

# Charged Lepton Flavor Violation:

$$\mu^- N \rightarrow e^- N$$

and

$$\mu^- N \rightarrow e^+ N$$

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R. Bernstein  
Fermilab

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San Juan, Puerto Rico

# Outline

- The search for muon-electron conversion:



- Experimental Technique:

- Mu2e, COMET, DeeMe

- Mu2e-II

- Very preliminary comments on



# What is Muon to Electron Conversion?

- Muon converts to an electron in the field of a nucleus:

$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A, Z) \rightarrow e^- + N(A, Z))}{\Gamma(\mu^- + N(A, Z) \rightarrow \text{all muon captures})}$$

- Effectively zero Standard Model background
  - a signal is new physics
- Experiments to be discussed can measure a signal with SES of  $\sim 3 \times 10^{-17}$ , discovery at  $2 \times 10^{-16}$ 
  - typical predictions for EWSB  $\sim 10^{-15}$

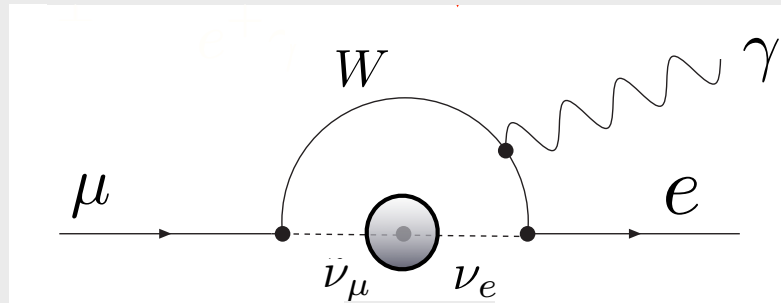
# How Small is Our SES?

- Pretty Rare: let us know if this happens to you!  
You can immediately join the experiment

Probability of...	
rolling a 7 with two dice	1.67E-01
rolling a 12 with two dice	2.78E-02
getting 10 heads in a row flipping a coin	9.77E-04
drawing a royal flush (no wild cards)	1.54E-06
getting struck by lightning in one year in the US	2.00E-06
winning Pick-5	5.41E-08
winning MEGA-millions lottery (5 numbers+megaball)	3.86E-09
your house getting hit by a meteorite this year	2.28E-10
drawing two royal flushes in a row (fresh decks)	2.37E-12
your house getting hit by a meteorite today	6.24E-13
getting 53 heads in a row flipping a coin	1.11E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next six months	1.14E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next three months	2.85E-17

# Neutrinos and Mu2e

- Neutrino Oscillations are the only Standard Model background (except neutrino oscillations are not in the Standard Model)



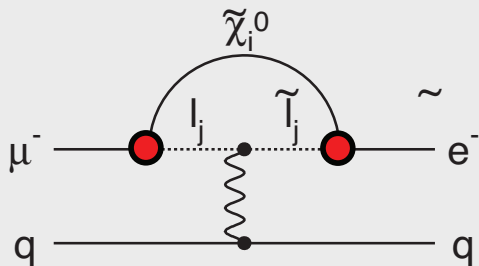
$$\text{BR}(\mu \rightarrow e \gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-54}$$

- nobody understood why  $\mu \rightarrow e \gamma$  wasn't  $10^{-4}$  until we hypothesized two neutrinos ( $\nu_\mu \neq \nu_e$ )
- linkage of CLFV to oscillations, see-saw, Majorana vs. Dirac are getting more interest

# Contributions to $\mu e$ Conversion

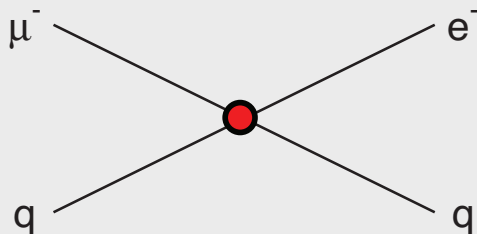
## Supersymmetry

rate  $\sim 10^{-15}$



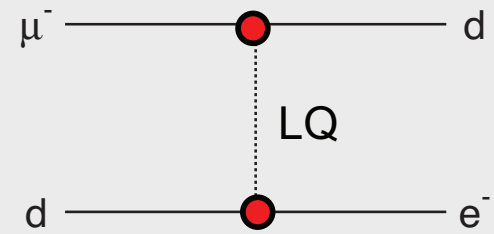
## Compositeness

$\Lambda_c \sim 3000 \text{ TeV}$



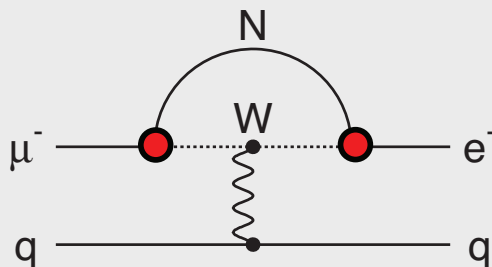
## Leptoquark

$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{ed})^{1/2} \text{ TeV}/c^2$



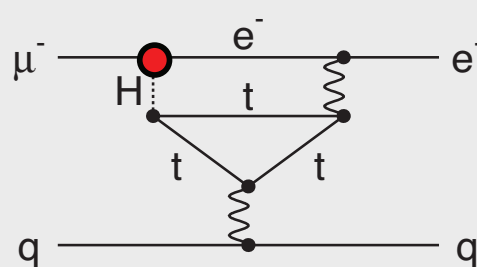
## Heavy Neutrinos

$|U_{\mu N} U_{eN}|^2 \sim 8 \times 10^{-13}$



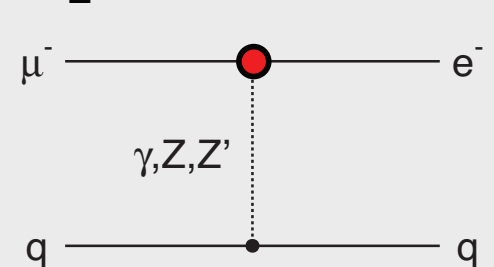
## Second Higgs Doublet

$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu\mu})$



## Heavy Z' Anomal. Z Coupling

$M_{Z'} = 3000 \text{ TeV}/c^2$



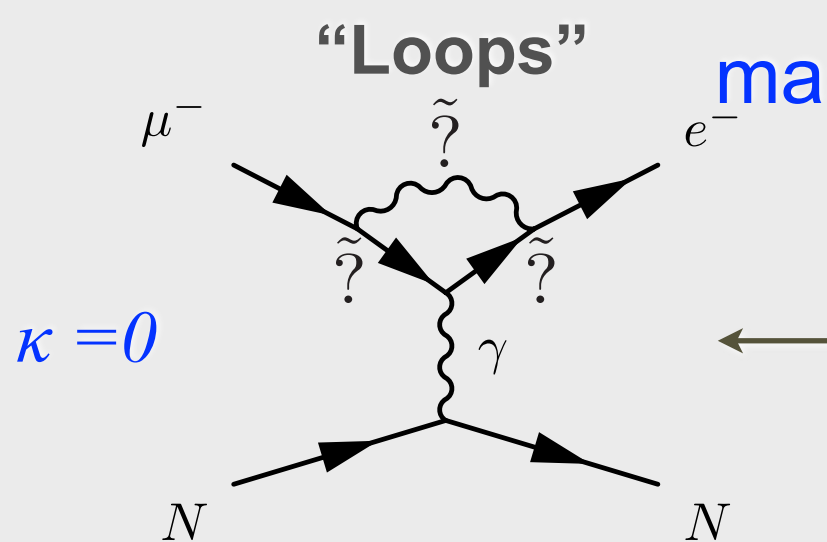
also see Flavour physics of leptons and dipole moments, [arXiv:0801.1826](https://arxiv.org/abs/0801.1826) ;

Marciano, Mori, and Roney, Ann. Rev. Nucl. Sci. 58, doi:[10.1146/annurev.nucl.58.110707.171126](https://doi.org/10.1146/annurev.nucl.58.110707.171126) ;

de Gouvea and Vogel, arXiv:1303.4097v2

# Effective Lagrangian

$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{u}_L \gamma_\mu u_L + \bar{d}_L \gamma_\mu d_L)$$



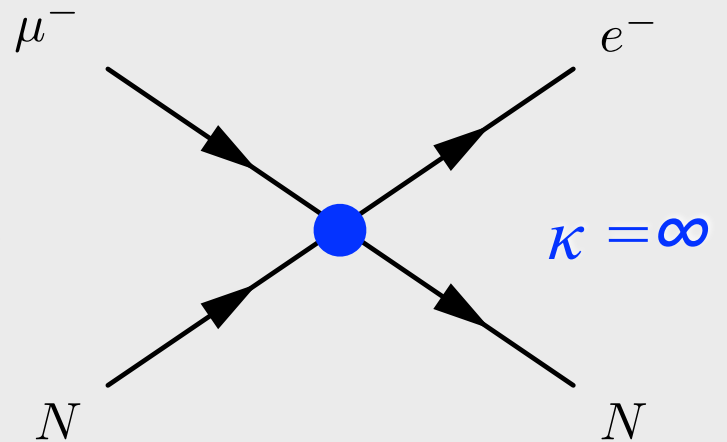
Supersymmetry and Heavy Neutrinos

**Contributes to  $\mu \rightarrow e\gamma$**   
(just imagine the photon is real)

mass scale  $\Lambda$

$\kappa$

**“Contact Terms”**



New Particles at High Mass Scale (leptoquarks, heavy Z,...)

**Does not produce  $\mu \rightarrow e\gamma$**

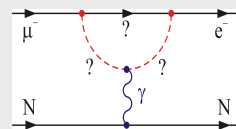
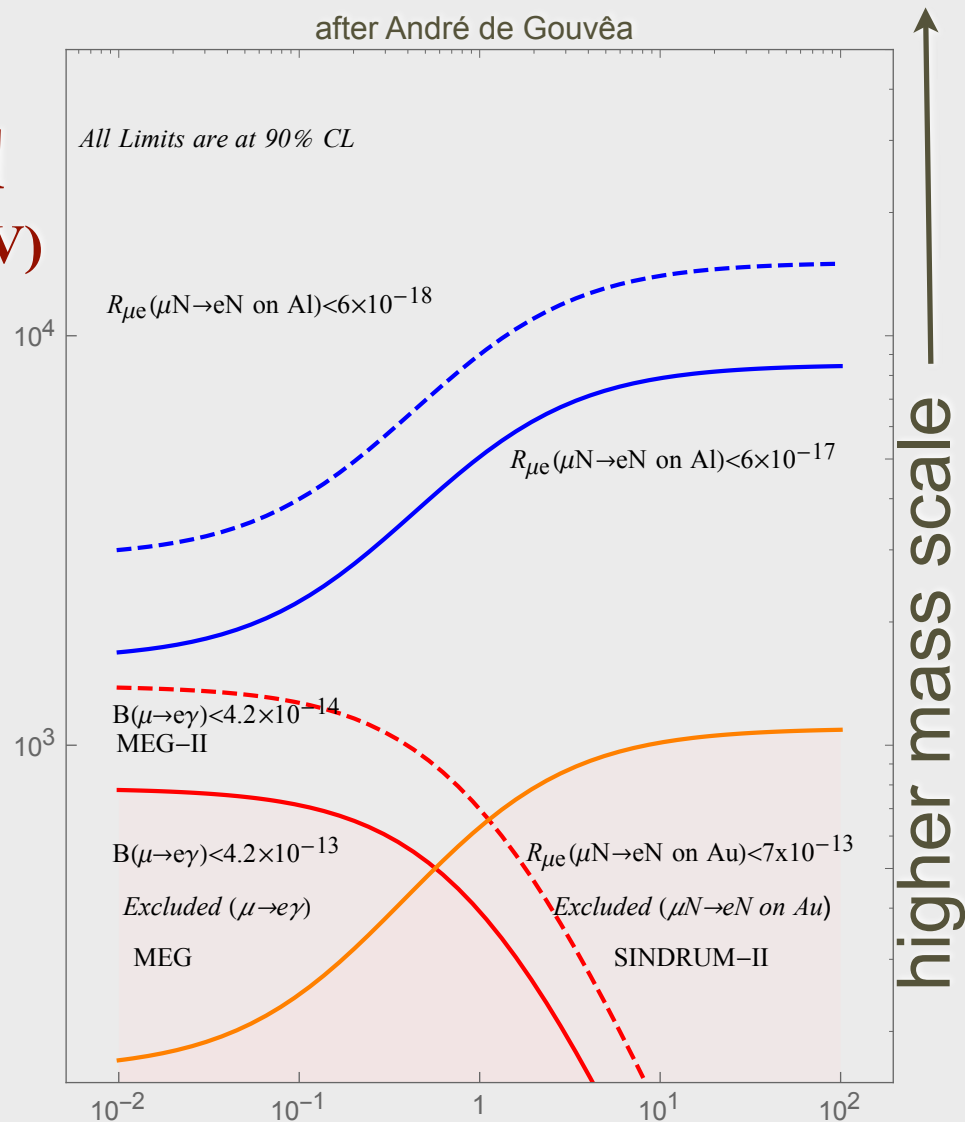
# $\mu e$ Conversion and $\mu \rightarrow e \gamma$

1) Mass Reach to  $\sim 10^4$  TeV for unit coupling, x10 in mass scale for dim-6

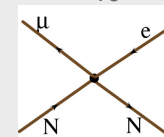
2) Mu2e and MEG-II are complementary in loop-dominated physics.

3) These are discovery experiments

$\Lambda$   
(TeV)



$\kappa$

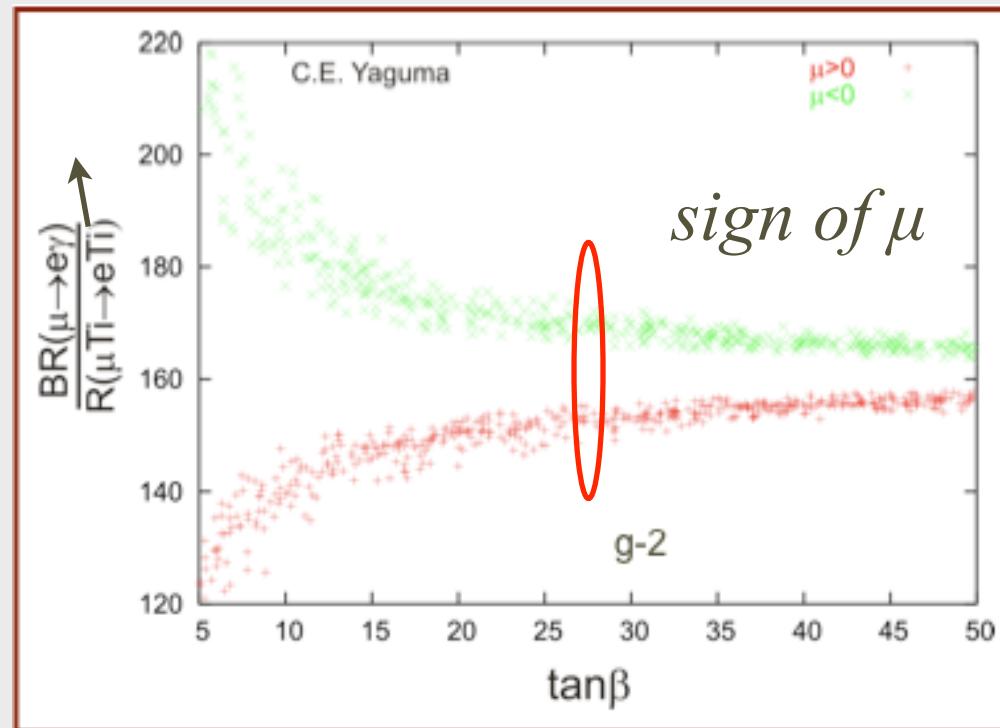




# Mu2e, g-2, and $\mu \rightarrow e\gamma$

Yaguna,  
hep-ph/0502014v2

MSSM w mSUGRA



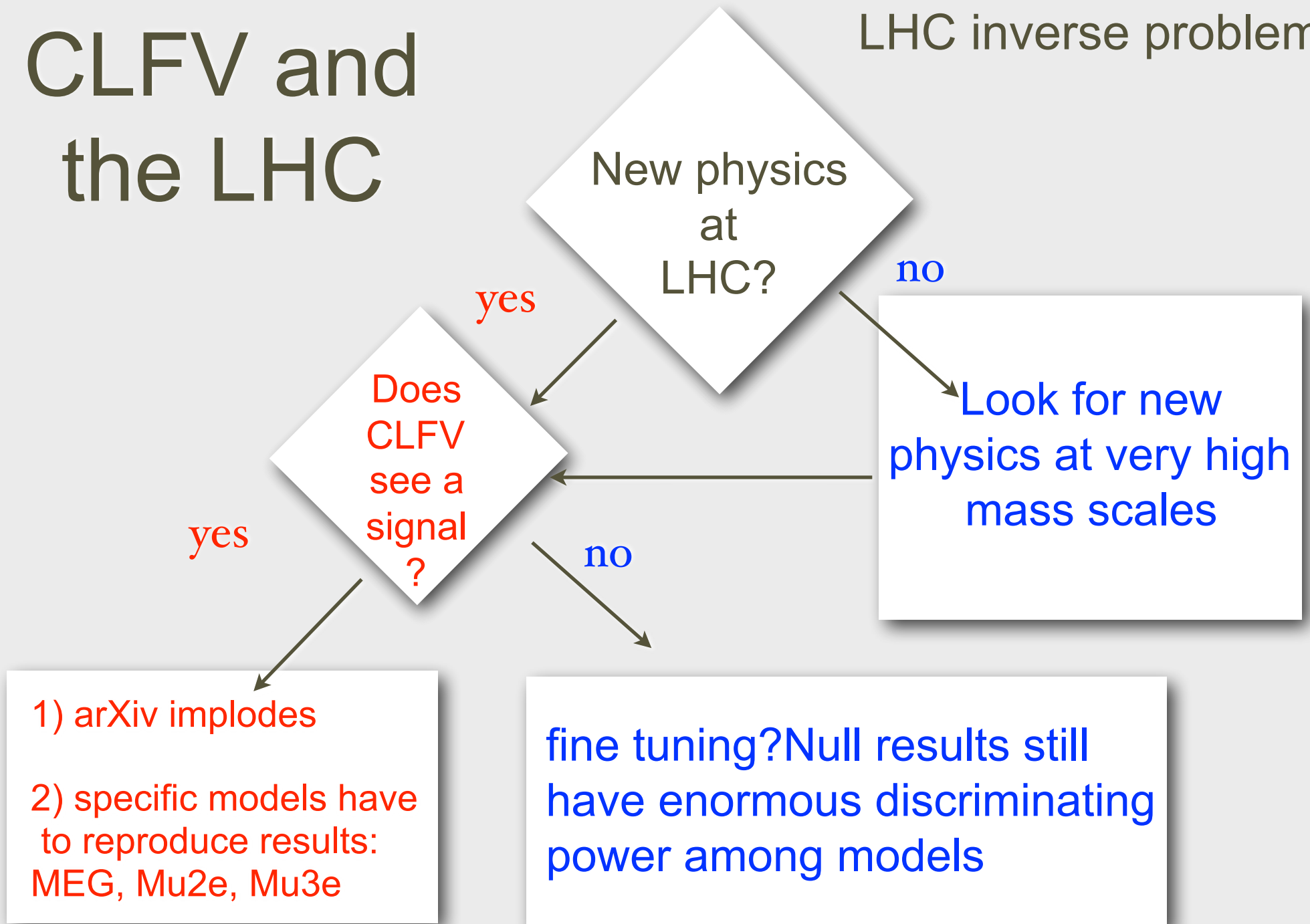
$$M_H = \sqrt{2} |\mu|$$

- Need:
  - observation of CLFV in more than one channel, and/or
  - evidence from LHC, g-2, or elsewhere

