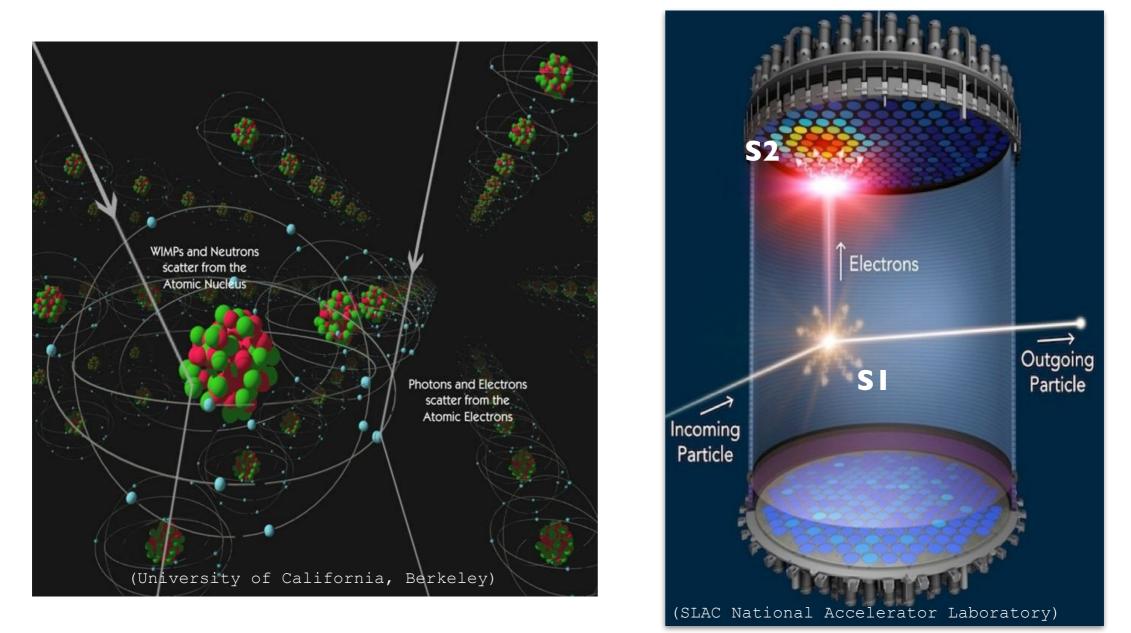
Projected WIMP sensitivity of the LUX-ZEPLIN dark matter experiment

UCLA DM 2018, 20-23rd February Jim Dobson, University College London

WIMP search with a LXe-TPC



Look for low energy WIMP-induced nuclear recoils



Requires: large target mass + low energy threshold + low background

LUX-ZEPLIN (LZ)

