

Resummation for transverse observables at hadron colliders

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Based on

1604.02191 with E. Re and P. Torrielli
and

1705.09127 with W. Bizon, E. Re, L. Rottoli, and P. Torrielli

+ ongoing work with

W. Bizon, X. Chen, Gehrmann-De Ridder, Gehrmann, Glover, A. Huss,
E. Re, L. Rottoli, and P. Torrielli

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Outline

- Theory precision at colliders:
 - fixed-order vs. all-order perturbation theory
- Factorisation theorems and semi-numerical resummation
- Momentum-space resummation for transverse observables
- Predictions for differential distributions at $N^3\text{LL}+\text{NNLO}$ at the LHC
 - Higgs production
 - Drell-Yan production
- Conclusions

Fixed-order vs. All-order

- Fixed-order calculations of radiative corrections are formulated in a well established way (technically challenging, but well posed problem):
 - compute amplitudes at a given order
 - provide an effective subtraction of IRC divergences
 - compute any IRC-safe observable

$$\Sigma(v) = \int_0^v \frac{1}{\sigma_{\text{Born}}} \frac{d\sigma}{dv'} dv' \sim \overset{\text{LO}}{1} + \overset{\text{NLO}}{\alpha_s} + \overset{\text{NNLO}}{\alpha_s^2} + \dots$$

- All-order calculations are still at an earlier stage of evolution
 - Each different observable has its own type of sensitivity to IRC physics, it is hard to formulate a general method that works for all at a generic perturbative order
 - Higher-order resummations are therefore often formulated in an observable-dependent way, for few well-behaved collider observables

$$\Sigma(v) = \int_0^v \frac{1}{\sigma_{\text{Born}}} \frac{d\sigma}{dv'} dv' \sim e^{\overset{\text{LL}}{\alpha_s^n L^{n+1}} + \overset{\text{NLL}}{\alpha_s^n L^n} + \overset{\text{NNLL}}{\alpha_s^n L^{n-1}} + \dots} \quad v \rightarrow 0$$

Factorisation of the observable

- Factorisation of the amplitude is not enough as the all-order radiation is tangled by the observable

$$\Sigma(v) = \int d\Phi_{\text{rad}} \sum_{n=0}^{\infty} |\mathcal{M}(k_1, \dots, k_n)|^2 \Theta(v - V(k_1, \dots, k_n))$$

- In order to perform an all-order calculation, one needs to **break** the observable too into hard, soft and collinear pieces. This can be done for some observables which treat the radiation rather **inclusively**

- e.g. transverse momentum of a massive singlet

[Parisi, Petronzio '79]
[Collins et al. '85]
[Bozzi et al. '05]
[Becher et al. '10+'12]

$$\delta^{(2)}(\vec{p}_t - (\vec{k}_{t1} + \dots + \vec{k}_{tn})) = \int \frac{d^2\vec{b}}{4\pi^2} e^{-i\vec{b} \cdot \vec{p}_t} \prod_{i=1}^n e^{i\vec{b} \cdot \vec{k}_{ti}},$$

$$\begin{aligned} \frac{d^2\Sigma(p_t)}{d\Phi_B dp_t} &= \sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \int b db p_t J_0(p_t b) \mathbf{f}^T(b_0/b) \mathbf{C}_{N_1}^{c_1; T}(\alpha_S(b_0/b)) H_{\text{CSS}}(M) \mathbf{C}_{N_2}^{c_2}(\alpha_S(b_0/b)) \mathbf{f}(b_0/b) \\ &\times \exp \left\{ - \sum_{\ell=1}^2 \int_{b_0/b}^M \frac{dk_t}{k_t} \mathbf{R}'_{\text{CSS}, \ell}(k_t) \right\}. \end{aligned}$$

[Catani, Grazzini '11][Catani et al. '12][Gehrmann, Luebbert, Yang '14][Davies, Stirling '84]
[De Florian, Grazzini '01][Becher, Neubert '10][Li, Zhu '16][Vladimirov '16]

Eluding observable factorisation

- Factorisation is a powerful tool, but limited to observables that have a simple analytic expression in the relevant limits or do not mix soft and collinear radiation (e.g. jet rates)
- Ultimately, we want to use the modern knowledge of IRC dynamics to make more accurate generators. At present a general framework to assess the accuracy of Parton Showers is missing
 - It is of primary importance to formulate a link between higher-order resummation and PS
- Can we devise a formulation without a factorisation formula ?
 - **recursive IRC safety**: simple set of criteria for the observable that allows one to formulate the resummation at NLL for global observables without the need for an explicit factorisation. [Banfi, Salam, Zanderighi '01-'04]
 - Most of modern global observables fall into this category.
 - The method can be reformulated and extended at higher logarithmic orders
 - [Banfi, McAslan, PM, Zanderighi '14-'16]
 - [PM, Re, Torrielli '16]
 - [Bizon, PM, Re, Rottoli, Torrielli '17]

A case study: transverse observables

- **Transverse and inclusive** observables in colour-singlet production offer a clean experimental and theoretical environment for precision physics:

$$V(\{\vec{p}\}, k) \equiv V(k) = d_\ell g_\ell(\phi) \left(\frac{k_t}{M} \right)^a$$

$$V(\{\vec{p}\}, k_1, \dots, k_n) = V(\{\vec{p}\}, k_1 + \dots + k_n)$$

- **SM measurements** (e.g. W, Z, photon,...): parton distributions, strong coupling, W mass,...
 - sensitivity to non-perturbative effects (hadronisation, intrinsic k_t) only through transverse recoil
 - very little/no sensitivity to multi-parton interactions
- **BSM measurements/constraints** (e.g. Higgs): light/heavy NP, Yukawa couplings,...
- **Theoretically interesting:**
 - clean environment to **test/calibrate exclusive generators** against high perturbative orders
 - **Two mechanisms compete** in the $p_t \rightarrow 0$ limit:
 - Sudakov (exponential) suppression when $k_{ti} \sim p_t \ll M$
 - Azimuthal cancellations (power suppression, dominant) when $p_t \ll k_{ti} \ll M$

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Can we build a more exclusive solution in momentum space ?

See also work in [Ebert, Tackmann '16][Kang, Lee, Vaidya '17]

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Direct space: virtual corrections

- Write all-order cross section as ($V(\{\tilde{p}\}, k_1, \dots, k_n) = |\vec{k}_{t1} + \dots + \vec{k}_{tn}|$)

$$\Sigma(v) = \int d\Phi_B \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \int \prod_{i=1}^n [dk_i] |M(\tilde{p}_1, \tilde{p}_2, k_1, \dots, k_n)|^2 \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_n))$$

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All-order form factor
e.g. [Dixon, Magnea, Sterman '08]

$$\mathcal{V}(\Phi_B) = \frac{\text{Tree} + \text{1-loop} + \text{2-loop} + \dots}{\text{Tree}^2}$$

Direct space: real radiation

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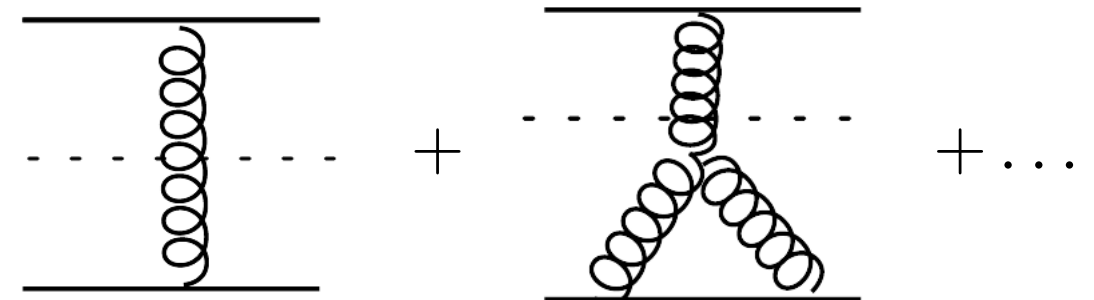
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Real emissions

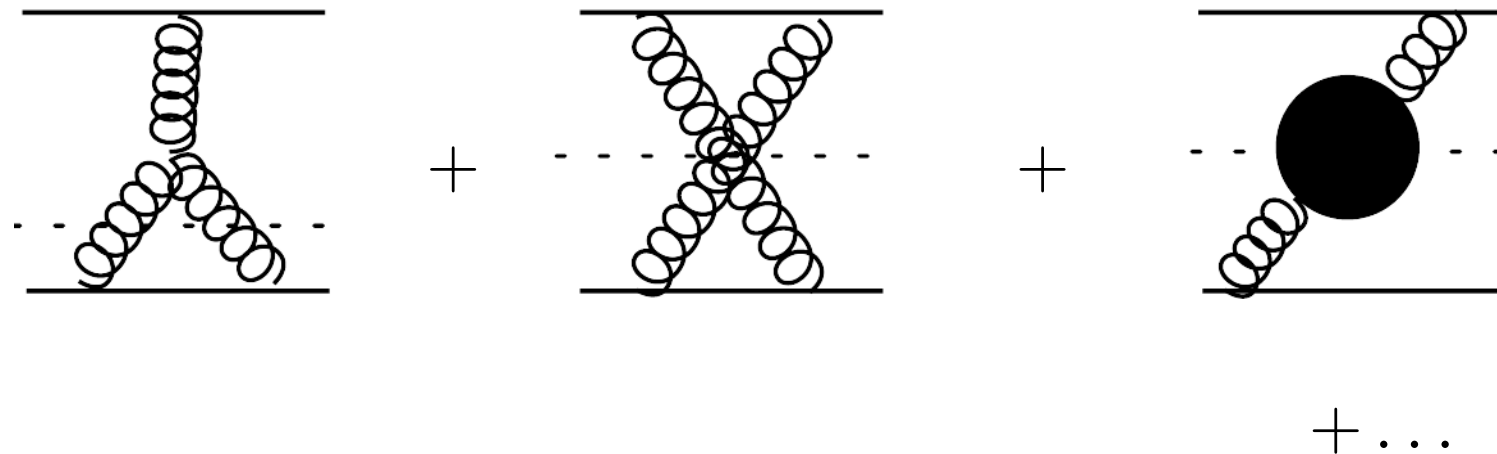
- Logarithmic counting: we need a logarithmic hierarchy in the squared amplitudes
(resummation means iteration of lower-order amplitudes)

e.g.
soft radiation (one log down in hard-collinear case)

$$|\tilde{M}(k_a)|^2 = \frac{|M(\tilde{p}_1, \tilde{p}_2, k_a)|^2}{|M_B(\tilde{p}_1, \tilde{p}_2)|^2} = |M(k_a)|^2$$



$$|\tilde{M}(k_a, k_b)|^2 = \frac{|M(\tilde{p}_1, \tilde{p}_2, k_a, k_b)|^2}{|M_B(\tilde{p}_1, \tilde{p}_2)|^2} - \frac{1}{2!} |M(k_a)|^2 |M(k_b)|^2 \rightarrow$$



