



# SUPERWEAK FORCE

---

Zoltán Trócsányi

Eötvös University and MTA-DE Particle Physics Research Group

FFK-2019, Tihany, 11 June 2019

---

---

# OUTLINE

---

1. Status of particle physics
2.  $U(1)_Z$  extension of SM
3. Constraints on the parameter space

---

## Status of particle physics: energy frontier

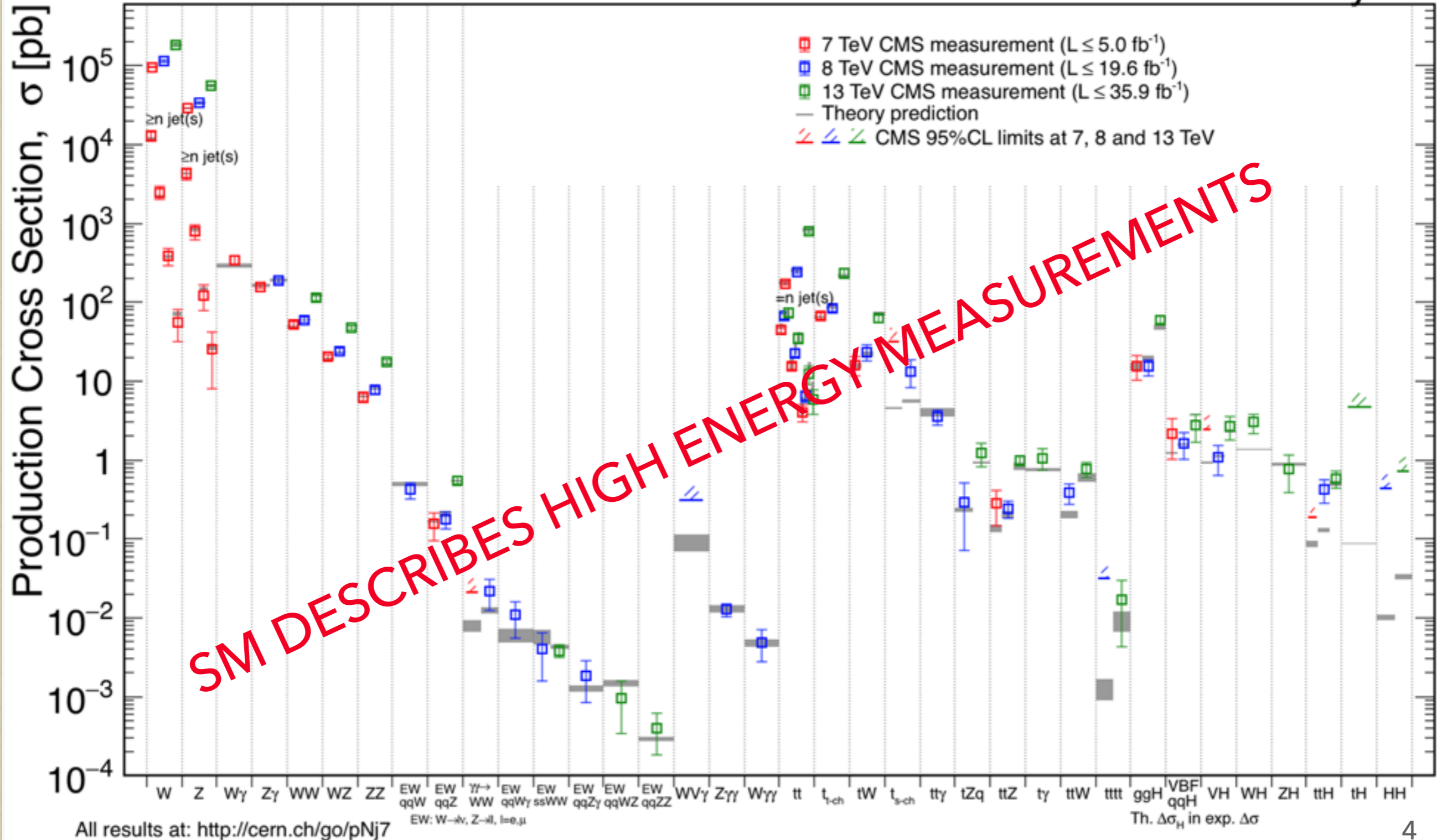
---

- LEP, LHC: SM describes final states of particle collisions precisely [remember talks of Monday afternoon]

# SM@LHC: theory vs. 36 measurements at CMS

July 2018

CMS Preliminary



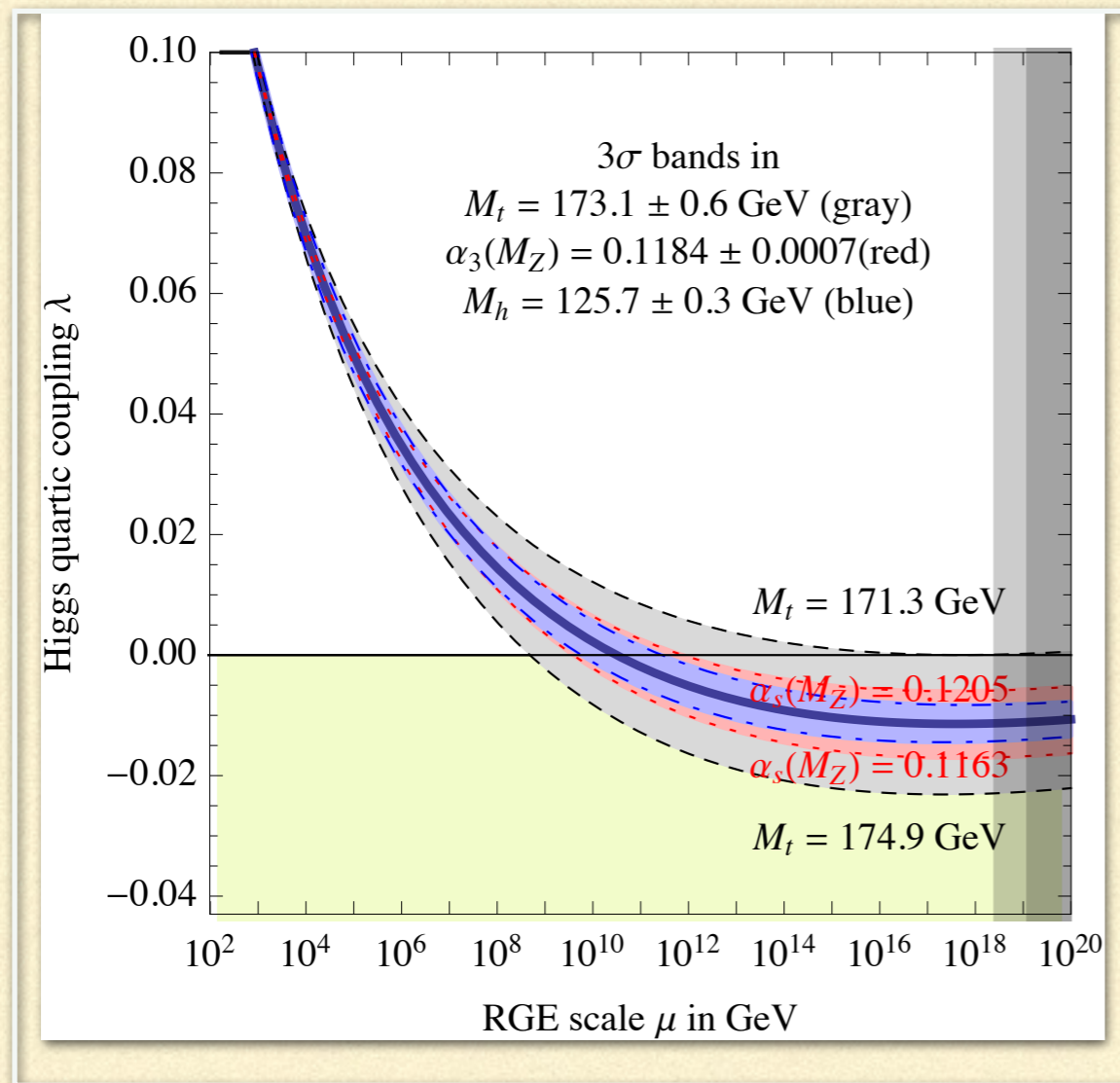
---

## Status of particle physics: energy frontier

---

- LEP, LHC: SM describes final states of particle collisions precisely [remember talks of Monday afternoon]
- **SM is unstable**

