

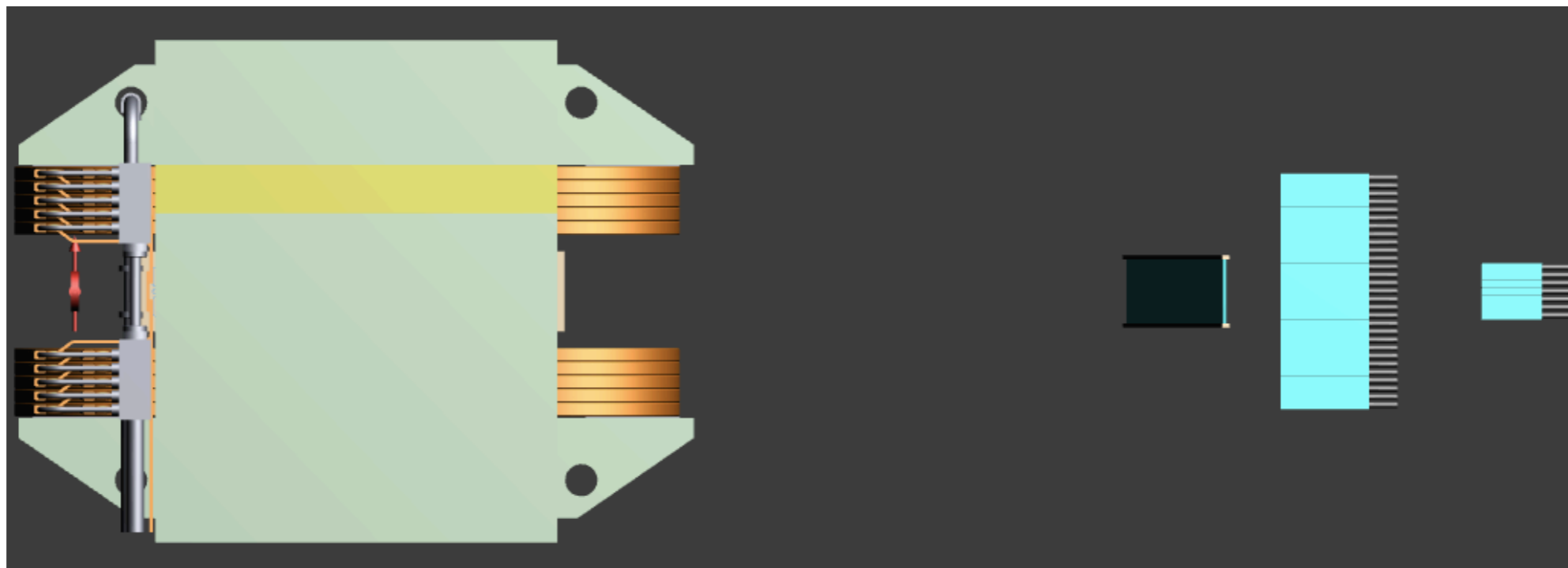


Istituto Nazionale di Fisica Nucleare



PADME electromagnetic calorimeter

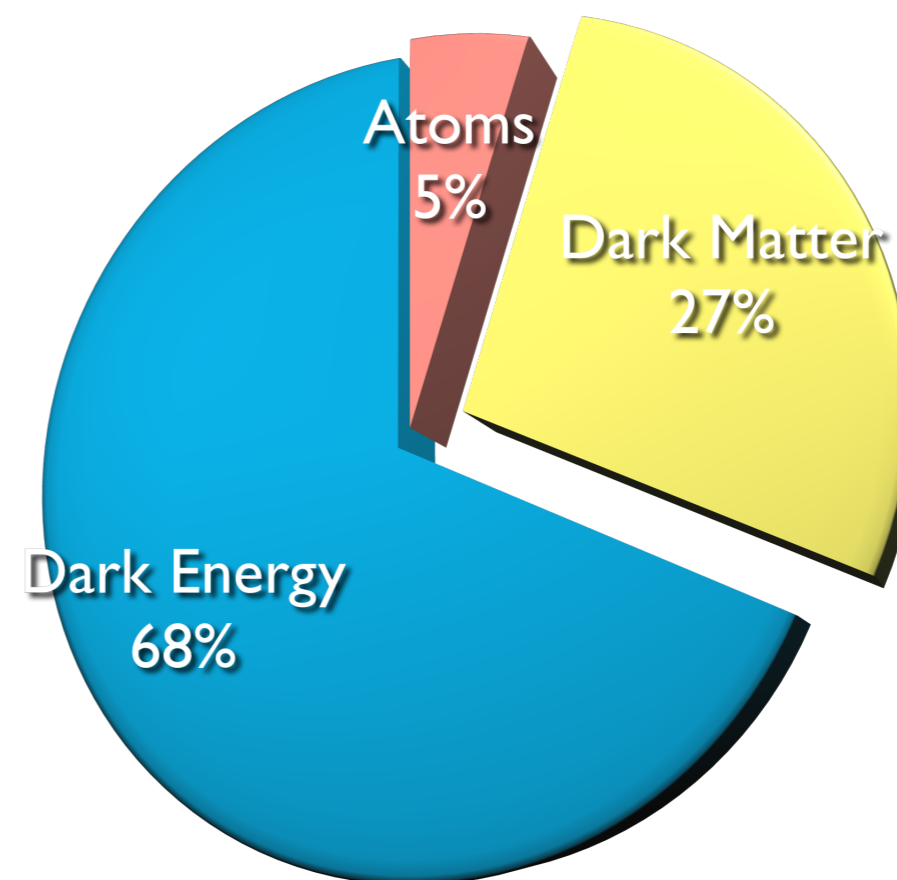
Gabriele Piperno



The Dark Matter problem

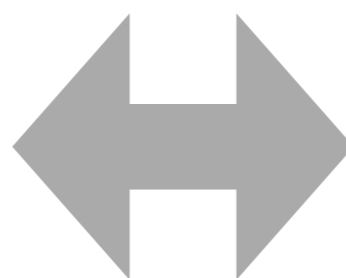
Evidences:

- spiral galaxies
- Cosmic Microwave Background
- gravitational lensing
- galaxy clusters
- Big Bang Nucleosynthesis
- large scale structures



Properties:

- stable (half life \sim universe age)
- cold (non relativistic)
- gravitational force
- non baryonic



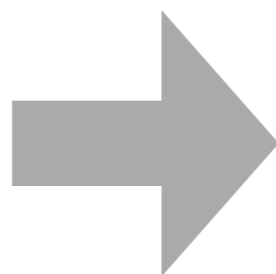
Open questions:

- DM nature
- interaction(s) w/ SM
- A whole new dark sector?
- dark sector forces?

Dark Photon

Possible solution to the DM elusiveness:
 DM does not interact directly w/ SM, but only by means of “portals”.

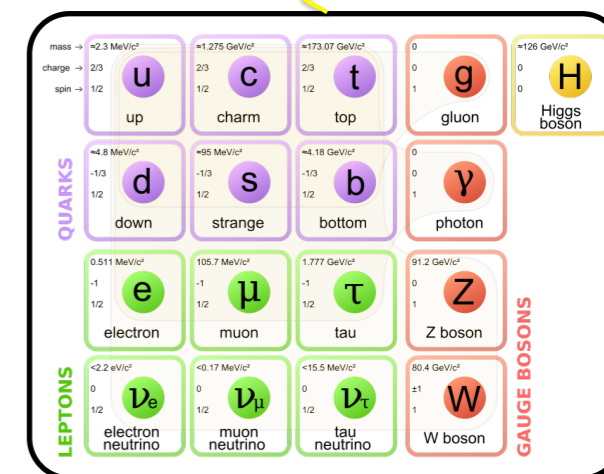
The simplest model adds a U(1) gauge symmetry and its boson: the Dark Photon A'



- SM particles are neutral under this symmetry
- new field couples to the SM w/ effective charge ϵq

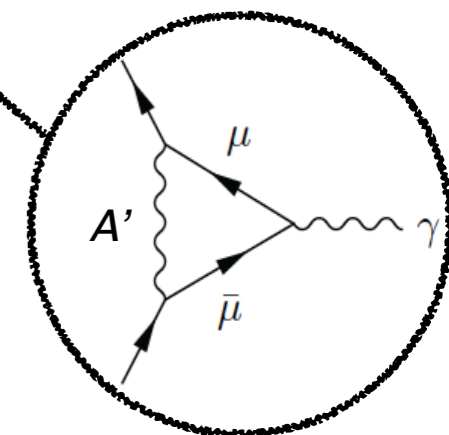
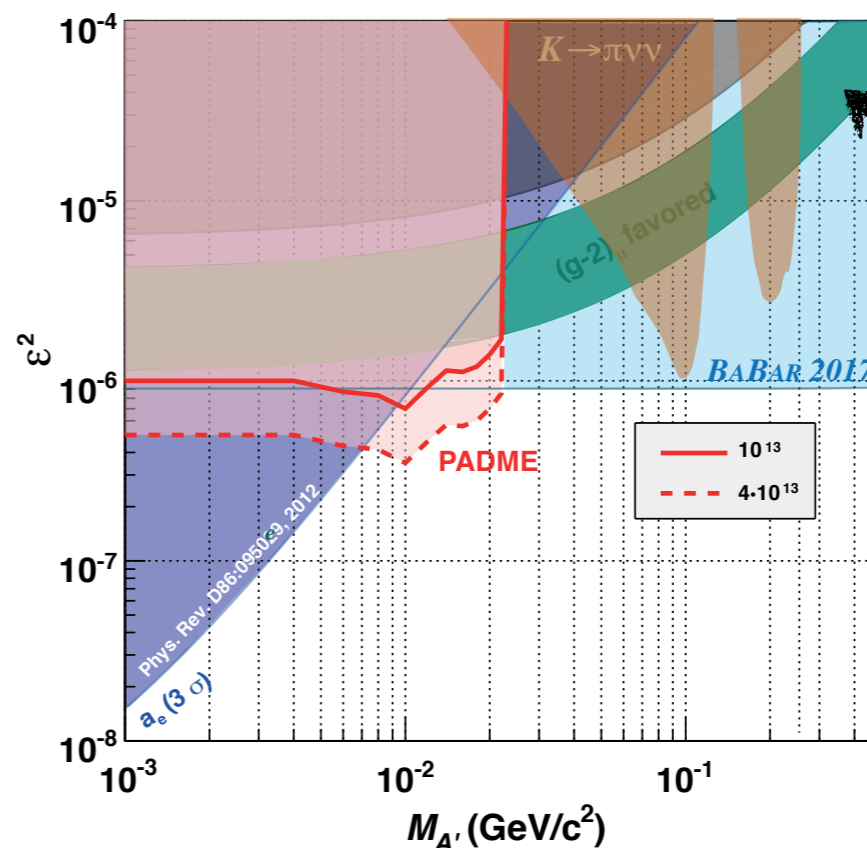
Dark Sector

Portals (A')



Depending on the model, the A' could (partially) explain the $(g-2)_\mu$ discrepancy and the ^8Be anomaly (see backup)

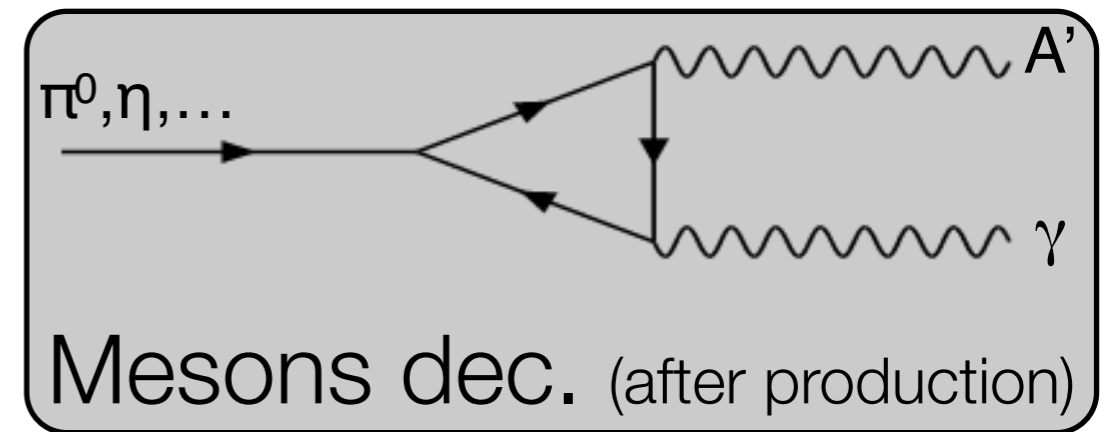
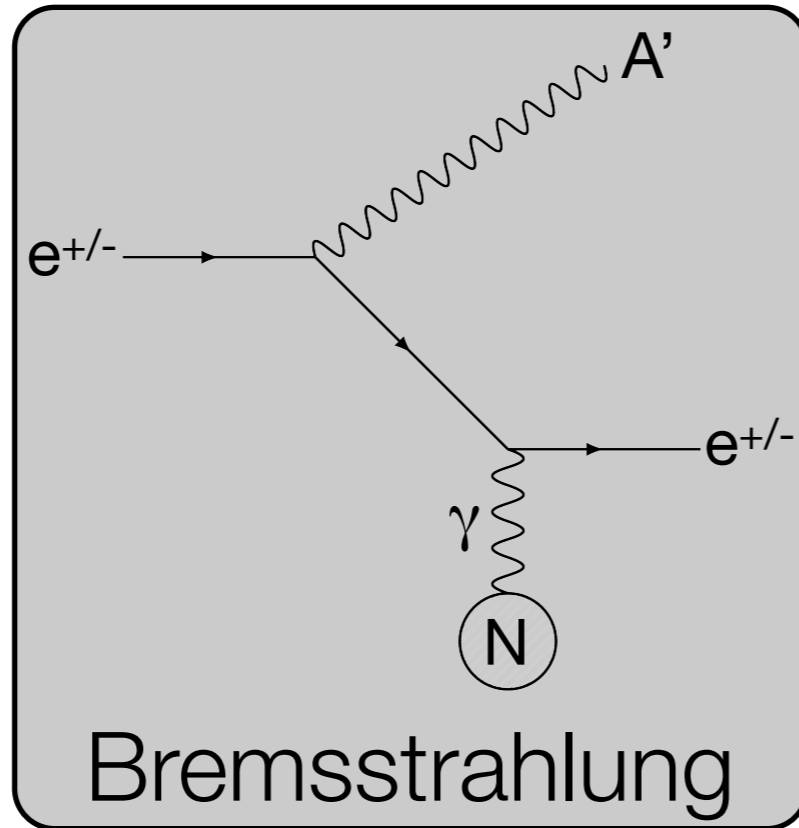
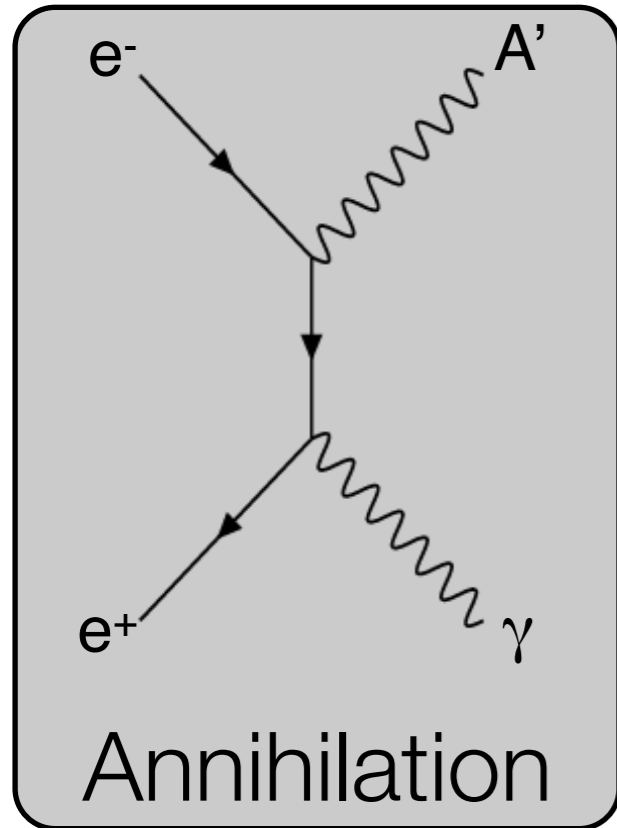
Exclusion plot assuming A' as A' model the one presented above



Excluded as only solution of the $(g-2)_\mu$ problem

Dark photon production and decays

In e^+/e^- collisions Dark Photon can be produced in 3 main ways:



Visible decays

If DM particles w/ $m_{\text{DM}} \leq m_{A'}/2$ do not exist:

- $A' \rightarrow \text{SM}$ (visible) decays
 - up to $2m_\mu$, $\text{BR}(e^+e^-) = 1$ (if $m_{A'} > 2m_e$)

A' lifetime proportional to:

$$1/(\alpha \varepsilon^2 m_{A'})$$

Invisible decays

If DM particles w/ $m_{\text{DM}} \leq m_{A'}/2$ exist:

- $A' \rightarrow \text{DM}$ (invisible) w/ (likely) $\text{BR} \approx 1$
- SM decays suppressed by a factor ε^2

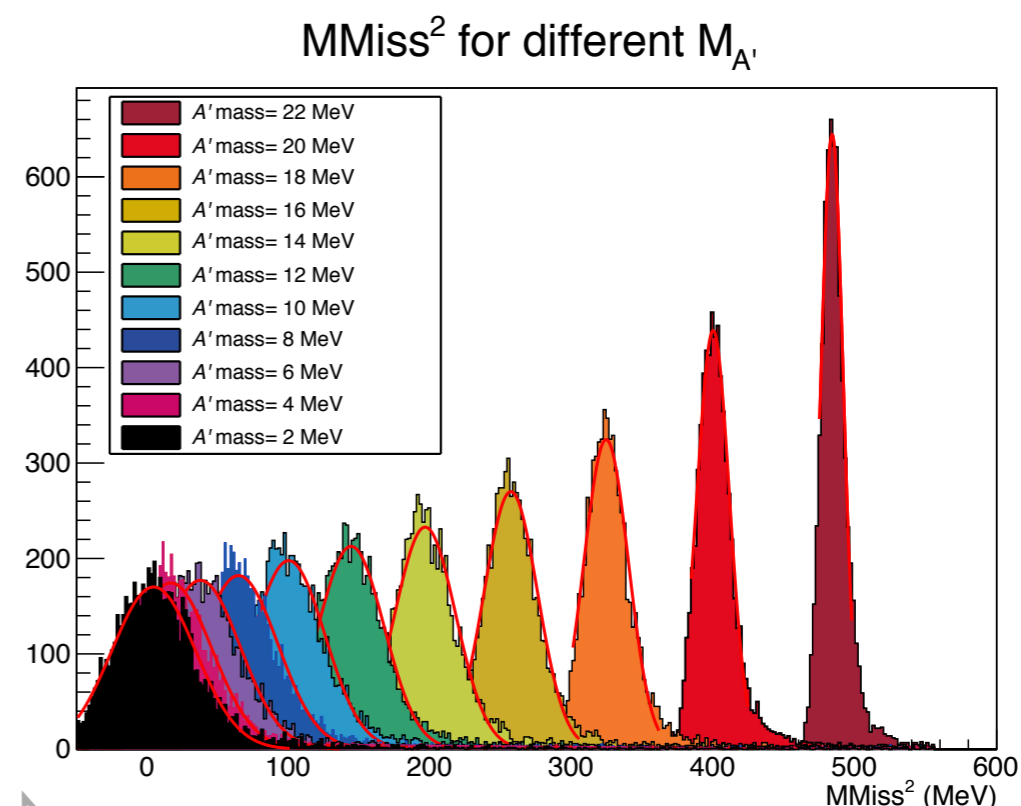
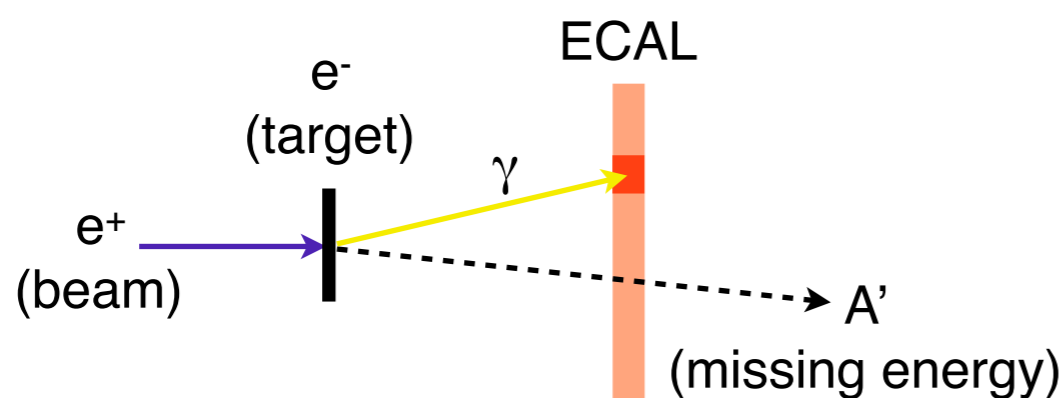
A' lifetime proportional to:

$$1/(\alpha_{\text{D}} m_{A'})$$

α_{D} : A' coupling constant to the Dark Sector

The PADME approach

A' search in e^+e^- annihilations looking for missing mass (invisible decay) in a kinematically constrained condition



- known beam energy and position
- measured photon energy and position

$$m_{\text{Miss}}^2 = (\mathbf{P}_{\text{beam}} + \mathbf{P}_e - \mathbf{P}_\gamma)^2$$

- minimal model dependent assumptions: A' couples to leptons
- can set limits on coupling of any new light particle that can be produced in e^+e^- annihilation: **Dark Photon, Axion Like Particles, Dark Higgs**

The detector

active target

- diamond (low z)
- 100 μm thickness
- info on beam time, spot size, e^+ number

(high energy) e^+/e^- veto

- plastic scintillator bars

small angle calorimeter

- 25 $3 \times 3 \times 14 \text{ cm}^3$ PbF_2
- 0-20 mrad ang. cov.
- fast: 3 ns Cher. light signals

electromagnetic calorimeter

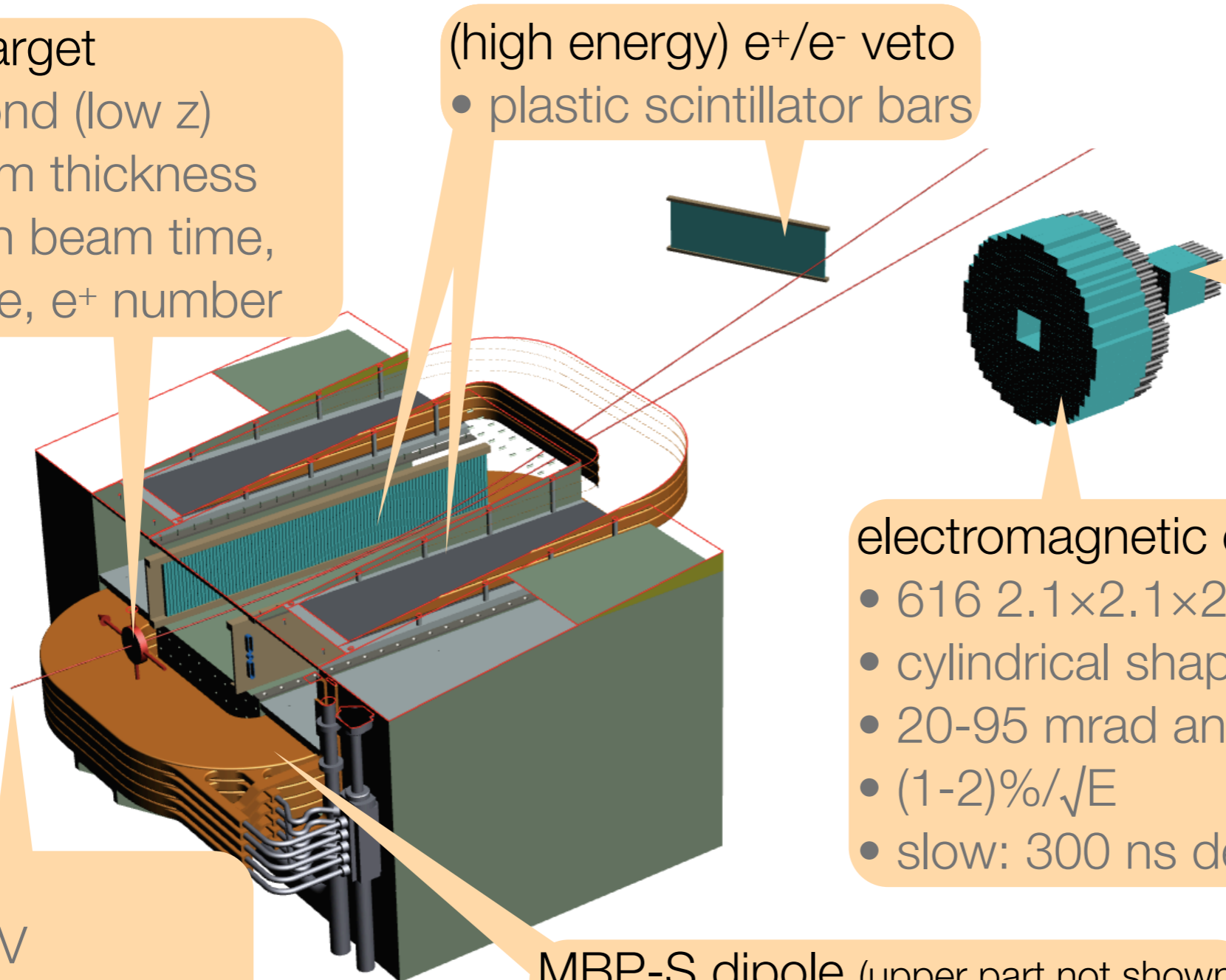
- 616 $2.1 \times 2.1 \times 23 \text{ cm}^3$ BGO
- cylindrical shape w/ central hole
- 20-95 mrad ang. cov.
- $(1-2)\%/\sqrt{E}$
- slow: 300 ns dec. time for scint. light

e^+ beam

- 550 MeV
- 5000 e^+ per bunch
- 40 ns bunch, every 20 ms

MBP-S dipole (upper part not shown)

