

Effective theories of the proton: neutrinos, atoms and dark matter

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UK and Fermilab

PIKIO meeting
University of Kentucky

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many thanks to Susan Gardner, Chris Crawford, Wolfgang Korsch and Michael Kovash for bringing us together !

thanks to many collaborators and colleagues, especially: J. Arrington, M. Betancourt, P. Kammel, G. Lee, W. Marciano, A. Meyer, G. Paz, R. Gran, A. Sirlin, M. Solon

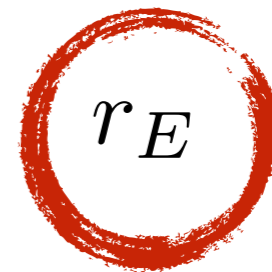
Overview

- Motivation
- Example 1: Muon Capture and neutrinos
- Example 2: universality in WIMP-nucleon scattering
- Example 3: Proton radius puzzle

Focus on 3 numbers


$$r_A$$

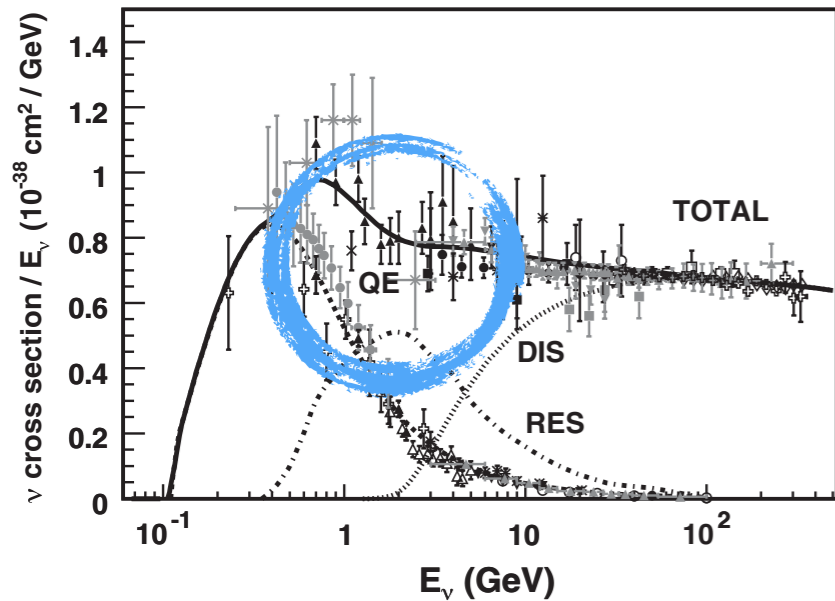

$$\sigma_{\chi N}$$


$$r_E$$

- r_A : *determines signal cross section for LBNE*
- $\sigma_{\chi N}$: *universal cross section for next-gen. DD*
- r_E : *5 σ shift in Rydberg (or something even more interesting)*

- neutrinos:

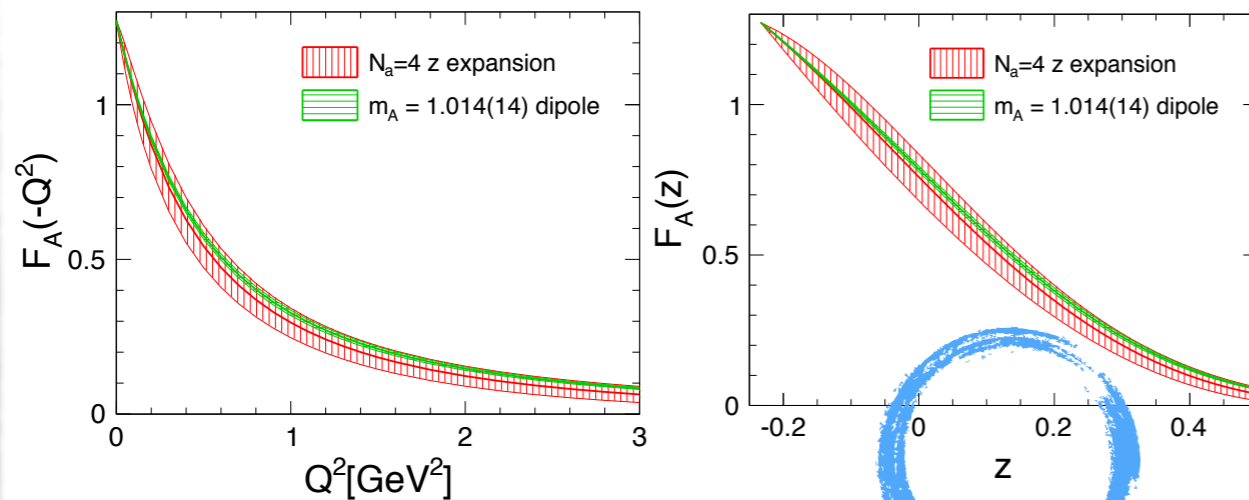
must confront large uncertainty in signal process of ν_e appearance at long baseline neutrino experiment



Quasielastic dominance

$$\langle p(p') | J_W^{+\mu} | n(p) \rangle \propto \bar{u}^{(p)}(p') \left\{ \gamma^\mu F_1(q^2) + \frac{i}{2m_N} \sigma^{\mu\nu} q_\nu F_2(q^2) + \gamma^\mu \gamma_5 F_A(q^2) + \frac{1}{m_N} q^\mu \gamma_5 F_P(q^2) \right\} u^{(n)}(p)$$

Unknown: axial-vector form nucleon form factor

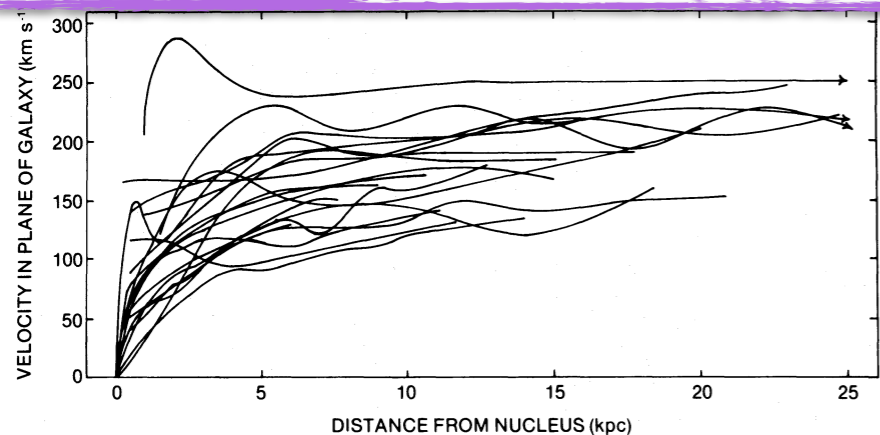


F_A linear in "z" variable

$$\rightarrow r_A^2 \equiv 6 \frac{d}{dq^2} F_A(q^2) \Big|_{q^2=0}$$

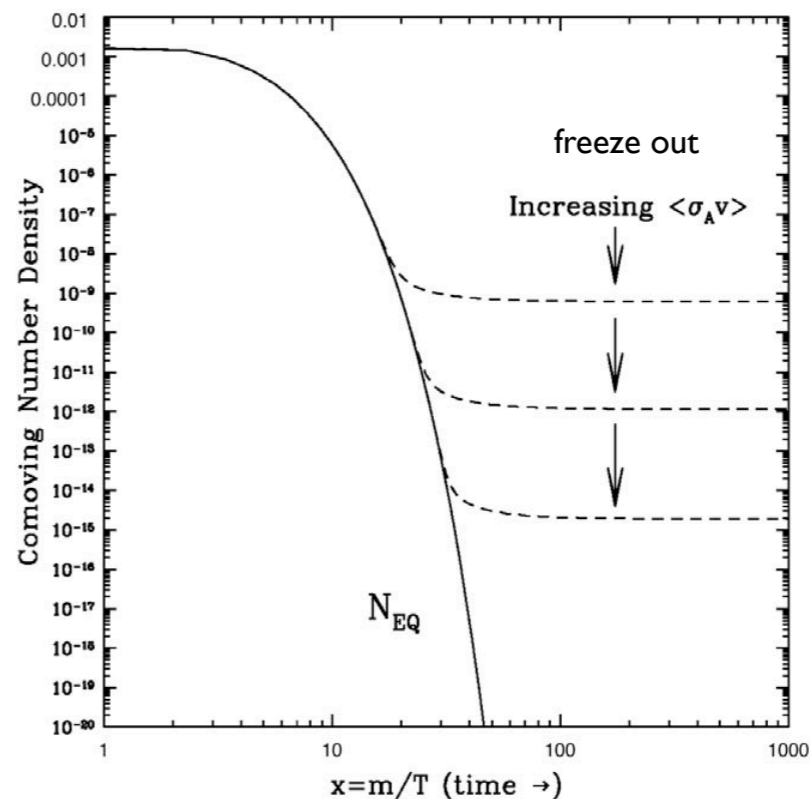
#1 nucleon axial radius

- dark matter:



It exists

+



It might be a thermal relic Weakly Interacting Massive Particle (WIMP)

ATLAS SUSY Searches - 95% CL Lower Limits

Model	$\epsilon, \lambda, T, \gamma$	Jets	L_{T}^{min} [L/10 ¹⁰]	Mass limit	Reference	
MSSM/CMSSM	0	2-6 jets	Yes	20.3	1.1 TeV	1405.315
$\tilde{g}, \tilde{q} \rightarrow \tilde{q} \tilde{g}$	0	2-6 jets	Yes	20.3	1.1 TeV	1405.315
$\tilde{g}, \tilde{g} \rightarrow \tilde{g} \tilde{g}$ (compressed)	1, μ	0-1 jet	Yes	19.9	1.1 TeV	1405.315
$\tilde{g}, \tilde{g} \rightarrow \tilde{g} \tilde{g}$	0	2-6 jets	Yes	20.3	1.2 TeV	1405.315
$\tilde{g}, \tilde{g} \rightarrow \tilde{g} \tilde{g} \tilde{g}$	1, μ	3-6 jets	Yes	20	1.2 TeV	1405.315
$\tilde{g}, \tilde{g} \rightarrow \tilde{g} \tilde{g} \tilde{g} \tilde{g}$	2, μ	3-3 jets	-	20	1.2 TeV	1405.315
GMSB (NLSB)	$1.2\gamma + 0.1\tilde{g}$	0-2 jets	Yes	20.3	1.2 TeV	1407.003
GGM (non-NLSB)	2, γ	-	Yes	20.3	1.2 TeV	1407.003
GGM (non-NLSB)	1, $\mu + \gamma$	-	Yes	1.8	819 GeV	ATLAS CONF 2014-001
GGM (Higgsportal NLSB)	γ	1-4	Yes	1.8	819 GeV	ATLAS CONF 2012-144
GGM (Higgsportal NLSB)	2, μ (Z)	0-3 jets	Yes	1.8	819 GeV	ATLAS CONF 2012-152
Gravitino LSP	0	mono-jet	Yes	20.3	945 GeV	1302.0518

→ $\sigma_{\chi N}$

#2 Universal cross section for heavy WIMP-nucleon scattering

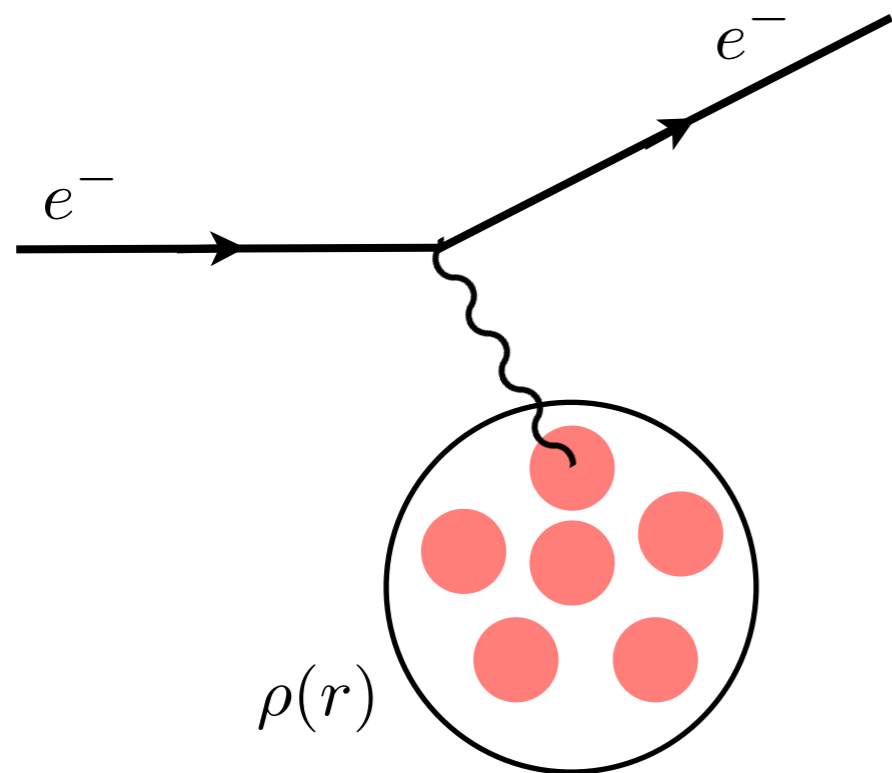
WIMP paradigm pushed to larger masses ($\gg m_W$)

- precision spectroscopy:

Most mundane resolution of the proton radius puzzle:

- change fundamental Rydberg constant by $\sim 5\sigma$
- revise inferences from several decades of both electron scattering and hydrogen spectroscopy

And the neutrino problem is harder (flux, nuclear effects, statistics).
So we need to get this right.

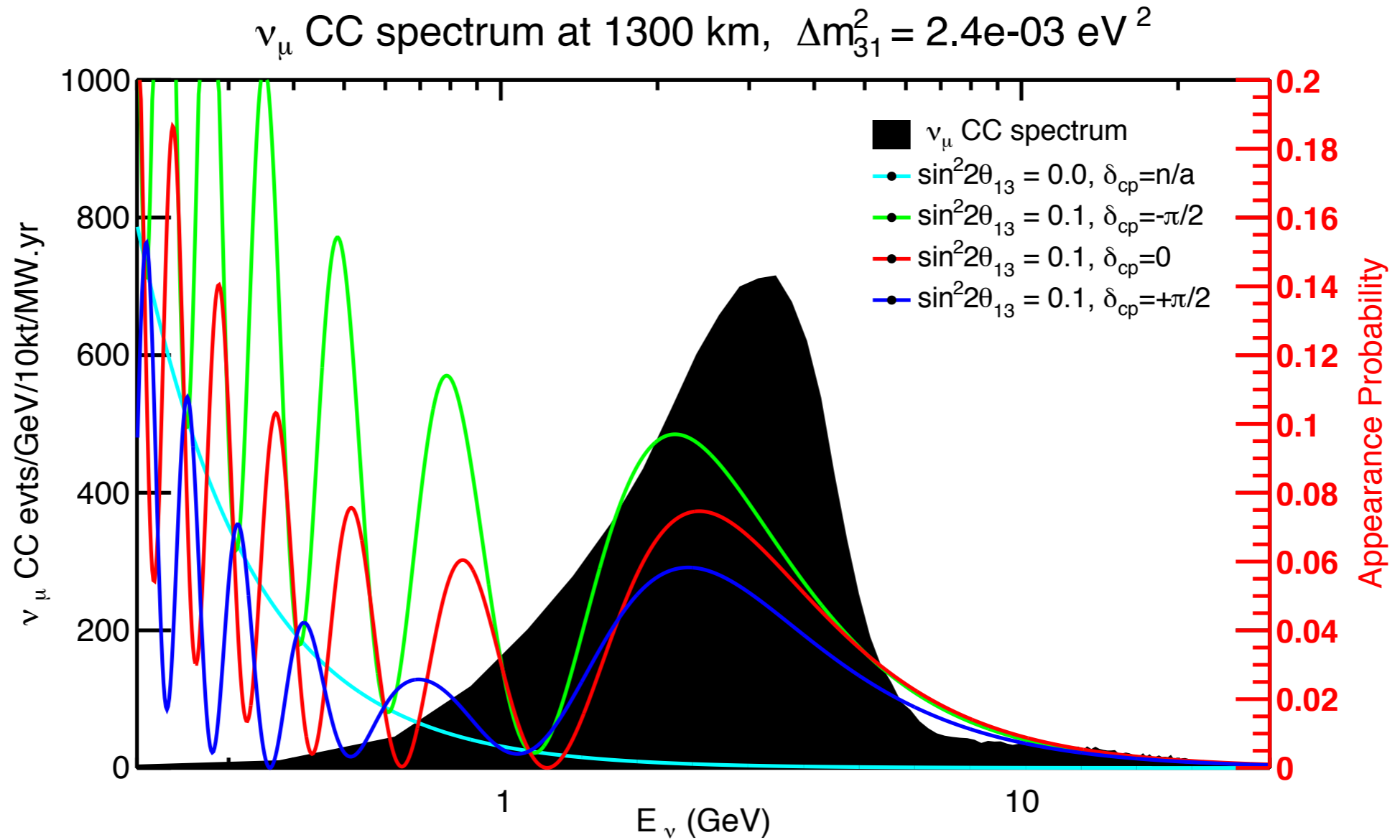


$$E_n \sim \frac{\text{Rydberg}}{n^2} \left(1 - \frac{r_E^2}{n^2} \right)$$

#3 proton charge radius

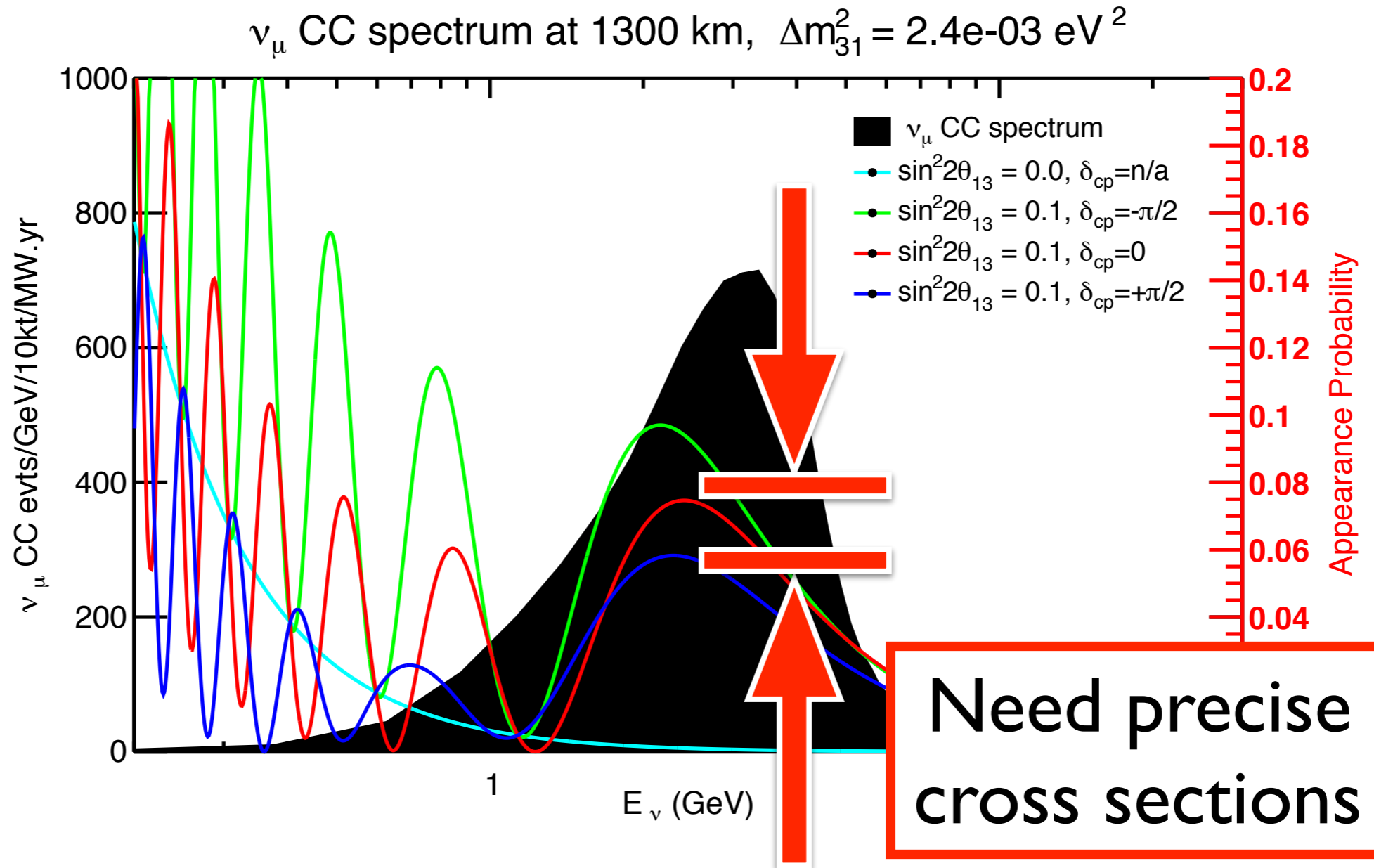
I. muon capture and neutrino cross sections

Recent neutrino discoveries (neutrino mixing) have set the stage for yet further discoveries (leptonic CP violation, ...)



LBNE, 1307.7335

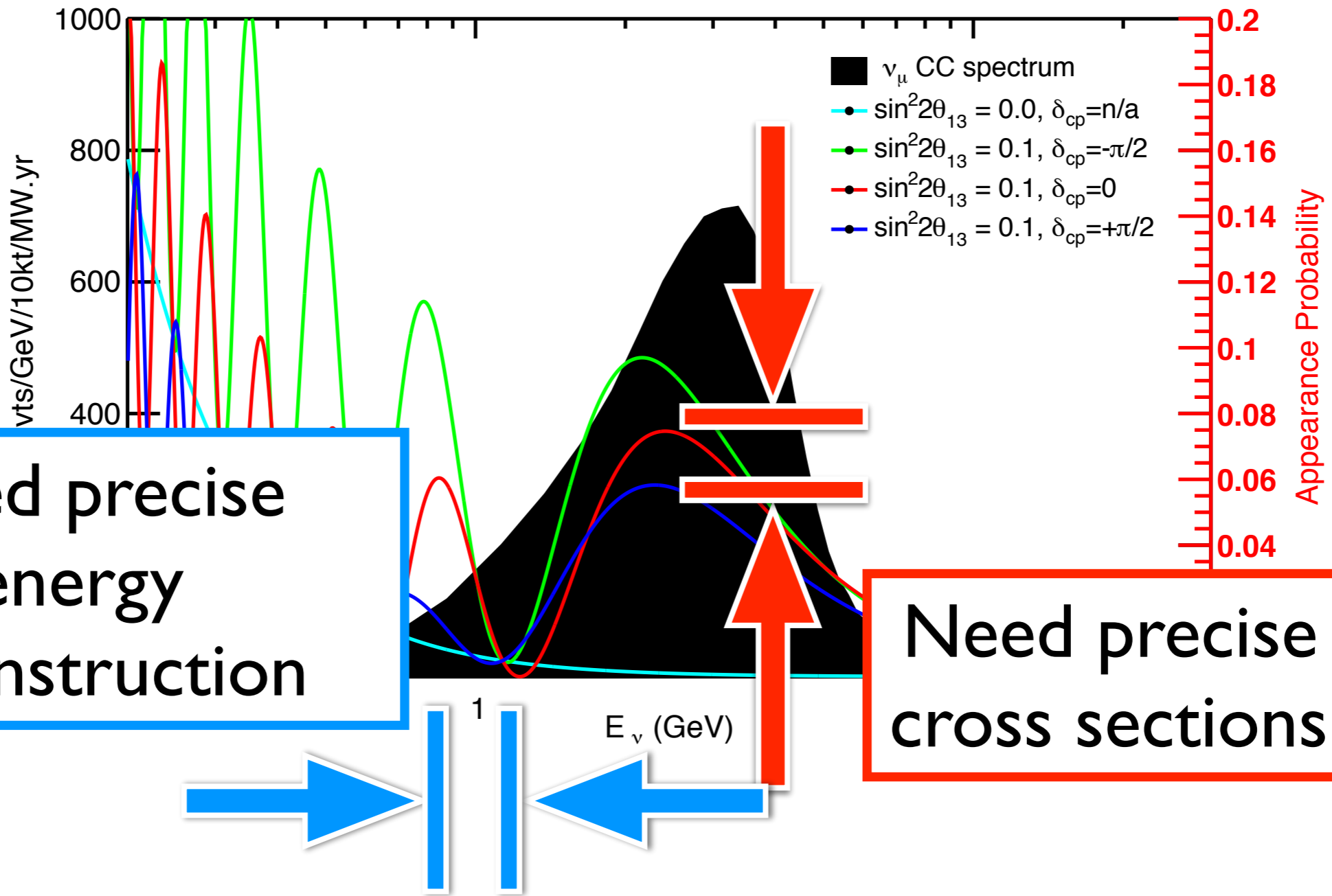
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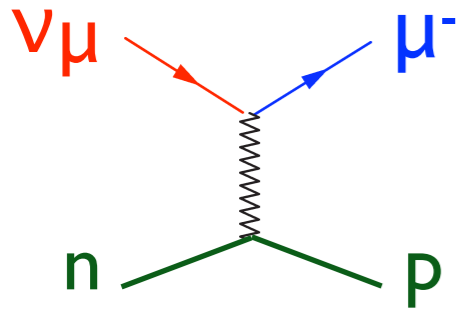
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ν_μ CC spectrum at 1300 km, $\Delta m_{31}^2 = 2.4e-03 \text{ eV}^2$



LBNE, 1307.7335

Start with the basic process



$$\sigma(\nu n \rightarrow \mu p) = |\cdots \cdot F_A(q^2) \cdots|^2$$

poorly known axial-vector form factor

A common ansatz for F_A has been employed for the last ~40 years:

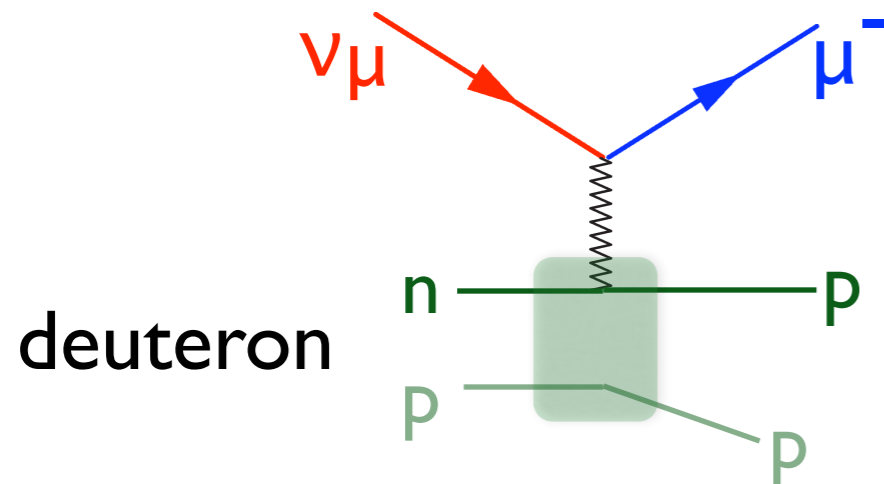
$$F_A^{\text{dipole}}(q^2) = F_A(0) \left(1 - \frac{q^2}{m_A^2}\right)^{-2}$$

Inconsistent with QCD.

Typically quoted uncertainties are (too) small (e.g. compared to proton charge form factor!)

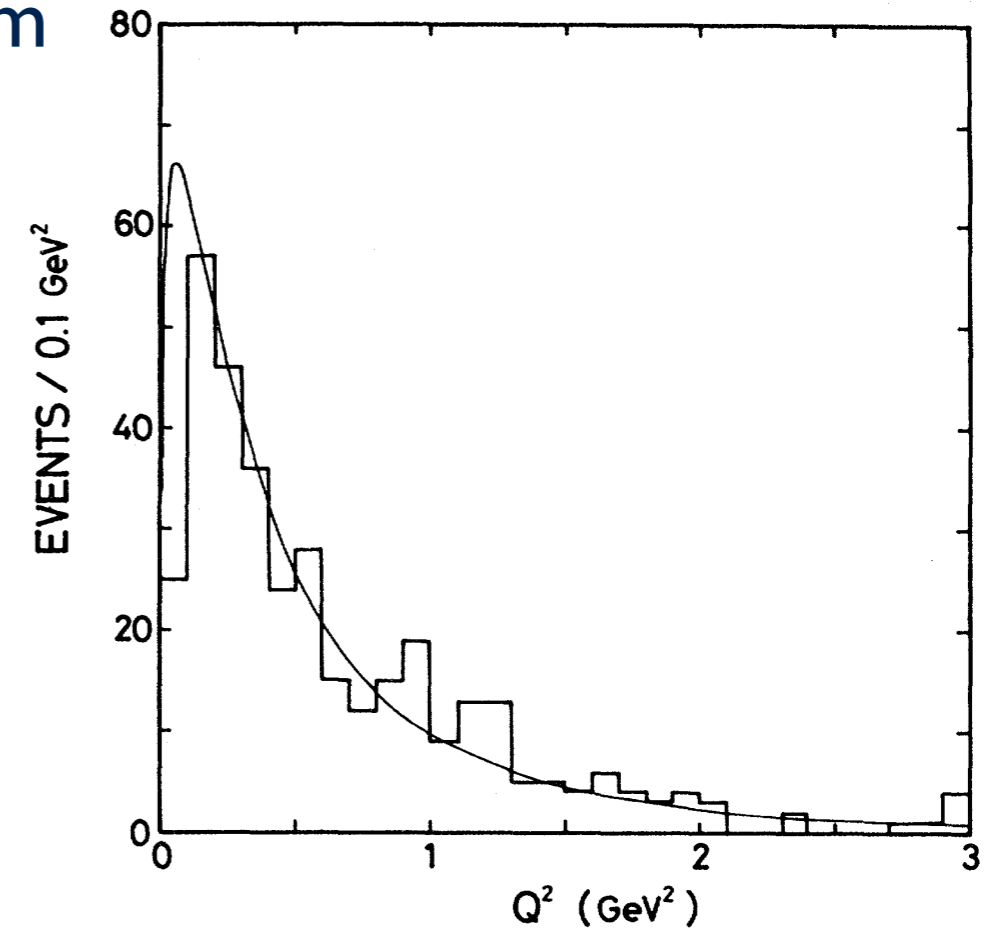
$$\frac{1}{F_A(0)} \left. \frac{dF_A}{dq^2} \right|_{q^2=0} \equiv \frac{1}{6} r_A^2 \quad r_A = 0.674(9) \text{ fm}$$

Best source of almost-free neutrons: deuterium



Deuterium bubble chamber data

- small(-ish) nuclear effects
- small(-ish) experimental uncertainties
- small statistics, ~ 3000 events in world data



Fermilab 15-foot deuterium bubble chamber, PRD 28, 436 (1983)

also:

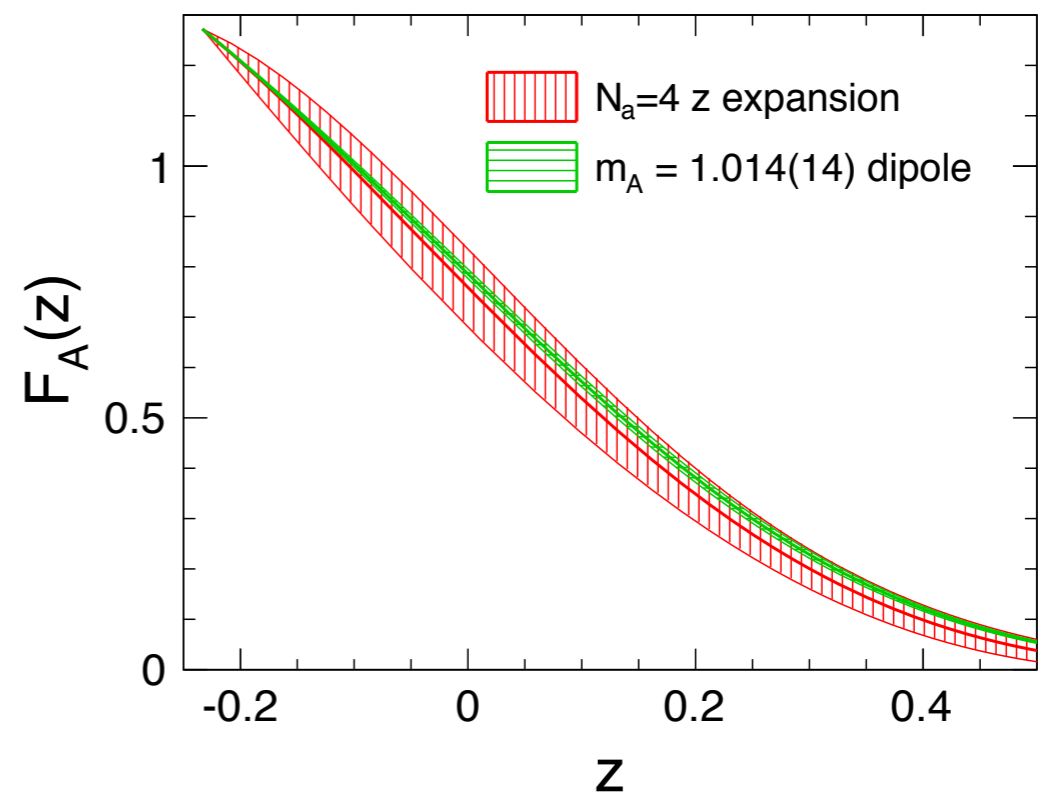
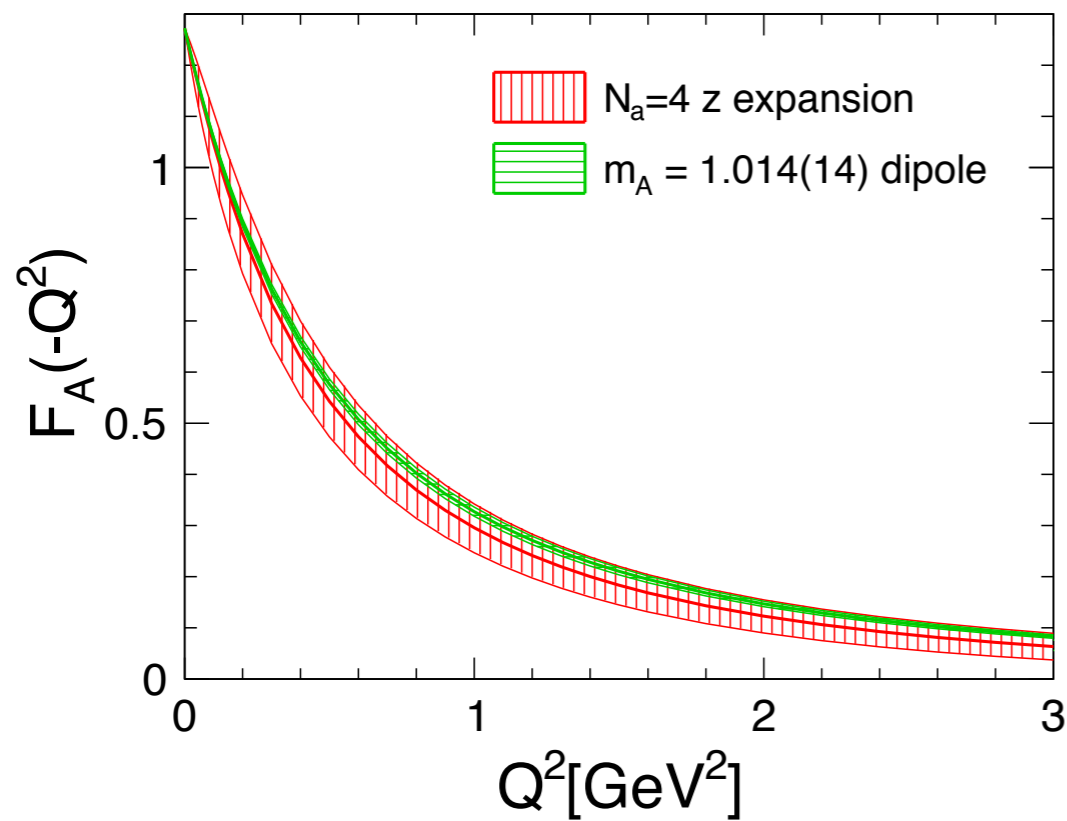
ANL 12-foot deuterium bubble chamber, PRD 26, 537 (1982)

