

Determination of Backgrounds for the LUX Experiment

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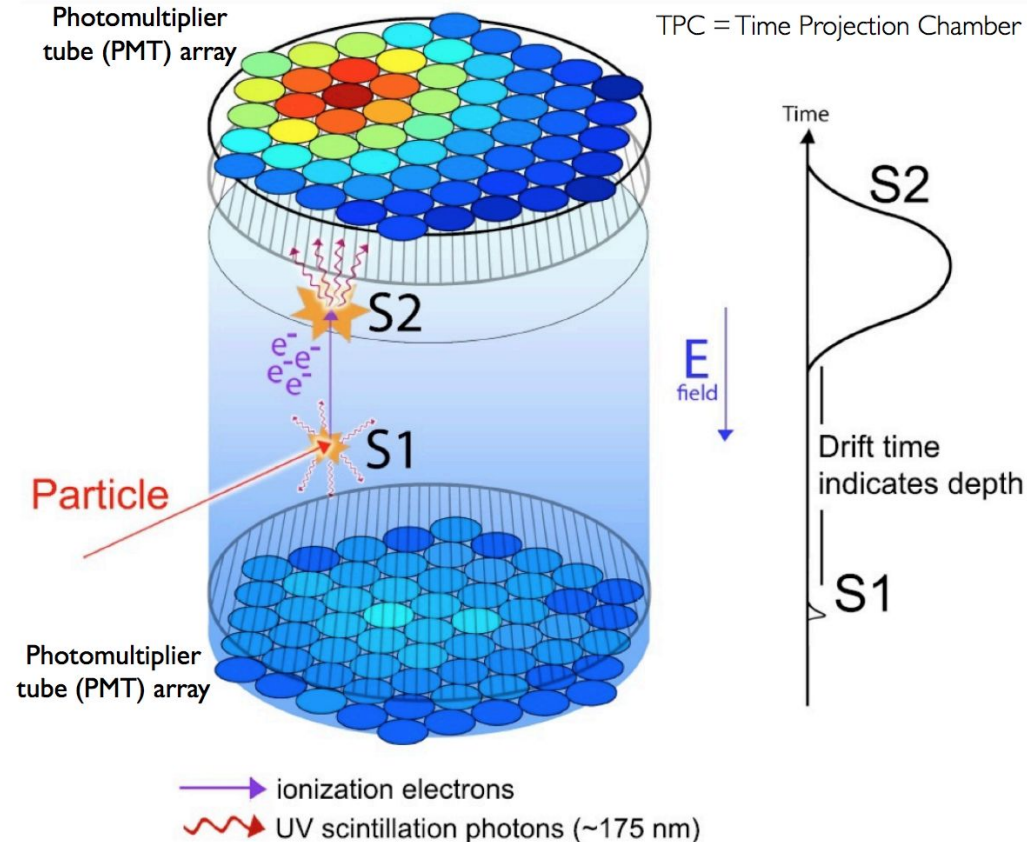
UCLA Dark Matter 2018

On behalf of the LUX collaboration

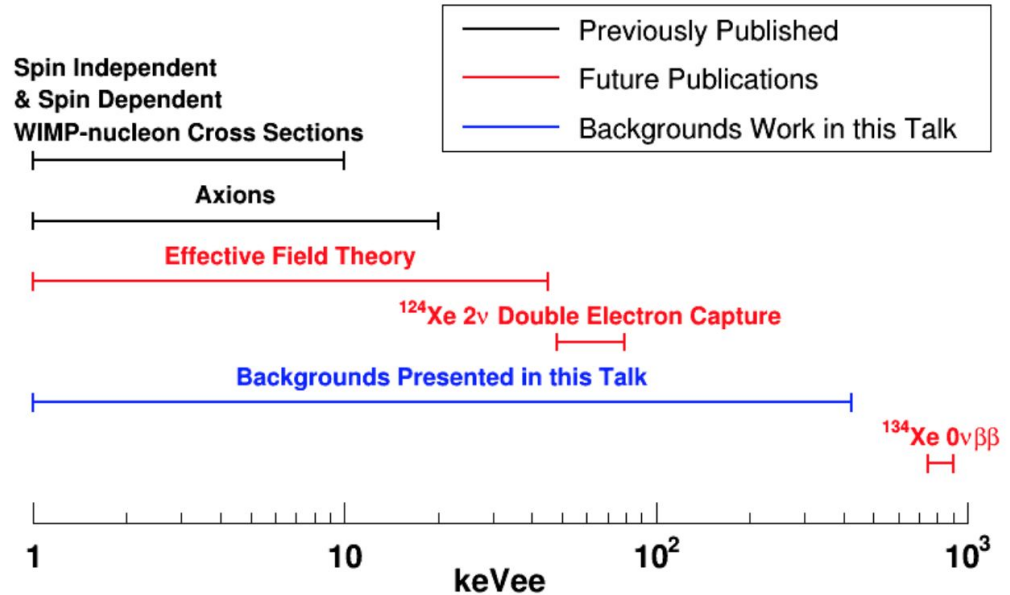
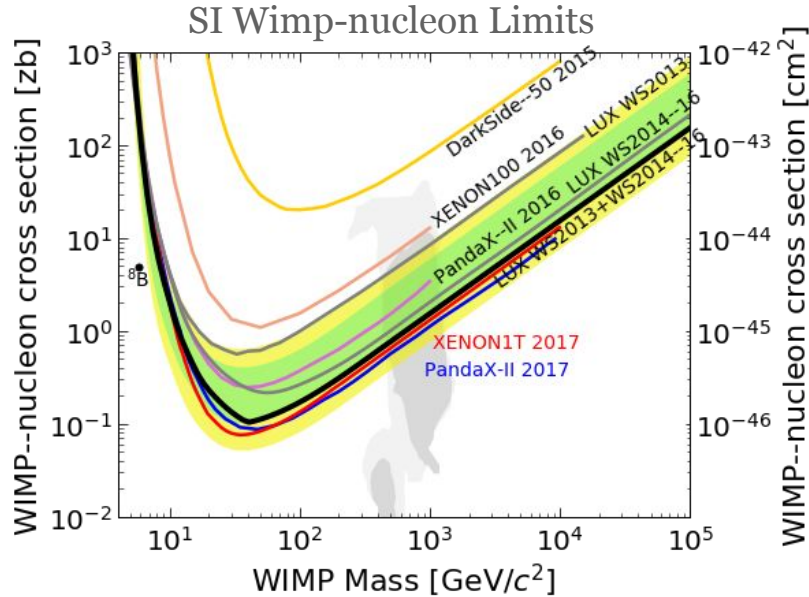