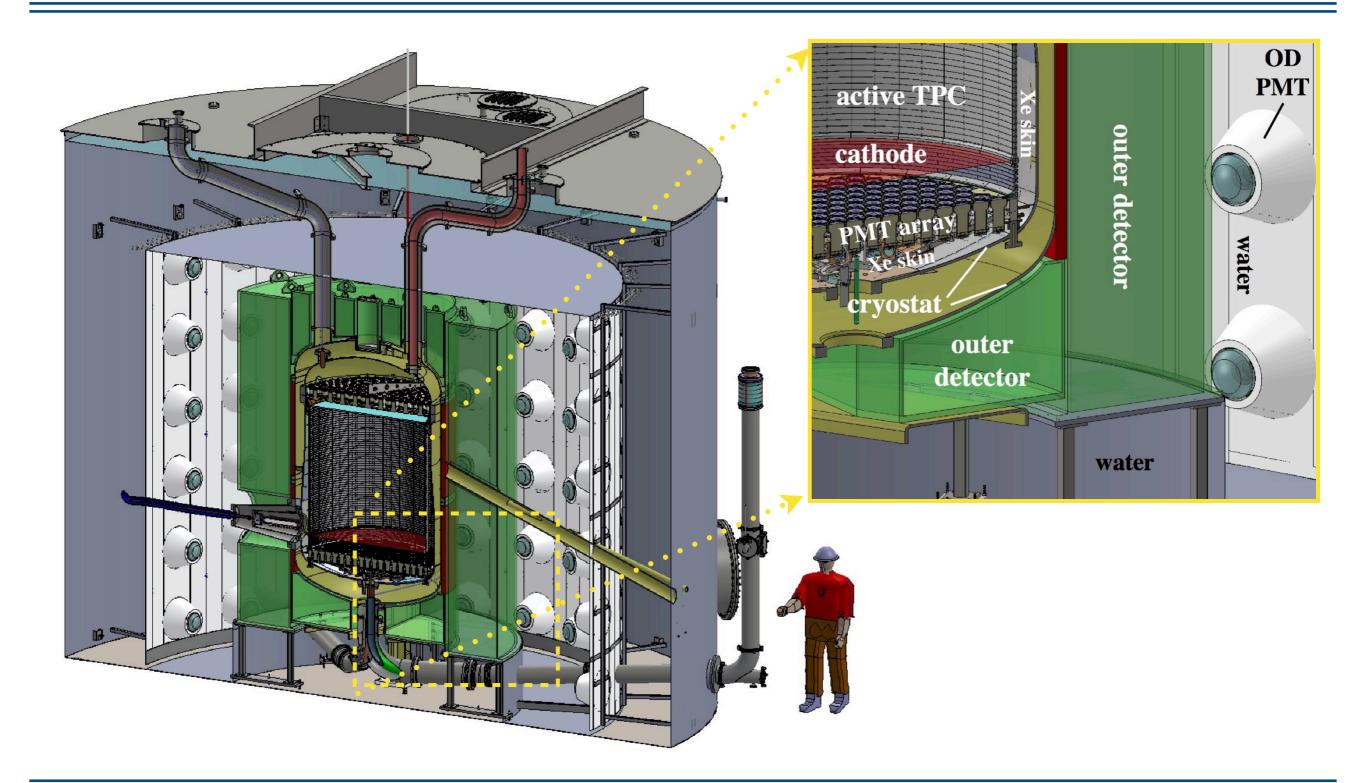


- LXe-TPC: ×50 scale up of LUX
- I mile underground (4300 m w.e.) at SURF
- Underground installation 2019
- Physics data taking 2020

LZ Total mass - 10 T WIMP Active Mass - 7 T WIMP Fiducial Mass - 5.6

LUX

A discovery instrument



Background strategy



• Xenon purification to remove ⁸⁵Kr and ³⁹Ar:

- Dedicated facility at SLAC: based on LUX demonstrated techniques
- Final natKr/Xe 0.015 ppt (g/g)

• Extensive radio-assay campaign for detector materials

- γ-screening, ICP-MS, NAA
- ~1000 assays so far, ~1000 to go

• Strict surface cleanliness program:

- Assembly in dedicated Rn-reduced cleanroom
 - Dust < 500 ng/cm² on all LXe wetted surfaces
 - Rn-daughter plate on TPC walls < 0.5 mBq/m²

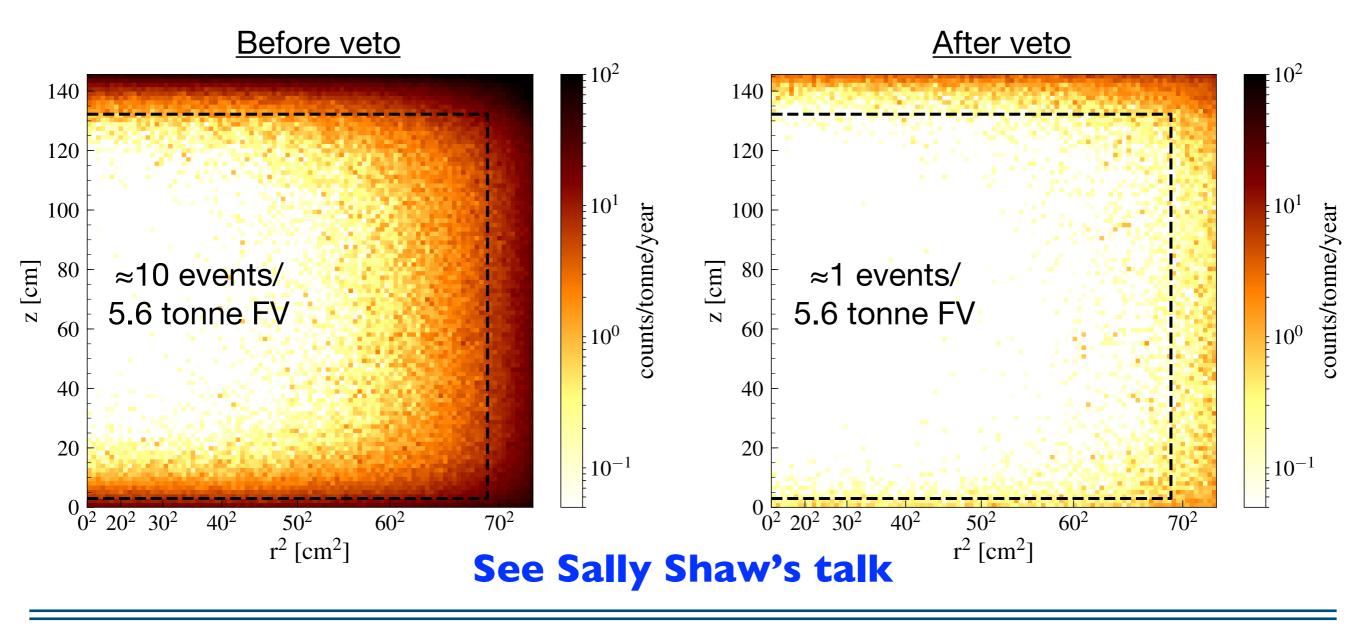
• Active veto to suppress and characterise backgrounds

