

# Search for exotic particles at NA62

Gaia Lanfranchi - LNF-INFN

on behalf of the NA62 Collaboration

See A. Ceccucci's and L. Bician's talks for a complete view of the NA62 physics reach

ALPS 2018 – Obergurgl – April 2018

# Preamble

---

With the **discovery at the LHC** of the Higgs boson, the main missing block for the experimental validation of the Standard Model is now in place.

An additional LHC result of great importance (and totally unexpected) is that a large new territory has been explored and no unambiguous signal of New Physics has been found.

After  $\sim 7$  years of excellent running of the LHC and its experiments some conclusions can be drawn.

# No New Physics in direct searches

(almost all the limits are above the TeV scale)

CERN DG's New Year speech, Jan 2018

## ATLAS Exotics Searches\* - 95% CL Upper Exclusion Limits

Status: July 2017

ATLAS Preliminary

$$\int \mathcal{L} dt = (3.2 - 37.0) \text{ fb}^{-1}$$

$$\sqrt{s} = 8, 13 \text{ TeV}$$

Model	$\ell, \gamma$	Jets <sup>†</sup>	$E_T^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Limit	Reference	
Extra dimensions	ADD $G_{KK} + g/q$	$0 e, \mu$	1-4 j	Yes	36.1	$M_D$ 7.75 TeV	$n = 2$
	ADD non-resonant $\gamma\gamma$	$2 \gamma$	-	-	36.7	$M_S$ 8.6 TeV	$n = 3$ HLZ NLO
	ADD QBH	-	2 j	-	37.0	$M_{\text{th}}$ 8.9 TeV	$n = 6$
	ADD BH high $\Sigma p_T$	$\geq 1 e, \mu$	$\geq 2 j$	-	3.2	$M_{\text{th}}$ 8.2 TeV	$n = 6, M_D = 3 \text{ TeV}$ , rot BH
	ADD BH multijet	-	$\geq 3 j$	-	3.6	$M_{\text{th}}$ 9.55 TeV	$n = 6, M_D = 3 \text{ TeV}$ , rot BH
	RS1 $G_{KK} \rightarrow \gamma\gamma$	$2 \gamma$	-	-	36.7	$G_{KK}$ mass 4.1 TeV	$k/M_{\text{pl}} = 0.1$
Bulk RS $G_{KK} \rightarrow WW \rightarrow qq\ell\nu$	$1 e, \mu$	1 J	Yes	36.1	$G_{KK}$ mass 1.75 TeV	$k/M_{\text{pl}} = 1.0$	
2UED / RPP	$1 e, \mu$	$\geq 2 b, \geq 3 j$	Yes	13.2	KK mass 1.6 TeV	Tier (1,1), $\mathcal{B}(A^{1,3} \rightarrow tt) = 1$	
Gauge bosons	SSM $Z' \rightarrow \ell\ell$	$2 e, \mu$	-	-	36.1	$Z'$ mass 4.5 TeV	
	SSM $Z' \rightarrow \tau\tau$	$2 \tau$	-	-	36.1	$Z'$ mass 2.4 TeV	
	Leptophobic $Z' \rightarrow bb$	-	2 b	-	3.2	$Z'$ mass 1.5 TeV	
	Leptophobic $Z' \rightarrow tt$	$1 e, \mu$	$\geq 1 b, \geq 1 J/2$	Yes	3.2	$Z'$ mass 2.0 TeV	$\Gamma/m = 3\%$
	SSM $W' \rightarrow \ell\nu$	$1 e, \mu$	-	Yes	36.1	$W'$ mass 5.1 TeV	
	HVT $V' \rightarrow WW \rightarrow qq\gamma\gamma$ model B	$0 e, \mu$	2 J	-	36.7	$V'$ mass 3.5 TeV	$g_V = 3$
HVT $V' \rightarrow WH/ZH$ model B	multi-channel	-	-	36.1	$V'$ mass 2.93 TeV	$g_V = 3$	
LRSM $W'_R \rightarrow tb$	$1 e, \mu$	2 b, 0-1 j	Yes	20.3	$W'$ mass 1.92 TeV		
LRSM $W'_R \rightarrow tb$	$0 e, \mu$	$\geq 1 b, 1 J$	-	20.3	$W'$ mass 1.76 TeV		
CI	CI $qqqq$	-	2 j	-	37.0	A 21.8 TeV	$\eta_{LL}$
	CI $\ell\ell qq$	$2 e, \mu$	-	-	36.1	A 40.1 TeV	$\eta_{LL}$
	CI $uutt$	$2(SS) \geq 3 e, \mu \geq 1 b, \geq 1 j$	Yes	20.3	A 4.9 TeV	$ C_{\text{rel}}  = 1$	
DM	Axial-vector mediator (Dirac DM)	$0 e, \mu$	1-4 j	Yes	36.1	$m_{\text{med}}$ 1.5 TeV	$g_S = 0.25, g_V = 1.0, m(\chi) < 400 \text{ GeV}$
	Vector mediator (Dirac DM)	$0 e, \mu, 1 \gamma$	$\leq 1 j$	Yes	36.1	$m_{\text{med}}$ 1.2 TeV	$g_S = 0.25, g_V = 1.0, m(\chi) < 480 \text{ GeV}$
	VV $\chi\chi$ EFT (Dirac DM)	$0 e, \mu$	1 J, $\leq 1 j$	Yes	3.2	M 700 GeV	$m(\chi) < 150 \text{ GeV}$
LQ	Scalar LQ 1 <sup>st</sup> gen	$2 e$	$\geq 2 j$	-	3.2	LQ mass 1.1 TeV	$\beta = 1$
	Scalar LQ 2 <sup>nd</sup> gen	$2 \mu$	$\geq 2 j$	-	3.2	LQ mass 1.05 TeV	$\beta = 1$
	Scalar LQ 3 <sup>rd</sup> gen	$1 e, \mu$	$\geq 1 b, \geq 3 j$	Yes	20.3	LQ mass 640 GeV	$\beta = 0$
Heavy quarks	VLQ $TT \rightarrow Ht + X$	$0$ or $1 e, \mu$	$\geq 2 b, \geq 3 j$	Yes	13.2	T mass 1.2 TeV	$\mathcal{B}(T \rightarrow Ht) = 1$
	VLQ $TT \rightarrow Zt + X$	$1 e, \mu$	$\geq 1 b, \geq 3 j$	Yes	36.1	T mass 1.16 TeV	$\mathcal{B}(T \rightarrow Zt) = 1$
	VLQ $TT \rightarrow Wb + X$	$1 e, \mu$	$\geq 1 b, \geq 1 J/2$	Yes	36.1	T mass 1.35 TeV	$\mathcal{B}(T \rightarrow Wb) = 1$
	VLQ $BB \rightarrow Hb + X$	$1 e, \mu$	$\geq 2 b, \geq 3 j$	Yes	20.3	B mass 700 GeV	$\mathcal{B}(B \rightarrow Hb) = 1$
	VLQ $BB \rightarrow Zb + X$	$2/\geq 3 e, \mu$	$\geq 2/\geq 1 b$	-	20.3	B mass 790 GeV	$\mathcal{B}(B \rightarrow Zb) = 1$
	VLQ $BB \rightarrow Wt + X$	$1 e, \mu$	$\geq 1 b, \geq 1 J/2$	Yes	36.1	B mass 1.25 TeV	$\mathcal{B}(B \rightarrow Wt) = 1$
VLQ $QQ \rightarrow WqVq$	$1 e, \mu$	$\geq 4 j$	Yes	20.3	Q mass 690 GeV		
Excited fermions	Excited quark $q^* \rightarrow qg$	-	2 j	-	37.0	$q^*$ mass 6.0 TeV	only $u^*$ and $d^*$ , $\Lambda = m(q^*)$
	Excited quark $q^* \rightarrow q\gamma$	$1 \gamma$	1 j	-	36.7	$q^*$ mass 5.3 TeV	only $u^*$ and $d^*$ , $\Lambda = m(q^*)$
	Excited quark $b^* \rightarrow bg$	-	1 b, 1 j	-	13.3	$b^*$ mass 2.3 TeV	
	Excited quark $b^* \rightarrow Wt$	1 or 2 $e, \mu$	1 b, 2-0 j	Yes	20.3	$b^*$ mass 1.5 TeV	$f_b = f_t = f_g = 1$
	Excited lepton $\ell^*$	$3 e, \mu$	-	-	20.3	$\ell^*$ mass 3.0 TeV	$\Lambda = 3.0 \text{ TeV}$
Excited lepton $\nu^*$	$3 e, \mu, \tau$	-	-	20.3	$\nu^*$ mass 1.6 TeV	$\Lambda = 1.6 \text{ TeV}$	
Other	LRSM Majorana $\nu$	$2 e, \mu$	2 j	-	20.3	$N^0$ mass 2.0 TeV	$m(W_R) = 2.4 \text{ TeV}$ , no mixing
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	$2, 3, 4 e, \mu$ (SS)	-	-	36.1	$H^{\pm\pm}$ mass 870 GeV	DY production
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\tau$	$3 e, \mu, \tau$	-	-	20.3	$H^{\pm\pm}$ mass 400 GeV	DY production, $\mathcal{B}(H^{\pm\pm} \rightarrow \ell\tau) = 1$
	Monotop (non-res prod)	$1 e, \mu$	1 b	Yes	20.3	spin-1 invisible particle mass 657 GeV	$\alpha_{\text{non-res}} = 0.2$
	Multi-charged particles	-	-	-	20.3	multi-charged particle mass 785 GeV	DY production, $ q  = 5e$
	Magnetic monopoles	-	-	-	7.0	monopole mass 1.34 TeV	DY production, $ g  = 1g_D$ , spin 1/2

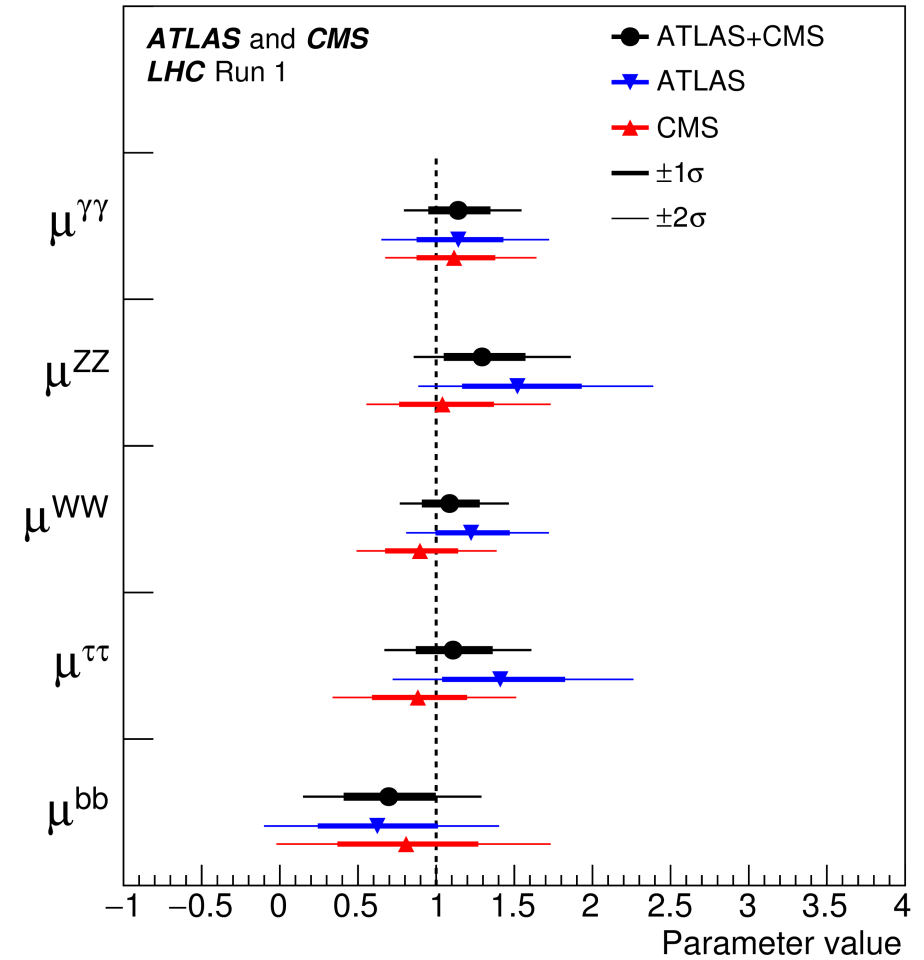
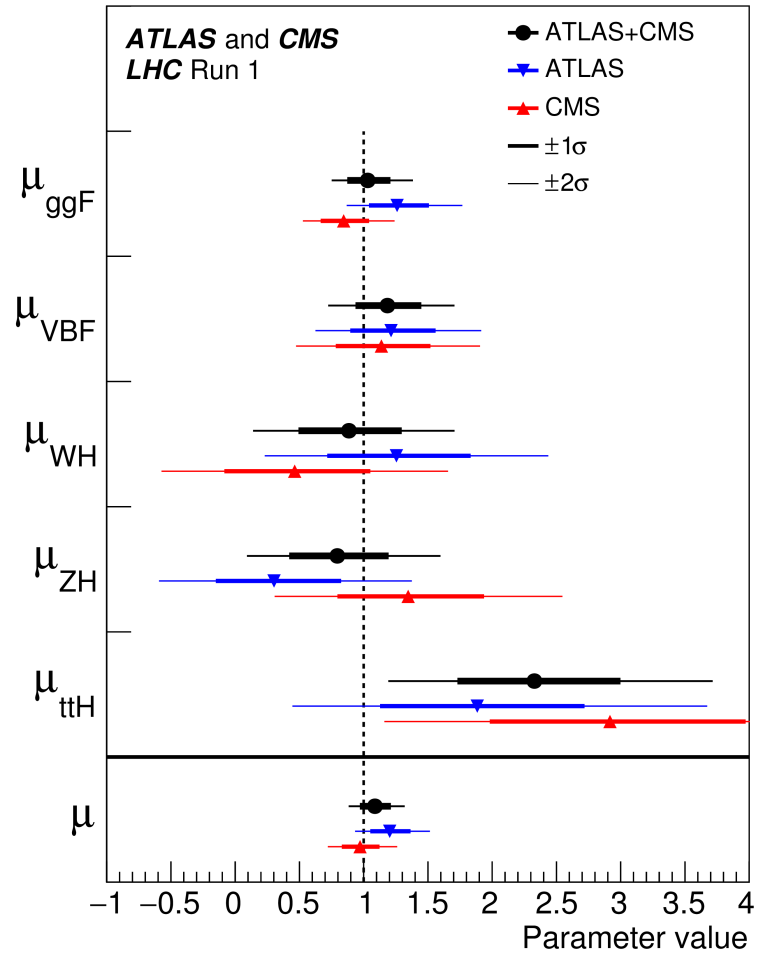
\*Only a selection of the available mass limits on new states or phenomena is shown.

