



Universität
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The Latest from the XENON Dark Matter Project

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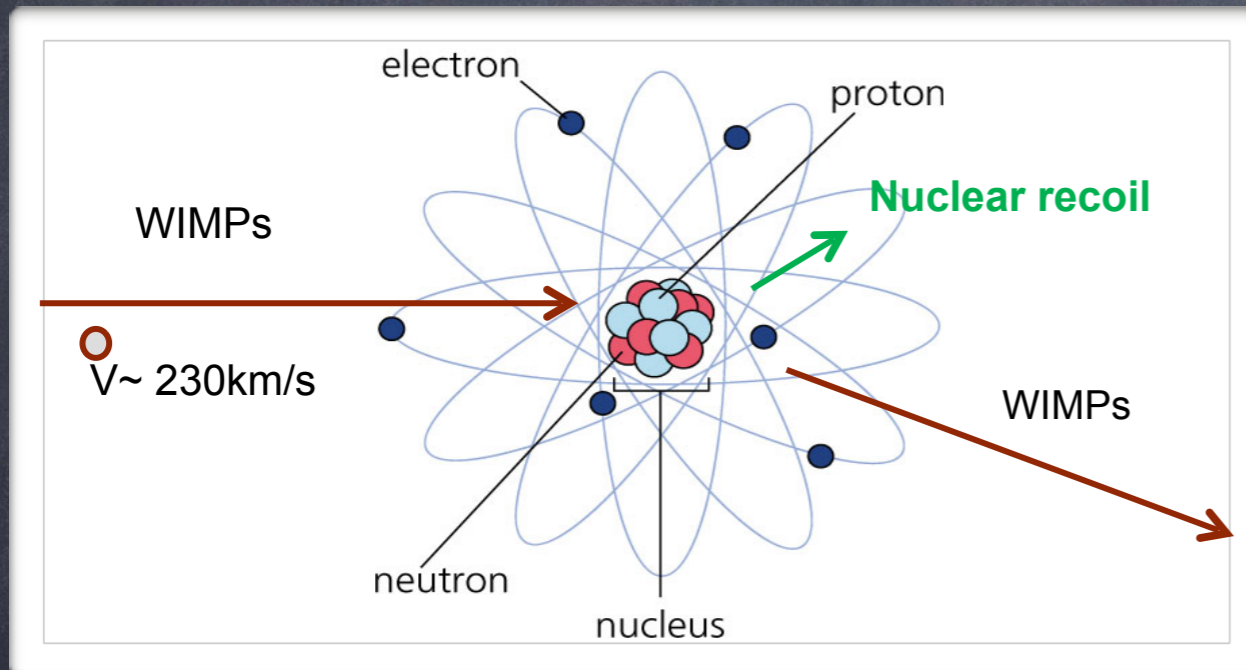
University of Zurich

For the XENON Collaboration

ALPS 2018

Direct Detection

Elastic scattering off nucleus

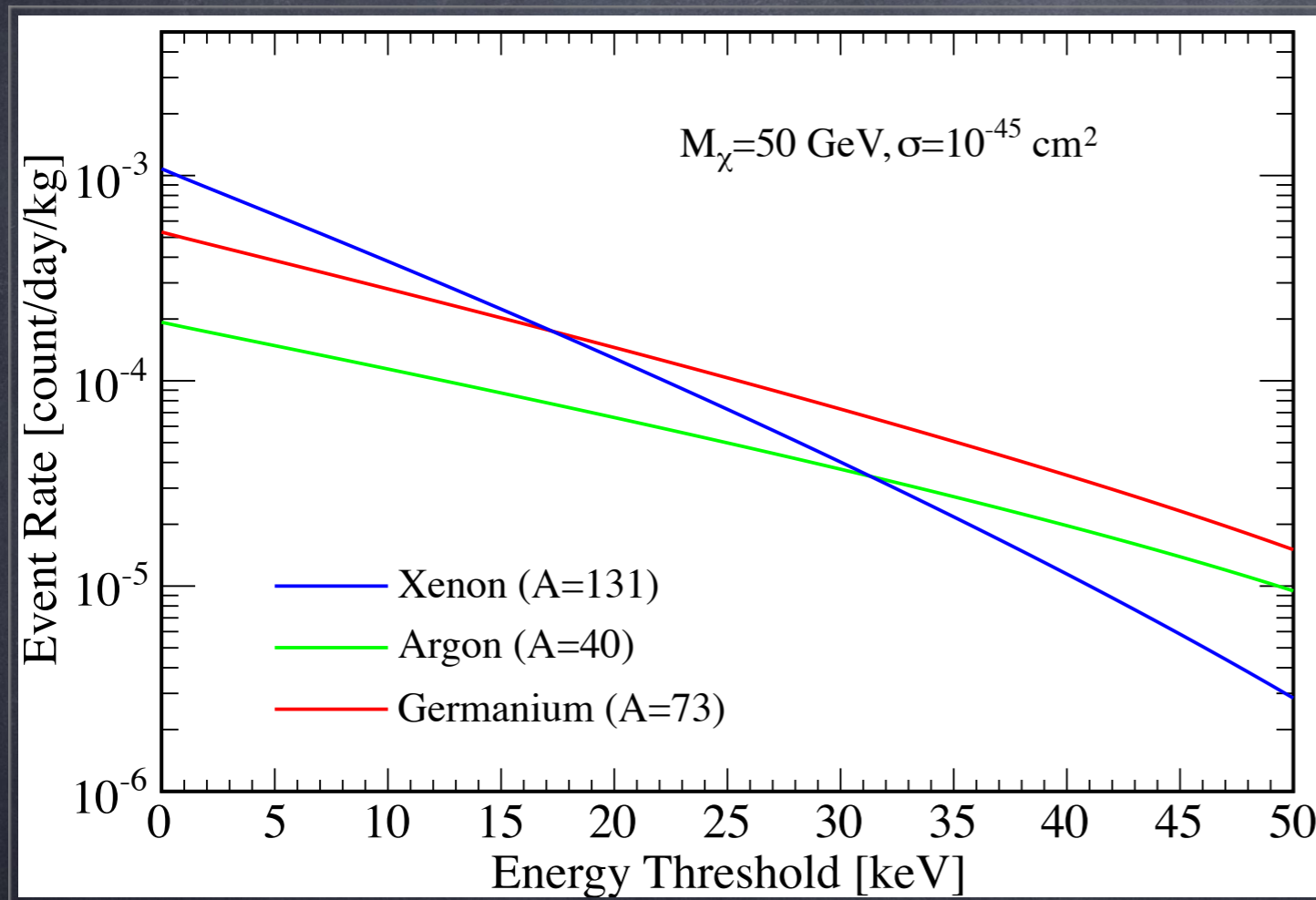


Recoils $\sim 10 \text{ keV}$

Requirements

- Large target mass
- Low energy threshold
- Ultra-low background

A Xenon Target



- Rate $\propto A^2$
- High purity
- “Easy” cryogenics at -95°C
- Large density
- Self-shielding
- Compact/scalable

Time Projection Chamber

- Energy

- Light signal (S1)

- Charge signal (S2)

