

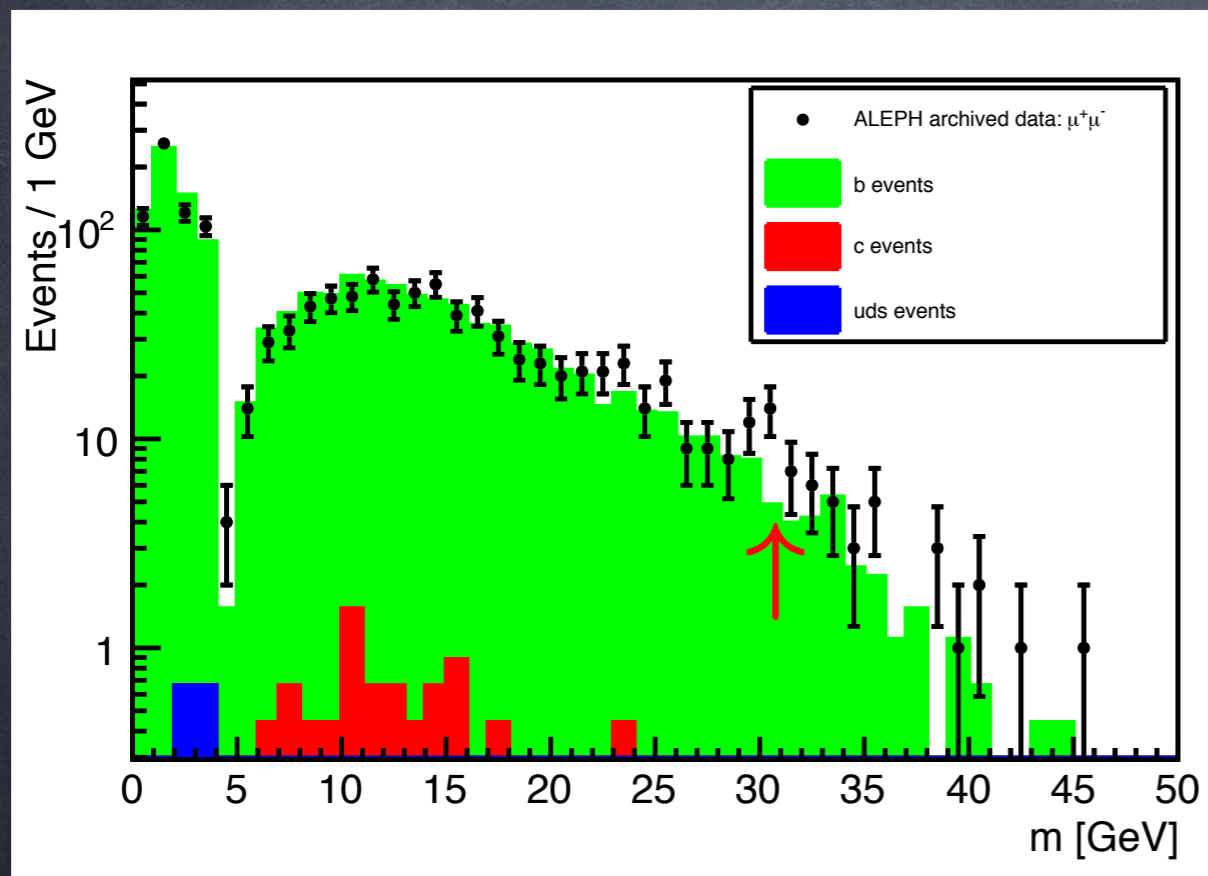
The 30 GeV Dimuon at ALEPH

Kenneth Lane, with Lukas Pritchett
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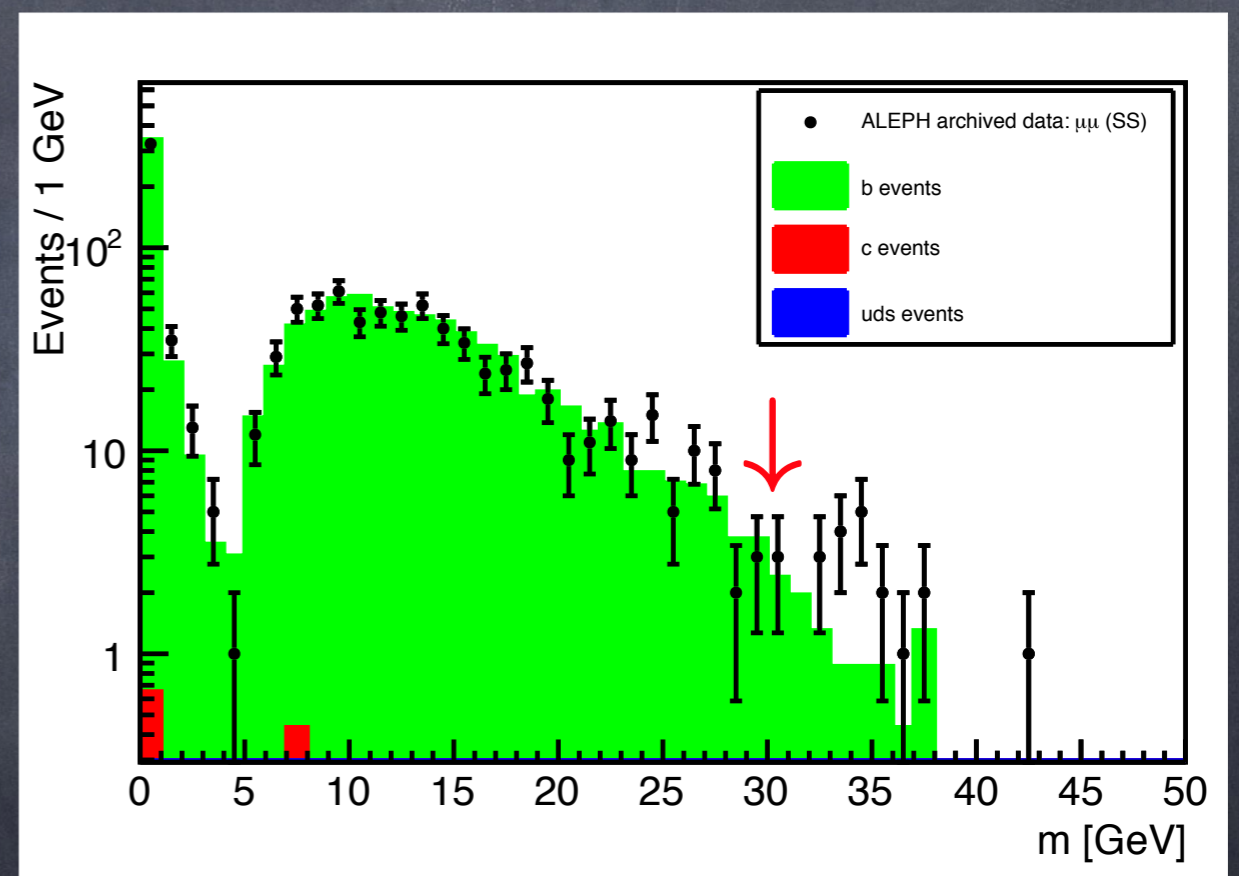
- The ALEPH data
- The 2HDM model
- The two options
- The predictions
- The fly in the ointment

The ALEPH data

From A. Heister, "Observation of an excess at 30 GeV in the opposite sign di-muon spectra of $Z \rightarrow b\bar{b} + X$ events by the ALEPH detector at LEP", arXiv:1610.06536



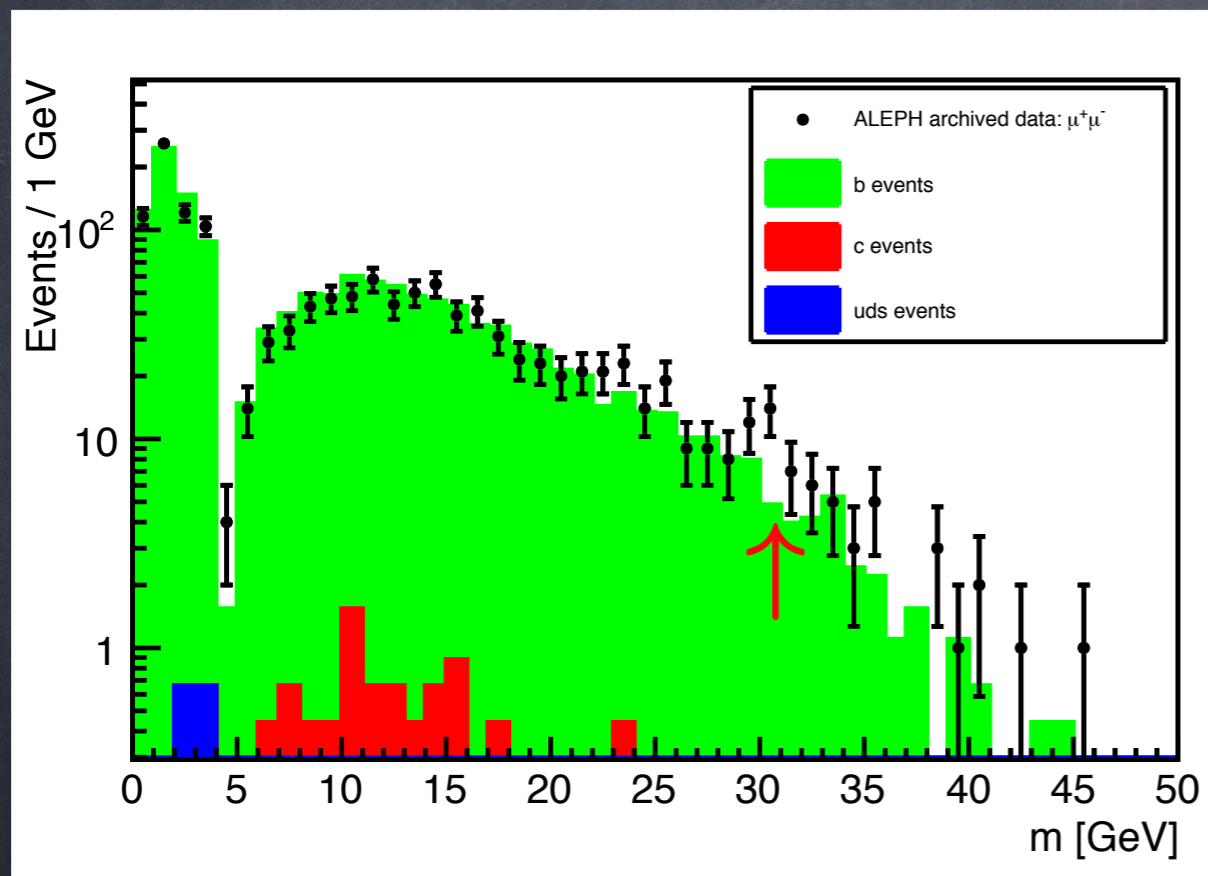
opposite-sign $M_{\mu\mu}$
(from arXiv:1610.06536)



same-sign $M_{\mu\mu}$
(from arXiv:1610.06536)