BNL 7-foot deuterium bubble chamber, PRD23, 2499 (1981)

- F_A with complete error budget:

$$[a_1, a_2, a_3, a_4] = [2.30(13), -0.6(1.0), -3.8(2.5), 2.3(2.7)]$$

$$C_{ij} = \begin{pmatrix} 1 & 0.350 & -0.678 & 0.611 \\ 0.350 & 1 & -0.898 & 0.367 \\ -0.678 & -0.898 & 1 & -0.685 \\ 0.611 & 0.367 & -0.685 & 1 \end{pmatrix}$$



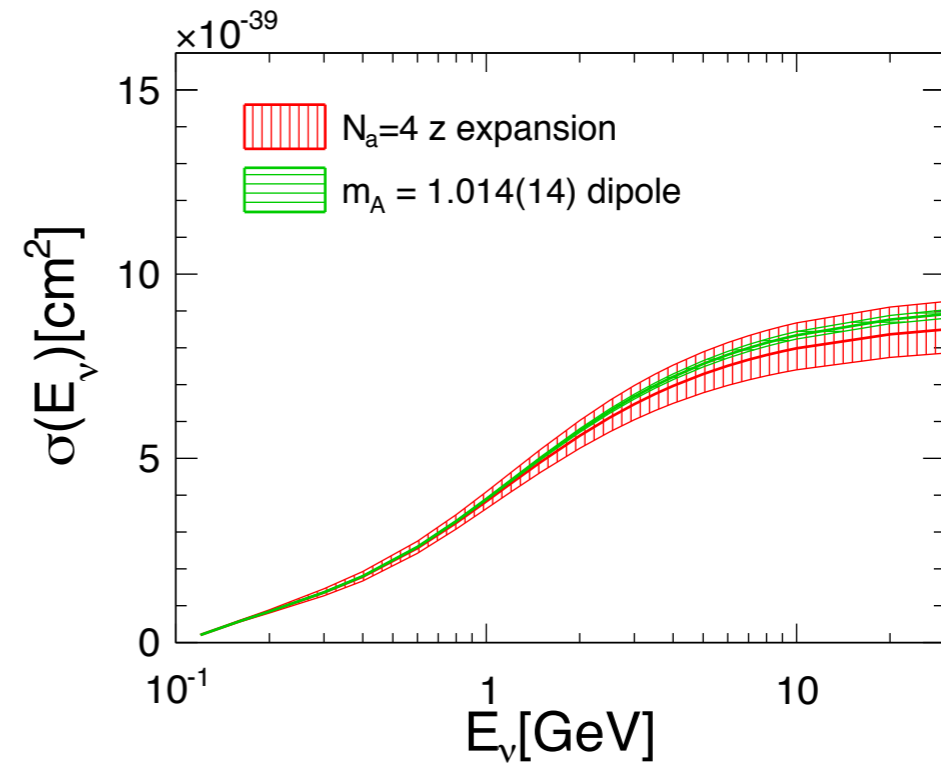
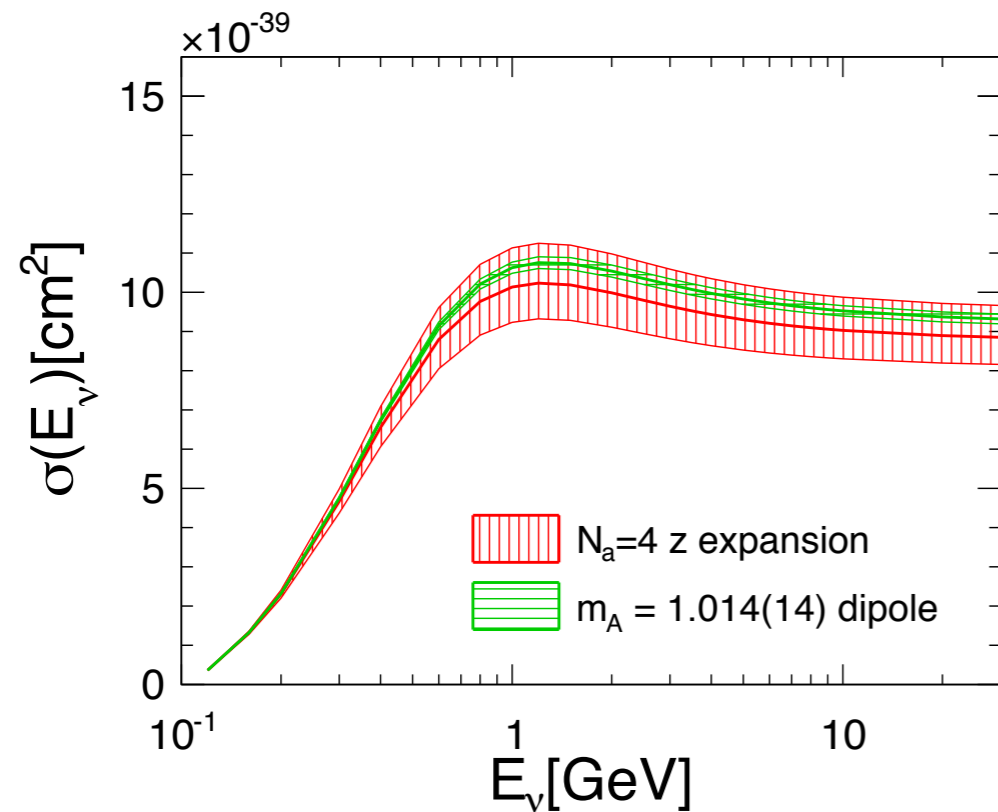
Derived observables: 1) axial radius

$$\frac{1}{F_A(0)} \left. \frac{dF_A}{dq^2} \right|_{q^2=0} \equiv \frac{1}{6} r_A^2$$

$$r_A^2 = 0.46(22) \text{ fm}^2$$

- order of magnitude larger uncertainty compared to historical dipole fits
- impacts comparison to other data, e.g. pion electroproduction, muon capture

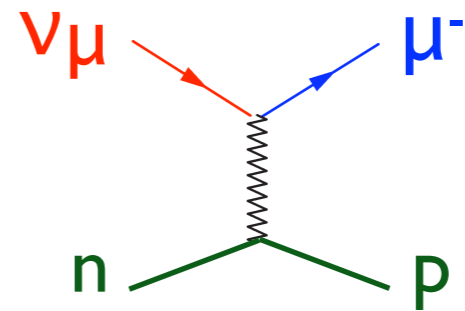
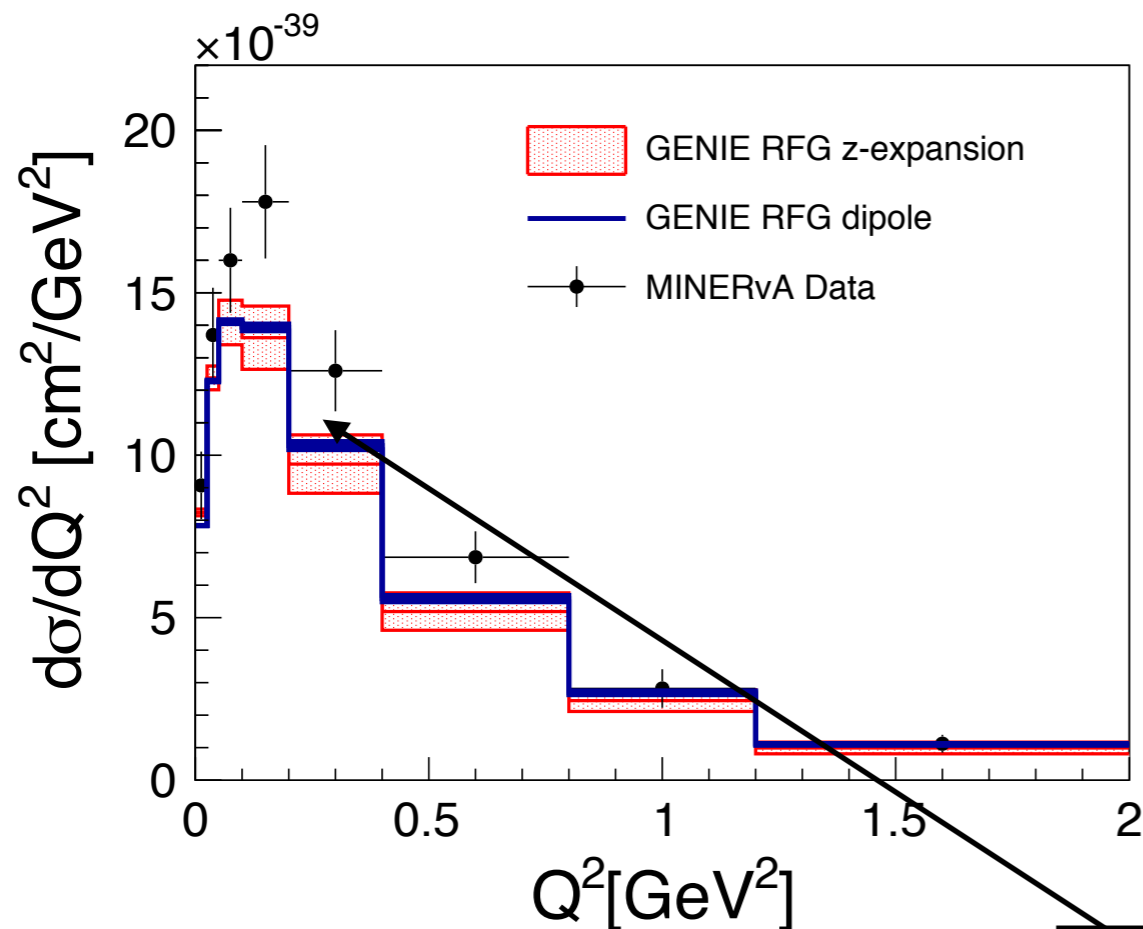
Derived observables: 2) neutrino-nucleon quasi elastic cross sections



$$\sigma_{\nu n \rightarrow \mu p}(E_\nu = 1 \text{ GeV}) = 10.1(0.9) \times 10^{-39} \text{ cm}^2$$

$$\sigma_{\nu n \rightarrow \mu p}(E_\nu = 3 \text{ GeV}) = 9.6(0.9) \times 10^{-39} \text{ cm}^2$$

discriminating nuclear models

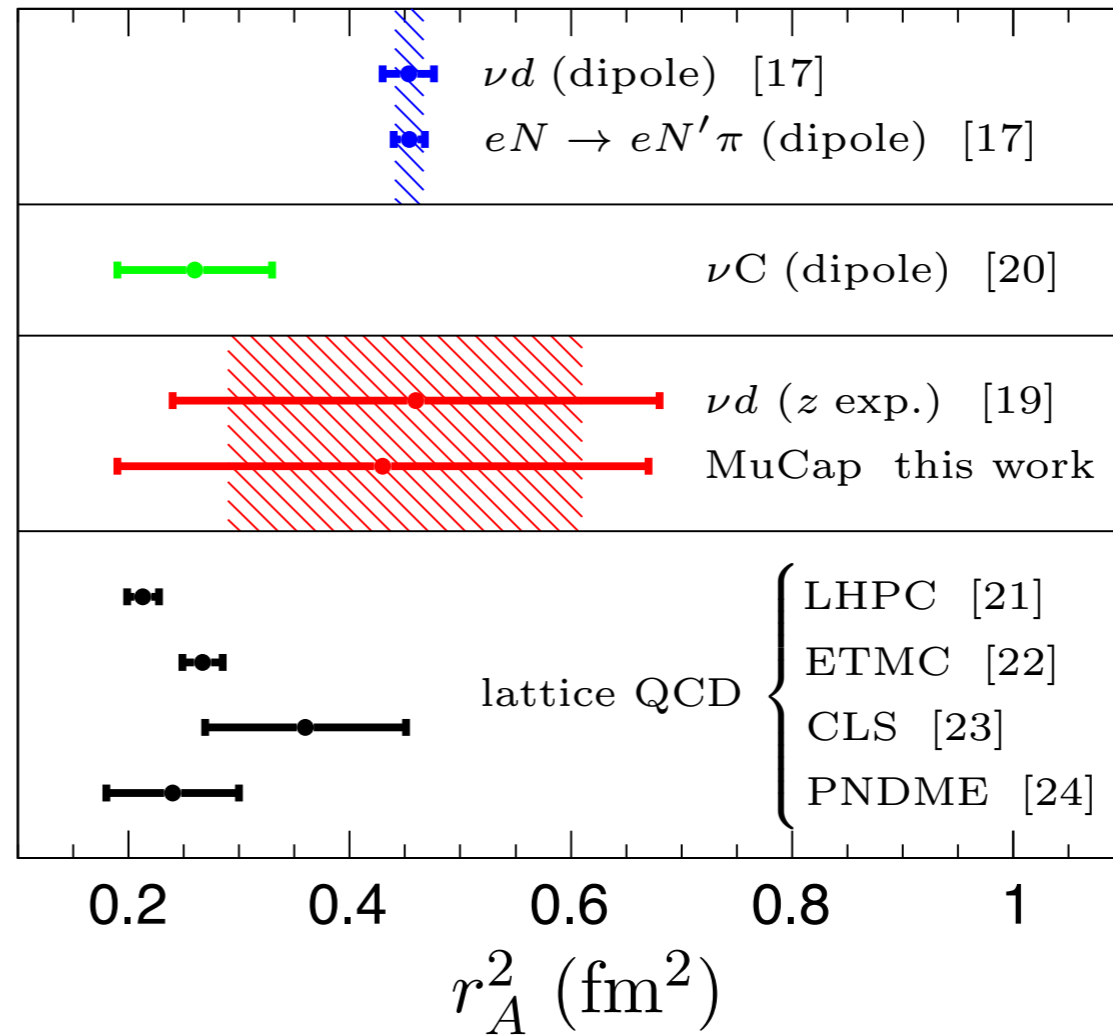
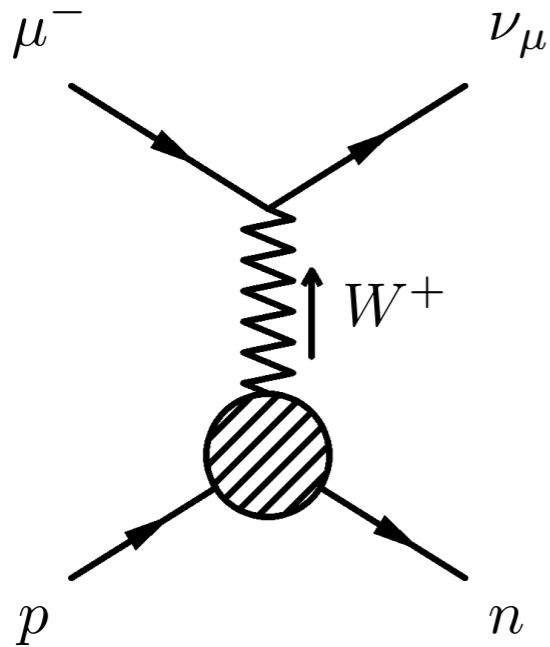


$$\sigma(\nu n \rightarrow \mu p) = |\dots \cdot F_A(q^2) \cdot \dots|^2$$

poorly known axial form factor

want to extract nuclear and flux effects from this comparison: but large nucleon level form factor uncertainty

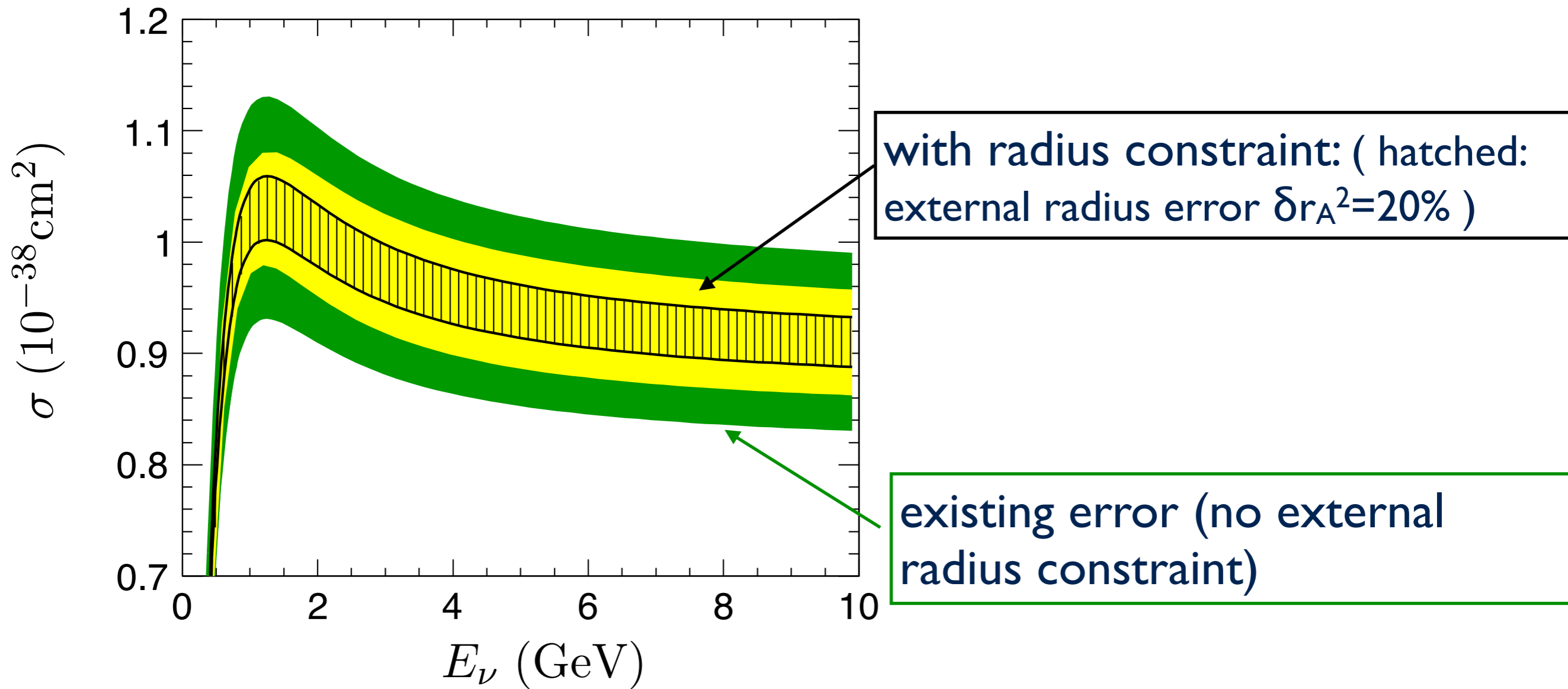
muon capture constraints



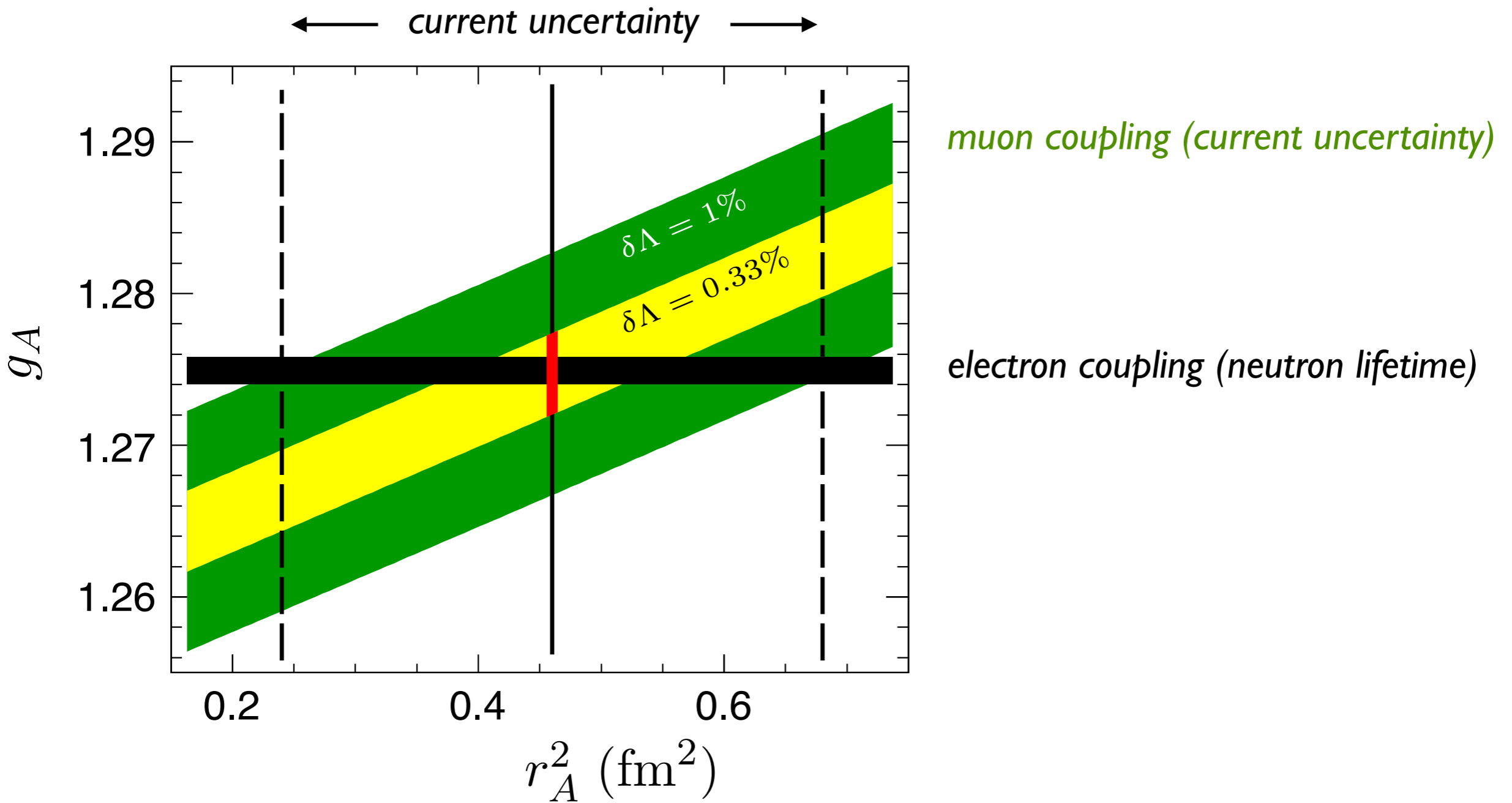
- potential factor ~ 3 improvement from next generation muon capture experiment

RJH, Kammel, Marciano, Sirlin / 708.08462

implications for quasielastic neutrino cross sections



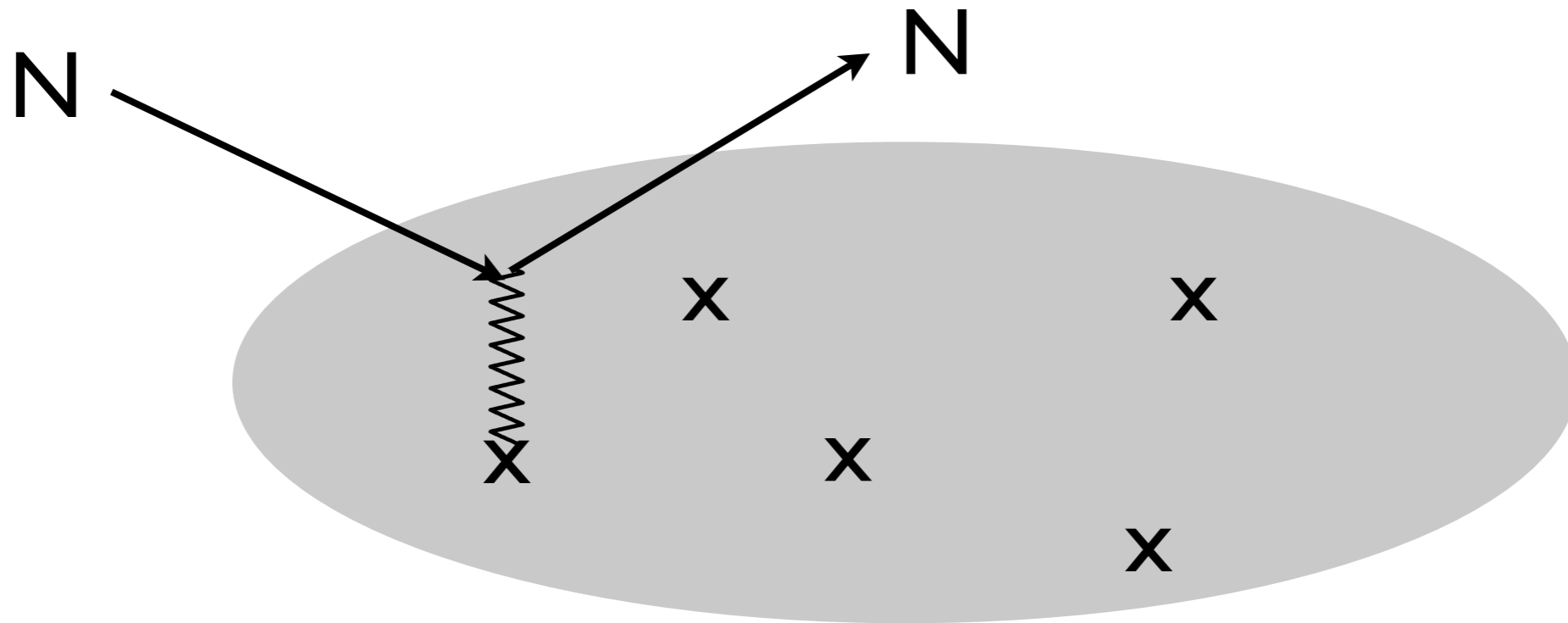
test of electron-muon universality



2. the universal WIMP-nucleon cross section

Mechanisms versus models

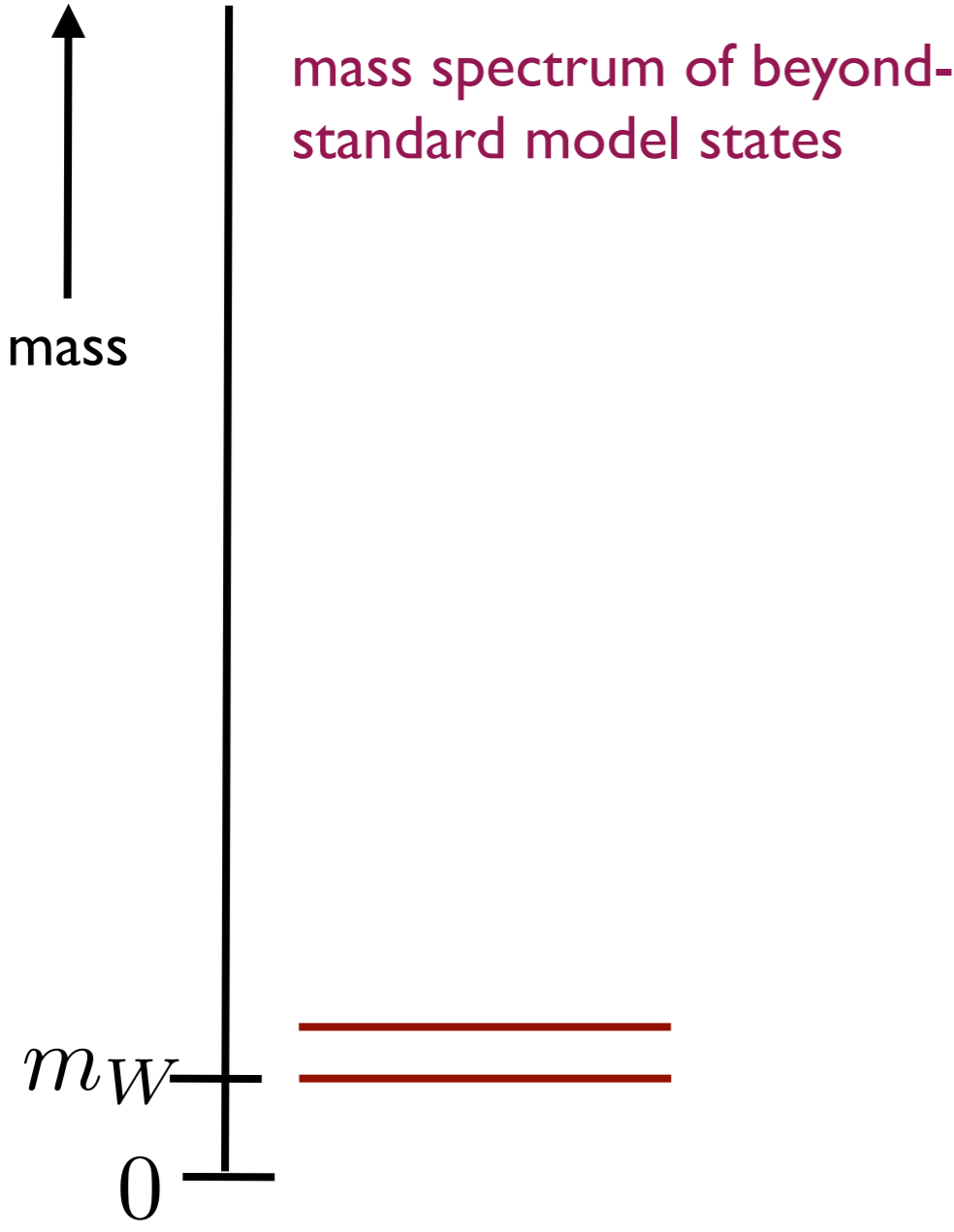
Electroweak charged WIMP Mechanism versus WIMP Model



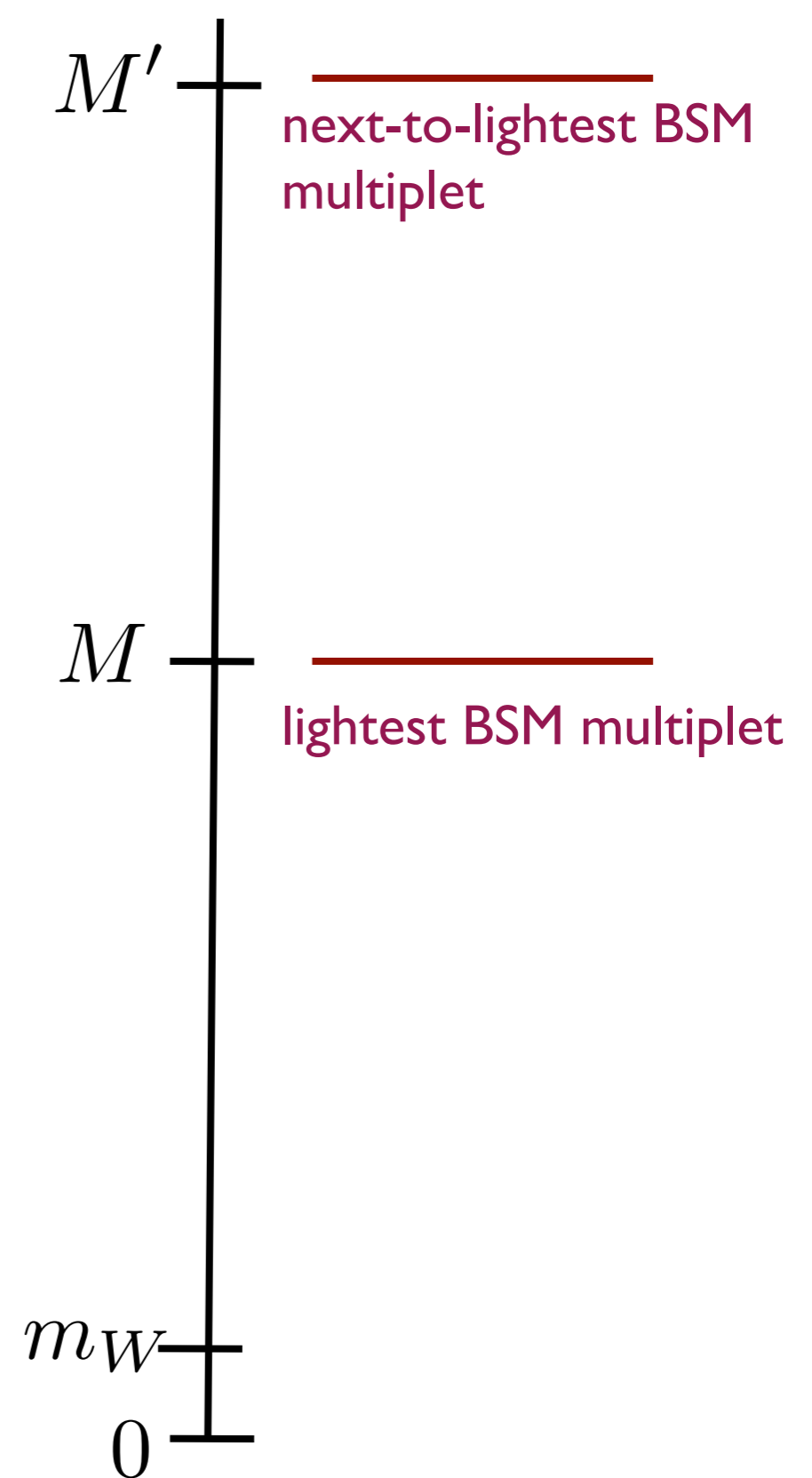
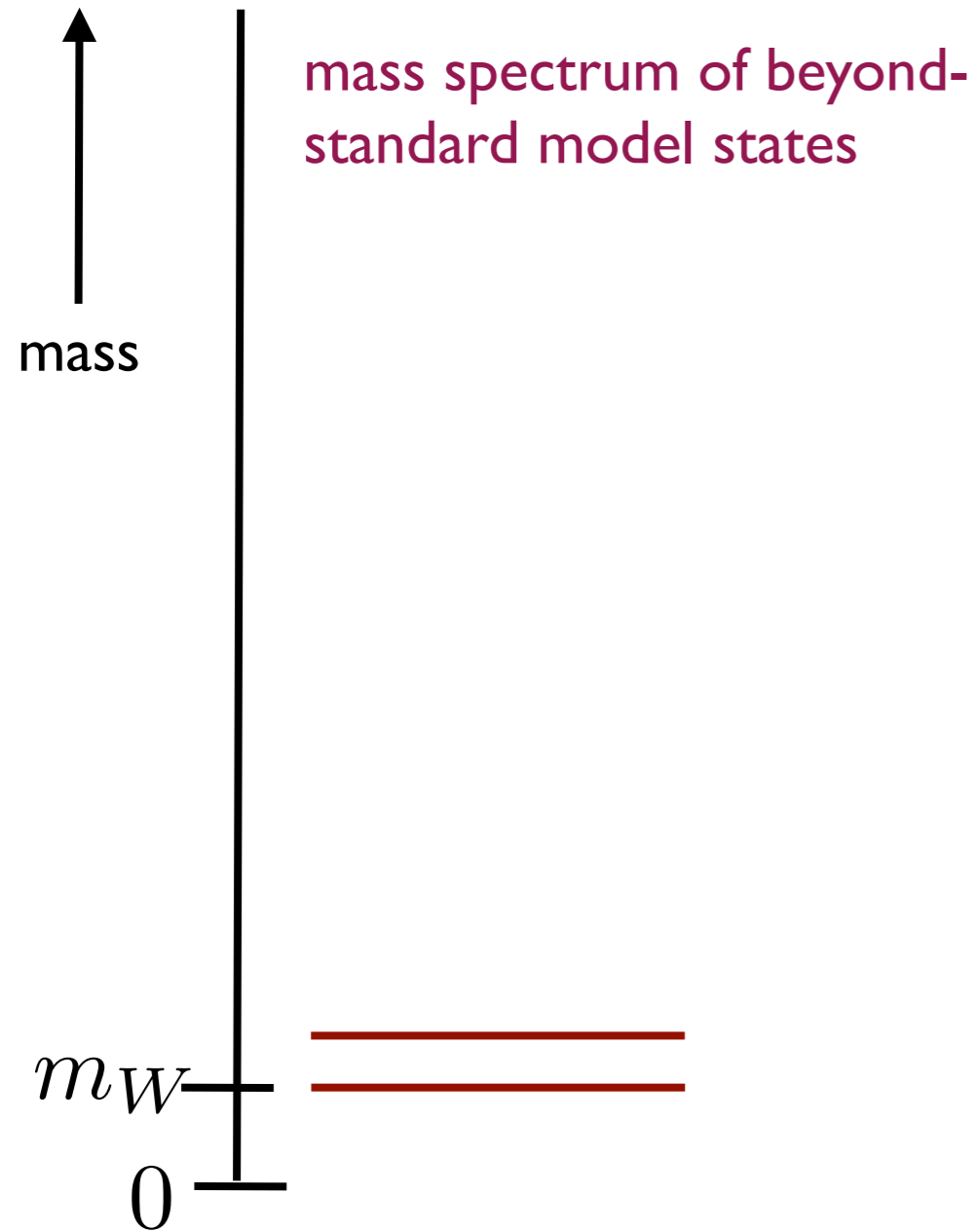
Focus on self-conjugate $SU(2)$ triplet. Could be:

- SUSY wino
- Weakly Interacting Stable Pion
- Minimal Dark Matter
- ...