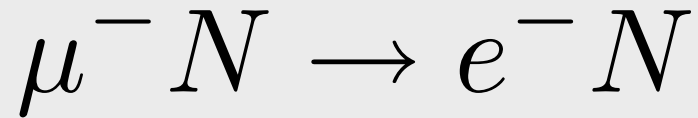
to allow discrimination among different models

# CLFV and the LHC

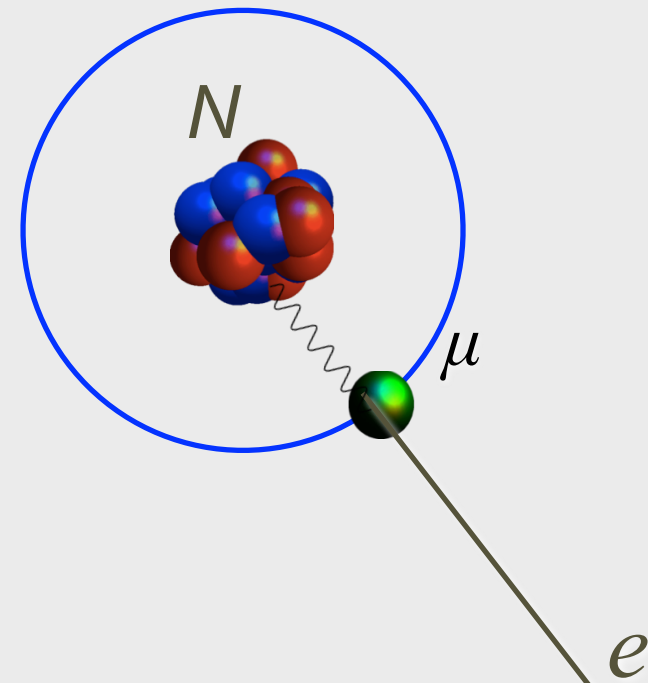
LHC inverse problem



# Experimental Signal



- A Single Mono-energetic Electron
  - muon falls into 1s state in “stopping target”
- Nucleus coherently recoils off outgoing electron, no breakup
- If  $N = \text{Al}$ ,  $E_e \sim 105. \text{ MeV} \sim m_\mu$ 
  - electron energy depends on  $Z$
  - physics in  $Z$ -dependence of rate



very low energies for  
HEP!

# Mu2e: x10,000 improvement

- World's most intense muon beam:
  - $10^{10}$   $\mu$ /sec at detector under construction
  - Superconducting solenoid system to form muon beam
    - Efficient muon collection; see  $\sim 10^{18}$  muons
- Have to reduce backgrounds to take advantage of statistics
  - Pulsed Beam reduces a background
  - And a high-resolution spectrometer to make muon-related backgrounds small

# Backgrounds and Design

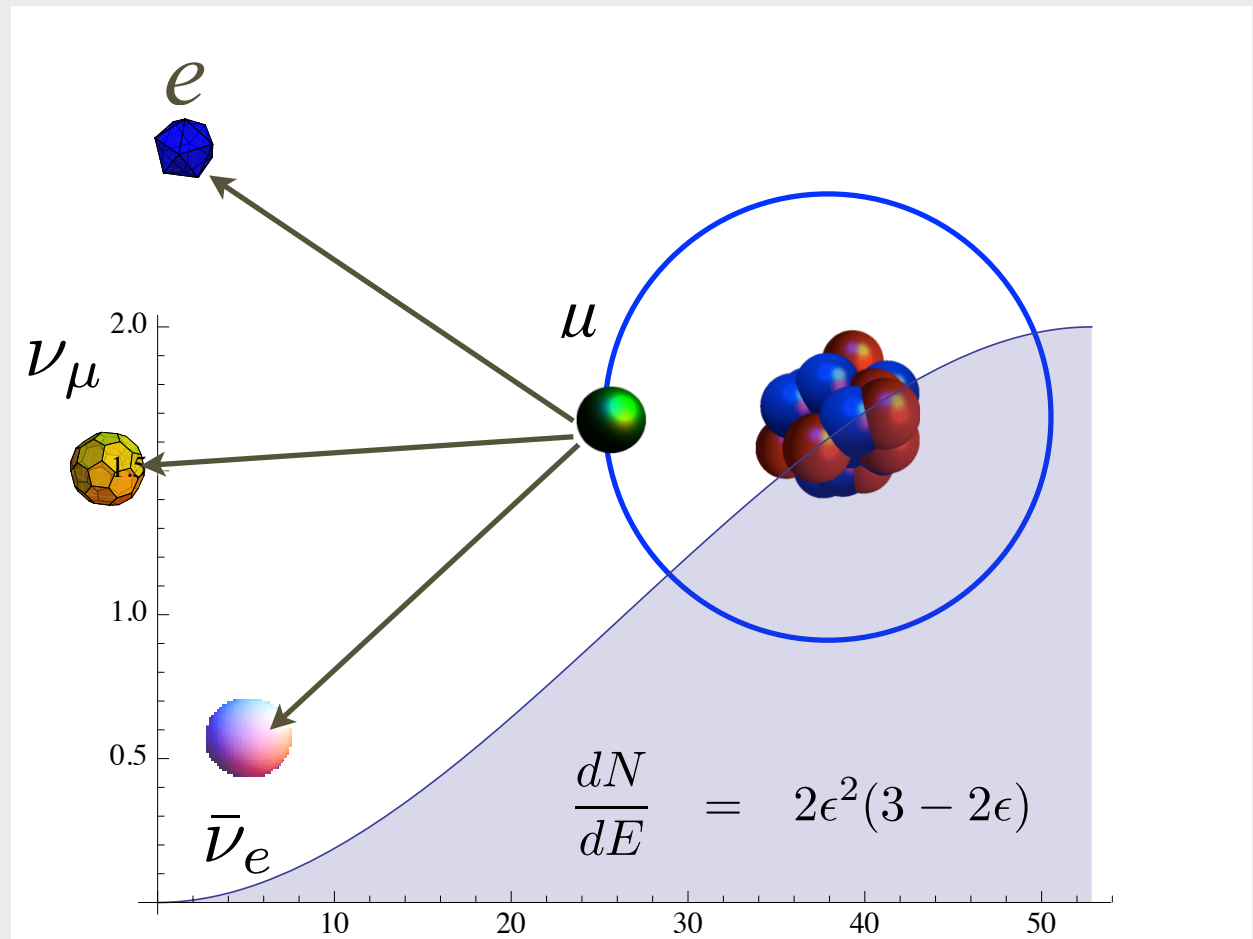
	<i>source</i>	<i>scales with</i>	<i>solution</i>
Intrinsic Backgrounds	decay-in-orbit	# of muons	detector resolution
Beam	radiative pion capture	closeness to beam pulse	pulsed beam
Running Time	cosmic ray	live time	veto system

# Decay-In-Orbit: usually not background

- Peak and Endpoint of free muon decay spectrum is at

$$E_{\text{max}} = \frac{m_{\mu}^2 + m_e^2}{2m_{\mu}} \approx 52.8 \text{ MeV}$$

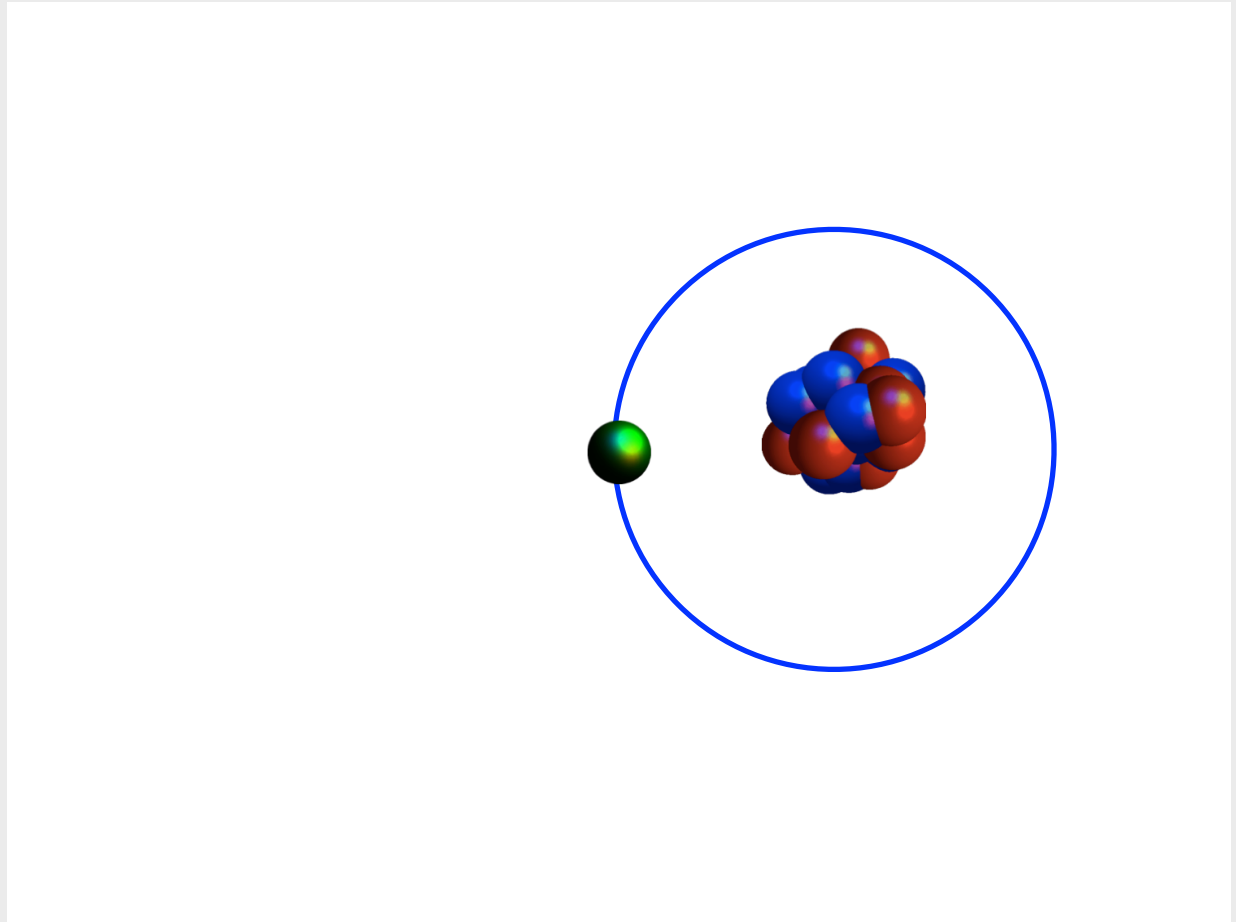
- Detector will be insensitive to electrons at this energy
- Recall *signal* at  $105 \text{ MeV} \gg 52.8 \text{ MeV}$



Michel=free muon decay

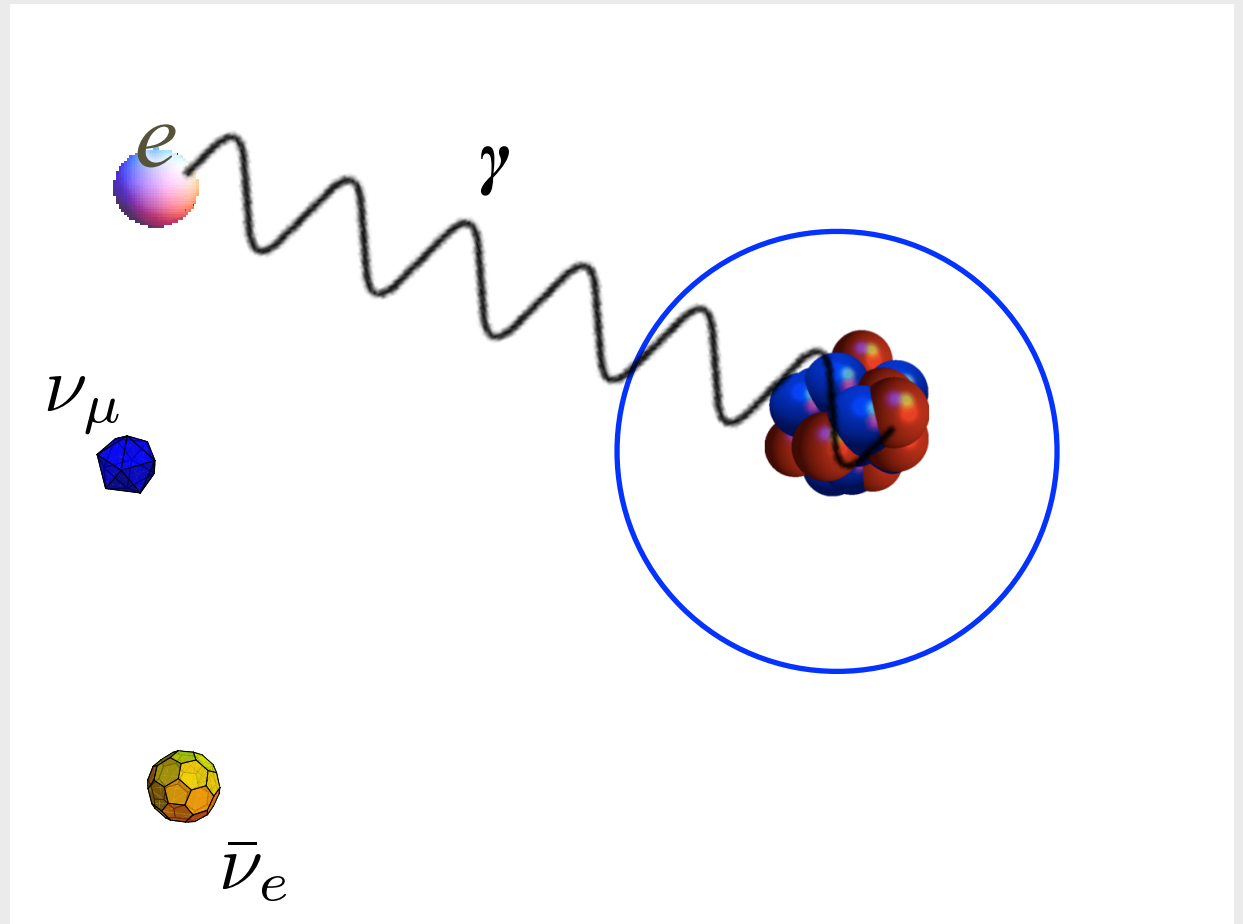
# Decay-In-Orbit Background

- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, **the DIO electron can be exactly at conversion energy** (up to neutrino mass)