Large Underground Xenon (LUX) Detector

- Two-phase xenon time projection chamber
- An interacting particle deposits energy in two channels
 - Excitation
 - Ionization
- Prompt scintillation (S1 signal) is immediately detected by PMTs
- Electrons are drifted upward and extracted into the gas phase region creating secondary scintillation detected by the PMTs (S2 signal)



Extending Energy



- In Jan 2017, LUX published final spin independent WIMP limit ($\sim 1\text{-}10$ keVee)
- Data from two runs: WS-2013 (95 live days) and WS-2014/16 (332 live days)
- Now we're looking at high energy physics processes
- Additional backgrounds from β , ϵ , and isomeric transition decays in the xenon
- This talk will present background search results from WS-2014/16 in the energy range 0 - 425 keVee

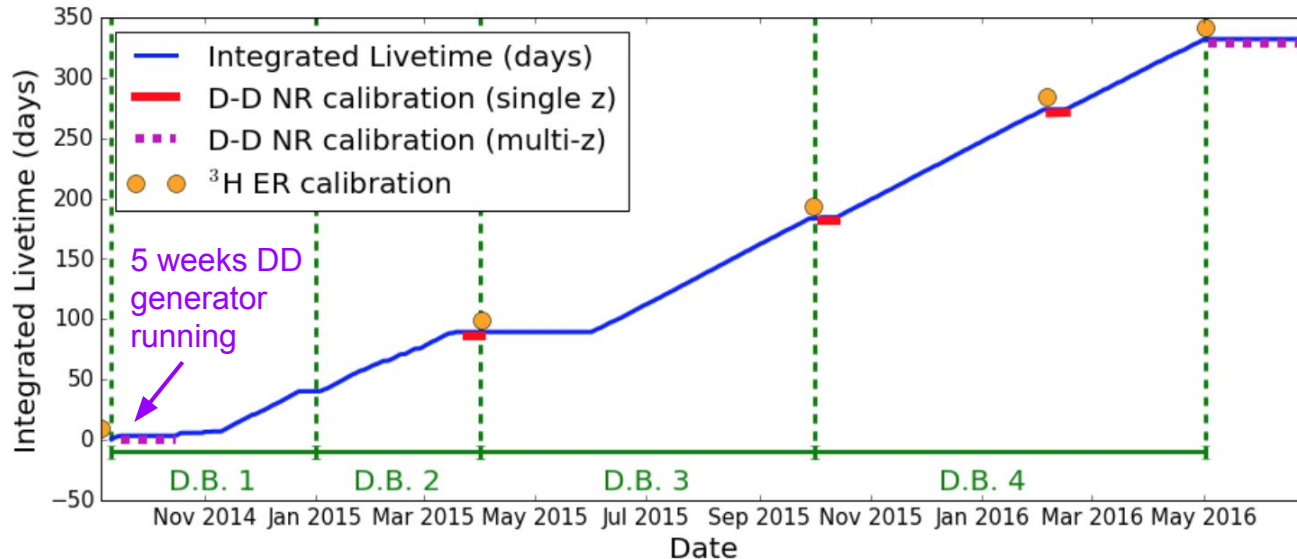
Outline



1. Calibration of WS-2014/16 electron recoil energy scale
2. Four sources backgrounds:
 - a. Short lived radioisotopes from activation of the xenon with a DD neutron generator calibrations
 - b. ^{210}Pb from ^{222}Rn daughter plate-out on detector surfaces
 - c. Detector effects, such as PMT afterpulsing, photoionization of grids and impurities, and electron trains
 - d. Neutron backgrounds from PMTs and PTFE

WS-2014/16 Calibration Data

- Every three months we perform:
 - Calibrations of the electron recoil (ER) band, light and charge yields with tritium source
 - Calibrations of the nuclear recoil (NR) band, light and charge yields with 2.45 MeV neutrons from a DD generator
 - **Calibrations of the energy of electron recoils in the detector using ^{131m}Xe , ^{129m}Xe , ^{125}Xe xenon activation lines from the DD neutron generator**

