

- motivating the model
- Lagrangian and symmetries
- anticipating the mass spectrum
- lattice methods
- lattice results
- implications for phenomenology

Baryonic matter gets about 99% of its mass from QCD.

The proton is stable due to an accidental $U(1)_B$ in the Standard Model.

Dark matter might also arise from a non-abelian gauge theory.

The lightest baryon could be stable due to an accidental $U(1)_B$.

What is the minimal such model?

- smallest non-abelian gauge theory: $SU(2)$
- fewest fermions: 1 dark quark
- smallest fermion representation: the fundamental
- no renormalizable couplings to the Standard Model.

Is the dark baryon stable enough to be dark matter?

In general, $SU(2)$ permits decay via $\mathcal{L} = \frac{1}{\Lambda} \bar{q}q|H|^2$.

As we'll see, the 1-flavor case is special, having no such problem.

The Lagrangian of this model,

$$\mathcal{L} = -\frac{1}{2} \text{Tr}(F_{\mu\nu} F^{\mu\nu}) + \mathcal{L}_{\text{fermion}} + \mathcal{L}_{\text{higher dim}},$$

has a global $U(2)$ symmetry because the gauge theory is pseudo-real.

$$\begin{aligned} \mathcal{L}_{\text{fermion}} &= \bar{q}(i\not{D} - m)q \\ &= \bar{Q}i\not{D}Q - \frac{m}{2} \left(Q^T i\sigma^2 C E Q + \bar{Q} i\sigma^2 C E \bar{Q}^T \right) \end{aligned}$$

where C is the charge conjugation matrix, σ^2 is a color Pauli, and

$$Q = \begin{pmatrix} q_L \\ -i\sigma^2 C \bar{q}_R^T \end{pmatrix}, \quad E = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}.$$

The global $U(2) = SU(2)_B \times U(1)_A$ has an anomalous axial $U(1)_A$.

The baryonic isospin $SU(2)_B$ remains unbroken.

Therefore there will be no Goldstone bosons.

Operators for scalar, pseudoscalar, vector and axial vector mesons are

$$\begin{aligned}\bar{q}q &= \frac{1}{2} \left(Q^T i\sigma^2 C E Q + \bar{Q} i\sigma^2 C E \bar{Q}^T \right) \\ \bar{q}\gamma_5 q &= -\frac{1}{2} \left(Q^T i\sigma^2 C E Q - \bar{Q} i\sigma^2 C E \bar{Q}^T \right) \\ \bar{q}\gamma^\mu q &= \bar{Q}\gamma^\mu \tau^3 Q \quad \Rightarrow \text{part of an isotriplet} \\ \bar{q}\gamma^\mu \gamma^5 q &= \bar{Q}\gamma^\mu Q\end{aligned}$$

where τ^3 is an isospin Pauli.

The dark matter candidate is the baryon, coupling to $\bar{Q}\gamma^\mu \tau^\pm Q$. It has no dimension 5 decay, so it is cosmologically allowed.*

More than 1 flavour would have given *spin-zero* baryons that can decay at dim 5, eg. $\mathcal{L} = \frac{1}{\Lambda} \bar{q}q |H|^2$, which is a cosmological problem.*

* Appelquist et al., PRD92 (2015) 075030

We use the Wilson action with $\beta = \frac{4}{g^2} = 2.2$

and $\kappa = (8 + 2m)^{-1} = \text{several values}$.

The main study is $12^3 \times 32$ with >1000 configurations per ensemble.

We also use $12^3 \times 48$ for some topics.

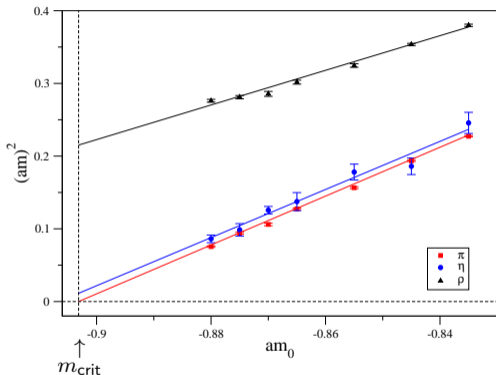
We use the RHMC algorithm from the HiRep code.

Disconnected quark loops use an unbiased stochastic estimator.

Del Debbio, Patella, Pica, PRD81 (2010) 094503

We use Z_2 stochastic wall sources (Z2SEMWall).

Boyle et al., JHEP 08 (2008) 086



π is not a physical particle but it identifies the massless limit by $m_\pi^2 = 0$.
We can define a shifted bare quark mass $m_q \equiv m_0 - m_{crit}$.

Topological susceptibility, χ , vanishes as $m_q \rightarrow 0$.

χ is a gauge field quantity in the theory's integration measure.

