

Baryon Asymmetry

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Baryon Asymmetry

Place

Location: The Hong Kong University of Science and Technology

Address: Hong Kong, China

Files

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Date: **28 Jun 19:30 - 21:30**

Description

Discussion Leader: Pasquale Di Bari (University of Southampton, United Kingdom)

[Timetable](#) | [Contribution List](#)

Wed 28/06

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19:00

Electroweak Baryogenesis

Thomas KONSTANDIN

The Hong Kong University of Science and Technology

19:30 - 19:45

20:00

Unifying Inflation with the Axion, Dark Matter, Baryogenesis and the Seesaw Mechanism

Carlos TAMARIT

Baryogenesis from Right-Handed Neutrino Oscillations

Takehiko ASAKA

The Hong Kong University of Science and Technology

20:10 - 20:25

Leptogenesis from Realistic Models

Chee Sheng FONG

The Hong Kong University of Science and Technology

20:30 - 20:45

21:00

Models on the Origin of Ordinary and Dark Matter

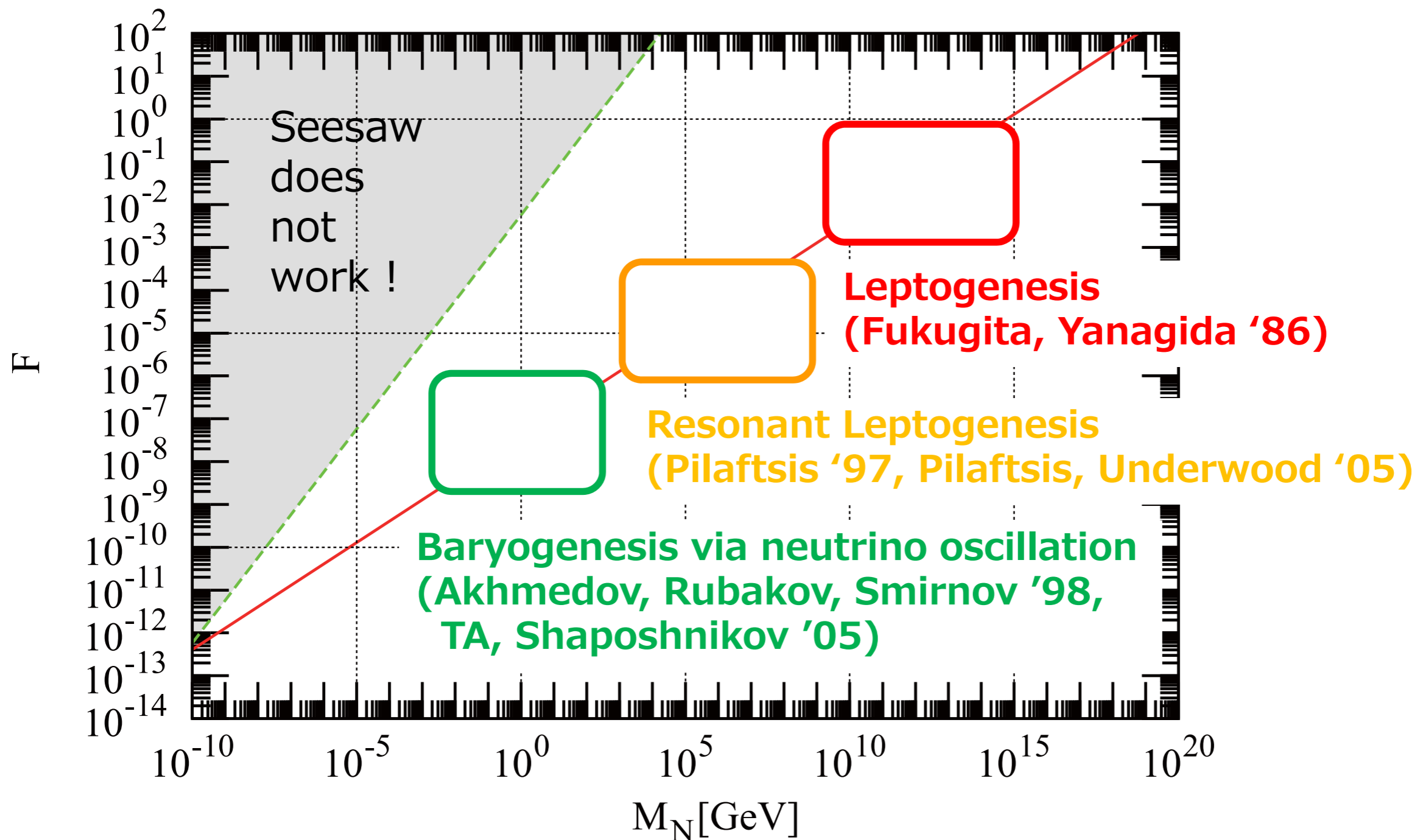
Peihong GU

The Hong Kong University of Science and Technology

20:50 - 21:05

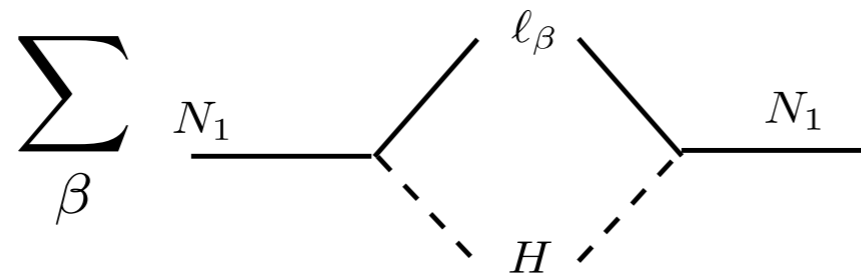
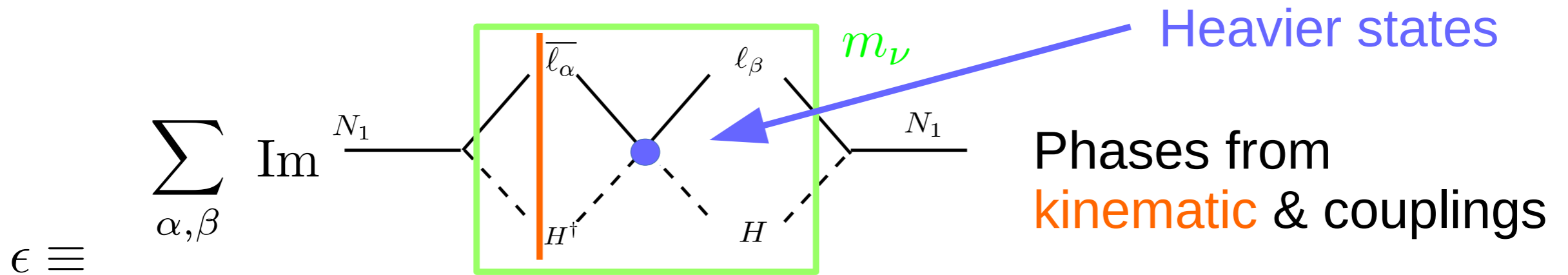
Yukawa coupling and Majorana mass

Explain baryon asymmetry of the universe too!

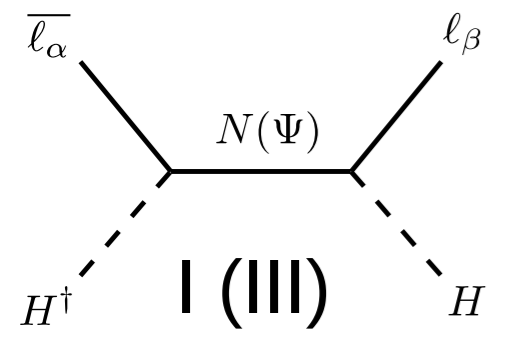


Restrictions from neutrino masses

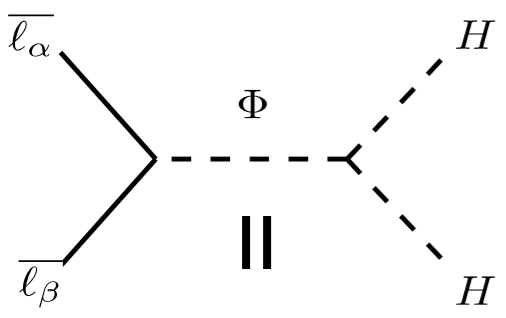
- CP violation is tied to neutrino mass scale



$$|\epsilon| \sim \frac{1}{16\pi} \frac{m_\nu M_1}{v^2}$$



$|\epsilon| \gtrsim 10^{-7} \implies M_1 \gtrsim 3 \times 10^9 \left(\frac{0.05 \text{ eV}}{m_\nu} \right) \text{ GeV}$



[Davidson & Ibarra (hep-ph/0202239)]

Caveats

$$M_i \sim M_j$$

- “Heavier states” not much heavier → **resonant enhancement**

Type I seesaw

$$\epsilon_i \equiv \sum_{\alpha, \beta} \text{Im} N_i$$

$$\sum_{\beta} N_i$$

[Pilaftsis (hep-ph/9707235)]

$$\epsilon_i \sim \frac{1}{16\pi} \frac{m_\nu M_i}{v^2} \times \sum_j \sin(\phi_i - \phi_j) \frac{(M_i^2 - M_j^2) M_j^2}{(M_i^2 - M_j^2)^2 + M_i^2 \Gamma_j^2}$$

Can go as low as possible i.e. $M_i > 132 \text{ GeV}$ (more later)

Doesn't apply on those which rely on different physical mechanisms...

Resonant leptogenesis

[Pilaftsis (hep-ph/9707235)] [Pilaftsis & Underwood (hep-ph/0309342)]

- The right-handed neutrinos are quasi-degenerate to enhance the CP violation from self energy corrections
- Can be probed at LHC & CLFV [Bray, Lee & Pilaftsis (hep-ph/0702294)]
- Quasi-degeneracy $M_2 - M_1 = \mu \ll M_{1,2} \approx M$ due to
 - Approximate family symmetry (ok)
 - Soft SUSY breaking terms (ok)
 - Approximate lepton number (*difficult*)

[Grossman et al. (hep-ph/0307081)]

[D' Ambrosio et al. (hep-ph/0308031)]

[Review (hep-ph/1107.5312)]

Difficult because mass degeneracy is tied to CP violation!

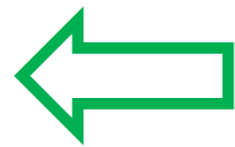
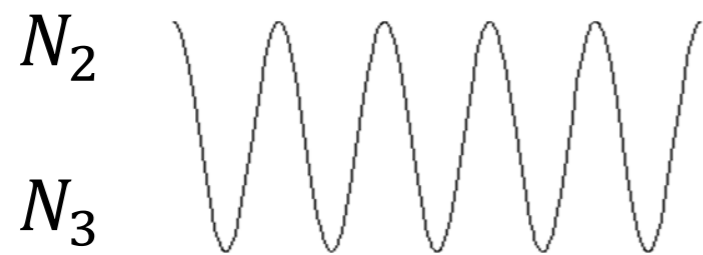
$$M_2 - M_1 = \mu \quad |\epsilon| \propto \mu$$

Baryogenesis via neutrino oscillation

- Oscillation of RH neutrinos can be a source of BAU

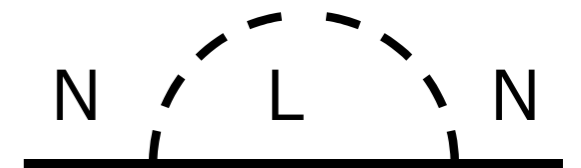
Akhmedov, Rubakov, Smirnov ('98) / TA, Shaposhnikov ('05)

- ▣ Oscillation starts at $T_{osc} \simeq (M_0 M_N \Delta M)^{1/3}$

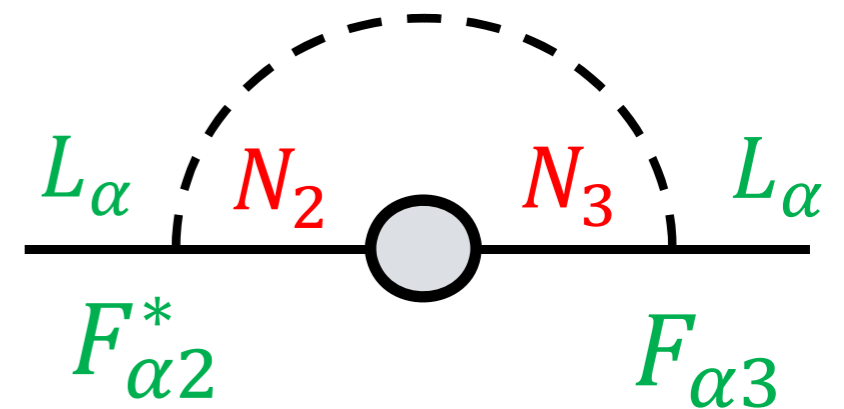
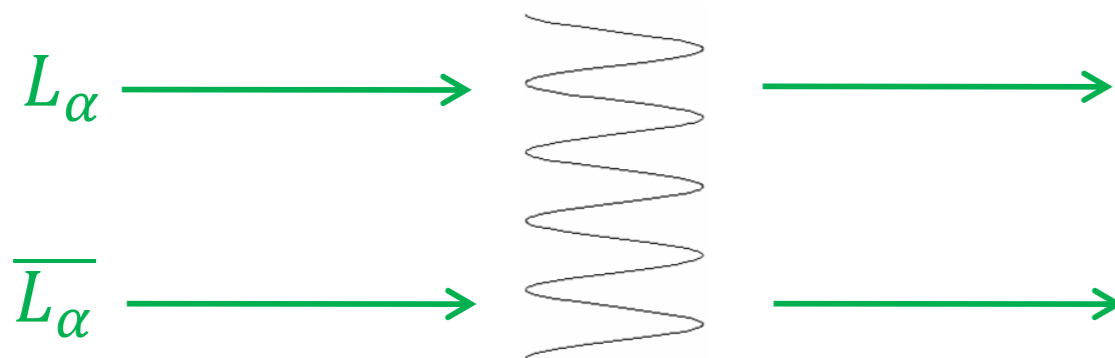


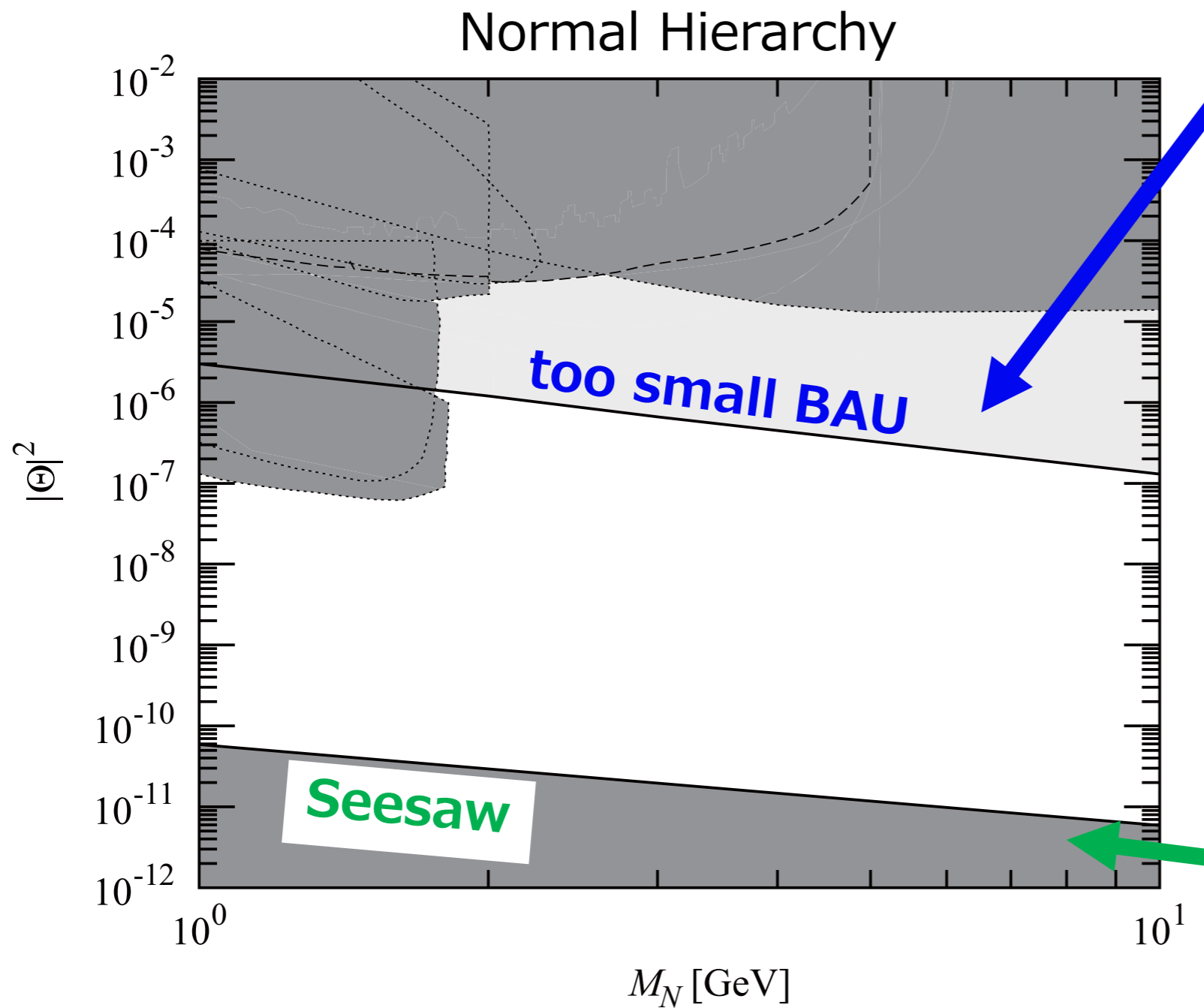
$$V_N = \frac{T^2}{8k} F^\dagger F$$

Medium effects



- ▣ Asymmetries are generated since evolution rates of L_α and \overline{L}_α are different due to CPV





Bound from BAU

to avoid strong washout

Canetti, Shaposhnikov '10
[arXiv:1006.0133]

Drewes, Garbrecht, Gueter,
Klaric '16 [arXiv:1609.09069]

TA, Eijima, Ishida, Minogawa,
Yoshii '17 [to appear]

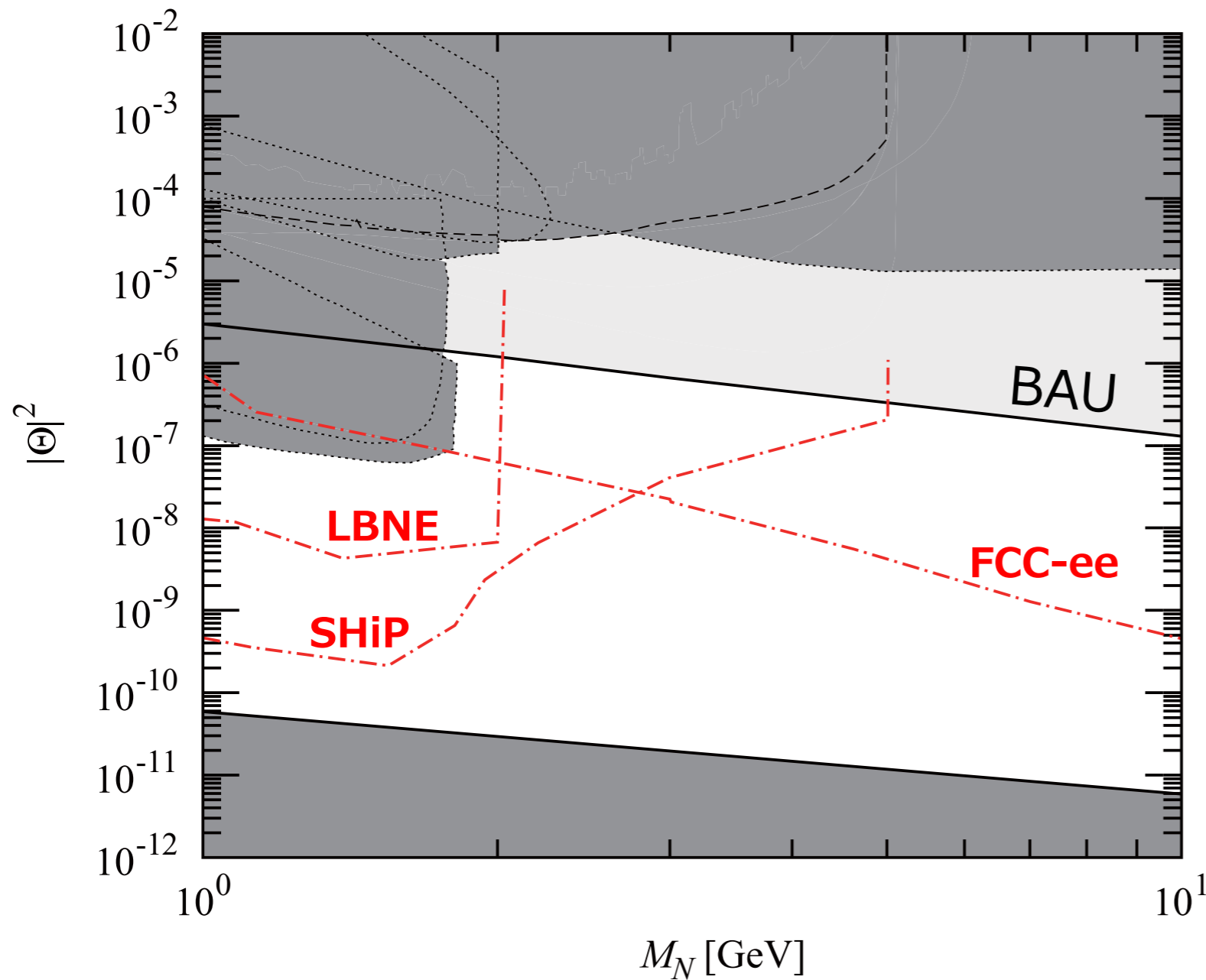
See also a poster #16
by Juraj Klaric

Bound from Seesaw

to explain neutrino masses

$$|\Theta|^2 > \frac{\sum m_i}{2 M_N}$$

Normal Hierarchy



Sensitivity for $|\theta_{\mu}|^2$

- **LBNE (DUNE)**
 N decay inside near detector
Adams et al '13 [arXiv:1307.7335]

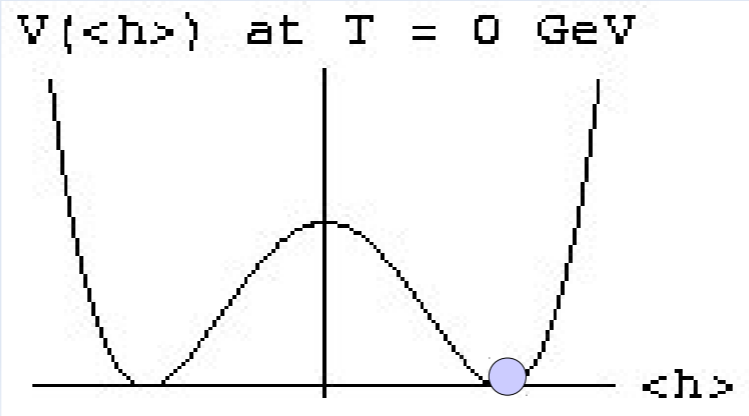
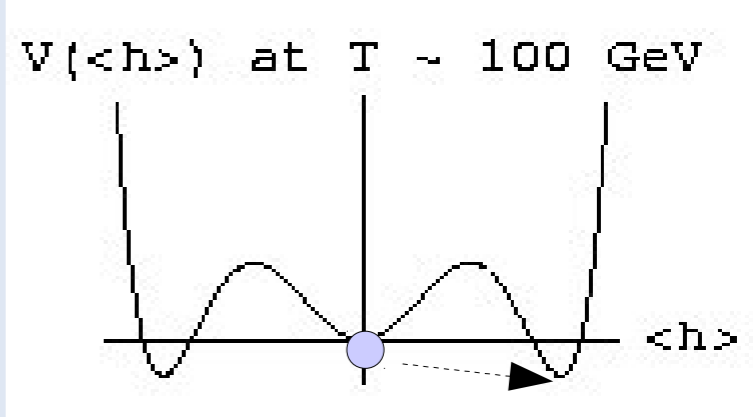
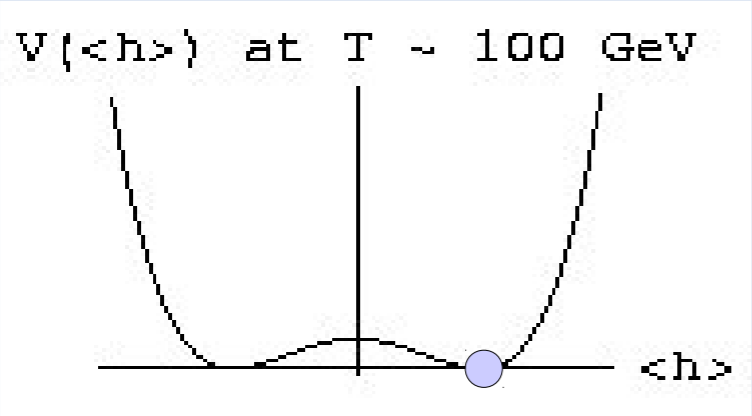
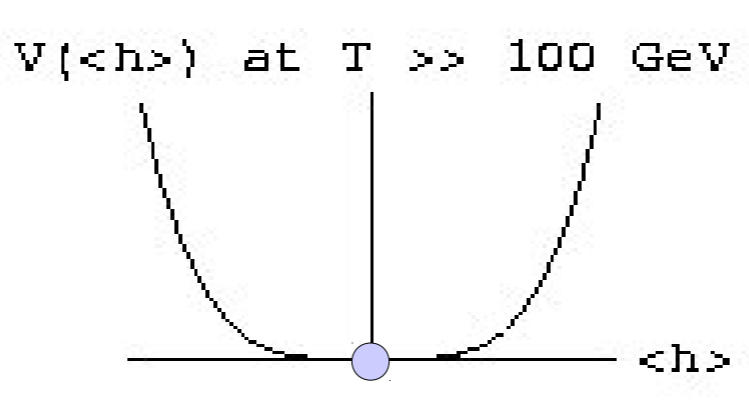
- **SHiP**
 beam dump exp.
Anelli et al '13 [arXiv:1504.04956]

- **FCC-ee at Z-pole**
 displaced vertex of N decay
Blondel, Graverini, Serra, Shaposhnikov (FCC-ee study team) '14 [arXiv:1411.5230]

eq

First-order phase transition

The free energy (as a function of the Higgs vev) decides the nature of the phase transition:

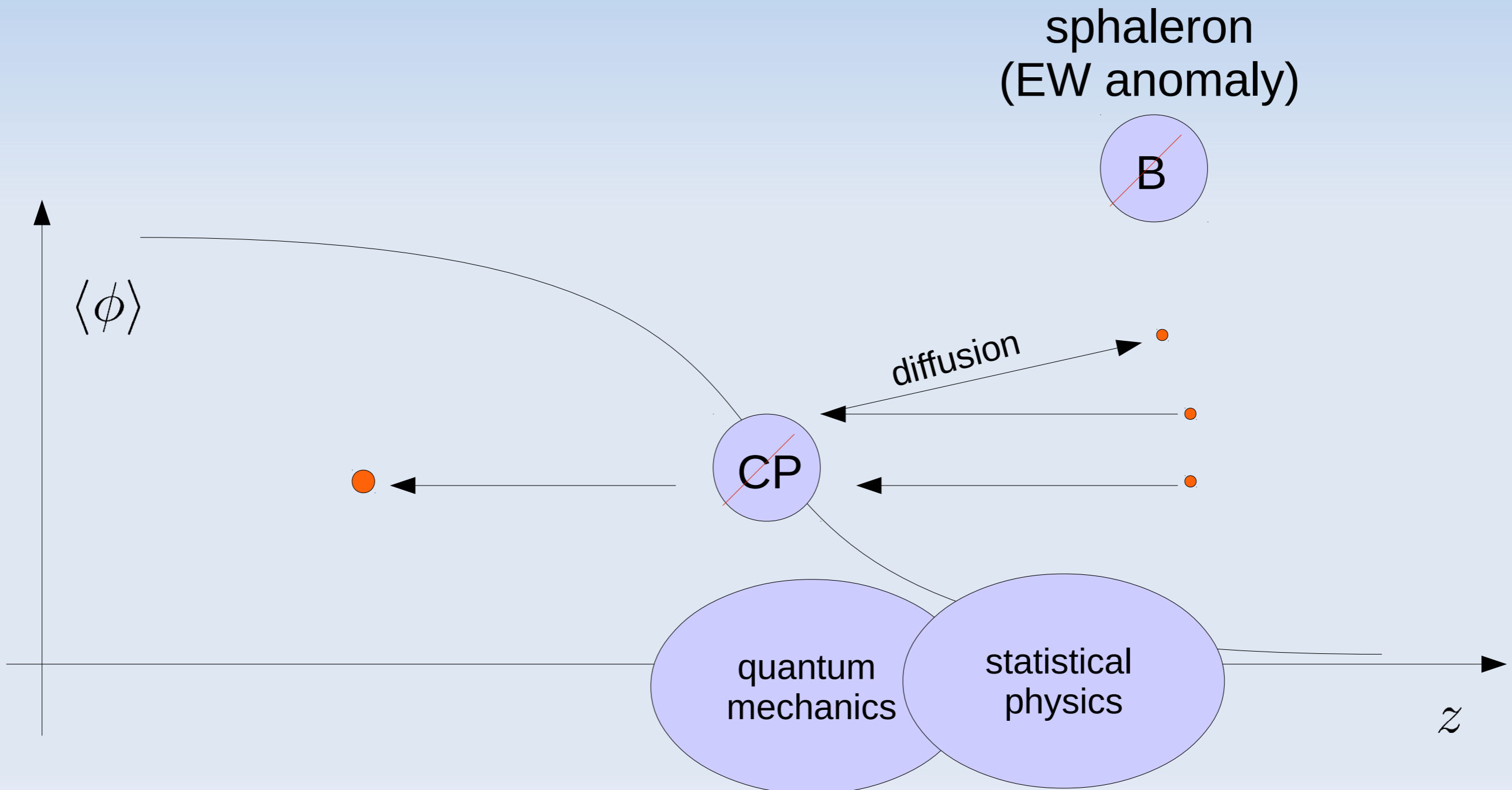


second-order PT
crossover

first-order PT

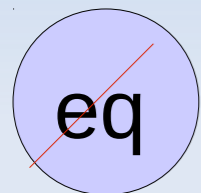
What are the challenges?