# SM is unstable



Degrassi et al., arXiv:1205.6497

---

# Status of particle physics: energy frontier

---

- LEP, LHC: SM describes final states of particle collisions precisely
- SM is unstable
- **No proven sign of new physics beyond SM** at colliders\*

\*There are some indications **below discovery significance** (such as muon anomalous magnetic moment, lepton flavor non-universality in meson decays)

---

# Status of particle physics: intensity frontier

---

- Universe at large scale described precisely by cosmological SM:  $\Lambda$ CDM ( $\Omega_m = 0.3$ )



---

# Status of particle physics: intensity frontier

---

- Universe at large scale described precisely by cosmological SM:  $\Lambda$ CDM ( $\Omega_m = 0.3$ )
- Neutrino flavours oscillate

---

## Status of particle physics: intensity frontier

---

- Universe at large scale described precisely by cosmological SM:  $\Lambda$ CDM ( $\Omega_m = 0.3$ )
- Neutrino flavours oscillate
- Existing baryon asymmetry cannot be explained by CP asymmetry in SM

---

## Status of particle physics: intensity frontier

---

- Universe at large scale described precisely by cosmological SM:  $\Lambda$ CDM ( $\Omega_m = 0.3$ )
- Neutrino flavours oscillate
- Existing baryon asymmetry cannot be explained by CP asymmetry in SM
- Inflation of the early, accelerated expansion of the present Universe

---

## Extension of SM

---

- There are many extensions proposed, mostly with the aim of **predicting some observable effect at the LHC** – but **there are none so far**, so may give up

---

## Extension of SM

---

- There are many extensions proposed, mostly with the aim of **predicting some observable effect at the LHC** – but **there are none so far**, so may give up
- SM is highly efficient – let us **stick to efficiency** the only exception of economical description is the relatively large number of Yukawa couplings

---

# Extension of SM

---

- **Neutrinos must play a key role**

with non-zero masses they must feel another force apart from the weak one, such as Yukawa coupling to a scalar, which requires the existence of right-handed neutrinos

---

# Extension of SM

---

- **Neutrinos must play a key role**

with non-zero masses they must feel another force apart from the weak one, such as Yukawa coupling to a scalar, which requires the existence of right-handed neutrinos

- **Simplest extension** of  $G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$  is to  $G = G_{SM} \times U(1)_Z$

---

# Extension of SM

---

- **Neutrinos must play a key role**

with non-zero masses they must feel another force apart from the weak one, such as Yukawa coupling to a scalar, which requires the existence of right-handed neutrinos

- **Simplest extension** of  $G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$  is to  $G = G_{SM} \times U(1)_Z$

renormalizable gauge theory without any other symmetry

- Fix Z-charges by requirement of



---

# Extension of SM

---

- **Neutrinos must play a key role**

with non-zero masses they must feel another force apart from the weak one, such as Yukawa coupling to a scalar, which requires the existence of right-handed neutrinos

- **Simplest extension** of  $G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$  is to  $G = G_{SM} \times U(1)_Z$

renormalizable gauge theory without any other symmetry

- Fix Z-charges by requirement of

- gauge and gravity **anomaly cancellation** and

---

# Extension of SM

---

- **Neutrinos must play a key role**

with non-zero masses they must feel another force apart from the weak one, such as Yukawa coupling to a scalar, which requires the existence of right-handed neutrinos

- **Simplest extension** of  $G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$  is to  $G = G_{SM} \times U(1)_Z$

renormalizable gauge theory without any other symmetry

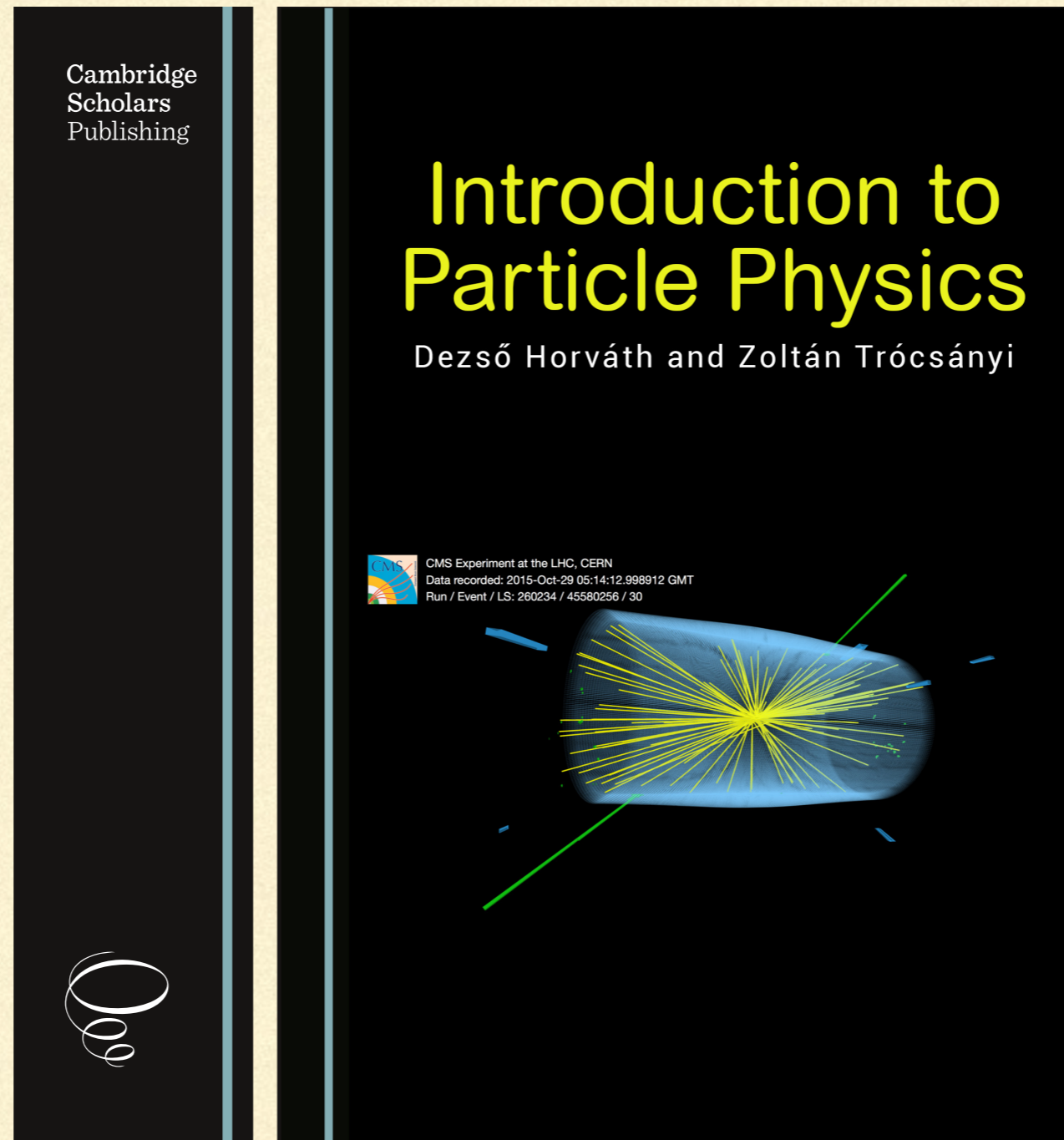
- Fix Z-charges by requirement of

- gauge and gravity **anomaly cancellation** and
- **gauge invariant Yukawa terms for neutrino** mass generation