<sup>†</sup>Small-radius (large-radius) jets are denoted by the letter j (J).

# No New Physics in Higgs couplings

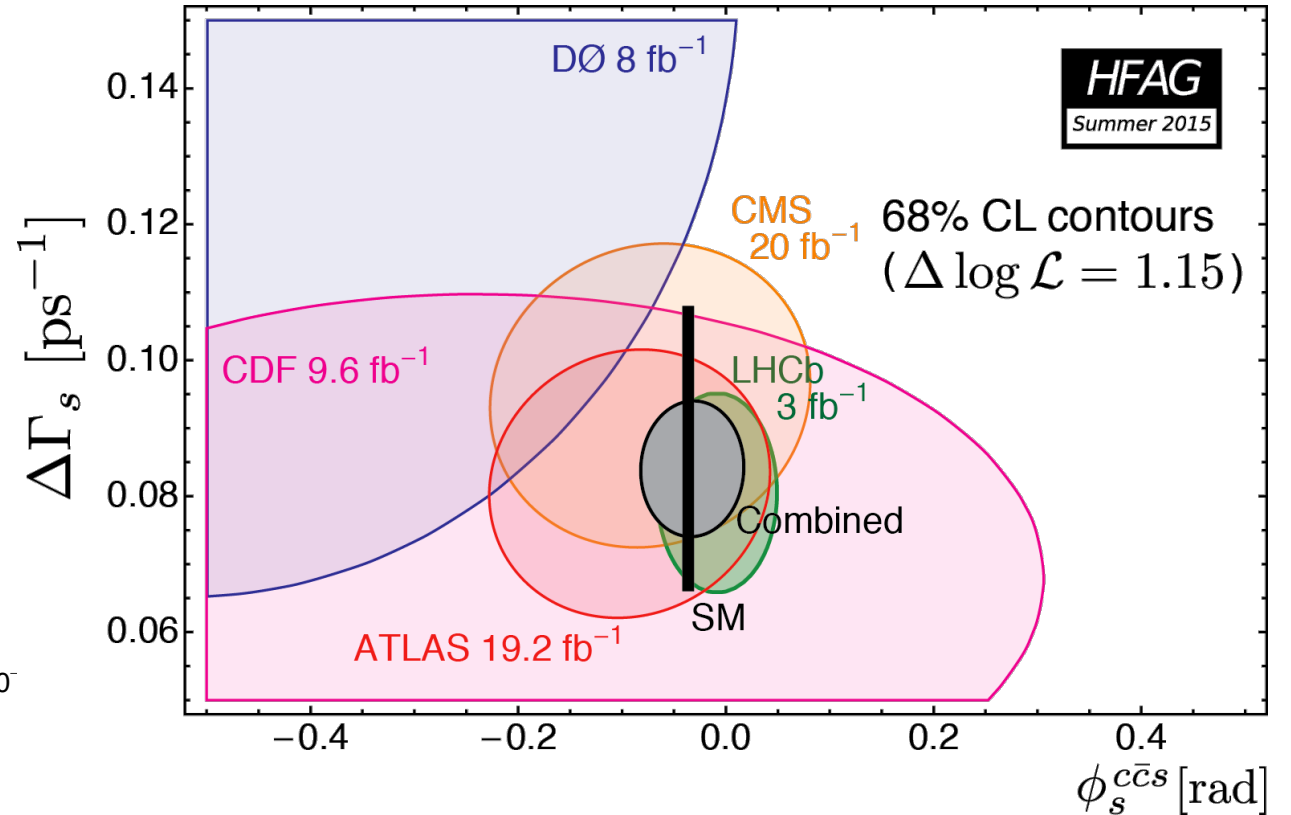
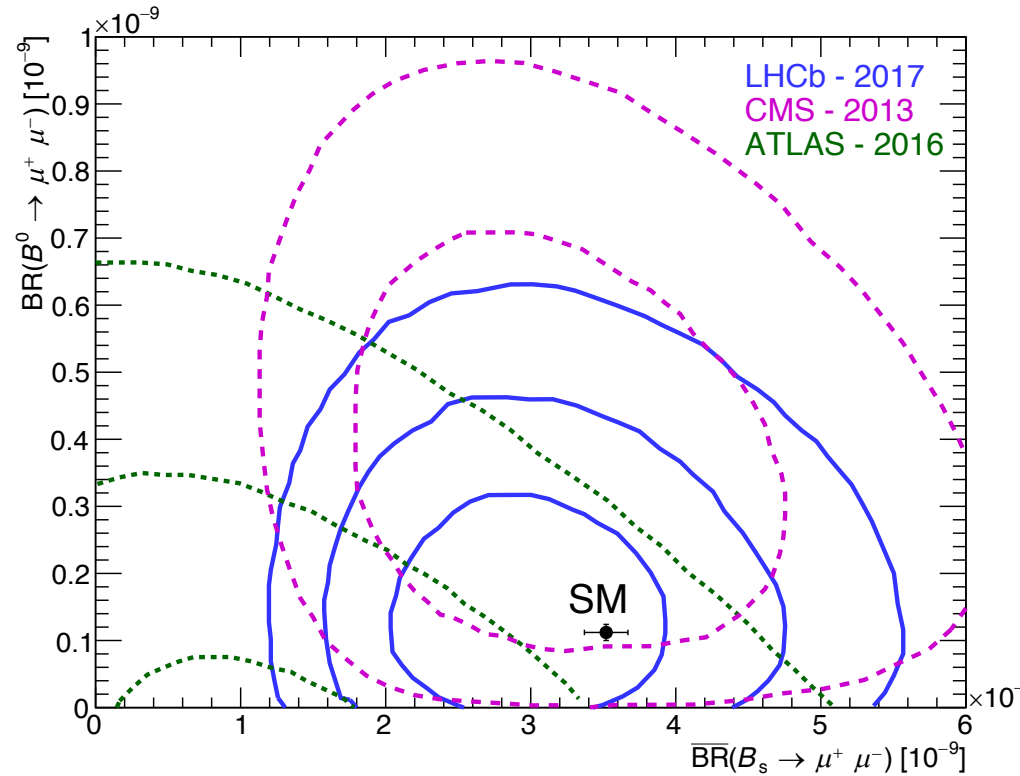
Production and decay rates compatible with the SM Higgs at 20-60 % accuracy

P. Meridiani EPS conference 2017

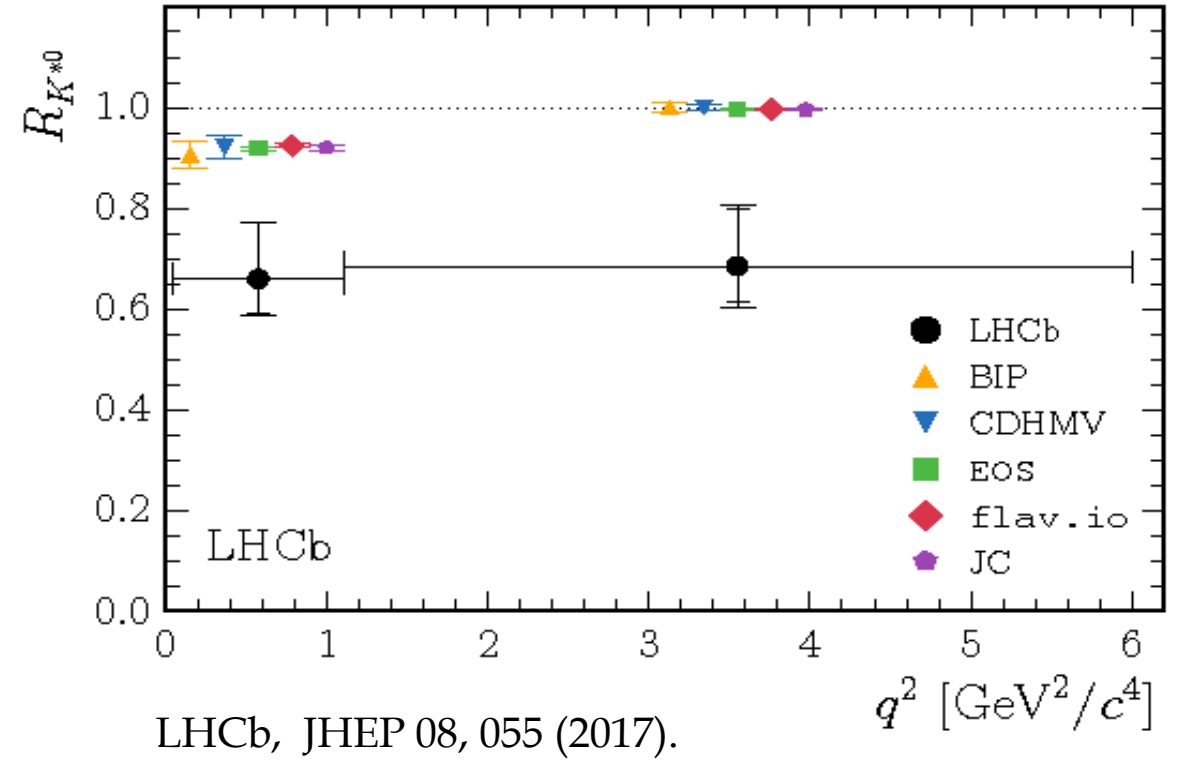
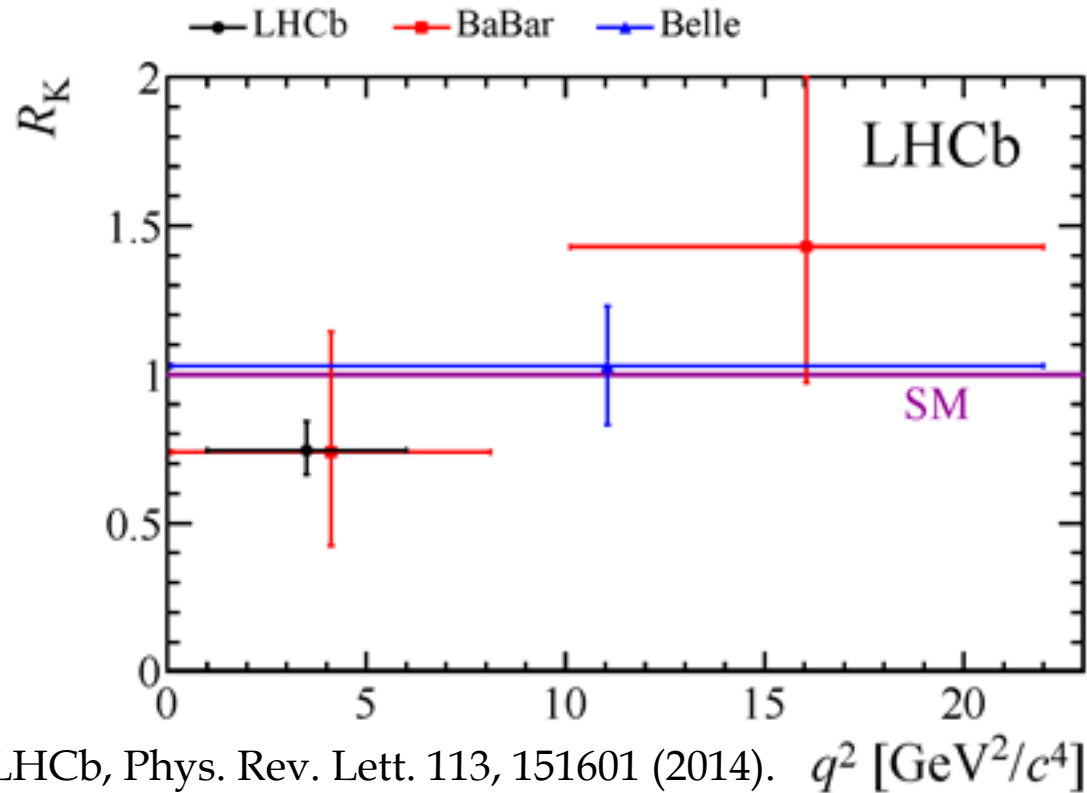


# No New Physics in Flavour ?

The  $B^0_{(s)} \rightarrow \mu^+ \mu^-$  branching fractions and  $B_s$  mixing phase in excellent agreement with SM predictions



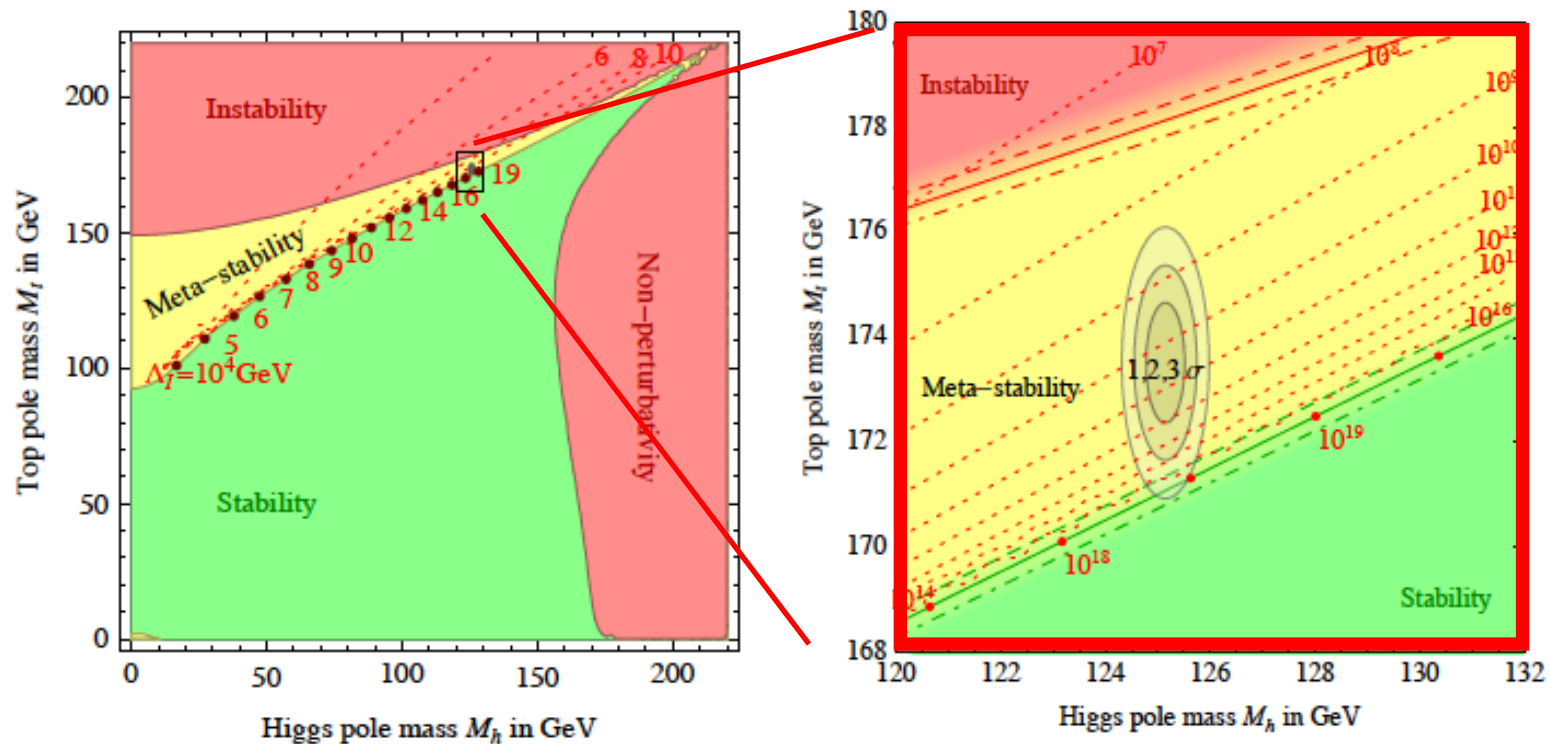
# Lepton Flavor Universality anomalies from LHCb



**An open issue, yet.**

# A “metastable” world

The main reason for NP at the TeV scale was to solve the fine tuning problem of the mass of the Higgs...