[Kuzmin, Rubakov, Shaposhnikov '85]
[Cohen, Kaplan, Nelson '93]



Pre LHC EWBG

Before LHC, the main focus was on supersymmetric models.



Strong first-order electroweak phase transition from light stops

$$m_{\tilde{t}} \lesssim m_t$$



NMSSM

[Menon, Morrisey, Wagner '04]

[Huber, TK, Prokopec, Schmidt '06]

U'(1) MSSM

[Kang, Langacker, Li, Liu '04]



CP violation
From the chargino sector

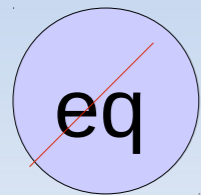
$$m_{\chi^\pm} < 200 \text{ GeV}$$



problematic, also
because of EDMs

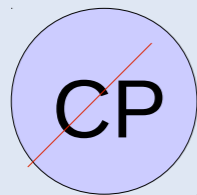
EWBG in the LHC era

After LHC run II, the focus in EWBG is more on minimal models:



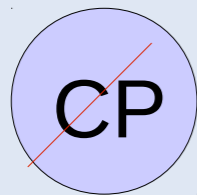
Strong first-order electroweak phase transition from extended scalar sector

Two Higgs doublet model

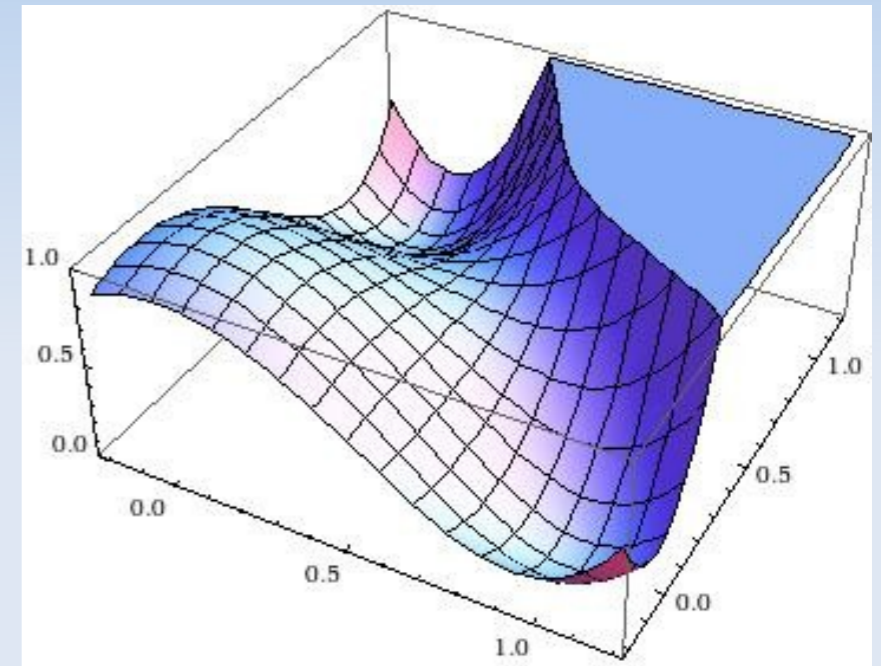


CP violation from the Higgs sector

Singlet extension with a low cutoff



CP violation from the dim-5 top-singlet operators



new dof
low cutoff
in principle testable

Composite Higgs models

The Higgs could be a Pseudo-Goldstone boson of a broken global symmetry

$$\text{QCD: } \frac{SU(2)_L \times SU(2)_R}{SU(2)_V} \rightarrow 3\pi$$

The broken symmetry will determine the light degrees of freedom and their quantum numbers

$$\frac{SO(5)}{SO(4)} \rightarrow H$$

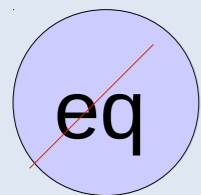
but also

$$\frac{SO(6)}{SO(5)} \rightarrow H + S \quad \frac{SO(6)}{SO(4) \times SO(2)} \rightarrow 2H$$

[Kaplan, Georgi '84]

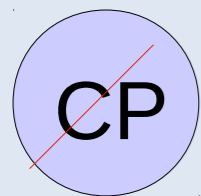
Ingredients

Two ingredients of baryogenesis are missing in the Standard Model. These are provided in models that have an **additional singlet** in the low energy **effective** description



Strong first-order electroweak
phase transition

$$V(s, h)$$



CP violation
from **dimension-five**
operators

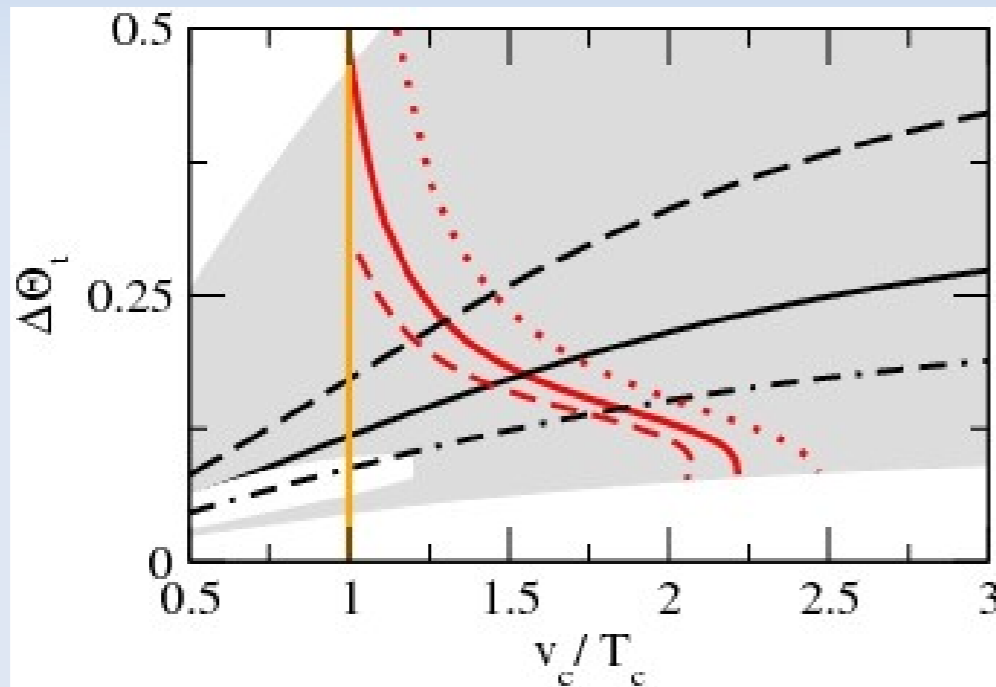
$$\mathcal{L} \ni y_t \bar{\psi}_Q H \psi_t + \frac{\tilde{y}_t}{f} S \bar{\psi}_Q H \psi_t + h.c.$$
$$\Im(y_t \tilde{y}_t^*) \neq 0$$

Baryogenesis

$$\Delta\theta_t \simeq \frac{\Im(y_t \tilde{y}_t^*)}{y_t y_t^*} \frac{\Delta s}{f} \quad m_s = 130 \text{ GeV}$$

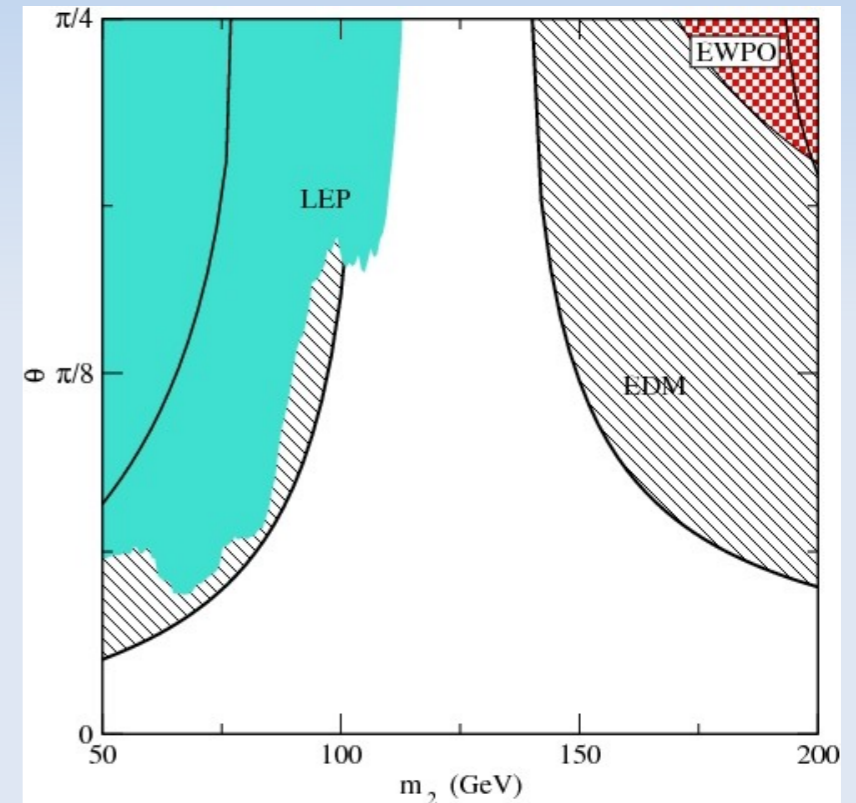
$$\Delta\theta_t \gtrsim 0.15 \quad m_h = 120 \text{ GeV}$$

strength of
CP violation



strength of the phase transition

Higgs-singlet mixing
CP violation



singlet mass

[Espinosa, Gripaio, TK, Riva '11]

CPV is mostly present during the phase transition and does not require sizable mixing the broken phase → nightmare scenario

Dark Matter

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Place

Location: The Hong Kong University of Science and Technology

Address: Hong Kong, China

Files

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Date: **29 Jun 09:00 - 12:35**

Description

Discussion Leader: Jianglai Liu (Shanghai Jiao Tong University, China)

Timetable | [Contribution List](#)

Thu 29/06

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09:00

Particle Theories of Dark Matter

Patrick FOX

The Hong Kong University of Science and Technology

09:00 - 09:25

Dark Matter Direct Detection Experiment

Wolfgang RAU

The Hong Kong University of Science and Technology

09:30 - 09:55

10:00

Dark Matter Searches in Space

Jin CHANG

The Hong Kong University of Science and Technology

10:00 - 10:25

11:00

Experimental Searches of Axions

Yannis SEMERTZIDIS

The Hong Kong University of Science and Technology

10:50 - 11:15

Dark Matter Searches in Colliders

Ning ZHOU

The Hong Kong University of Science and Technology

11:20 - 11:45

12:00

Recent Results from the XENON1T Experiment

Shingo KAZAMA

The Hong Kong University of Science and Technology

11:50 - 12:05

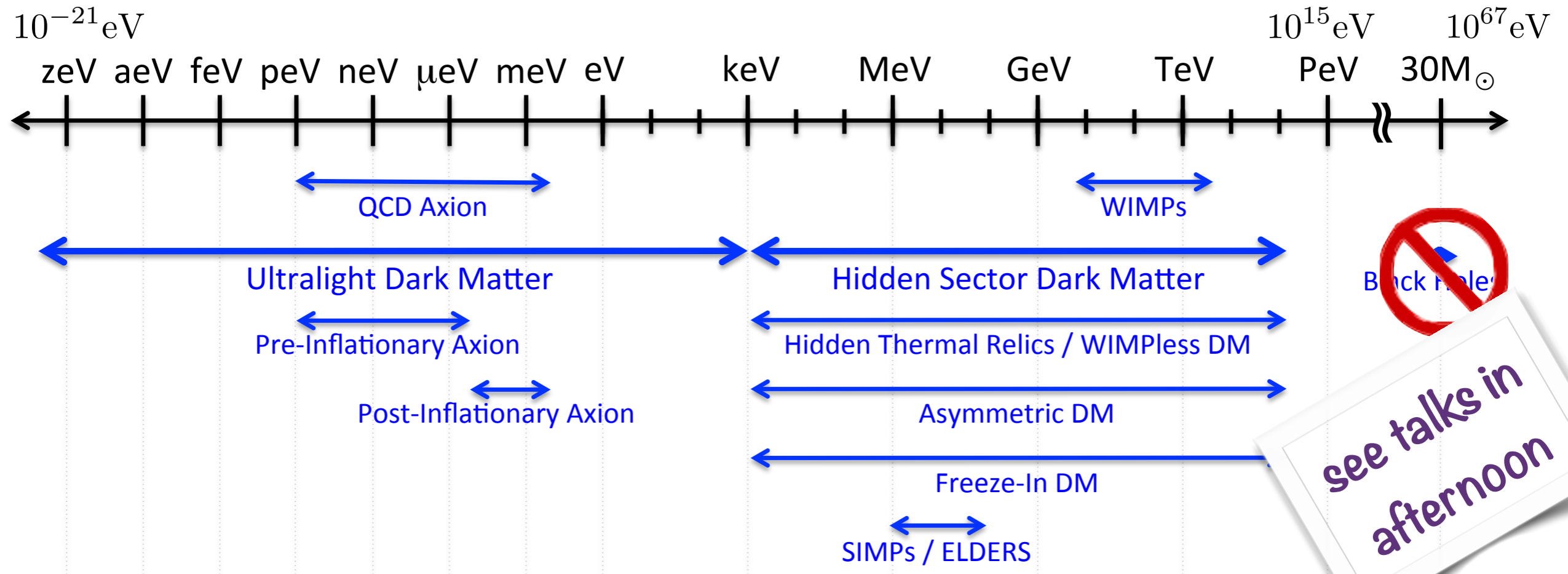
Diversity Problem of Galaxies and Self-Interacting Dark Matter

Hai-Bo YU

The Hong Kong University of Science and Technology

12:10 - 12:25

Particle theories

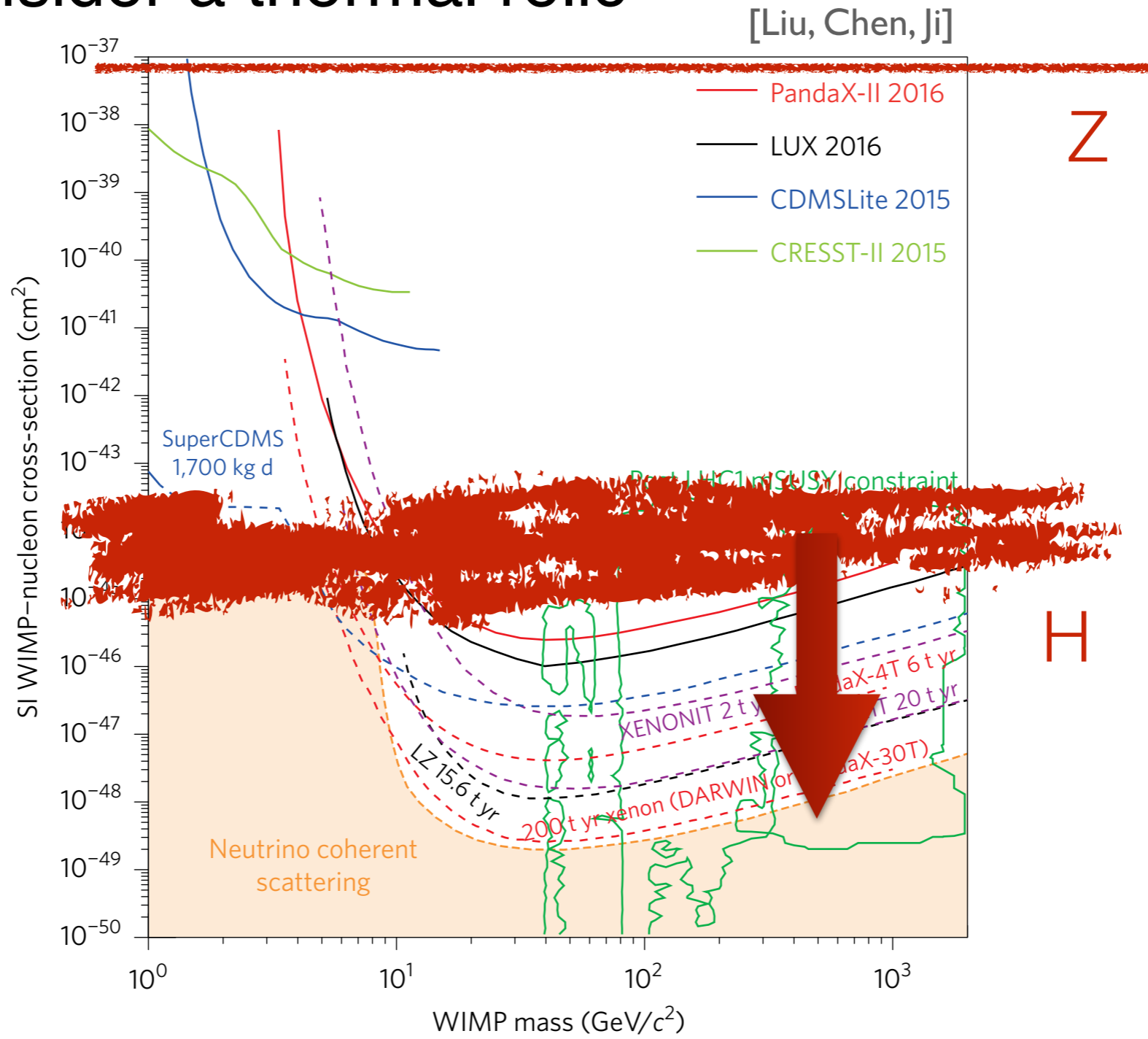


see talks in afternoon

[Feng-US Cosmic Visions White papers]

