



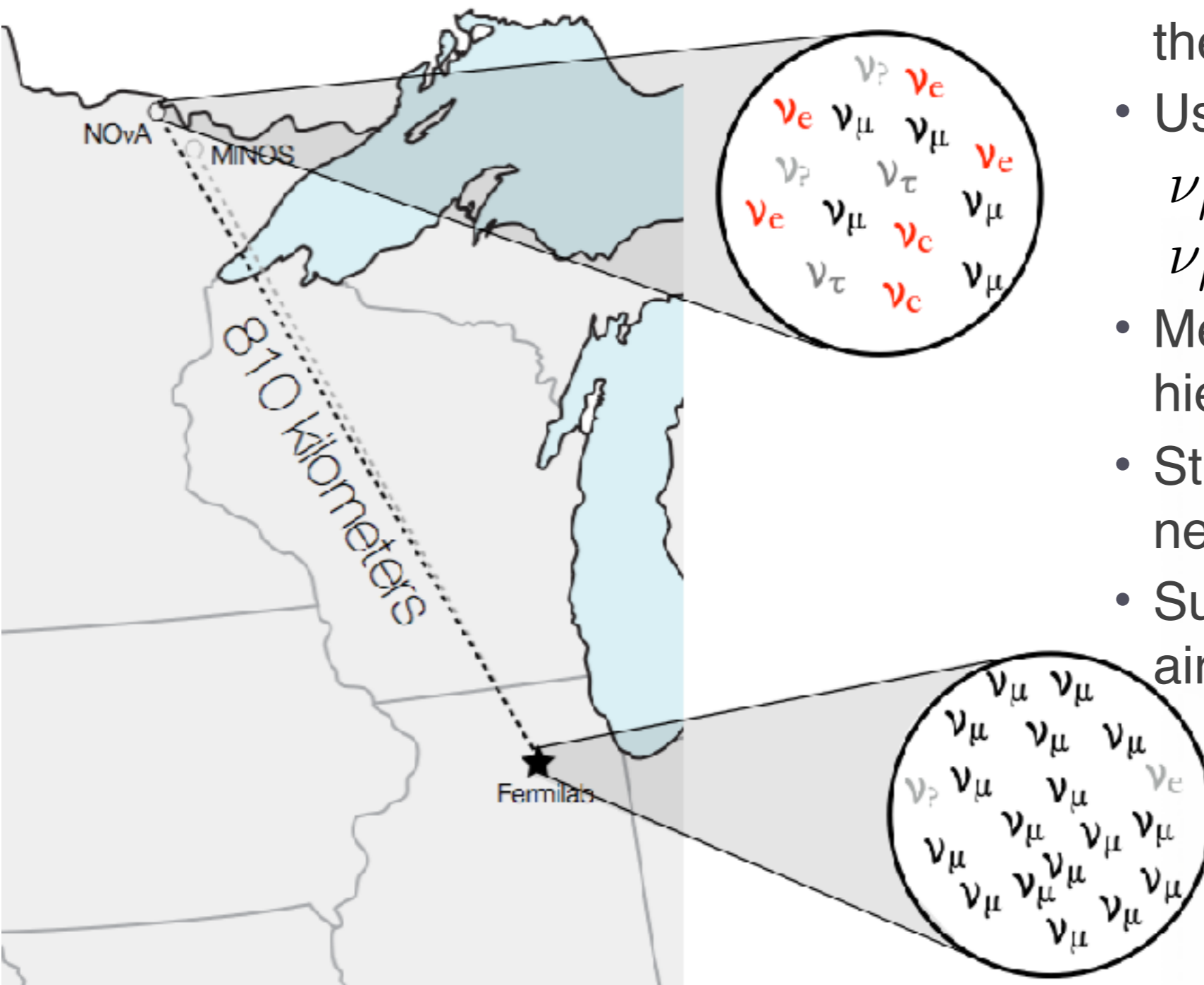
# NOvA: Latest Results and Future Plans

Evan Niner, Fermilab

NNN 2017

27 October 2017

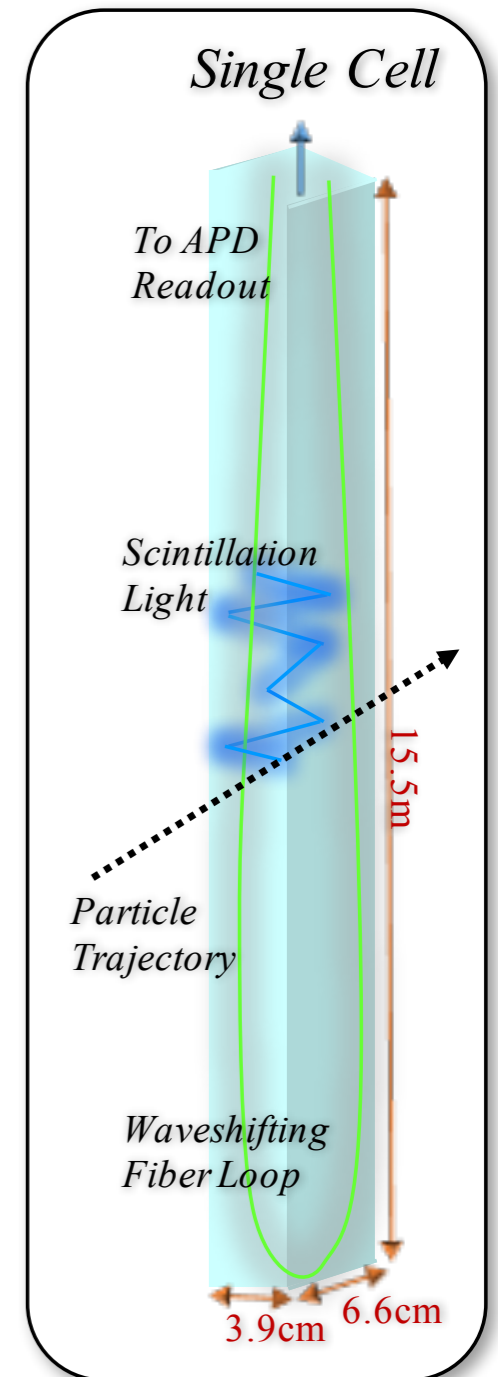
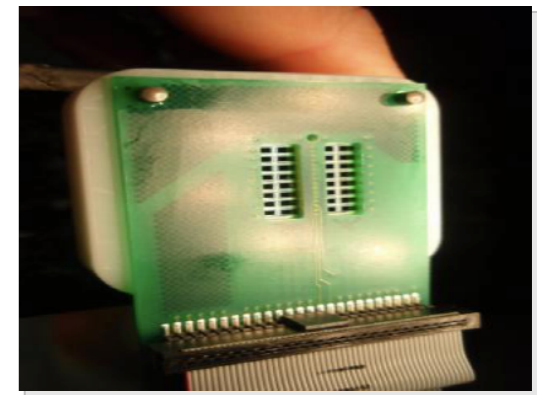
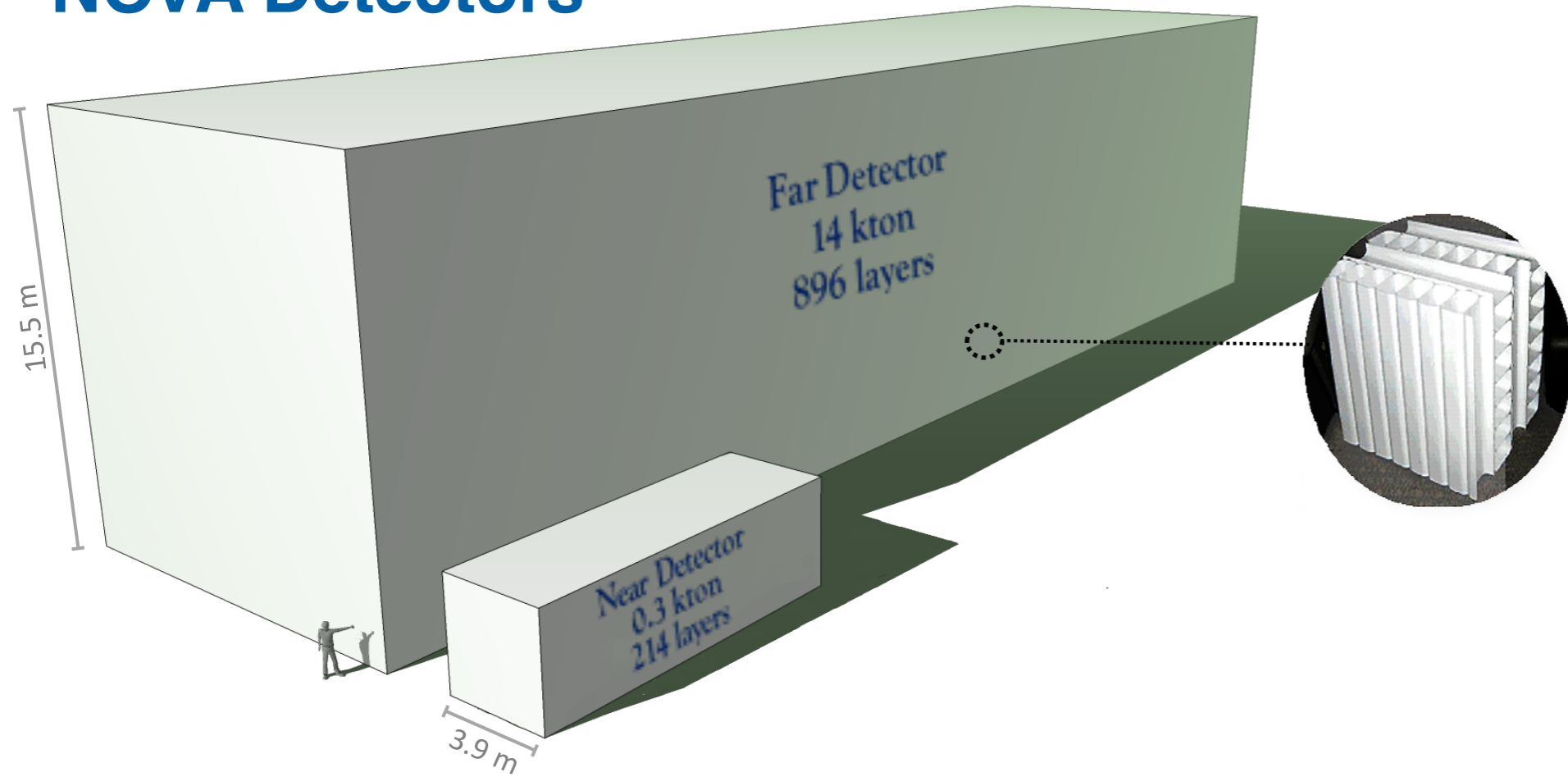
# The NOvA Experiment



- Observe neutrinos from NuMI neutrino beam line at Fermilab
- Two functionally identical detectors
- 810 km baseline, the longest in the world
- Uses four oscillation channels:
 
$$\nu_\mu \rightarrow \nu_\mu \quad \bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$$

$$\nu_\mu \rightarrow \nu_e \quad \bar{\nu}_\mu \rightarrow \bar{\nu}_e$$
- Measure  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m^2_{32}$ , mass hierarchy, and  $\delta_{cp}$
- Sterile neutrino searches, neutrino cross sections
- Supernova, monopoles, cosmic air showers, exotic physics

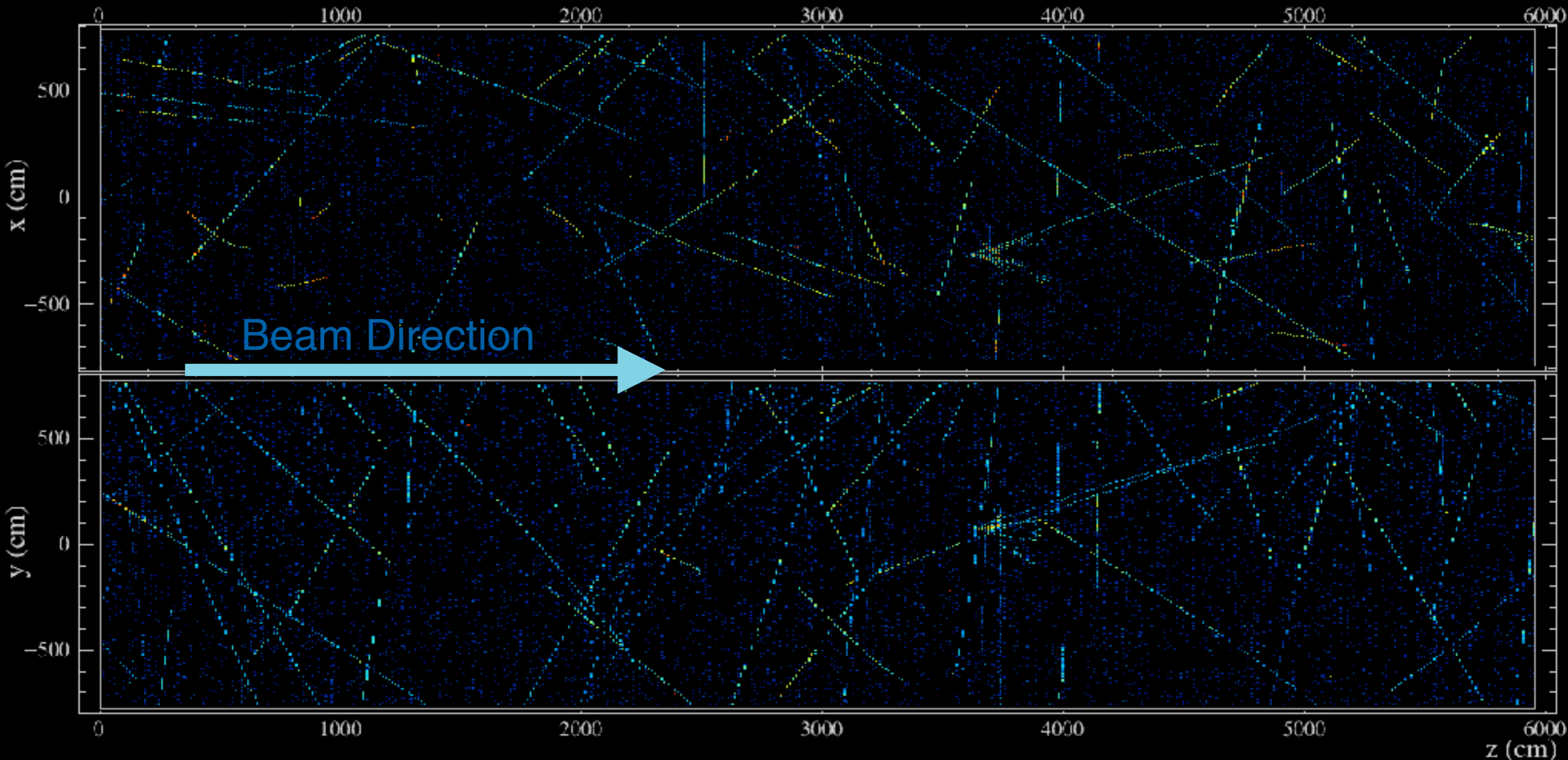
# NOvA Detectors



- Two functionally identical detectors
- Extruded plastic cells alternating vertical and horizontal orientation filled with liquid scintillator

# Far Detector 550 $\mu\text{s}$ Readout Window

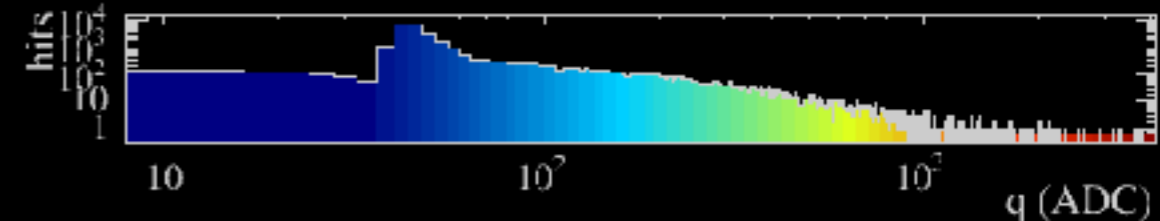
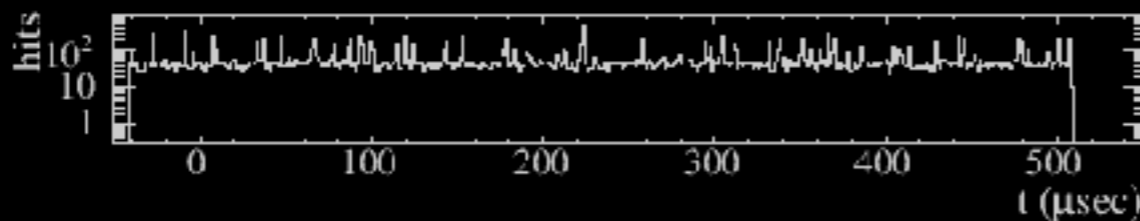
Cell hits colored by charge deposition



NOvA - FNAL E929

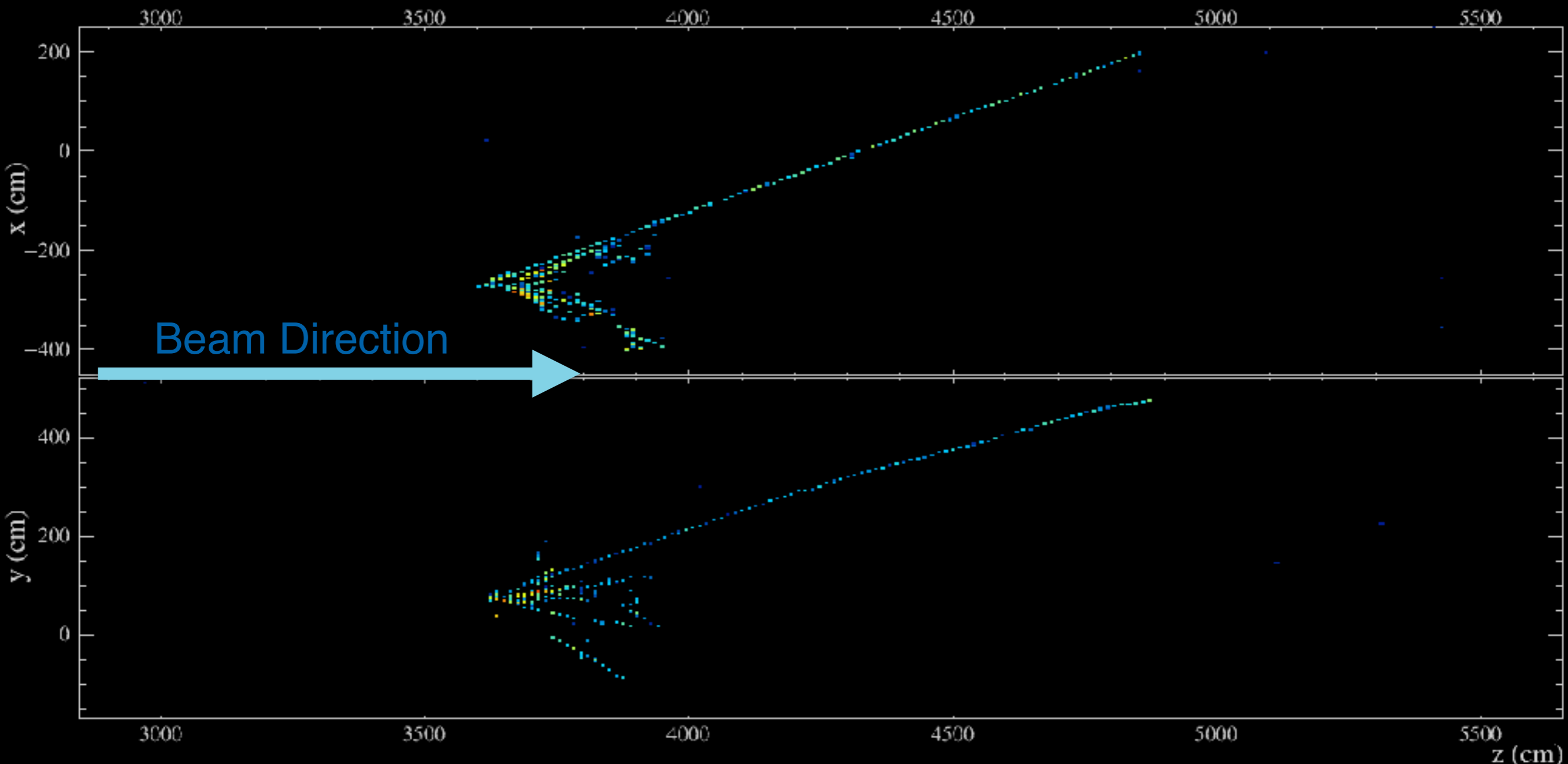
Run: 18520 / 13  
Fvent: 178402 / --

UTC Fri Jan 9, 2015  
00:13:53.087341608



# Far Detector Neutrino Interaction

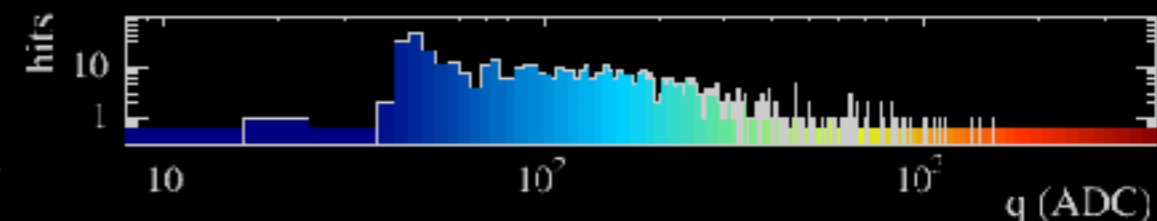
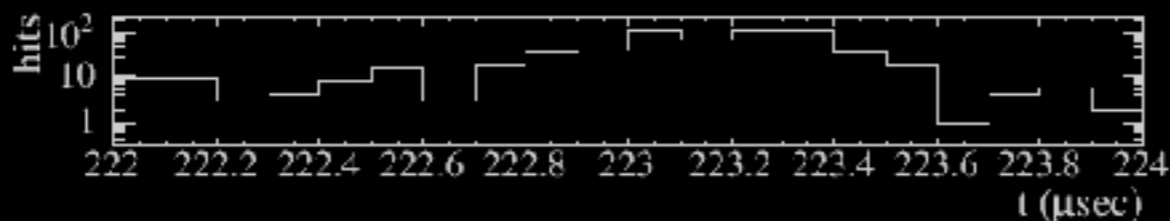
Cell hits colored by charge deposition



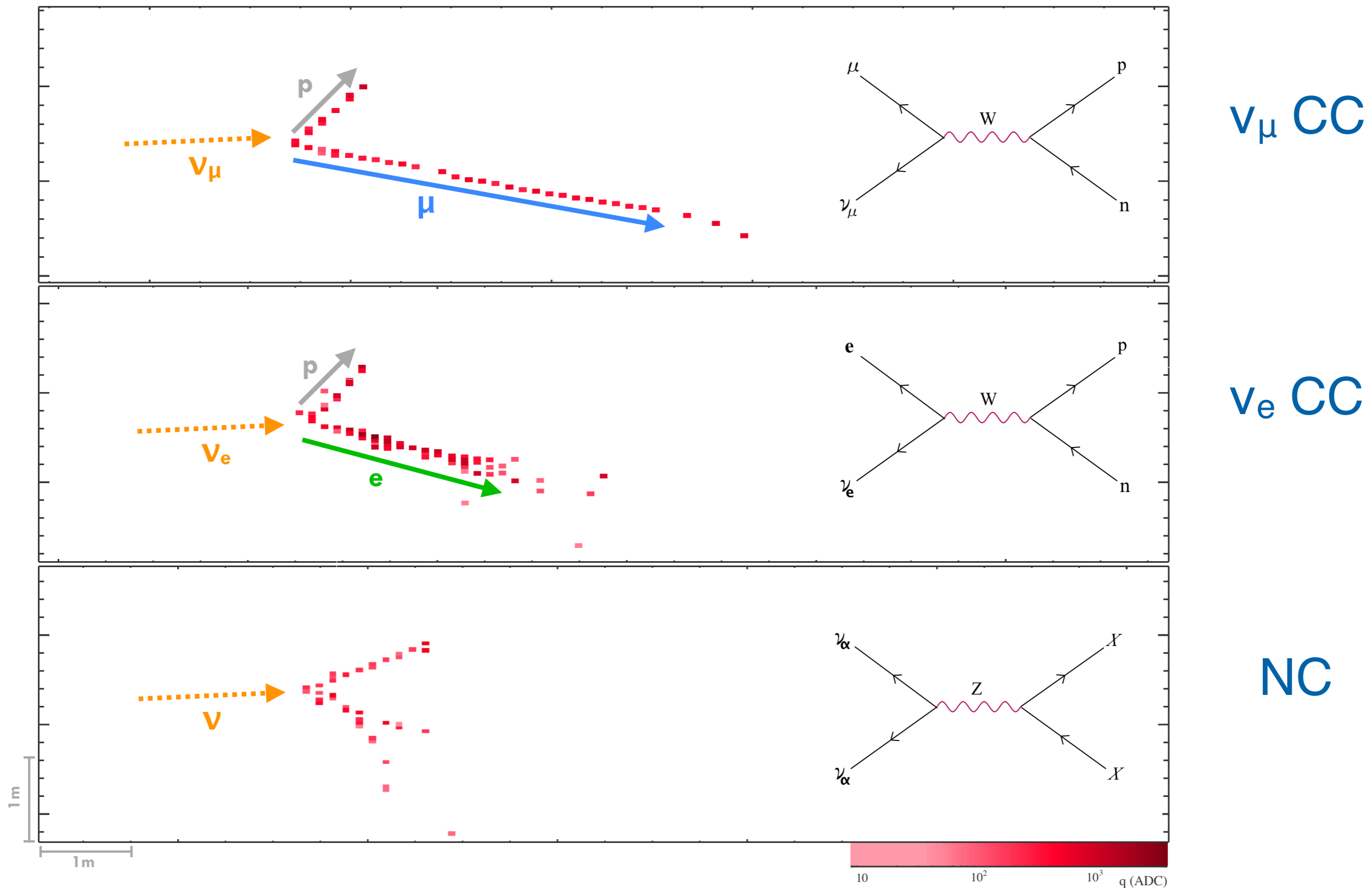
NOvA - FNAL E929

Run: 18520 / 13  
Fvent: 178402 / --

UTC Fri Jan 9, 2015  
00:13:53.087341608



# Event Topologies



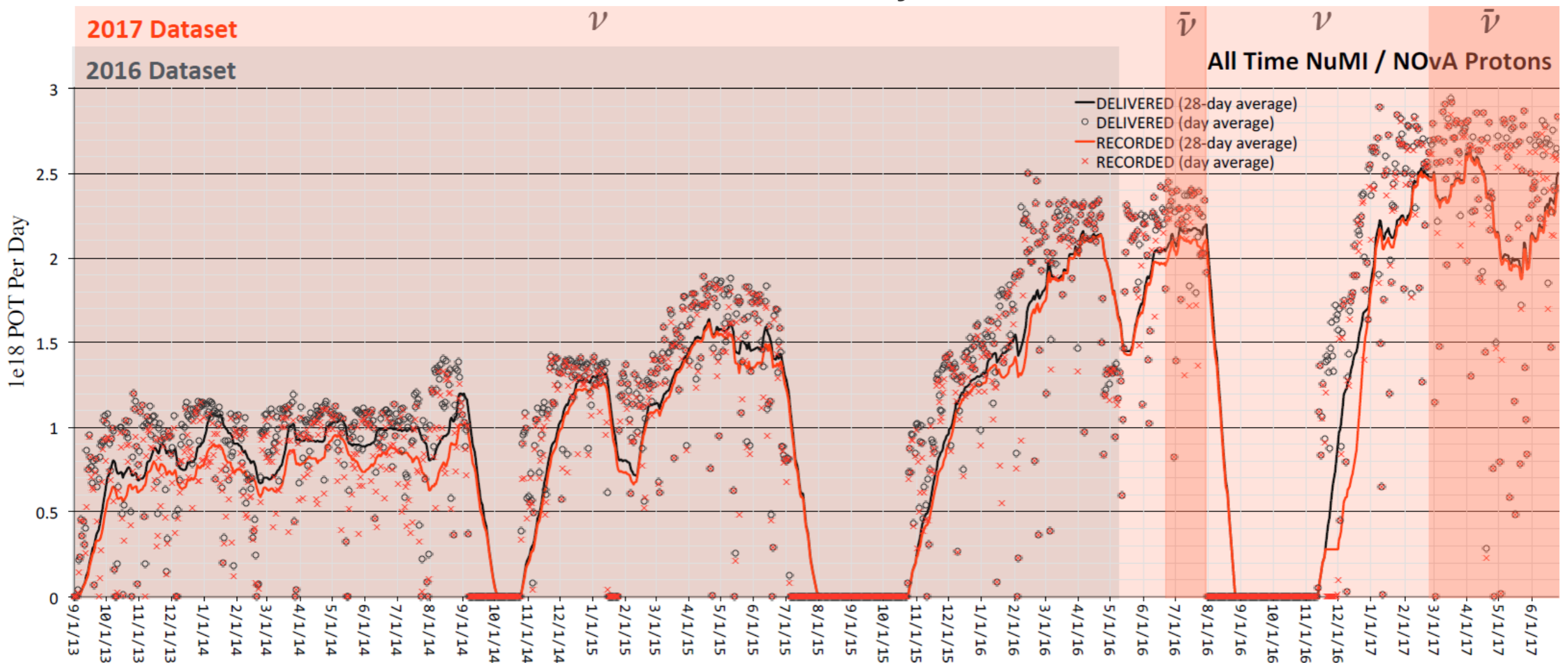
$\nu_\mu$  CC

$\nu_e$  CC

NC

# NuMI Beam Performance

- $6.05 \times 10^{20}$  full detector equivalent POT in 2016 dataset
  - $\nu_\mu$  disappearance and  $\nu_e$  appearance results
- $8.85 \times 10^{20}$  full detector equivalent POT in 2017 dataset
  - neutral current disappearance result
- Beam in anti-neutrino mode now, routinely above 650 kW