- 0.5 T
- 1 m length. \times 23 cm gap



ECAL overview

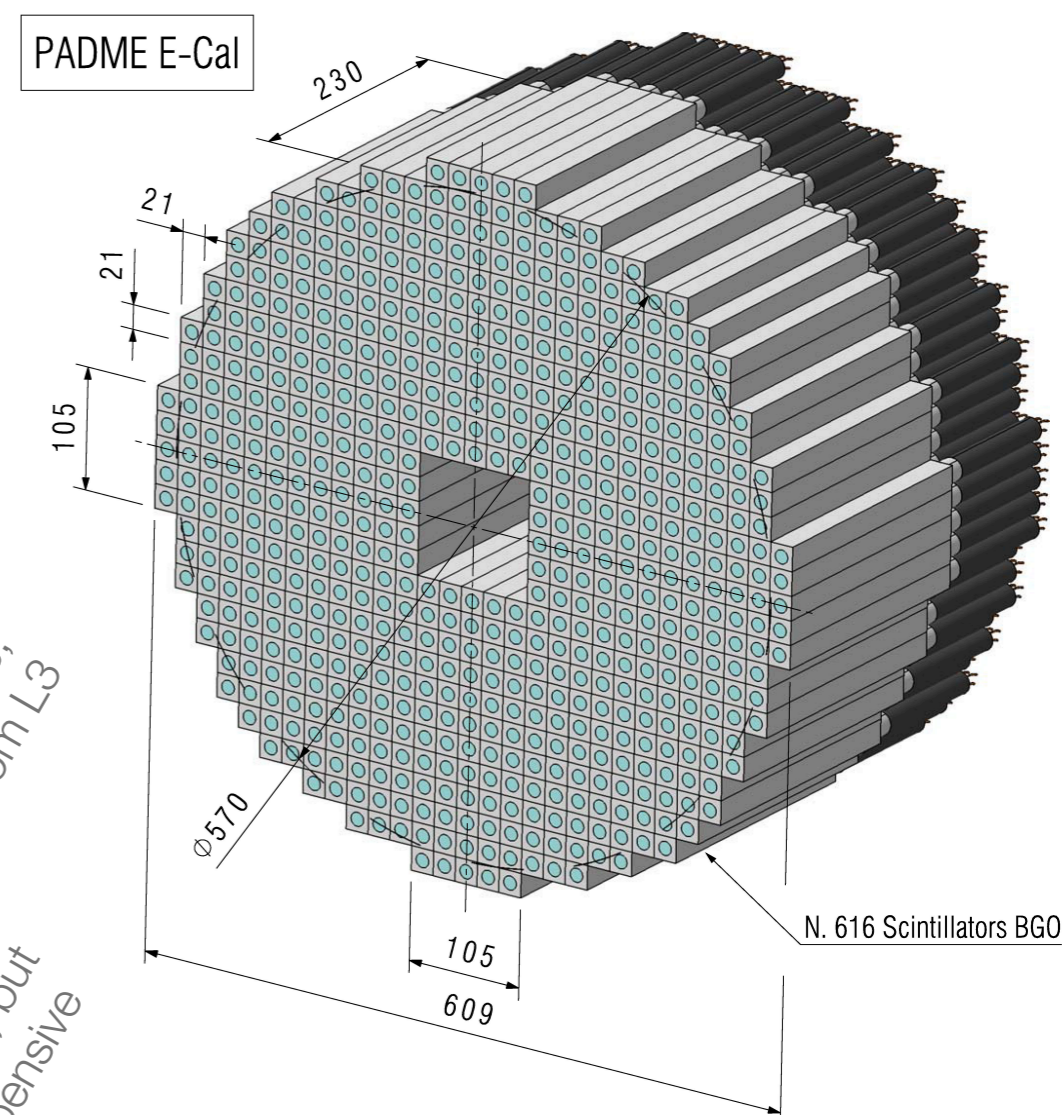
Required characteristics:

- $\sigma_E \approx (1-2)\%/\sqrt{E(\text{GeV})}$
 - good light yield
 - containment
- cluster time resolution < 1 ns
- angular resolution ≈ 2 mrad
- angular coverage: $[20,93]$ mrad
- angular acceptance: $[26,83]$ mrad
- central hole for brems. to SAC (faster)

616 BGO $2.1 \times 2.1 \times 23$ cm³
@ 3m from the target

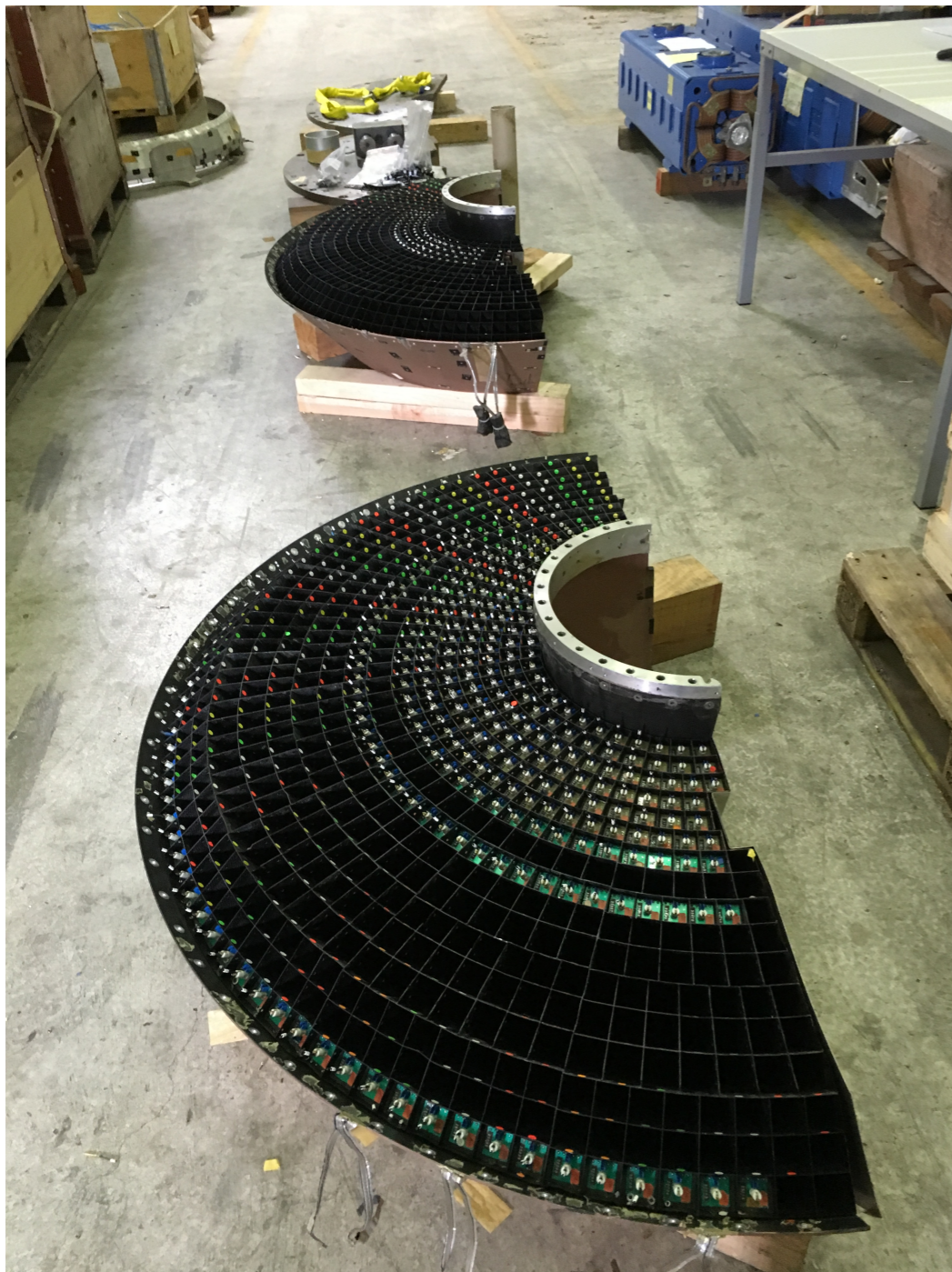
Parameter:	ρ	MP	X_0^*	R_M^*	dE^*/dx	λ_I^*	τ_{decay}	λ_{max}	n^{\ddagger}	Relative output [†]	Hygro-scopic?	$d(\text{LY})/dT$
Units:	g/cm ³	°C	cm	cm	MeV/cm	cm	ns	nm				%/°C [‡]
NaI(Tl)	3.67	651	2.59	4.13	4.8	42.9	245	410	1.85	100	yes	-0.2
BGO	7.13	1050	1.12	2.23	9.0	22.8	300	480	2.15	21	no	-0.9
BaF ₂	4.89	1280	2.03	3.10	6.5	30.7	650 ^s	300 ^s	1.50	36 ^s	no	-1.9 ^s
							0.9 ^f	220 ^f		4.1 ^f		0.1 ^f
CsI(Tl)	4.51	621	1.86	3.57	5.6	39.3	1220	550	1.79	165	slight	0.4
CsI(pure)	4.51	621	1.86	3.57	5.6	39.3	30 ^s	420 ^s	1.95	3.6 ^s	slight	-1.4
							6 ^f	310 ^f		1.1 ^f		
PbWO ₄	8.3	1123	0.89	2.00	10.1	20.7	30 ^s	425 ^s	2.20	0.3 ^s	no	-2.5
							10 ^f	420 ^f		0.077 ^f		
LSO(Ce)	7.40	2050	1.14	2.07	9.6	20.9	40	402	1.82	85	no	-0.2
LaBr ₃ (Ce)	5.29	788	1.88	2.85	6.9	30.4	20	356	1.9	130	yes	0.2

2° best choice, for free from L3
best choice, but very expensive



Crystal procurement

L3 half-endcaps where crystals are...



...taken



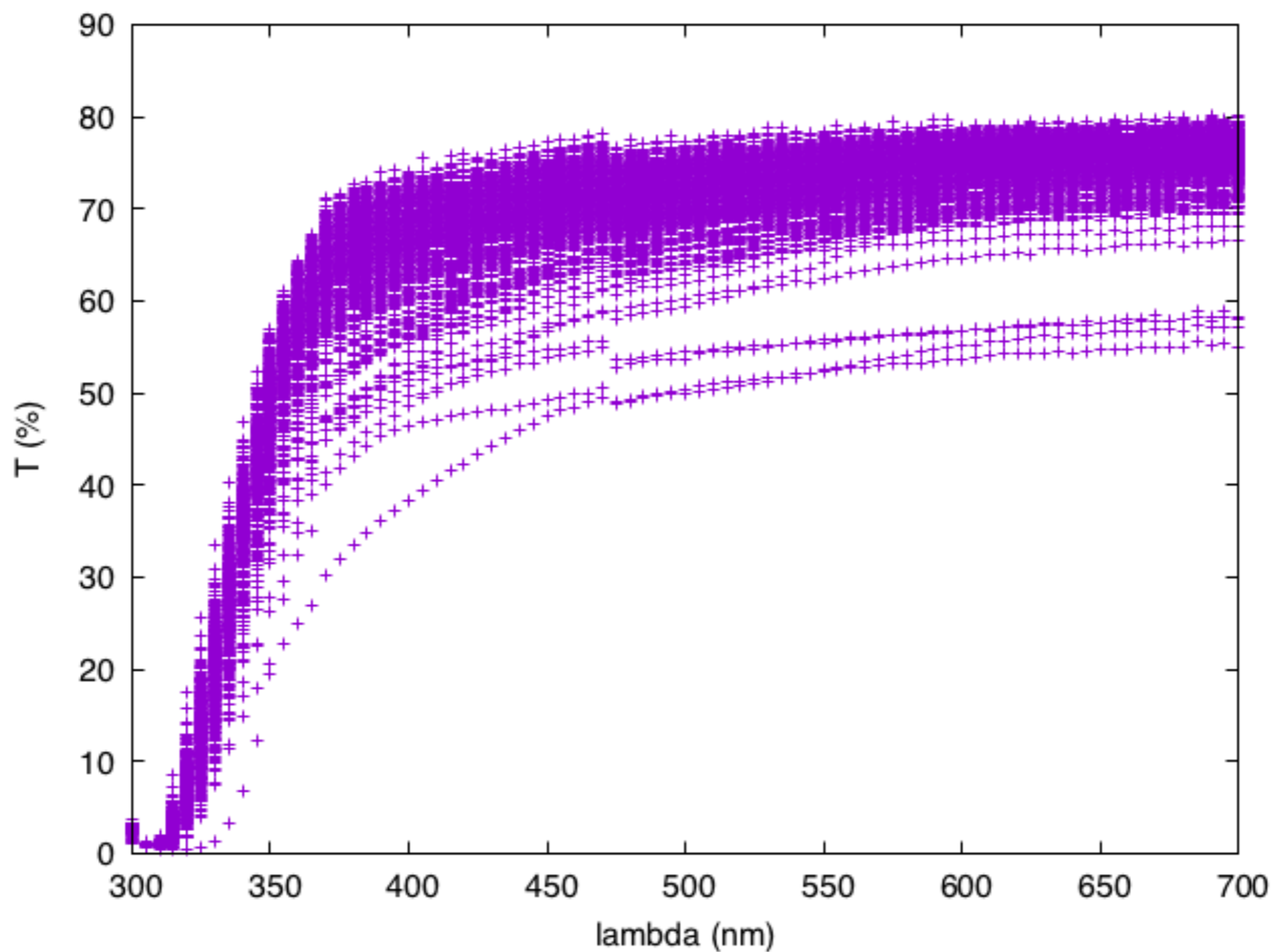
Crystal optical properties

After crystals selection the following steps are executed:

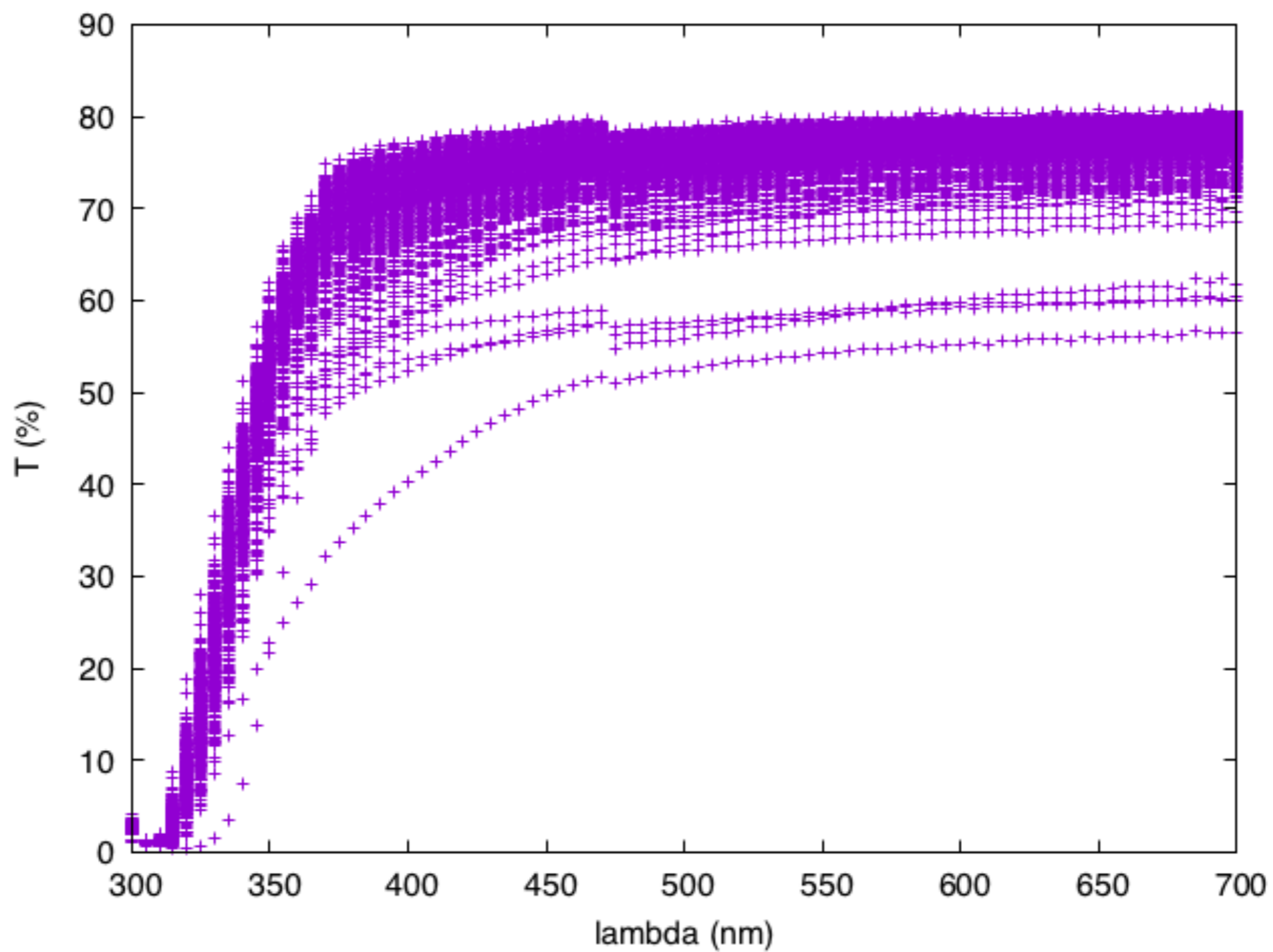
- Photosensor removal (mechanically after 48h in acetone)
- Paint removal (with water)
- Transmittance measurement
- Annealing
 - $T_{\text{amb}} \rightarrow 200 \text{ }^\circ\text{C}$ in 3 h
 - 200 $^\circ\text{C}$ for 6 h
 - 200 $^\circ\text{C} \rightarrow T_{\text{amb}}$ “natural”
- Transmittance measurement

Everything is performed at CERN at LAB27

Transmittance before annealing

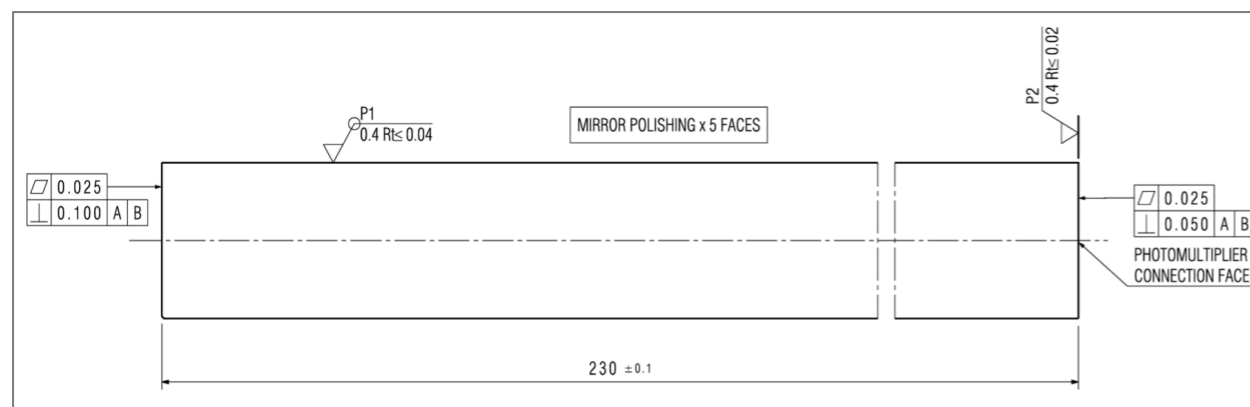
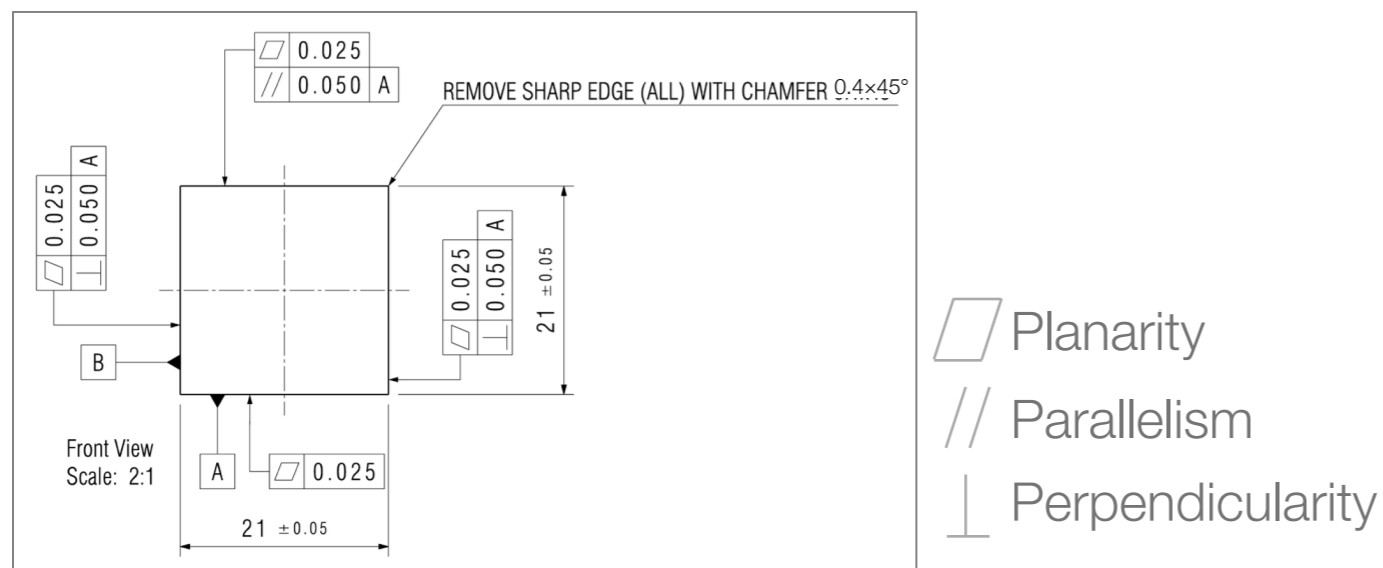


Transmittance after annealing



Crystals cut and polished at SILO (Italy)

They produced identical parallelepipeds starting from different truncated pyramid shapes (L3 endcaps geometry was pointing)

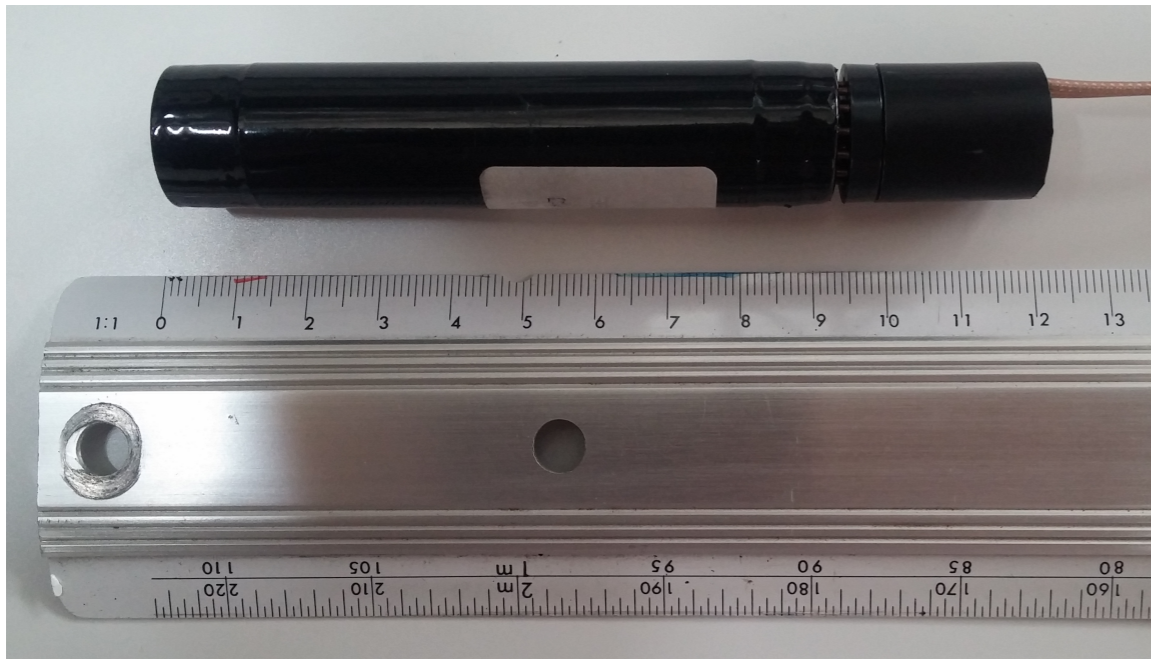


Mechanical tolerances (more stringent limits are set for the square shape)