# Decay-In-Orbit Background

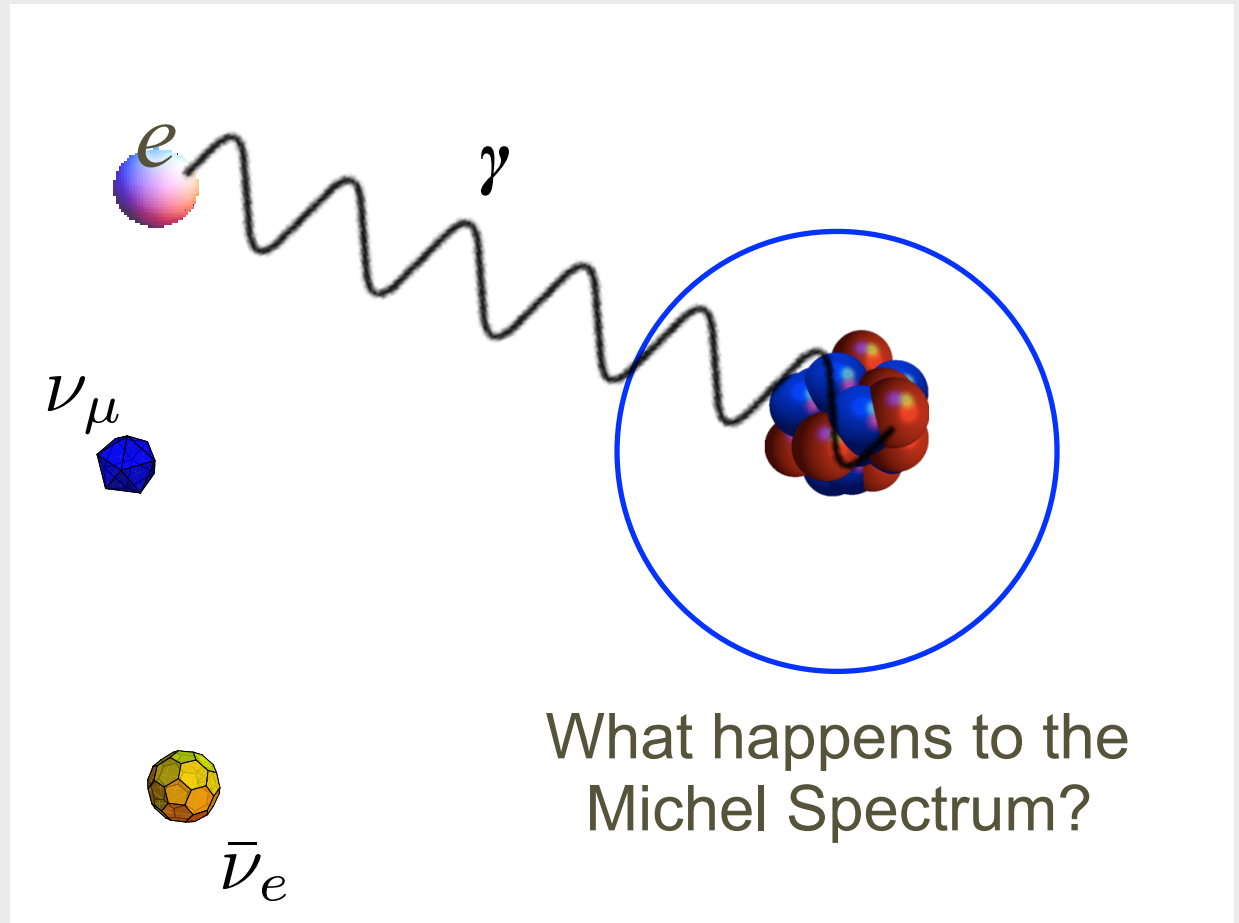
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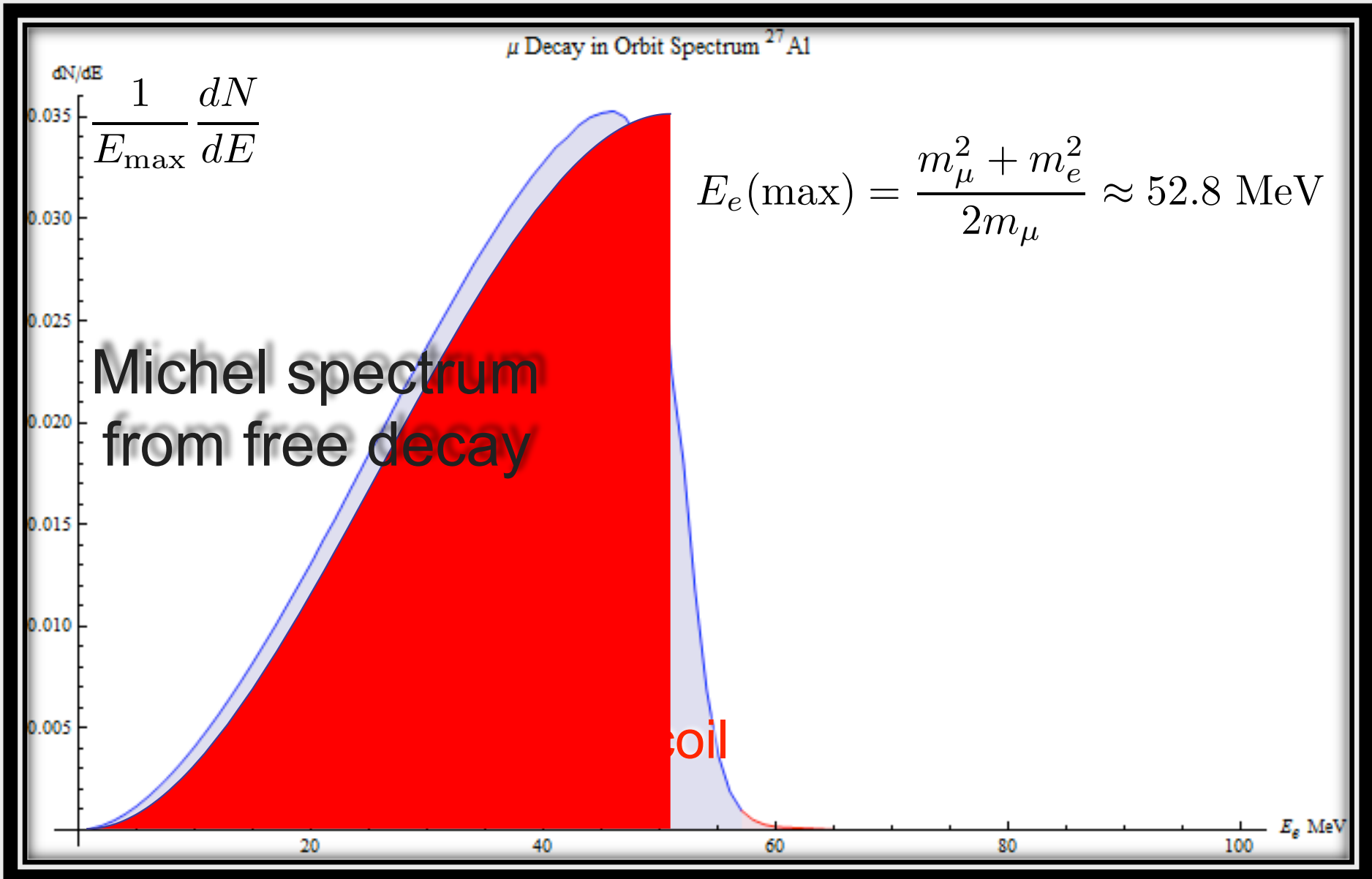
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# Decay-in-Orbit Shape

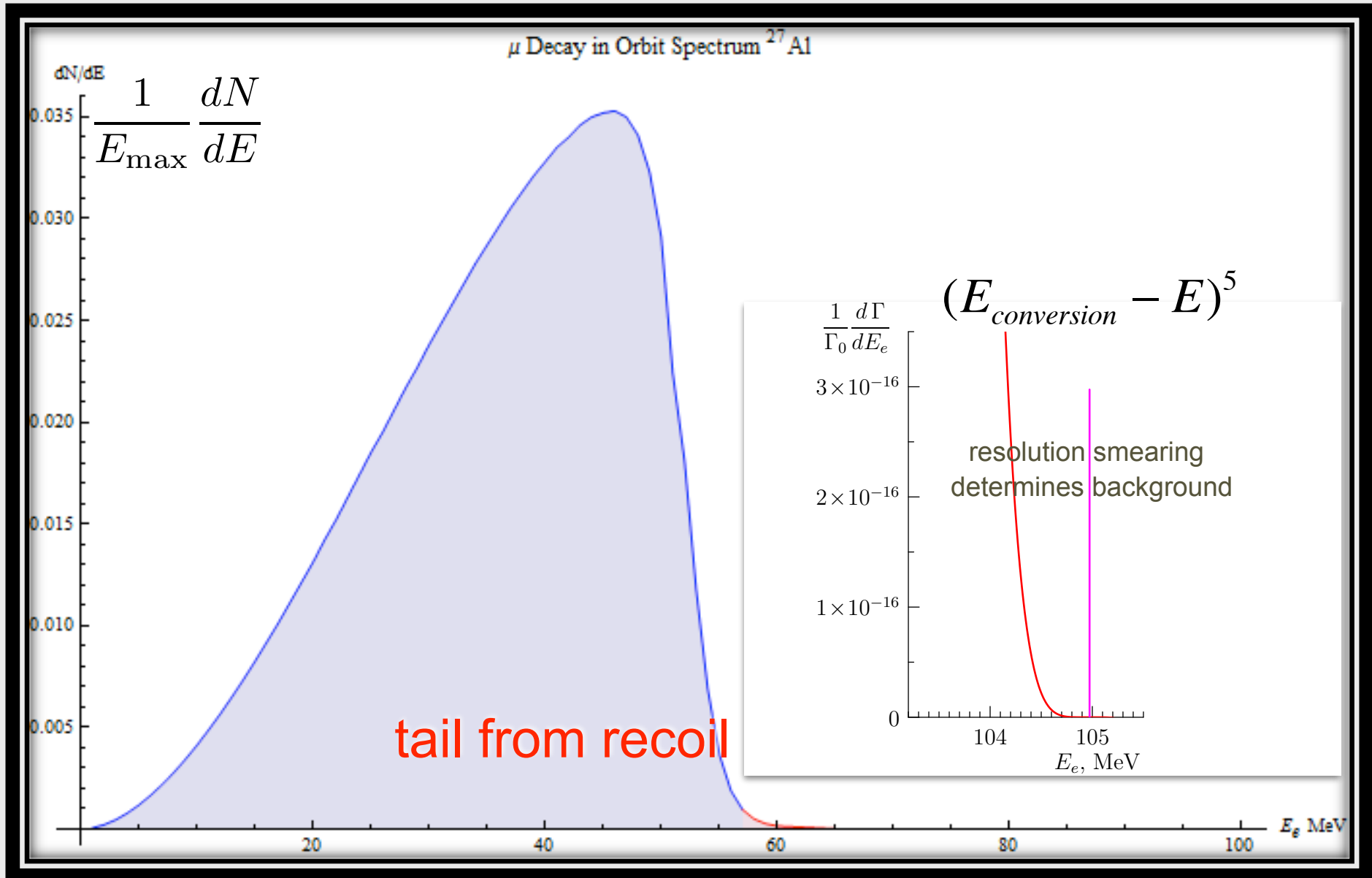
Szafron 10.5506/APhysPolB.46.2279: Radiative Corrections



Czarnecki: 10.1016/j.physletb.2015.12.008 many other papers

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# Backgrounds and Design

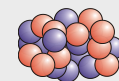
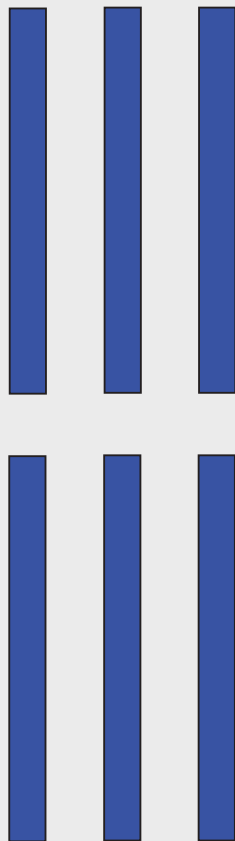
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<b>Beam</b>	<b>radiative pion capture</b>	<b>closeness to beam pulse</b>	<b>pulsed beam</b>
Running Time	cosmic ray	live time	veto system

# Radiative Pion Capture

- Need  $\pi$ 's to make  $\mu$ 's
- Sometimes those  $\pi$ 's live long enough to make it to the stopping target
- They can then undergo radiative pion capture (RPC):
  - $\pi N \rightarrow \gamma N'$  and the  $\gamma$  converts in the Al.
  - $\pi N \rightarrow \gamma N'$  and the  $\gamma$  internally converts
- Sometimes the  $e^-$  is at the signal energy

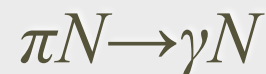
# Prompt Backgrounds and Pulsed Beam

Al target foils: muon converts here



= muons, electrons, pions

Radiative Pion Capture:

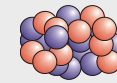
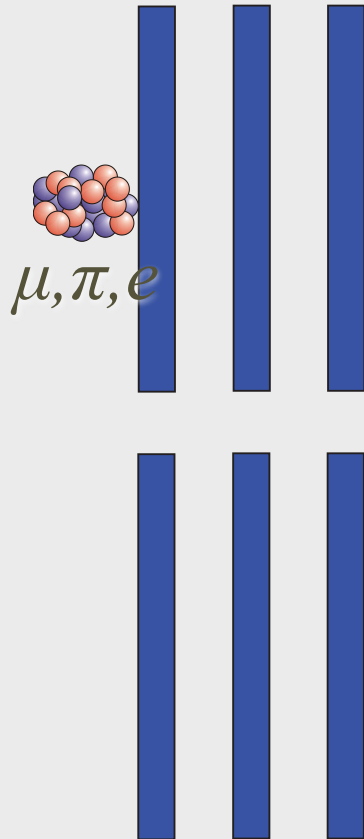


$\gamma \rightarrow e^+ e^-$  in foils

pulsed beam lets us wait (1) until after beam flash, and (2) RPC background decays away

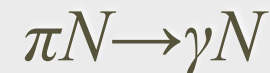
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Radiative Pion Capture:

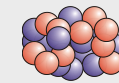


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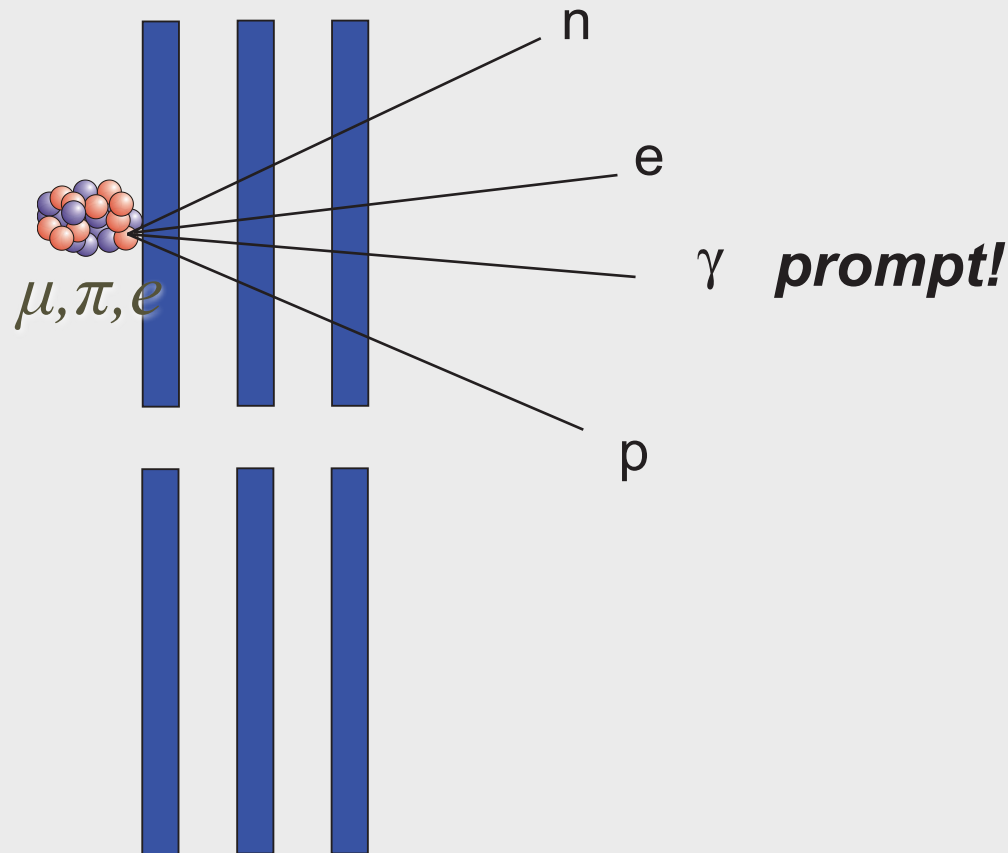
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# Prompt Backgrounds and Pulsed Beam

All target foils: muon converts here



= muons, electrons, pions



Radiative Pion Capture:

$$\pi N \rightarrow \gamma N$$

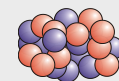
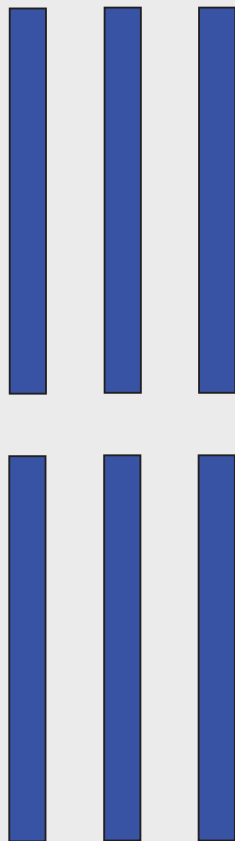
$$\gamma \rightarrow e^+ e^- \text{ in foils}$$

pulsed beam lets us wait (1) until after beam flash, and (2) RPC background decays away



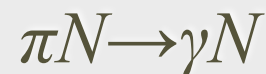
# Prompt Backgrounds and Pulsed Beam

Al target foils: muon converts here



= muons, electrons, pions

Radiative Pion  
Capture:

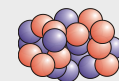


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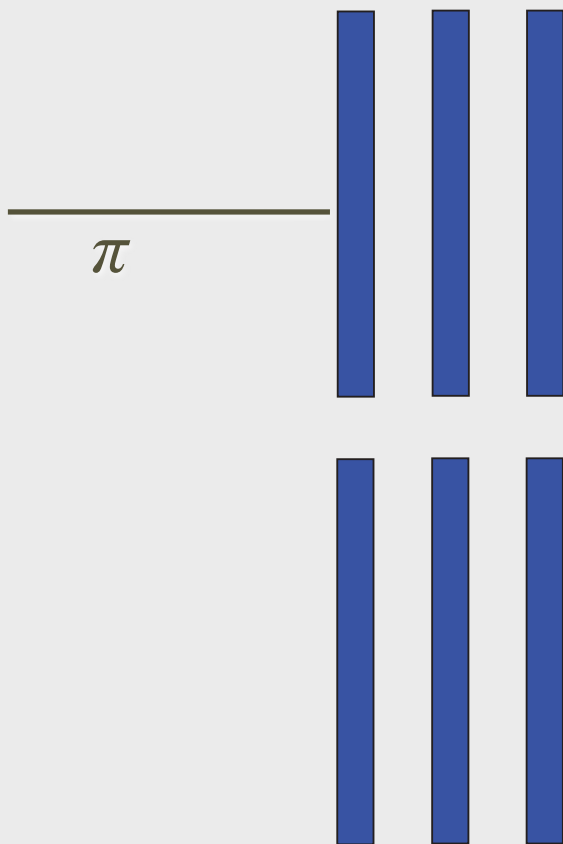
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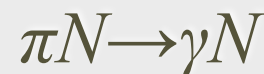
Al target foils: muon converts here