Buttazzo et al.  
JHEP 1312 (2013) 089;

A.V. Bednyakov,  
Phys.Part.Nucl. 48  
(2017) no.5, 698

.....However the values of the Higgs ( $m_H \sim 125.1$  GeV) and top mass ( $m_{top} = 173.1$  GeV) the Nature has chosen make the Standard Model a self-consistent (effective field) theory all the way up to the quantum gravity Planck scale, **even in absence of NP at the TeV scale.**

# An attractive possibility

---

- **The Standard Model, as it is, is renormalizable and predictive up to the Planck scale.**
- **However we do have experimental evidence of physics beyond SM:**
  - Dark Matter
  - Neutrino masses and oscillations;
  - Baryogenesis
  - Cosmological inflation
  - Accelerated expansion of the Universe
- **An attractive possibility:**
  - there could be no new scale between the EW and the Planck scale:**
    - **all the unresolved problems could be explained by NP below the EW scale, that does not perturb the smooth running of the Higgs self-coupling.**



# Interaction between DM and SM mediated by Gauge-invariant operators

---

$B_{\mu\nu}$  - vector portal, dimension 2: dark photon

$H^\dagger H$  - scalar portal, dimension 2: new scalars

$H^T L$  - neutrino portal, dimension 5/2: new leptons, HNLs

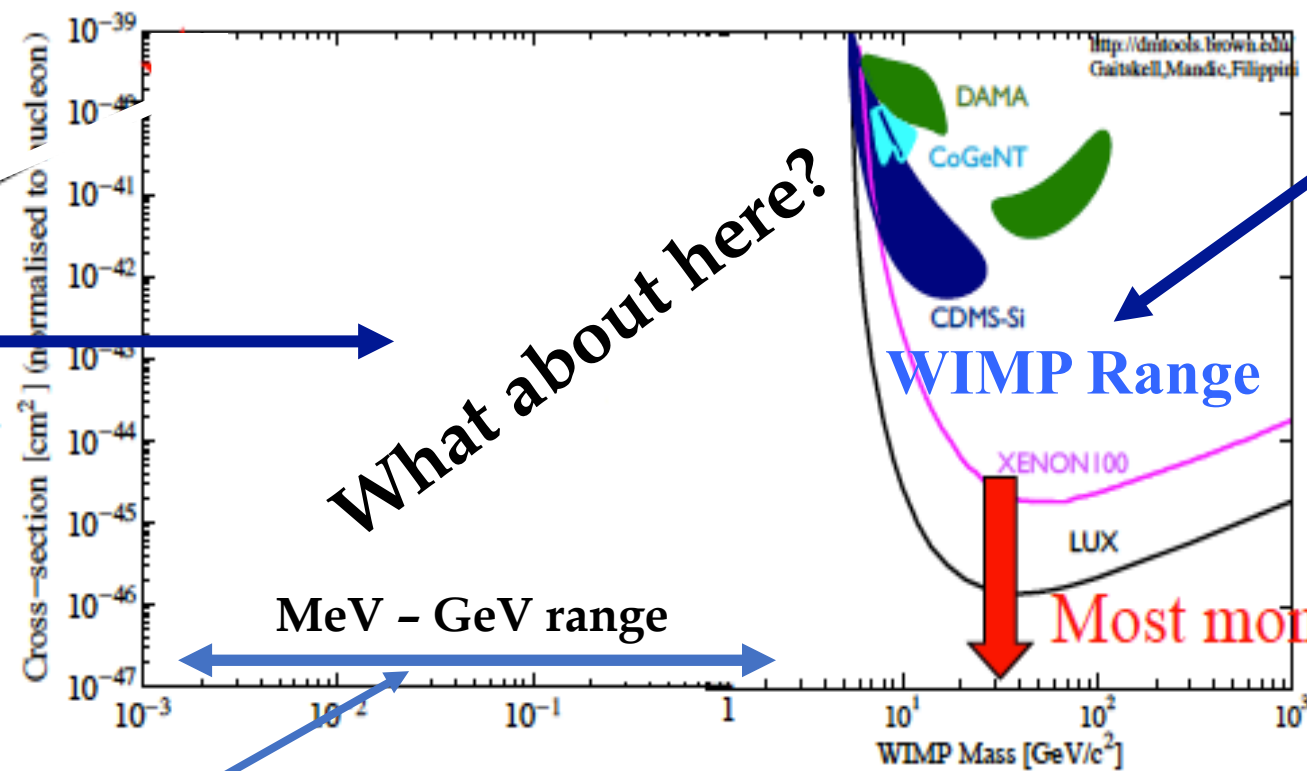
$G_{\mu\nu} \tilde{G}^{\mu\nu}$  - axion portal, dimension 4, new pseudo-scalars

...

B-hypercharge field, H - Higgs field, L- leptonic doublet

# MeV-GeV region for Light Dark Matter with thermal origin

Light Dark Matter  
with light mediators  
(hence new forces)  
that couple very weakly to SM  
and are therefore *hidden*  
or *dark*

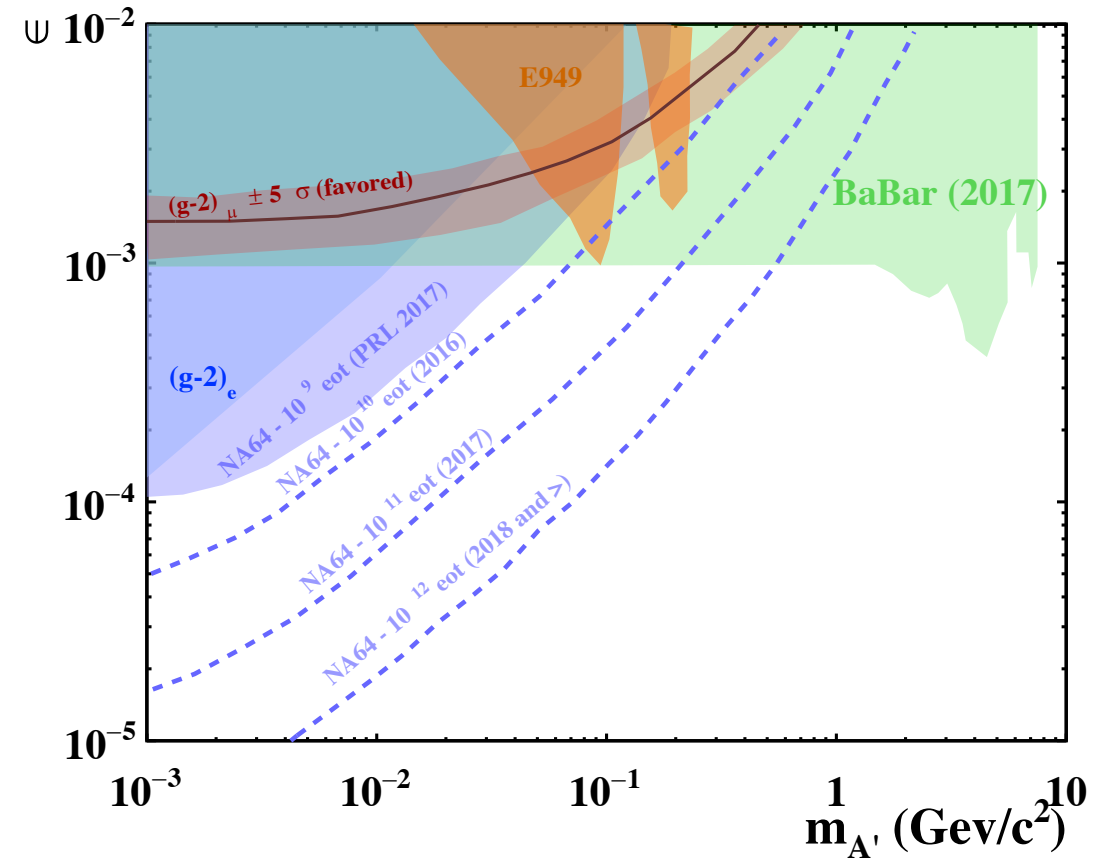
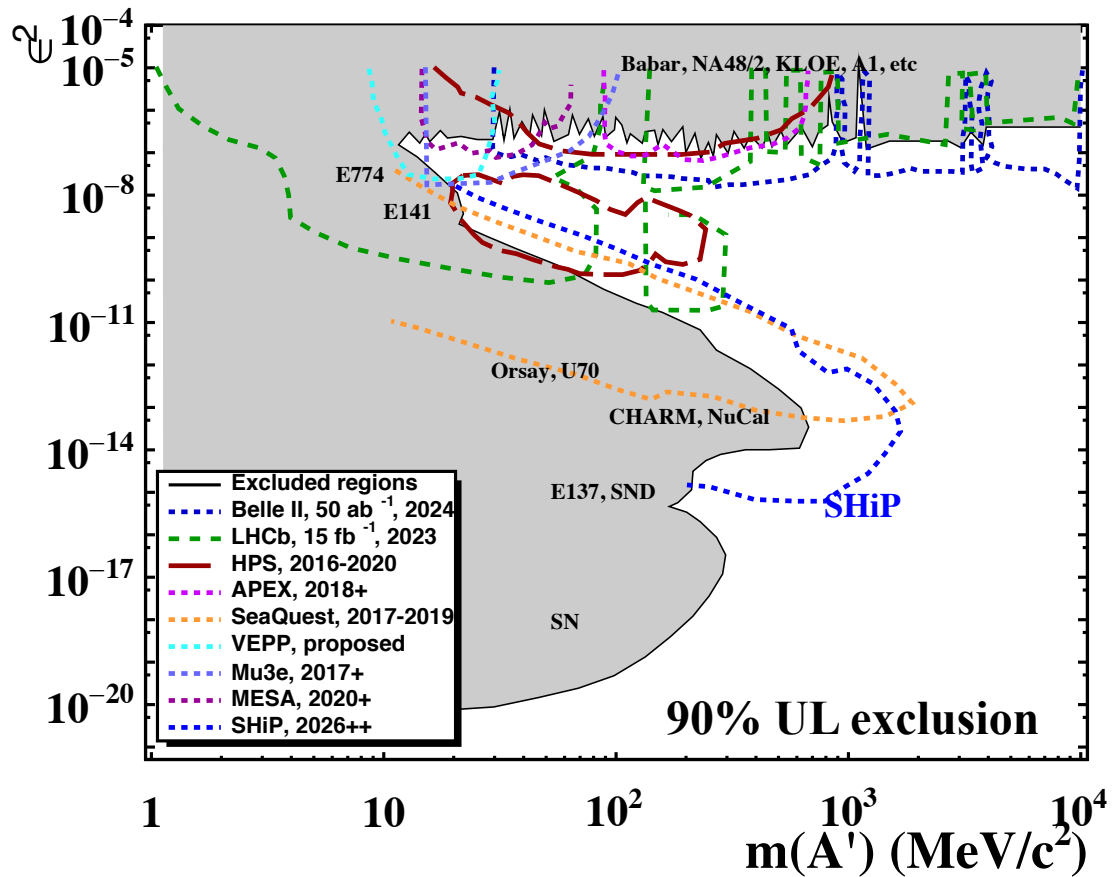


traditional WIMP at TeV  
scale with Z mediator  
(excluded by current limits)

