

# Light Nonthermal Dark Matter: A Minimal Model & Detection Prospects

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# Outline:

- Introduction
- Minimal model (DM coupled to RH quarks)
- Detection prospects (direct, indirect, LHC)
- Minimal model 2 (DM coupled to LH quarks, monotop chirality)
- SUSY version (new opportunities, multicomponent DM)
- Outlook

Based on recent work:

PRD 88, 023525 (2013)    PRL 111, 051302 (2013)    PRD 89, 127305 (2014)  
PRD 91, 055033 (2015)    JHEP 1612, 046 (2016)

# Introduction:

Energy budget of the universe according to observations:

Two big problems to address:

## 1) Dark Matter (DM)

What is the nature of DM?

How was it produced?

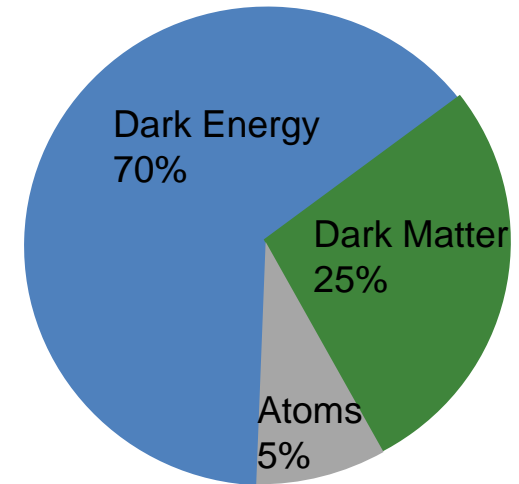
## 2) Baryon Asymmetry of Universe (BAU)

Why is it nonzero?

How was it generated?

Also, a possible coincidence puzzle:

Why the DM and baryons have comparable energy densities?



## A Minimal Model:

$B$  and  $L$  are accidental symmetries of SM at the perturbative level.

We adopt a bottom-up approach and consider a minimal extension of the SM with renormalizable  $B$  interactions:

R.A., B. Dutta PRD 88, 023525 (2013)

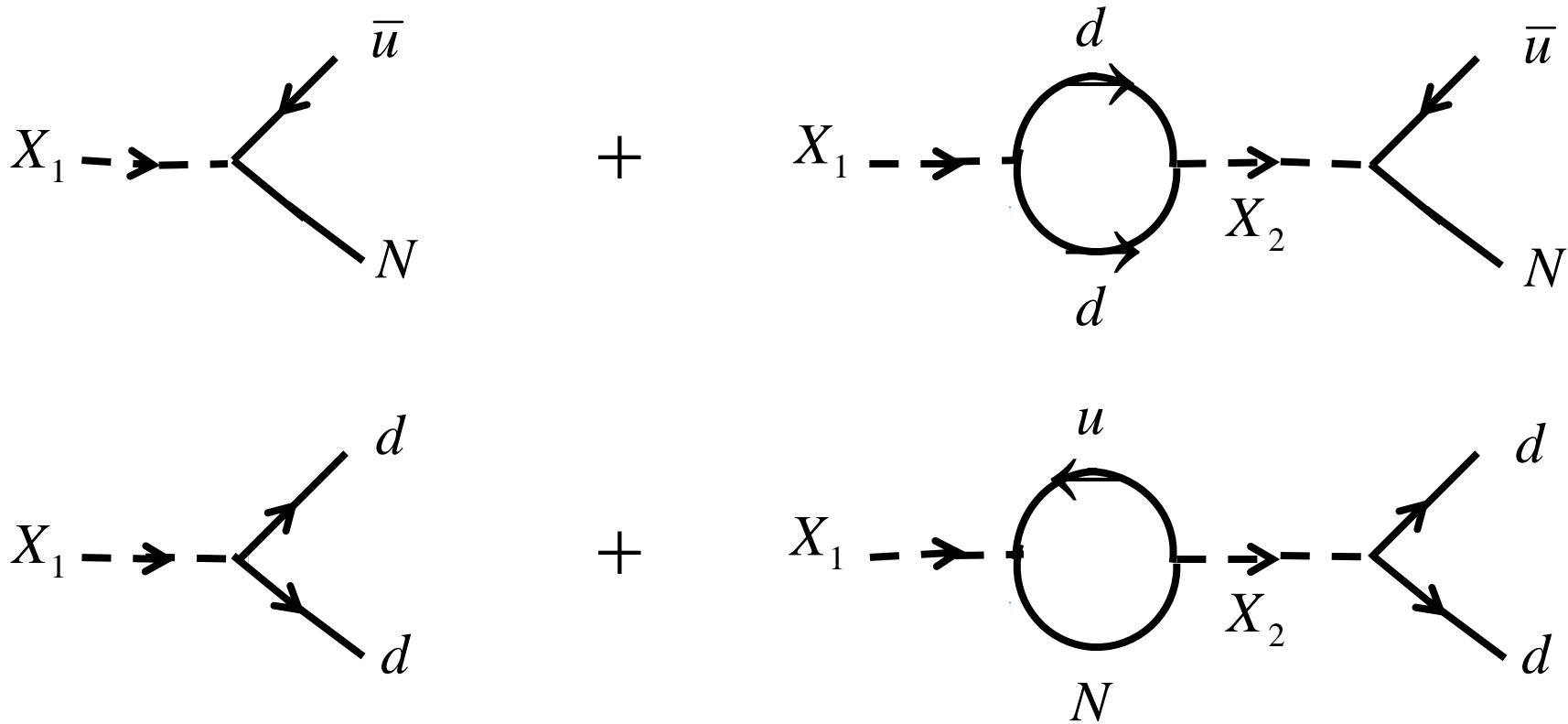
$$L_{new} = \lambda'_{\alpha ij} X_{\alpha} d_i^c d_j^c + \lambda_{\alpha i} N X_{\alpha}^* u_i^c + m_{\alpha}^2 |X_{\alpha}|^2 + \frac{m_N}{2} N N \\ + h.c. + kinetic terms$$

$X_{1,2}$  : **Iso-singlet** color-triplet scalars  $Y = +4/3$

$N$  : Singlet fermion

This is the minimum field content that is required to generate a nonzero baryon asymmetry via out-of-equilibrium decay of  $X$ .