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Radiative Pion Capture:

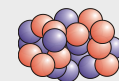


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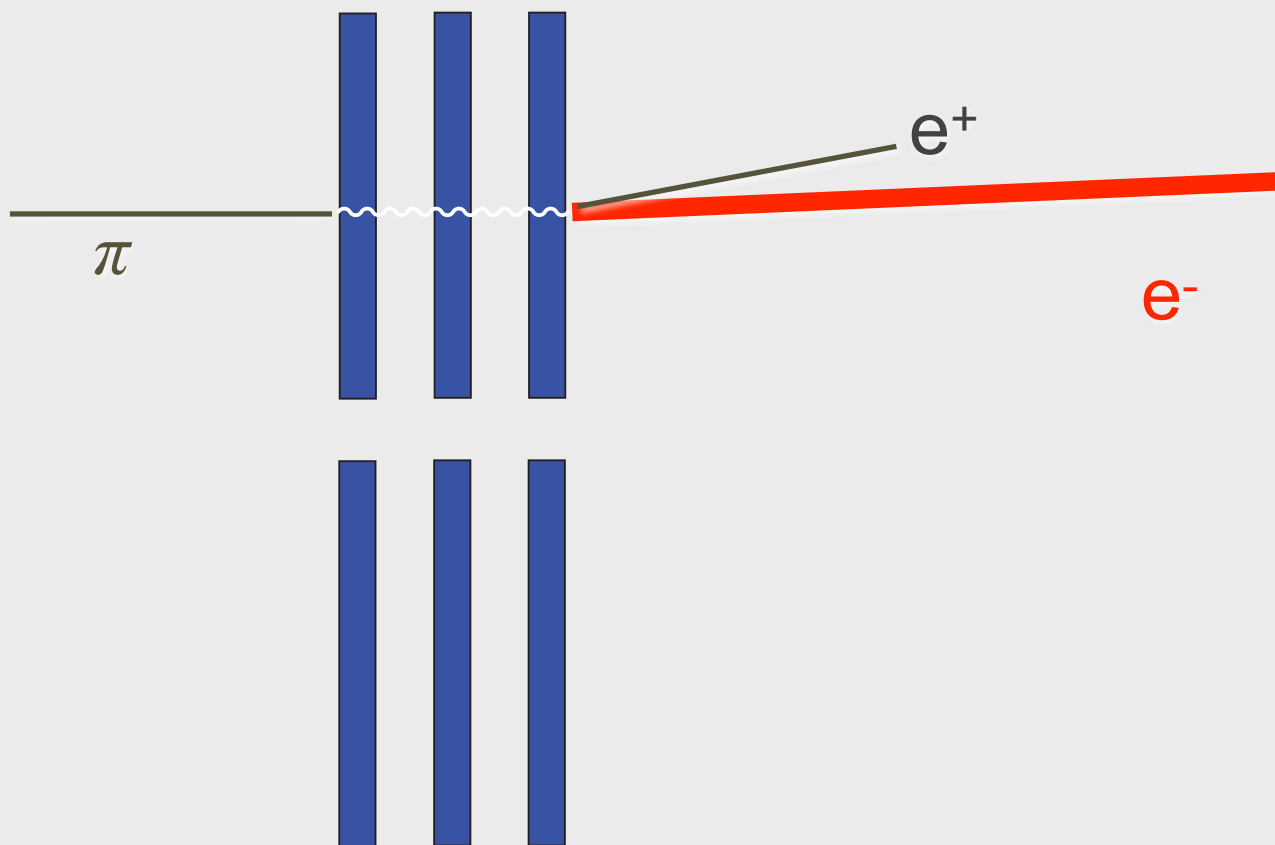
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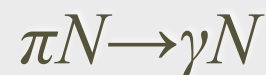
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= muons, electrons, pions



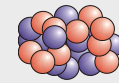
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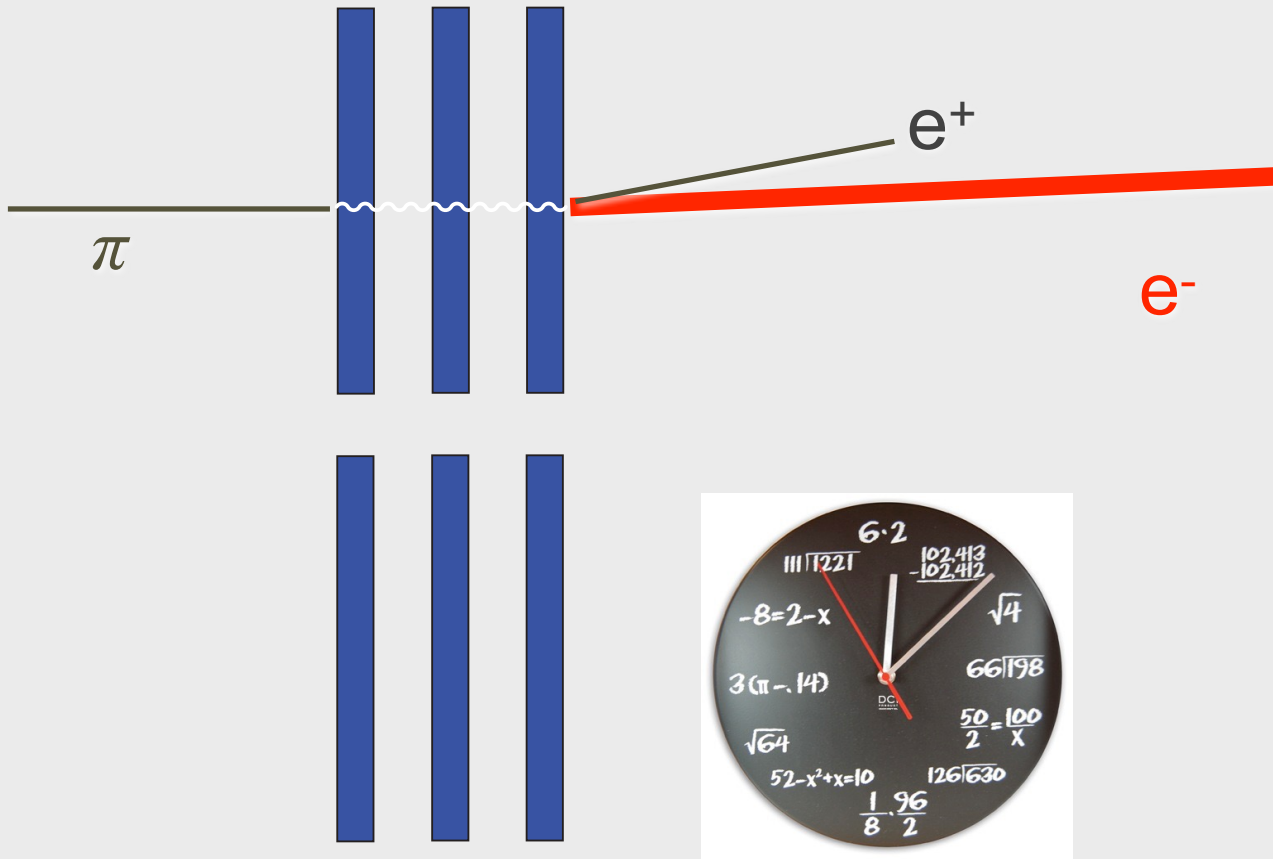
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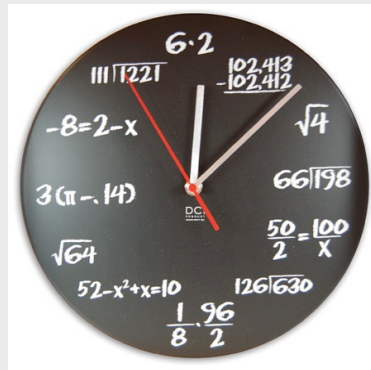
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Radiative Pion Capture:

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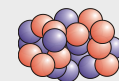
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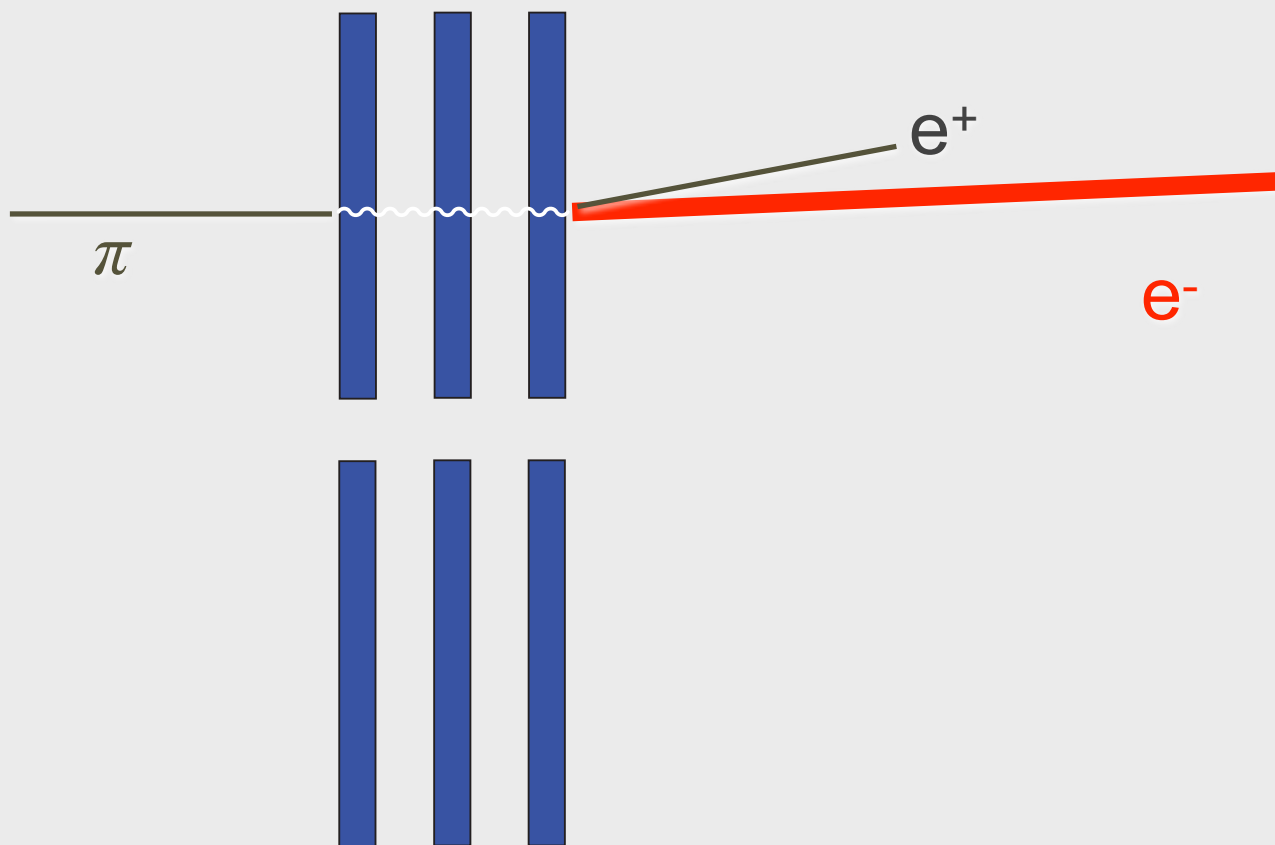
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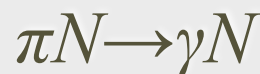
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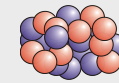
Radiative Pion Capture:



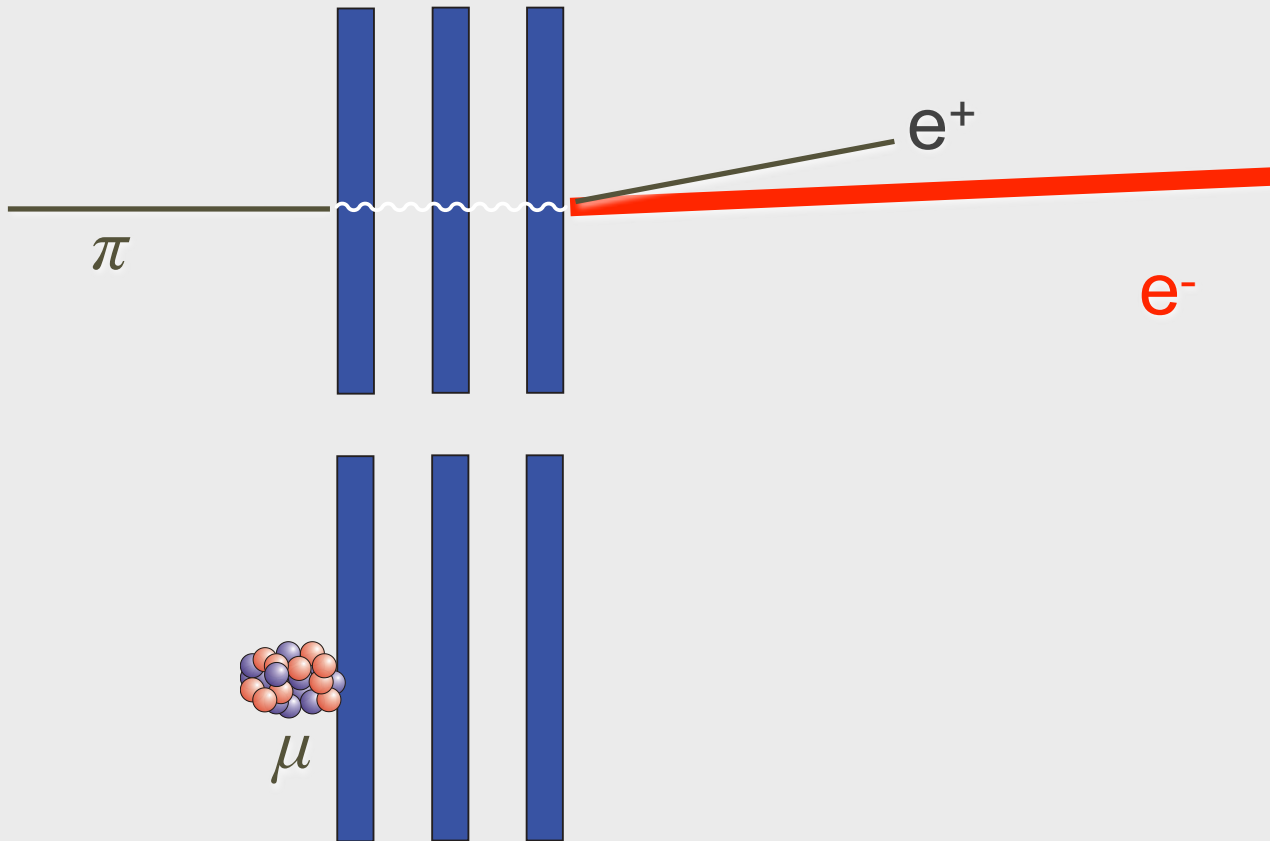
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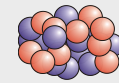
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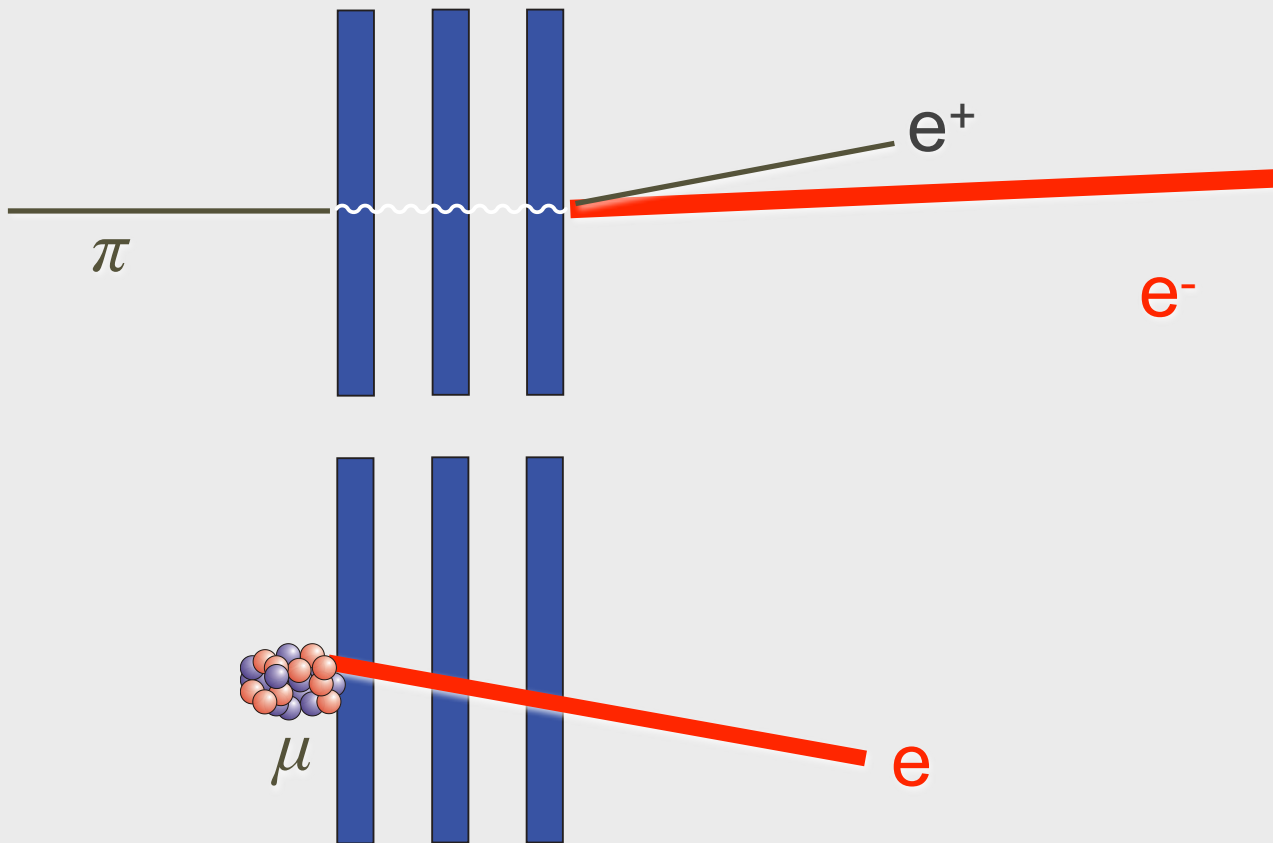
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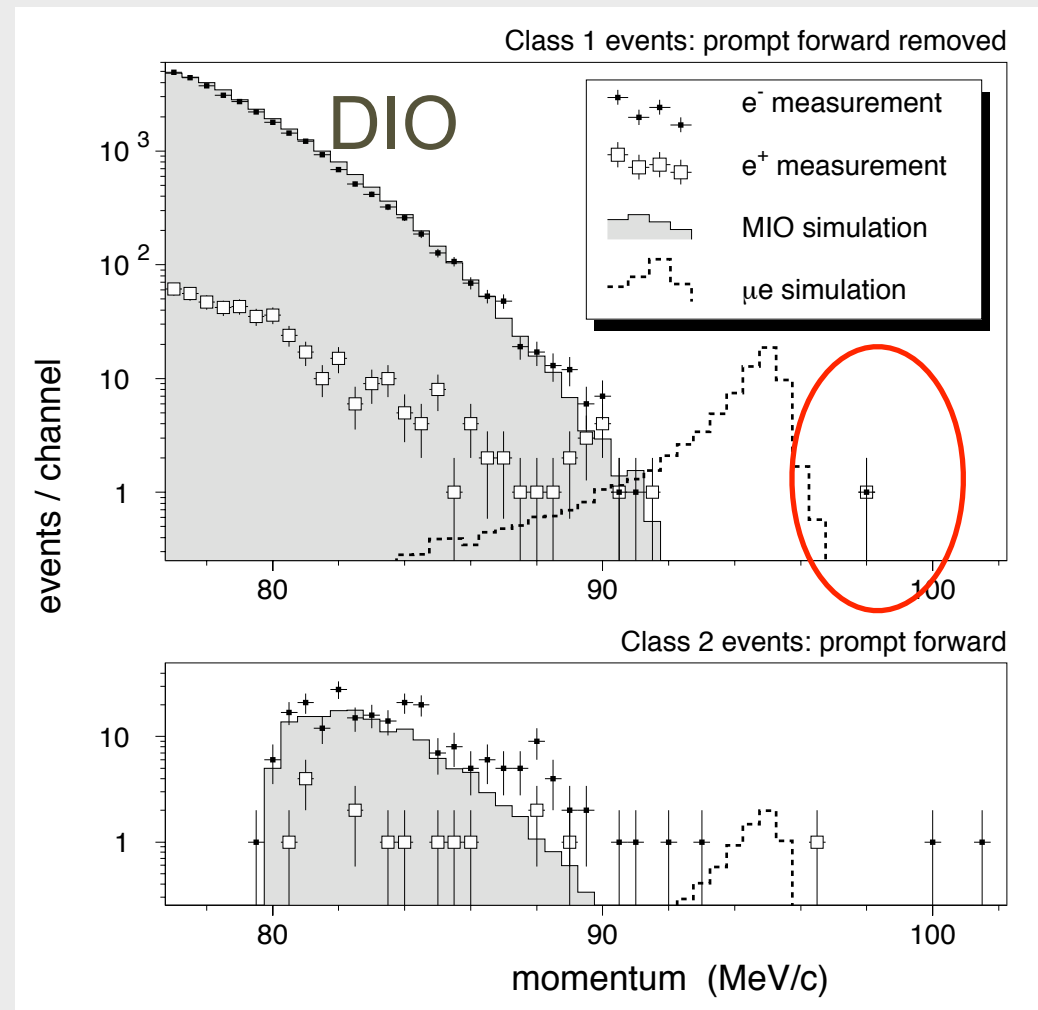
now  $\mu$  is only source of  
105 MeV electron