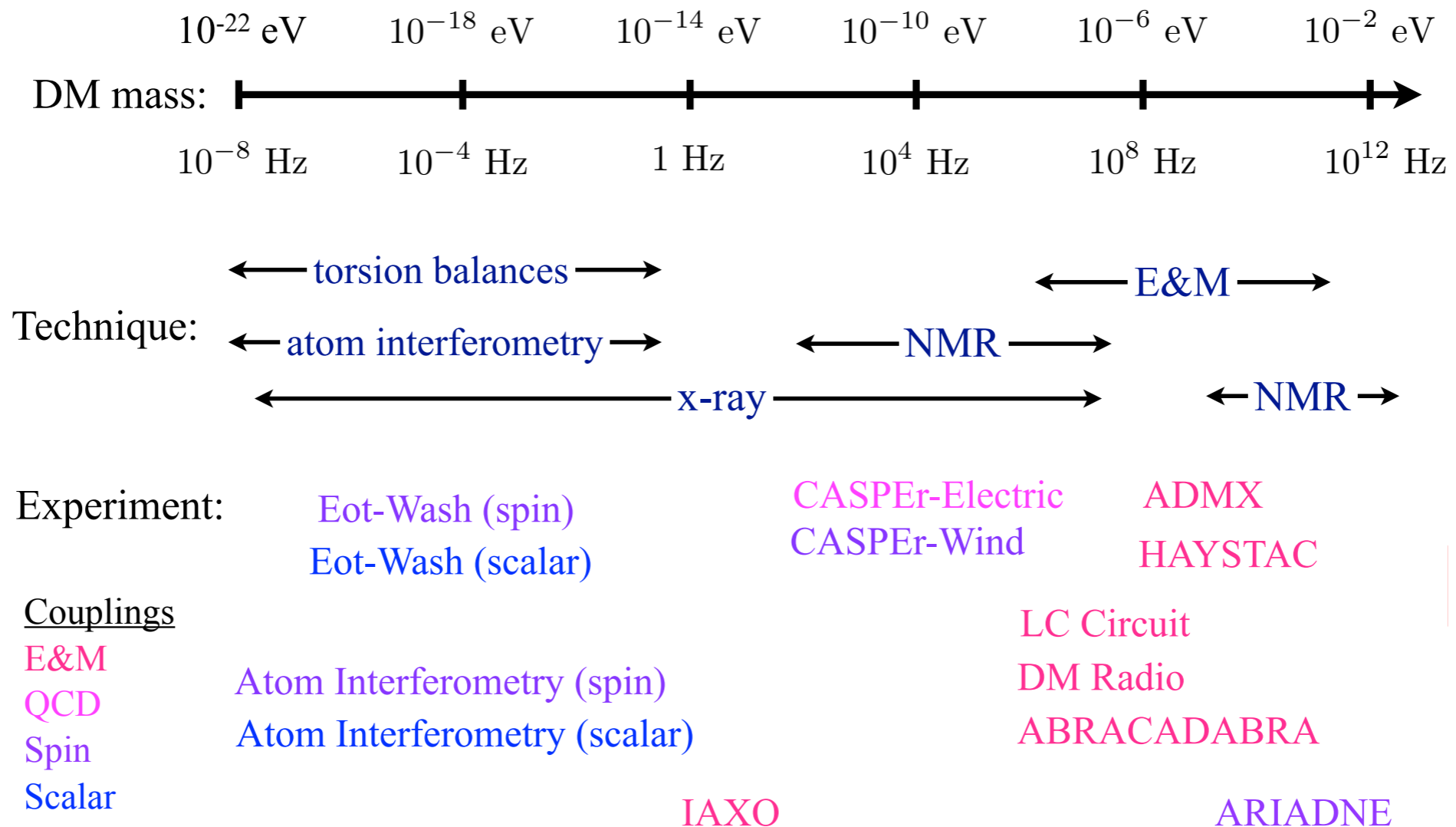
WIMP

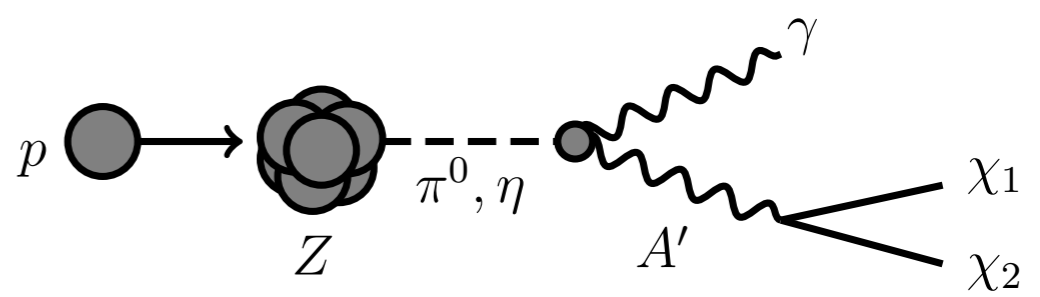
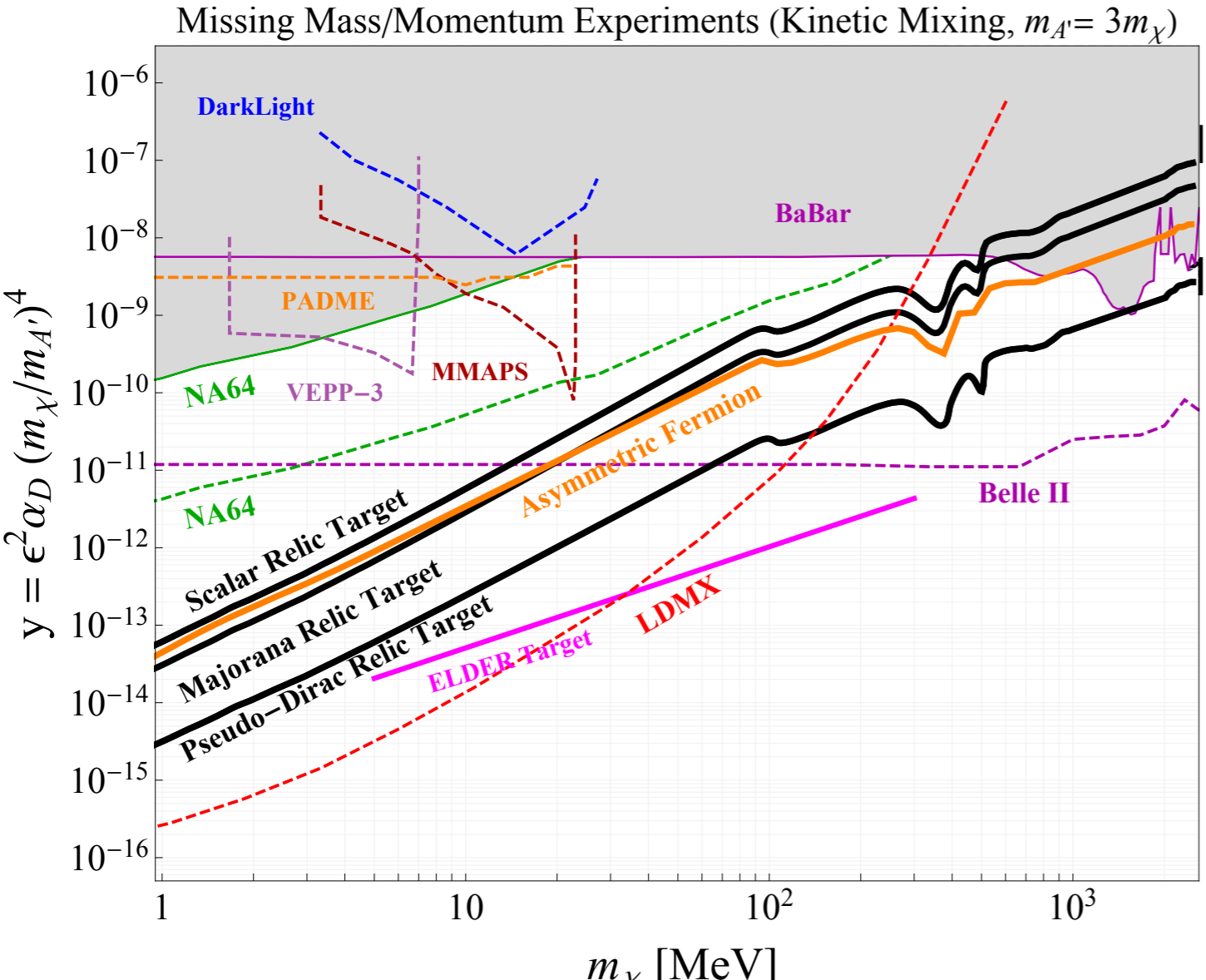
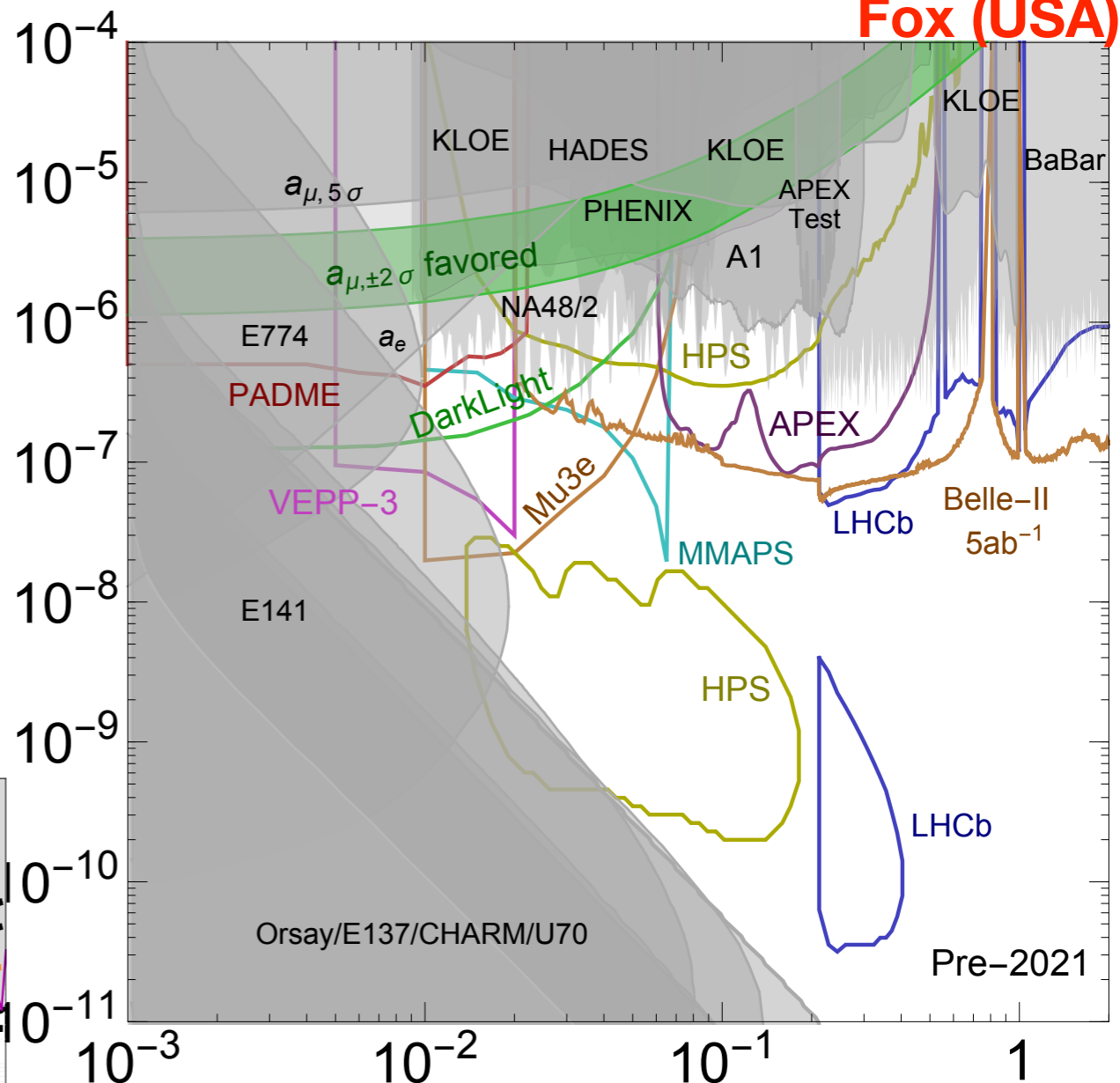
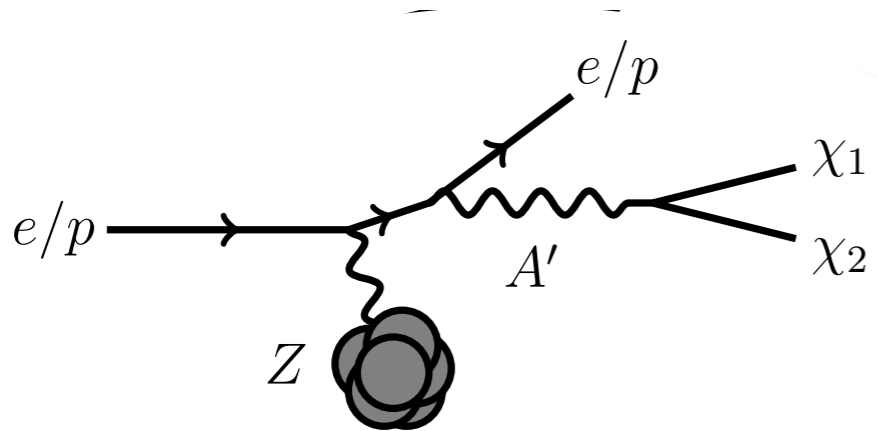
- DM interacts through weak (or weak scale) couplings
- Lee-Weinberg and Unitarity constrain mass range
 - ~ 1 GeV – ~ 10 TeV
- Usually consider a thermal relic



sub-keV DM

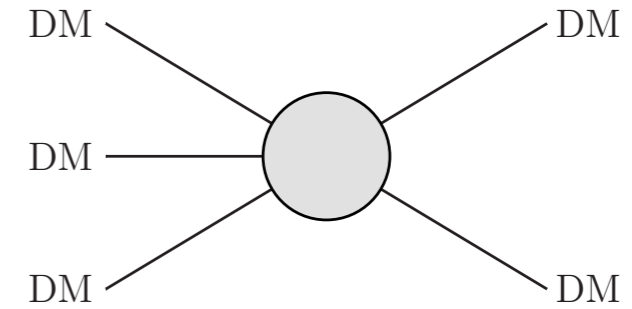
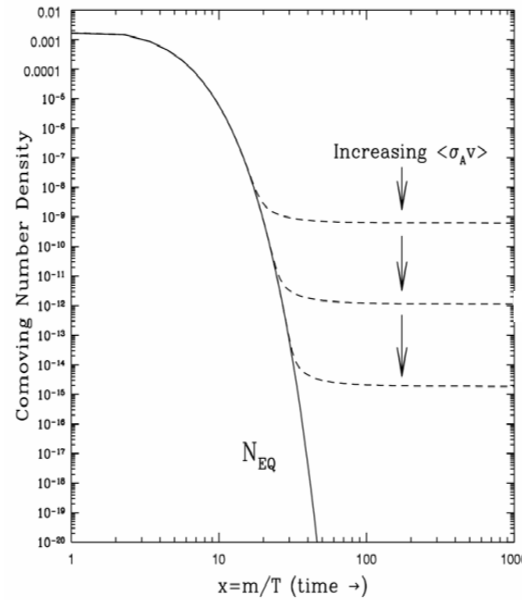
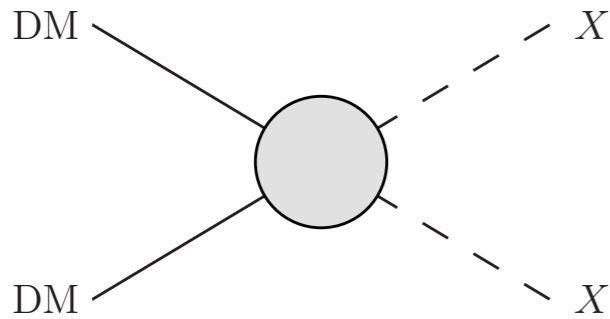
- Very light DM is bosonic
- Heavier than 10^{-22} eV
- More appropriately thought of as semiclassical wave, large n
- Or, absorption of DM, linear coupling to matter





Hidden sector DM—thermal relics

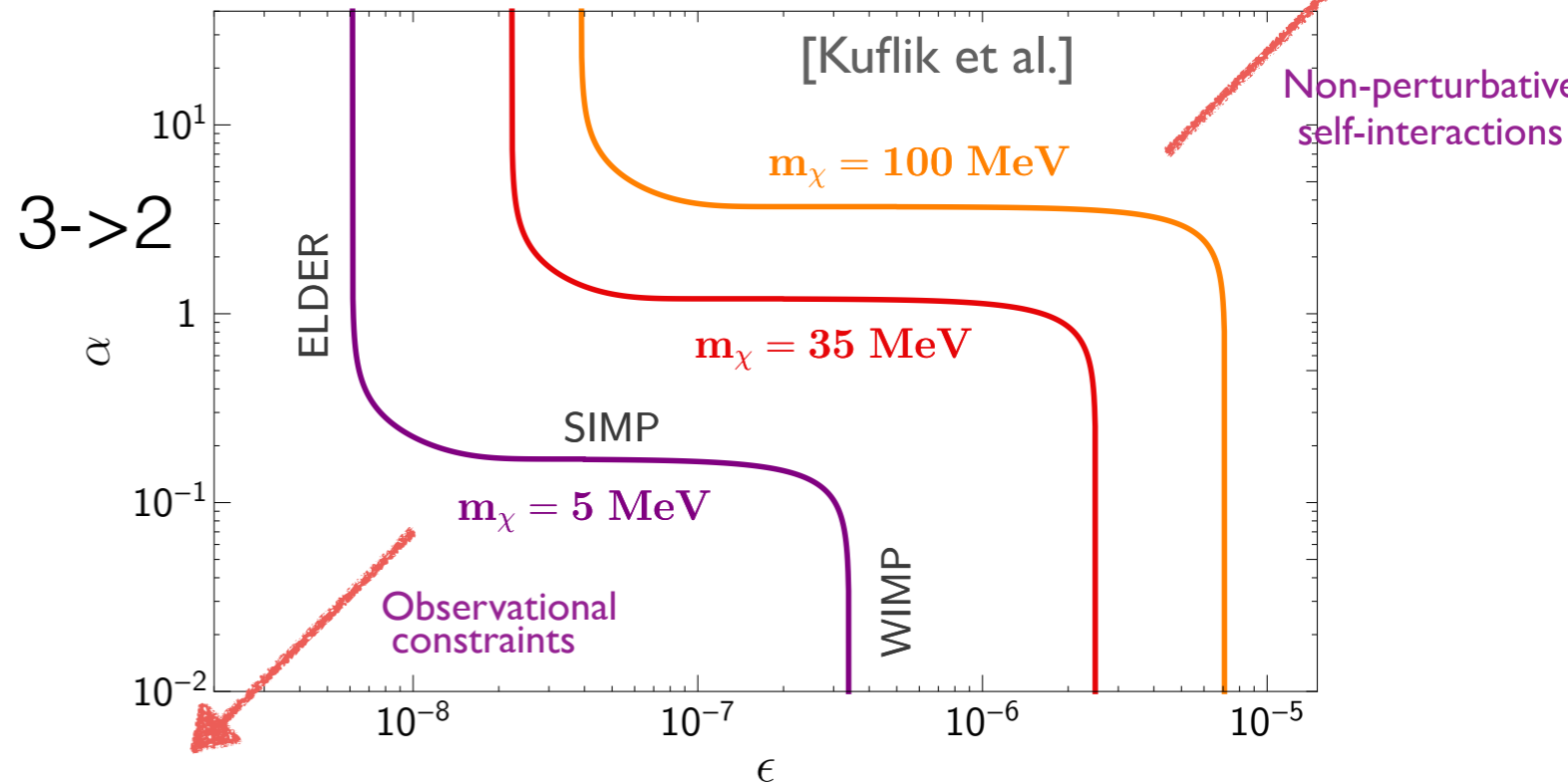
Leads to interesting changes in cosmology



cannibalization

WIMPless-miracle
(1-100 MeV scale)
SIMP-miracle
ELDER...

all smoothly connected
in parameter space



DM-SM elastic scatter

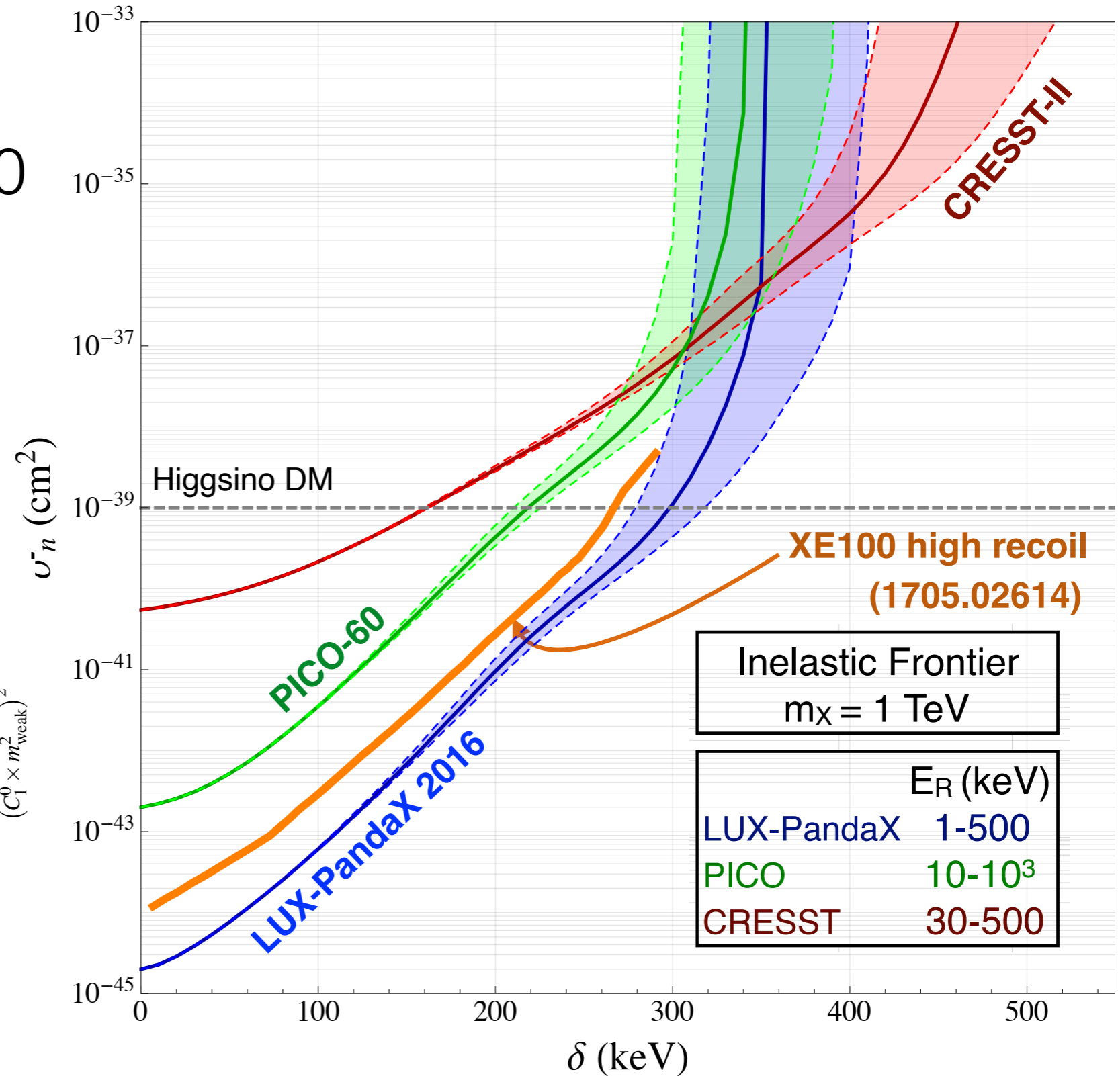
Crossing the Inelastic Frontier—Xenon 100

Reanalysis

7.6 ton-days, Xe100

6.6-240 keVnr

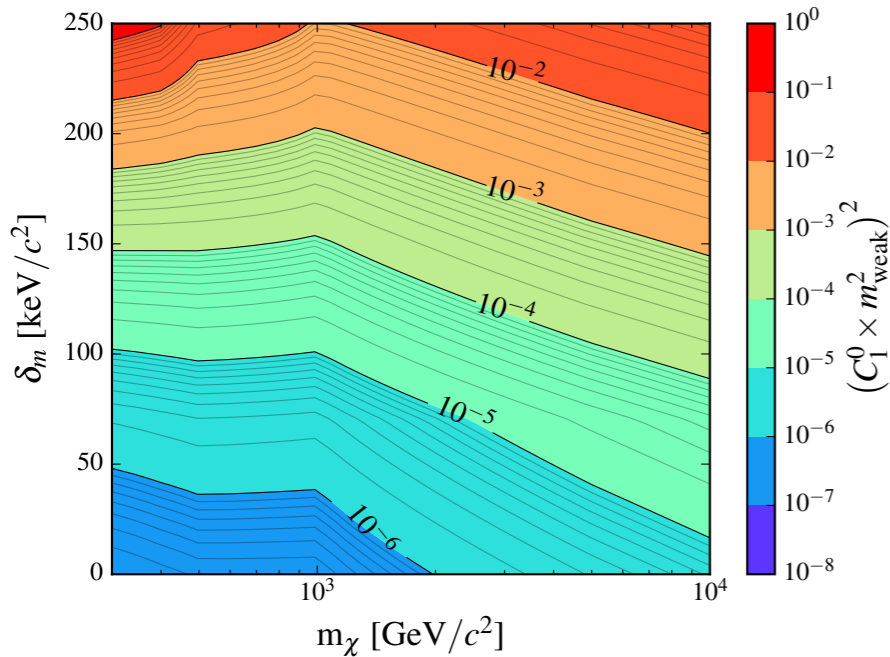
Inelastic analysis



**XE100 high recoil
(1705.02614)**

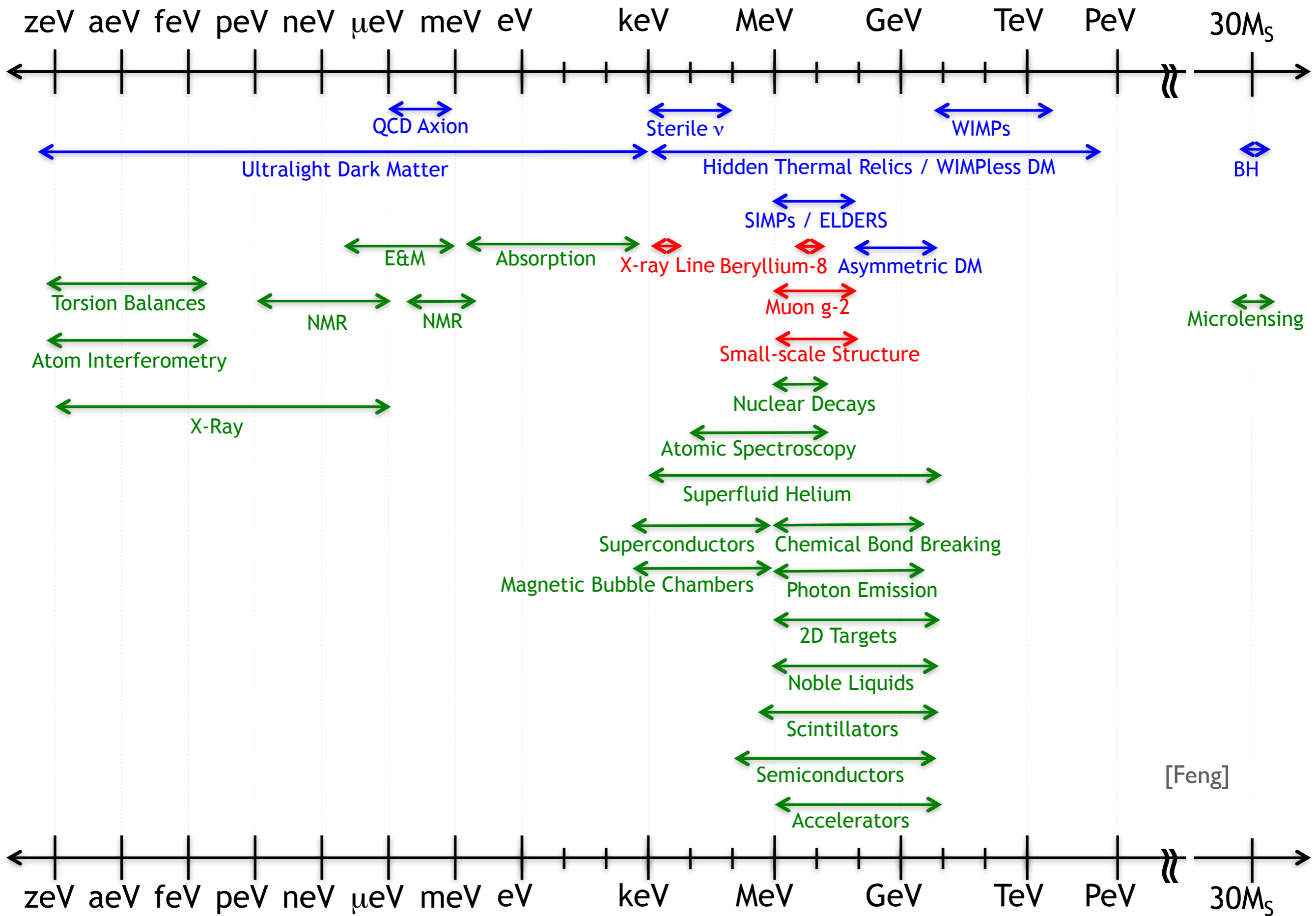
**Inelastic Frontier
 $m_\chi = 1 \text{ TeV}$**

	E_R (keV)
LUX-PandaX	1-500
PICO	10-10 ³
CRESST	30-500



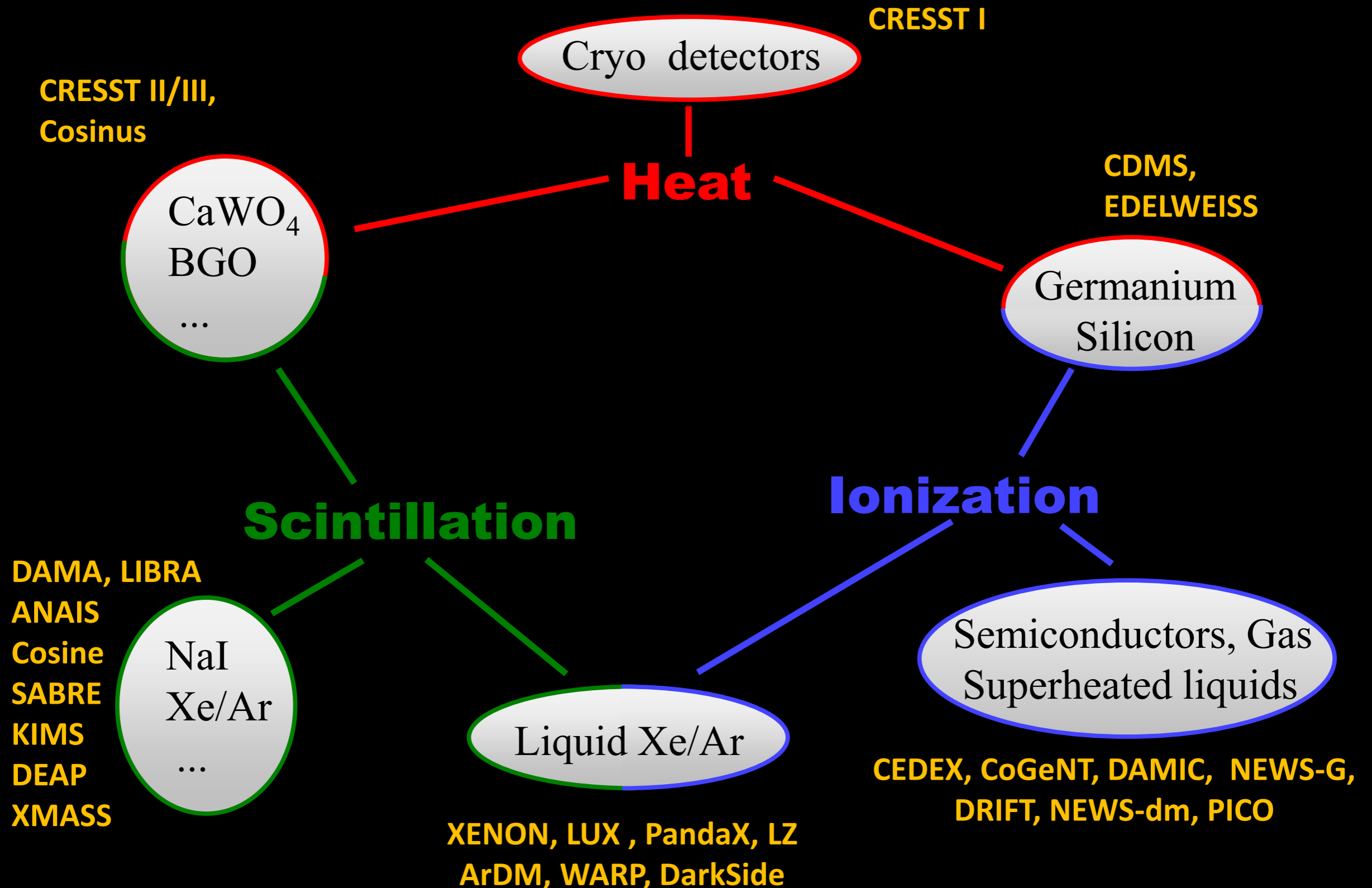
Dark Sector Candidates, Anomalies, and Search Techniques

Fox (USA)

