

The Galactic Center



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In collaboration with: Francesca Calore, Emma Storm & Christoph Weniger



Overview

1. GCE Analysis with SkyFACT
2. Standard template fitting \leftrightarrow Skyfact
3. Can we see a GCE in the disk?

The Galactic Centre Excess

Goodenough & Hooper 2009, Vitale+ (Fermi coll.) 2009, Hooper & Goodenough 2011, Hooper & Linden 2011, Boyarsky+ 2011 (no signal), Abazajian & Kaplinghat 2012, Hooper & Slatyer 2013, Huang+ 2013, Gordon & Macias 2013, Macias & Gordon 2014, Zhou+2014, Abazajian+ 2014, Daylan+2014, Calore+ 2014, Gaggero + 2015, Carlson+ 2015, Fermi-LAT 2016, 2017

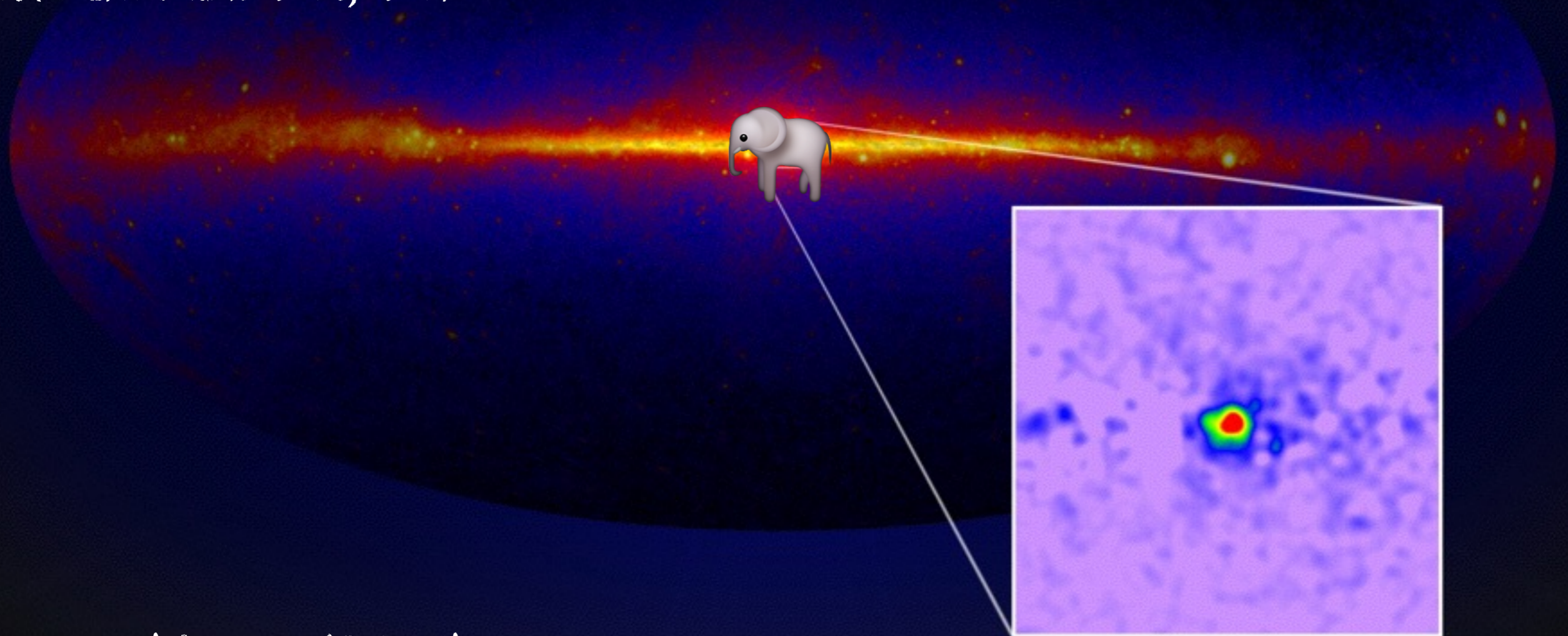


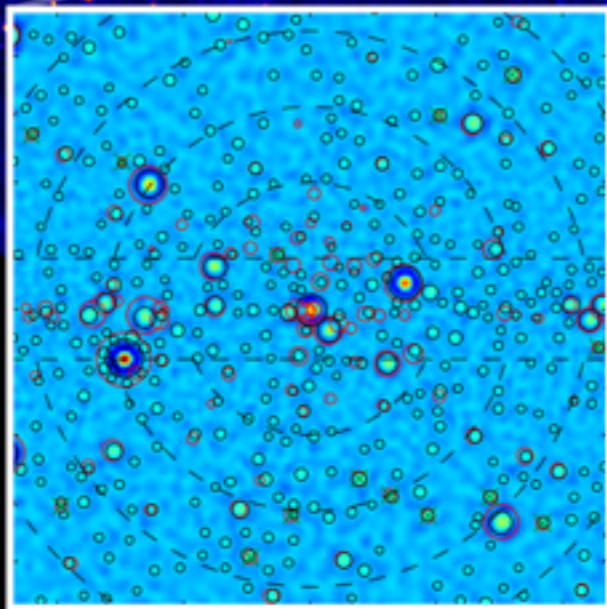
Image credit: NASA/T. Linden

The Galactic Centre Excess

Speckled gamma-ray emission from inner Galaxy points to a new source population

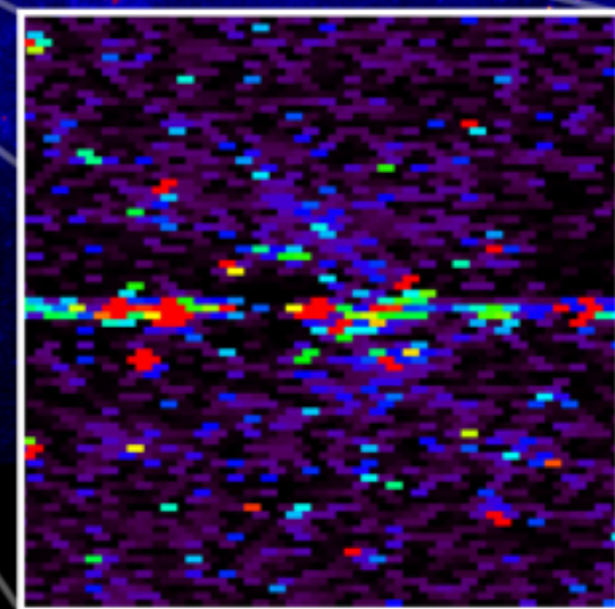


Wavelet transformation



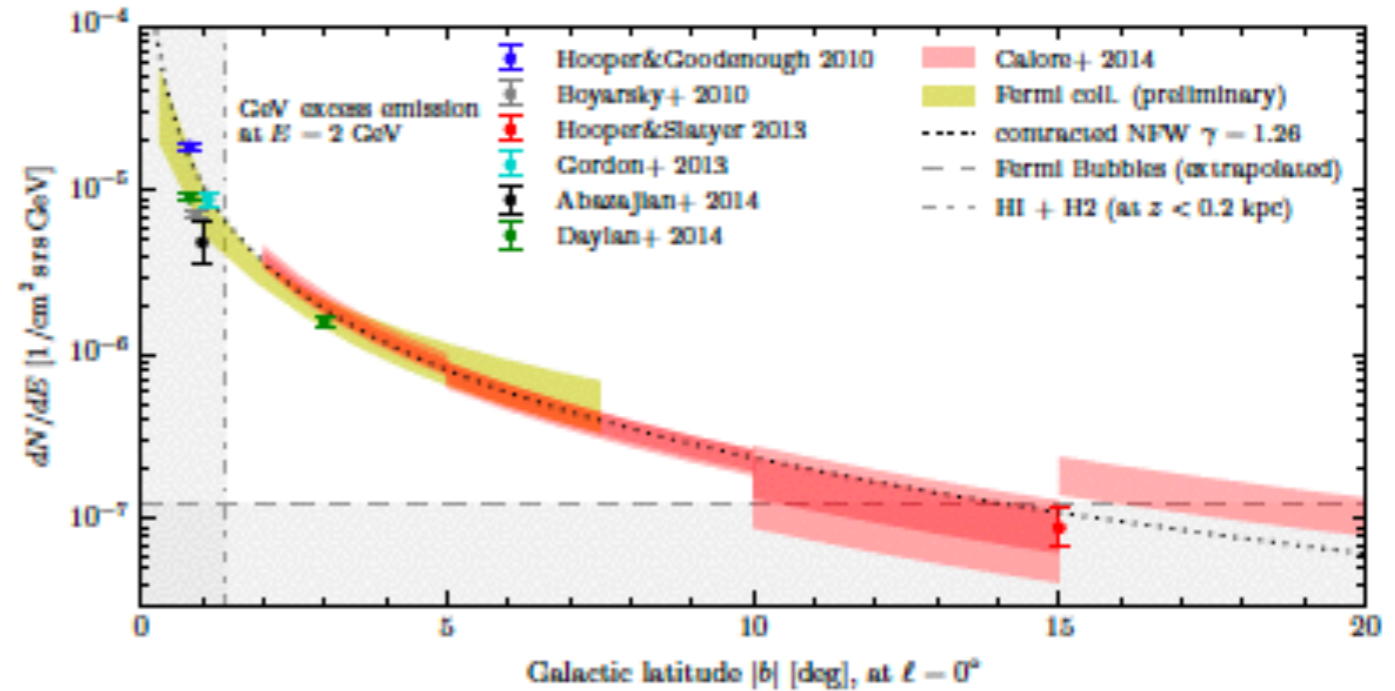
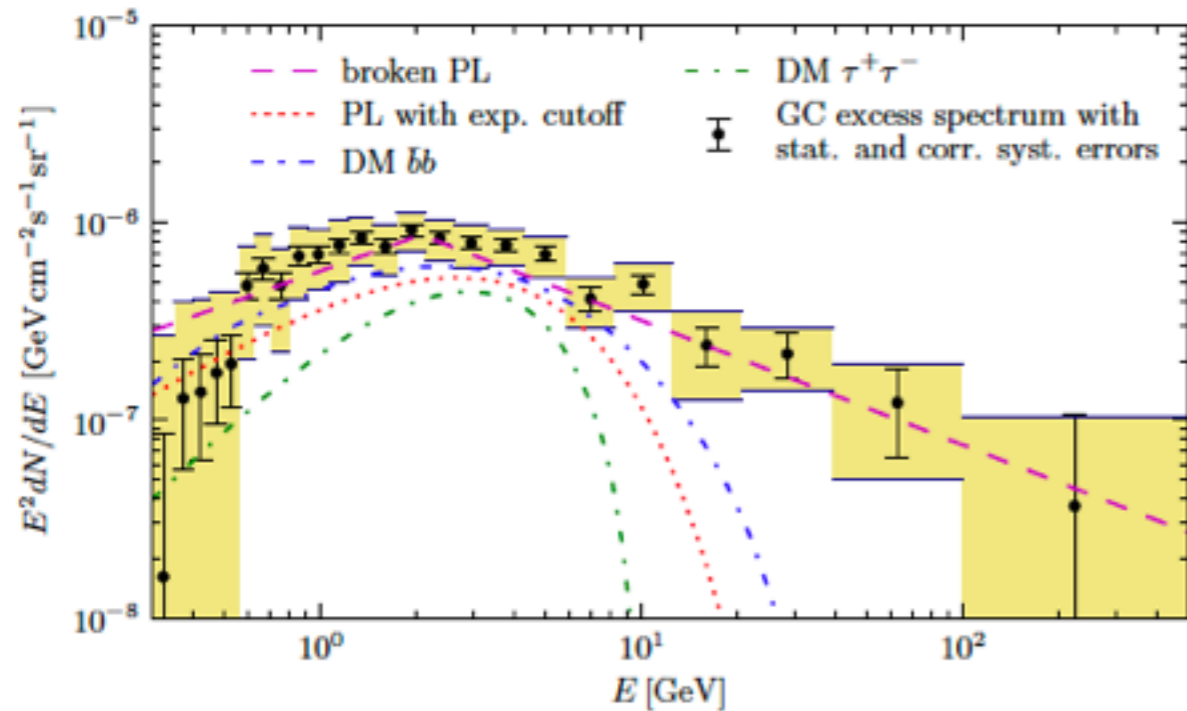
Bartels et al. 2016

Non-Poissonian noise



Lee et al. 2016

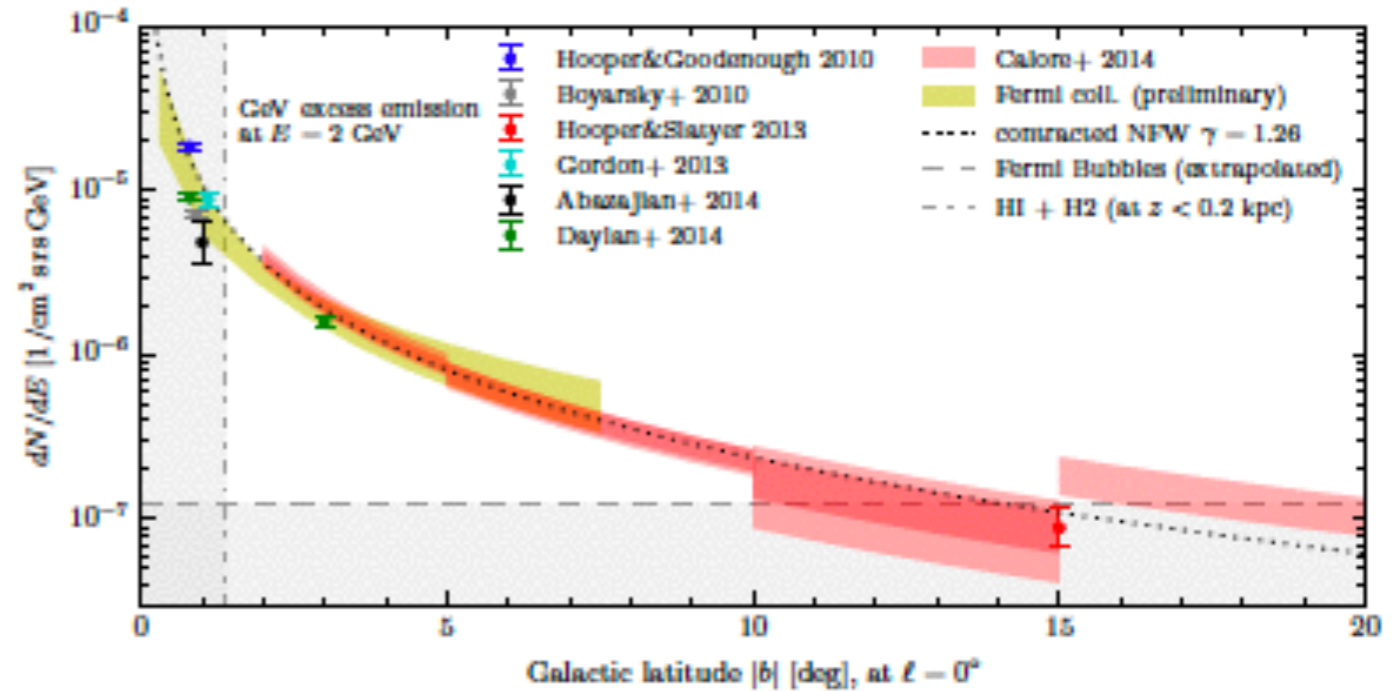
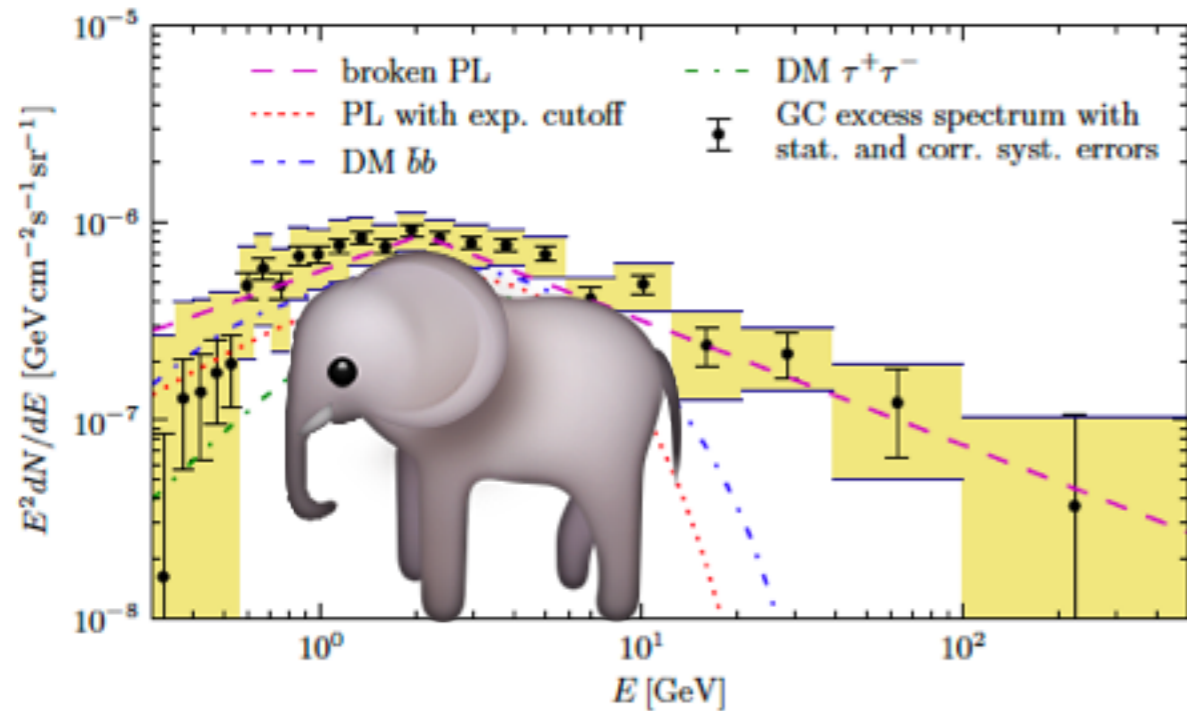
GCE features