pulsed beam lets us wait (1) until after beam flash, and (2) RPC background decays away

# SINDRUM-II Results

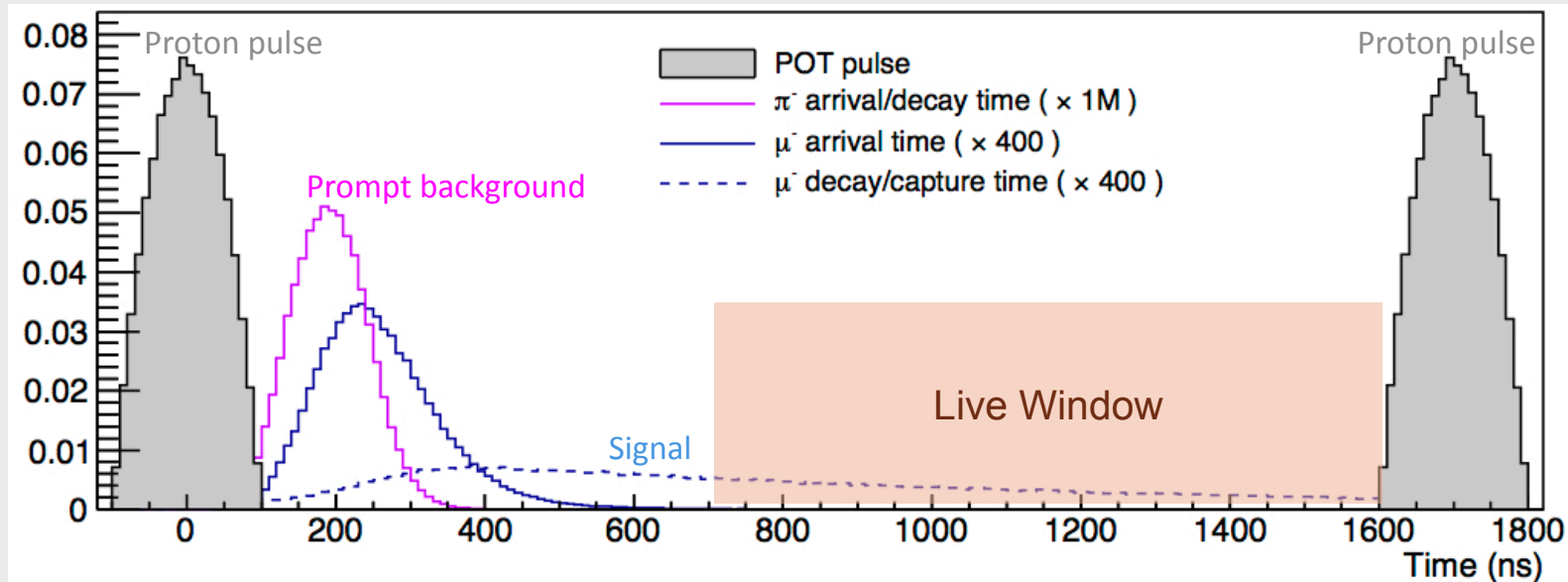
- Effectively constant beam
  - 51 MHz (**~20 nsec**) repetition rate, ~0.3 nsec pulse
- Small time separation between signal and prompt pion backgrounds
  - bottom plot is first half of **20 nsec**, top plot is 2nd half
  - delay lowers background

$$B_{\mu e}^{\text{Au}} < 7 \times 10^{-13} \text{ @ 90\% CL}$$





# Pulsed Beam Structure

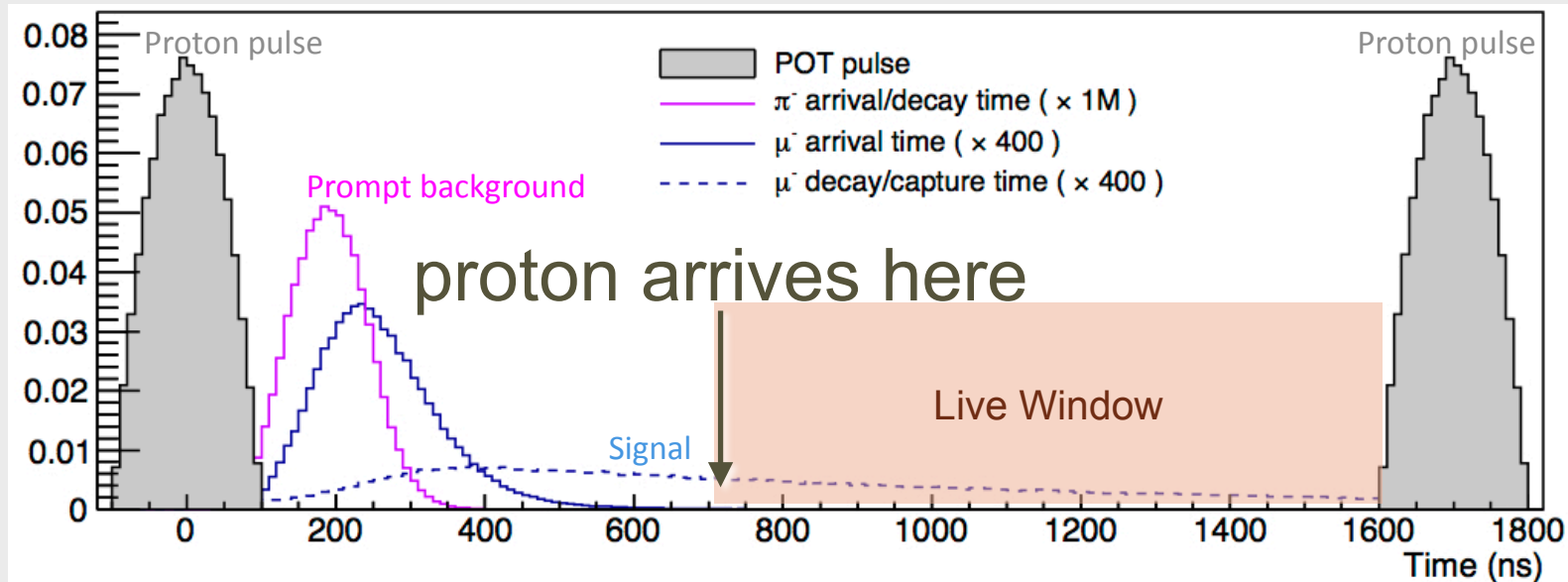


$\pi$  backgrounds

signal window

- Wait out pion backgrounds, then open signal window
- Can't “restart the clock” with out-of-pulse protons
- Need  $10^{-10}$  in-pulse/out-of-pulse ratio (**extinction**)
  - will measure and monitor during experiment

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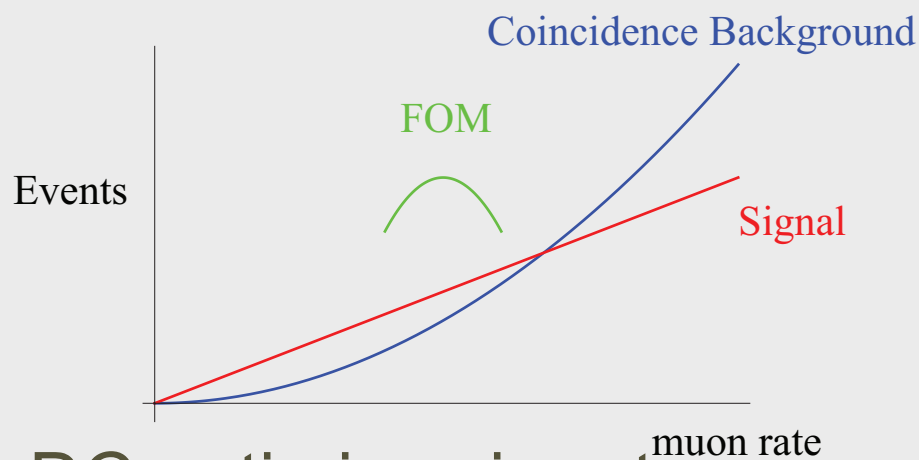
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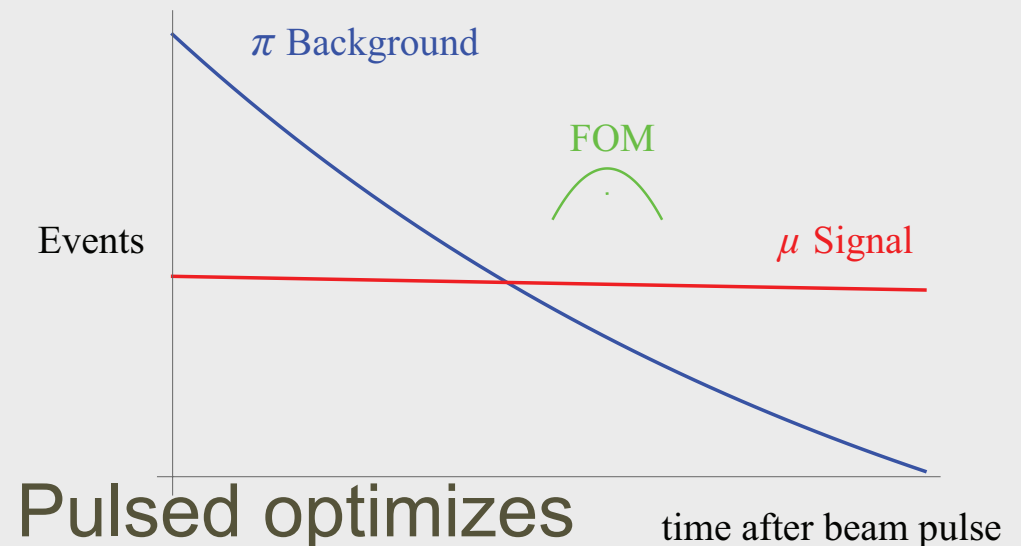
# Pulsed vs. “Steady” Beam

- Decay expts,  $\mu \rightarrow e\gamma$  and  $\mu \rightarrow eee$  want a steady beam: PSI
  - need to minimize backgrounds from coincidences of two decays
  - $(\text{Rate})^2$  bkg vs  $\text{Rate}(\text{signal})$
- muon-electron conversion wants a pulsed beam: FNAL/J-PARC
  - Many pion-induced backgrounds after proton pulse
  - Take advantage of 26 nsec lifetime to “wait it out”



DC optimizes in rate

R. Bernstein (FNAL)

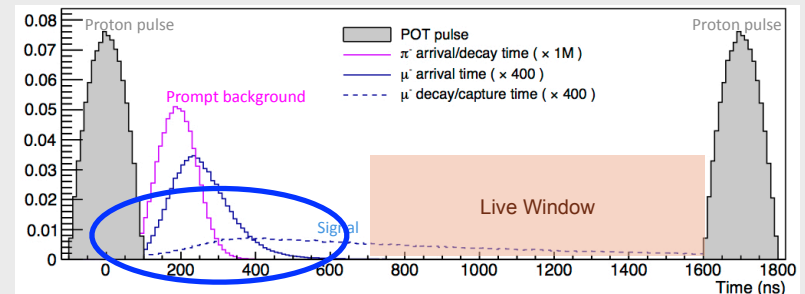


Pulsed optimizes  
in time

COFI 2018

# Beam Flash

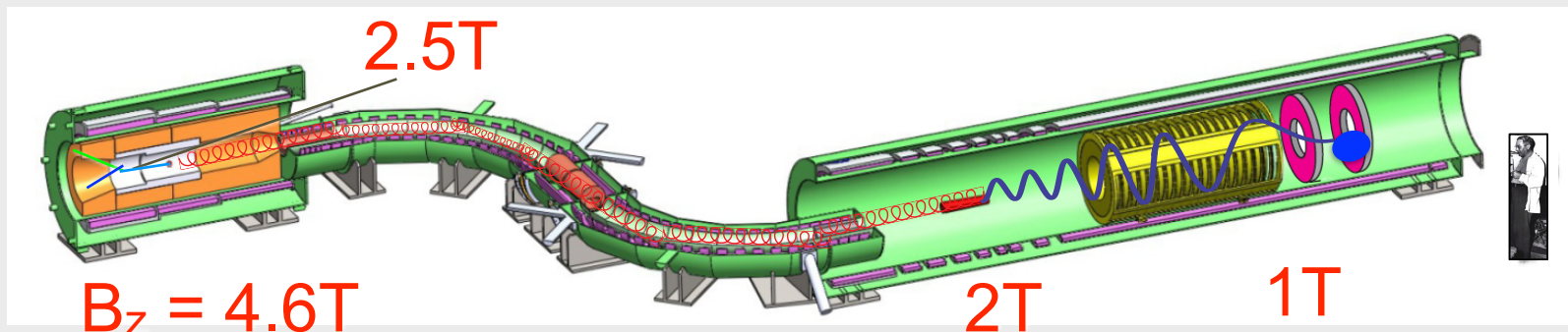
- After proton beam strikes target, “flash” of photons and electrons, and many electrons hit stopping target and detector
- Large neutron flux as well that would deaden detectors
- By waiting:
  - activity from beam flash is over
  - And pion backgrounds have decayed by  $\sim 10^{11}$
- This early flash has implications for Mu2e and especially for potential upgrades



# Actual Experiments: Mu2e/COMET

[mu2e.fnal.gov](http://mu2e.fnal.gov)