---

Focus only on addition to the SM,  
find SM in this new book:

---



# Fermions

- fermion fields:

$$\psi_{q,1}^f = \begin{pmatrix} U^f \\ D^f \end{pmatrix}_L \quad \psi_{q,2}^f = U_R^f, \quad \psi_{q,3}^f = D_R^f$$

$$\psi_{l,1}^f = \begin{pmatrix} \nu^f \\ \ell^f \end{pmatrix}_L \quad \psi_{l,2}^f = \nu_R^f, \quad \psi_{l,3}^f = \ell_R^f$$

where

$$\psi_{L/R} \equiv \psi_{\mp} = \frac{1}{2} (1 \mp \gamma_5) \psi \equiv P_{L/R} \psi$$

( $\nu_L$  can  $\nu_R$  can also be Majorana neutrinos, embedded into different Dirac spinors)

- covariant derivatives:

$$D_j^\mu = \partial^\mu + ig_L \mathbf{T} \cdot \mathbf{W}^\mu + ig_Y y_j B'^\mu + i(g'_Z z_j - g'_Y y_j) Z'^\mu$$

# Anomaly free charge assignment

field	$SU(3)_c$	$SU(2)_L$	$y_j$	$z_j$	$z_j$	$r_j = z_j/z_\phi - y_j$
$U_L, D_L$	3	2	$\frac{1}{6}$	$z_1$	$\frac{1}{6}$	0
$U_R$	3	1	$\frac{2}{3}$	$z_2$	$\frac{7}{6}$	$\frac{1}{2}$
$D_R$	3	1	$-\frac{1}{3}$	$2z_1 - z_2$	$-\frac{5}{6}$	$-\frac{1}{2}$
$\nu_L, \ell_L$	1	2	$-\frac{1}{2}$	$-3z_1$	$-\frac{1}{2}$	0
$\nu_R$	1	1	0	$z_2 - 4z_1$	$\frac{1}{2}$	$\frac{1}{2}$
$\ell_R$	1	1	-1	$-2z_1 - z_2$	$-\frac{3}{2}$	$-\frac{1}{2}$
$\phi$	1	2	$\frac{1}{2}$	$z_\phi$	1	$\frac{1}{2}$
$\chi$	1	1	0	$z_\chi$	-1	-1

essentially from neutrino-scalar interactions

# Anomaly free charge assignment

re-parametrization

field	$SU(3)_c$	$SU(2)_L$	$y_j$	$z_j$	$z_j$	$r_j = z_j/z_\phi - y_j$
$U_L, D_L$	3	2	$\frac{1}{6}$	$z_1$	$\frac{1}{6}$	0
$U_R$	3	1	$\frac{2}{3}$	$z_2$	$\frac{7}{6}$	$\frac{1}{2}$
$D_R$	3	1	$-\frac{1}{3}$	$2z_1 - z_2$	$-\frac{5}{6}$	$-\frac{1}{2}$
$\nu_L, \ell_L$	1	2	$-\frac{1}{2}$	$-3z_1$	$-\frac{1}{2}$	0
$\nu_R$	1	1	0	$z_2 - 4z_1$	$\frac{1}{2}$	$\frac{1}{2}$
$\ell_R$	1	1	-1	$-2z_1 - z_2$	$-\frac{3}{2}$	$-\frac{1}{2}$
$\phi$	1	2	$\frac{1}{2}$	$z_\phi$	1	$\frac{1}{2}$
$\chi$	1	1	0	$z_\chi$	-1	-1

$$D_j^\mu = \partial^\mu + ig_L \mathbf{T} \cdot \mathbf{W}^\mu + iy_j g_Y B'^\mu + i(r_j g'_Z + y_j g'_{ZY}) Z'^\mu$$

---

# Scalars

---

- Standard  $\phi$  complex  $SU(2)_L$  doublet and new  $\chi$  complex singlet:

$$\mathcal{L}_{\phi,\chi} = [D_{\mu}^{(\phi)} \phi]^* D^{(\phi)\mu} \phi + [D_{\mu}^{(\chi)} \chi]^* D^{(\chi)\mu} \chi - V(\phi, \chi)$$

---

# Scalars

---

- Standard  $\phi$  complex  $SU(2)_L$  doublet and new  $\chi$  complex singlet:

$$\mathcal{L}_{\phi,\chi} = [D_{\mu}^{(\phi)} \phi]^* D^{(\phi)\mu} \phi + [D_{\mu}^{(\chi)} \chi]^* D^{(\chi)\mu} \chi - V(\phi, \chi)$$

- with scalar potential

$$V(\phi, \chi) = V_0 - \mu_{\phi}^2 |\phi|^2 - \mu_{\chi}^2 |\chi|^2 + (|\phi|^2, |\chi|^2) \begin{pmatrix} \lambda_{\phi} & \frac{\lambda}{2} \\ \frac{\lambda}{2} & \lambda_{\chi} \end{pmatrix} \begin{pmatrix} |\phi|^2 \\ |\chi|^2 \end{pmatrix}$$



# Scalars

- Standard  $\phi$  complex  $SU(2)_L$  doublet and new  $\chi$  complex singlet:

$$\mathcal{L}_{\phi,\chi} = [D_{\mu}^{(\phi)} \phi]^* D^{(\phi)\mu} \phi + [D_{\mu}^{(\chi)} \chi]^* D^{(\chi)\mu} \chi - V(\phi, \chi)$$

- with scalar potential

$$V(\phi, \chi) = V_0 - \mu_{\phi}^2 |\phi|^2 - \mu_{\chi}^2 |\chi|^2 + (|\phi|^2, |\chi|^2) \begin{pmatrix} \lambda_{\phi} & \frac{\lambda}{2} \\ \frac{\lambda}{2} & \lambda_{\chi} \end{pmatrix} \begin{pmatrix} |\phi|^2 \\ |\chi|^2 \end{pmatrix}$$

- After SSB,  $G \rightarrow SU(3)_c \times U(1)_{QED}$ :

$$\phi = \frac{1}{\sqrt{2}} e^{i\mathbf{T} \cdot \boldsymbol{\xi}(x)/v} \begin{pmatrix} 0 \\ v + h'(x) \end{pmatrix} \quad \& \quad \chi(x) = \frac{1}{\sqrt{2}} e^{i\eta(x)/w} (w + s'(x))$$

---

# Fermion-scalar interactions

---

- Standard Yukawa terms:

$$\mathcal{L}_Y = - \left[ c_D (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} D_R + c_U (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(0)*} \\ -\phi^{(+)*} \end{pmatrix} U_R + c_\ell (\bar{\nu}_\ell, \bar{\ell})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} \ell_R \right] + \text{h.c.}$$

- lead to fermion masses after SSB:

$$\mathcal{L}_Y = - \left( 1 + \frac{h(x)}{v} \right) [\bar{D}_L M_D D_R + \bar{U}_L M_U U_R + \bar{\ell}_L M_\ell \ell_R] + \text{h.c.}$$

# Fermion-scalar interactions

- Standard Yukawa terms:

$$\mathcal{L}_Y = - \left[ c_D (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} D_R + c_U (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(0)*} \\ -\phi^{(+)*} \end{pmatrix} U_R + c_\ell (\bar{\nu}_\ell, \bar{\ell})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} \ell_R \right] + \text{h.c.}$$

- lead to fermion masses after SSB:

$$\mathcal{L}_Y = - \left( 1 + \frac{h(x)}{v} \right) [\bar{D}_L M_D D_R + \bar{U}_L M_U U_R + \bar{\ell}_L M_\ell \ell_R] + \text{h.c.}$$

- Neutrino Yukawa terms ( $z_\chi = -2z_{\nu_R}$ ):

$$\mathcal{L}_Y^\nu = - \sum_{i,j} \left( (c_\nu)_{ij} \bar{L}_{i,L} \cdot \tilde{\phi} \nu_{j,R} + \frac{1}{2} (c_R)_{ij} \overline{\nu_{i,R}^c} \nu_{j,R} \chi \right) + \text{h.c.}$$

