Baryon Asymmetry

Overview	_		Edit PDF iCal
Timetable	Baryon	Asymmetry	
Contribution List	Place Location: T	The Hong Kong University of Science and Technology	
My Conference	Address: H	ong Kong, China	
Registration	Files		Edit files
Registration Form	Date: 28 .	Jun 19:30 - 21:30	
	Descriptic Discussion Kingdom)	on Leader: Pasquale Di Bari (University of Southampton, United	
	Timetable	e Contribution List	
	Wed 2	28/06	
		Print	Full screen Filter
	19:00		
		Electroweak Baryogenesis	Thomas KONSTANDIN
		The Hong Kong University of Science and Technology	19:30 - 19:45
	20:00	Unifying Inflation with the Axion, Dark Matter, Baryogenesis and the Mechanism	Seesaw Carlos TAMARIT
		Baryogenesis from Right-Handed Neutrino Oscillations	Takehiko ASAKA
		The Hong Kong University of Science and Technology	20:10 - 20:25
		Leptogenesis from Realistic Models	Chee Sheng FONG
		The Hong Kong University of Science and Technology	20:30 - 20:45
		Models on the Origin of Ordinary and Dark Matter	Peihong GU
	21:00	The Hong Kong University of Science and Technology	20:50 - 21:05

Explain baryon asymmetry of the universe too!



Restrictions from neutrino masses

• CP violation is tied to neutrino mass scale



Caveats

 $M_i \sim M_j$

• "Heavier states" not much heavier → *resonant enhancement*



Leptogenesis (realistic models) - C. S. Fong

Resonant leptogenesis

[Pilaftsis (hep-ph/9707235)] [Pilaftsis & Underwood (hep-ph/0309342)]

- The right-handed neutrinos are guasi-degenerate to enhance the CP violation from self energy corrections
- Can be probed at LHC & CLFV [Bray, Lee & Pilaftsis (hep-ph/0702294)]
- Quasi-degeneracy $M_2 M_1 = \mu \ll M_{1,2} \approx M$ due to

 - Approximate family symmetry (ok)
 Soft SUSY breaking terms (ok)
 [Grossman et al. (hep-ph/0307081)]
 [D' Ambrosio et al. (hep-ph/0308031)]
 [Review (hep-ph/1107.5312)]
 - Approximate lepton number (difficult)

Difficult because mass degeneracy is tied to CP violation!

$$M_2 - M_1 = \mu \qquad \qquad |\epsilon| \propto \mu$$

Baryogenesis via neutrino oscillation

- Oscillation of RH neutrinos can be a source of BAU Akhmedov, Rubakov, Smirnov ('98) / TA, Shaposhnikov ('05)
 - Oscillation starts at $T_{osc} \simeq (M_0 M_N \Delta M)^{1/3}$

Medium effects

NÍL

- **\square** Asymmetries are generated since evolution rates of L_{α} and
 - $\overline{L_{\alpha}}$ are different due to CPV





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2017/06/28

Baryogenesis region



2017/06/28

Sensitivities by future searches



Normal Hierarchy

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2017/06/28

First-order phase transition

The free energy (as a function of the Higgs vev) decides the nature of the phase transition:



What are the challenges?



Pre LHC EWBG

Before LHC, the main focus was on supersymmetric models.



Strong first-order electroweak phase transition from light stops

 $m_{\tilde{t}} \lesssim m_t$

NMSSM

[Menon, Morrisey, Wagner '04] [Huber, TK, Prokopec, Schmidt '06] U'(1) MSSM [Kang, Langacker, Li, Liu '04]



CP violation From the chargino sector

 $m_{\chi^{\pm}} < 200 \,\mathrm{GeV}$



problematic, also because of EDMs

EWBG in the LHC era

0.5

0.0

0.0

0.5

After LHC run II, the focus in EWBG is more on minimal models:



Strong first-order electroweak phase transition from extended scalar sector

Two Higgs doublet model



CP violation from the Higgs sector

Singlet extension with a low cutoff



CP violation from the dim-5 top-singlet operators new dof low cutoff in principle testable

0.0

Composite Higgs models

The Higgs could be a Pseudo-Goldstone boson of a broken global symmetry

QCD:
$$\frac{SU(2)_L \times SU(2)_R}{SU(2)_V} \to 3\pi$$

The broken symmetry will determine the light degrees of freedom and their quantum numbers

$$\frac{SO(5)}{SO(4)} \to H$$

but also

$$\frac{SO(6)}{SO(5)} \to H + S$$

$$\frac{SO(6)}{SO(4) \times SO(2)} \to 2H$$

[Kaplan, Georgi '84]

Ingredients

Two ingredients of baryogenesis are missing in the Standard Model. These are provided in models that have an additional singlet in the low energy effective description



Strong first-order electroweak phase transition

CP violation from dimension-five operators

$$\mathcal{L} \ni y_t \bar{\psi}_Q H \psi_t + \frac{\tilde{y}_t}{f} S \bar{\psi}_Q H \psi_t + h.c.$$
$$\Im(y_t \tilde{y}_t^*) \neq 0$$

Baryogenesis



CPV is mostly present during the phase transition and does not require sizable mixing the broken phase \rightarrow nightmare scenario