4.6T  $\longrightarrow$  B-field gradient  $\longrightarrow$  1T



Mu2e

Production

Transport

Detector

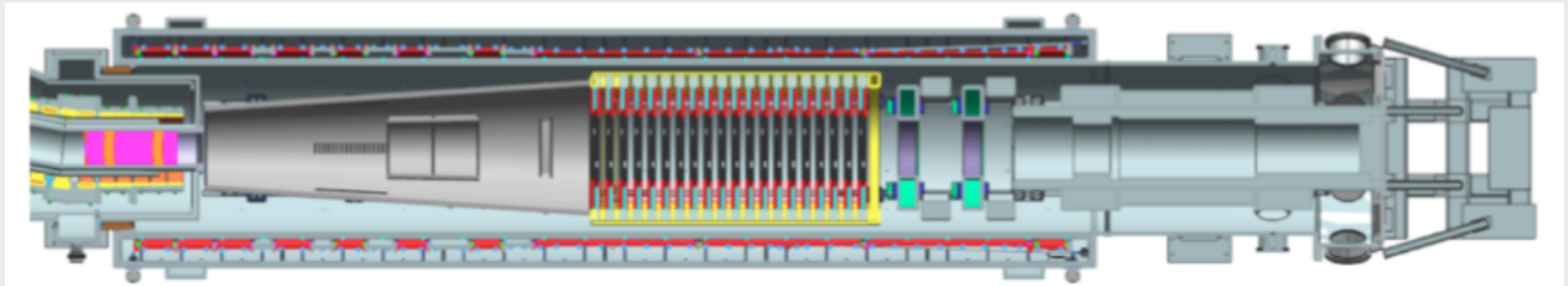
- Target protons at 8 GeV inside superconducting solenoid at 8kW: shielding a challenge
- Capture muons and guide through S-shaped region to Al stopping target

# Detector Layout

finding electron track and momentum determination: highest precision, needed to reject decay-in-orbit

tracker

beam stop

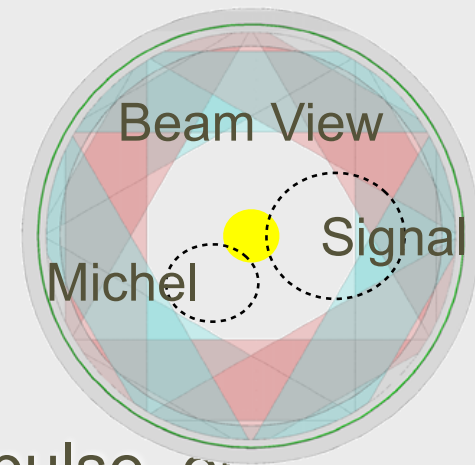


stopping  
target

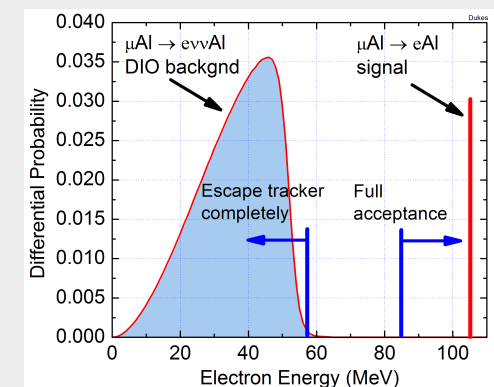
calorimeter

- 1) PID separation of  $\pi/\mu/e$
- 2) Seed for Track Finder
- 3) Standalone Trigger

# Why Annular?

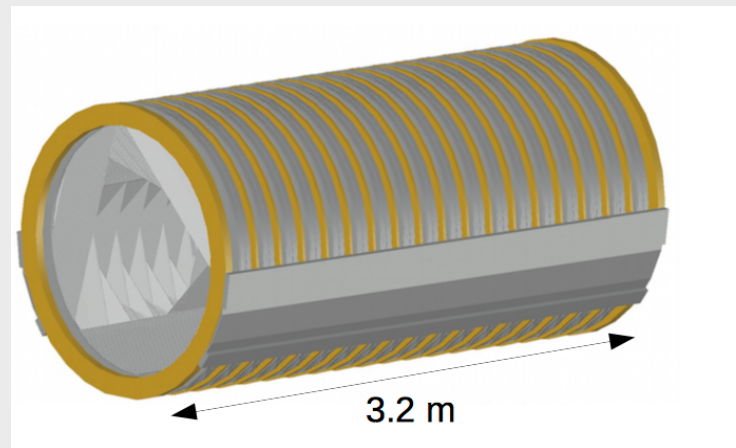


- Reduces beam activity and a background
  - don't want initial "flash" of activity from proton pulse, or remnant muon beam in detector (only  $\sim 1/2$  of muon beam stops in target)
  - eliminates all but a tiny number of decay-in-orbit background
  - $p \sin\theta \sim qBR$  in solenoid, Michel  $e^-$  at small radius
  - so we're not rejecting  $10^{18}$ ; only  $\sim 10^5$  even reconstructed
- COMET uses a different method; not annular
  - advantages and disadvantages



# Straw Tube Tracker

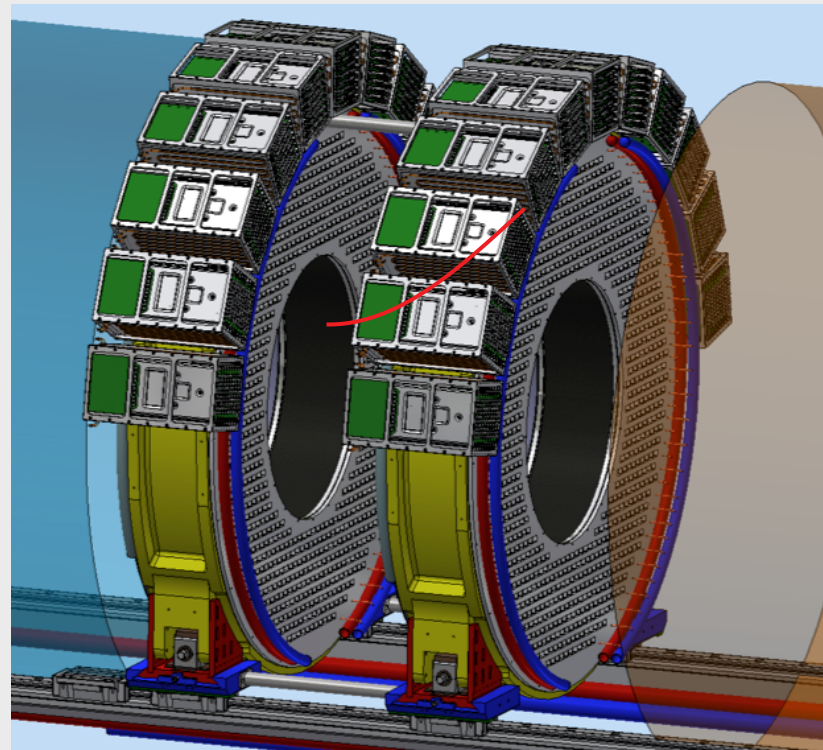
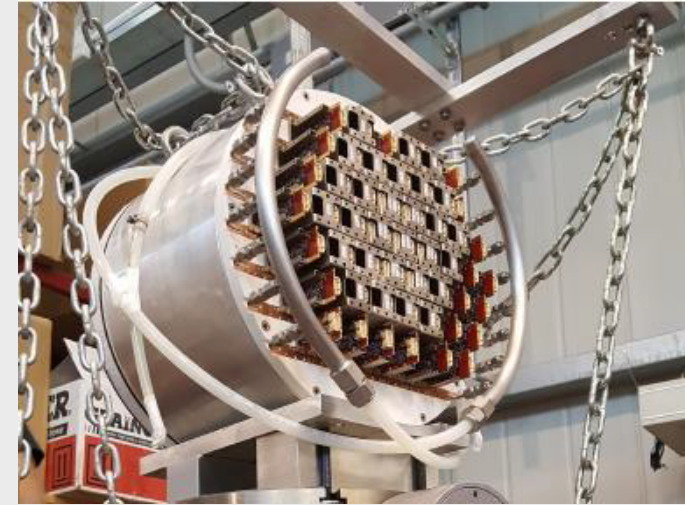
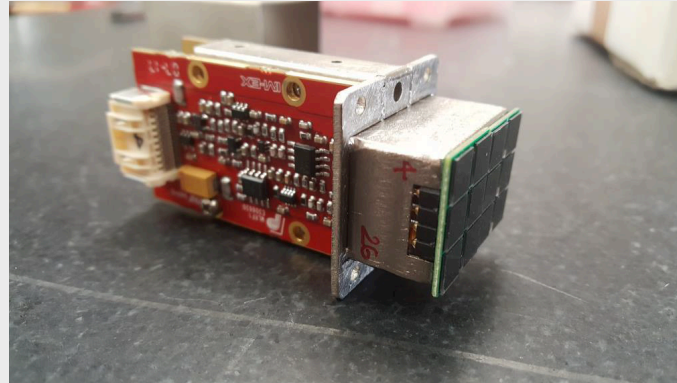
- 15 microns of Kapton (human hair = 10–80 microns)
- ~1 meter long straws, 5 mm diameter with a central gold-plated tungsten wire
- ~20,000 straws in a vacuum of  $10^{-4}$  Torr





# Calorimeter

- Two CsI crystal disks
  - hole, as in tracker, for passage of muon beam
  - distance chosen such that if signal electron goes down hole, hits next disk



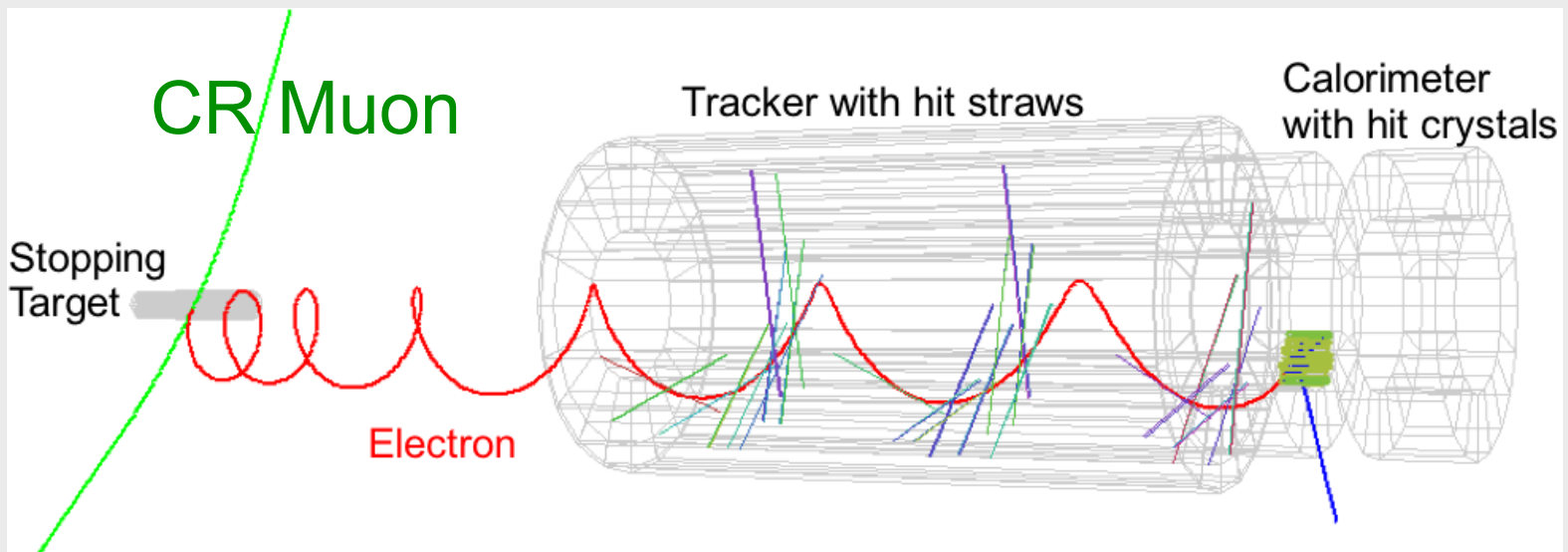
“module 0”:  
51 crystals

# Backgrounds and Design

	<i>source</i>	<i>scales with</i>	<i>solution</i>
Intrinsic Backgrounds	decay-in-orbit	# of muons	detector resolution
Beam	radiative pion capture	closeness to beam pulse	pulsed beam
Running Time	cosmic ray	live time	veto system

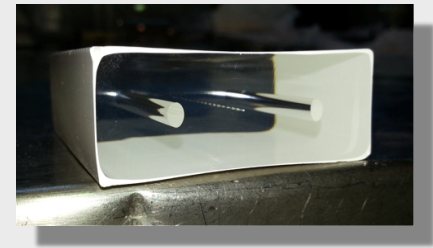
# Cosmic Ray Background

- At  $10^{-17}$  there's a lot of rare backgrounds; here's one that is surprising but not too rare: 1/day!

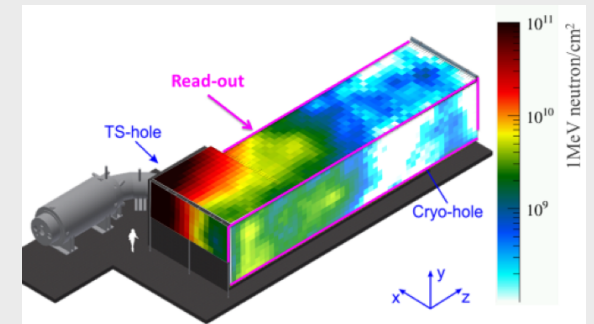


- the experiment needs a 99.99% efficient veto (3/4 layers)
- in an environment with  $10^{10}$  neutrons/cm<sup>2</sup>/sec from proton pulse

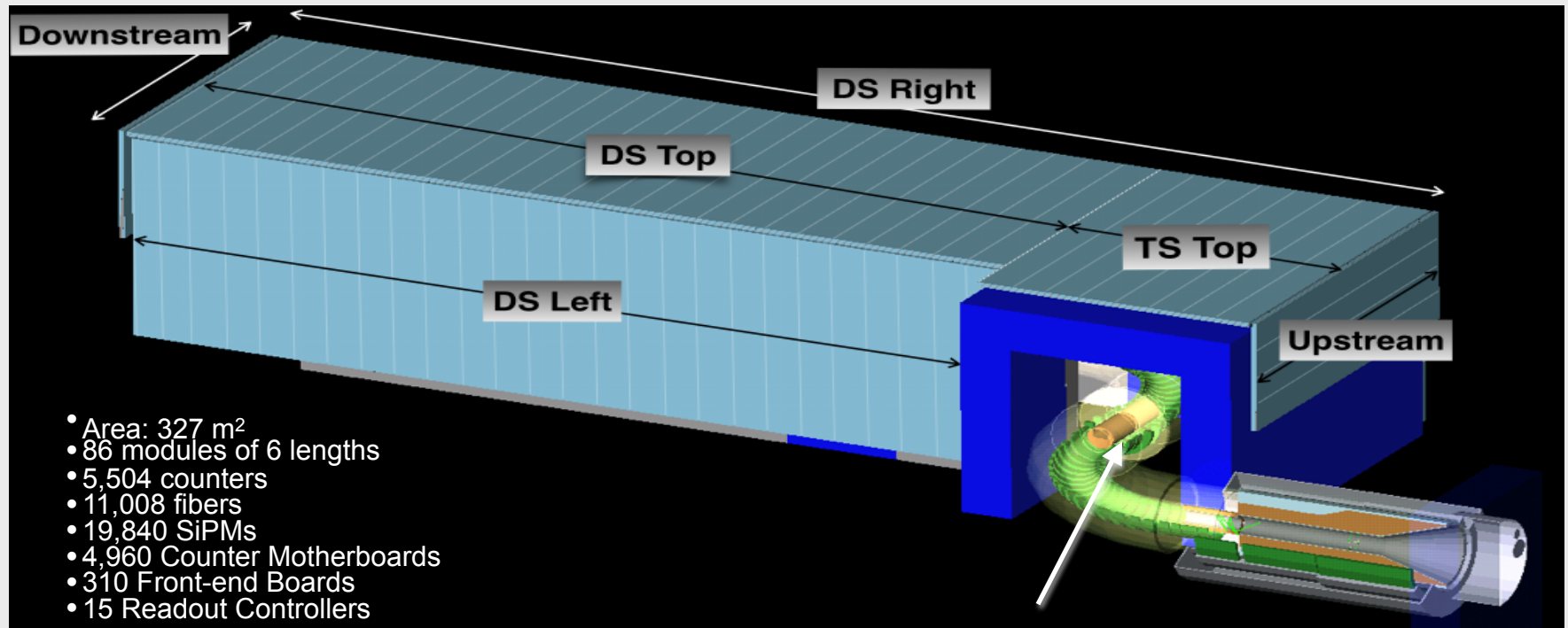
# CR Veto Details



Neutron Rates



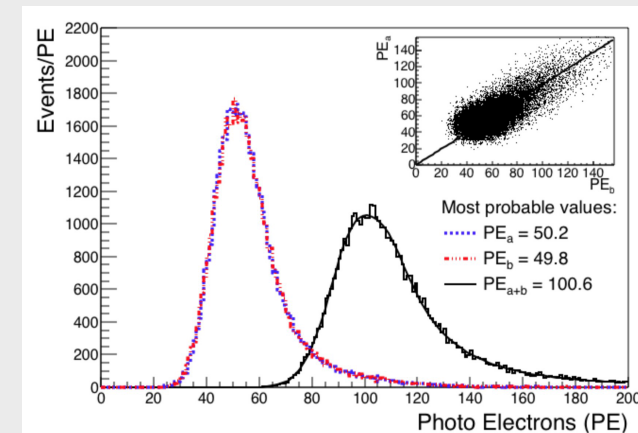
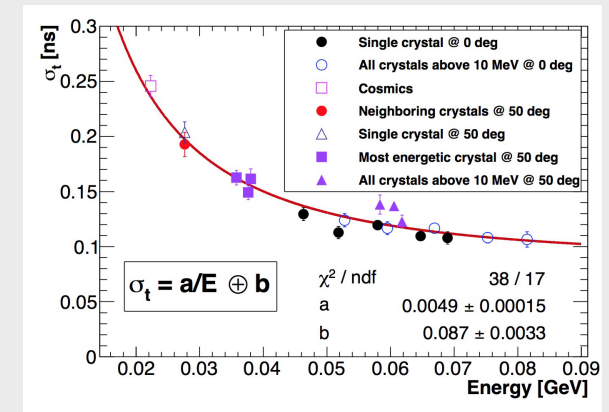
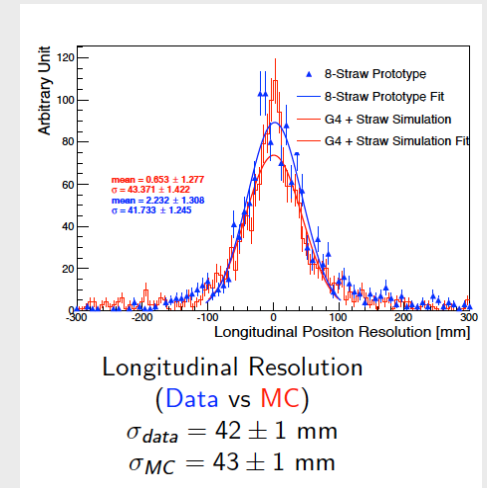
- Veto systems are never totally hermetic
  - even with 26 km of veto counters, area of a large house (~3000 ft<sup>2</sup>) (extrusions completed)
- Simulated 250 experiment live-times in targeted regions