Courtesy of M. Pospelov

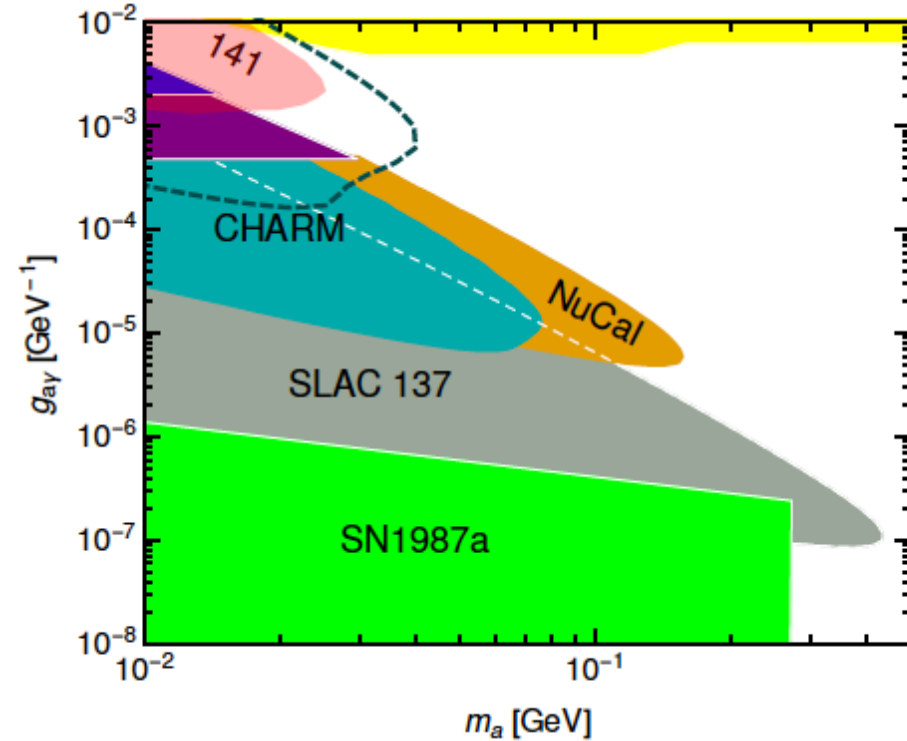
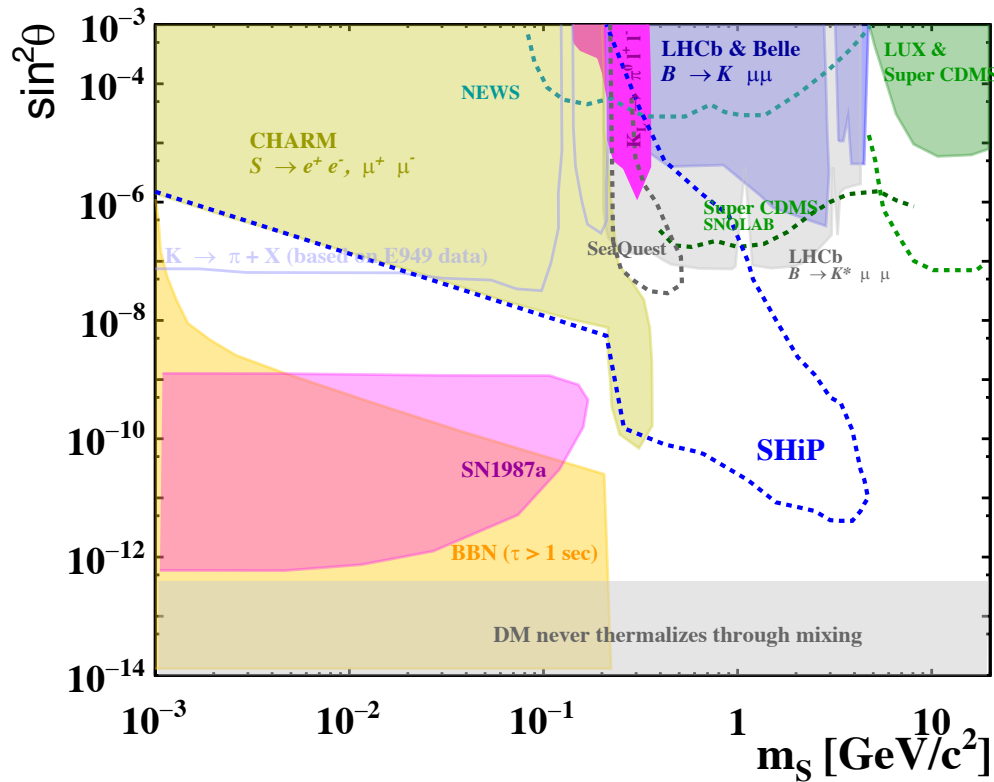
Range compatible with the relic density

# Current exclusion limits and prospects for Dark Photon into visible and invisible (DM) decays



Dark Photon produced via kinetic mixing with ordinary photons

# Current exclusion limits and prospects for Dark Scalar and ALPS with photon coupling

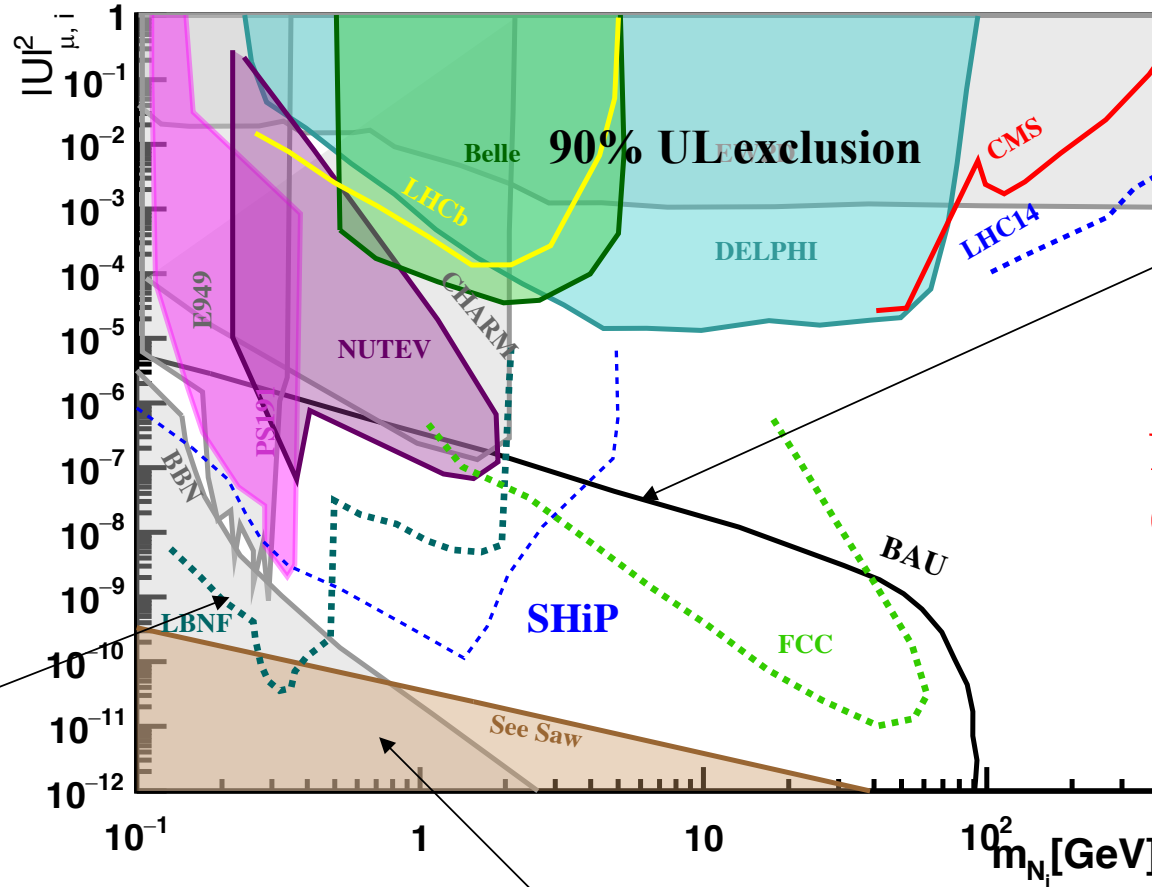


Scalars produced via kinetic mixing with the Higgs, either real or virtual  
(via FCNC loops in kaon and b decays)

arXiv:1708.05776

# Current exclusion limits and prospects for Heavy Neutral Leptons

$U_e^2 \cdot U_\mu^2 \cdot U_\tau^2 = 1:16:3.8$   
 Normal hierarchy of active  
 neutrino masses



**BAU generation** requires out of equilibrium: mixing angle of  $N_{2,3}$  to active neutrinos cannot be too large

**Low couplings: very long lifetimes (10-10000 km at the SPS energies)**

**BBN:** decays of  $N_{2,3}$  must not spoil Big Bang Nucleosynthesis ( $T \sim 1$  MeV)

**See-saw:** mixing angle of  $N_{2,3}$  to active neutrinos cannot be too small

# How to produce Hidden Particles in the MeV-GeV range?

---

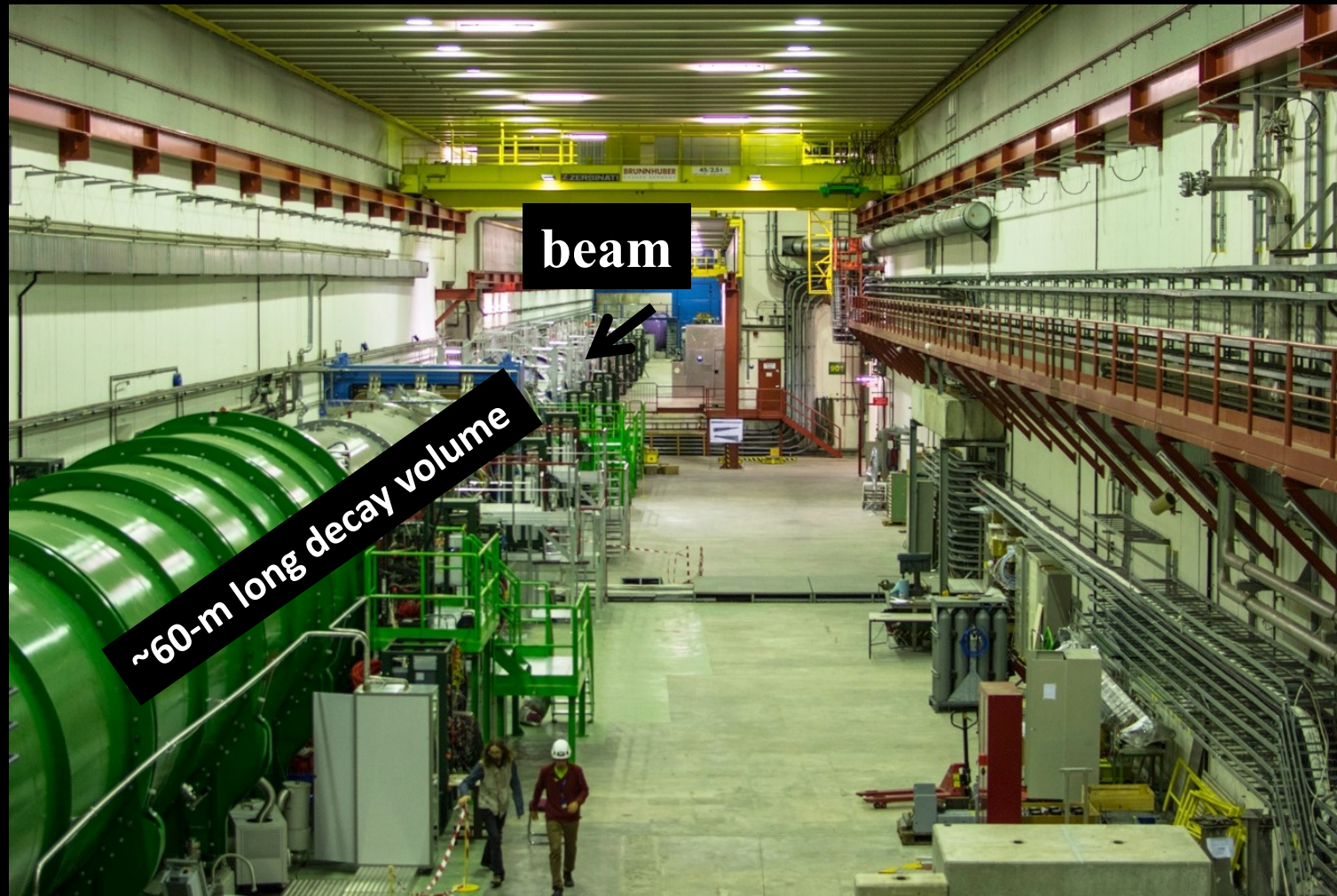
1) Light and feebly-interacting particles can be **originated by the decay of beauty, charm and strange hadrons and by photons produced in the interaction of protons with a target**. As the charm and beauty cross-sections increase steeply with the energy, **a high-intensity, high-energy proton beam is required to improve over the current results**:

→ *To date the world best line to produce high intensity fluxes of beauty and charm hadrons and photons through the interactions of protons on a high-Z target is a 400 GeV/c proton beam line extracted from the CERN SPS .*

2) The smallness of the couplings implies that the hidden sector mediators are also **very long-lived** (up to several km at the SPS energies) compared to the bulk of the SM particles:

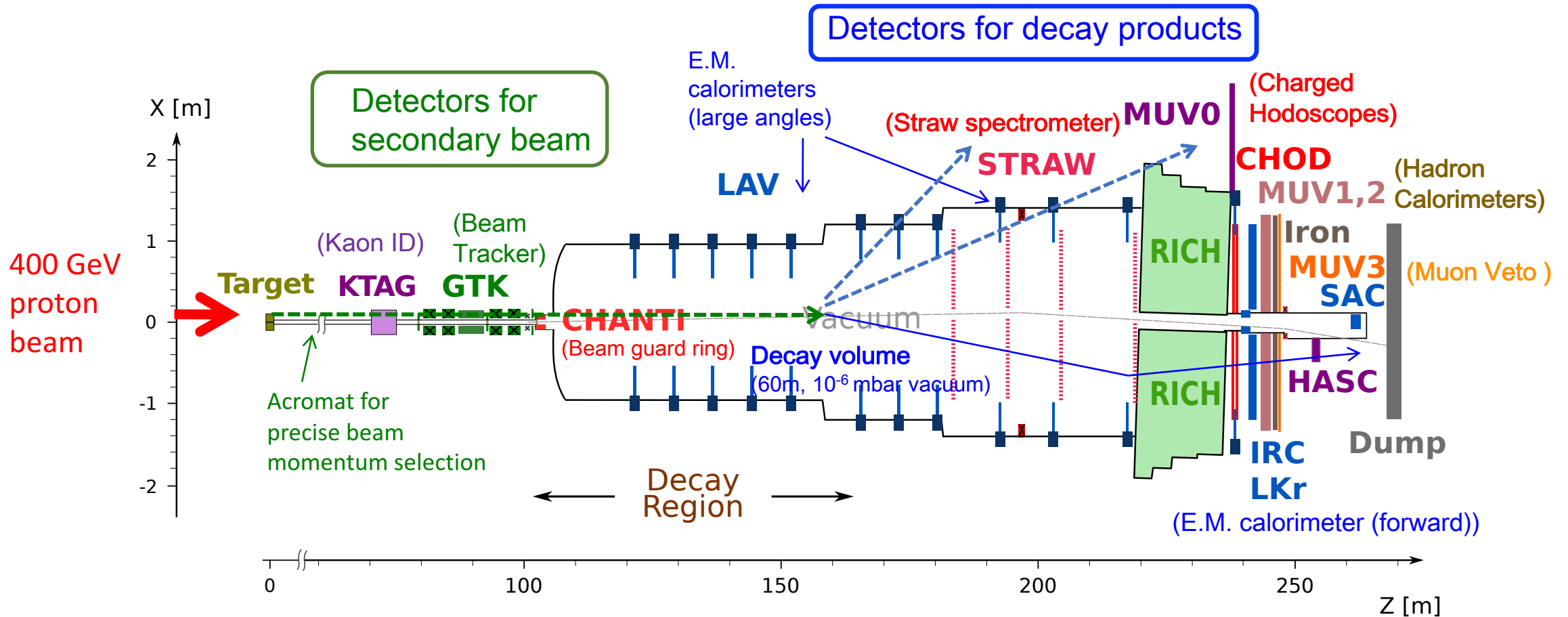
→ *The decays to SM particles can optimally be detected only using an experiment with decay volume tens of meters long followed by a spectrometer with particle identification capabilities.*