Calore, Cholis & Weniger (2014)

Calore, Cholis, McCabe & Weniger (2015)

GCE features



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1. SKYFACT

EMMA STORM
SUN 12:00
arXiv:1705.04065

Summary:

- Hybrid between image reconstruction & template fitting

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This means:
freedom in spectral and
spatial templates to adapt
to errors/incompleteness!

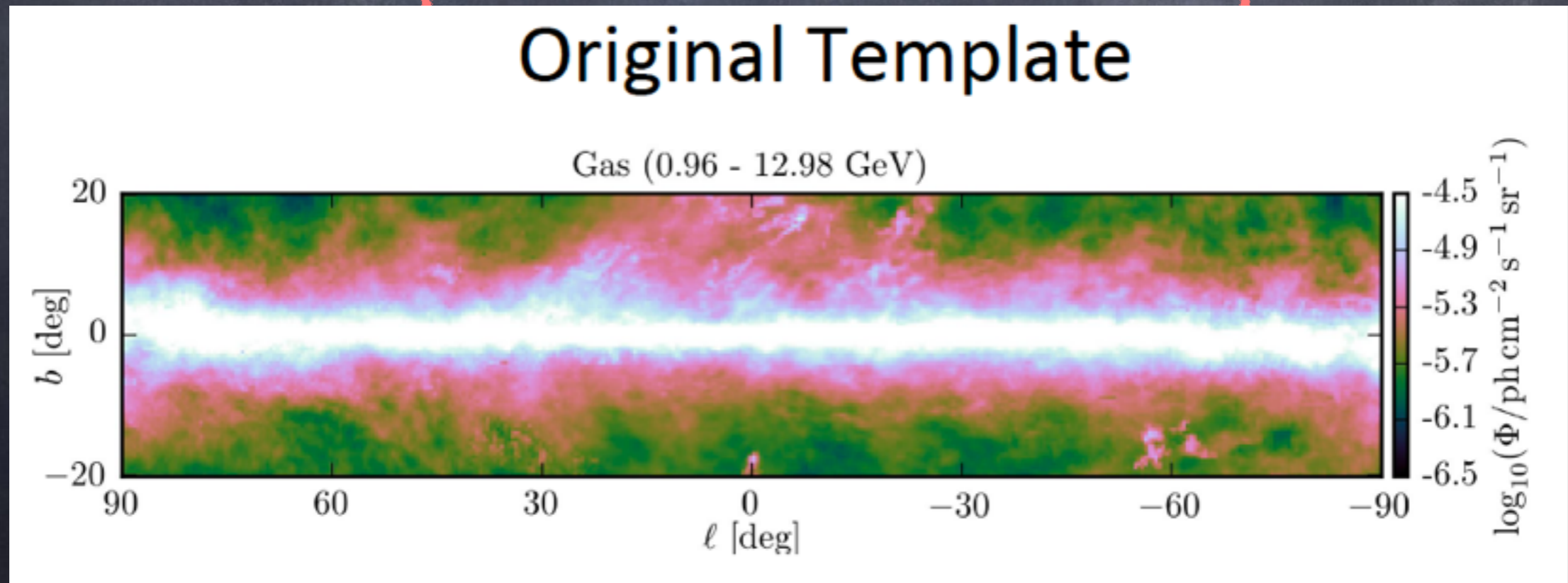
e.g. "dark gas"

1. SKYFACT

EMMA storm
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arXiv:1705.04065

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- Hybrid between image reconstruction & template fitting