E. Kolb, S. Wolfram NPB 172, 224 (1980); Erratum-ibid 195, 542 (1982)



$$\varepsilon_1 = \frac{1}{8\pi} \frac{\sum_{i,j,k} \text{Im}(\lambda_{1k}^* \lambda_{2k} \lambda'_{1ij} \lambda'_{2ij})}{\sum_{i,j} |\lambda'_{1ij}|^2 + \sum_k |\lambda_{1k}|^2} \frac{m_1^2}{m_1^2 - m_2^2}$$

$$\varepsilon_2 = \varepsilon_1 (1 \leftrightarrow 2)$$

$|\lambda_1 \lambda'_{12}|$  severely constrained by  $\Delta B = 2$  ,  $\Delta S = 2$  processes:

1)  $n - \bar{n}$  oscillations.

2) Double proton decay  $pp \rightarrow K^+ K^+$  .

For  $m_N \sim O(\text{GeV})$  ,  $m_X \sim O(\text{TeV})$  we must have:

$$|\lambda_1 \lambda'_{12}| < 10^{-6}$$

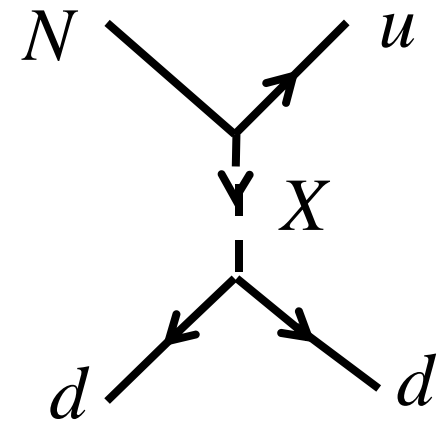
Experimental bounds on  $K_s^0 - \bar{K}_s^0$  and  $B_s^0 - \bar{B}_s^0$  oscillations are satisfied too.

Successful baryogenesis then needs nontrivial flavor structure of  $\lambda_i, \lambda'_{ij}$  or degeneracy in  $m_{X_1}, m_{X_2}$  .

R.A., B. Dutta, K. Sinha PRD 82, 035004 (2010)

$X$  mediates a 4-fermion interaction:

$$\frac{\lambda\lambda'}{m_X^2} N u_i^c d_j^c d_k^c$$



This operator results in the following decays:

$$m_N > m_p + m_e : N \rightarrow p + e^- + \bar{\nu}_e, \quad \bar{p} + e^+ + \nu_e$$

$$m_N < m_p - m_e : p \rightarrow N + e^+ + \nu_e, \quad N + e^- + \bar{\nu}_e$$

$N$  is stable and becomes a viable DM candidate if:

$$m_p - m_e \leq m_N \leq m_p + m_e$$

The condition is stable against radiative corrections for:

$$\lambda < O(10^{-1})$$

**Stability of DM candidate is tied to the stability of proton.**

No additional symmetry, like R-parity, is invoked.

**Odd & even number of DM particles produced from SM particles.**

$N$  quanta produced from/annihilate to quarks in the early universe:

$$m_N < T \ll m_X : \Gamma \sim (|\lambda|^4 + |\lambda|^2 |\lambda'|^2) \frac{T^5}{m_X^4}$$

$$H \sim T^2 / M_P \Rightarrow T_{dec} \sim (m_X^4 / \lambda^4 M_P)^{1/3}$$

$$|\lambda|, |\lambda'| < 1, m_X \sim O(\text{TeV}) :$$

$$T_{dec} > 10 \text{ MeV}$$

Thermal freeze-out results in overproduction of DM.

B. Lee, S. Weinberg PRL 39, 165 (1977)

**Nonthermal mechanism needed to obtain DM relic abundance.**



# Detection Prospects:

Direct detection:

Effective interaction:

$$\frac{1}{m_X^2 - (p_N - p'_u)^2} (\bar{\psi}_N P_L \psi_N) (\bar{\psi}_u P_R \psi_u)$$

Spin-independent piece:

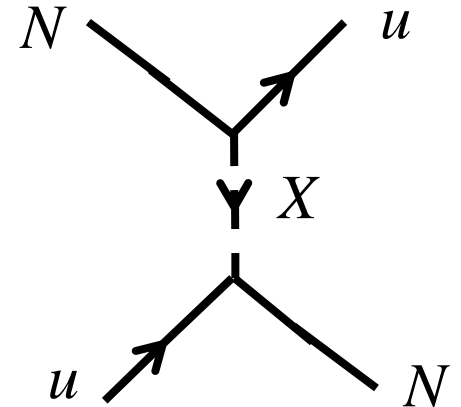
$$\frac{1}{m_X^4} (\bar{\psi}_N \gamma^\mu \partial^\nu \psi_N) \left[ (\bar{\psi}_u \gamma_\mu \partial_\nu \psi_u) - (\partial_\nu \bar{\psi}_u \gamma_\mu \psi_u) \right]$$

Spin-dependent piece:

