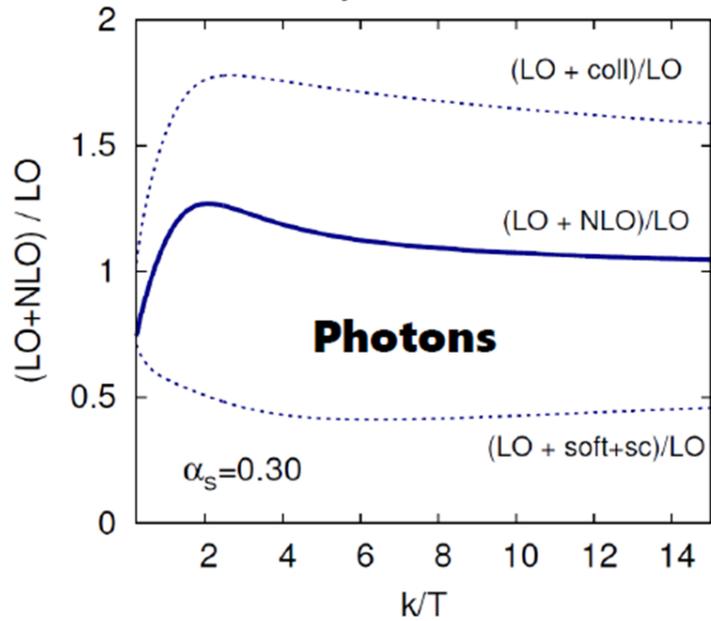
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Perturbative rate at LO and NLO in g_s

QM2017

(Ref.: Ghiglieri, Hong, Kurkela, Lu, Moore & Teaney, JHEP, 2013)

NLO (g_s^3) correction to photon rate is small

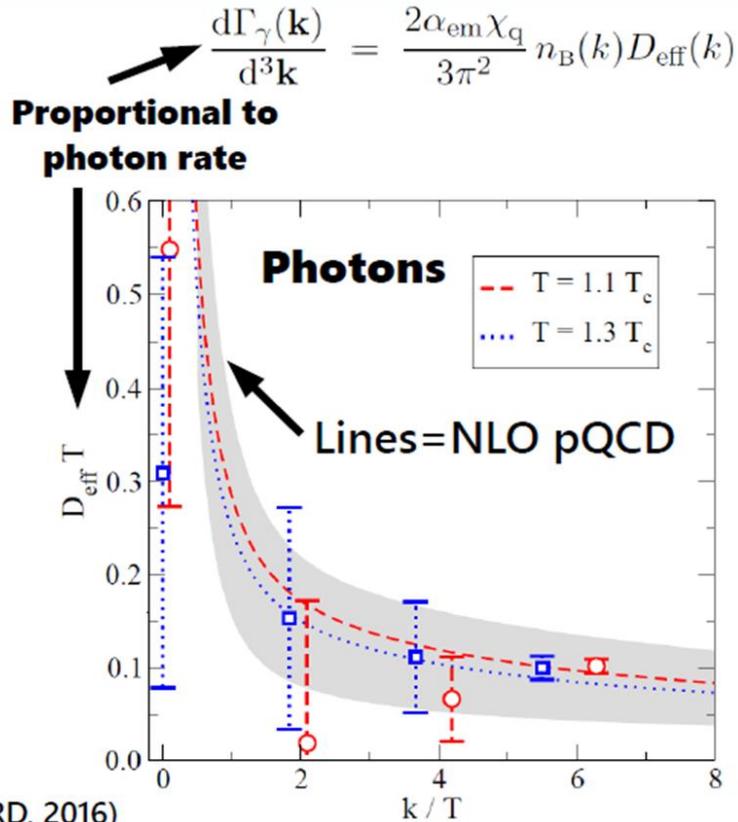(Unlike e.g. heavy quark energy loss and shear viscosity)

Graph axes: y-axis "(LO+NLO) / LO", x-axis "k/T". Curves labelled (LO + coll)/LO, (LO + NLO)/LO, (LO + soft+sc)/LO. $\alpha_s = 0.30$. "Photons".