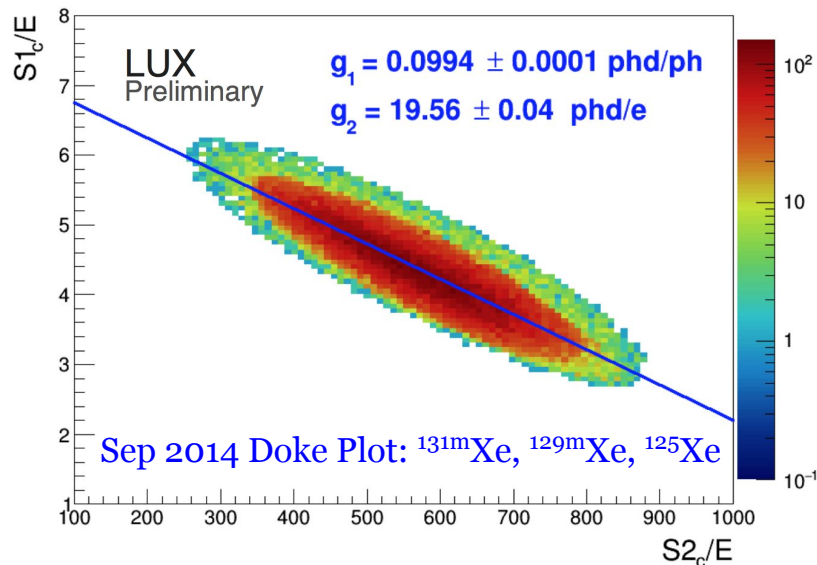


- The **data is divided into four date bins** and specific energy reconstruction parameters are applied to each bin

LUX ER Energy Calibration

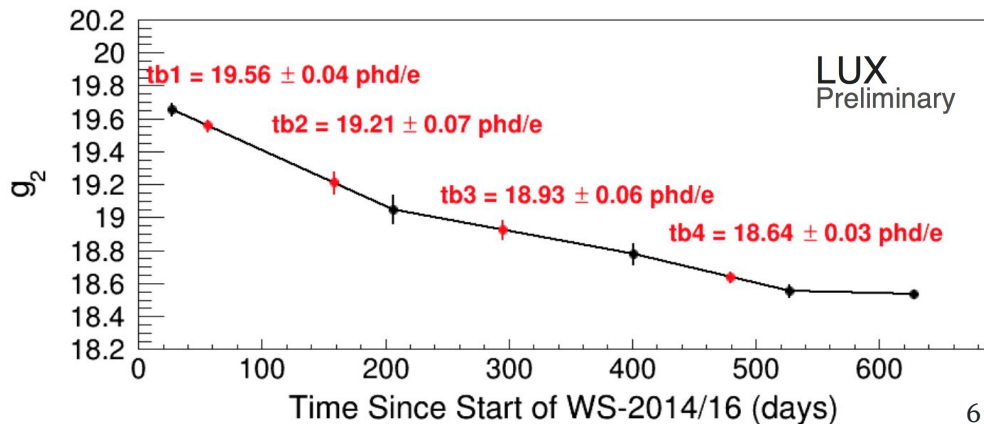
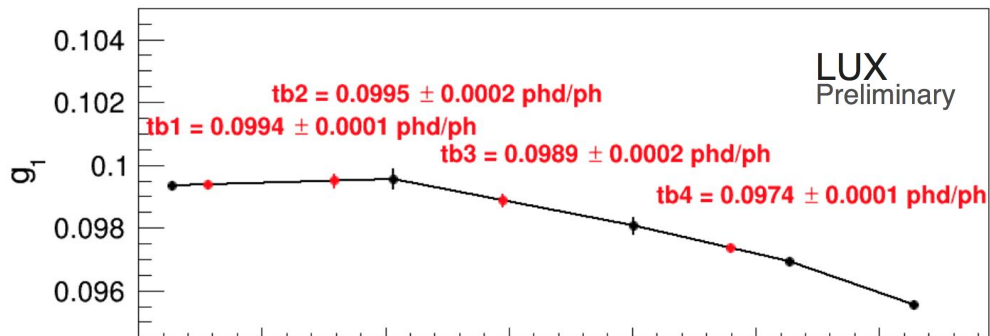
$$E = W (S_1/g_1 + S_2/g_2)$$

g_1 = efficiency for detection of a prompt scintillation photons
 g_2 = efficiency/gain for detection of electron signal
 W = average energy to produce a single excited or ionized atom

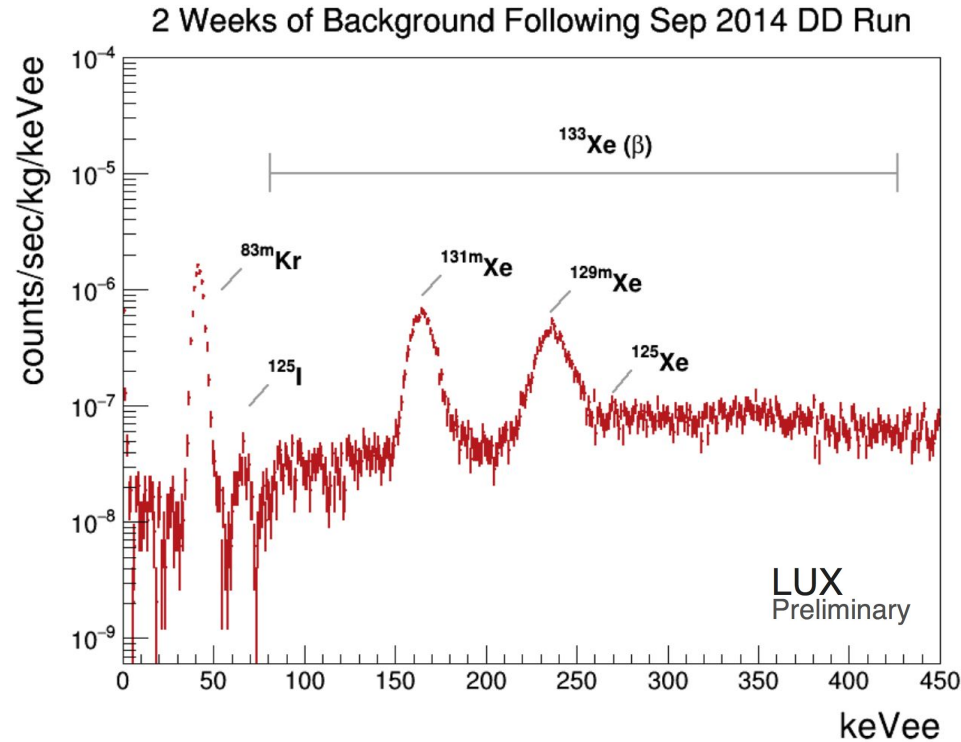


Error on g_1 and g_2 includes both statistical and systematic components. The systematic component comes from the elliptical cuts used to select the calibration sources.