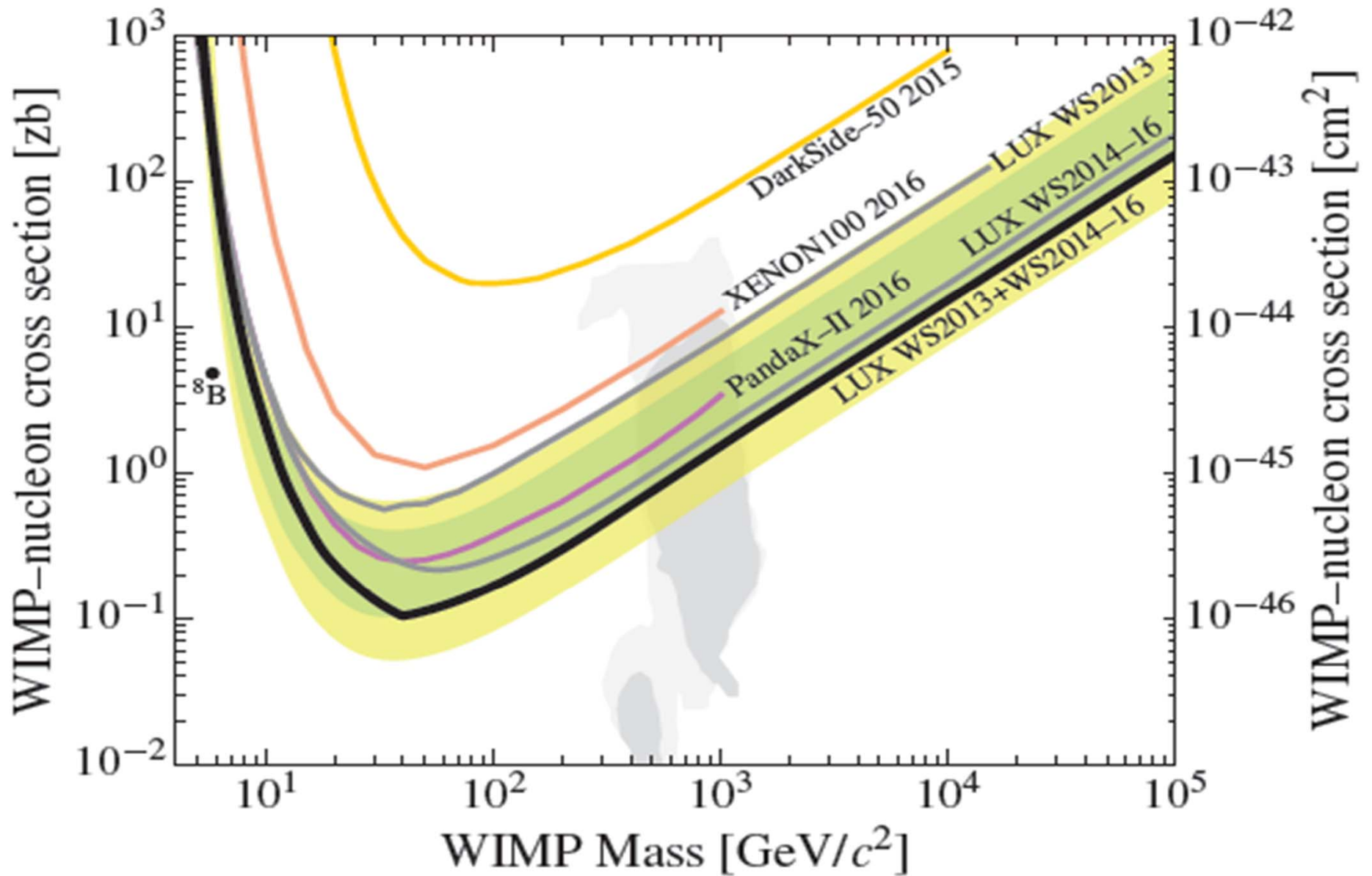
$$\frac{1}{m_X^2} (\bar{\psi}_N \gamma^\mu \gamma^5 \psi_N) (\bar{\psi}_u \gamma_\mu \gamma^5 \psi_u)$$