- Position

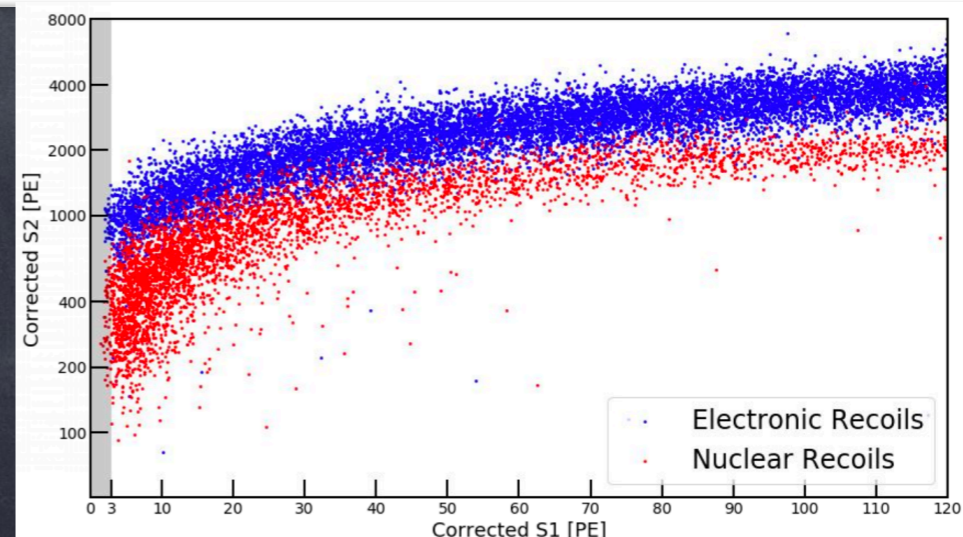
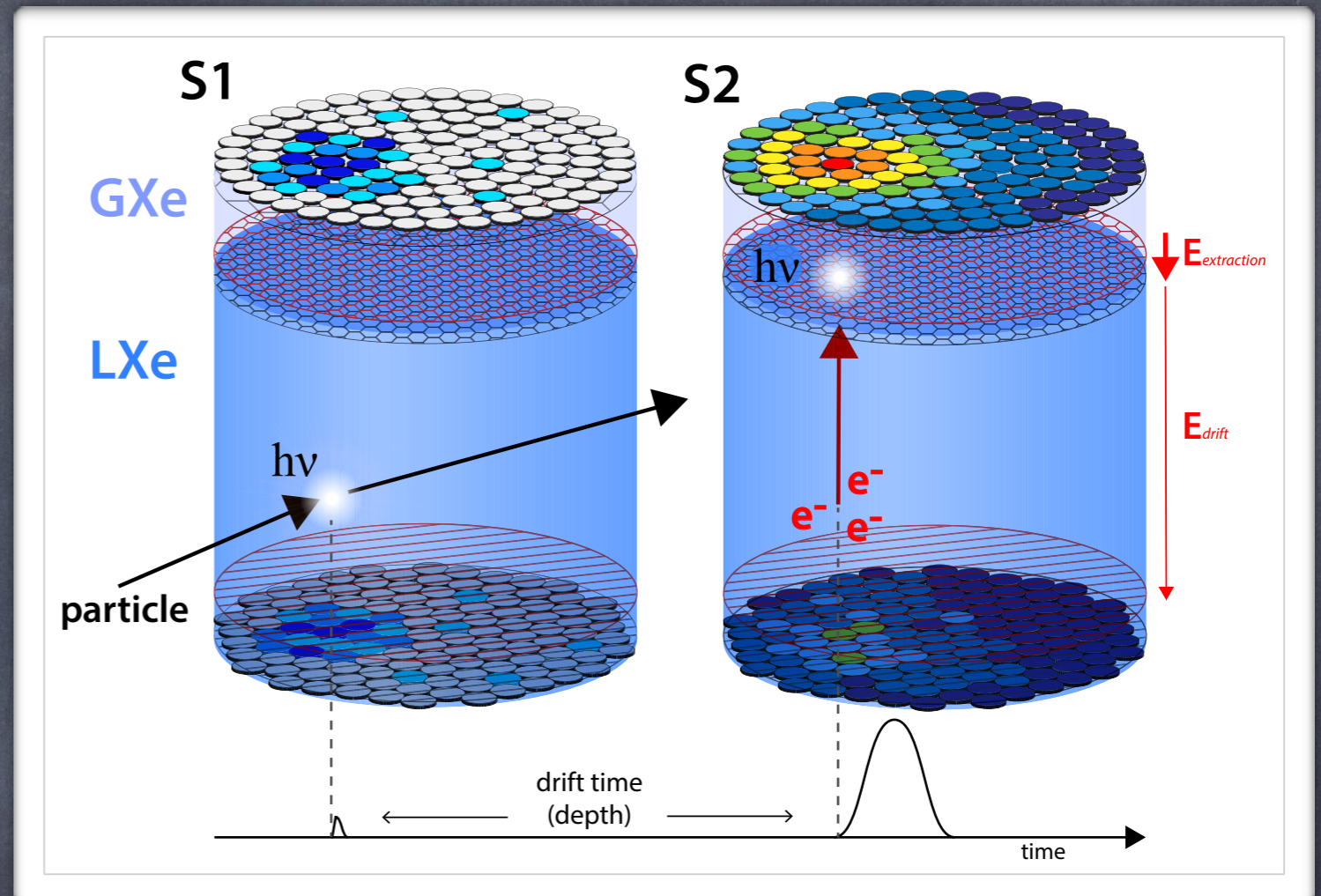
- X,Y from top array

- Z from drift time

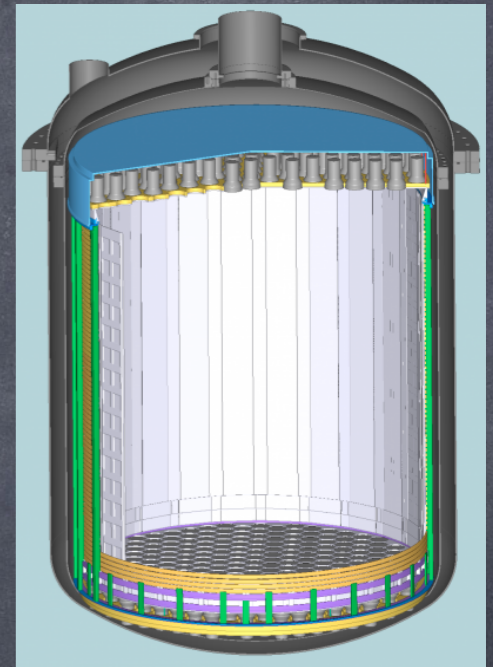
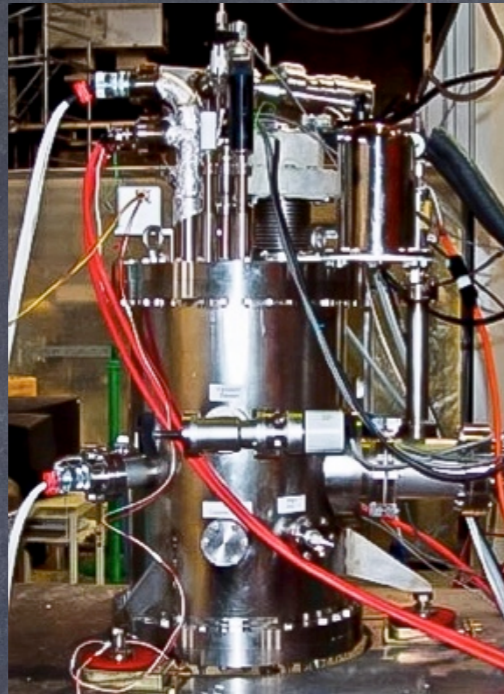
- Discrimination