See Hugh Lippincott's talk

Active veto system



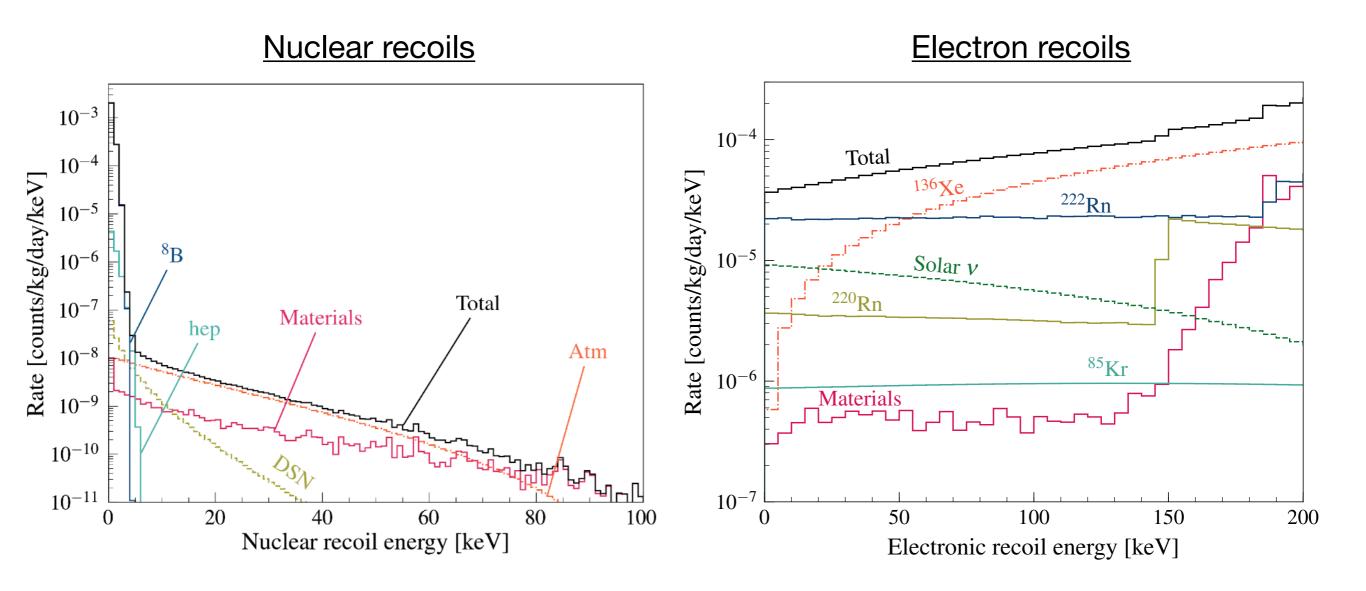
- WIMP-like nuclear recoil backgrounds in 6-30 keV region of interest
- Before/after application of outer detector + skin vetos



Projected background rates



- Counts/kg/day/keV in 5.6 tonne fiducial volume
- Single scatter events with no veto signal