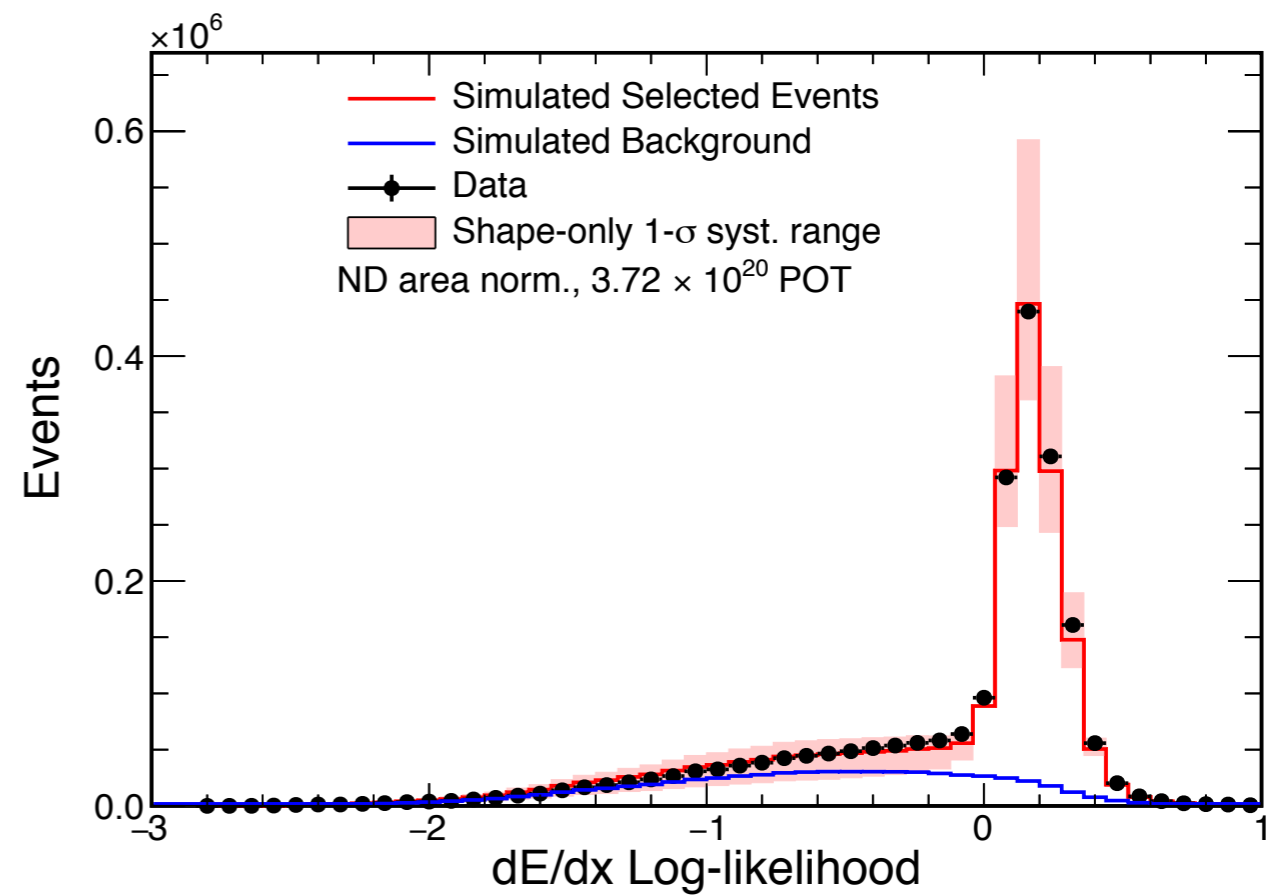


# Muon Neutrino Disappearance Analysis

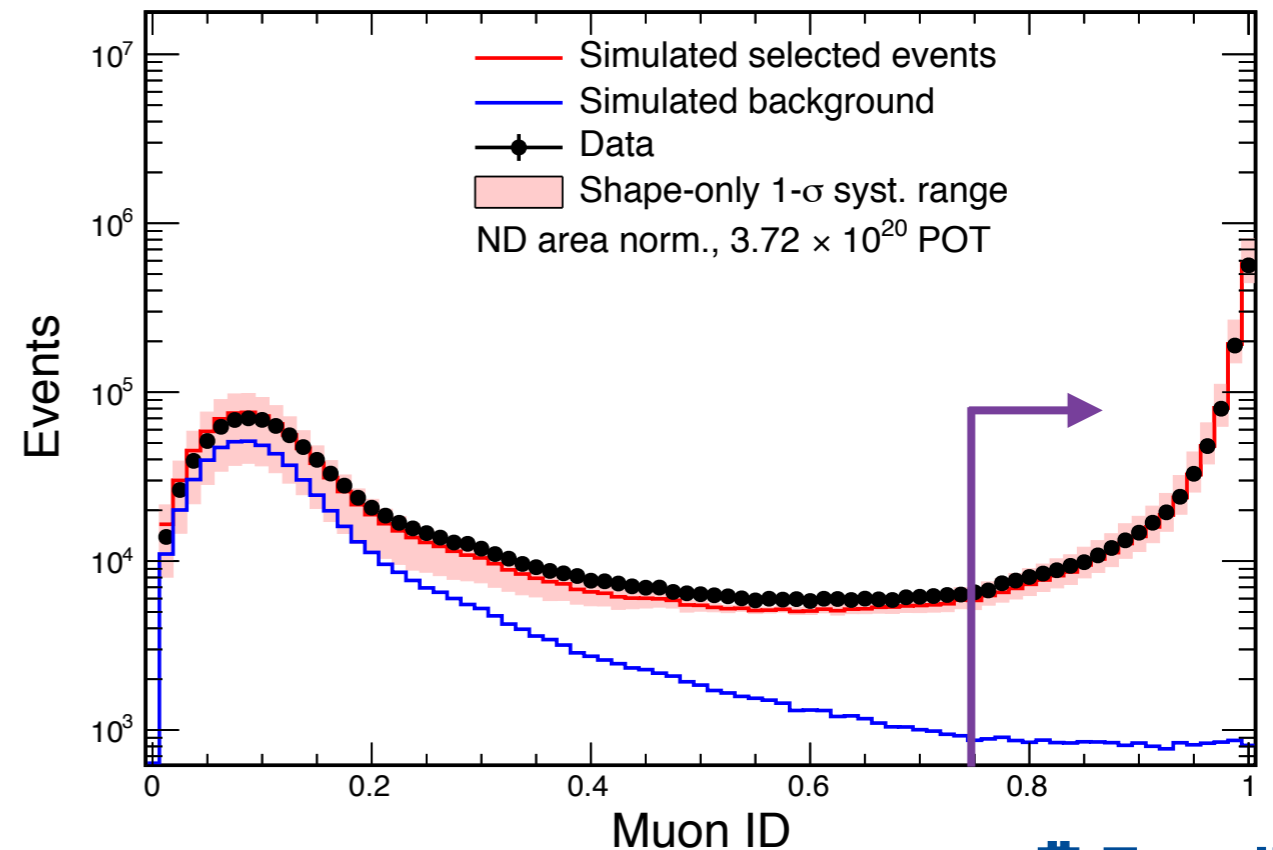


# Muon Neutrino Selection

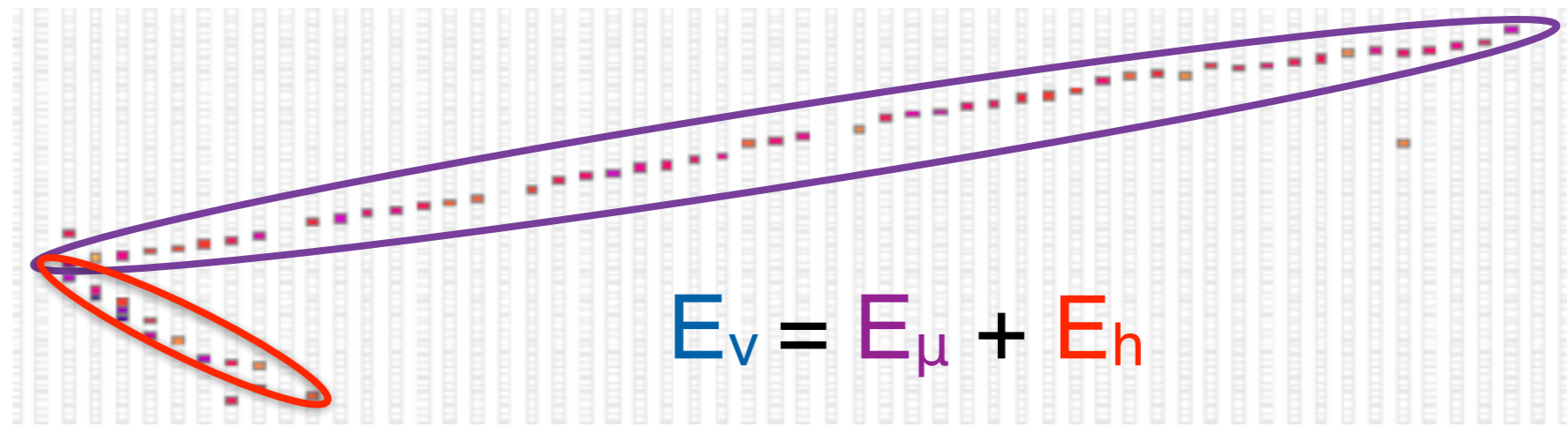
- Separate  $\nu_\mu$  CC interactions from NC and cosmic-ray backgrounds
- Containment cuts remove activity near walls
- Four variable k-Nearest Neighbor algorithm to select muons
  - track length
  - $dE/dx$  along track
  - scattering along track
  - track-only plane fraction
- Selection is 81% efficient and 91% pure



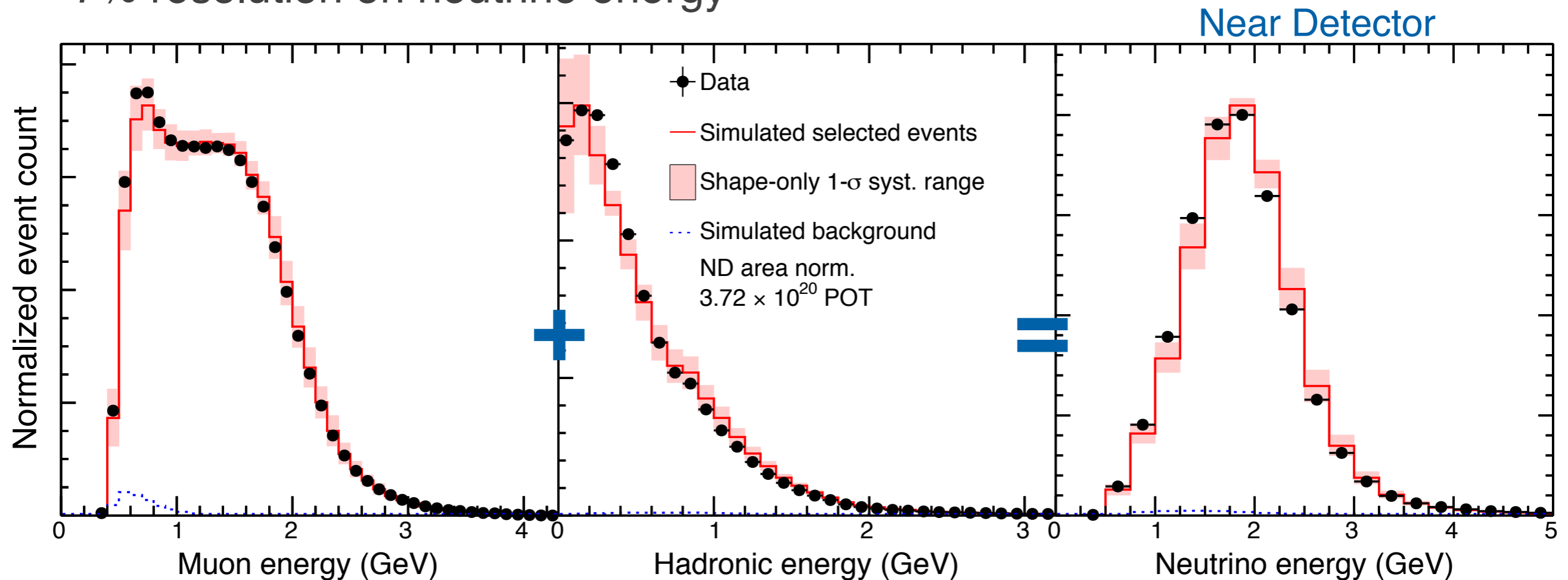
NOvA Preliminary



# Energy Estimation

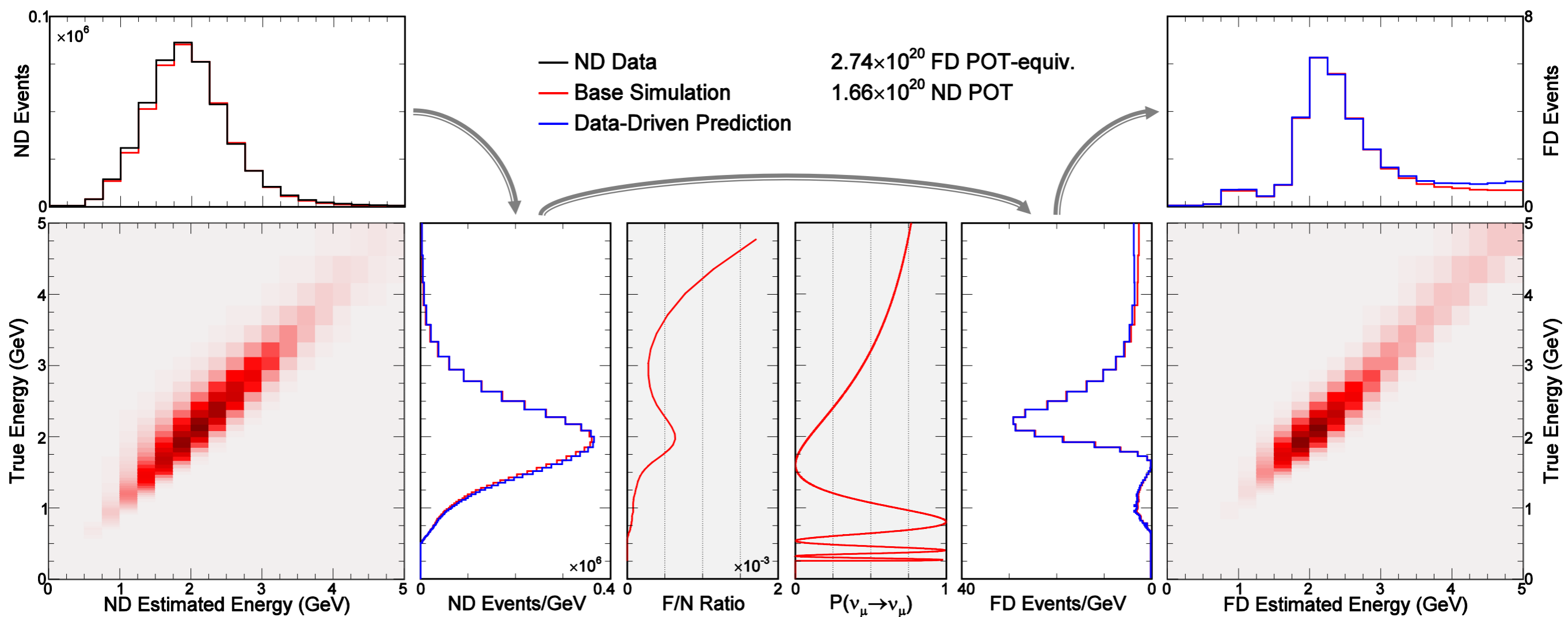


- Muon energy estimated from muon length
- Hadronic energy estimated calorimetrically from off-track hits
- $\sim 7\%$  resolution on neutrino energy



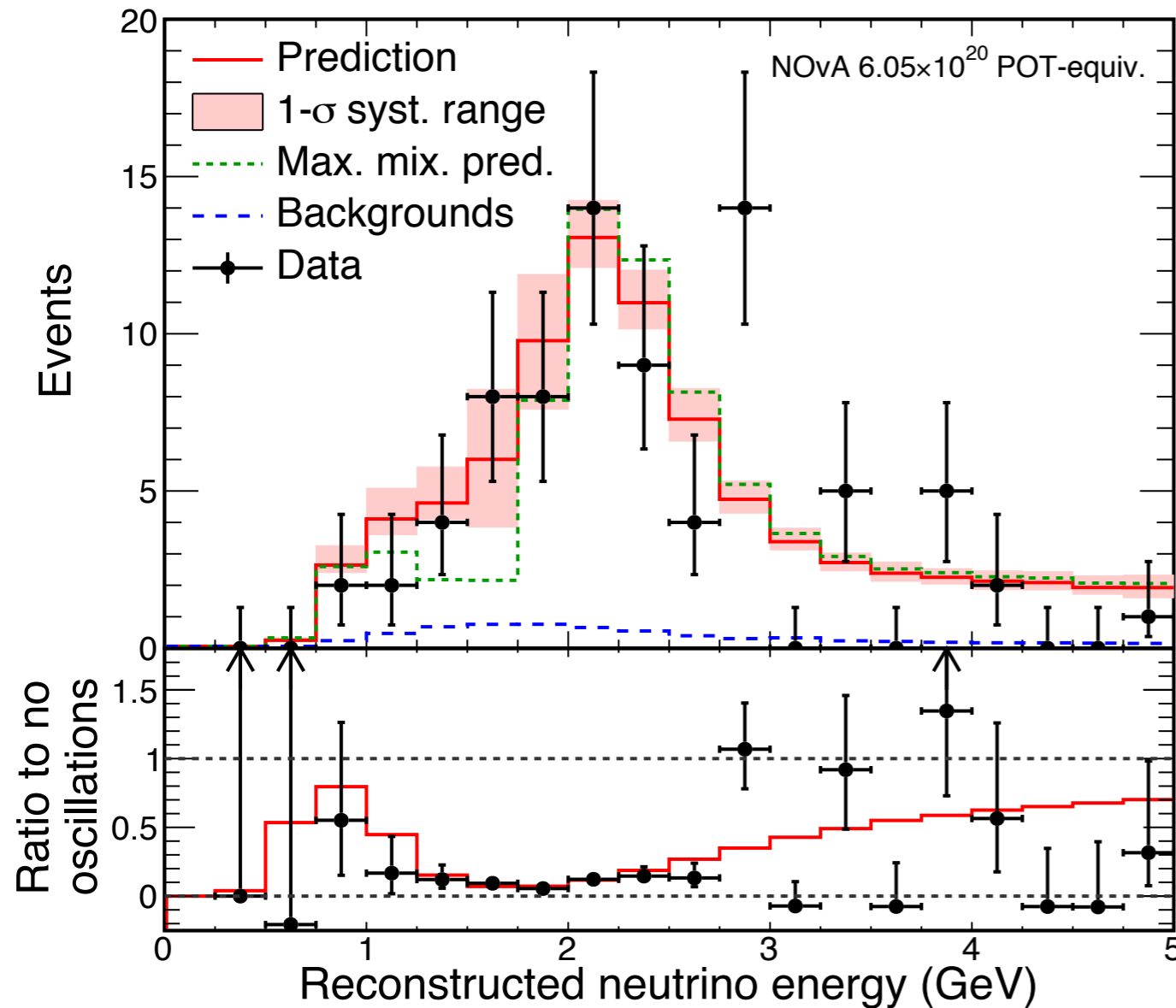
# Extrapolation

- Use high statistics ND data/MC to adjust prediction at FD
  - Translate ND data/MC observation to true energy
  - Oscillate ratio to the FD
  - Smear back into reconstructed energy
- Reduces systematic uncertainties

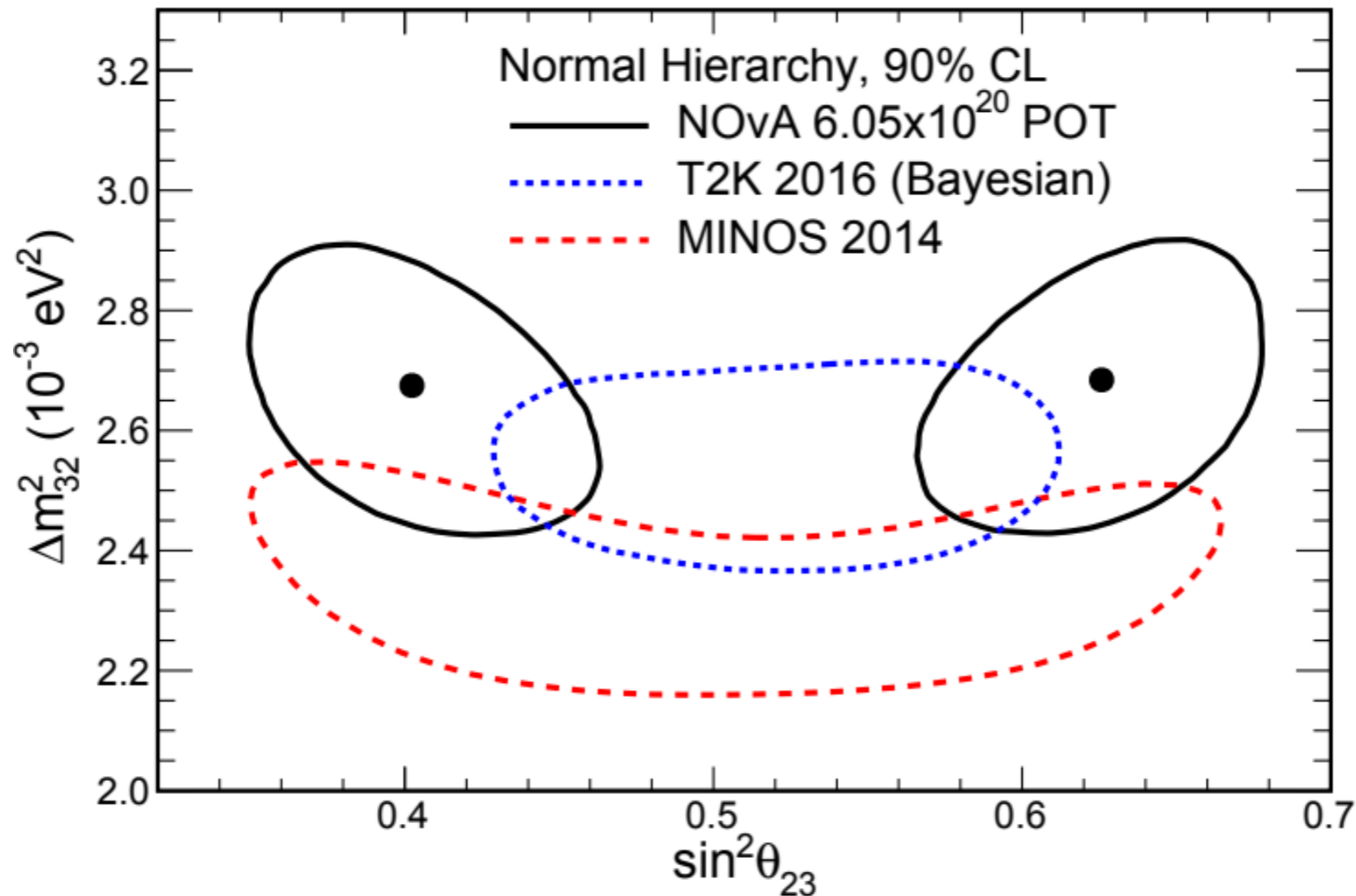


# $\nu_\mu \rightarrow \nu_\mu$ Oscillation Results

PRL 118, 151802



- 473 +/- 30 events predicted in the absence of oscillations
- Observed 78 events
- 82 events predicted at the best fit point including 3.7 beam background and 2.9 cosmic induced events
- Maximal mixing disfavored at 2.6 sigma



Best fit (in NH):

$$|\Delta m_{32}^2| = 2.67 \pm 0.11 \times 10^{-3} \text{eV}^2$$

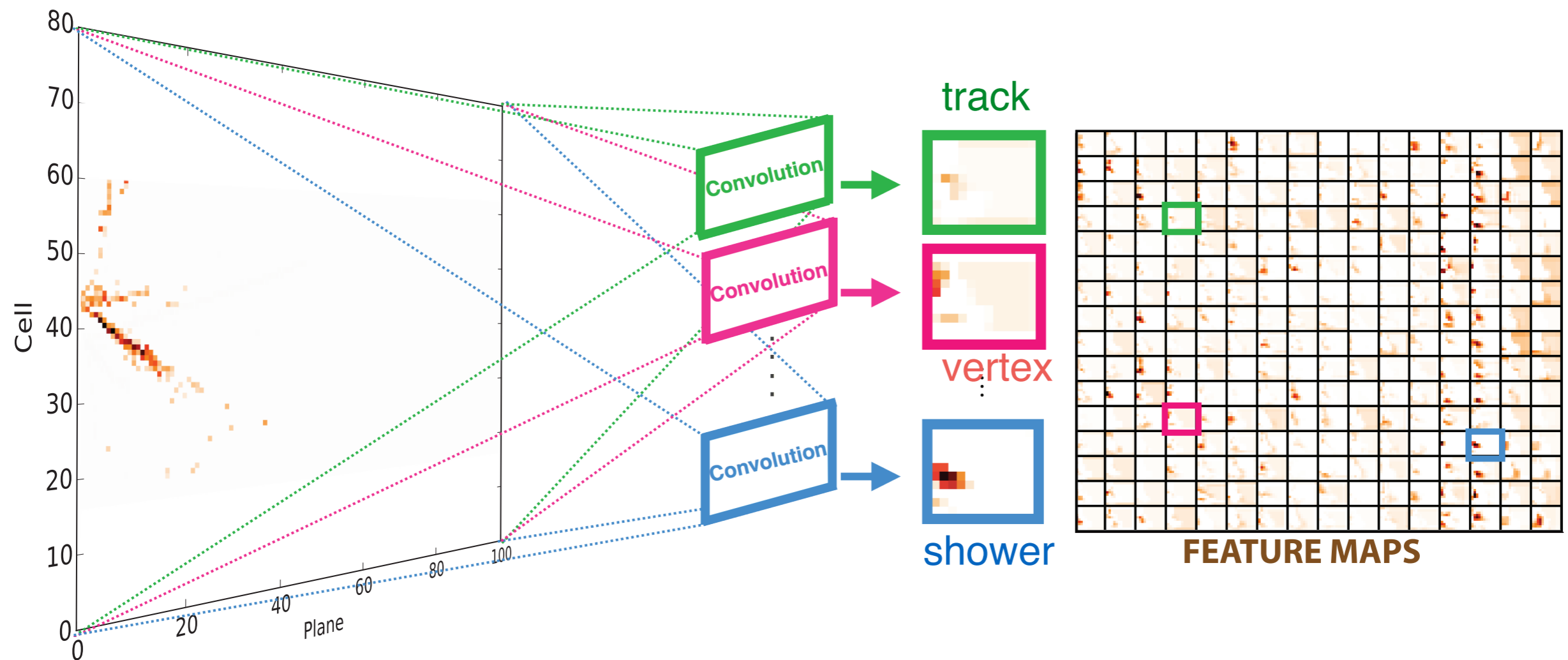
$$\sin^2 \theta_{23} = 0.404_{-0.022}^{+0.030} (0.624_{-0.030}^{+0.022})$$