- Charge-to-light ratio

- Higher for **ER** than **NR**



Phases of XENON



	XENON10	XENON100	XENON1T	XENONnT
Era	2005–2007	2008–2016	2012–2018	2019–2023
Mass	25 kg	161 kg	3200 kg	~8000 kg
Limit (Sensitivity)	$8.8 \times 10^{-44} \text{ cm}^2$	$1.1 \times 10^{-45} \text{ cm}^2$	$7.7 \times 10^{-47} \text{ cm}^2$ ($1.6 \times 10^{-47} \text{ cm}^2$)	($\sim 10^{-48} \text{ cm}^2$)

The XENON Collaboration



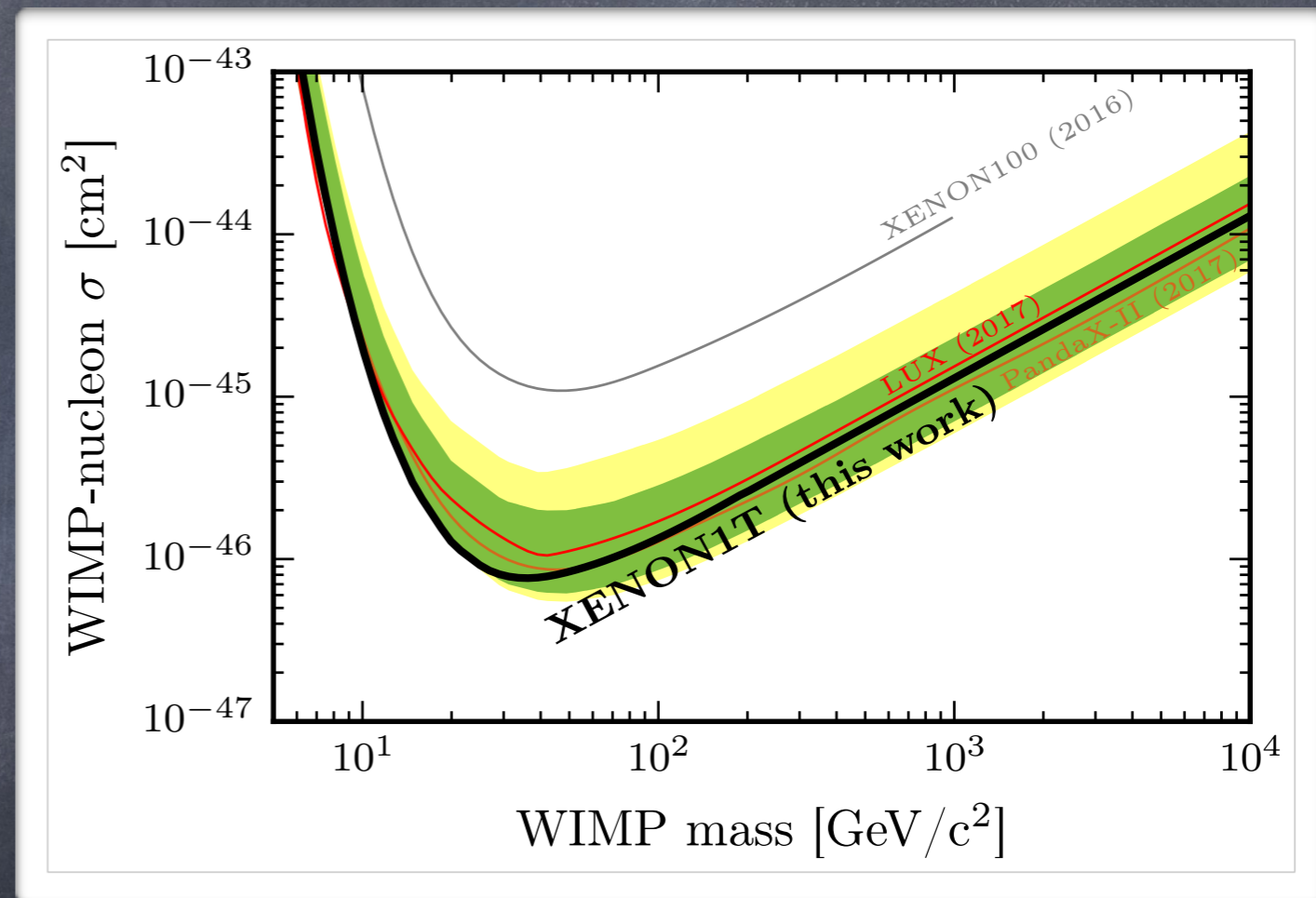
LNGS



XENON1T: First Results

- 34 live days of data
- October 2016 – January 2017
- No evidence of dark matter
- Set upper limit
- 247 more live days since then

PRL 119, 181301 (2017)



ER Backgrounds

JCAP04(2016)027

Source	Background [(kg · day · keV) ⁻¹]	Background [y ⁻¹]	Fraction [%]
Materials	$(7.3 \pm 0.7) \cdot 10^{-6}$	29 ± 3	4.1
²²² Rn	$(1.54 \pm 0.15) \cdot 10^{-4}$	620 ± 60	85.4
⁸⁵ Kr	$(7.7 \pm 1.5) \cdot 10^{-6}$	31 ± 6	4.3
¹³⁶ Xe	$(2.3 \pm 1.1) \cdot 10^{-6}$	9 ± 4	1.4
Solar neutrinos	$(8.9 \pm 0.2) \cdot 10^{-6}$	36 ± 1	4.9
Total	$(1.80 \pm 0.15) \cdot 10^{-4}$	720 ± 60	100

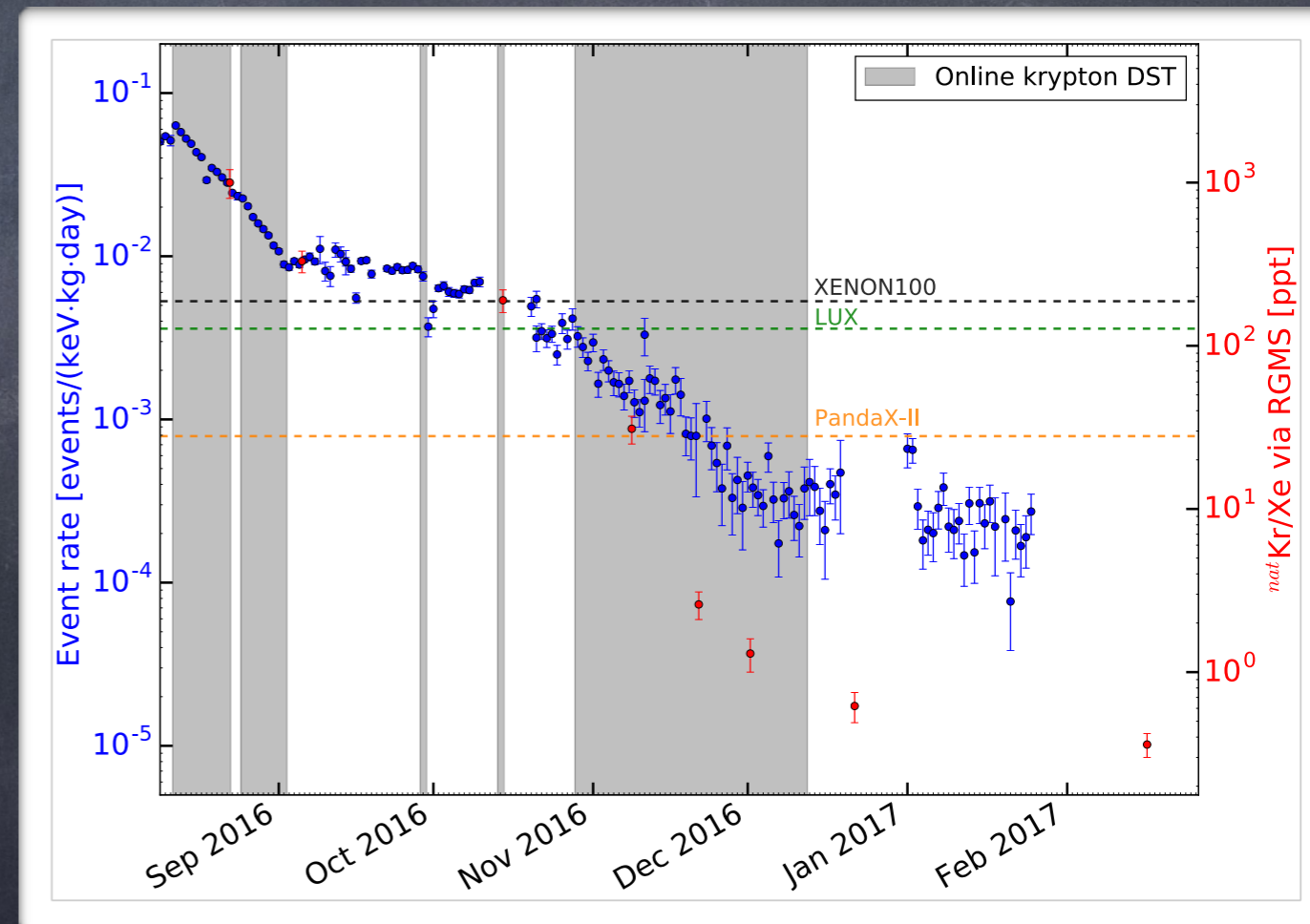
expectations in 2–12 keV, 1T fiducial volume,
before ER/NR discrimination