$$\sigma_{SI} \sim |\lambda|^4 \frac{O(\text{GeV})^6}{m_X^8}$$

$$\sigma_{SD} \sim |\lambda|^4 \frac{O(\text{GeV})^4}{m_X^4}$$

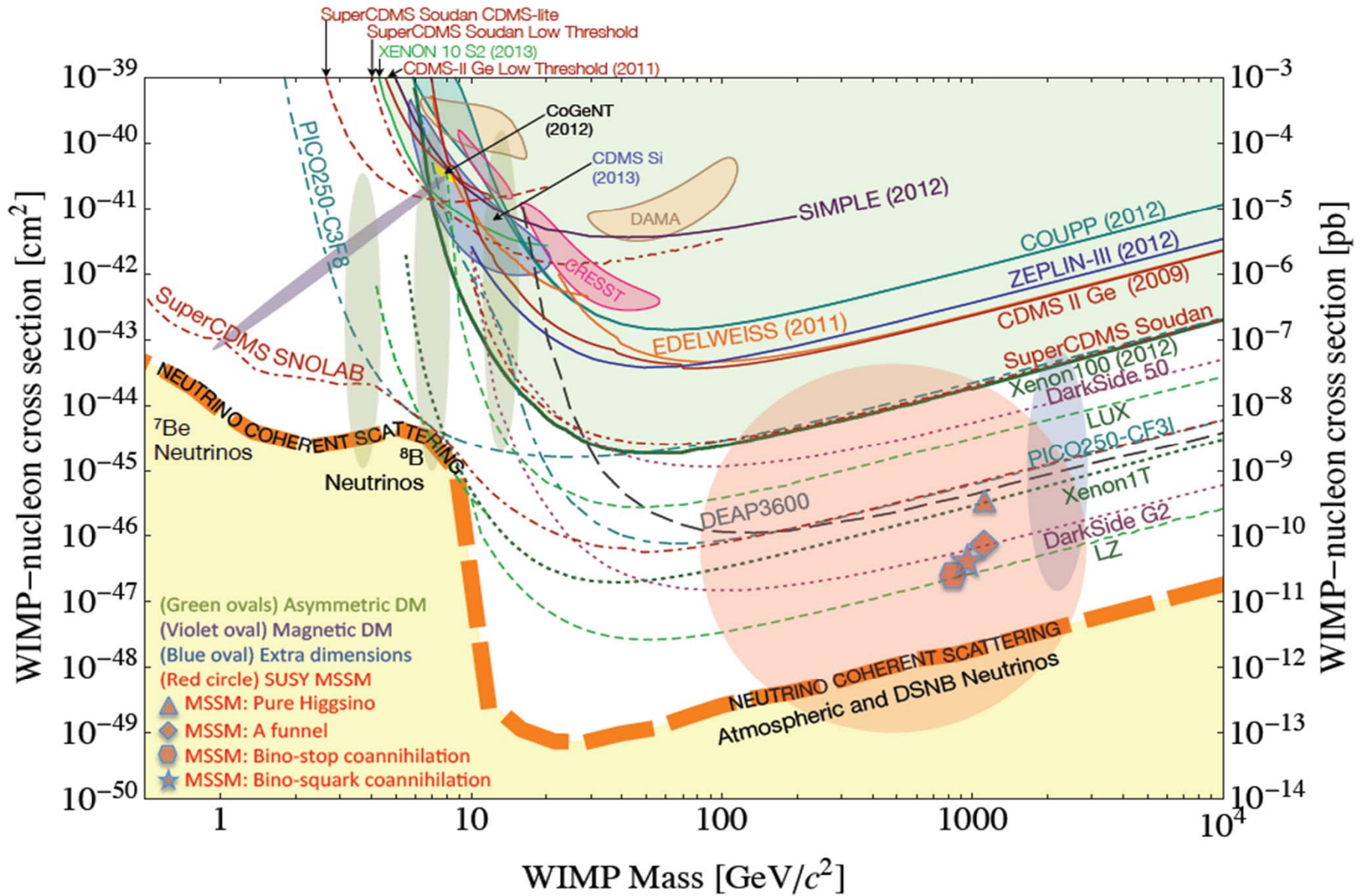


LUX Collaboration PRL 118, 021303 (2017)