# The NA62 experiment in the ECN3 experimental hall



The NA62 detector, served by the P42 400 GeV proton line and designed to measure the branching fraction of the rare decay  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ , is perfectly suited to search for hidden particles

# The NA62 experiment: conceptual scheme

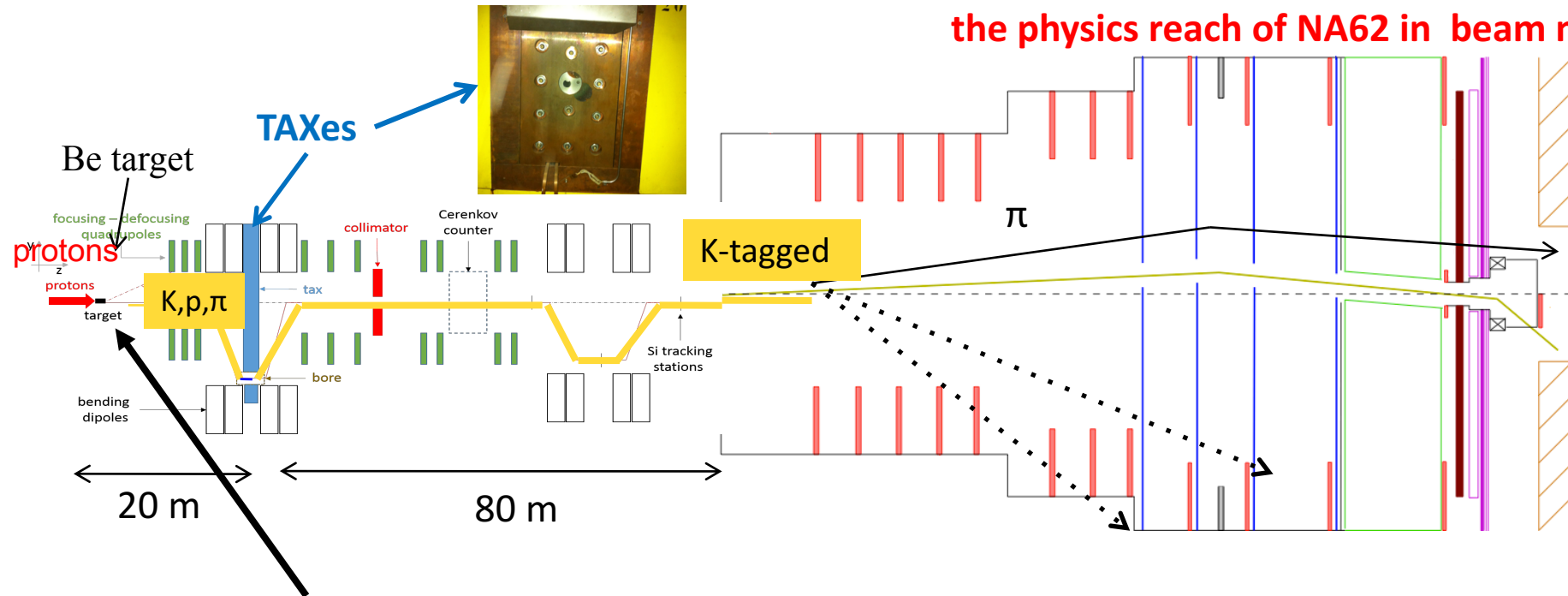


**NA62 ~ (2.0-3.0)x10<sup>18</sup> protons-on-target/year @ 400 GeV/c**



# NA62 in beam operation mode

See A. Ceccucci and L. Bician's talks for a wide view of the physics reach of NA62 in beam mode

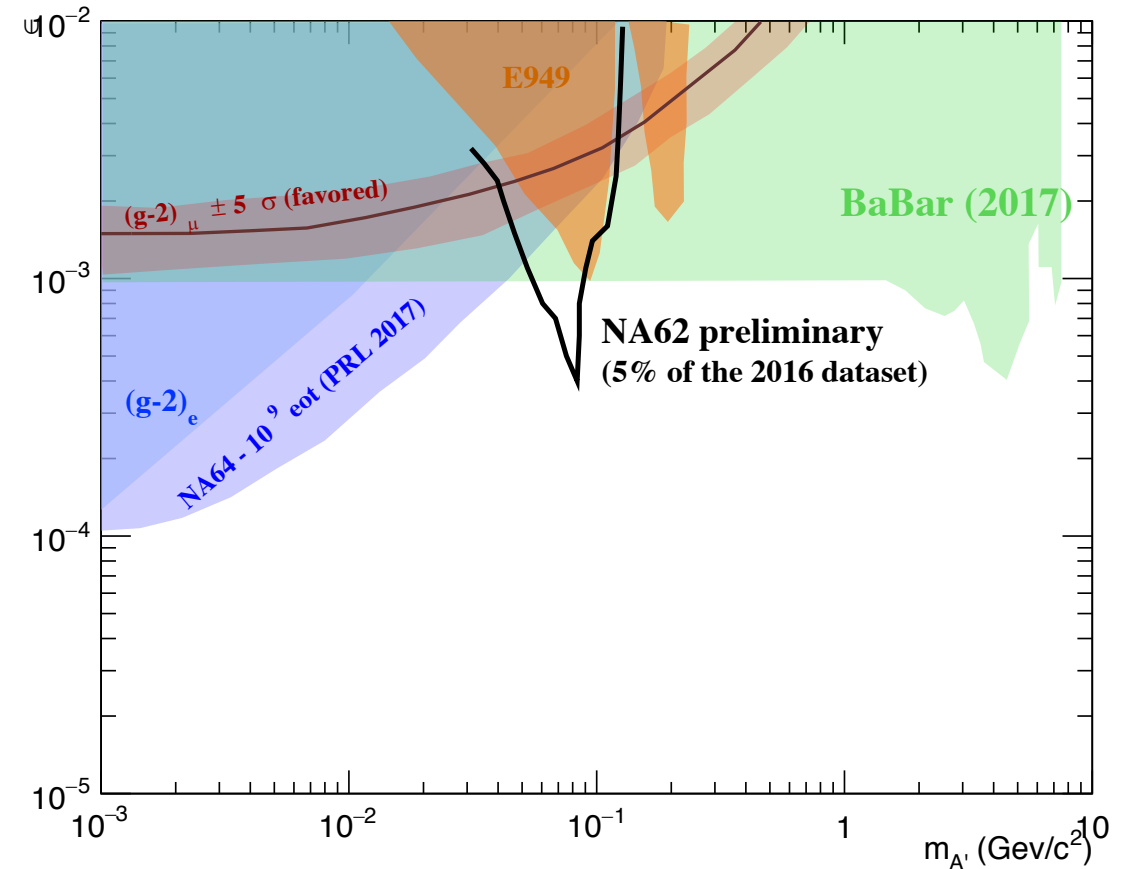
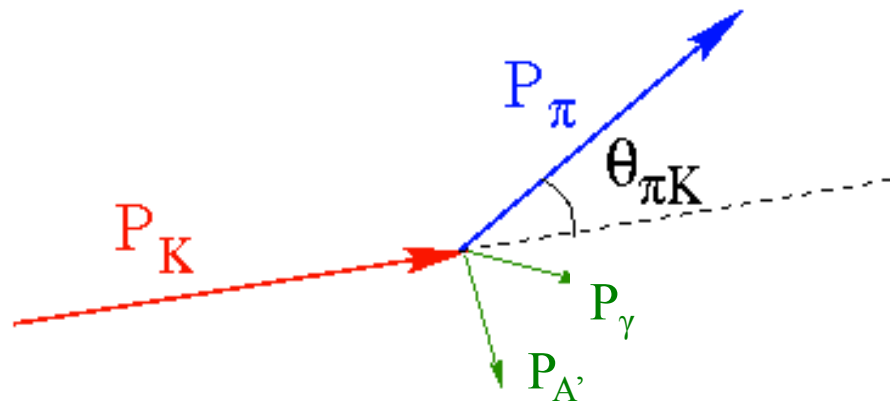


The 400 GeV/c proton beam impinges on a Be-target producing a mixed beam of K, pions, protons. A magnetic system and two 1.6-m long, water-cooled, copper collimators (TAXes) select a monochromatic beam of 75 GeV/c momentum. Kaons are tagged via a Cerenkov detector.

# Search for Hidden Particles from kaon decays: Dark Photon $\rightarrow$ invisible

Use the decay  $K^+ \rightarrow \pi^+ \pi^0, \pi^0 \rightarrow \gamma A'$  BR  $\sim 20\%$

- Reconstruct  $K^+, \pi^+$ , one photon
- Constrain  $A'$  using the  $\pi^0$  mass
- Limit improves linearly with  $N(K^+)$ ;
- Analysis with the full data set in progress.

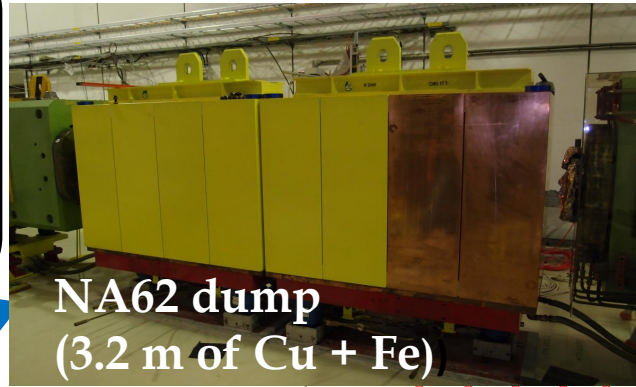


Search for Hidden Particles from Kaon decays limits the reach below the kaon mass

# NA62 in “dump” operation mode

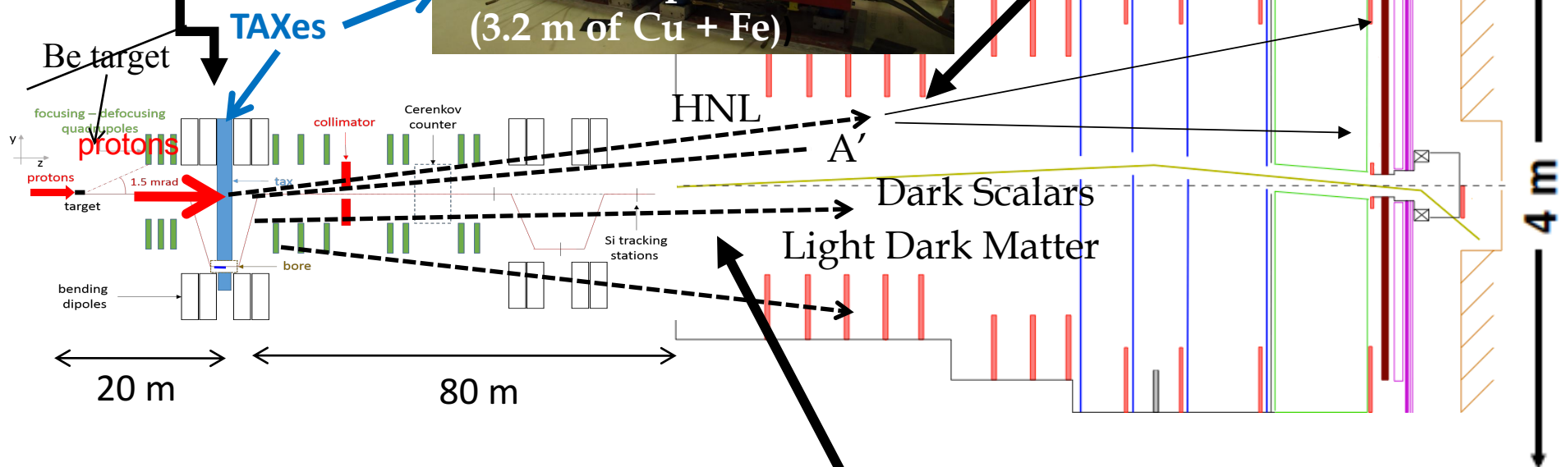
In dump mode the target can be moved away from the beam and the beam let impinging on the copper. Hence: the TAXes can act as a dump ( $2 \times 10.7 \lambda_I$ ).

→ this operation is easy, quick (15 minutes) and fully reversible.



**Signal signature:**

a vertex appearing in the decay volume and nothing else

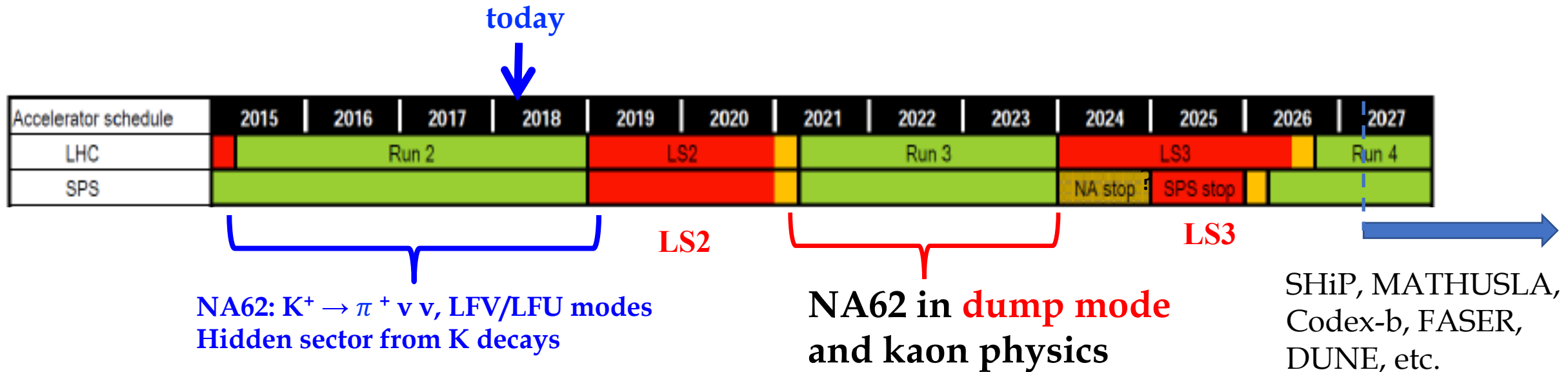


Heavy Neutral Leptons, Dark Photons, Dark scalars, and ALPS can be originated by charm, beauty and photons produced in the interaction of protons with the dump.