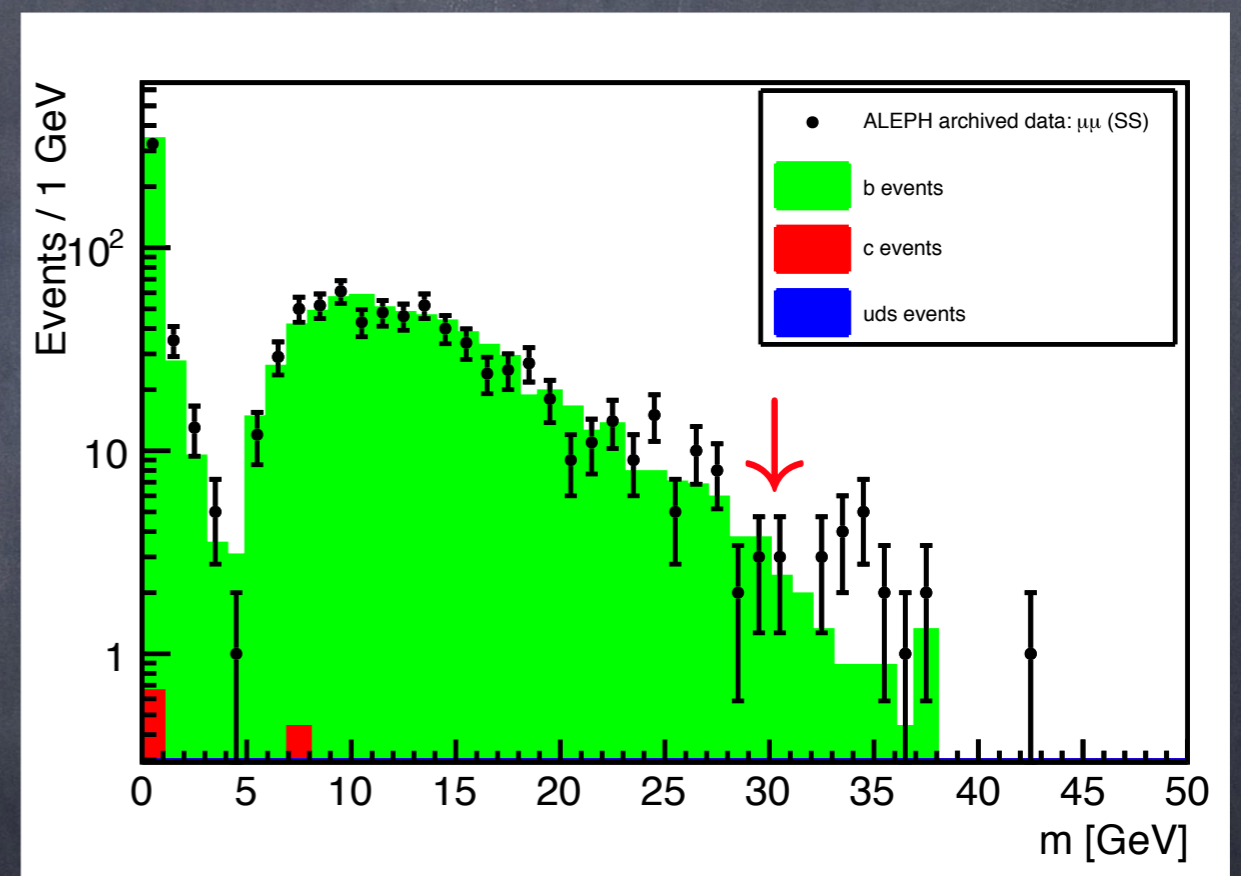
The ALEPH data

From A. Heister, "Observation of an excess at 30 GeV in the opposite sign di-muon spectra of $Z \rightarrow b\bar{b} + X$ events by the ALEPH detector at LEP", arXiv:1610.06536



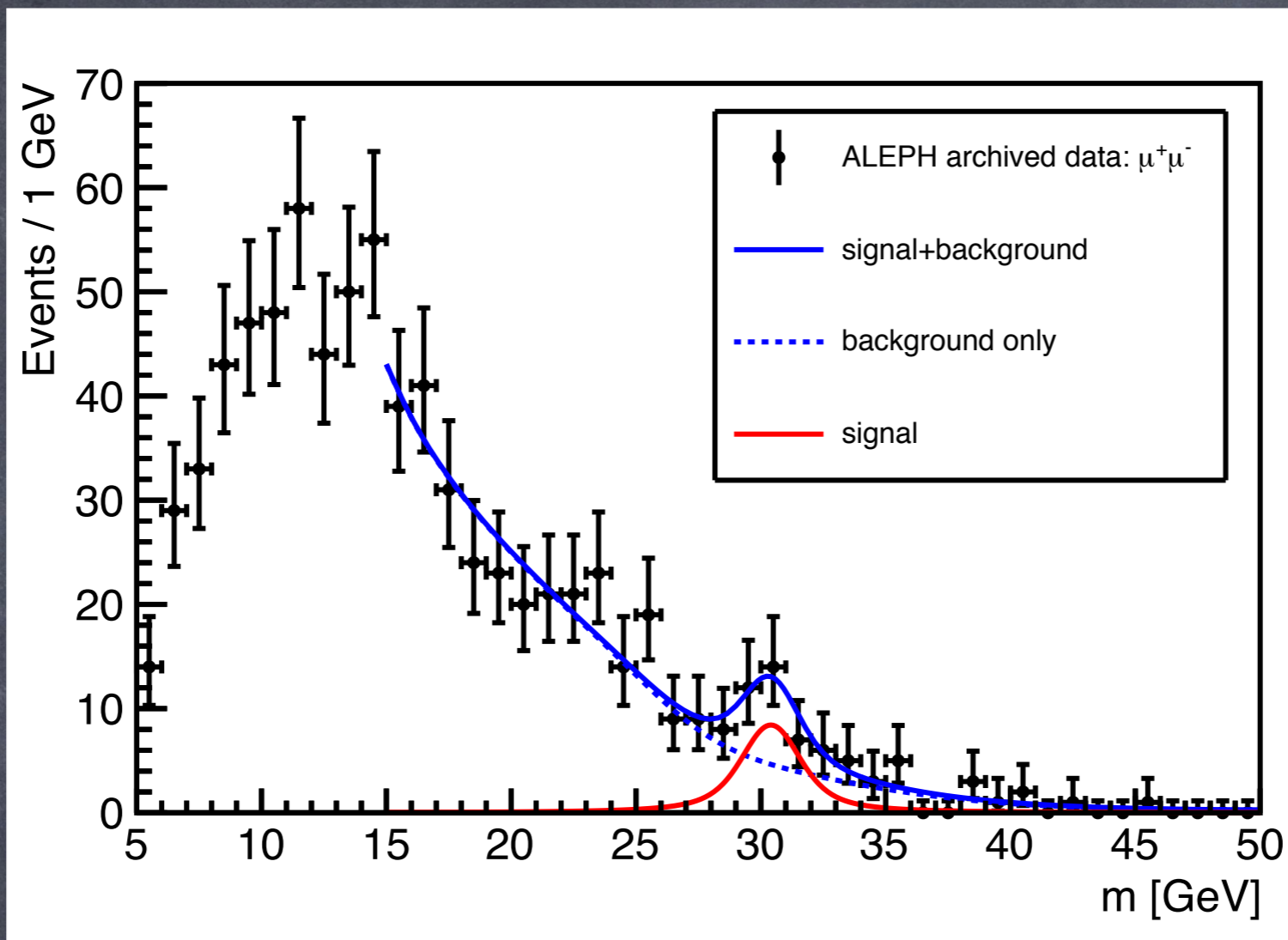
opposite-sign $M_{\mu\mu}$

NO MC reproduces this excess!



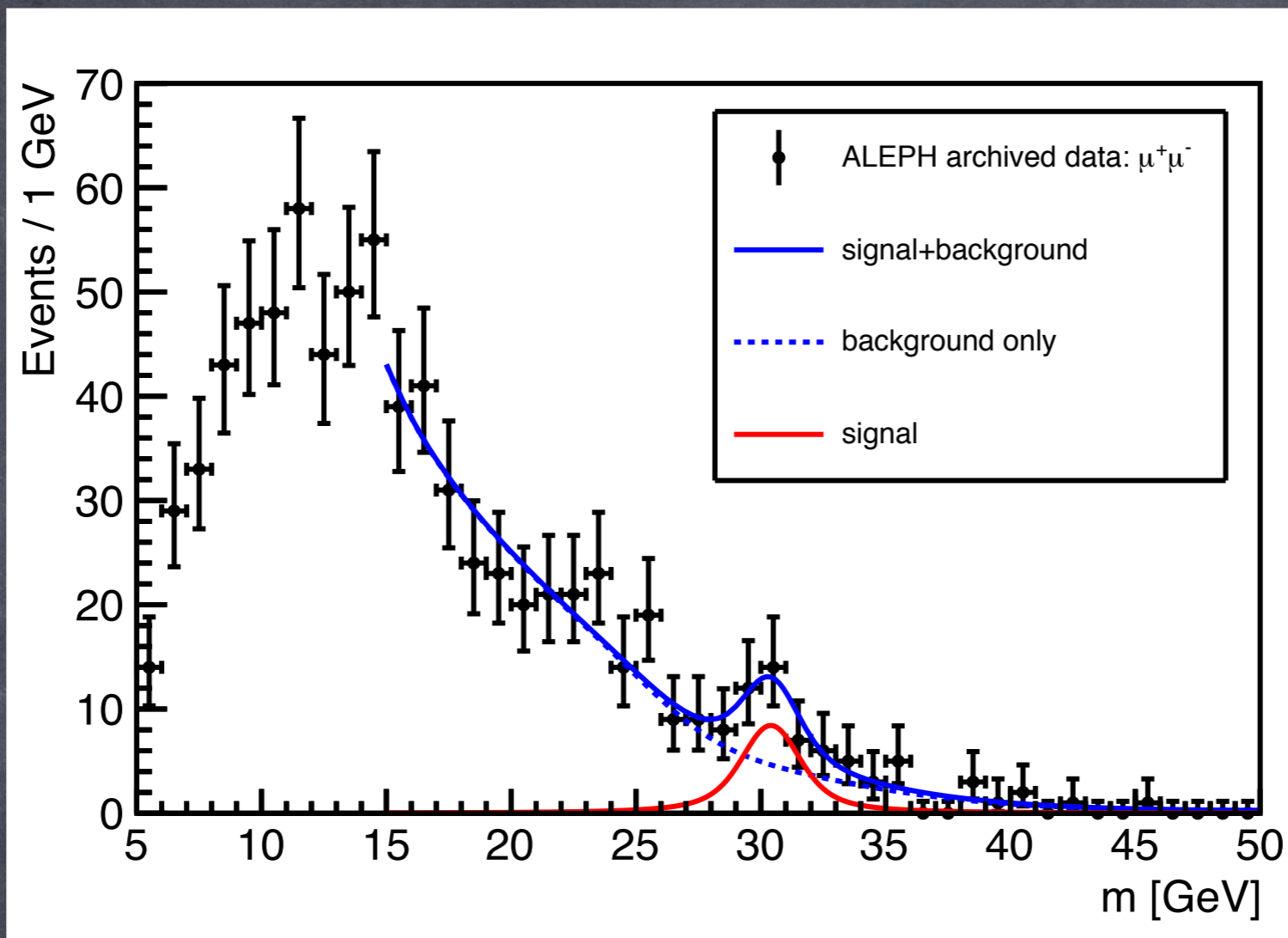
same-sign $M_{\mu\mu}$

(from arXiv:1610.06536)