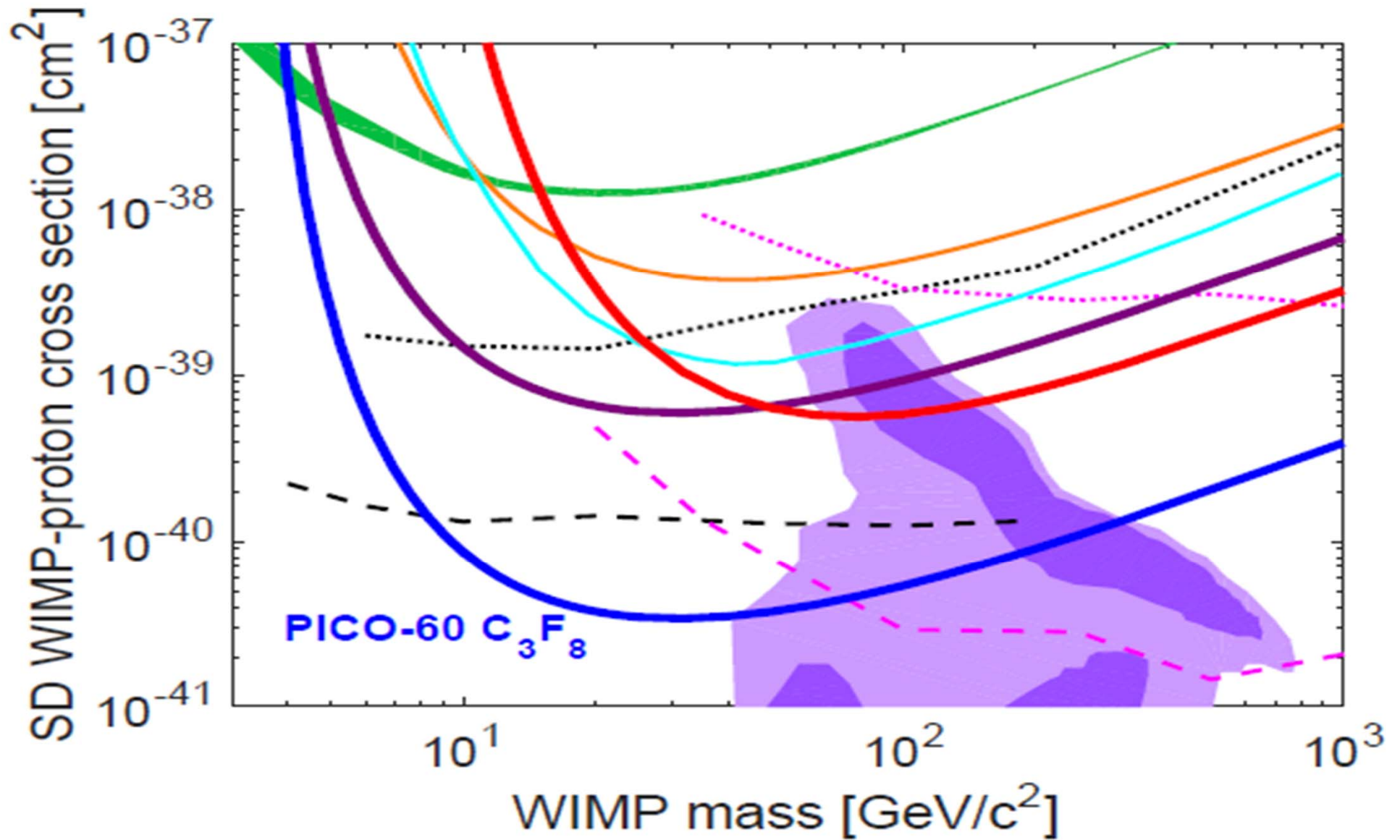


$$m_X \sim O(\text{TeV}) \Rightarrow \sigma_{SI} < 10^{-52} \text{ cm}^2$$

# Snowmass CF1 Summary (2013)

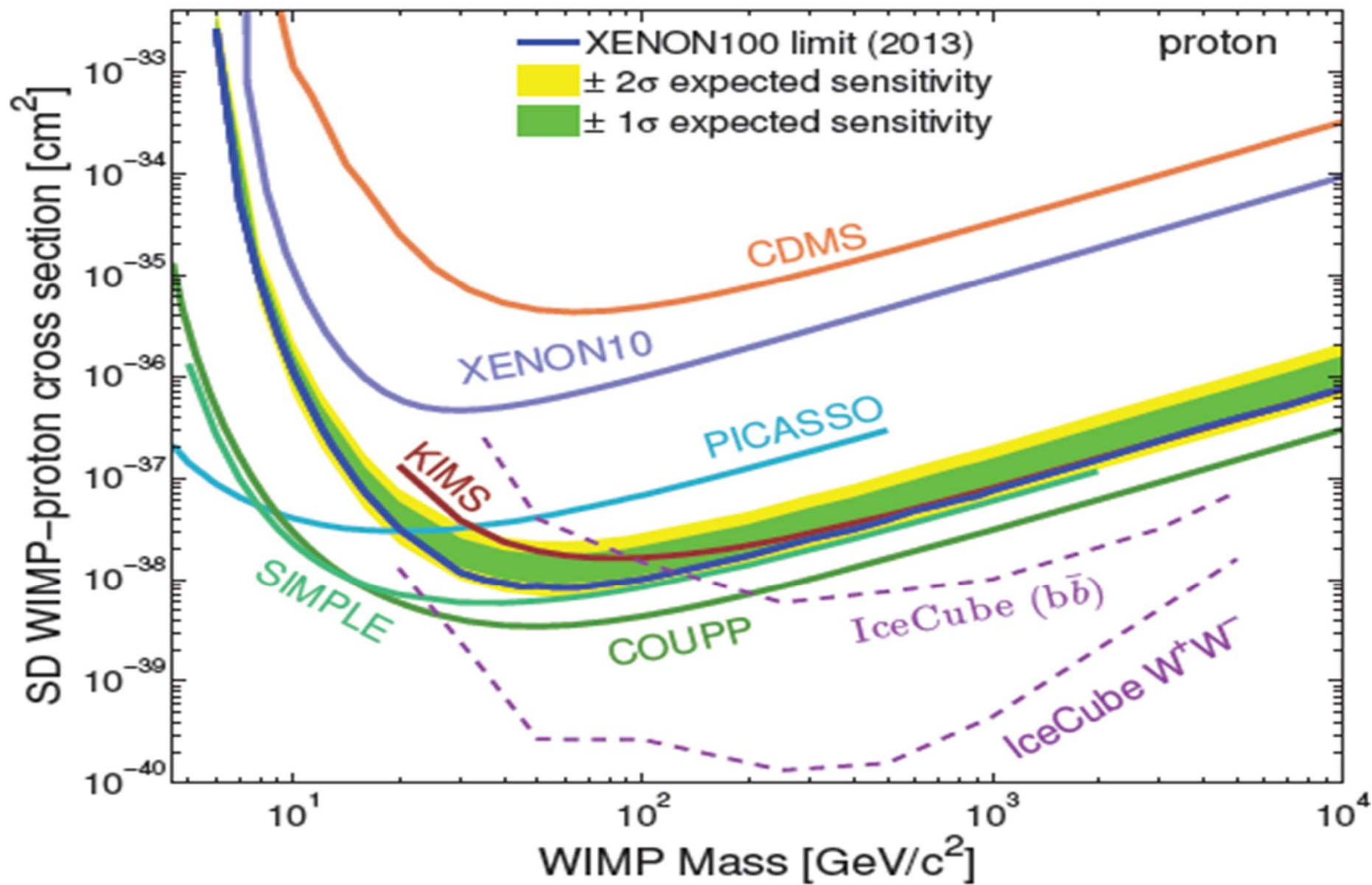


$$m_X \sim O(\text{TeV}) \Rightarrow \sigma_{SI} < 10^{-52} \text{ cm}^2$$

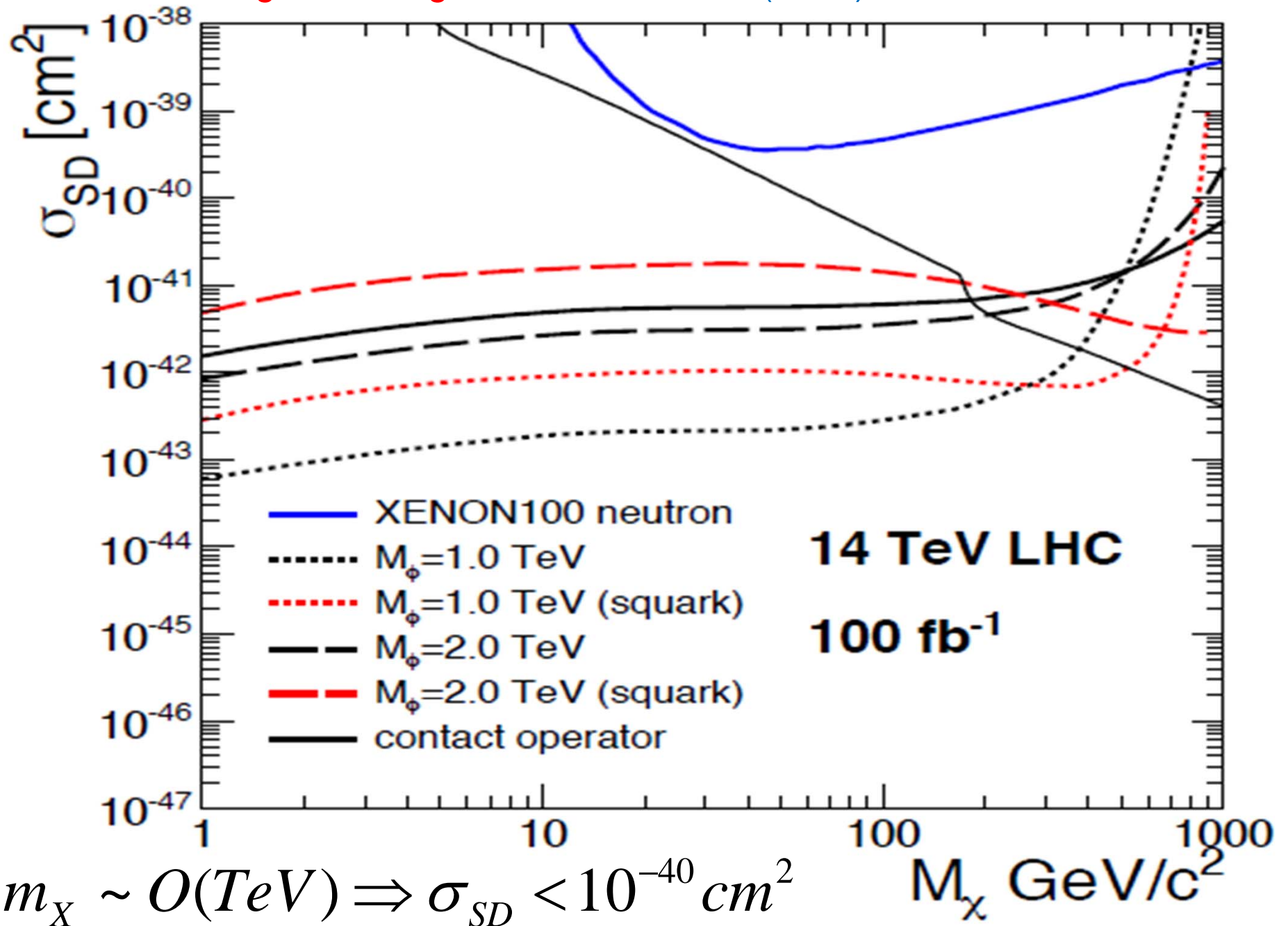


$$m_X \sim O(\text{TeV}) \Rightarrow \sigma_{SD} < 10^{-40} \text{ cm}^2$$

# Snowmass CF1 Summary (2013)

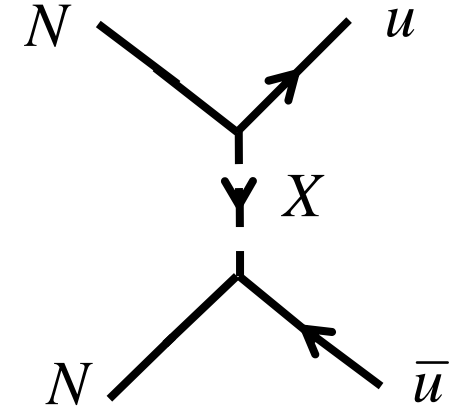


$$m_X \sim O(\text{TeV}) \Rightarrow \sigma_{SD} < 10^{-40} \text{ cm}^2$$



## Indirect detection:

$$\langle \sigma_{ann} v \rangle \sim |\lambda|^4 \frac{|\vec{p}|^2}{m_X^4}$$