☉ ⁸⁵Kr/natKr ~ 2 × 10⁻¹¹

☉ Online distillation:
natKr/Xe ~ 0.62 ppt

²²² Rn	3.8 d	α ↓ 5.5 MeV
²¹⁸ Po	3.05 min	α ↓ 6.0 MeV
²¹⁴ Pb	26.8 min	β ↓
²¹⁴ Bi	19.9 min	β ↓
²¹⁴ Po	164 μs	α ↓ 7.7 MeV
²¹⁰ Pb	22.3 a	β ↓
²¹⁰ Bi	5.0 d	β ↓
²¹⁰ Po	138 d	α ↓ 5.3 MeV
²⁰⁶ Pb	stable	

- ☉ Detector materials screened for radio-purity with HPGe detectors
- ☉ Krypton now subdominant
- ☉ Lowest background level achieved:
(62 ± 11) events/
(tonne · year · keV)

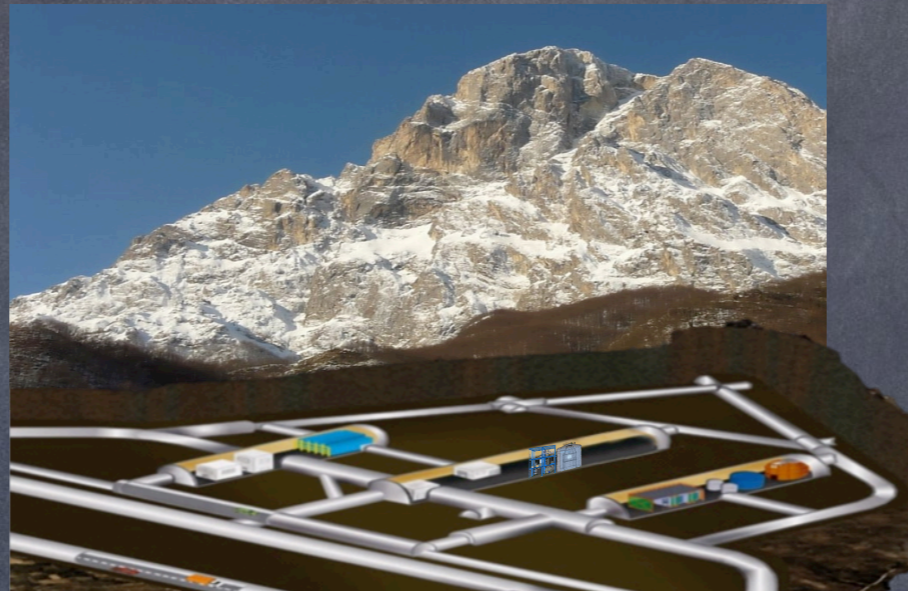


EPJC (2017) 77: 276

NR Backgrounds

Muon-induced Neutrons

- 3600 m.w.e. overburden
- 700 m³ demineralized water
- Water Cherenkov Muon Veto
- 0.012 events/(tonne·year)



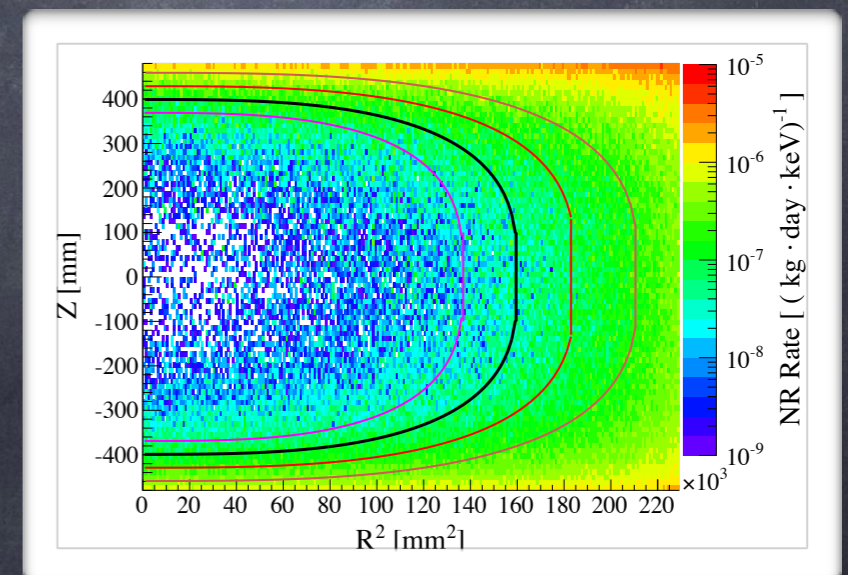
Radiogenic

- Spontaneous fission and (α, n) reactions from U and Th chains
- Reduced with material selection
- O(1) event/(tonne·year)