Signal + background model used to extract signal parameters (from arXiv:1610.06536)

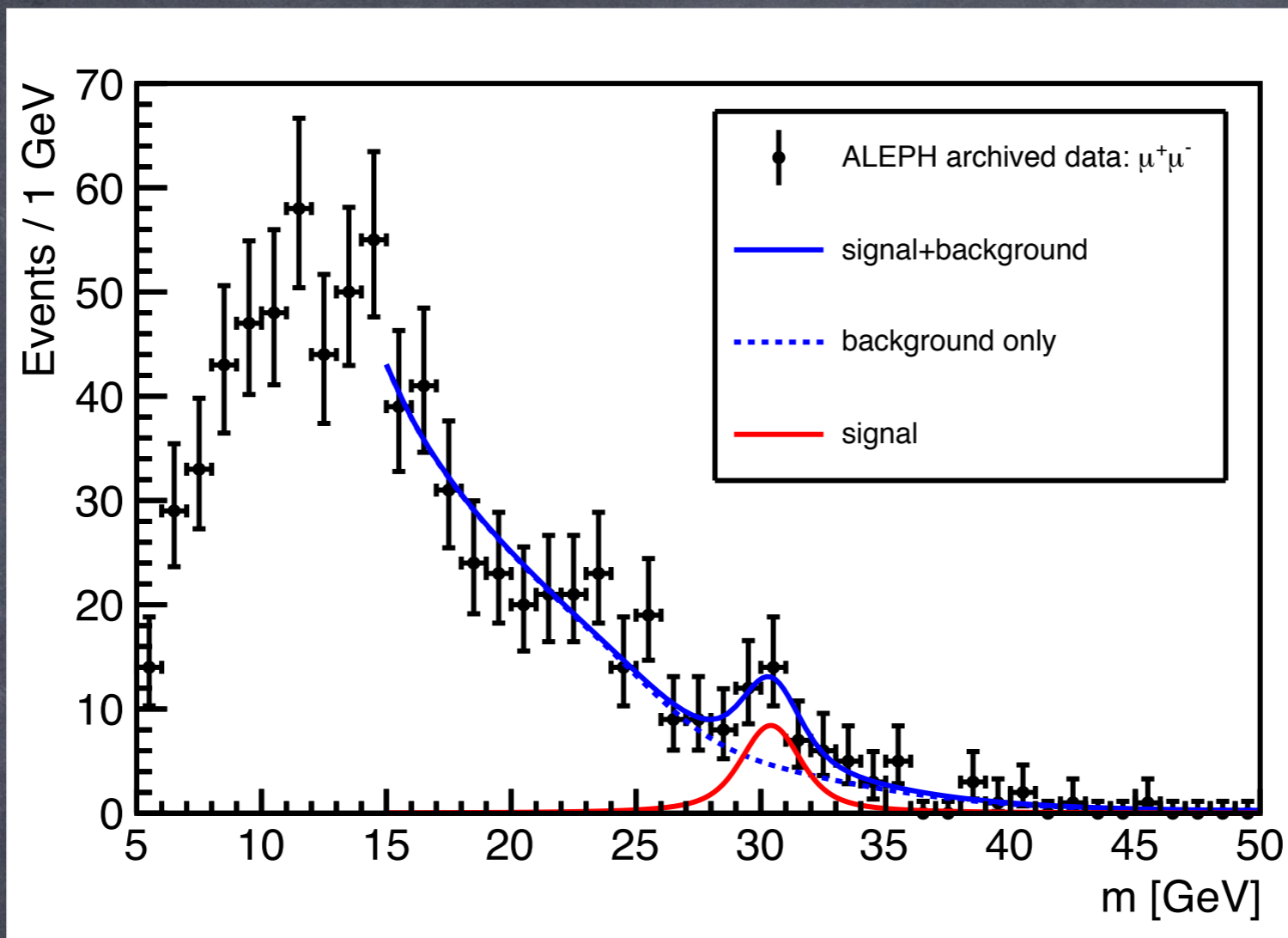
32 ± 11 events, significance $\simeq 3\sigma$ (from statistical tools)



Signal + background model used to extract signal parameters (from arXiv:1610.06536)

32 ± 11 events, significance $\simeq 3\sigma$

Natural width is (1.78 ± 1.14) GeV, IOW, consistent with zero (again, see 1610.06536)



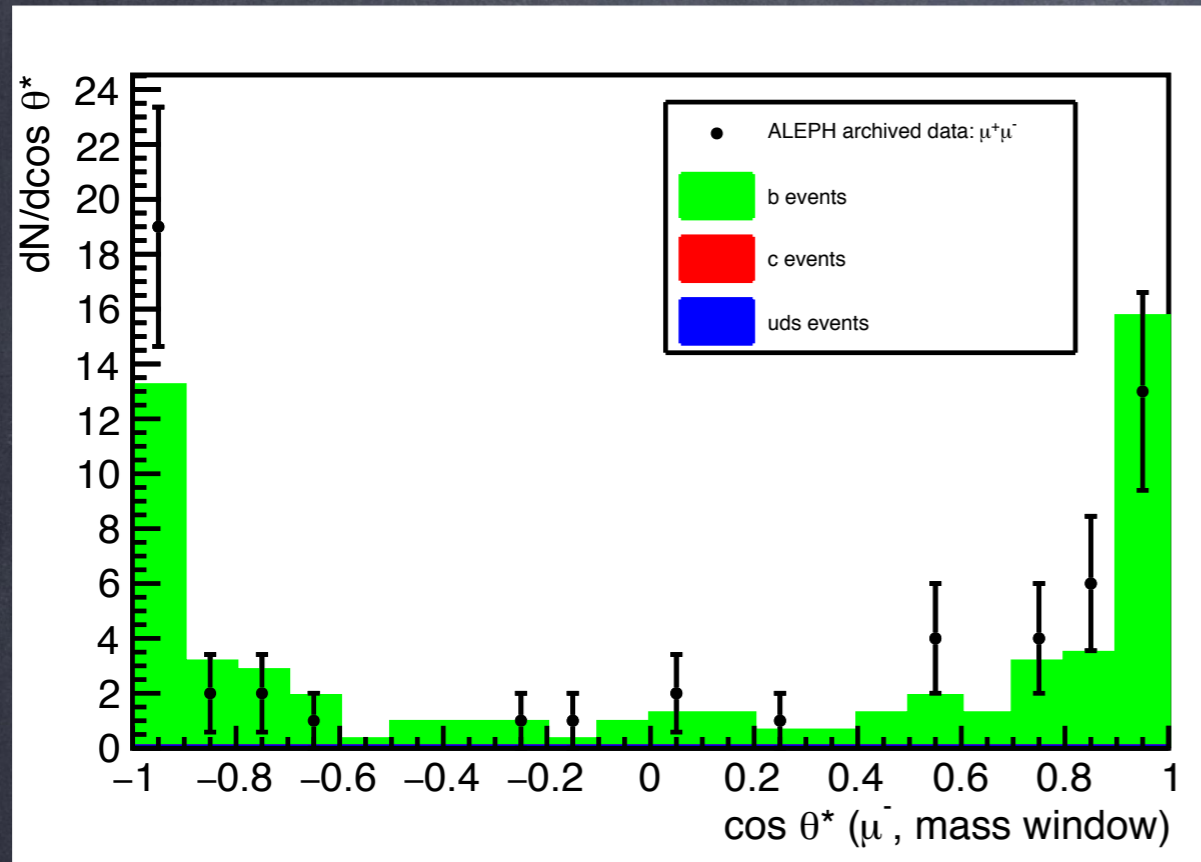
Signal + background model used to extract
signal parameters (from arXiv:1610.06536)