Update coming in fall 2017 with  
50% more data

# Electron Neutrino Appearance Analysis

# Electron Neutrino Selection

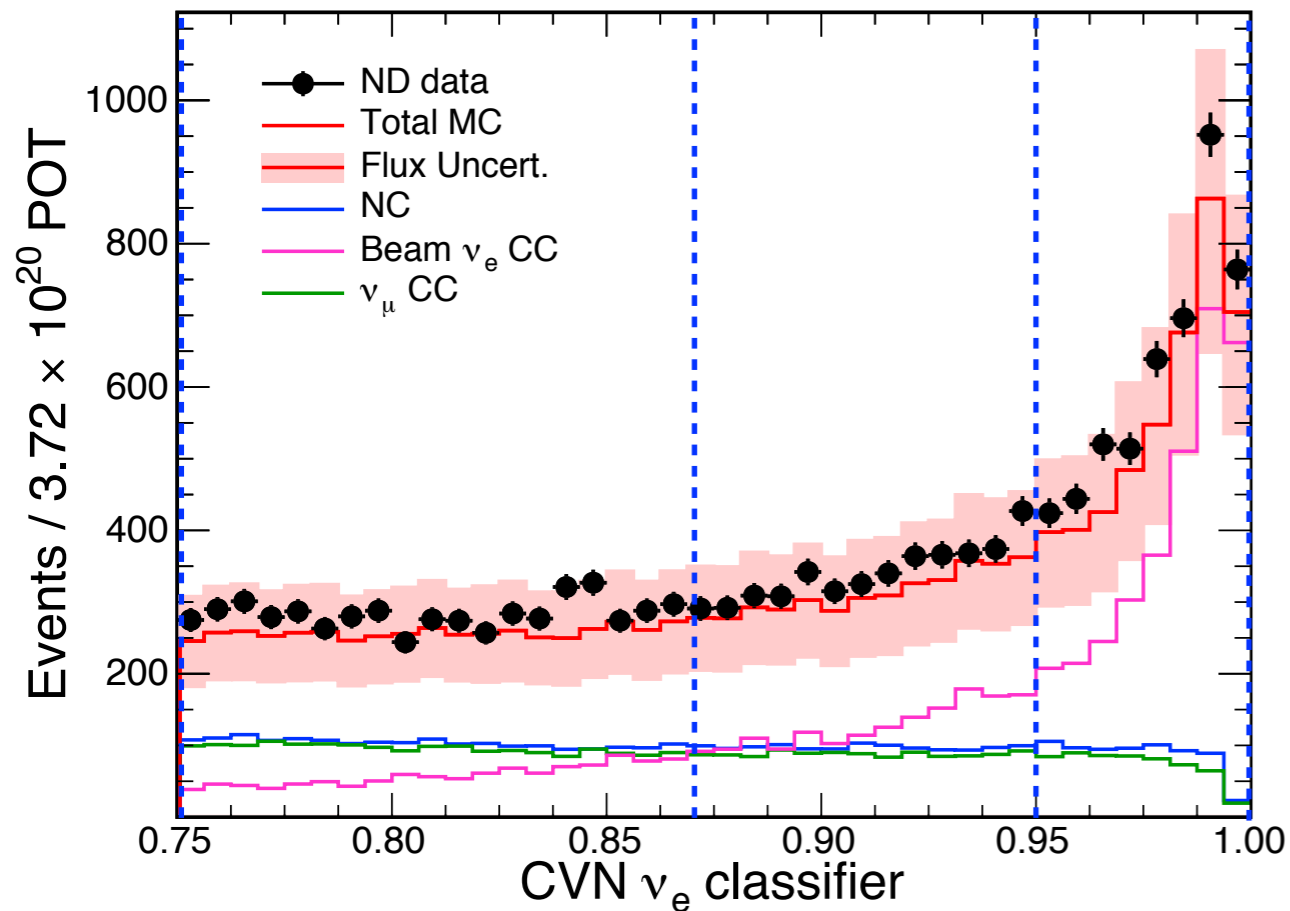
- Implemented a deep convolutional neural network
- Input 2D pixel maps of NOvA events
- Output classification as  $\nu_\mu$ ,  $\nu_e$ ,  $\nu_\tau$ , NC, or cosmic
- Below: an electron neutrino interaction and the first layer of feature maps extracted from the convolutional kernels



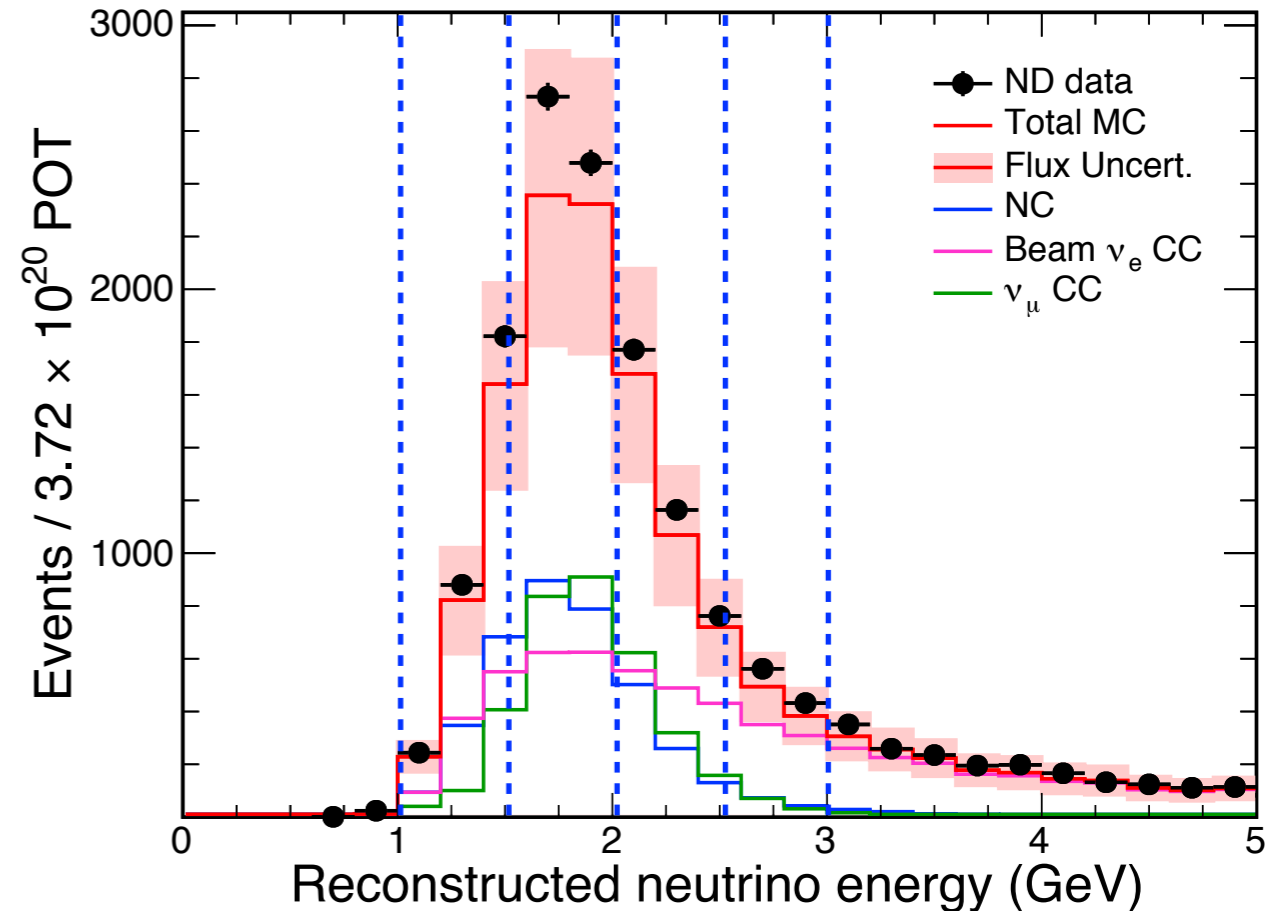
A. Aurisano and A. Radovic and D. Rocco et. al, JINST 11 P09001 (2016)

# Electron Neutrino Selection

NOvA Preliminary



NOvA Preliminary

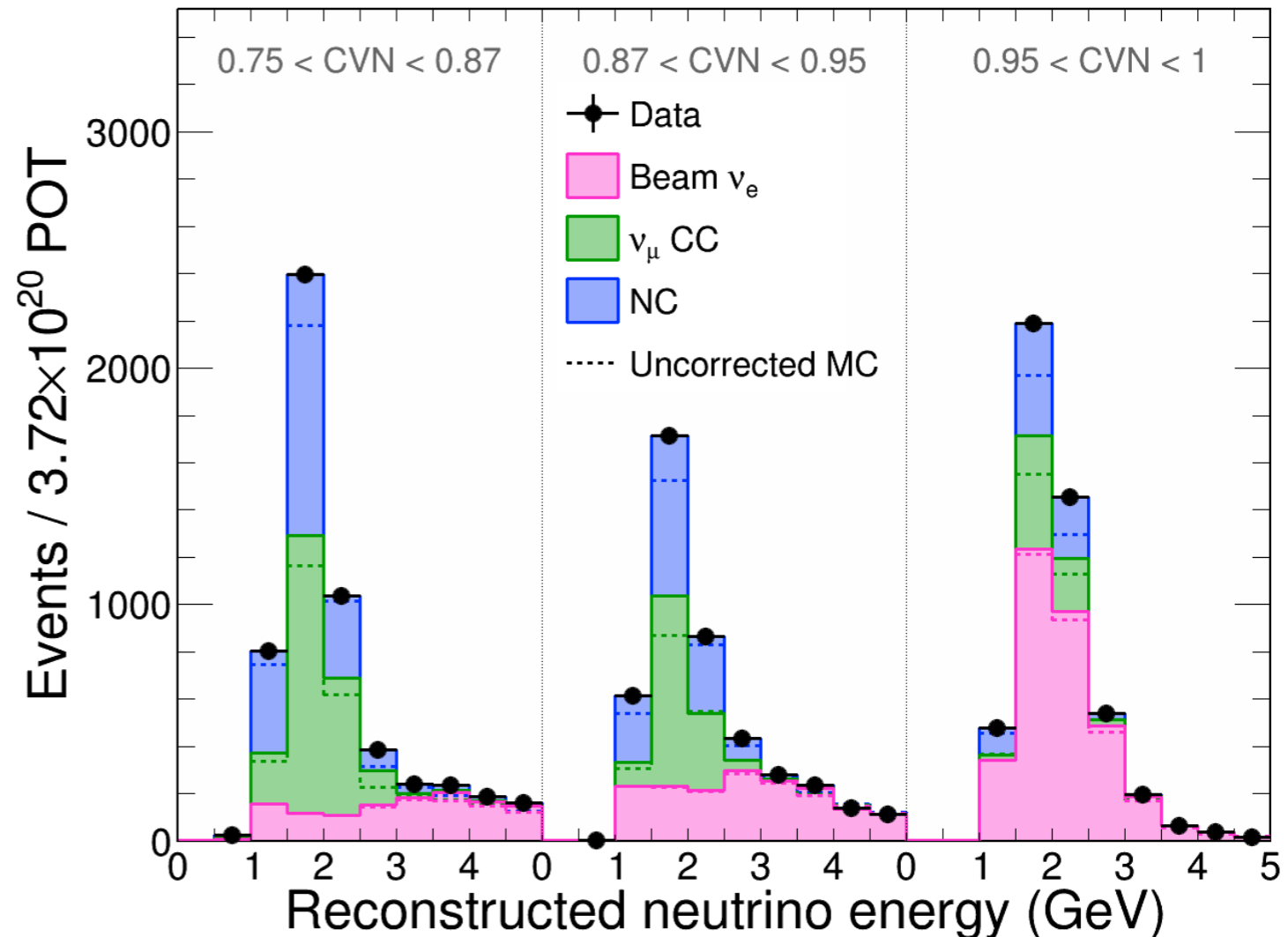


- 73%  $\nu_e$  CC selection efficiency, 76% purity with CVN classifier
- Good ND Data/MC agreement
- Equivalent to 30% increase in exposure compared to conventional IDs
- Bin analysis in four bins of energy and three of CVN



# Data Driven Background Corrections

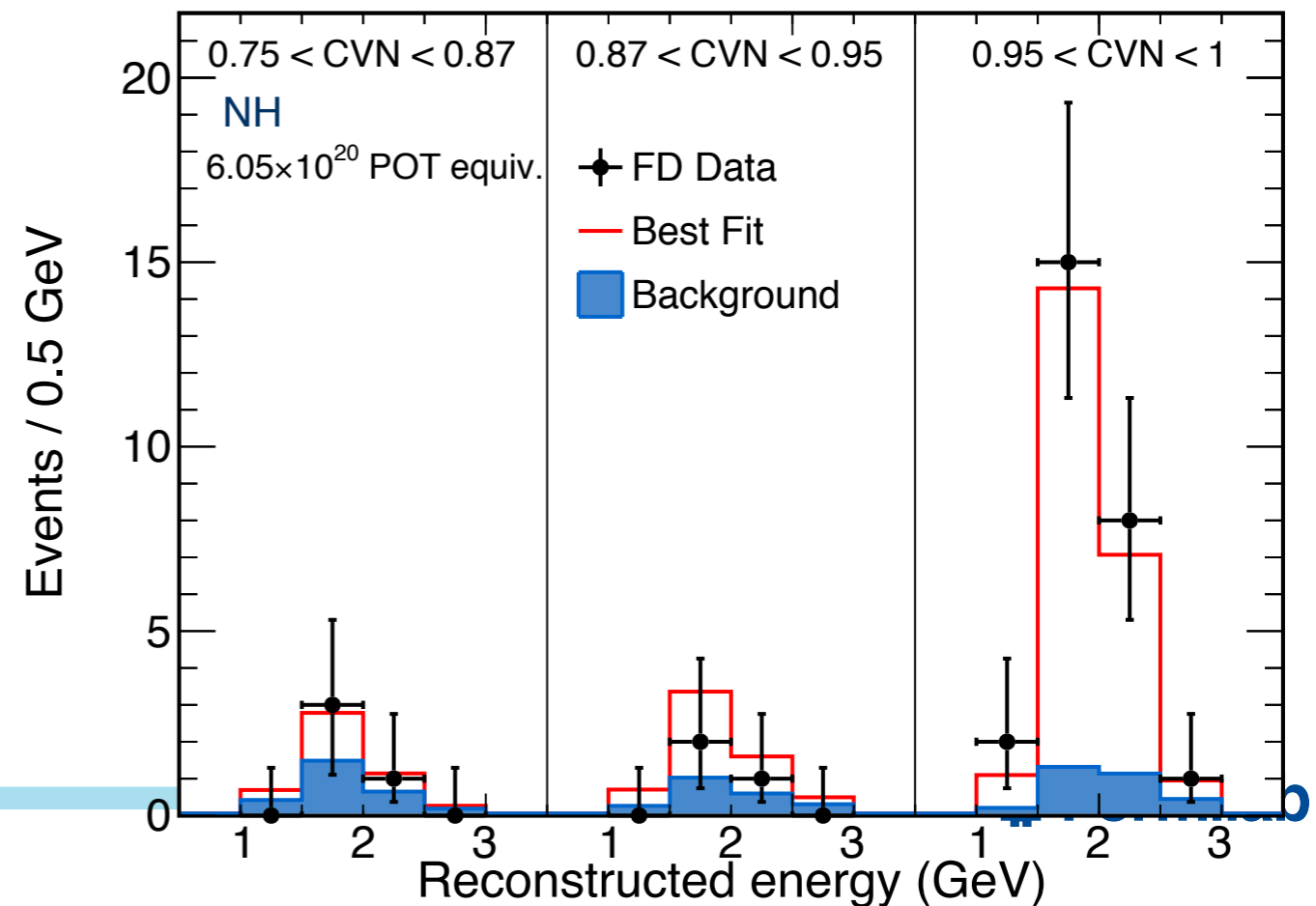
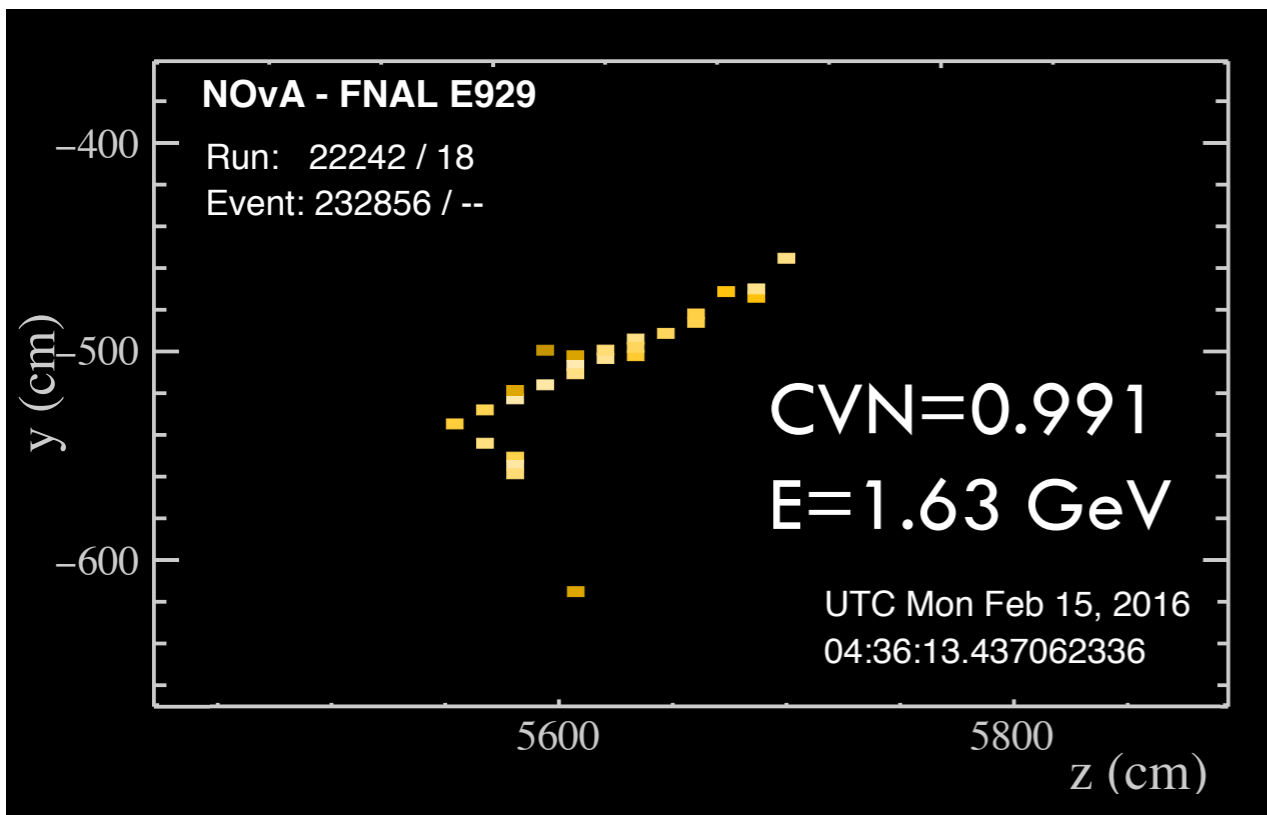
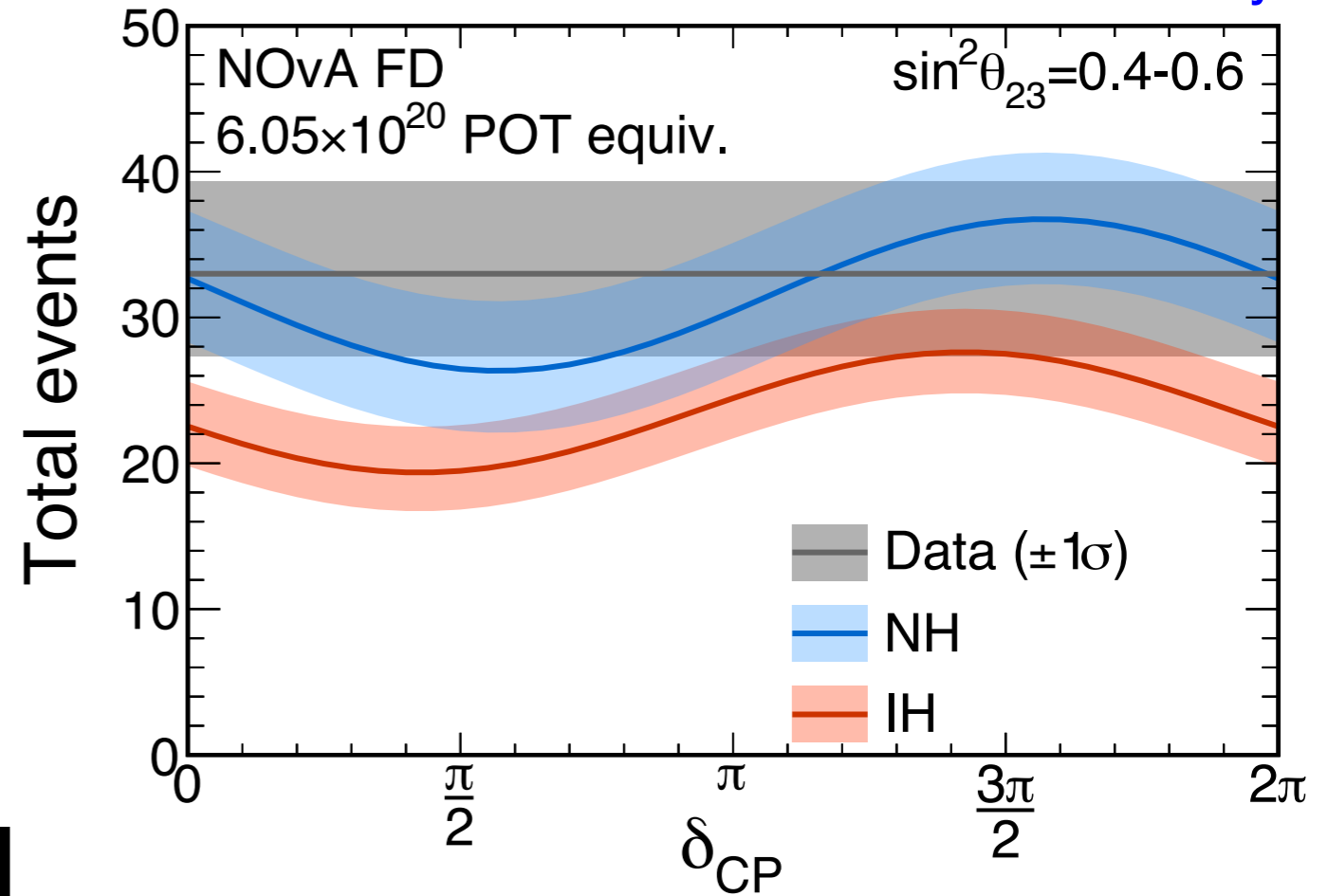
- $\nu_e$  CC selection in the ND picks out FD backgrounds
  - beam  $\nu_e$  CC
  - $\nu_\mu$  CC
  - NC
- $\sim 10\%$  excess of data over MC in the ND
- Extrapolate data/MC differences to adjust FD prediction
- Data-driven methods to decompose data into components
  - beam  $\nu_e$  CC adjusted by parent  $\pi^+$  and  $k^+$
  - Michel electron templates used to adjust NC and  $\nu_\mu$  CC



# $\nu_\mu \rightarrow \nu_e$ Oscillation Results

- Observe 33 events on background of  $8.2 \pm 0.8$  events

NOvA Preliminary



# $\nu_\mu \rightarrow \nu_e$ Oscillation Results

- Fit for hierarchy,  $\delta_{CP}$ ,  $\sin^2\theta_{23}$ 
  - Constrain  $\sin^2 2\theta_{13} = 0.085 \pm 0.005$  from reactor experiments
  - Simultaneous fit with NOvA disappearance data

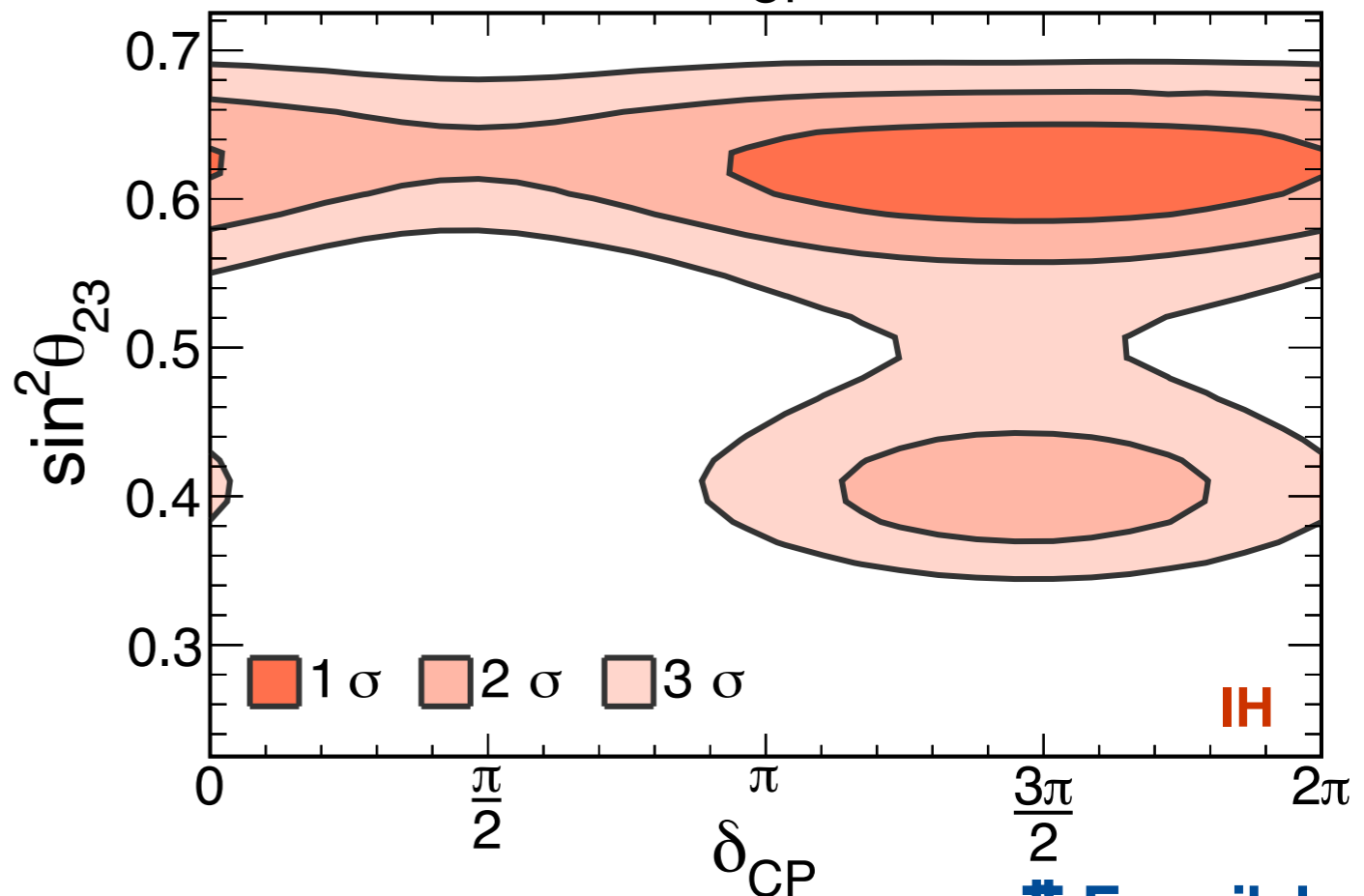
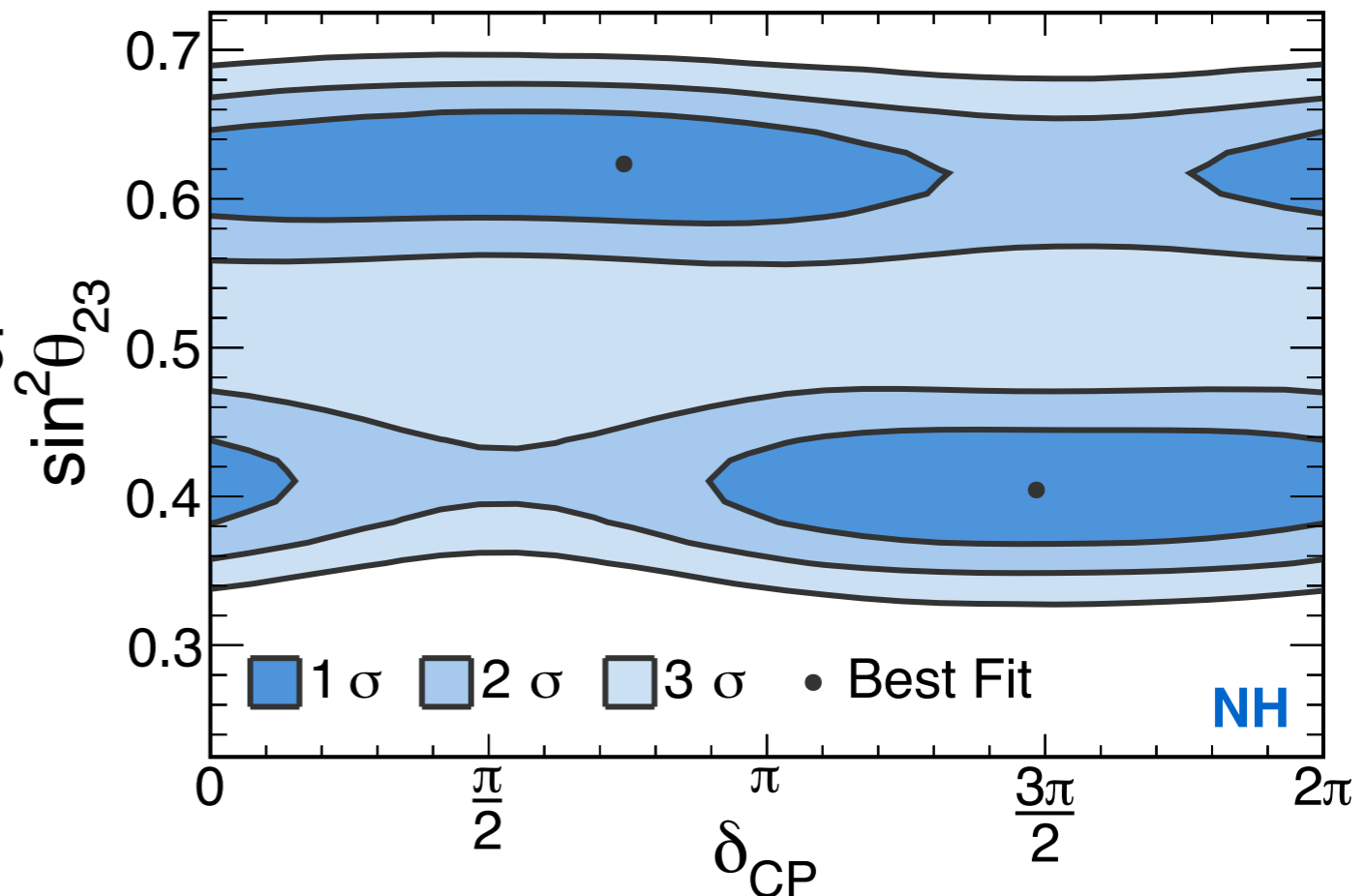
- Global best fit, two degenerate points in Normal Hierarchy

$$\delta_{cp} = 1.48\pi, \sin^2(\theta_{23}) = 0.404$$

$$\delta_{cp} = 0.74\pi, \sin^2(\theta_{23}) = 0.623$$

- best fit IH-NH,  $\Delta\chi^2=0.47$
- Lower octant, IH is disfavored at greater than 93% C.L for all values of  $\delta_{CP}$