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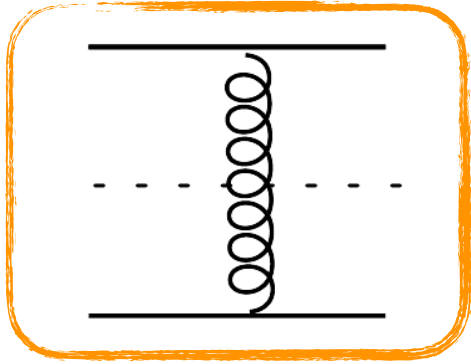
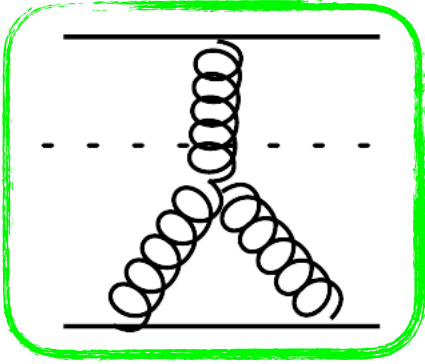
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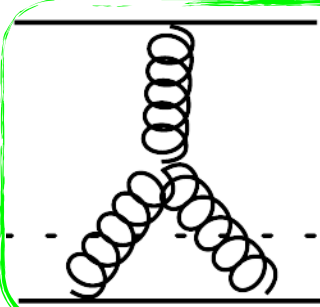
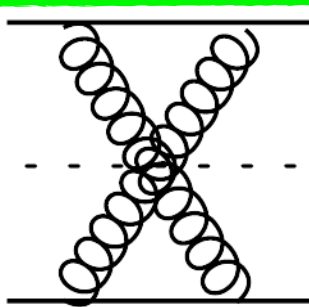
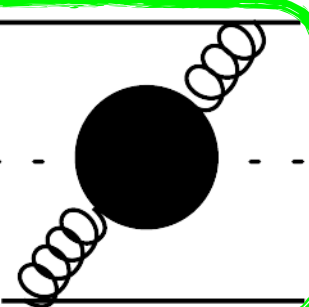
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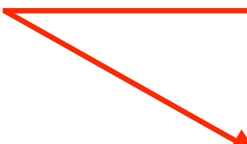
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 $+$

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All-order subtraction of IRC singularities

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Subtraction of the
IRC poles and
computation of
the observable

All-order subtraction of IRC singularities

- Subtraction of the IRC poles between $\sum_{n=0}^{\infty} \int \prod_{i=1}^n [dk_i] |M(\tilde{p}_1, \tilde{p}_2, k_1, \dots, k_n)|^2$ and $\mathcal{V}(\Phi_B)$:
 - introduce a phase-space resolution scale (**slicing parameter**) $Q_0 = \epsilon k_{t1}$
 - real correlated blocks with total transverse momentum $k_{ti} < \epsilon k_{t1}$ (**unresolved**) do not modify the observable, and can be *ignored* in the measurement function
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$$\begin{aligned}
 \sum_{n=0}^{\infty} |M(\tilde{p}_1, \tilde{p}_2, k_1, \dots, k_n)|^2 &\longrightarrow |M_B(\tilde{p}_1, \tilde{p}_2)|^2 \\
 &\times \sum_{n=0}^{\infty} \frac{1}{n!} \left\{ \prod_{i=1}^n \left(|M(k_i)|^2 + \int [dk_a][dk_b] |\tilde{M}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \right. \right. \\
 &\left. \left. + \int [dk_a][dk_b][dk_c] |\tilde{M}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots \right) \right\}
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$$\prod_{i=1}^n \int [dk_i] \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \frac{1}{n!} \left\{ \prod_{i=1}^n \left(|M(k_i)|^2 + \int [dk_a][dk_b] |\tilde{M}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \right. \right. \\ \left. \left. + \int [dk_a][dk_b][dk_c] |\tilde{M}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots \right) \Theta(\epsilon k_{t1} - k_{ti}) \right\}$$

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$$\begin{aligned}
 & \prod_{i=1}^n \int [dk_i] \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \frac{1}{n!} \left\{ \prod_{i=1}^n \left(|M(k_i)|^2 + \int [dk_a][dk_b] |\tilde{M}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \right. \right. \\
 & \quad \left. \left. + \int [dk_a][dk_b][dk_c] |\tilde{M}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots \right) \Theta(\epsilon k_{t1} - k_{ti}) \right\} \\
 & \quad \propto \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(\epsilon k_{t1})} R'(k_{t1})
 \end{aligned}$$

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$$R(\epsilon k_{t1}) \equiv \sum_{\ell=1}^2 \int_{\epsilon k_{t1}}^M \frac{dk_t}{k_t} R'_\ell(k_t) = \sum_{\ell=1}^2 \int_{\epsilon k_{t1}}^M \frac{dk_t}{k_t} \left(A_\ell(\alpha_s(k_t)) \ln \frac{M^2}{k_t^2} + B_\ell(\alpha_s(k_t)) \right)$$

Anomalous dimensions
start differing from b-
space ones at N³LL

$$\begin{aligned} & \prod_{i=1}^n \int [dk_i] \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \frac{1}{n!} \left\{ \prod_{i=1}^n \left(|M(k_i)|^2 + \int [dk_a][dk_b] |\tilde{M}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \right. \right. \\ & \quad \left. \left. + \int [dk_a][dk_b][dk_c] |\tilde{M}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots \right) \Theta(\epsilon k_{t1} - k_{ti}) \right\} \\ & \propto \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(\epsilon k_{t1})} R'(k_{t1}) \end{aligned}$$

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$$\hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v) = \left[\mathbf{C}_{N_1}^{c_1; T}(\alpha_s(\mu_0)) H(\mu_R) \mathbf{C}_{N_2}^{c_2}(\alpha_s(\mu_0)) \right] \int_0^M \frac{dk_{t1}}{k_{t1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} \times e^{-\mathbf{R}(\epsilon k_{t1})} \exp \left\{ - \sum_{\ell=1}^2 \left(\int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \frac{\alpha_s(k_t)}{\pi} \mathbf{\Gamma}_{N_\ell}(\alpha_s(k_t)) + \int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \mathbf{\Gamma}_{N_\ell}^{(C)}(\alpha_s(k_t)) \right) \right\}$$

Sudakov radiator:
integral of single
inclusive block.

DGLAP anomalous dims

RGE evolution of
coeff. functions

All-order subtraction of IRC singularities

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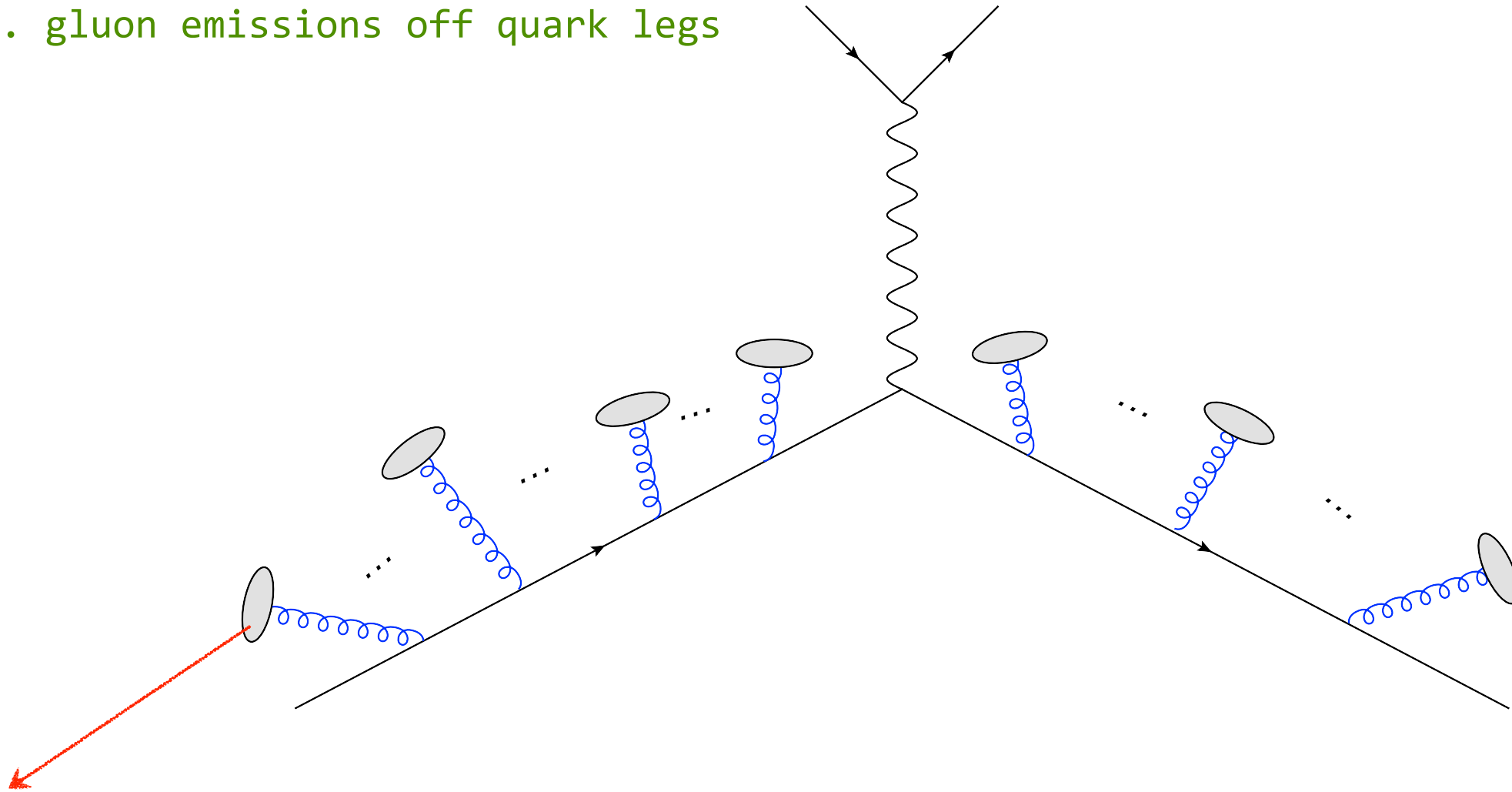
$$\begin{aligned}
 \hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v) = & \left[\mathbf{C}_{N_1}^{c_1; T}(\alpha_s(\mu_0)) H(\mu_R) \mathbf{C}_{N_2}^{c_2}(\alpha_s(\mu_0)) \right] \int_0^M \frac{dk_{t1}}{k_{t1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi} \quad \text{DGLAP anomalous dims} \\
 & \times e^{-\mathbf{R}(\epsilon k_{t1})} \exp \left\{ - \sum_{\ell=1}^2 \left(\int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \frac{\alpha_s(k_t)}{\pi} \mathbf{\Gamma}_{N_\ell}(\alpha_s(k_t)) + \int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \mathbf{\Gamma}_{N_\ell}^{(C)}(\alpha_s(k_t)) \right) \right\} \quad \text{RGE evolution of coeff. functions} \\
 & \sum_{\ell_1=1}^2 \left(\mathbf{R}'_{\ell_1}(k_{t1}) + \frac{\alpha_s(k_{t1})}{\pi} \mathbf{\Gamma}_{N_{\ell_1}}(\alpha_s(k_{t1})) + \mathbf{\Gamma}_{N_{\ell_1}}^{(C)}(\alpha_s(k_{t1})) \right) \\
 & \times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} \sum_{\ell_i=1}^2 \left(\mathbf{R}'_{\ell_i}(k_{ti}) + \frac{\alpha_s(k_{ti})}{\pi} \mathbf{\Gamma}_{N_{\ell_i}}(\alpha_s(k_{ti})) + \mathbf{\Gamma}_{N_{\ell_i}}^{(C)}(\alpha_s(k_{ti})) \right) \\
 & \times \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})),
 \end{aligned}$$