32 ± 11 events, significance $\simeq 3\sigma$

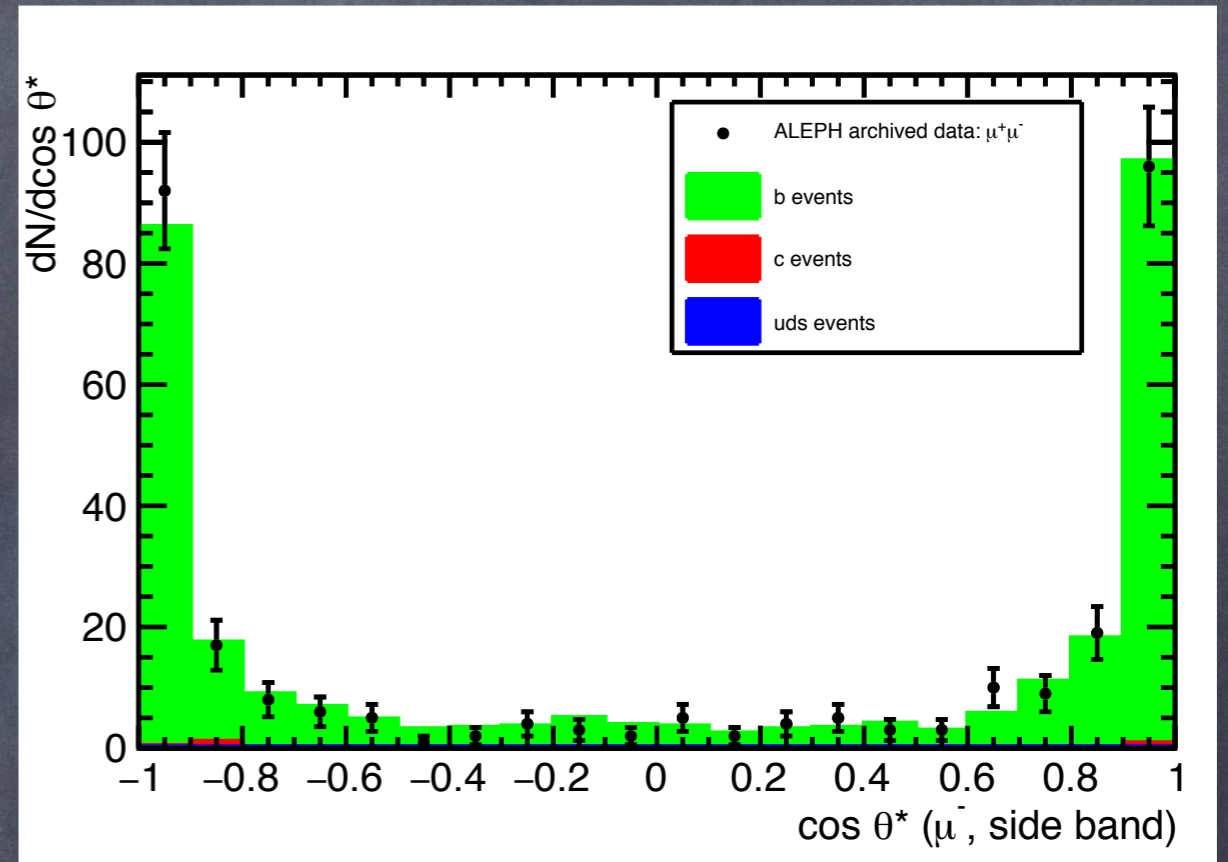
$$\Rightarrow B(Z \rightarrow \bar{b}bX(\rightarrow \mu^+\mu^-)) = (2.77 \pm 0.95) \times 10^{-4}$$

μ^- angular distribution in dimuon rest frame (from arXiv:1610.06536)

$$\cos \theta^* = \hat{p}_{\mu^-} \cdot \hat{p}_{\mu^+ \mu^-}$$



$$M_{\mu^+\mu^-} = (30.40 \pm 3.85) \text{ GeV}$$



$$M_{\mu^+\mu^-} = 15 - 50 \text{ GeV}$$

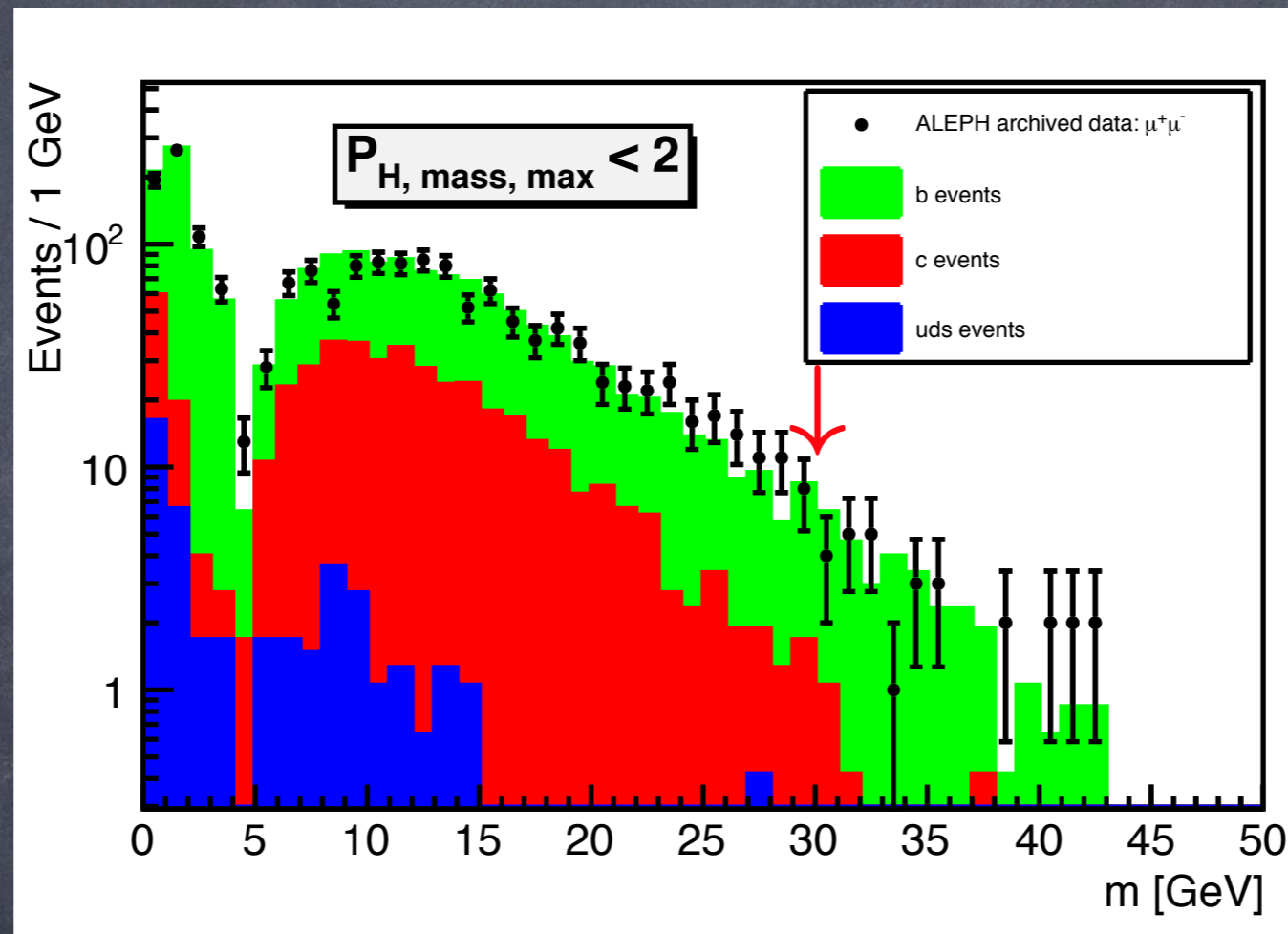
excluding signal region

The “horns” at $\cos \theta^* \cong \pm 1$ are mostly $b \rightarrow c\mu\nu$

There is a weak (~ 1 -sigma) e^+e^- near 30 GeV.

I'm going to Fuhgeddaboutit.

Finally, inverting the b-tag shows no evidence of a dimuon excess near 30 GeV:



==> Hypothesis: The dimuon excess near 30 GeV is associated with Z-decays and $\bar{b}b$ production.

The 2HDM model

Impose softly-broken $U(1)_\phi$ symmetry: $(Y_{EW} = \frac{1}{2})$

$$Y_\phi(\phi_1) = 0, \quad Y_\phi(\phi_2) = 1;$$

$$Y_\phi(q_{Lk}) = Y_\phi \begin{pmatrix} u_{Lk} \\ d_{Lk} \end{pmatrix} = Y_\phi(u_{Rk}) = Y_\phi(d_{Rk}) = 0; \quad (k = 1, 2, 3)$$

$$Y_\phi(L_{Lk}) = Y_\phi \begin{pmatrix} \nu_{Lk} \\ \ell_{Lk} \end{pmatrix} = \frac{1}{2}, \quad Y_\phi(\ell_{Rk}) = -\frac{1}{2}; \quad (k = 1, 2)$$

$$Y_\phi(L_{L3}) = Y_\phi(\ell_{R3}) = 0.$$

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$$Y_\phi(L_{L3}) = Y_\phi(\ell_{R3}) = 0.$$

$$\Gamma_1^\ell = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \times \end{pmatrix}, \quad \Gamma_2^\ell = \begin{pmatrix} \times & \times & 0 \\ \times & \times & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$\Gamma_1^{u,d} = \begin{pmatrix} \times & \times & \times \\ \times & \times & \times \\ \times & \times & \times \end{pmatrix}, \quad \Gamma_2^{u,d} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

Impose softly-broken $U(1)_\phi$ symmetry: $(Y_{EW} = \frac{1}{2})$

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$$Y_\phi(L_{L3}) = Y_\phi(\ell_{R3}) = 0.$$

to maintain:

- $\phi_{10} \cong H(125)$ with \cong SM couplings
- $\phi_{20} \cong h$, $\phi_{23} \cong \eta_A$ couple to $\mu^+ \mu^-$, $\bar{b}b$ (weakly, thru mixing)
- $\phi_2^\pm \cong h^\pm$ couple to $\mu^\pm \nu_\mu$

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$$Y_\phi(L_{L3}) = Y_\phi(\ell_{R3}) = 0.$$

$$V(\phi_1, \phi_2) = -\mu_1^2 \phi_1^\dagger \phi_1 - \mu_2^2 \phi_2^\dagger \phi_2 - \mu_3^2 (\phi_1^\dagger \phi_2 + \phi_2^\dagger \phi_1) + \lambda_1 (\phi_1^\dagger \phi_1)^2 \\ + \lambda_2 (\phi_2^\dagger \phi_2)^2 + 2\lambda_3 (\phi_1^\dagger \phi_1)(\phi_2^\dagger \phi_2) + 2\lambda_4 (\phi_1^\dagger \phi_2)(\phi_2^\dagger \phi_1).$$

Impose softly-broken $U(1)_\phi$ symmetry: $(Y_{EW} = \frac{1}{2})$

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$$V(\phi_1, \phi_2) = -\mu_1^2 \phi_1^\dagger \phi_1 - \mu_2^2 \phi_2^\dagger \phi_2 - \mu_3^2 (\phi_1^\dagger \phi_2 + \phi_2^\dagger \phi_1) + \lambda_1 (\phi_1^\dagger \phi_1)^2 \\ + \lambda_2 (\phi_2^\dagger \phi_2)^2 + 2\lambda_3 (\phi_1^\dagger \phi_1)(\phi_2^\dagger \phi_2) + 2\lambda_4 (\phi_1^\dagger \phi_2)(\phi_2^\dagger \phi_1).$$

$$\langle \phi_i \rangle_0 = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_i \end{pmatrix} \quad \text{for } \mu_{1,2,3}^2 > 0, \quad \lambda_{1,2} > 0, \quad \lambda_{3,4} \text{ real.}$$