**PRL 118, 231801**



# $\nu_\mu \rightarrow \nu_e$ Oscillation Results

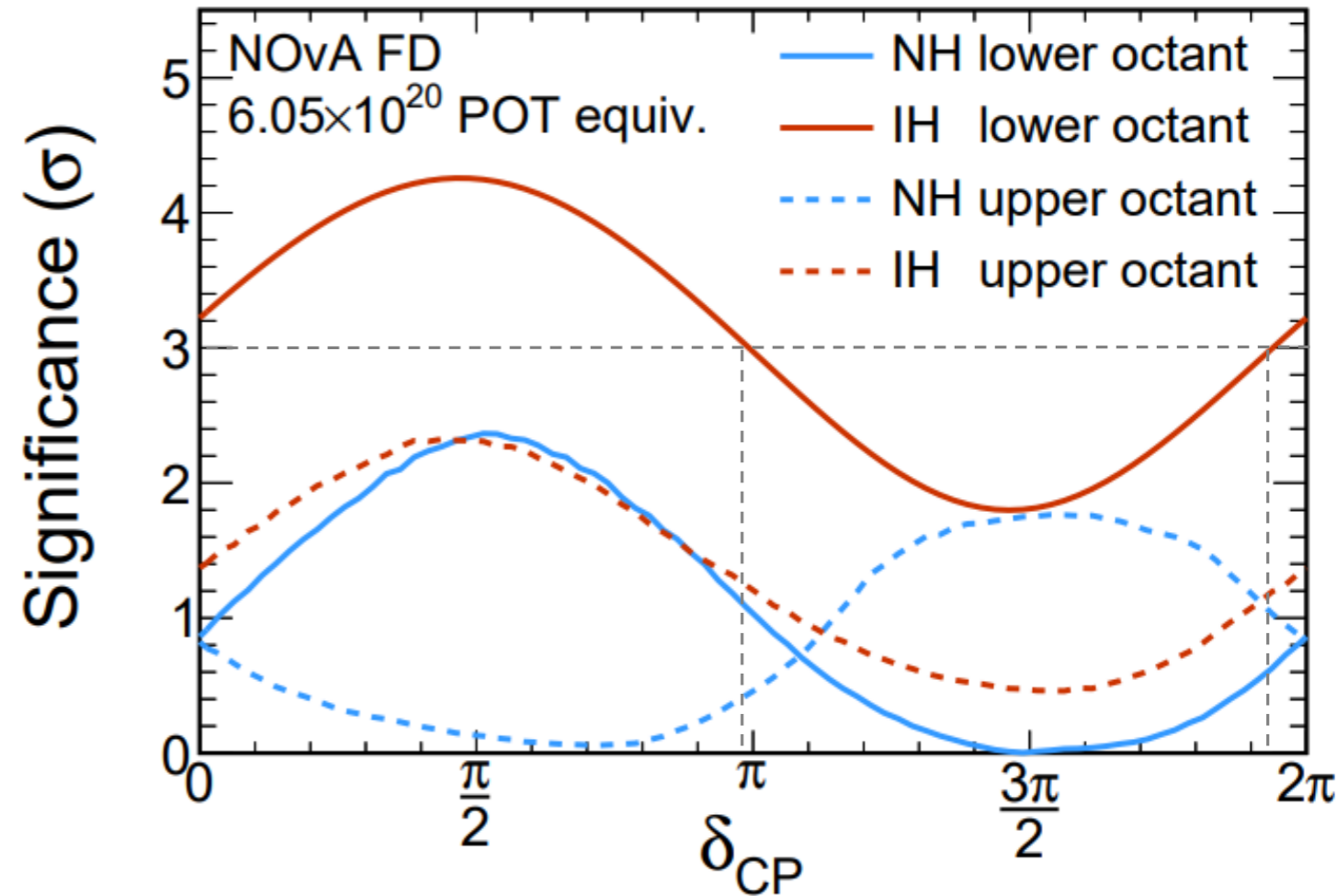
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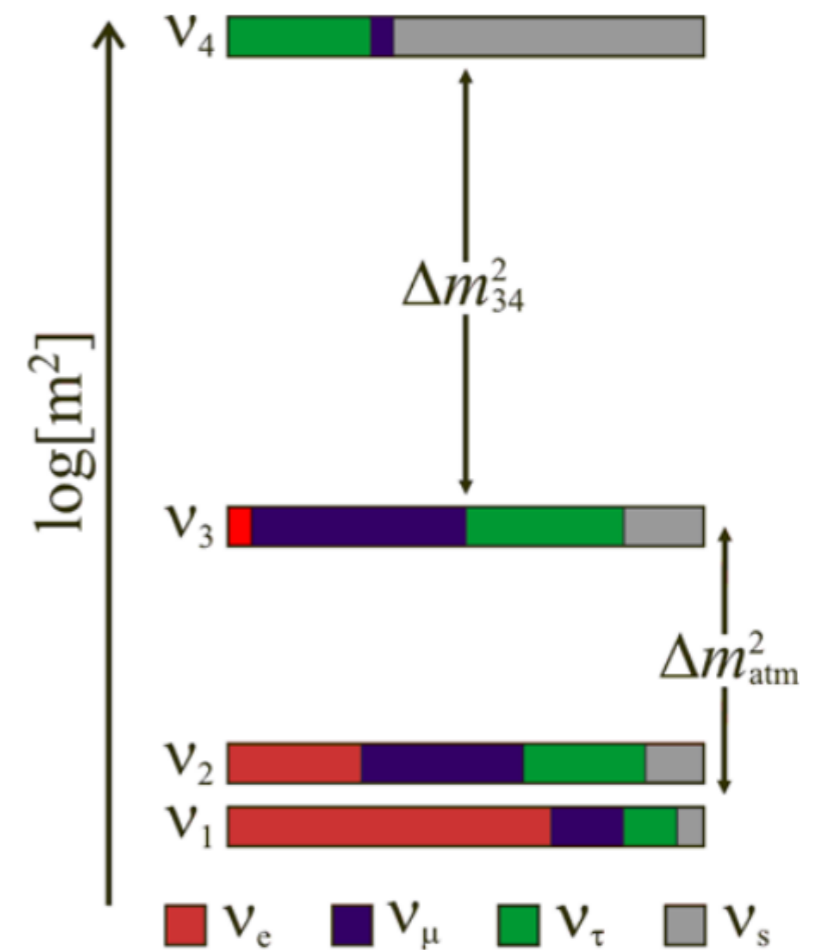
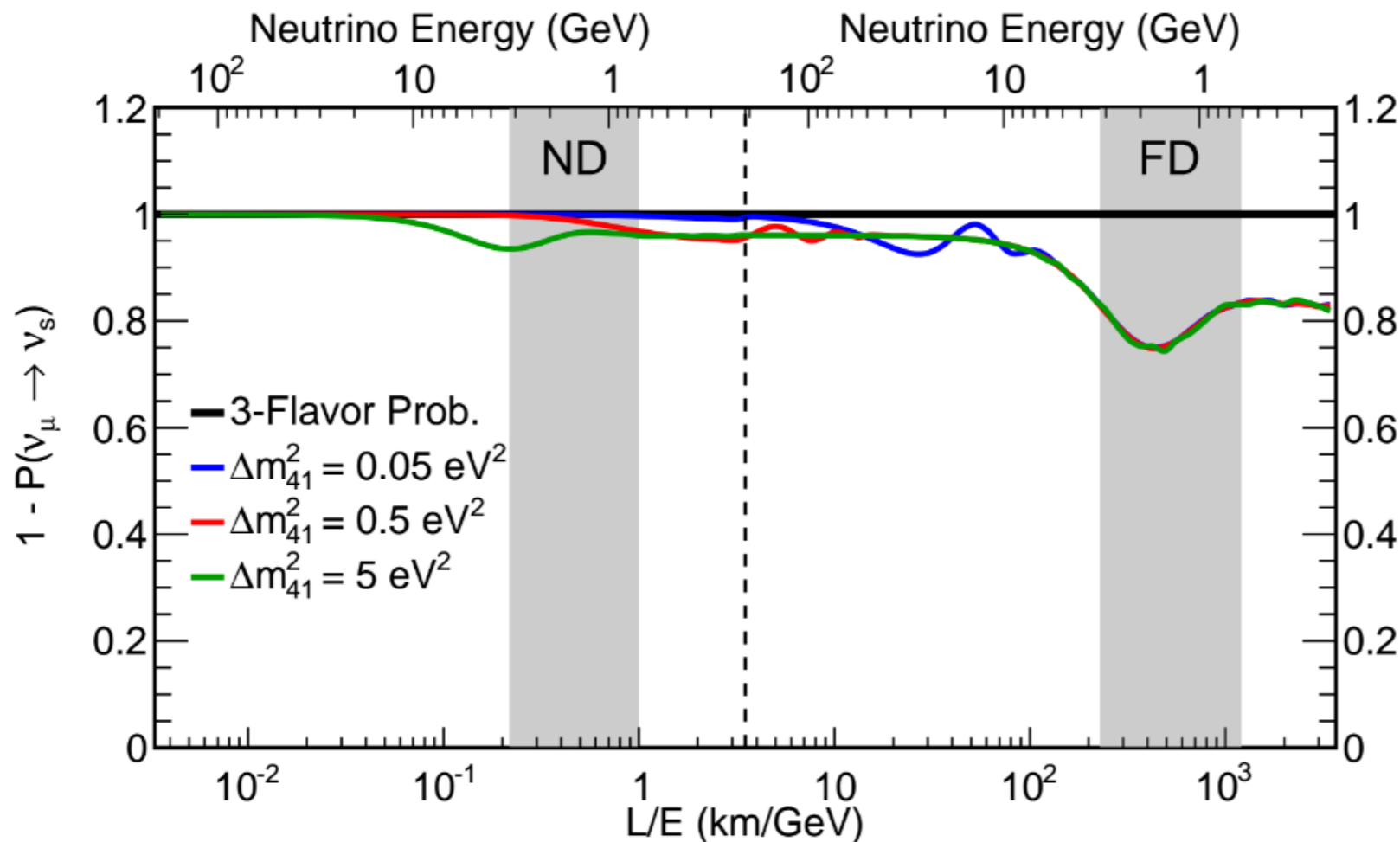
Update coming in fall 2017 with 50% more data

**PRL 118, 231801**

# Neutral Current Disappearance Analysis Updated for 2017

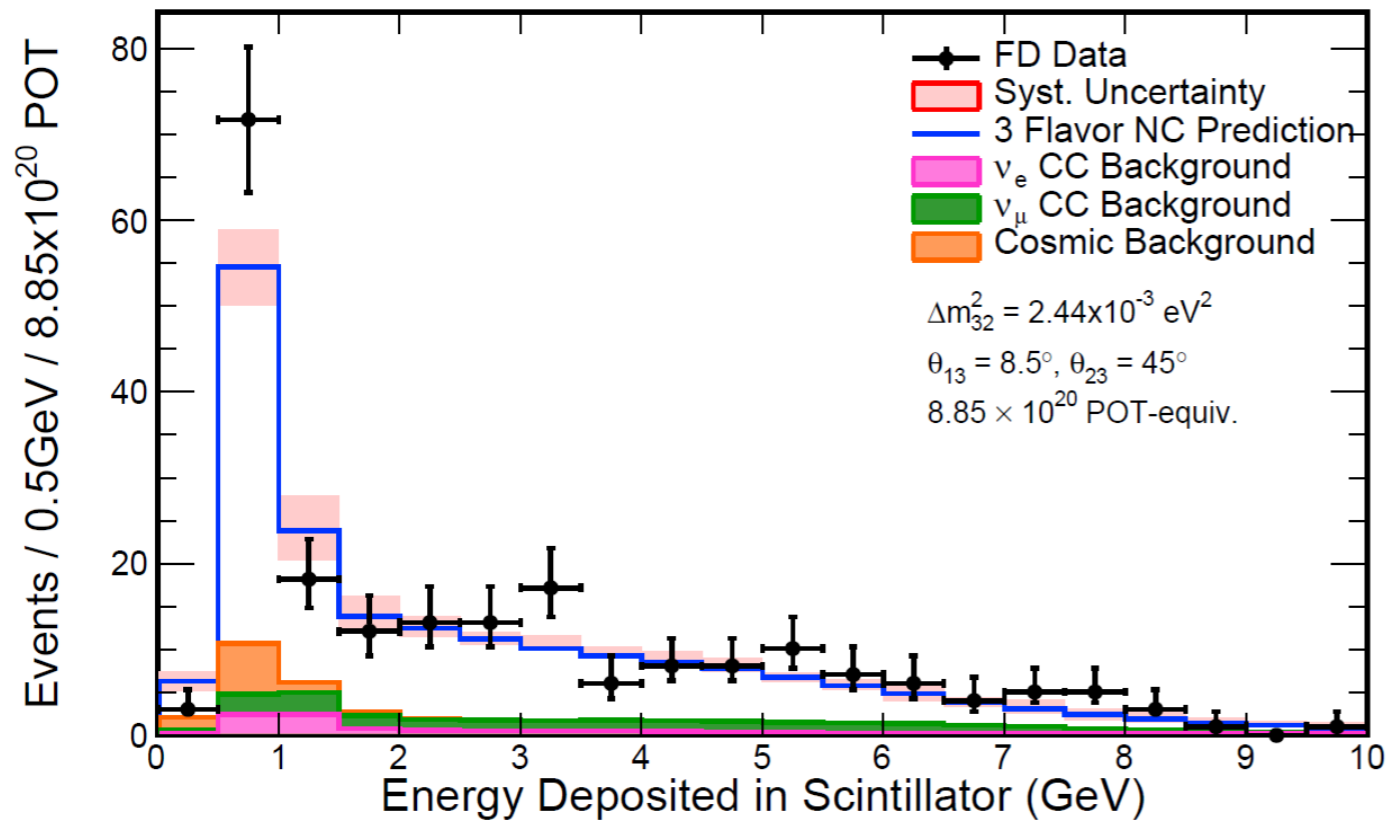
# NC Disappearance Measurement

- NC spectrum is unaffected by oscillation of active flavors
- Search for depletion of NC events in FD to interpret in a 3+1 model
- 2017 analysis updated with 50% exposure increase, simulation improvements, improvements to event selection, cosmic rejection



# NC Disappearance Results

NOvA Preliminary



**3 flavor prediction:  $191.16 \pm 13.82(\text{stat}) \pm 21.99 (\text{syst})$**

**Measured 214 NC candidates**

**No evidence for a depletion of NC events**

$$R_{\text{NC}} \equiv \frac{F^{\text{data}} - \sum F^{\text{pred}}(\text{bkg})}{F^{\text{pred}}(\text{NC})}$$

No NC disappearance  $\rightarrow R = 1$

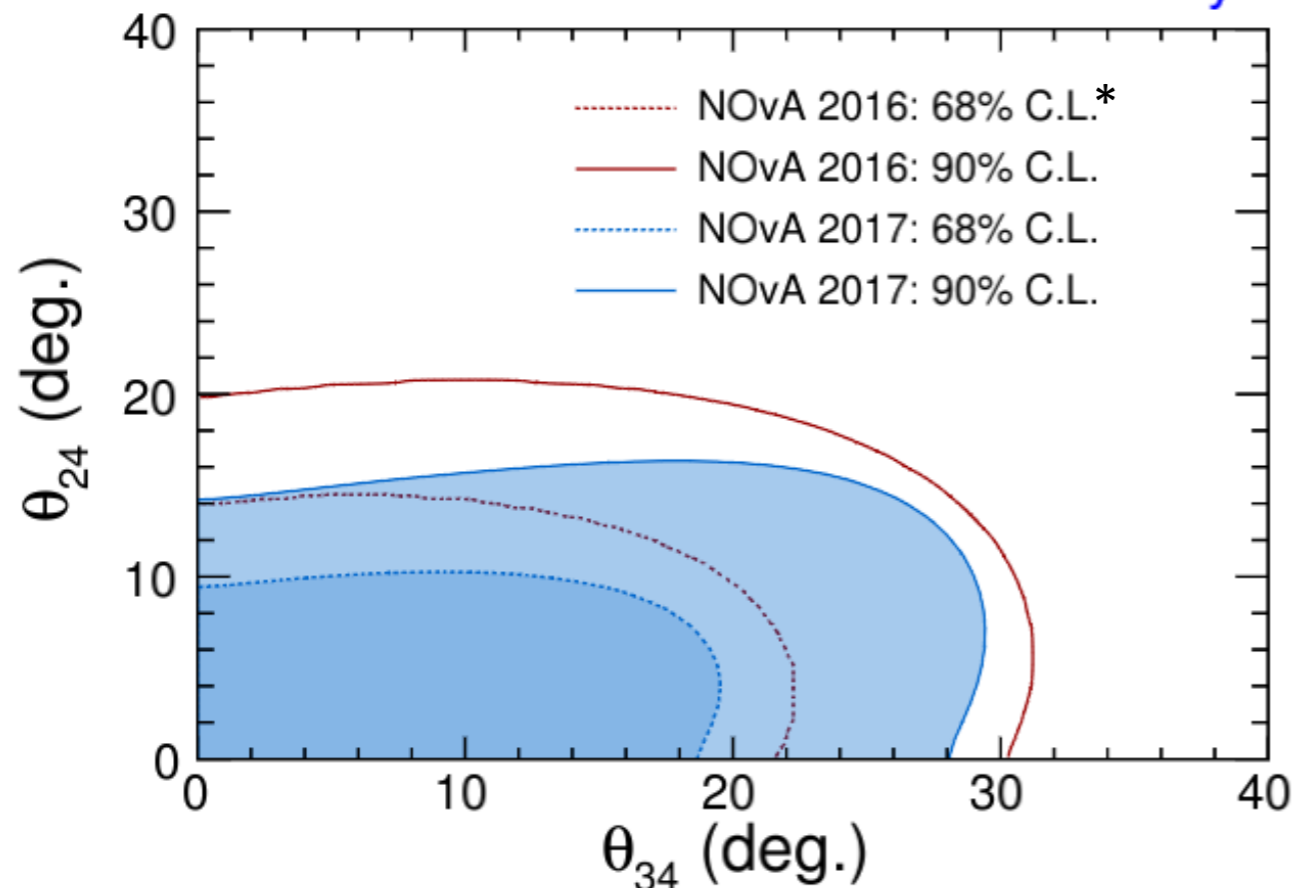
Using **NOvA**'s two degenerate best fit points for  $\sin^2(\theta_{23})$ ,  $|\Delta m_{32}^2|$ , and  $\delta_{CP}$  (NH)

R-values	0 – 2.5 GeV
$\theta_{23} = 45$ (2016)	<b>1.190</b> $\pm 0.160$ ( <i>stat.</i> ) $^{+0.080}_{-0.130}$ ( <i>syst.</i> )
$\theta_{23} = 45$ (2017)	<b>1.190</b> $\pm 0.123$ ( <i>stat.</i> ) $^{+0.143}_{-0.124}$ ( <i>syst.</i> )
$\theta_{23} < 45$	<b>1.179</b> $\pm 0.123$ ( <i>stat.</i> ) $^{+0.142}_{-0.124}$ ( <i>syst.</i> )
$\theta_{23} > 45$	<b>1.176</b> $\pm 0.123$ ( <i>stat.</i> ) $^{+0.142}_{-0.124}$ ( <i>syst.</i> )

# NC Disappearance Results

- Shape-based fit of  $\theta_{24}$  and  $\theta_{34}$ , no FC corrections
- Constrain with NOvA's degenerate best fit points for  $\sin^2(\theta_{23})$ ,  $|\Delta m_{32}^2|$ , and  $\delta_{CP}$  (NH)
- Profile over  $\sin^2(\theta_{23})$  and  $\delta_{24}$

NOvA Preliminary



	$\theta_{24}$	$\theta_{34}$	$ U_{\mu 4} ^2$	$ U_{\tau 4} ^2$
NOvA 2016	$20.8^\circ$	$31.2^\circ$	0.126	0.268
NOvA 2017	$16.2^\circ$	$29.8^\circ$	0.078	0.228
MINOS	$7.3^\circ$	$26.6^\circ$	0.016	0.20
SuperK	$11.7^\circ$	$25.1^\circ$	0.041	0.18
IceCube	$4.1^\circ$	-	0.005	-
IceCube-DeepCore	$19.4^\circ$	$22.8^\circ$	0.11	0.15

In a 3+1 analysis, for  $\Delta m_{41}^2 = 0.5 \text{ eV}^2$ :

$$\theta_{24} < 16.2 \text{ at } 90\% \text{ C.L.}$$

$$\theta_{34} < 29.8 \text{ at } 90\% \text{ C.L.}$$

\* 2016 fit is rate only and applies maximal mixing constraint

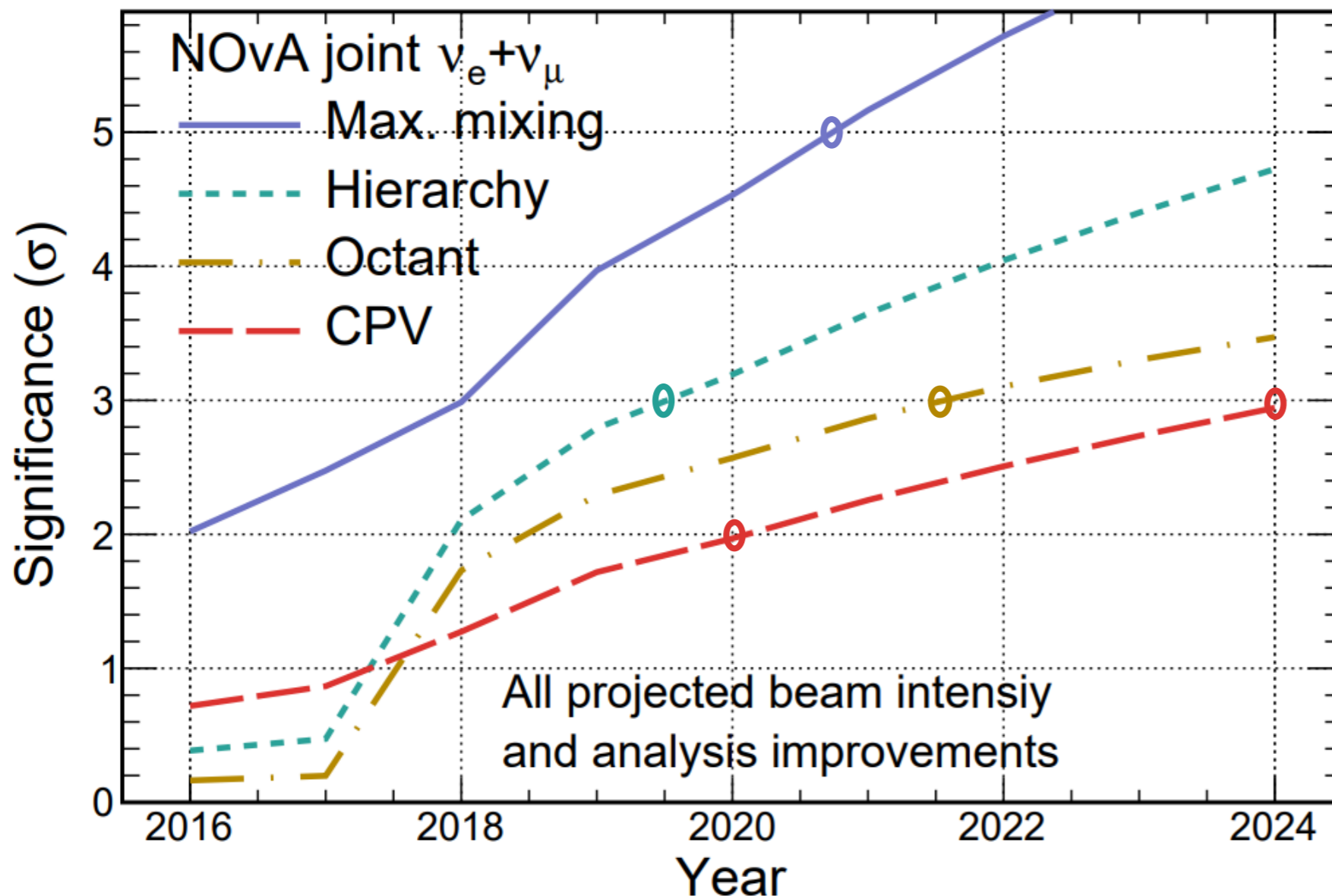


# Looking Forward

- Switched to anti-neutrino running in February 2017
- Run 50% neutrino, 50% anti-neutrino after 2018

Normal  $\delta_{CP}=3\pi/2$ ,  $\sin^2\theta_{23}=0.403$   
 $\Delta m_{32}^2=2.5\times 10^{-3}\text{eV}^2$ ,  $\sin^2\theta_{13}=0.022$

## NOvA Simulation



- improved systematics
- 25% exposure increase from analysis improvements
- 40 weeks of beam starting 2018
- 800 kW in 2019, 900 kW in 2021

# Conclusions

- A broad program of NOvA physics
- Best fit to  $\nu_\mu$  disappearance data disfavors maximal mixing at  $2.6\sigma$
- Joint fit of NOvA appearance and disappearance data
  - Weak preference for normal hierarchy
  - Inverted hierarchy, lower octant is excluded at  $> 93\%$  C.L.
- Updated NC disappearance analysis with  $8.85 \times 10^{20}$  POT sees no evidence for sterile neutrino mixing
- Updated  $\nu_\mu$  disappearance and  $\nu_e$  appearance with  $8.85 \times 10^{20}$  POT coming this fall
- Anti-neutrino running underway with results in summer 2018

# Questions



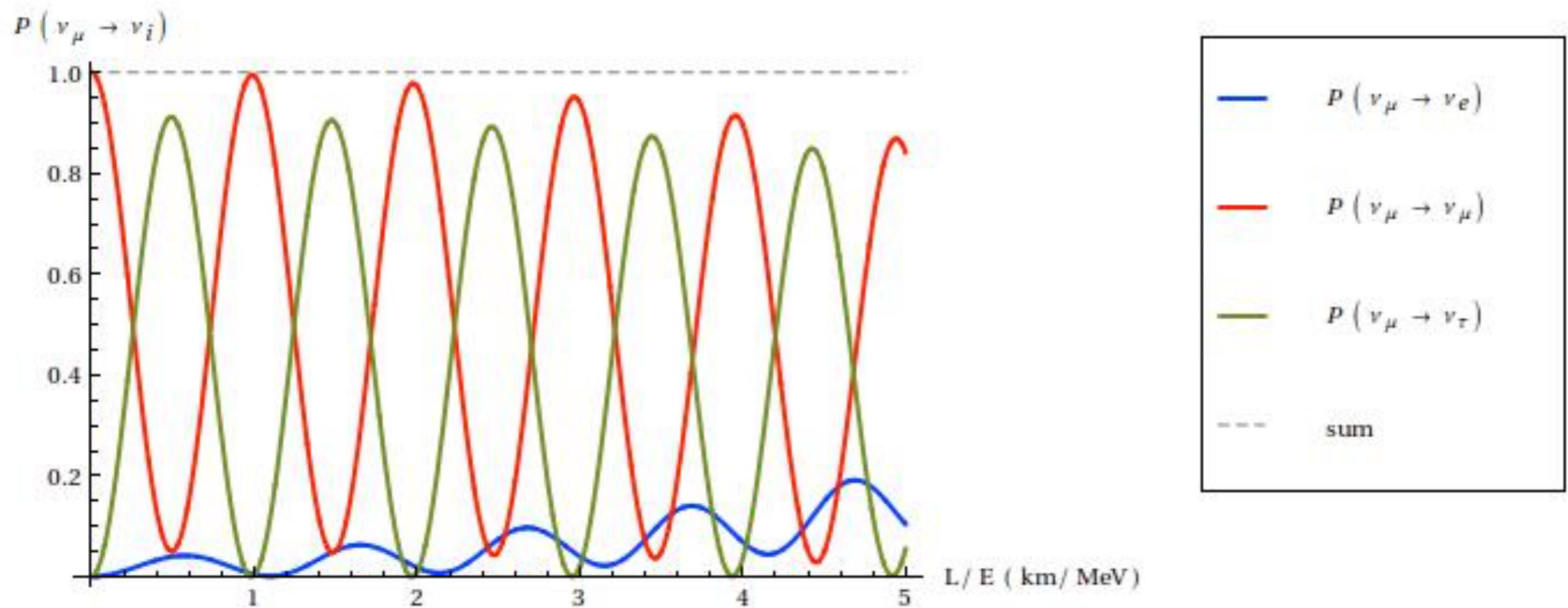


Image from: <http://invisibles.eu/outreach/entry/ceaseless-transformation-three-neutrinos>