# NA62 in kaon and dump modes: scientific scheduling

NA62 has the main goal of measuring the  $\text{BR}(\text{K}^+ \rightarrow \pi^+ \nu \bar{\nu})$  with 10% accuracy;

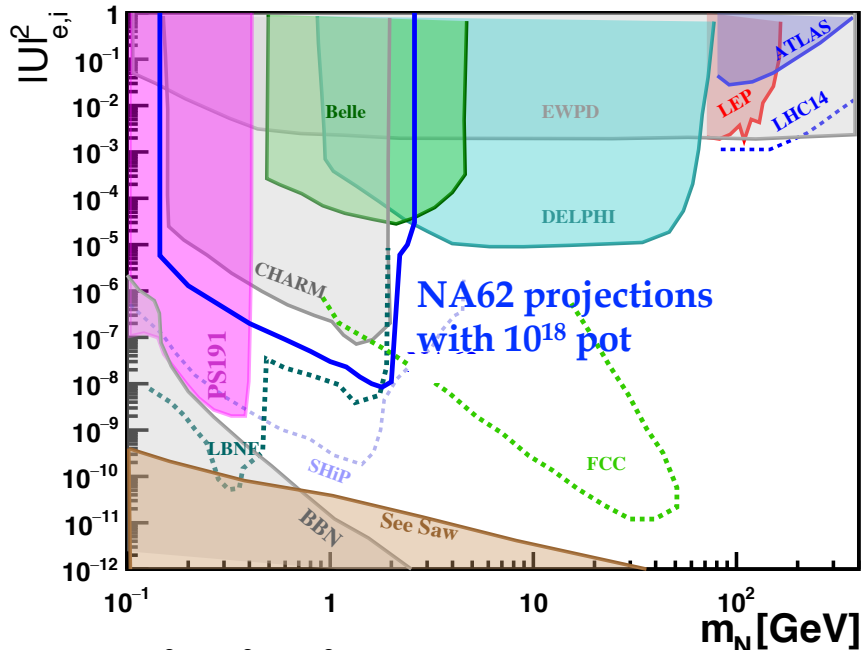
- **Before LS2 (2017-2018)** many searches in the hidden sector will be performed using the kaon beam.
- **After LS2 (2021++)** there is a window of opportunity to run NA62 in beam-dump mode to collect at least  $10^{18}$  pot to search for hidden particles from charm and beauty decays, and photons.



**Goal: integrate at least  $\sim 10^{18}$  pot in dump mode by 2023**  
(corresponding to  $\sim 3$  months of dedicated data taking in 2021-2023)

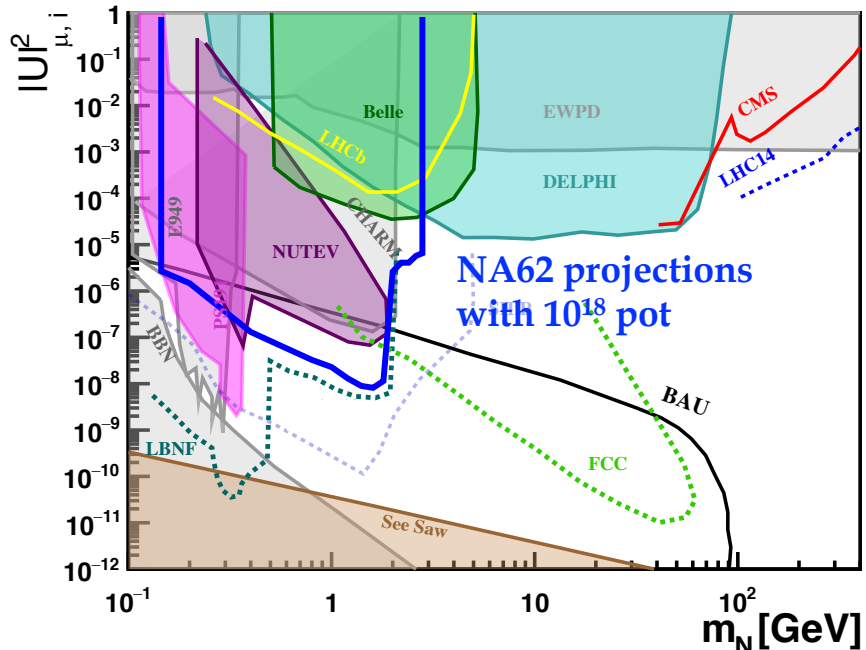
# NA62 sensitivity with $10^{18}$ pot in dump mode: Heavy Neutral Leptons

Scenario 1



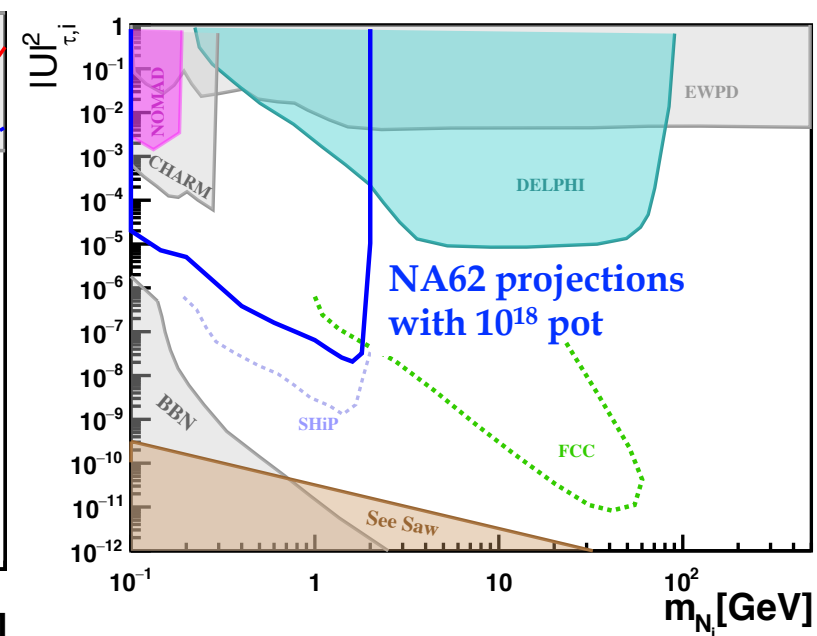
$U^2_e : U^2_\mu : U^2_\tau = 52 : 1 : 1$   
 Normal hierarchy of active  $\nu$  masses

Scenario 2



$U^2_e : U^2_\mu : U^2_\tau = 1 : 16 : 3.8$   
 Normal hierarchy of active  $\nu$  masses

Scenario 3

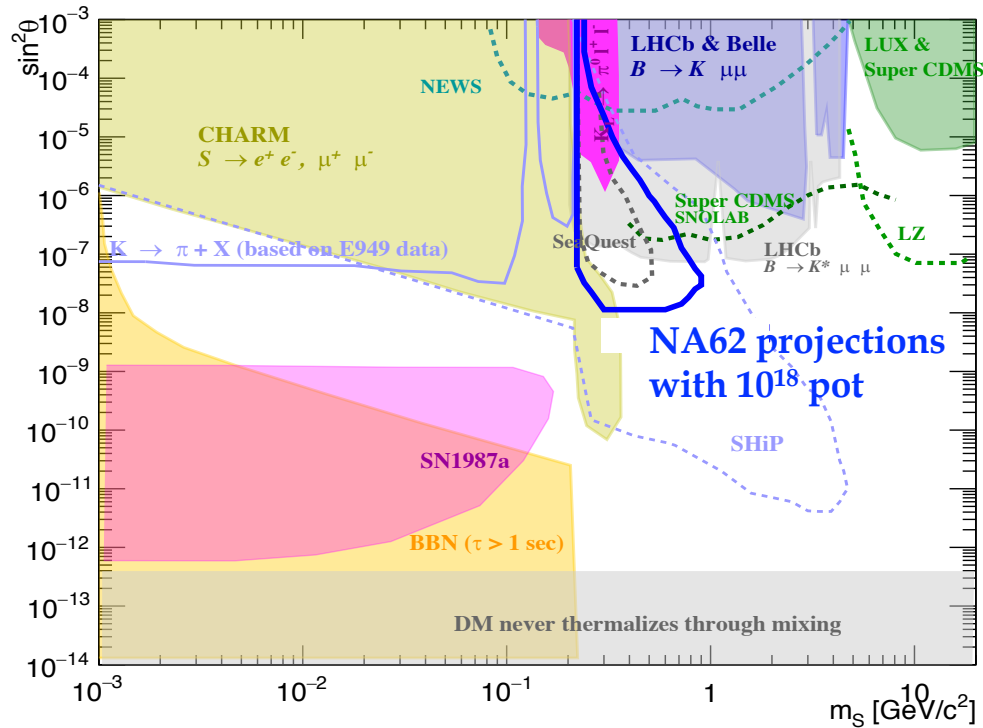


$U^2_e : U^2_\mu : U^2_\tau = 0.061 : 1 : 4.3$   
 Normal hierarchy of active  $\nu$  masses

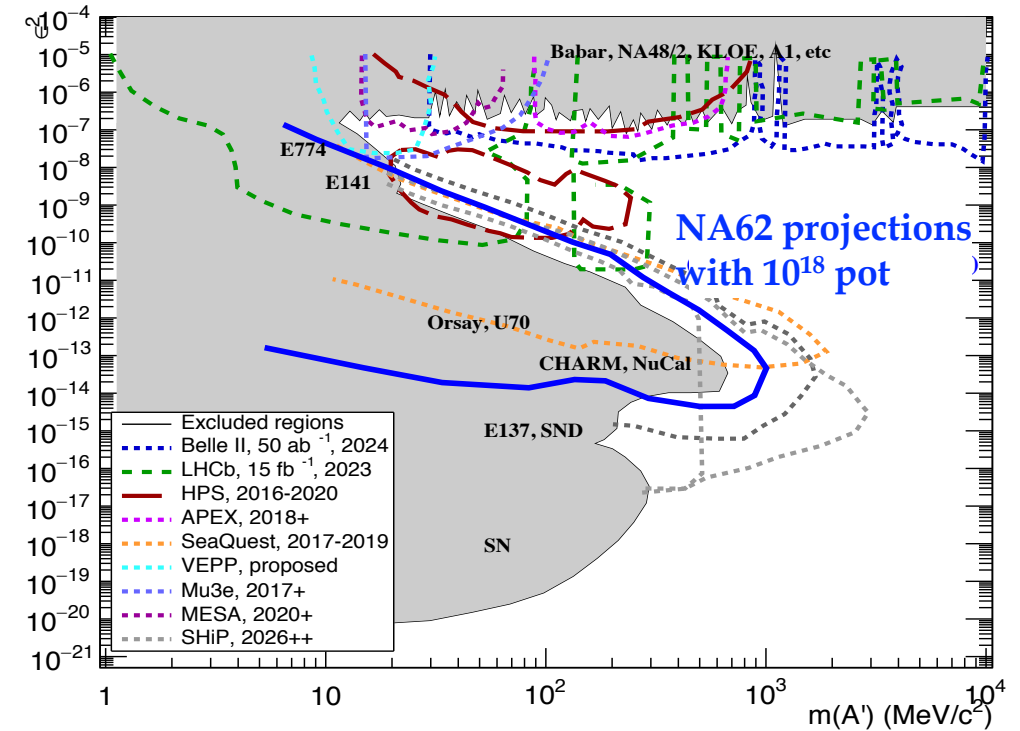
These sensitivities assume to detect all 2-track final states, including open channels, and zero background.

# NA62 sensitivity with $10^{18}$ pot in dump mode

## Dark Scalar:



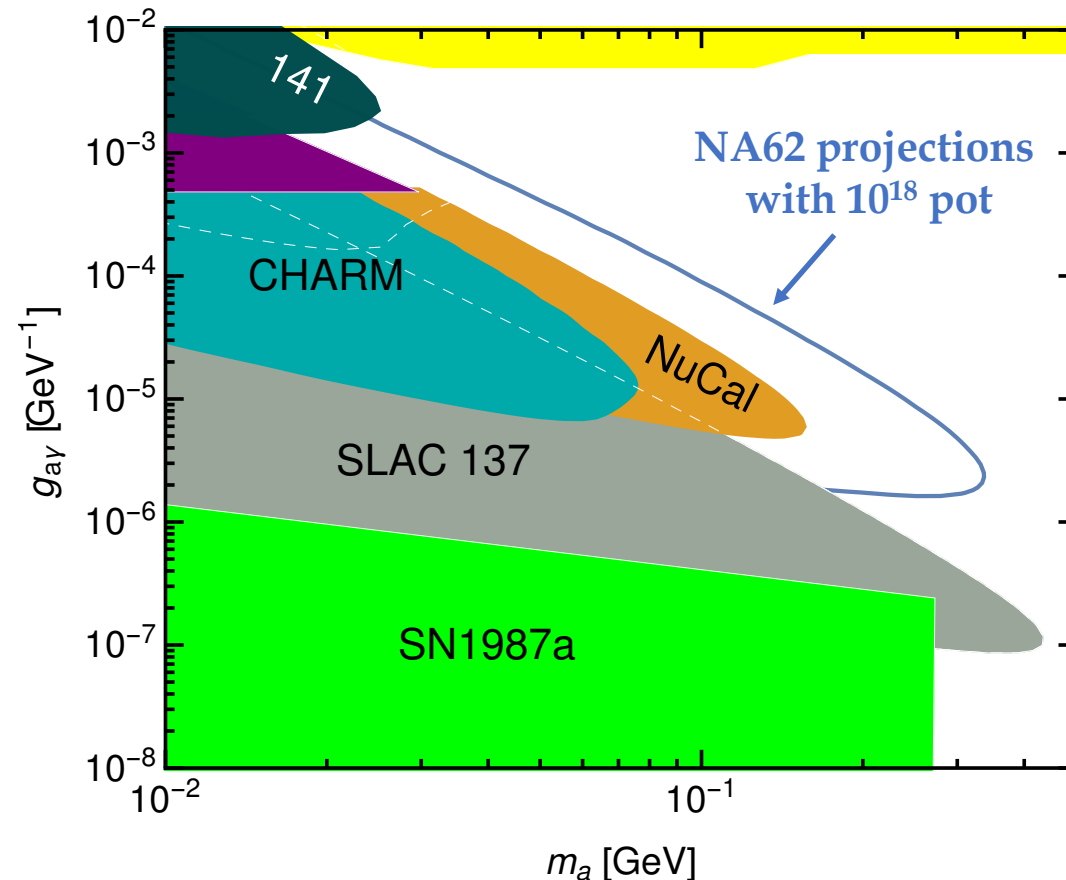
## Dark Photon:



The Dark Scalar sensitivity plot assume all 2-track final states, with zero background  
 The Dark Photon projections are for di-muon final state only and (still) miss the inclusion of the two dominant production processes (QED, QCD): the curve is very conservative!