Counts/1000 days: WIMP-search ROI

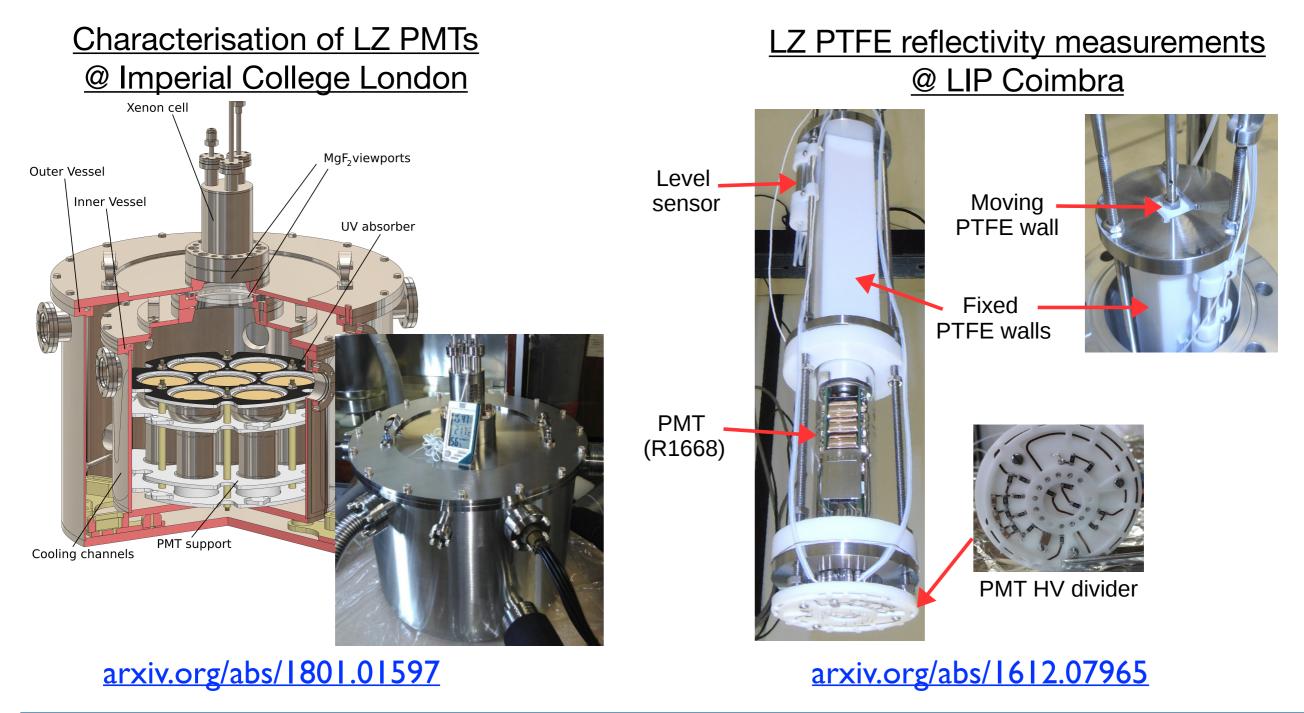


5.6 ton fiducial, 1000 live-days ~1.5 - 6.5 keV, single scatters, no coincident veto

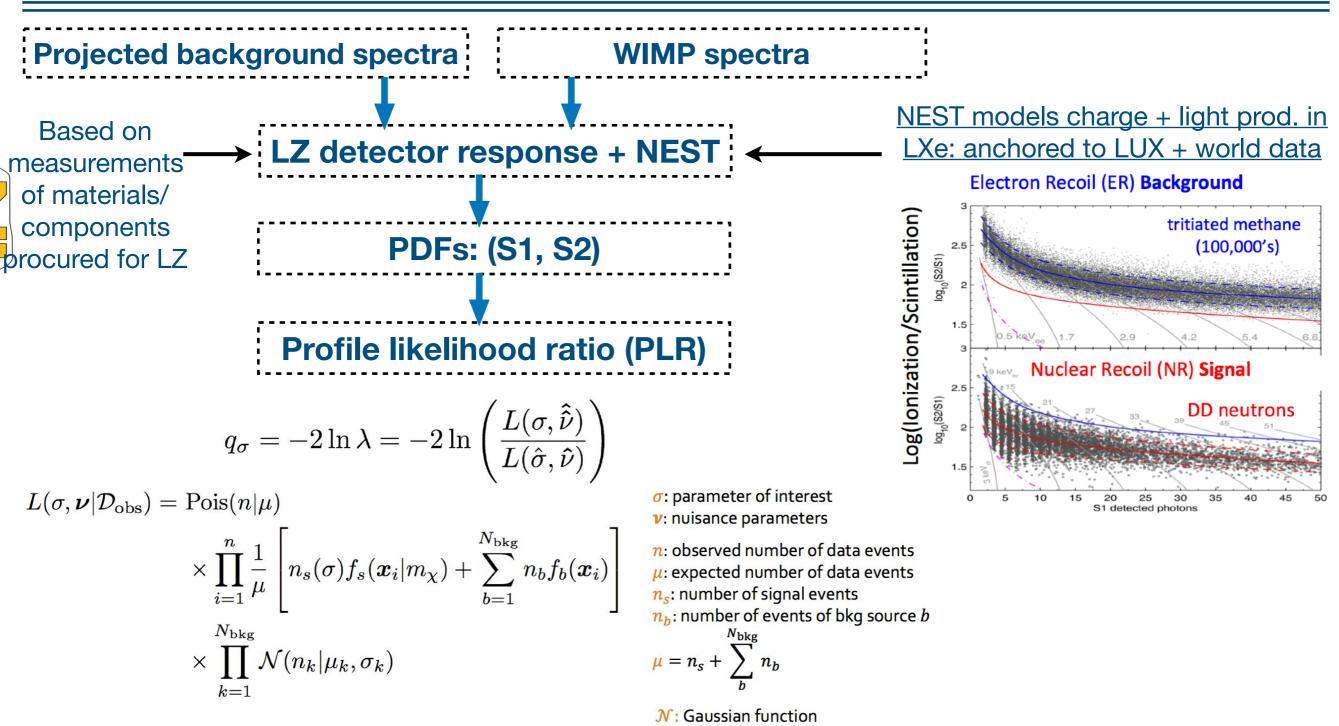
Background Source	ERs	NRs
Detector Components	9	0.07
Dispersed Radionuclides — Rn, Kr, Ar	816	
Laboratory and Cosmogenics	5	0.06
Surface Contamination and Dust	40	0.39
Physics Backgrounds — 2β decay, neutrinos*	322	0.51
Total (after 99.5% discrimination and 50% NR efficiency)	6.4	48
	* n	ot including