A Euclidean lattice definition is

$$\chi = \frac{\langle Q^2 \rangle}{L^3 T}, \quad Q = \sum_x q(x), \quad q(x) = -\frac{1}{32\pi^2} \epsilon_{\mu\nu\rho\sigma} \text{Tr} \left[F_{\mu\nu}(x) F_{\rho\sigma}(x) \right]$$

Statistical noise is reduced
by the “slab method”.

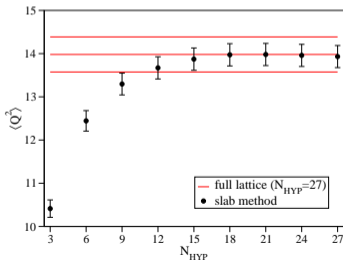
Bietenholz, deForcrand, Gerber, JHEP12(2015)070
JLQCD(Aoki et al.) PTEP2018(2018)043B07

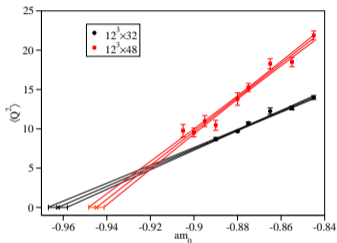
Temporal slabs: $\frac{T}{8} < \text{width} < \frac{T}{2}$.

Spatial slabs: $\frac{L}{3} < \text{width} < \frac{2L}{3}$.

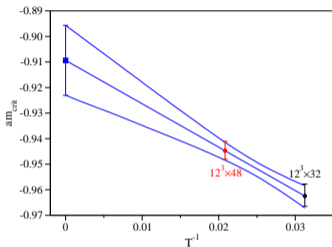
We also use HYP smearing.

Hasenfratz, Knechtli, PRD64(2001)034504





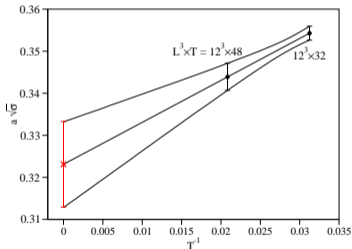
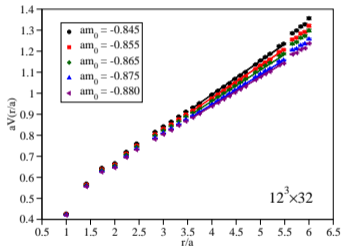
↑ ↑
 m_{crit} for each lattice



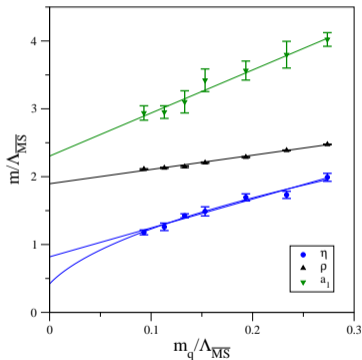
↑
 extrapolated m_{crit}

This agrees with the result from method #1.

We calculate the static potential, extract the string tension, then obtain the scale: $\Lambda_{\overline{MS}} = 0.7712\sqrt{\sigma}$. Bali Phys.Rep.343(2001)1



Conclusion: $a\sqrt{\sigma} = 0.323(10) \Rightarrow a\Lambda_{\overline{MS}} = 0.249(8)$.

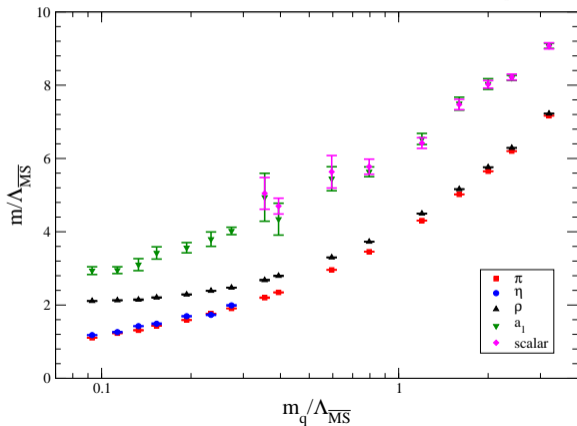


Masses are obtained from standard correlator fits.

ρ is the dark matter candidate.
 It has no disconnected diagrams.
 Lattice data for ρ are precise.
 For a light quark, $m_\rho \sim 2\Lambda_{\overline{MS}}$.

η fits are $m_\eta = c_0 + c_1 m_q$ and
 $m_\eta^2 = c_0 + c_1 m_q$.