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\nu_\mu \rightarrow \nu_\mu$$

$$\nu_\mu \rightarrow \nu_\tau$$

atmospheric and  
long-baseline

$$\nu_e \rightarrow \nu_e$$

$$\nu_\mu \rightarrow \nu_e$$

reactor and  
long-baseline

$$\nu_e \rightarrow \nu_e$$

$$\nu_e \rightarrow \nu_\mu + \nu_\tau$$

solar and  
reactor

## $\nu_\mu \rightarrow \nu_e$ Appearance channel

$$P(\nu_\mu \rightarrow \nu_e) \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}} \right|^2$$

$$= P_{atm} + P_{sol} + 2\sqrt{P_{atm}P_{sol}} (\cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta)$$

$$\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} - aL)}{\Delta_{31} - aL} \Delta_{31}$$

Depends on relative sign of "a" and  $\Delta_{31}$

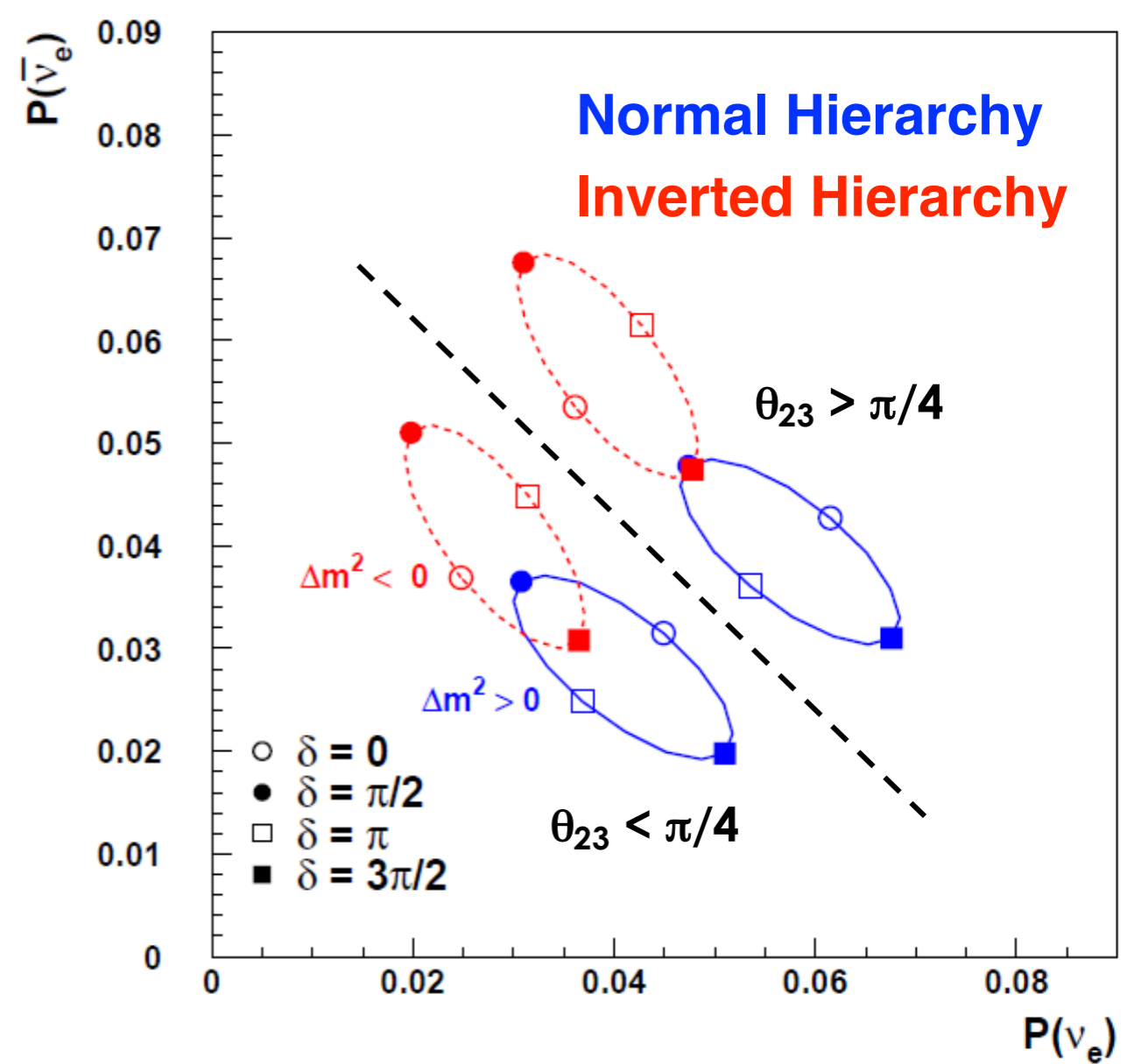
$$\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{aL} \Delta_{21}$$

$$a = \frac{G_F N_e}{\sqrt{2}} \approx \frac{1}{3500 \text{ km}}$$

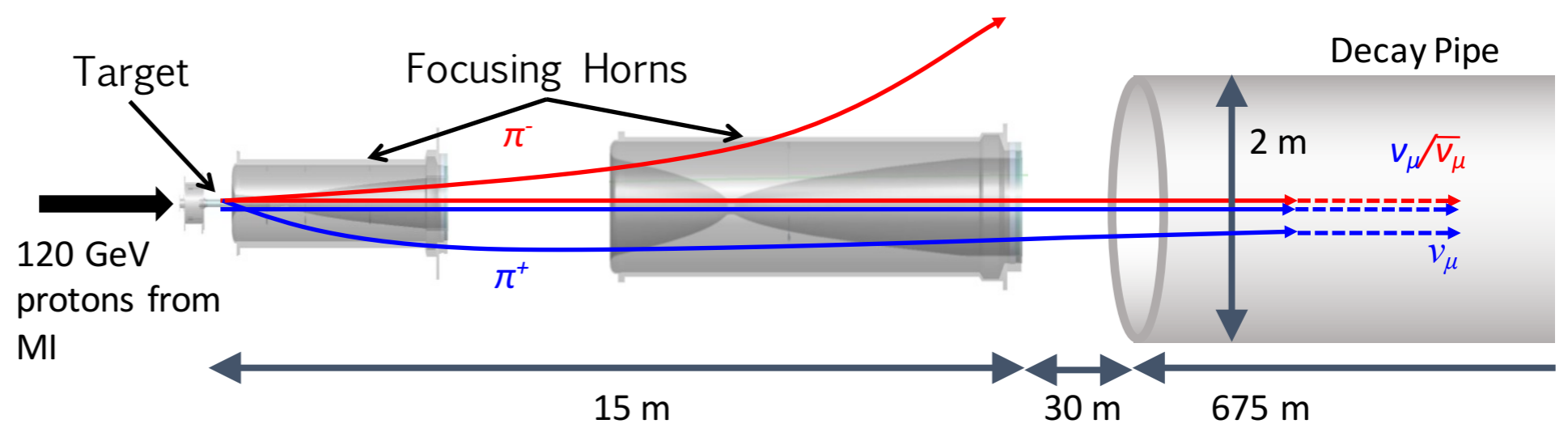
$aL=0.08$  for  $L=295\text{km}$  T2K baseline

$aL=0.23$  for  $L=810\text{km}$  NOvA baseline

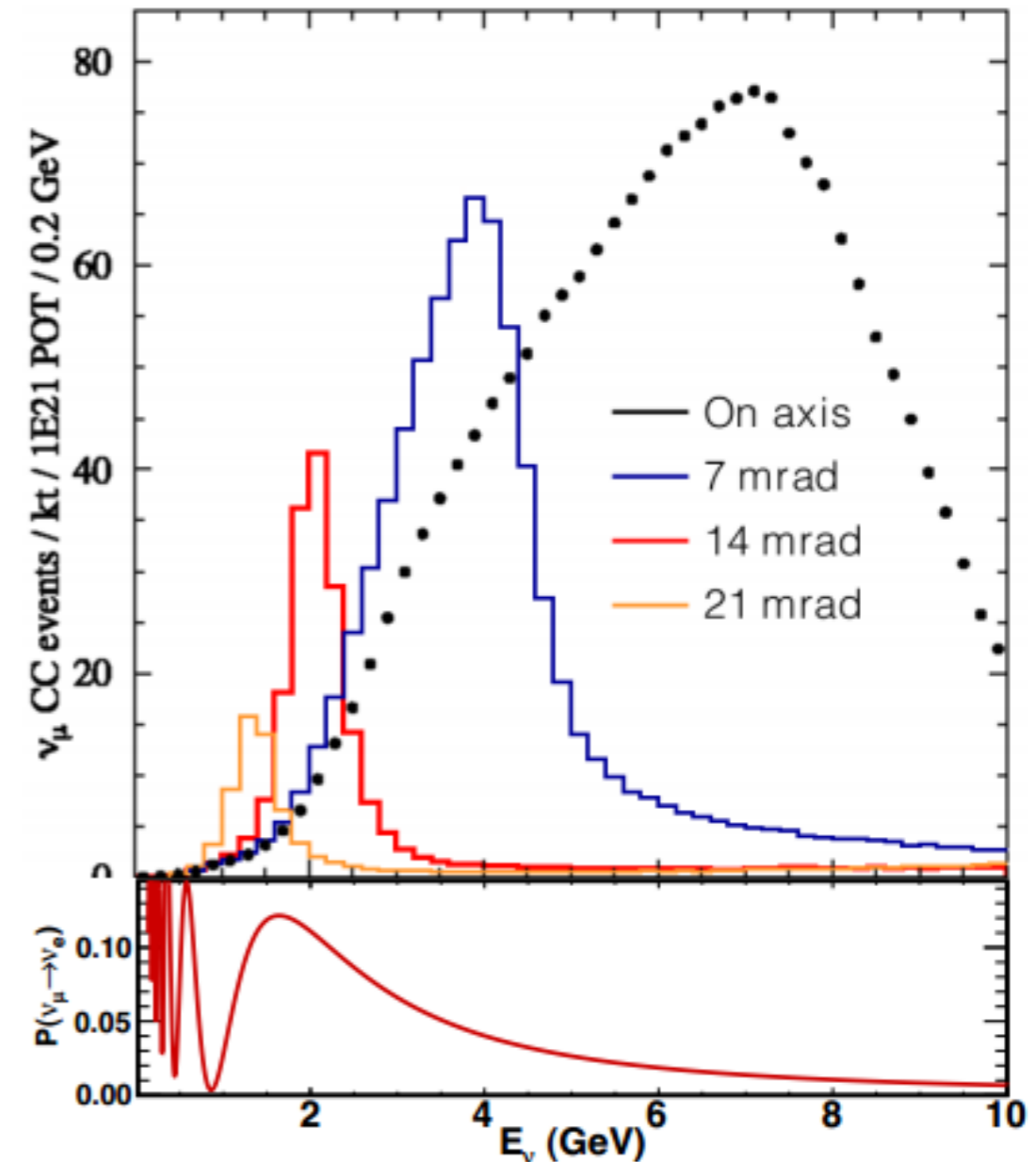
Oscillation probability is sensitive to: **mass ordering**, **CP violating phase**, and  $\theta_{23}$  octant.



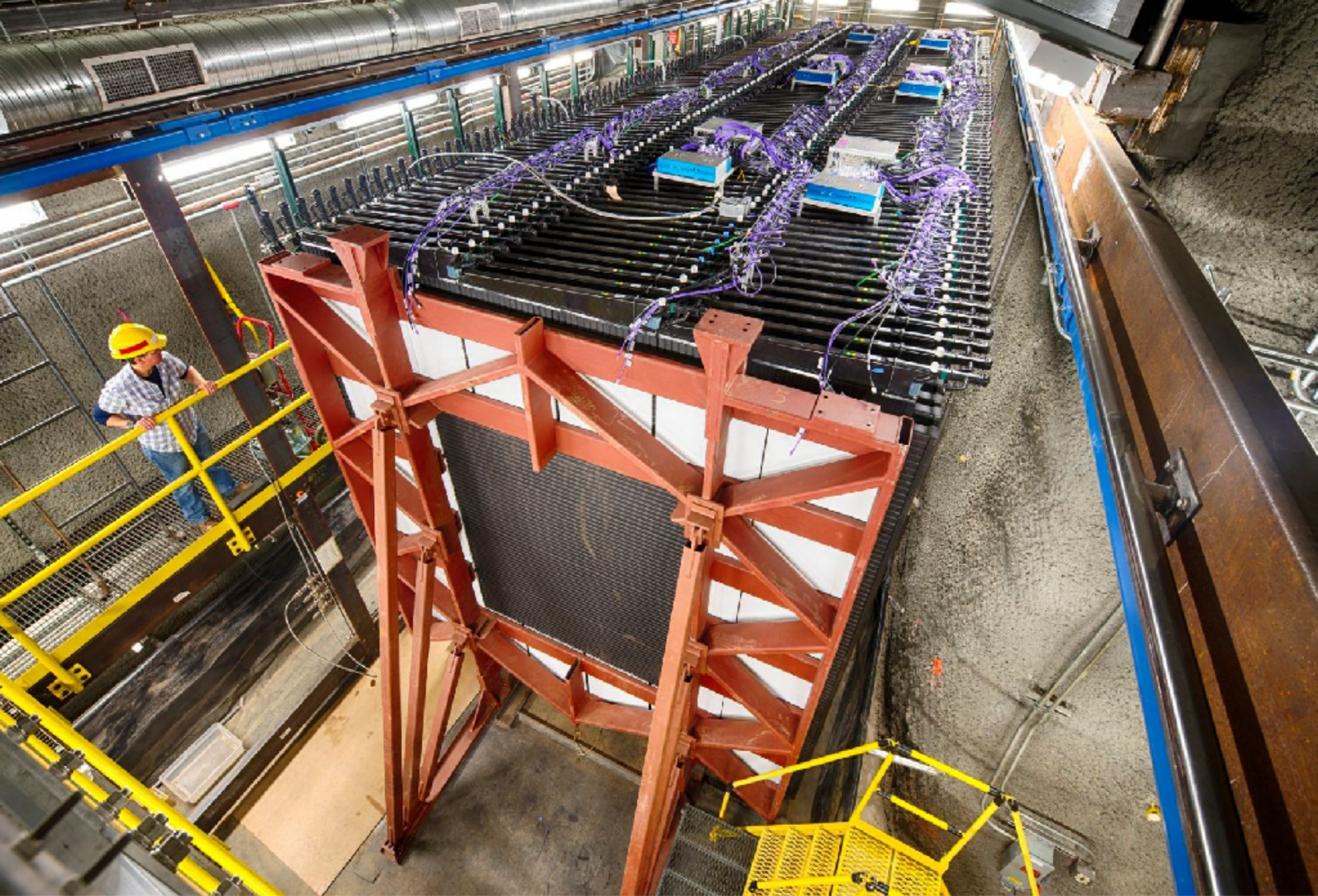
# NuMI Beam



- 120 GeV protons extracted from the Main Injector at Fermilab in  $10 \mu\text{s}$  spills
- Magnetic focusing horns allow selection of charge sign for selecting a neutrino or anti-neutrino beam
- 14.6 milli-radians off-axis, narrow beam around oscillation maximum
- Beam 97.5%  $\nu_\mu$  with 0.7%  $\nu_e$  and 1.8% wrong-sign contamination







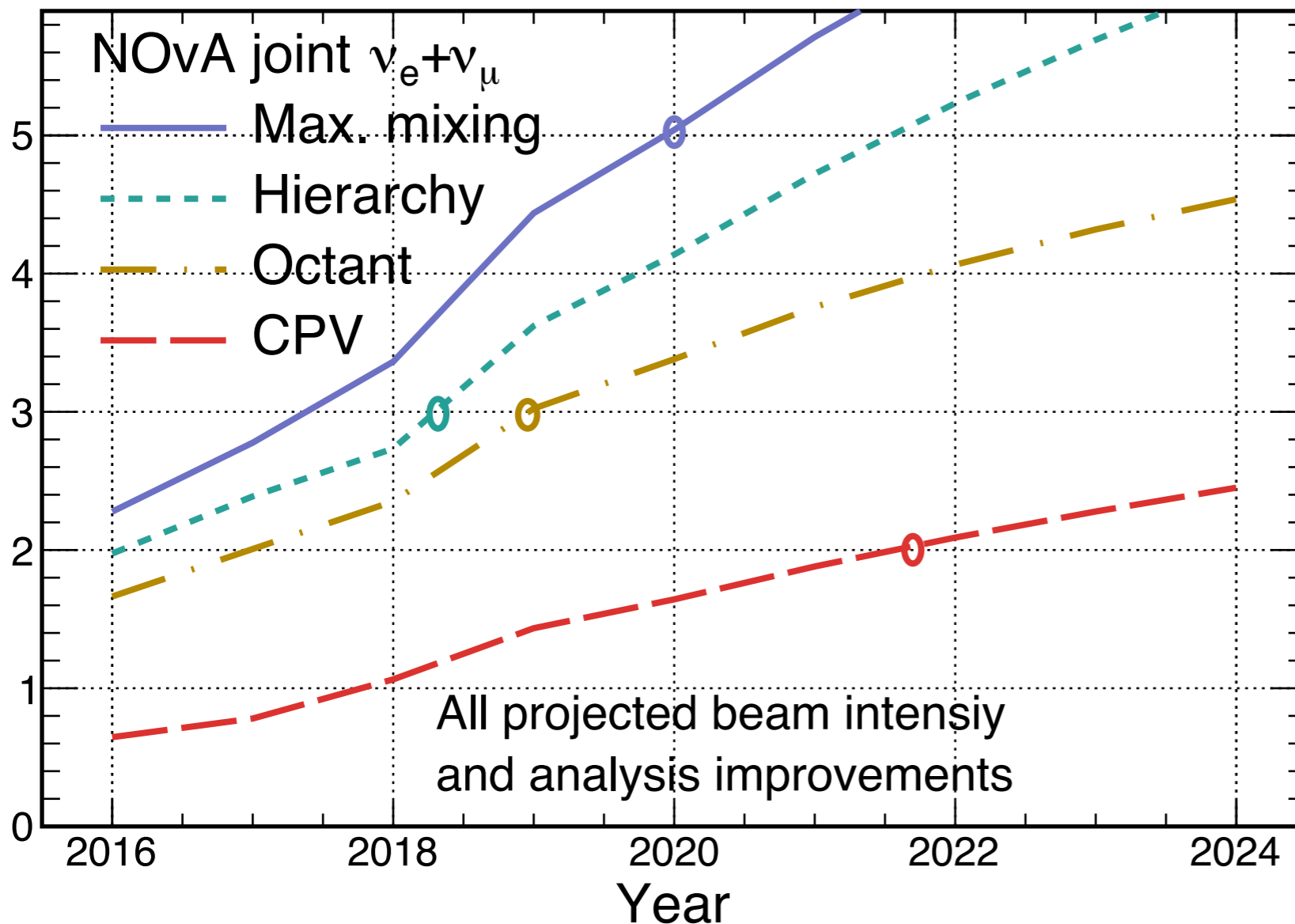


# Looking Forward

- Switched to anti-neutrino running in February 2017
- Run 50% neutrino, 50% anti-neutrino after 2018

Normal  $\delta_{CP}=3\pi/2$ ,  $\sin^2\theta_{23}=0.625$   
 $\Delta m_{32}^2=2.5\times 10^{-3} eV^2$ ,  $\sin^2\theta_{13}=0.022$

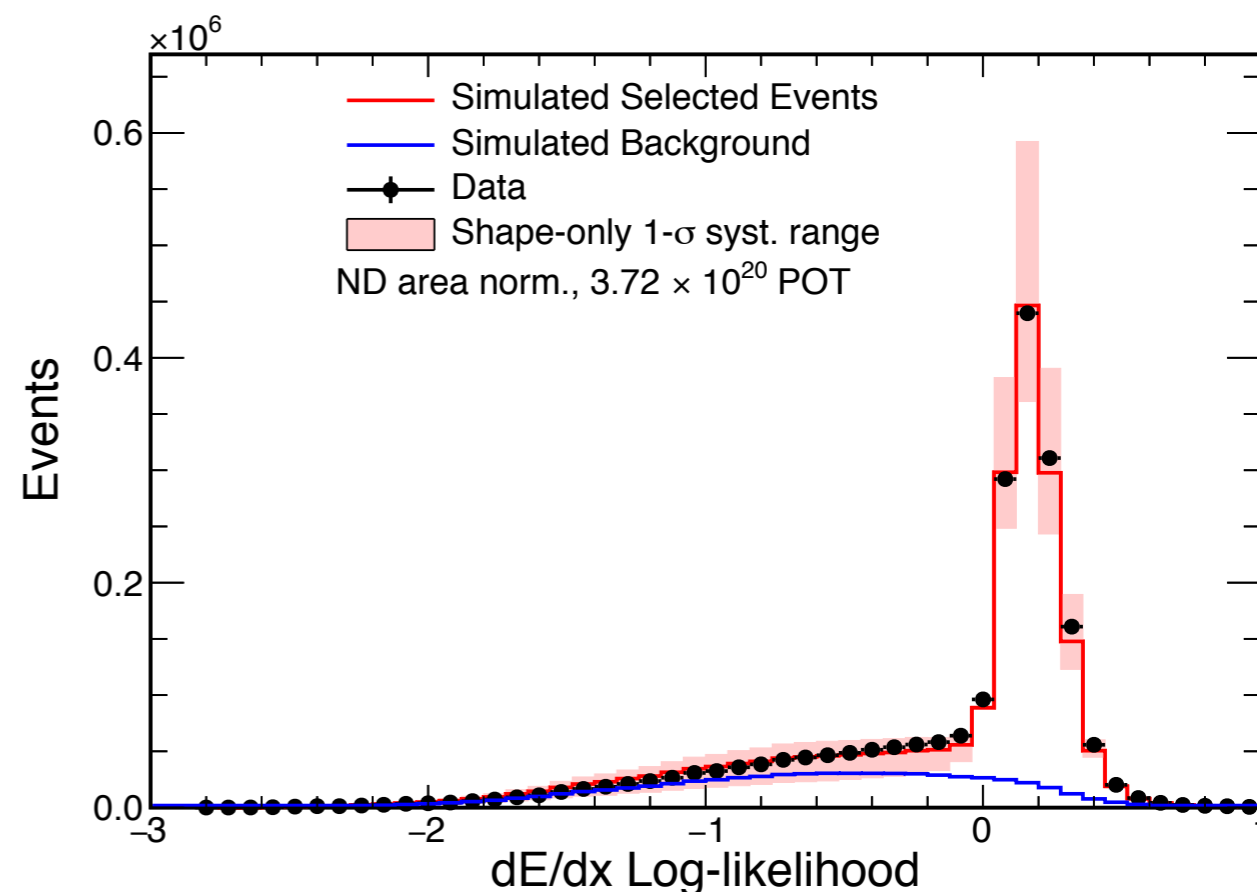
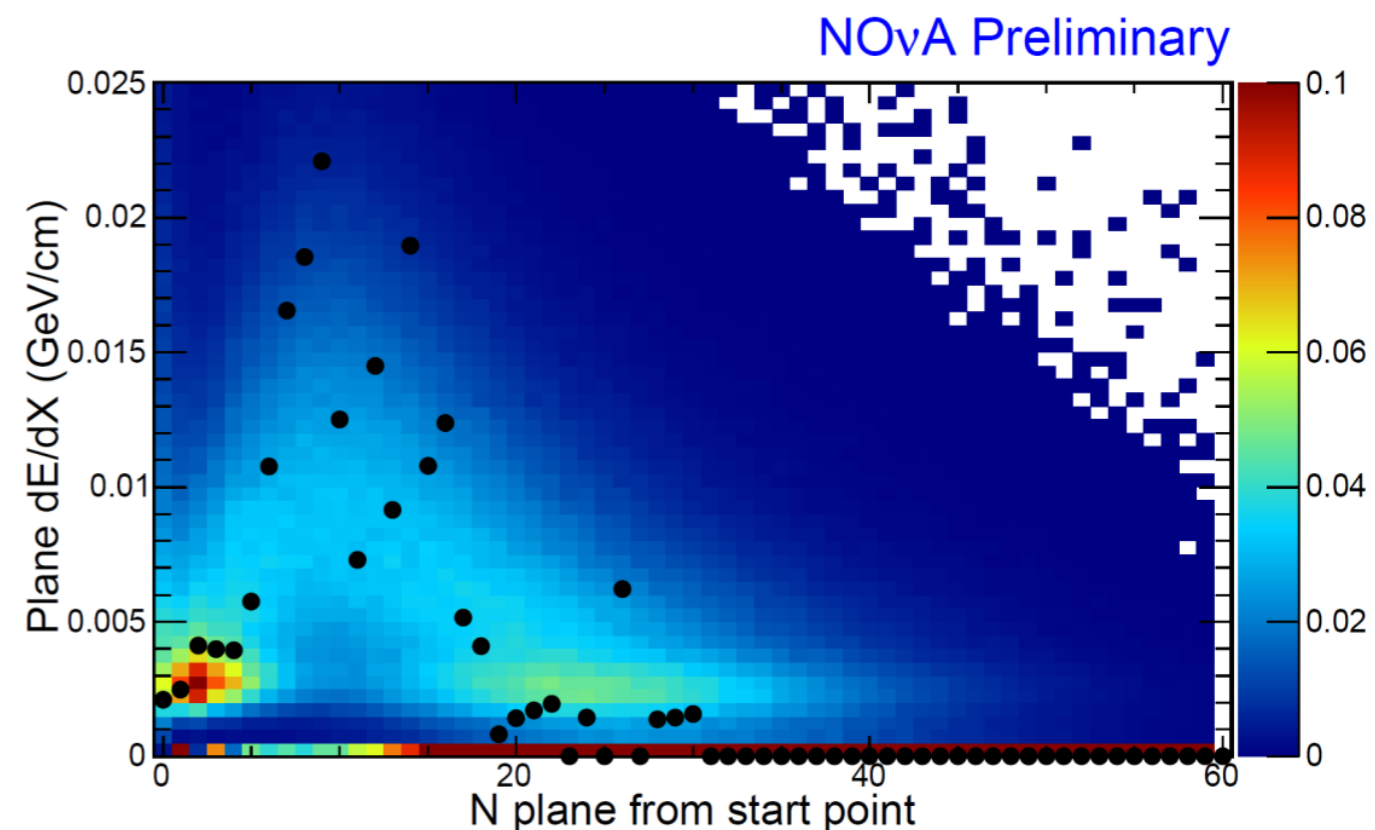
## NOvA Simulation



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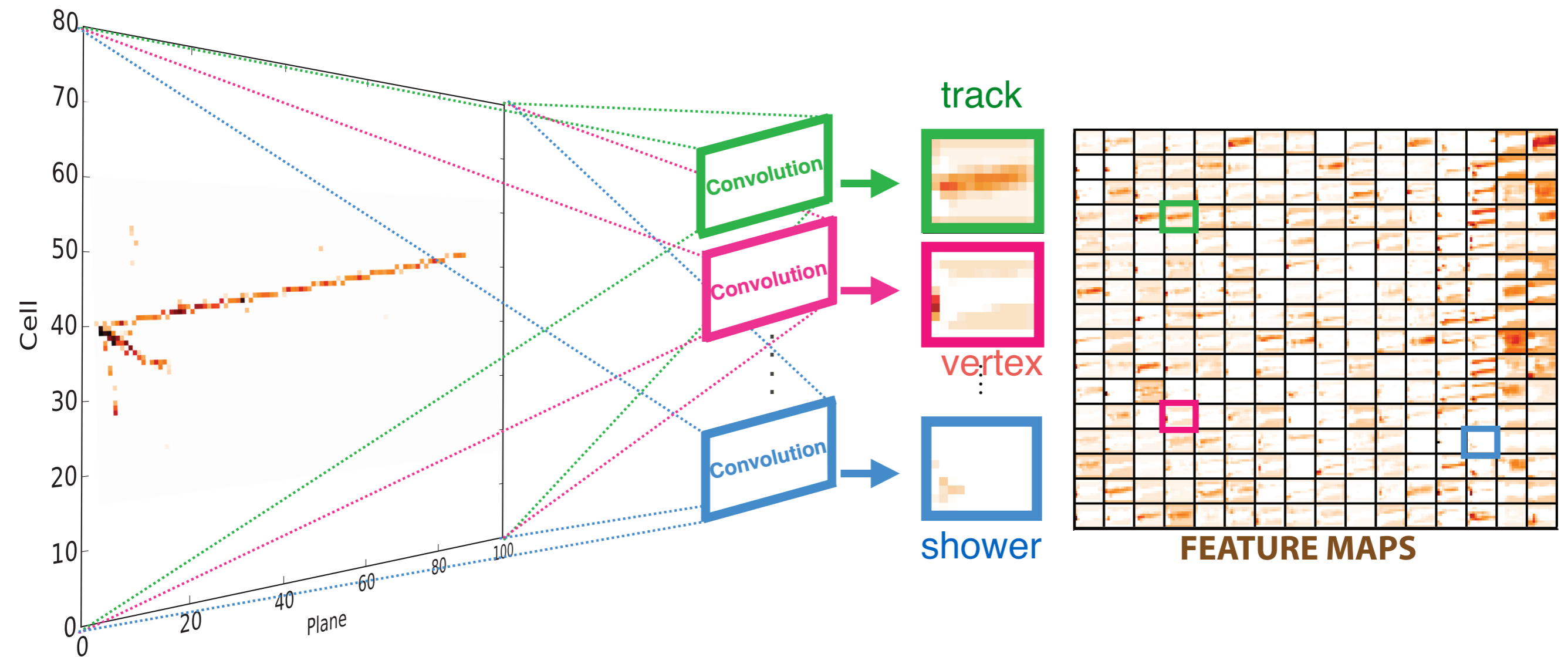
# Event Classification before Deep Learning