- compute **resolved** (reals only) in 4 dim. with $\epsilon \rightarrow 0$ (**MC events !**)

Physical picture: MC generator

- This is, essentially, a *quasi-exclusive generator* with higher logarithmic accuracy

➡ e.g. gluon emissions off quark legs



$$\begin{aligned}
 |M(k_i)|^2 &+ \int [dk_a][dk_b] |\tilde{M}(k_a, k_b)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} - \vec{k}_{ti}) \delta(Y_{ab} - Y_i) \\
 &+ \int [dk_a][dk_b][dk_c] |\tilde{M}(k_a, k_b, k_c)|^2 \delta^{(2)}(\vec{k}_{ta} + \vec{k}_{tb} + \vec{k}_{tc} - \vec{k}_{ti}) \delta(Y_{abc} - Y_i) + \dots
 \end{aligned}$$

Small transverse momentum limit

- CSS result recovered by simply transforming observable into b-space and integrating over radiation (see backup material)
- Clear physical picture of the dynamics of azimuthal cancellations at small transverse momentum

e.g. NLL with $\mathcal{L}(k_{t1}) = 1$ for simplicity

$$\frac{d^3\Sigma(p_t)}{d^2p_t d\Phi_B} = \sigma^{(0)}(\Phi_B) \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} R'(k_{t1}) \int d\mathcal{Z}[\{R'(k_{t1}), k_i\}] \delta^{(2)}(\vec{p}_t - \sum_{i=1}^{n+1} \vec{k}_{ti})$$

- Transition from exponential to a power-like suppression at small transverse momentum

$$\frac{d^2\Sigma(p_t)}{dp_t d\Phi_B} \simeq 4p_t \sigma^{(0)}(\Phi_B) \int_{\Lambda_{\text{QCD}}}^M \frac{dk_{t1}}{k_{t1}^3} e^{-R(k_{t1})} \simeq 2p_t \sigma^{(0)}(\Phi_B) \left(\frac{\Lambda_{\text{QCD}}^2}{M^2} \right)^{\frac{16}{25} \ln \frac{41}{16}}$$

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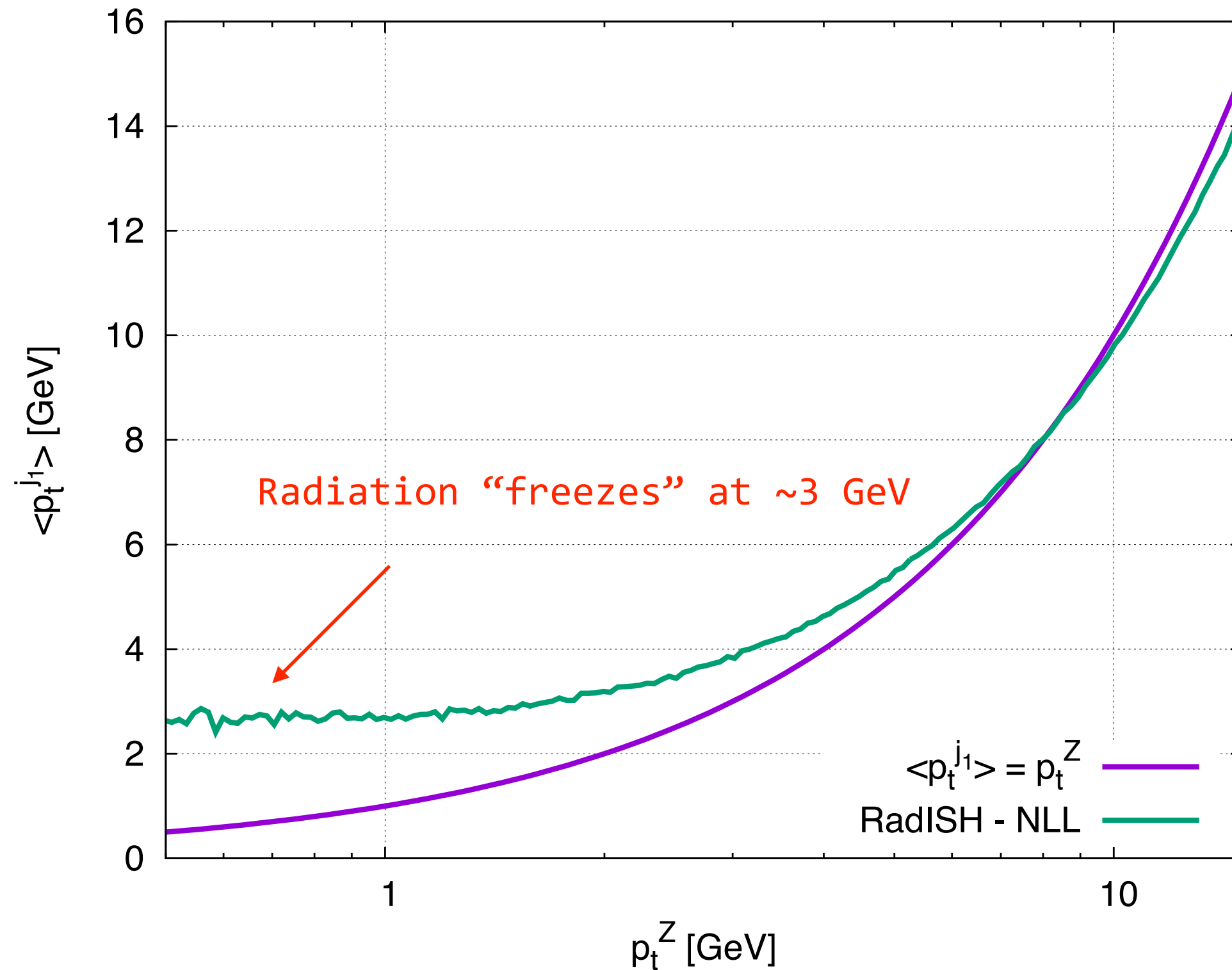
Random azimuthal orientation of momenta
leads to scaling $\propto 1/k_{t1}^2$

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Small transverse momentum limit

➡ e.g. Z production at 14 TeV



Matching to Fixed Order

- Implementation in a MC code (**RadISH**) up to N³LL
 - fully differential in Born kinematics
 - matching to fixed order cumulative distribution, e.g. Higgs:

[Anastasiou et al. '15-'16]

[Boughezal et al. '15]

[Caola et al. '15]

[Chen et al. '16]

$$\sigma_{pp \rightarrow H}^{\text{N}^3\text{LO}} - \Sigma_{1\text{-jet}}^{\text{NNLO}}(p_t^H)$$

- Additive vs. multiplicative schemes

OLD CHOICE :

$$\Sigma_{\text{MAT}}(p_t) = (\Sigma_{\text{RES}}(p_t))^Z \frac{\Sigma_{\text{FO}}(p_t)}{(\Sigma_{\text{EXP}}(p_t))^Z}$$

$$Z = \left(1 - \left(\frac{p_t}{Q_{\text{match}}}\right)\right)^h \Theta(Q_{\text{match}} - p_t)$$

R - SCHEME :

$$\Sigma_{\text{MAT}}(p_t) = \Sigma_{\text{RES}}(p_t) + \Sigma_{\text{FO}}(p_t) - \Sigma_{\text{EXP}}(p_t)$$

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Higher-order (in a logarithmic sense) constants from FO in the multiplicative scheme. No extra parameters needed