The 2HDM masses

CP-even:

$$M^2(\phi_{10}, \phi_{20}) = \begin{pmatrix} \mu_3^2 v_2/v_1 + 2\lambda_1 v_1^2 & -\mu_3^2 + 2(\lambda_3 + \lambda_4)v_1 v_2 \\ -\mu_3^2 + 2(\lambda_3 + \lambda_4)v_1 v_2 & \mu_3^2 v_1/v_2 + 2\lambda_2 v_2^2 \end{pmatrix},$$

$$H = \phi_{10} \cos \alpha + \phi_{20} \sin \alpha, \quad h = -\phi_{10} \sin \alpha + \phi_{20} \cos \alpha,$$

$$\text{where } \tan 2\alpha = \frac{2(2(\lambda_3 + \lambda_4)v_1 v_2 - \mu_3^2)}{\mu_3^2(v_2/v_1 - v_1/v_2) + 2(\lambda_1 v_1^2 - \lambda_2 v_2^2)};$$

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$$M^2(H, h) = \frac{1}{2v_1 v_2} \left\{ \mu_3^2 v^2 + 2(\lambda_1 v_1^2 + \lambda_2 v_2^2)v_1 v_2 \right. \\ \pm \left[[\mu_3^2 v^2 + 2(\lambda_1 v_1^2 + \lambda_2 v_2^2)v_1 v_2]^2 - 8[\mu_3^2(\lambda_1 v_1^4 + \lambda_2 v_2^4)v_1 v_2 \right. \\ \left. \left. + 2\lambda_1 \lambda_2 v_1^4 v_2^4 + 2\mu_3^2(\lambda_3 + \lambda_4)v_1^3 v_2^3 - 2(\lambda_3 + \lambda_4)^2 v_1^4 v_2^4] \right]^{\frac{1}{2}} \right\}.$$

CP-even:

$$M^2(\phi_{10}, \phi_{20}) = \begin{pmatrix} \mu_3^2 v_2/v_1 + 2\lambda_1 v_1^2 & -\mu_3^2 + 2(\lambda_3 + \lambda_4)v_1 v_2 \\ -\mu_3^2 + 2(\lambda_3 + \lambda_4)v_1 v_2 & \mu_3^2 v_1/v_2 + 2\lambda_2 v_2^2 \end{pmatrix},$$

$$H = \phi_{10} \cos \alpha + \phi_{20} \sin \alpha, \quad h = -\phi_{10} \sin \alpha + \phi_{20} \cos \alpha,$$

$$\text{where } \tan 2\alpha = \frac{2(2(\lambda_3 + \lambda_4)v_1 v_2 - \mu_3^2)}{\mu_3^2(v_2/v_1 - v_1/v_2) + 2(\lambda_1 v_1^2 - \lambda_2 v_2^2)};$$

$$M^2(H, h) = \frac{1}{2v_1 v_2} \left\{ \mu_3^2 v^2 + 2(\lambda_1 v_1^2 + \lambda_2 v_2^2)v_1 v_2 \right.$$


$$\pm \left[\left[\mu_3^2 v^2 + 2(\lambda_1 v_1^2 + \lambda_2 v_2^2)v_1 v_2 \right]^2 - 8 \left[\mu_3^2 (\lambda_1 v_1^4 + \lambda_2 v_2^4)v_1 v_2 \right. \right.$$

$$\left. \left. + 2\lambda_1 \lambda_2 v_1^4 v_2^4 + 2\mu_3^2 (\lambda_3 + \lambda_4)v_1^3 v_2^3 - 2(\lambda_3 + \lambda_4)^2 v_1^4 v_2^4 \right] \right]^{\frac{1}{2}} \left. \right\}.$$

CP-odd:

$$\pi_A = \phi_{13} \cos \beta + \phi_{23} \sin \beta, \quad \eta_A = \phi_{13} \sin \beta - \phi_{23} \cos \beta,$$


$$\text{where } \tan \beta = \frac{v_2}{v_1};$$

$$M_{\pi_A}^2 = 0, \quad M_{\eta_A}^2 = \mu_3^2 \frac{v^2}{v_1 v_2}.$$


CP-odd:


$$\pi_A = \phi_{13} \cos \beta + \phi_{23} \sin \beta, \quad \eta_A = \phi_{13} \sin \beta - \phi_{23} \cos \beta,$$

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$$M_{\pi_A}^2 = 0, \quad M_{\eta_A}^2 = \mu_3^2 \frac{v^2}{v_1 v_2}.$$


Charged:

$$\pi^\pm = \phi_1^\pm \cos \beta + \phi_2^\pm \sin \beta, \quad h^\pm = \phi_1^\pm \sin \beta - \phi_2^\pm \cos \beta,$$

$$M_{\pi^\pm}^2 = 0, \quad M_{h^\pm}^2 = \left(\frac{\mu_3^2}{v_1 v_2} - \lambda_4 \right) v^2.$$


The 2HDM interactions

Leptons:

$$\begin{aligned} \mathcal{L}_{Yl} = & - \sum_{\ell_k=e,\mu} \frac{m_{\ell_k}}{v \sin \beta} \bar{\ell}_k [v \sin \beta + H \sin \alpha + h \cos \alpha - i\eta_A \cos \beta \gamma_5] \ell_k \\ & - \frac{m_\tau}{v \cos \beta} \bar{\tau} [v \cos \beta + H \cos \alpha - h \sin \alpha + i\eta_A \sin \beta \gamma_5] \tau \\ & + h^+ \left[\sum_{k=e,\mu} \frac{\sqrt{2} m_{\ell_k} \cot \beta}{v} \bar{\nu}_{kL} \ell_{kR} - \frac{\sqrt{2} m_\tau \tan \beta}{v} \bar{\nu}_{\tau L} \tau_R \right] + \text{h.c.} \end{aligned}$$

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 & - \frac{m_\tau}{v \cos \beta} \bar{\tau} [v \cos \beta + H \cos \alpha - h \sin \alpha + i\eta_A \sin \beta \gamma_5] \tau \\
 & + h^+ \left[\sum_{k=e,\mu} \frac{\sqrt{2} m_{\ell_k} \cot \beta}{v} \bar{\nu}_{kL} \ell_{kR} - \frac{\sqrt{2} m_\tau \tan \beta}{v} \bar{\nu}_{\tau L} \tau_R \right] + \text{h.c.}
 \end{aligned}$$

N.B.: These interactions induce no detectable charged-lepton flavor violation!

Quarks:

$$\begin{aligned}
 \mathcal{L}_{Y_q} = & - \sum_{d_k=d,s,b} \frac{m_{d_k}}{v \cos \beta} \bar{d}_k (v \cos \beta + H \cos \alpha - h \sin \alpha - i\gamma_5 \eta_A \sin \beta) d_k \\
 & - \sum_{u_k=u,c,t} \frac{m_{u_k}}{v \cos \beta} \bar{u}_k (v \cos \beta + H \cos \alpha - h \sin \alpha + i\gamma_5 \eta_A \sin \beta) u_k \\
 & - \frac{\sqrt{2} \tan \beta}{v} \sum_{k,l=1}^3 [\bar{u}_{kL} (V \mathcal{M}_d)_{kl} h^+ d_{lR} - \bar{d}_{kL} (V^\dagger \mathcal{M}_u)_{kl} h^- u_{lR}] + \text{h.c.}
 \end{aligned}$$

$$\implies \frac{B(\eta_A \rightarrow \bar{b}b)}{B(\eta_A \rightarrow \mu^+ \mu^-)} = \frac{3m_b^2 \tan^4 \beta}{m_\mu^2}$$

EW bosons:

$$\begin{aligned}\mathcal{L}_{EW} = & \frac{e}{\sin 2\theta_W} \left[(h \cos(\beta - \alpha) - H \sin(\beta - \alpha)) \overleftrightarrow{\partial}_\mu \eta_A \right] Z^\mu \\ & + \frac{e}{2 \sin \theta_W} \left[(\eta_A \pm ih \cos(\beta - \alpha) \mp iH \sin(\beta - \alpha)) \overleftrightarrow{\partial}_\mu h^\pm \right] W^\mp{}^\mu \\ & + \left[\frac{2e^2}{\sin^2 2\theta_W} Z^\mu Z_\mu + \frac{e^2}{\sin^2 \theta_W} W^{+\mu} W_\mu^- \right] [H \cos(\beta - \alpha) + h \sin(\beta - \alpha)] .\end{aligned}$$

The two options

To be close to SM and consistent with LHC H-data require:

- small α for weak $\phi_{10} - \phi_{20}$ mixing
- small β for weak $\pi_A - \eta_A$ mixing

So...

- $v^2 \cong v_1^2 \gg v_2^2$
- $(\mu_3^2 - 2\lambda_1 v_1 v_2)^2 \gg 8(\lambda_1 \lambda_2 - (\lambda_3 + \lambda_4)^2) v_2^4$

The two options

To be close to SM and consistent with LHC H-data require:

- small α for weak $\phi_{10} - \phi_{20}$ mixing
- small β for weak $\pi_A - \eta_A$ mixing

$$\implies M^2(H, h) \cong \begin{cases} \max(2\lambda_1 v_1^2, \mu_3^2 v^2 / v_1 v_2) \\ \min(2\lambda_1 v_1^2, \mu_3^2 v^2 / v_1 v_2) \end{cases}$$

i.e., either

$$(1) M_H^2 \cong 2\lambda_1 v^2 \text{ and } M_h^2 \cong M_{\eta_A}^2 \cong \mu_3^2 v^2 / v_1 v_2$$

or

$$(2) M_H^2 \cong M_{\eta_A}^2 \cong \mu_3^2 v^2 / v_1 v_2 \text{ and } M_h^2 \cong 2\lambda_1 v^2$$

Option 1:

$Z \rightarrow h\eta_A \rightarrow \mu^+\mu^-\bar{b}b$ plus “Higgsstrahlung” :

$Z \rightarrow Z^*h$ with $Z^* \rightarrow \bar{b}b$, $h \rightarrow \mu^+\mu^-$

or $Z \rightarrow \bar{b}b$ with $b(\bar{b}) \rightarrow b(\bar{b}) + h/\eta_A \rightarrow b(\bar{b})\mu^+\mu^-$

Option 2:

Higgsstrahlung only, with $\mathcal{O}(10^{-10})$
contribution to $Z \rightarrow \bar{b}b + \text{dimuon}$.