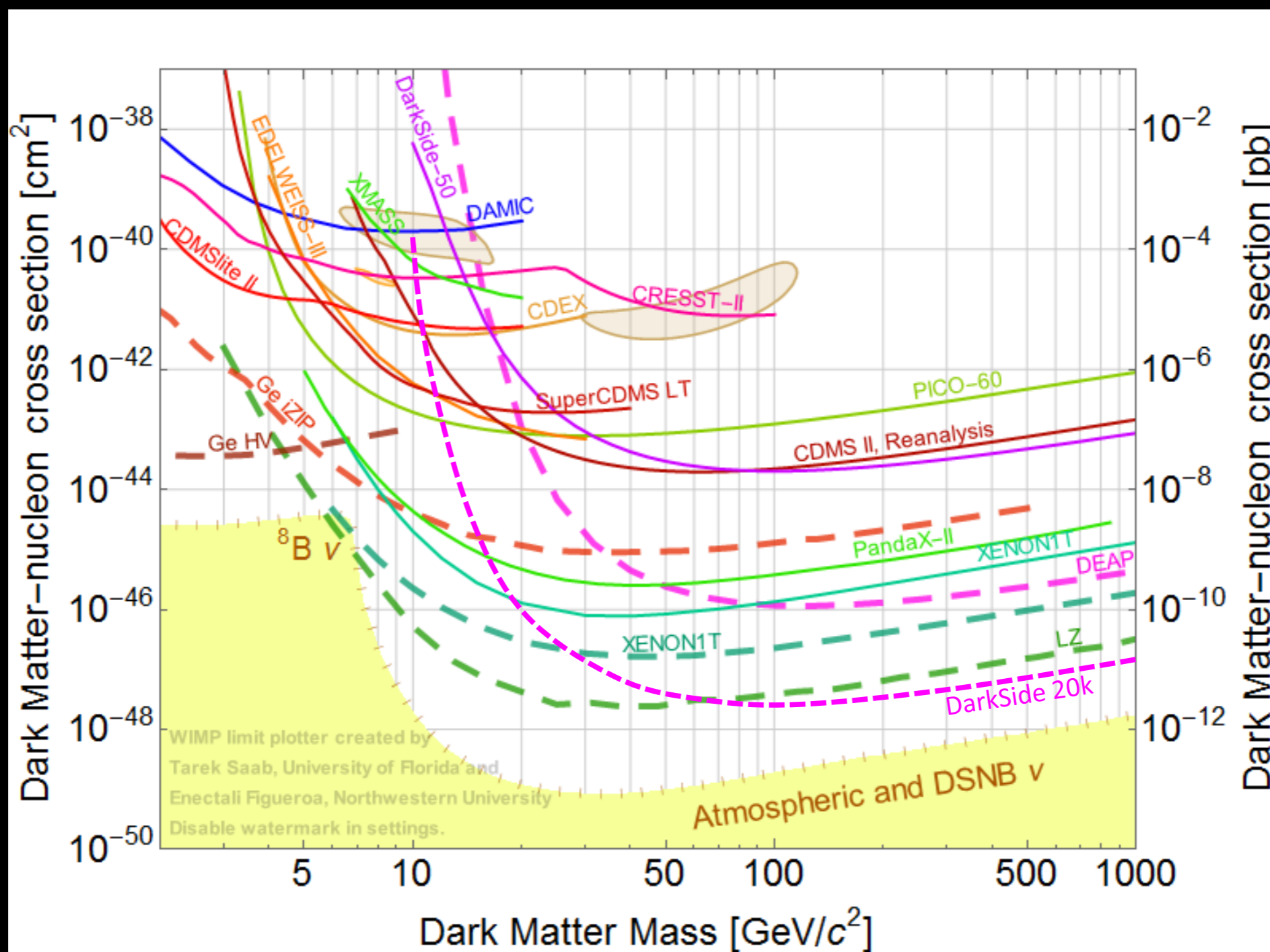


[Feng]

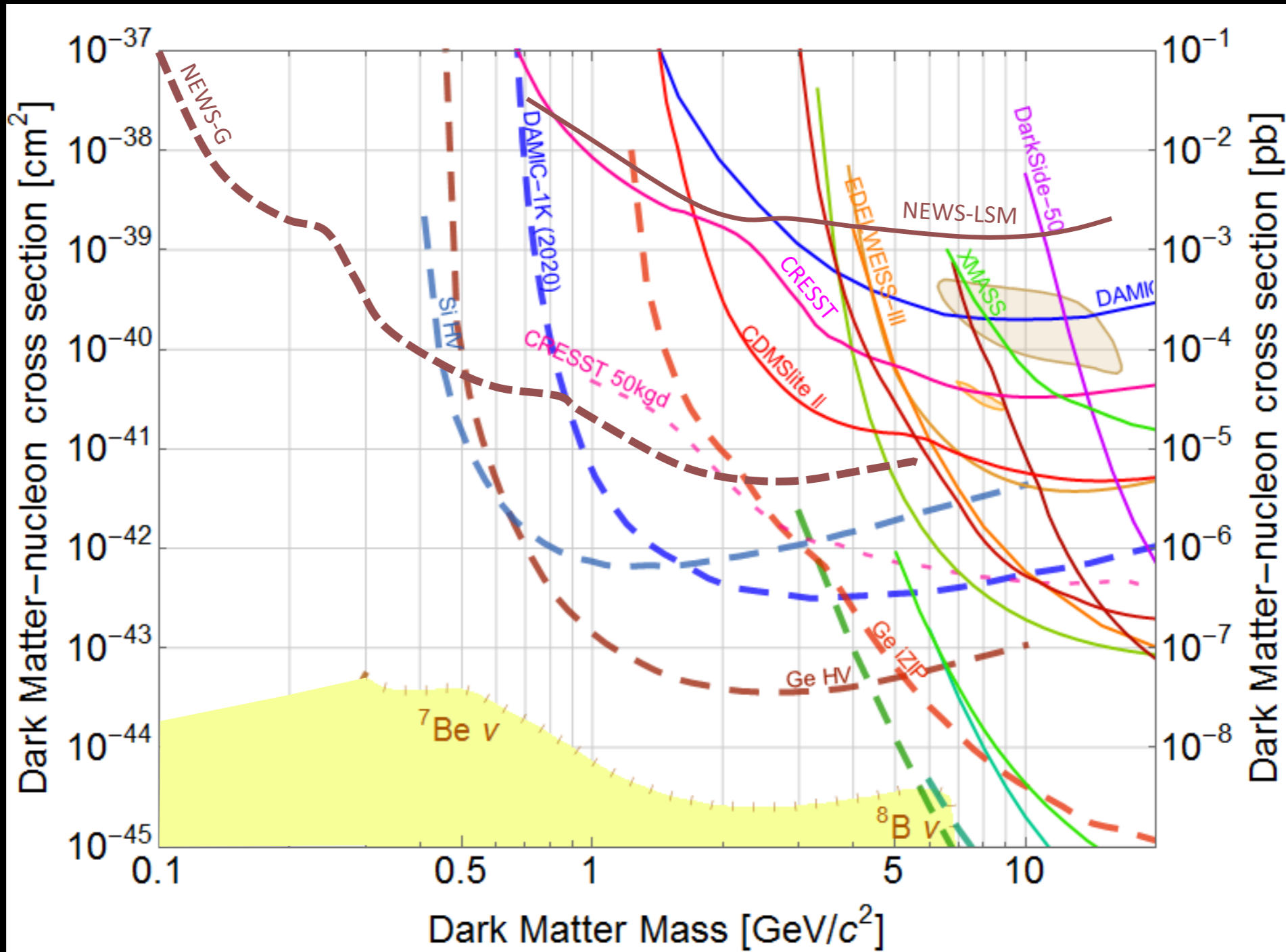
Direct Detection Triangle



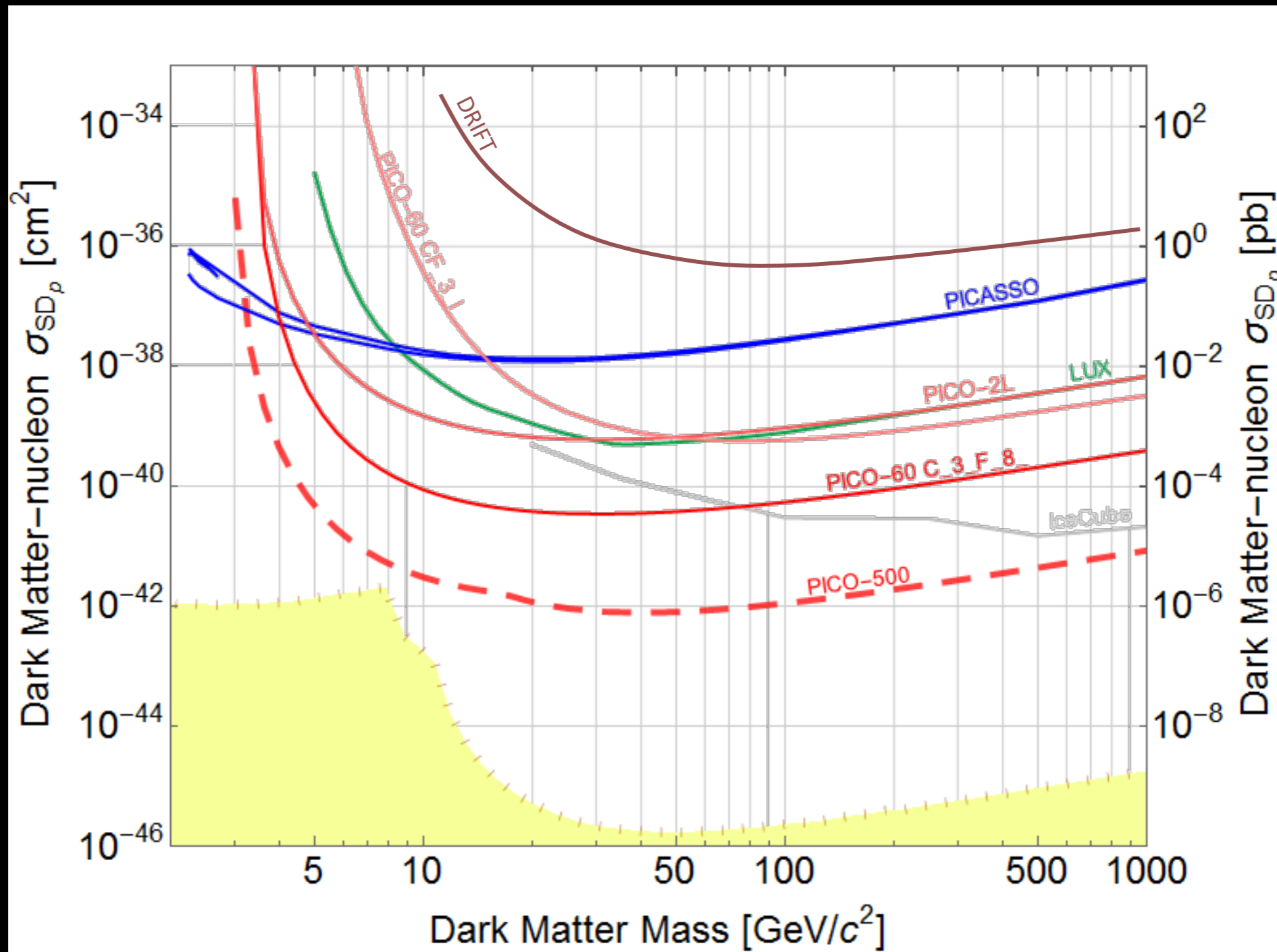
Results and Projections



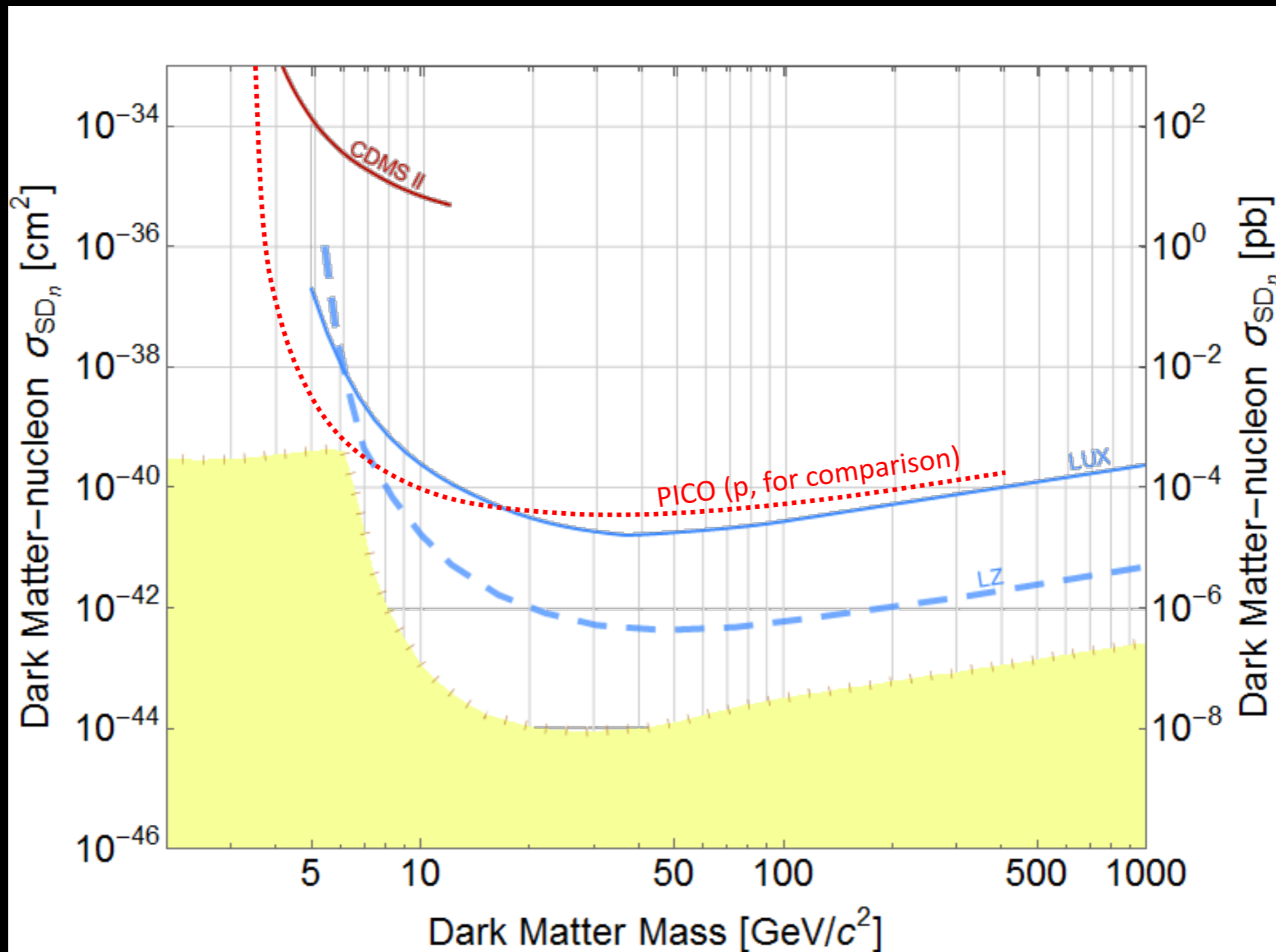
Results and Projections



Results and Projections



Results and Projections



After LHC

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Physics After LHC

Place

Location: The Hong Kong University of Science and Technology

Address: Hong Kong, China

Files

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Date: **28 Jun 11:00 - 14:30**

Description

Discussion Leader: Michelangelo Mangano (CERN, Switzerland)

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Wed 28/06

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11:00

The Unique Role of Future High-Energy pp Colliders

Andrea WULZER

The Hong Kong University of Science and Technology

11:05 - 11:25

The Unique Role of Future High-Energy ee Colliders

Tao HAN

The Hong Kong University of Science and Technology

11:30 - 11:50

12:00

The Technology Landscape of Future Accelerators

Lyn EVANS

The Hong Kong University of Science and Technology

11:55 - 12:20

13:00

Detectors for Future Collider Experiments

Lucie LINSSEN

The Hong Kong University of Science and Technology

12:55 - 13:15

Searches for Unconventional Signatures at Future Accelerators

David CURTIN

The Hong Kong University of Science and Technology

13:20 - 13:40

14:00

Future Experiments to Complement High-E Colliders in the Search and Study of BSM

Eder IZAGUIRRE

The Hong Kong University of Science and Technology

13:45 - 14:05

Ideology

HEP before the LHC



HEP before the F.C.



Particle physics is not **validation** anymore, rather it is **exploration of unknown territories** *

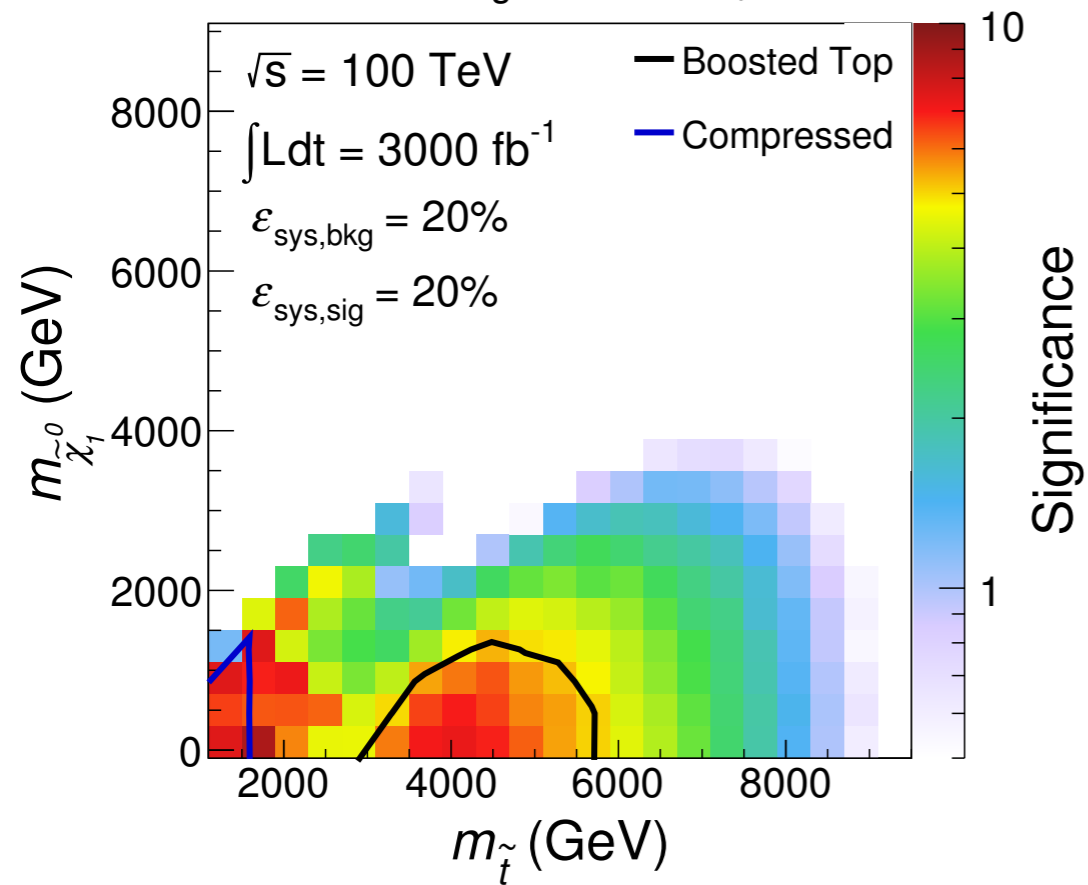
* Not necessarily a bad thing. Columbus left for his trip just because he had no idea of where he was going !!

Naturalness

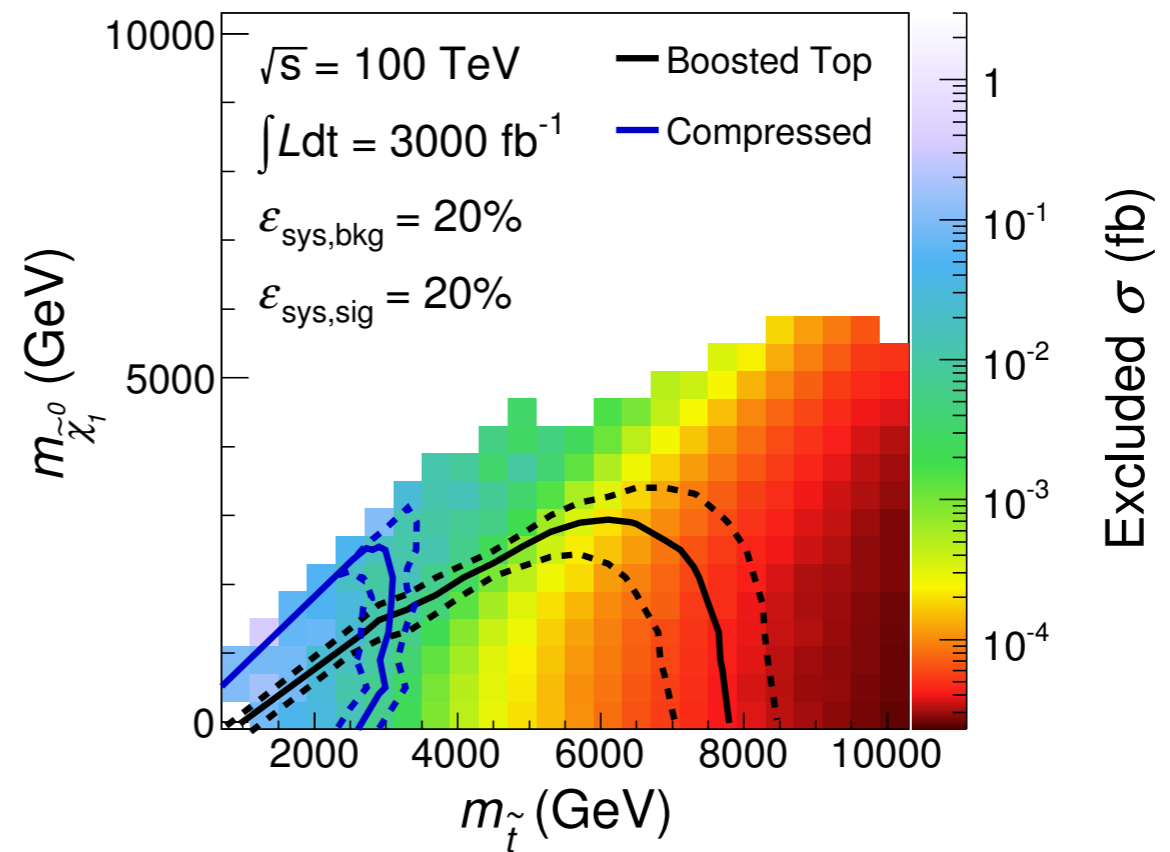
Stop to top + Neutralino:

[arXiv:1406.4512]

CL_s Discovery



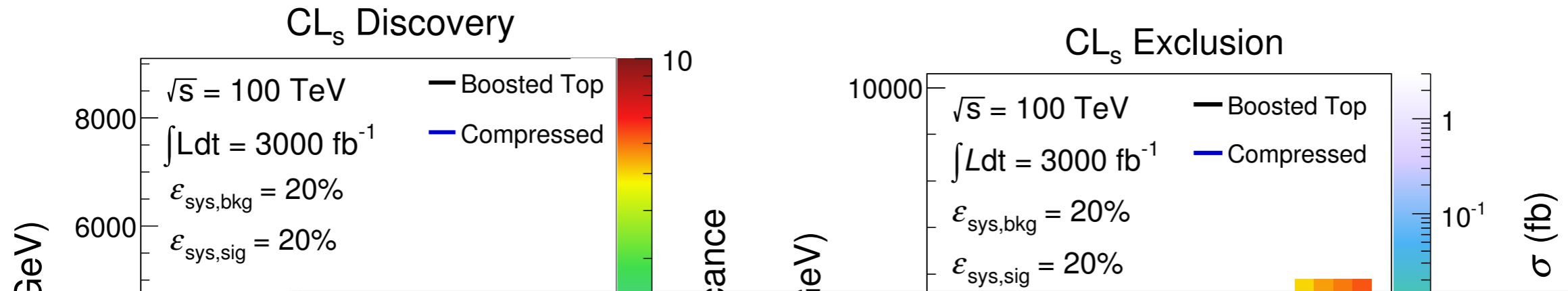
CL_s Exclusion



Naturalness

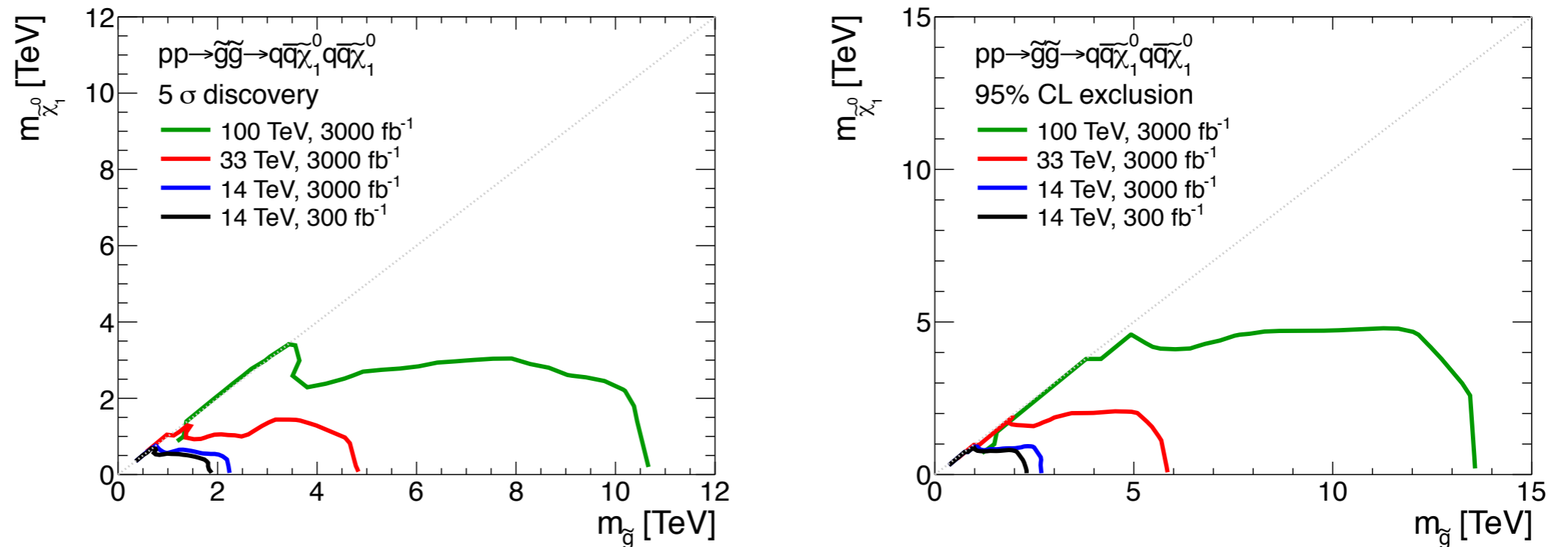
Stop to top + Neutralino:

[arXiv:1406.4512]



Glucino to Neutralino + q:

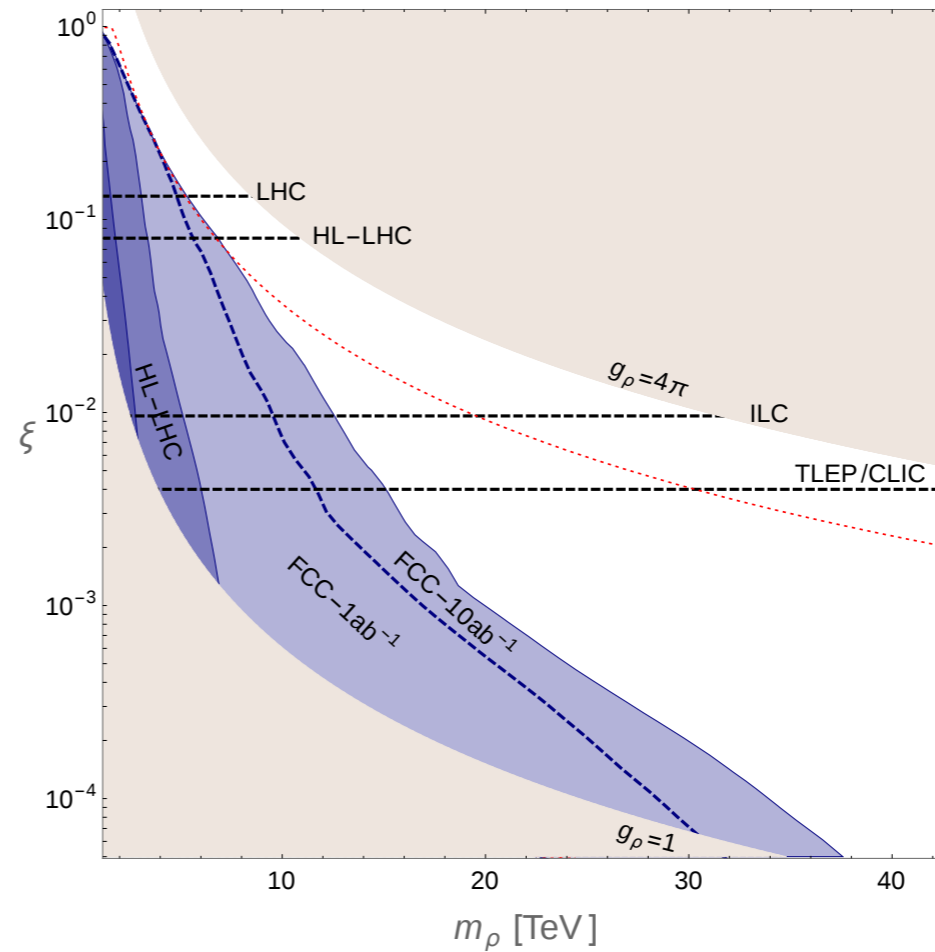
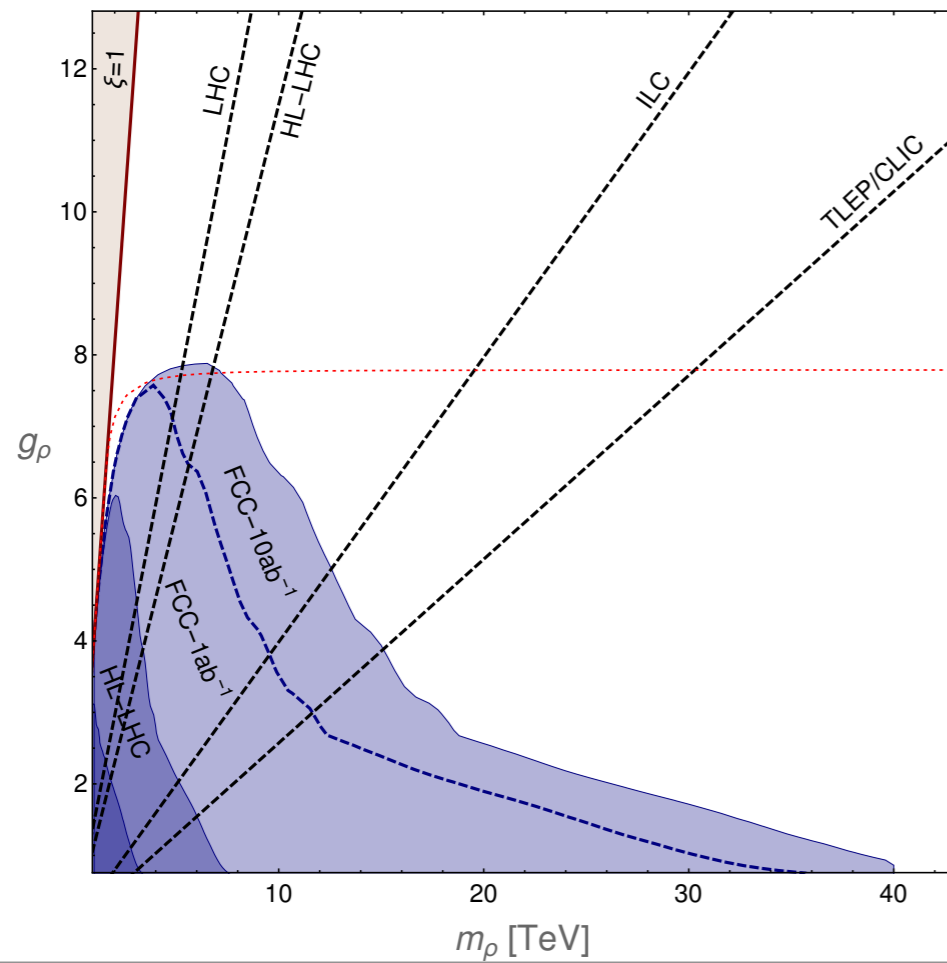
[arXiv:1606.00947]



Naturalness

Spin-1 CH resonances (or V'):

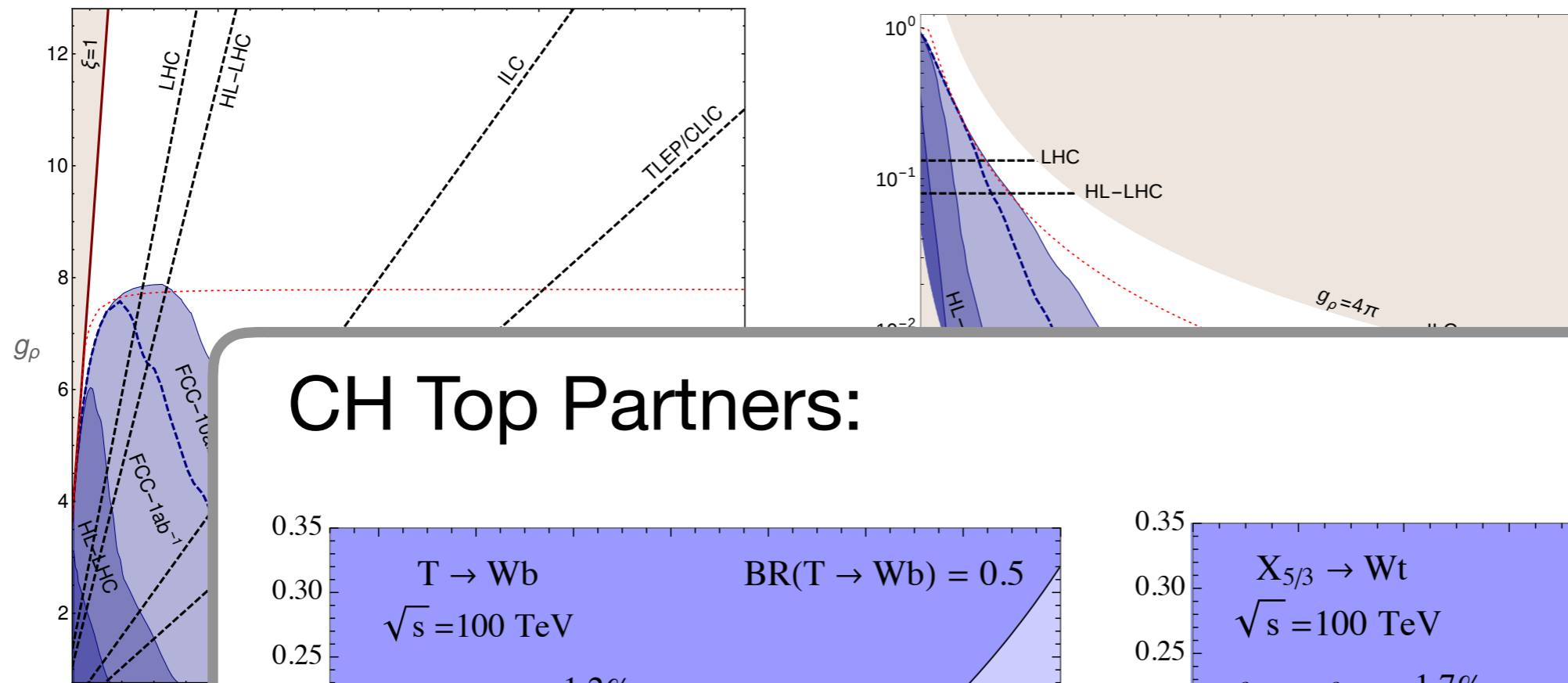
[arXiv:1502.01701]



Naturalness

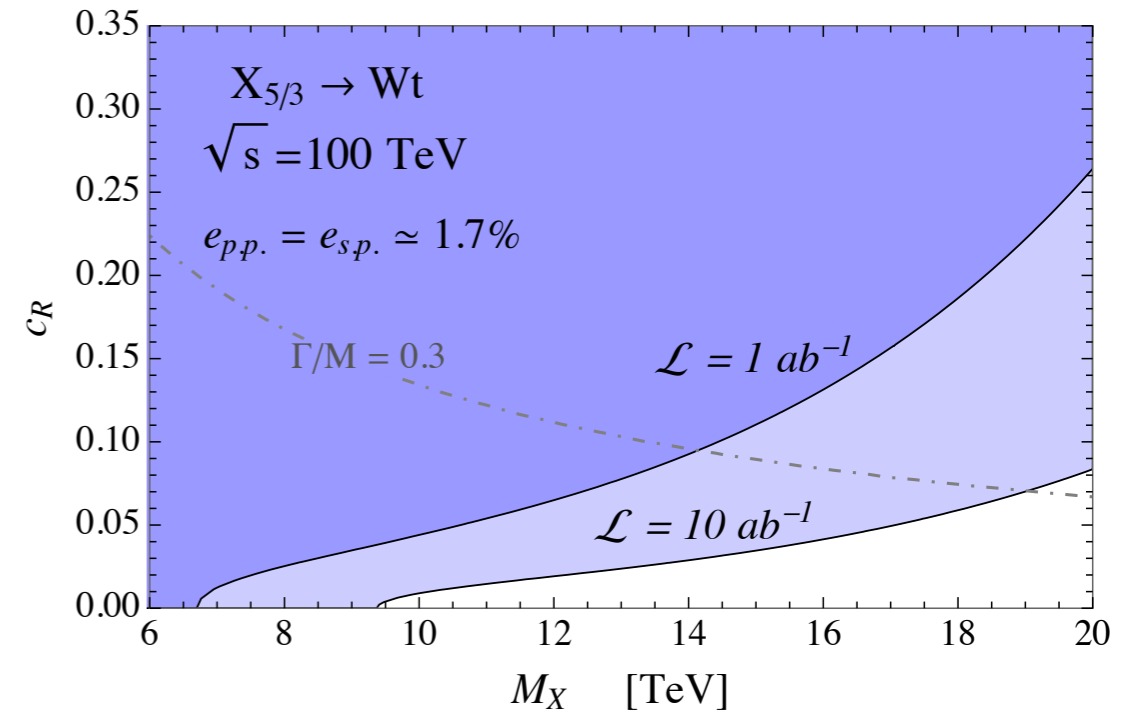
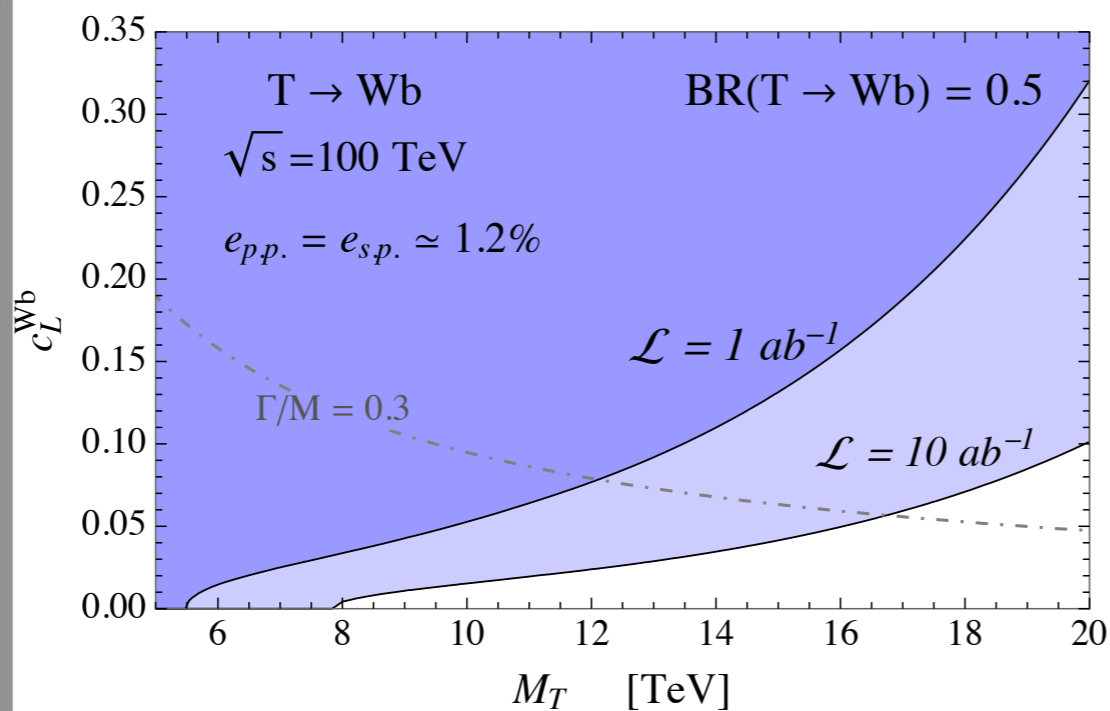
Spin-1 CH resonances (or V'):

[arXiv:1502.01701]



CH Top Partners:

[arXiv:1409.0100]

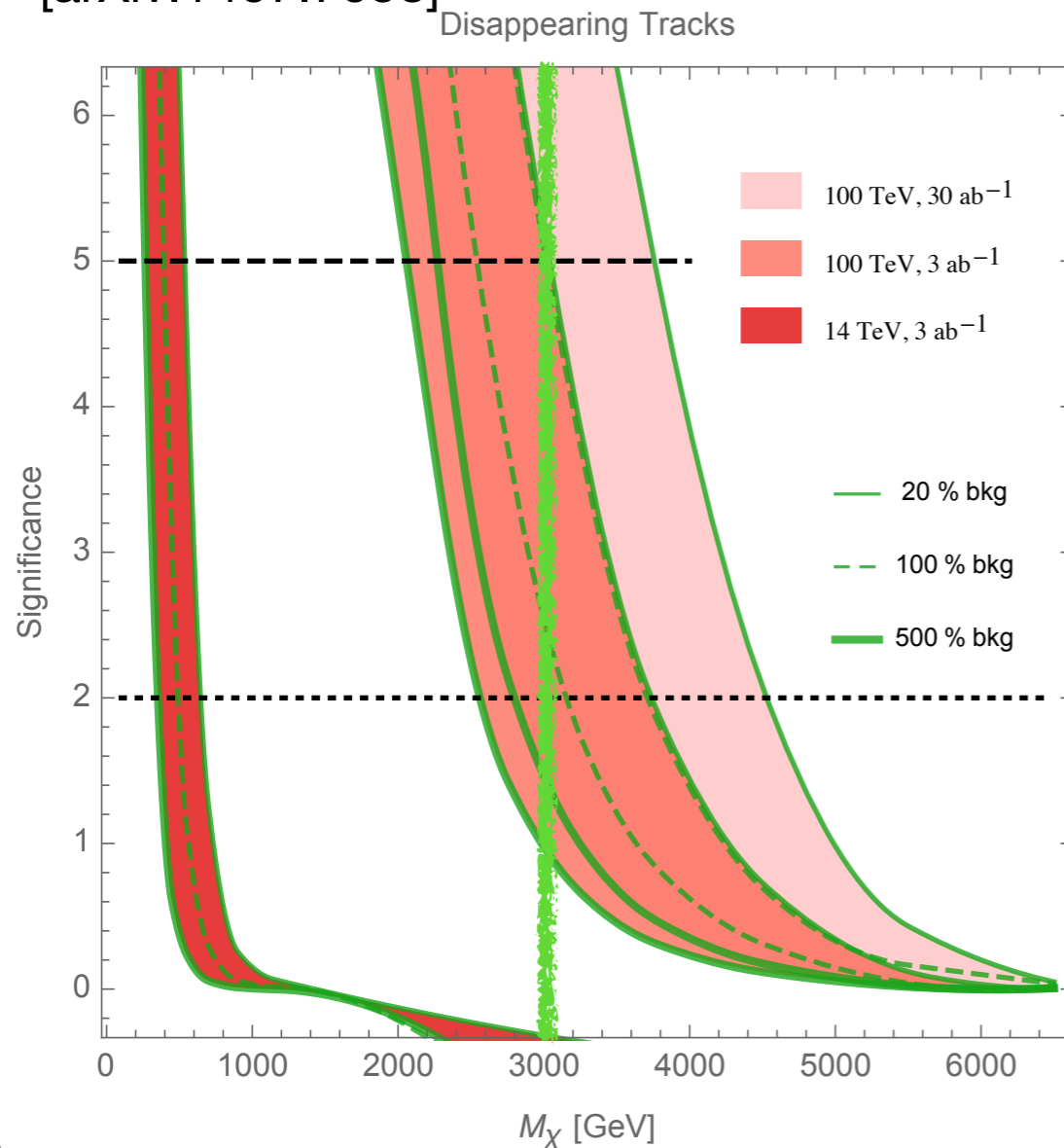


Dark Matter

100 TeV can probe WIMP DM inaccessible to current and future Direct Detection

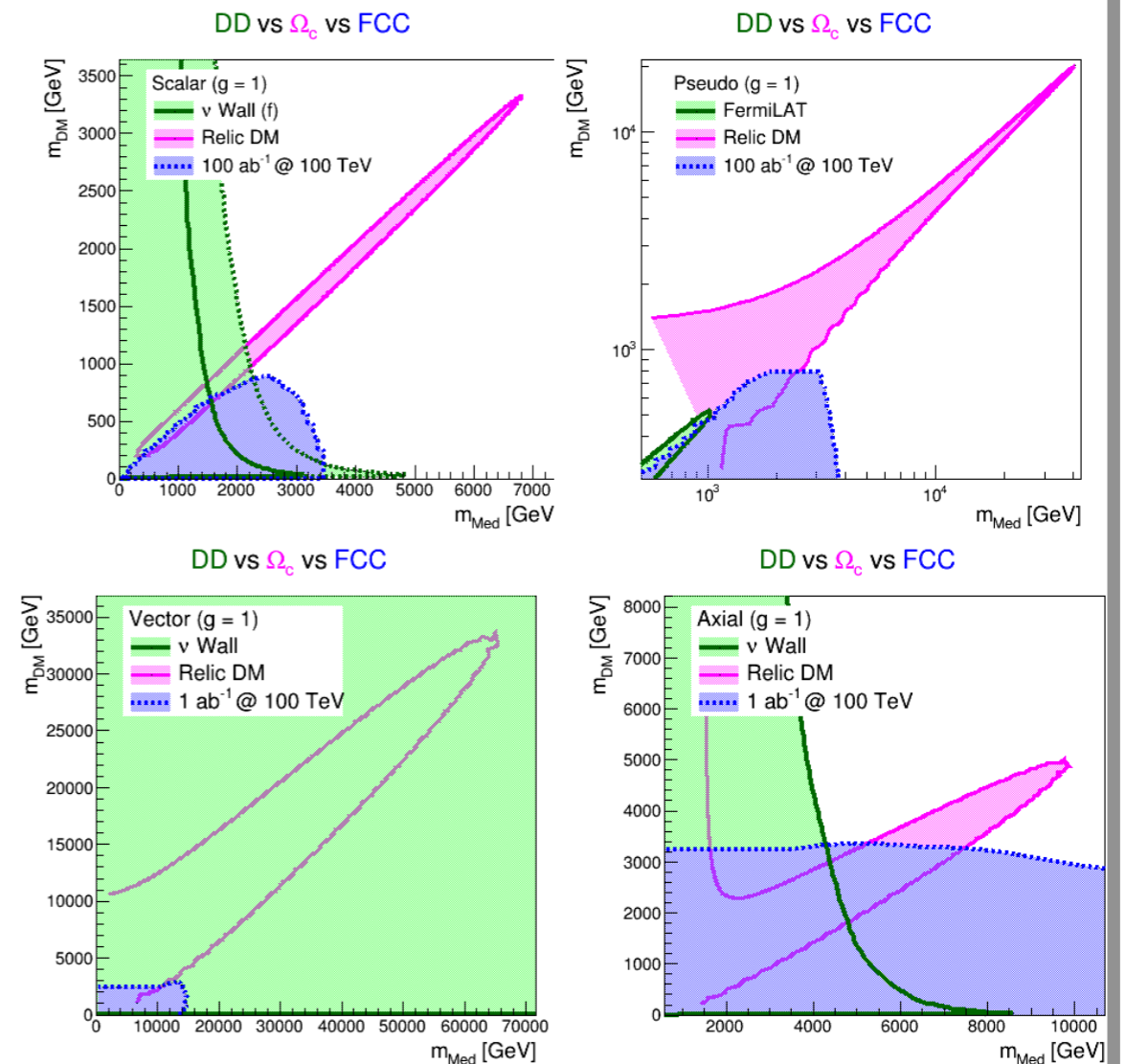
Wino DM. (disappearing tracks:)

[arXiv:1407.7058]



Crossing neutrino wall:

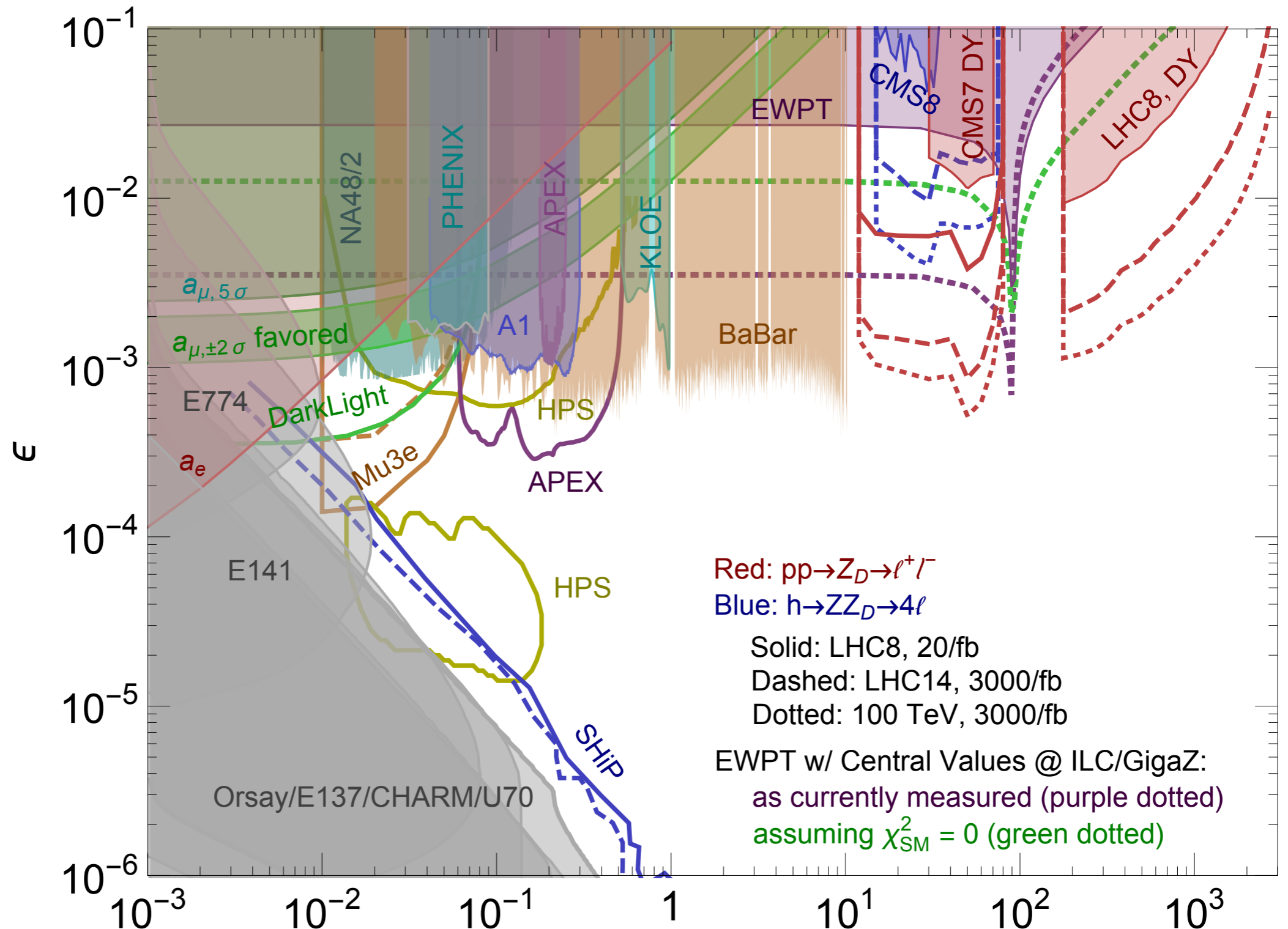
[arXiv:1606.00947]



Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...