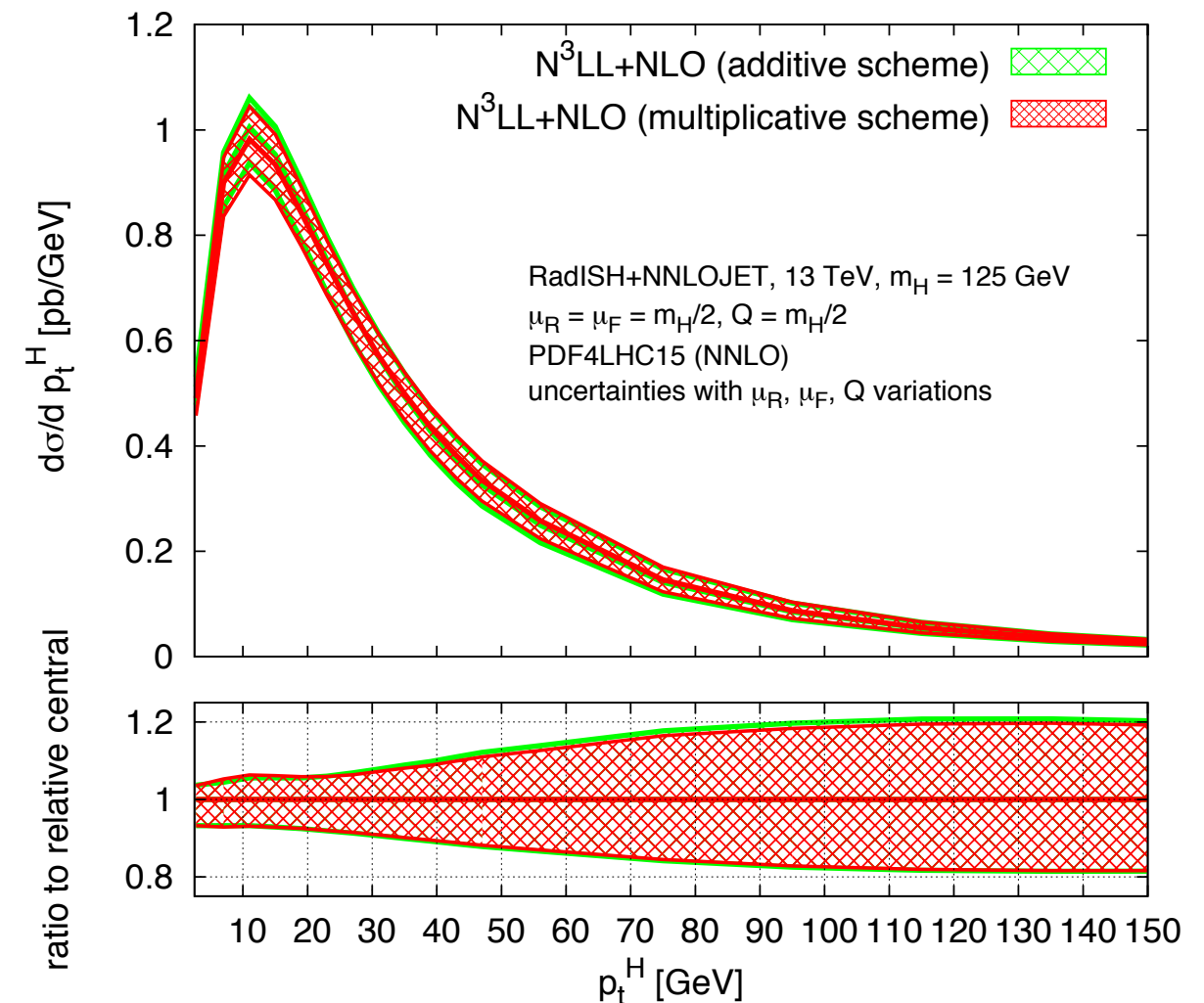
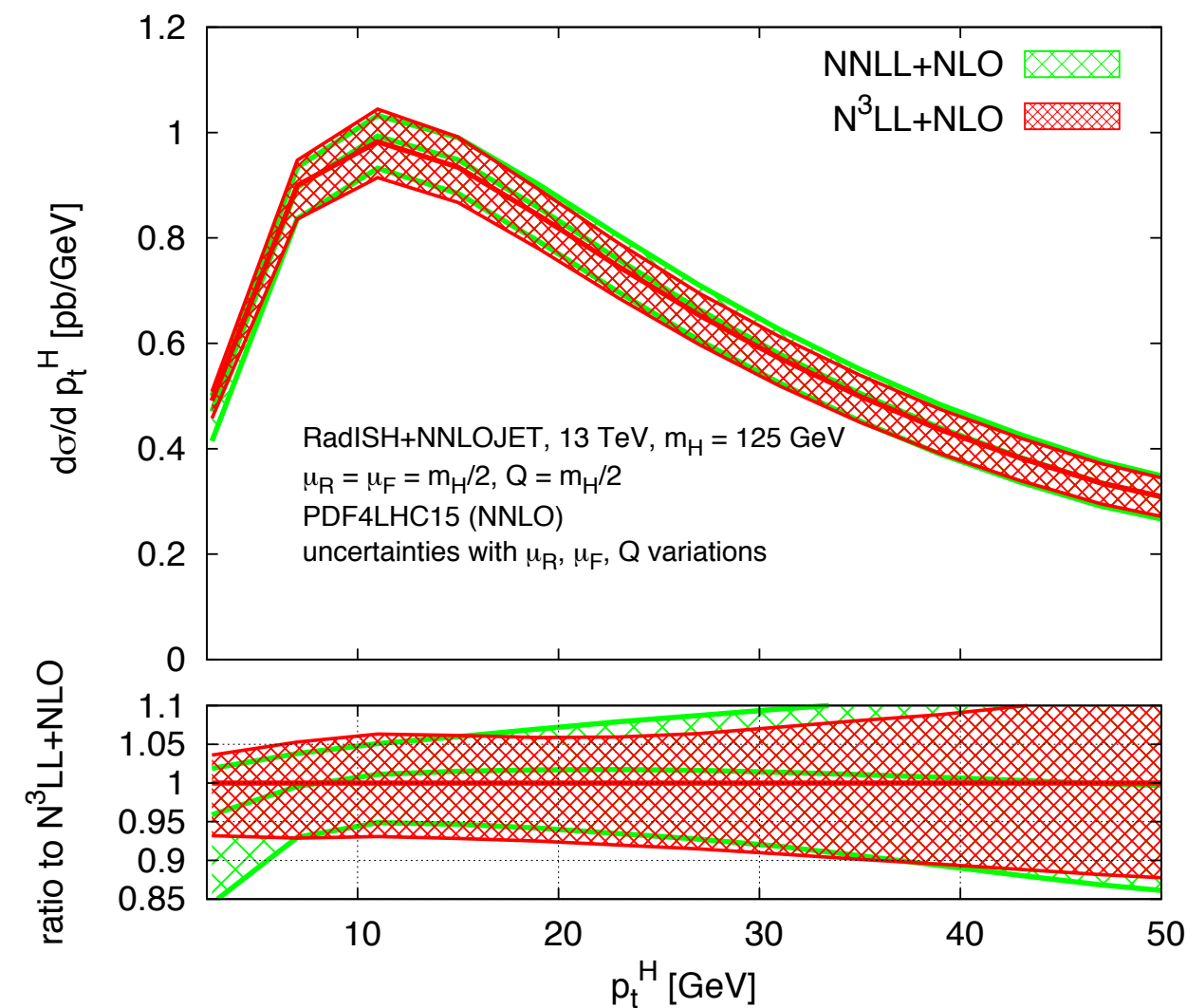


An example: Higgs p_T spectrum

- Implementation in a MC code (**RadISH**) up to N^3LL
 - fully differential in Born kinematics

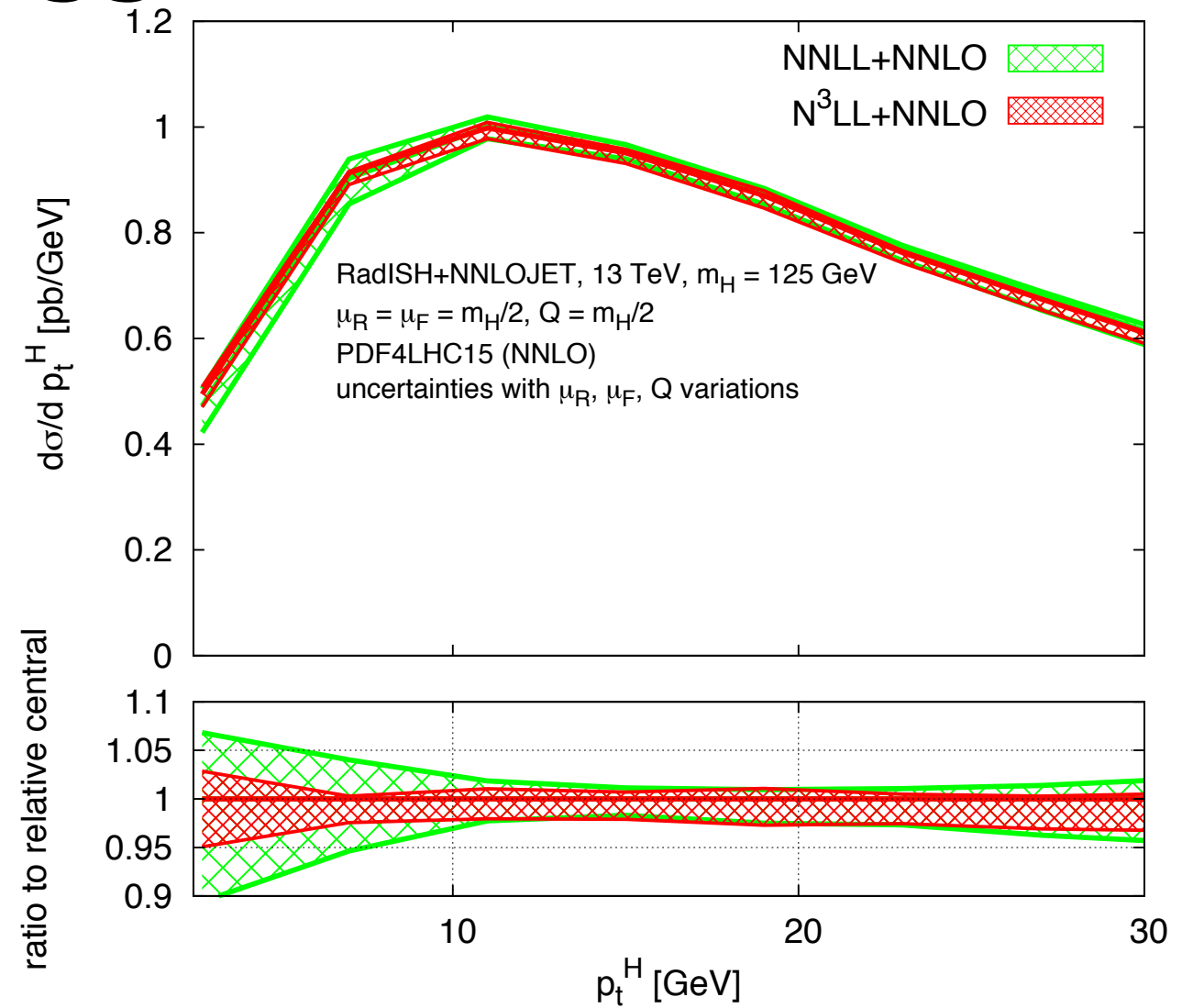
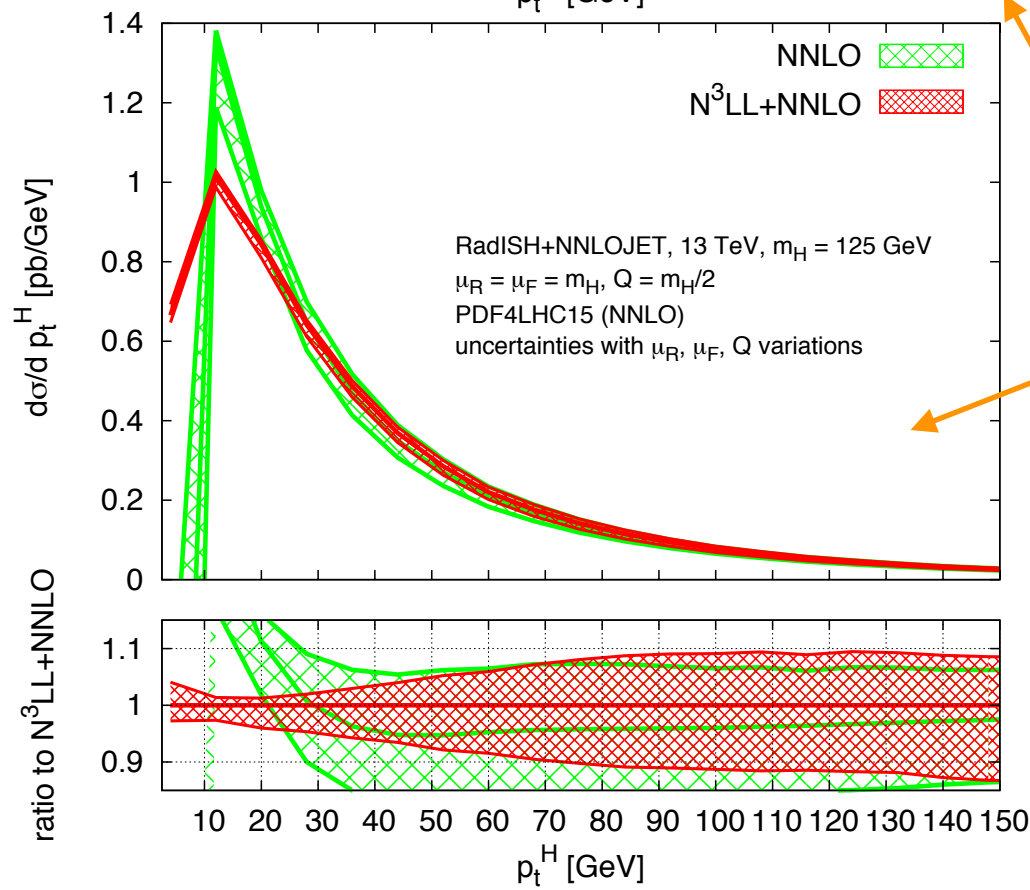
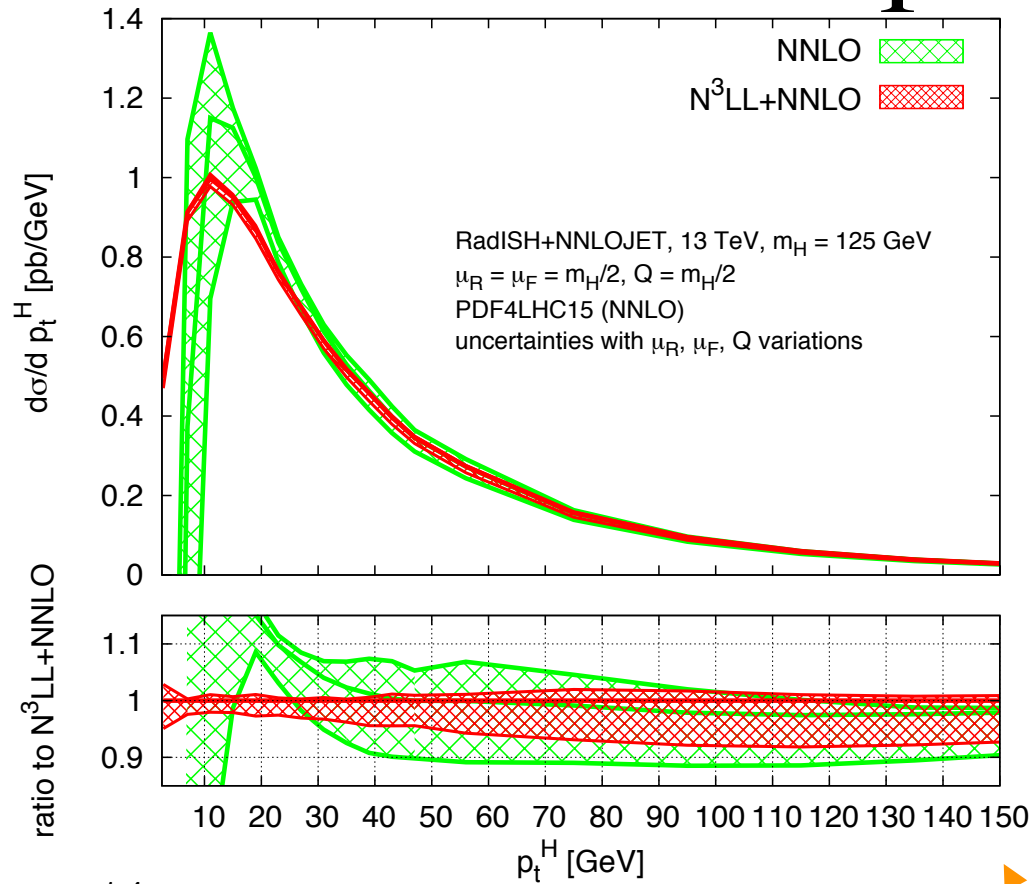
→ N^3LL corrections moderate, reduction of uncertainty at small p_t