So, only option 1 is available:

$$B(Z \rightarrow h\eta_A \rightarrow \bar{b}b\mu^+\mu^-) \cong B(Z \rightarrow h\eta_A) [B(h \rightarrow \bar{b}b)B(\eta_A \rightarrow \mu^+\mu^-) + (h \leftrightarrow \eta_A)]$$

where

$$B(Z \rightarrow h\eta_A) = \frac{2\alpha_{EM} p^3}{3M_Z^2 \sin^2 2\theta_W \Gamma_Z} \cos^2(\beta - \alpha)$$

$$= 0.0141 \cos^2(\beta - \alpha) \text{ for } M_H = M_{\eta_A} = 30 \text{ GeV}$$

So, only option 1 is available:

$$B(Z \rightarrow h\eta_A \rightarrow \bar{b}b\mu^+\mu^-) \cong B(Z \rightarrow h\eta_A) [B(h \rightarrow \bar{b}b)B(\eta_A \rightarrow \mu^+\mu^-) + (h \leftrightarrow \eta_A)]$$

where

$$B(Z \rightarrow h\eta_A) = \frac{2\alpha_{EM} p^3}{3M_Z^2 \sin^2 2\theta_W \Gamma_Z} \cos^2(\beta - \alpha)$$

$$= 0.0141 \cos^2(\beta - \alpha) \text{ for } M_H = M_{\eta_A} = 30 \text{ GeV}$$

Wait for it!

But... $H \rightarrow hh, \eta_A \eta_A, h^+ h^-$??

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- Require $\lambda_4 < 0$ so that

$$M_{h^\pm} > \frac{1}{2} M_H \cong 63 \text{ GeV} \text{ or}$$

$$M_{h^\pm} > (M_{\tilde{\mu}^\pm})_{\text{limit}} \cong 95 \text{ GeV}$$

- Require $|\lambda_3 + \lambda_4| < 5.44 \times 10^{-3}$

$$\text{so that } \Gamma(H \rightarrow hh), \Gamma(H \rightarrow \eta_A \eta_A) < \frac{1}{2} \text{ MeV}$$

But... $H \rightarrow hh, \eta_A \eta_A, h^+ h^-$??

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 $M_{h^\pm} > \frac{1}{2} M_H \cong 63 \text{ GeV}$ or
 $M_{h^\pm} > (M_{\tilde{\mu}^\pm})_{\text{limit}} \cong 95 \text{ GeV}$
- Require $|\lambda_3 + \lambda_4| < 5.44 \times 10^{-3}$
so that $\Gamma(H \rightarrow hh), \Gamma(H \rightarrow \eta_A \eta_A) < \frac{1}{2} \text{ MeV}$
(Yes, it's fine tuning, but who cares?)

The predictions

Tables' inputs:

- $M_{h,\eta_A} = 30 \text{ GeV}$, $M_H = 125 \text{ GeV}$, $M_{h^\pm} = 65 \text{ GeV}$
- $\lambda_4 = (M_{\eta_A}^2 - M_{h^\pm}^2)/v^2 = -5.49 \times 10^{-2}$
- $\lambda_3 + \lambda_4 = -3.00 \times 10^{-3}$;
results insensitive to $|\lambda_3 + \lambda_4| < 5 \times 10^{-3}$.

| v_2 (GeV) | β | α | μ_3^2 (GeV ²) | λ_1 | λ_2 |
|-------------|---------|-------------------------|-------------------------------|-------------|------------------------|
| 10 | 0.04066 | -0.348×10^{-2} | 36.6 | 0.1293 | 0.833×10^{-2} |
| 12.5 | 0.05084 | -0.435×10^{-2} | 45.7 | 0.1294 | 0.833×10^{-2} |
| 15 | 0.06101 | -0.522×10^{-2} | 54.8 | 0.1295 | 0.833×10^{-2} |
| 20 | 0.08139 | -0.695×10^{-2} | 72.9 | 0.1299 | 0.832×10^{-2} |

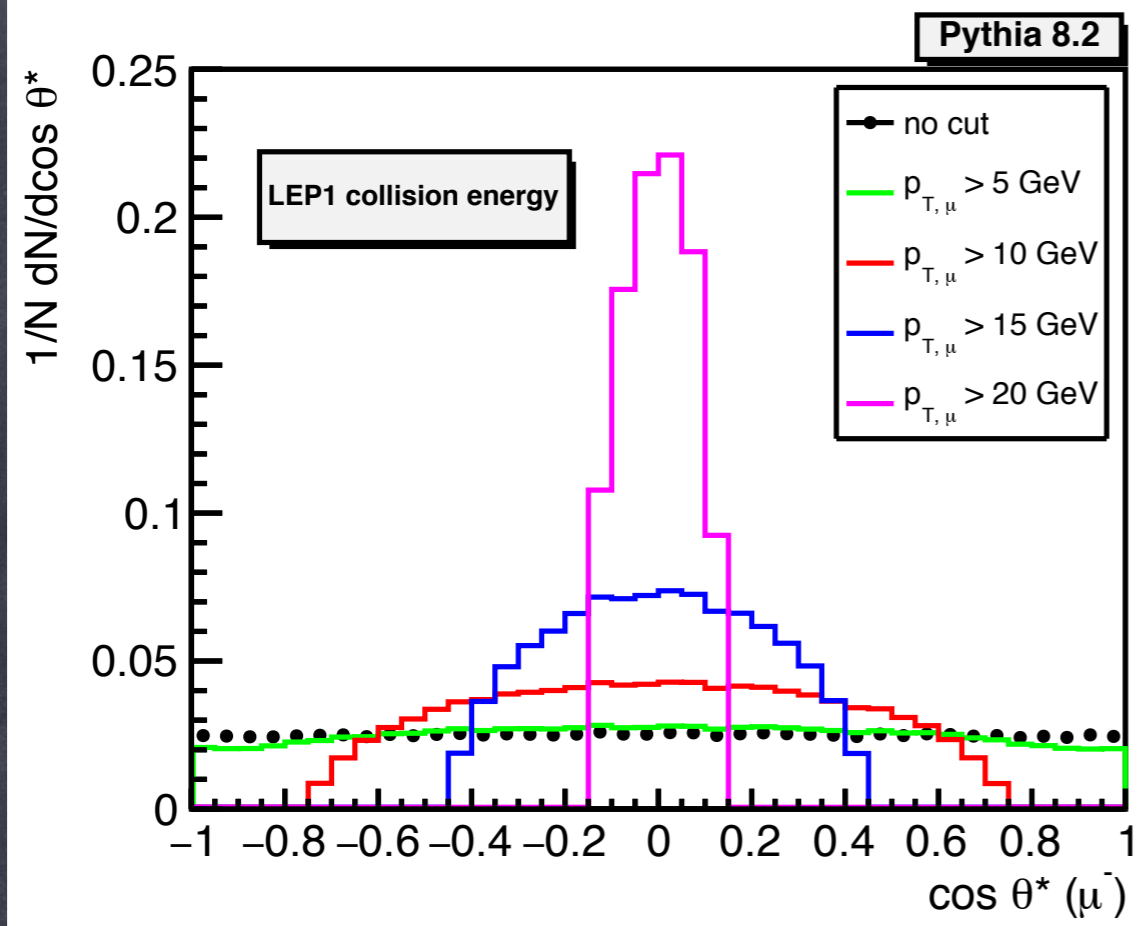
| v_2 (GeV) | $B(h \rightarrow \mu^+ \mu^-)$ | $B(h \rightarrow \bar{b}b)$ | $B(\eta_A \rightarrow \mu^+ \mu^-)$ | $B(\eta_A \rightarrow \bar{b}b)$ | $B(Z \rightarrow \bar{b}b\mu\mu)$ |
|-------------|--------------------------------|-----------------------------|-------------------------------------|----------------------------------|-----------------------------------|
| 10 | 0.9999 | 0.444×10^{-4} | 0.9923 | 0.667×10^{-2} | 0.950×10^{-4} |
| 12.5 | 0.9998 | 1.085×10^{-4} ↑ | 0.9814 | 1.613×10^{-2} ↑ | 2.297×10^{-4} ↑ |
| 15 | 0.9997 | 2.249×10^{-4} | 0.9622 | 3.286×10^{-2} | 4.673×10^{-4} |
| 20 | 0.9991 | 7.105×10^{-4} | 0.8889 | 9.652×10^{-2} | 13.67×10^{-4} |

The predictions

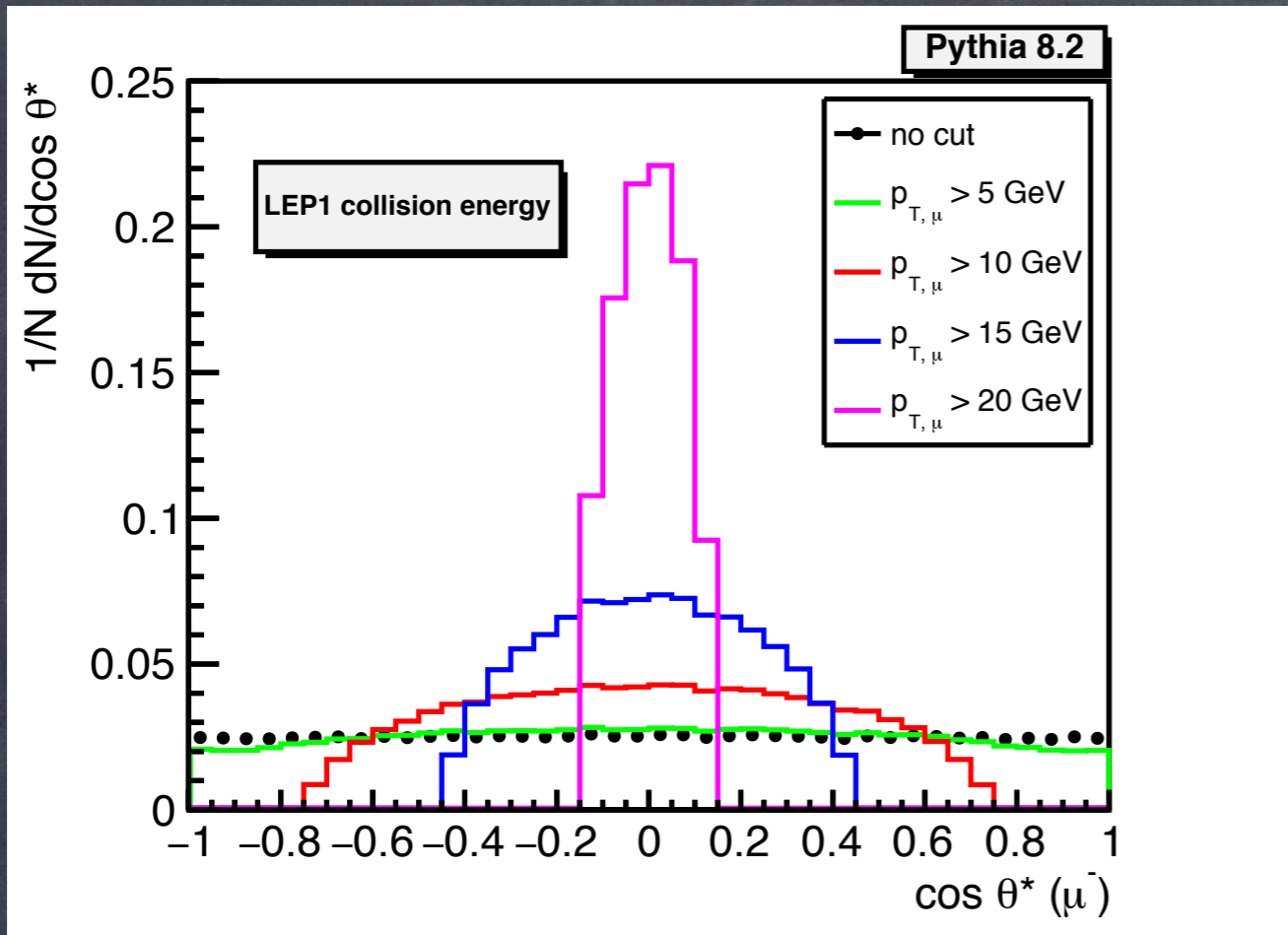
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- 3.) Signal dimuons have an isotropic $\cos\theta^*$ distribution. Its flat shape is modified to an inverted \sim parabola in $\theta_T < \theta^* < \pi - \theta_T$ when there is a $p_T \gtrsim 7 \text{ GeV}$ cut on the muons. (θ_T increases with the p_T cut.)
All the signal lies in $|\cos\theta^*| < \cos\theta_T$. (See figures.)