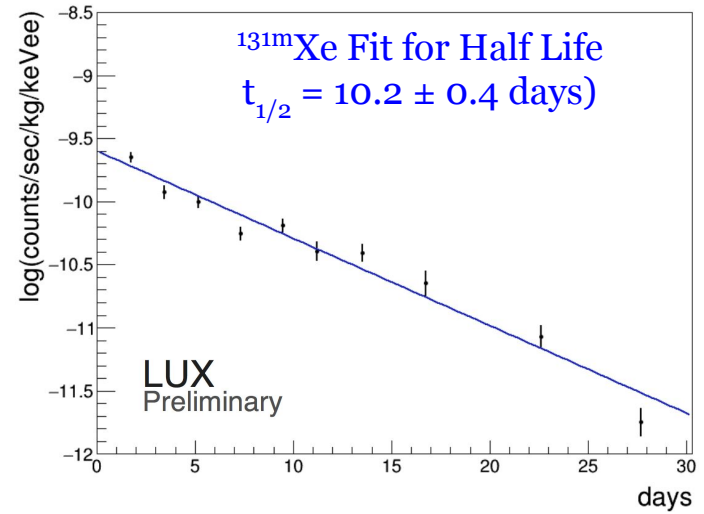
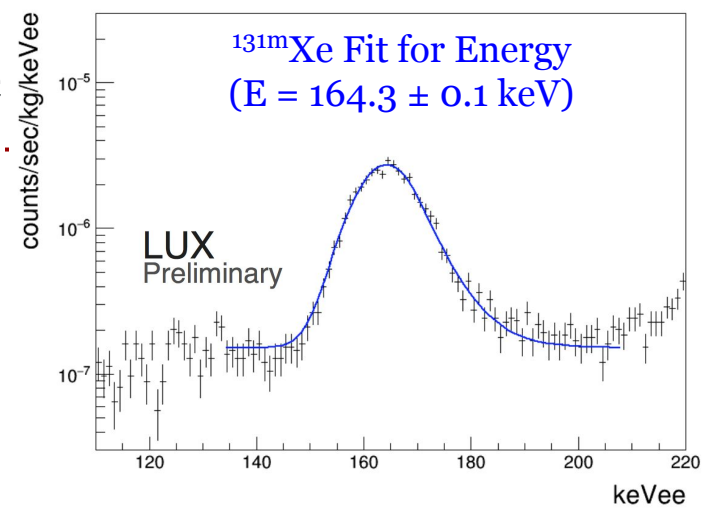
- g_1 and g_2 calculated at each DD calibration time
- g_1 and g_2 interpolated to the midpoint of each date bin



Short Lived DD-n Activation of Xe



- In addition to activation peaks, $^{83\text{m}}\text{Kr}$ (continuously being injected for calibration) is visible
- **Data on this slide follows the multiple neutron calibration runs over 5 weeks at different z-depths that occurred in Sep 2014**



Determination of Short Lived Xe Activation Isotopes

	Decay Mode	Energy *	Measured Energy	Half Life *	Measured Decay Constant
^{125}I	ϵ	67.3 keV **	67.1 \pm 0.4 keV	59.4 d	5.4 \pm 0.4 d ***
^{133}Xe	$\beta + \text{IT}$	$Q_{\text{max}} = 346.4 + \text{IT } 81 \text{ keV}$		5.2475 d	5.0 \pm 0.3 d
$^{131\text{m}}\text{Xe}$	IT	163.9 keV	163.6 \pm 0.1 keV	11.84 d	10.7 \pm 0.4 d
$^{129\text{m}}\text{Xe}$	IT	236.1 keV	235.3 \pm 0.1 keV	8.88 d	9.1 \pm 0.3 d
^{125}Xe	ϵ	275 keV **	275.5 \pm 0.6 keV	16.9 h	

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Uncertainties in table are statistical uncertainty in the fit for energy or half life