The expected degeneracies emerge at large m_q .



f_η is defined by

$$Z_A \langle 0 | A_t(0) | P \rangle = f_\eta m_\eta$$

Fitting lattice data to

$$C_{A_t P}(t) = AB \left(e^{-m_\eta t} - e^{-m_\eta(T-t)} \right)$$

$$C_{PP}(t) = B^2 \left(e^{-m_\eta t} + e^{-m_\eta(T-t)} \right)$$

gives the decay constant

$$f_\eta = Z_A A (2\kappa)^2 \sqrt{\frac{2}{m_\eta L^3}}$$

f_ρ is defined by

$$Z_V \langle 0 | V_i(0) | V \rangle = f_\rho m_\rho \epsilon_i$$

Fitting lattice data to

$$C_{\sigma_t V}(t) = CD \left(e^{-m_\rho t} - e^{-m_\rho(T-t)} \right)$$

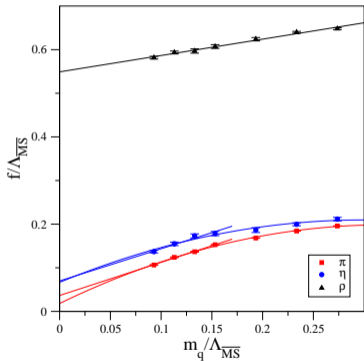
$$C_{VV}(t) = D^2 \left(e^{-m_\rho t} + e^{-m_\rho(T-t)} \right)$$

gives the decay constant

$$f_\rho = Z_V D (2\kappa)^2 \sqrt{\frac{2}{m_\rho L^3}}$$

One-loop perturbative expressions for Z_V and Z_A are taken from

Del Debbio et al. JHEP06(2008)007.



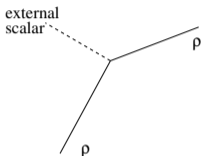
The data for f_ρ are linear.

Here, 2 fit options are shown:

$$f = c_0 + c_1 m_0 \text{ for small } m_0,$$

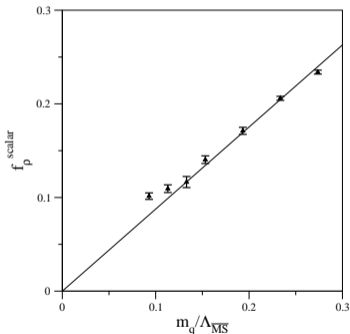
$$f = c_0 + c_1 m_0 + c_2 m_0^2.$$

The dark matter candidate and Higgs may interact by a scalar coupling.



The Feynman-Hellman theorem leads to a dimensionless ratio:

$$f_{\rho}^{\text{scalar}} = \frac{m_q}{m_{\rho}} \frac{\partial m_{\rho}}{\partial m_q}$$



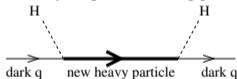
See, for example,

Junnarkar, Walker-Loud, PRD87(2013)114510

Detmold, McCullough, Pochinsky, PRD90(2014)114506

- The dark hadrons that are **stable under dark SU(2)** are
 - η = lightest meson
 - ρ = lightest baryon
- Dark hadrons have no renormalizable couplings to the Higgs.
 - Higgs $\rightarrow \eta\eta$ might occur if $2m_\eta < m_H$. (Also Higgs $\rightarrow \rho\rho$.)
 - However, the coupling to the Higgs **should be indirect**.

Example:



- Dark hadron decay rates:
 - η can decay at **dim 5**, for example $\mathcal{L} = \frac{1}{\Lambda} \bar{q}q|H|^2$.
 - This allows $\tau_\eta \lesssim 1$ second, evading BBN constraints.
 - ρ can decay at **dim 6**, for example $\mathcal{L} = \frac{1}{\Lambda^2} \bar{q}\gamma_\mu q H \nabla^\mu H$.
 - This allows $\tau_\rho \gtrsim 10^{24}$ years, making ρ a dark matter candidate.

This model has another nice feature:

The dark matter relic density is set naturally, through freeze-out.

At early times (high temperatures): $\rho\bar{\rho} \rightleftharpoons \eta\eta$

At late times (low temperatures): $\rho\bar{\rho} \rightarrow \eta\eta$

Recall $\tau_{\eta} \lesssim 1$ second.

This freeze-out mechanism

- works because $m_{\rho} > m_{\eta}$, as the lattice results show.
- is special to the 1-flavour theory because ρ is the dark matter.

Is dark matter the lightest hadron in a non-abelian gauge theory?

Summary:

- The minimal model is $SU(2)$ with 1 fundamental fermion.
- With only 1 fermion, the dark matter particle is spin 1, not spin 0.
- Lattice calculations give hadron masses, decay constants, scalar couplings, and the scale $\Lambda_{\overline{MS}}$.
- This minimal model can satisfy phenomenological constraints.
- Dark matter freeze-out is natural.

Possible future work:

- extended lattice study, with multiple lattice spacings and volumes.
- study of dark hadron scattering and dark nuclear physics.

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