Directed flow due to the initial source tilt and density asymmetry in CutAu and AutAu collisions at STAR

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-0.5

-1

0

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Motivation

SIAR



No models can reproduce directed flow, v_1 , dependence on (pseudo)rapidity, p_T , $\sqrt{s_{NN}}$, and particle species, simultaneously. Still missing an important piece in the picture of heavy-ion collisions, 0.5 e.g. vorticity and/or 3D initial condition.

$$\frac{dN}{d\phi} \propto 1 + 2v_n \sum_n \cos n(\phi - \Psi_n)$$



0.5

-0.5



Origin of directed flow

Contributions to v_1 :

- Initial source tilt [1,2] (1)
- Initial density asymmetry at non-zero rapidity [3] (2)
- Initial density asymmetry due to fluctuations [4] (3)

In addition, for asymmetric Cu+Au collisions:

(4) Intrinsic density asymmetry due to the geometry ($N_{part}^{Au} > N_{part}^{Cu}$) (5) $N_{part}^{Au} > N_{part}^{Cu}$ leads to a rapidity shift of v_1

* Contributions from (2) and (3) are called "dipole flow"; named after "dipole-like" density asymmetry.

Cu+Au provides a unique opportunity to study the role of the different mechanisms in v_1 .

Results and Discussion

 v_1 was measured relative to two spectator planes Ψ_{SP} as done in ALICE [5] and was decomposed into "conventional" (1)+(2) and "fluctuation" (3) components. * Ψ_{SP} resolution was estimated by 3-subevent method with ZDCSMD ($|\eta| > 6.3$) and BBC (3.3< $|\eta| < 5$) $v_1^{\text{conv}}(v_1^{\text{odd}}) = (v_1\{\Psi_{\text{SP}}^p\} - v_1\{\Psi_{\text{SP}}^t\})/2$ $v_1 = \langle \cos(\phi - \Psi_{\rm SP}) \rangle$ $\langle p_{\rm x} \rangle = \langle p_{\rm T} \cos(\phi - \Psi_{\rm SP}) \rangle$ $v_1^{\text{fluc}}(v_1^{\text{even}}) = (v_1\{\Psi_{\text{SP}}^p\} + v_1\{\Psi_{\text{SP}}^t\})/2$



- \Box Similar v₁^{conv(odd)} slopes in Au+Au and Cu+Au but larger than at the LHC
 - The source tilt likely depends on the collision energy but not on the system size
- \Box Cu+Au V₁^{conv} is shifted upward relative to Au+Au
 - as expected from the intrinsic density asymmetry
- \Box Cu+Au <p_x>^{conv} is shifted toward Au-going direction relative to Au+Au
 - as expected from asymmetric participants > ^{0.001} star

- out the direction of the density asymmetry
- \Box Relative contribution from the source tilt to v₁ slope, r^{tilt}
- at RHIC ~2/3 $r^{\text{tilt}} = \frac{(dv_1/d\eta)^{\text{tilt}}}{dv_1/d\eta} \approx \frac{2}{3} \frac{(d\langle p_x \rangle/d\eta)/\langle p_T \rangle}{dv_1/d\eta}$
- at LHC ~1/3 [5]

(smaller source tilt due to baryon transparency)

- \Box v₁^{fluc(even)} nearly rapidity-independent and $< p_{y} > fluc(even)$ close to zero
- \Box Very weak centrality dependence of $v_1^{fluc(even)}$
 - Similar dipole-like density fluctuations for all centralities
- \Box <p_x>^{even} ~ 0 in symmetric systems
 - Feature of the dipole flow due to the momentum conservation



The STAR Collaboration

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Conclusions

- □ The results are consistent with a picture of the directed flow originating from the initial source tilt and the initial density asymmetry.
- \Box Relative contribution to v₁^{odd} slope from the initial source tilt is ~2/3 in mid-central collisions at RHIC and the rest comes from the rapiditydependent density asymmetry

References

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