targeted simulations here, for example

# Tracker/Calorimeter/ CRV Test Beam Results

- Energy and time response of 3X3 CsI crystal array
- PE yields of CRV scintillation counters read out by SiPMs
- Tracker Resolution measured and modeled at level of charge cluster formation
- *Measured resolutions and properties in the simulations*



arXiv: 1709.06587, submitted to NIM

DOI: 10.1088/1748-0221/12/05/P05007

# Beam Under Construction

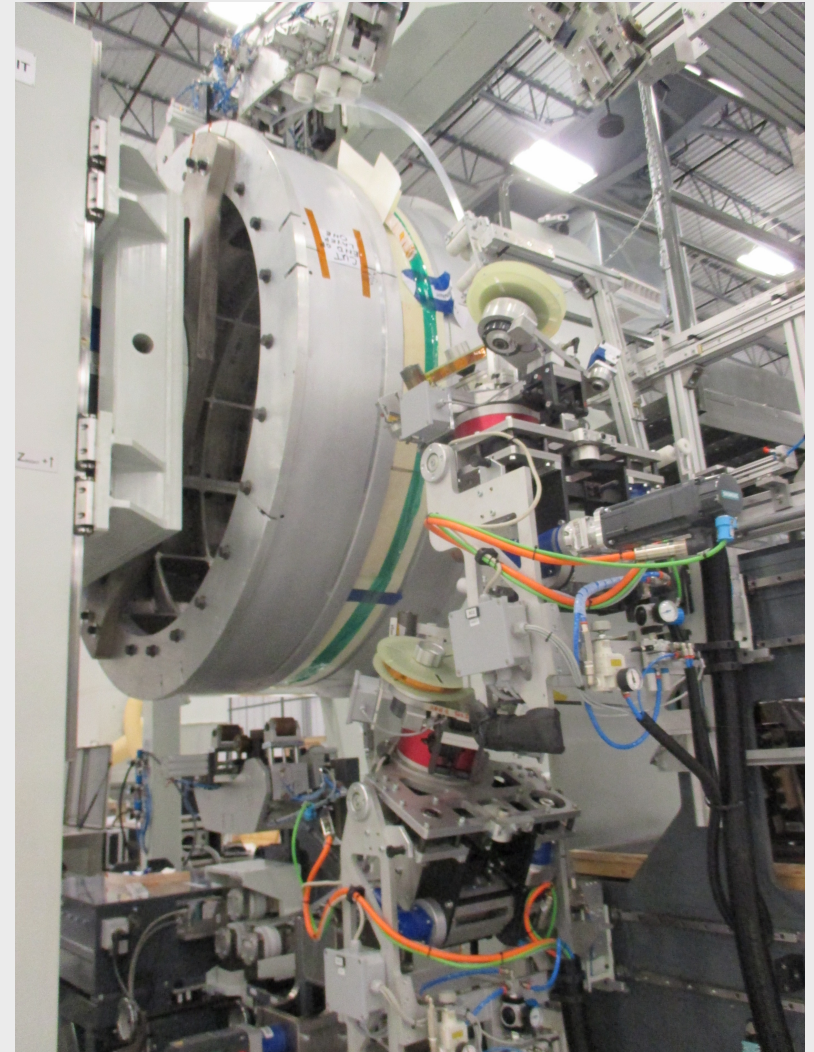
diagnostic absorber near g-2/Mu2e split:  
is the beam as expected and does the  
extinction work?



checking polarity now!  
(before we turn on)

# Mu2e Solenoid

Coil modules at ASG



~1 of 75  
km of  
cable and  
spares  
(cable  
complete)