# Fermion-scalar interactions

- Standard Yukawa terms:

$$\mathcal{L}_Y = - \left[ c_D (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} D_R + c_U (\bar{U}, \bar{D})_L \begin{pmatrix} \phi^{(0)*} \\ -\phi^{(+)*} \end{pmatrix} U_R + c_\ell (\bar{\nu}_\ell, \bar{\ell})_L \begin{pmatrix} \phi^{(+)} \\ \phi^{(0)} \end{pmatrix} \ell_R \right] + \text{h.c.}$$

- lead to fermion masses after SSB:

$$\mathcal{L}_Y = - \left( 1 + \frac{h(x)}{v} \right) [\bar{D}_L M_D D_R + \bar{U}_L M_U U_R + \bar{\ell}_L M_\ell \ell_R] + \text{h.c.}$$

- Neutrino Yukawa terms ( $z_\chi = -2z_{\nu_R}$ ):

$$\mathcal{L}_Y^\nu = - \sum_{i,j} \left( (c_\nu)_{ij} \bar{L}_{i,L} \cdot \tilde{\phi} \nu_{j,R} + \frac{1}{2} (c_R)_{ij} \overline{\nu_{i,R}^c} \nu_{j,R} \chi \right) + \text{h.c.}$$

---

## After SSB neutrino mass terms appear

---

$$\mathcal{L}_Y^\nu = -\frac{1}{2} \sum_{i,j} \left[ (\overline{\nu}_L, \overline{\nu}_R^c)_i M(h, s)_{ij} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}_j + \text{h.c.} \right]$$

where

$$M(h, s)_{ij} = \begin{pmatrix} 0 & m_D \left(1 + \frac{h}{v}\right) \\ m_D \left(1 + \frac{h}{v}\right) & M_M \left(1 + \frac{s}{w}\right) \end{pmatrix}_{ij}$$

**6x6 symmetric matrix** ( $m_D$  complex,  $M_M$  real)

---

## After SSB neutrino mass terms appear

---

$$\mathcal{L}_Y^\nu = -\frac{1}{2} \sum_{i,j} \left[ (\overline{\nu}_L, \overline{\nu}_R^c)_i M(h, s)_{ij} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}_j + \text{h.c.} \right]$$

where

$$M(h, s)_{ij} = \begin{pmatrix} 0 & m_D \left(1 + \frac{h}{v}\right) \\ m_D \left(1 + \frac{h}{v}\right) & M_M \left(1 + \frac{s}{w}\right) \end{pmatrix}_{ij}$$

**6x6 symmetric matrix** ( $m_D$  complex,  $M_M$  real)

in diagonal: Majorana mass terms (so  $\nu_L$  **massless!**)

---

## After SSB neutrino mass terms appear

---

$$\mathcal{L}_Y^\nu = -\frac{1}{2} \sum_{i,j} \left[ (\overline{\nu}_L, \overline{\nu}_R^c)_i M(h, s)_{ij} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}_j + \text{h.c.} \right]$$

where

$$M(h, s)_{ij} = \begin{pmatrix} 0 & m_D \left(1 + \frac{h}{v}\right) \\ m_D \left(1 + \frac{h}{v}\right) & M_M \left(1 + \frac{s}{w}\right) \end{pmatrix}_{ij}$$

**6x6 symmetric matrix** ( $m_D$  complex,  $M_M$  real)

in diagonal: Majorana mass terms (so  $\nu_L$  **massless!**)

but  $\nu_L$  and  $\nu_R$  have the same q-numbers,  
**can mix, leading to type-I see-saw**

---

## Effective light neutrino masses

---

If  $m_i \ll M_j$ , can integrate out the heavy neutrinos

$$\mathcal{L}_{\text{dim-5}}^\nu = -\frac{1}{2} \sum_i m_{\text{M},i} \left(1 + \frac{h}{v}\right)^2 \left(\overline{\nu'_{i,L}} \nu'_{i,L} + \text{h.c.}\right)$$

where  $m_{\text{M},i} = \frac{m_i^2}{M_i}$  are Majorana masses



---

## Effective light neutrino masses

---

If  $m_i \ll M_j$ , can integrate out the heavy neutrinos

$$\mathcal{L}_{\text{dim-5}}^\nu = -\frac{1}{2} \sum_i m_{\text{M},i} \left(1 + \frac{h}{v}\right)^2 \left(\overline{\nu'_{i,L}} \nu'_{i,L} + \text{h.c.}\right)$$

where  $m_{\text{M},i} = \frac{m_i^2}{M_i}$  are Majorana masses

if  $m_i \sim \text{O}(100\text{keV})$  and  $M_j \sim \text{O}(100\text{GeV})$ , then

$$m_{\text{M},i} \sim \text{O}(0.1\text{eV})$$

---

# Mixing in the neutral gauge sector

---

$$\begin{pmatrix} W_{\mu}^3 \\ B'_{\mu} \\ Z'_{\mu} \end{pmatrix} = \underline{M}(\sin \theta_W, \sin \theta_T) \begin{pmatrix} Z_{\mu}^0 \\ T_{\mu} \\ A_{\mu} \end{pmatrix}$$

- **QED** current remains **unchanged**:

$$\mathcal{L}_{\text{QED}} = -e A_{\mu} J_{\text{em}}^{\mu}, \quad J_{\text{em}}^{\mu} = \sum_{f=1}^3 \sum_{j=1}^3 e_j \left( \bar{\psi}_{q,j}^f(x) \gamma^{\mu} \psi_{q,j}^f(x) + \bar{\psi}_{l,j}^f(x) \gamma^{\mu} \psi_{l,j}^f(x) \right)$$

---

# Neutral current interactions

---

$$\mathcal{L}_{Z^0} = -eZ_\mu^0 \left( \cos \theta_T J_{Z^0}^\mu + \sin \theta_T J_T^\mu \right) = -eZ_\mu^0 J_{Z^0}^\mu + \mathcal{O}(\theta_T)$$

$$\mathcal{L}_T = -eT_\mu \left( \sin \theta_T J_{Z^0}^\mu + \cos \theta_T J_T^\mu \right) = -eT_\mu J_T^\mu + \mathcal{O}(\theta_T)$$

---

# Neutral current interactions

---

$$\mathcal{L}_{Z^0} = -e Z_\mu^0 \left( \cos \theta_T J_{Z^0}^\mu + \sin \theta_T J_T^\mu \right) = -e Z_\mu^0 J_{Z^0}^\mu + \mathcal{O}(\theta_T)$$

$$\mathcal{L}_T = -e T_\mu \left( \sin \theta_T J_{Z^0}^\mu + \cos \theta_T J_T^\mu \right) = -e T_\mu J_T^\mu + \mathcal{O}(\theta_T)$$

- **current with  $Z^0$  remains unchanged:**

$$J_{Z^0}^\mu = \sum_{f=1}^3 \sum_{j=1}^3 \frac{T_3 - \sin^2 \theta_W e_j}{\sin \theta_W \cos \theta_W} \left( \bar{\psi}_{q,j}^f(x) \gamma^\mu \psi_{q,j}^f(x) + \bar{\psi}_{l,j}^f(x) \gamma^\mu \psi_{l,j}^f(x) \right)$$

# Neutral current interactions

$$\mathcal{L}_{Z^0} = -eZ_\mu^0 \left( \cos \theta_T J_{Z^0}^\mu + \sin \theta_T J_T^\mu \right) = -eZ_\mu^0 J_{Z^0}^\mu + \mathcal{O}(\theta_T)$$

$$\mathcal{L}_T = -eT_\mu \left( \sin \theta_T J_{Z^0}^\mu + \cos \theta_T J_T^\mu \right) = -eT_\mu J_T^\mu + \mathcal{O}(\theta_T)$$

- **current with  $Z^0$  remains unchanged:**

$$J_{Z^0}^\mu = \sum_{f=1}^3 \sum_{j=1}^3 \frac{T_3 - \sin^2 \theta_W e_j}{\sin \theta_W \cos \theta_W} \left( \bar{\psi}_{q,j}^f(x) \gamma^\mu \psi_{q,j}^f(x) + \bar{\psi}_{l,j}^f(x) \gamma^\mu \psi_{l,j}^f(x) \right)$$

- **but mixes with new current of new couplings:**

$$J_T^\mu = \sum_{f=1}^3 \sum_{j=1}^3 \frac{\gamma'_Z r_j + \gamma'_{ZY} y_j}{\sin \theta_W} \left( \bar{\psi}_{q,j}^f(x) \gamma^\mu \psi_{q,j}^f(x) + \bar{\psi}_{l,j}^f(x) \gamma^\mu \psi_{l,j}^f(x) \right)$$

---

# Possible consequences with 5 new parameters

---

- The **massive T vector boson** is a natural **candidate for WIMP** dark matter if it is sufficiently stable (mass of  $\sim 1$  MeV: super weak new force).