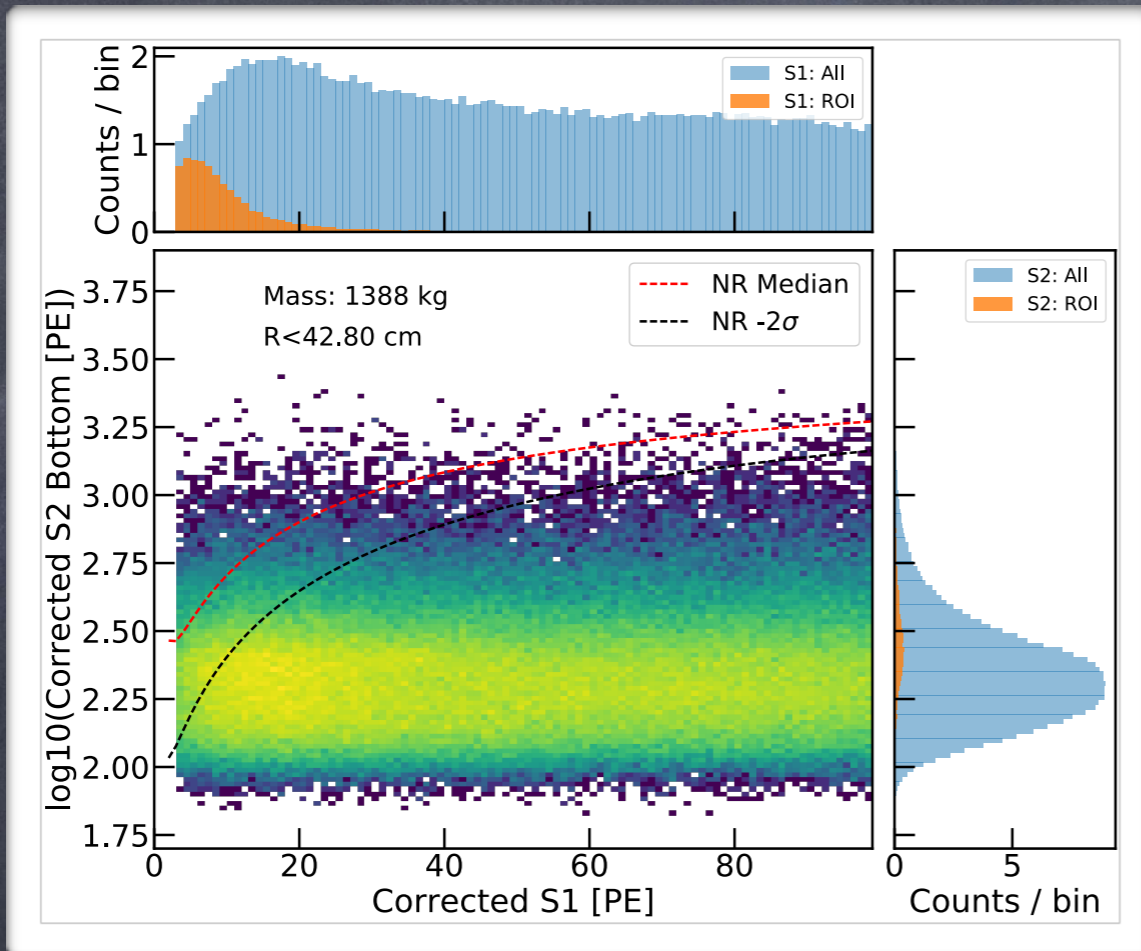
CEVNS

- Irreducible
- Almost no contribution above 5 keV

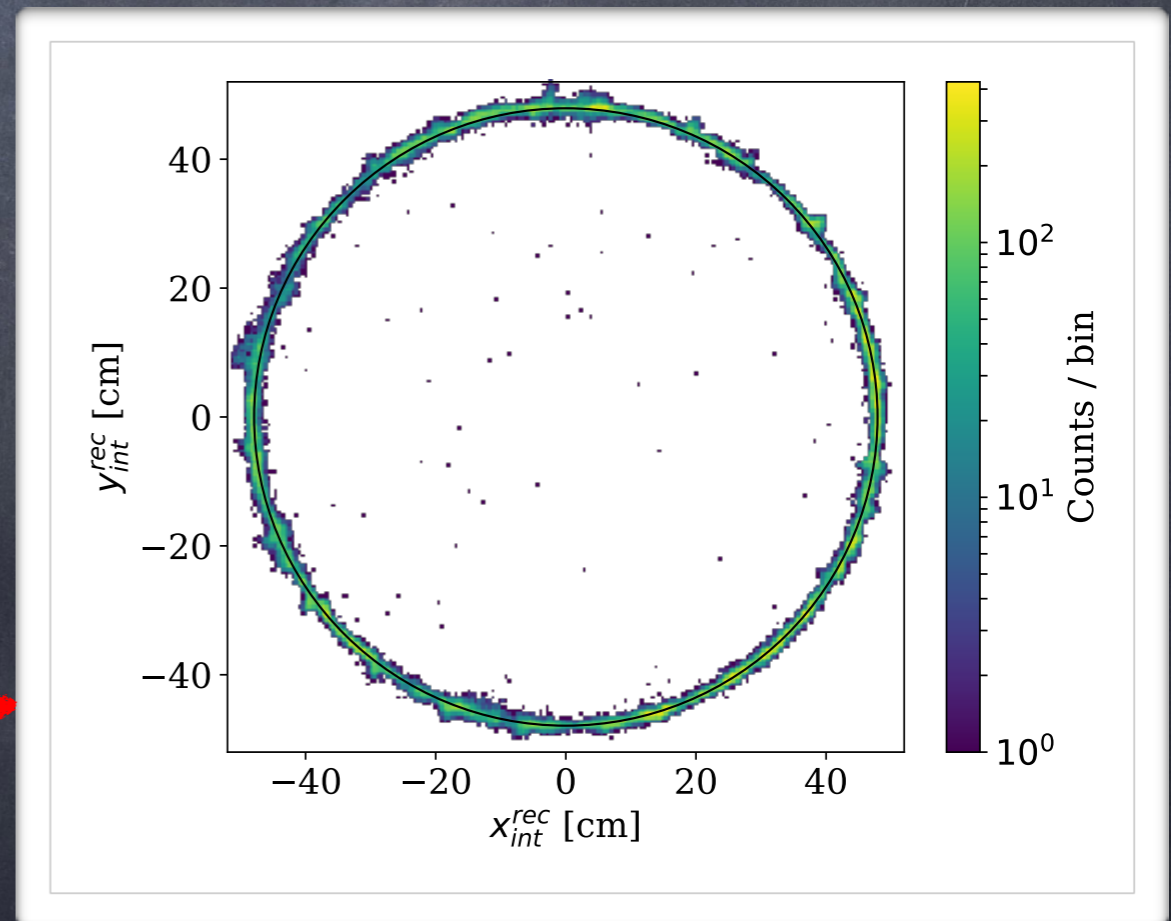


Surface Events

^{222}Rn	3.8 d
α	5.5 MeV
^{218}Po	3.05 min
α	6.0 MeV
^{214}Pb	26.8 min
β	
^{214}Bi	19.9 min
β	
^{214}Po	164 μs
α	7.7 MeV
^{210}Pb	22.3 a
β	
^{210}Bi	5.0 d
β	
^{210}Po	138 d
α	5.3 MeV
^{206}Pb	stable