- Likelihood Identifier (LID)
  - Compare longitudinal and transverse  $dE/dx$  in leading shower to templates for different particle hypotheses
  - Build neural net from these inputs and reconstructed quantities.
  - Identifies electron neutrinos
- ReMID
  - Build a KNN classifier from four reconstructed quantities related to muons (length,  $dE/dx$ , scattering)
  - Identifies muon neutrinos



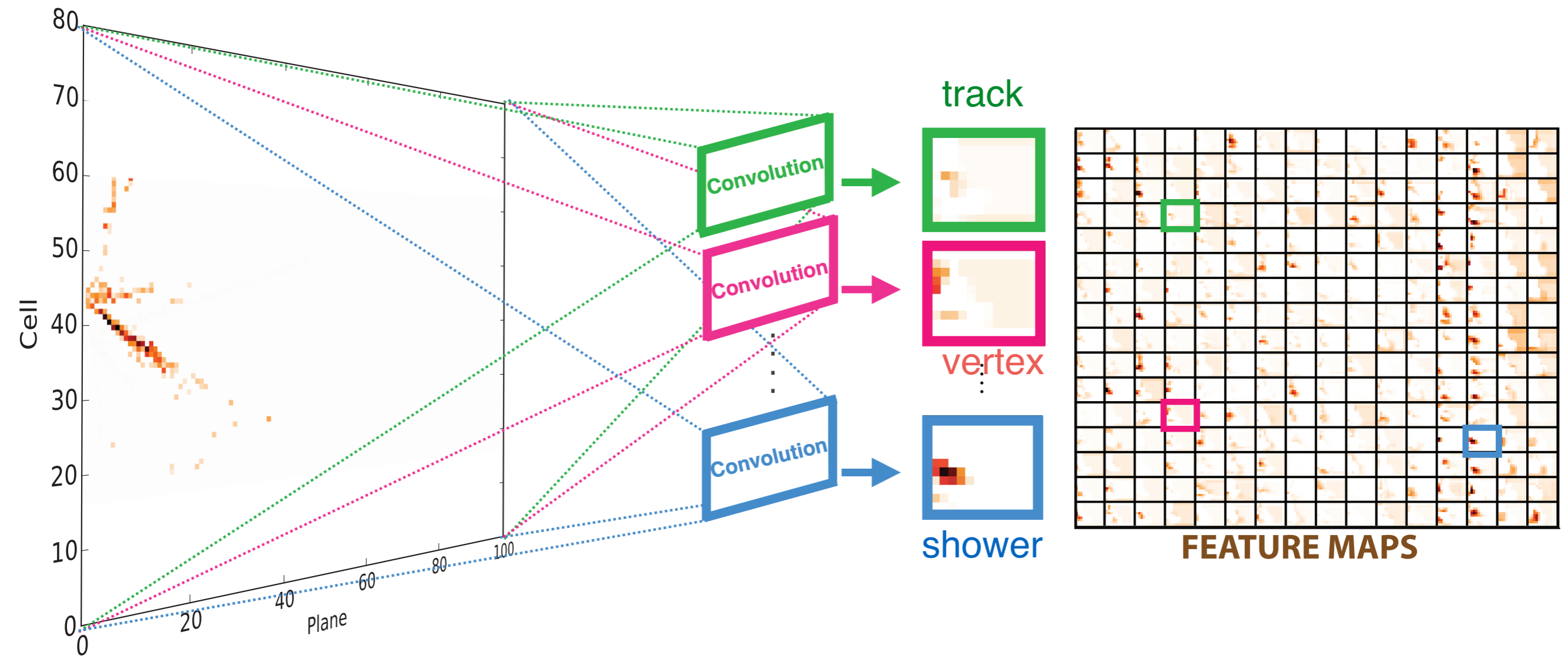
# Convolutional Visual Network (CVN) Selection

- Showing a muon neutrino interaction and the first layer of feature maps extracted from the convolutional kernels



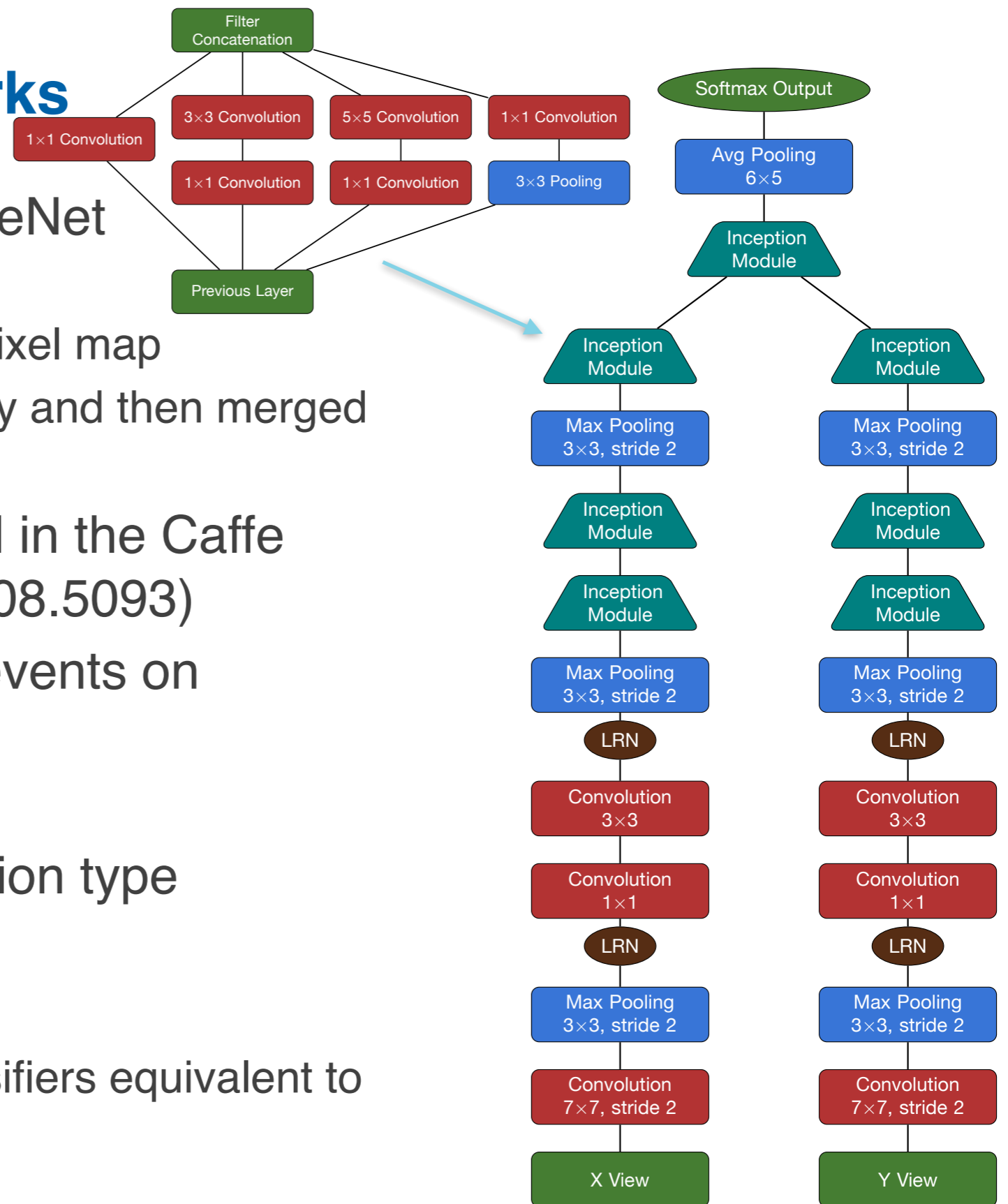
# Convolutional Visual Network (CVN) Selection

- Showing a electron neutrino interaction and the first layer of feature maps extracted from the convolutional kernels
- The strong features extracted are the shower as opposed to the muon track



# Convolutional Neural Networks

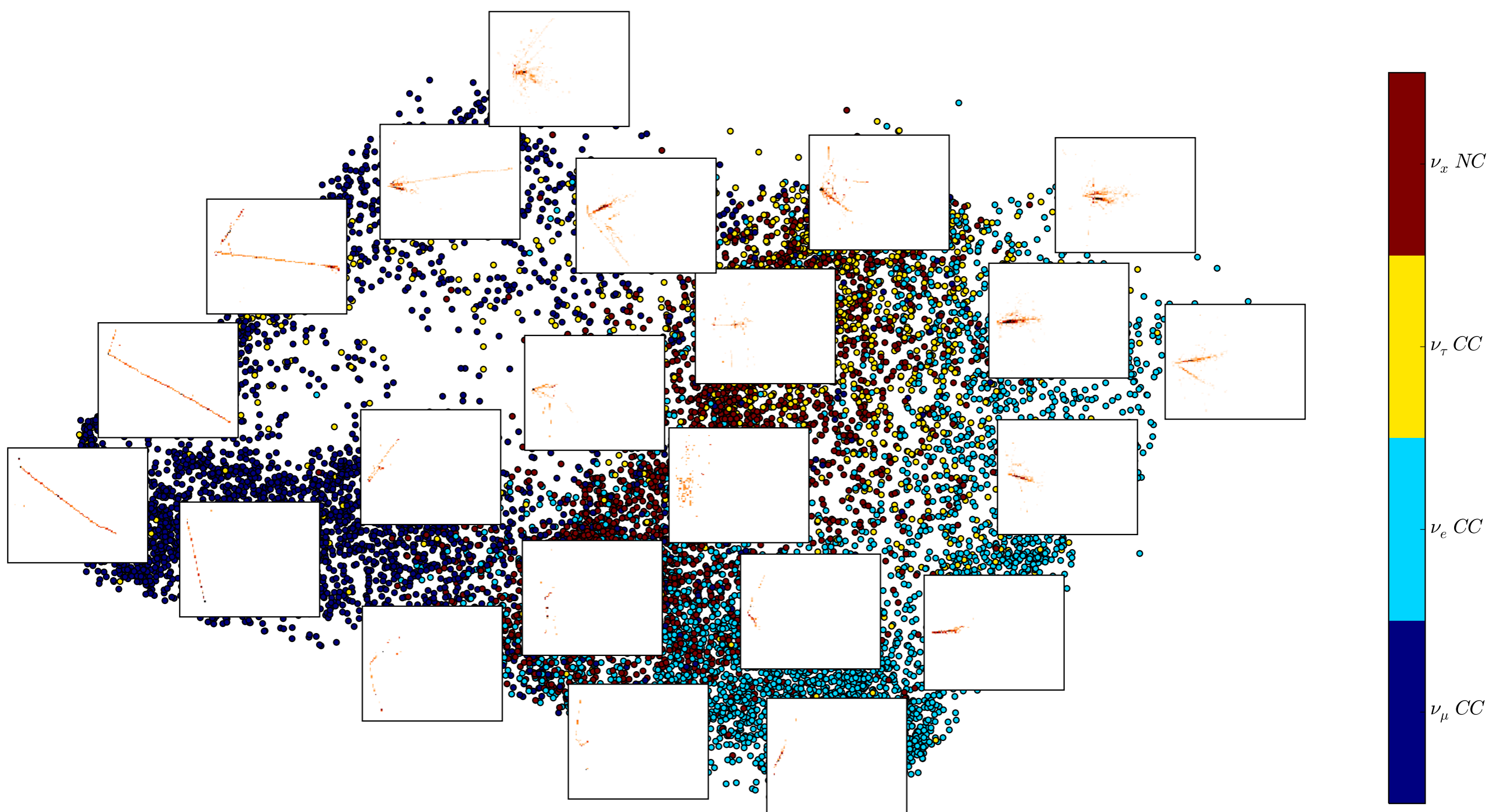
- Architecture adapted from GoogLeNet
  - C. Szegedy et al., arXiv:1409.4842
  - Input is 80 cell x 200 plane detector pixel map
  - Each event view processed separately and then merged
- Network implemented and trained in the Caffe Framework (Y. Jia et al., arXiv:1408.5093)
- Trained on 4.7 million simulated events on Fermilab GPU cluster
- Output classifies neutrino interaction type ( $\nu_\mu, \nu_\tau, \nu_e, NC$ )
- Used in appearance analysis.
  - Performance gain over previous classifiers equivalent to adding 30% more detector mass



A. Aurisano and A. Radovic and D. Rocco et. al, JINST 11 P09001 (2016)



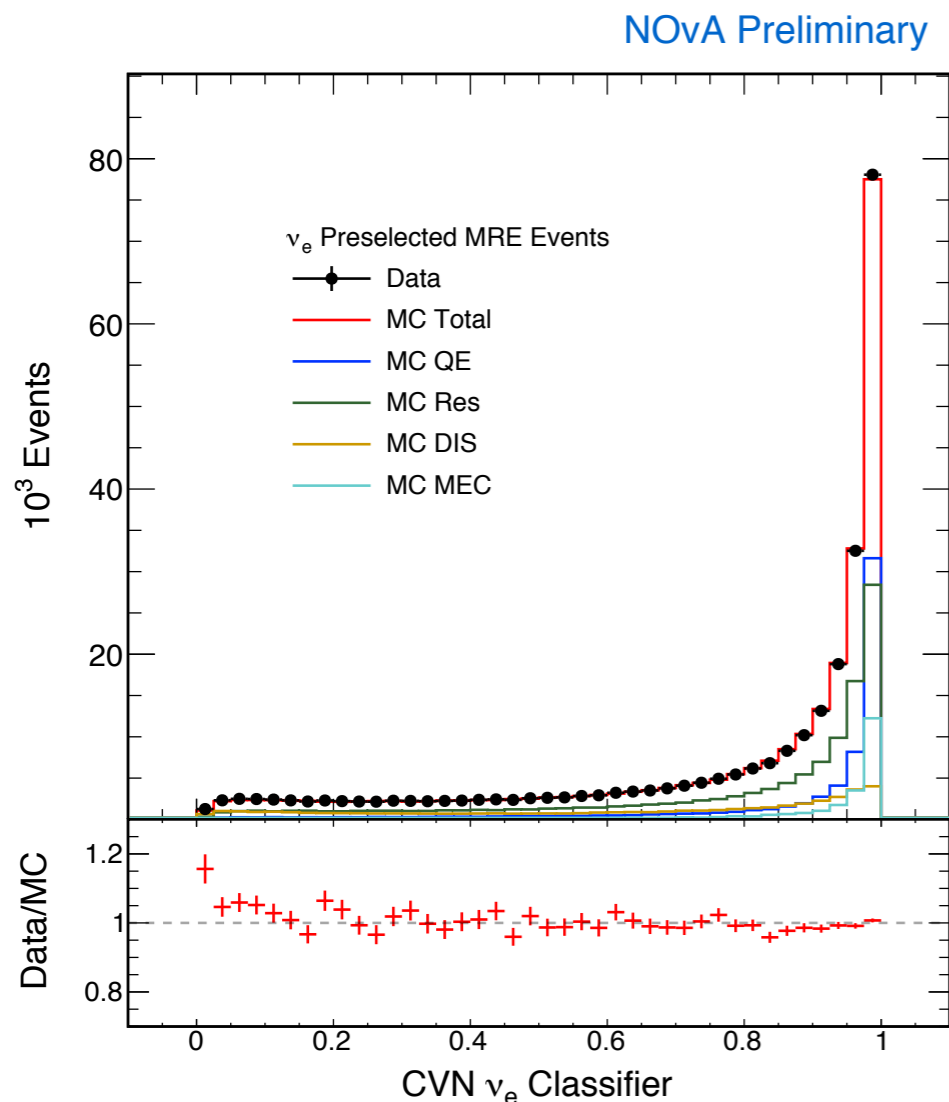
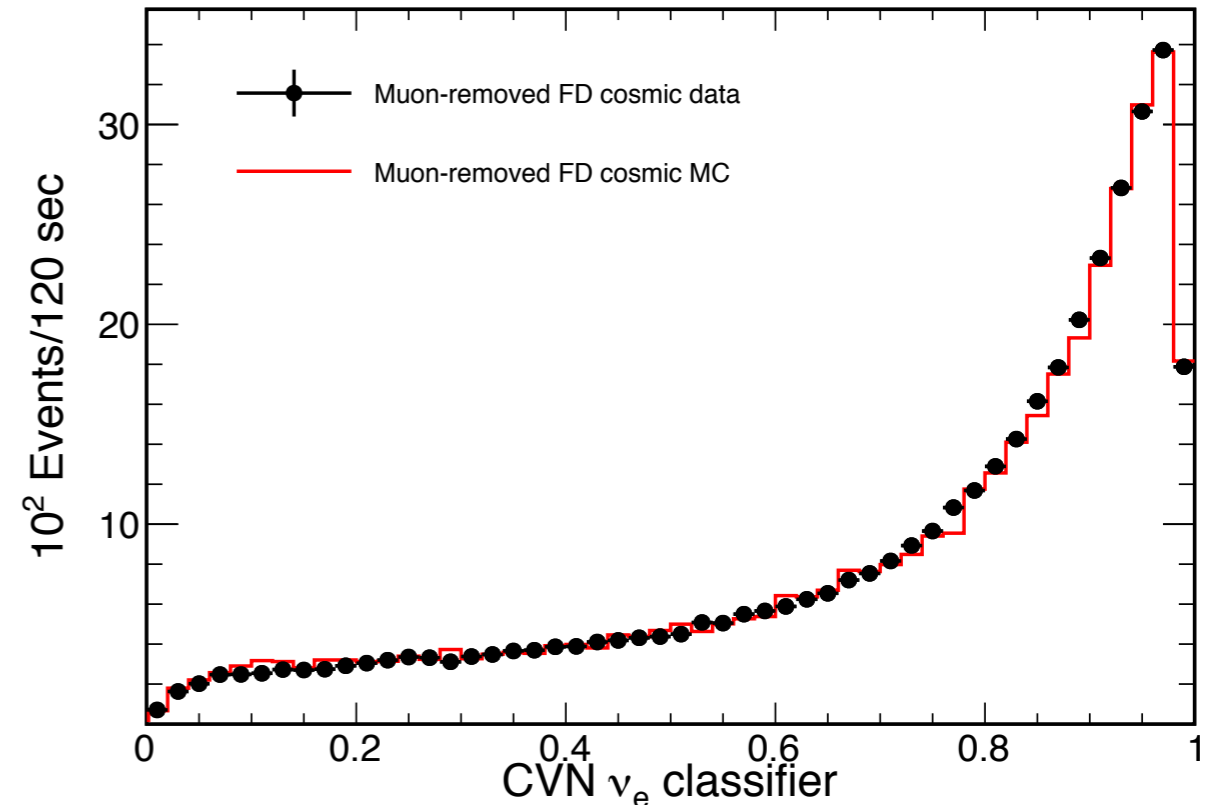
t-SNE representation of CVN classification. Truth labels shown for the training sample.



t-SNE representation of CVN classification. Truth labels shown for the training sample.

# Evaluating Signal Efficiency

- Remove cosmic ray muon from FD events in data and simulation
- Apply selection to remaining bremsstrahlung shower to benchmark simulation of electron selection



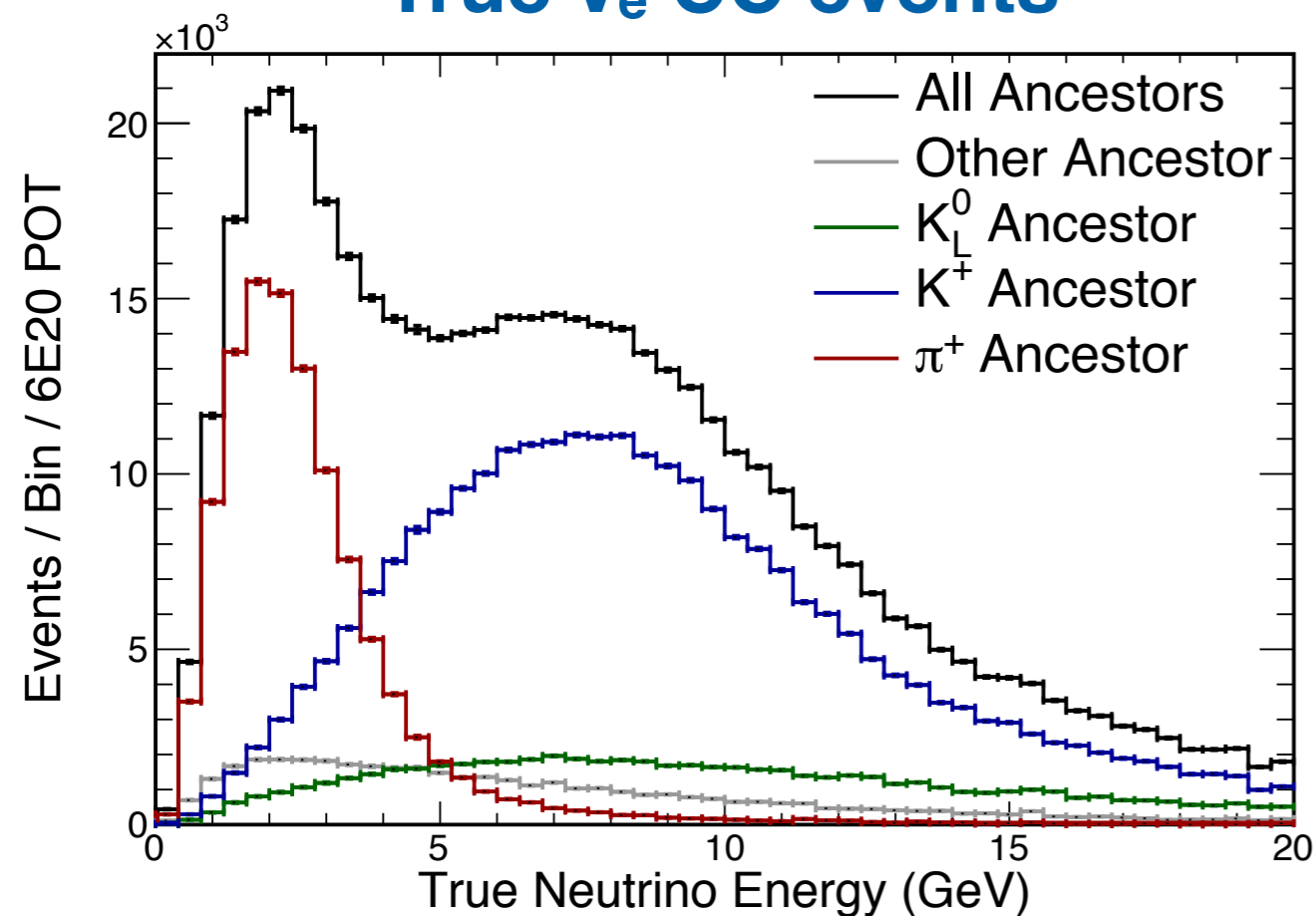
- EM showers should be well modeled, check if selection efficiency differences come from hadronic side
- Remove reconstructed muons from selected  $\nu_\mu$  events, replace with simulated electron (MRE)
- better than 1% agreement between efficiency for selecting data MRE events and efficiency for selecting MC MRE events



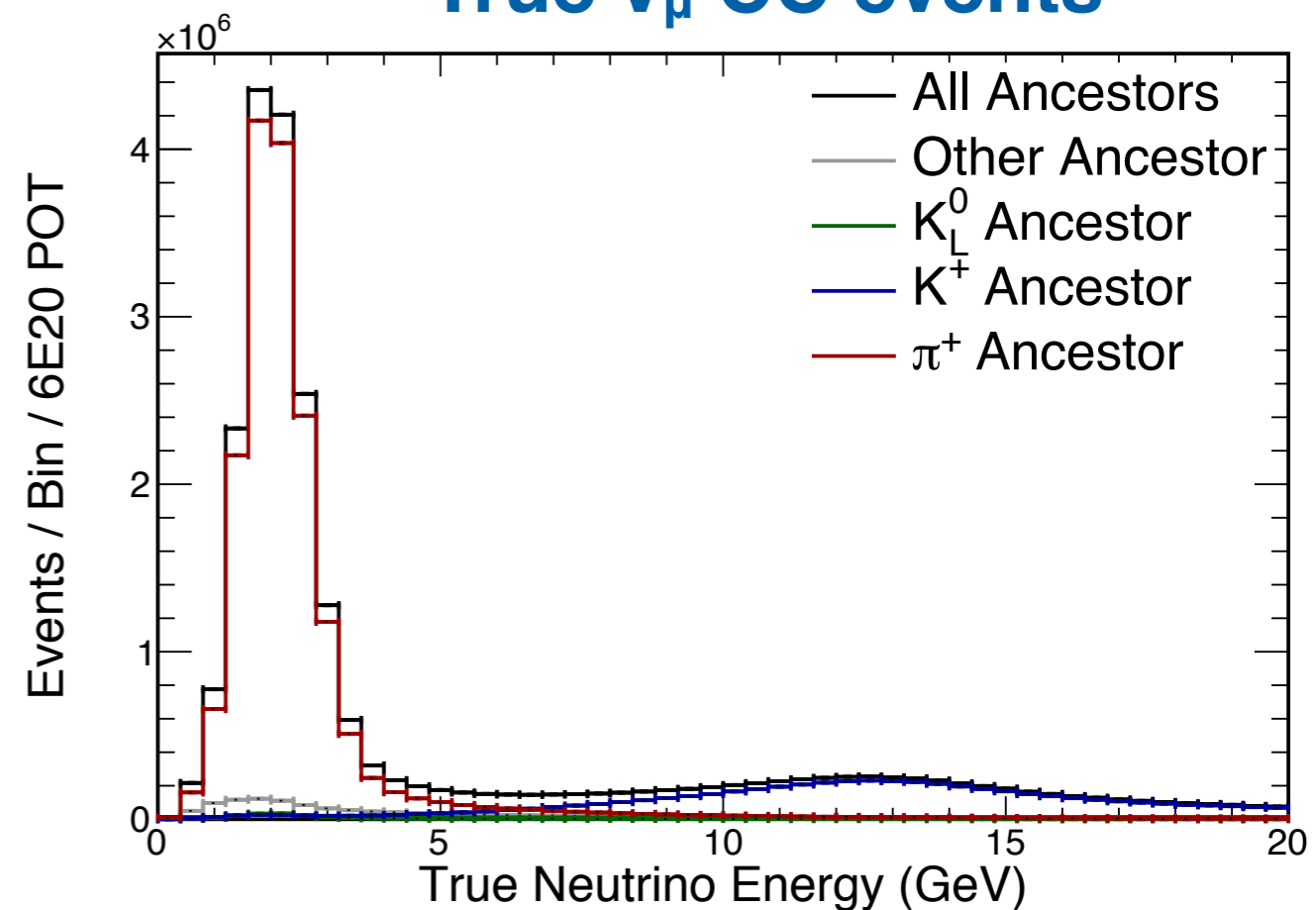
# ND Data Decomposition: Beam $\nu_e$ CC

- Low energy  $\nu_\mu$  and  $\nu_e$  trace back to the same  $\pi$  ancestors
- Use  $\nu_\mu$  at lower energy to reweight decaying pions in  $(p_T, p_z)$  space
- Decreases  $\nu_e$  with  $\pi^+$  parent 3-4%
- Weight  $\nu_e$  with  $K^+$  parents up 17% based on  $\nu_\mu$  high-E tail
- Overall effect is 1% increase in 1-3 GeV range in intrinsic beam  $\nu_e$  CC events