---

# Possible consequences with 5 new parameters

---

- The **massive T vector boson** is a natural **candidate for WIMP** dark matter if it is sufficiently stable (mass of  $\sim 1$  MeV: super weak new force).
- **Majorana neutrino mass terms** are generated **by the SSB of the scalar fields**, providing the origin of neutrino masses and oscillations.

---

# Possible consequences with 5 new parameters

---

- The **massive T vector boson** is a natural **candidate for WIMP** dark matter if it is sufficiently stable (mass of  $\sim 1$  MeV: super weak new force).
- **Majorana neutrino mass terms** are generated **by the SSB of the scalar fields**, providing the origin of neutrino masses and oscillations.
- **Diagonalization** of neutrino mass terms **leads to the PMNS matrix**, which in turn can be the source of lepto-baryogenesis.



---

# Possible consequences with 5 new parameters

---

- The **massive T vector boson** is a natural **candidate for WIMP** dark matter if it is sufficiently stable (mass of  $\sim 1$  MeV: super weak new force).
- **Majorana neutrino mass terms** are generated **by the SSB of the scalar fields**, providing the origin of neutrino masses and oscillations.
- **Diagonalization** of neutrino mass terms **leads to the PMNS matrix**, which in turn can be the source of lepto-baryogenesis.
- The **vacuum of the  $\chi$  scalar is charged** ( $z_j = -1$ ) that may be a **source of accelerated expansion** of the universe as seen now.

---

# Possible consequences with 5 new parameters

---

- The **massive T vector boson** is a natural **candidate for WIMP** dark matter if it is sufficiently stable (mass of  $\sim 1$  MeV: super weak new force).
- **Majorana neutrino mass terms** are generated **by the SSB of the scalar fields**, providing the origin of neutrino masses and oscillations.
- **Diagonalization** of neutrino mass terms **leads to the PMNS matrix**, which in turn can be the source of lepto-baryogenesis.
- The **vacuum of the  $\chi$  scalar is charged** ( $z_j = -1$ ) that may be a **source of accelerated expansion** of the universe as seen now.
- The second scalar together with the established BEH field may be the source of **hybrid inflation**.

---

## Credibility requirement

---

Is there any region of the parameter space of the model that is not excluded by experimental results, both established in standard model phenomenology and elsewhere?

---

## Credibility requirement

---

Is there any region of the parameter space of the model that is not excluded by experimental results, both established in standard model phenomenology and elsewhere?

Answer is not immediate, extensive studies are needed

# Contribution of the new gauge boson to the anomalous magnetic moment of the muon

using the new neutral currents:

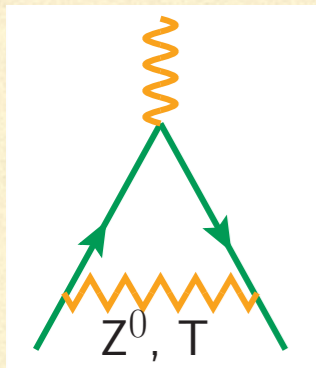
$$\Delta a_\mu = a_\mu^{(\text{T+SM})} - a_\mu^{(\text{SM})} = a_\mu^{(Z^0)}(h_f, \theta_T) - a_\mu^{(Z^0)}(0, 0) + a_\mu^{(T^0)}(h_f, \theta_T)$$

where

$$a_\mu^{(X)}(h_f, \theta_T) = \frac{G_F m_\mu^2}{6\sqrt{2}\pi^2} \left[ 3C_X^+ C_X^- - (C_X^+)^2 - (C_X^-)^2 \right]$$

with chiral couplings:

$$V_\alpha \bar{f}_i f_j: -ie\gamma_\alpha (C^- P_- + C^+ P_+)$$

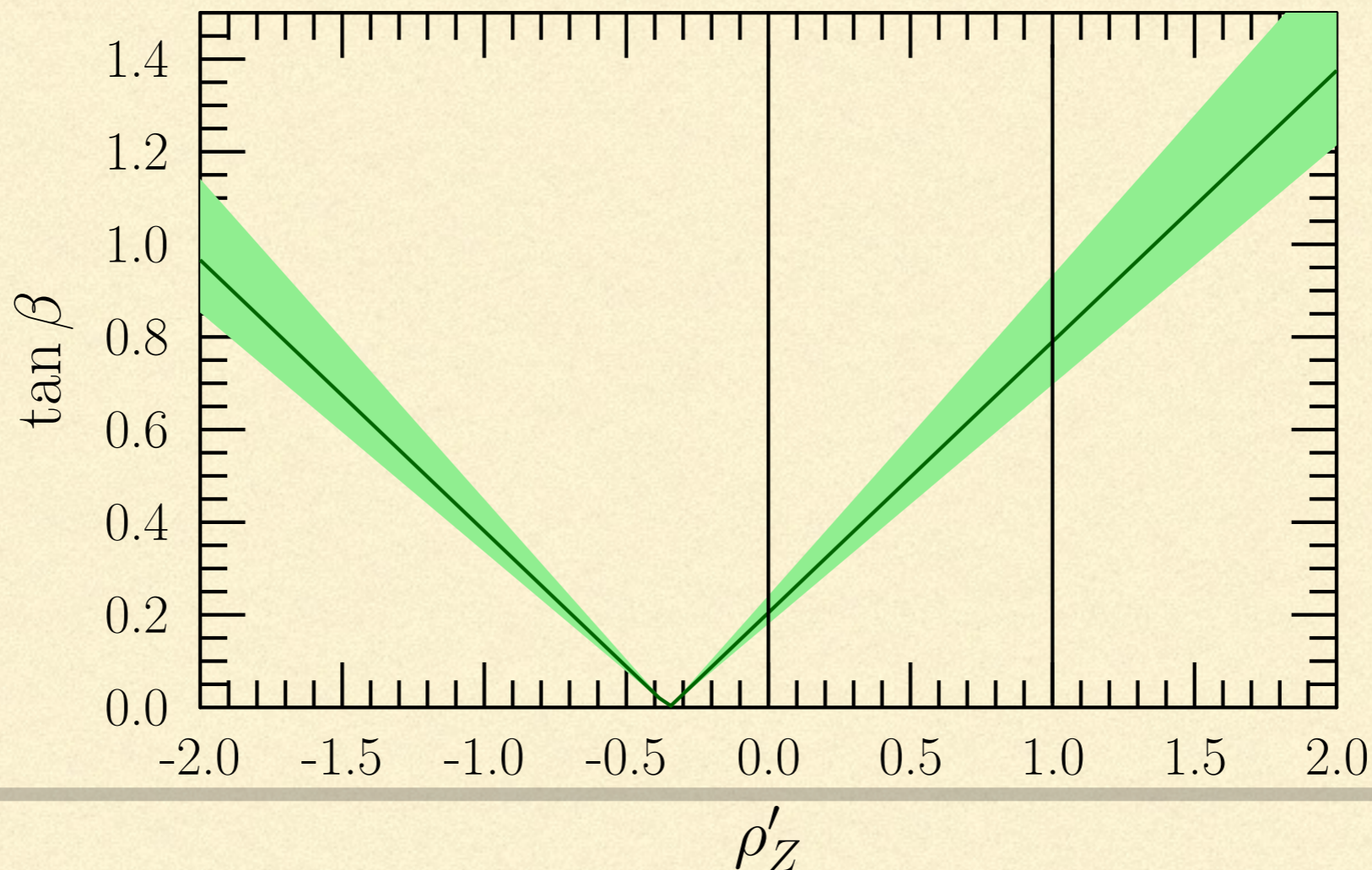


$V \bar{f}_i f_j$	$C^+$	$C^-$
$\gamma \bar{f}_i f_j$	$e_f \delta_{ij}$	$e_f \delta_{ij}$
$Z \bar{f}_i f_j$	$(g_f^+ \cos \theta_T - h_f^+ \sin \theta_T) \delta_{ij}$	$(g_f^- \cos \theta_T - h_f^- \sin \theta_T) \delta_{ij}$
$T \bar{f}_i f_j$	$(g_f^+ \sin \theta_T + h_f^+ \cos \theta_T) \delta_{ij}$	$(g_f^- \sin \theta_T + h_f^- \cos \theta_T) \delta_{ij}$