Dark photon:
[arXiv:1412.0018]



Energy and Accuracy Frontier @ 100 TeV

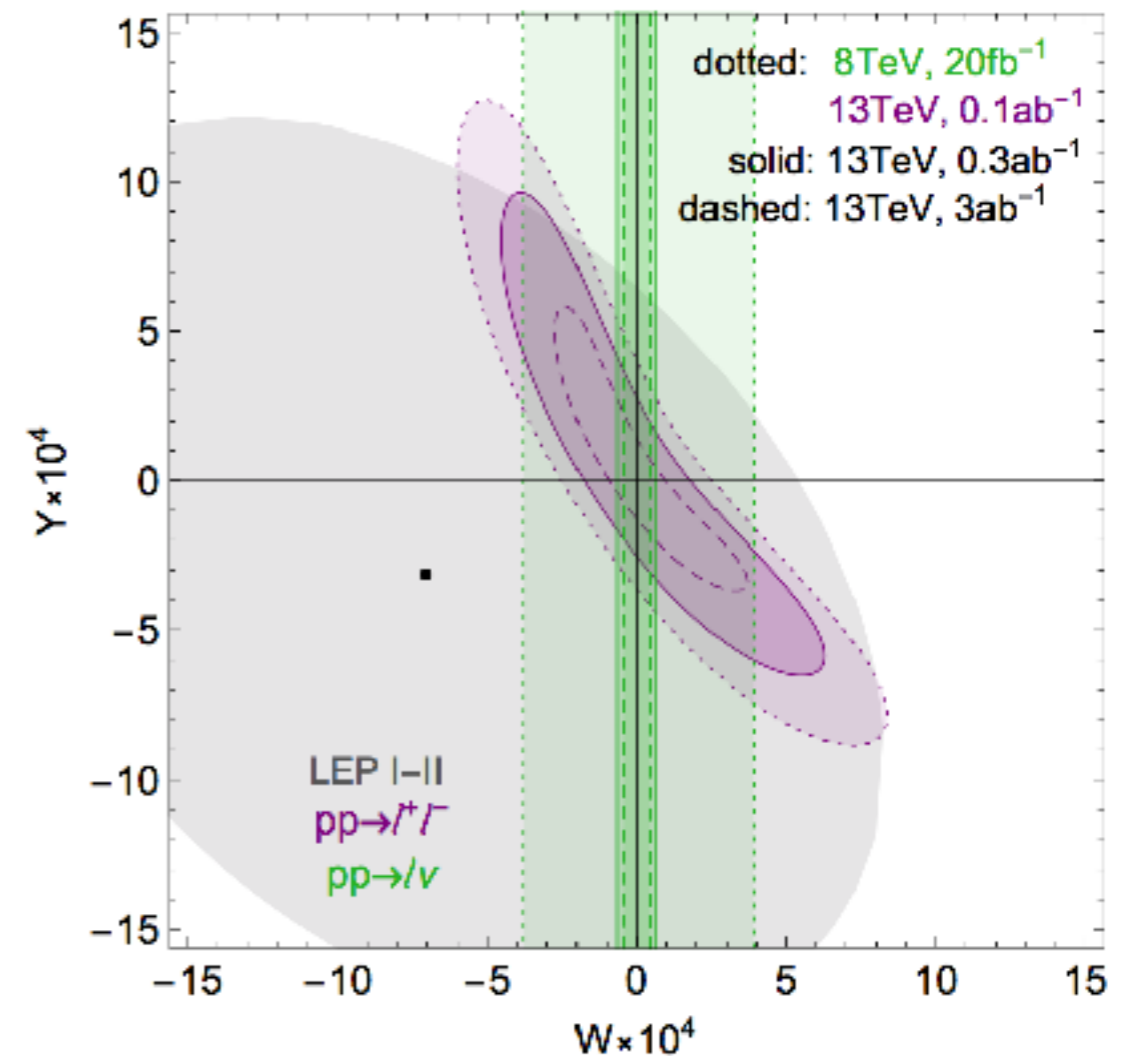
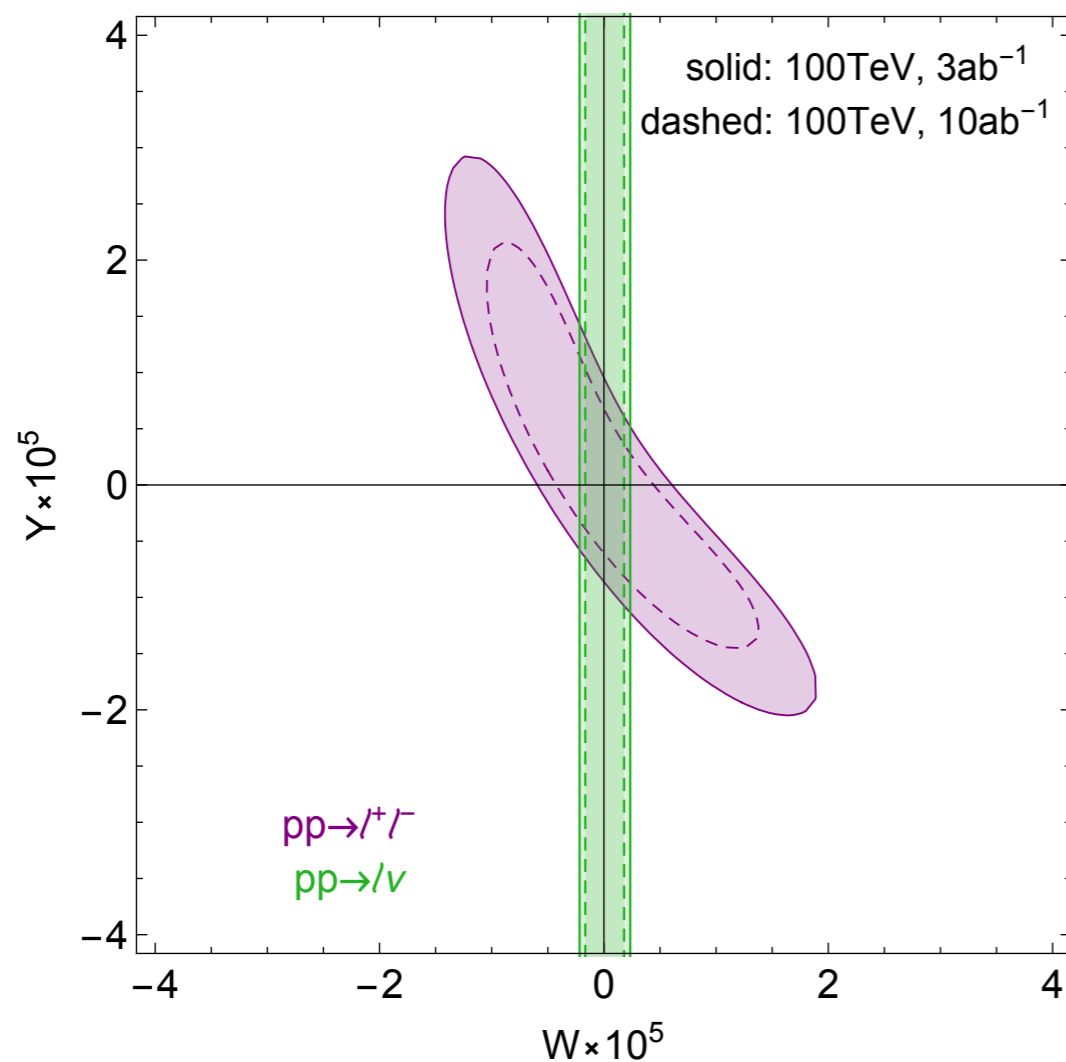
Enhanced indirect NP effects in high mass tails

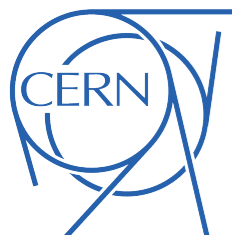


No need of extreme accuracy for indirect NP probe

EWPT @ hadron colliders: (W and Y oblique par.s)

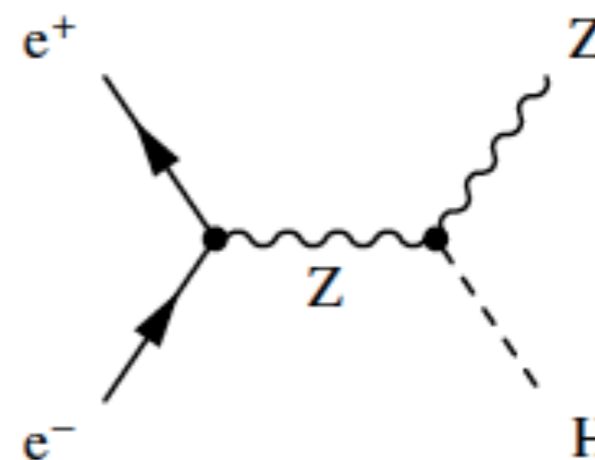
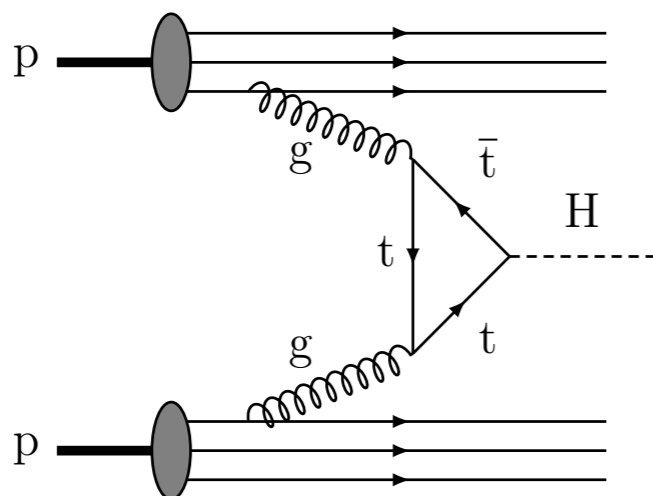
[arXiv:1609.08157]



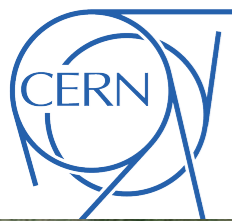


pp collisions / e^+e^- collisions

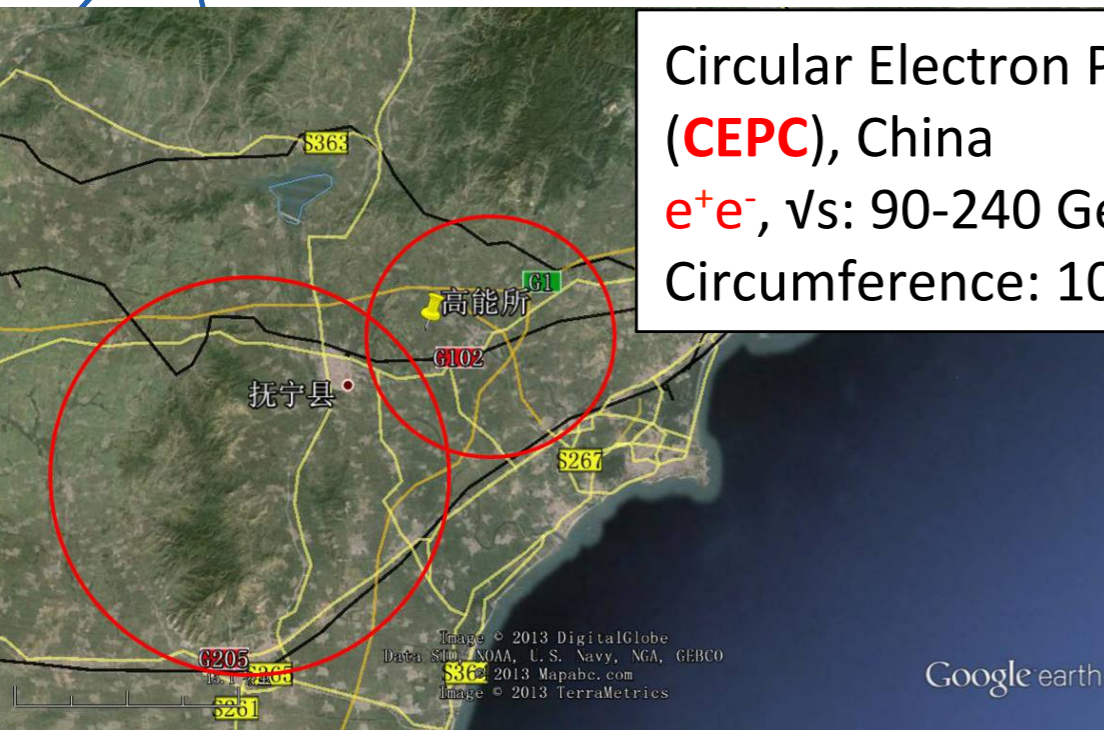
to tackle the open questions in particle physics



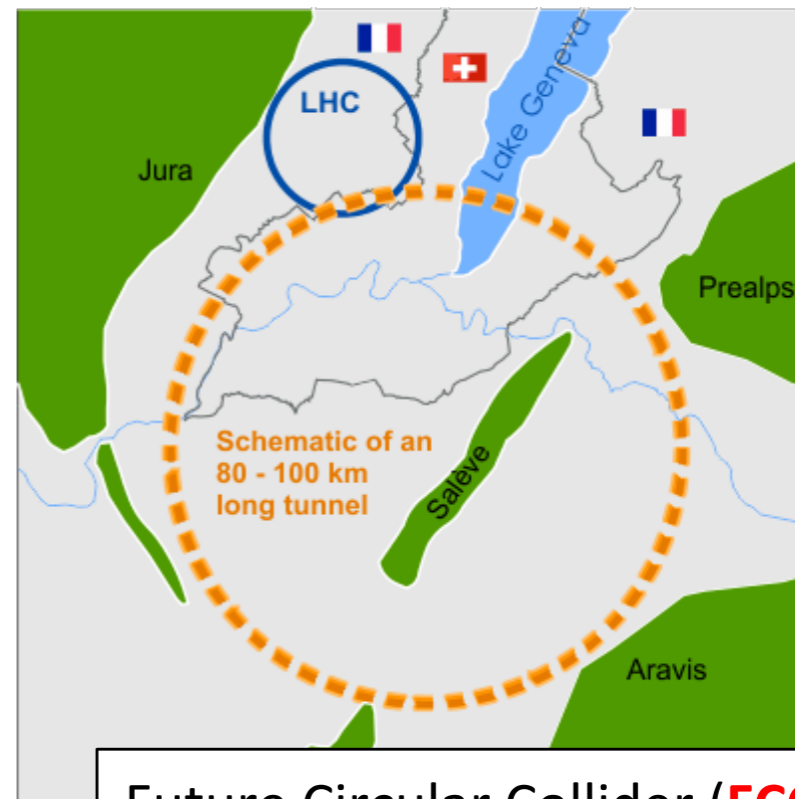
p-p collisions	e^+e^- collisions
<p>Proton is compound object</p> <ul style="list-style-type: none"> → Initial state unknown → Limits achievable precision 	<p>e^+/e^- are point-like</p> <ul style="list-style-type: none"> → Initial state well defined (\sqrt{s} / opt: polarisation) → High-precision measurements ←
<p>High rates of QCD backgrounds</p> <ul style="list-style-type: none"> → Complex triggering schemes ← → High levels of radiation ← 	<p>Cleaner experimental environment</p> <ul style="list-style-type: none"> → Less / no need for triggers → Lower radiation levels
High cross-sections for colored-states	Superior sensitivity for electro-weak states
Very high-energy circular pp colliders feasible	High energies ($>\approx 350$ GeV) require linear collider



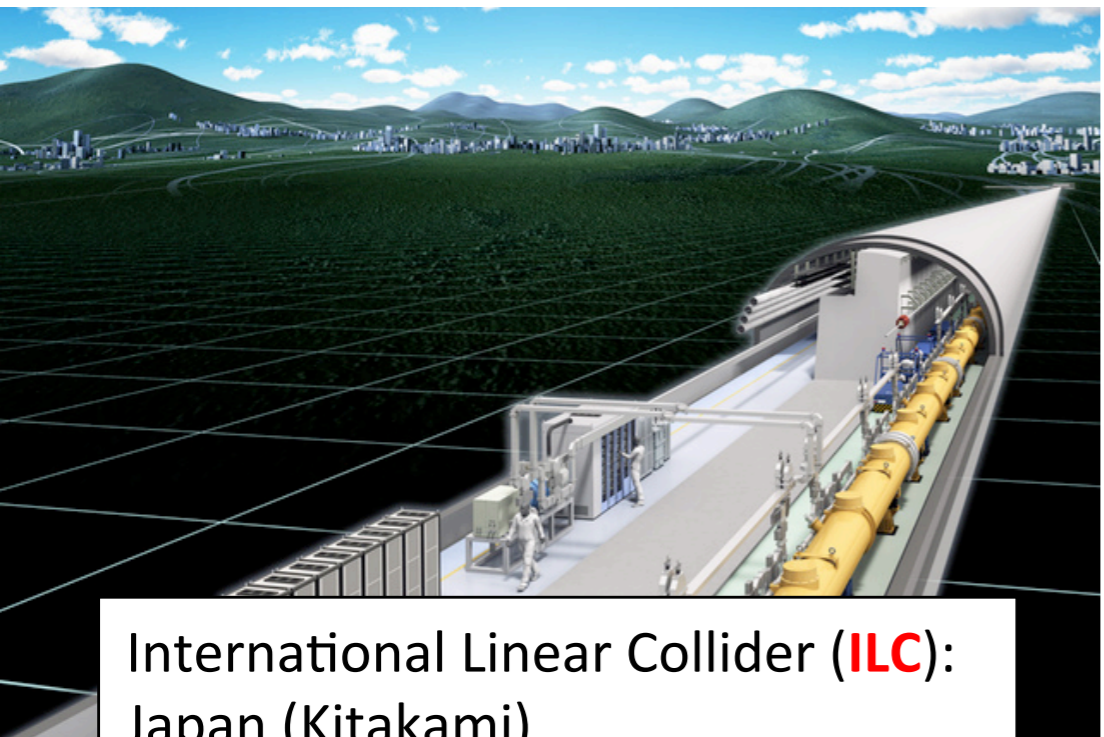
studies of high-energy e^+e^- colliders



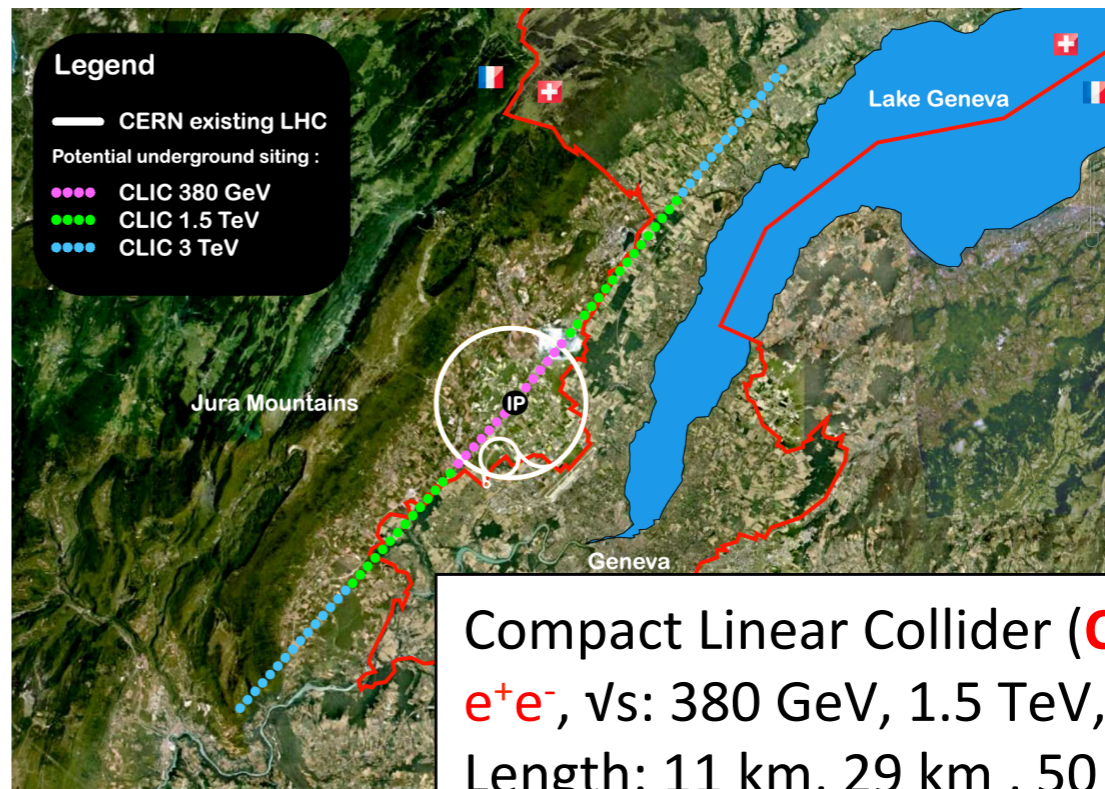
Circular Electron Positron Collider (**CEPC**), China
 e^+e^- , \sqrt{s} : 90-240 GeV; SPPC pp,
Circumference: 100 km



Future Circular Collider (**FCC-ee**): CERN
 e^+e^- , \sqrt{s} : 90 - 350 GeV; FCC-hh pp
Circumference: 97.75 km



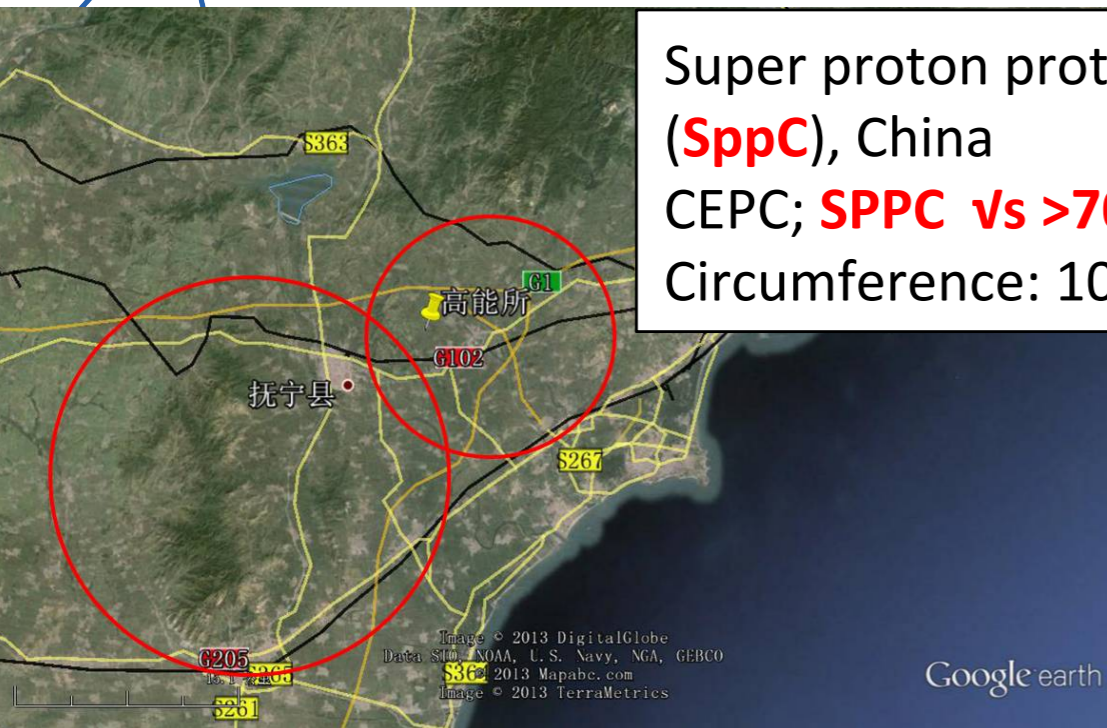
International Linear Collider (**ILC**):
Japan (Kitakami)
 e^+e^- , \sqrt{s} : 250 – 500 GeV (1 TeV)
Length: 17 km, 31 km (50 km)



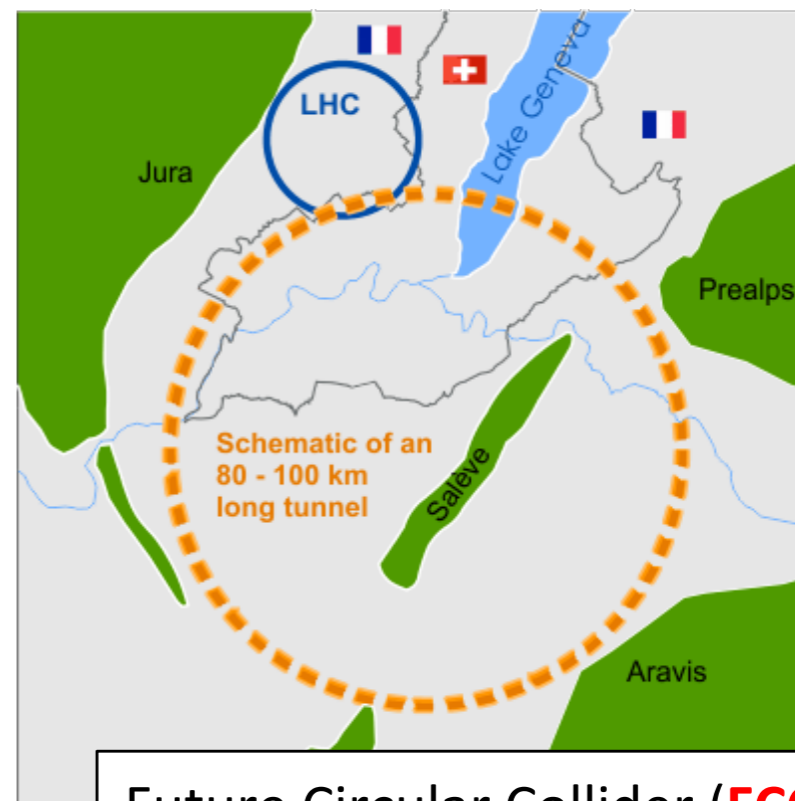
Compact Linear Collider (**CLIC**): CERN
 e^+e^- , \sqrt{s} : 380 GeV, 1.5 TeV, 3 TeV
Length: 11 km, 29 km, 50 km



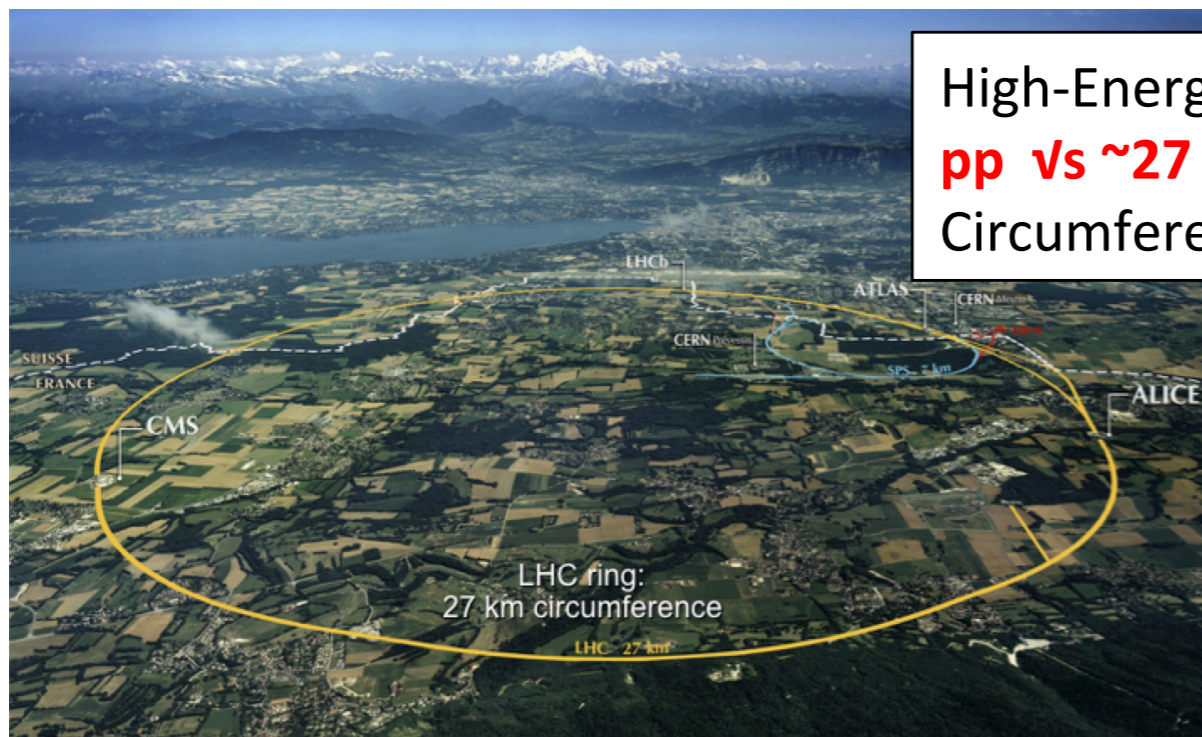
studies of high-energy pp colliders



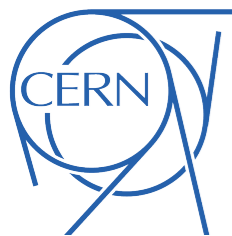
Super proton proton Collider (**SppC**), China
CEPC; **SPPC vs >70 TeV**
Circumference: 100 km