Present null results of direct detection and collider searches may indicate large WIMP/New Physics mass scale



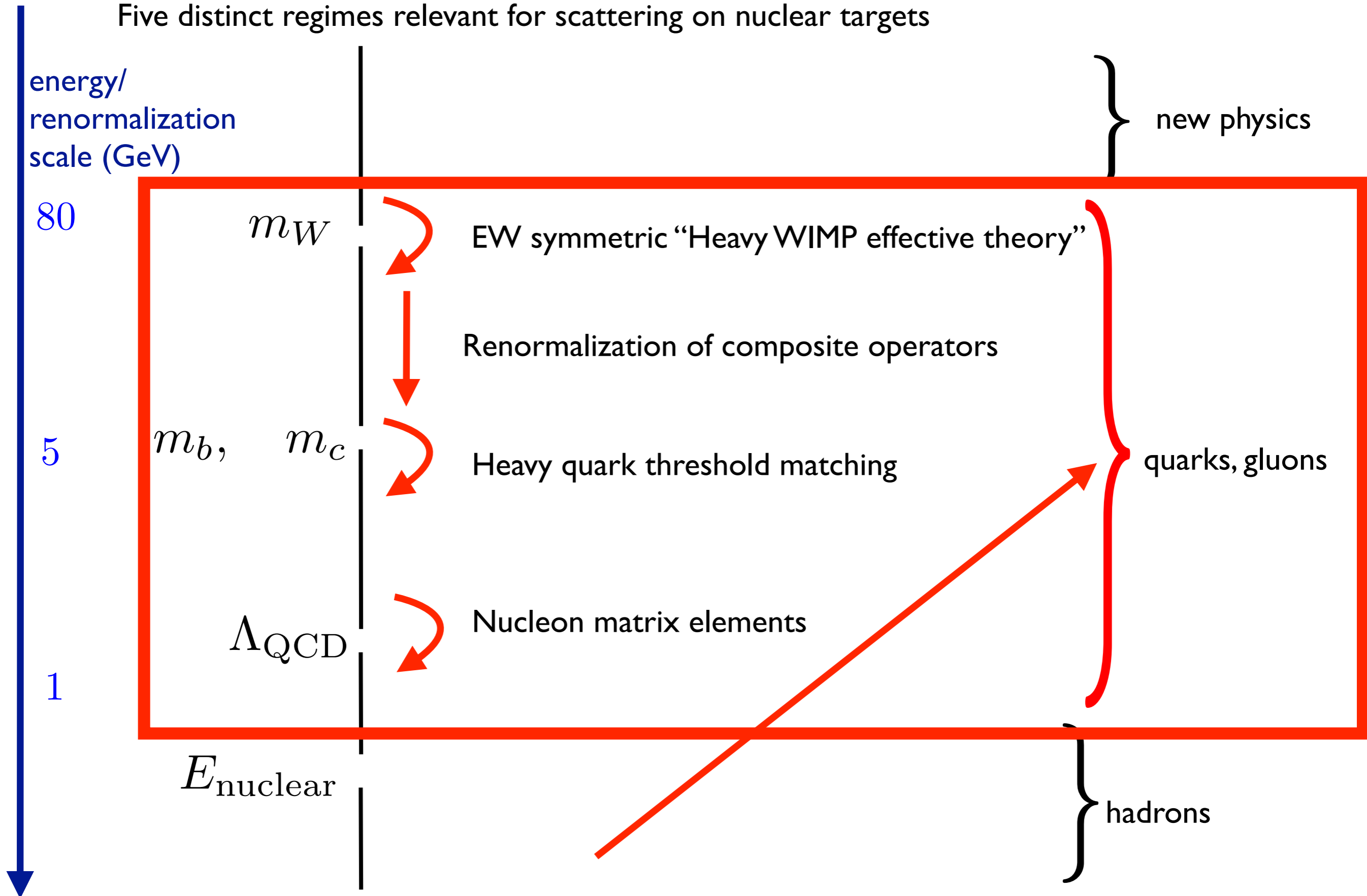
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If WIMP mass $M \gg m_W$, isolation ($M'-M \gg m_W$) becomes generic. Expand in m_W/M , $m_W/(M'-M)$

Large WIMP mass regime is a focus of future experiments in direct, indirect and collider probes

Five distinct regimes relevant for scattering on nuclear targets



“SM anatomy” of interactions between weak and hadronic scales

Direct detection

Many manifestations of heavy particle symmetry:

- hydrogen/deuterium spectroscopy

$$E_n(H) = -\frac{1}{2}m_e(Z\alpha)^2 + \dots \quad (m_e Z\alpha) \ll m_e$$

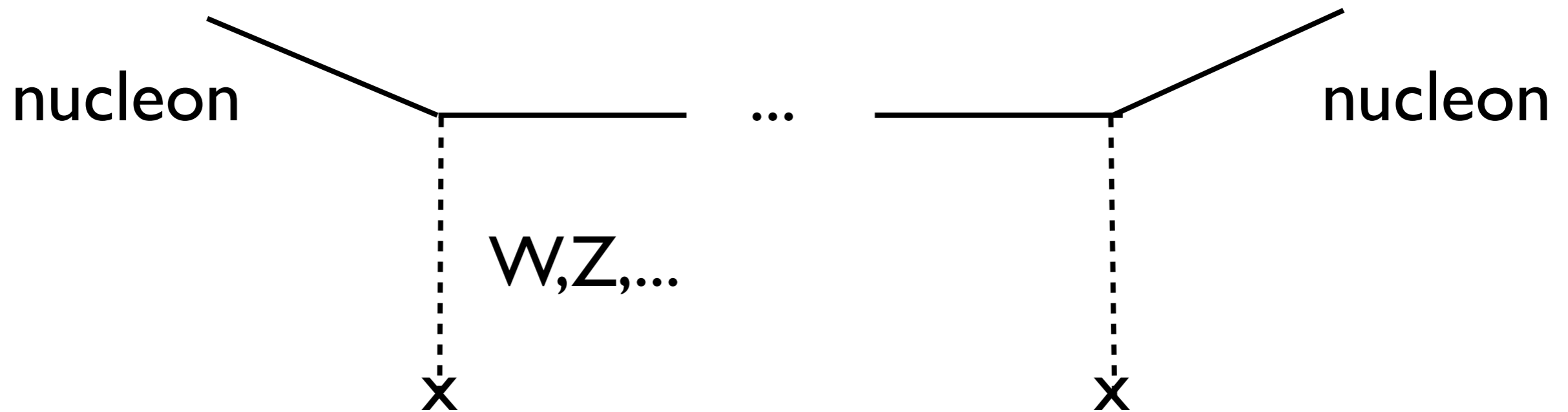
- heavy meson B/B* transitions

$$F^{B \rightarrow D}(v' = v) = 1 + \dots \quad \Lambda_{\text{QCD}} \ll m_{b,c}$$

- DM interactions

$$\sigma(\chi N \rightarrow \chi N) = ?$$

$$m_W \ll m_\chi$$



basic problem in SM physics: scattering of nucleon from SU(2)xU(1) source

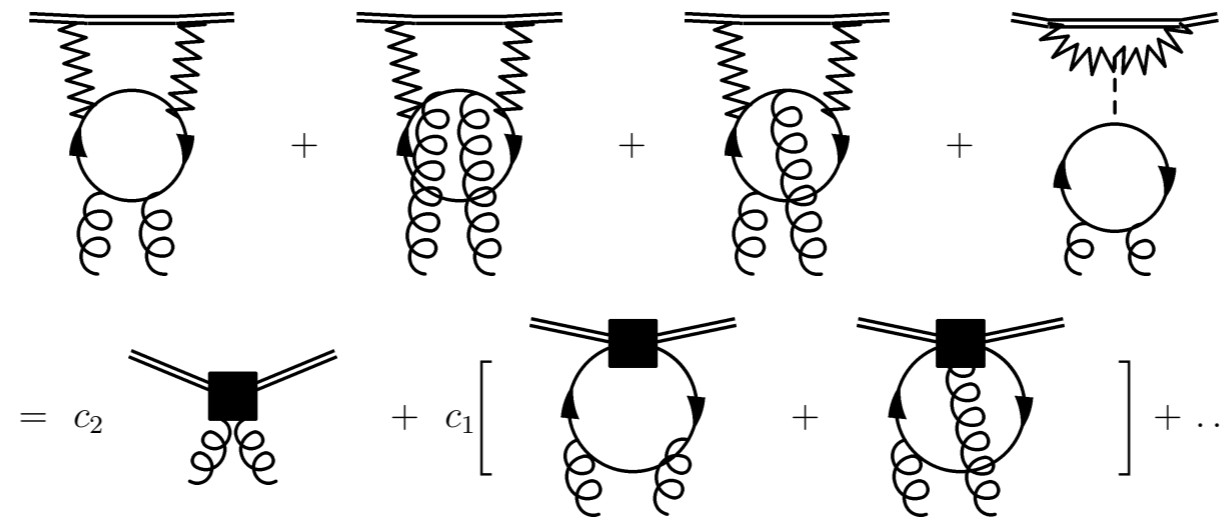
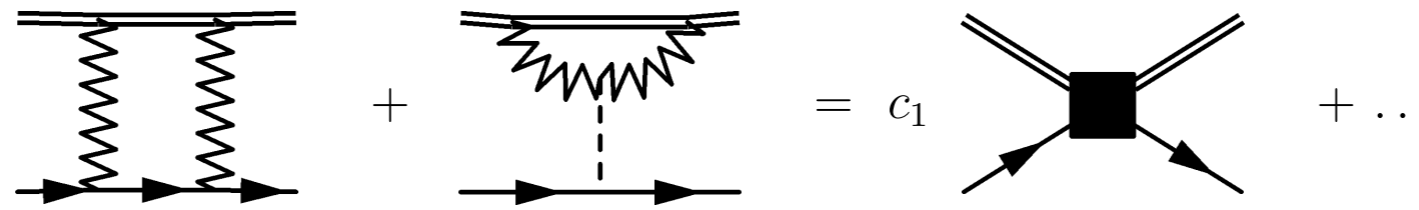
Scale separation:

dark sector
d.o.f.

SM
d.o.f.

params.
(beyond mass)

M	$\chi^{(+,-,0)}$	$Q, A_\mu^a, W_\mu^i, B_\mu$	0
m_W	$\chi_v^{(+,-,0)}$	$Q, A_\mu^a, W_\mu^i, B_\mu$	0
m_b, m_c	$\chi_v^{(0)}$	u, d, s, c, b, A_μ^a	12
Λ_{QCD}	$\chi_v^{(0)}$	u, d, s, A_μ^a	8
m_π	$\chi_v^{(0)}$	N, π	3
$1/R_{nucleus}$	$\chi_v^{(0)}$	n, p	2
	$\chi_v^{(0)}$	\mathcal{N}	1




- the heavy lifting is necessary: large gluon matrix element, amplitude cancellations

Scale separation:

dark sector
d.o.f.

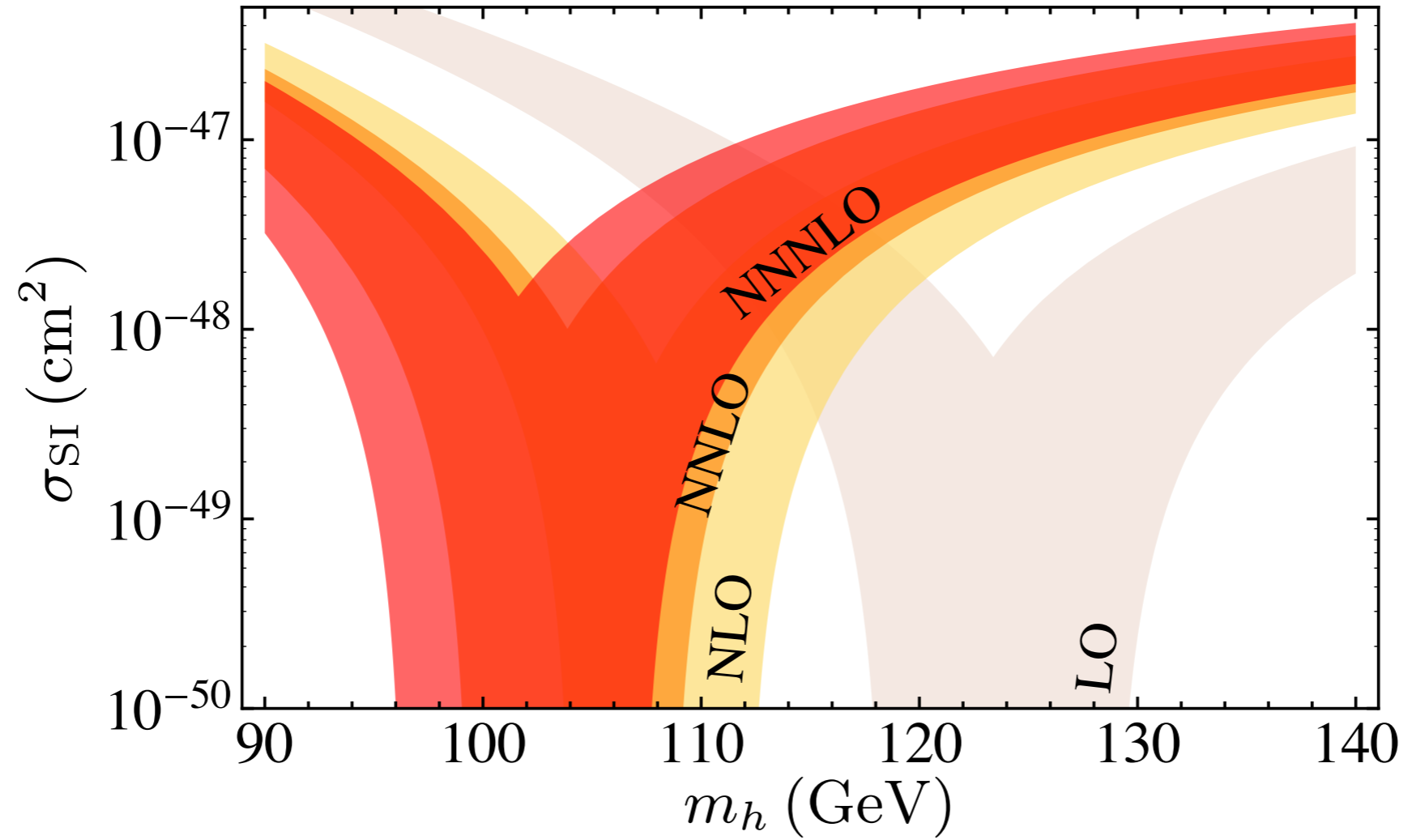
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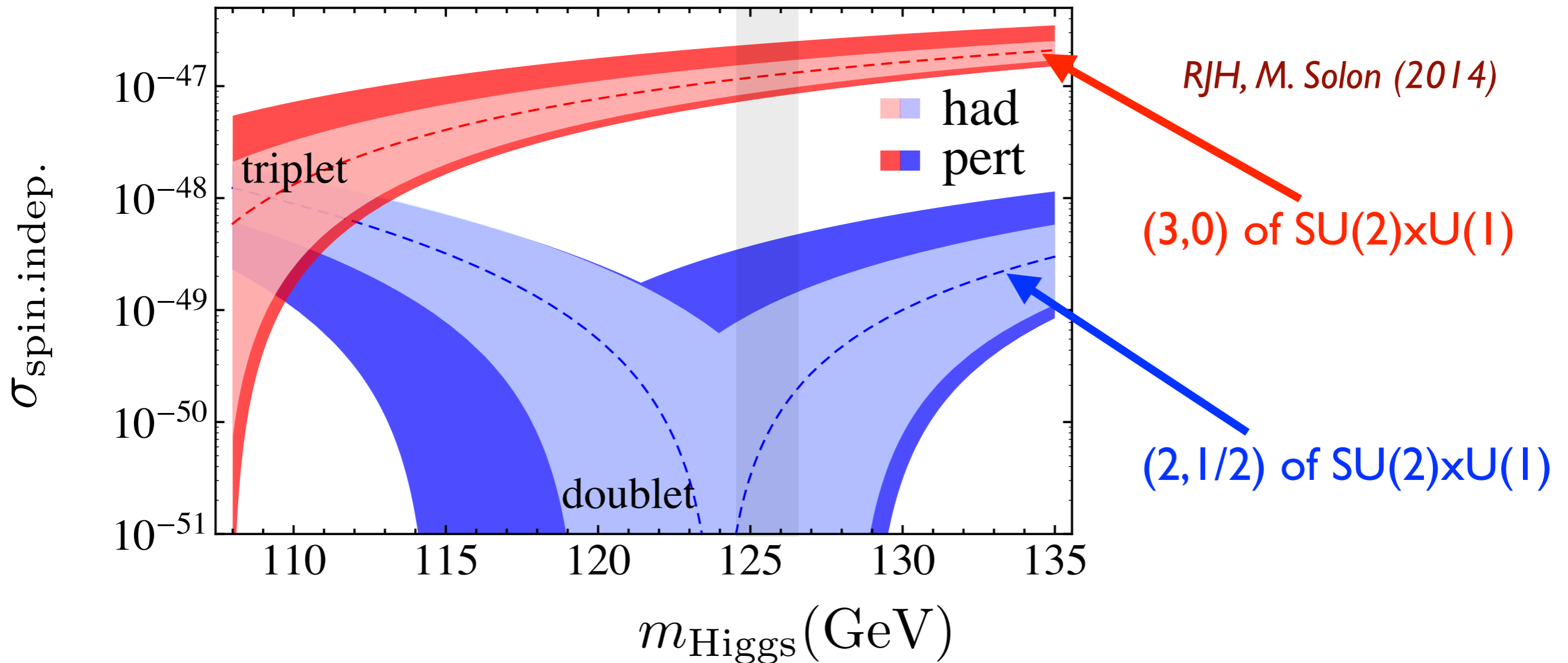


M	$\chi^{(+,-,0)}$	$Q, A_\mu^a, W_\mu^i, B_\mu$	0
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- the heavy lifting is necessary



Benchmarks: large mass, low velocity limit

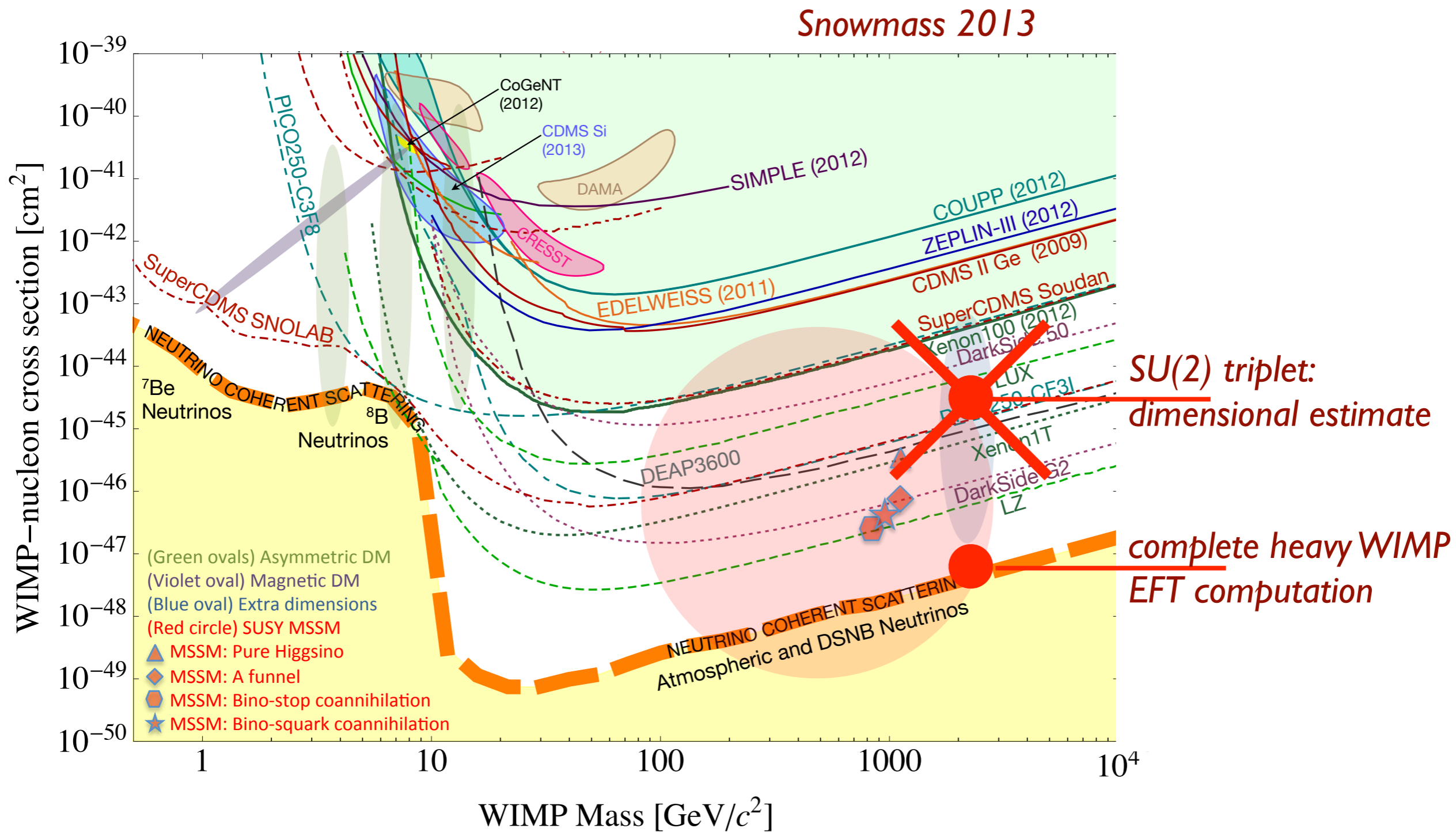


- Suppressed versus dimensional estimate ($\sim 10^{-45} \text{cm}^2$)

- $1/M$ power corrections under investigation

C.-Y. Chen, RjH, M. Solon, A. Wijangco, to appear

- the heavy lifting is necessary



3. the proton radius puzzle

Some facts about the proton radius puzzle

1) It has generated a lot of attention and controversy



The New York Times

2) The *most mundane* resolution necessitates:

- $>5\sigma$ shift in fundamental Rydberg constant
- discarding or revising decades of results in e-p scattering and hydrogen spectroscopy

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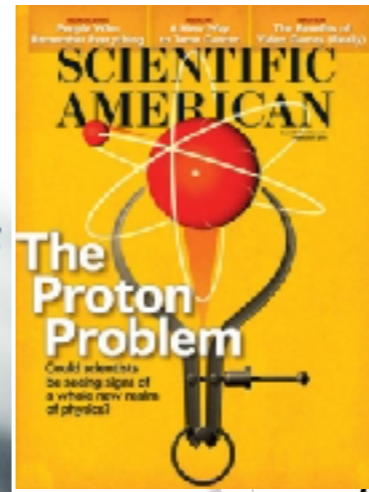
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“The good news is that it’s not my problem”

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This is everybody's problem (HEP, NP, AMO,...):

3) E.g. systematic effects in electron-proton scattering impact neutrino-nucleus scattering, *at a level large compared to long baseline precision requirements*



“The good news is that it’s not my problem”

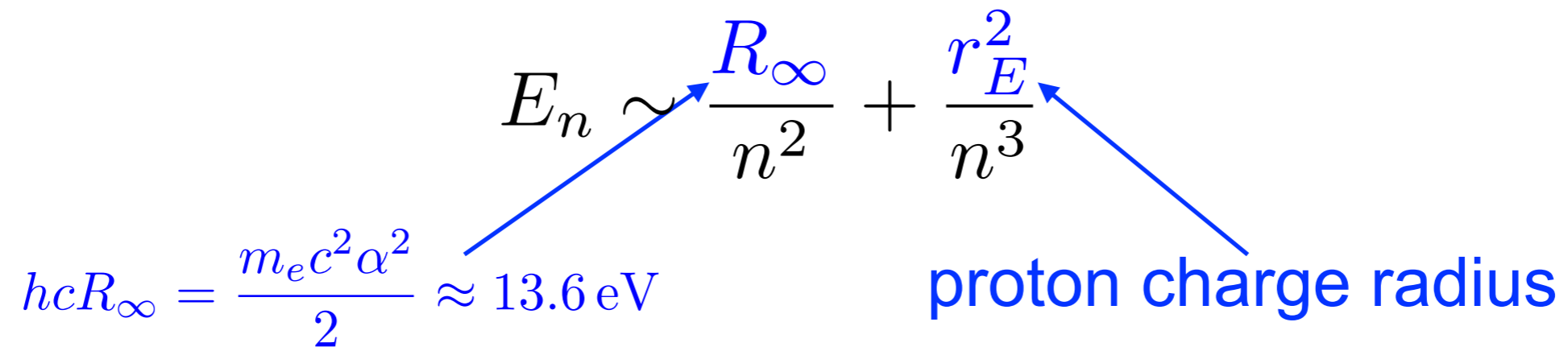
the puzzle

Recall hydrogen spectrum:

$$E_n \sim \frac{R_\infty}{n^2} + \frac{r_E^2}{n^3}$$

$hcR_\infty = \frac{m_e c^2 \alpha^2}{2} \approx 13.6 \text{ eV}$

proton charge radius

The diagram shows the energy level equation $E_n \sim \frac{R_\infty}{n^2} + \frac{r_E^2}{n^3}$. A blue arrow points from the R_∞ term to the definition $hcR_\infty = \frac{m_e c^2 \alpha^2}{2} \approx 13.6 \text{ eV}$. Another blue arrow points from the r_E^2 term to the text "proton charge radius".

Disentangle 2 unknowns, R_∞ and r_E , using well-measured 1S-2S hydrogen transition *and*

electron-based
measurements

muon-based
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electron-based measurements

- another hydrogen interval

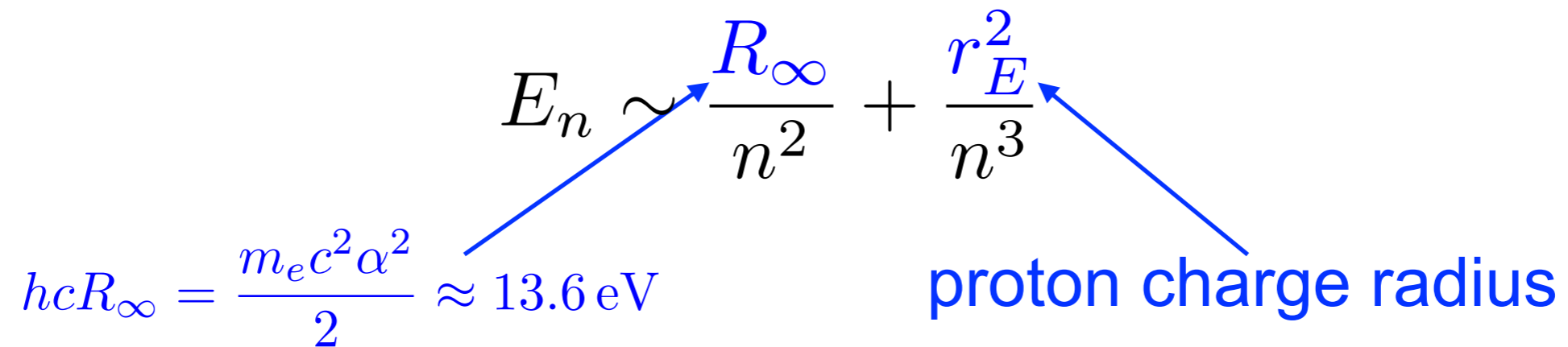
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- another hydrogen interval
- electron-proton scattering determination of r_E

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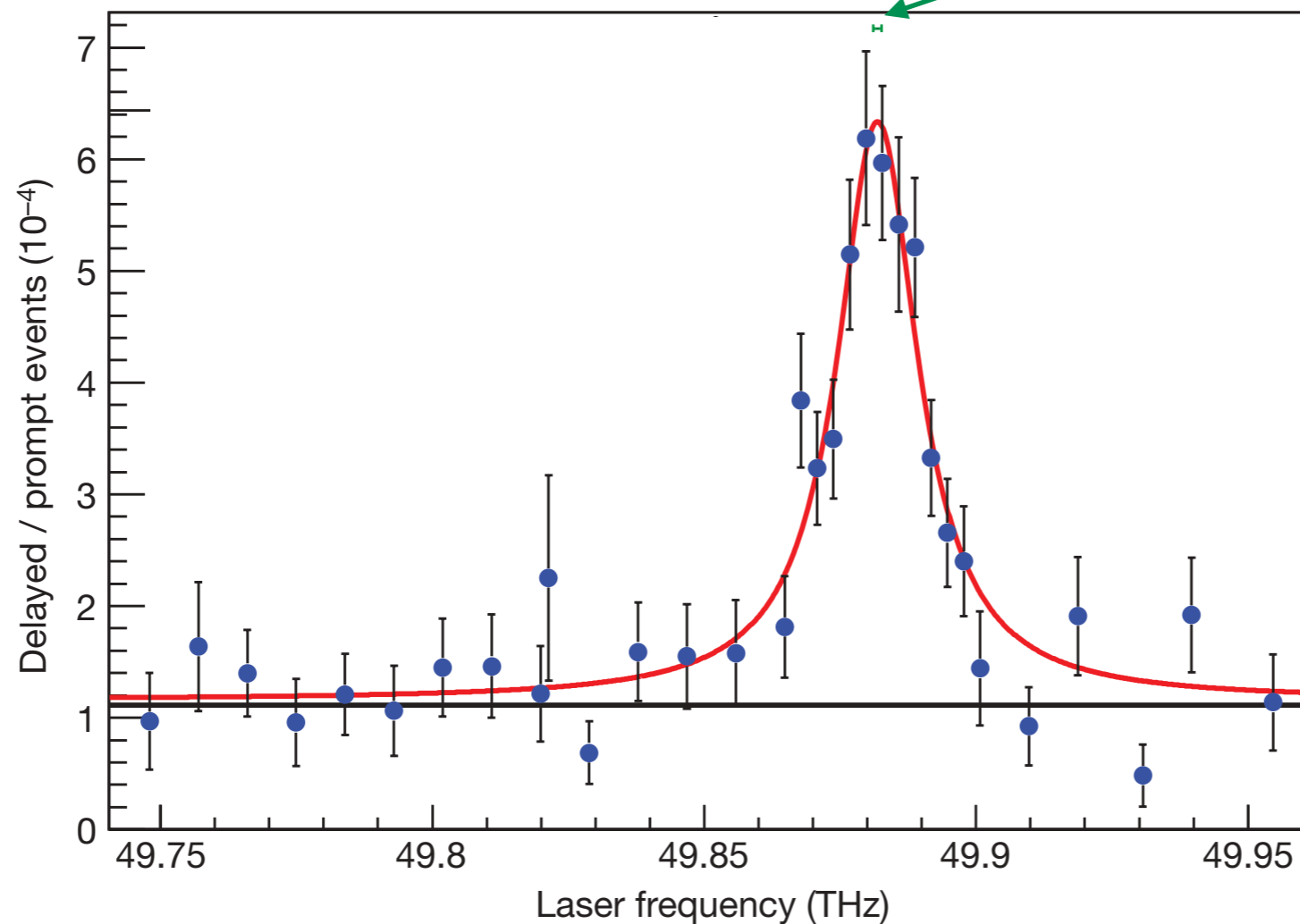
- a muonic hydrogen interval

7σ discrepancy between electron-based versus muon-based measurements

muonic hydrogen Lamb shift measurement

measured frequency of
2S-2P transition in muonic H

Pohl et al. (CREMA collaboration), Nature 466, 213 (2010)