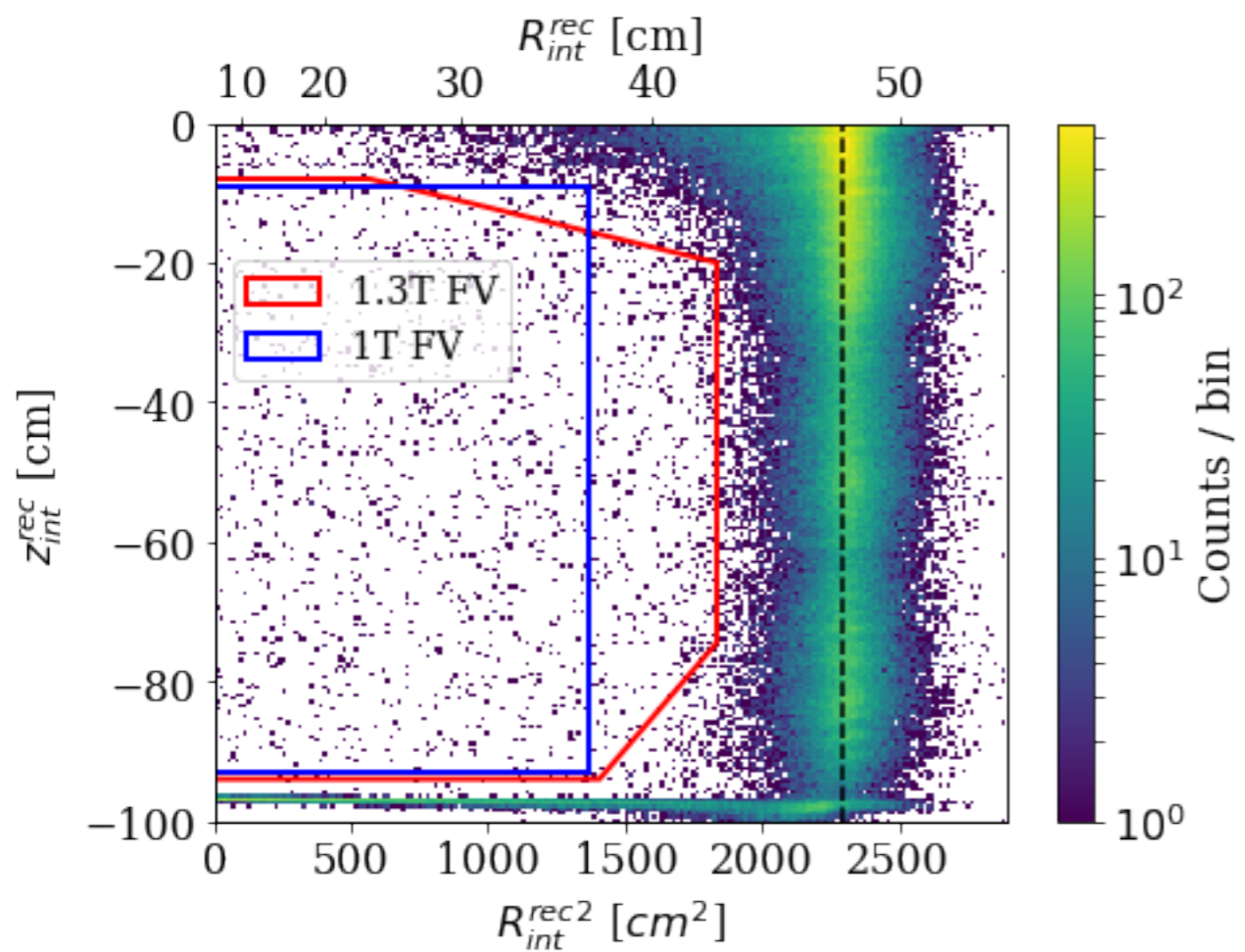
- Charge loss on PTFE wall surface
- Events incorrectly reconstructed in signal region



Suppress with fiducial volume

Data-driven model

Fiducial Volume



- Exploit self-shielding
- First Result: 1 tonne
- This time: +300 kg
- Better position reconstruction
- Improved corrections for field distortion and wall charge-up
- Spatial dimension in statistical interpretation

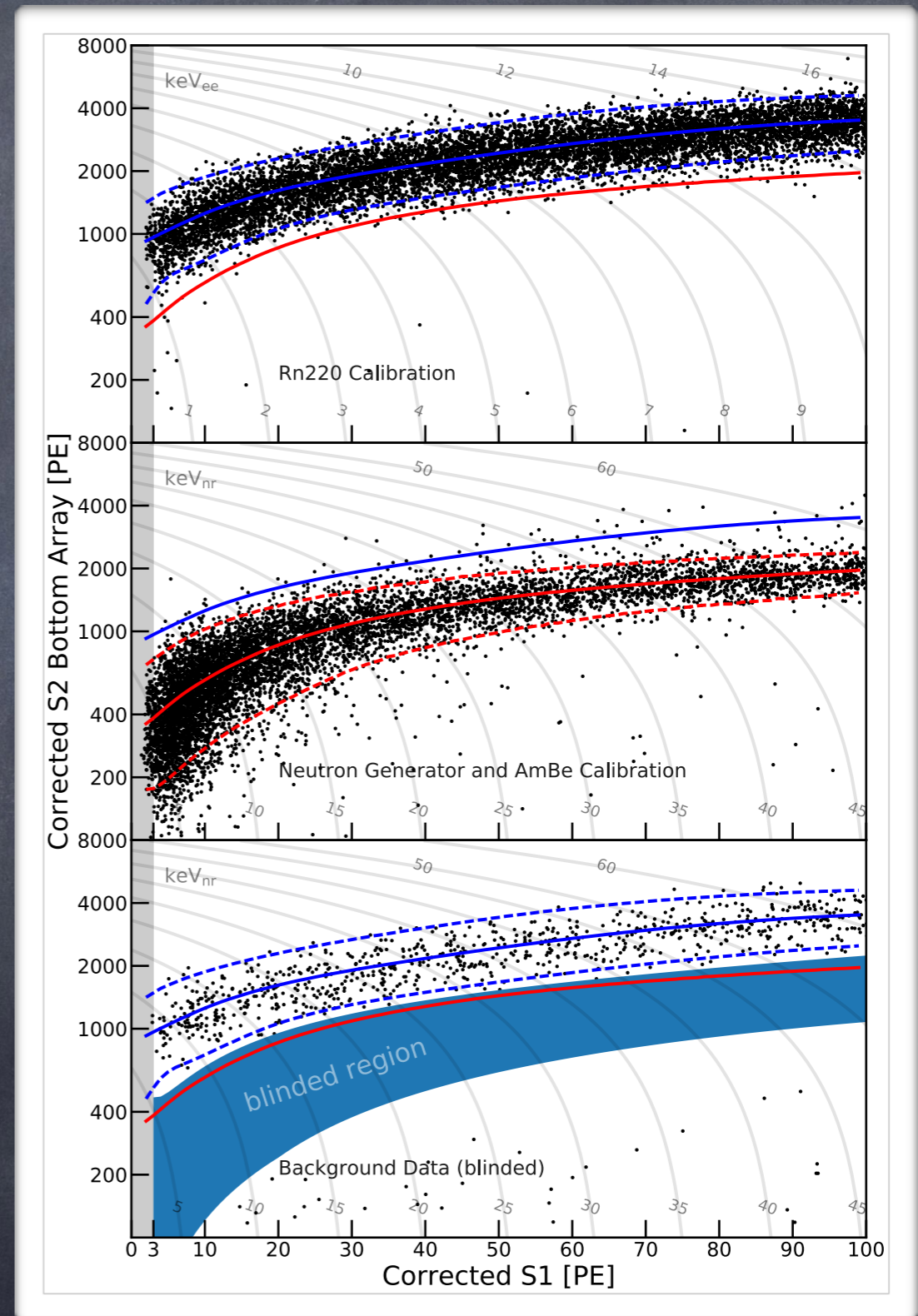
Calibration

ER

- Inject ^{220}Rn
- Daughter ^{212}Pb mimics ^{214}Pb background
- Spreads uniformly