Future Circular Collider (**FCC-hh**): CERN
FCC-ee; **FCC-hh vs ~100 TeV**
Circumference: 97.75 km

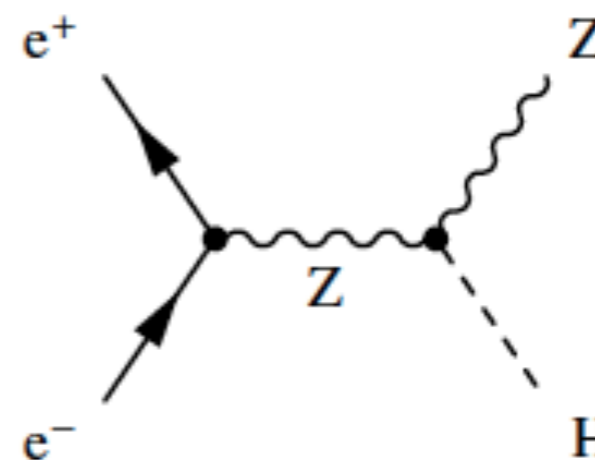
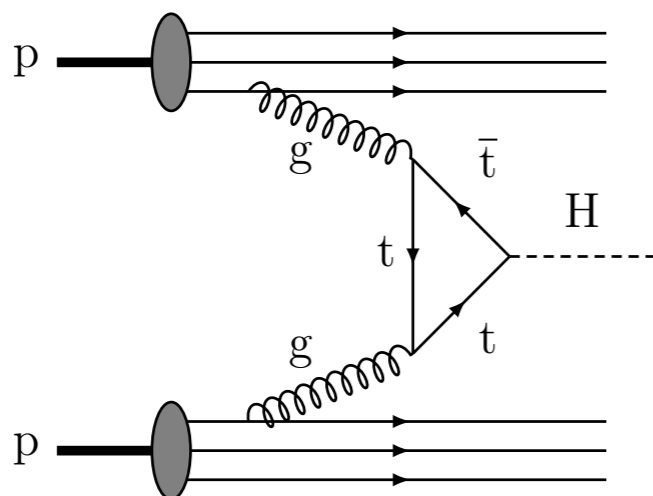


High-Energy LHC (**HE-LHC**): CERN
pp vs ~27 TeV
Circumference: 27 km

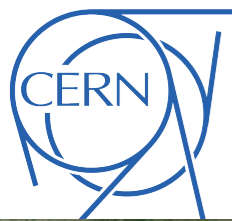


pp collisions / e^+e^- collisions

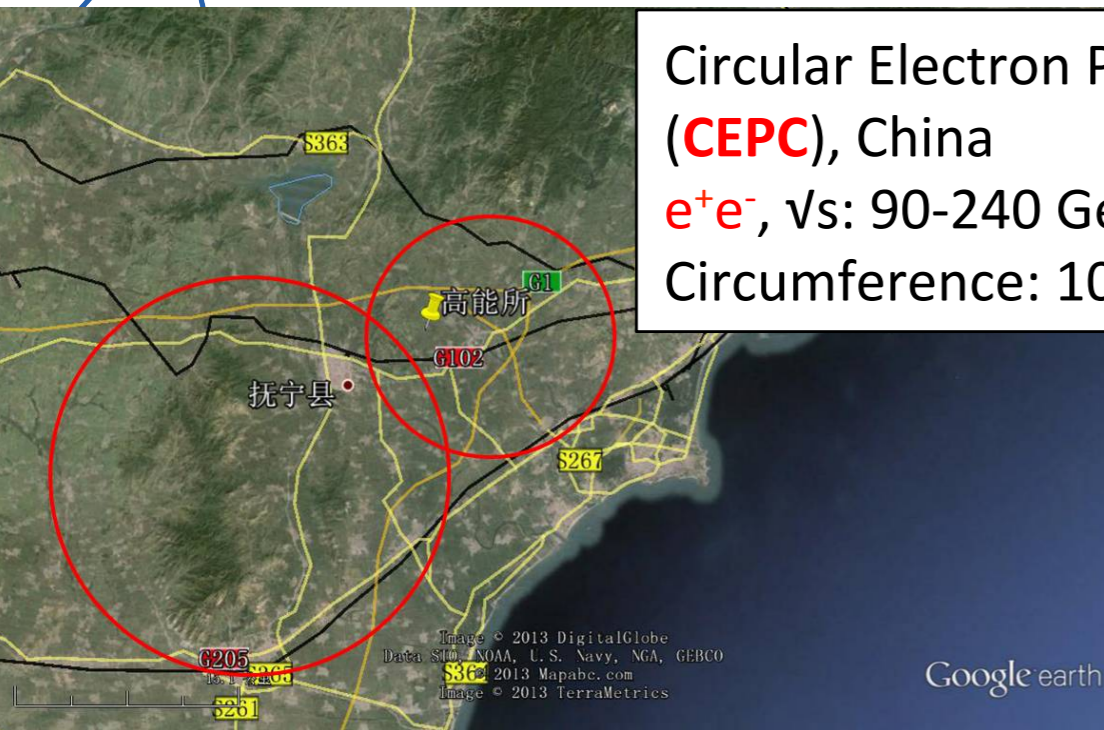
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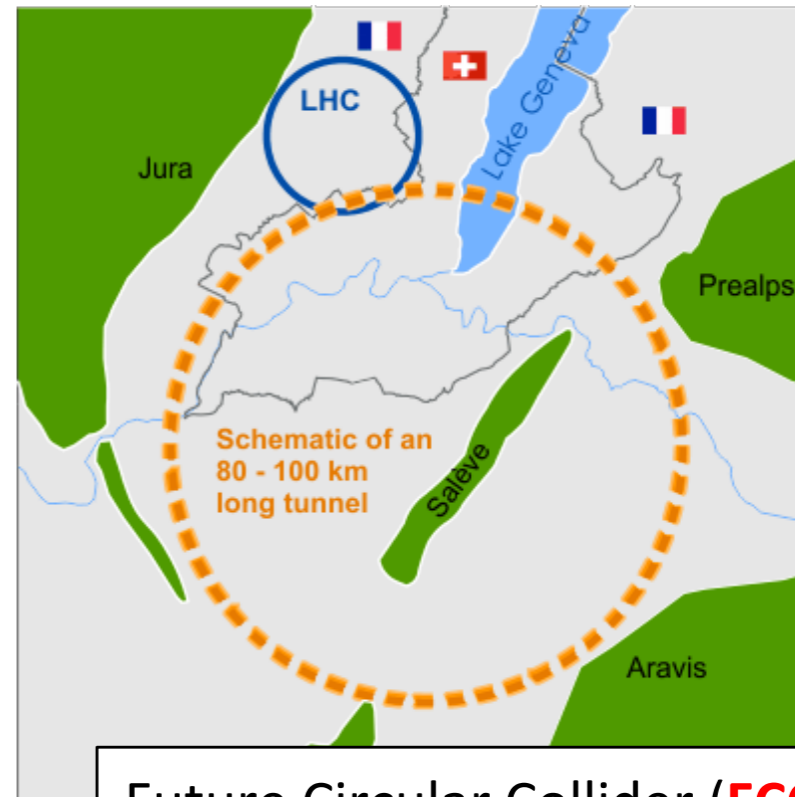
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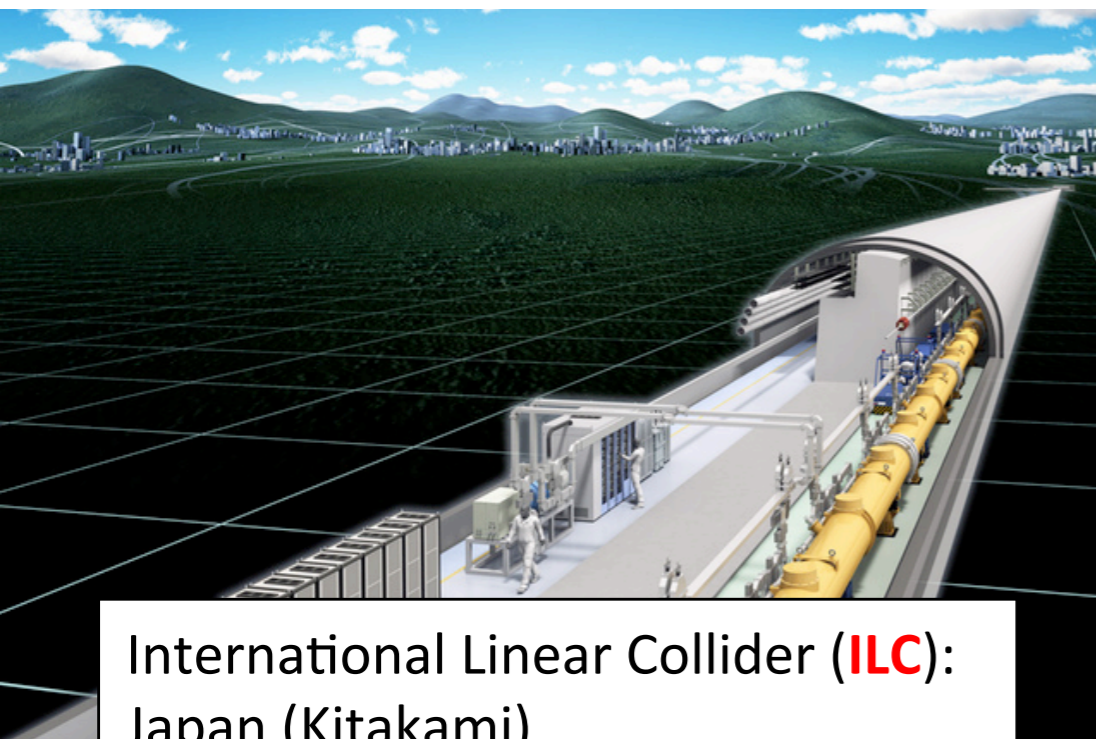
studies of high-energy e^+e^- colliders



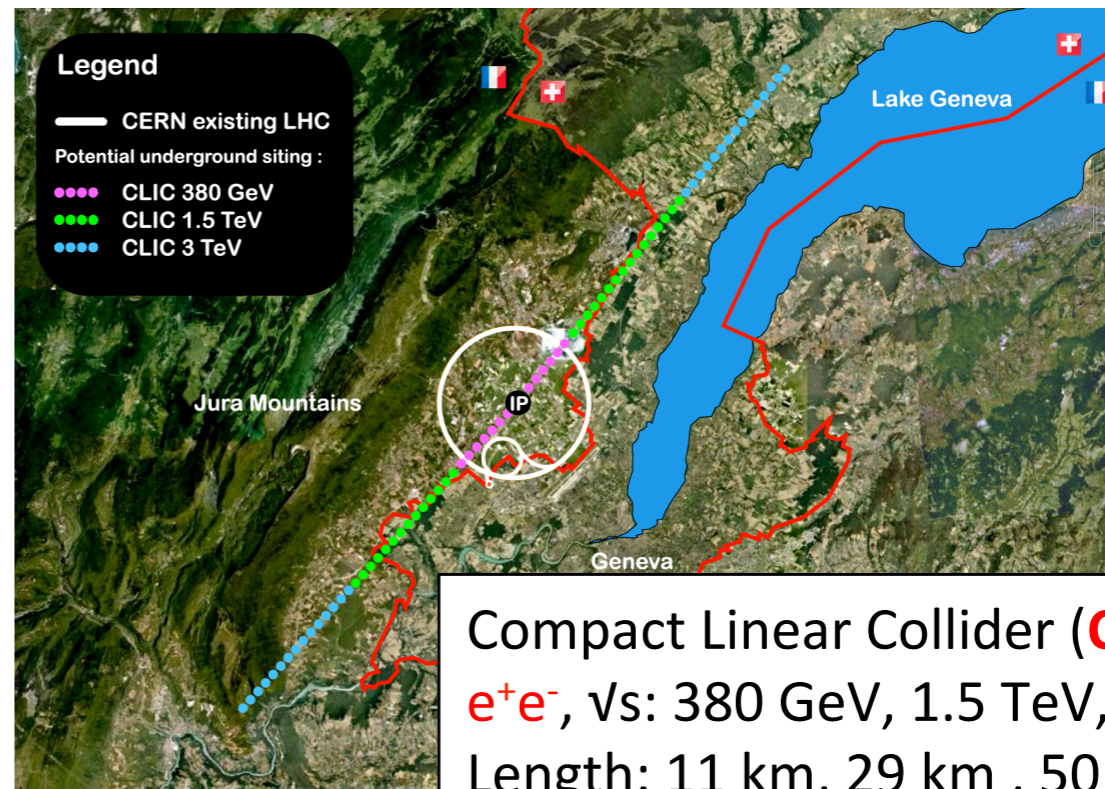
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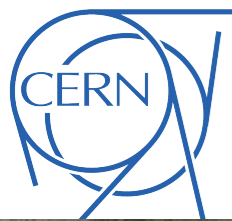
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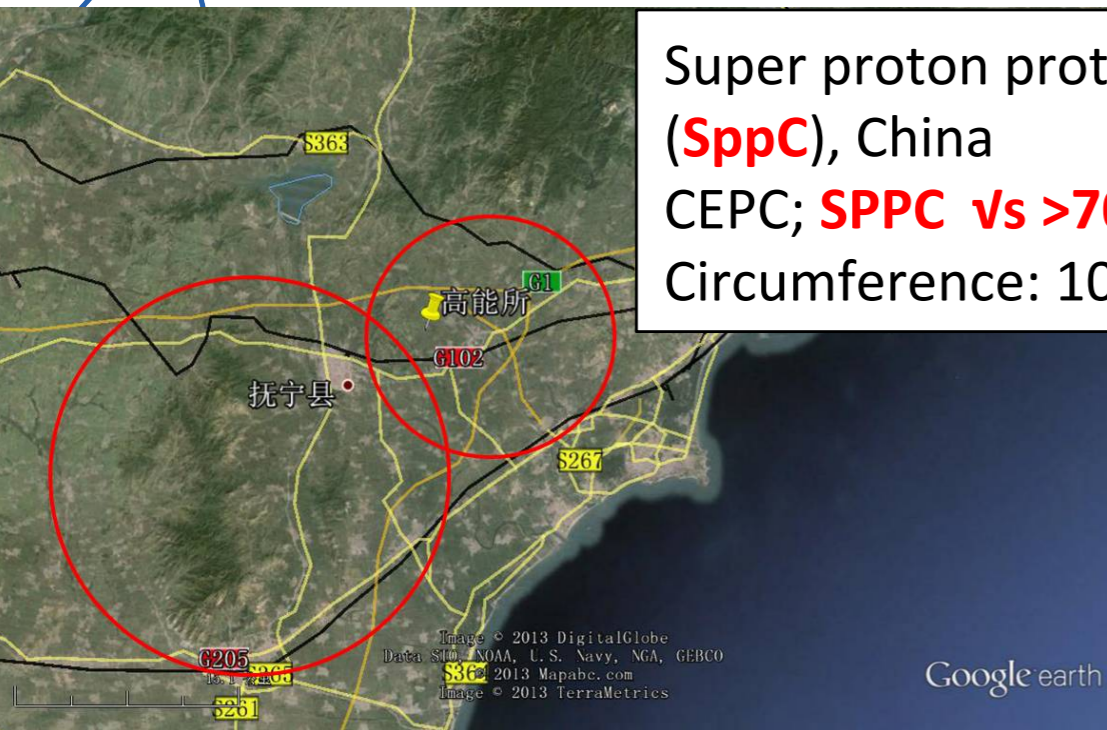
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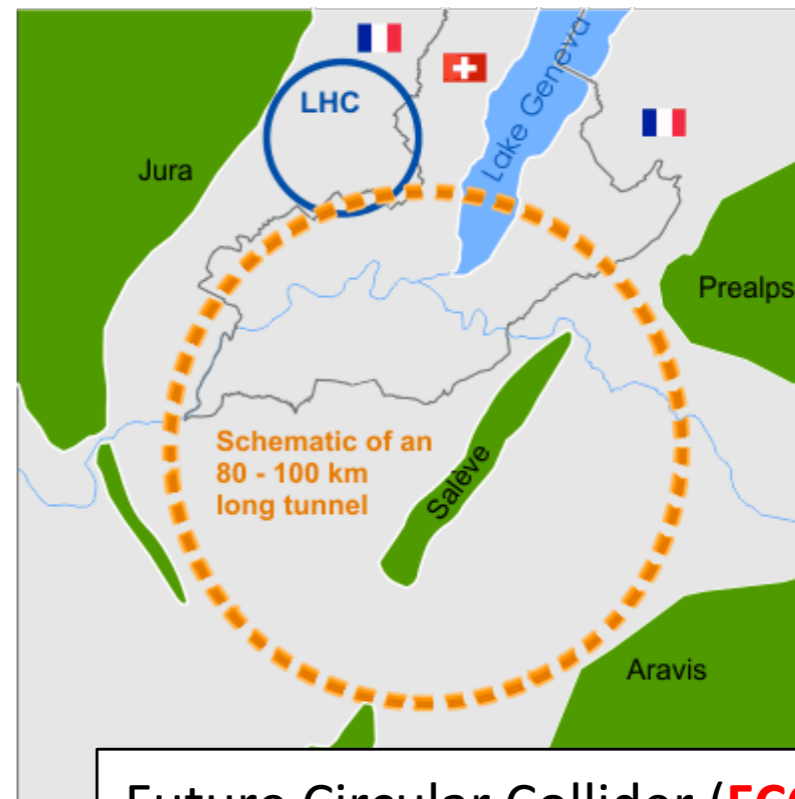
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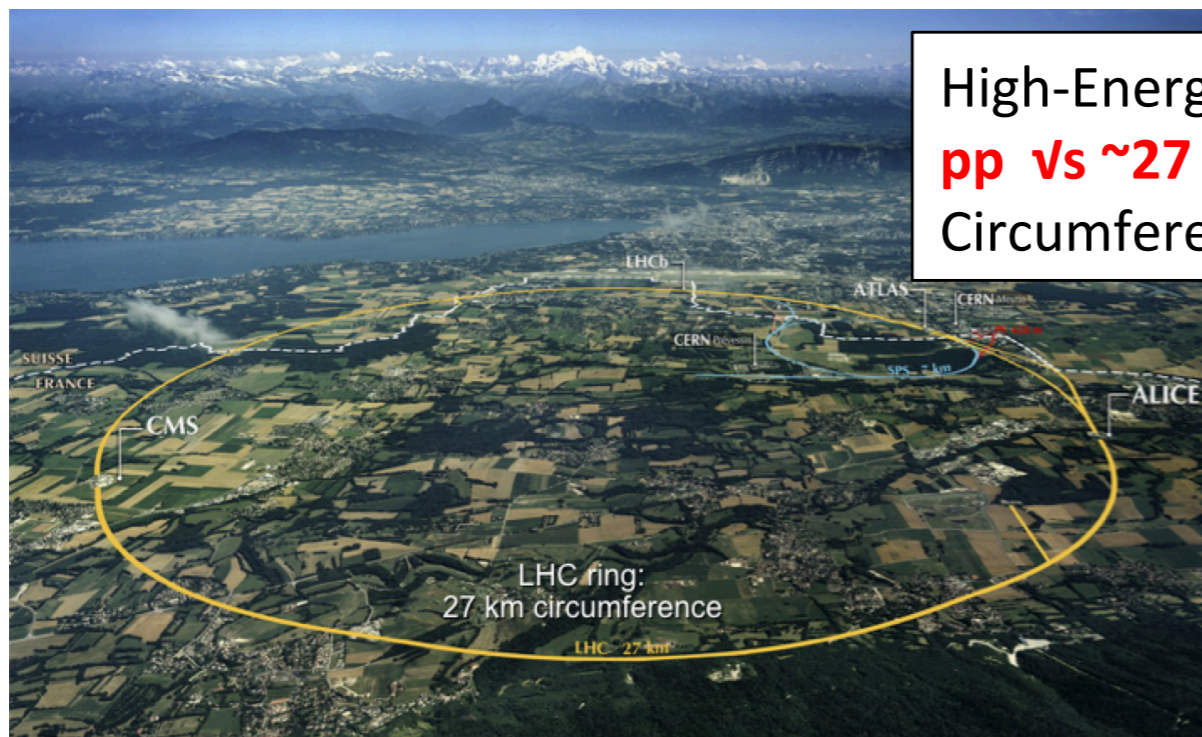
studies of high-energy **pp** colliders



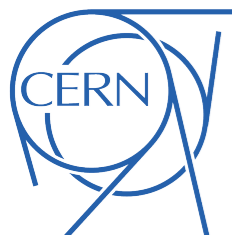
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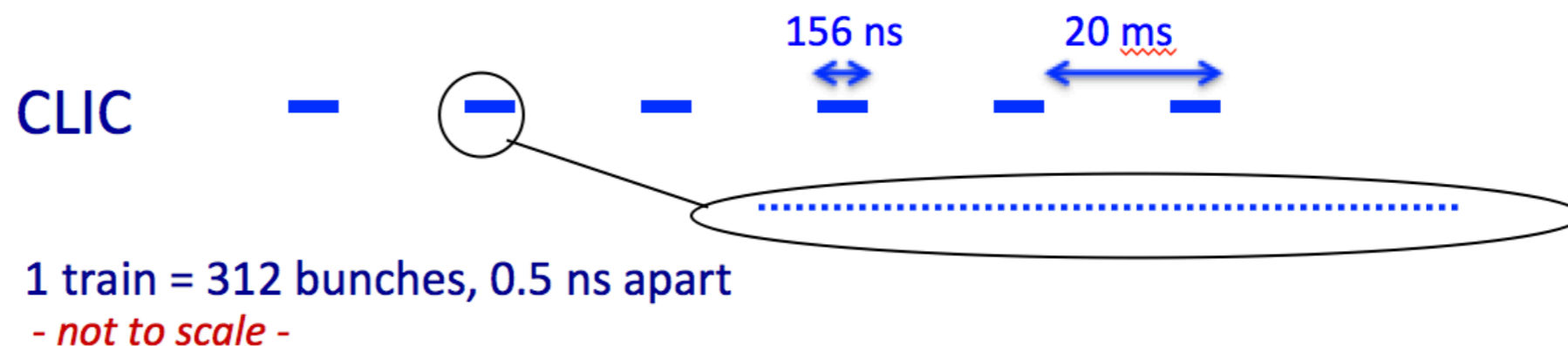


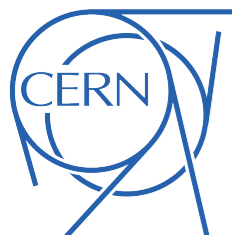
High-Energy LHC (**HE-LHC**): CERN
pp vs ~27 TeV
Circumference: 27 km



linear e^+e^- accelerator parameters

Parameter	ILC		CLIC		
	250 GeV (next stage)	500 GeV	380 GeV	1.5 TeV	3 TeV
Luminosity \mathcal{L} ($10^{34}\text{cm}^{-2}\text{sec}^{-1}$)	1.5	1.8	1.5	3.7	5.9
\mathcal{L} above 99% of \sqrt{s} ($10^{34}\text{cm}^{-2}\text{sec}^{-1}$)	1.3	1.0	0.9	1.4	2.0
Accelerator gradient (MV/m)	31.5	31.5	72	72/100	72/100
Site length (km)	~17	31	11.4	29	50
Repetition frequency (Hz)	10	5	50	50	50
Bunch separation (ns)	554	554	0.5	0.5	0.5
Number of bunches per train	1312	1312	352	312	312
Beam size at IP σ_x/σ_y (nm)	729/7.7	474/5.9	150/2.9	~60/1.5	~40/1
Beam size at IP σ_z (μm)	300	300	70	44	44





circular e^+e^- collider parameters

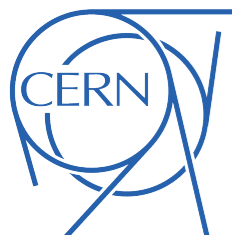


FCC-ee parameters:

parameter	Z	W	H (ZH)	ttbar
\sqrt{s} [GeV]	91	160	240	350
Beam current [mA]	1400	147	29	6.4
Number of bunches	71000	7500	740	62
Bunch intensity [10^{11}]	0.4	0.4	0.8	2.1
Bunch spacing [ns]	2.5 / 5.0	40	400	5000
SR energy loss / turn [GeV]	0.036	0.34	1.71	7.72
Total RF voltage [GV]	0.25	0.8	3.0	9.5
Long. damping time [turns]	1280	235	70	23
Bunch length with SR & BS [mm]	4.1	2.3	2.2	2.9
Luminosity / IP [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	130	16	5	1.4

Note on CEPC:

- pre-CDR 2015, 54 km ring
- CDR expected in 2017, 100 km ring → parameters @ H (HZ), W, Z under study *(see next slide)*

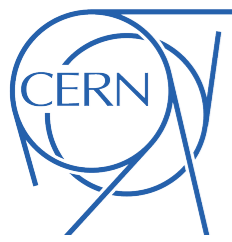


FCC-hh, HE-LHC, HL-LHC, LHC parameters

New tunnel

LHC tunnel

parameter	New tunnel		LHC tunnel		
		FCC-hh	HE-LHC	HL-LHC	LHC
\sqrt{s} [TeV]		100	27	14	14
Dipole field [T]		16	16	8.33	8.33
Circumference [km]		97.75	26.7	26.7	26.7
Beam current [A]		0.5	1.12	1.12	0.58
Bunch intensity [10^{11}]	1	1 (0.2)	2.2 (0.44)	2.2	1.15
Bunch spacing [ns]	25	25 (5)	25 (5)	25	25
Synchr. rad. power / ring [kW]		2400	101	7.3	3.6
SR power / length [W/m/ap.]		28.4	4.6	0.33	0.17
Long. emit. damping time [h]		0.54	1.8	12.9	12.9
Peak luminosity [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	5	30	25	5	1
events/bunch crossing	170	~1000 (200)	~800 (160)	135	27