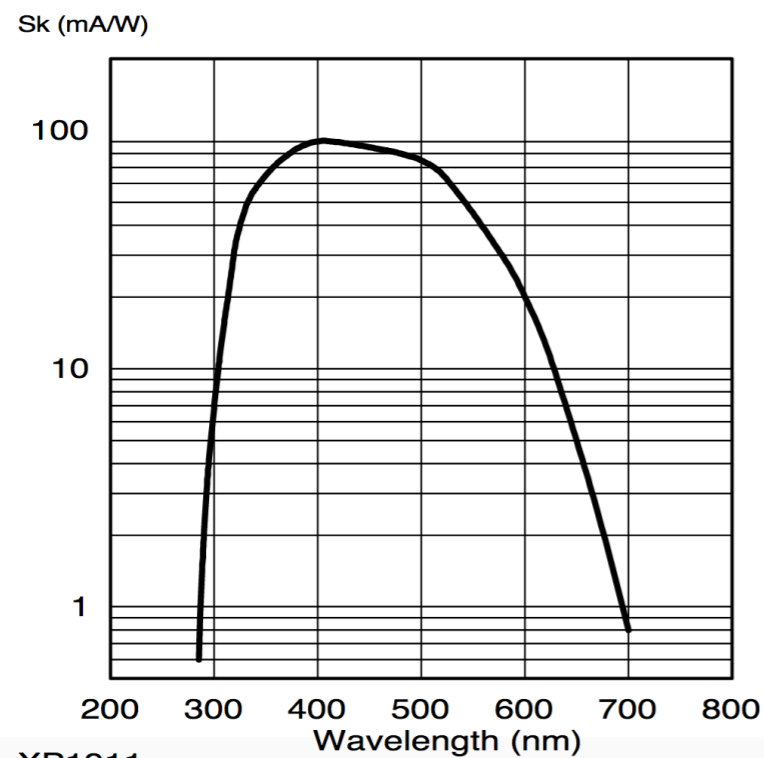
We performed a quality check at LNF on some crystals, to verify that dimensions are within specification, w/ positive results

HZC XP1911

We modified the mechanical design

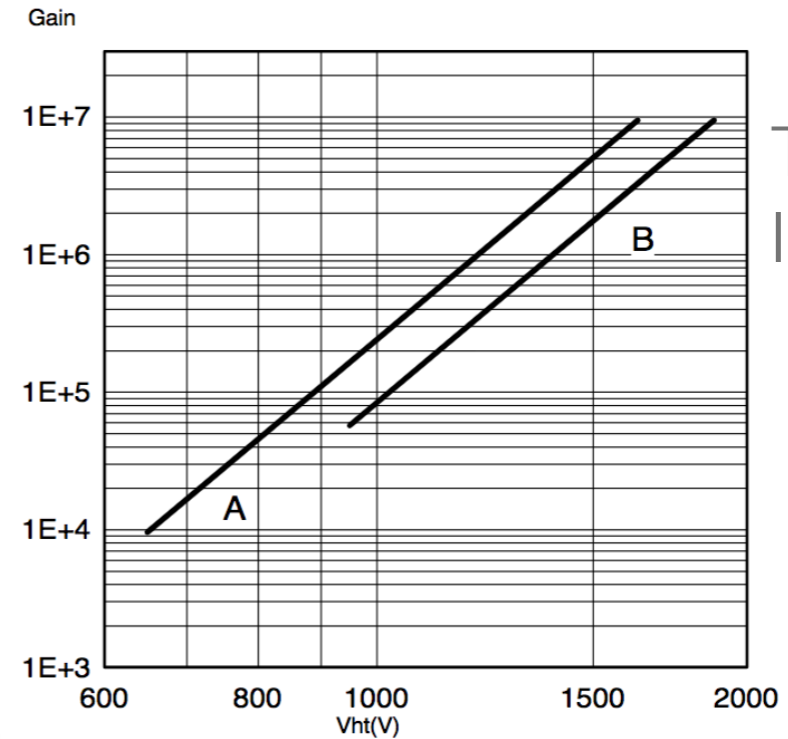


Typical spectral characteristics



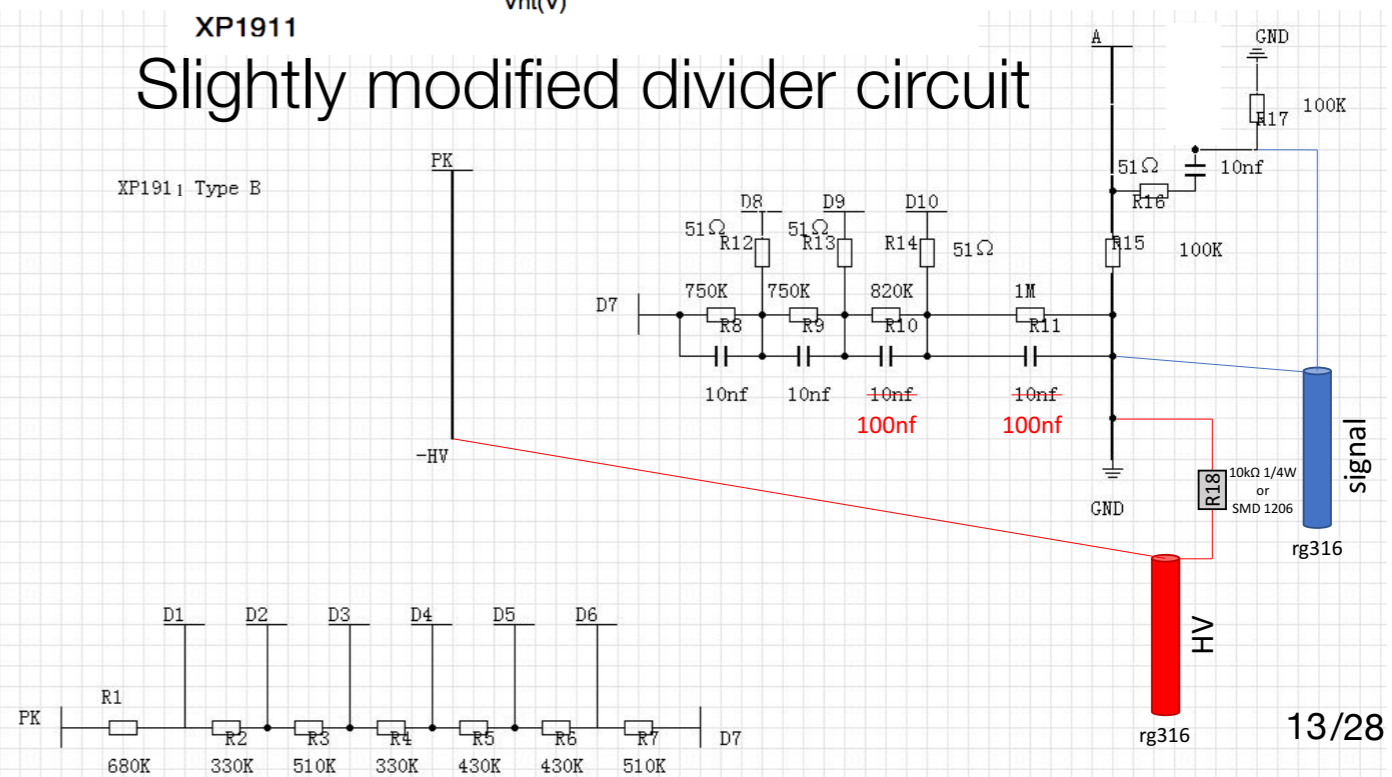
XP1911

Typical gain curve



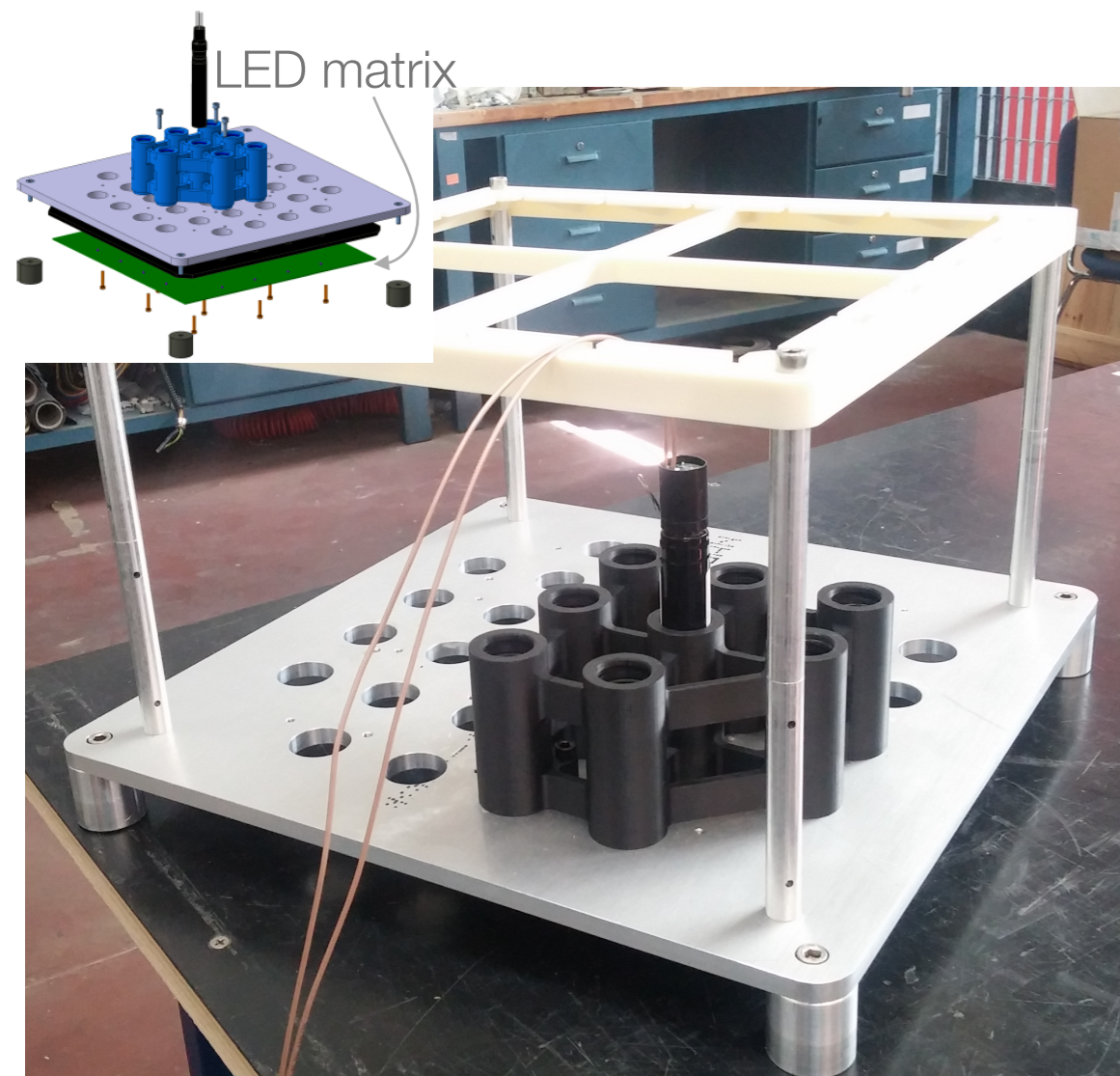
Type B: higher linearity

Slightly modified divider circuit

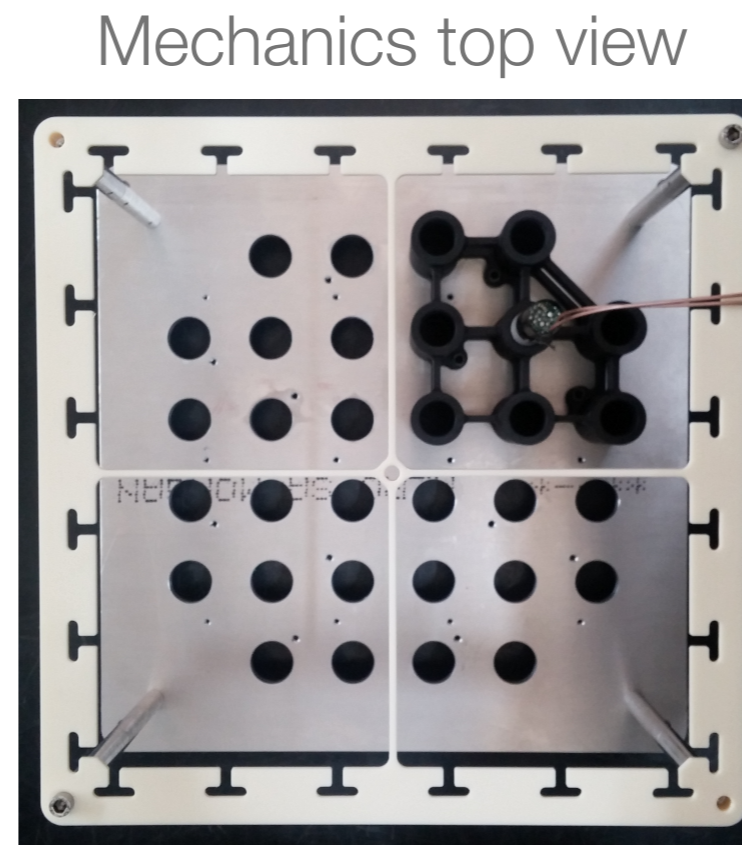


PMTs test

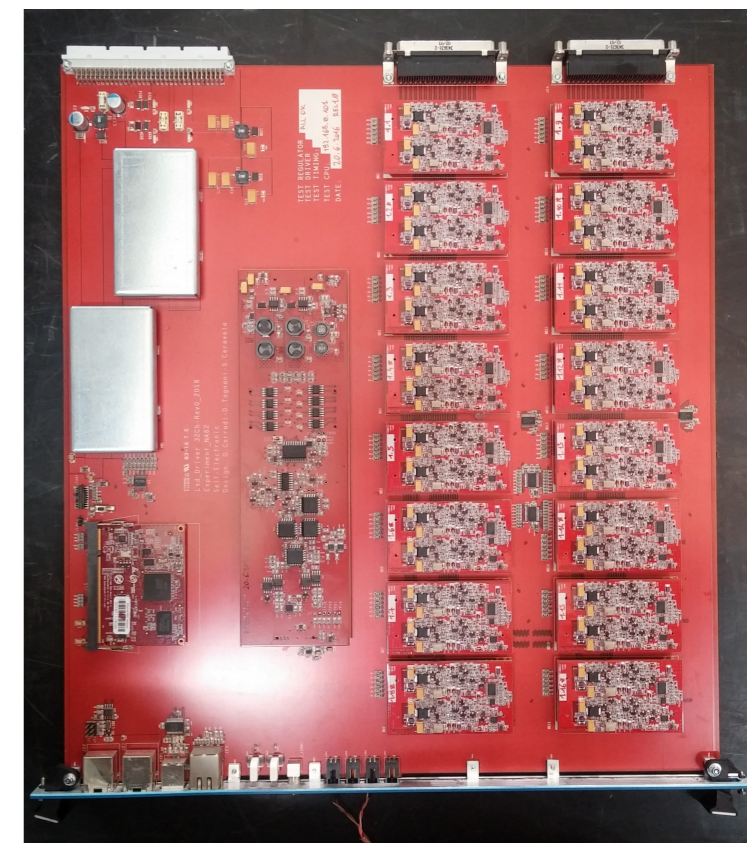
32 PMTs at a time were tested with a LED matrix (one per tube): pulsing the LEDs we see if the PMT works and its response to the light. If results are good, tubes are sent to SILO for gluing.



Mechanics for PMTs test



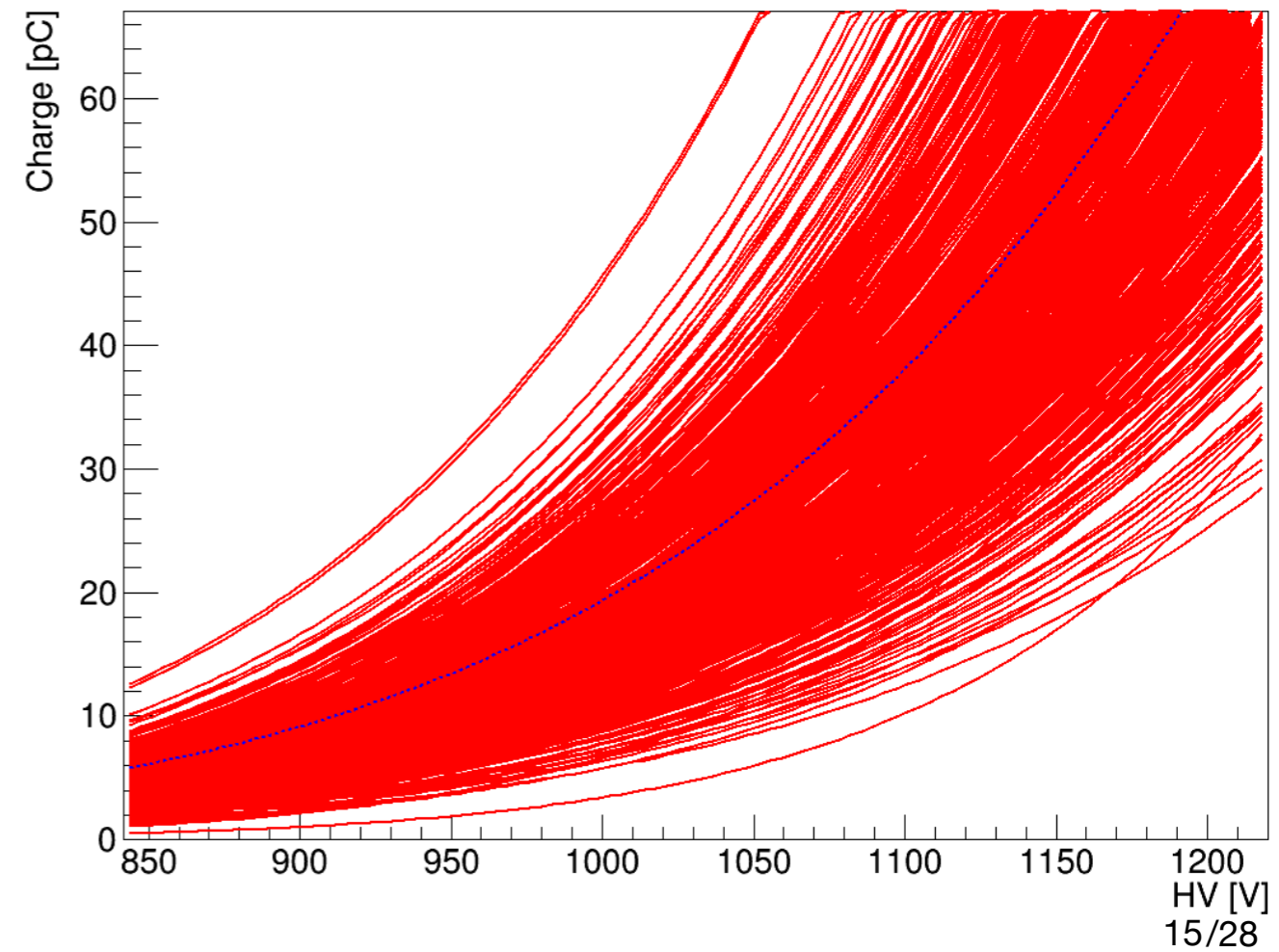
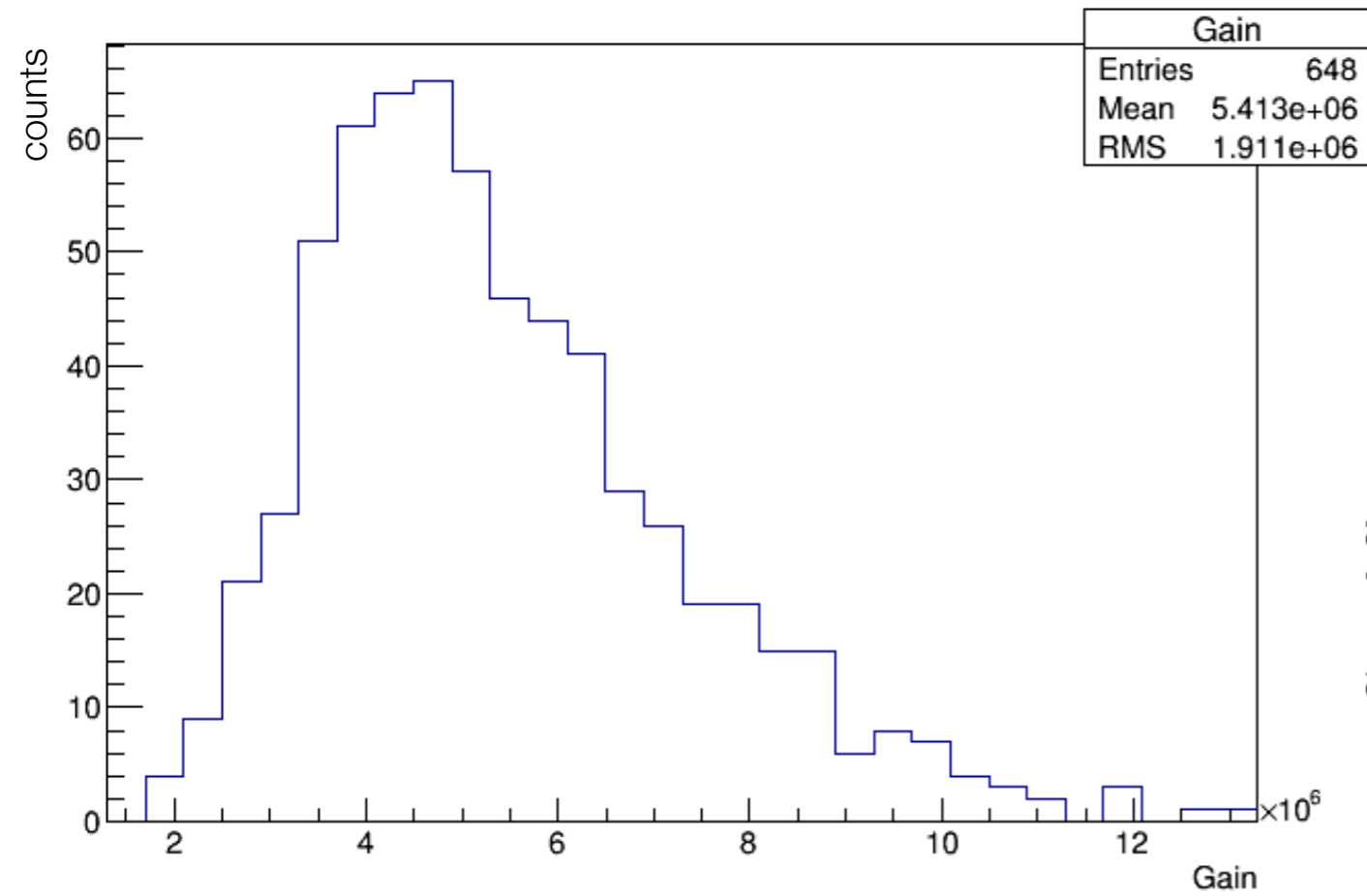
Mechanics top view