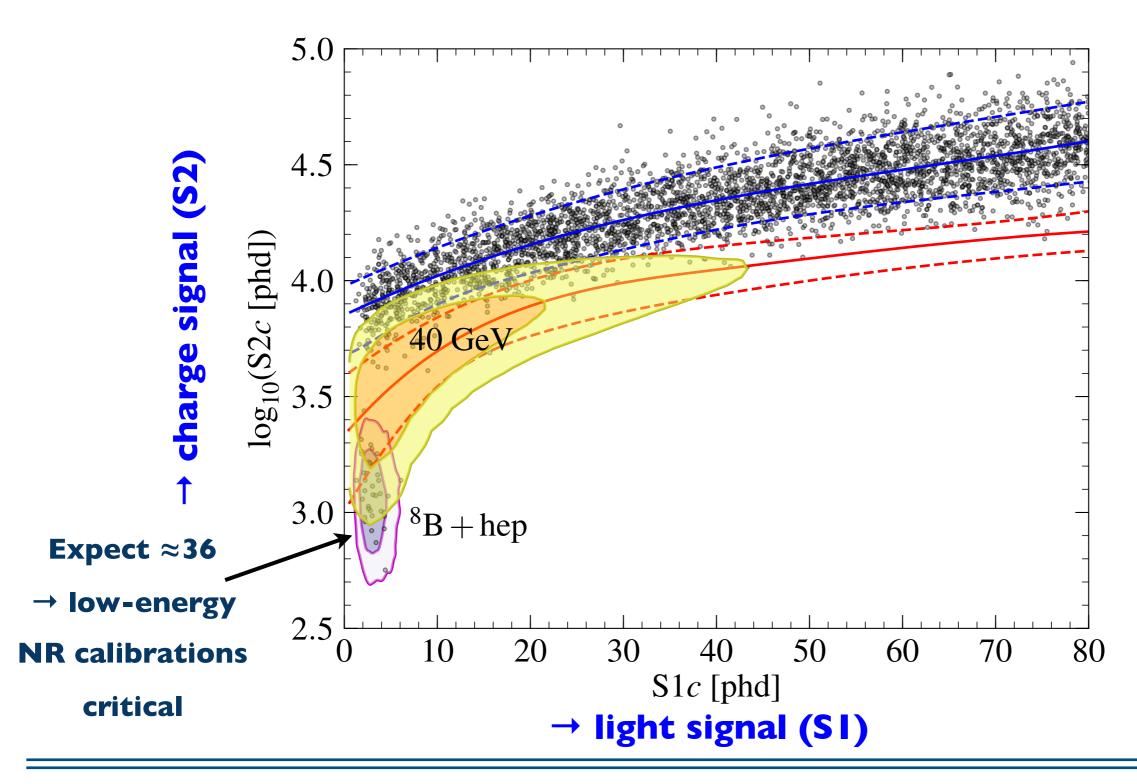
TPC light collection efficiency now 11.9% (cf. 7.5% TDR baseline)