→ Good agreement between different matching schemes, choose multiplicative solution at higher order





An example: Higgs p_T spectrum



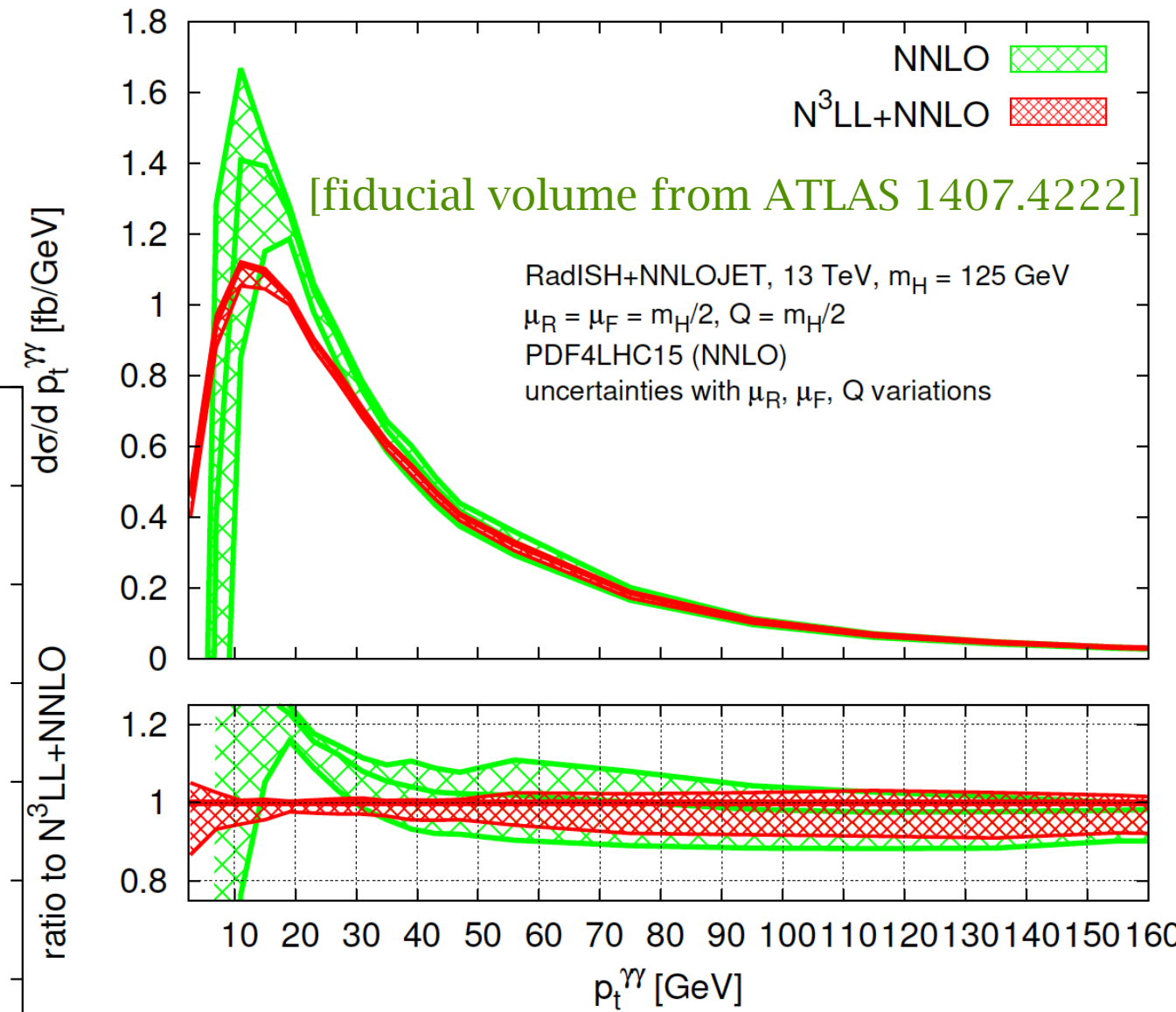
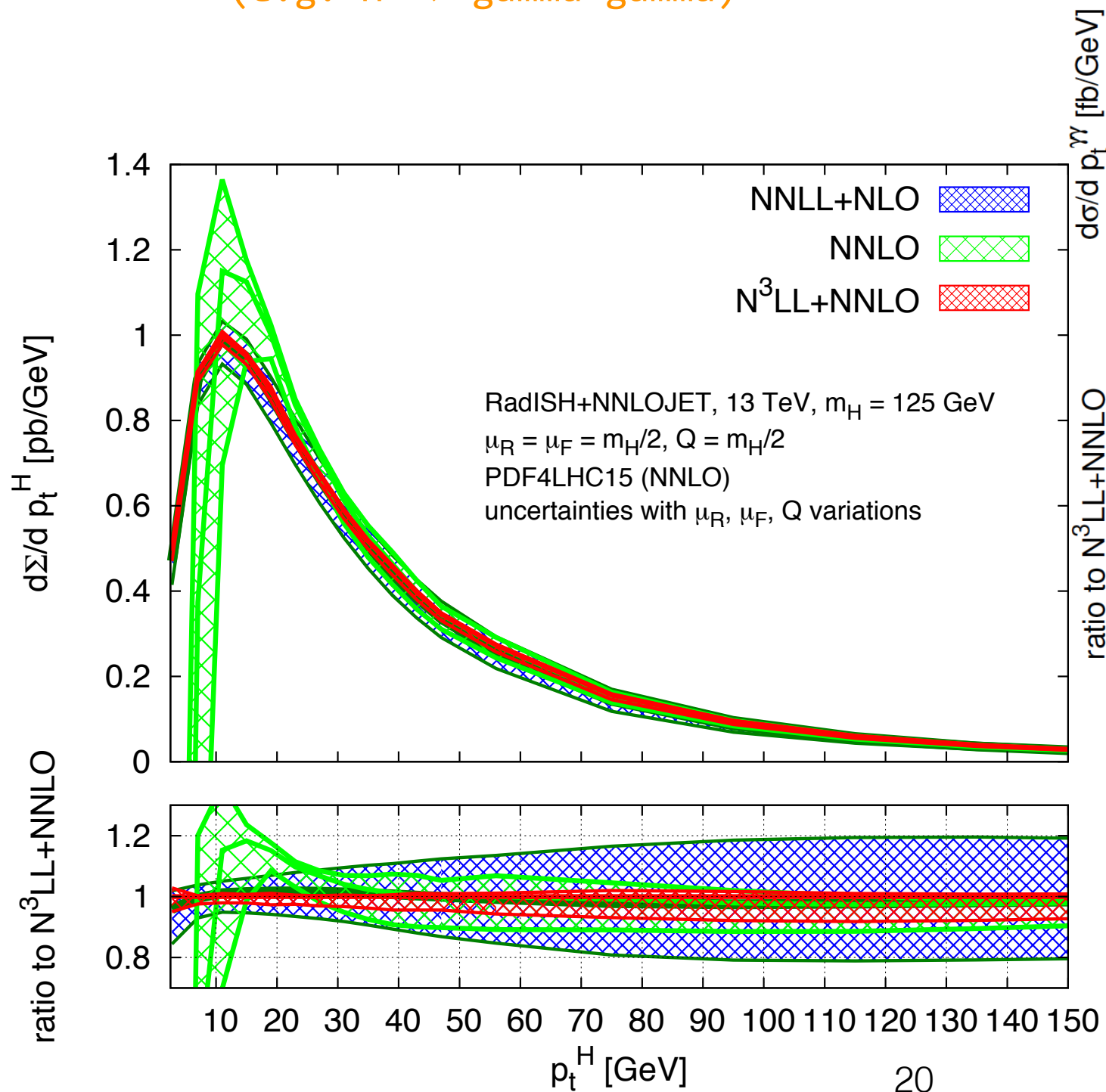
→ Important cancellations at $m_H/2$ (!), uncertainties likely underestimated at this scale (long known problem)