LED driver board

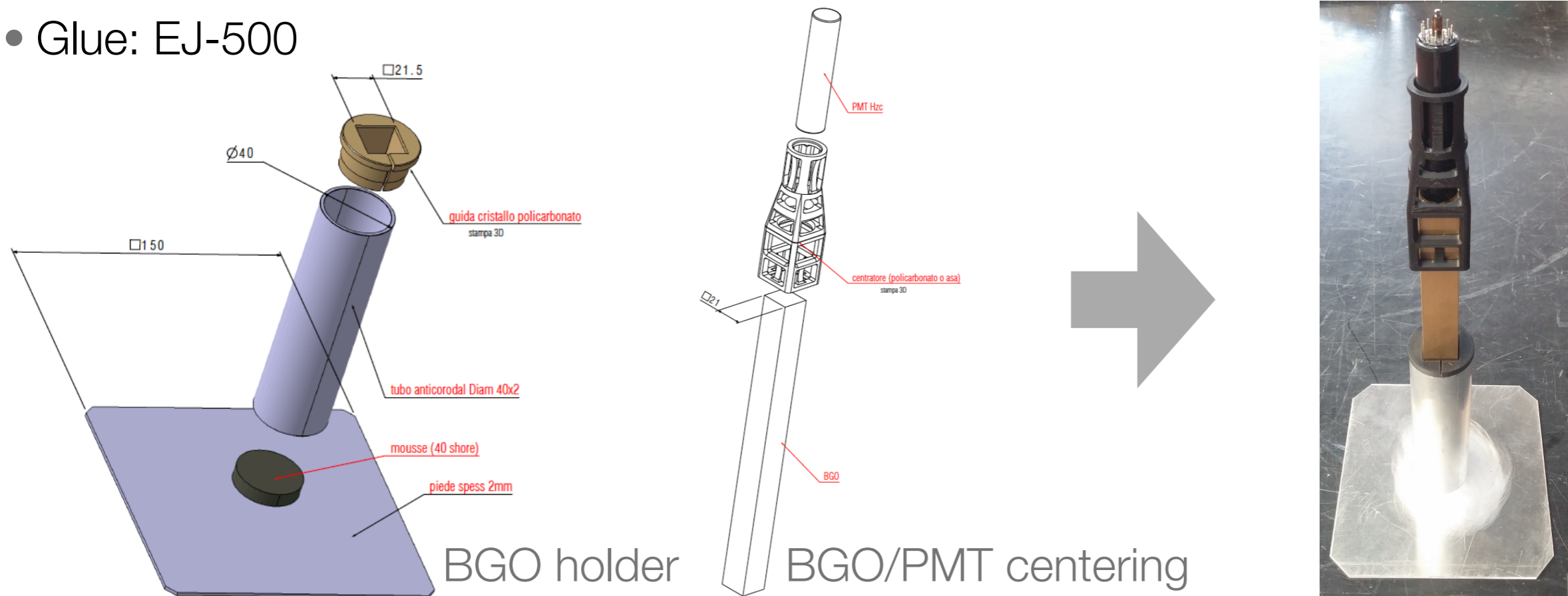
Global PMT results

Gain at 1680.0V

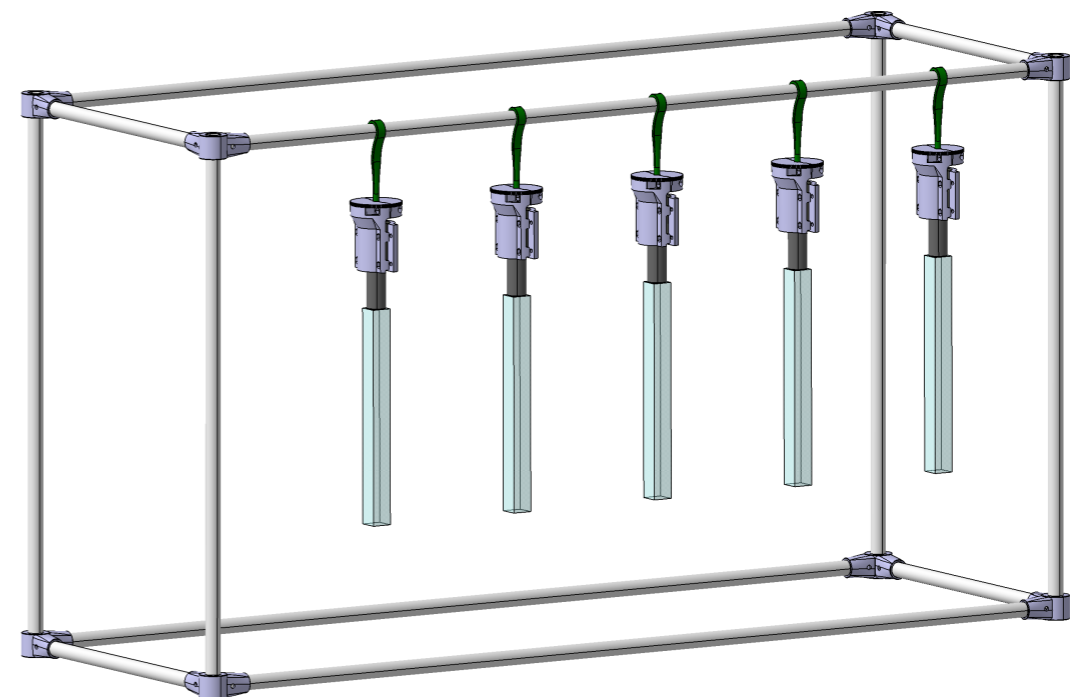


Gluing and painting at SILO

- Glue: EJ-500

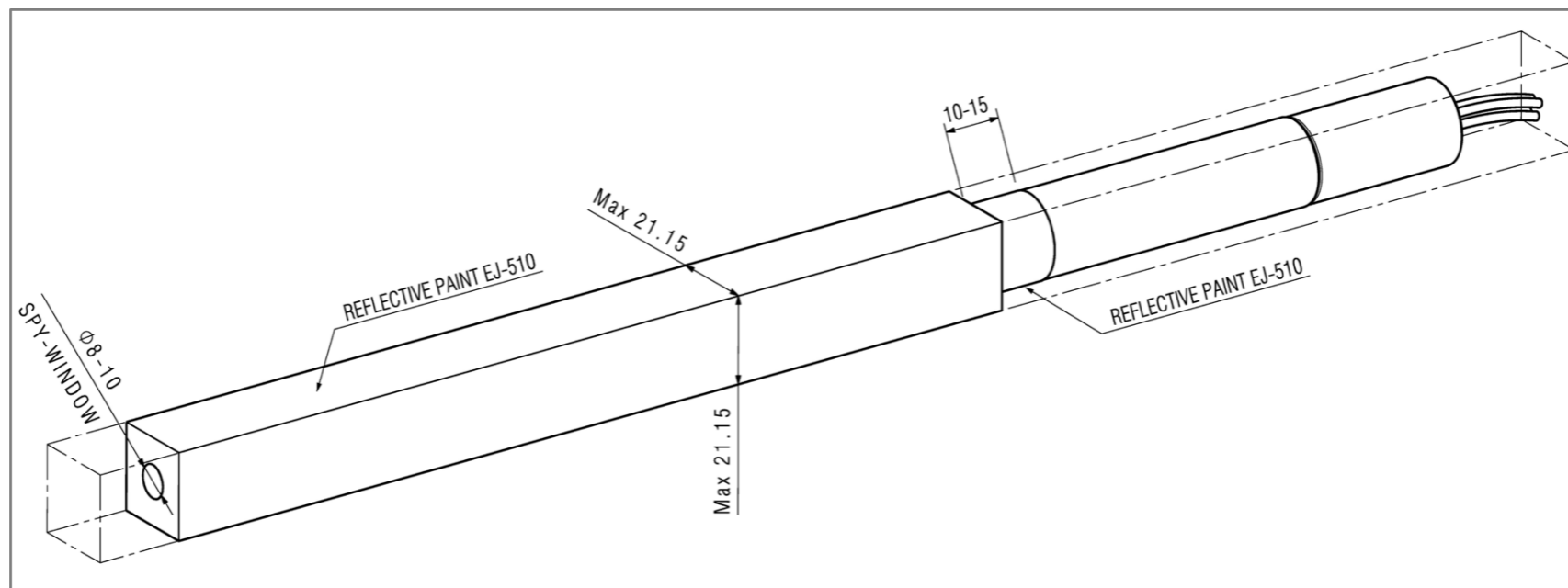


- Paint: EJ-510
- 3 layers of white paint ($\approx 100\mu\text{m}$)



Scintillating Units (SU)

Maximum footprint after painting



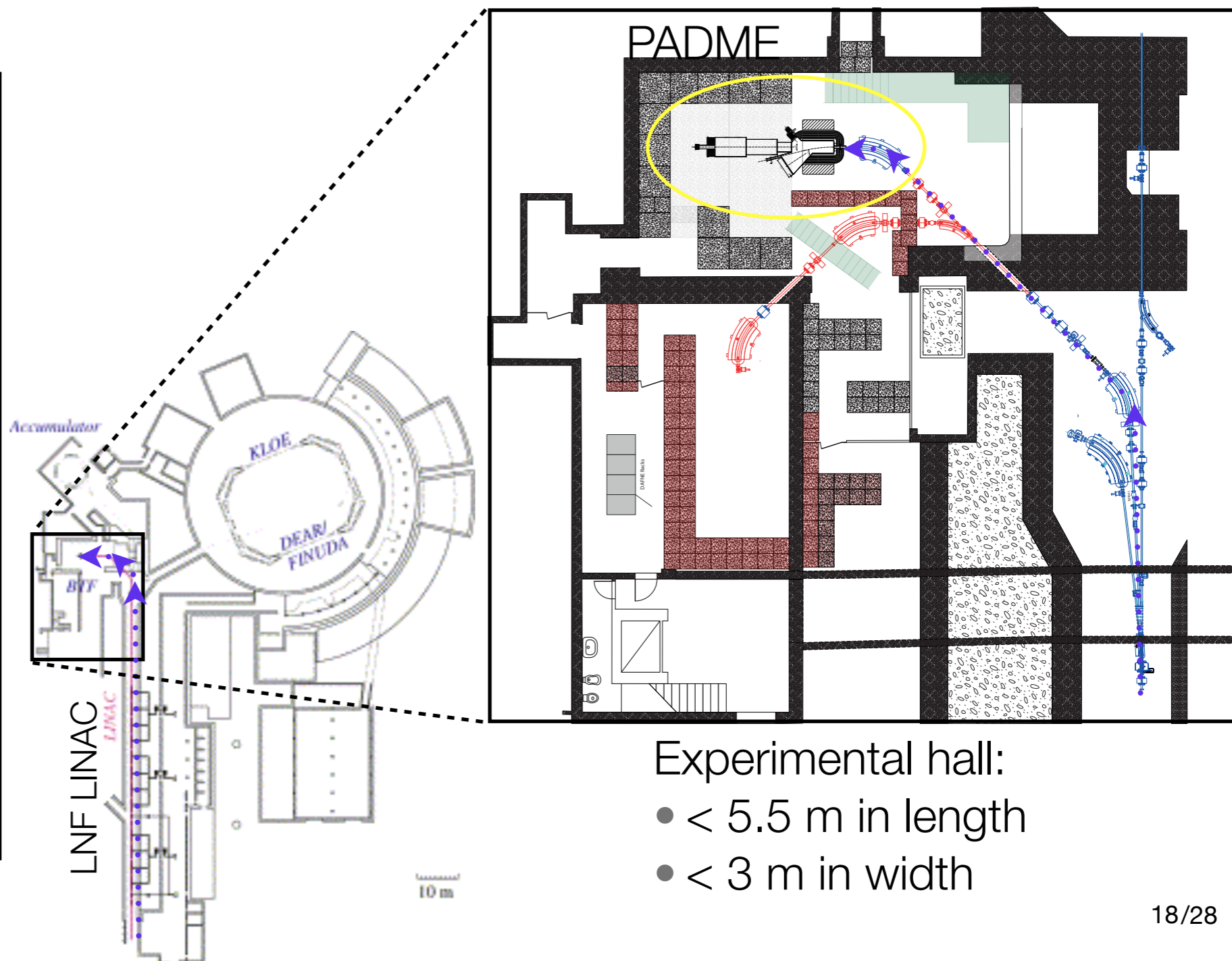
Currently we have 580 $2.1 \times 2.1 \times 23\text{cm}^3$ painted and glued units at LNF



The LNF Beam Test Facility (BTF)

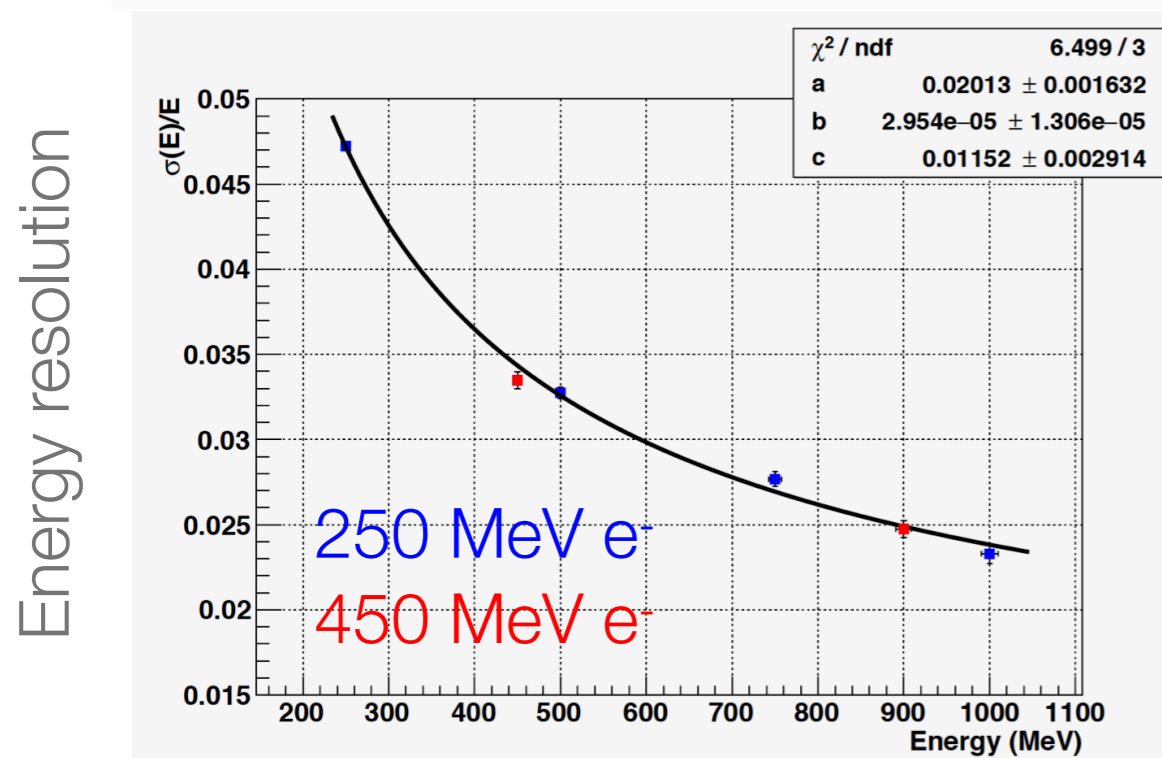
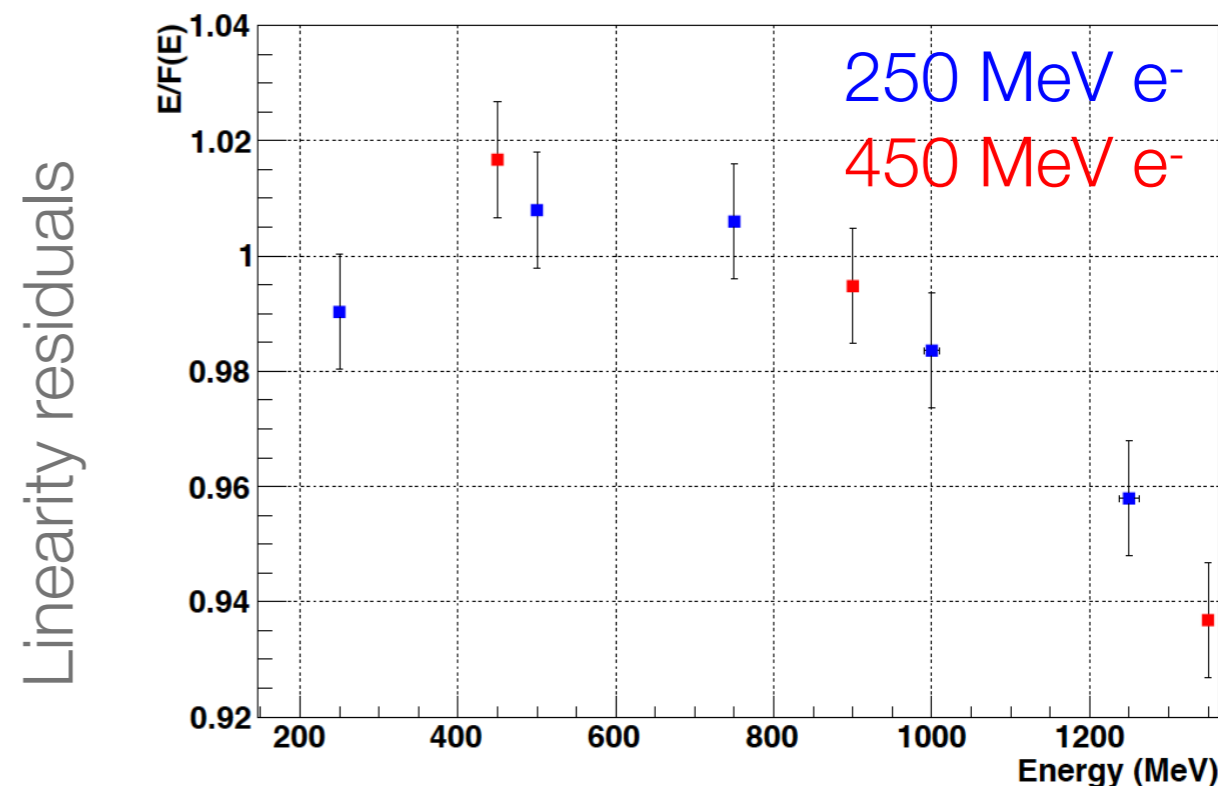
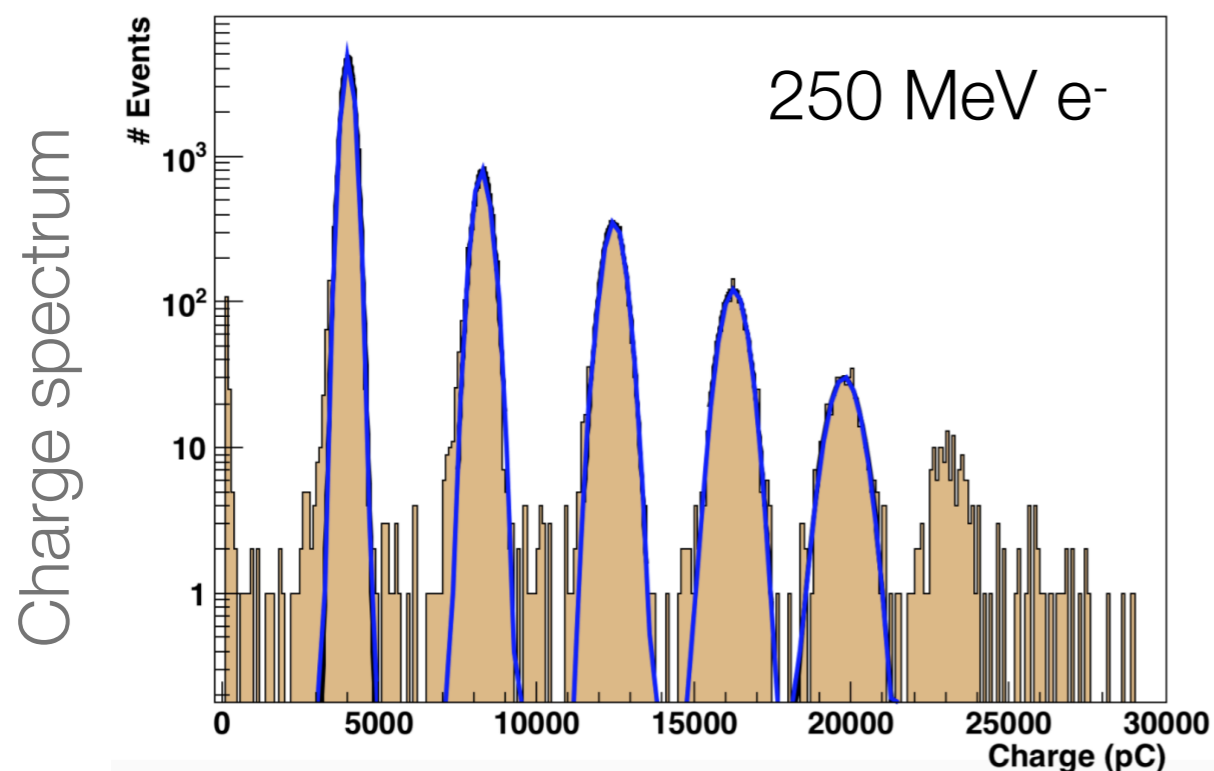
PADME experimental hall be is the Beam Test Facility of the Laboratori Nazionali di Frascati (~Rome, IT), the same place where the test beams have been performed.

	Dedicated mode	
	W/ target	W/o target
Particle species	e ⁺ /e ⁻ selectable by user	
Energy [MeV]	25-700 (e ⁺) 25-700 (e ⁻)	250-730 (e ⁺) 250-530 (e ⁻)
Energy spread	1%	
Rep. rate [Hz]	1-49 selectable by user	
Pulse duration [ns]	1.5-40 selectable by user	
Intensity [particles/bunch]	1-10 ⁵ depending on energy	10 ³ -3 · 10 ¹⁰
Max average flux	3.125 · 10 ¹⁰ particles/s	
Spot size [mm]	0.5-25 (y) × 0.6-55 (x)	
Divergence [mrad]	1-1.5	



- Experimental hall:
- < 5.5 m in length
 - < 3 m in width

Calorimeter prototype performance @ BTF



Linearity is within 2% up to 1GeV
(gain 5×10^5)

Energy resolution is within the expectation, with reference to the L3 experience

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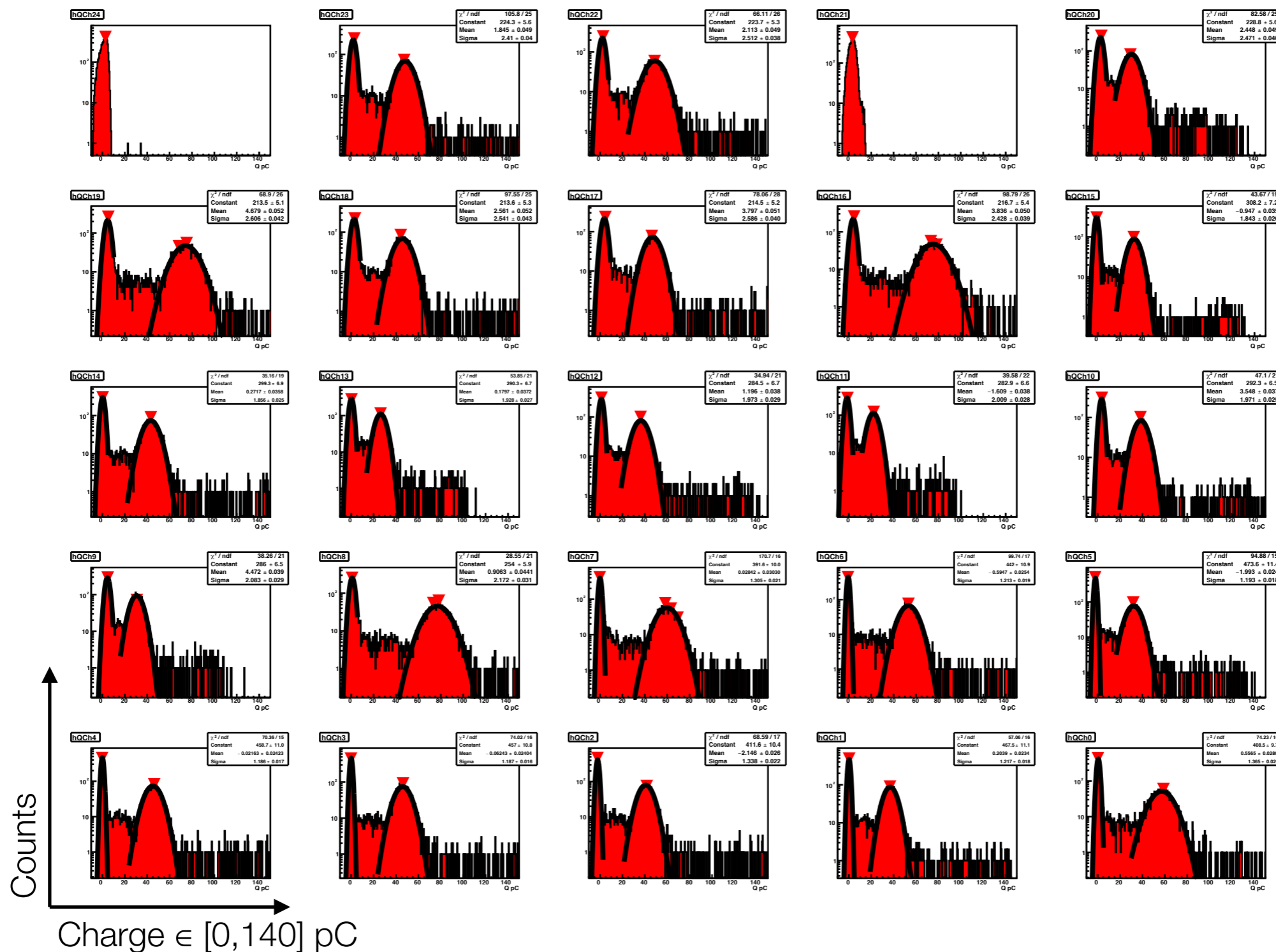
^{22}Na setup

- A $3\times 3\times 20\text{ mm}^3$ LYSO crystal read by a SiPM is used as trigger
- ^{22}Na source faced to each crystal, to exploit its γ back-to-back emission: one in the trigger, one in the SU
- 10 HV tested on PMTs: from 1100V to 1550V in steps of 50V



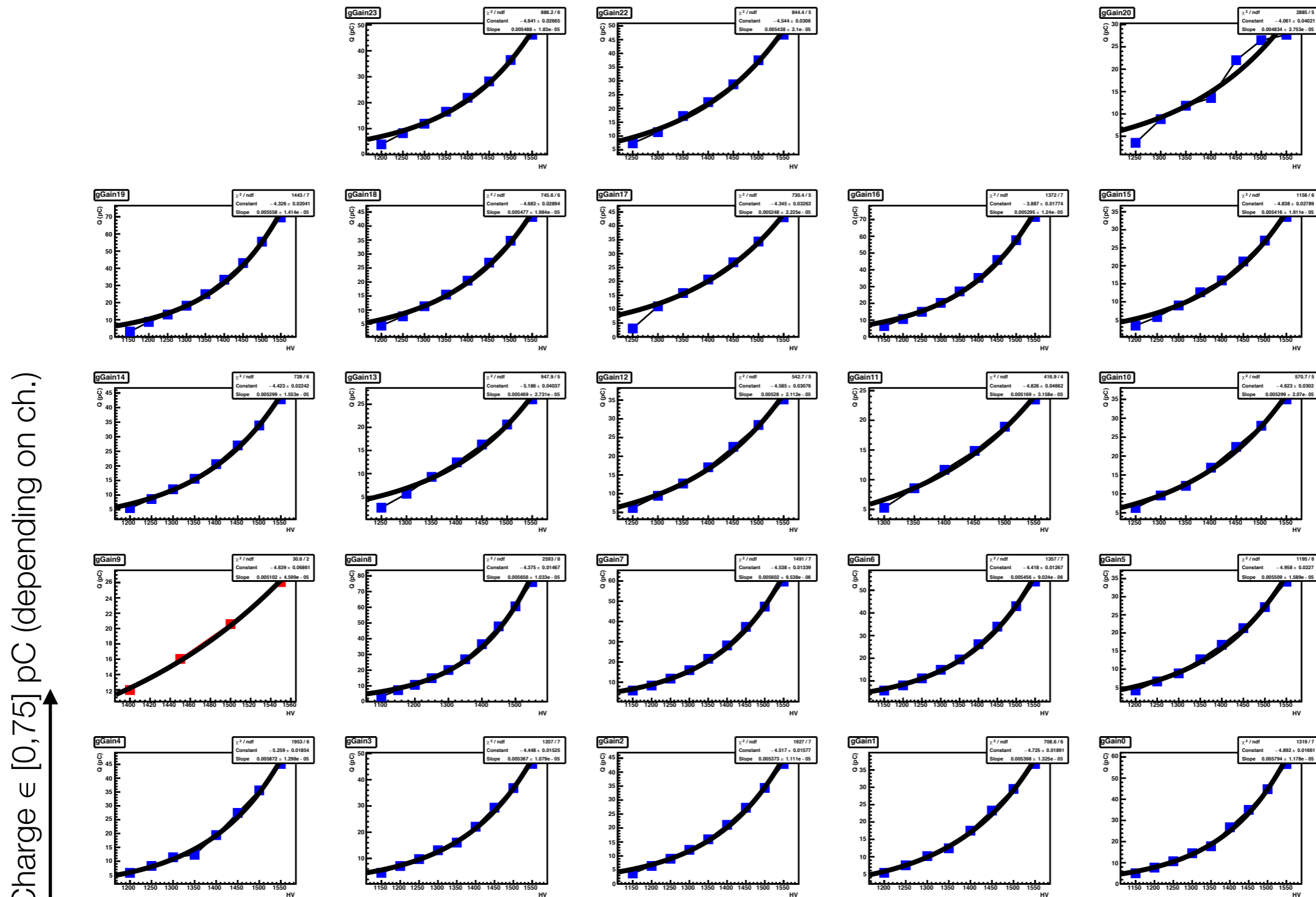
Reconstructed spectra w/ ^{22}Na source

Charge is given by an integration window of $1\ \mu\text{s}$ ($\approx 150\ \text{ns}$ pre-pulse), sampled at $1\ \text{GS/s}$



Example of ^{22}Na spectra at 1550 V

Gain curves w/ ^{22}Na 511 keV peak



Charge $\in [0, 75]$ pC (depending on ch.)