## True $\nu_e$ CC events

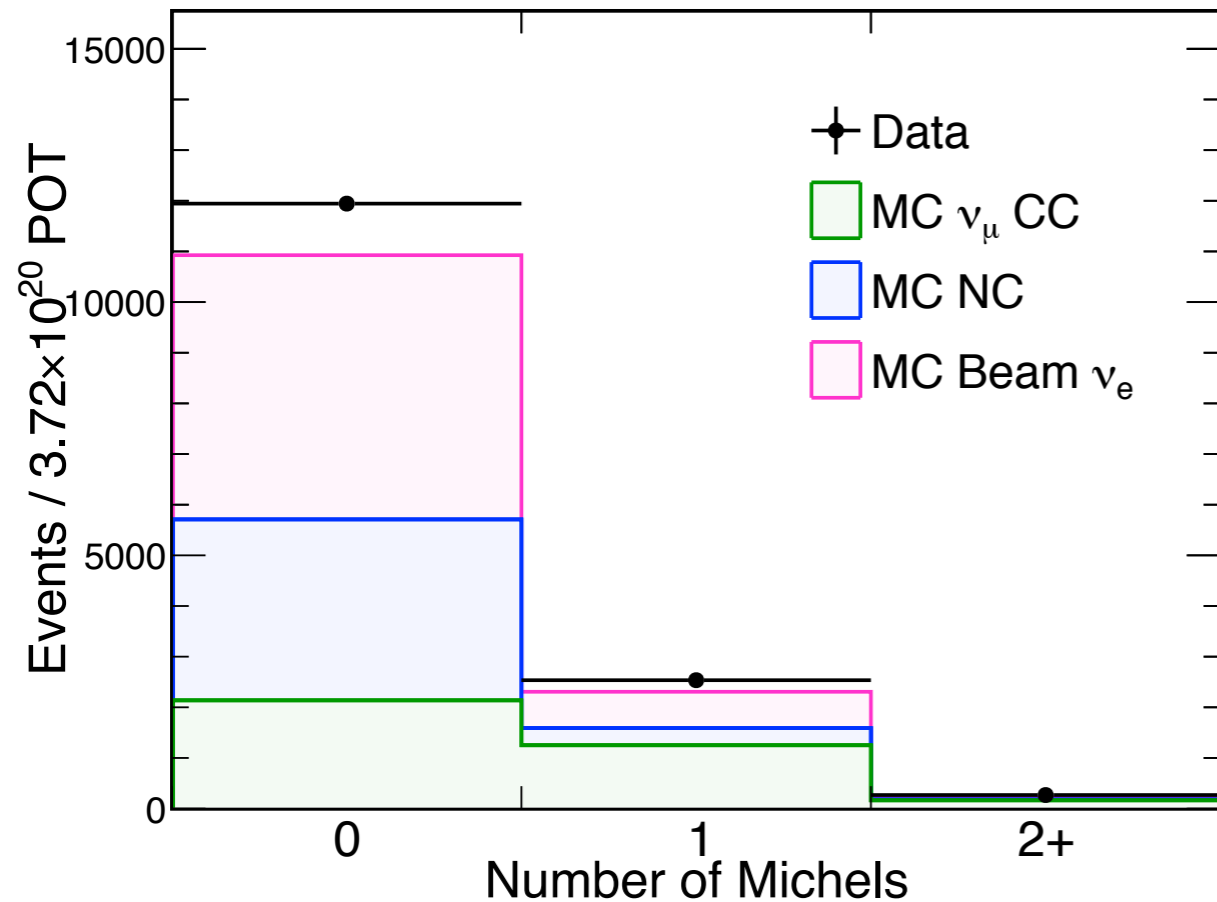


## True $\nu_\mu$ CC events

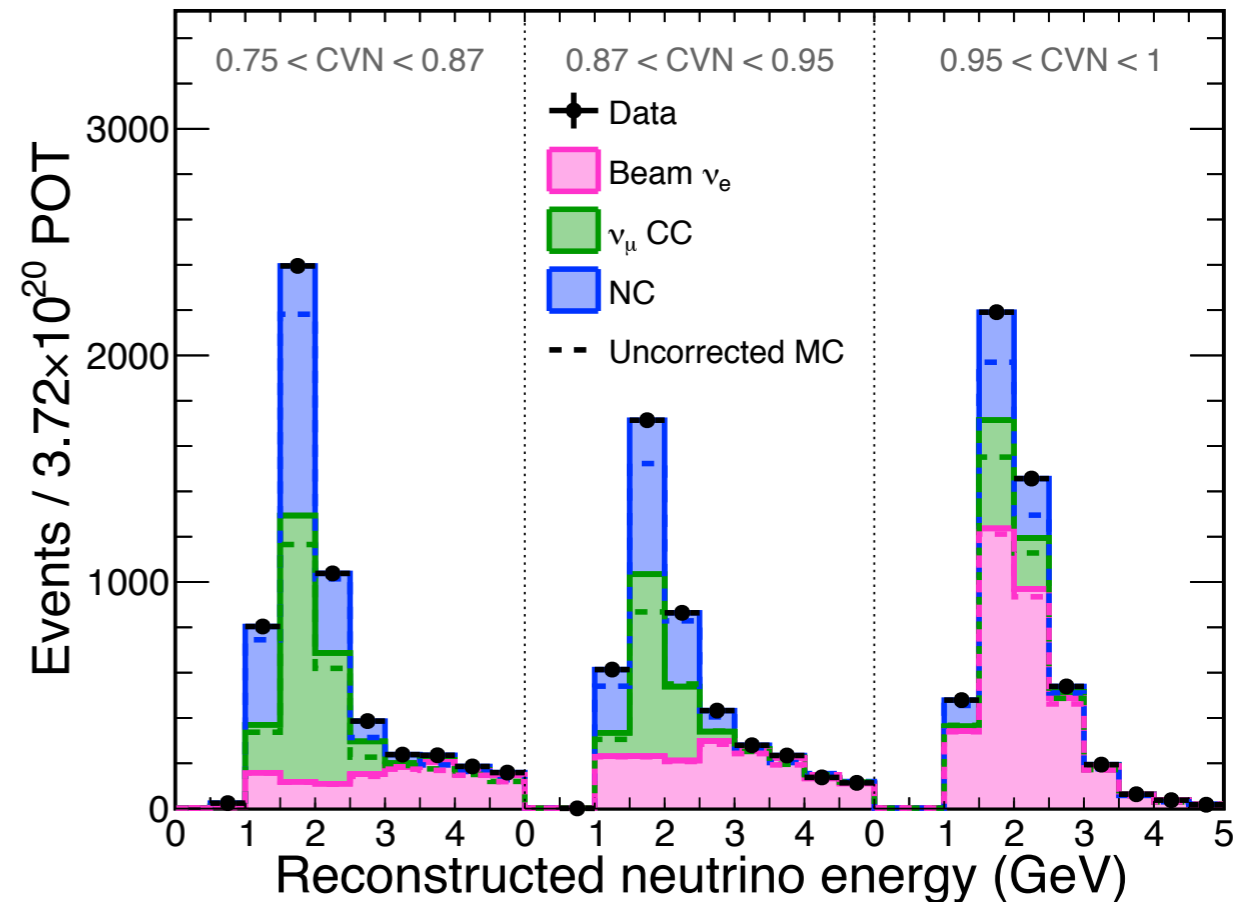


# ND Data Decomposition: Michel Electrons

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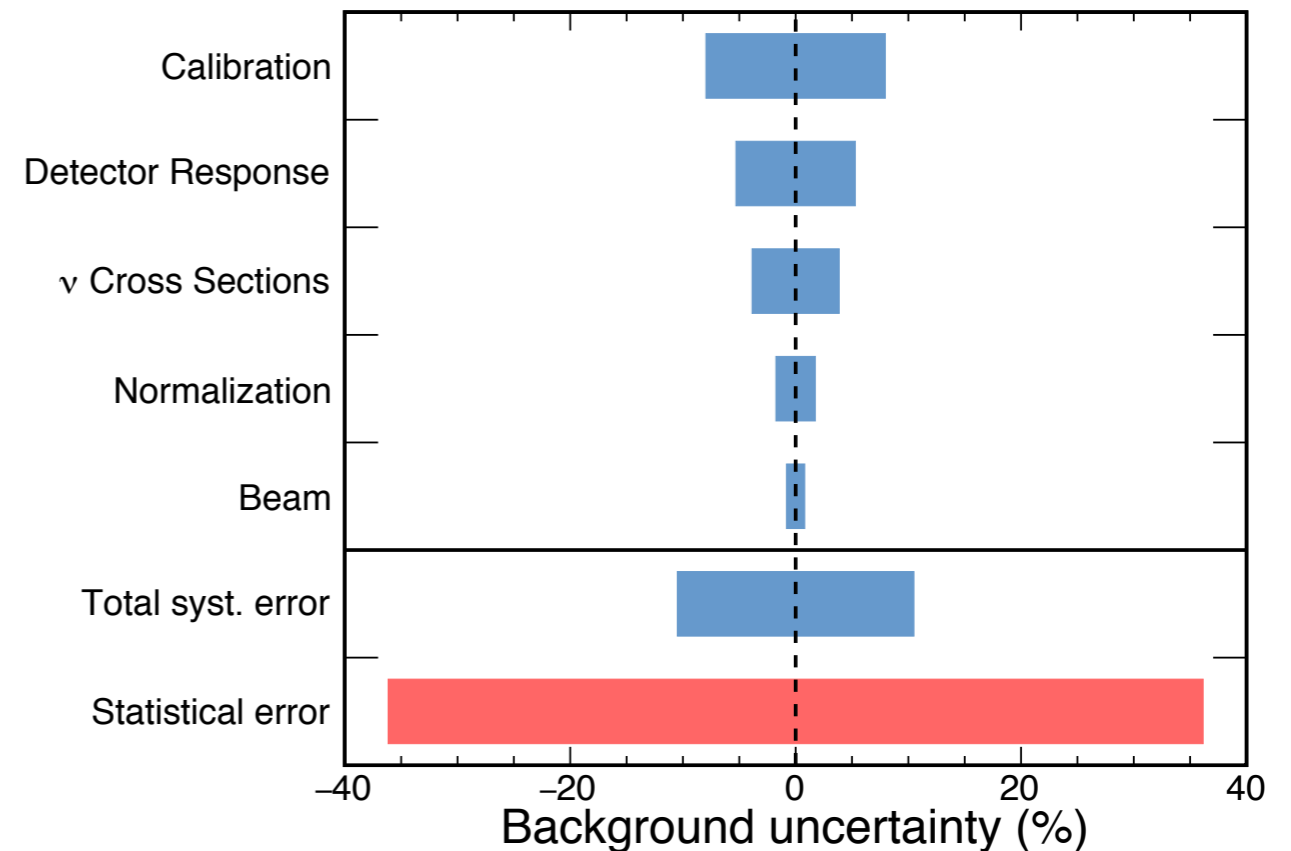
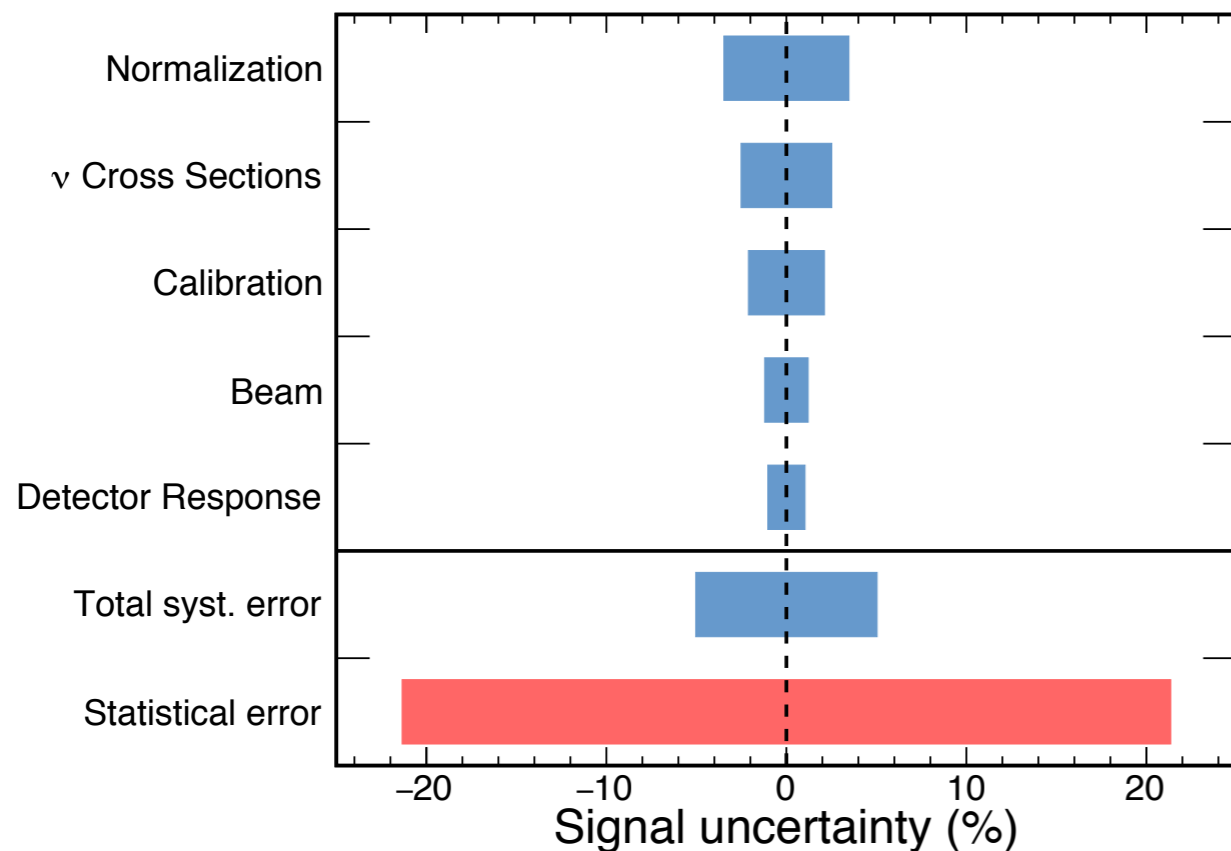


NOvA Preliminary



- $\nu_\mu$  CC events contain Michel electron from muon decay
- $\sim 1$  more Michel in  $\nu_\mu$  events than  $\nu_e$  or NC
- Fit observed number of Michels in each bin of energy and PID by adjusting  $\nu_\mu$ /NC ratio
- Data excess assigned between NC (+17%) and  $\nu_\mu$  CC (+10%)

# Systematic Uncertainties



- Multiple sources of systematic error considered
- Extrapolate FD predictions with special MC samples for each effect.
- Uncertainty quoted as difference between shifted and nominal predictions
- Fit nuisance parameters as pull terms
- Statistical uncertainties dominate

# $\nu_\mu \rightarrow \nu_e$ Oscillation Prediction

- Prediction dependent on oscillation parameters

Signal events

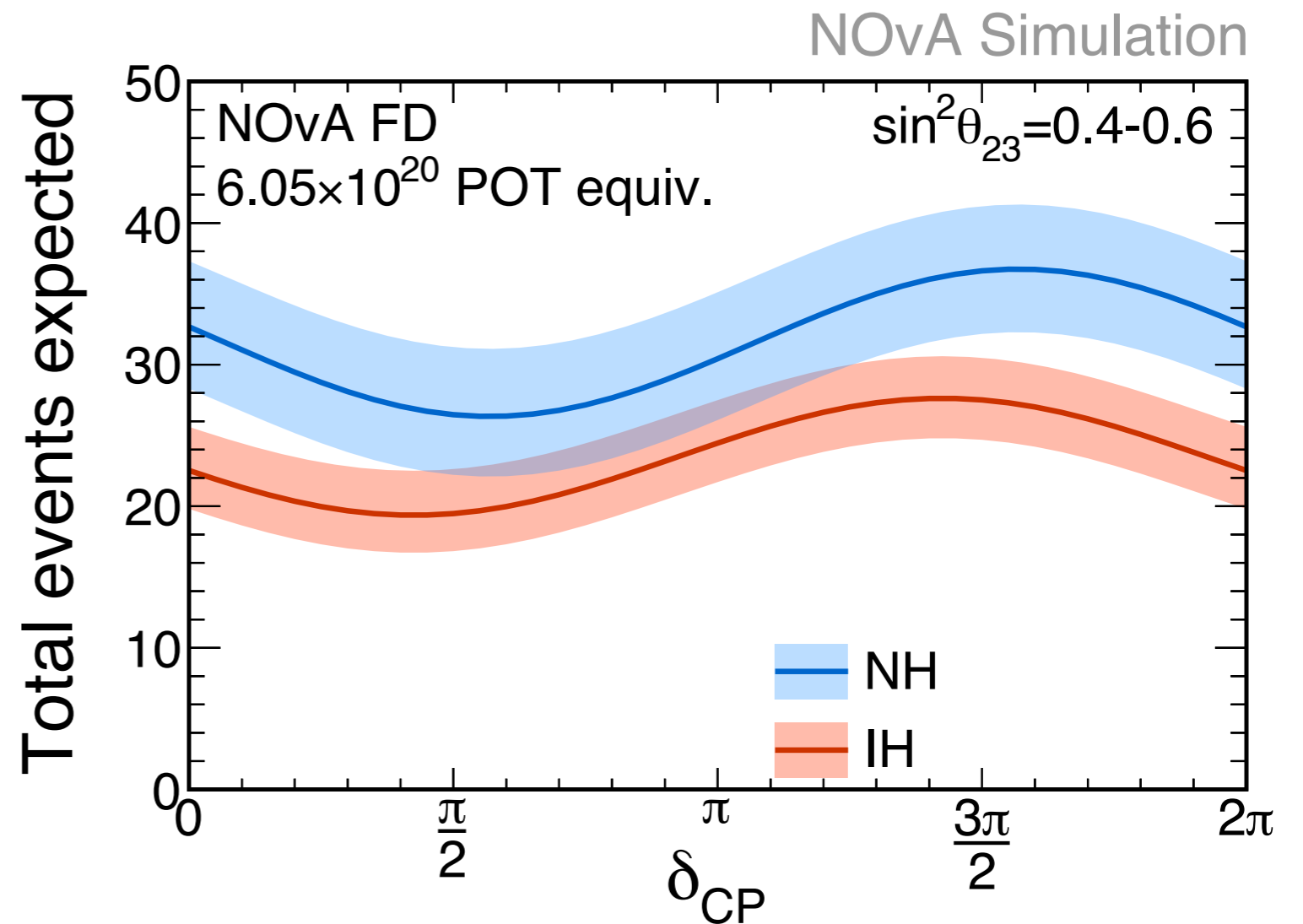
( $\pm 5\%$  systematic uncertainty):

NH, $3\pi/2,$	IH, $\pi/2,$
28.2	11.2

Background by component

( $\pm 10\%$  systematic uncertainty):

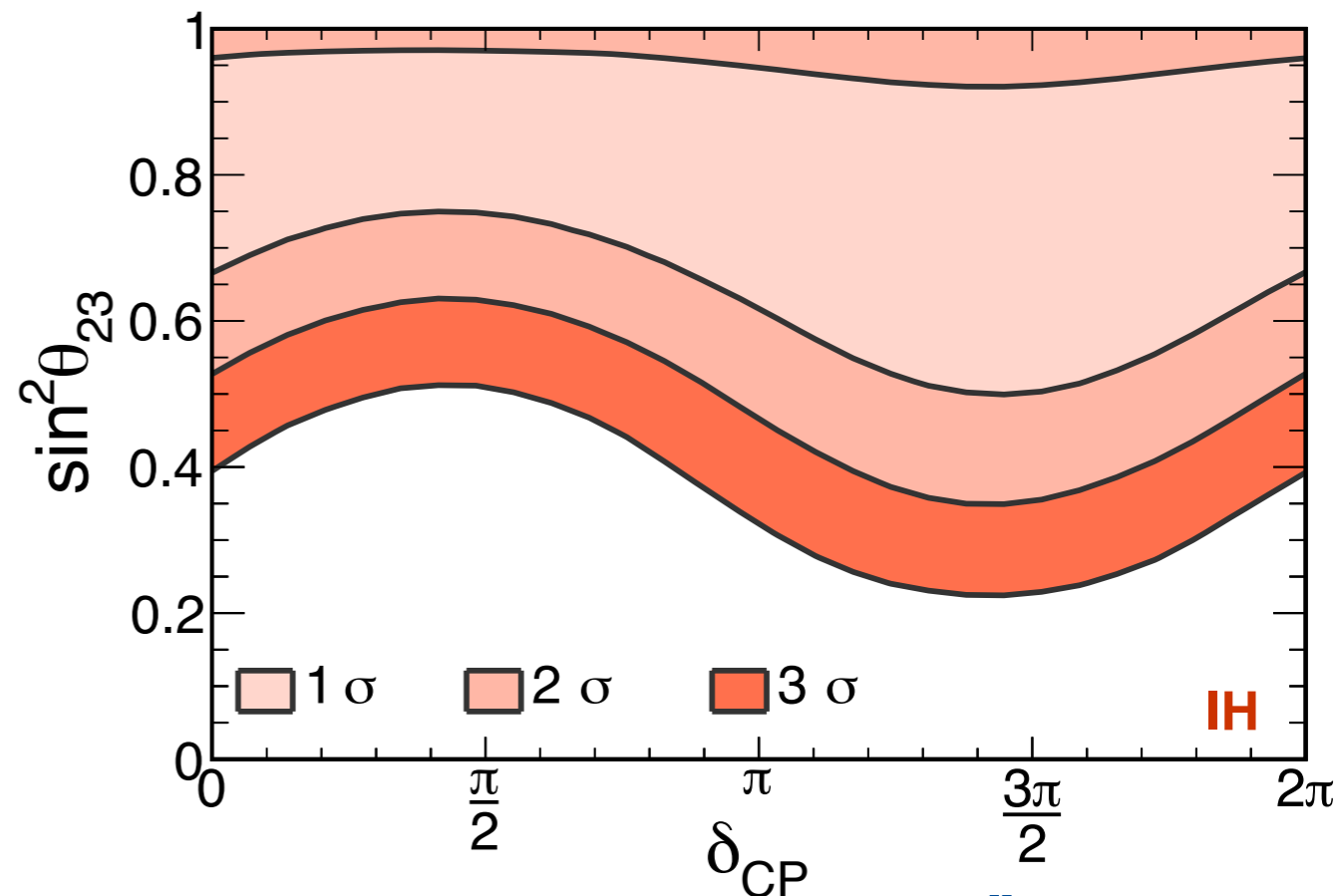
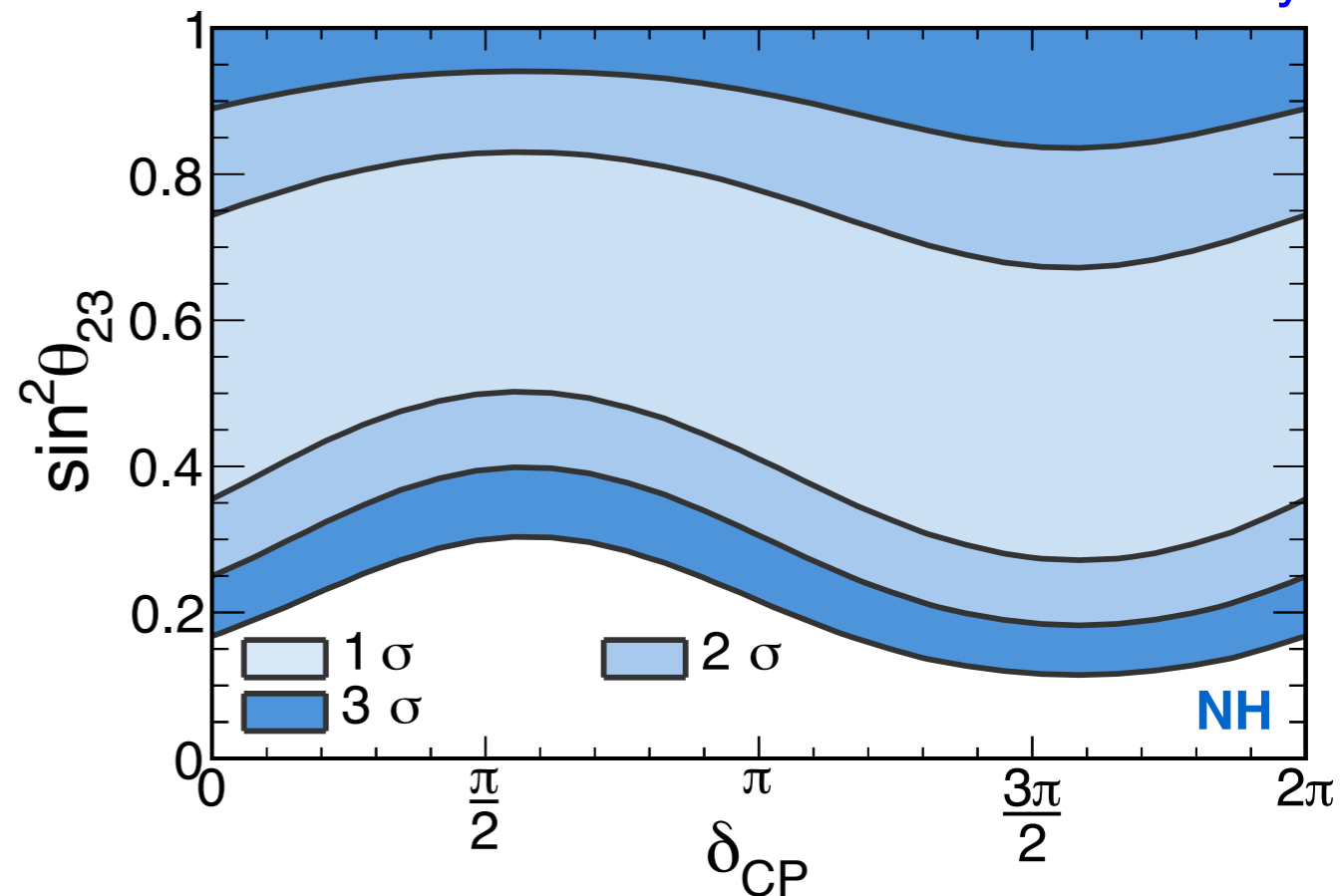
Total BG	NC	Beam $\nu_e$	$\nu_\mu$ CC	$\nu_\tau$ CC	Cosmics
8.2	3.7	3.1	0.7	0.1	0.5



## $\nu_\mu \rightarrow \nu_e$ Oscillation Results

- Fit for hierarchy,  $\delta_{CP}$ ,  $\sin^2\theta_{23}$ 
  - Constrain  $\sin^2(2\theta_{13})=0.085\pm 0.05$
  - Constrain  $\Delta m^2=2.44\pm 0.06\times 10^{-3}$  eV<sup>2</sup>, NH
  - $(-2.49\pm 0.06\times 10^{-3}$  eV<sup>2</sup>, IH)

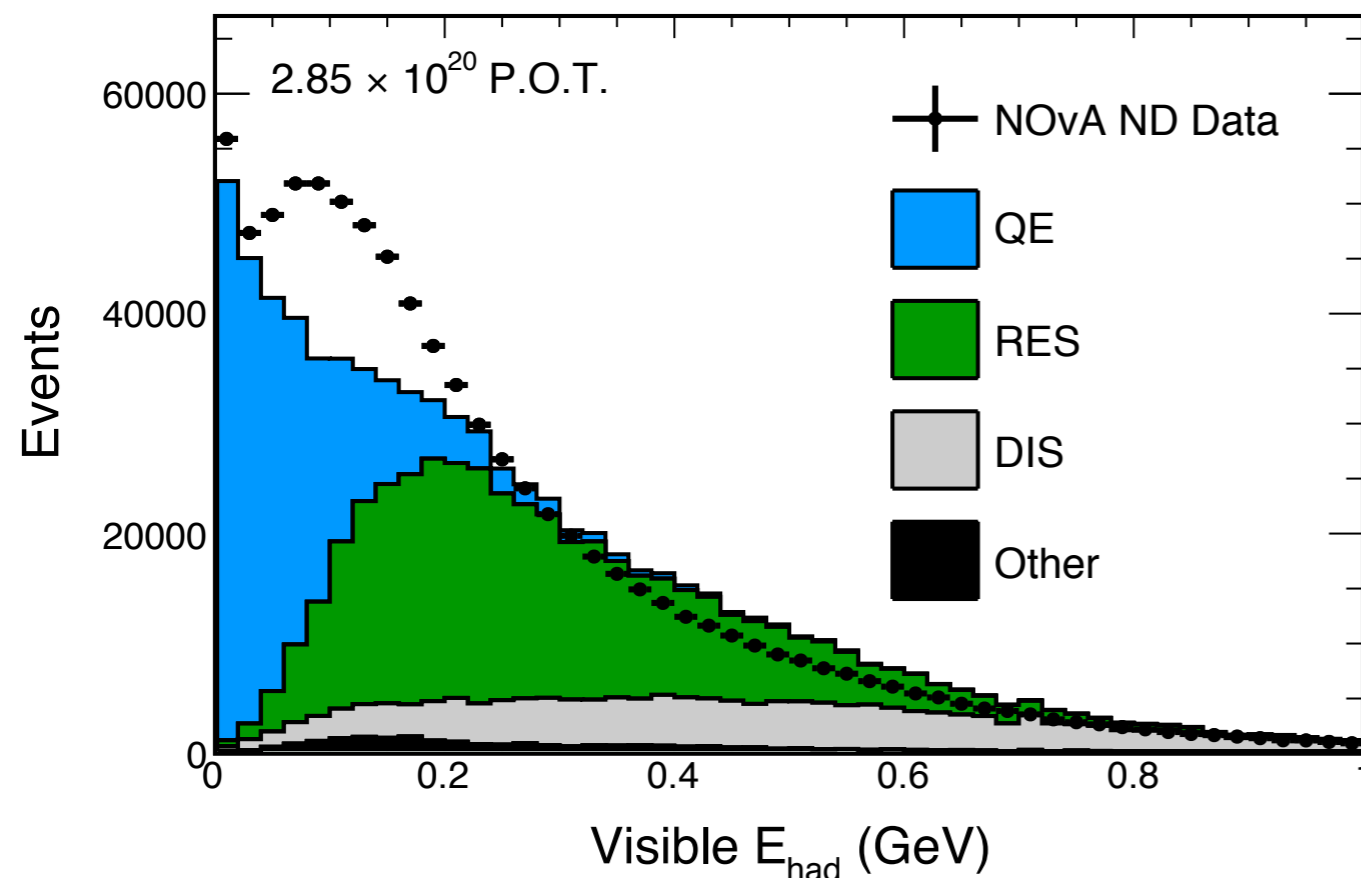
- Systematic effects included as nuisance parameters (normalization, flux, calibration, cross section, and detector response effects)



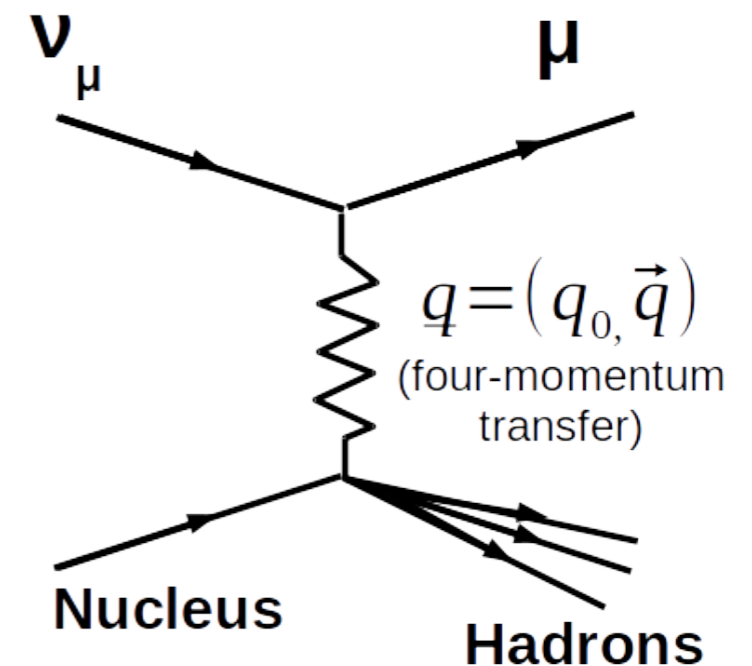
# Nuclear Model Corrections

Near Detector hadronic energy distribution suggests unsimulated process between quasi-elastic and delta production

