

# Sub-GeV mass dark matter with SuperCDMS

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on behalf of the SuperCDMS collaboration

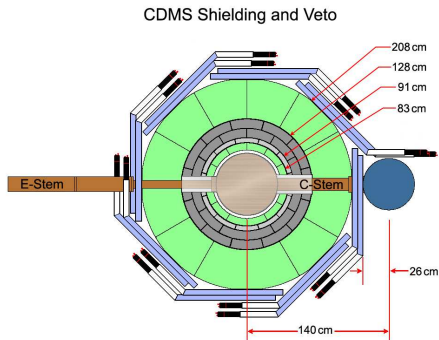
February 23, 2018



SMU®

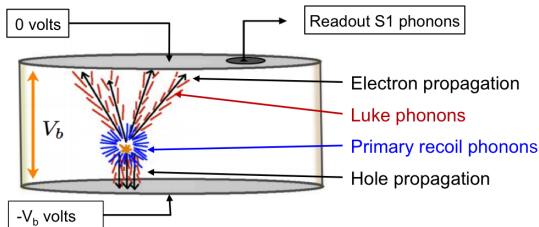


# SuperCDMS Soudan



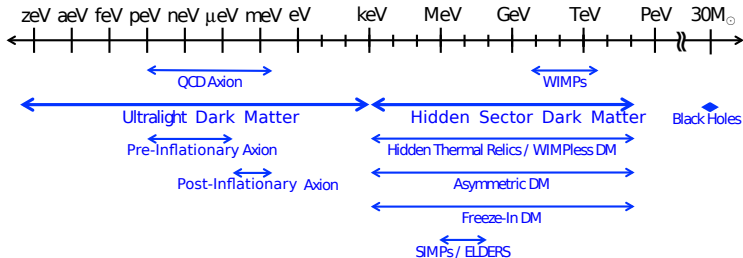
- Located in Soudan Underground Lab,  $\approx \frac{1}{2}$  mile underground with 2090 M.W.E. of overburden
- Utilizes the same shielding and cryostat from CDMS-II experiment
- Collected about 2500 kg-days of raw exposure over experiment lifetime
- Data taking ended in 2015, collaboration shifting focus to SNOLAB

## HV Biasing

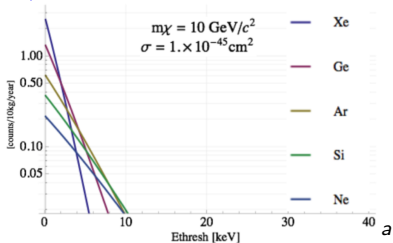


- Phonons are created from charges passing through a crystal through Neganov-Trofimov-Luke effect
- The contribution to total phonon energy goes as  $N_{e/h}eV_b$  : proportional to bias voltage  $V_b$
- High bias voltage allows us to measure small amount of charges through phonon signal (CDMSlite mode/HV)
- Trade-off: no separate measurement of primary phonon signal, sacrifices ER/NR discrimination

# Low mass dark matter Battaglieri et al. arXiv:1707.04591

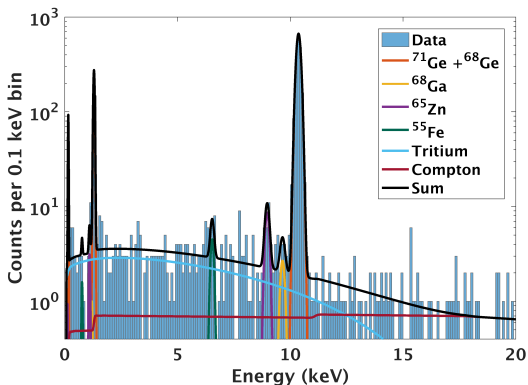


- WIMPs are just one possibility
- Kinematics for light masses disfavorable, sensitivity driven by threshold
- Need new ideas and approaches to probe these low masses



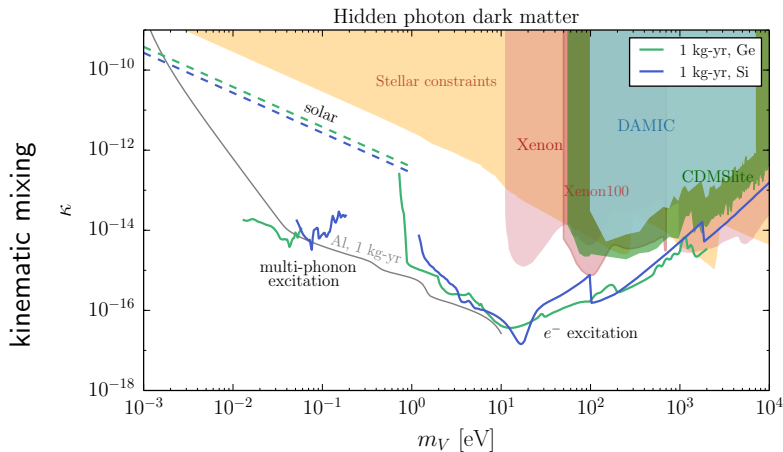
<sup>a</sup>Mirabolfathi - arXiv:1308.0044