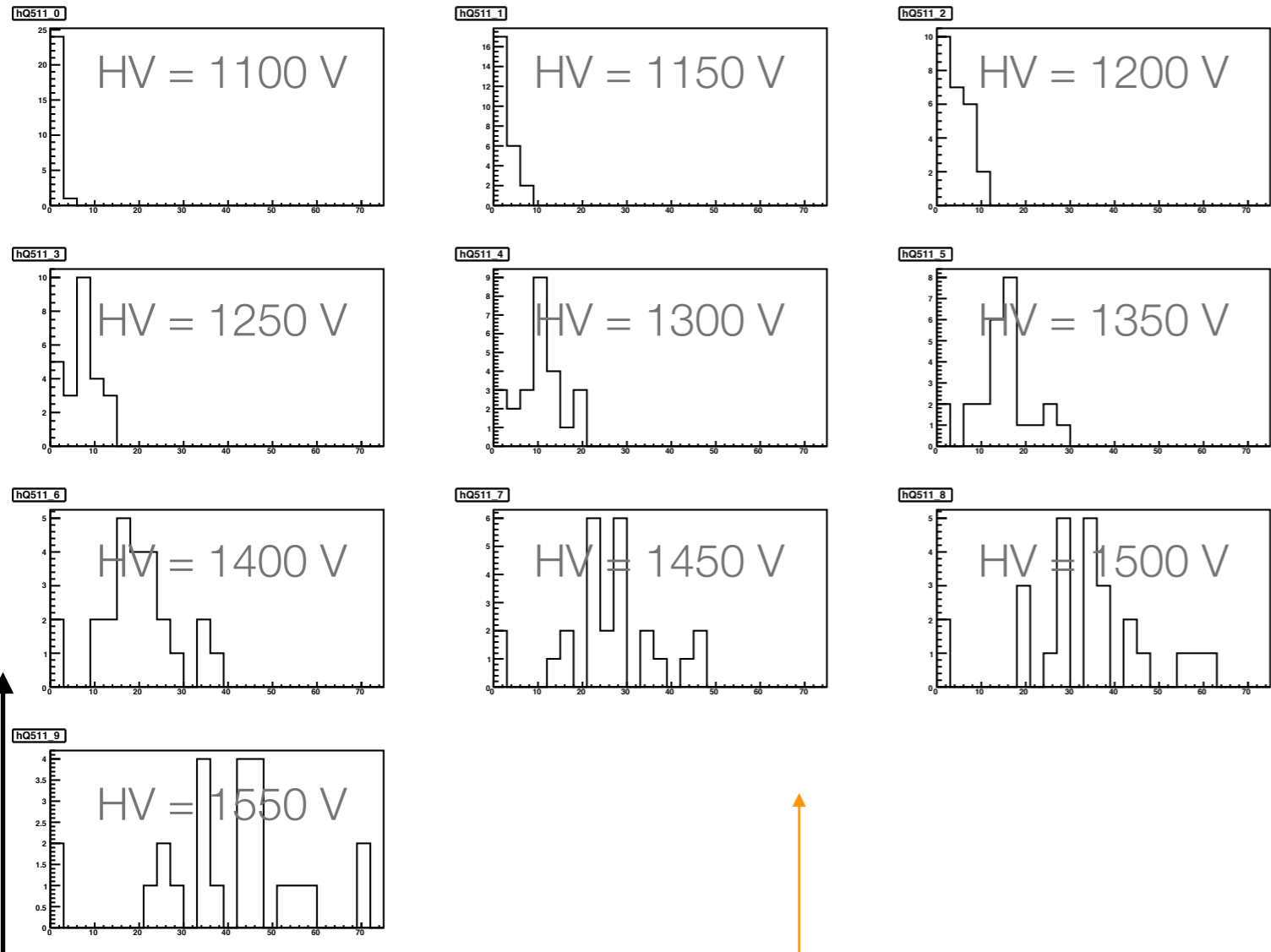
HV $\in [1100, 1550]$ V depending on ch.

Fitting function: $Q = \text{const} \cdot (\text{HV})^{\text{slope}}$

fit parameters

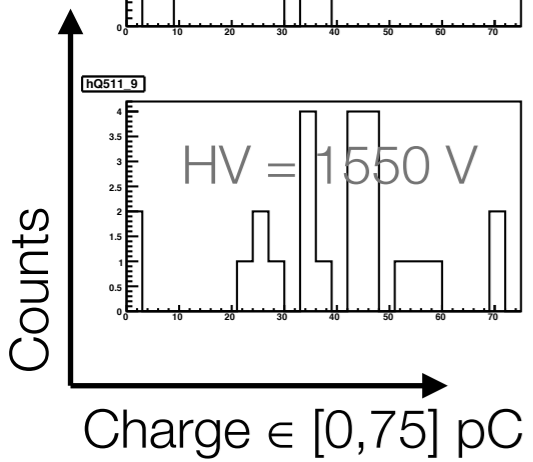
Gain equalization

511 keV peak mean distributions

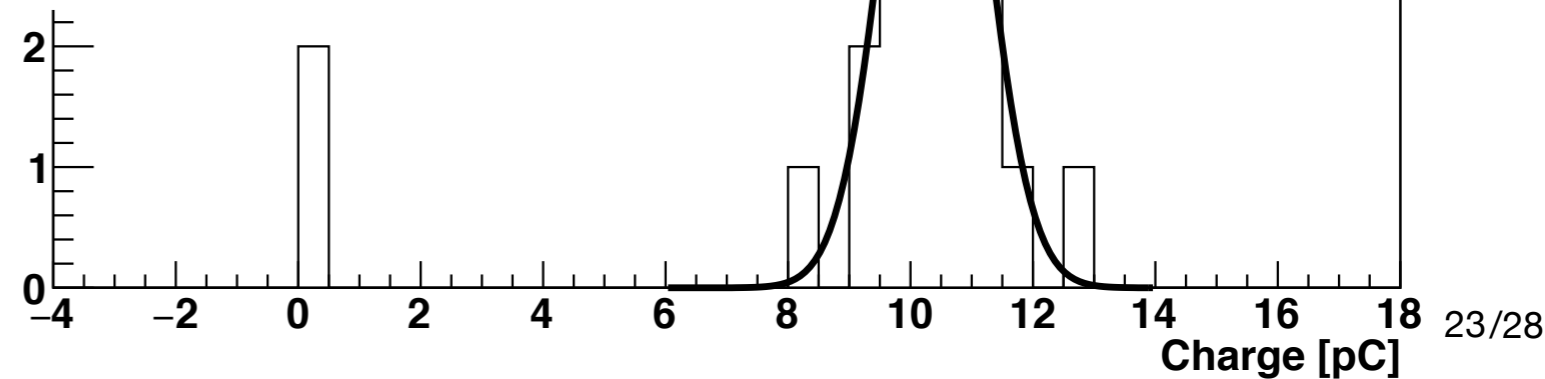


Equalization at 20 pC/MeV

χ^2 / ndf	2.345 / 5
Constant	5.565 ± 2.128
Mean	10.4 ± 0.2
Sigma	0.7731 ± 0.2982



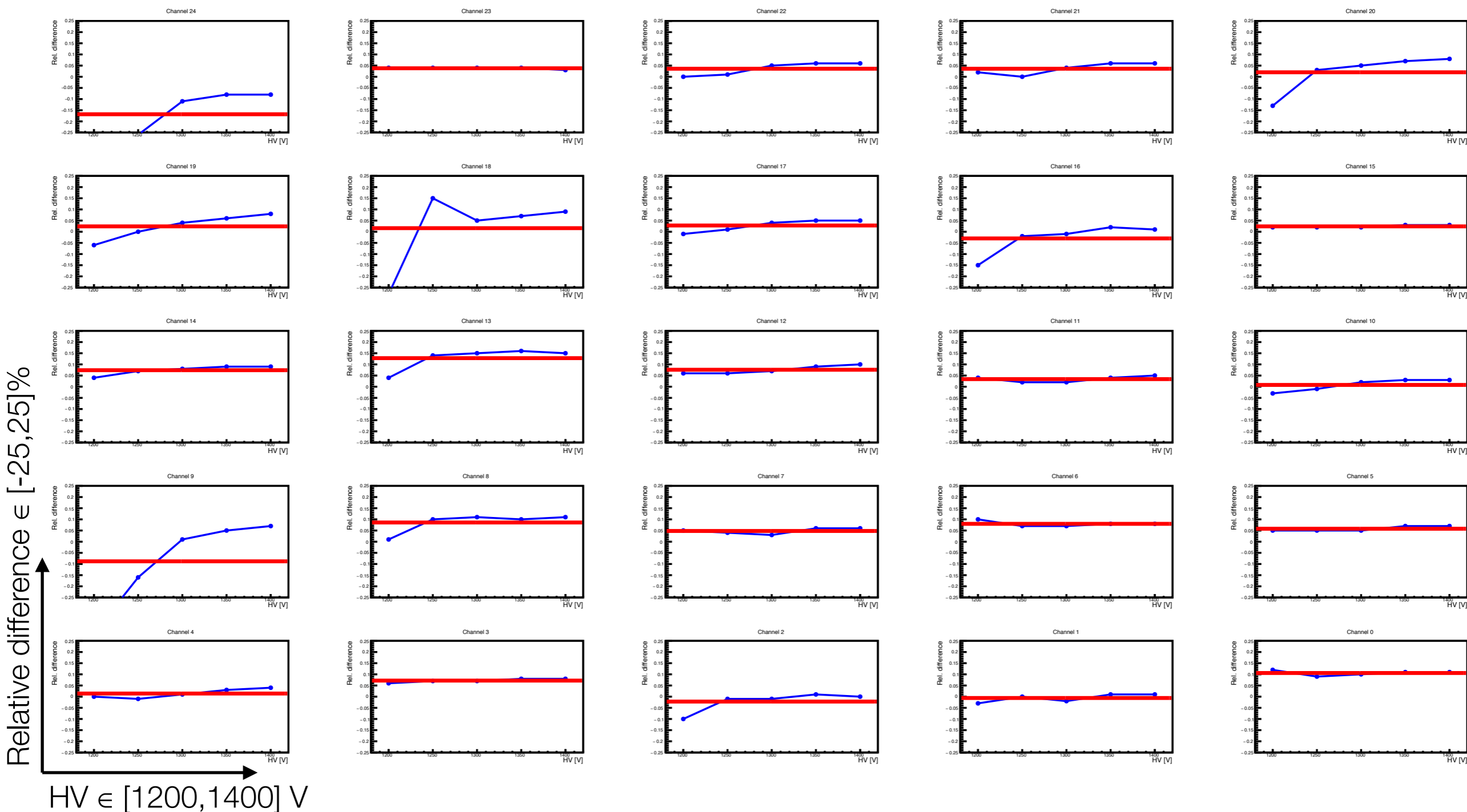
Peaks distribution for fixed HVs



Reproducibility

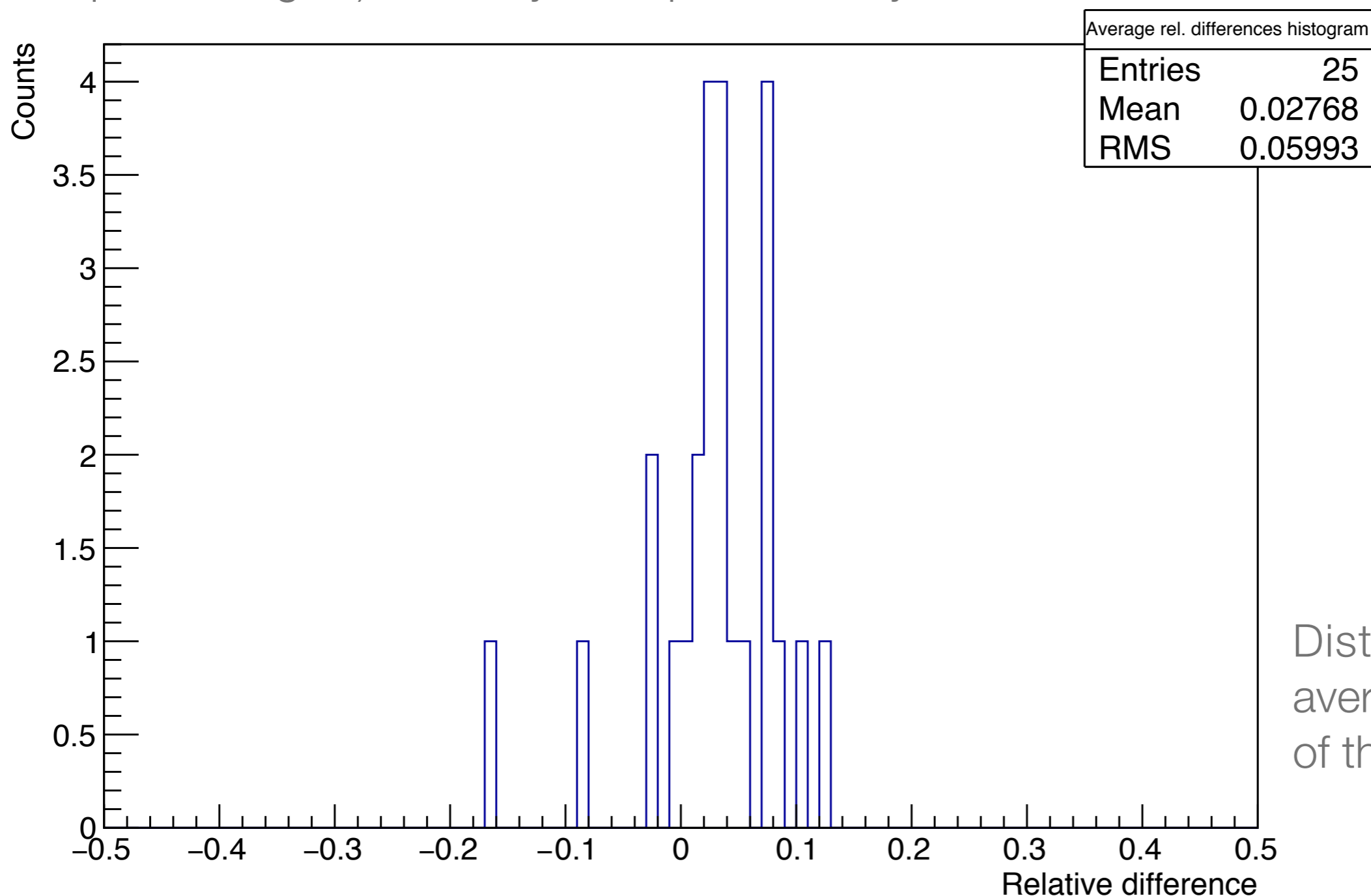
For each SU of a group of 25 we performed 2 times the same HV scan

Relative difference of the 511 keV peak charge as a function of the HV



Charge relative differences distribution

- Larger relative variations are due to small absolute values
- Measurements have been done in different conditions (daylight, black cover positioning,...) that may have produced systematic variations



Distribution of the averages (red lines) of the previous slide

Calorimeter mechanical design

ECAL (BGO + filler)

^{22}Na movement
(calibration & transparency)

SAC

ECAL support

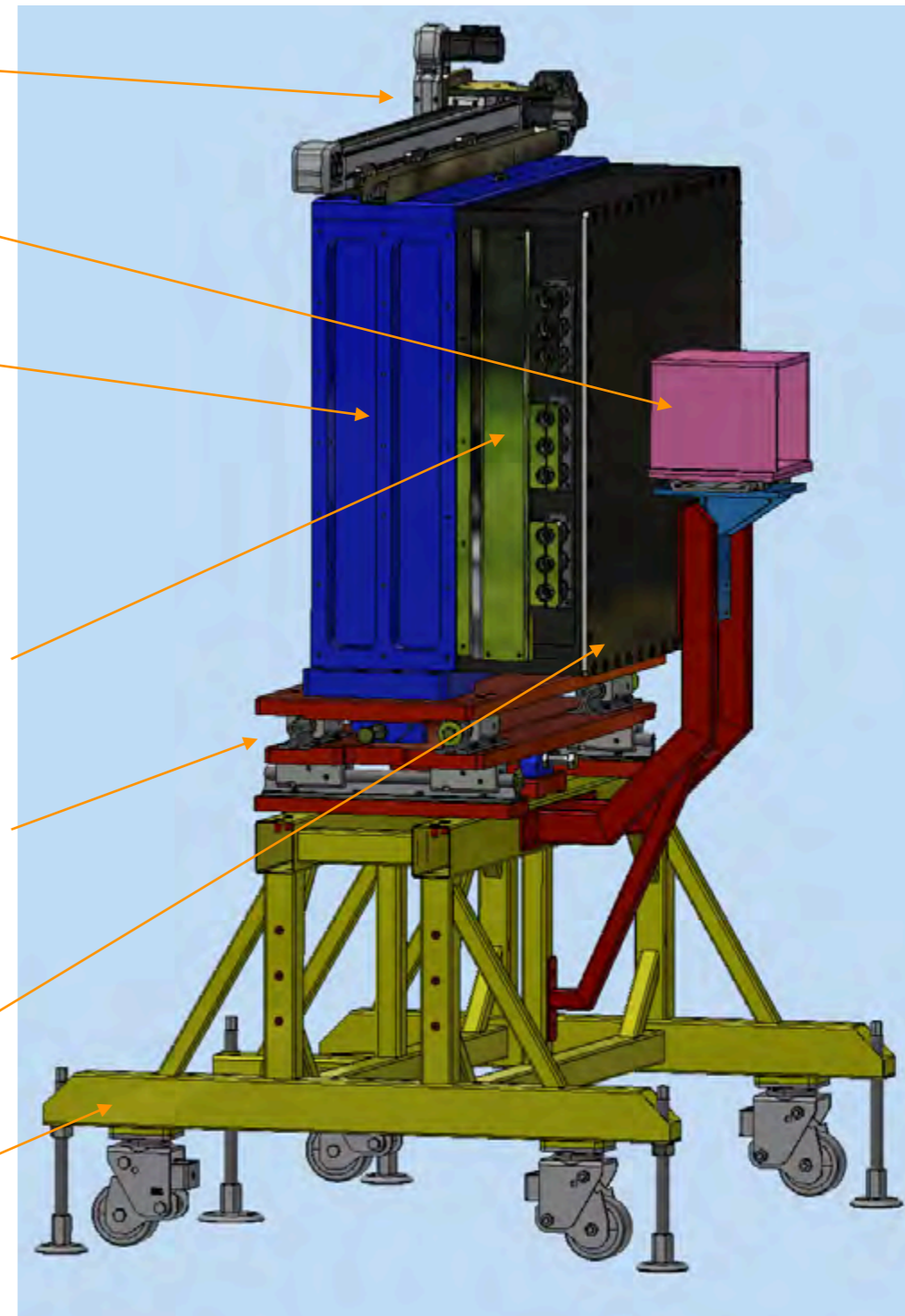
PMT enclosure

Support table

Inner support

Front/rear panels
(light tightness)

Support structure



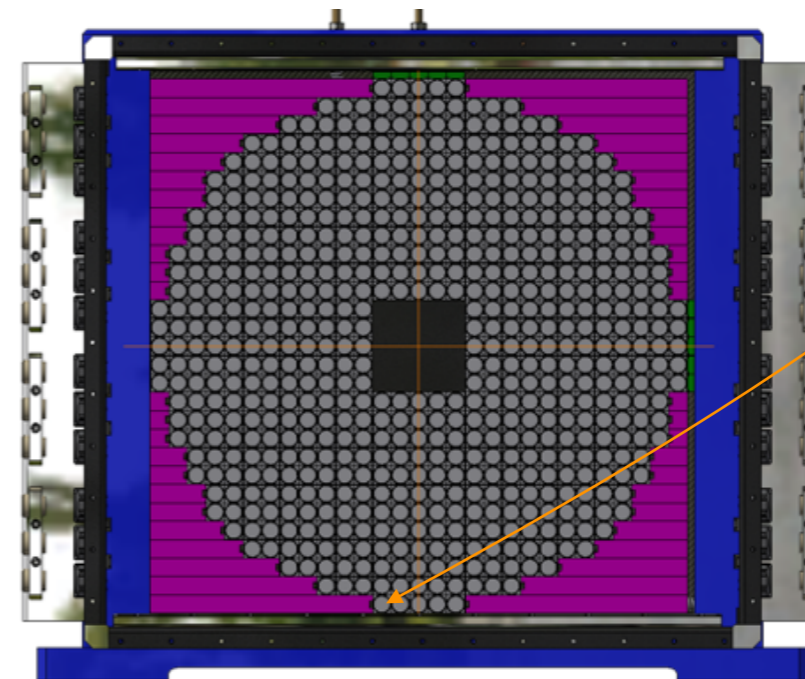
ECAL assembly procedure



Procedure:

for each layer

- first crystal bottom left
- complete first layer
- block layer with locking screws
- equalize for different SU heights
- go to next layer



front view

Ready for the assembly!

Conclusions

- Dark Photon is predicted by many physics models, that could explain different experimental results: Dark Matter, $(g-2)_\mu$, ^8Be anomaly
- PADME is an experiment hosted at the Laboratori Nazionali di Frascati searching for invisible Dark Photon decays
- The electromagnetic calorimeter is one of the most important components of the detector and is currently under construction
- Calorimeter readout: 616 HZC XP1911 (PMTs) w/ $\approx 5\%$ gain uniformity at nominal HV
- Scintillating units
 - very low threshold (≈ 0.5 MeV)
 - good reproducibility w/ variations $< 3\%$
- ECAL prototype
 - energy resolution is compatible with the L3 results: $2\%/\sqrt{E(\text{GeV})}$
 - good charge reconstruction linearity w/ variations $< 2\%$ up to 1 GeV