NLO="LO"+"collinear (coll)"+"soft"+"semi-collinear (sc)"

Perturbative photon rate vs lattice

$$\frac{d\Gamma_\gamma(\mathbf{k})}{d^3\mathbf{k}} = \frac{2\alpha_{em}\chi_q}{3\pi^2} n_B(k) D_{eff}(k)$$

Proportional to photon rate

For T~T_c^+, perturbative QGP photon rate ~ consistent with lattice
(via vector correlation functions)

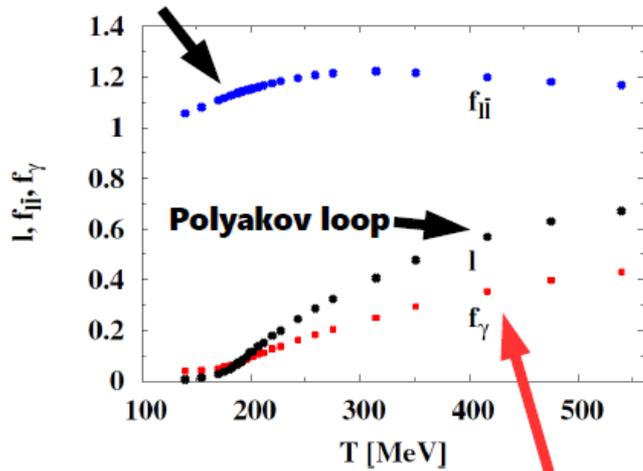(note: quarks are quenched in both the lattice and pQCD calculations)

(Ref.: Ghiglieri, Kaczmarek, Laine & Meyer, PRD, 2016)

Photons

Lines=NLO pQCD

Legend:
- – – T = 1.1 T_c
- ⋯⋯ T = 1.3 T_c

Axes: D_{eff} T (vertical), k / T (horizontal)

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Effective QGP model

Semi-QGP: Mean field with suppressed Polyakov loop

Enhancement factor for dilepton rate [compared to pQCD]

Dilepton enhancement also seen in other calculations
(Lee and Zahed, PRC, 2014; Islam et al, JHEP 2015)

$$k\frac{d^3\Gamma_\gamma^{LO}}{d\boldsymbol{k}} = k\frac{d^3\Gamma_\gamma^{2\to2}}{d\boldsymbol{k}} + k\frac{d^3\Gamma_\gamma^{collinear}}{d\boldsymbol{k}}$$

Suppression factor for "2→2" part of photon rate
(possibly **less suppression** for **collinear** rate)

(Refs.: Gale et al, PRL, 2015; Hidaka, Lin, Pisarski & Satow, JHEP, 2015)

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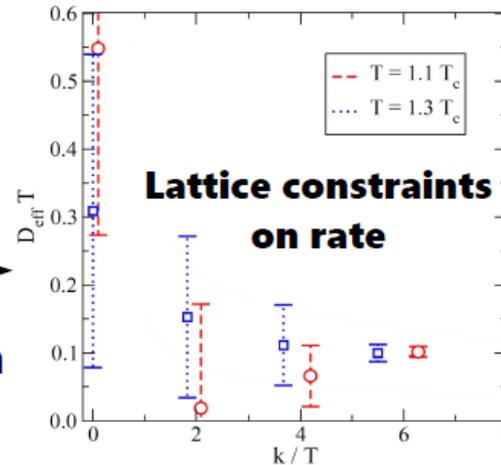
Thermal $\gamma/\ell^+\ell^-$ rate around T_c?