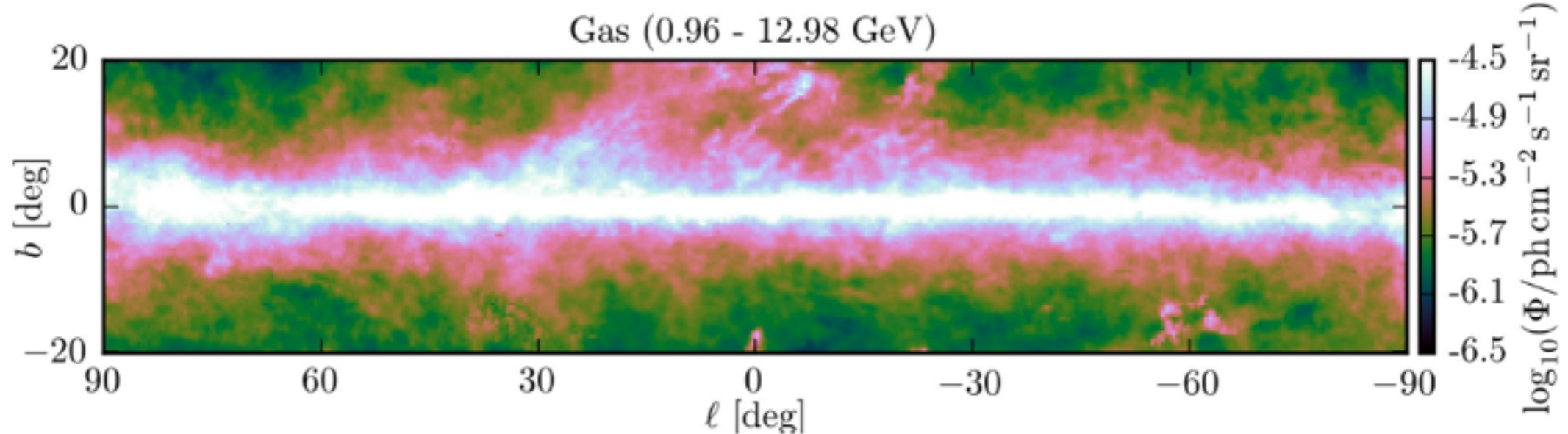
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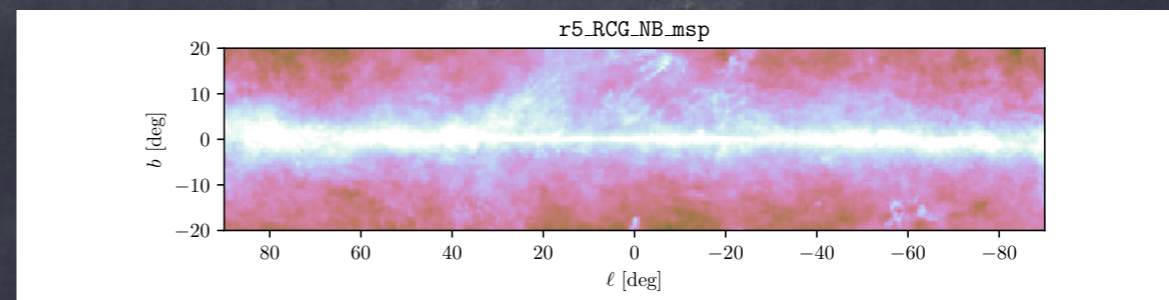
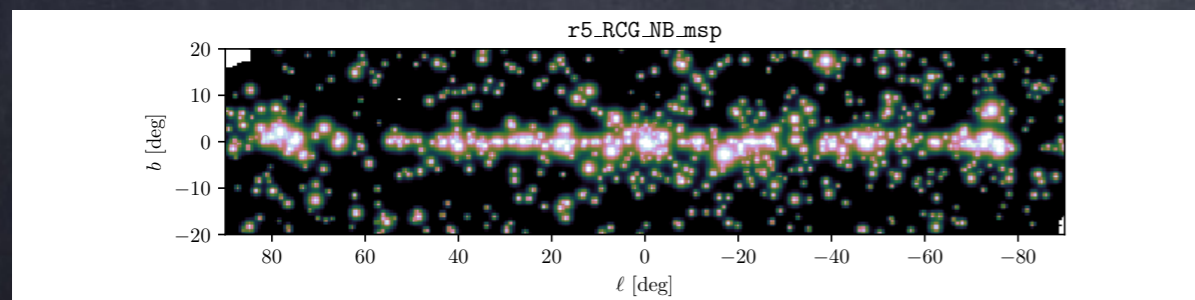
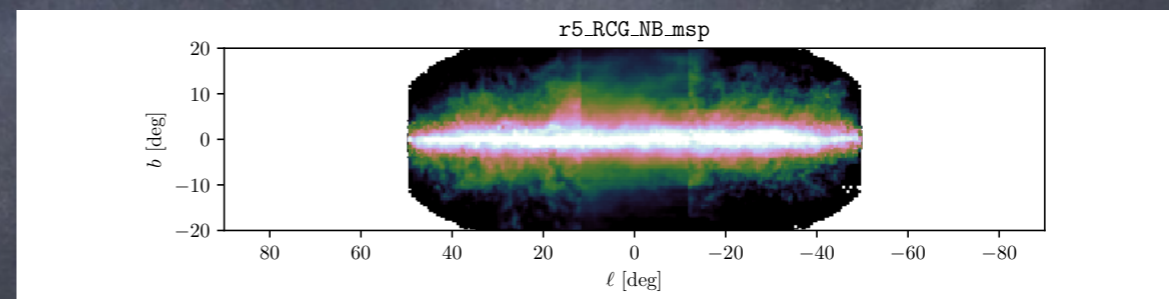
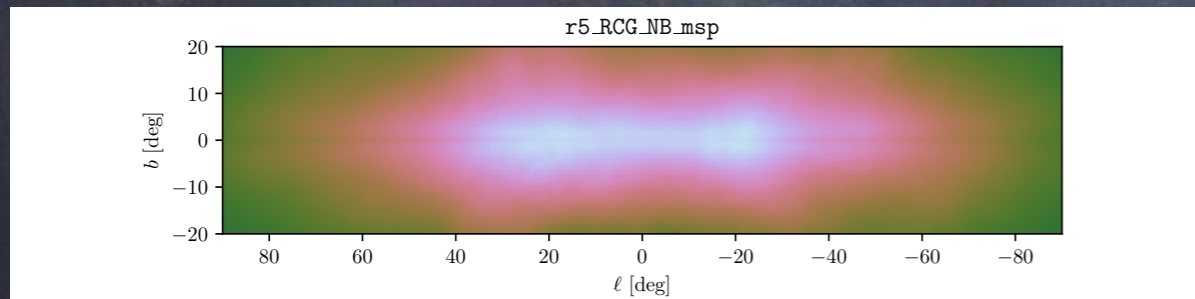
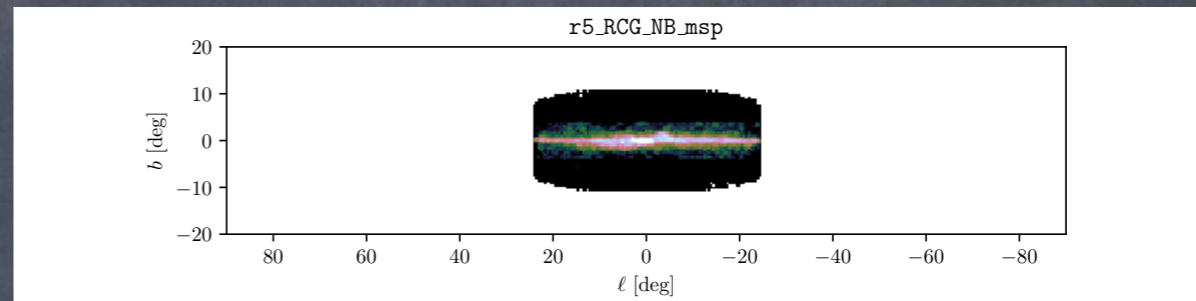
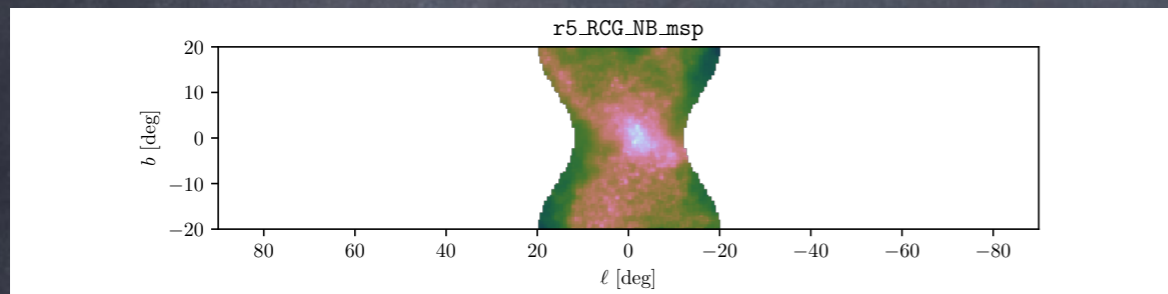
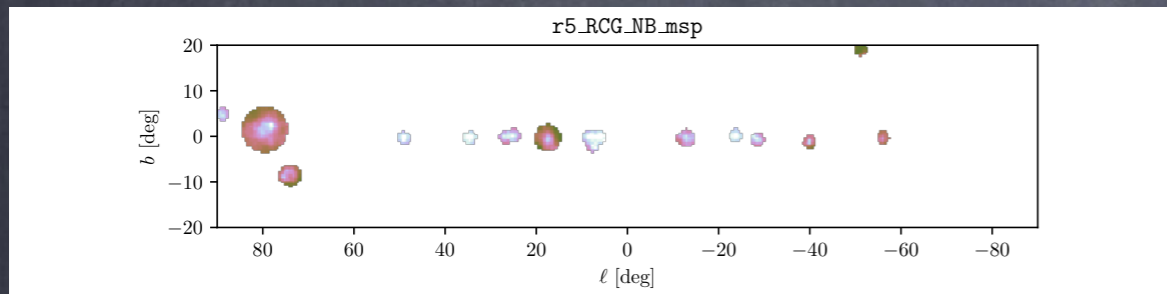
Summary:

- Hybrid between image reconstruction & template fitting

Best-fit Template



Fore-Background



GCE Morphology

- Reanalysis using SkyFACT

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DM

cNFW



Bartels+ in prep.