# Contribution of the new gauge boson to $a_\mu$

$$a_\mu^{(\text{T+SM})} - a_\mu^{(\text{SM})} = \frac{G_F m_\mu^2}{6\sqrt{2}\pi^2} \left( \frac{(1 + \rho'_Z) \cos^2 \theta_W - \frac{1}{2}}{\tan \beta} + \mathcal{O}(\theta_T, \gamma'_Z) \right)^2 \quad \rho'_Z = \frac{\gamma'_{ZY}}{\gamma'_Z}$$

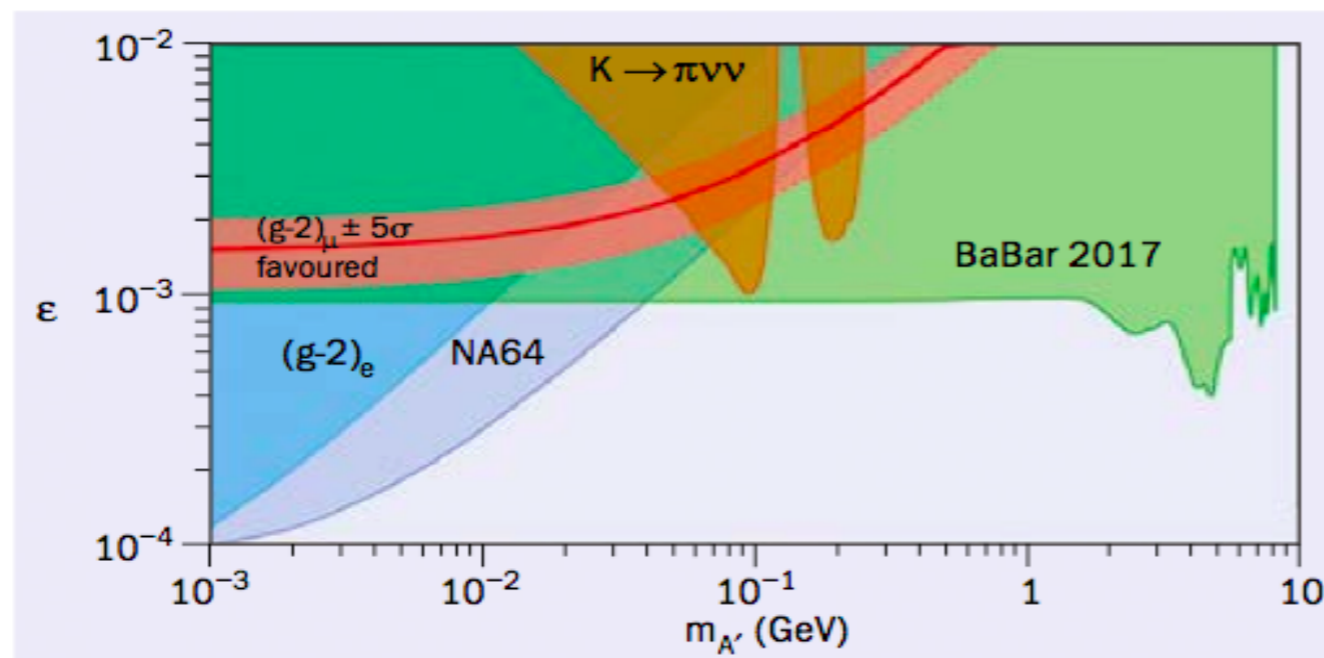
experimentally:  $a_\mu^{(\text{exp})} - a_\mu^{(\text{SM})} = 268(76) \cdot 10^{-11}$



# A' explanation of the anomaly ruled out?

CERN Courier April 2017

## News



*Regions of the dark-photon parameter space (mixing strength versus mass) excluded by BaBar (green) compared with the previous constraints. The new analysis rules out dark-photon coupling as the explanation for the muon (g-2) anomaly and places stringent constraints on dark-sector models.*

of Caltech, who has worked on dark-photon models. “In contrast to massless dark photons, which are analogous to ordinary photons, this experiment constrains a slightly different idea of dark force-carrying particles that are associated with a broken symmetry, which therefore get a mass and

then can decay. They are more like ‘dark Z bosons’ than dark photons.”