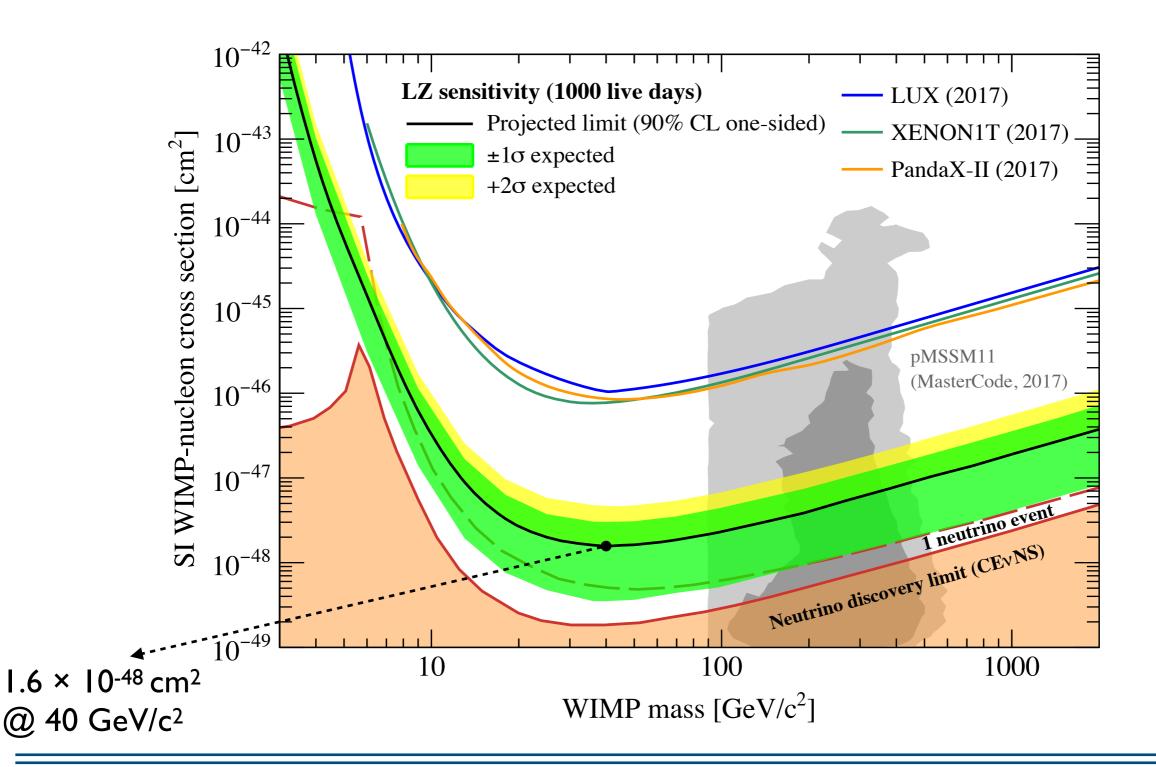
Sensitivity estimates



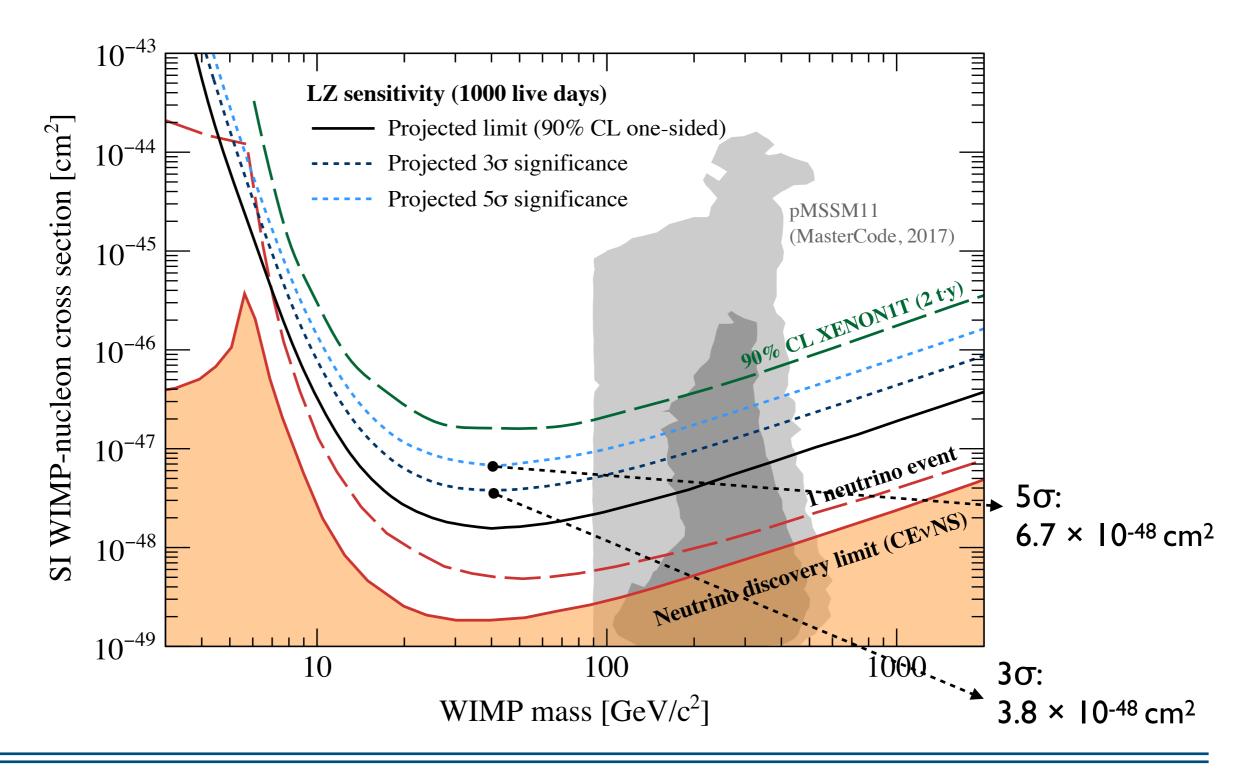
N_{bkg}: number of bckg sources



Projected WIMP sensitivity (1000 live-days 5.6 tonne fiducial)



3σ and 5σ discovery potential (1000 live-days 5.6 tonne fiducial)







• LZ is optimised for WIMP discovery

- 7-tonne active mass + low energy threshold
- Extensive radio-assay and surface cleanliness \rightarrow BG control
- Near-hermetic active veto system suppresses remaining NR backgrounds
- Order of magnitude sensitivity improvement beyond running experiments
 - → exploring new WIMP parameter space
 - $\rightarrow 5\sigma$ discovery potential
- Will have sensitivity to other signals:
 - Astrophysical neutrinos, ALPs, $0\nu\beta\beta$, ...
- Underground installation 2019 → physics data 2020

LZ projected sensitivity paper: arxiv.org/abs/1802.06039

Backups

	Background Source	Mass (kg)	U early (mBq/kg)	U late (mBq/kg)	Th early (mBq/kg)	Th late (mBq/kg)	Co60 (mBq/kg)	K40 (mBq/kg)	n/yr	ER (cts)	NR (cts)
	Detector Components									9	0.07
External Backgrounds	PMT Structures R11410 3" PMTs R8778 2" PMTs R8520 Skin 1" PMTs PMT Bases PMT Cabling TPC PTFE Grid Wires and Rings Field Shaping Rings TPC Sensors and Thermometers PMT Conduits, HX and Tubing HV Conduits and Cables Cryostat Outer Detector	122 92 6 2 3 83 184 96 92 5 215 138 2778 22950	3.89 71.6 138 62.2 359 6.19 0.02 7.39 5.49 21.8 