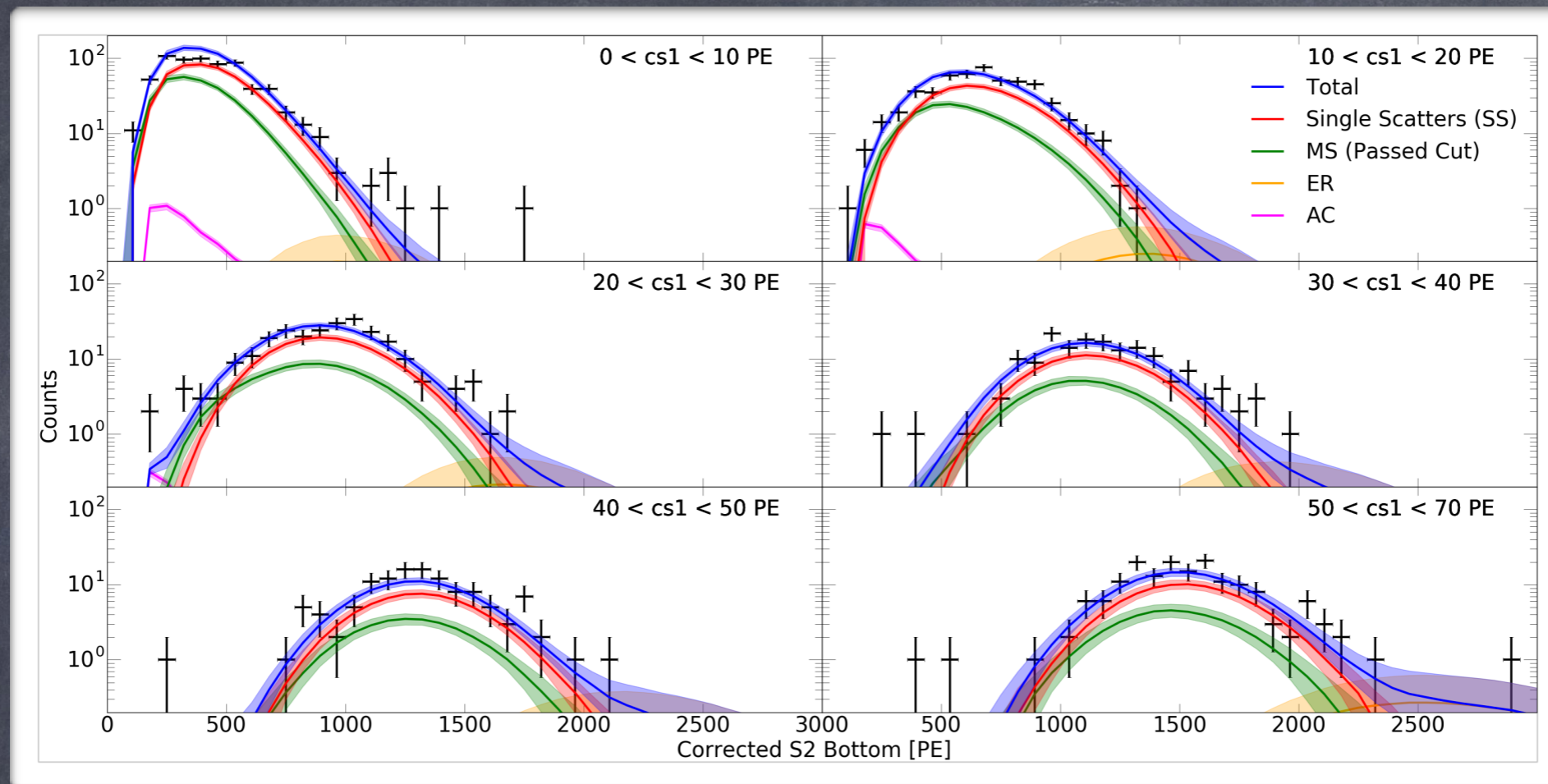
NR

- External DD Neutron Generator and AmBe source



NIM, 879:31-38 (2017) PRD 95, 072008 (2017)

Parametrizing Signal and Background



- WIMP signal, ER/NR background models derived through agreement between calibration and Monte Carlo
- Surface background and accidental coincidence described with empirical models