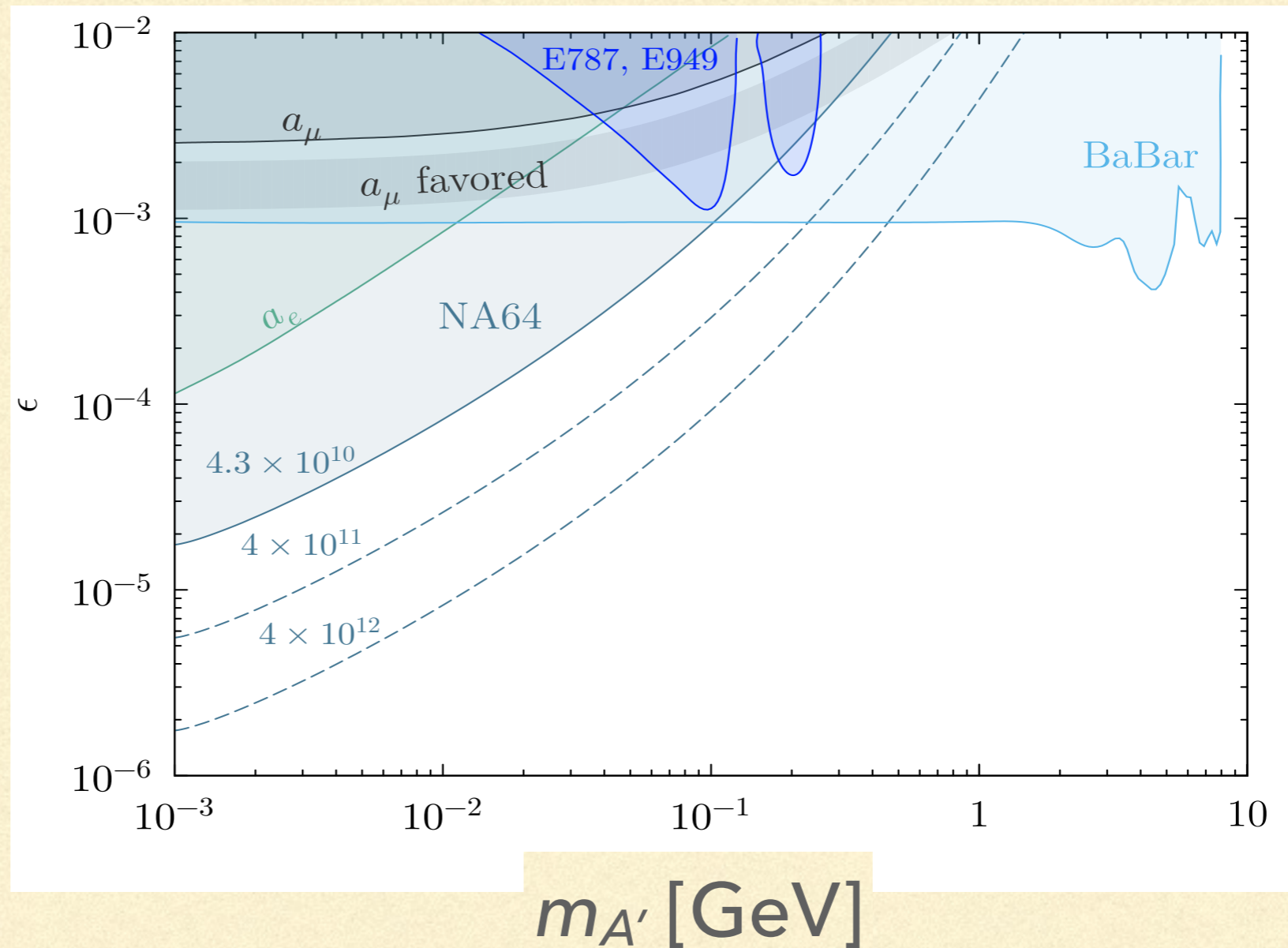
### ● Further reading

BaBar Collaboration 2017 arXiv:1702.03327.  
NA64 Collaboration 2017 *Phys. Rev. Lett.* **118** 011802.

# Searches for invisibly decaying, light, neutral gauge bosons

best exclusion limits by BaBar and NA64

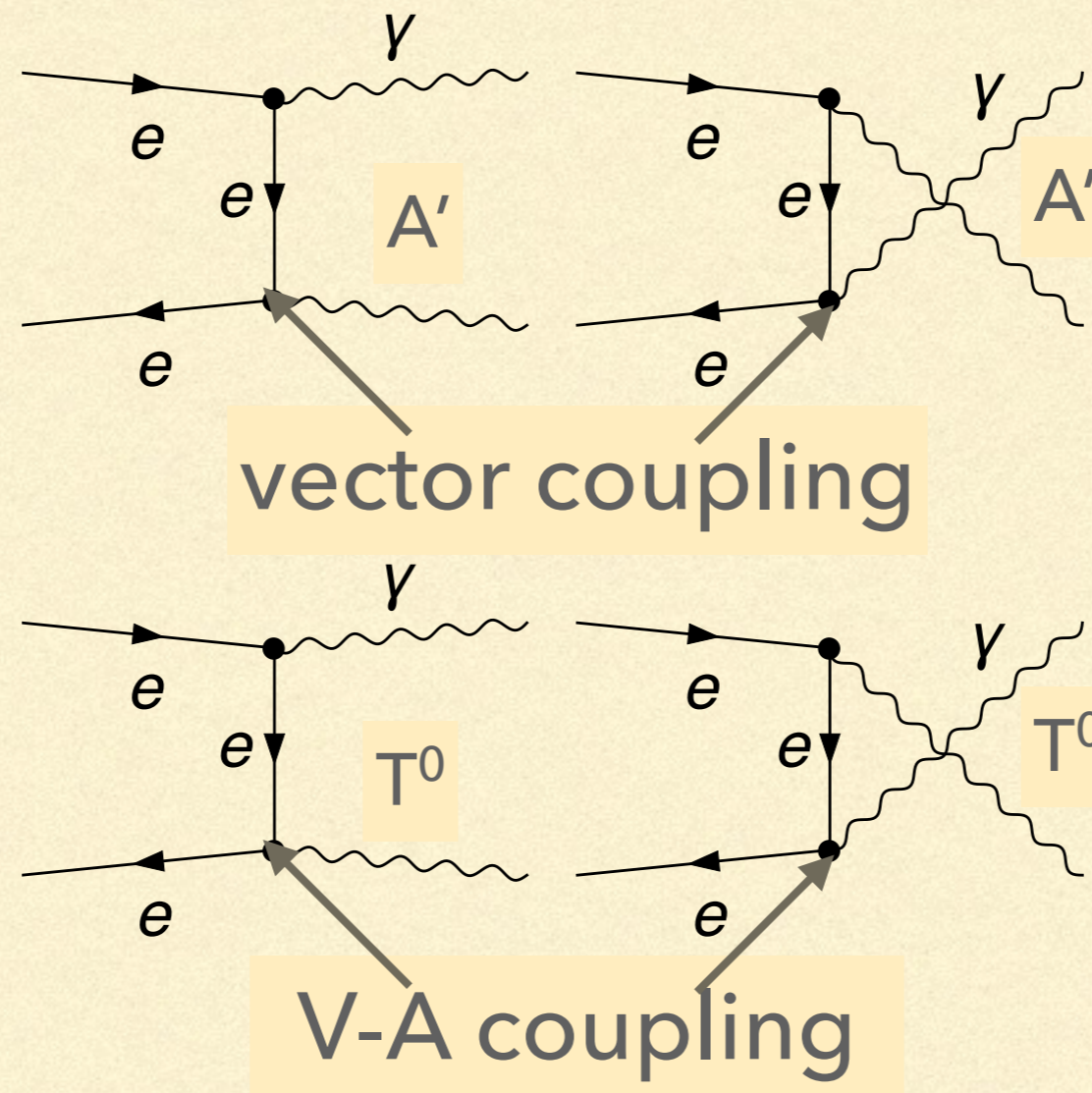
kinetic mixing parameter  $\epsilon$  (with vector coupling of new boson to fermions)





# Effective kinetic mixing for BaBar production

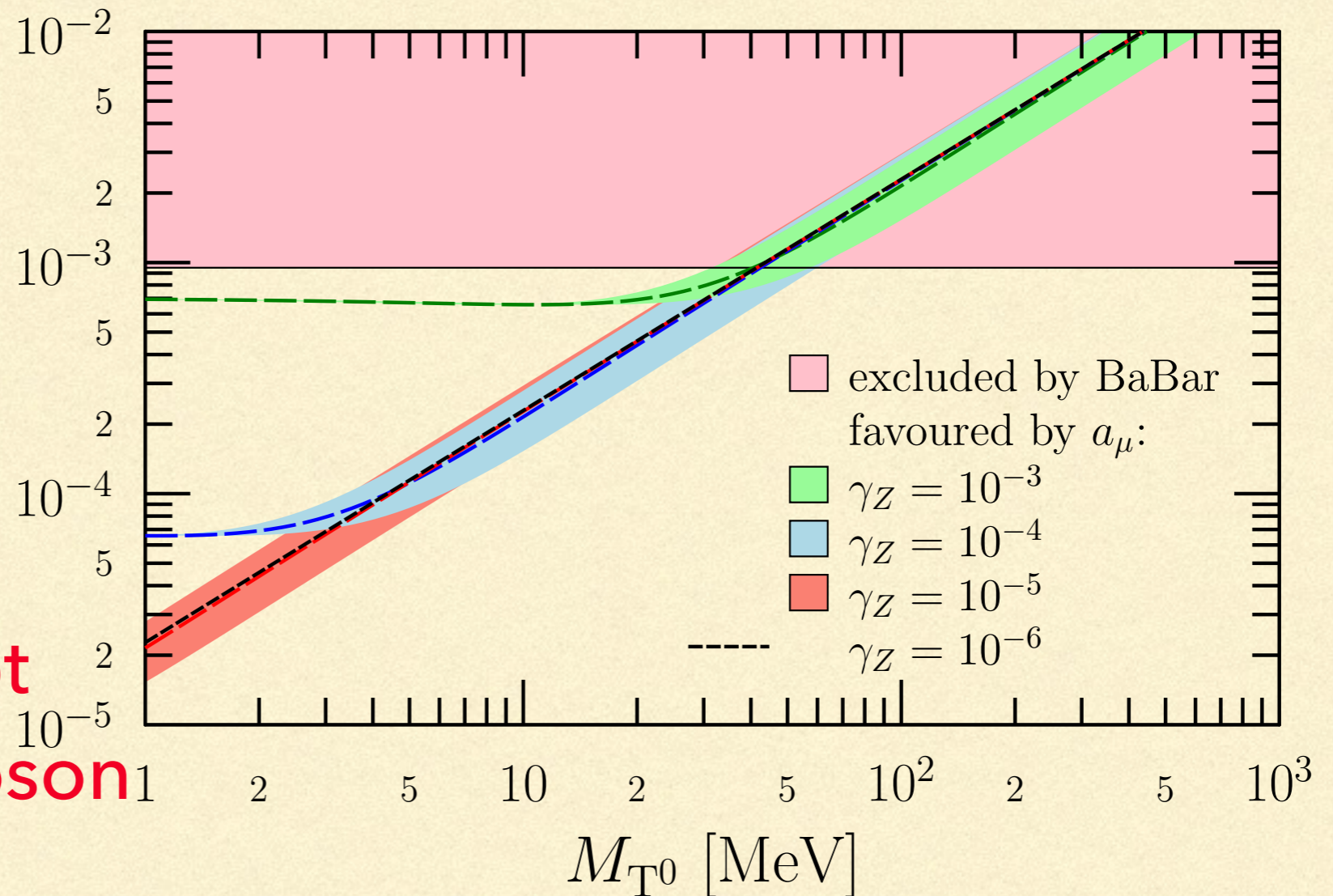
Production channel:



# Favoured region by $a_\mu$ vs. BaBar exclusion limit

$$\epsilon_{\text{eff}} = \sqrt{\frac{\sigma(e^+e^- \rightarrow \gamma T^0)}{\sigma(e^+e^- \rightarrow \gamma A')/\epsilon^2}} = \sqrt{\left(v_e^{(T)}\right)^2 + \left(a_e^{(T)}\right)^2} = \frac{\gamma'_Z}{2 \sin \theta_W} \sqrt{\frac{5}{2} \rho'_Z{}^2 + \rho'_Z + \frac{1}{2}}$$

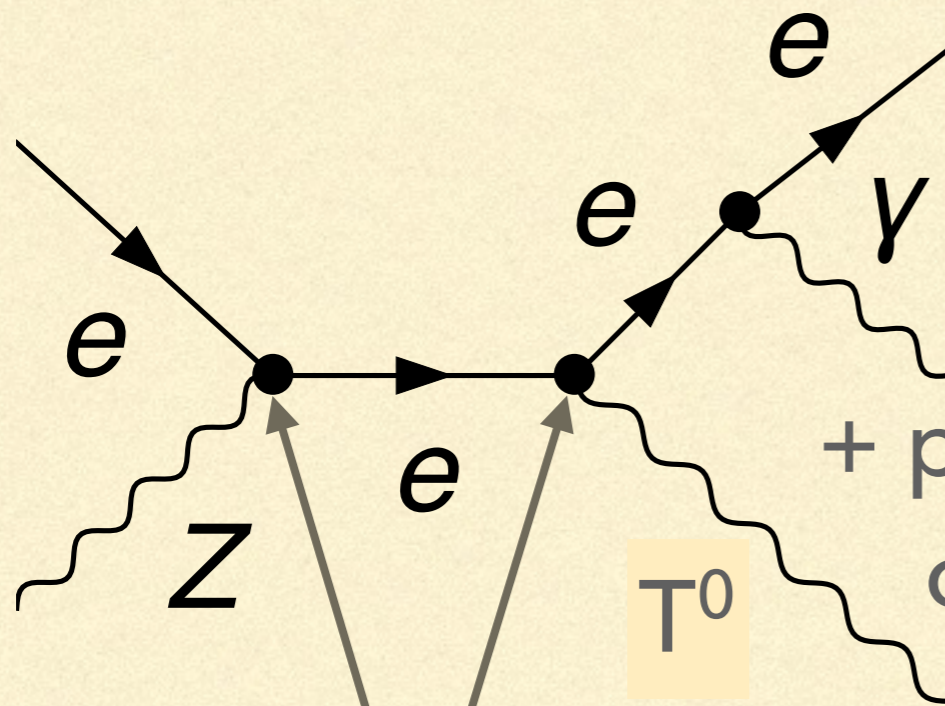
effective  
kinetic mixing  
parameter  $\rightarrow \epsilon_{\text{eff}}$



**BaBar does not  
exclude the  $T^0$  boson**

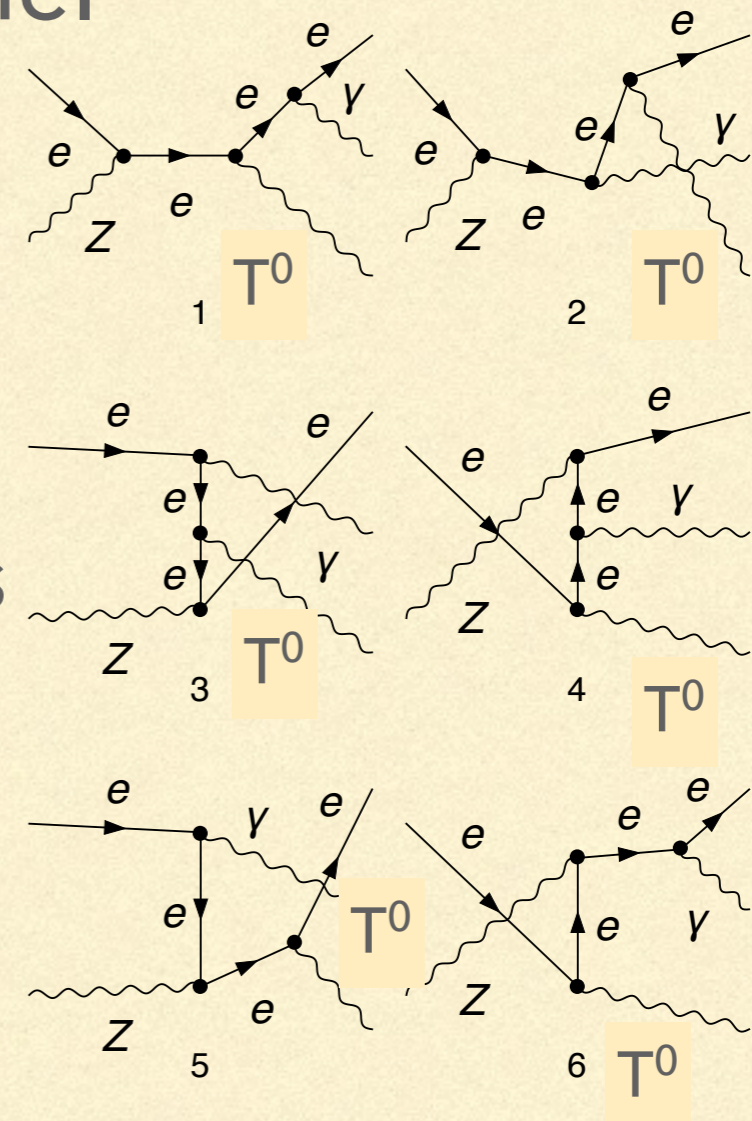
# Searches for invisibly decaying, light, neutral gauge bosons, NA64 limit

## Production channel



V-A couplings

+ permutations of bosons:



Expect result at  
the EPSHEP 2019

---

# Conclusions

---

- Established observations do not suggest a rich BSM physics

---

# Conclusions

---

- Established observations do not suggest a rich BSM physics
- $U(1)_Z$  extension has the potential of explaining all known results

---

# Conclusions

---

- Established observations do not suggest a rich BSM physics
- $U(1)_Z$  extension has the potential of explaining all known results
- Anomaly cancellation and neutrino mass generation mechanism are used to fix the Z-charges up to reasonable assumptions

---

# Conclusions

---

- Established observations do not suggest a rich BSM physics
- $U(1)_Z$  extension has the potential of explaining all known results
- Anomaly cancellation and neutrino mass generation mechanism are used to fix the Z-charges up to reasonable assumptions
- Parameter space can be constrained from existing experimental results

---

UV behavior

with Zoltán Péli

---