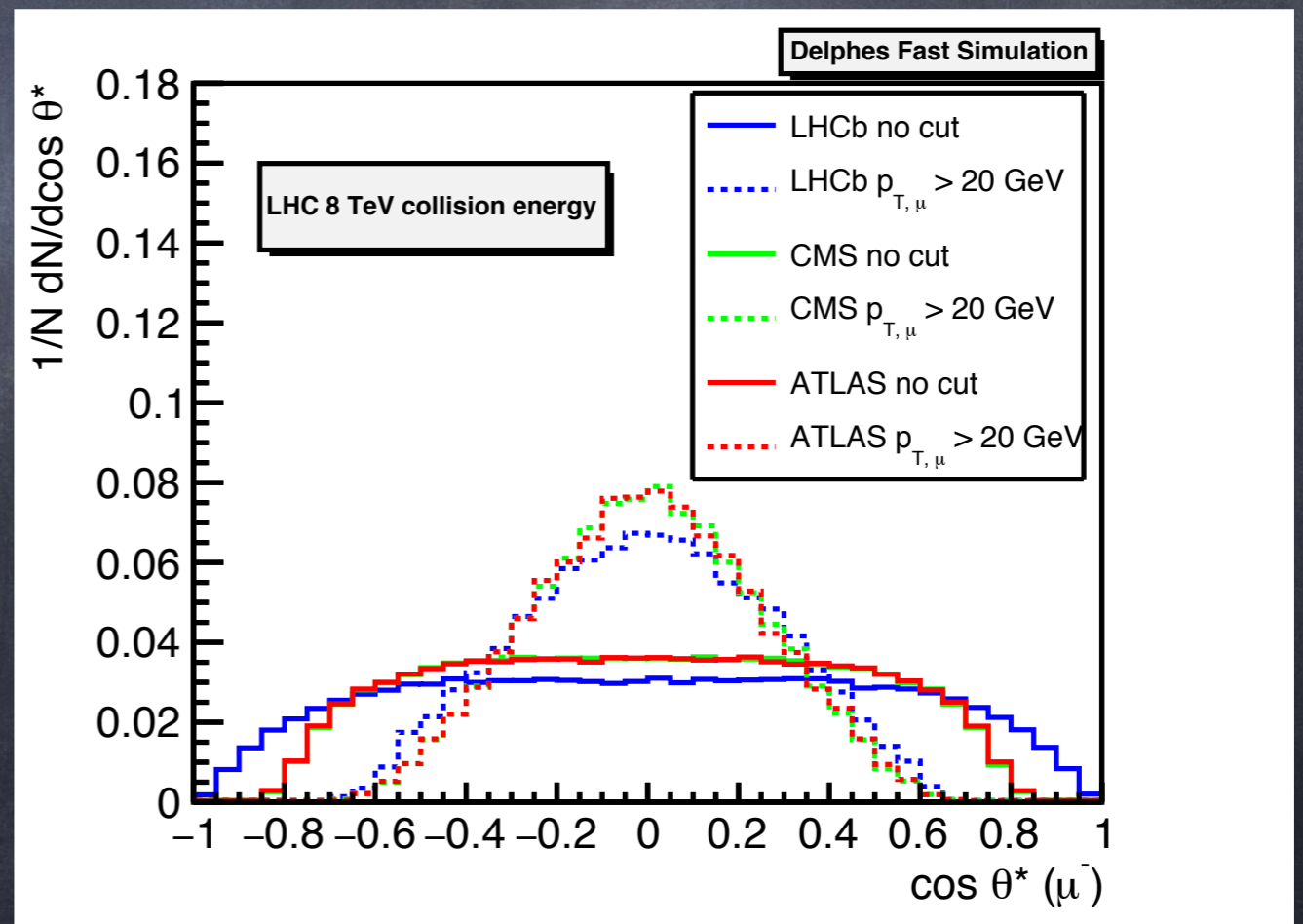
Z's at rest

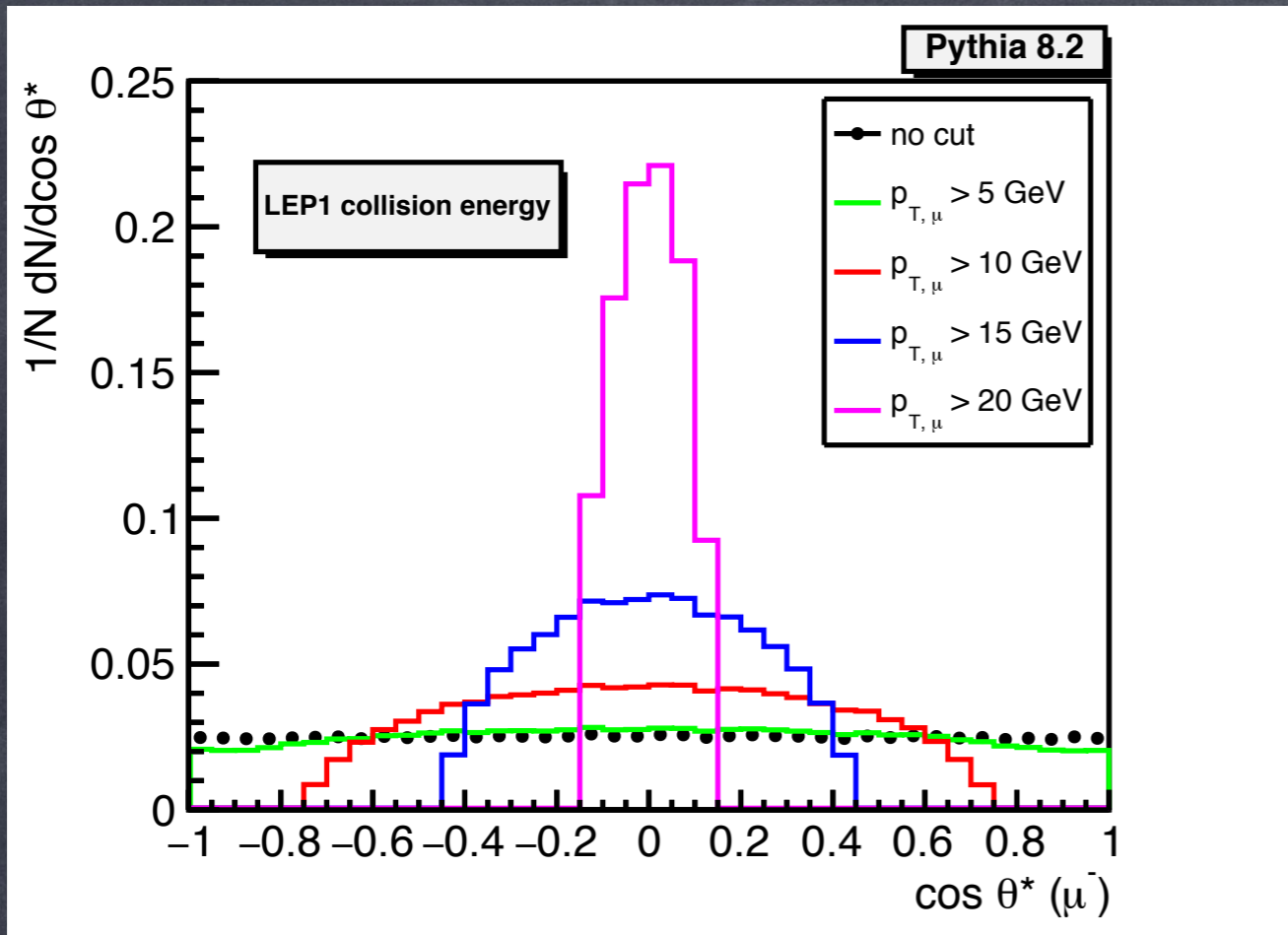


Z's at rest

Note: Normalized to unit area!