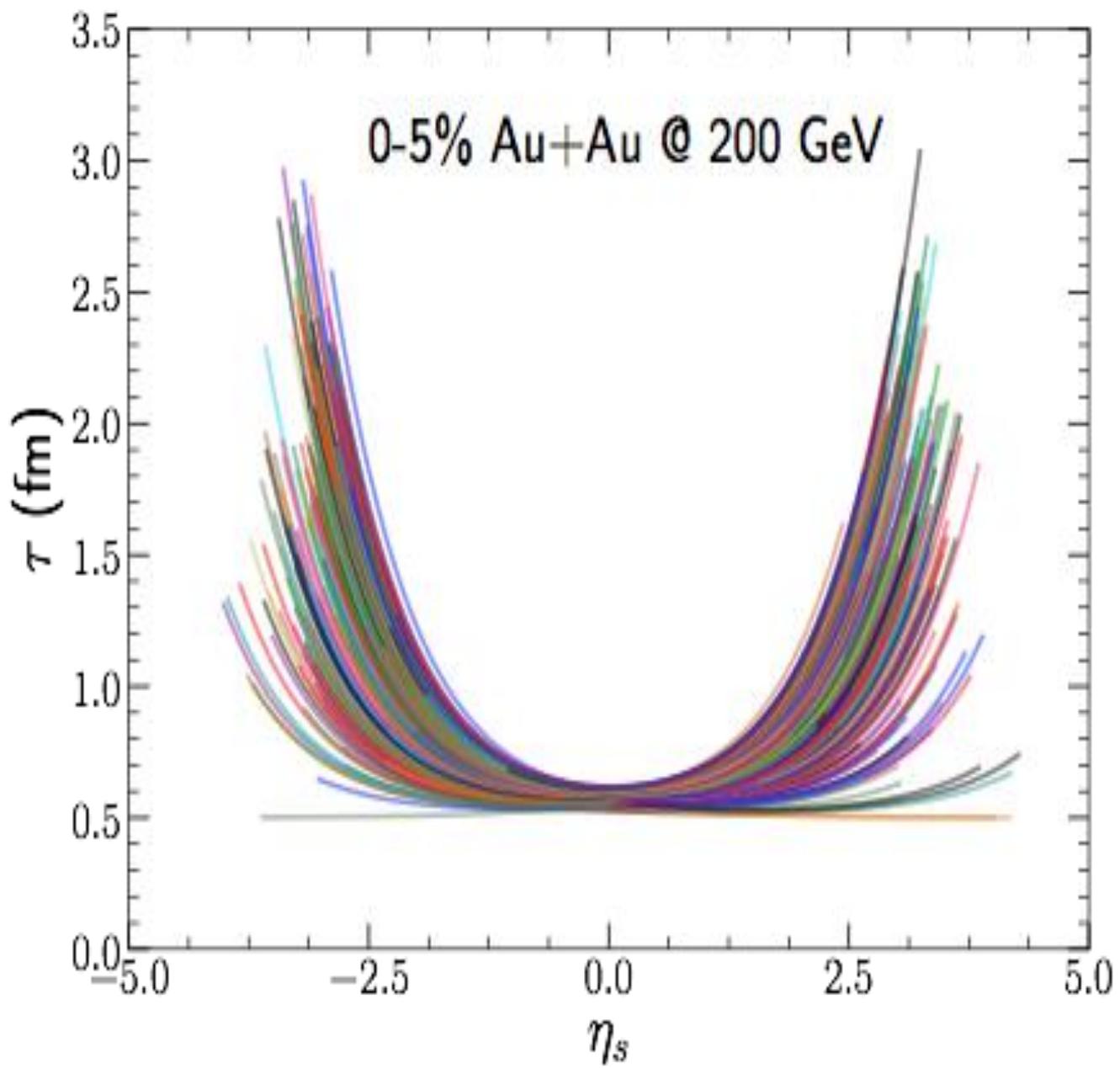## Is the perturbative photon rate reliable at T as low as 200-300 MeV?

>> Mixed signal from different approaches <<

Outlook

- Rate from **bottom-up holography**?
  e.g. Yang et al, arXiv:1609.07208 [poster Monday]
       Finazzo & Rougemont, PRD, 2016;