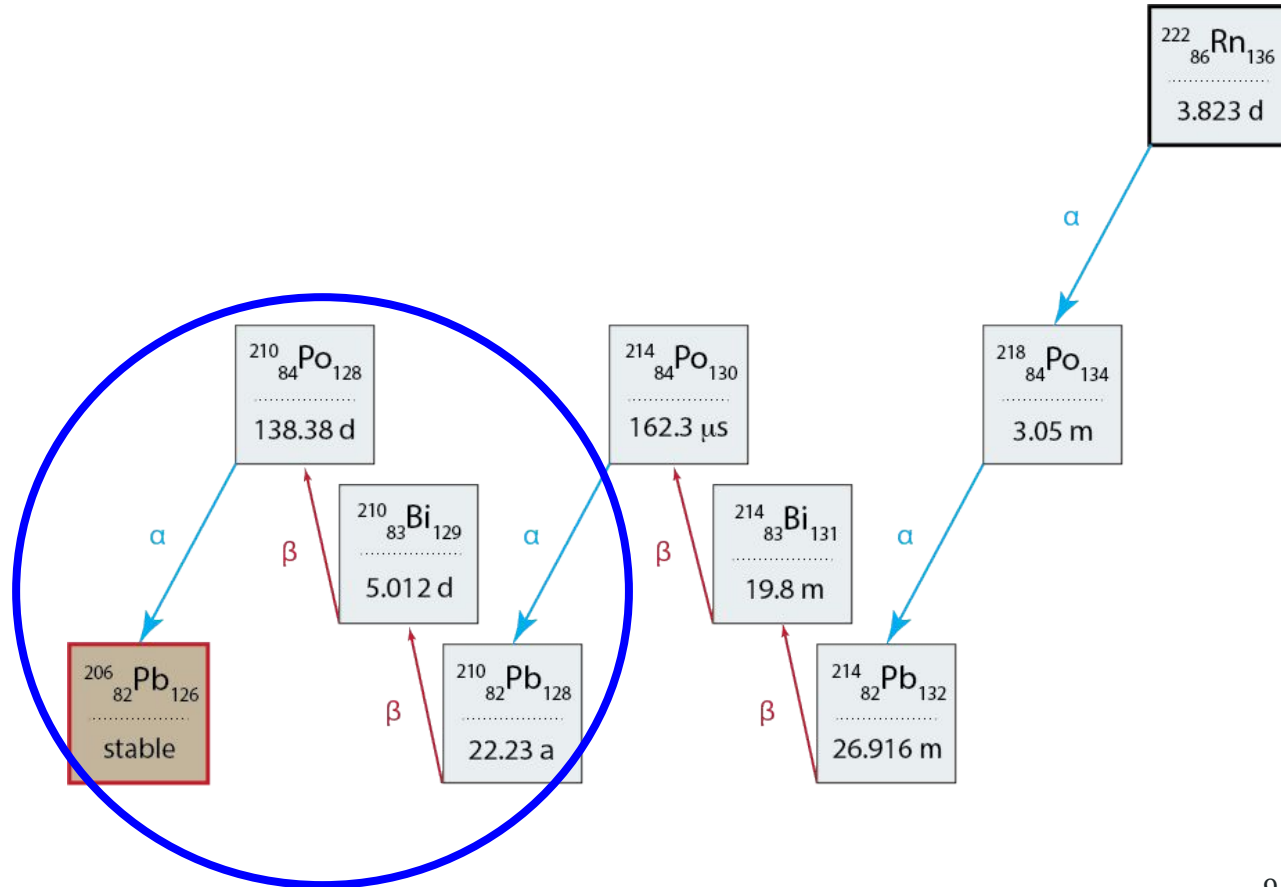
*Energy and half life measurements from National Nuclear Data Center website (<http://www.nndc.bnl.gov/chart/chartNuc.jsp>)

**An estimate of energy deposited in electron recoils in LUX from a K-shell electron capture

***Effective half life. Represents rate at which ^{125}I is removed from the xenon by the getter. Energy is the close the two neutrino Double Electron Capture ^{124}Xe Energy (63.6 keV)

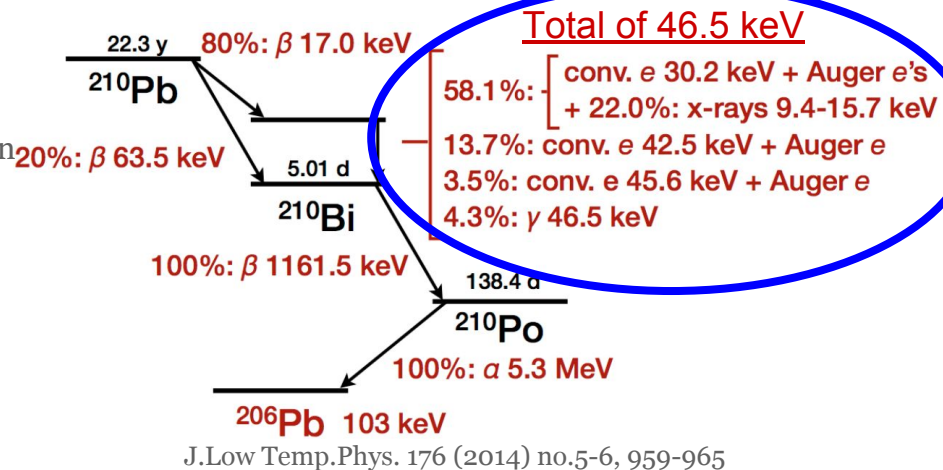
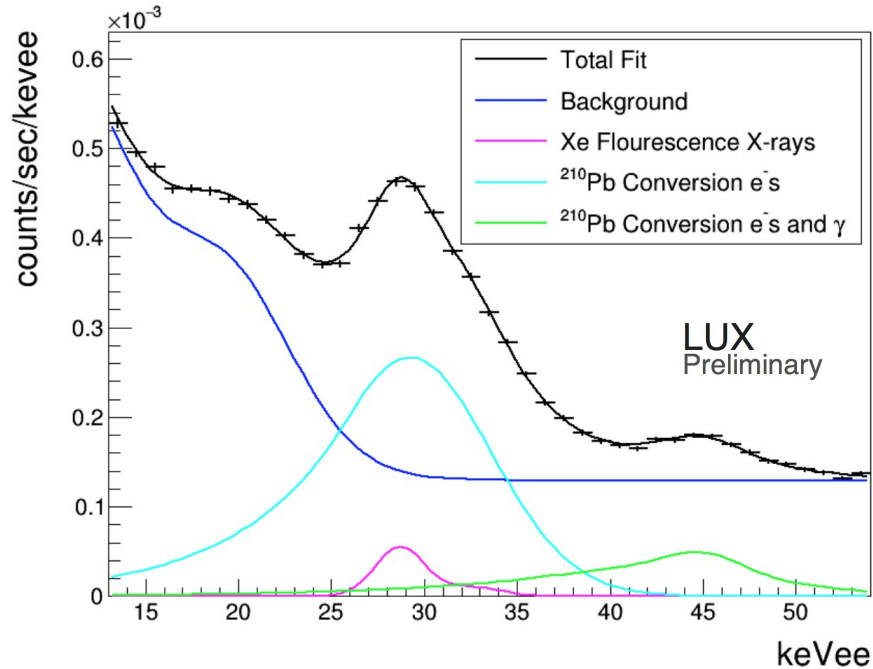
^{210}Pb on Detector surfaces