# NA62 sensitivity for ALPs $\rightarrow \gamma\gamma$ in dump mode

- study ALP production via Primakoff effect (JHEP 1602 (2016) 018) at target;
- search for ALP  $\rightarrow \gamma\gamma$  in NA62 fiducial volume, account for geometrical acceptance
- assume zero-background, evaluate expected 90%-CL exclusion plot



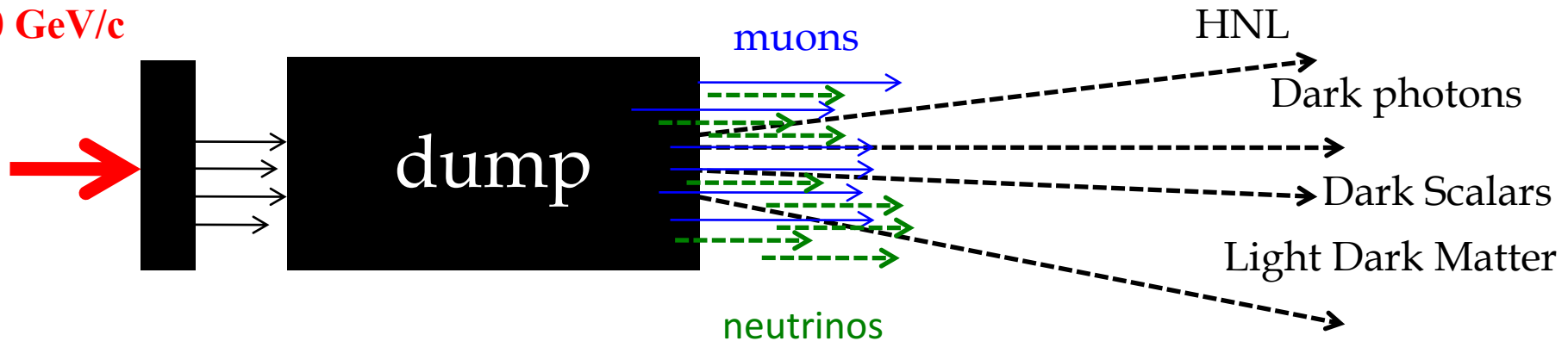
See also Dobrich et al., JHEP 02 (2016) 018 for details.

# ....Background, background, background.....

---

A dump with suitable length stops all beam-induced backgrounds but neutrinos and muons:

**SPS protons**  
**400 GeV/c**

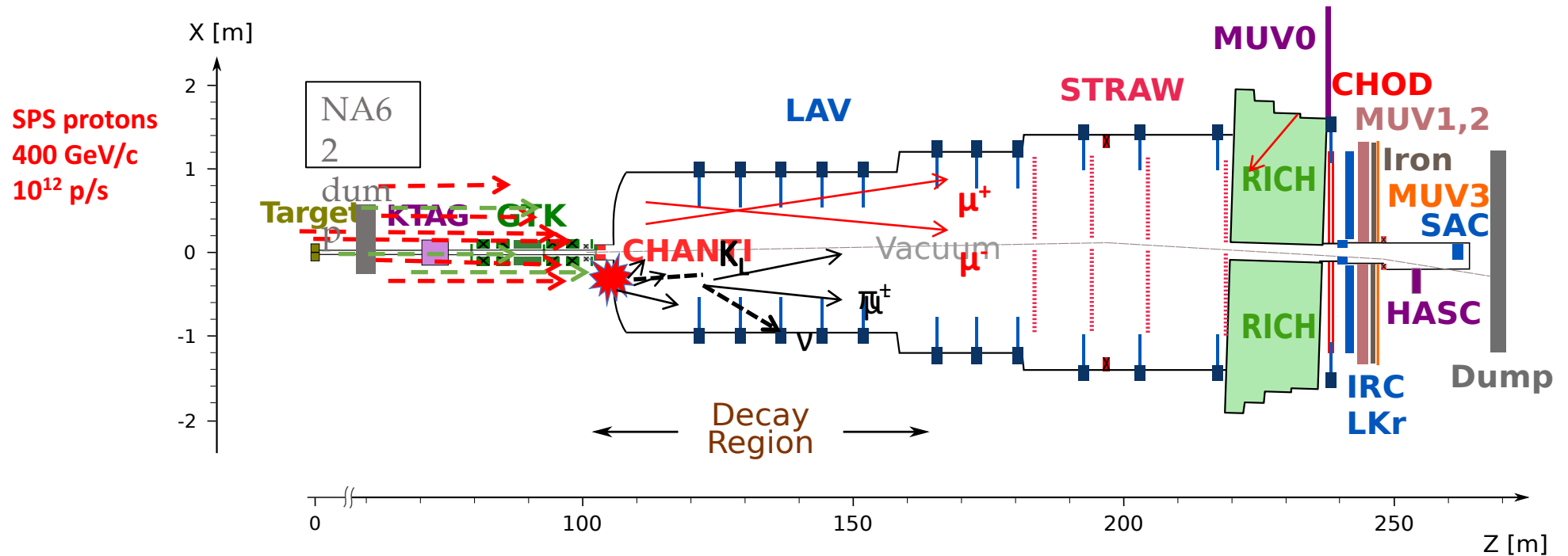


Any kind of feebly-interacting  
long-lived particle or LDM  
(put here your favored model);

Background is the name of the game.



# Main Backgrounds in dump mode



## Muon halo & neutrino halo

- In *beam mode* about  $\sim 5$  MHz of  $\mu^+$  and 150 kHz  $\mu^-$  are present due to early decays in flight of K and  $\pi$  in the beam;
- In *dump mode*, the muon halo is reduced by (at least) 2 orders of magnitudes
- **Muons produce inelastic interactions and combinatorial background**
- **Neutrinos can produce inelastic interactions in the material surrounding the FV.**

# NA62-DUMP: data driven background estimate

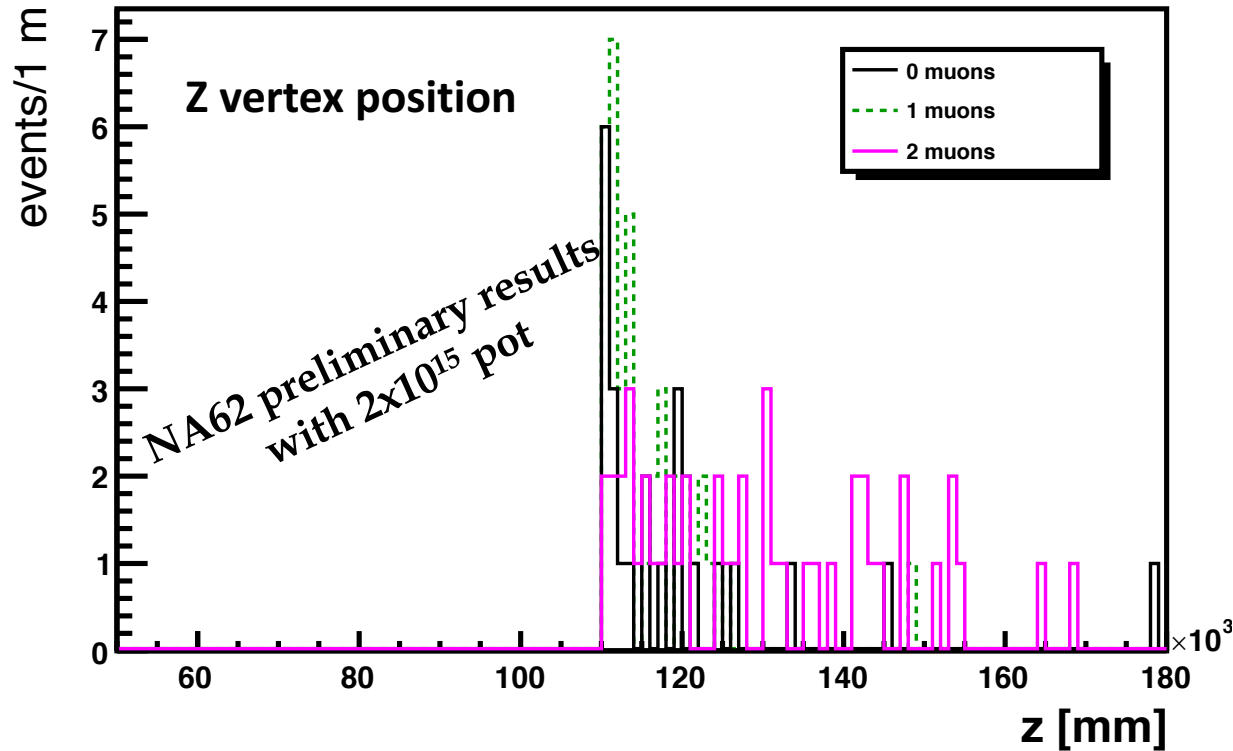
---

In November 2016, few hours long run taken in dump mode,  $\sim 2 \times 10^{15}$  pot collected for ALPs search and in preparation for the longer physics runs in dump mode in 2021-2023:

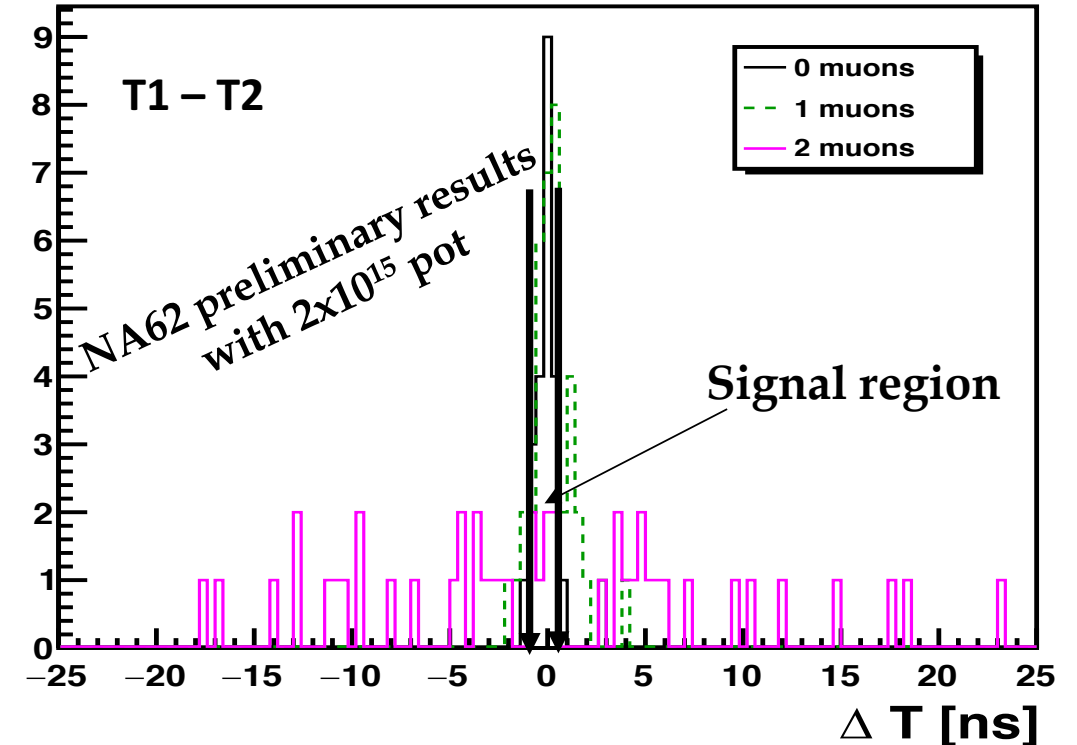
- ✓ - This small dataset already provides crucial information on the background level;
- ✓ - About x5 more data in dump mode has been collected in 2017 and is currently being analyzed.
- ✓ - Further runs planned for 2018 to complete the study of beam-dump operation before LS2.