- **Comparing** holography/semi-QGP **with lattice** calculations? ➡
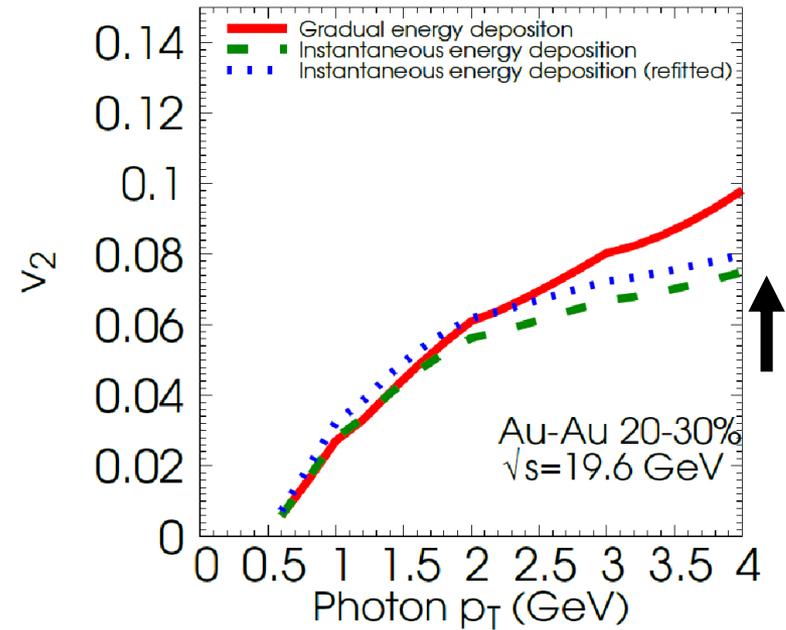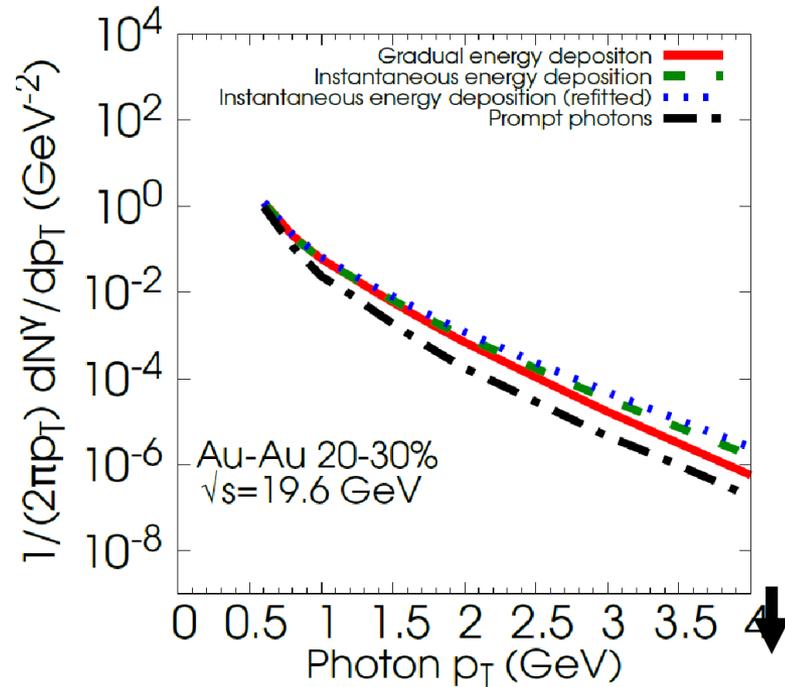
- Holography vs perturbative rate at high temperature?

Lattice constraints on rate

Legend:
- $T = 1.1\ T_c$
- $T = 1.3\ T_c$

(Axes: $D_{eff}\,T$ vs k/T)

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Energy deposition at high rapidity – sqrt(s)=200 GeV