- During construction ^{222}Rn progeny plate out on the inner PTFE walls of the detector
- All short lived isotopes decay away leaving ^{210}Pb , ^{210}Bi , and ^{210}Po
- These isotopes can be absorbed off of the walls into the xenon



^{210}Pb on Walls

Fit requires contributions from ^{210}Pb conversion electrons in the range 30.1-33.1 keV, ^{210}Pb conversion electrons and gamma ray in range 42.6-46.5 keV, and xenon fluorescence in range 29.5-34.5 keV. Lower limit is set assuming LUX is capable of seeing a fraction of the ^{210}Pb decay products on the wall.

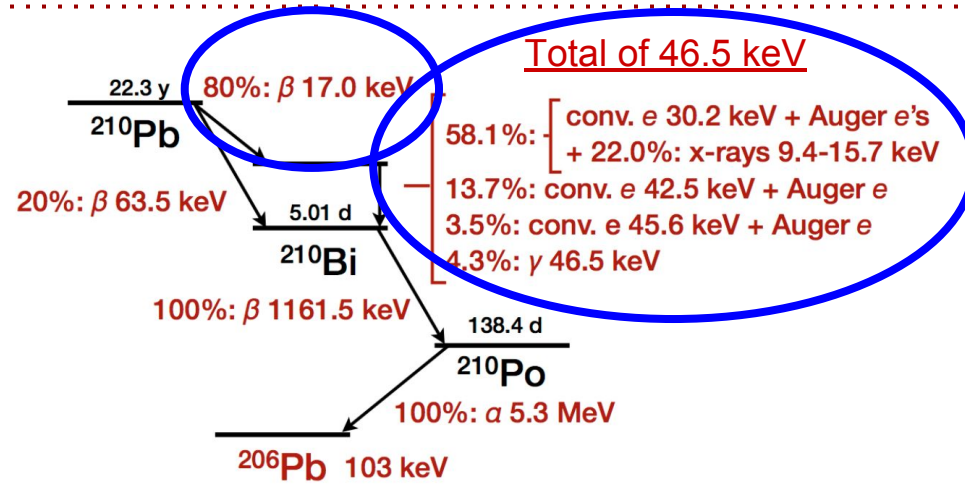
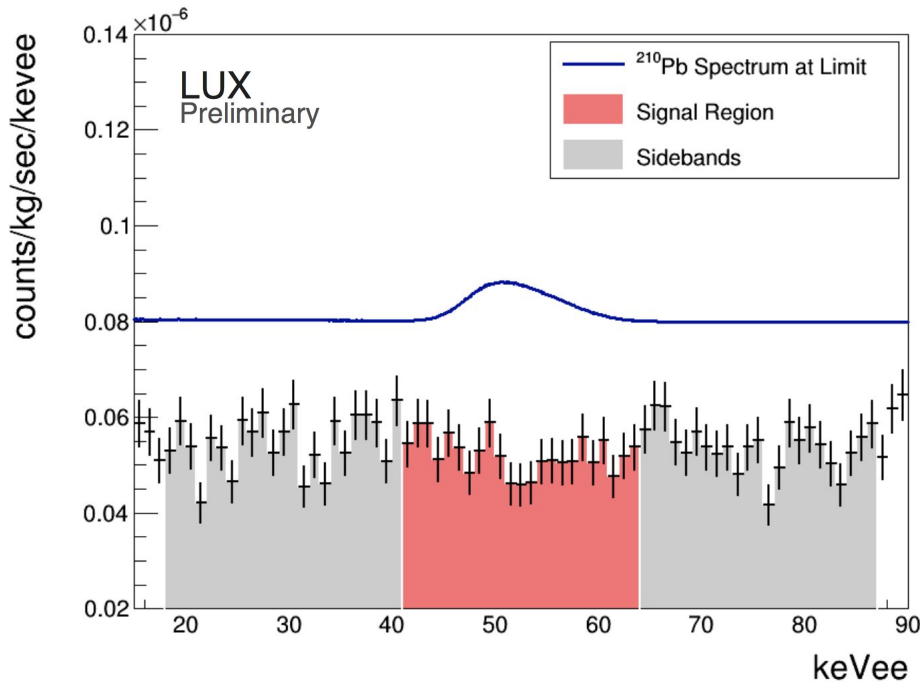


**Activity of ^{210}Pb on wall
in the fiducial volume
drift range for
WS-2014/16
> 9.6 ± 0.6 mBq**

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^{210}Pb in the Fiducial Volume

Background rate is estimated in 17 keV beta region using sidebands. Limit is set at 90% confidence interval as defined in: Applied Radiation and Isotopes Vol 53, Issues 1–2, 15 July 2000, P 45-50.