GCE Morphology

- Reanalysis using SkyFACT

DM

Stellar Mass distribution

cNFW

Boxy bulge

Nuclear bulge

Launhardt+ 2002

Dwek+ 1995; Cao+ 2013

Bartels+ in prep.

Results

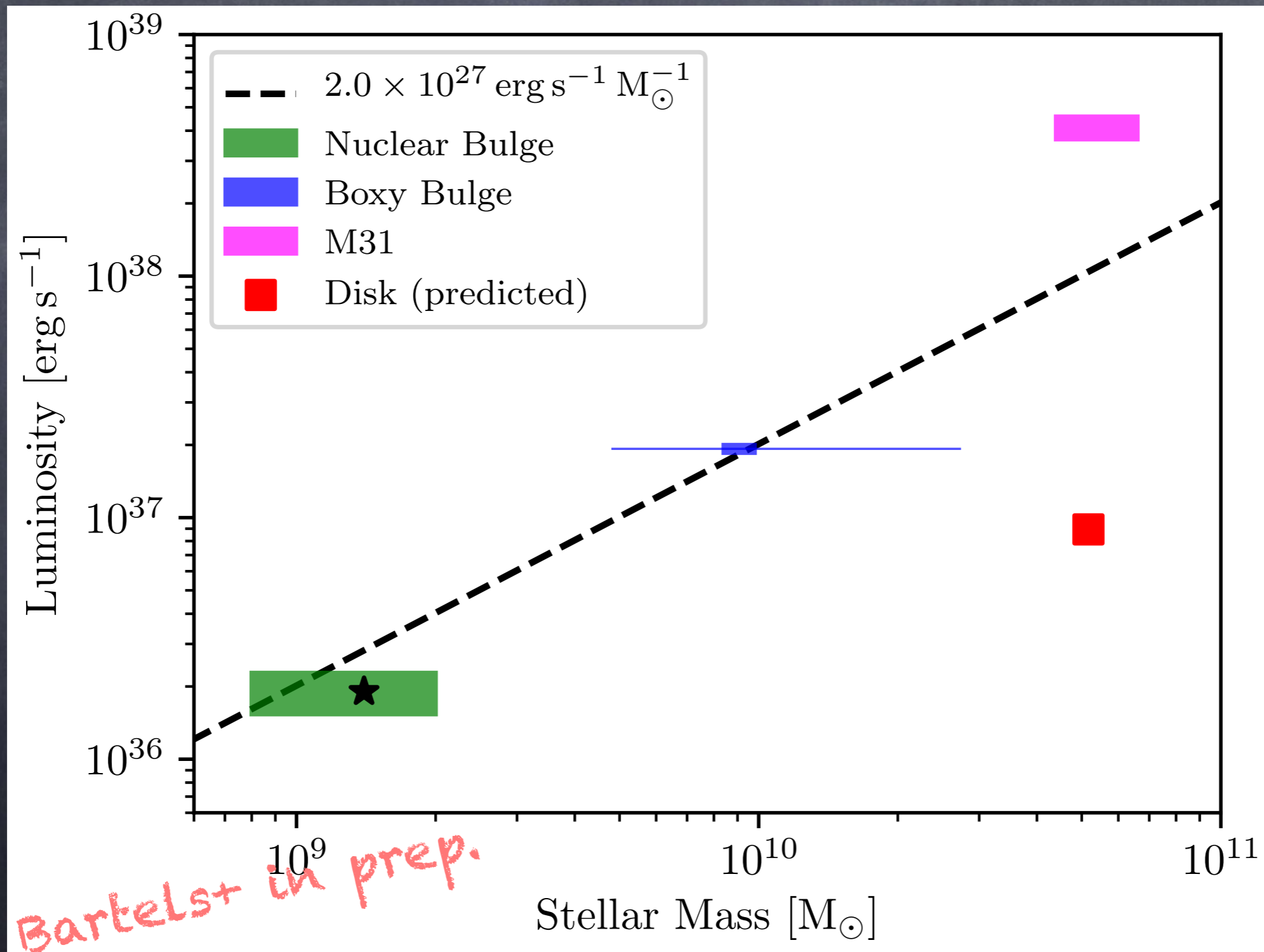
Run	$-2 \ln \mathcal{L}$	
	free spectrum	MSP spectrum
r5_RCG_NB_X	647808.1	648020.2
r5_RCG_NB	647831.2	648027.5
r5_RCG	647884.7	648061.7
r5_BulgeGC	647916.5	648140.3
r5_Einasto	647961.4	648188.6
r5_NFW126	648021.8	648242.4
r5_NFW100	648049.8	648278.6

Results

TS \sim 150 difference
between bulge and DM
template

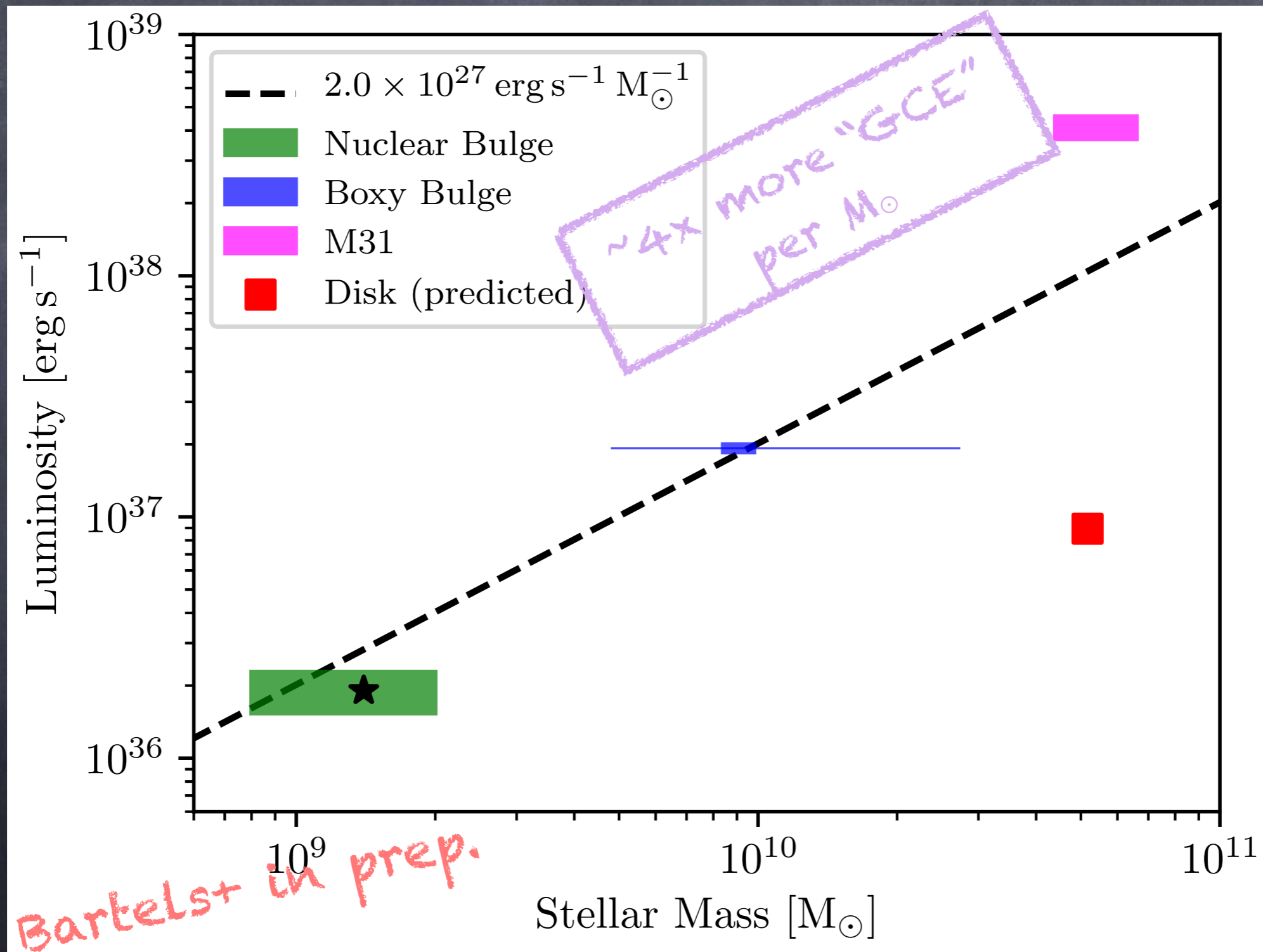
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GCE vs. Stellar Mass