→ N^3LL corrections amount to a few-% at small p_t , reduction of band below 10 GeV consistently with NLO matching



An example: Higgs p_T spectrum

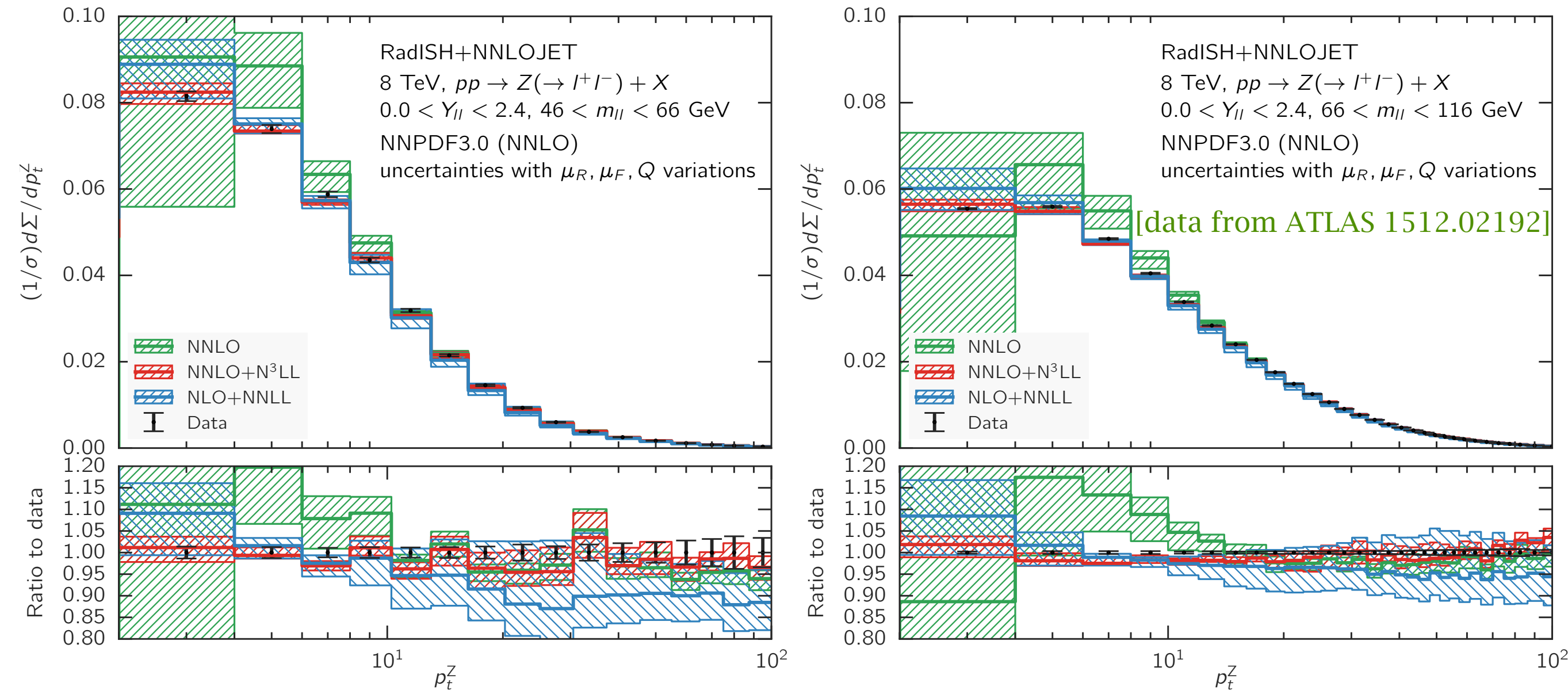
→ Simulation of fiducial distributions
(e.g. $H \rightarrow \gamma\gamma$)



→ Good convergence across different perturbative orders



An example: DY distributions (pT)



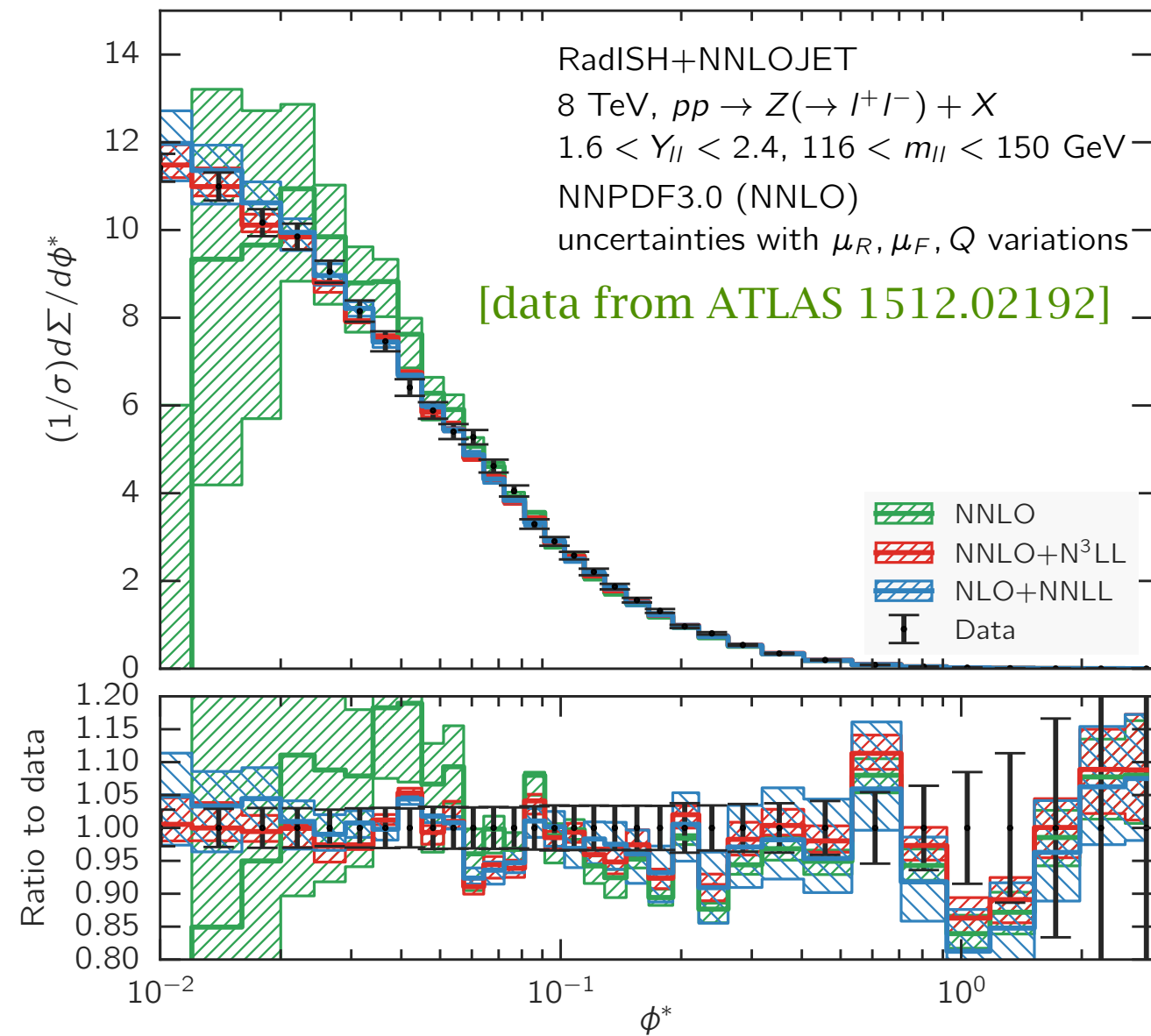
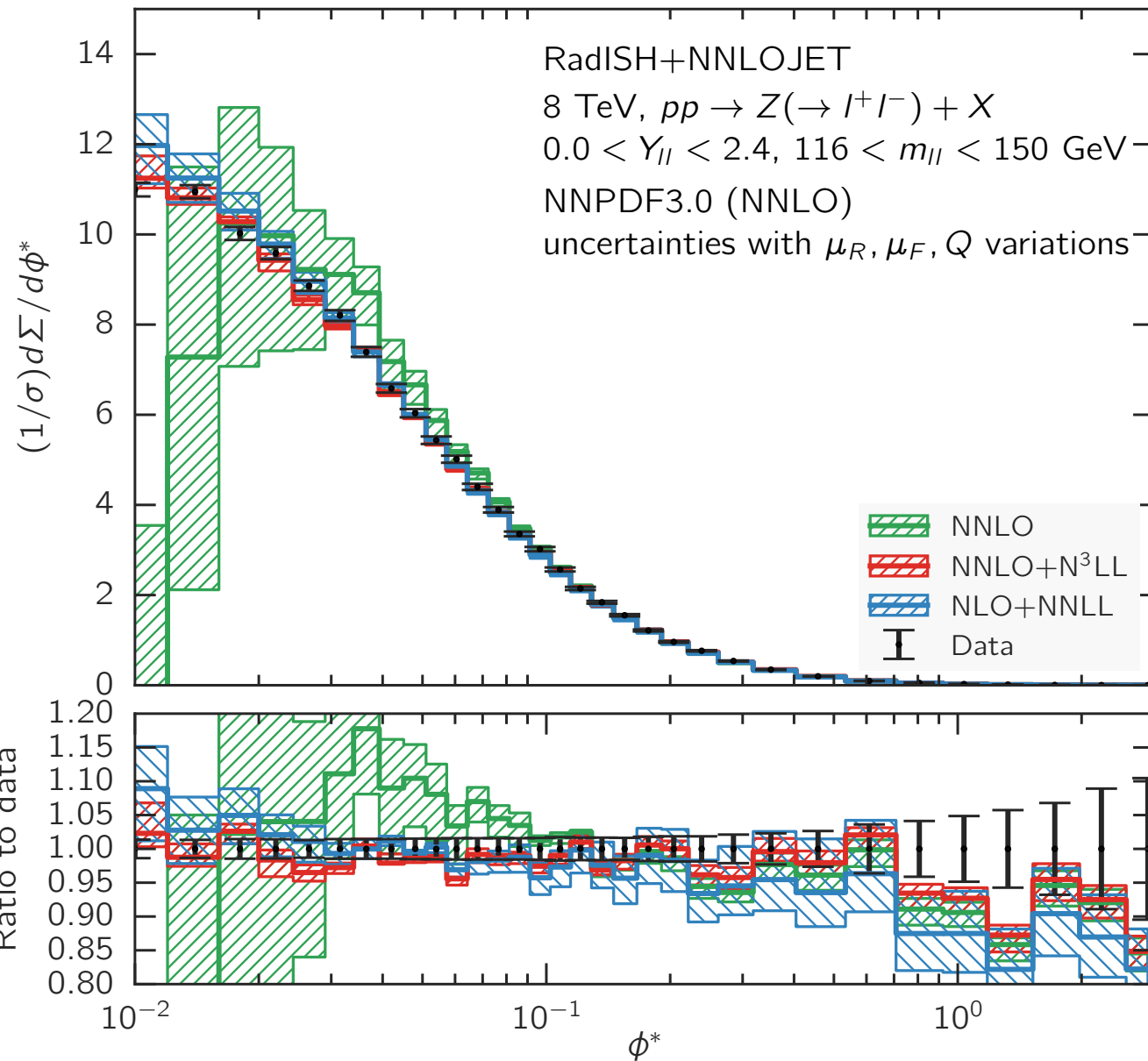
➡ Matching to differential NNLO from NNLOJET, assume N³LO correction to total XS is zero (i.e. no as^3 constant term included)