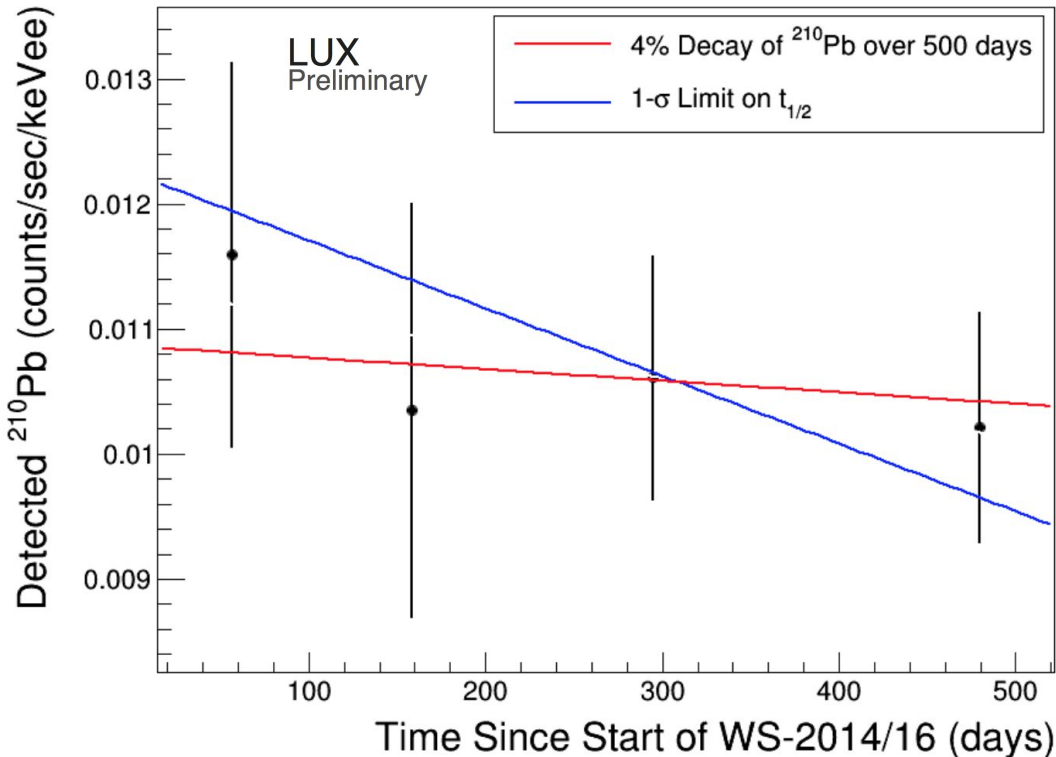


J.Low Temp.Phys. 176 (2014) no.5-6, 959-965

**Activity of ^{210}Pb
 in fiducial volume
 < 0.099 $\mu\text{Bq/kg}$**

LUX Preliminary

Limit on Leaching of ^{210}Pb into Xenon



- Activity of ^{210}Pb was measured for each date bin
- If there is no leaching, ^{210}Pb activity will decay by 4% over length of WS-2014/16
- Limit on decay constant for leaching of ^{210}Pb from detector walls is given as the fit value less 1- σ , correcting for 4% ^{210}Pb decay

**$t_{1/2}$ of ^{210}Pb leaching off wall
> 1.6×10^3 days**

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Detector Effects: High Energy Event Classification

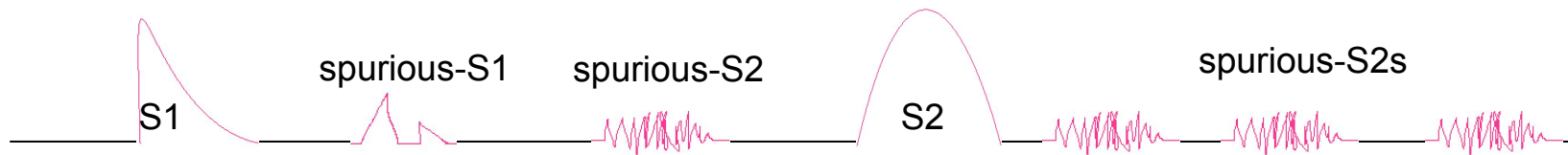
Following Large S1s:

- PMT afterpulsing
 - During PMT calibrations the probability of afterpulsing was measured
 - The results of this calibration were not folded into the pulse/event classification algorithm for high energy data
- Photo-ionization of the grids and impurities in the xenon