3.18 3.61 2.88 6.13	0.95 3.20 59.4 5.29 78.0 7.06 0.02 2.76 1.14 5.82 0.46 2.30 0.63 4.74	0.72 3.12 16.9 4.91 39.1 1.34 0.03 2.49 0.72 2.29 0.46 0.61 0.48 3.78	0.65 2.99 16.9 4.85 33.4 1.67 0.03 2.28 0.65 1.88 0.56 0.76 0.51 3.71	0.23 2.91 16.2 24.4 1.06 0.01 0.00 10.0 0.00 1.32 1.23 1.23 1.4 0.31 0.33	3.28 15.4 413 337 55.4 6.45 0.12 28.0 2.00 61.0 1.39 2.5 2.62 13.8	13.6 81.8 53.0 53.7 28.9 17.5 22.5 16.3 41.0 6.75 5.87 26.5 323 8061	0.31 1.27 0.05 0.02 0.28 0.89 0.04 3.64 0.65 0.06 0.03 0.02 1.27 0.62	0.002 0.011 0.006 0.005 0.002 0.001 0.006 0.005 0.011 0.001 0.001 0.001 0.018 0.001
	Surface Contamination									40	0.39
	Dust (intrinsic activity, 500 ng/cm2) Plate-out (PTFE panels, 50 nBq/cm2) 210Bi mobility (0.1 µBq/kg) Ion-misreconstruction (50 nBq/cm2) 210Pb (in bulk PTFE, 10 mBq/kg)									0.2 - 40.0 - -	0.05 0.05 - 0.16 0.12
	Laboratory and Cosmogenics									5	0.06
	Laboratory Rock Walls Muon Induced Neutrons Cosmogenic Activation	-								4.6 - 0.2	0.00 0.06 -
Internal Backgrounds	Xenon Contaminants									816	0
	222Rn (1.81 _µ Bq/kg) 220Rn (0.09 _µ Bq/kg) natKr (0.015 ppt g/g) natAr (0.45 ppb g/g)									678 111 24.5 2.5	- - -
	Physics									322	0.51
	136Xe 2vββ Solar neutrinos (pp+7Be+13N) Diffuse supernova neutrinos Atmospheric neutrinos									67 255 0 0	0 0 0.05 0.46
	Total									1192	1.03
Total (with 99.5% ER discrimination, 50% NR efficiency)							5.96 6 .	0.51 . 48			