[Gehrmann-De Ridder, T. Gehrmann, E.W.N. Glover, A. Huss, T.A. Morgan '16]

➡ (sub-)percent precision in data, theory can reach ~3-5% accuracy...
Other effects important (QED, PDFs, quark masses, hadronisation)

➡ Relevant for W-mass studies

An example: DY distributions (ϕ^*)



➡ Similar conclusions for angular distributions

Conclusions

- Higher-order resummation can be formulated directly in momentum space without the need for a factorisation for the considered observable
- The approach I briefly outlined is generalised to any rIRC safe observable in two-scale problems
 - Systematic extension to any logarithmic order
 - Efficient implementation in a computer code: e.g. ARES, RadISH
 - Analytic resummation formulated in a language closer to parton showers
- Differential distributions at $N^3\text{LL}+\text{NNLO}$
 - Higgs: uncertainties in the 5%-10% range - consistent inclusion of quark-mass effects necessary at this order of accuracy (ongoing study)
 - DY: uncertainties reduced to $\sim 5\%$ across the whole spectrum - good agreement with data in the large-invariant mass bins (study low invariant mass in progress)
 - Improving on this requires the assessment of several effects: NP corrections, quark-mass corrections, QED, theory uncertainties in PDFs, ...

Thank you for listening

Squared amplitude decomposition

- Write all-order cross section as ($V(\{\tilde{p}\}, k_1, \dots, k_n) = |\vec{k}_{t1} + \dots + \vec{k}_{tn}|$)

$$\Sigma(v) = \int d\Phi_B \mathcal{V}(\Phi_B) \sum_{n=0}^{\infty} \int \prod_{i=1}^n [dk_i] \underline{|M(\tilde{p}_1, \tilde{p}_2, k_1, \dots, k_n)|^2} \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_n))$$

Real emissions

- Recast all-order squared ME for n real emissions as iteration of correlated blocks
- Scaling of the observable in the presence of radiation *must* preserve the above hierarchy

e.g. soft radiation (analogous considerations for hard-collinear)

$$\begin{aligned} |M(\tilde{p}_1, \tilde{p}_2, k_1, \dots, k_n)|^2 &= |M_B(\tilde{p}_1, \tilde{p}_2)|^2 \left\{ \left(\frac{1}{n!} \prod_{i=1}^n |M(k_i)|^2 \right) + \right. \\ &\left[\sum_{a>b} \frac{1}{(n-2)!} \left(\prod_{\substack{i=1 \\ i \neq a, b}}^n |M(k_i)|^2 \right) |\tilde{M}(k_a, k_b)|^2 + \right. \\ &\left. \sum_{a>b} \sum_{\substack{c>d \\ c, d \neq a, b}} \frac{1}{(n-4)!2!} \left(\prod_{\substack{i=1 \\ i \neq a, b, c, d}}^n |M(k_i)|^2 \right) |\tilde{M}(k_a, k_b)|^2 |\tilde{M}(k_c, k_d)|^2 + \dots \right] \\ &\left. + \left[\sum_{a>b>c} \frac{1}{(n-3)!} \left(\prod_{\substack{i=1 \\ i \neq a, b, c}}^n |M(k_i)|^2 \right) |\tilde{M}(k_a, k_b, k_c)|^2 + \dots \right] + \dots \right\}, \end{aligned}$$

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In addition to this counting, requiring that the observable is recursively IRC safe allows one to construct a (simpler) all-order subtraction scheme

Monte Carlo formulation

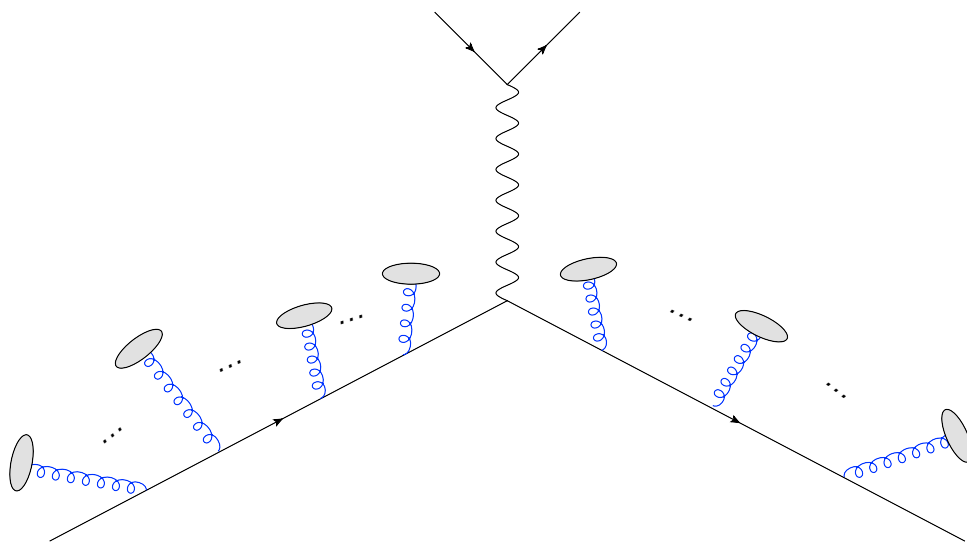
- One great simplification: choice of the resolution variable such that correlated blocks entering at $N^k\text{LL}$ in the unresolved radiation only contribute at $N^{k+1}\text{LL}$ in the resolved case
- i.e. we can expand out the cutoff dependence and retain in the Radiator only the terms necessary to cancel the singularities in the resolved radiation

$$R(\epsilon k_{t1}) = R(k_{t1}) + R'(k_{t1}) \ln \frac{1}{\epsilon} + \frac{1}{2} R''(k_{t1}) \ln^2 \frac{1}{\epsilon} + \dots$$

$$R'(k_{ti}) = R'(k_{t1}) + R''(k_{t1}) \ln \frac{k_{t1}}{k_{ti}} + \dots$$

Expansion is safe since
in the resolved
radiation
 $k_{t1}/k_{ti} \sim 1$

e.g. at NLL



$$\simeq \int \frac{dk_{t1}}{k_{t1}} \partial_L \left(-e^{-R(\mathbf{k}_{t1})} \mathcal{L}(\mathbf{k}_{t1}) \right) \int d\mathcal{Z}[\{R'(\mathbf{k}_{t1}), k_i\}] \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}))$$

$$\int d\mathcal{Z}[\{R'(k_{t1}), k_i\}] = \epsilon^{R'(\mathbf{k}_{t1})} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon k_{t1}}^{k_{t1}} \frac{dk_{ti}}{k_{ti}} R'(\mathbf{k}_{t1})$$

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- One great simplification: choice of the resolution variable such that correlated blocks entering at $N^k\text{LL}$ in the unresolved radiation only contribute at $N^{k+1}\text{LL}$ in the resolved case
- i.e. we can expand out the cutoff dependence and retain in the Radiator only the terms necessary to cancel the singularities in the resolved radiation

$$R(\epsilon k_{t1}) = R(k_{t1}) + R'(k_{t1}) \ln \frac{1}{\epsilon} + \frac{1}{2} R''(k_{t1}) \ln^2 \frac{1}{\epsilon} + \dots$$

$$R'(k_{ti}) = R'(k_{t1}) + R''(k_{t1}) \ln \frac{k_{t1}}{k_{ti}} + \dots$$

Expansion is safe since
in the resolved
radiation
 $k_{t1}/k_{ti} \sim 1$

- Corrections beyond NLL are obtained as follows
 - Add subleading effects in the Sudakov radiator and constants
 - Correct *a fixed number* of the NLL resolved emissions:

- only one at NNLL
- two at N³LL
- ...

e.g. at NNLL see:
[Banfi, PM, Salam, Zanderighi '12]
[Banfi, McAslan, PM, Zanderighi '14-'16]

Numerical implementation: RadISH

- Since the **transverse momenta of the *resolved* reals are of the same order**, we can expand the whole integrand about $k_{ti} \sim k_{t1}$ up to the desired logarithmic accuracy
- This expansion allows us to compute higher-order corrections to the NLL *resolved* reals by simply including one correction at a time

e.g. expansion up to NLL

$$\frac{d\Sigma(v)}{d\Phi_B} = \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R(k_{t1})} \mathcal{L}_{\text{N}^3\text{LL}}(k_{t1}) \right) \int dZ[\{R', k_i\}] \Theta(v - V(\{\bar{p}\}, k_1, \dots, k_{n+1}))$$

$$\begin{aligned} \mathcal{L}_{\text{N}^3\text{LL}}(k_{t1}) = & \sum_{c,c'} \frac{d|M_B|_{cc'}^2}{d\Phi_B} \sum_{i,j} \int_{x_1}^1 \frac{dz_1}{z_1} \int_{x_2}^1 \frac{dz_2}{z_2} f_i\left(k_{t1}, \frac{x_1}{z_1}\right) f_j\left(k_{t1}, \frac{x_2}{z_2}\right) \\ & \left\{ \delta_{ci}\delta_{c'j}\delta(1-z_1)\delta(1-z_2) \left(1 + \frac{\alpha_s(\mu_R)}{2\pi} H^{(1)}(\mu_R) + \frac{\alpha_s^2(\mu_R)}{(2\pi)^2} H^{(2)}(\mu_R) \right) \right. \\ & + \frac{\alpha_s(\mu_R)}{2\pi} \frac{1}{1-2\alpha_s(\mu_R)\beta_0 L} \left(1 - \alpha_s(\mu_R) \frac{\beta_1}{\beta_0} \frac{\ln(1-2\alpha_s(\mu_R)\beta_0 L)}{1-2\alpha_s(\mu_R)\beta_0 L} \right) \\ & \times \left(C_{ci}^{(1)}(z_1)\delta(1-z_2)\delta_{c'j} + \{z_1 \leftrightarrow z_2; c, i \leftrightarrow c', j\} \right) \\ & + \frac{\alpha_s^2(\mu_R)}{(2\pi)^2} \frac{1}{(1-2\alpha_s(\mu_R)\beta_0 L)^2} \left(\left(C_{ci}^{(2)}(z_1) - 2\pi\beta_0 C_{ci}^{(1)}(z_1) \ln \frac{M^2}{\mu_R^2} \right) \delta(1-z_2)\delta_{c'j} \right. \\ & + \{z_1 \leftrightarrow z_2; c, i \leftrightarrow c', j\} \left. \right) + \frac{\alpha_s^2(\mu_R)}{(2\pi)^2} \frac{1}{(1-2\alpha_s(\mu_R)\beta_0 L)^2} \left(C_{ci}^{(1)}(z_1)C_{c'j}^{(1)}(z_2) + G_{ci}^{(1)}(z_1)G_{c'j}^{(1)}(z_2) \right) \\ & \left. + \frac{\alpha_s^2(\mu_R)}{(2\pi)^2} H^{(1)}(\mu_R) \frac{1}{1-2\alpha_s(\mu_R)\beta_0 L} \left(C_{ci}^{(1)}(z_1)\delta(1-z_2)\delta_{c'j} + \{z_1 \leftrightarrow z_2; c, i \leftrightarrow c', j\} \right) \right\} \end{aligned}$$