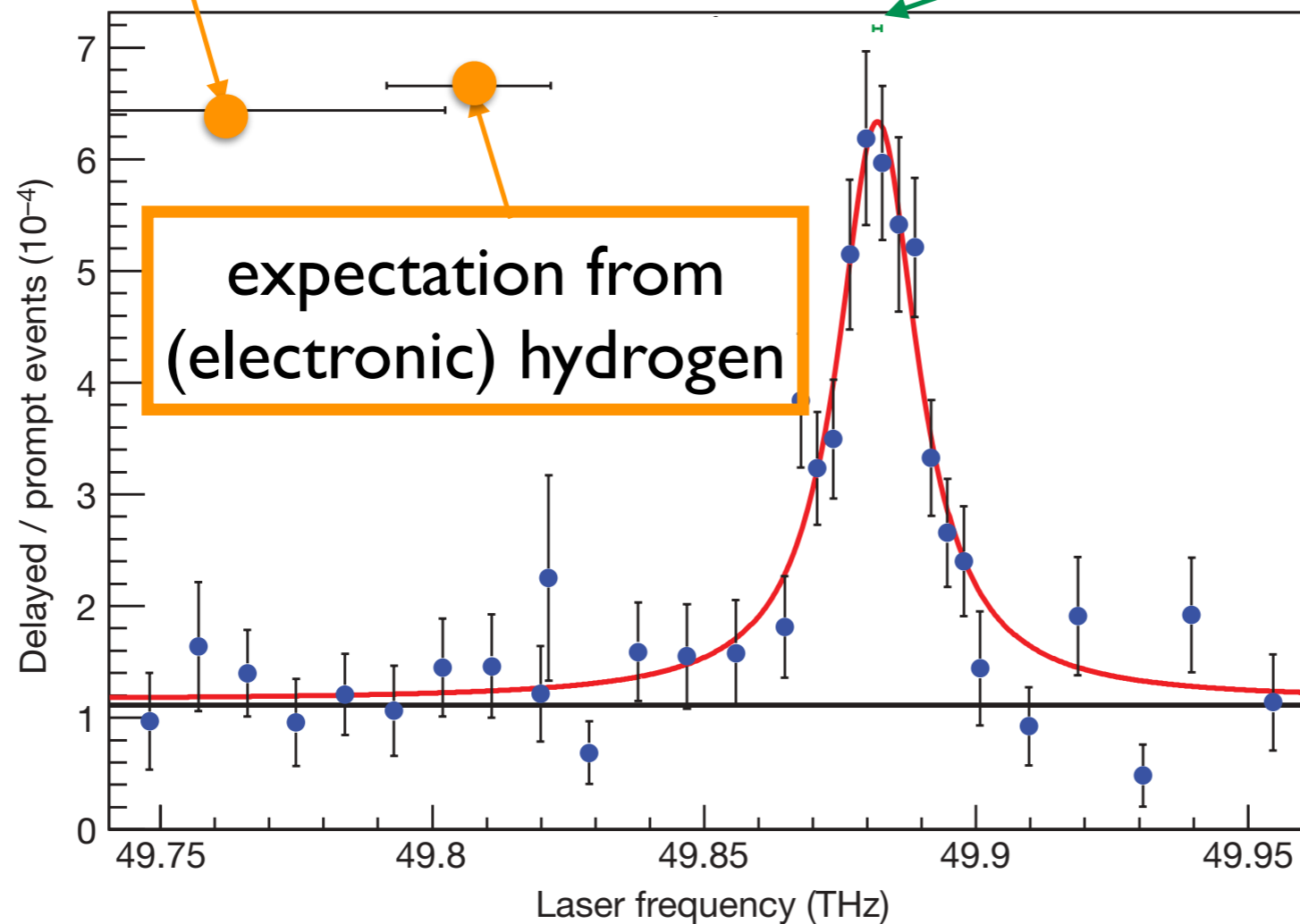
new experimental capabilities: surprises and new insight ?

muonic hydrogen Lamb shift measurement

expectation from
e-p scattering

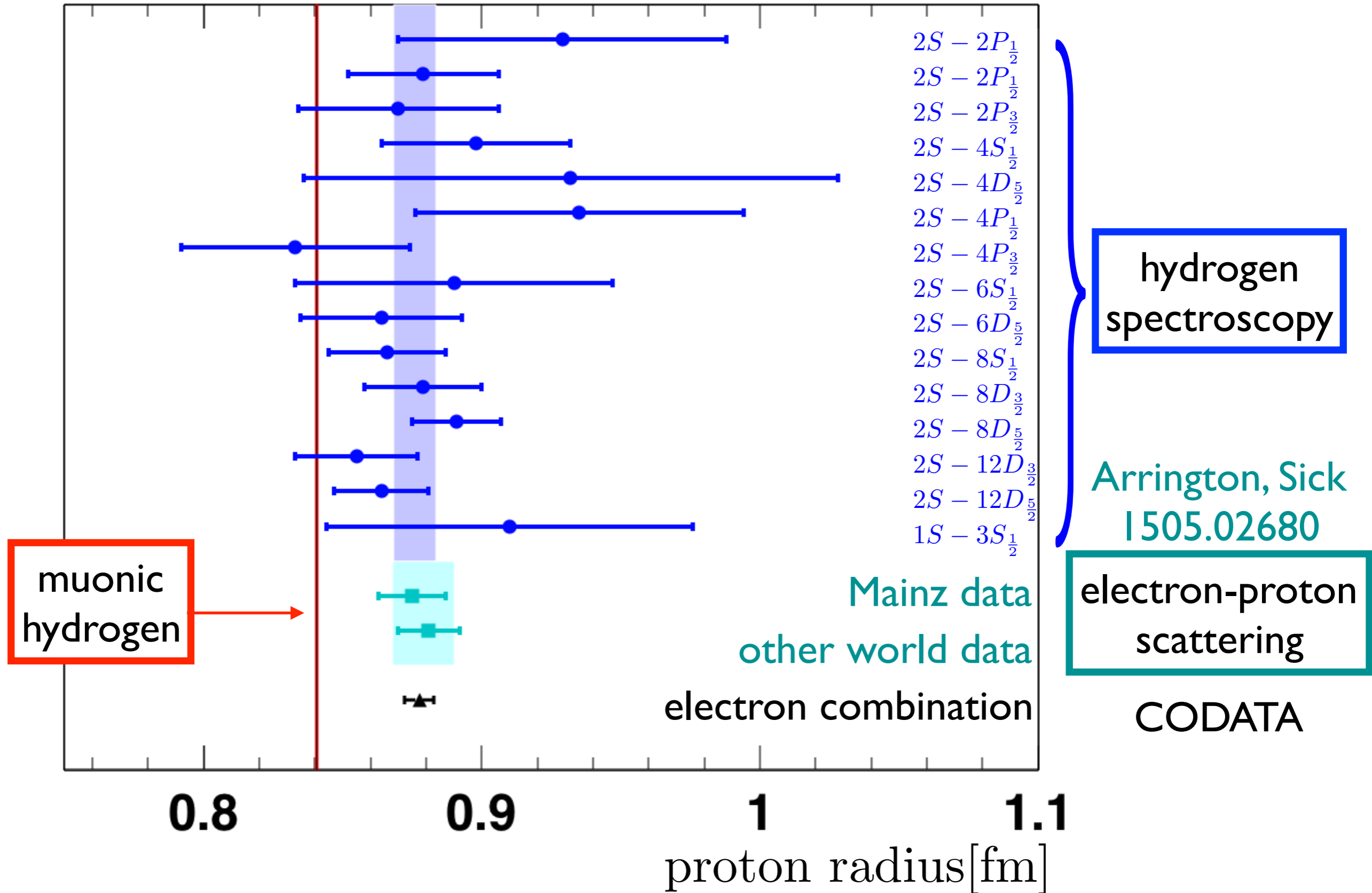
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new experimental capabilities: surprises and new insight ?

summary of electron- and muon- based measurements



status of some theory issues

electron-proton scattering: theory issues

radius is defined as slope of form factor

i) what are the constraints on nonlinearities?

radiative corrections impact radius extraction and can be large (~30%)

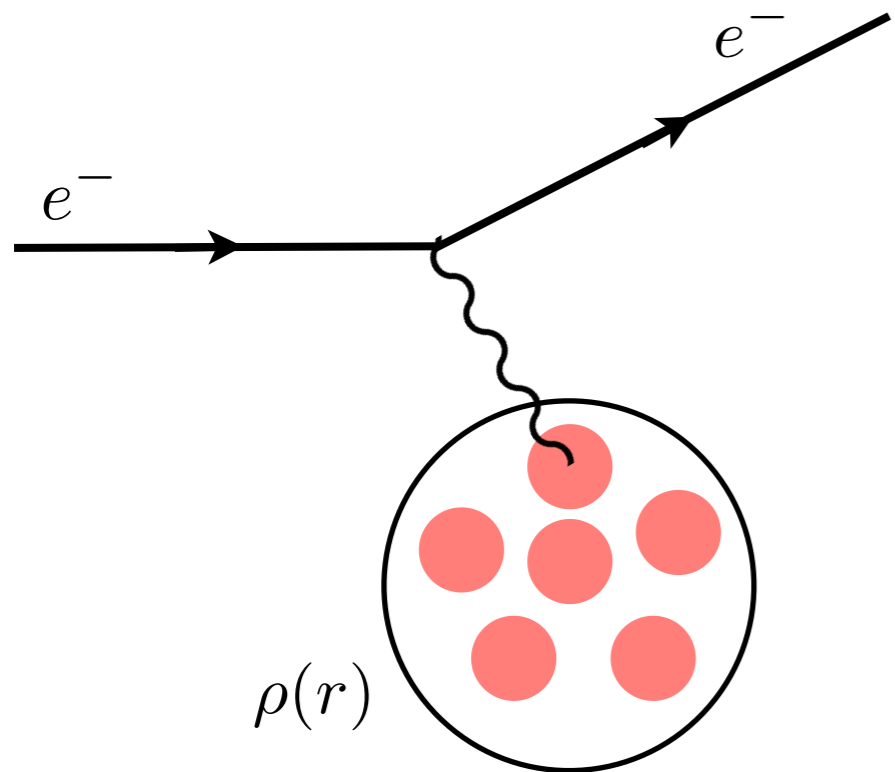
ii) are radiative corrections controlled at the sub percent level?

i) what are the constraints on nonlinearities?

recall scattering from extended classical charge distribution:

$$\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega} \right)_{\text{pointlike}} |F(q^2)|^2$$

$$\begin{aligned} F(q^2) &= \int d^3r e^{i\mathbf{q}\cdot\mathbf{r}} \rho(\mathbf{r}) \\ &= \int d^3r \left[1 + i\mathbf{q}\cdot\mathbf{r} - \frac{1}{2}(\mathbf{q}\cdot\mathbf{r})^2 + \dots \right] \rho(\mathbf{r}) \\ &= 1 - \frac{1}{6}\langle r^2 \rangle \mathbf{q}^2 + \dots \end{aligned}$$



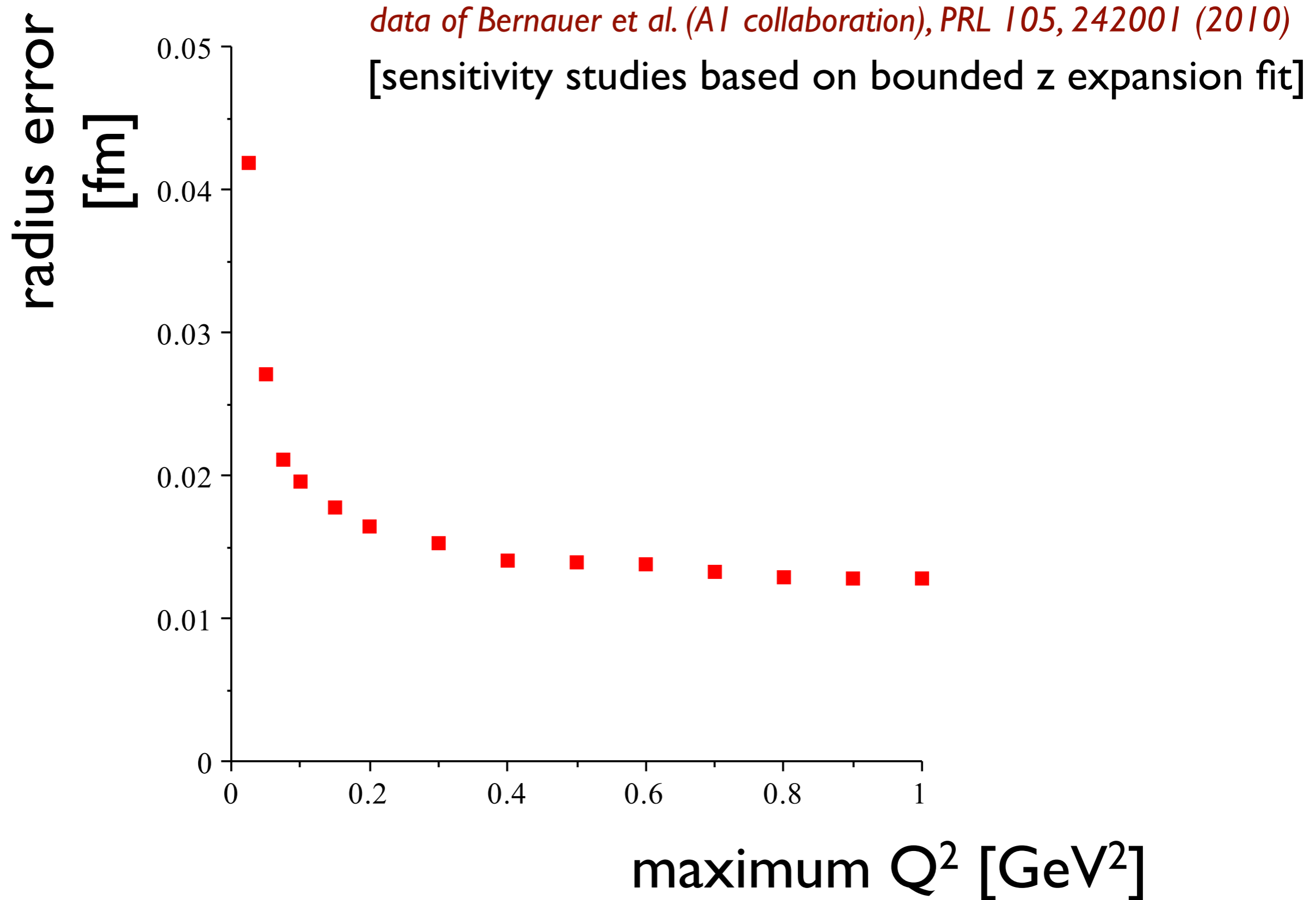
for the relativistic, QM, case, define radius as slope of form factor

$$\begin{aligned} \langle J^\mu \rangle &= \gamma^\mu F_1 + \frac{i}{2m_p} \sigma^{\mu\nu} q_\nu F_2 \\ G_E &= F_1 + \frac{q^2}{4m_p^2} F_2 \quad G_M = F_1 + F_2 \end{aligned}$$

$$r_E^2 \equiv 6 \frac{d}{dq^2} G_E(q^2) \Big|_{q^2=0}$$

(up to radiative corrections)

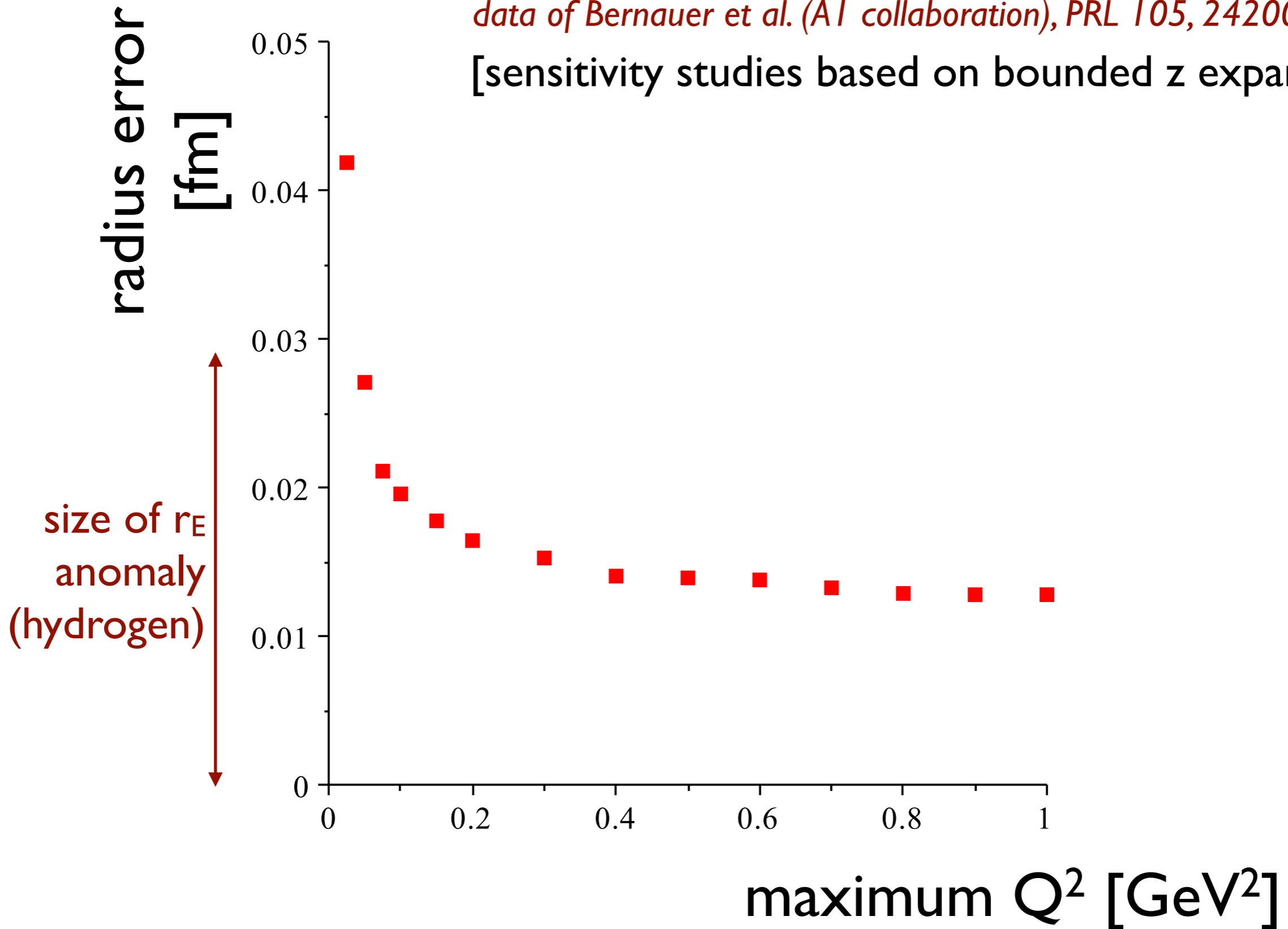
Radius extraction requires data over a Q^2 range where a simple Taylor expansion of the form factor is invalid



Radius extraction requires data over a Q^2 range where a simple Taylor expansion of the form factor is invalid

data of Bernauer et al. (A1 collaboration), PRL 105, 242001 (2010)

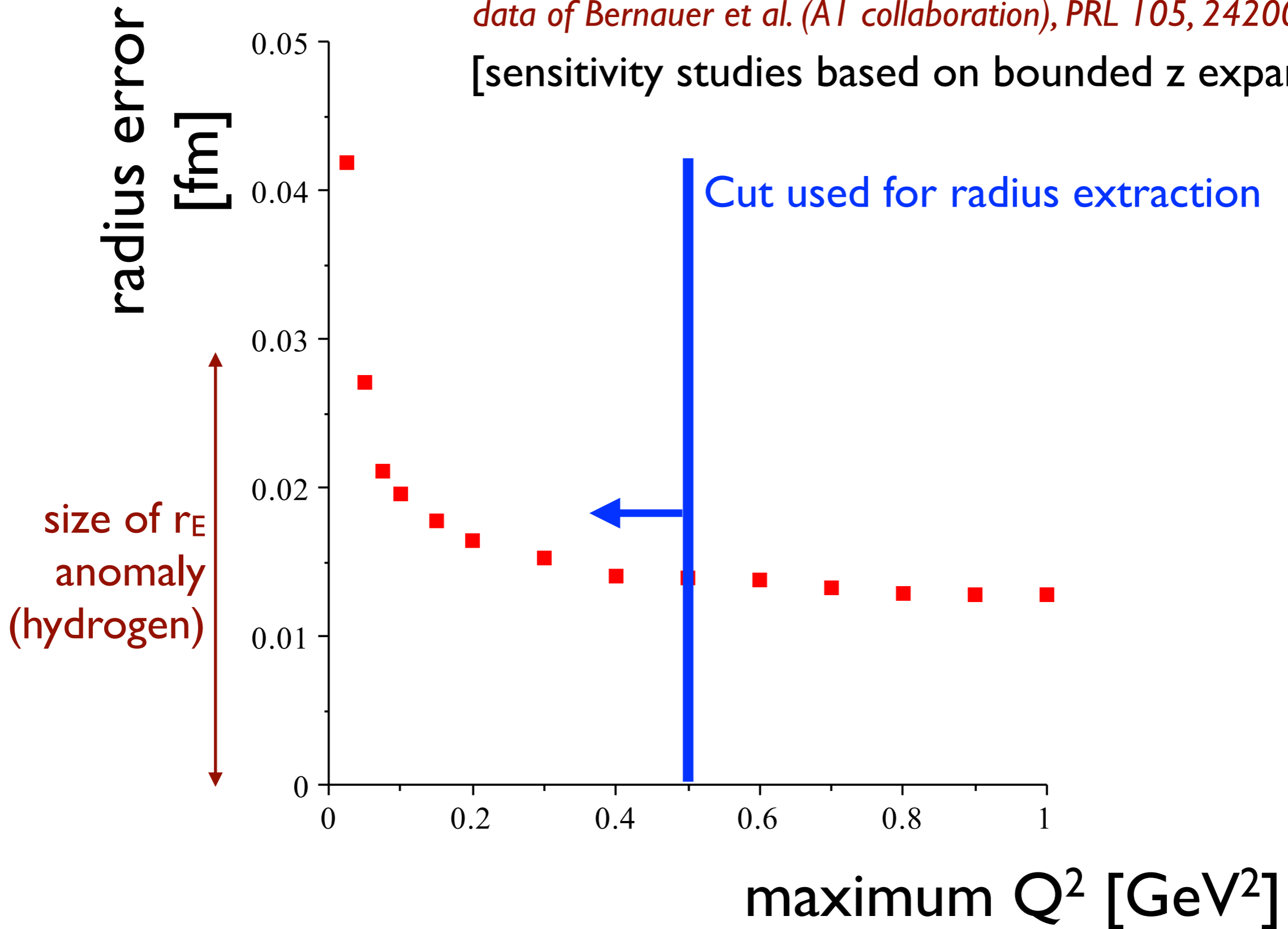
[sensitivity studies based on bounded z expansion fit]



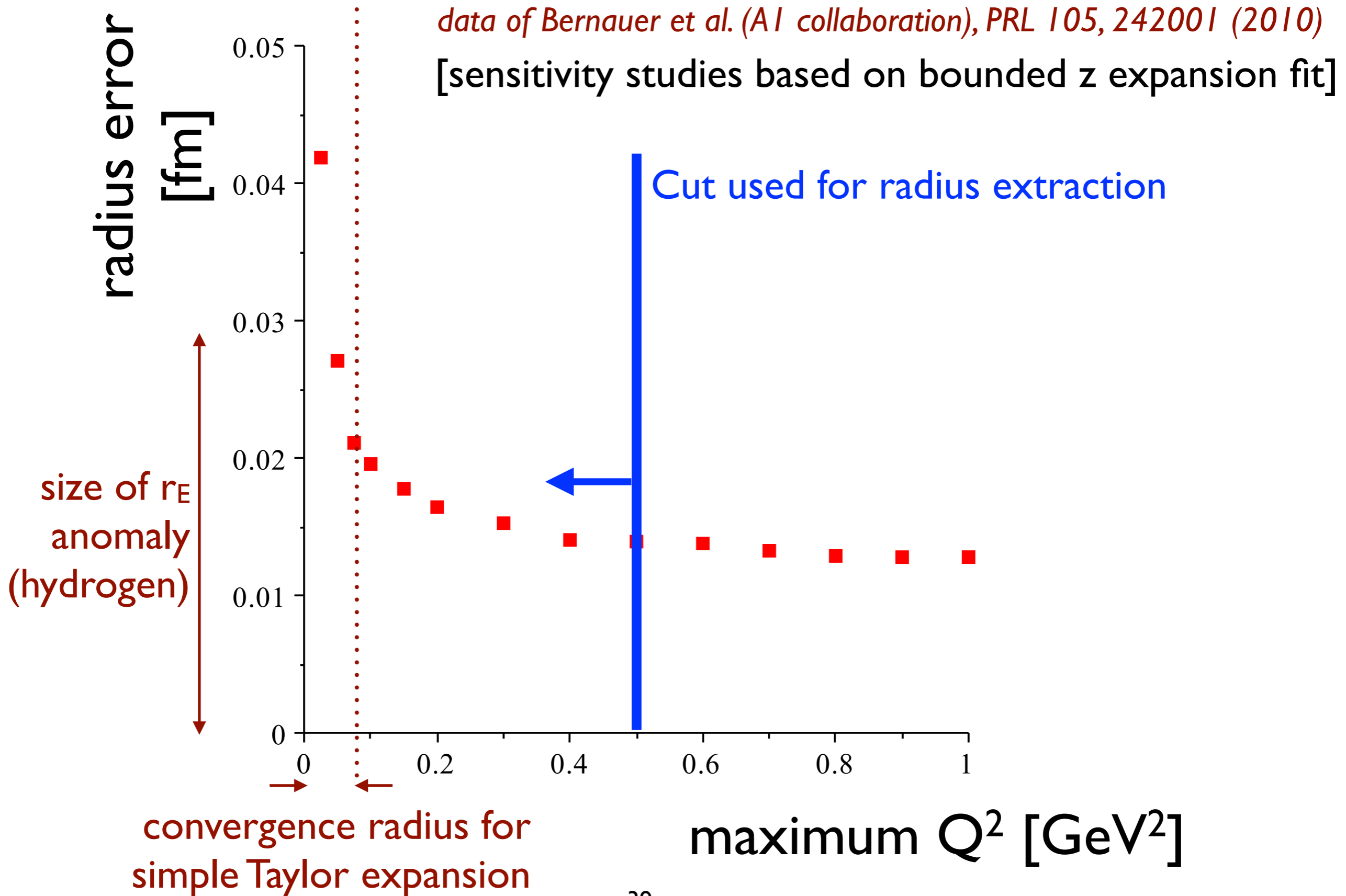
Radius extraction requires data over a Q^2 range where a simple Taylor expansion of the form factor is invalid

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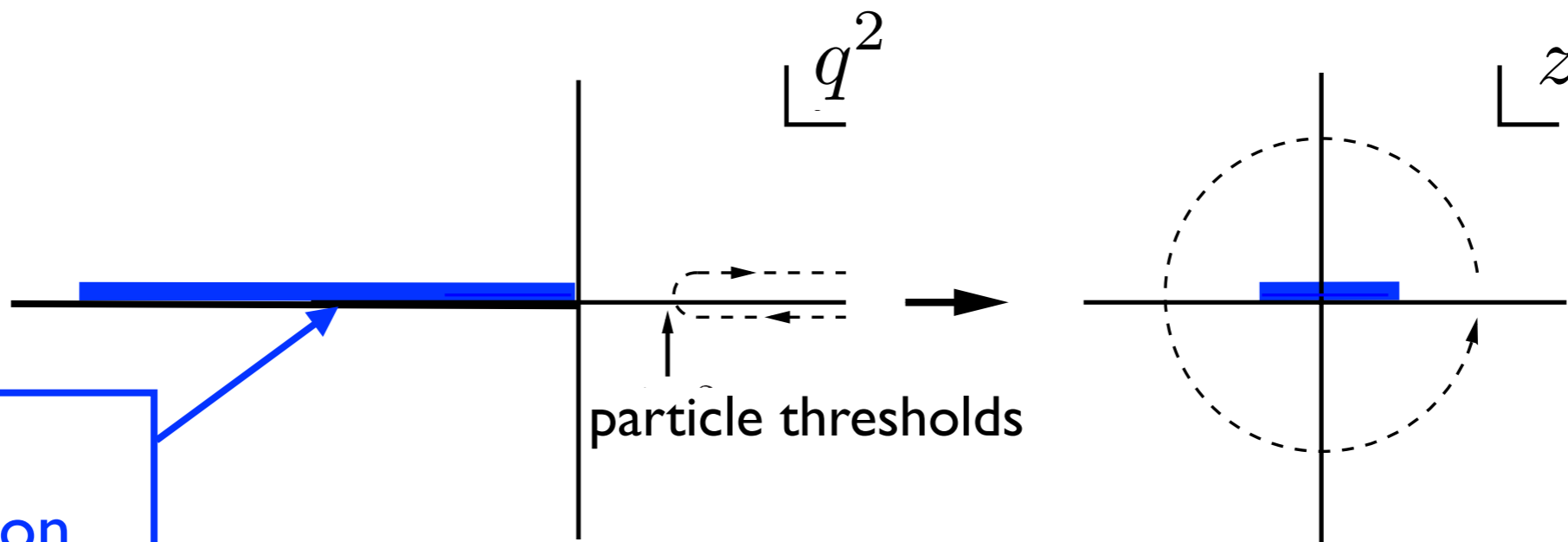
[sensitivity studies based on bounded z expansion fit]



Radius extraction requires data over a Q^2 range where a simple Taylor expansion of the form factor is invalid



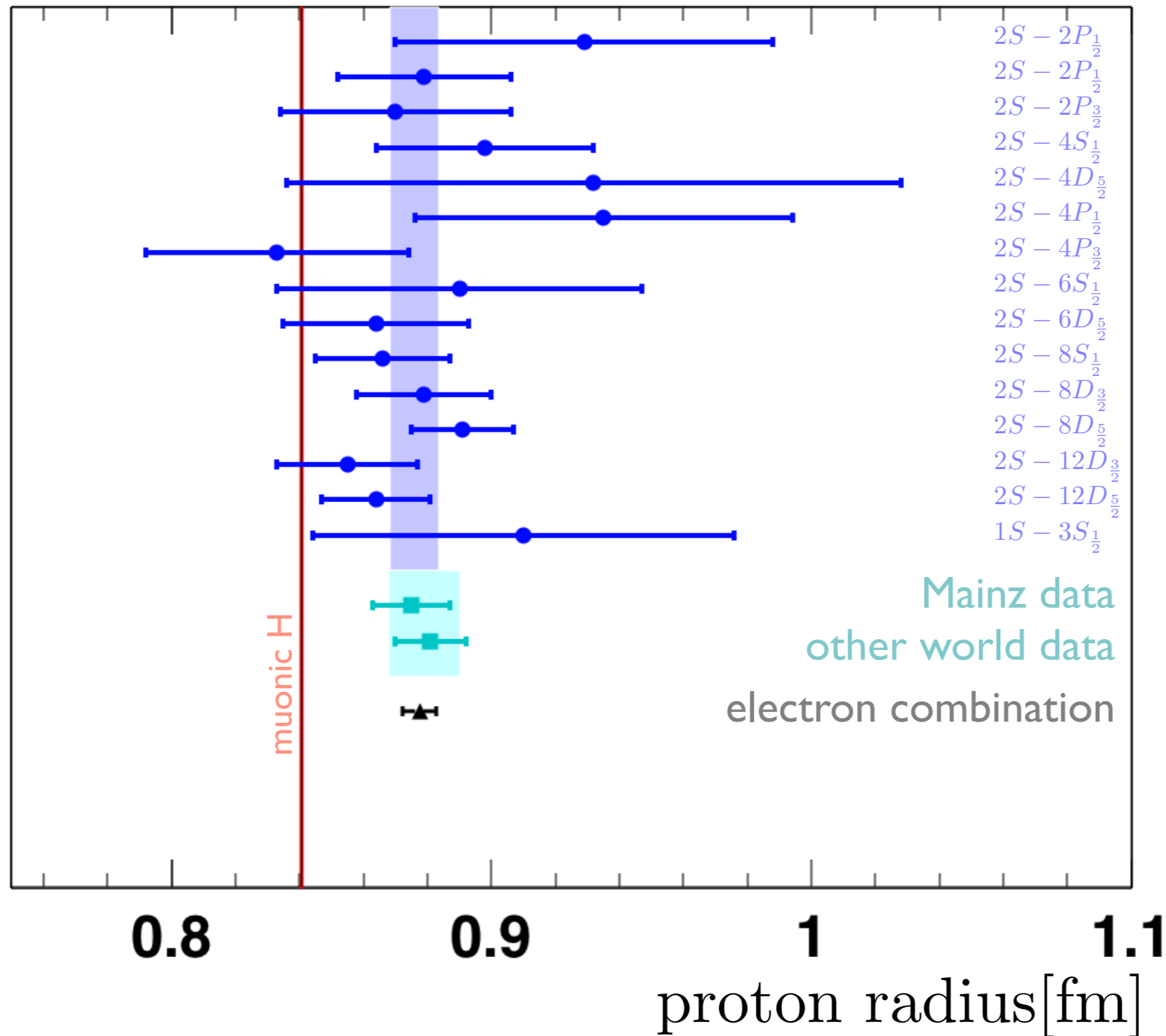
That's ok: underlying QCD tells us that Taylor expansion of form factor in appropriate variable is convergent



$$F(q^2) = \sum_k a_k [z(q^2)]^k$$

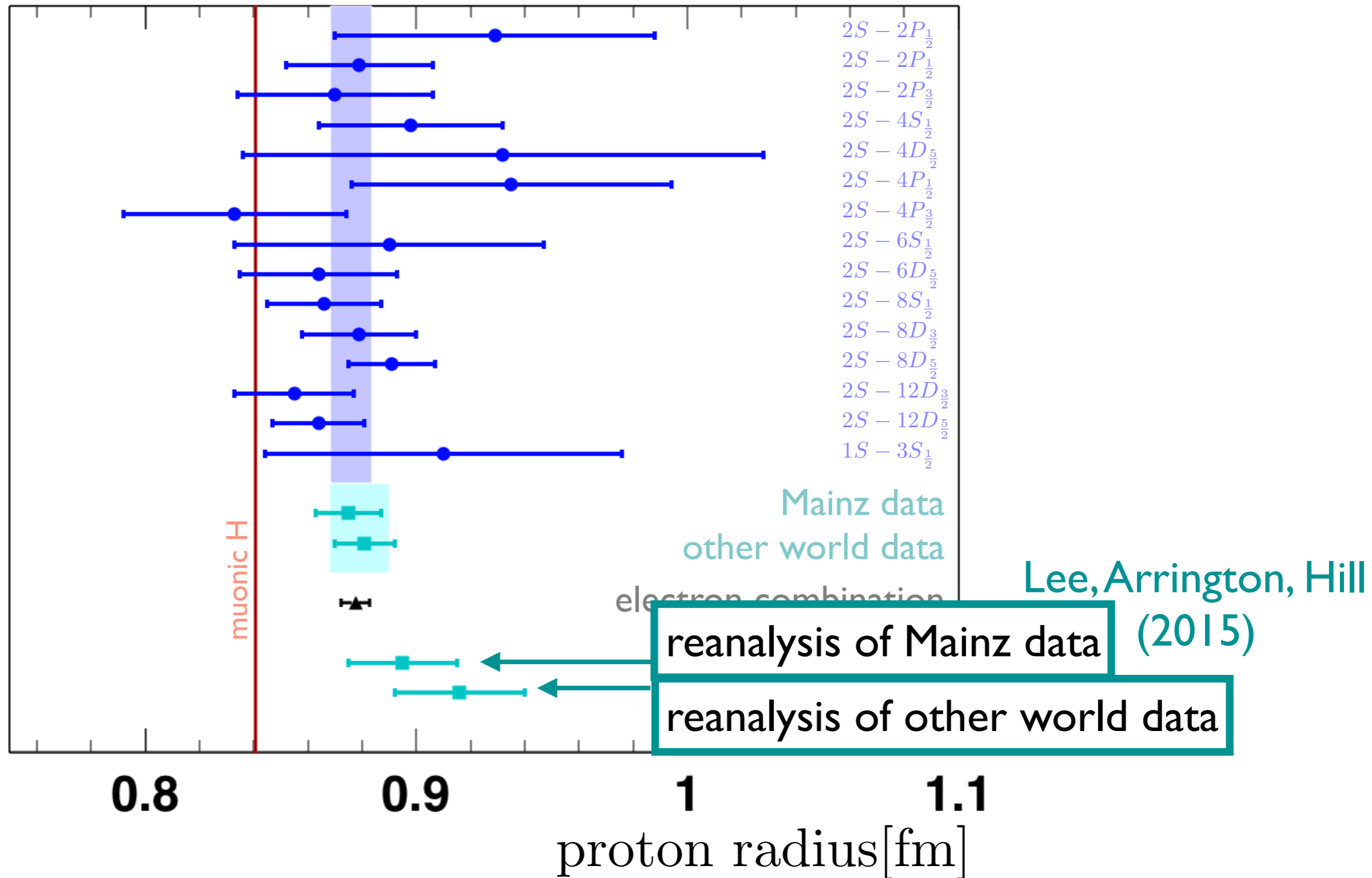
coefficients in rapidly convergent expansion encode nonperturbative QCD