0-5% Au+Au @ 200 GeV

Instantaneous energy deposition - refitting

Probing the early plasma: low $\sqrt{s_{NN}}$

Photon v_2: before prompt dilution

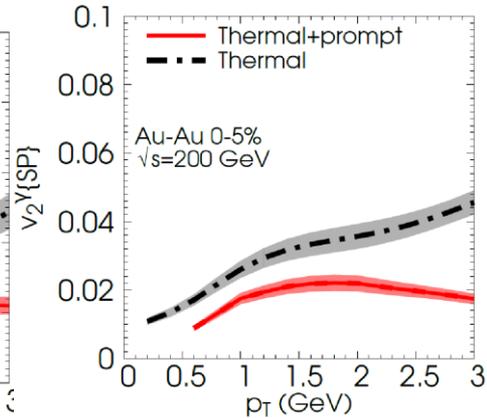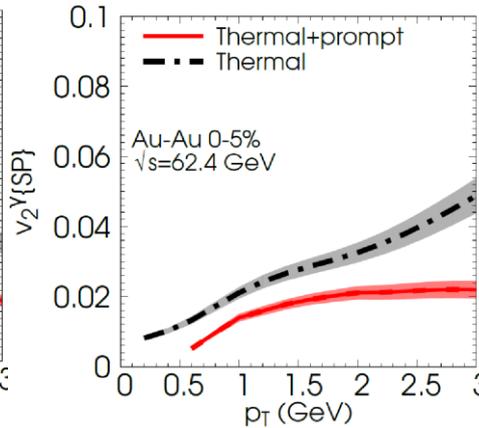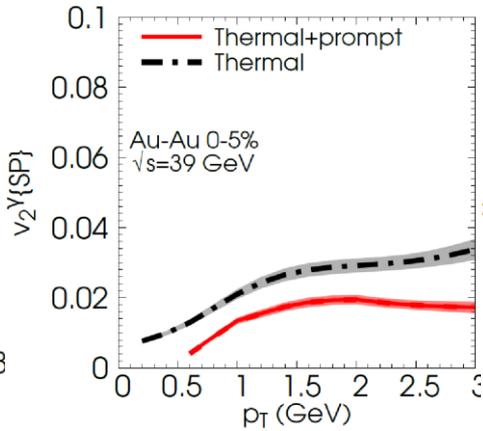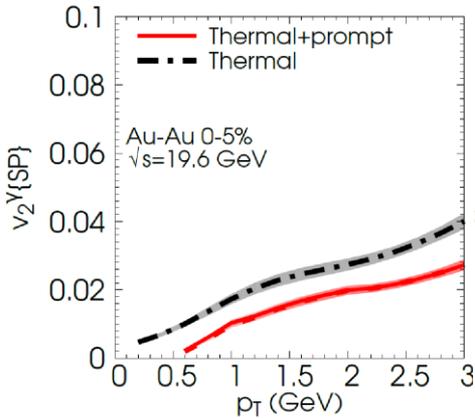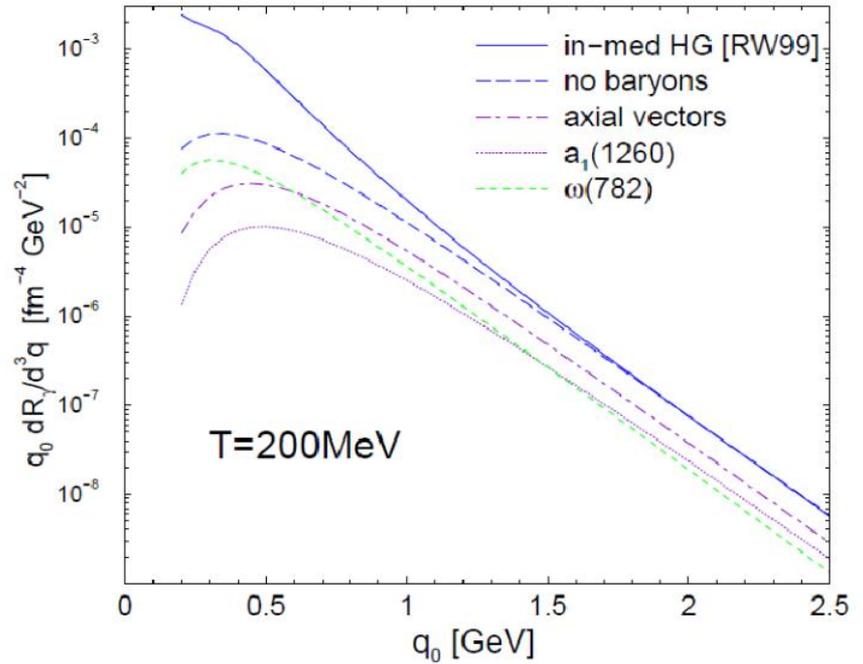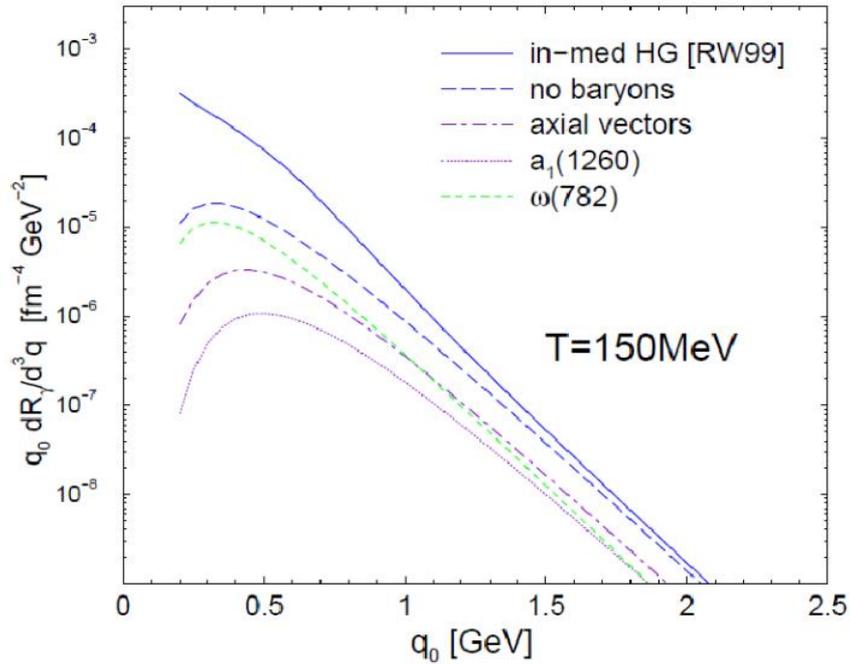# Photon v_2 in low $\sqrt{s_{NN}}$ collisions