• Coefficient functions and hard-virtual corrections absorbed into effective parton luminosities

Numerical implementation: RadISH

- Since the **transverse momenta of the *resolved* real emissions are of the same order**, we can expand the whole integrand about $k_{ti} \sim k_{t1}$ up to the desired logarithmic accuracy
- This expansion allows us to compute higher-order corrections to the NLL *resolved* real emissions by simply including one correction at a time

e.g. expansion up to NLL

$$\frac{d\Sigma(v)}{d\Phi_B} = \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R(k_{t1})} \mathcal{L}_{\text{N}^3\text{LL}}(k_{t1}) \right) \int dZ[\{R', k_i\}] \Theta(v - V(\{\bar{p}\}, k_1, \dots, k_{n+1}))$$

$$k_{ti}/k_{t1} = \zeta_i = \mathcal{O}(1)$$

$$\int dZ[\{R', k_i\}] G(\{\bar{p}\}, \{k_i\}) = \epsilon^{R'(k_{t1})} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} R'(k_{t1}) G(\{\bar{p}\}, k_1, \dots, k_{n+1})$$

• Coefficient functions and hard-virtual corrections absorbed into effective parton luminosities

• The ensemble of NLL real emissions dZ is generated as a parton shower. Fast numerical evaluation with Monte-Carlo methods.

Numerical implementation: RadISH

- Since the **transverse momenta of the *resolved* real** are of the same order, we can expand the whole integrand about $k_{ti} \sim k_{t1}$ up to the desired logarithmic accuracy
- This expansion allows us to compute higher-order corrections to the NLL *resolved* real by simply including one correction at a time

e.g. expansion up to NNLL

$$\begin{aligned} \frac{d\Sigma(v)}{d\Phi_B} = & \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R(k_{t1})} \mathcal{L}_{\text{NLL}}(k_{t1}) \right) \int dZ[\{R', k_i\}] \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \\ & + \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} \int dZ[\{R', k_i\}] \int_0^1 \frac{d\zeta_s}{\zeta_s} \frac{d\phi_s}{2\pi} \left\{ \left(R'(k_{t1}) \mathcal{L}_{\text{NNLL}}(k_{t1}) - \partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) \right) \right. \\ & \times \left(R''(k_{t1}) \ln \frac{1}{\zeta_s} + \frac{1}{2} R'''(k_{t1}) \ln^2 \frac{1}{\zeta_s} \right) - R'(k_{t1}) \left(\partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) - 2 \frac{\beta_0}{\pi} \alpha_s^2(k_{t1}) \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \ln \frac{1}{\zeta_s} \right) \\ & \left. + \frac{\alpha_s^2(k_{t1})}{\pi^2} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \right\} \left\{ \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_s)) - \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \right\} \end{aligned}$$

• Coefficient functions and hard-virtual corrections absorbed into effective parton luminosities

• The ensemble of NLL real emissions dZ is generated as a parton shower. Fast numerical evaluation with Monte-Carlo methods.

Numerical implementation: RadISH

- Since the **transverse momenta of the *resolved* real** are of the same order, we can expand the whole integrand about $k_{ti} \sim k_{t1}$ up to the desired logarithmic accuracy
- This expansion allows us to compute higher-order corrections to the NLL *resolved* real by simply including one correction at a time

e.g. expansion up to N³LL

$$\begin{aligned}
 \frac{d\Sigma(v)}{d\Phi_B} &= \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} \partial_L \left(-e^{-R(k_{t1})} \mathcal{L}_{\text{N}^3\text{LL}}(k_{t1}) \right) \int dZ[\{R', k_i\}] \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \\
 &+ \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} \int dZ[\{R', k_i\}] \int_0^1 \frac{d\zeta_s}{\zeta_s} \frac{d\phi_s}{2\pi} \left\{ \left(R'(k_{t1}) \mathcal{L}_{\text{NNLL}}(k_{t1}) - \partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) \right) \right. \\
 &\times \left(R''(k_{t1}) \ln \frac{1}{\zeta_s} + \frac{1}{2} R'''(k_{t1}) \ln^2 \frac{1}{\zeta_s} \right) - R'(k_{t1}) \left(\partial_L \mathcal{L}_{\text{NNLL}}(k_{t1}) - 2 \frac{\beta_0}{\pi} \alpha_s^2(k_{t1}) \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \ln \frac{1}{\zeta_s} \right) \\
 &\left. + \frac{\alpha_s^2(k_{t1})}{\pi^2} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \right\} \left\{ \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_s)) - \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \right\} \\
 &+ \frac{1}{2} \int \frac{dk_{t1}}{k_{t1}} \frac{d\phi_1}{2\pi} e^{-R(k_{t1})} \int dZ[\{R', k_i\}] \int_0^1 \frac{d\zeta_{s1}}{\zeta_{s1}} \frac{d\phi_{s1}}{2\pi} \int_0^1 \frac{d\zeta_{s2}}{\zeta_{s2}} \frac{d\phi_{s2}}{2\pi} R'(k_{t1}) \\
 &\times \left\{ \mathcal{L}_{\text{NLL}}(k_{t1}) (R''(k_{t1}))^2 \ln \frac{1}{\zeta_{s1}} \ln \frac{1}{\zeta_{s2}} - \partial_L \mathcal{L}_{\text{NLL}}(k_{t1}) R''(k_{t1}) \left(\ln \frac{1}{\zeta_{s1}} + \ln \frac{1}{\zeta_{s2}} \right) \right. \\
 &\left. + \frac{\alpha_s^2(k_{t1})}{\pi^2} \hat{P}^{(0)} \otimes \hat{P}^{(0)} \otimes \mathcal{L}_{\text{NLL}}(k_{t1}) \right\} \\
 &\times \left\{ \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s1}, k_{s2})) - \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s1})) - \right. \\
 &\left. \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1}, k_{s2})) + \Theta(v - V(\{\tilde{p}\}, k_1, \dots, k_{n+1})) \right\} + \mathcal{O} \left(\alpha_s^n \ln^{2n-6} \frac{1}{v} \right)
 \end{aligned}$$

• Coefficient functions and hard-virtual corrections absorbed into effective parton luminosities

• The ensemble of NLL real emissions dZ is generated as a parton shower. Fast numerical evaluation with Monte-Carlo methods.

Equivalence to CSS formula

- Hard-collinear emissions off initial-state legs require some care in the treatment of kinematics. Final result reads

$$\frac{d\Sigma(v)}{dp_t d\Phi_B} = \int_{\mathcal{C}_1} \frac{dN_1}{2\pi i} \int_{\mathcal{C}_2} \frac{dN_2}{2\pi i} x_1^{-N_1} x_2^{-N_2} \sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \mathbf{f}_{N_1}^T(\mu_0) \frac{d\hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v)}{dp_t} \mathbf{f}_{N_2}(\mu_0)$$

$$\hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v) = \left[\mathbf{C}_{N_1}^{c_1; T}(\alpha_s(\mu_0)) H(\mu_R) \mathbf{C}_{N_2}^{c_2}(\alpha_s(\mu_0)) \right] \int_0^M \frac{dk_{t1}}{k_{t1}} \int_0^{2\pi} \frac{d\phi_1}{2\pi}$$

$$\times e^{-\mathbf{R}(\epsilon k_{t1})} \exp \left\{ - \sum_{\ell=1}^2 \left(\int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \frac{\alpha_s(k_t)}{\pi} \mathbf{\Gamma}_{N_\ell}(\alpha_s(k_t)) + \int_{\epsilon k_{t1}}^{\mu_0} \frac{dk_t}{k_t} \mathbf{\Gamma}_{N_\ell}^{(C)}(\alpha_s(k_t)) \right) \right\}$$

$$\sum_{\ell_1=1}^2 \left(\mathbf{R}'_{\ell_1}(k_{t1}) + \frac{\alpha_s(k_{t1})}{\pi} \mathbf{\Gamma}_{N_{\ell_1}}(\alpha_s(k_{t1})) + \mathbf{\Gamma}_{N_{\ell_1}}^{(C)}(\alpha_s(k_{t1})) \right)$$

$$\times \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=2}^{n+1} \int_{\epsilon}^1 \frac{d\zeta_i}{\zeta_i} \int_0^{2\pi} \frac{d\phi_i}{2\pi} \sum_{\ell_i=1}^2 \left(\mathbf{R}'_{\ell_i}(k_{ti}) + \frac{\alpha_s(k_{ti})}{\pi} \mathbf{\Gamma}_{N_{\ell_i}}(\alpha_s(k_{ti})) + \mathbf{\Gamma}_{N_{\ell_i}}^{(C)}(\alpha_s(k_{ti})) \right)$$

$$\times \Theta(v - V(\{\vec{p}\}, k_1, \dots, k_{n+1})),$$

- Formulation equivalent to b-space result, up to a scheme change. Using the delta representation for the distribution one finds

$$\delta^{(2)}(\vec{p}_t - (\vec{k}_{t1} + \dots + \vec{k}_{tn})) = \int \frac{d^2 \vec{b}}{4\pi^2} e^{-i\vec{b} \cdot \vec{p}_t} \prod_{i=1}^n e^{i\vec{b} \cdot \vec{k}_{ti}}$$

$$\frac{d\Sigma(v)}{dp_t d\Phi_B} = \int_{\mathcal{C}_1} \frac{dN_1}{2\pi i} \int_{\mathcal{C}_2} \frac{dN_2}{2\pi i} x_1^{-N_1} x_2^{-N_2} \sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \mathbf{f}_{N_1}^T(\mu_0) \frac{d\hat{\Sigma}_{N_1, N_2}^{c_1, c_2}(v)}{dp_t} \mathbf{f}_{N_2}(\mu_0) =$$

$$\sum_{c_1, c_2} \frac{d|M_B|_{c_1 c_2}^2}{d\Phi_B} \int b db p_t J_0(p_t b) \mathbf{f}^T(b_0/b) \mathbf{C}_{N_1}^{c_1; T}(\alpha_s(b_0/b)) H(M) \mathbf{C}_{N_2}^{c_2}(\alpha_s(b_0/b)) \mathbf{f}(b_0/b)$$

$$\times \exp \left\{ - \sum_{\ell=1}^2 \int_0^M \frac{dk_t}{k_t} \mathbf{R}'_{\ell}(k_t) (1 - J_0(bk_t)) \right\}.$$

$$(1 - J_0(bk_t)) \simeq \Theta(k_t - \frac{b_0}{b}) + \frac{\zeta_3}{12} \frac{\partial^3}{\partial \ln(Mb/b_0)^3} \Theta(k_t - \frac{b_0}{b}) + \dots$$