Reanalysis of scattering data reveals strong influence of shape assumptions



Errors larger, but discrepancy remains

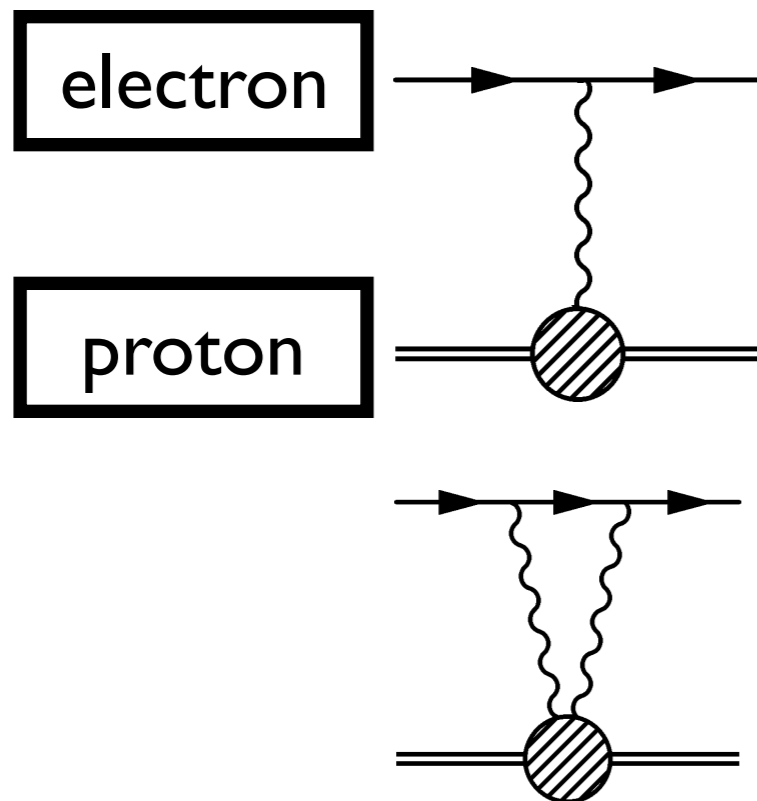
Reanalysis of scattering data reveals strong influence of shape assumptions



Errors larger, but discrepancy remains

muonic hydrogen spectroscopy: theory issues

muonic atoms more sensitive to radius, but also more sensitive to other proton structure

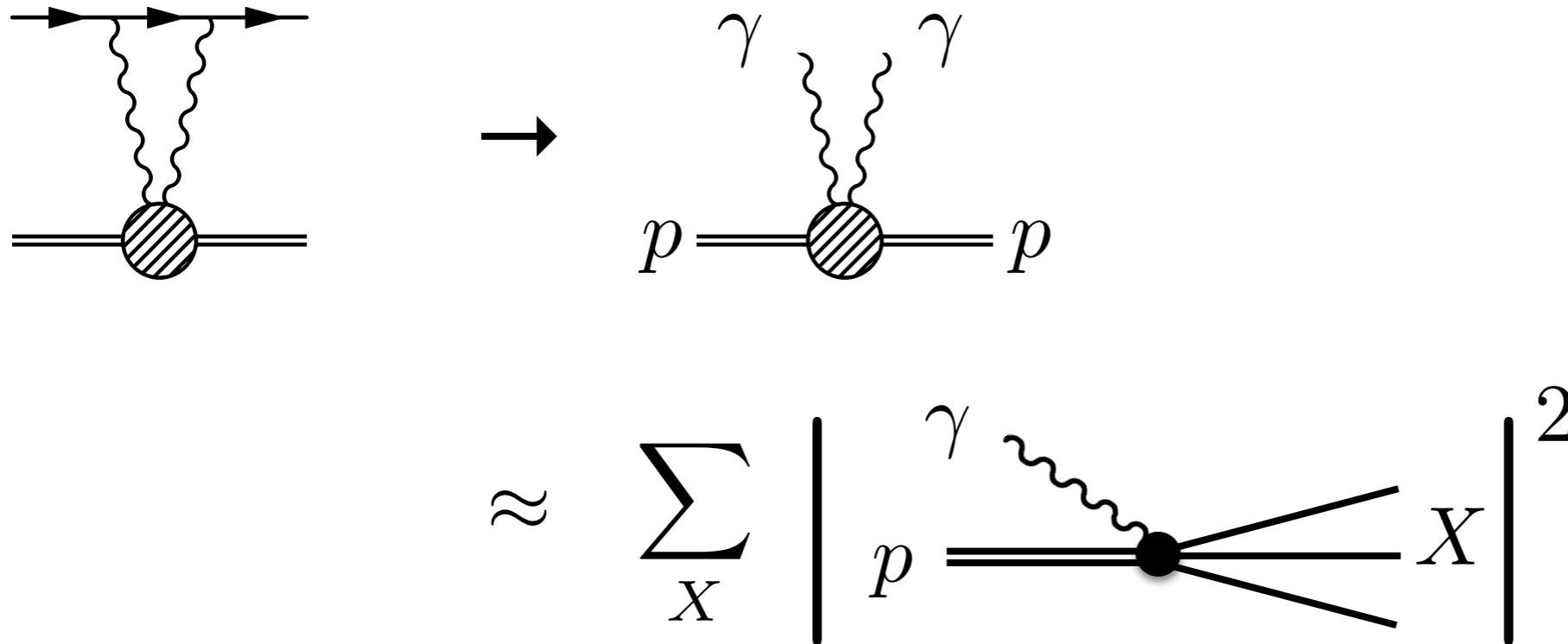


one-photon exchange → radius “signal”

two-photon exchange → “background”

- are subleading proton structure effects under control?

Optical theorem for two-photon exchange in muonic hydrogen



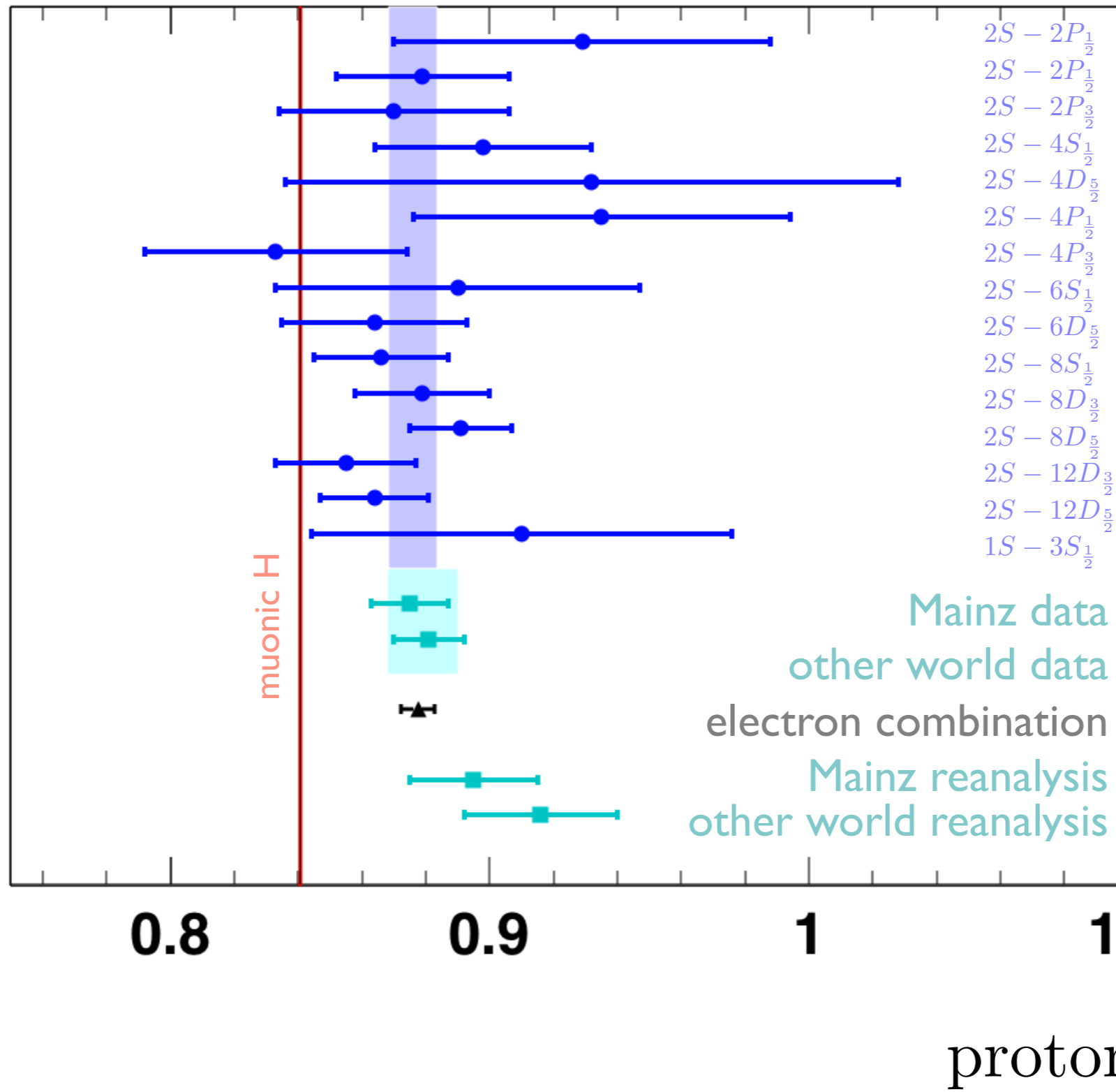
If a dispersion relation is valid, contribution completely determined by measurable quantities in electron-proton scattering. But:

$$W_1(\nu, 0) = -2 + \mathcal{O}(\nu^2) = \frac{1}{\pi} \int d\nu'^2 \frac{\text{Im}W_1(\nu', 0)}{\nu'^2 - \nu^2} > 0 \quad ??$$

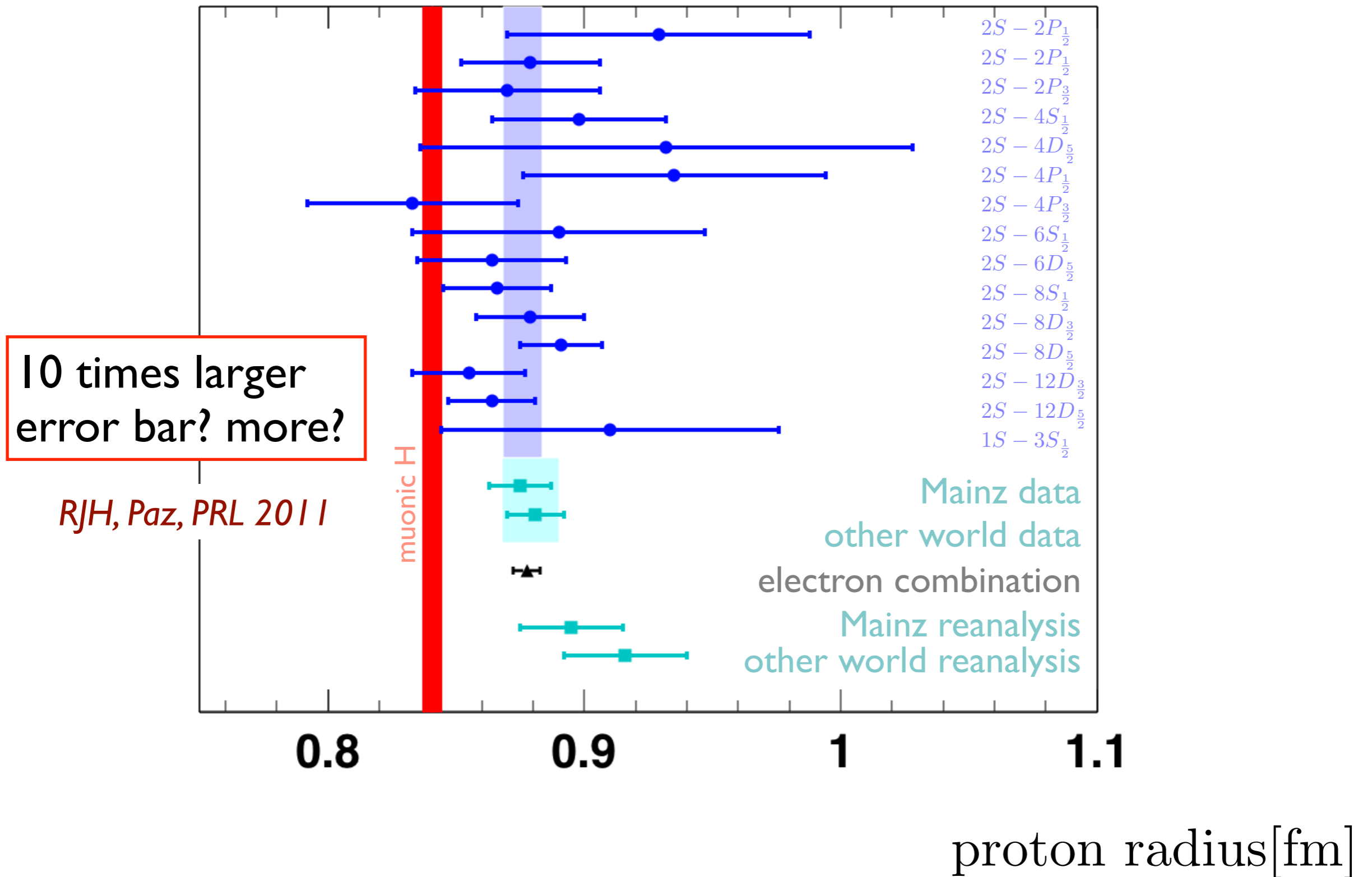
$$\Rightarrow W_1(\nu, Q^2) = W_1(0, Q^2) + \frac{\nu^2}{\pi} \int d\nu'^2 \frac{\text{Im}W_1(\nu', 0)}{\nu'^2(\nu'^2 - \nu^2)}$$

new hadronic function

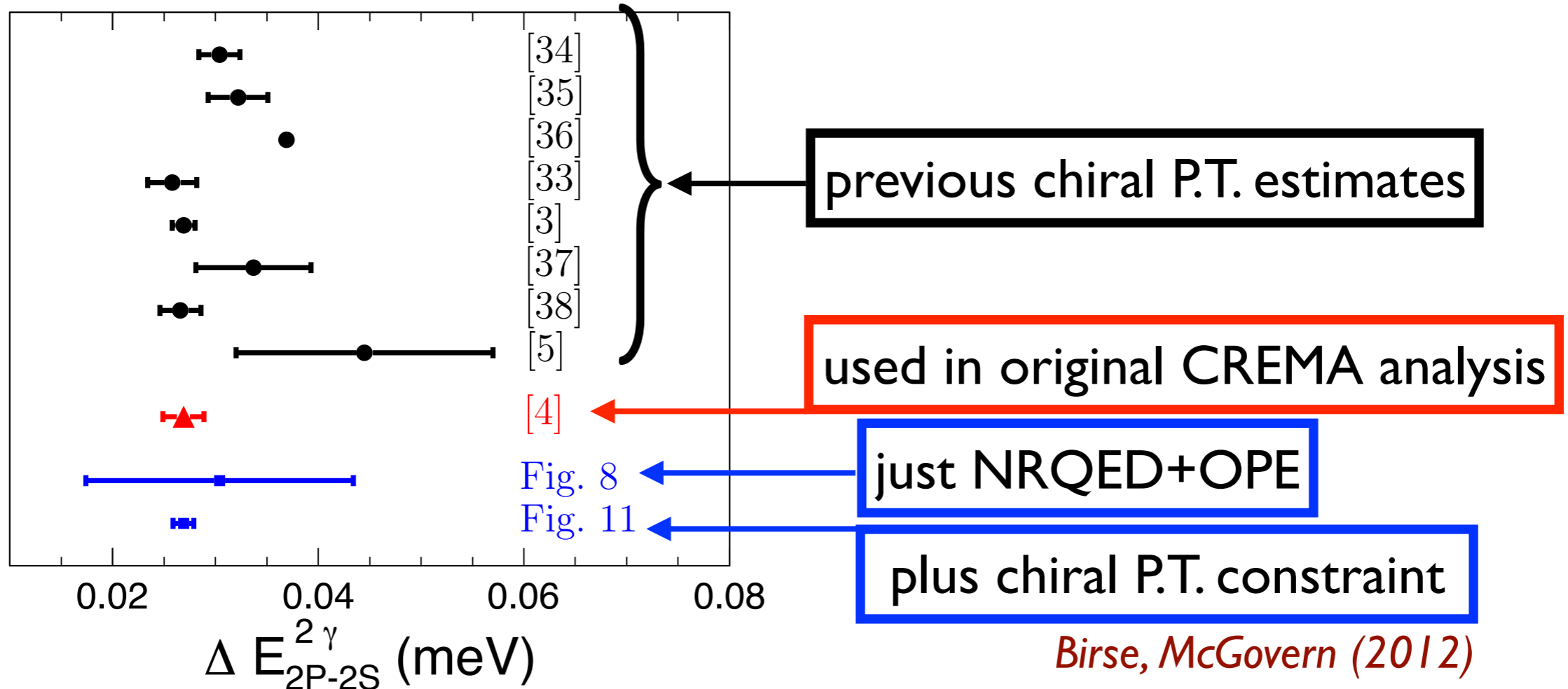
Model-dependent assumptions on two-photon exchange



Model-dependent assumptions on two-photon exchange

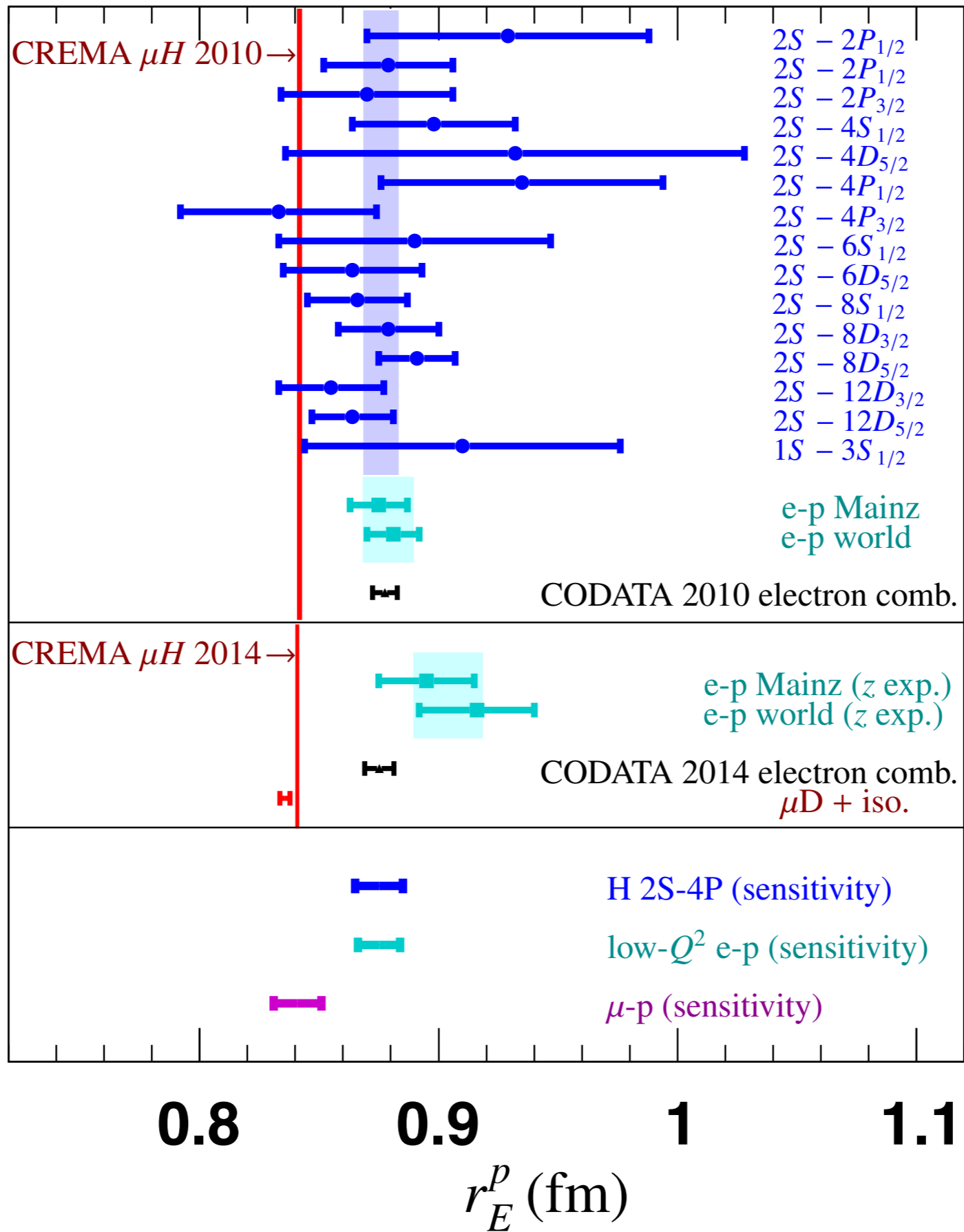


from RJH, Paz (2016)



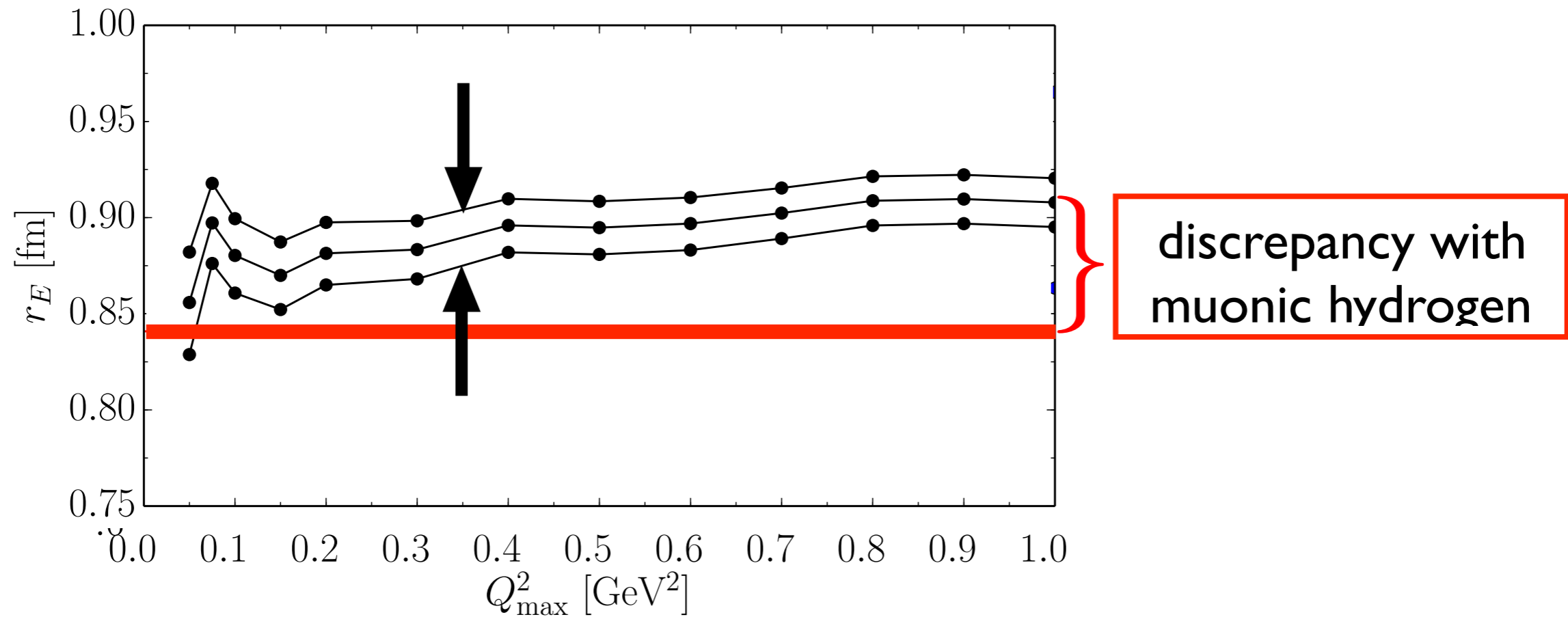
- *TPE remains dominant uncertainty in mu-H Lamb shift, rE_p , Rydberg (most precise fundamental constant). But not a solution to proton radius puzzle*

status and prospects



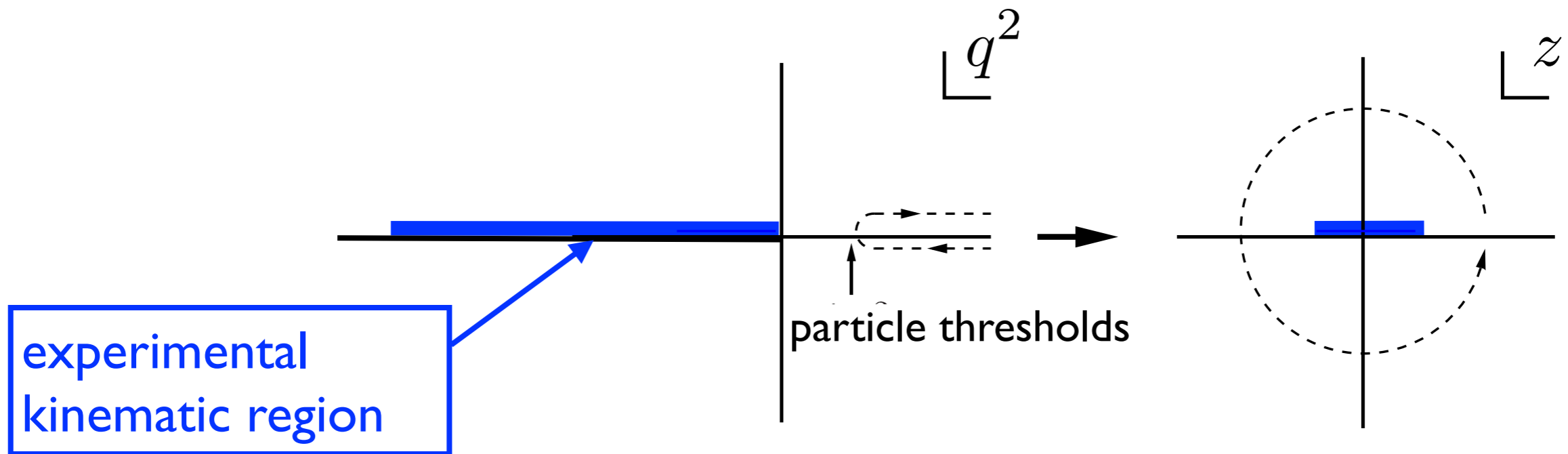
something else?

New physics under our noses?



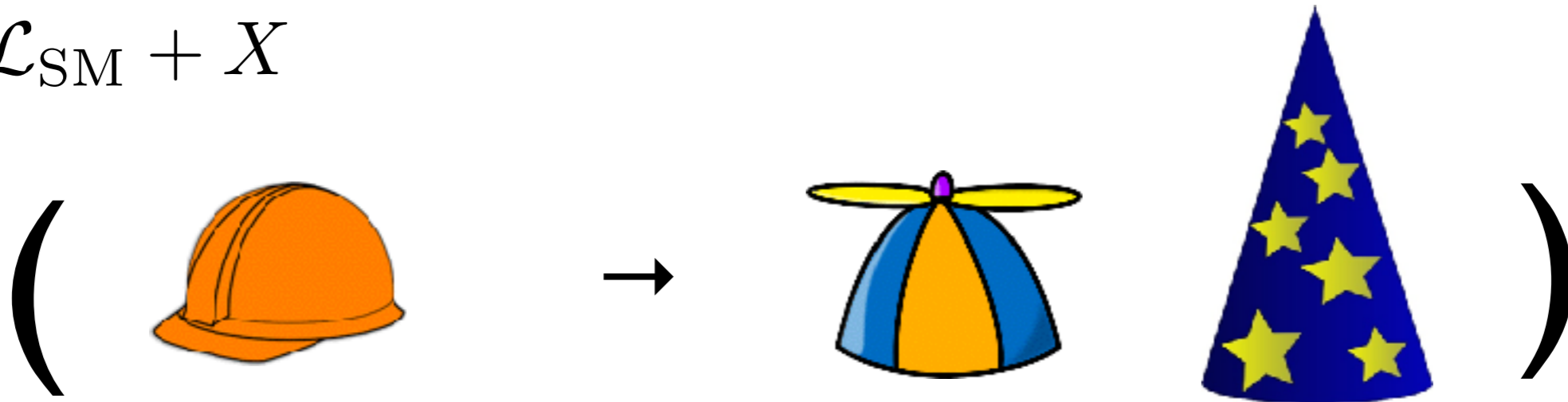
A new particle would violate the assumed analytic structure, and generate momentum-dependent effect

New physics under our noses?



A new particle would violate the assumed analytic structure, and generate momentum-dependent effect

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + X$$

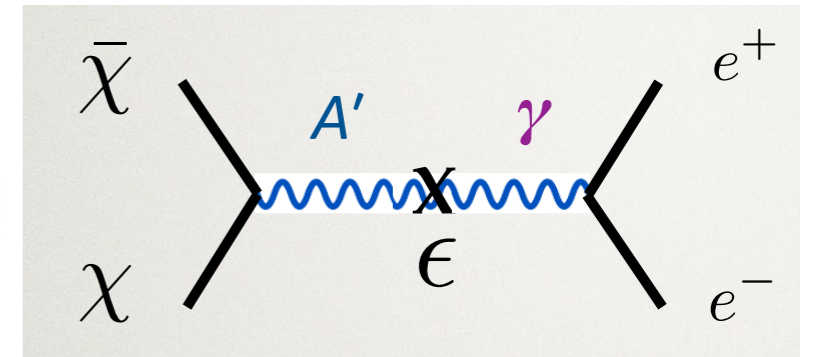
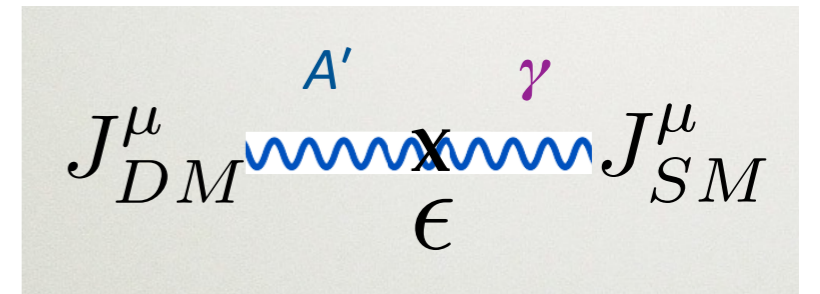
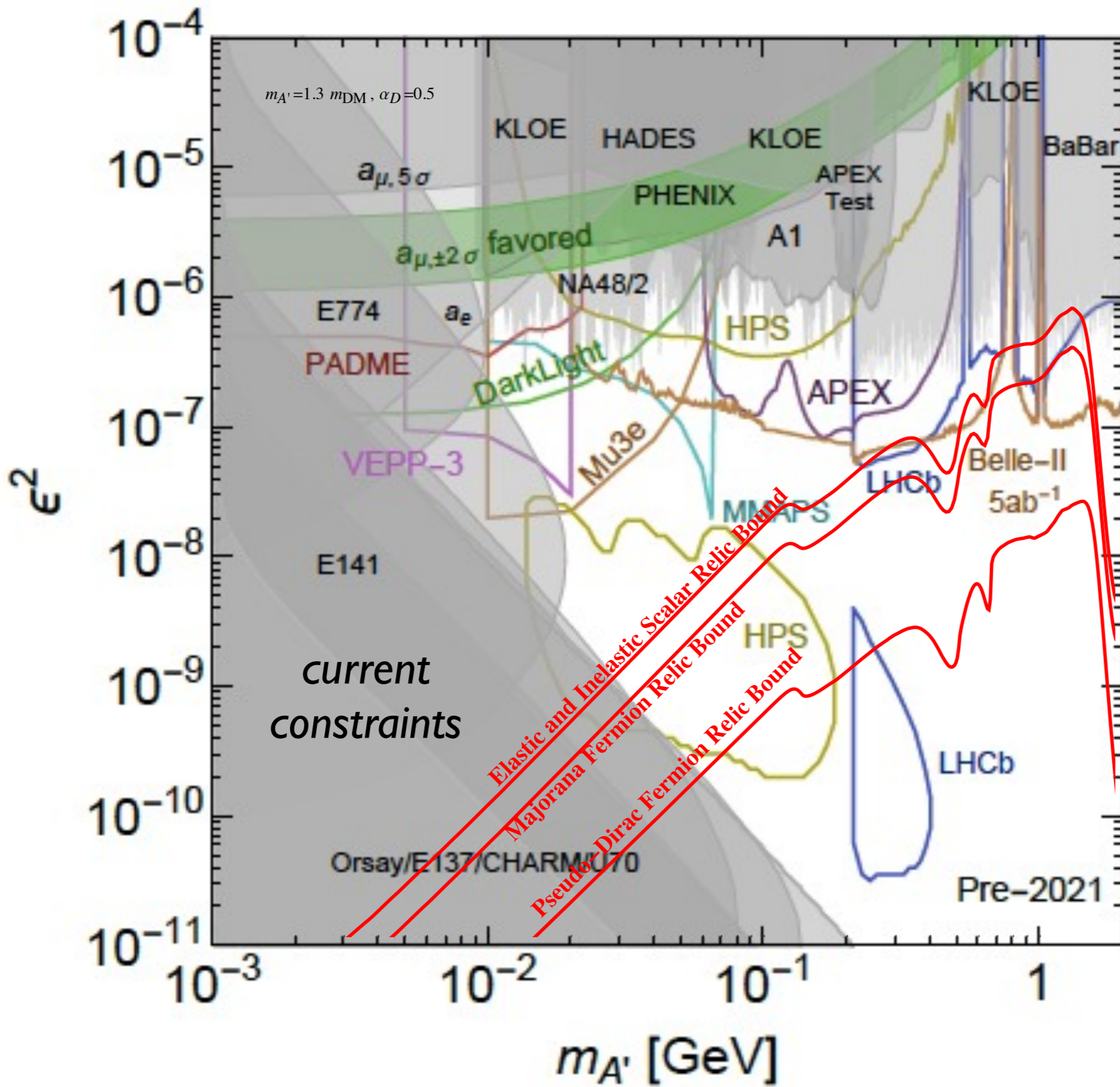