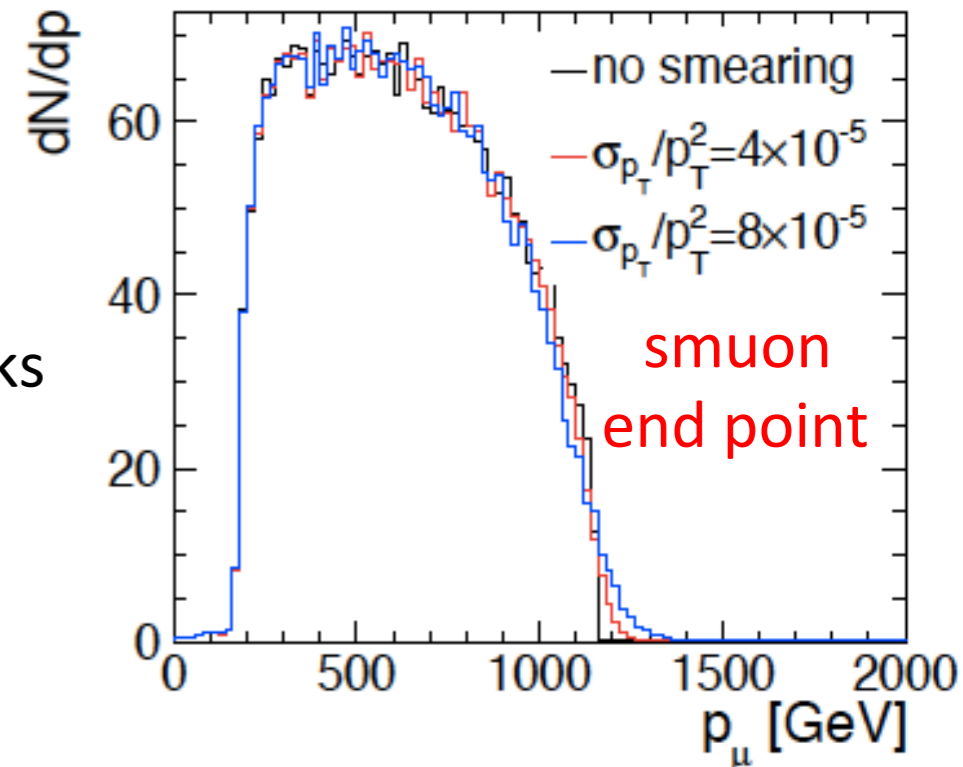


e^+e^- detector requirements (from physics)

★ momentum resolution:

e.g, ZH with $Z \rightarrow \mu\mu$, Smuon endpoint

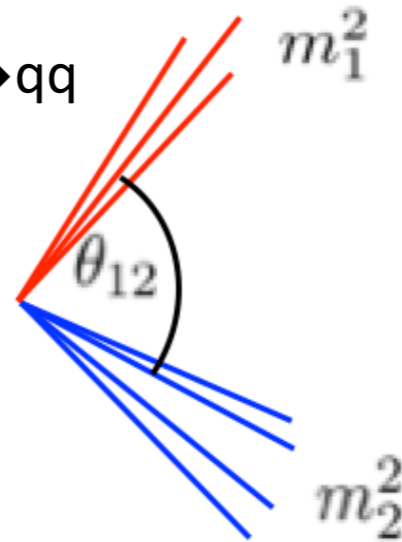
$$\sigma_{p_T} / p_T^2 \sim 2 \times 10^{-5} \text{ GeV}^{-1} \text{ for high } p_T \text{ tracks}$$



★ jet energy resolution:

e.g. W/Z/H di-jet mass separation, ZH with $Z \rightarrow qq$

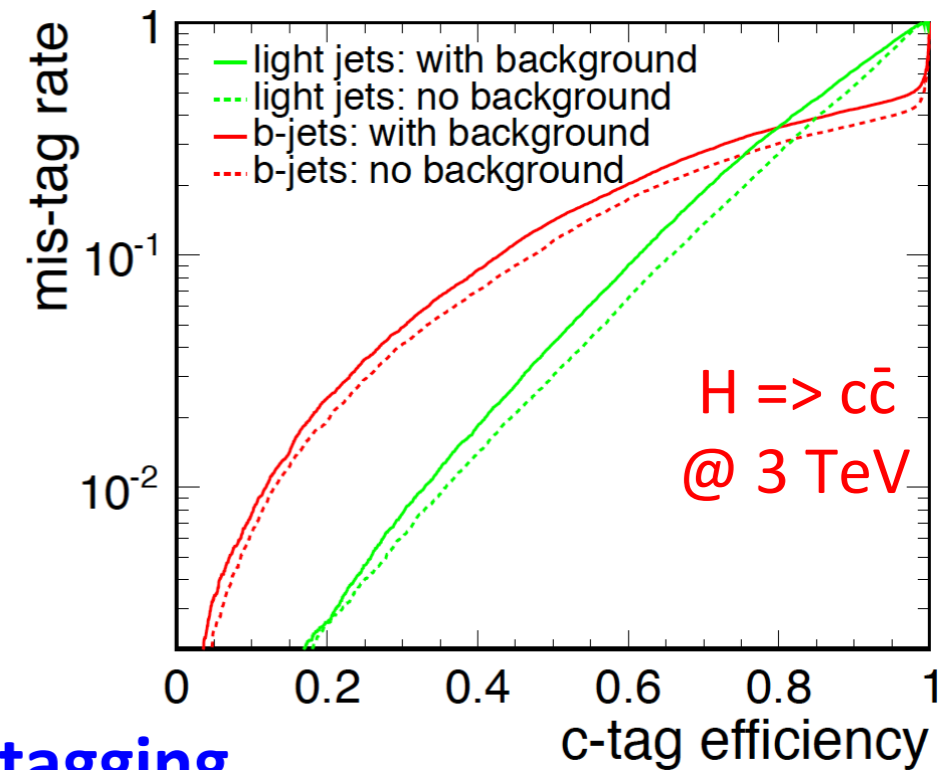
$$\frac{\sigma_E}{E} \sim 3.5 - 5 \% \text{ (for high-E jets, light quarks)}$$



★ impact parameter resolution:

e.g. c/b-tagging, Higgs BR

$$\sigma_{r\phi} = 5 \oplus 15 / (p[\text{GeV}] \sin^{\frac{3}{2}} \theta) \mu\text{m}$$



★ angular coverage, very forward electron/photon tagging

+ requirements from CLIC experimental conditions

calorimetry and PFA

Jet energy resolution + background suppression for optimal detector design

=> => fine-grained calorimetry + Particle Flow Analysis (PFA)

What is PFA?

Typical jet composition:
60% charged particles
30% photons
10% neutral hadrons



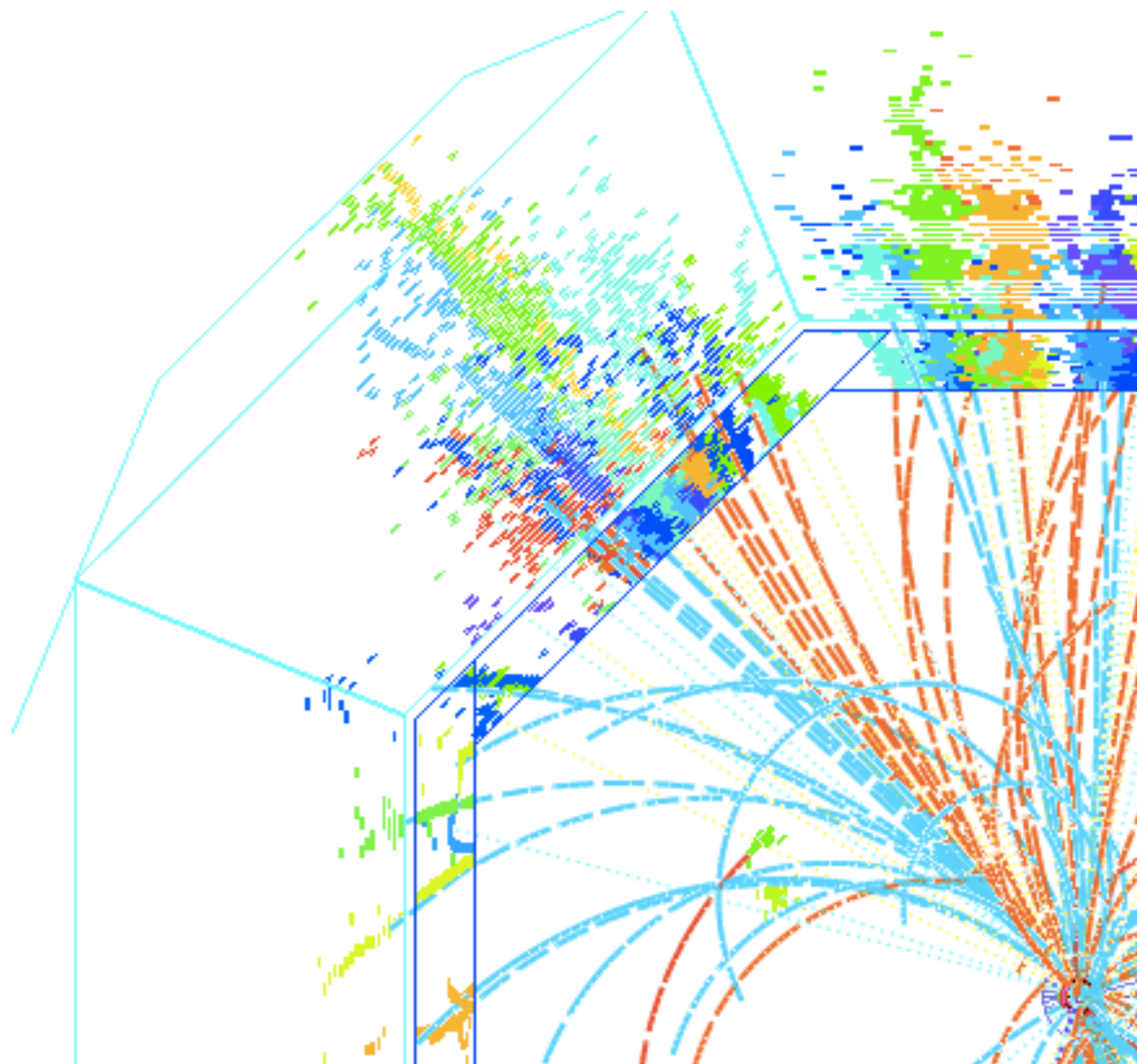
Always use the best info you have:

60% => tracker 😊 😊

30% => ECAL 😊

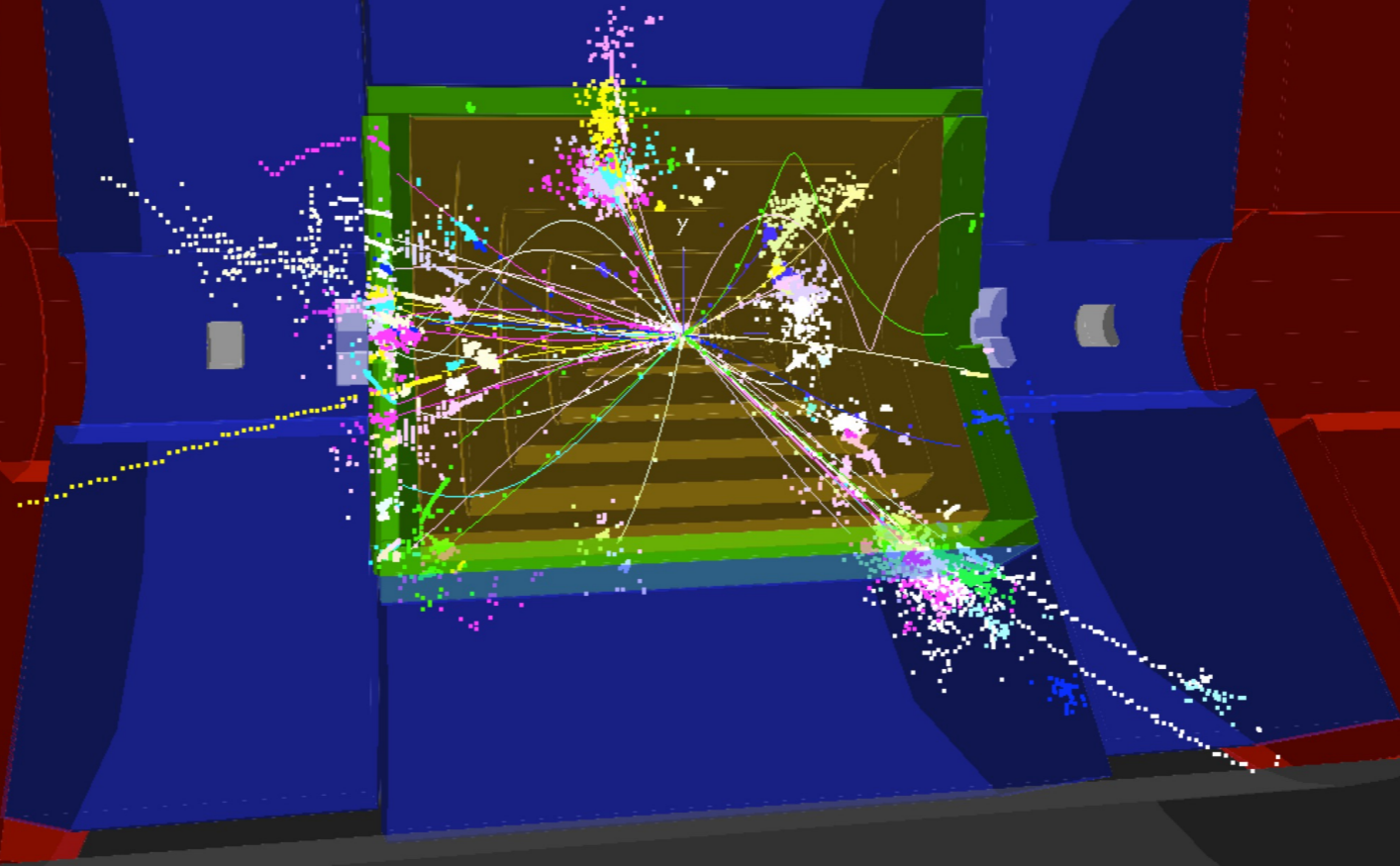
10% => HCAL 😞

Hardware + software !

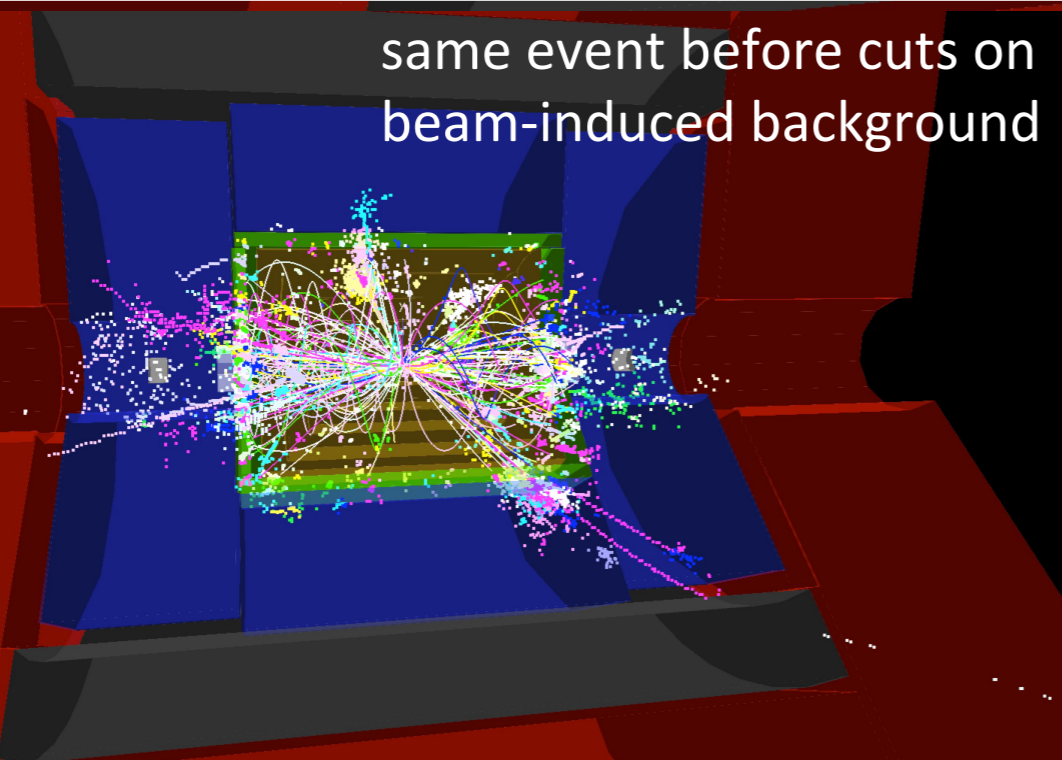


$e^+e^- \rightarrow t\bar{t}H \rightarrow WbW\bar{b}H \rightarrow q\bar{q}b \tau\nu\bar{b} b\bar{b}$

CLIC 1.4 TeV



same event before cuts on beam-induced background



Highly granular calorimetry + precise hit timing



Very effective in suppressing backgrounds for fully reconstructed particles



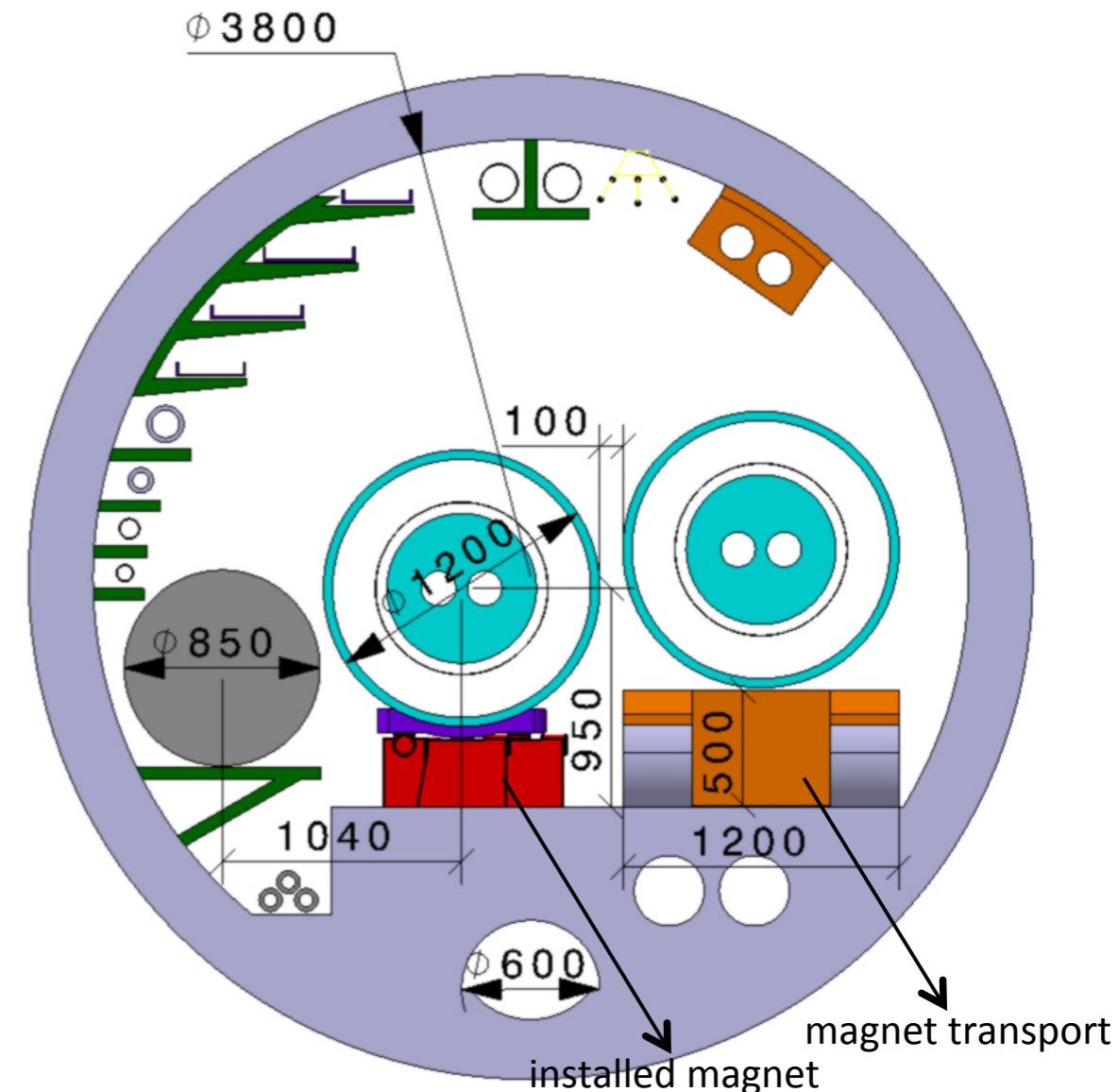
General trend for e^+e^- and pp options (e.g. CMS endcap calorimetry for HL-LHC)

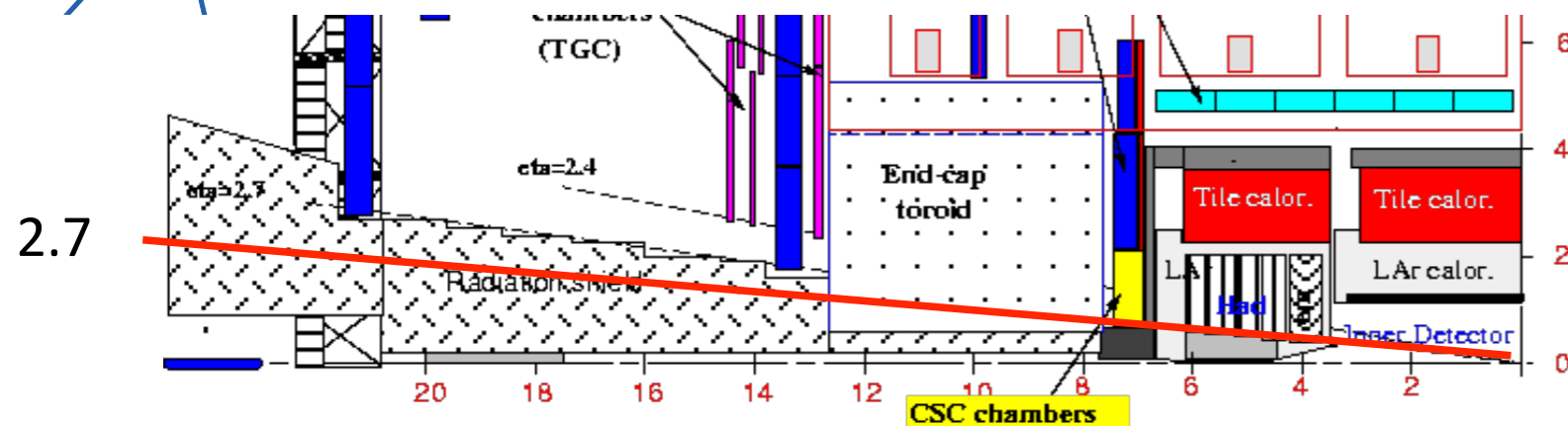
Use the FCC-hh magnet technology for a proton-proton collider in the LHC tunnel

- **$\sqrt{s}=27$ TeV** (=14 TeV * 16 T / 8.33 T)
- **Luminosity 4 times higher than HL-LHC** ($1/E^2$)
- Constraint on external diameter of magnet cryostat, 1.2 m, for LHC tunnel compatibility

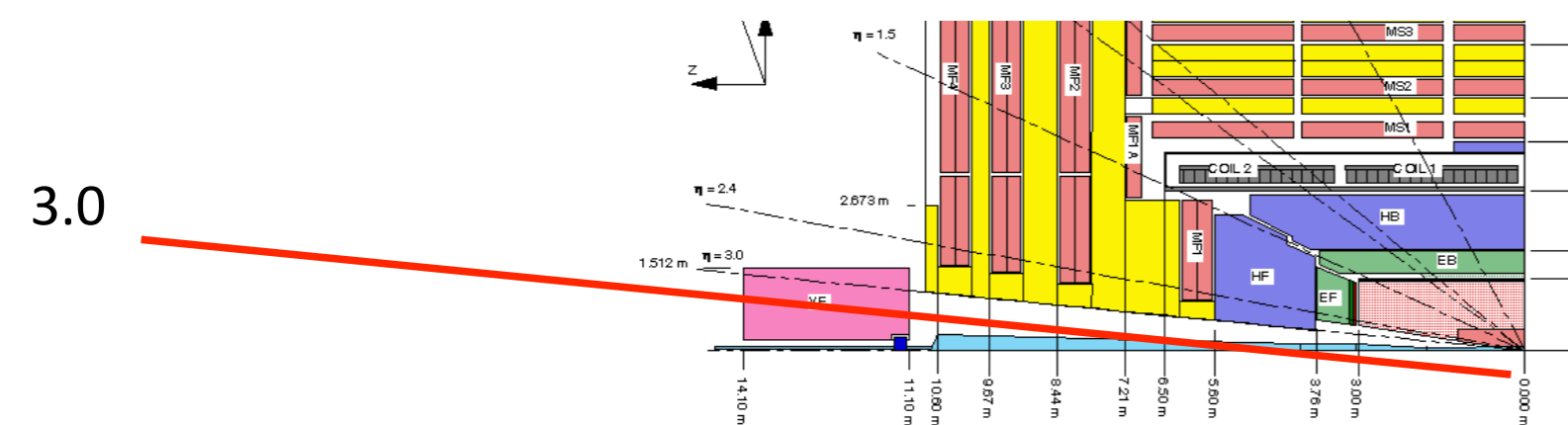
Key ingredients:

- FCC-hh magnet technology
- FCC-hh vacuum system
- HL-LHC crab waist scheme
- HL-LHC electron lens
- HL-LHC/LIU beam parameters (25 ns bunch structure, 5 ns option)

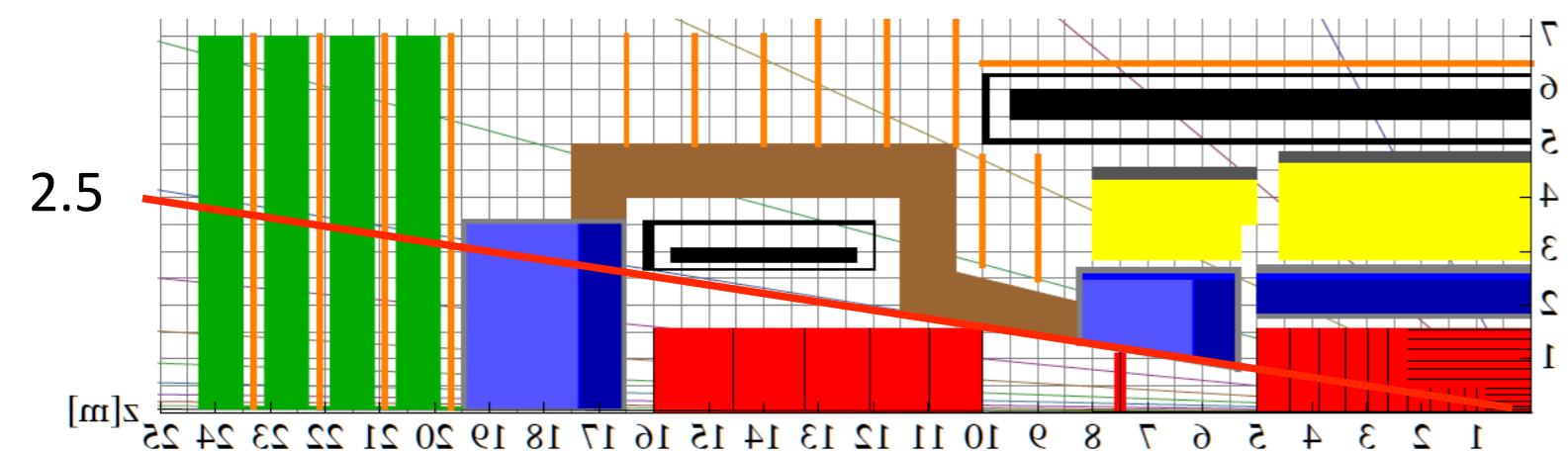




ATLAS



CMS



FCC-hh
(cavern length of 70 m required)

Compared to ATLAS / CMS, the forward calorimeters are moved far out in order to reduce radiation load and increase granularity.

→ A large shielding (brown) needed to stop neutrons from escaping to cavern and muon syst.