Backup

Dark Photon searches



BGO emission spectrum

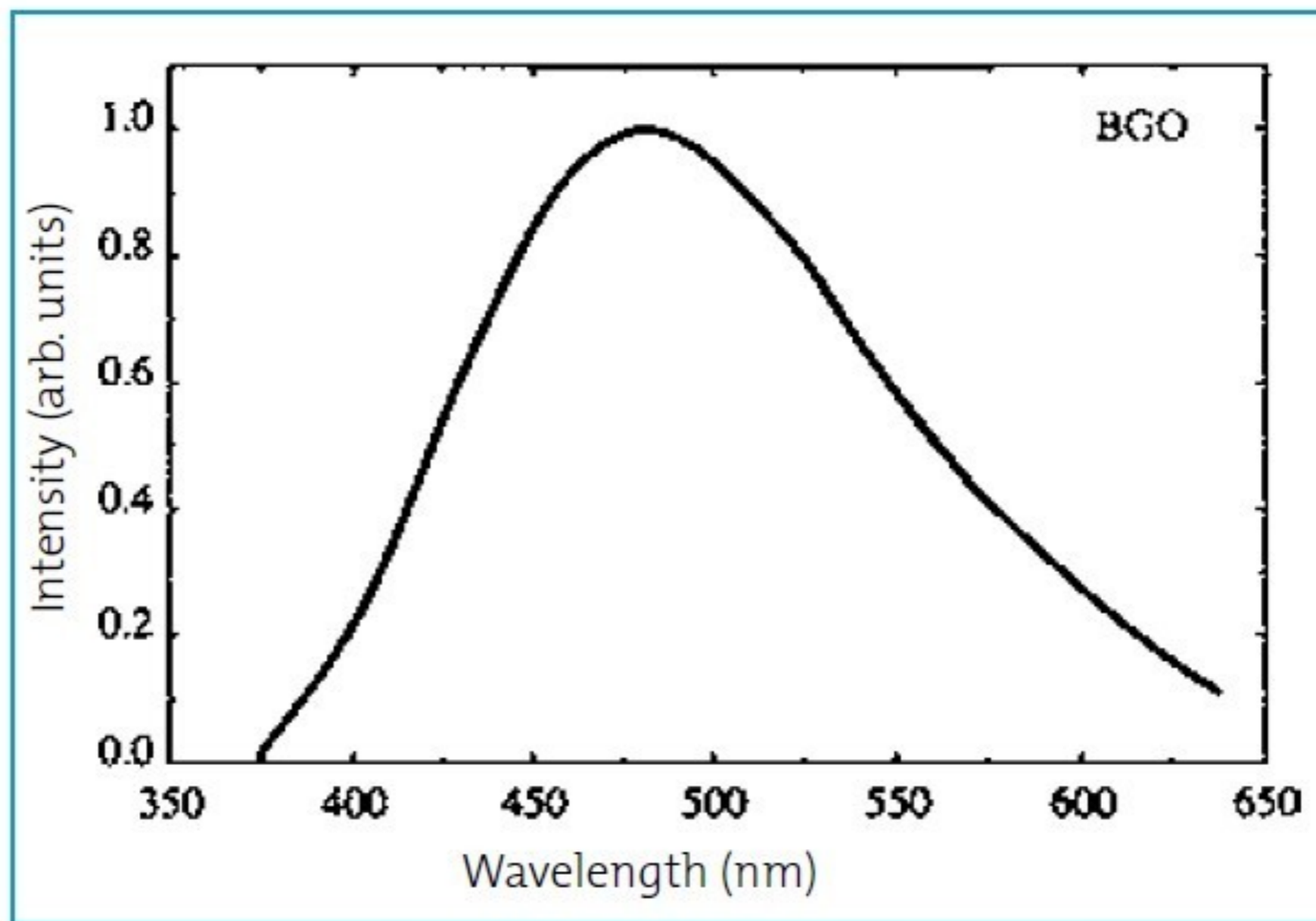
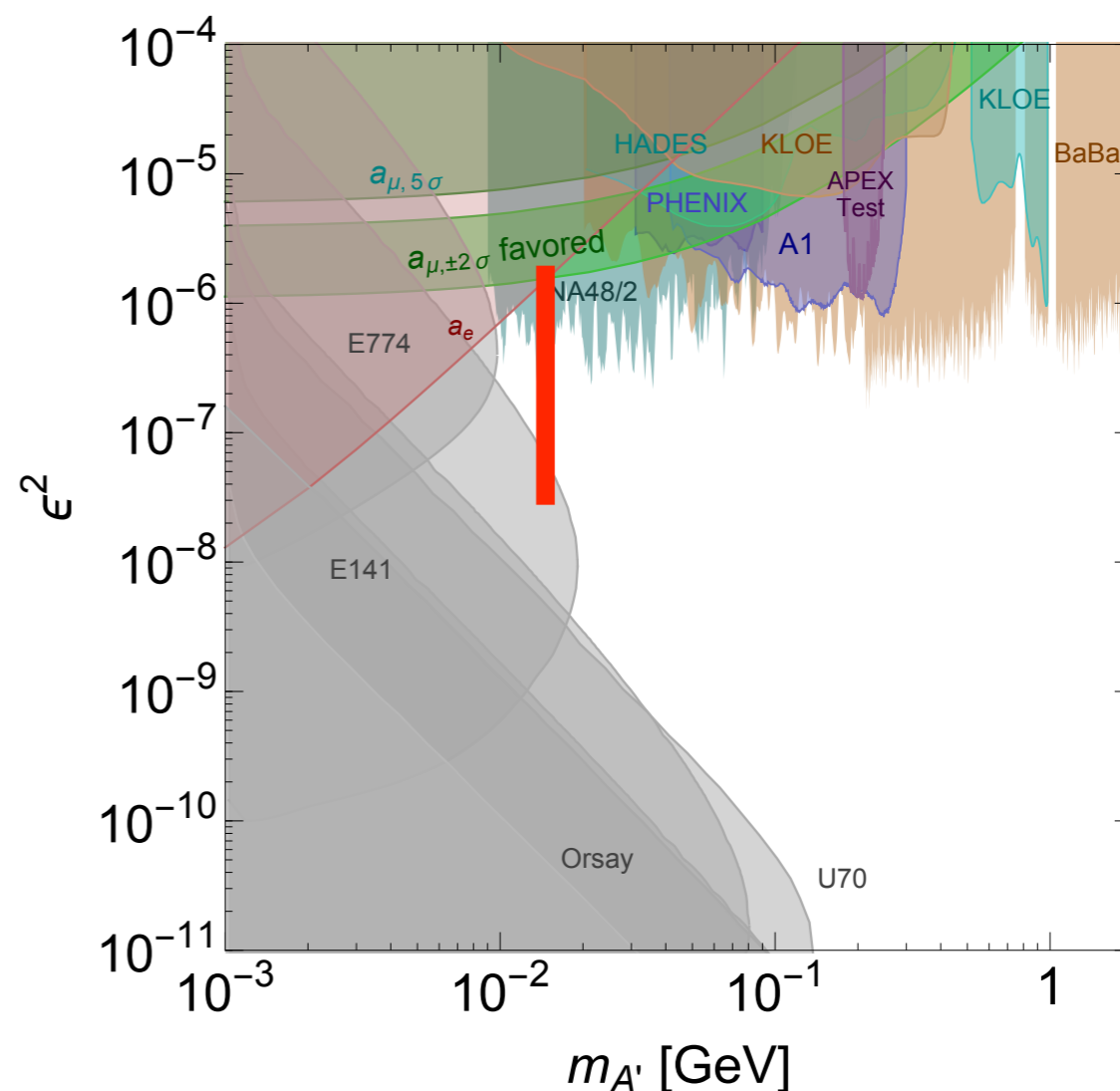
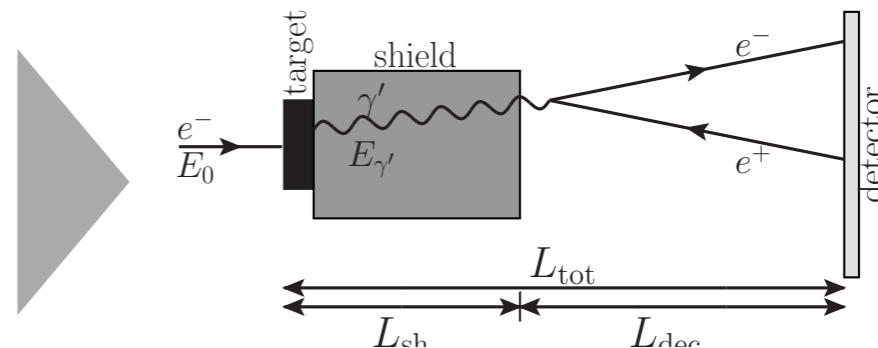


Figure 1. Scintillation emission spectrum of BGO

Visible search status

Techniques:

- beam dump (bremsstrahlung)
 - A' decay products detection after high z target (A' production) + shield (SM absorption)
- fixed target (bremsstrahlung, annihilation)
 - bump hunt in invariant mass spectrum, displaced vertices
- meson decay
 - only if A' couples w/ quarks
 - old experiments reanalysis

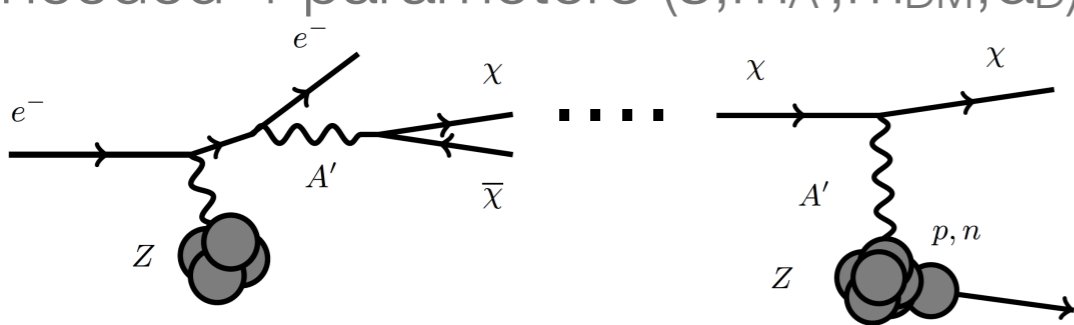


$(g-2)_\mu$ excluded in the simplest model, but still a lot of interest. In particular the ^8Be anomaly.

Invisible search status

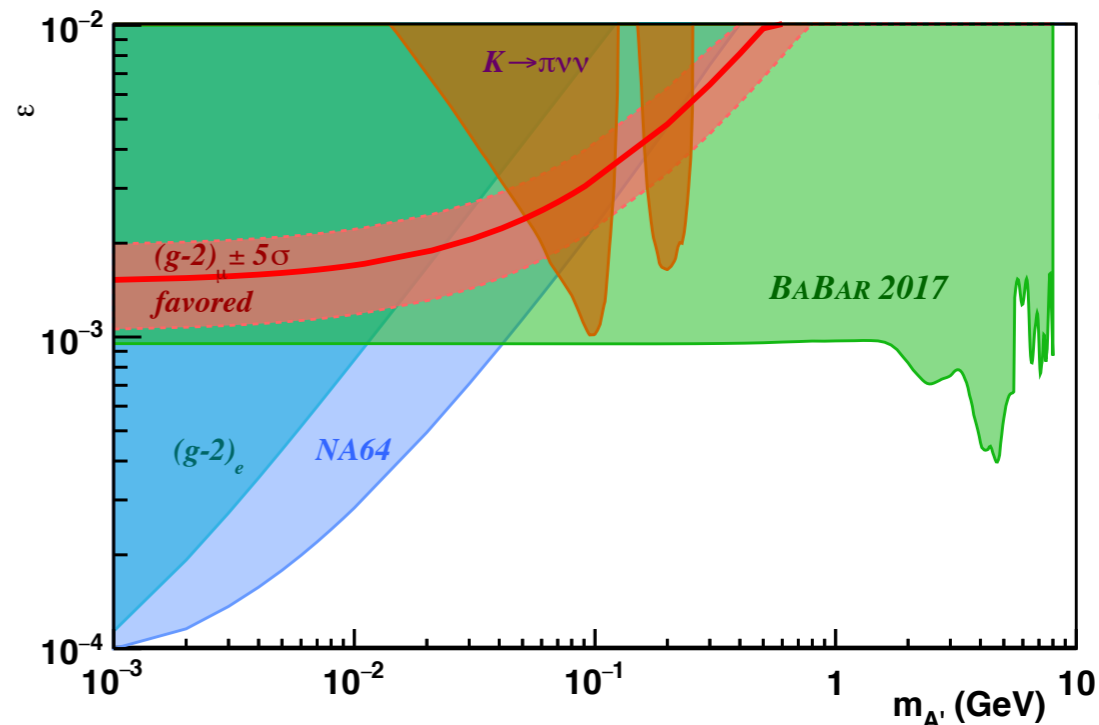
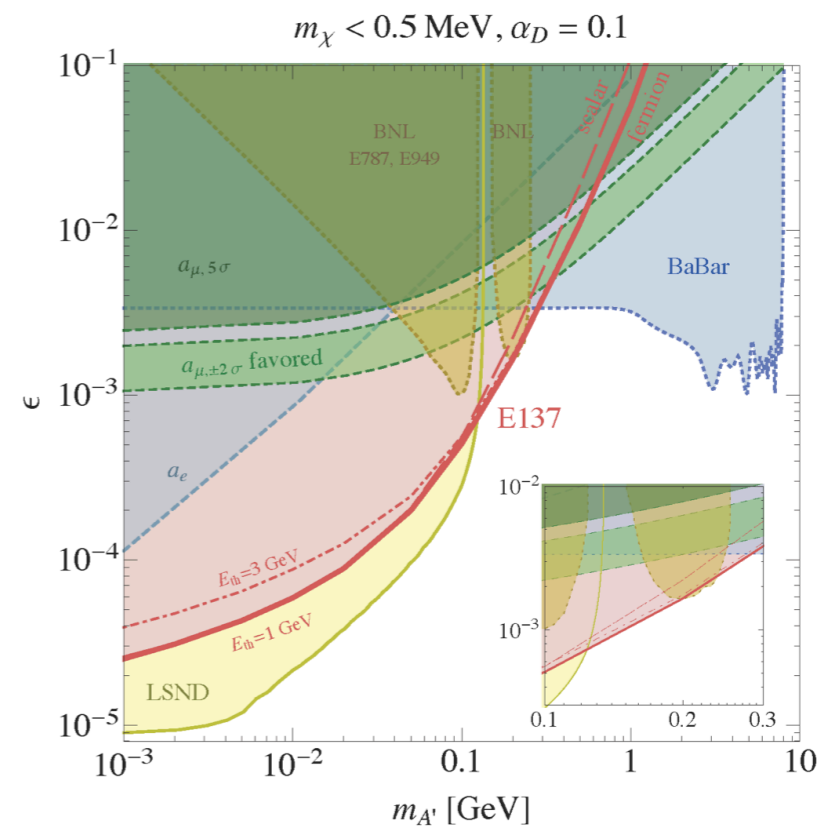
Techniques:

- DM scattering (bremsstrahlung)
 - detect the produced DM by scattering
 - needed 4 parameters ($\epsilon, m_{A'}, m_{\text{DM}}, \alpha_D$)



- missing energy/momentum search (bremsstrahlung)
 - not kinematically constrained process
 - observed energy/momentum smaller than expected

- missing mass search (annihilation)
 - kinematically constrained process
 - no assumption on A' decay chain



Not directly comparable

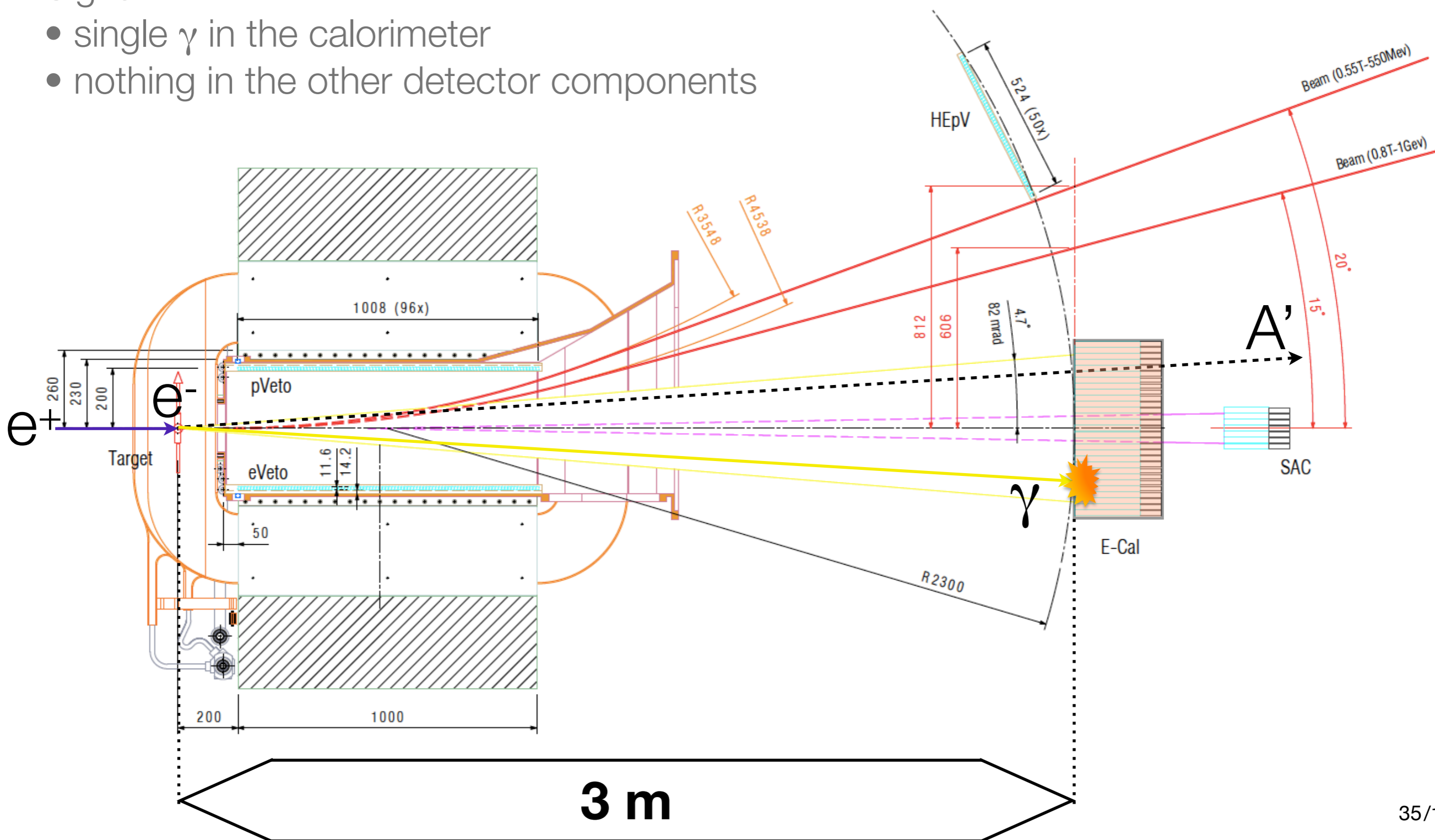
Beam Test Facility parasitic and dedicated modes

	Parasitic mode (DAΦNE working)		Dedicated mode	
	W/ target	W/o target	W/ target	W/o target
Particle species	e ⁺ /e ⁻ selectable by user	e ⁺ /e ⁻ depending on DAΦNE mode	e ⁺ /e ⁻ selectable by user	
Energy [MeV]	25-500	510	25-700 (e ⁺) 25-700 (e ⁻)	250-730 (e ⁺) 250-530 (e ⁻)
Energy spread	1% @ 500 MeV	1%	1%	
Rep. rate [Hz]	10-49 depending on DAΦNE mode		1-49 selectable by user	
Pulse duration [ns]	10		1.5-40 selectable by user	
Intensity [particles/bunch]	1-10 ⁵ depending on energy	10 ⁷ -1.5 · 10 ¹⁰	1-10 ⁵ depending on energy	10 ³ -3 · 10 ¹⁰
Max average flux	3.125 · 10 ¹⁰ particles/s			
Spot size [mm]	0.5-25 (y) × 0.6-55 (x)			
Divergence [mrad]	1-1.5			

Detector top view (w/ signal)

Signal:

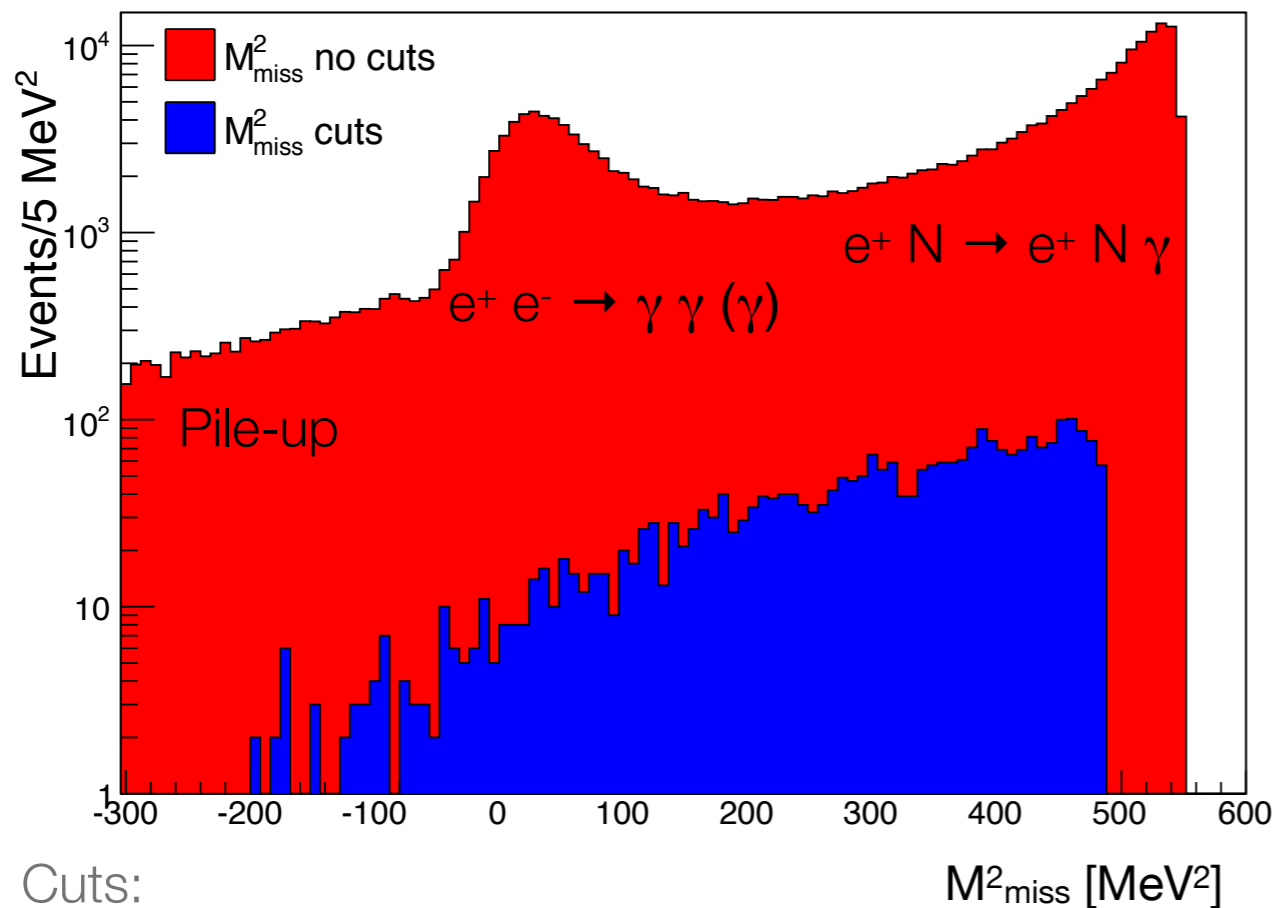
- single γ in the calorimeter
- nothing in the other detector components



Backgrounds

Largest backgrounds:

- $e^+ e^- \rightarrow \gamma \gamma (\gamma)$
- $e^+ N \rightarrow e^+ N \gamma$
- pile-up

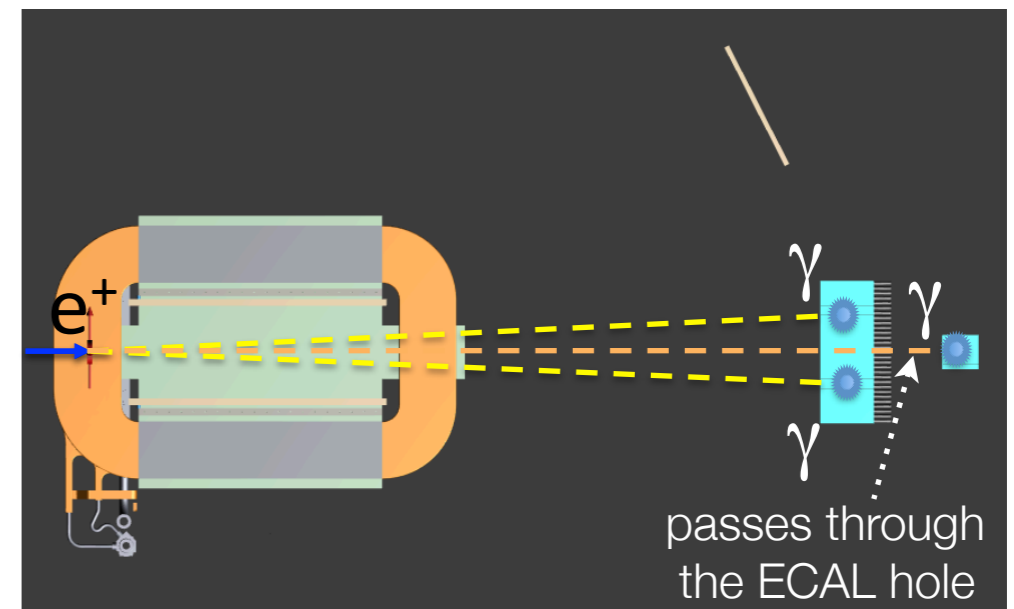


Cuts:

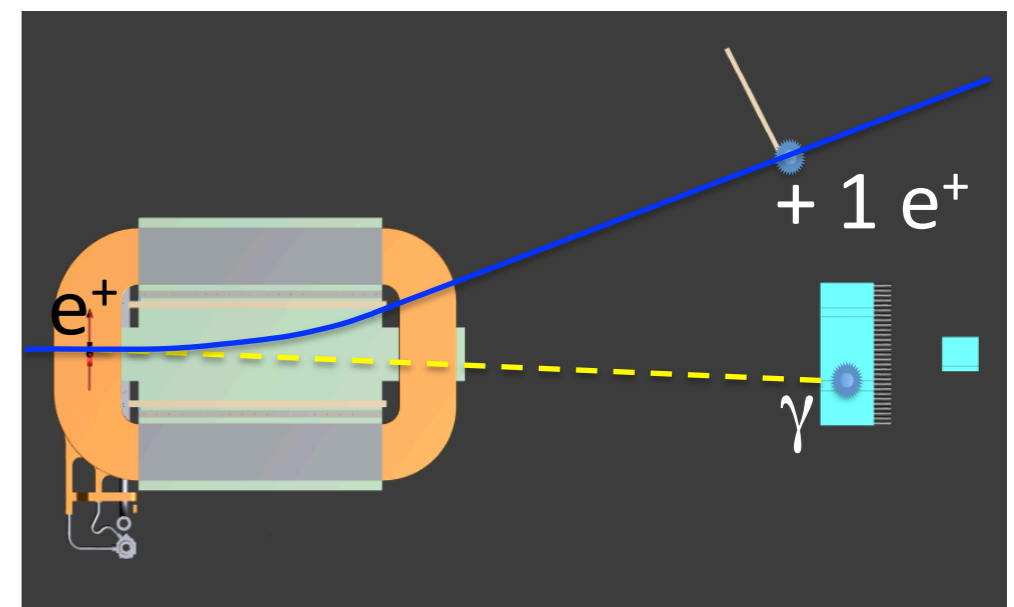
- 1 cluster in ECAL fiducial volume
- no hits in vetoes
- no γ in the SAC w/ $E_\gamma > 50$ MeV
- $20-150 \text{ MeV} < E_\gamma < 120-350 \text{ MeV}$ (depending on $m_{A'}$)

Backgrounds geometry

Annihilation (+ISR): $e^+ e^- \rightarrow \gamma \gamma (\gamma)$

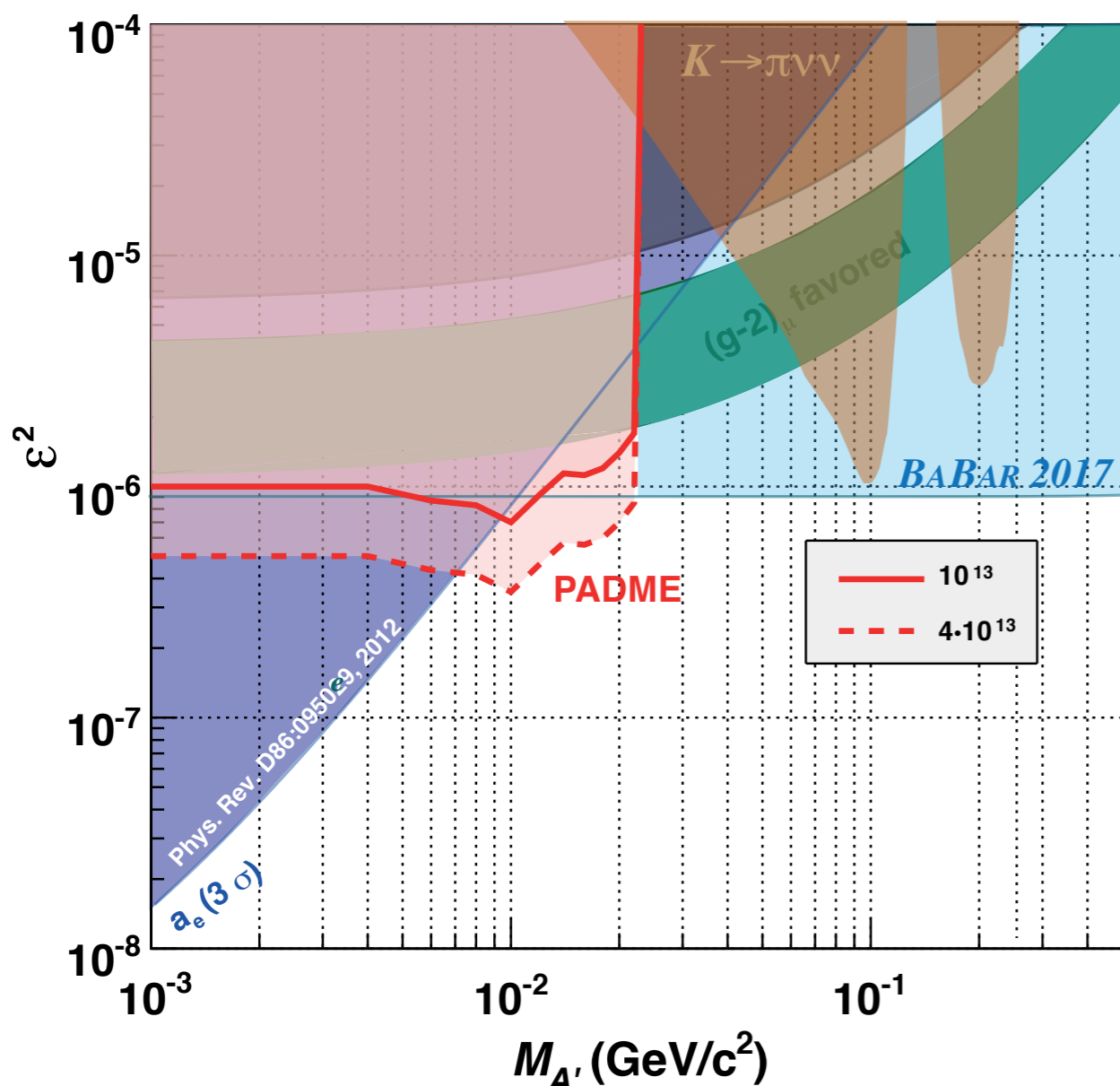


Bremsstrahlung: $e^+ N \rightarrow e^+ N \gamma$



Sensitivity

Based on $2.5 \cdot 10^{10}$ fully GEANT4 simulated 550 MeV e^+ on target events.
Number of BG events is extrapolated to 10^{13} e^+ on target.