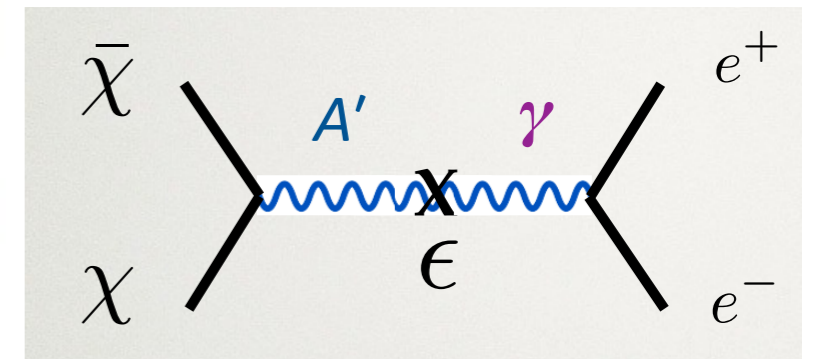
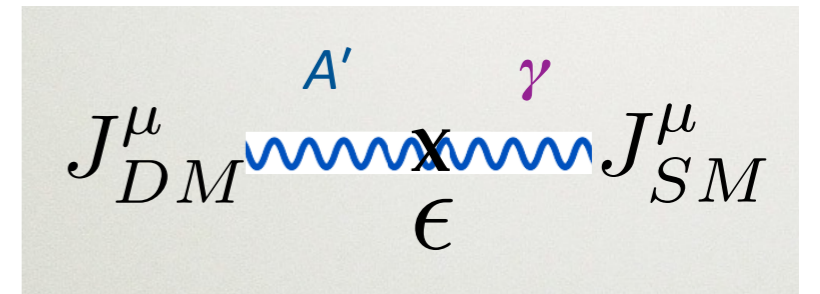
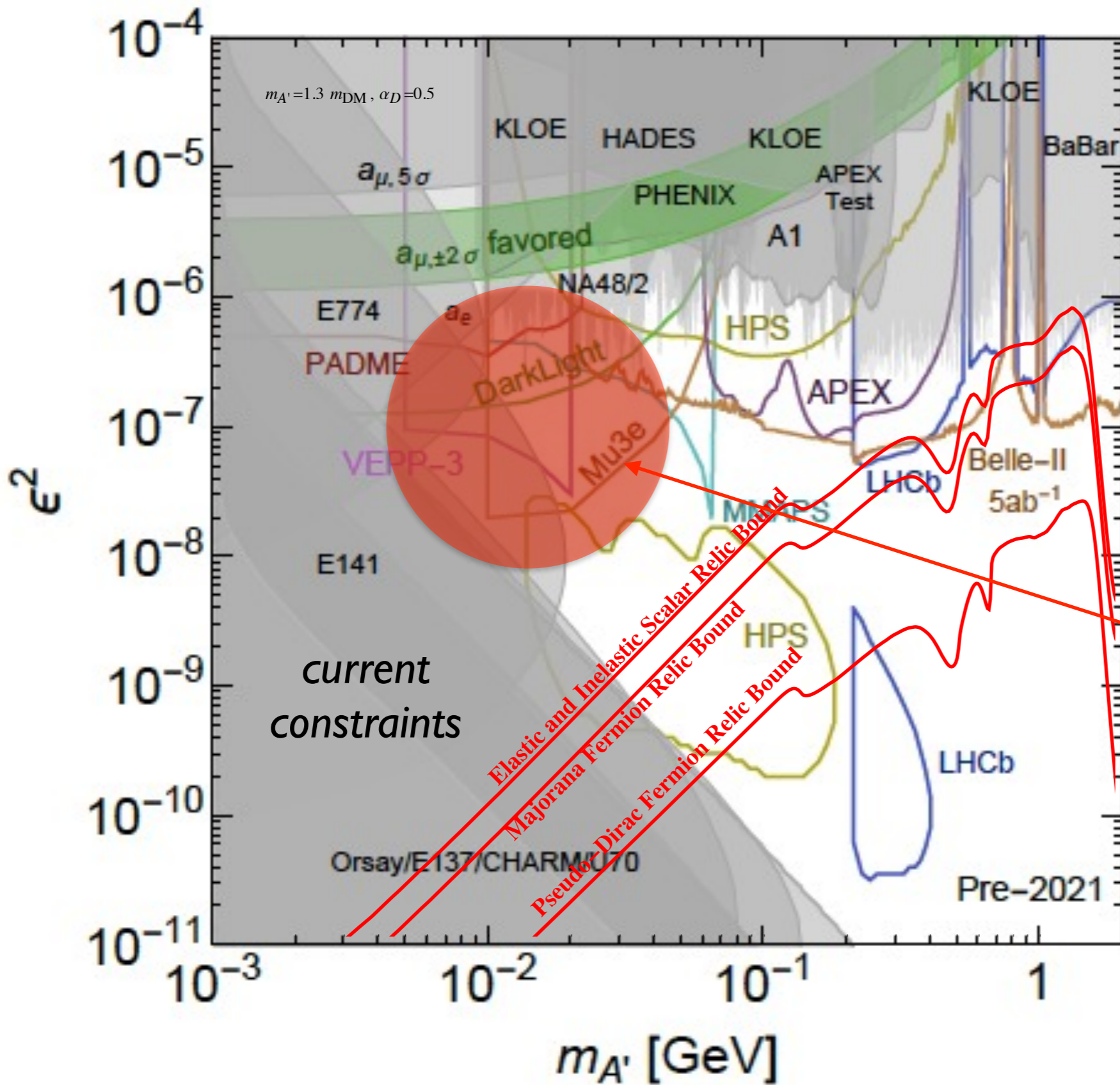
one possibility: $X = \text{dark photon}$

- depending on mass, consistent with $r_{eH} \sim r_{\mu H} < r_{e-p}$

$$-\frac{e^2}{Q^2} F(Q^2) \rightarrow -\frac{e^2}{Q^2} F(Q^2) \mp \frac{g^2}{Q^2 + m_V^2}$$

- especially interesting to see new eH results





interesting region for proton radius puzzle

summary

- proton radius puzzle, neutrinos, dark matter: important particle, nuclear, atomic overlap
- impossible not to cross boundaries
- Operator product expansion: Heavy WIMP direct detection \leftrightarrow two-photon effects in Lamb shifts
- Sudakov resummation: Heavy WIMP annihilation \leftrightarrow e-p, ν -N scattering
- need for precision in current and next generation experiments
- opportunity to develop and exploit modern tools and technology

Dark matter - Standard Model interactions

$$\mathcal{L} = \frac{1}{\Lambda^n} O_{\text{DM}} \times O_{\text{SM}}$$

d	Fermion	d	Scalar	d	Heavy particle
3	$\bar{\psi}[1, i\gamma_5, \gamma^\mu\gamma_5, \{\gamma^\mu, \sigma^{\mu\nu}\}]\psi$	2	$ \phi ^2$	3	$\bar{\chi}_v[1, \{\sigma_{\perp}^{\mu\nu}\}]\chi_v$
4	$\bar{\psi}[\{1, i\gamma_5, \gamma^\mu\gamma_5\}, \gamma^\mu, \sigma^{\mu\nu}]i\partial_{\perp}^{\rho}\psi$	3	$\{\phi^*i\partial_{\perp}^{\mu}\phi\}$	4	$\bar{\chi}_v[\{1\}, \sigma_{\perp}^{\mu\nu}]i\partial_{\perp}^{\rho}\chi_v$

d	QCD operator basis
3	$V_q^{\mu} = \bar{q}\gamma^{\mu}q$ $A_q^{\mu} = \bar{q}\gamma^{\mu}\gamma_5q$
4	$T_q^{\mu\nu} = im_q\bar{q}\sigma^{\mu\nu}\gamma_5q$ $O_q^{(0)} = m_q\bar{q}q, \quad O_g^{(0)} = G_{\mu\nu}^A G^{A\mu\nu}$ $O_{5q}^{(0)} = m_q\bar{q}i\gamma_5q, \quad O_{5g}^{(0)} = \epsilon^{\mu\nu\rho\sigma}G_{\mu\nu}^A G_{\rho\sigma}^A$ $O_q^{(2)\mu\nu} = \frac{1}{2}\bar{q}\left(\gamma^{\{\mu}iD_{\perp}^{\nu\}} - \frac{g^{\mu\nu}}{4}i\not{D}_{\perp}\right)q, \quad O_g^{(2)\mu\nu} = -G^{A\mu\lambda}G^{A\nu}_{\lambda} + \frac{g^{\mu\nu}}{4}(G_{\alpha\beta}^A)^2$ $O_{5q}^{(2)\mu\nu} = \frac{1}{2}\bar{q}\gamma^{\{\mu}iD_{\perp}^{\nu\}}\gamma_5q$

complete
QCD basis
for $d \leq 7$

Renormalization and matching (sample):

$$\mathcal{L}_{\phi_0, \text{SM}} = \frac{1}{m_W^3} \phi_v^* \phi_v \left\{ \sum_q \left[c_{1q}^{(0)} O_{1q}^{(0)} + c_{1q}^{(2)} v_\mu v_\nu O_{1q}^{(2)\mu\nu} \right] + c_2^{(0)} O_2^{(0)} + c_2^{(2)} v_\mu v_\nu O_2^{(2)\mu\nu} \right\} + \dots$$

$m_q \bar{q}q$: $G_{\mu\nu}^A G^{A\mu\nu}$

focus on spin-0 (evaluate spin-2 at weak scale)

Renormalization group evolution from weak scale to hadronic scales, with perturbative corrections at heavy quark mass thresholds

$$c_i(\mu_Q) = M_{ij}(\mu_Q) c'_j(\mu_Q).$$

$$M(\mu_Q) = \left(\begin{array}{ccc|cc} & & & M_{qQ} & M_{qg} \\ & \mathbb{1}(M_{qq} - M_{qq'}) + \mathcal{J}M_{qq'} & & \vdots & \vdots \\ & & & M_{qQ} & M_{qg} \\ \hline M_{gq} & \dots & M_{gq} & M_{gQ} & M_{gg} \end{array} \right)$$

Can show that:

$$M_{qq} \equiv 1, \quad M_{qq'} \equiv 0, \quad M_{gq} \equiv 0$$

M_{gQ} and M_{qQ} known through

3 loops:

Chetyrkin et al. (1997)

New results for gluon-induced decoupling relations

$$M_{gg}^{(2)} = \frac{11}{36} - \frac{11}{6} \log \frac{\mu_Q}{m_Q} + \frac{1}{9} \log^2 \frac{\mu_Q}{m_Q}$$

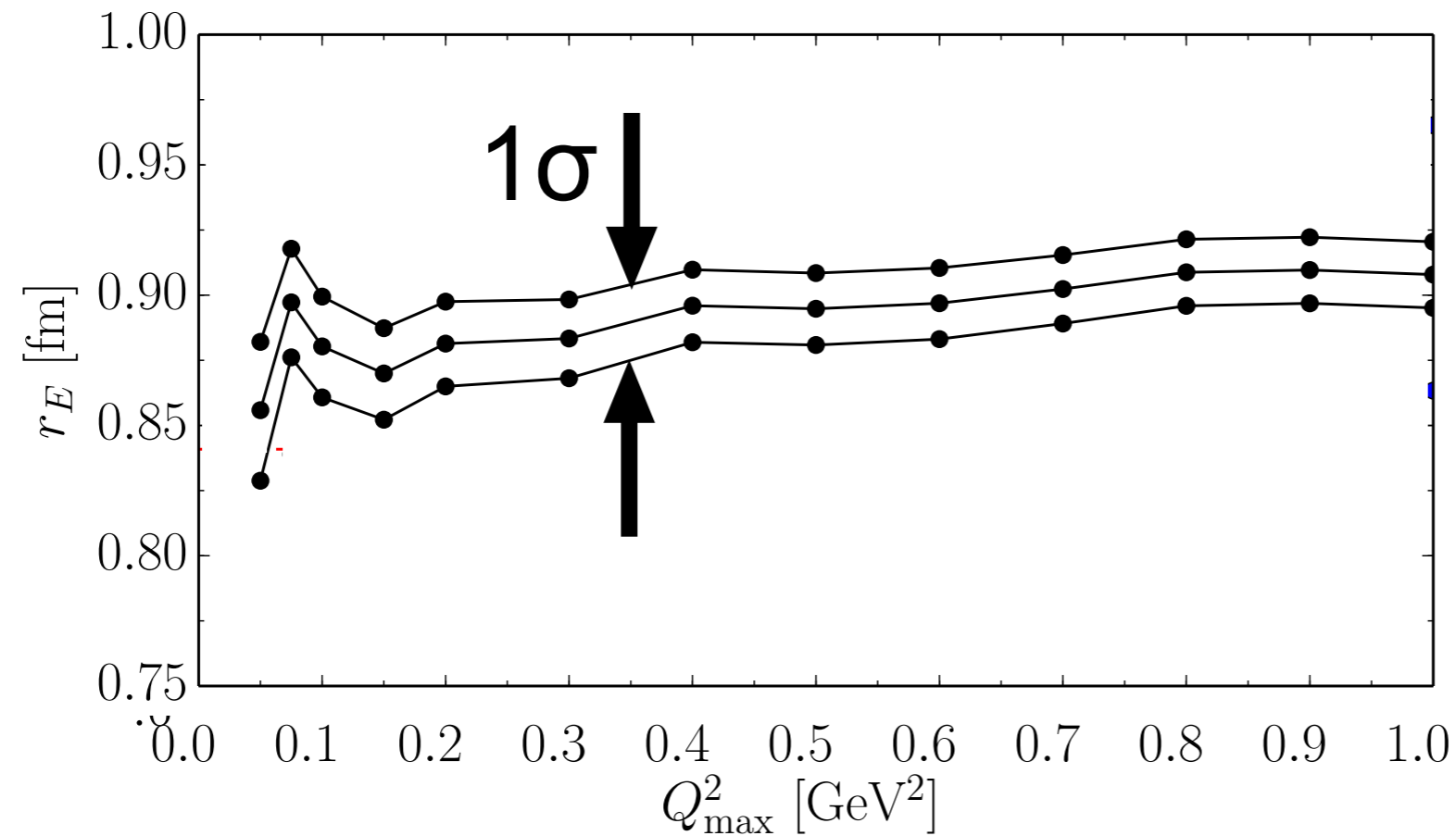
$$M_{gg}^{(3)} = \frac{564731}{41472} - \frac{2821}{288} \log \frac{\mu_Q}{m_Q} + \frac{3}{16} \log^2 \frac{\mu_Q}{m_Q} - \frac{1}{27} \log^3 \frac{\mu_Q}{m_Q} - \frac{82043}{9216} \zeta(3) \\ + n_f \left[-\frac{2633}{10368} + \frac{67}{96} \log \frac{\mu_Q}{m_Q} - \frac{1}{3} \log^2 \frac{\mu_Q}{m_Q} \right],$$

$$M_{qg}^{(2)} = -\frac{89}{54} + \frac{20}{9} \log \frac{\mu_Q}{m_Q} - \frac{8}{3} \log^2 \frac{\mu_Q}{m_Q}.$$

Hill, Solon (2014)



Reanalysis of scattering data also reveals potential dependence of radius on chosen Q^2 range



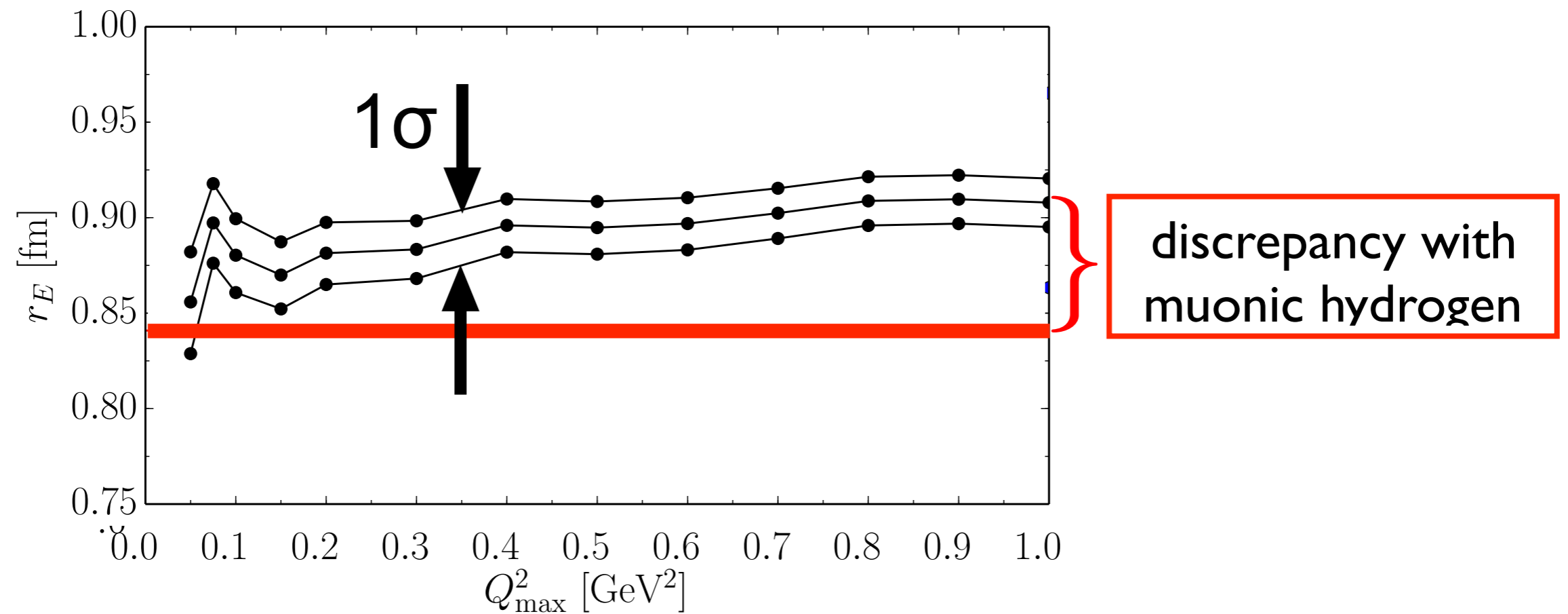
To reconcile e-p scattering with muonic hydrogen, could:

- consider only small Q^2 data (less data \Rightarrow larger error)
- overrule scattering data with other data or constraints

These options could avoid, but not resolve, the puzzle from electron scattering.

*An unaccounted effect impacting especially large Q^2 data?
(radiative corrections: another talk. new physics: another talk)*

Reanalysis of scattering data also reveals potential dependence of radius on chosen Q^2 range



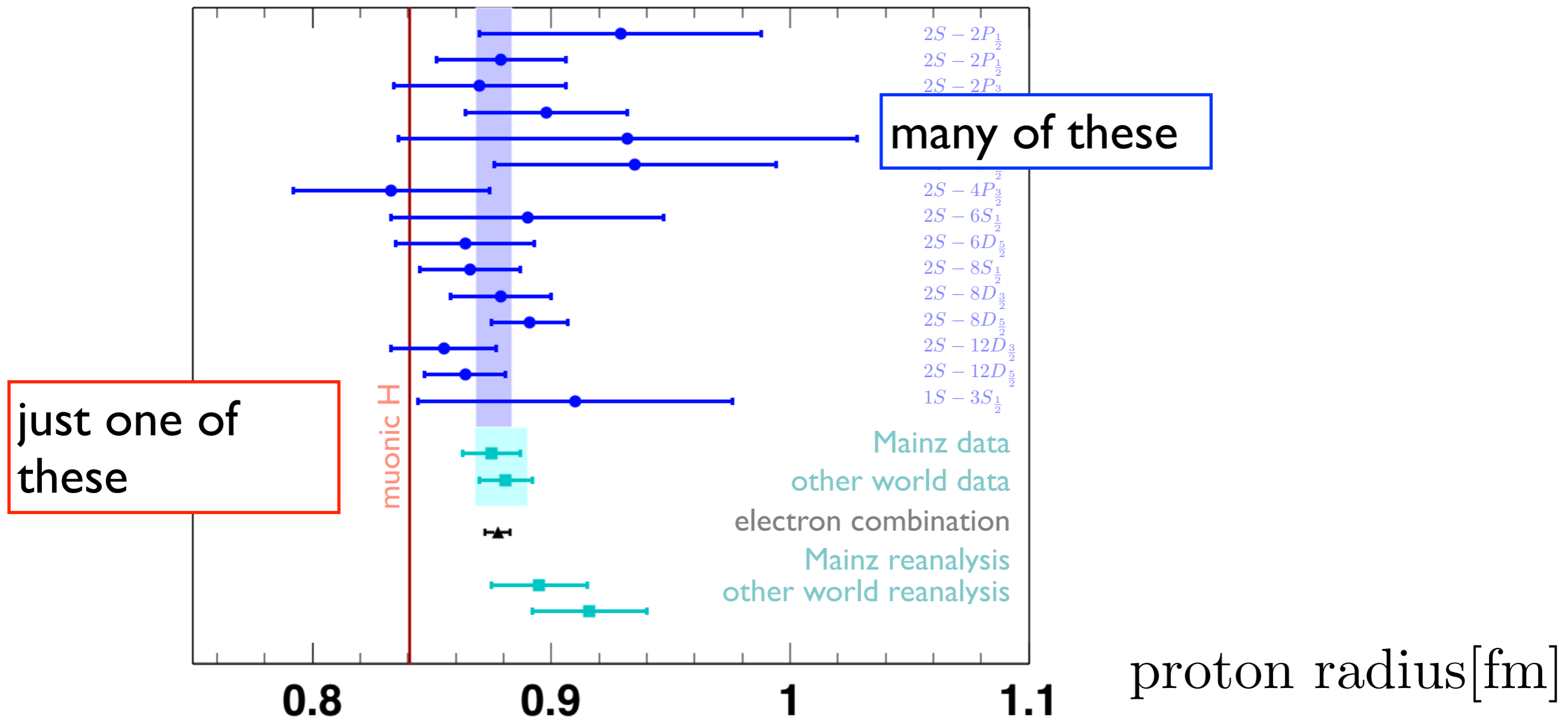
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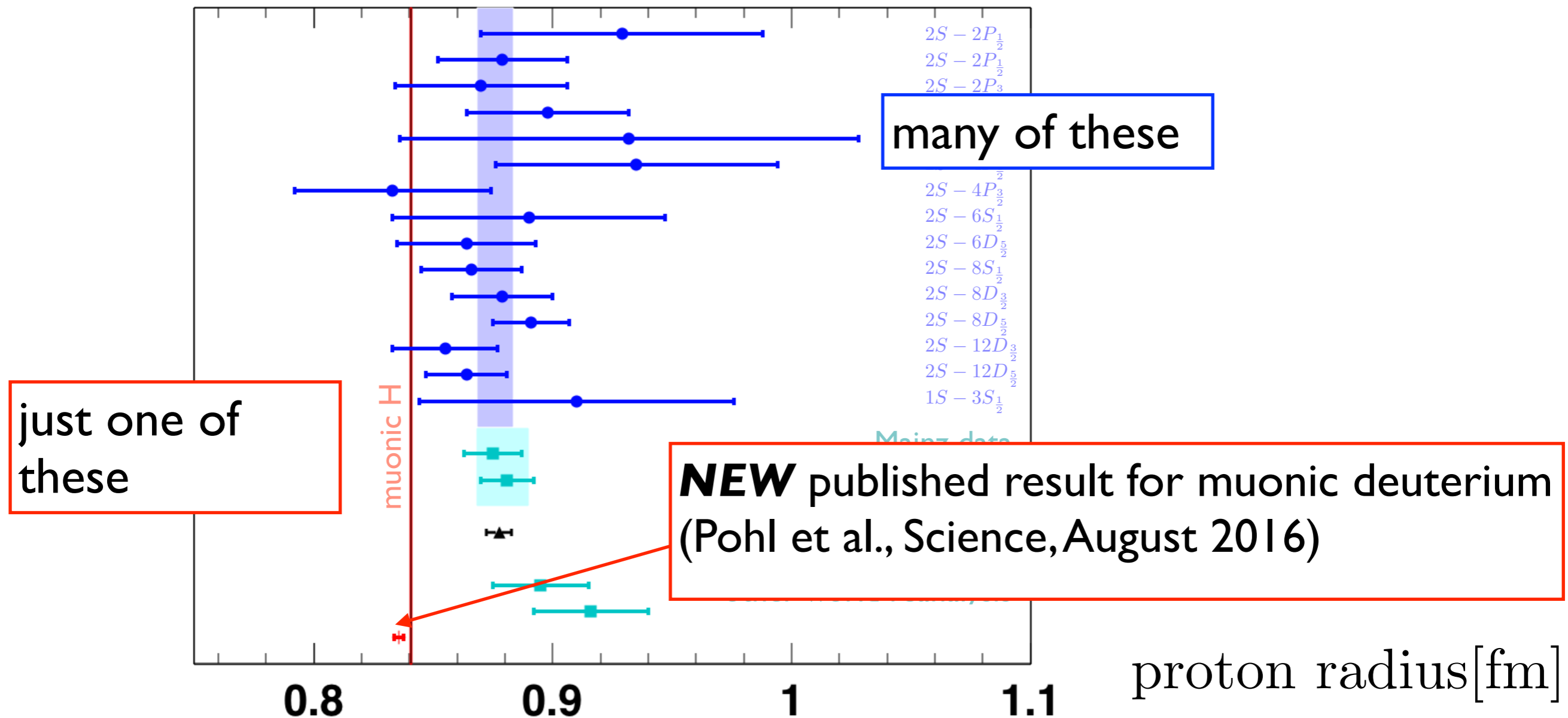
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spectroscopy of other light muonic atoms: D, He

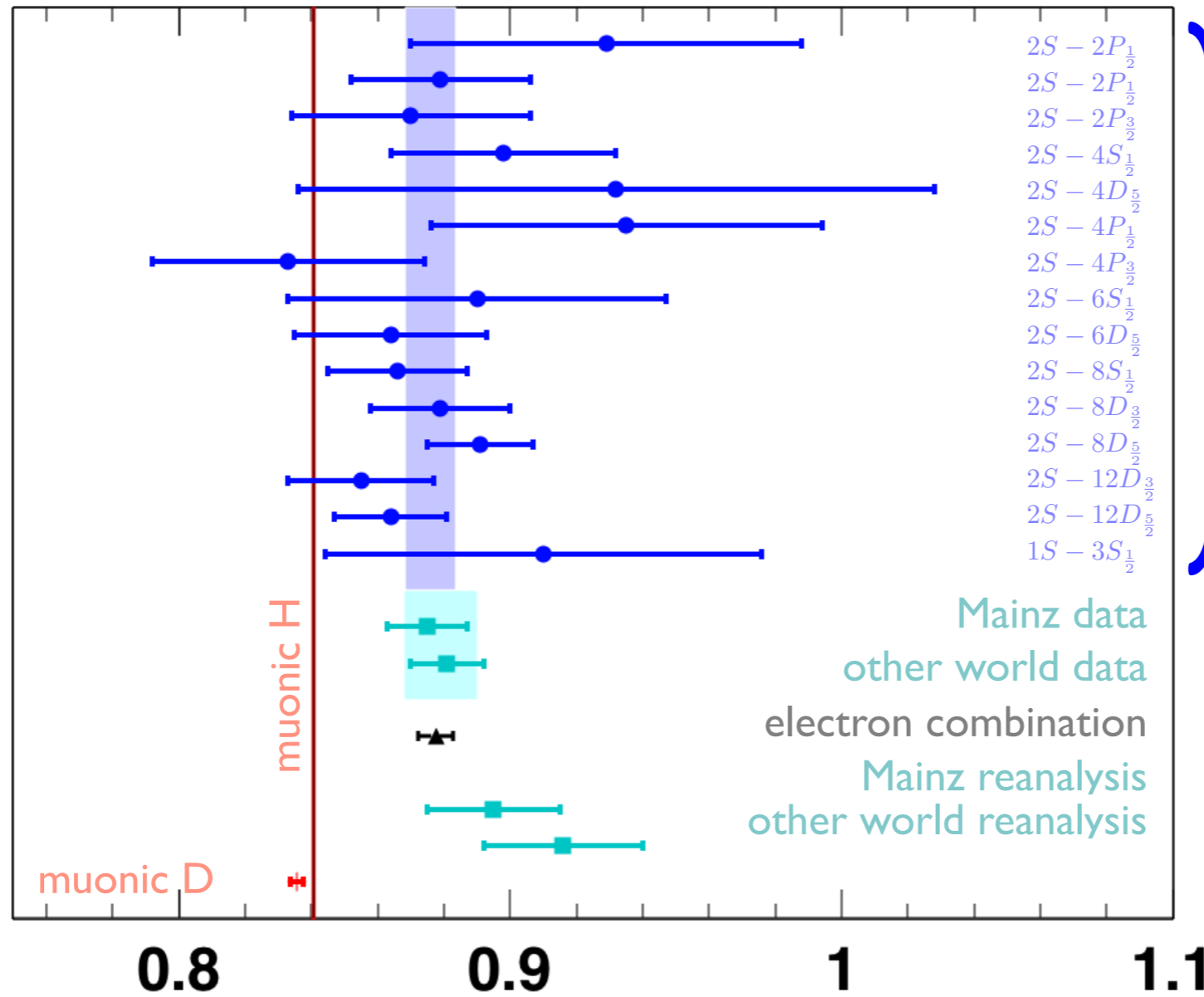


spectroscopy of other light muonic atoms: D, He



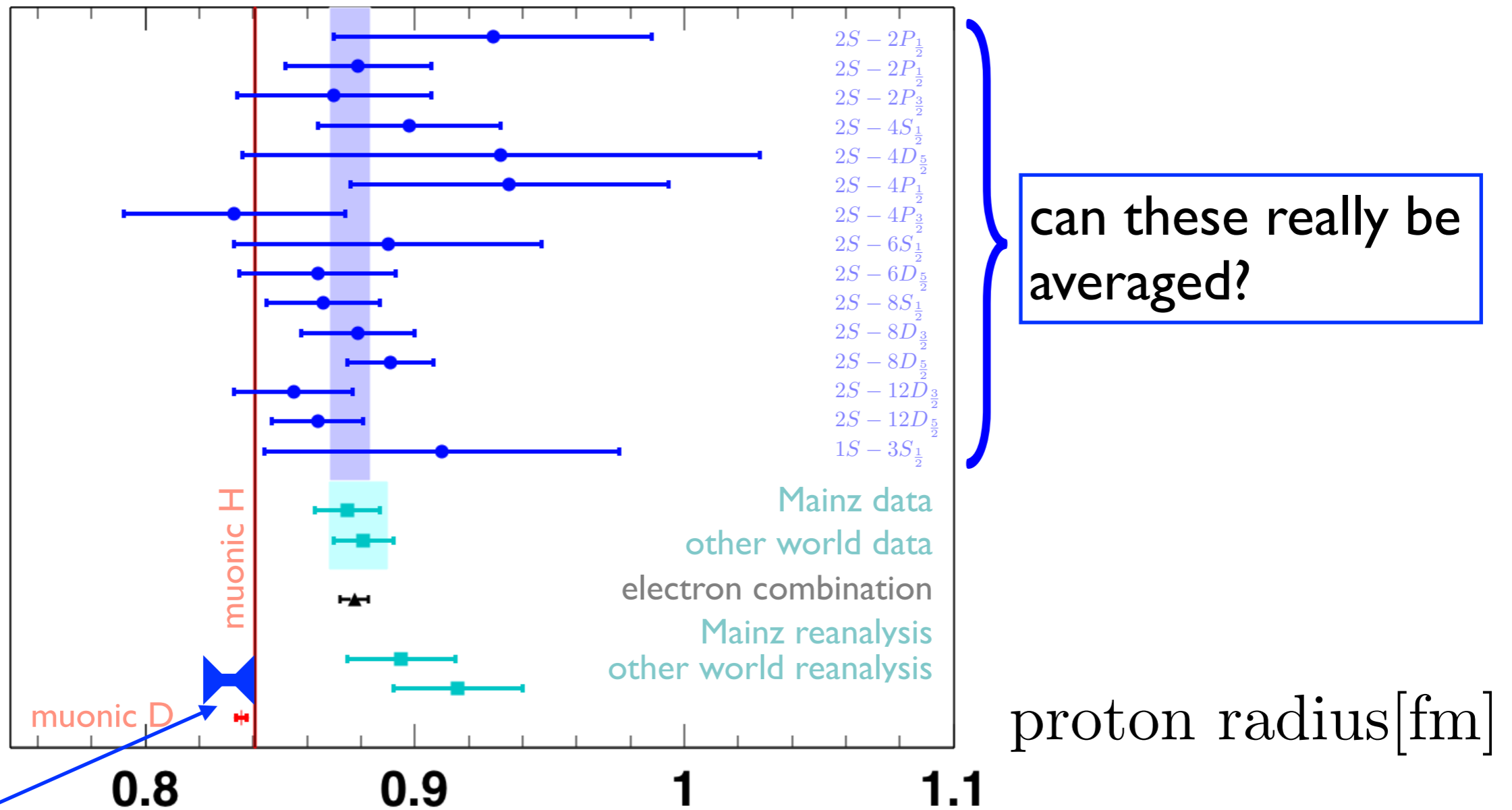
- small discrepancy with muonic hydrogen
- large discrepancy with hydrogen-only radius
- new results also anticipated with muonic helium: *theory improvement needed for nuclear structure corrections*

new (preliminary) hydrogen spectroscopy results



can these really be averaged?