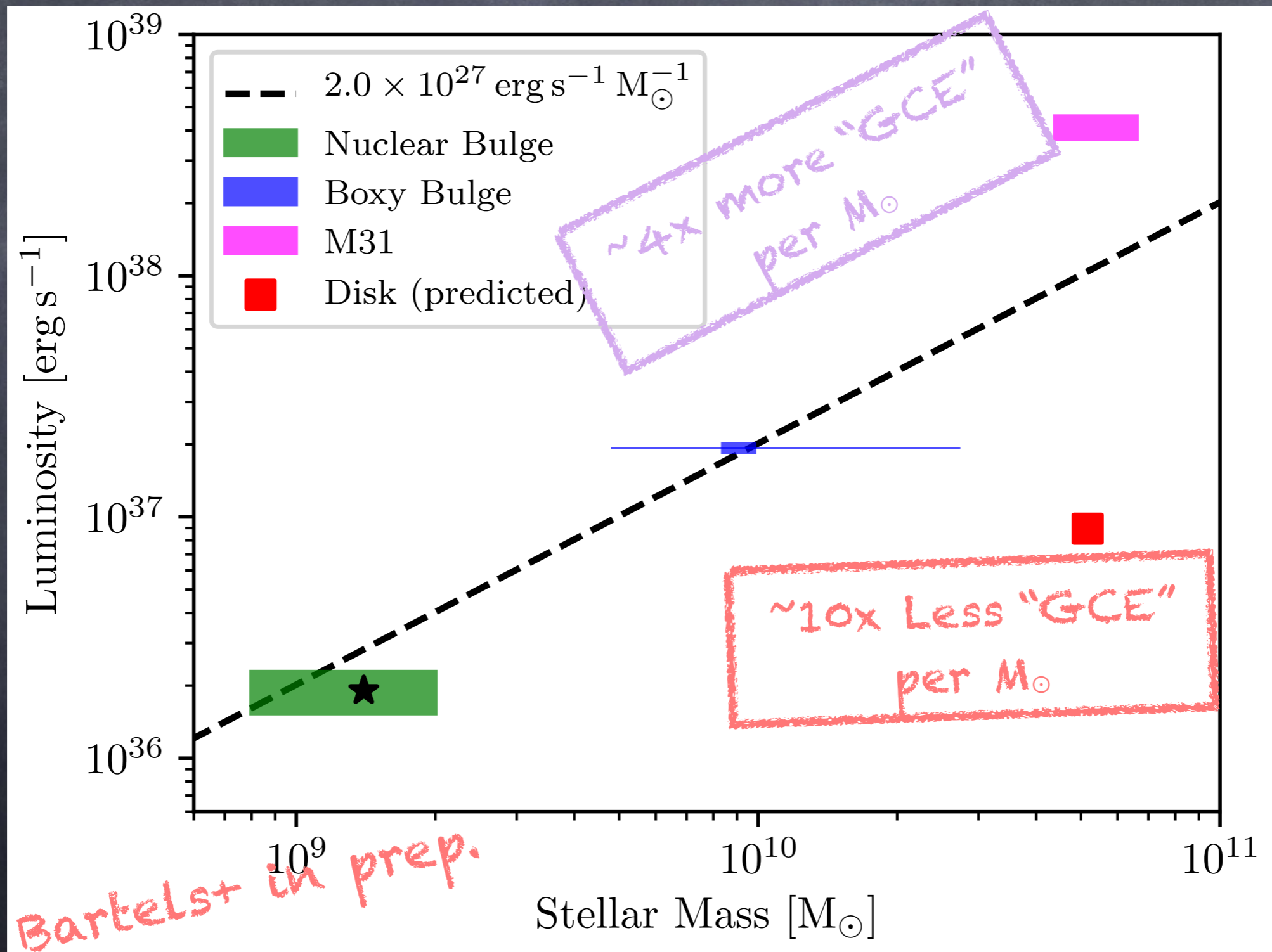


Bartels+ in prep.

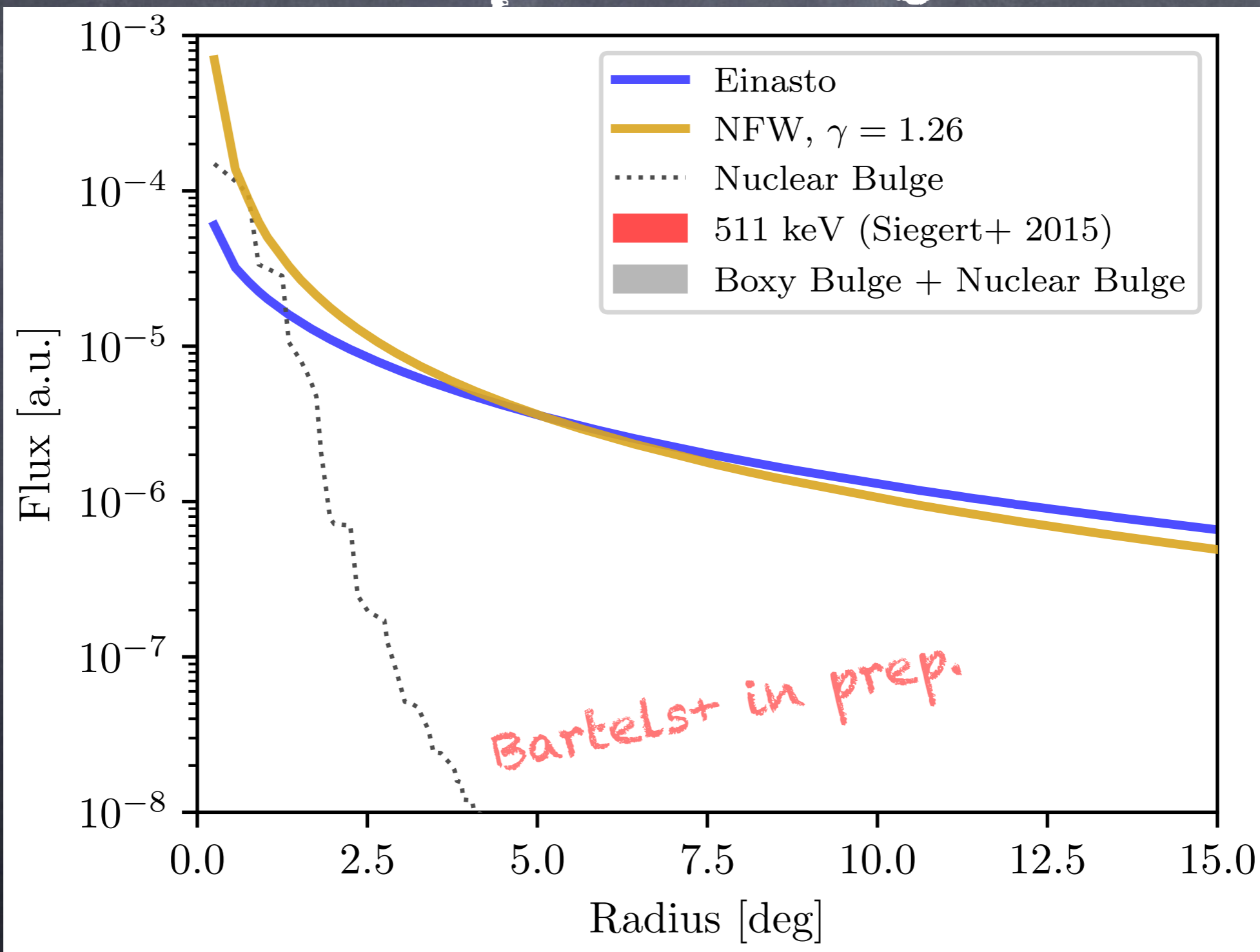
GCE vs. Stellar Mass



GCE vs. Stellar Mass



Surprising?