## Dark photon absorption with CDMSlite data



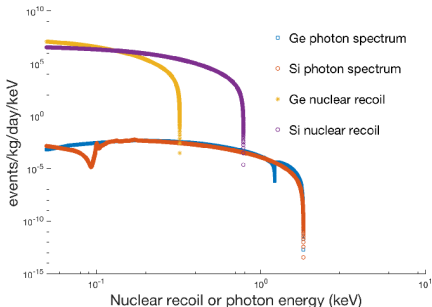
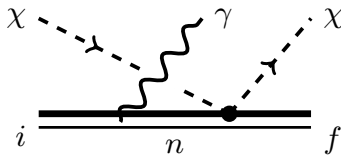
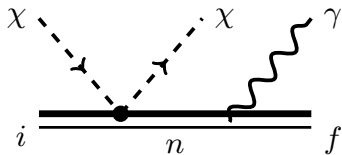
- Mediator between standard model and dark matter with finite mass ( $m_{A'}$ )
- Absorption rate,  $R \simeq \frac{\rho_{DM}}{m_{A'}c^2} \epsilon^2 \sigma_{p.e.}(E_{\text{gamma}} = m_{A'})c$
- Signal is mono-energetic electron with  $E = m_{A'}$
- Search strategy is a bump-hunt in our spectrum!

# Dark photons - Hochberg, Lin and Zurek Phys. Rev. D 95, 023013



- Sensitivity extends down to band gap (Ge .7 eV, Si 1.1 eV)
- We can expect to cover a large amount of parameter space at SNOLAB

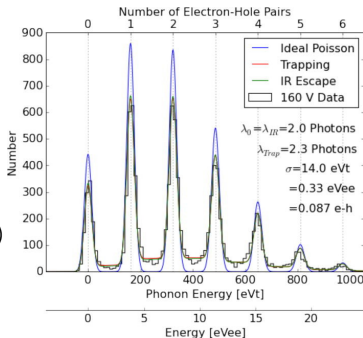
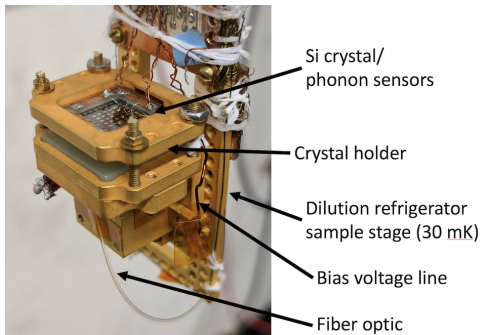
# Bremsstrahlung - Kouvaris and Pradler Phys. Rev. Lett. 118, 031803



1 GeV WIMP

- Recoiling nucleus can emit a photon as it slows down in material
- Energy of this photon can extend to higher energies than NR
- Photon acts as probe to low mass WIMP scattering below detector threshold

## Single e/h device - Romani et al. Appl. Phys. Lett. 112, 043501



- Single e/h-pair sensitivity has been recently demonstrated in 0.93g Si crystal
- Single e/h-pair resolution goal of SuperCDMS SNOLAB
- Such devices will have sensitivity to a variety of sub-GeV DM models with  $\sim g \cdot d$  exposures







# Summary



- SuperCDMS Soudan WIMP searches are nearly complete see B. Loer's talk
- Theorists have provided us with many new channels to explore
- Current CDMSlite data can be used to search for these signals
- Future is looking equally bright with new technological achievements and the planned SuperCDMS SNOLAB experiment ! see B. Loer's talk

# Backup

## Backup Slides

# SuperCDMS collaboration



California Inst. of Tech.



CNRS-LPN\*



Durham University



FNAL



NISER

NIST

NIST\*



Northwestern



PNNL



Queen's University



Santa Clara University



SLAC



South Dakota SM&T



SMU



SNOLAB



Stanford University



Texas A&M University



TRIUMF



U. British Columbia



U. California, Berkeley



U. Colorado Denver



U. Evansville



U. Florida



U. Montréal



U. Minnesota



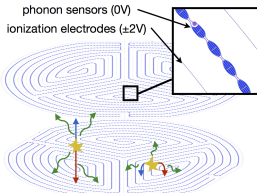
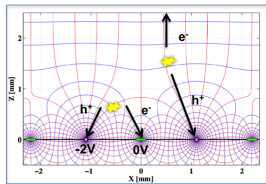
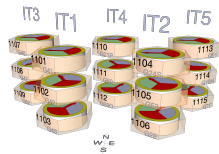
U. South Dakota



U. Toronto

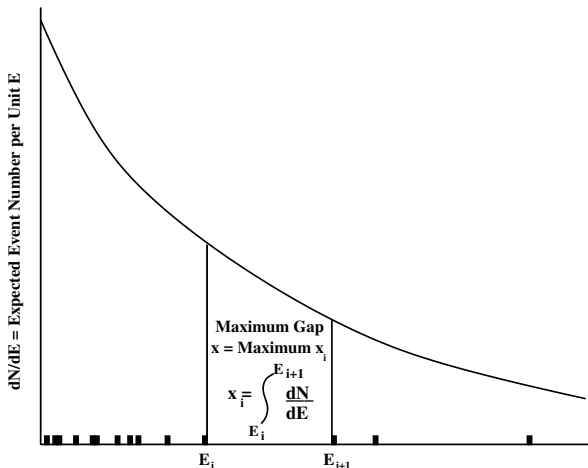
\* Associate members

# interleaved Z-sensitive Ionization and Phonon (iZIP) Detectors



- Detector array: 15 Ge iZIP detectors (0.6 kg each) operating around 50 mK
- 4 phonon and 2 charge channels on each detector face
- Phonon channels are grounded, charge channels are biased at  $\pm 2$  V
- Field configuration causes events near surface to have charge collection localized to one side

## Yellin Optimal Interval



Conceptually similar to the optimal gap method except that it allows for  $N$  events to fall into region.