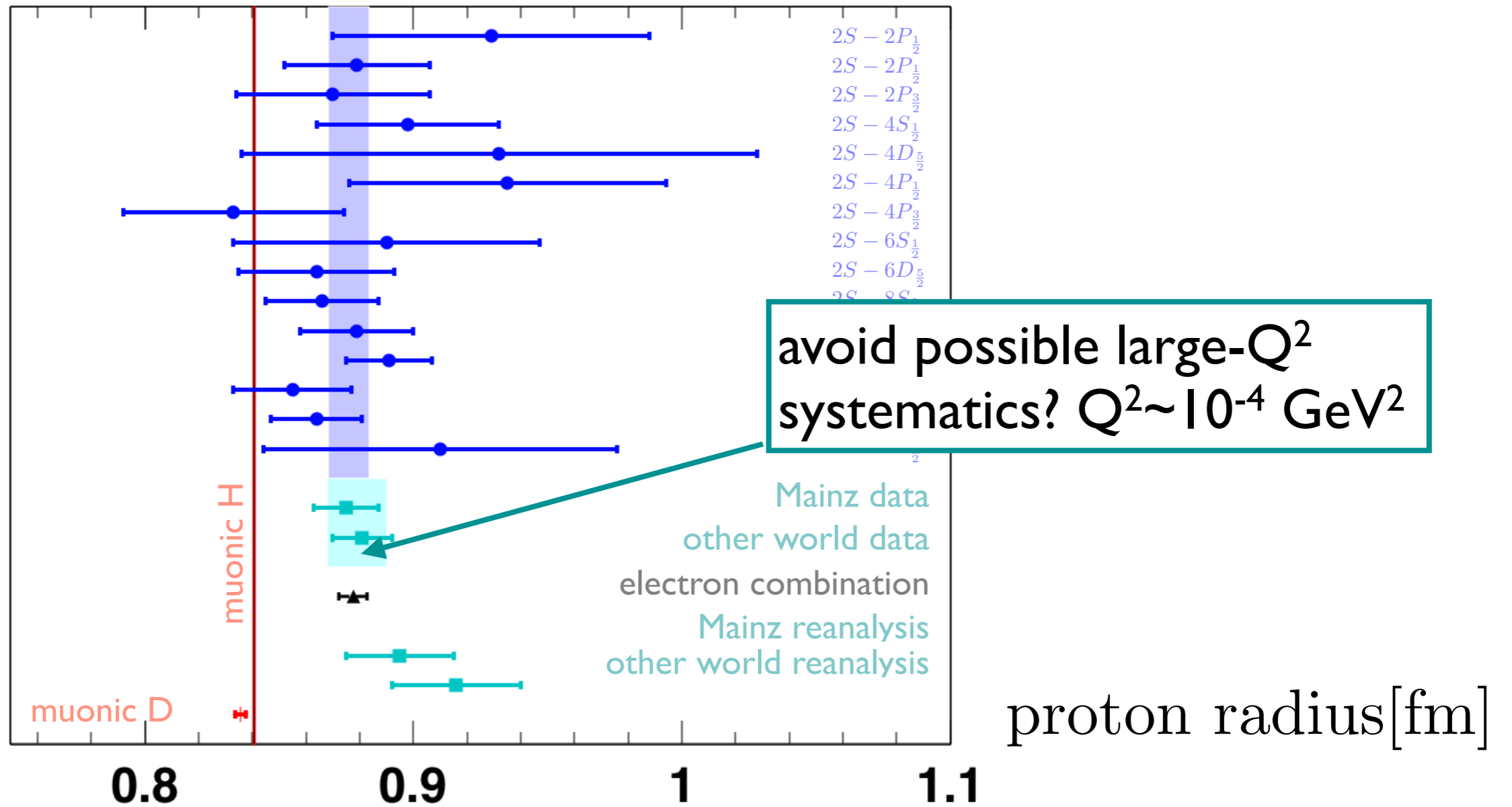
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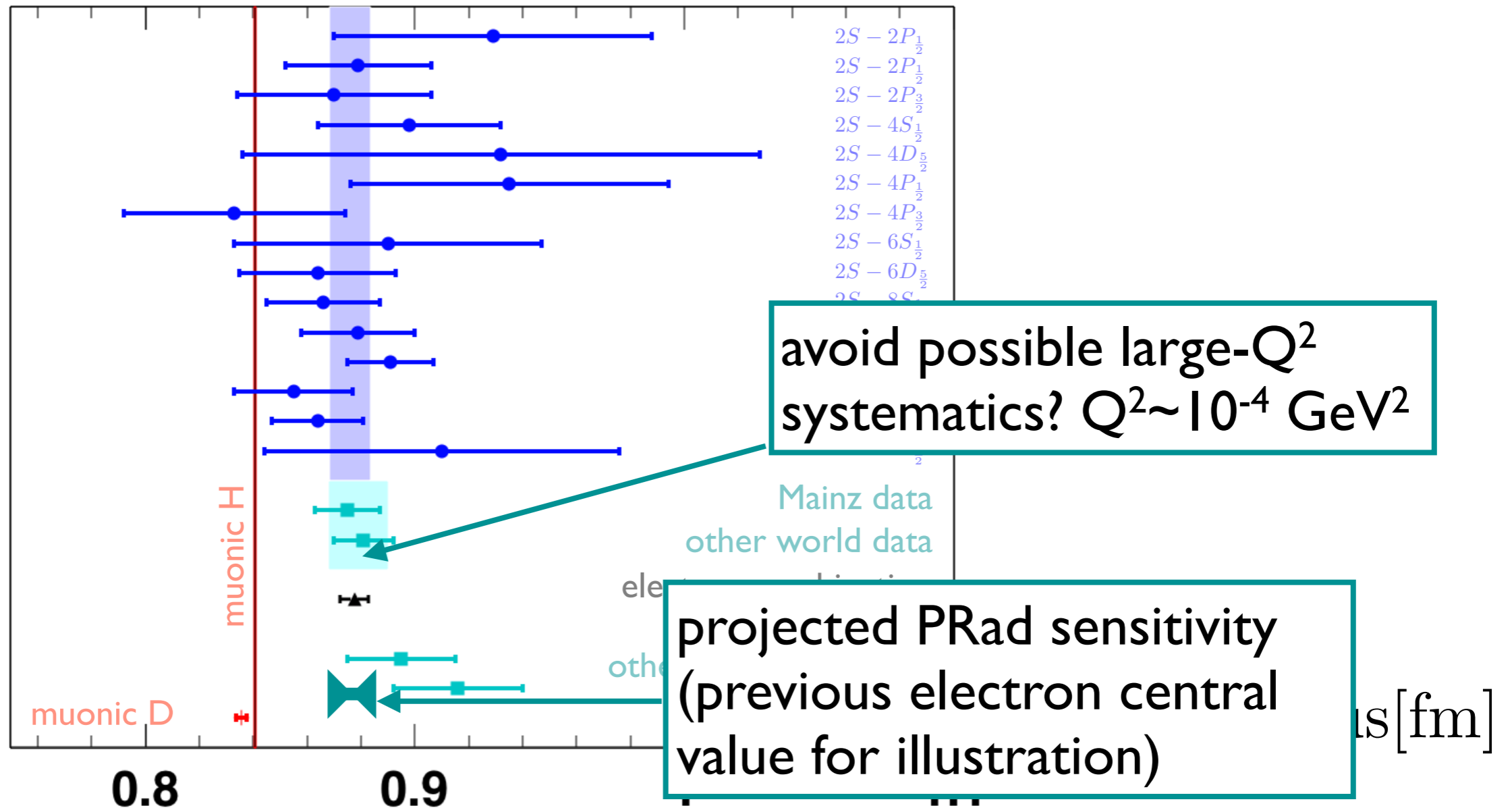
- Beyer, Maisenbacher, Matveev et al. (Garching): result for 2S-4P (submitted). *Error comparable to previous hydrogen average, central value consistent with muonic hydrogen (PRELIMINARY)*

- future new results anticipated from 2S-2P (York), 1S-3S (Paris), others

low- Q^2 electron-proton scattering: PRad at JLab



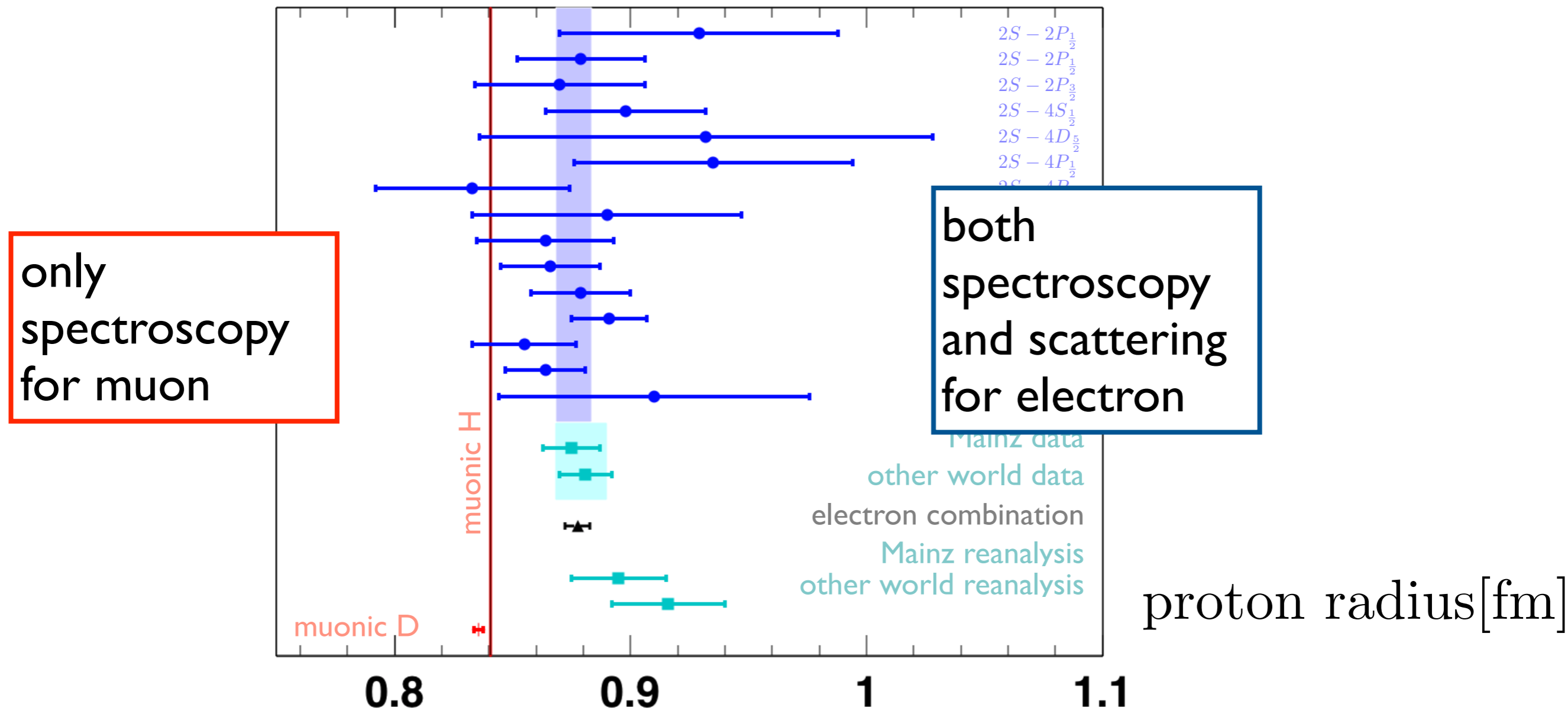
low- Q^2 electron-proton scattering: PRad at JLab



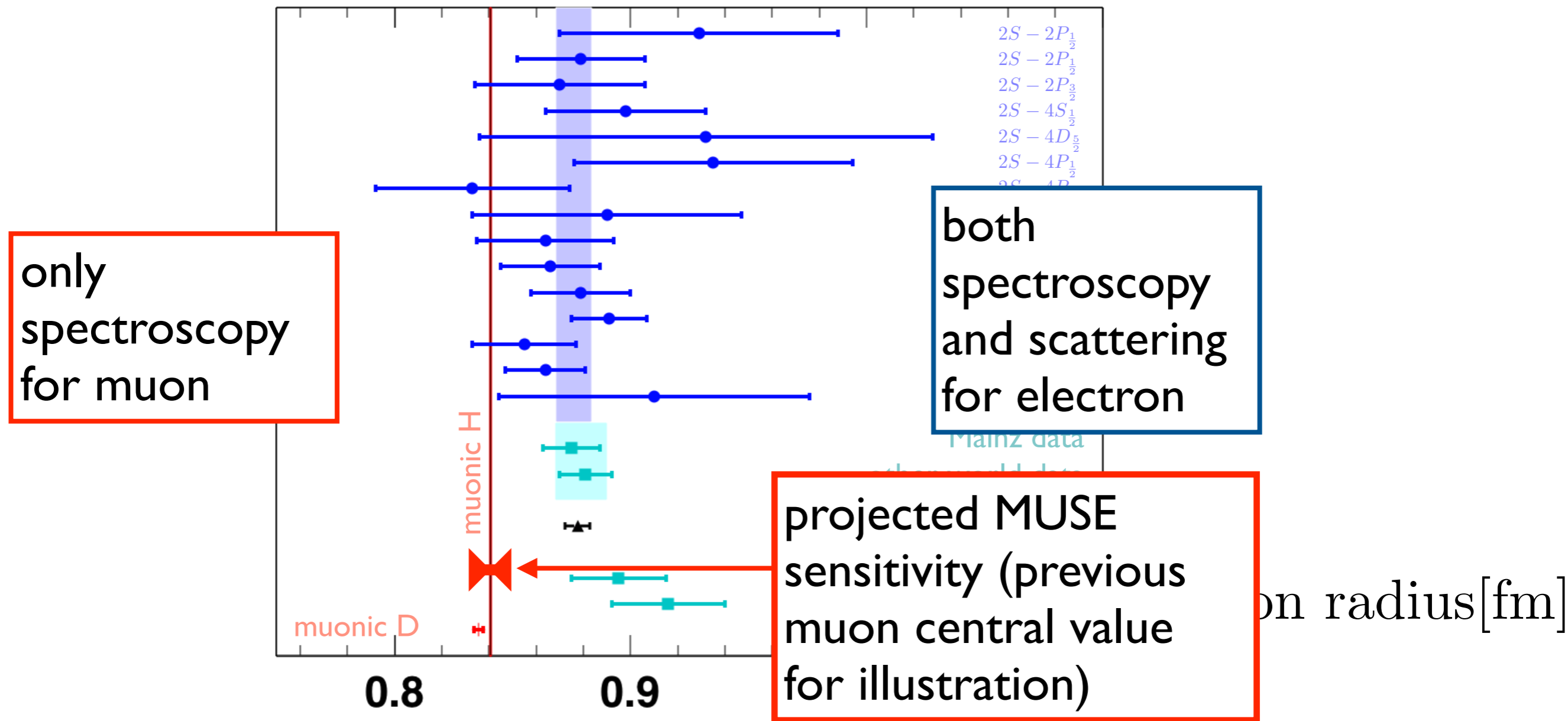
- non-magnetic spectrometer
- simultaneous calibration with e^-e^- (Moller) scattering
- windowless target

data collected in May/June 2016. first analysis

muon-proton scattering: MUSE at PSI



muon-proton scattering: MUSE at PSI



- measurement of e^+, e^-, μ^+, μ^-
- cancellation of systematics & direct two-photon sensitivity

production data-taking scheduled

- The proton radius puzzle has important implications

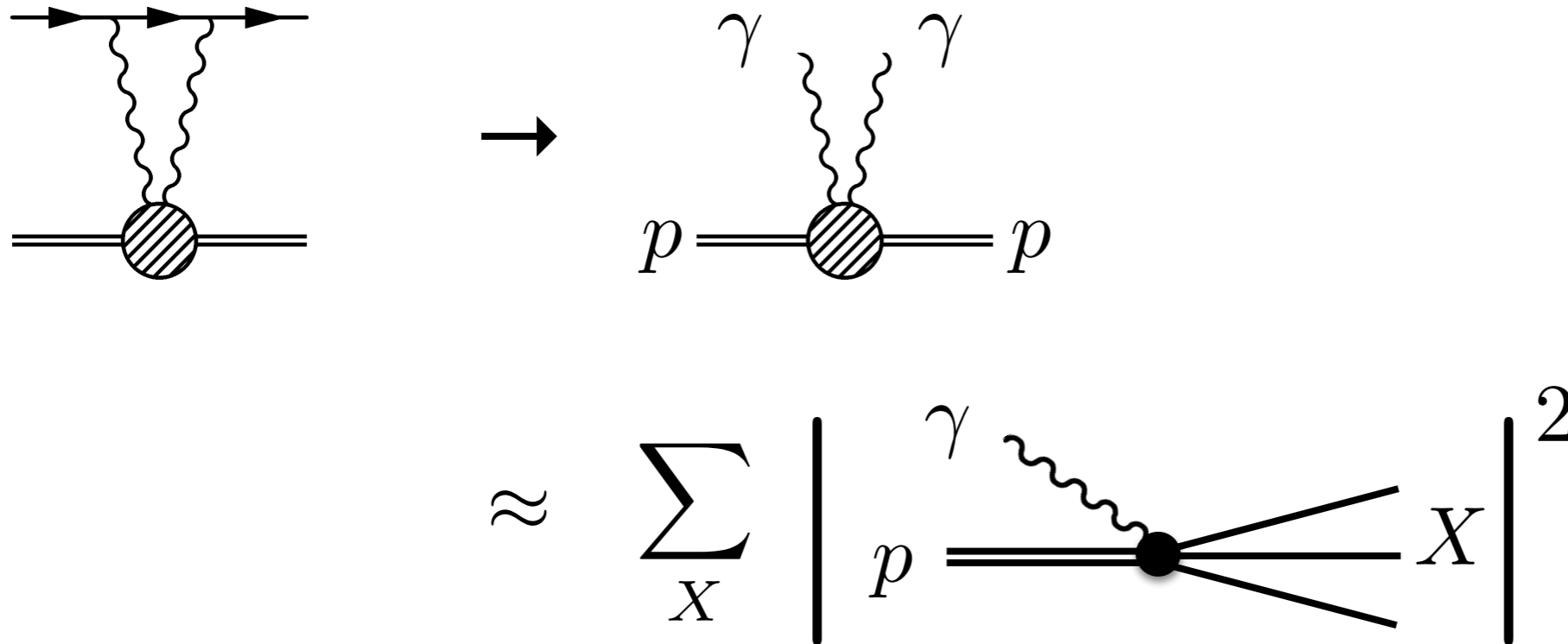
- dramatic shift in fundamental constants, and revising our theoretical understanding of many processes in atomic, nuclear and particle physics

or

- new physics?

- Improved analysis of muonic hydrogen disfavors enhanced two-photon exchange contribution as explanation
- Radiative corrections reanalyzed so far do not reconcile electron scattering results with muonic hydrogen. Work remains.
- The puzzle is motivating many new experiments
- The puzzle has driven important theoretical developments

Optical theorem for two-photon exchange in muonic hydrogen



If a dispersion relation is valid, contribution completely determined by measurable quantities in electron-proton scattering. But:

$$W_1(\nu, 0) = -2 + \mathcal{O}(\nu^2) = \frac{1}{\pi} \int d\nu'^2 \frac{\text{Im}W_1(\nu', 0)}{\nu'^2 - \nu^2} > 0 \quad ??$$

$$\Rightarrow W_1(\nu, Q^2) = W_1(0, Q^2) + \frac{\nu^2}{\pi} \int d\nu'^2 \frac{\text{Im}W_1(\nu', 0)}{\nu'^2(\nu'^2 - \nu^2)}$$

new hadronic function

Two theoretical tools to analyze this new hadronic function:

- Heavy particle effective field theory
- Operator product expansion

Adapt techniques developed for dark matter direct detection cross sections.

RJH, M.P. Solon, PRL (2014)

RJH, Paz (2016)

interesting saga: erratum ~40 years after publication

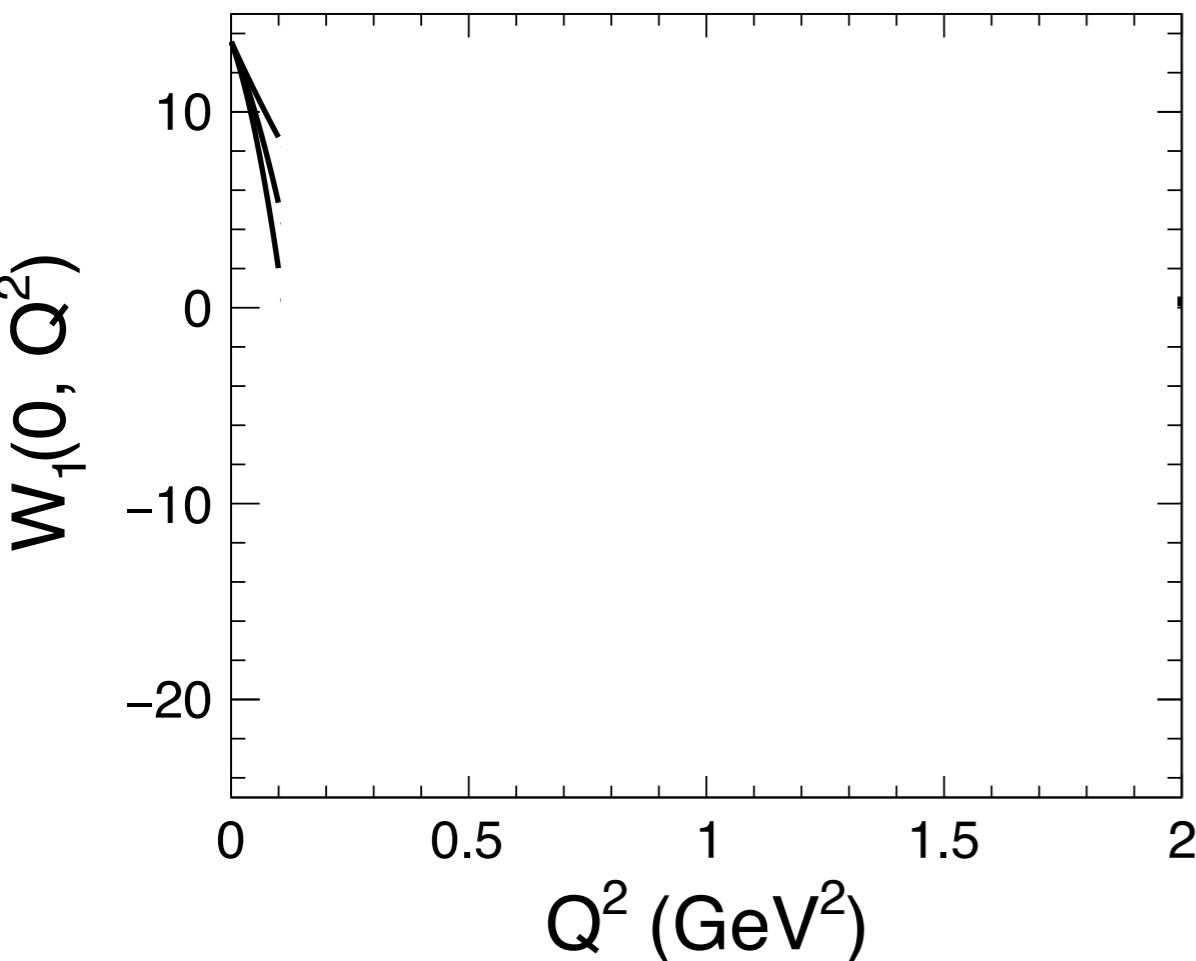
Collins, NPB 149, 90 (1979); erratum ibid 2017

- Heavy particle expansion

Determine interaction coefficients from measurable low-energy processes

$$\mathcal{L} = \psi^\dagger \left(i\partial_t + \dots \right) \psi$$

Apply the thus-determined Lagrangian to predict low- Q^2 expansion of our function

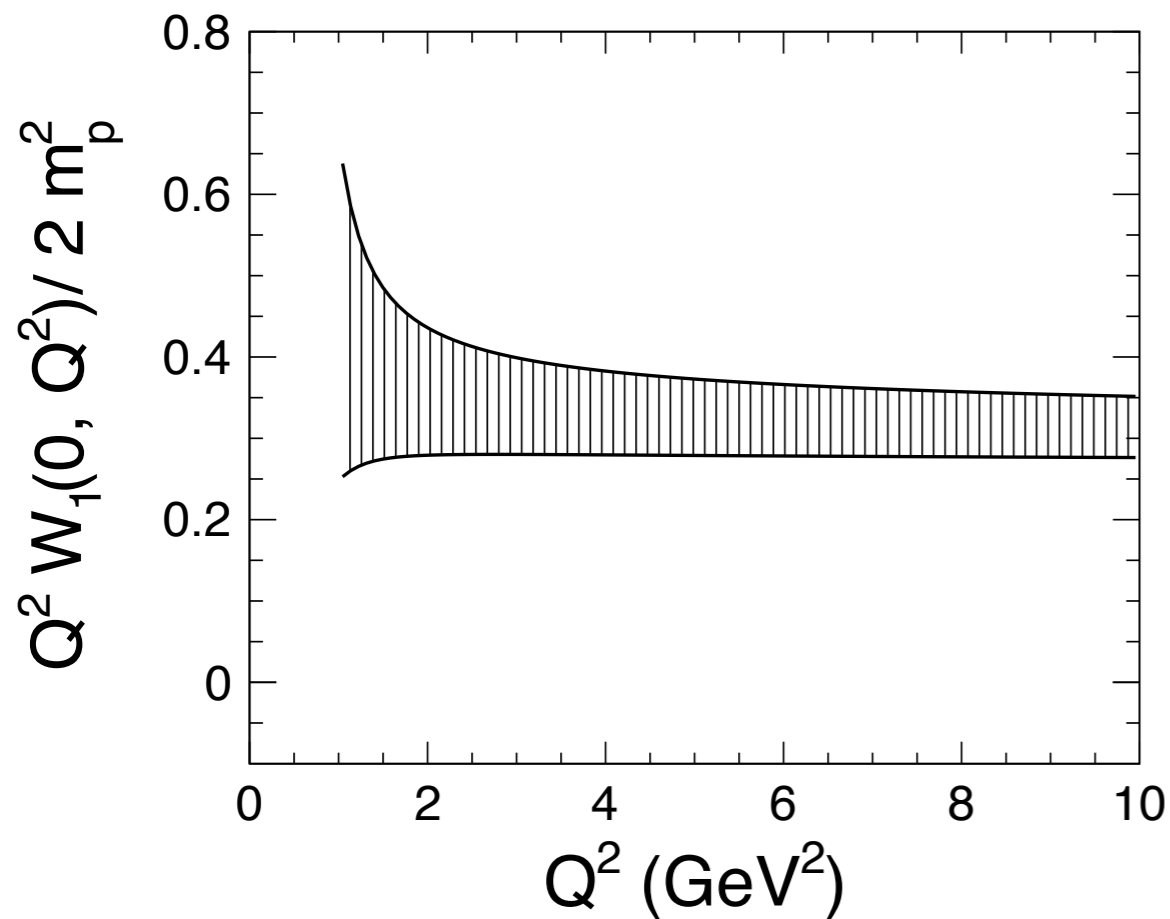


$$W_1(0, Q^2) = 2a_p(2 + a_p)$$

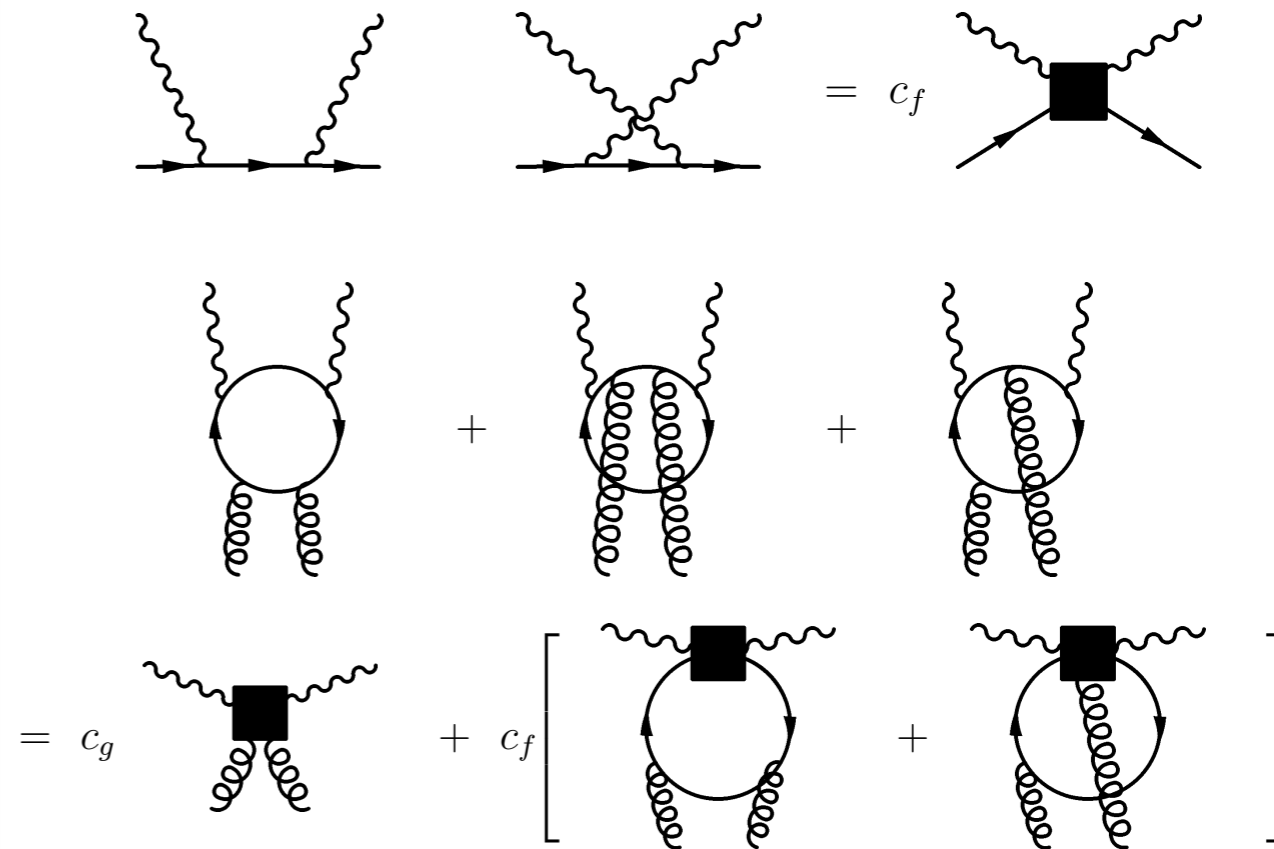
$$+ \left\{ \frac{2m_p^3 \bar{\beta}}{\alpha} - a_p - \frac{2}{3} \left[(1 + a_p)^2 m_p^2 (r_M^p)^2 - m_p^2 (r_E^p)^2 \right] \right\} \frac{Q^2}{m_p^2} + \dots$$

- Operator product expansion

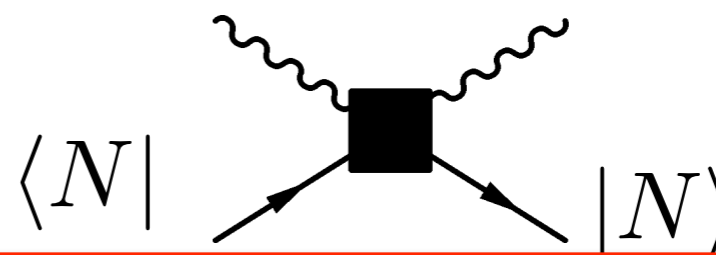
$$W_1(0, Q^2) = \frac{1}{Q^2} \sum_i c_i \langle O_i \rangle$$



c_i short distance matching:
perturbative QCD

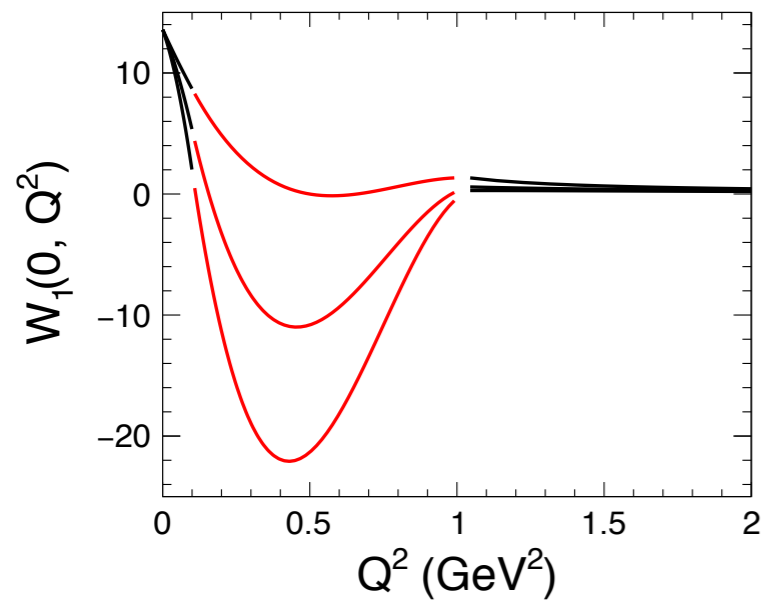


$\langle O_i \rangle$ local operator matrix
elements: nonperturbative
QCD

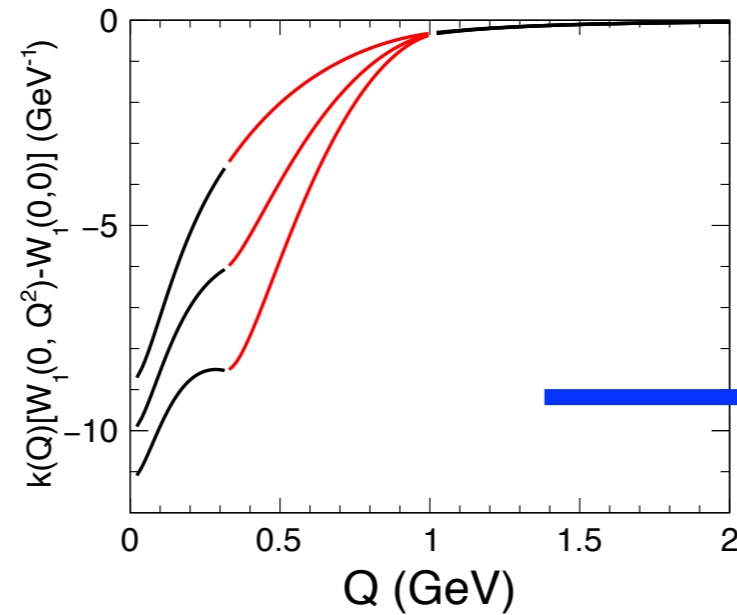


- Put these pieces together:

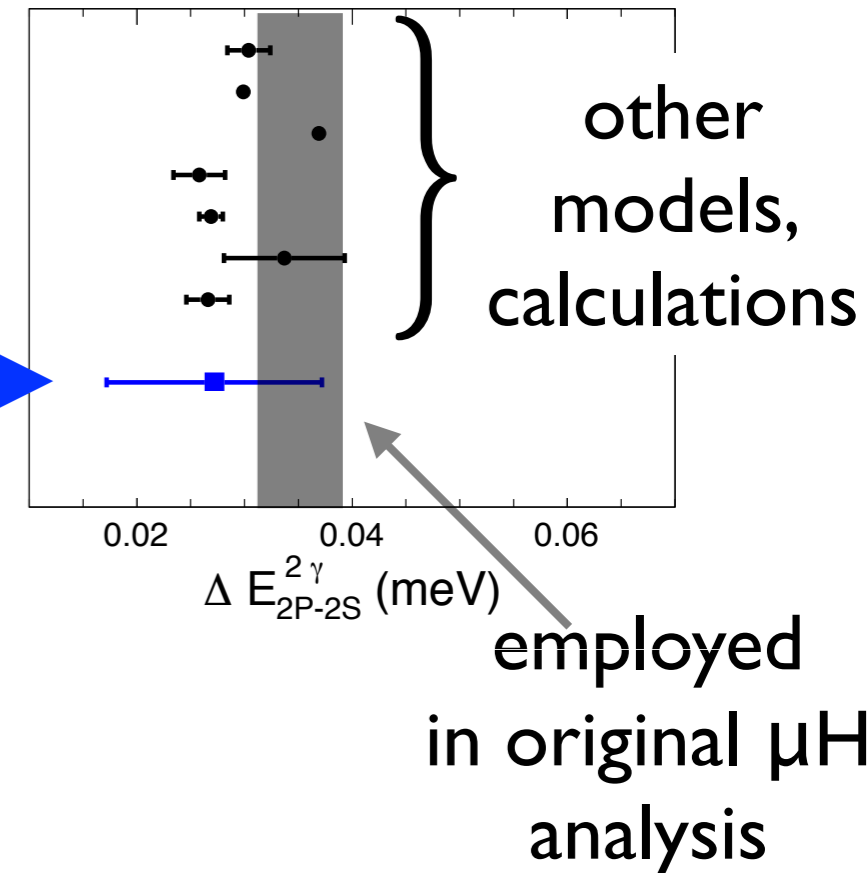
interpolation:



include weighting function:

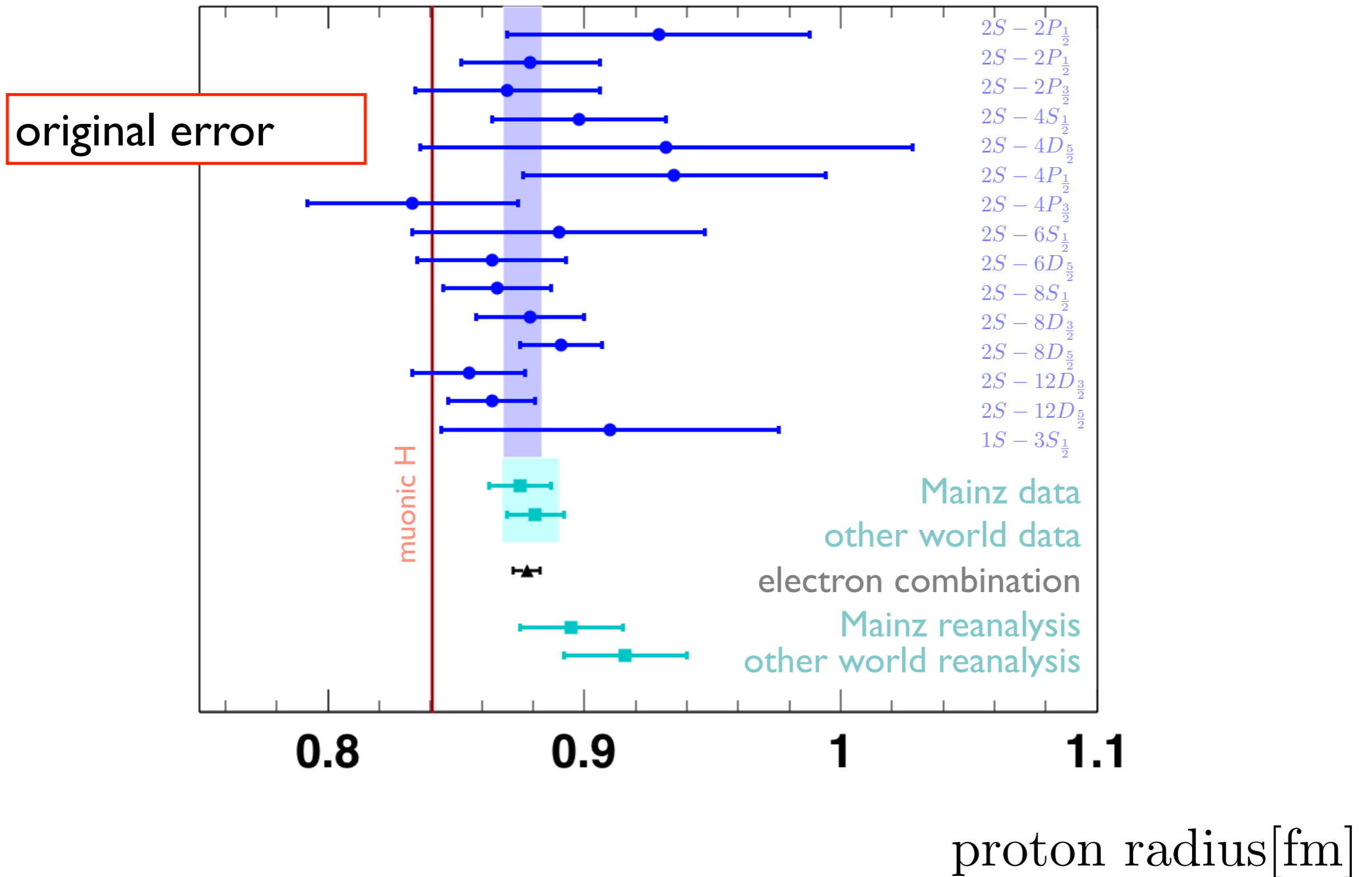


muonic hydrogen
Lamb shift

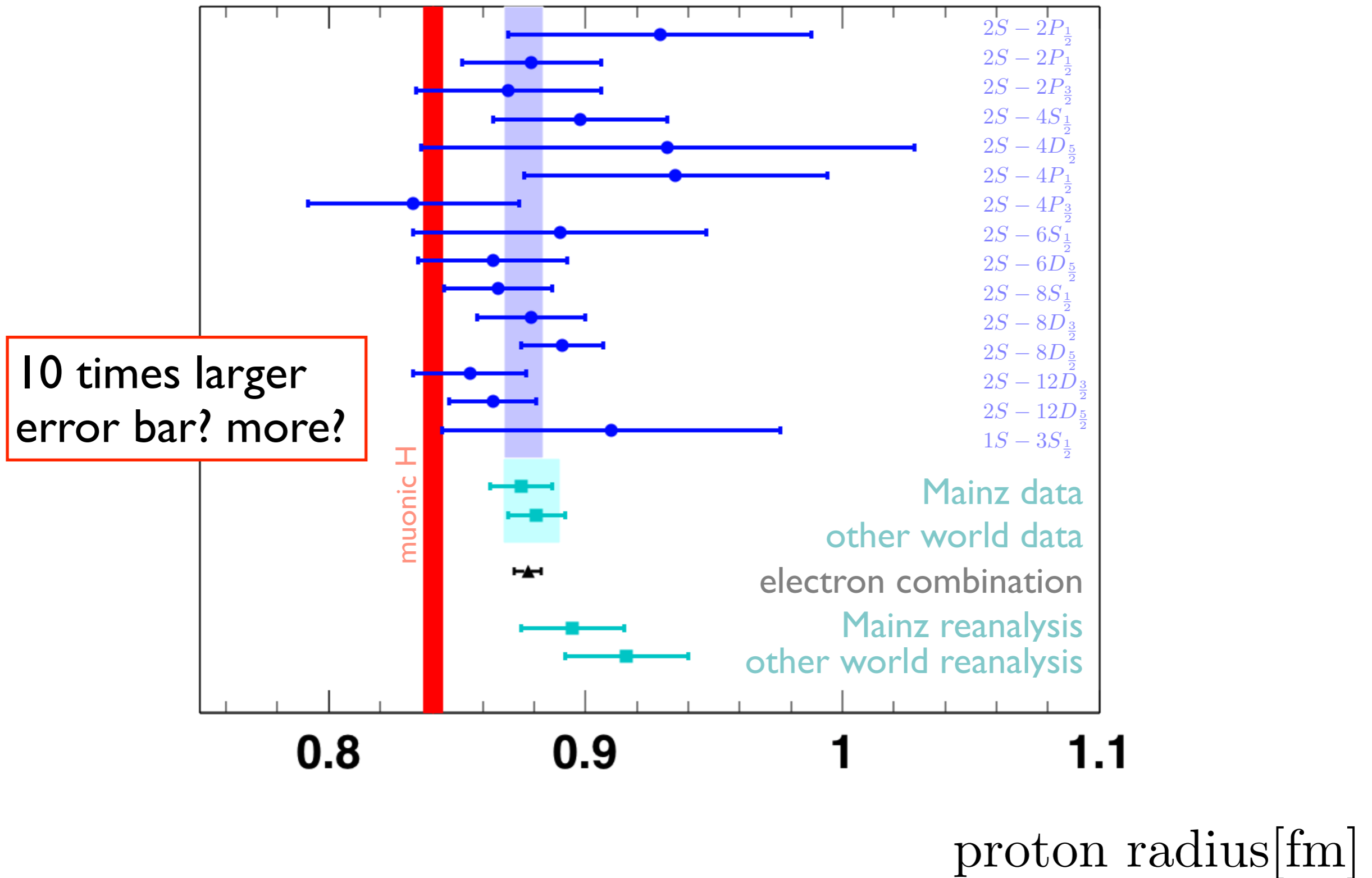


- OPE constraint: turns extrapolation into interpolation
- remains dominant theoretical error for muonic hydrogen

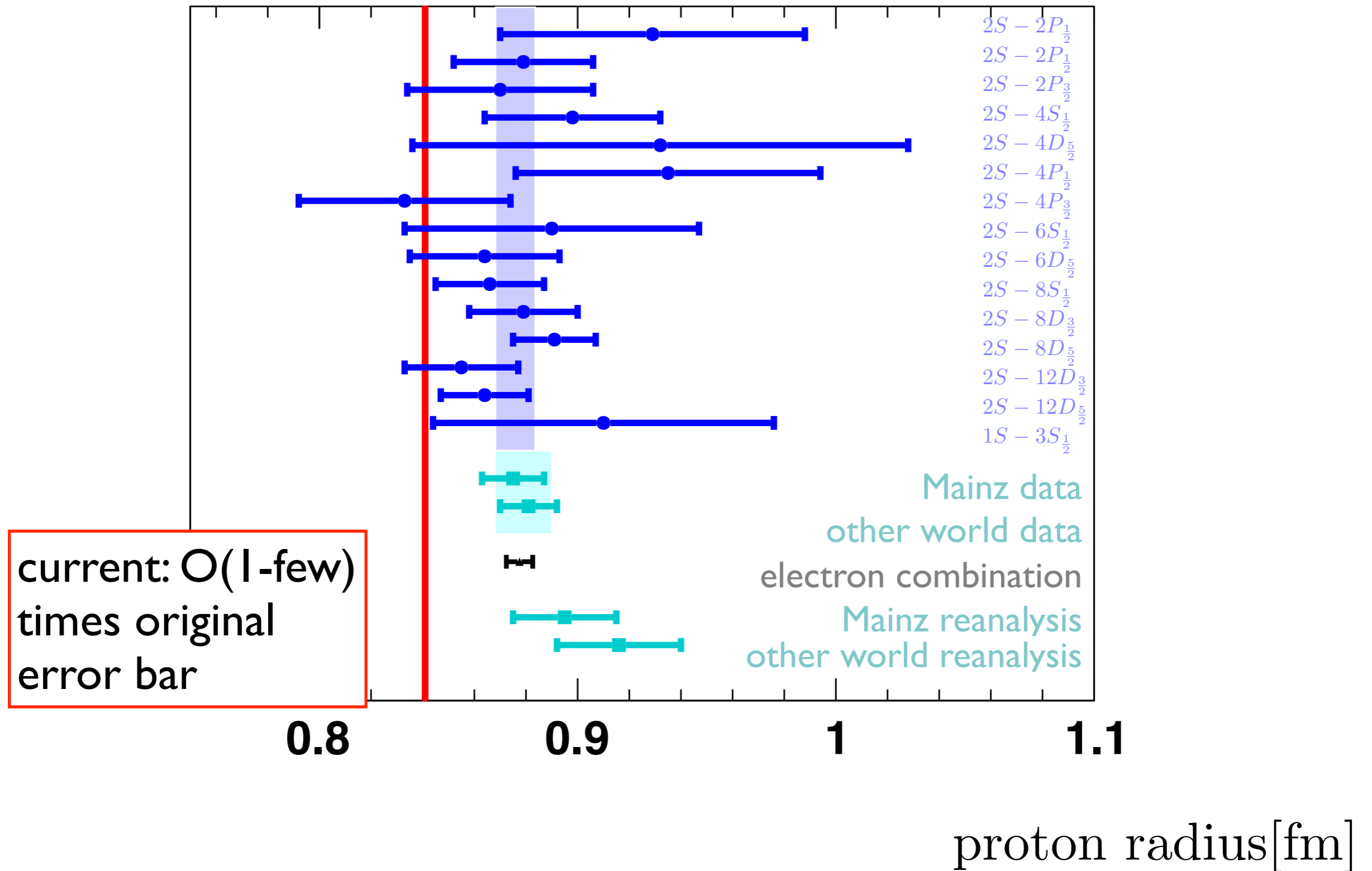
Model-dependent assumptions on two-photon exchange



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Model-dependent assumptions on two-photon exchange



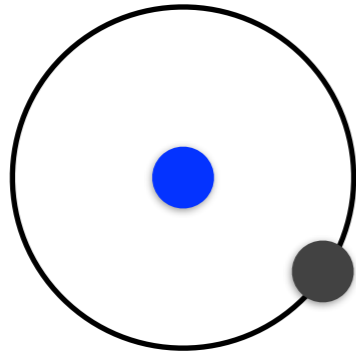
Many manifestations of heavy particle symmetry:

prediction:

small parameter:

- hydrogen/deuterium spectroscopy

$$E_n(H) = -\frac{1}{2}m_e(Z\alpha)^2 + \dots \quad (m_e Z\alpha) \ll m_e$$



--

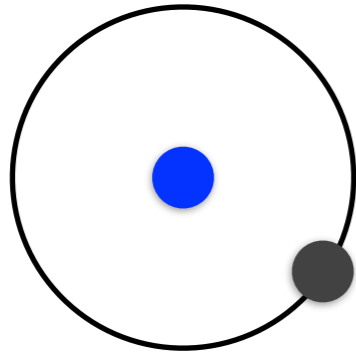
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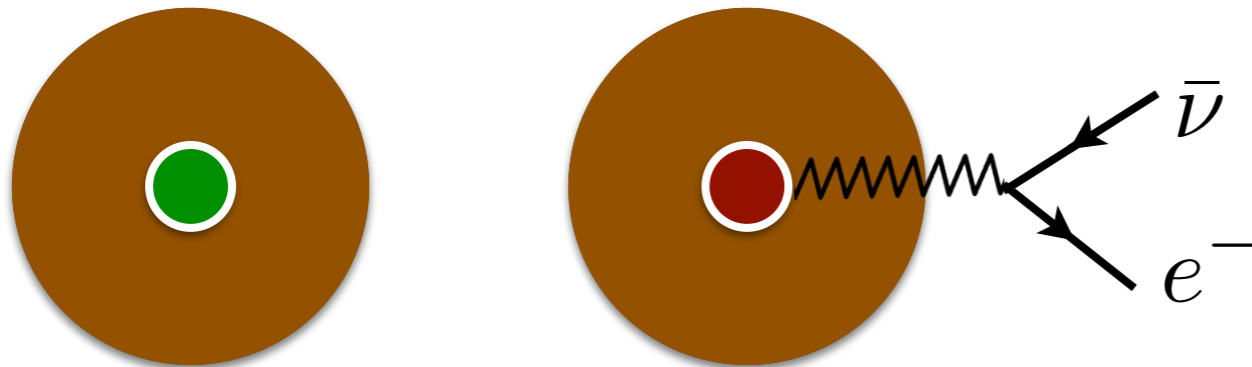
$$E_n(H) = -\frac{1}{2}m_e(Z\alpha)^2 + \dots \quad (m_e Z\alpha) \ll m_e$$



- heavy meson transitions

$$F^{B \rightarrow D}(v' = v) = 1 + \dots$$

$$\Lambda_{\text{QCD}} \ll m_{b,c}$$



--

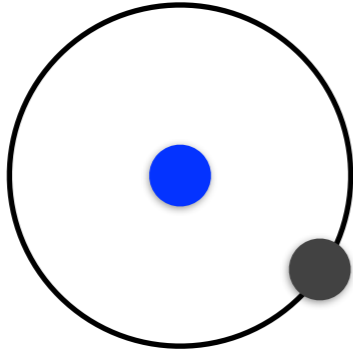
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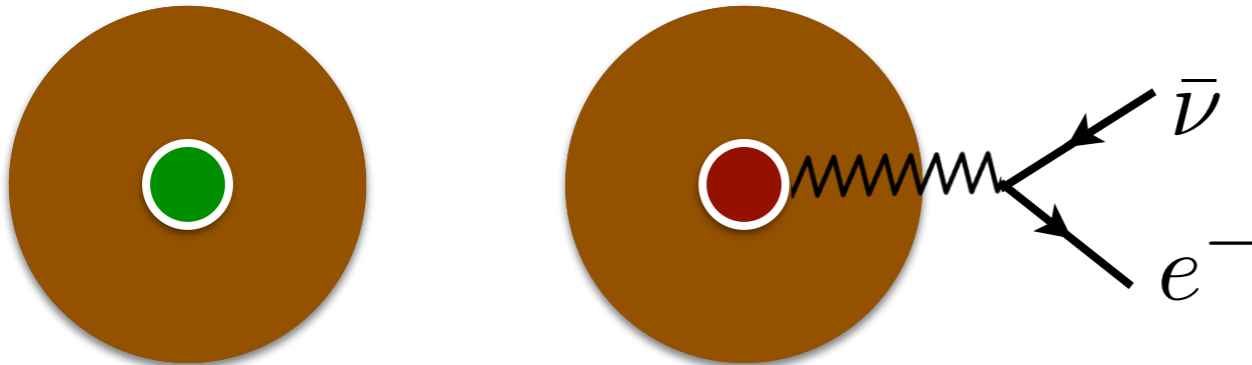
$$E_n(H) = -\frac{1}{2}m_e(Z\alpha)^2 + \dots \quad (m_e Z\alpha) \ll m_e$$



- heavy meson transitions

$$F^{B \rightarrow D}(v' = v) = 1 + \dots$$

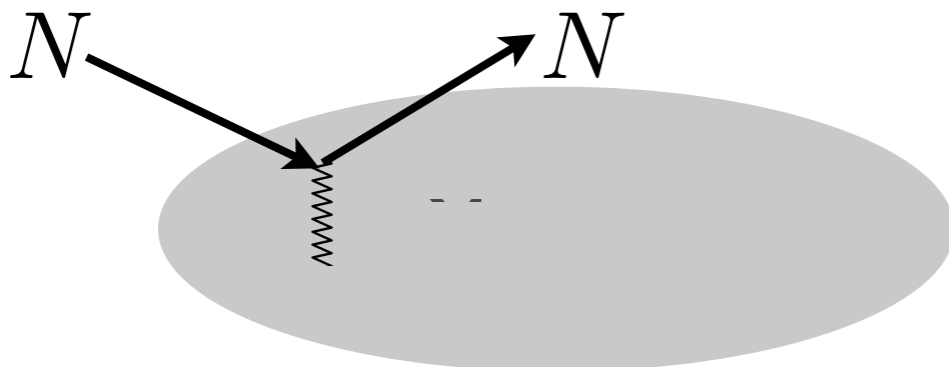
$$\Lambda_{\text{QCD}} \ll m_{b,c}$$



- DM interactions

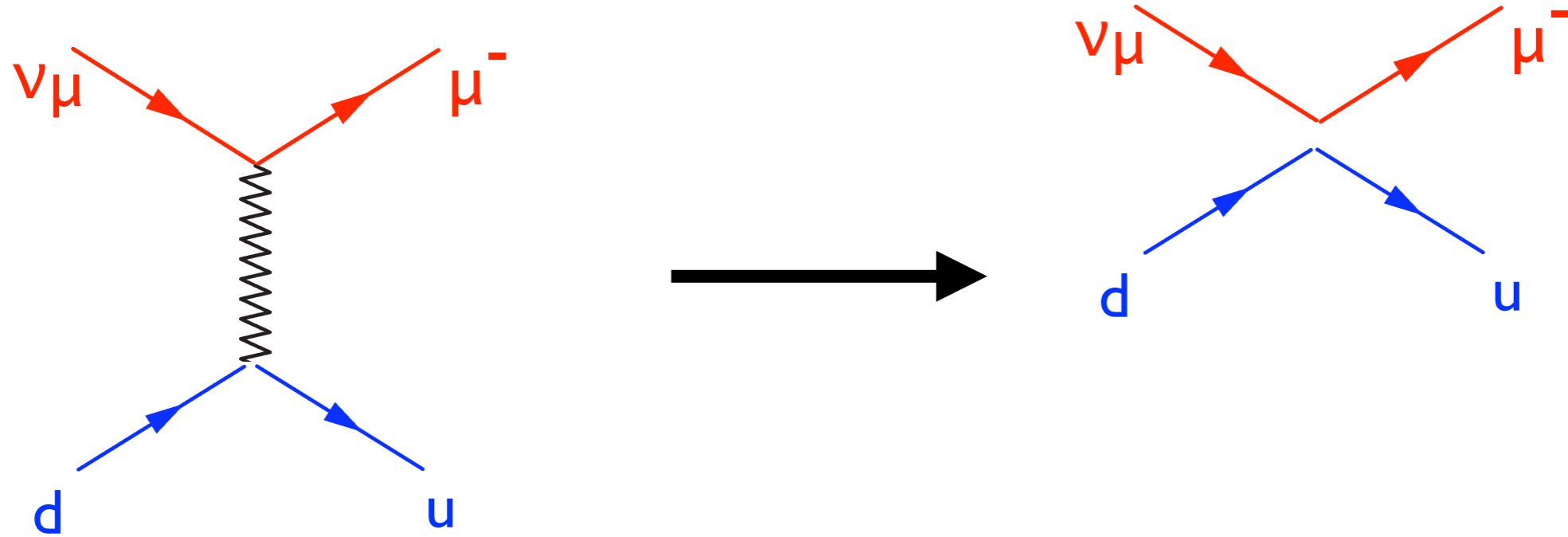
$$\sigma(\chi N \rightarrow \chi N) = ?$$

$$m_W \ll m_\chi$$



Many applications of operator product expansion:

E.g., Fermi theory of weak interactions:



$$[\bar{u}d]_{V-A} \frac{g^2}{\partial^2 - m_W^2} [\bar{\mu}\nu]_{V-A}$$

short distance coefficient

$$G_F \underbrace{[\bar{u}d]_{V-A} [\bar{\mu}\nu]_{V-A}} + \dots$$

local operator

Renormalization analysis for log-enhanced radiative corrections

$$d\sigma = \underbrace{H(M)}_{\text{hadron structure}} \times \underbrace{\frac{H(\mu)}{H(M)} \times J(\mu) \times S(\mu)}_{\text{radiative correction}}$$

hadron
structure

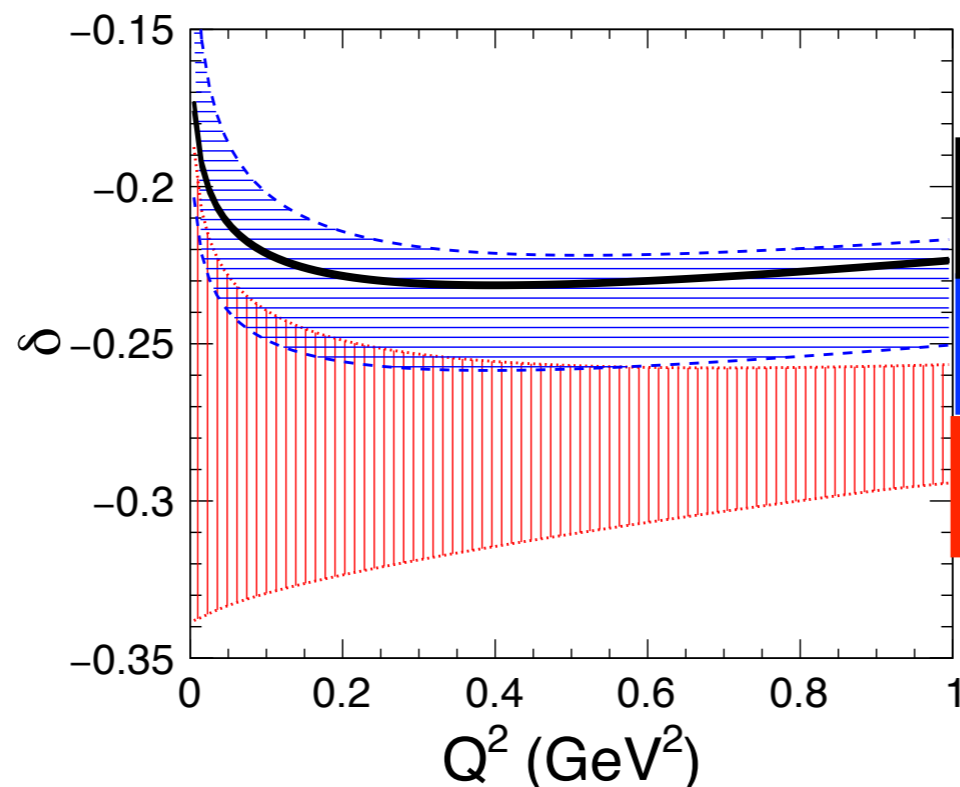
radiative
correction

numerically: $\alpha L^2 = \alpha \log^2 \frac{Q^2}{m^2} \sim 1 \quad \Rightarrow \quad \alpha L \sim \alpha^{\frac{1}{2}}, \text{ etc.}$

electron energy: $E = 1 \text{ GeV}$

electron energy loss cut: $\Delta E = 5 \text{ MeV}$

total radiative
correction



NLO
NLL
LL

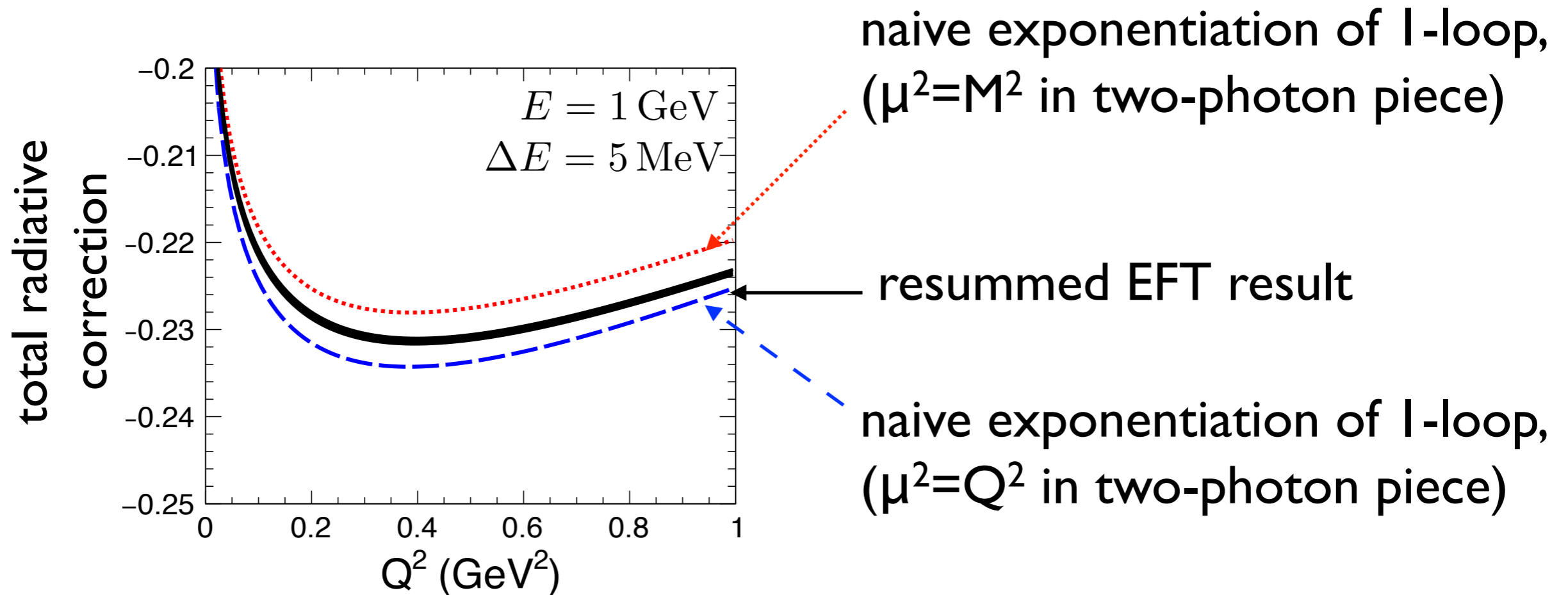
correct
through:

$\mathcal{O}(\alpha)$

$\mathcal{O}(\alpha^{\frac{1}{2}})$

$\mathcal{O}(1)$

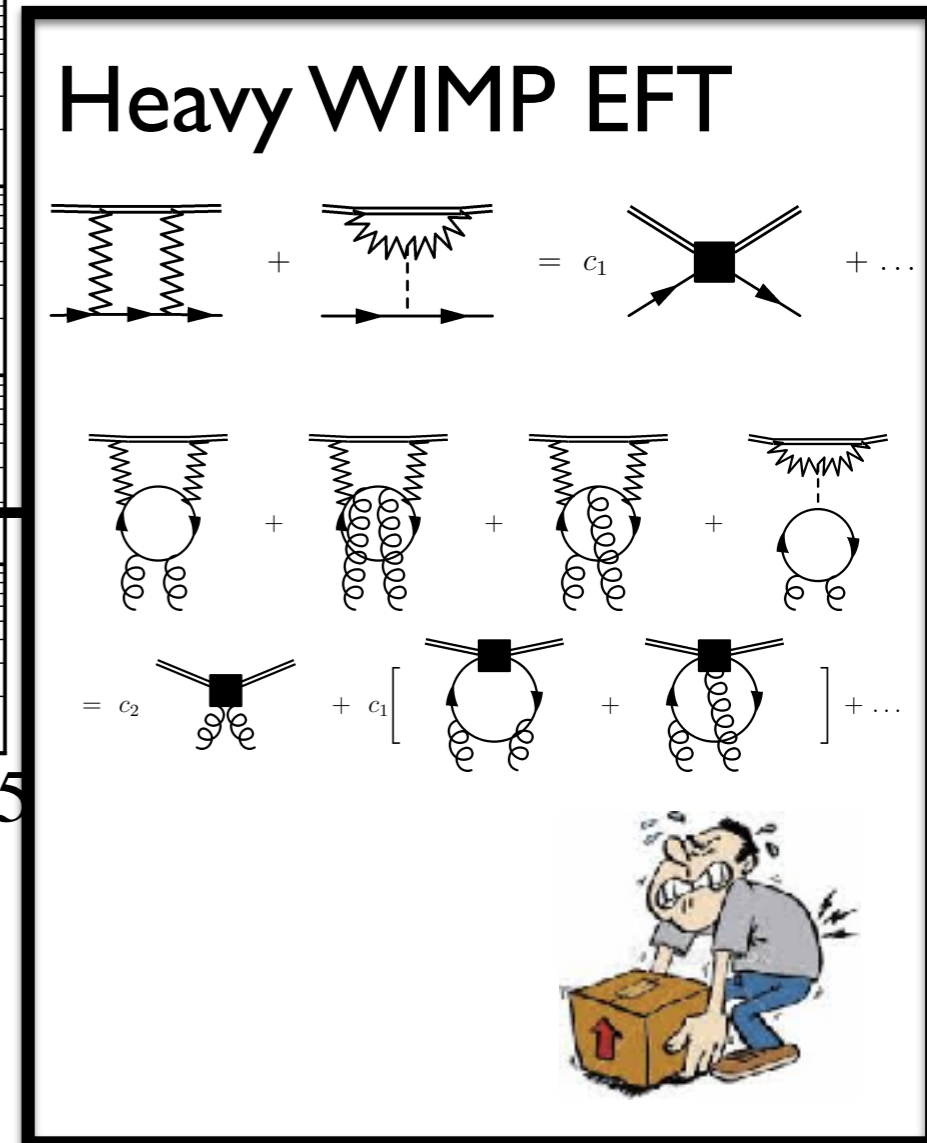
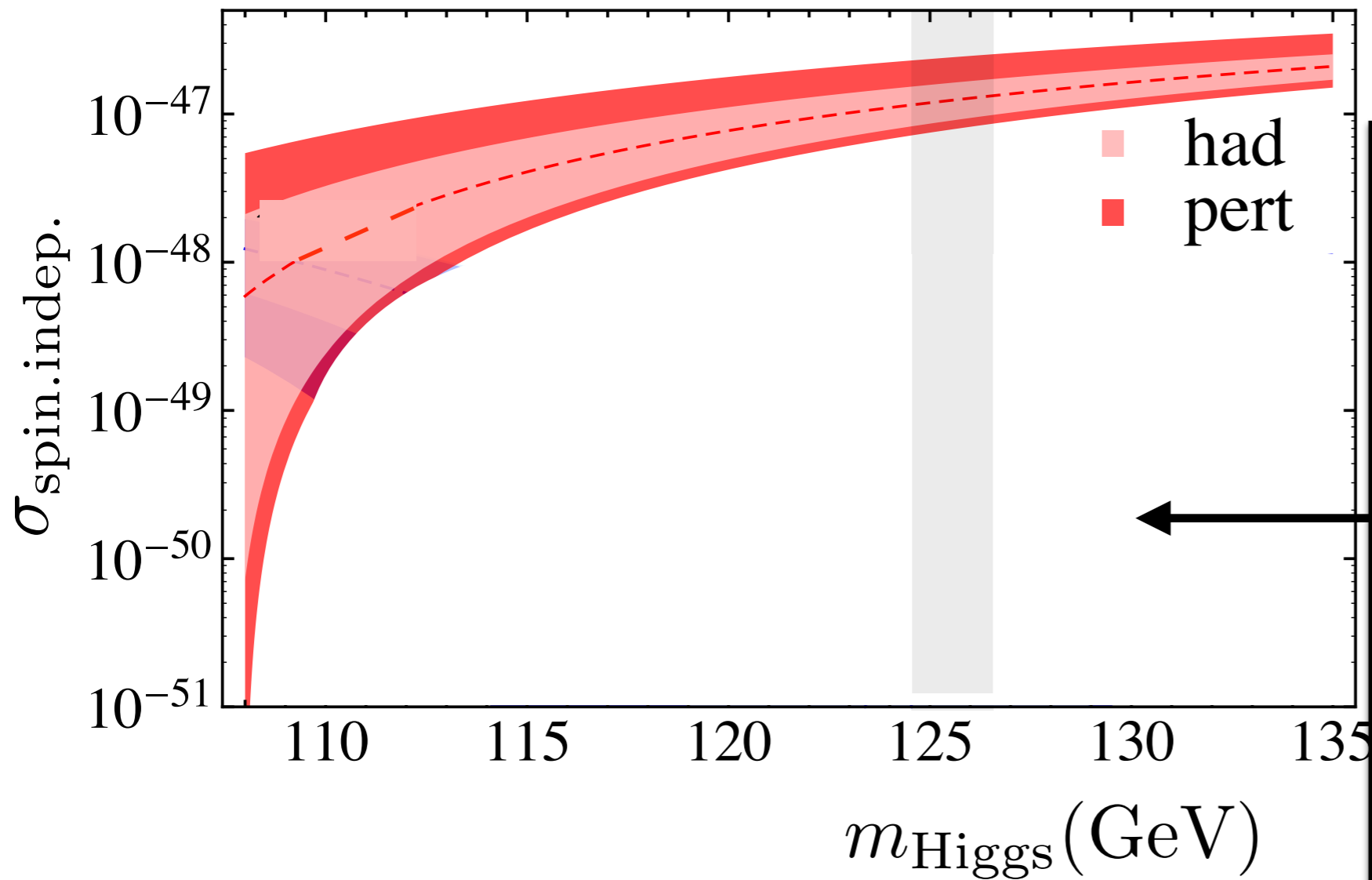
Comparison to previous implementations of radiative corrections, e.g. in AI analysis of electron-proton scattering data



- discrepancies at 0.5-1% compared to currently applied radiative correction models (cf. 0.2-0.5% systematic error budget of AI experiment)
- should be implemented directly in analysis, but doesn't appear to resolve anomaly (floating normalizations)
- model dependence in hard two-photon exchange remains

Model independent prediction for heavy WIMP scattering

RJH, M.P. Solon, PRL (2014)



- generally, expect WIMP scattering cross section to depend on mass, spin, and electroweak quantum numbers

- heavy WIMP regime: universal prediction for given quantum numbers (here consider electroweak triplet)