2. From Traditional Template Fitting to SkyFACT

Why was this not
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- Past studies checked for non-sphericity: negative results.

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- Past studies checked for non-sphericity: negative results.
(Daylan+ 2014, Calore+ 2014, but also see Linden+ 2016 & Fermi-LAT Collaboration 2017)
- We connected SkyFACT analysis to Calore+ 2014 analysis (implementing differences one by one)

Template Fitting ↔ Skyfact

- old
- ↑
- A1 Reproduction of Calore et al. (2015b), with spatial DM template.
 - A2 Reproduction of Calore et al. (2015b), with spatial stellar mass template.
 - B1 A1 gas/ICS templates changed to RUN5 templates (1 gas template).
 - B2 A2 gas/ICS templates changed to RUN5 templates (1 gas template).
 - C1 B1 with longitude extended to $|l| \leq 90^\circ$.
 - C2 B2 with longitude extended to $|l| \leq 90^\circ$.
 - D1 C1 with modulation as in RUN5.
 - D2 C2 with modulation as in RUN5.
 - E1 D1 with disk unmasked
 - E2 D2 with disk unmasked
 - F1 E1 with two gas rings used.
 - F2 E2 with two gas rings used.
 - G1 F1 with extended sources added in.
 - G2 F2 with extended sources added in.
 - H1 G1 with spectral bubble template.
 - H2 G2 with spectral bubble template.
- New
- ↓

Template Fitting ↔ Skyfact

Old



New

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DM preferred

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DM preferred

Bulge preferred

Template Fitting \leftrightarrow Skyfact

Old


 New

Template	χ^2		$\Delta\chi^2$
	(1) NFW $\gamma = 1.26$	(2) RCG_NB	
A	147174.1	147486.0	-311.9
B	165359.2	167419.6	-2060.4
C	718013.4	721344.0	-3330.6
D	562568.4	562995.2	-426.8
E	655669.2	654782.6	886.6
F	655113.2	654947.4	165.8
F	651279.8	651022.9	256.9
F	648635.9	648484.5	151.4

- G2 F2 with extended sources added in.
- H1 G1 with spectral bubble template.
- H2 G2 with spectral bubble template.

Template Fitting \leftrightarrow Skyfact

Old

 New

A
A
B
B
C
C
D
D
E
E
F
F
G
G2
H1
H2

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Difference arises when:

- Modulation is turned on
- Unmasking the disk

3. "Disk-GCE"

Estimate of GCE in the disk

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2. Simulate a disk population of MSPs (with a Lorimer distribution)
3. Estimate completeness of flux from local (< 3 kpc) volume assuming some luminosity function.
4. Get total disk flux & luminosity!

Results

Bulge

• Flux

$$1.8 \times 10^{-09} \text{ erg cm}^{-2} \text{ s}^{-1}$$

• Lum. vs. Mass

$$2 \times 10^{27} \text{ erg s}^{-1} M_{\odot}$$

Disk

• Flux

$$3 \times 10^{-09} \text{ erg cm}^{-2} \text{ s}^{-1}$$

• Lum. vs. Mass

$$1.8 \times 10^{26} \text{ erg s}^{-1} M_{\odot}$$

B/D ~ 0.8
(but the bulge is twice
as luminous)

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Disk

• Flux

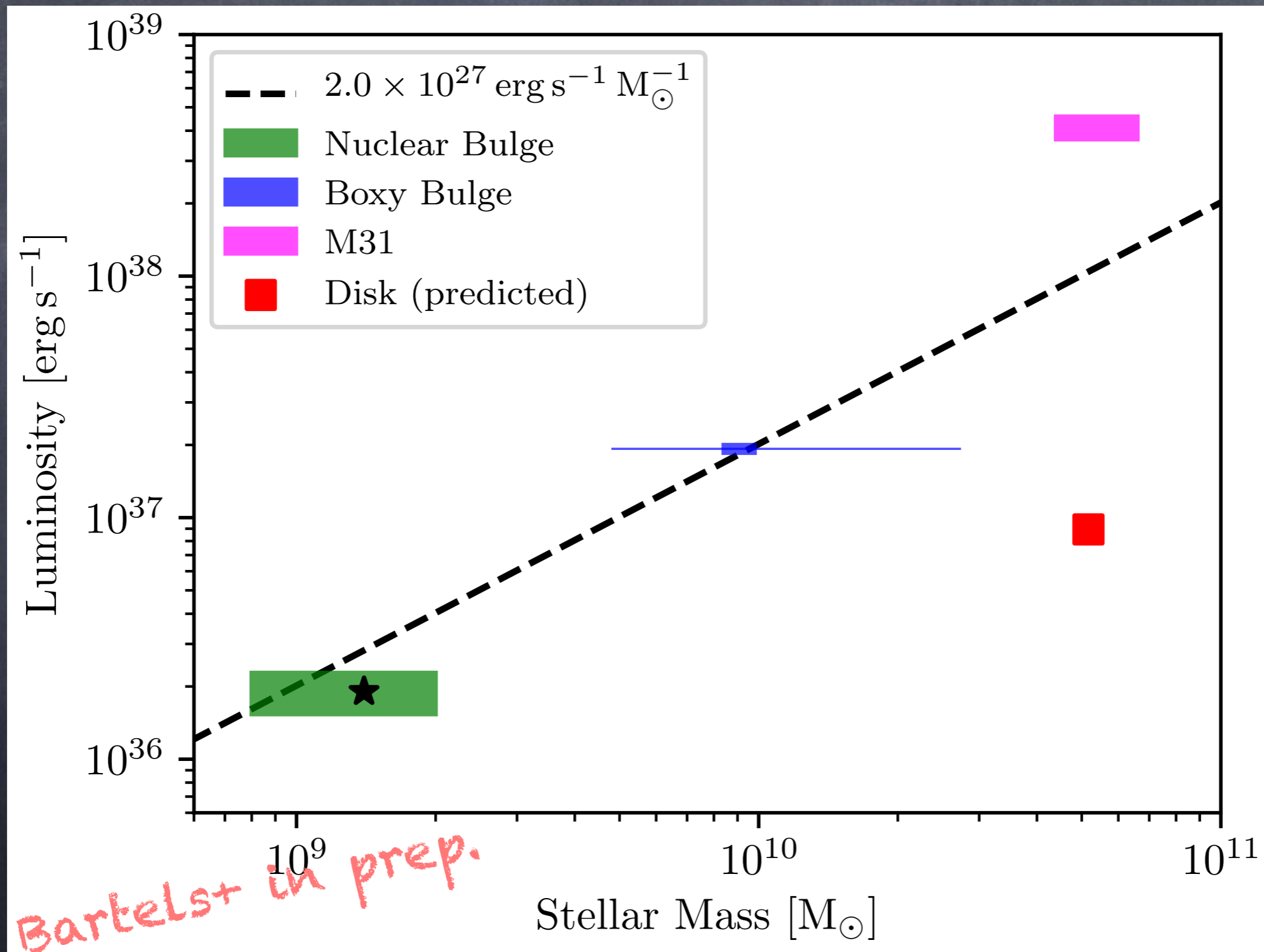
$$3 \times 10^{-09} \text{ erg cm}^{-2} \text{ s}^{-1}$$