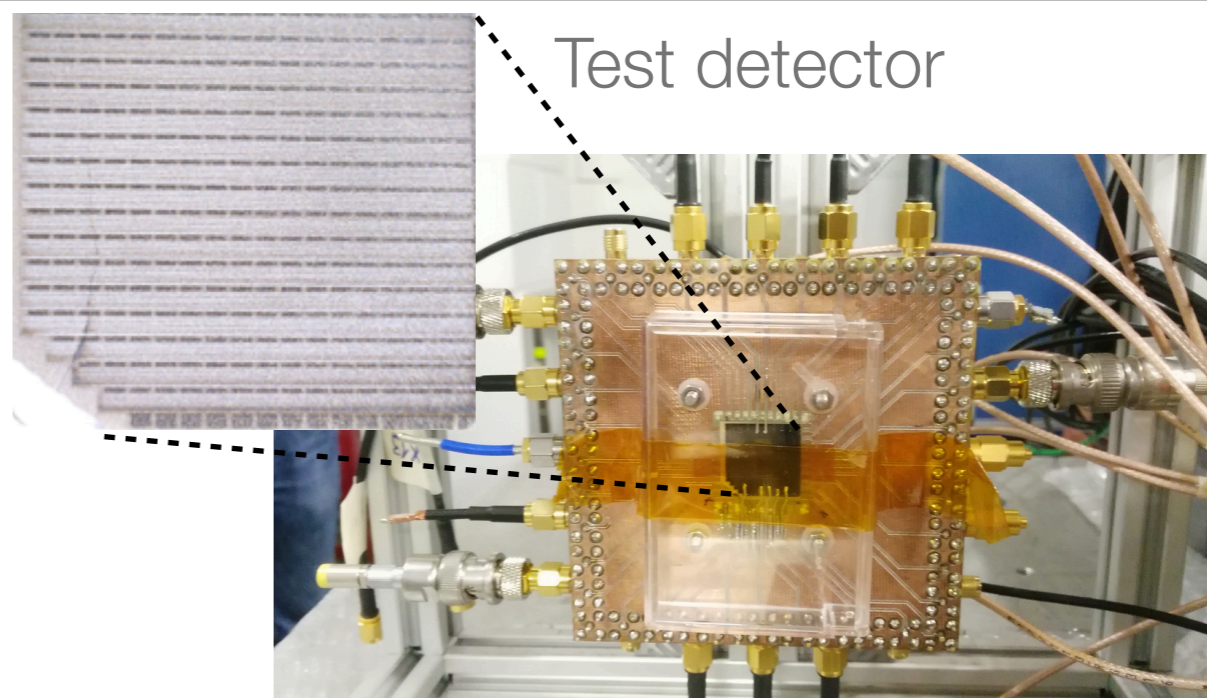
PADME can explore in a model-independent way the region down to $\epsilon \approx 10^{-3}$ w/:

- $m_{A'} < 23.7$ MeV ($E_{\text{beam}} = 550$ MeV)
- $m_{A'} < 27.7$ MeV ($E_{\text{beam}} = 750$ MeV)
- $m_{A'} < 32$ MeV ($E_{\text{beam}} = 1$ GeV)

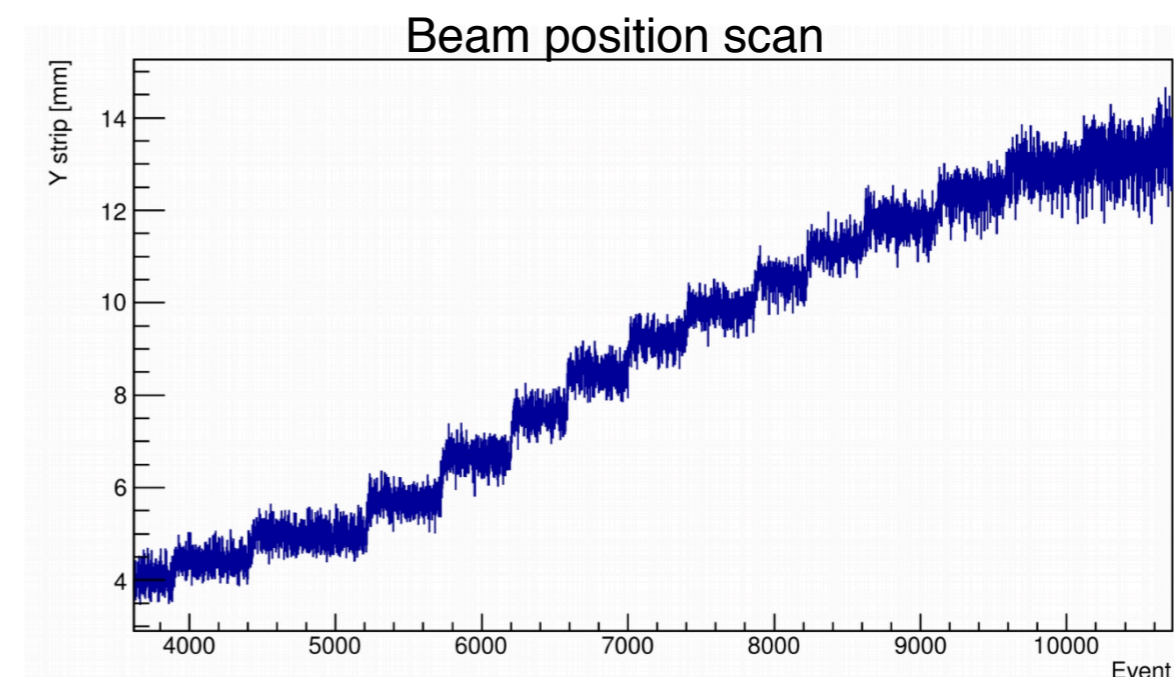
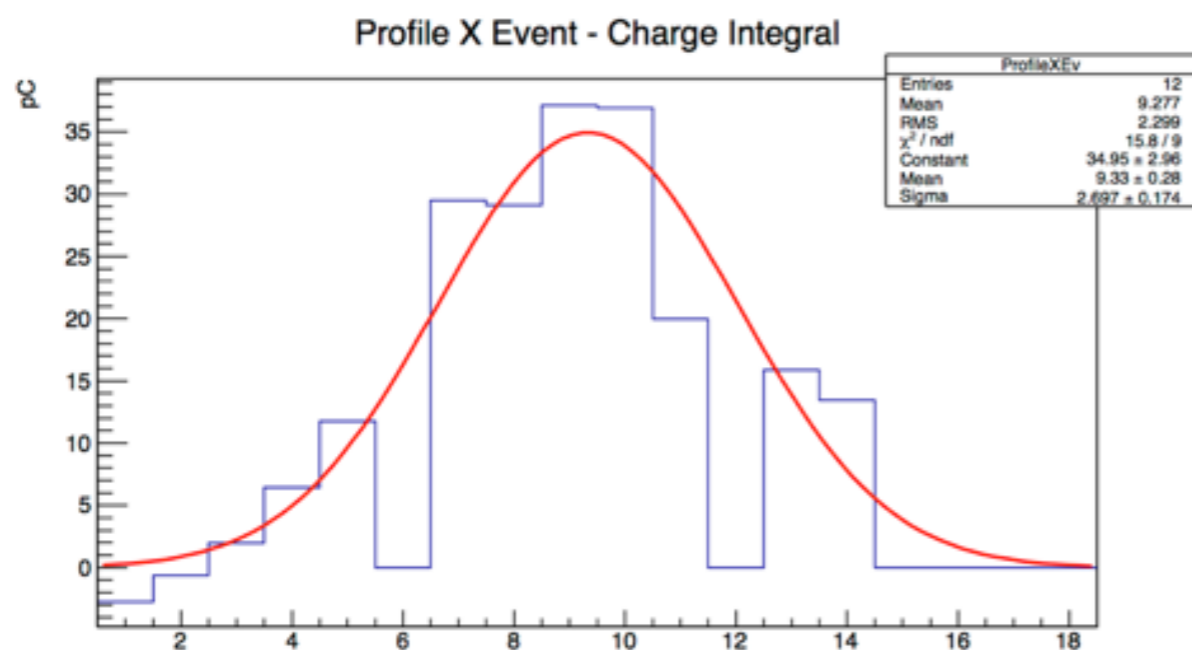
Active target

Features:

- Diamond (low z, reduced brems.)
- Dim.: $20 \times 20 \times 0.1 \text{ mm}^3$
- 19 horiz. \times 19 vert. active graphitic strips (average informations on beam)
- σ_{x-y} (beam position) $< 2 \text{ mm}$
- in vacuum w/ movement system



Test detector results (18 h. \times 18 v. active strips)

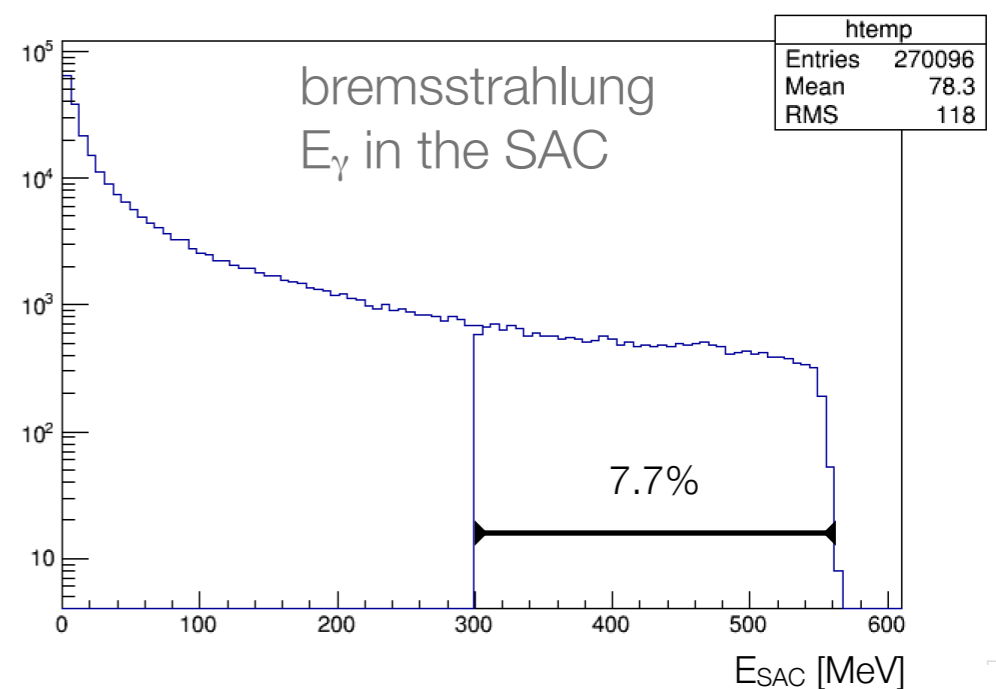
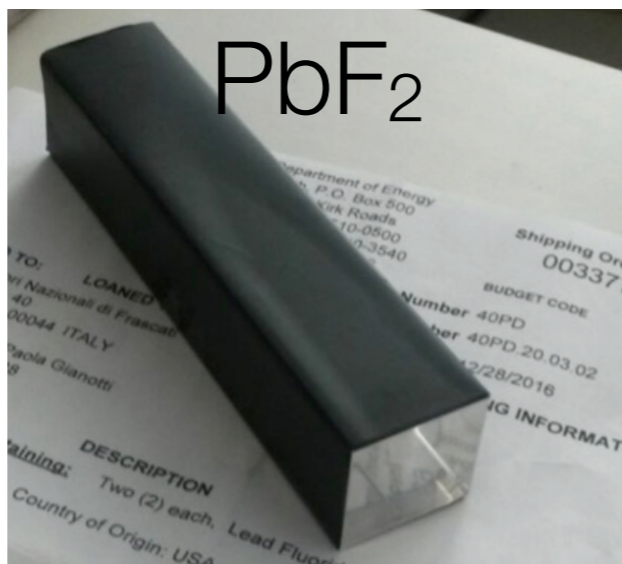
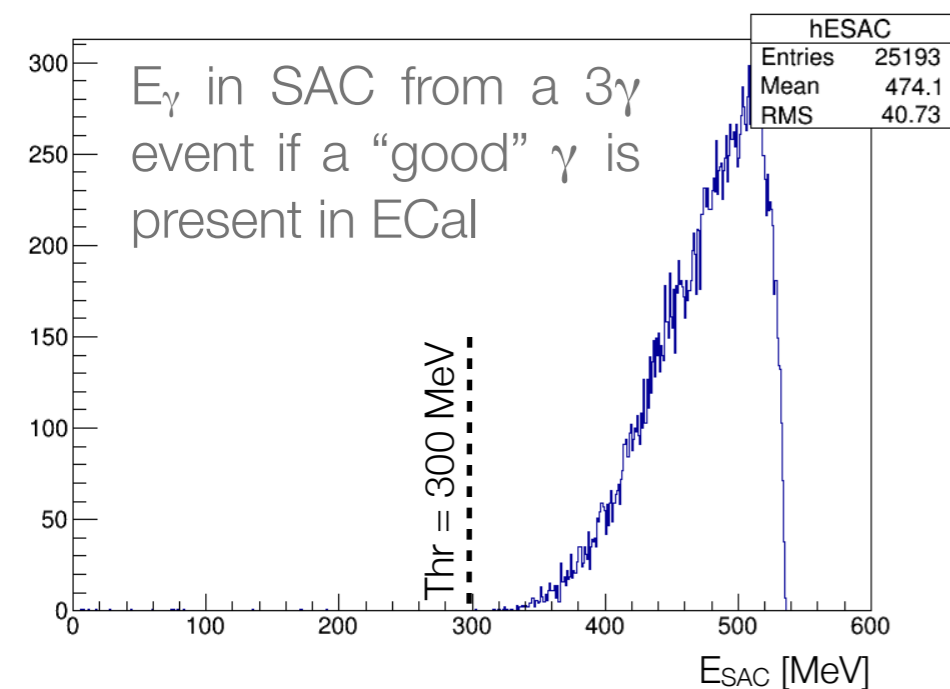
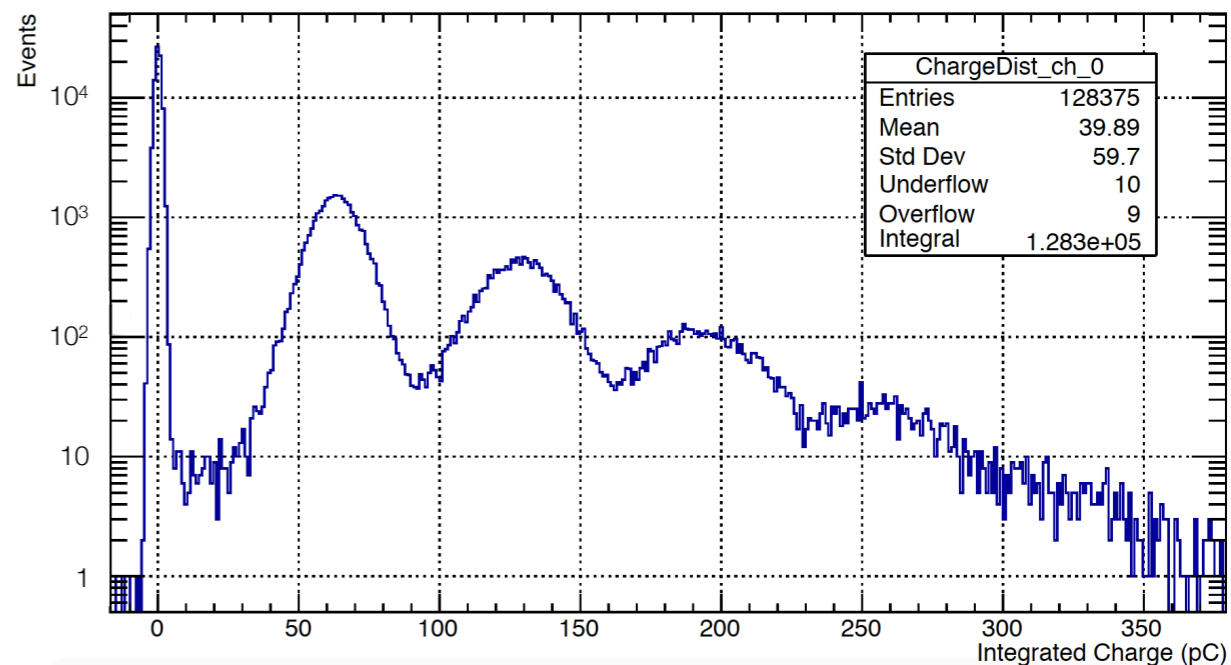


Spatial resolution: 0.2 mm (x-axis) \times 0.3 mm (y-axis)

Small Angle Calorimeter (SAC)

Characteristics:

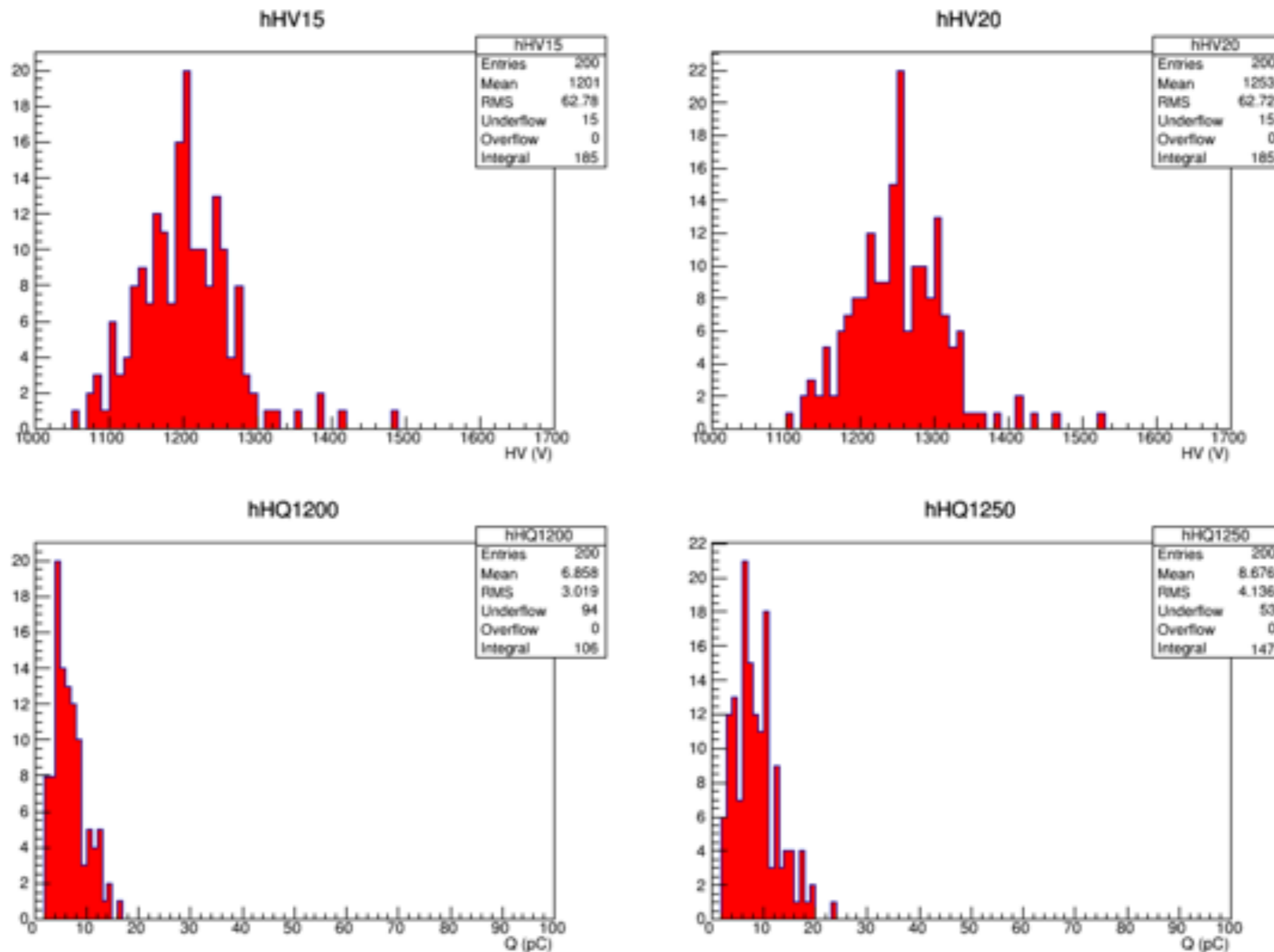
- $\sigma_E \approx 10\%$
- Cherenkov \rightarrow 3-4 ns signals
- angular coverage: $[0,20]$ mrad
- crystal wrapped w/ tedlar (only direct light)