$$m_X \sim O(\text{TeV}) \Rightarrow \langle \sigma_{ann} v \rangle \ll 10^{-31} \text{ cm}^3 / \text{s}$$

Much smaller than limits on galactic/extragalactic DM annihilation into gamma-rays and neutrinos.

Neutrino signal from solar DM annihilation depends on  $\sigma_{SD,SI}$  and  $\sigma_{ann}$ . However, it is negligible too since:

- 1) Both capture and annihilation are suppressed.
- 2) Evaporation dominates for  $O(\text{GeV})$  DM mass.

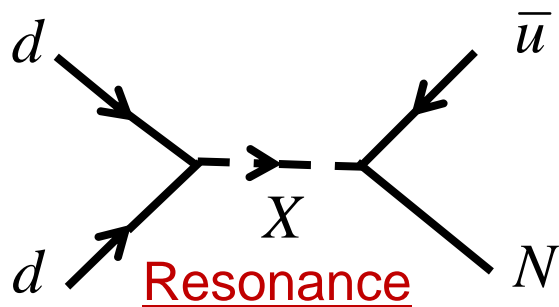
## Collider signals: (see the talk by S. Undleeb)

Both **odd** & even number of DM particles are produced from the interactions of the SM particles:

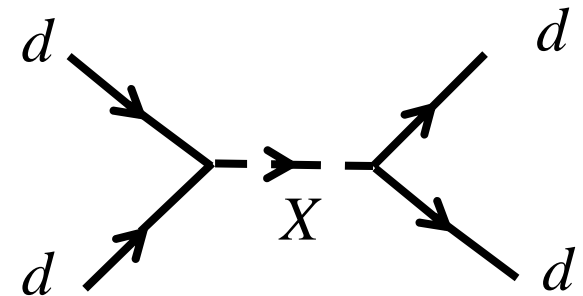
**Monojets** (including **monotops**) & dijets plus missing energy.