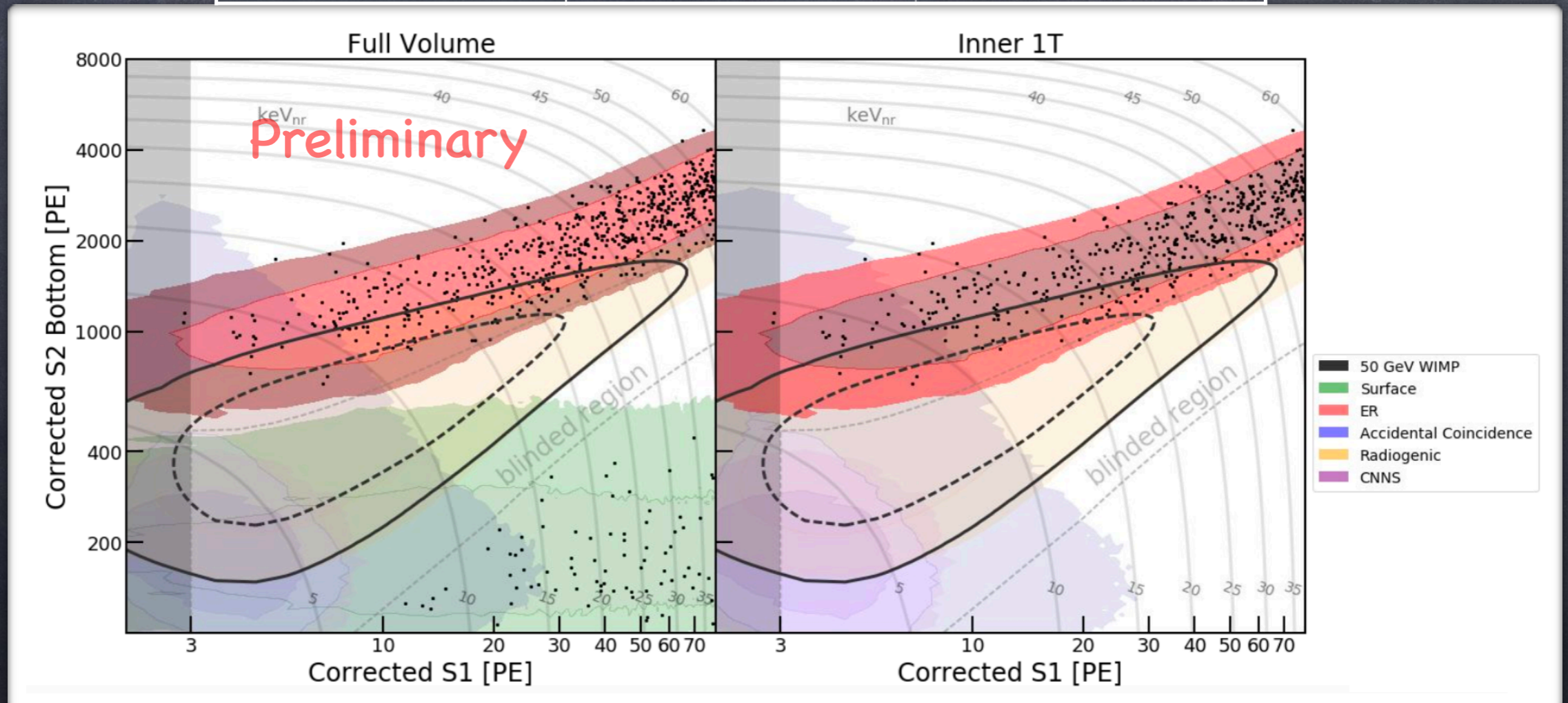
A Complete Background Model

Counts between
NR median and -2σ
quantile

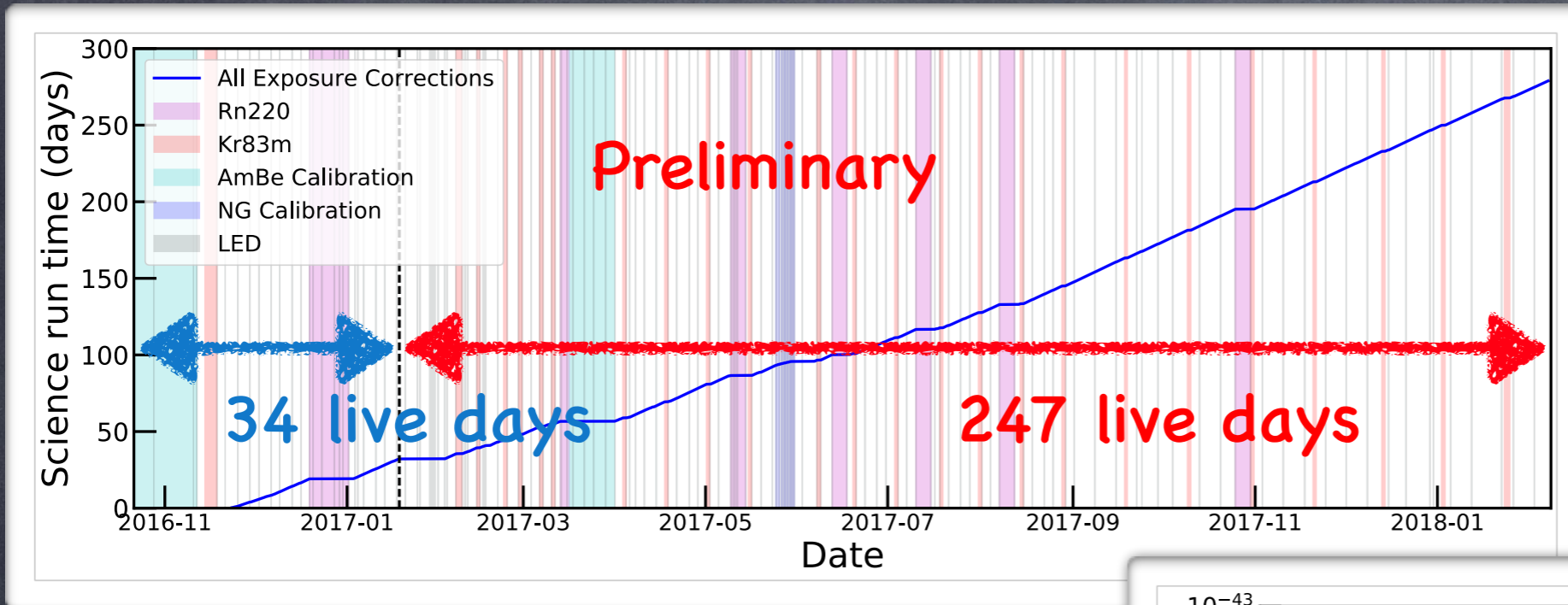
Source	1.3T FV	Inner 1T
ER	1.8 ± 0.2	1.4 ± 0.2
Radiogenic	0.6 ± 0.3	0.4 ± 0.2
CEVNS	0.04 ± 0.01	0.03 ± 0.01
Accidental	0.2 ± 0.1	0.1
Surface	6.1 ± 0.3	0.1
Total	8.7 ± 0.5	2.0 ± 0.3
WIMP (50 GeV, 10^{-47} cm^2)	0.8 ± 0.1	0.6 ± 0.1

Expectations in
3-70 PE, 247 days

Preliminary

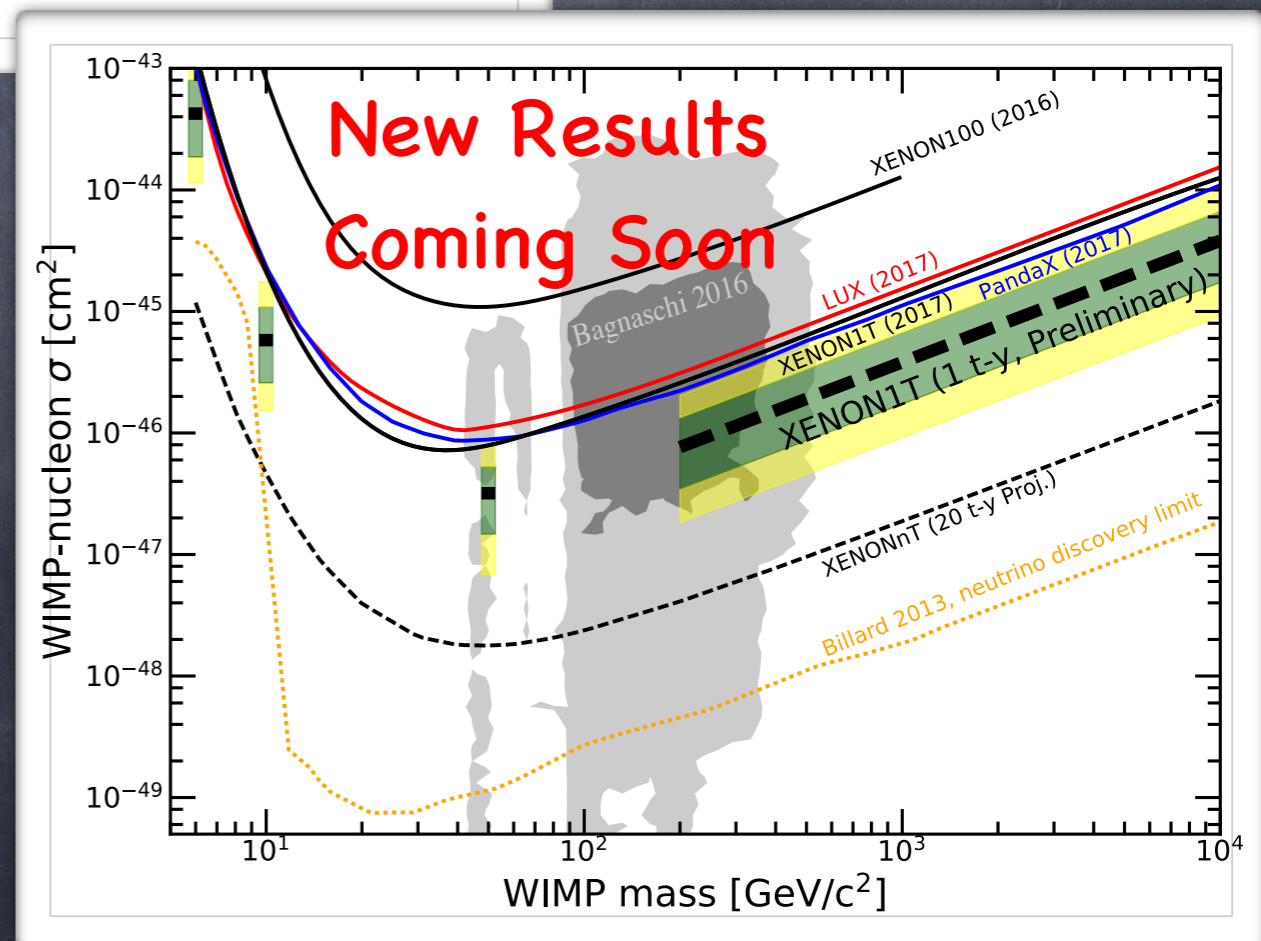


Upcoming Results



- >1 year of stable data taking
- Factor ~3 improvement in sensitivity

Data blinded and salted



Take-Home Message

- Sensitivity ~3 times better:
 - Livetime ~8 times longer
 - Larger fiducial volume (deeper understanding of detector)
- Unprecedented low background
- More rare event searches
 - Low-mass WIMPs (S2-only)
 - ER (axions, SuperWIMPs, ...)
 - Annual Modulation
- XENONnT in 2019