SAC must be sensible to photons over 300 MeV and blind under 100 MeV

HV (given charge) and charge (given HV) histo

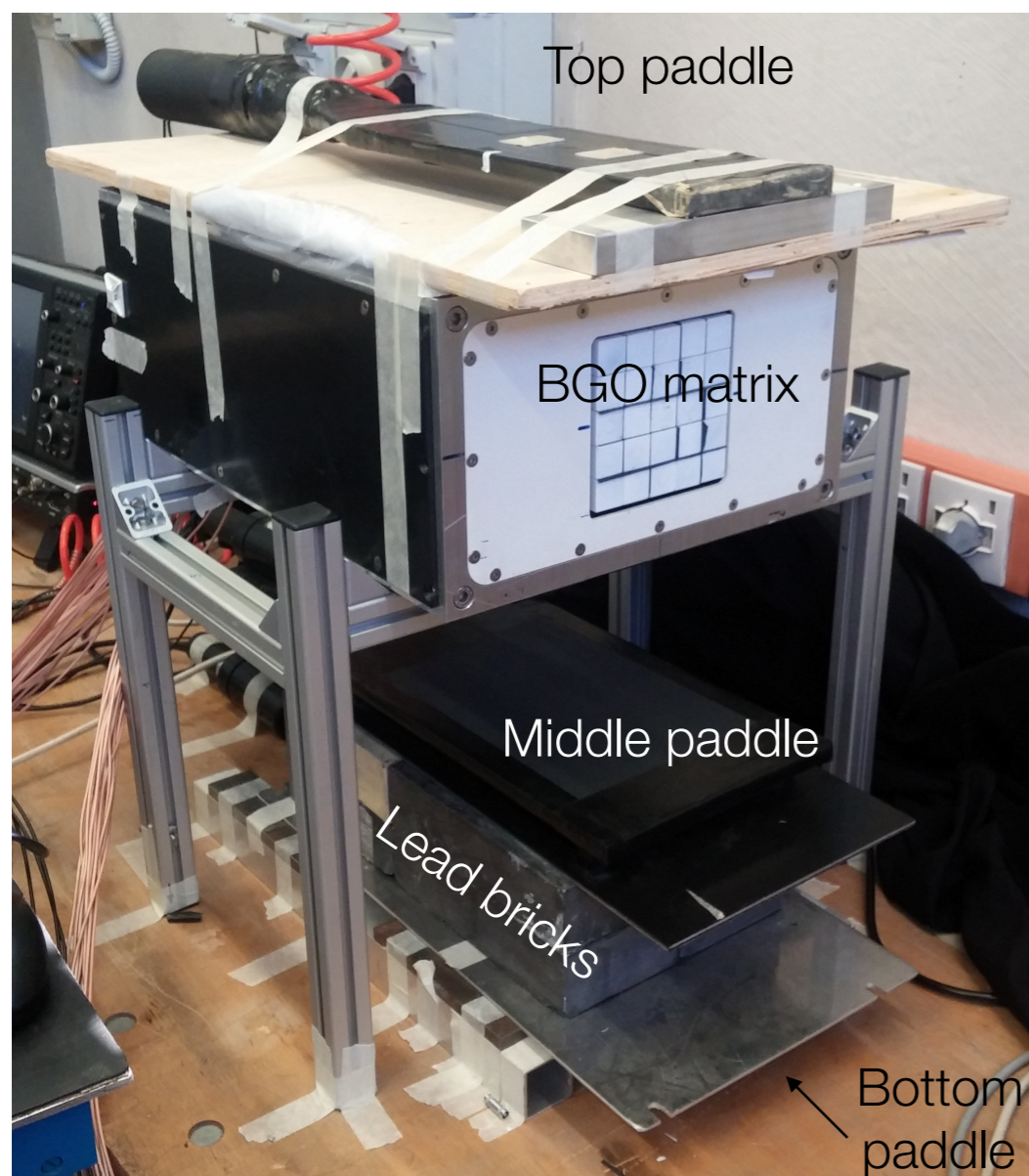
From ^{22}Na source measurements



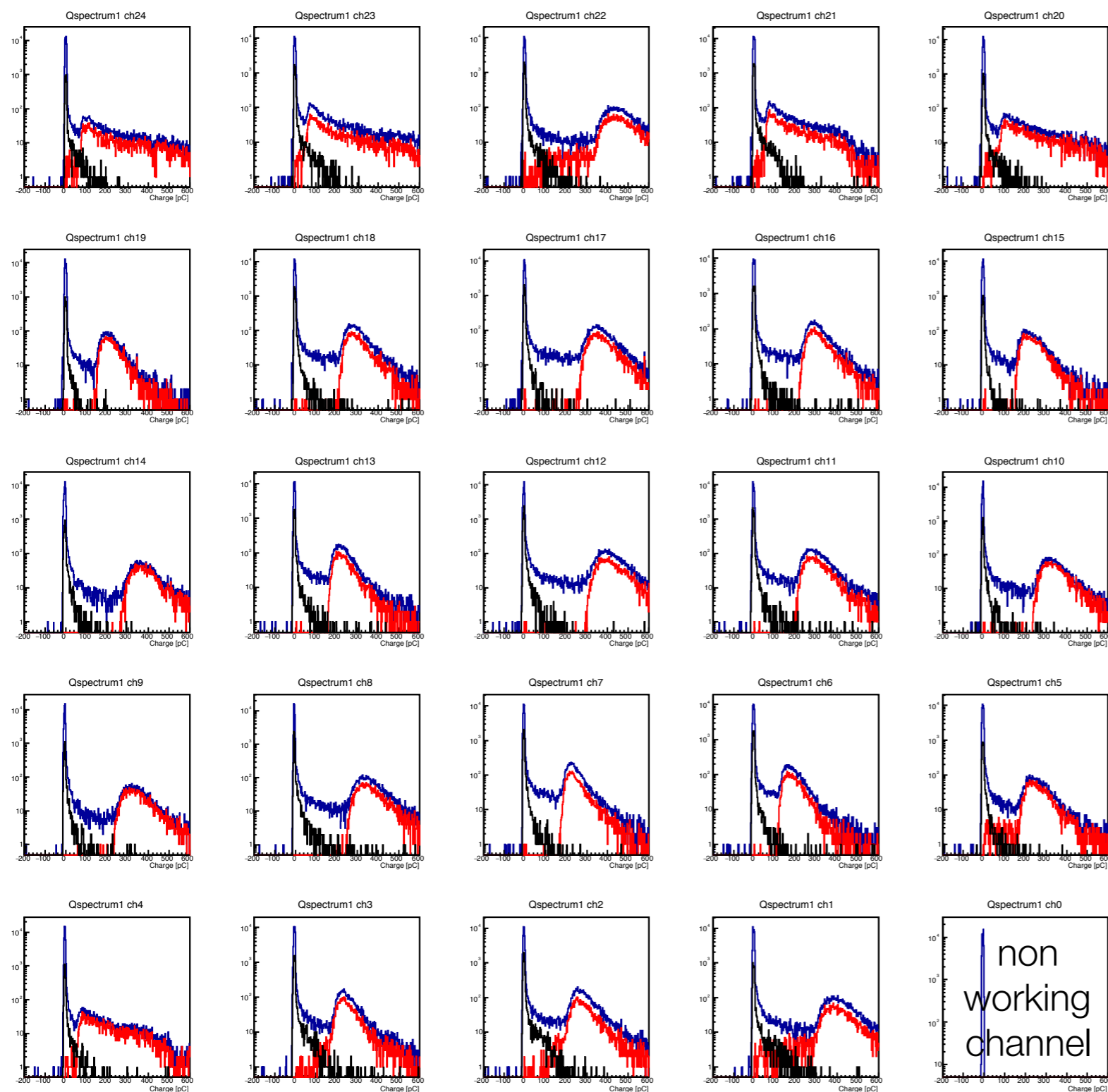
Cosmic ray setups

We performed CR runs with 2 different setups:

- 4x3 matrix
- 5x5 matrix with 50 μ m tedlar foils between crystals (see next slides)



Cosmic rays charge spectra (5x5 matrix)



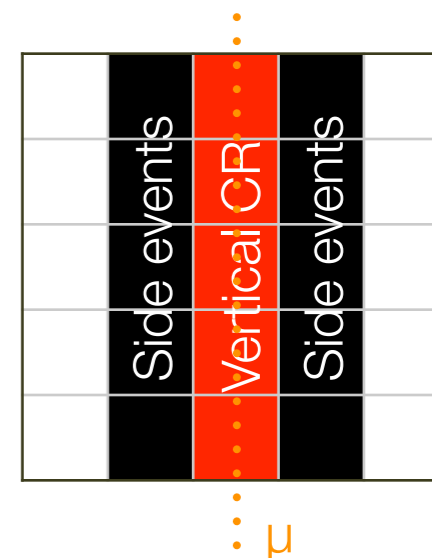
All events

Vertical CR

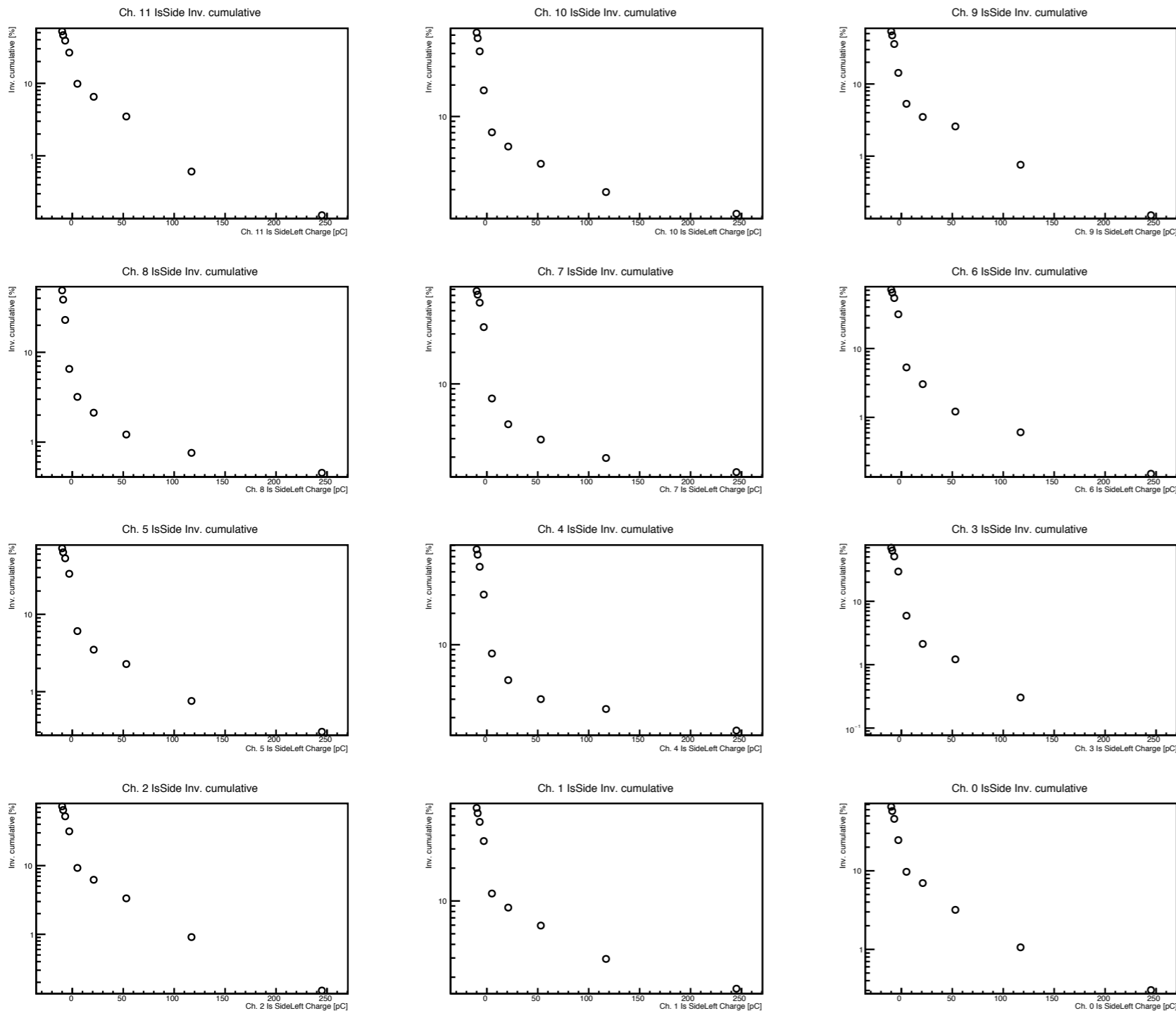
Side events (vertical CR passes through a crystal on the side)

Verticality is obtained requiring that the 5 largest signals are in column

Example:
 μ passing through central column



Optical crosstalk without tedlar (4x3 matrix)



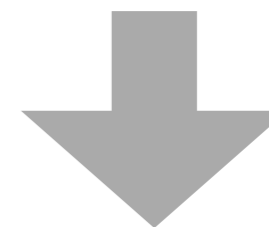
Inverse cumulative of the Side events without tedlar

1% is reached at ≈ 100 pC

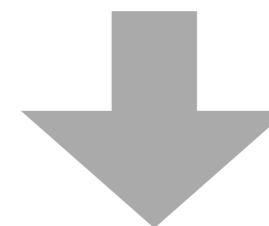
Optical crosstalk with tedlar (4x3 matrix)

Inverse cumulative of the Side events with tedlar

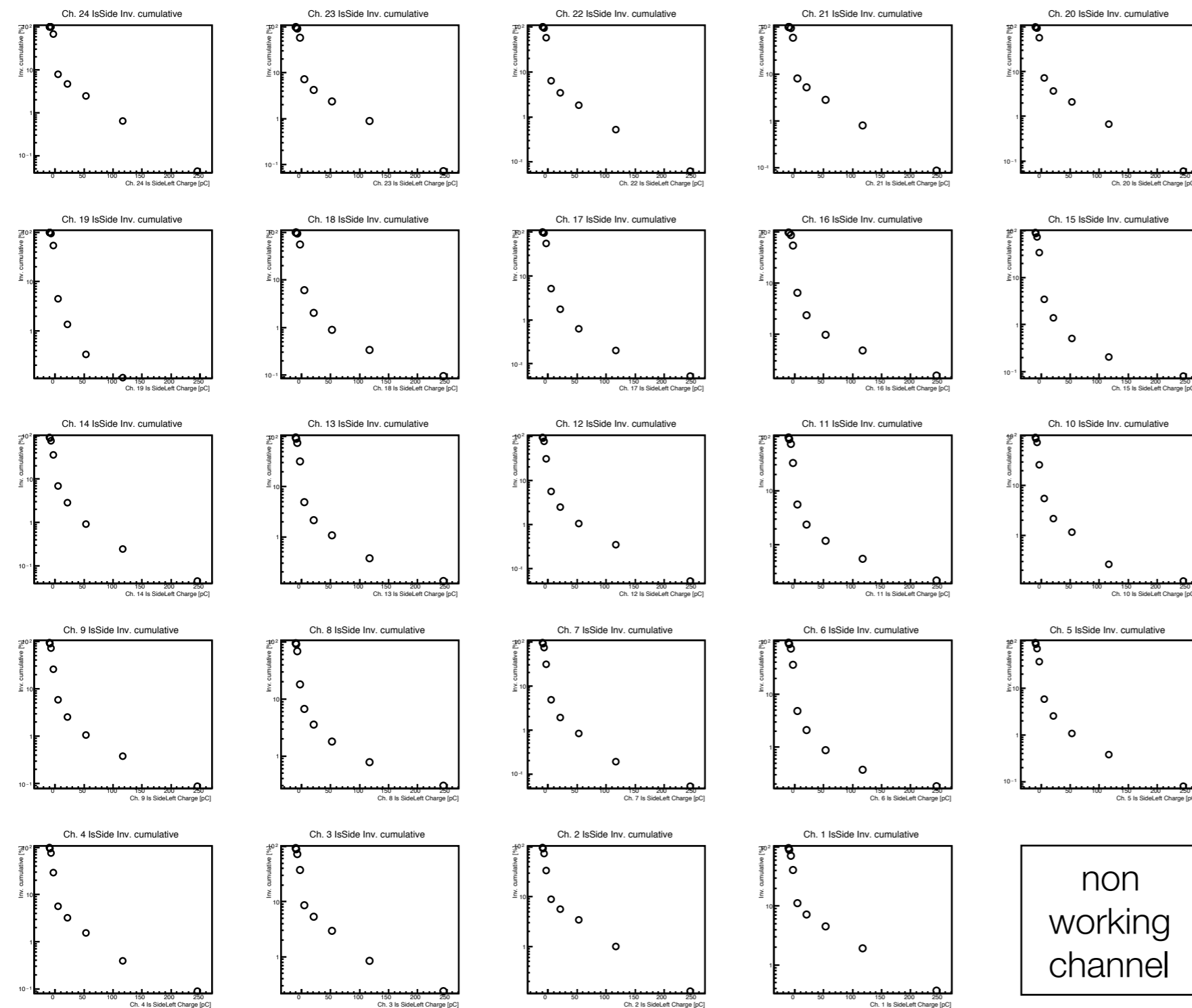
1% is reached at ~50pC



Tedlar is effective in preventing optical crosstalk

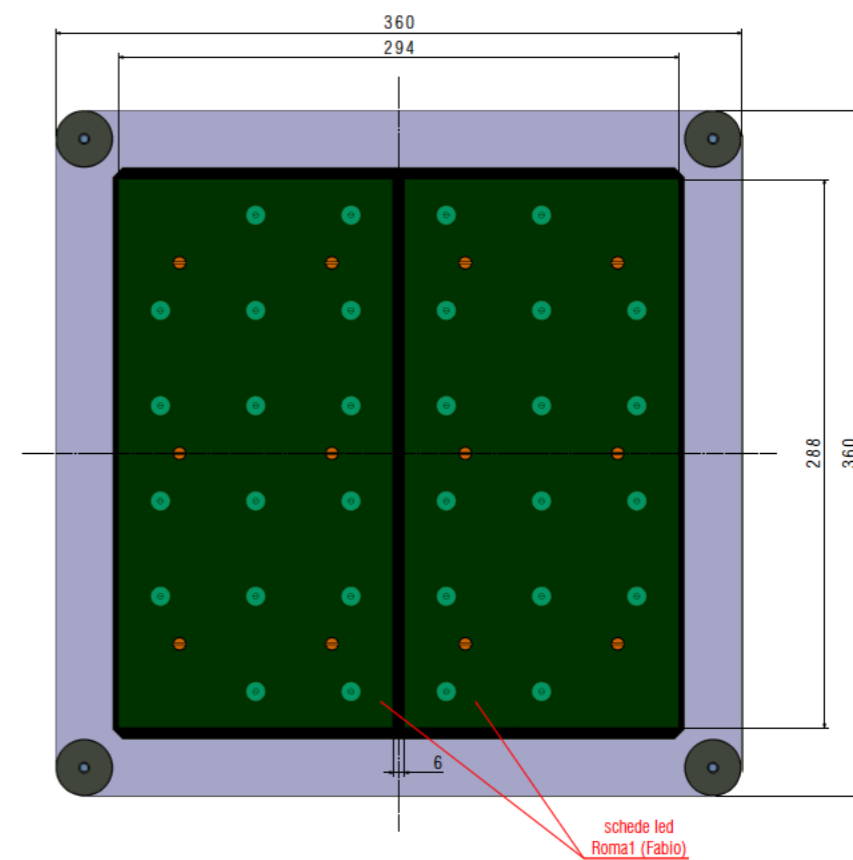
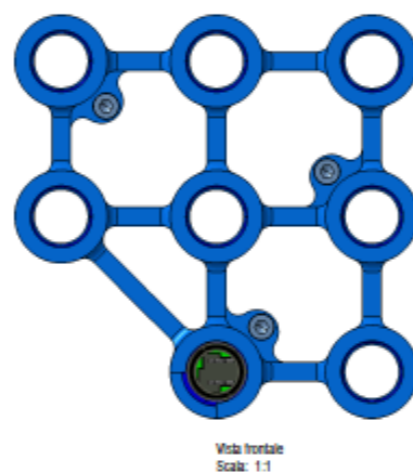
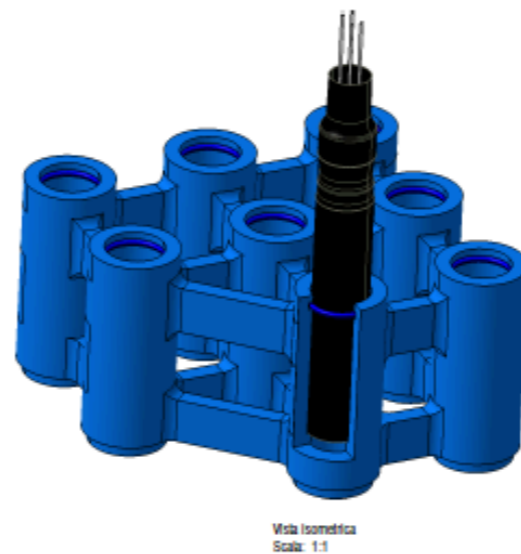
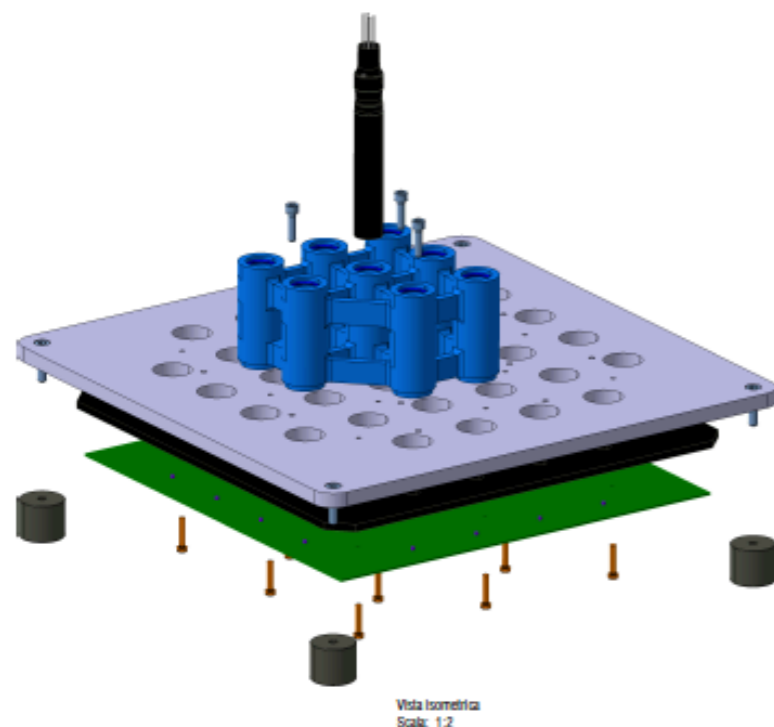


Tedlar will be used for the ECAL assembly (it also accommodates different SU heights)

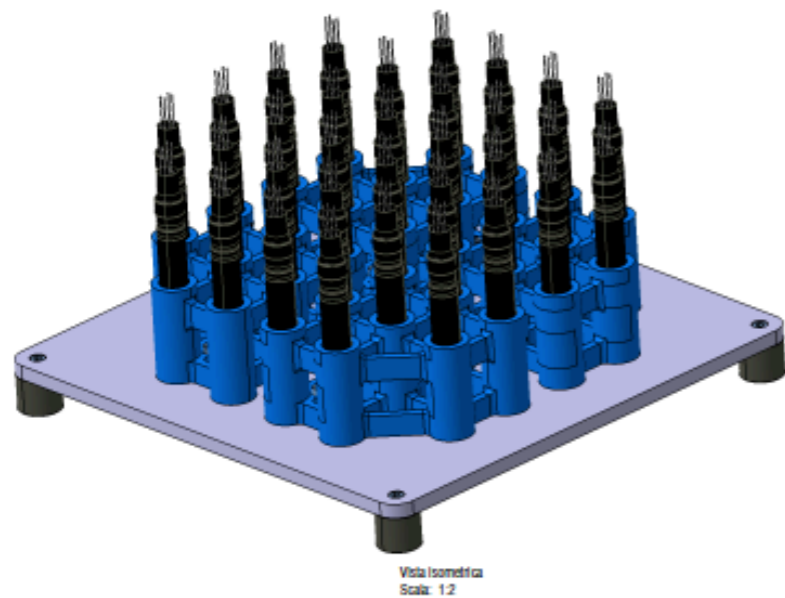


non working channel

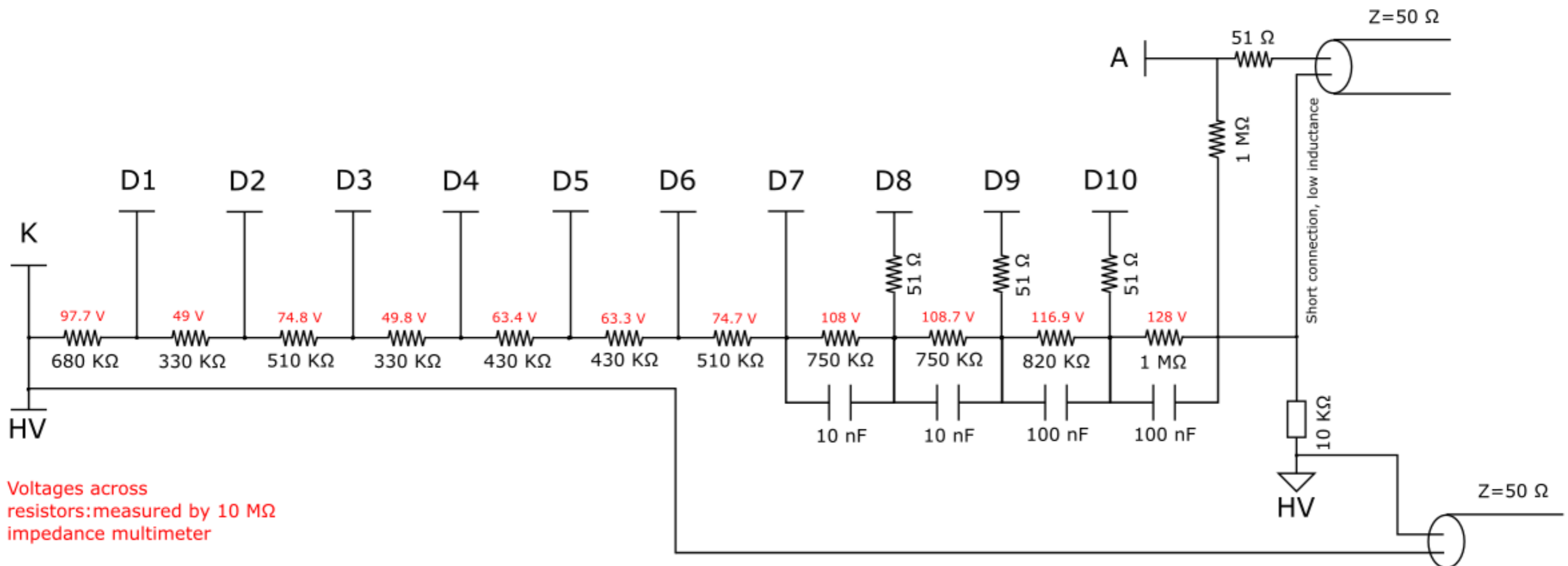
PMTs test station



Seen from below



XP1911 divider new design



PMTs dimensions test



DIMENSIONAL TEST WITH "TUBE" ID=20,4 H8

GOOD CASE:
Photomultiplier
pass through



BAD CASE:
Photomultiplier
does not pass
through