B. Dutta, Y. Gao, T. Kamon PRD 89, 096009 (2014)

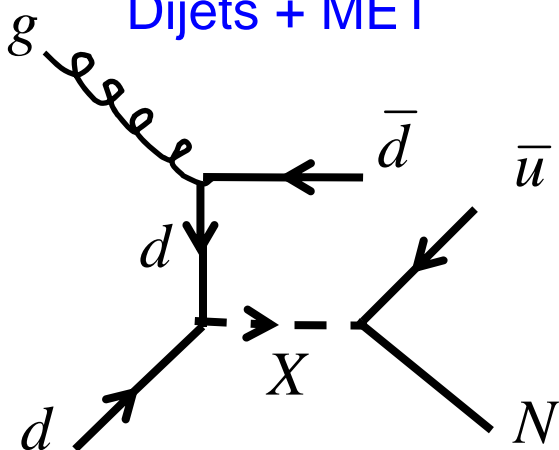
### Monojets + MET



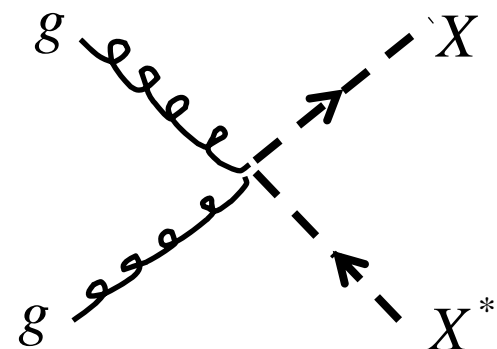
### Dijets



### Dijets + MET

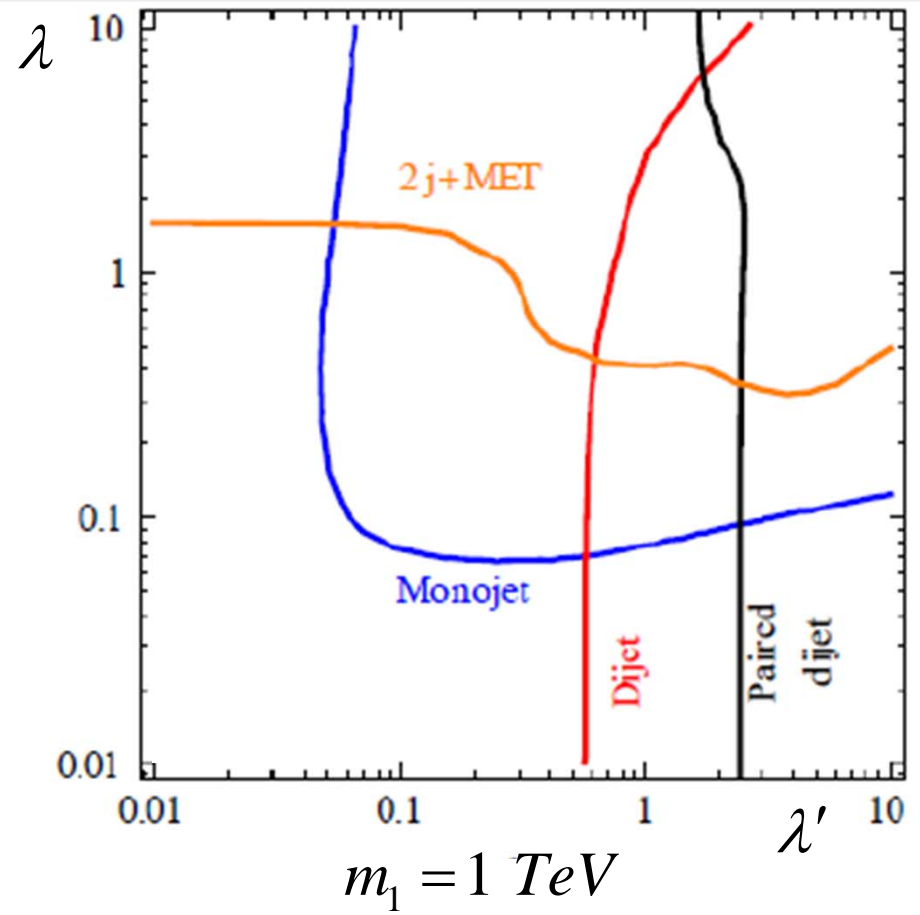
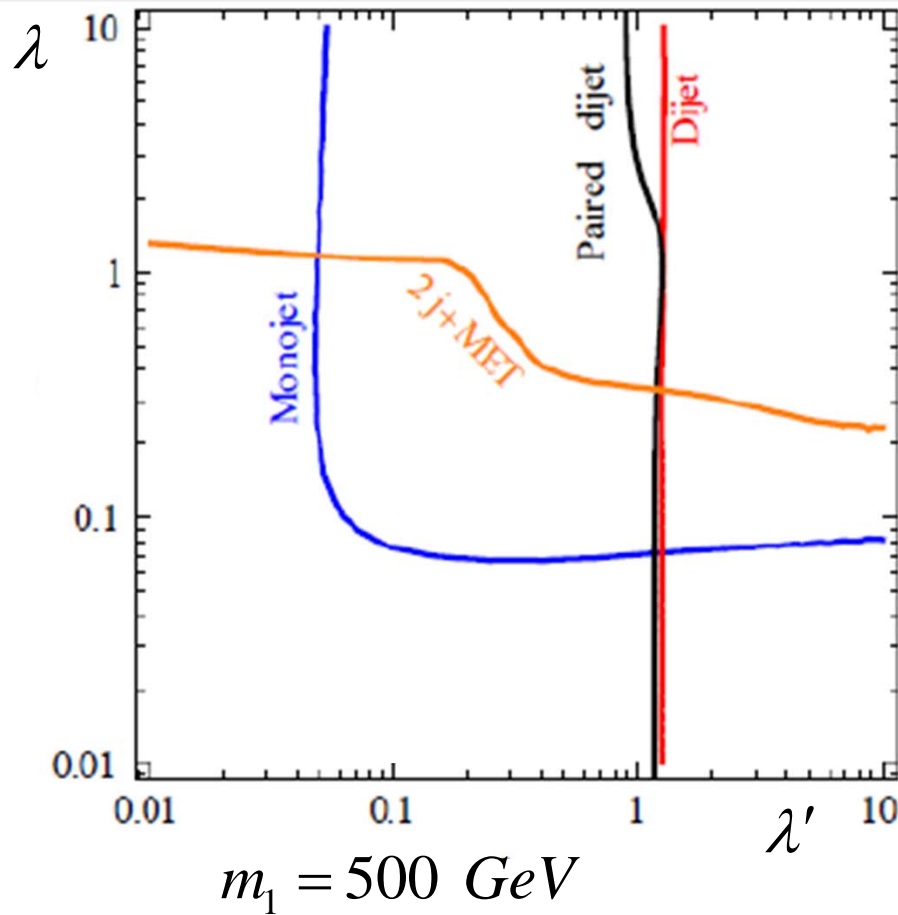


### Paired dijets + MET





Combined collider bounds (assuming single values for  $\lambda$  and  $\lambda'$ ):



B. Dutta, Y. Gao, T. Kamon PRD 89, 096009 (2014)

## Another Minimal Model:

In this model the DM is coupled to **LH quarks**.  
The model has a slightly larger field content.