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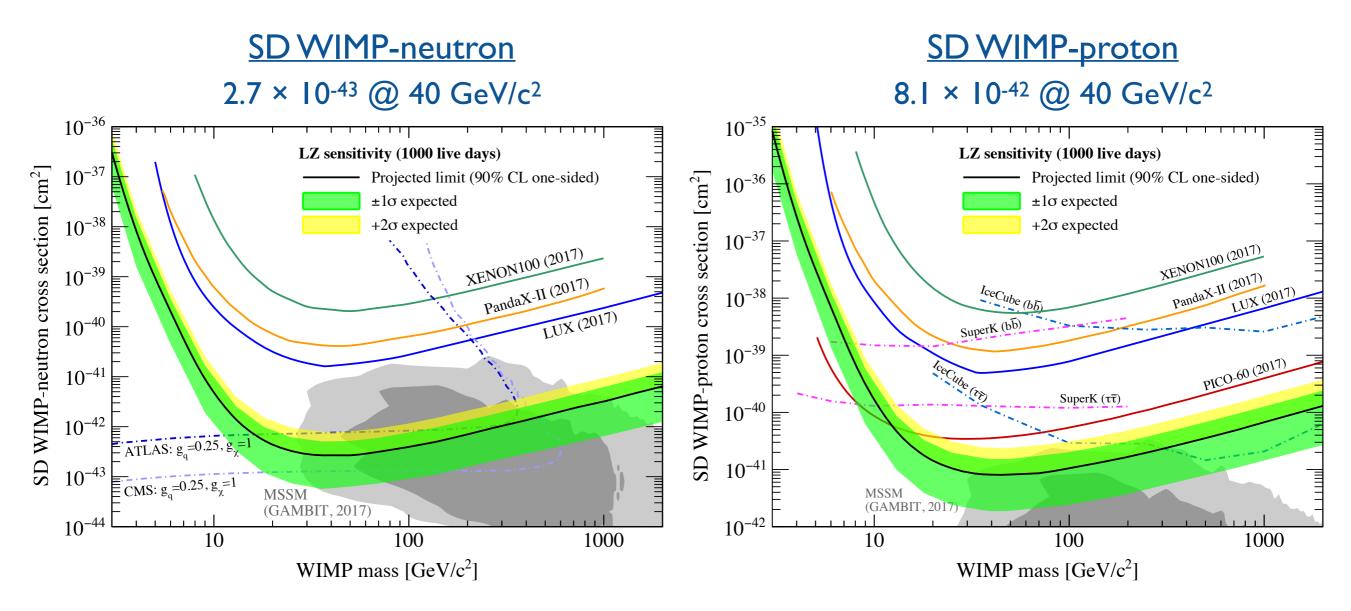
Key detector parameters

 Level L

Detector Parameter	Value
Photon Detection Efficiency (PDE)	_
PDE in liquid (g_1) [phd/ph]	0.119
PDE in gas $(g_{1,gas})$ [phd/ph]	0.102
Single electron size [phd]	83
Effective charge gain (g_2) [phd/e]	79
PTFE-LXe reflectivity	0.977
LXe photon absorption length [m]	100
PMT efficiency at 175 nm	0.269
Other Key Parameters	
Single phe trigger efficiency	0.95
Single phe relative width (Gaussian)	0.38
S1 coincidence level	3-fold
S2 electron extraction efficiency	0.95
Drift field $[V cm^{-1}]$	310
Electron lifetime [µs]	850

Spin-dependent sensitivity

Naturally occurring Xe: ~50% odd neutron isotopes (26.4% 129Xe and 21.2% 131Xe by mass)



Spin structure functions from Klos et al, Phys. Rev. D89, 029901 (2014) + SHM

Sensitivity vs Rn-level