# NA62-DUMP: data driven background estimate

Distribution of the z-coordinate of the vertex



Distribution of the 2-track time difference



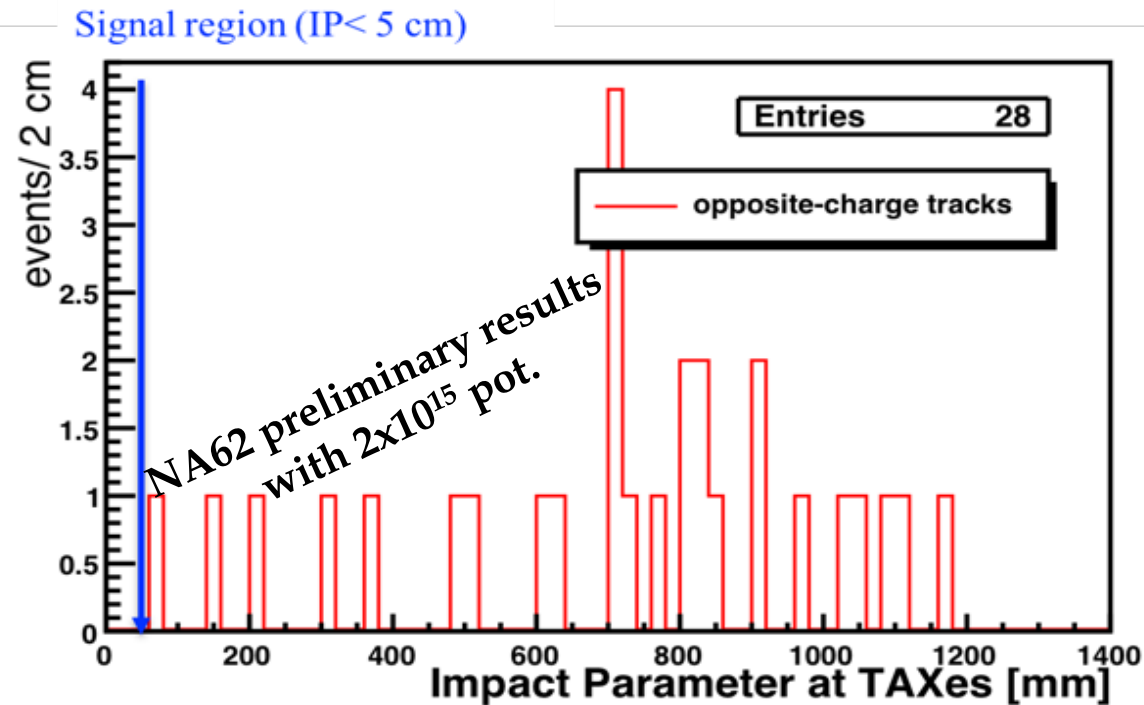
Sample divided in 2-, 1- and 0-muon categories:

- 2- $\mu$  sample has vertices spread along the FV and tracks mostly out-of-time:  
→ combinatorial background
- 0- or 1- $\mu$  samples are concentrated at the beginning of the FV, and tracks are mostly in-time:  
→ background from inelastic interactions in the last  $\lambda_1$  of the final collimator

# NA62-DUMP: data driven background estimate

Consider only opposite charge events with two tracks in time within 1 ns:

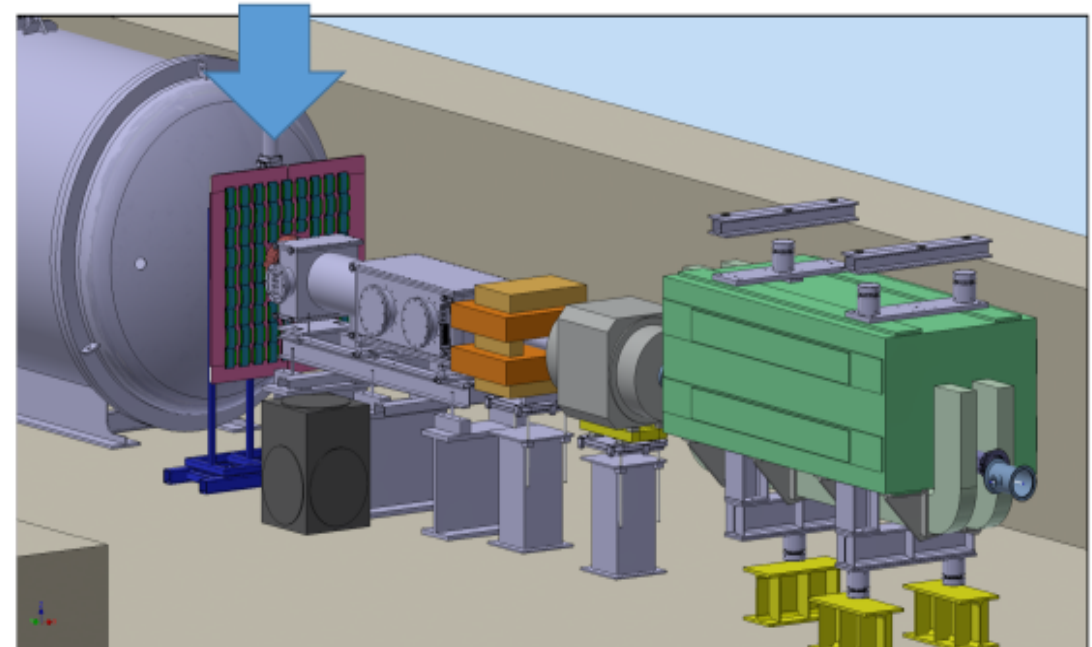
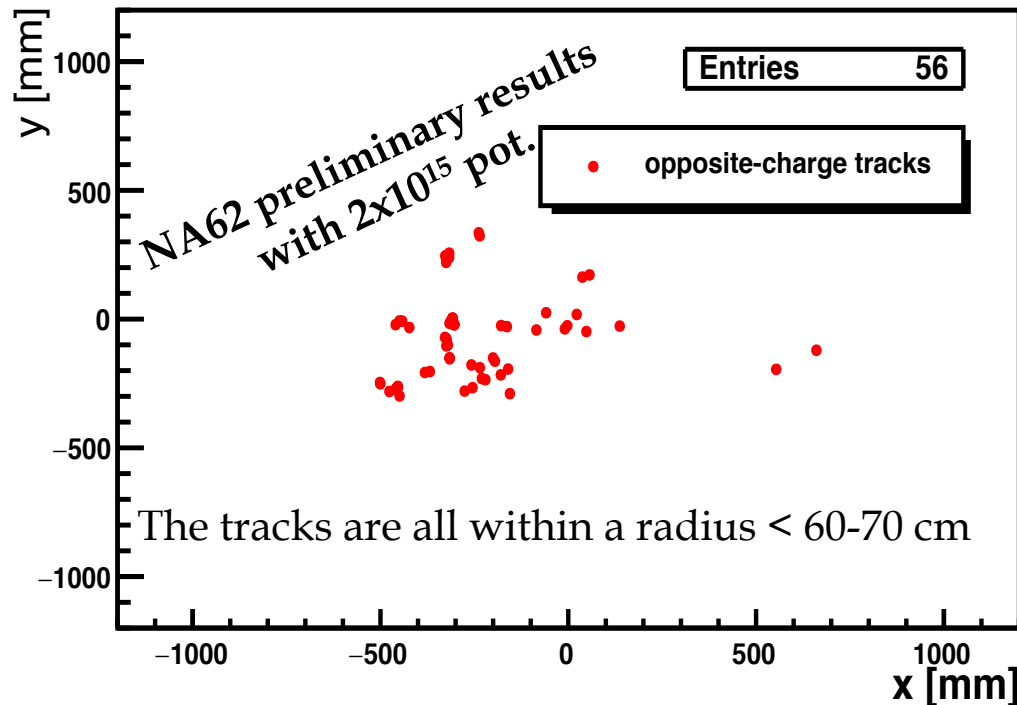
- 28 events are left out of  $2 \times 10^{15}$  pot
- All of them do not point backwards to TAXes (as expected)
- Requiring a mild pointing cut  $IP < 5$  cm we can easily reject all of them.....



..... but this implies we reject also partially reconstructed final states (eg: most of HNL decays)

# NA62-DUMP: extrapolation in front of the decay vessel

Extrapolation of the 2-tracks of the remaining 28 events at the beginning of the decay vessel:  
→ they are all concentrated either in the “empty” zone not covered by any detector in NA62



An Upstream Veto in front of the decay vessel with  $< 10^{-4}$  detector inefficiency (currently missing in the current setup) could veto most (all) of them, and largely broaden the physics reach of NA62.

# Conclusions

---

- **NA62 is successfully running in ECN3 with the main goal of measuring the BR( $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ ) with 10% accuracy.**
- **NA62 can be a major player in the search for hidden sector particles in Run 3:**
  - in Run 2 & 3: by searching for hidden particles in kaon decays;
  - in Run 3 : by collecting (at least)  $10^{18}$  pot in dump mode by 2023.
- **Preliminary studies with data taken in beam-dump modes show that the background can be kept under control.**

More data collected in 2017 and 2018 will allow us to consolidate this result.  
Further improvements in the setup are currently being investigated.

Thank you for your attention!



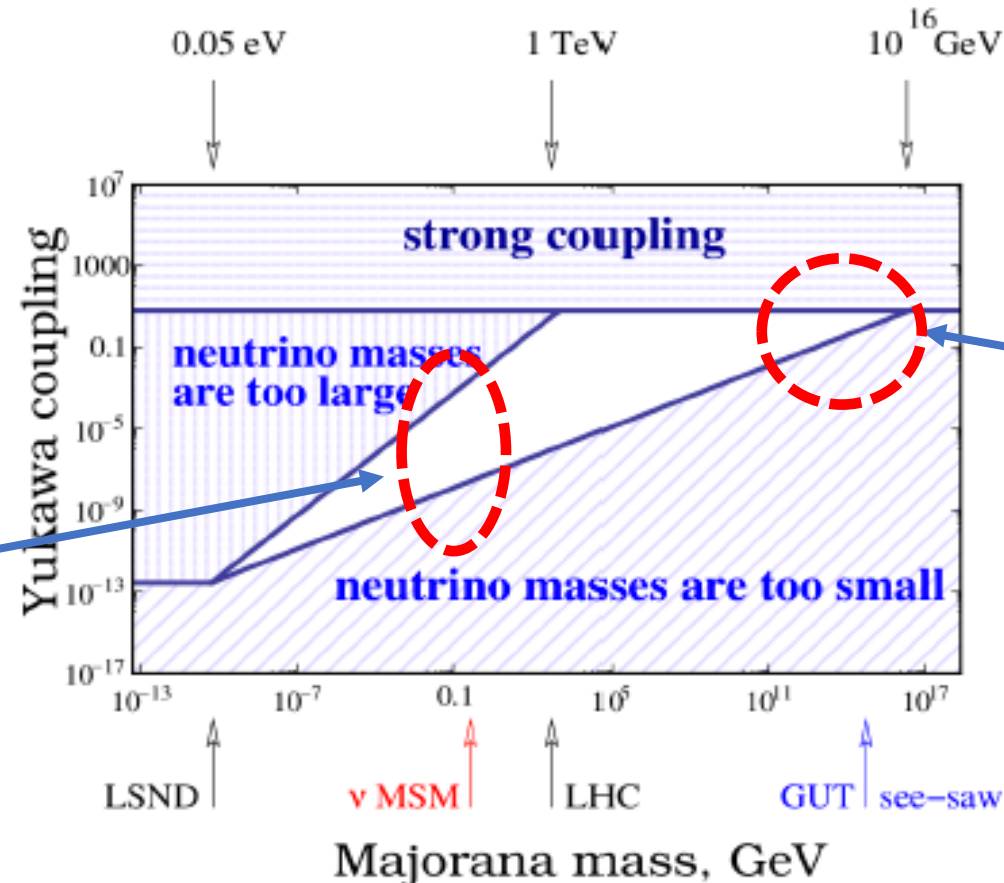
SPARES



# Heavy neutrino Yukawa couplings versus Majorana mass

It is “natural” to assume that the masses of RH neutrinos are similar to all the other masses (~EW scale):

$$F \sim \frac{\sqrt{m_{atm} M_N}}{v} \sim (10^{-6} - 10^{-13}),$$



It “natural” to assume that the Yukawa couplings of the RH neutrinos are similar to top Yukawa:

$$M_N \simeq \frac{F^2 v^2}{m_{atm}} \simeq 6 \times 10^{14} \text{ GeV}$$

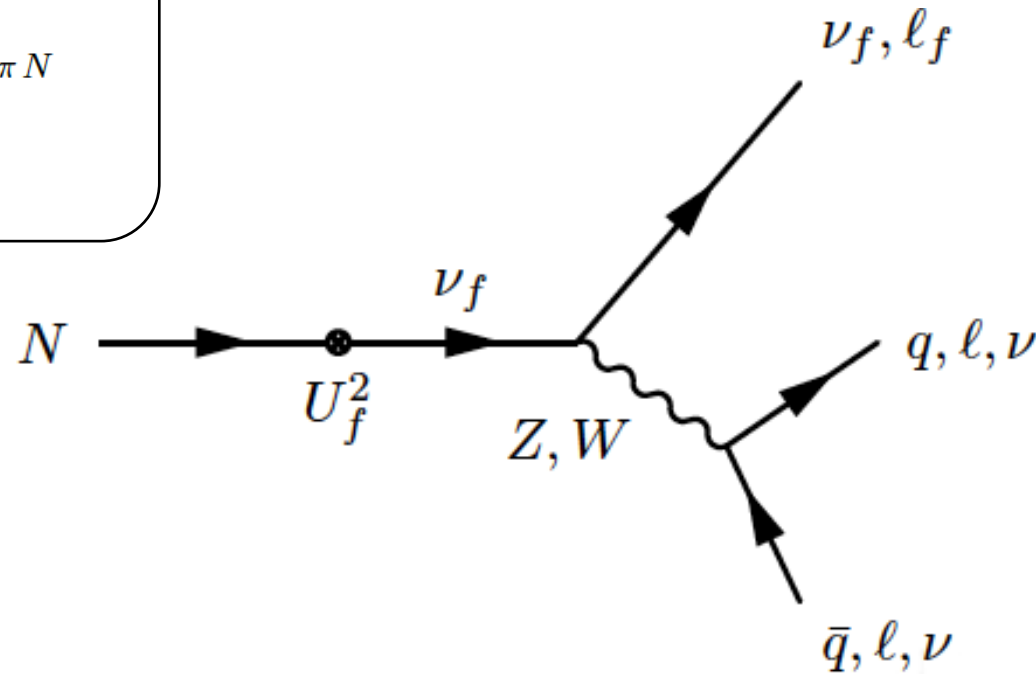
Any number in the white area works to be consistent with the observed pattern of neutrino masses and oscillations

# Production and decay of HNLs with mass in the MeV - GeV range

HNLs can be produced in decays where a neutrino is replaced by a  $N$  (kinetic mixing,  $U^2$ );  
**Main neutrino sources at the SPS energies: c and b mesons.**

## Production processes

- $D \rightarrow K \ell N$
- $D_s \rightarrow \ell N$
- $D_s \rightarrow \tau \nu_\tau$  followed by  $\tau \rightarrow \mu \nu N$  or  $\tau \rightarrow \pi N$
- $B \rightarrow \ell N$
- $B \rightarrow D \ell N$
- $B_s \rightarrow D_s \ell N$



## Decay channels

- $N \rightarrow H^0 \nu$ , with  $H^0 = \pi^0, \rho^0, \eta, \eta'$
- $N \rightarrow H^\pm \ell^\mp$ , with  $H = \pi, \rho$
- $N \rightarrow 3\nu$
- $N \rightarrow \ell_i^\pm \ell_j^\mp \nu_j$
- $N \rightarrow \nu_i \ell_j^\pm \ell_j^\mp$

They can then decay again to SM particles through mixing ( $U^2$ ) with a SM neutrino. This (now massive) neutrino can decay to a large amount of final states through emission of a  $Z^0$  or  $W$  boson