R.A., *et al* JHEP 1612, 046 (2016)

$$L_{new} = y_{1\alpha j} N X_{\alpha}^* Q_i + y_{2\alpha i} X_{\alpha} \bar{Y} d_i^c + y_{3\alpha i} X_{\alpha} Y d_i^c + m_Y Y \bar{Y} \\ + m_{\alpha}^2 |X_{\alpha}|^2 + \frac{m_N}{2} N N + h.c. + \text{kinetic terms}$$

$X_{1,2}$  : **Iso-doublet** color-triplet scalars  $Y = +1/3$

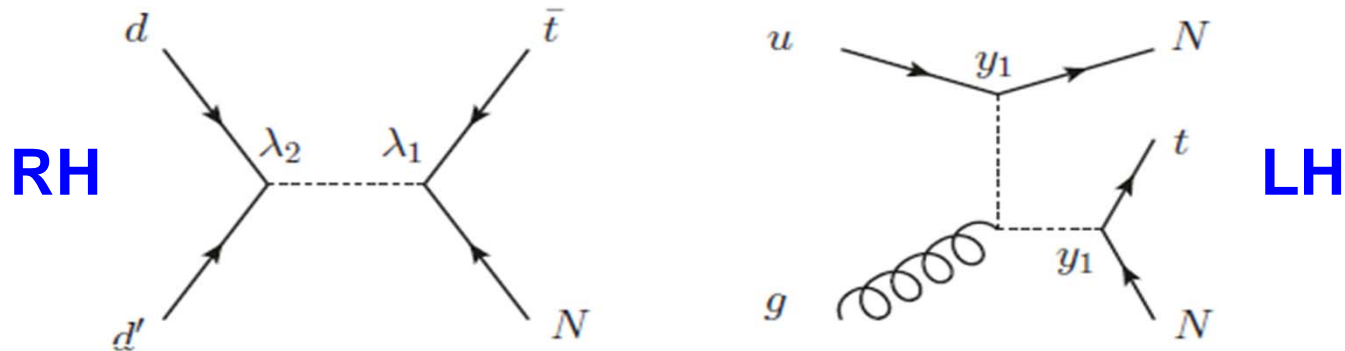
$Y, \bar{Y}$  : **Iso-doublet** fermions  $Y = +1, -1$   $m_Y > 45 \text{ GeV}$

$N$  : **Singlet** fermion

$N$  is the DM candidate if  $m_N < m_Y$ .

This model predicts **LH** monojet/monotop events.

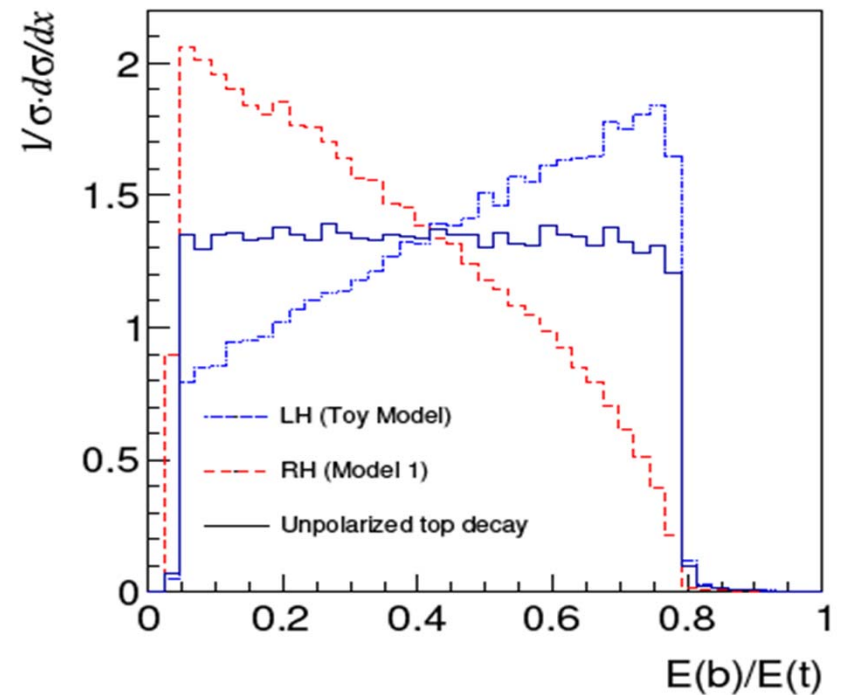
This is different from most of the simplified DM models.



Top chirality can be used as a discriminator between the models:

$$t_{R,L} \rightarrow W^+ + b_L$$

R.A., et al JHEP 1612, 046 (2016)



# Supersymmetric Version:

Extension to supersymmetry is straightforward:

$$W_{new} = \lambda'_{\alpha ij} X_{\alpha} d_i^c d_j^c + \lambda_{\alpha i} N \bar{X}_{\alpha} u_i^c + m_{\alpha} X_{\alpha} \bar{X}_{\alpha} + \frac{m_N}{2} N N$$

$X_{\alpha}, \bar{X}_{\alpha}$  : Iso-singlet color-triplet superfields  $Y = +4/3, Y = -4/3$

$N$  : Singlet superfield

The model can lead to thermal and non-thermal baryogenesis.

Babu, Mohapatra, Nasri PRL 98, 161301 (2007)

R.A., B. Dutta, K. Sinha PRD 82, 035004 (2010)

It also has a real scalar DM candidate  $\tilde{N}$  protected by R-parity:

$$m_{\tilde{N}}^2 = m_N^2 + \tilde{m}^2 \pm B m_N$$

The lighter of the two components of  $\tilde{N}$  is a DM candidate.

$m_{\tilde{N}}$  can have any value irrespective of  $m_N$  .

The prospect for direct detection of  $\tilde{N}$  is good.

R.A., B. Dutta, R. N. Mohapatra, K. Sinha PRL 111, 051302 (2013)

The model allows natural realization of the multicomponent DM scenario if  $m_N \approx 1 \text{ GeV}$  .

Two components of DM arise from the same superfield, both of which can in principle be detected.

Prospects for direct and indirect detection of  $\tilde{N}$  are good.

$N$  may be detected at the LHC through monojet/monotop events.

Both components of DM are produced in nonthermal fashion.

For example, from invisible decay of the gravitinos.

R.A., B. Dutta, F. S. Queiroz, L. E. Strigari, M-Y Wang PRD 91, 055033 (2015)

## Outlook:

- A minimal SM extension for TeV scale baryogenesis presented.
- It can accommodate an  $O(GeV)$  fermionic DM candidate.
- DM coupled to RH quarks, abundance produced non-thermally.
- Distinct monojet/monotop signals detectable at the LHC.
- A second model with DM coupling to LH quarks presented.
- Monotop chirality can be used to distinguish the two models.
- SUSY version has improved direct/indirect detection prospects.
- It can lead to a natural realization of multicomponent DM.