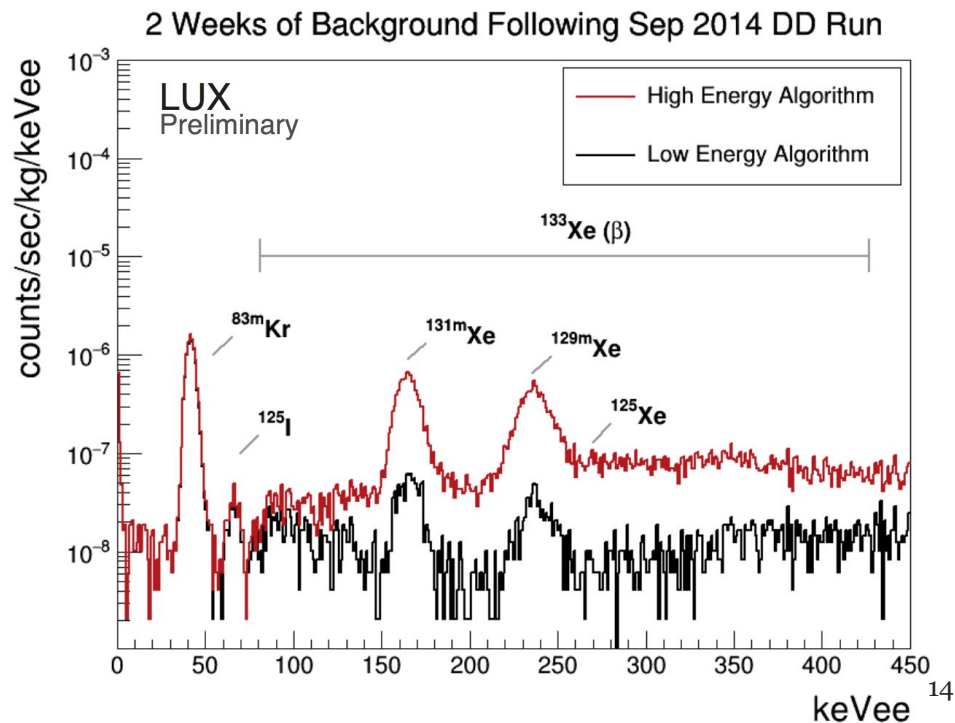
Following Large S2s:

- Electron trains
 - Trails of electrons caused by:
 - Photo-ionization of impurities
 - Emission of thermalized electrons not emitted in the primary S2 signal
 - Etc.

Spurious Pulse Classification Algorithm



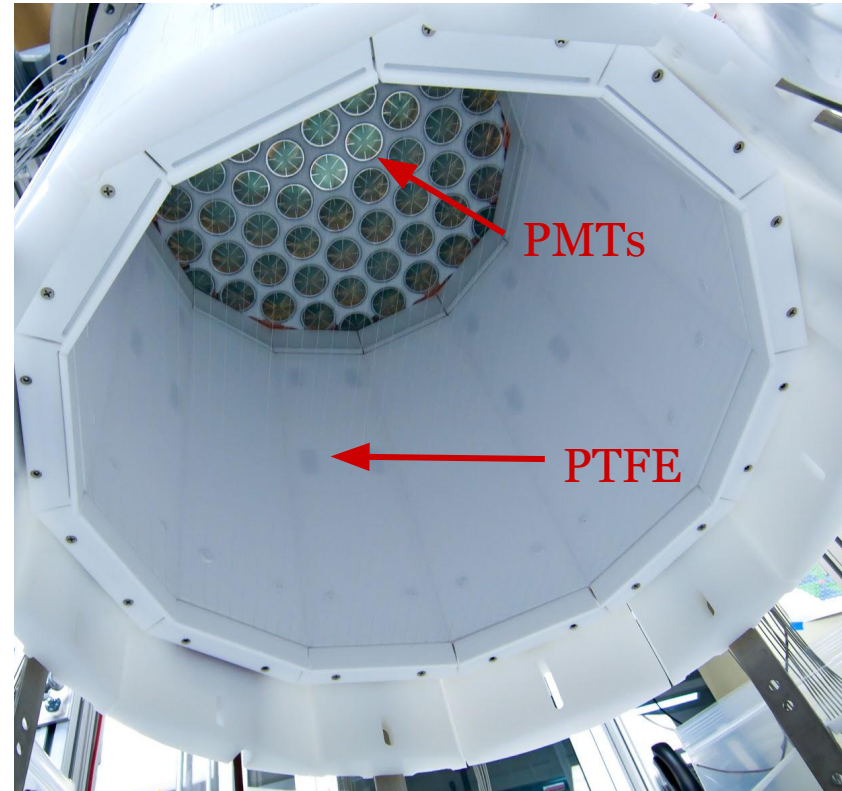
- Above ~ 45 keVee, the detector effects described on previous slide can introduce spurious pulses that have the topology of an S1 or an S2
- A new algorithm was developed to identify and classify spurious pulses in high energy events
- In the $^{131\text{m}}\text{Xe}$ data, the resulting acceptance of single S1, S2 events increases by a factor of 10 (efficiency of $\sim 100\%$)
- **Does not affect existing publications**



Simulation of Neutron Backgrounds

- PMTs
 - Neutrons from (α ,n) from ^{238}U -chain α 's
 - Neutrons from (α ,n) from ^{232}Th -chain α 's
 - Neutrons from ^{235}U -chain fission
- PTFE
 - Neutrons from (α ,n) from ^{210}Po (^{238}U late) chain α 's
- LUXSim was used to simulate energy depositions and libNEST was used to simulate the detector response
- Applied relevant data quality cuts
- Results (WIMP search ROI during WS-2014/16, 332 live days)
 - 0.16 events from PMTs
 - 0.016 events from PTFE

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View of LUX TPC from below

Conclusion

- Short lived activation products from DD neutron generator (**including ^{125}I with effective decay constant 5.4 ± 0.4 days**)
- ^{210}Pb on the detector wall and in the fiducial volume (**1600 day $t_{1/2}$ for leaching from the walls into the fiducial volume**)
- Detector effects, such as PMT afterpulsing, photoionization of grids and impurities, and electron trains (**10x acceptance increase for $^{131\text{m}}\text{Xe}$**)
- Neutron background from PMTs (**0.16 events during WS-2014/16**) and PTFE (**0.016 events during WS-2014/16**)

Thanks go to: the LUX Collaboration

Special thanks go to: Shaun Alsum, Vetri Velan, Sergey Burdin, Elizabeth Boulton, Rachel Mannino, Quentin Riffard, Scott Kravitz, Evan Pease, Cláudio Pascoal, and Kevin Lesko