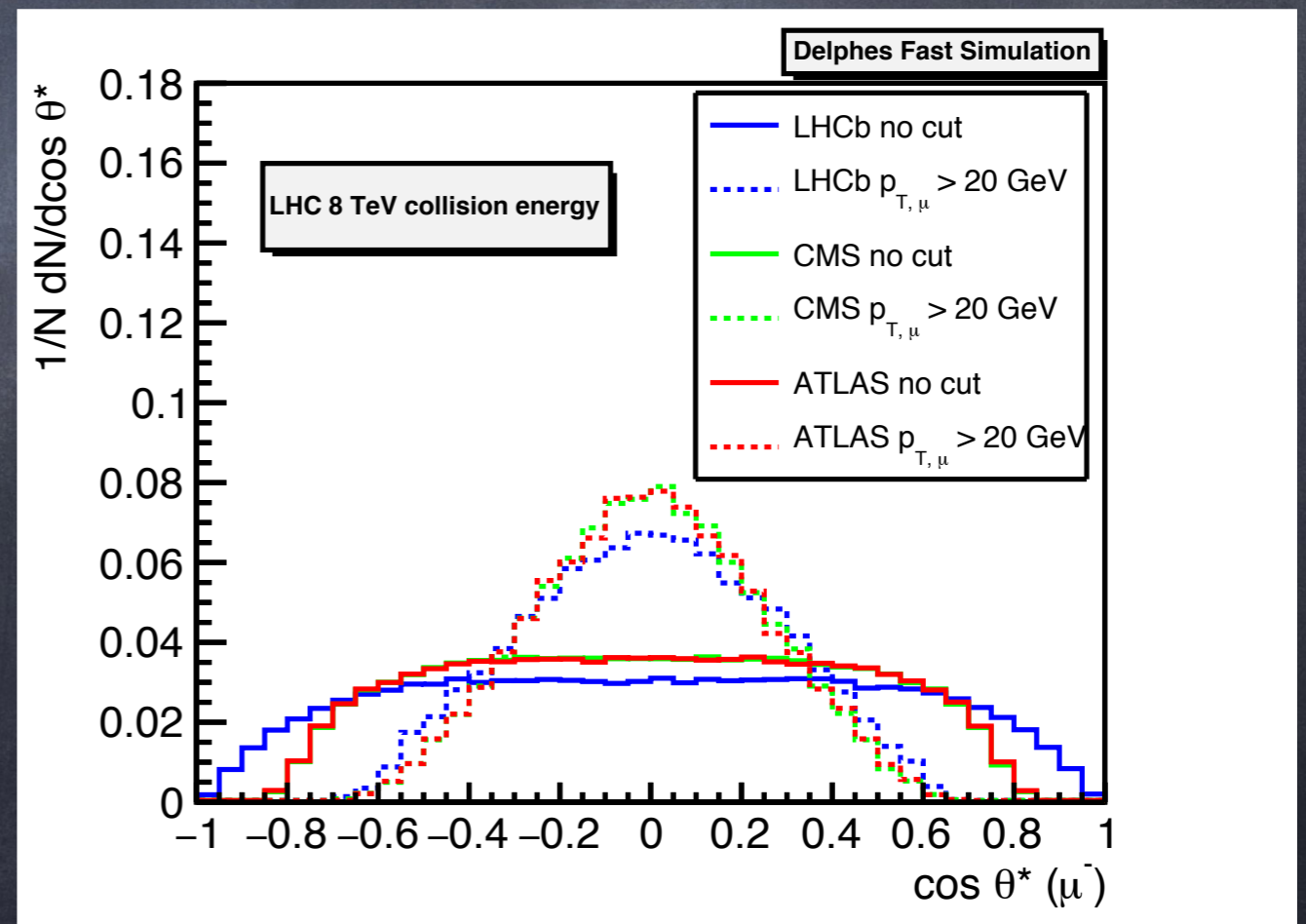
Z's boosted along beams





Low- p_T muons have large rapidity & escape

Lose backward-going muons for h at 90°



more predictions

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- 8.) And *no* $H(125) \rightarrow \mu^+ \mu^-$ if $(\alpha/\beta)^2 \ll 1$.

The fly in the ointment

(what you waited for!)

The fly in the ointment

$$B(h \rightarrow \mu^+ \mu^-) \cong B(\eta_A \rightarrow \mu^+ \mu^-) \cong 1$$

$$B(Z \rightarrow h\eta_A) = 0.014 \cos^2(\beta - \alpha) \text{ with } \beta, \alpha \text{ small}$$

$$\implies B(Z \rightarrow h\eta_A \rightarrow 4\mu) \cong 0.014$$

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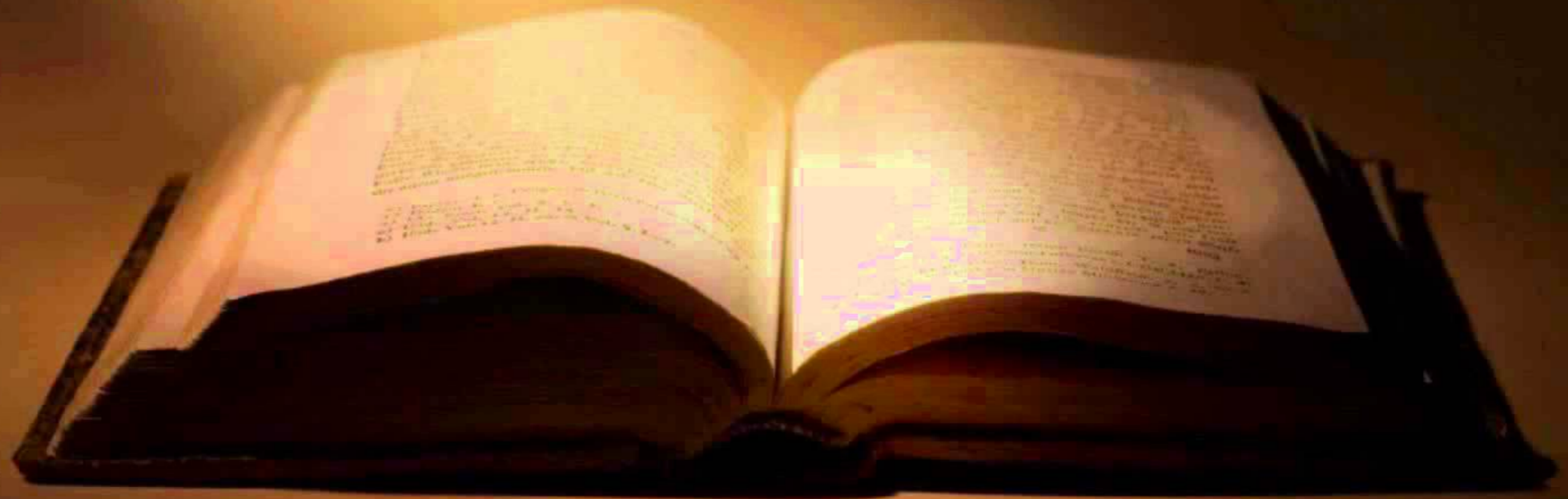
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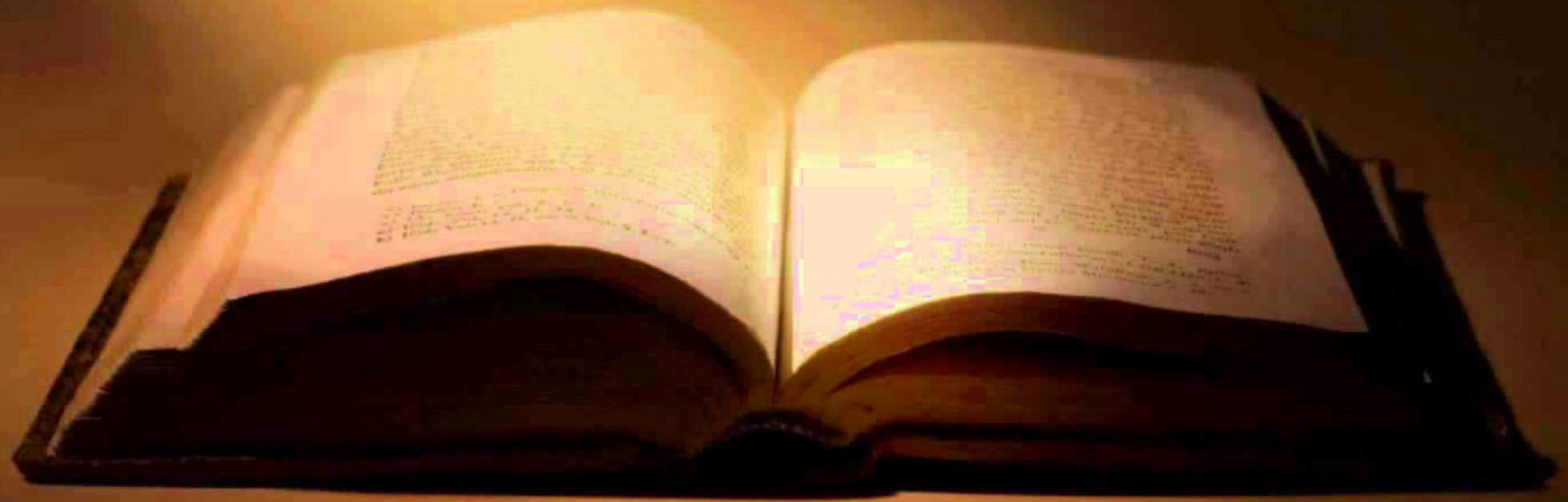
$$3300 \text{ times measured } B(Z \rightarrow 4\mu) = 4.2 \times 10^{-6}!!$$

Dead flies cause the ointment of the apothecary to send forth a stinking savour: so doth a little folly him that is in reputation for wisdom and honour. Ecclesiastes 10: 1



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Who? Me?



Some things we tried:

- 1.) Cannot raise M_{η_A} or alter
 $(B(h \rightarrow \mu^+ \mu^-) B(\eta_A \rightarrow \bar{b}b) \propto \cos^2 \alpha / \cos^2 \beta$
nor $B(h \rightarrow \bar{b}b) B(\eta_A \rightarrow \mu^+ \mu^-) \propto \sin^2 \alpha / \sin^2 \beta$
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and/or greatly increasing $B(H \rightarrow hh, \eta_A \eta_A)$.
- 2.) Use the Branco-Grimus-Lavoura (BGL)
mechanism to dilute $B(h, \eta_A \rightarrow \mu^+ \mu^-)$.

Here's what we tried for BGL...

BGL: Allow FCNC in 3rd generation
of up or down-Yukawa's; 4 possibilities:

FCNC:

$$\Gamma_1^{(u \text{ or } d)} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ \times & \times & \times \end{pmatrix}, \quad \Gamma_2^{(u \text{ or } d)} = \begin{pmatrix} \times & \times & \times \\ \times & \times & \times \\ 0 & 0 & 0 \end{pmatrix};$$

(or $1 \leftrightarrow 2$)

no FCNC:

$$\Gamma_1^{(d \text{ or } u)} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \times \end{pmatrix}, \quad \Gamma_2^{(d \text{ or } u)} = \begin{pmatrix} \times & \times & 0 \\ \times & \times & 0 \\ 0 & 0 & 0 \end{pmatrix}.$$

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for u -textures:

$$\Gamma_{h, \eta_A \bar{t}t} / \Gamma_{H \bar{t}t} \simeq \cot \beta \implies h, \eta_A \rightarrow gg (!!)$$

for d -textures:

$$B(h, \eta_A \rightarrow \bar{b}b) / B(h, \eta_A \rightarrow \mu^+ \mu^-) = 3(m_b/m_\mu)^2$$

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2.) For displayed FCNC textures in u -sector:

$$B(h, \eta_A \rightarrow \bar{c}c) \simeq 0.99 \gg B(h, \eta_A \rightarrow \mu^+ \mu^-; \bar{b}b)$$

inconsistent with no signal in inverted b -tag data;

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A no-go theorem?

bj to KL (SLAC, 1977):

There are no no-go theorems!

A challenge to theorists —

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Especially if the dimuon is also
seen in ATLAS, CMS, LHCb, L3...