Center-of-mass energy of collision

19.6 GeV   39 GeV   62.4 GeV   200 GeV
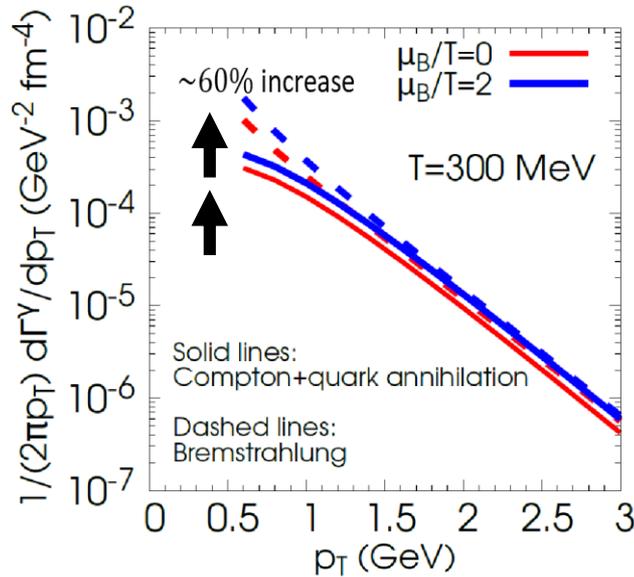

Significant v_2 across $\sqrt{s_{NN}}$

Photon rates: extra details

Hadronic photon rate at zero μ_B : from Turbide. Rapp and Gale (2004)
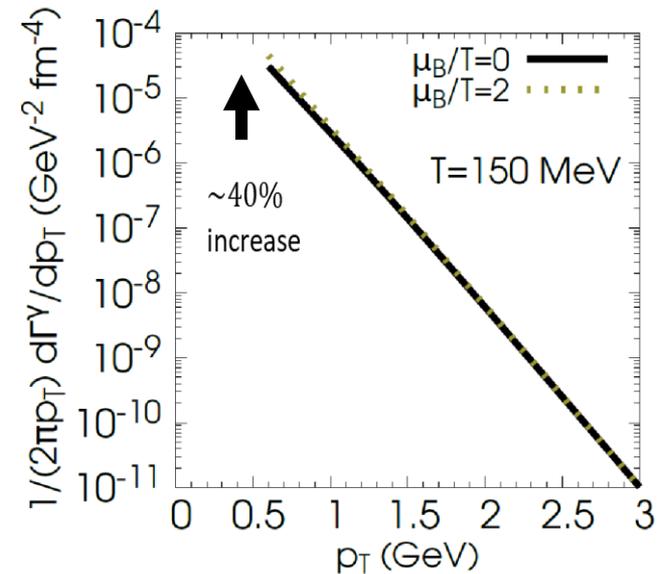
Effect of finite μ_B on thermal photon rate

QGP photon rate

Hadronic photon rate

~60% increase

$\mu_B/T=0$
$\mu_B/T=2$

T=300 MeV

Solid lines:
Compton+quark annihilation

Dashed lines:
Bremsstrahlung

$\mu_B/T=0$
$\mu_B/T=2$

~40% increase

T=150 MeV

⬆ Baryon chemical potential μ_B **increases the photon rate at** <u>**low momenta**</u>

$$e^{-\left(\frac{k-\mu_B}{T}\right)}$$