NOvA Preliminary



Similar conclusions from MINERvA data reported in P.A. Rodrigues et al., PRL 116 (2016) 071802

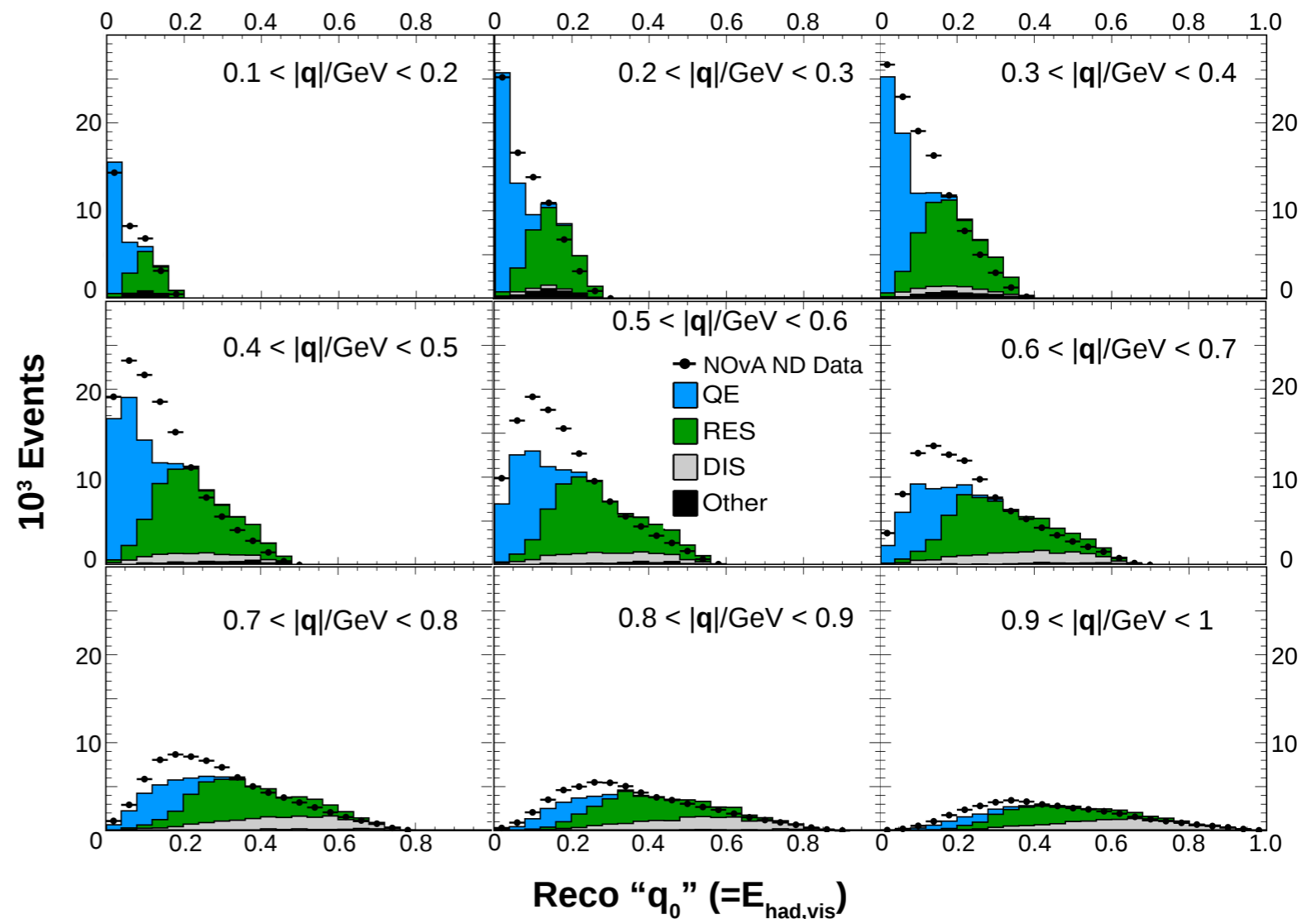
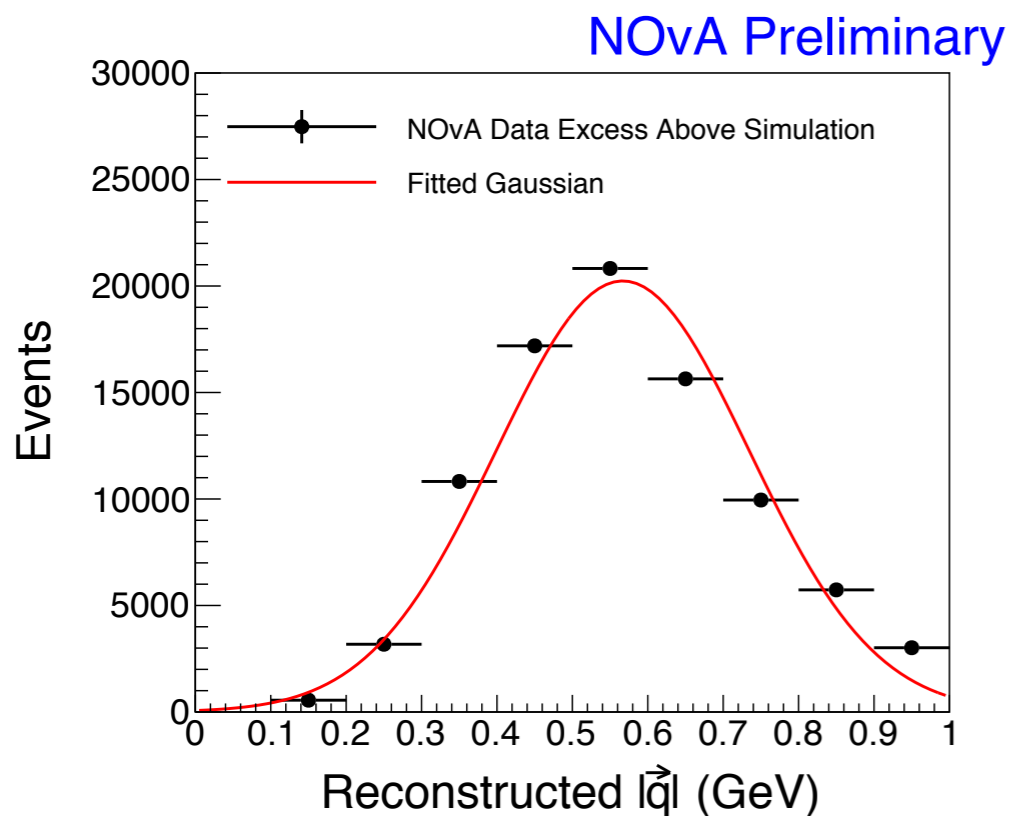


Solution: 2-particle, 2-hole (2p2h) events where neutrino is scattering off a nucleon-nucleon pair

# Nuclear Model Corrections

- Enable GENIE's empirical Meson Exchange Current model<sup>1</sup>
- Reweight to matched observed excess as a function of momentum transfer
- Weight single non-resonant pion production down by effectively 50%<sup>2</sup>

NOvA Preliminary



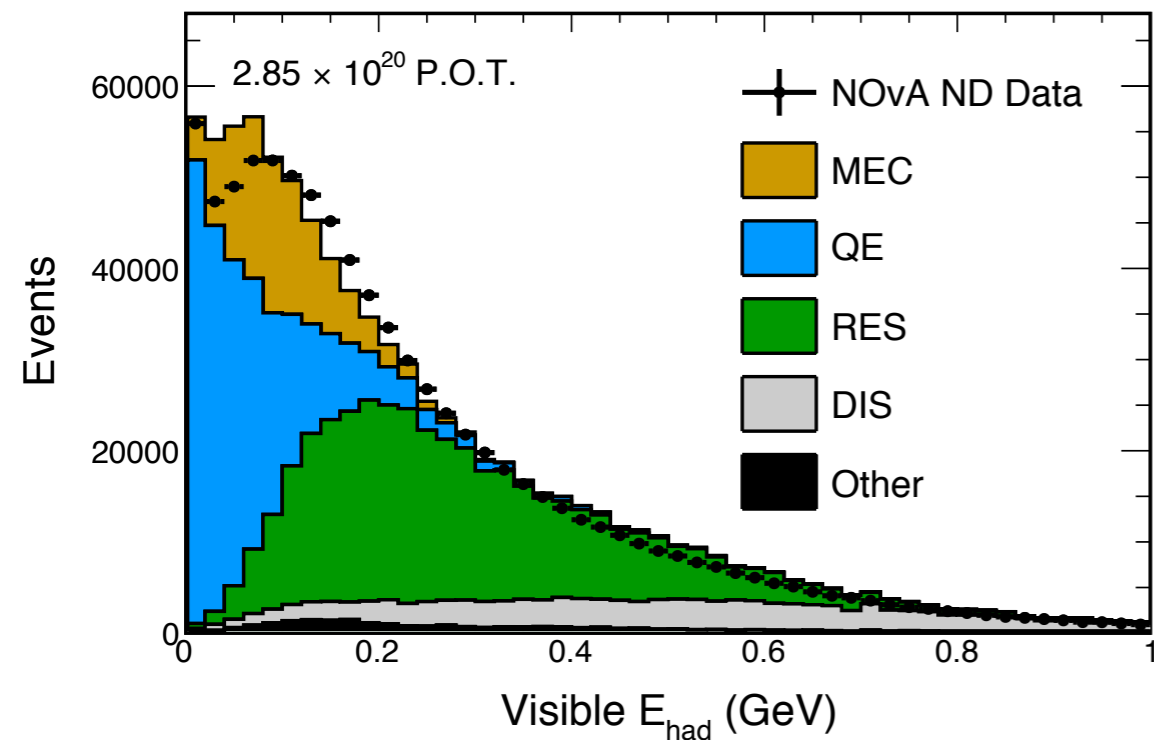
<sup>1</sup>S. Dytman, based on J. W. Lightbody, J. S. OConnell, *Comp. in Phys.* 2 (1988) 57

<sup>2</sup>P.A. Rodrigues et al., arXiv:1601.01888

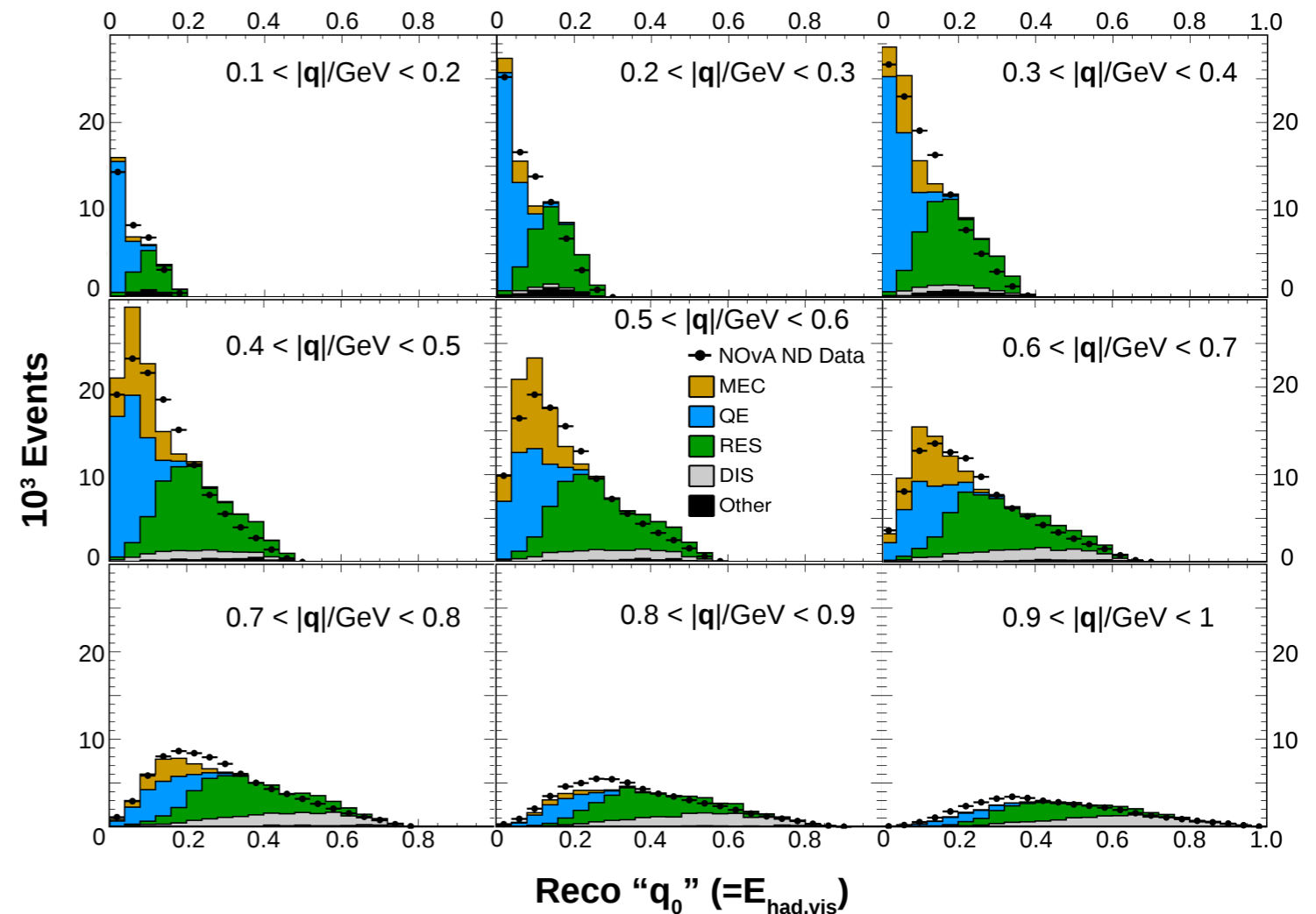
# Nuclear Model Corrections

- Take 50% systematic uncertainty on MEC component
- Reduces hadronic energy scale and quasi-elastic cross section systematic uncertainties

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NOvA Preliminary

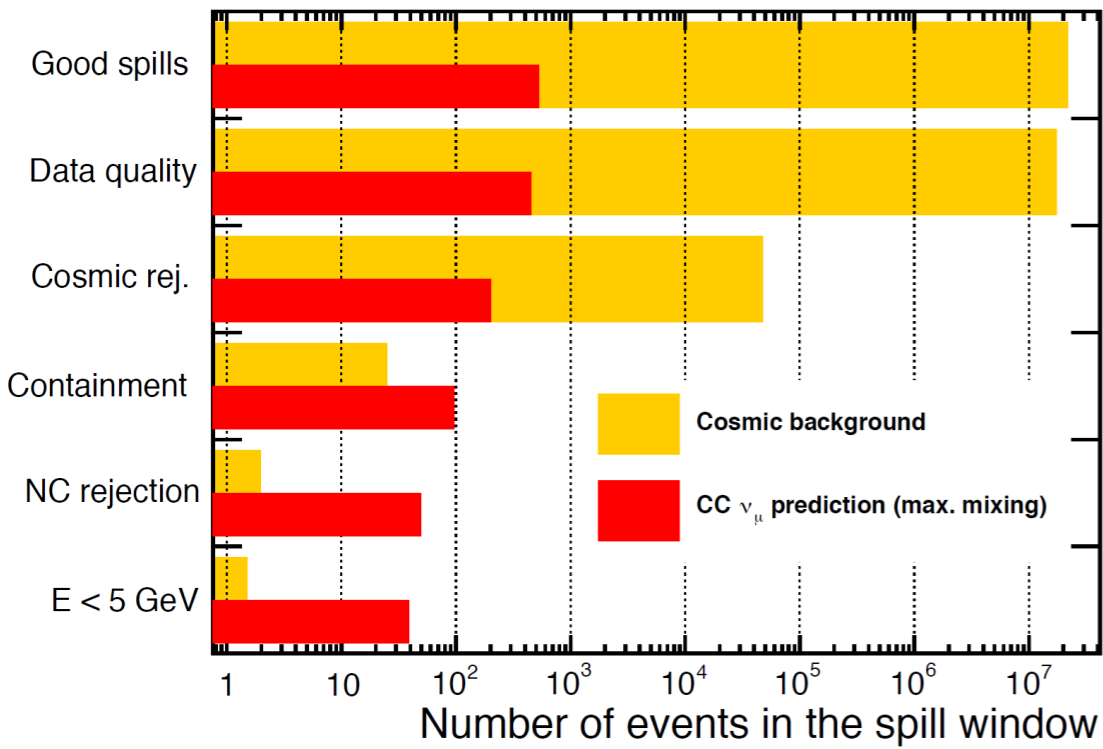


<sup>1</sup>S. Dytman, based on J. W. Lightbody, J. S. OConnell, *Comp. in Phys.* 2 (1988) 57

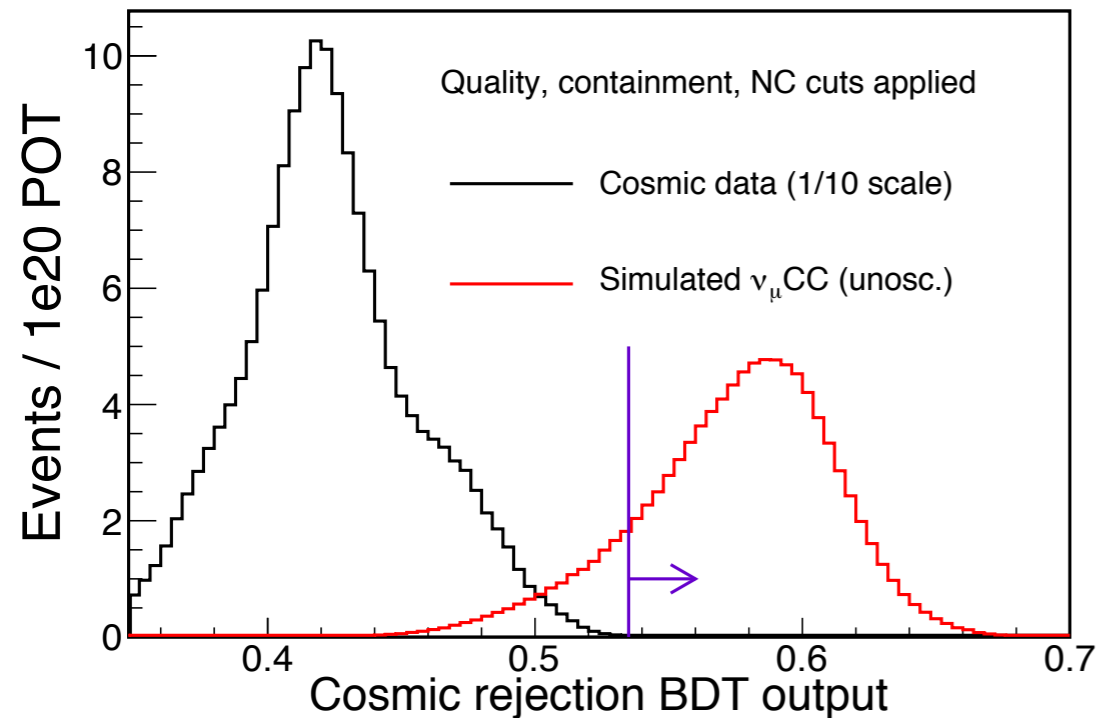
<sup>2</sup>P.A. Rodrigues et al., arXiv:1601.01888



# Cosmic Rejection



NOvA Preliminary



- Far Detector sees 150 kHz of cosmic induced events
- 10  $\mu$ s beam window at a rate of  $\sim 0.8$  Hz reduces background by  $10^5$
- Additional factor of  $10^7$  rejection achieved from event topology and a boosted decision tree (BDT) based on:
  - track direction
  - start/end points of track
  - track length
  - energy
  - number of hits
- Predict 2.7 cosmic background events

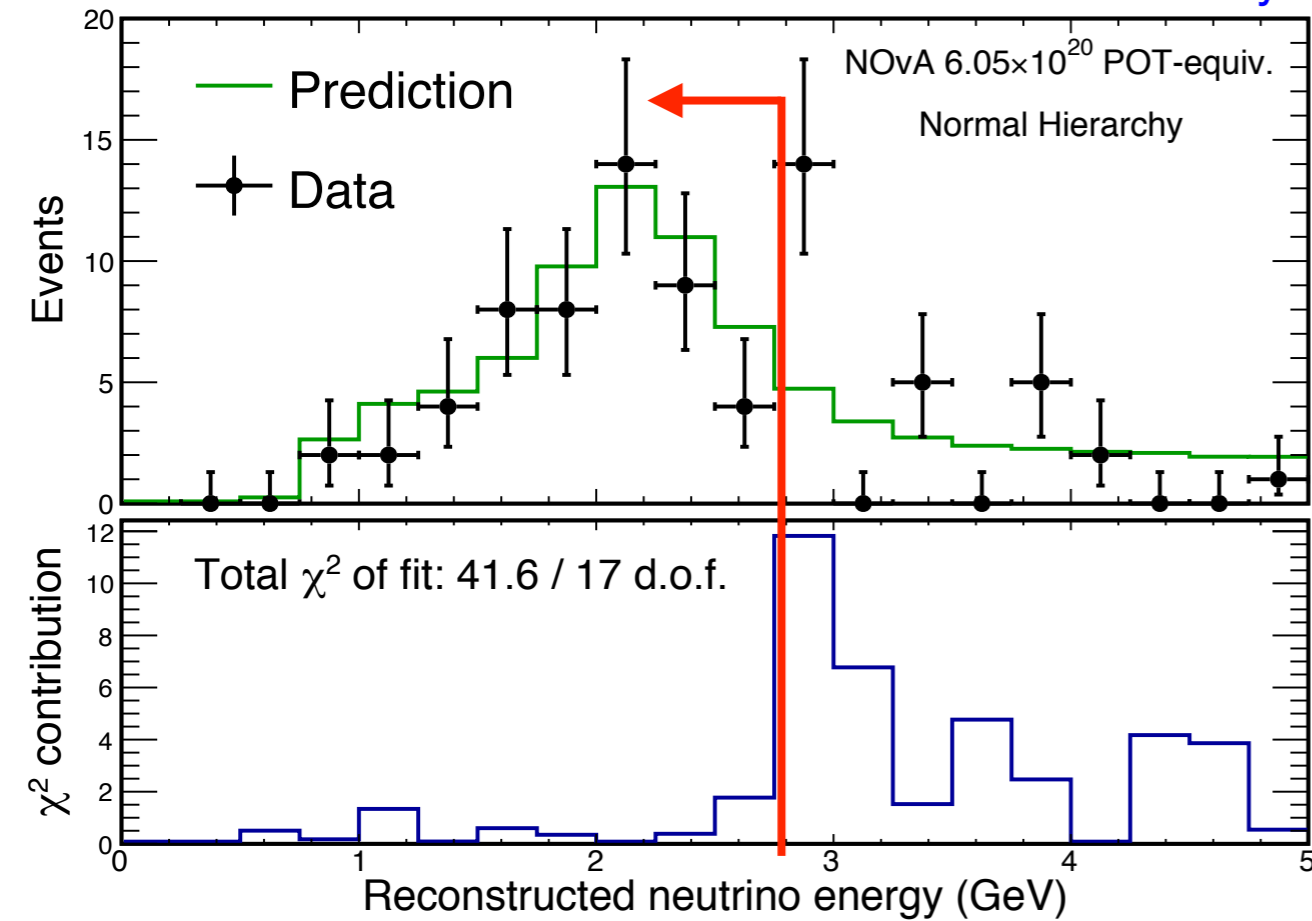
# Systematic Uncertainties

- Various sources of systematic uncertainty considered
- Propagate the effect of each through the extrapolation with specially modified MC samples
- Include as pull terms in fit
- Table shows increase in quadrature of measurement uncertainty

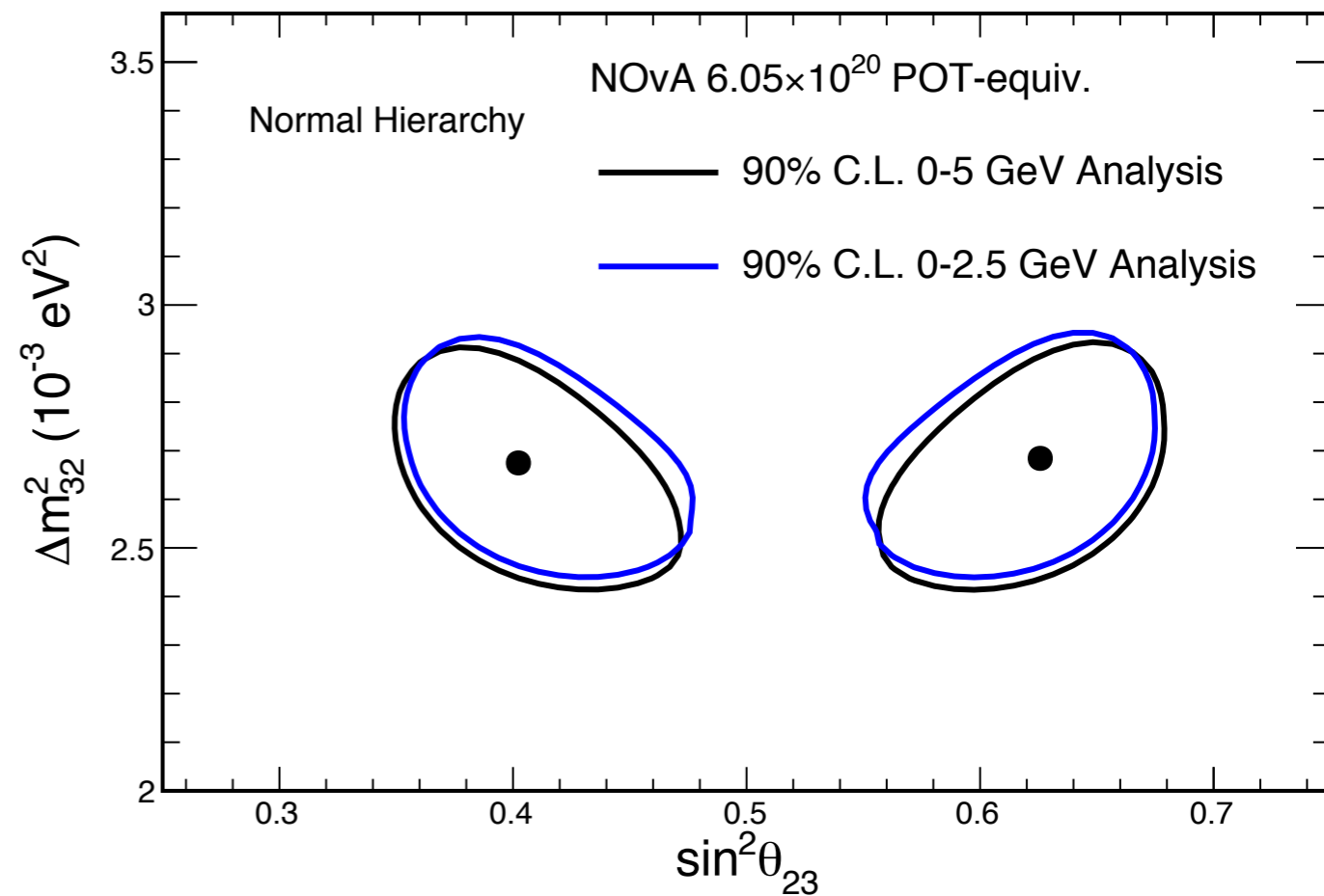
Systematic	Effect on $\sin^2(\theta_{23})$	Effect on $\Delta m^2_{32}$
Normalisation	$\pm 1.0\%$	$\pm 0.2\%$
Muon E scale	$\pm 2.2\%$	$\pm 0.8\%$
Calibration	$\pm 2.0\%$	$\pm 0.2\%$
Relative E scale	$\pm 2.0\%$	$\pm 0.9\%$
Cross sections + FSI	$\pm 0.6\%$	$\pm 0.5\%$
Osc. parameters	$\pm 0.7\%$	$\pm 1.5\%$
Beam backgrounds	$\pm 0.9\%$	$\pm 0.5\%$
Scintillation model	$\pm 0.7\%$	$\pm 0.1\%$
<b>All systematics</b>	<b><math>\pm 3.4\%</math></b>	<b><math>\pm 2.4\%</math></b>
<b>Stat. Uncertainty</b>	<b><math>\pm 4.1\%</math></b>	<b><math>\pm 3.5\%</math></b>

# $\nu_\mu \rightarrow \nu_\mu$ Oscillation Results

NOvA Preliminary



NOvA Preliminary



- Best fit  $\chi^2/\text{DOF} = 41.5/17$  is driven by the high energy tail
- There is no pull in the oscillation fit from the tail

# NC Disappearance Results

- Events classified using CVN
- Analysis Upgrades:
  - Increase of 50% more data
  - Improved cosmic rejection
  - Energy fit
  - 50% improvement to NC selection efficiency
  - Improved cross-section modeling
  - Improved detector response modeling

Background	Estimate
Total Bkgd	42.8
$\nu_\mu$ CC	22.2
$\nu_e$ CC	9.9
$\nu_\tau$ CC	2.8
Cosmic	7.9

Total Signal Prediction: **148.3**