• Lum. vs. Mass

$$1.8 \times 10^{26} \text{ erg s}^{-1} M_{\odot}$$

$\sim 10x$ more MSPs per M_{\odot}

GCE vs. Stellar Mass



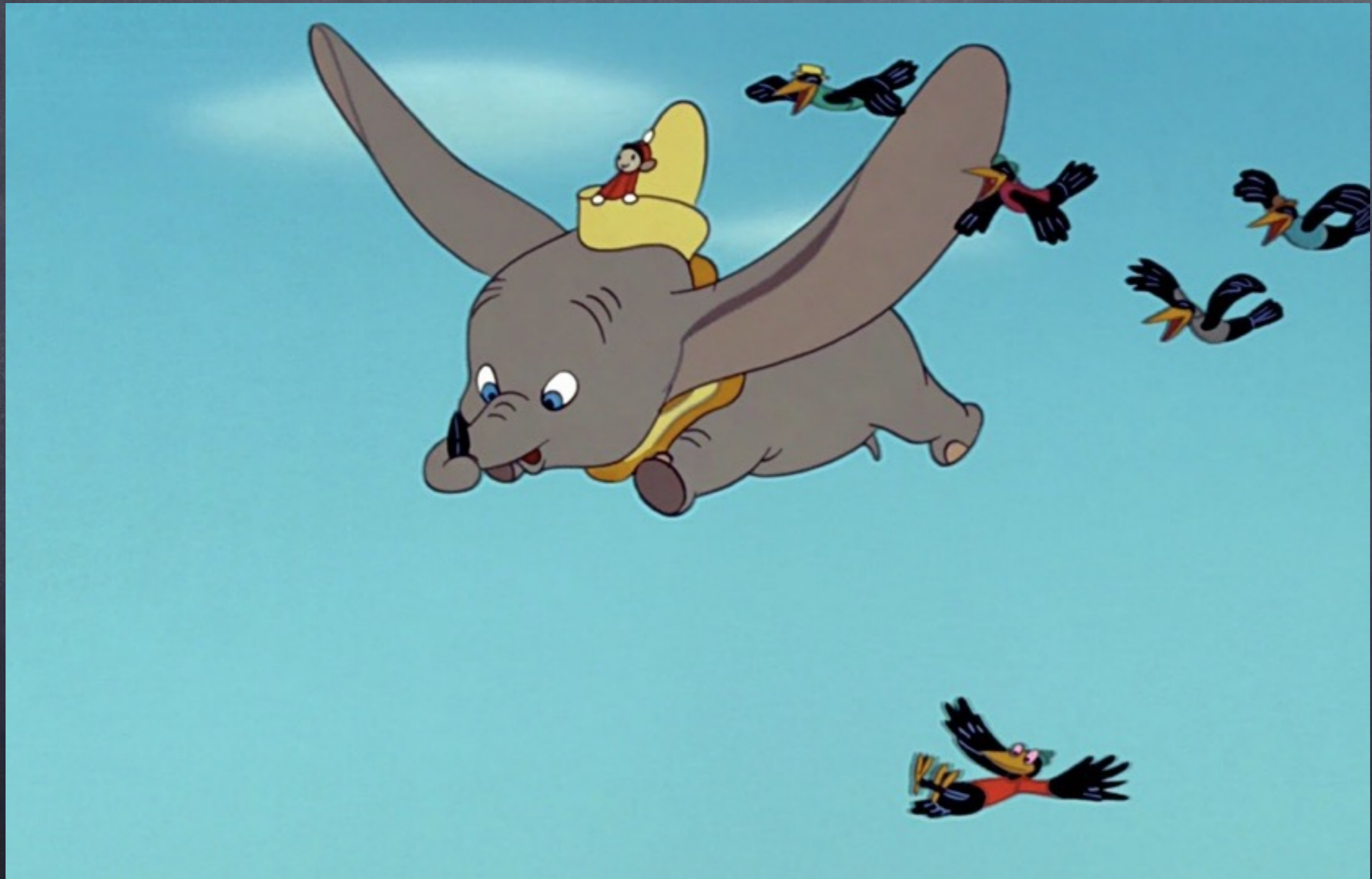
Can we detect it?

- 8% of ICS in our ROI ($180^\circ \times 41^\circ$)
- 2% of π^0
- This will be hard ...

Conclusion

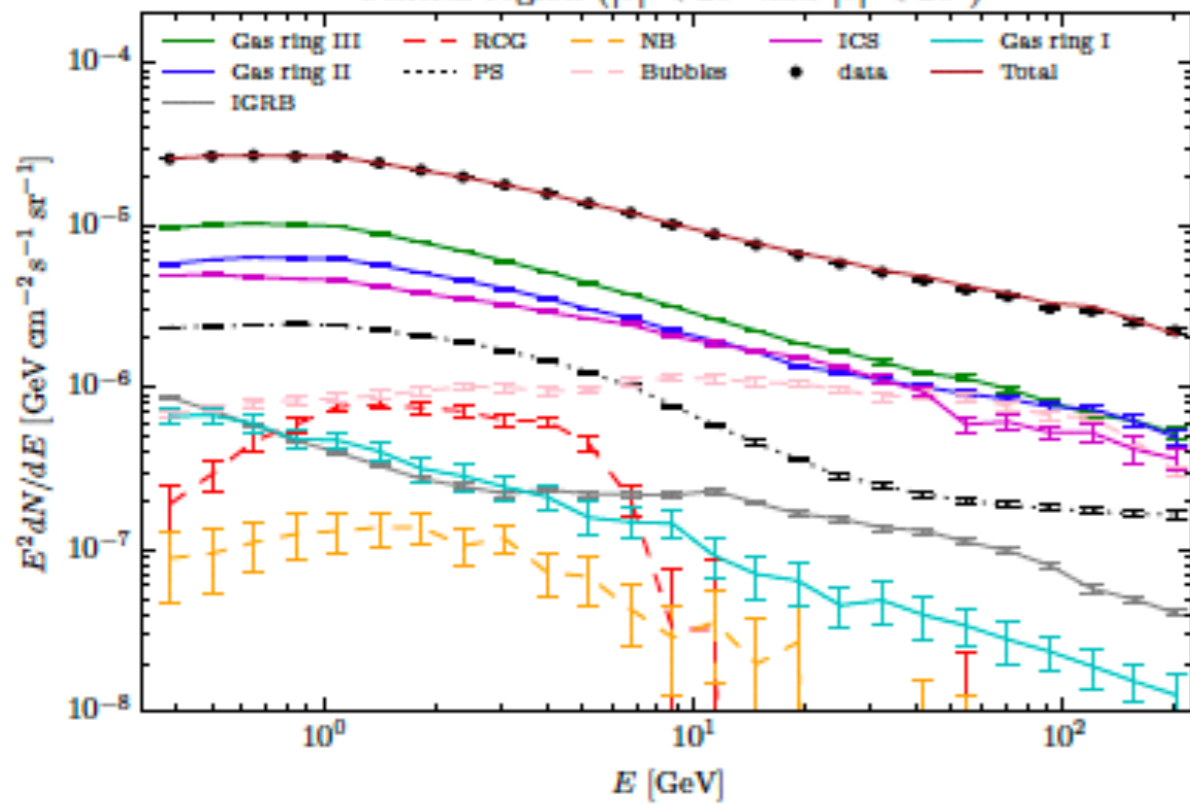
- Reanalysed the GCE with SkyFACT
- We find a correspondence with the Bulge + nuclear bulge
- GCE appears to trace stellar mass

THANK YOU :)



Backup: spectra

Central region ($|b| < 20^\circ$ and $|\ell| < 20^\circ$)



Central region ($|b| < 20^\circ$ and $|\ell| < 20^\circ$)

