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The QCD equation of state at finite density, from the known to the unknown

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The theory of quantum chromodynamics (QCD) is expected to have a rich phase structure at finite chemical potential and temperature. Its study is a central topic of high energy nuclear physics. Theoretical studies employing lattice QCD methods have already established that the transition from hadrons to quarks proceeds as a smooth crossover in the case of vanishing net baryon number density. At very large net baryon densities and low temperatures, astrophysical observations have become more important in constraining the QCD Equation of State. Nuclear matter ground-state properties as well as properties of compact stars and their violent mergers will serve to determine the equation of state at several times nuclear ground-state density.

In this talk I will present the first calculation of the QCD phase structure and thermodynamics which is shown to be consistent with lattice QCD results at small barychemical potential as well as nuclear matter properties and known constraints from compact star observations [1,2].

In this context I will discuss the most relevant properties and constraints which should be satisfied by any model which attempts to predict the QCD phase structure.

Furthermore, I will present results on the baryon number susceptibilities calculated with this model and discuss how nuclear interactions may strongly influence the measured baryon number fluctuations in nuclear collisions at low beam energies [1].

Finally, I will also address similarities and differences between the matter created in heavy ion collisions and in mergers of compact stars and how both can be described in a unified framework [3].

[1] A.Mukherjee, JS and S.Schramm, Phys. Rev. C 96, no. 2, 025205 (2017)

[2] A.Mukherjee, S.Schramm, JS and V. Dexheimer, arXiv:1706.09191 [nucl-th].(accepted for publication in Astronomy & Astrophysics)

[3] M.Hanuske, JS et al., J. Phys. Conf. Ser. 878, no. 1, 012031 (2017).

Content type

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Collaboration

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