at General Atomics  
splice between solenoids

# Beginning Production

- Ramping up production and cold tests

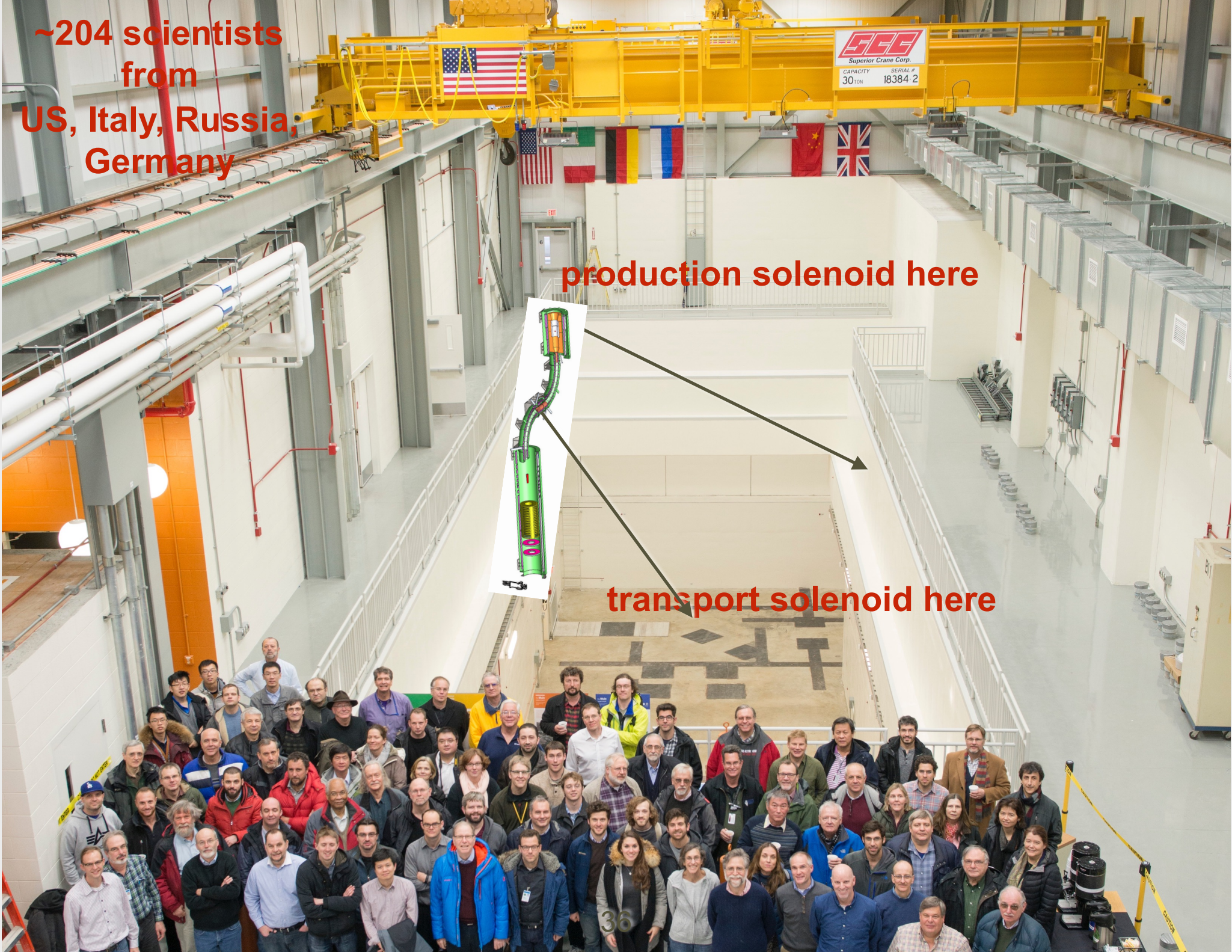




~204 scientists  
from  
US, Italy, Russia,  
Germany

production solenoid here

transport solenoid here

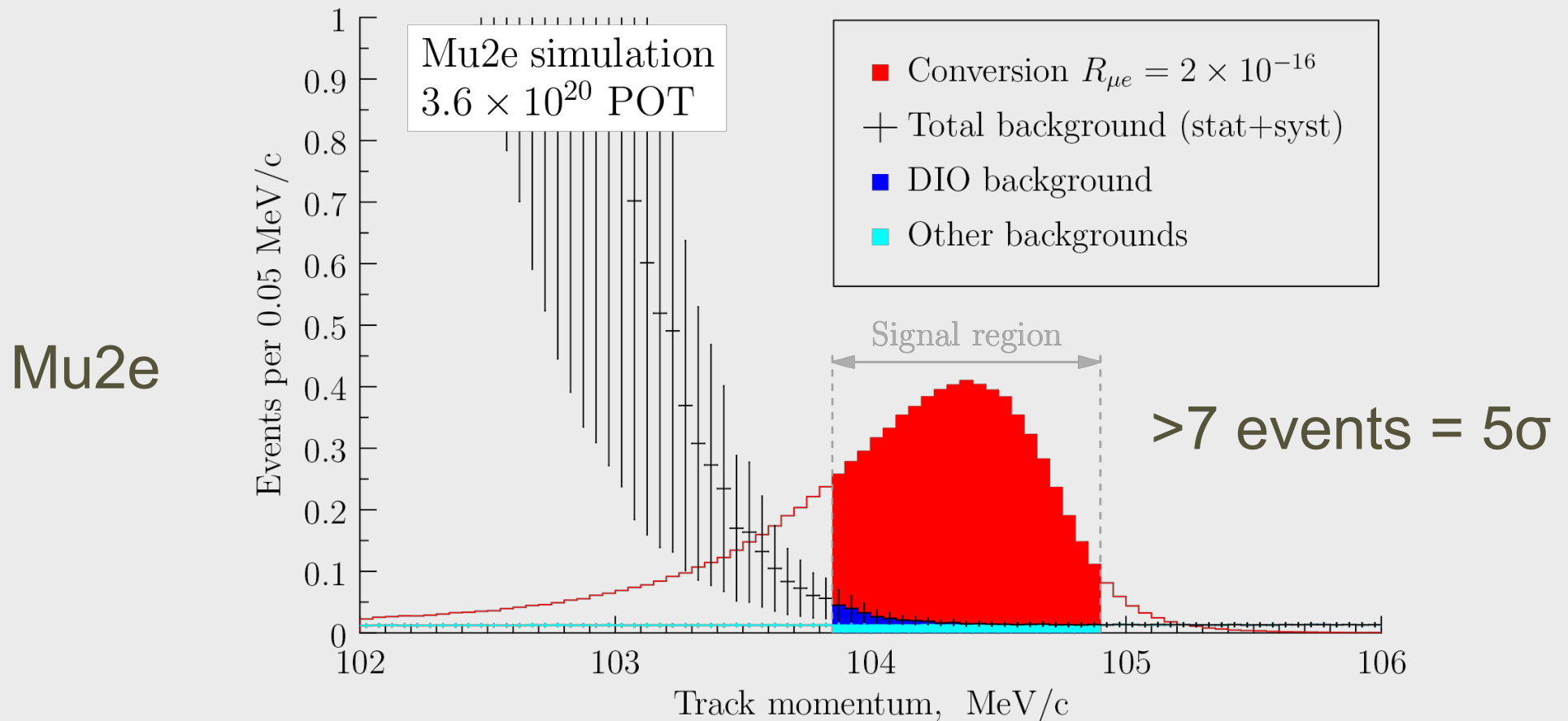


# Mu2e Expected

90% CL  $\sim 8 \times 10^{-17}$

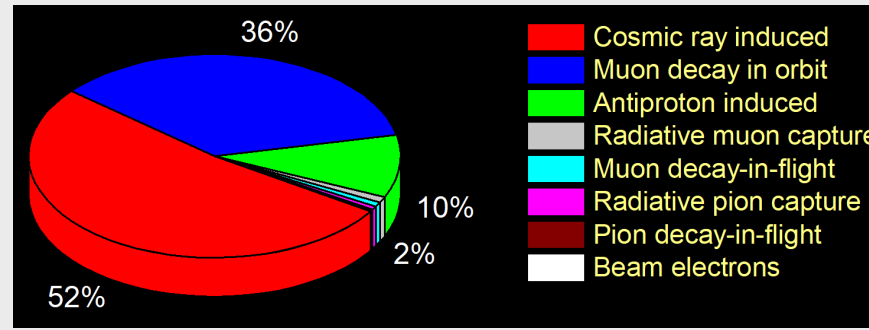
$5\sigma \sim 2 \times 10^{-16}$

x10000 better than SINDRUM-II



**typical SUSY at  $10^{-15}$ : 40 events vs 0.4 bkg**

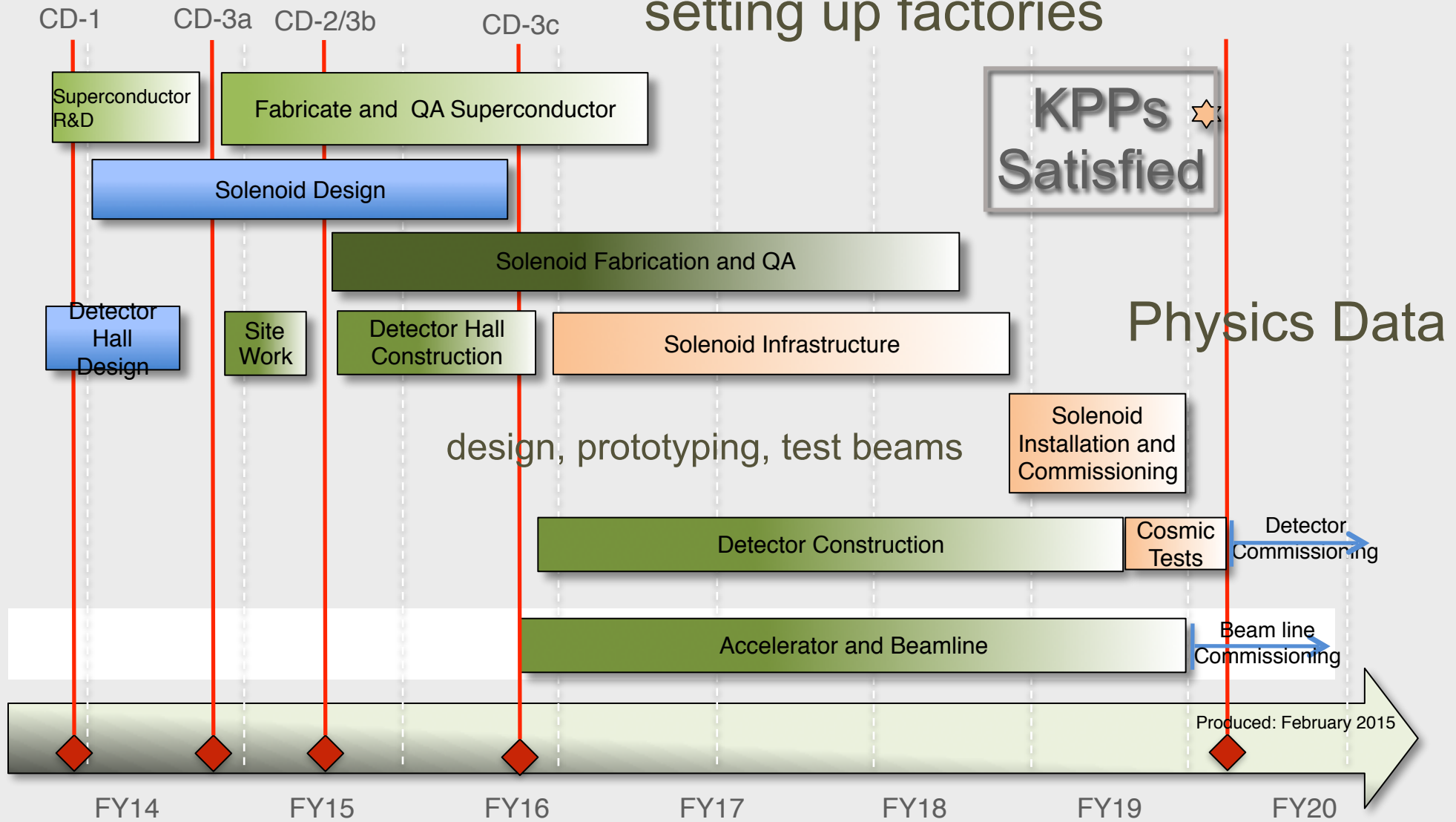
# Background Estimates



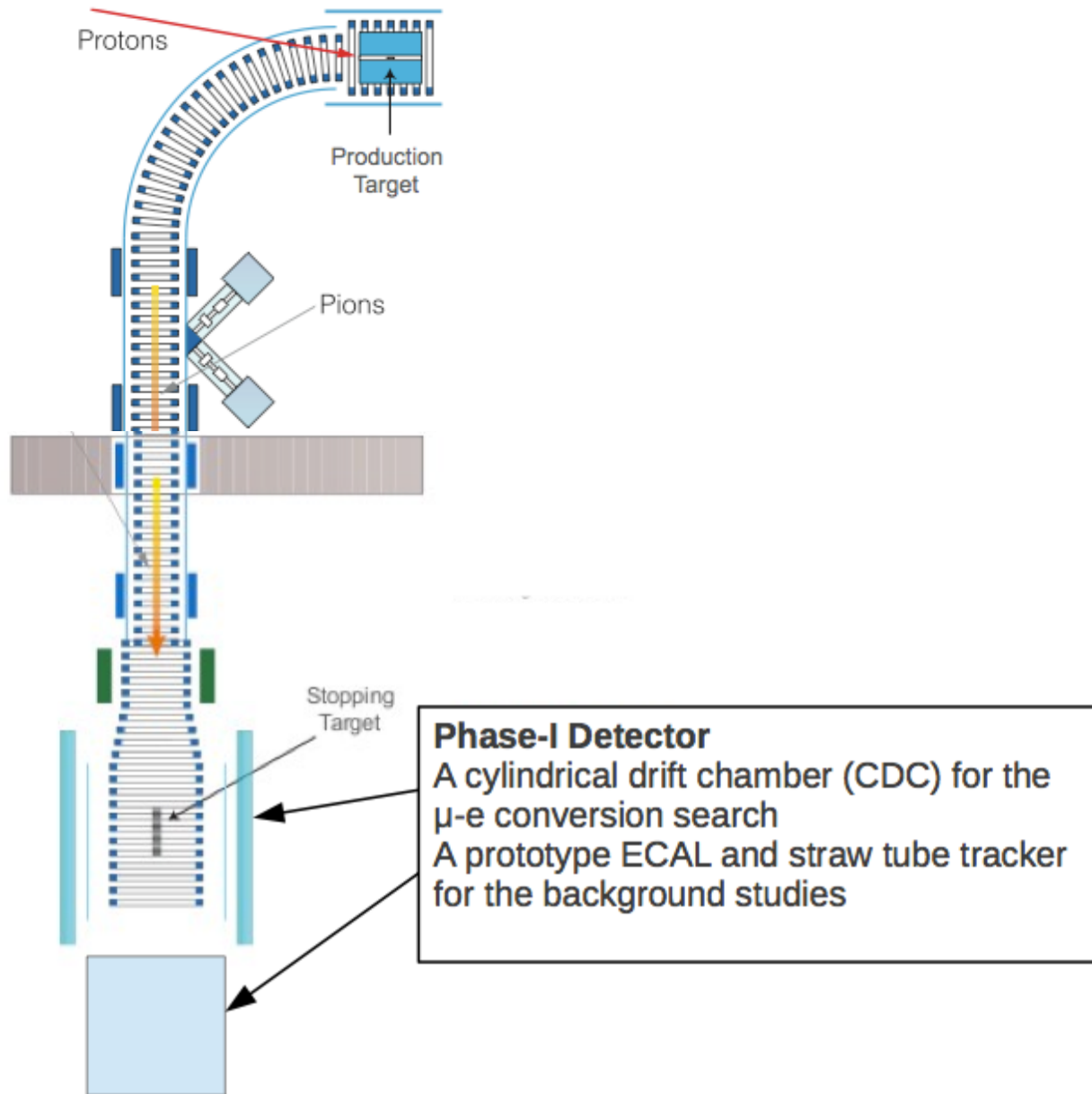
Process	Expected Number
Cosmic Ray Muons	$0.25 \pm 0.026$
DIO	$0.14 \pm 0.09$
RPC	$0.025 \pm 0.003$
Antiprotons	$0.047 \pm 0.024$
Muon DIF	$< 0.003$
Pion DIF	$0.001 \pm < 0.001$
Beam Electrons	$< 5 \times 10^{-4}$
Total	$0.46 \pm 0.03$

# Mu2e Schedule

Construction underway now;  
setting up factories

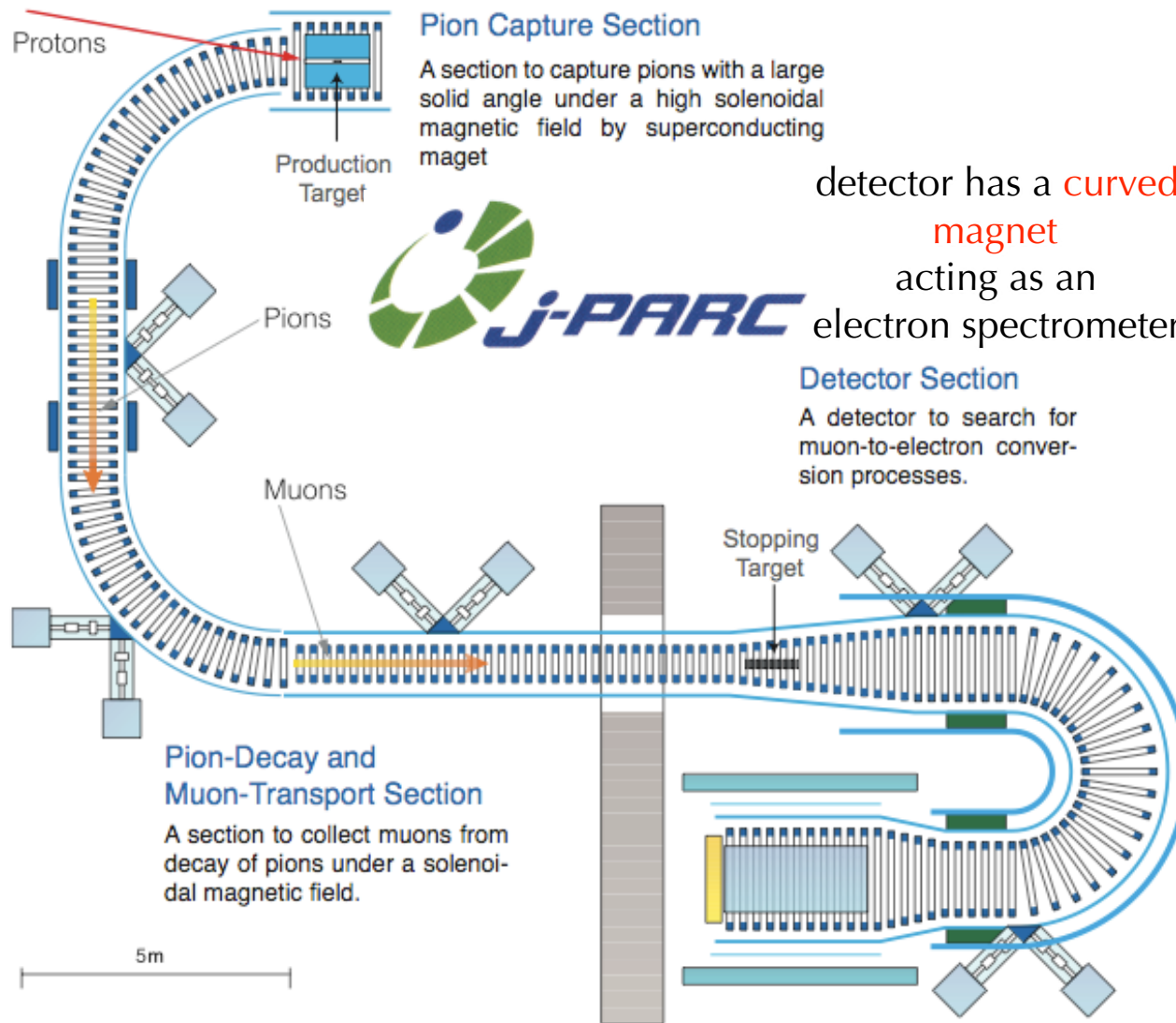


# COMET



- Stage I: (2021)
  - x100 better than SINDRUM-II
- Stage II
  - x10,000 better
  - same reach as Mu2e
- Important differences:
  - C- vs S- for transport
  - 2nd bend in detector solenoid

# COMET

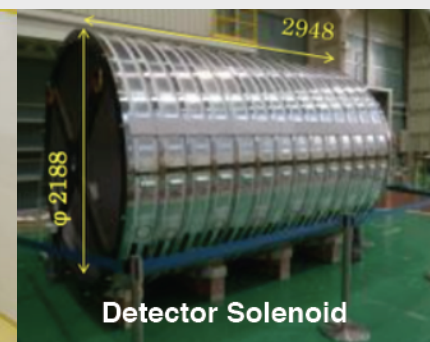
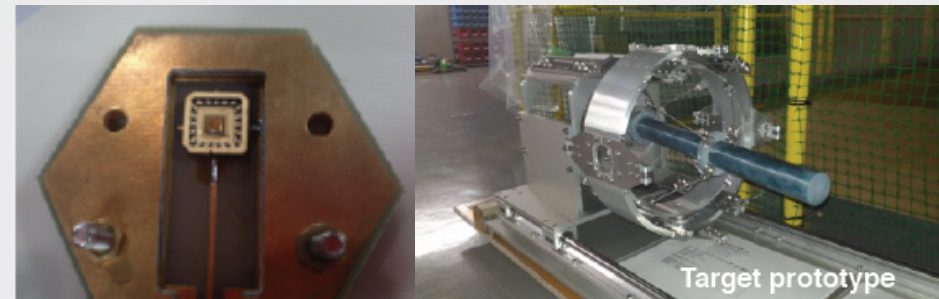
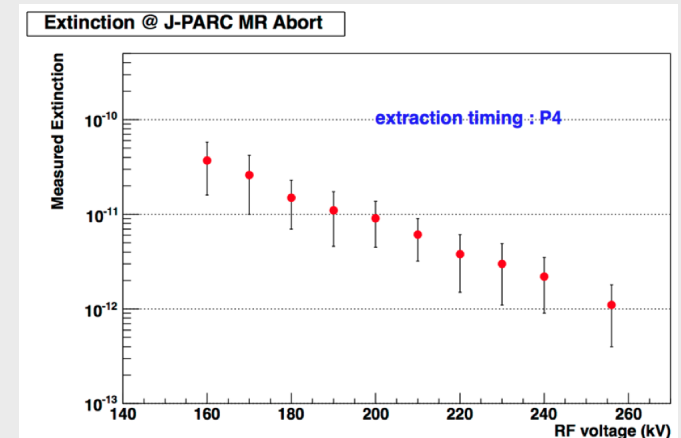


- Stage I: (2021)
  - x100 better than SINDRUM-II
- Stage II
  - x10,000 better
  - same reach as Mu2e
- Important differences:
  - C- vs S- for transport
  - 2nd bend in detector solenoid

# COMET Progress

Chen Wu, Tau 2016  
and  
Y. Kuno, priv. comm

- Extinction measured  $\sim 10^{-11}$
- Prototype Target
- Coil winding for capture solenoid
- Transport Solenoid in Hall
- 14 Coils of Detector Solenoid wound



# COMET Remaining

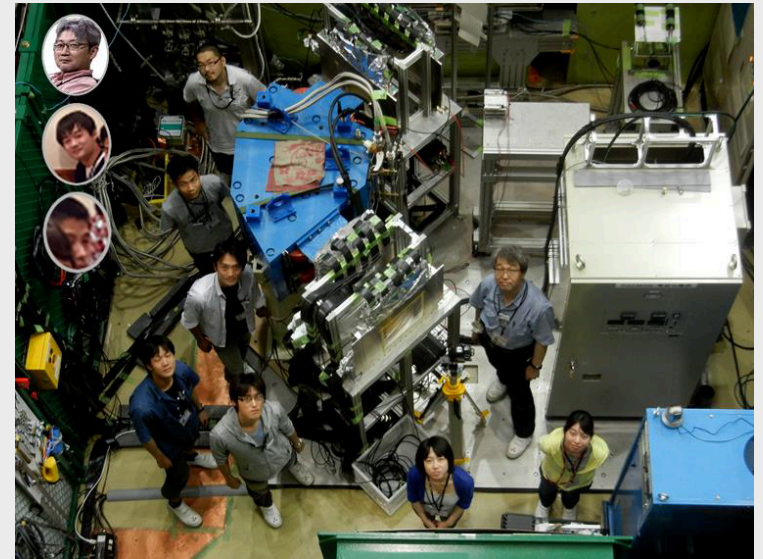
- Production Solenoid cable procured, but one has to then design a solenoid shielding system that will withstand heat and radiation
  - dumping up to 56 kW (Stage II) in a superconducting solenoid
  - design still underway
  - then needs construction
- Negotiation over running time with T2K
  - ~150 days live-time needed for Stage 1 (x100 improvement)
- Available schedules out-of-date, but Stage 1 ~ 2021

[http://j-parc.jp/researcher/Hadron/en/pac\\_1707/PAC24thMinutes\\_final\\_draft.pdf](http://j-parc.jp/researcher/Hadron/en/pac_1707/PAC24thMinutes_final_draft.pdf)

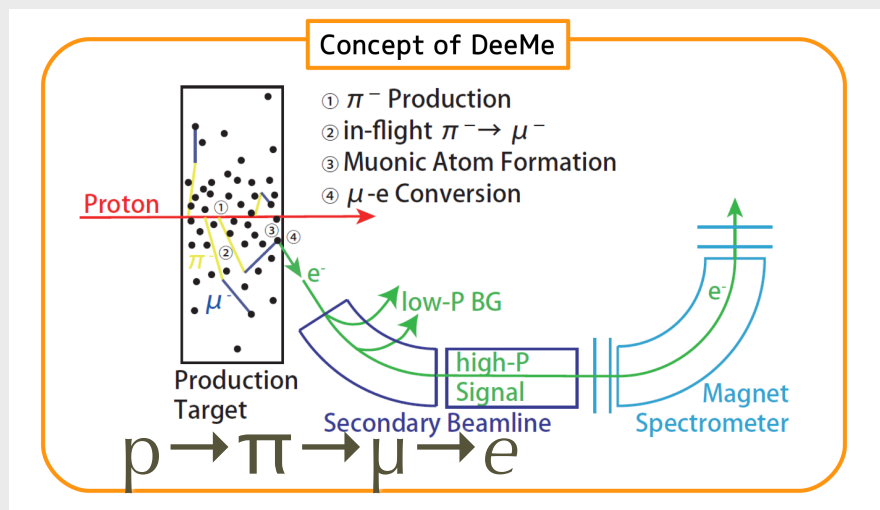
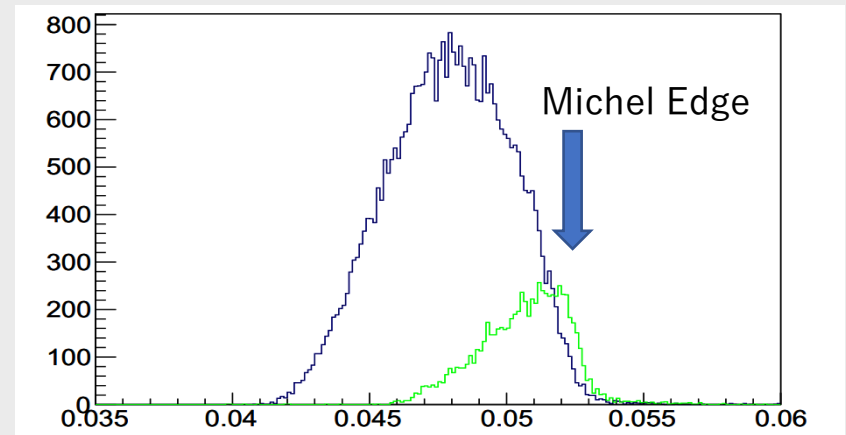


# DeeMe

- DeeMe at J-PARC  $\mu N \rightarrow e N$  with a  $2 \times 10^{-14}$  SES, x10 improvement
- **production target and conversion target are the same: not like Mu2e or COMET**



Calibration Spectra ( $\mu^+$  data)



# DeeMe HV Switching

- Since the detector is so “close” to the production target, the detector has to be turned off near the beam flash
- HV switching on a MWPC

## HV-Switching MWPC

- Expected detector rate
  - Prompt burst: 70 GHz/mm<sup>2</sup>
  - Delayed e<sup>-</sup> signal or μ-e conversion
- Turn-off only during the prompt burst
  - Development has completed.
  - Four MWPCs already fabricated.
  - All read-out circuits including amp., digitizer are ready.

**PTEP** Prog. Theor. Exp. Phys. 2017, 023C01 (11 pages)  
DOI: 10.1093/ptep/ptw195

**A fast high-voltage switching multiwire proportional chamber**

I. Natori<sup>1,2</sup>, N. Teshima<sup>3</sup>, M. Aoki<sup>4</sup>, H. Nishiguchi<sup>1</sup>, T. D. Nguyen<sup>5</sup>, Y. Takezaki<sup>2</sup>, Y. Furuya<sup>6</sup>, S. Ito<sup>2,4</sup>, S. Mihara<sup>1</sup>, D. Nagai<sup>3</sup>, Y. Nakatsugawa<sup>4</sup>, T. M. Nguyen<sup>5</sup>, Y. Senyo<sup>2</sup>, K. Shimizu<sup>2</sup>, and K. Yamamoto<sup>2</sup>

<sup>1</sup>High Energy Accelerator Research Organization (KEK), Beam-3 305-0861, Japan  
<sup>2</sup>Osaka City University, Gushu 328-8583, Japan  
<sup>3</sup>Osaka University, Osaka 565-0848, Japan  
<sup>4</sup>Present Address: Institute for Basic Science (IBS), Daejeon 30537, Republic of Korea  
<sup>5</sup>Present Address: Obayashi University, Okazaki 464-8556, Japan  
<sup>6</sup>Present Address: Institute of High Energy Physics (IHEP), Beijing 100049, China  
\*E-mail: natori@hep.kek.jp

Received October 21, 2016; Revised December 11, 2016; Accepted December 26, 2016; Published February 14, 2017

Next Talk by N. Teshima

# DeeMe Status

- Depends on J-PARC H-line construction (different from COMET, DeeMe in Materials Science section)

## Summary

- There is a competitive merit of physics in searching for  $\mu$ -e conversion at sensitivity of  $10^{-14}$  in timely manner.
- **It is important to maximize the potential of major discovery at J-PARC.**
- DeeMe, yet another mu-e conversion search with totally different method from COMET and Mu2e, creates harmonious diversity for J-PARC.
- DeeMe has already acquired **Stage-2 Approval from muon-PAC of KEK/IMSS.**
- Construction of **detector system has completed** with Grant-in-Aid for Scientific Research of Japan (Basic Science S, 2012–2016).
- An **integration test of the whole detector system** including DAQ had **successfully** performed. A **DIO measurement** had done at the same time, and analysis is ongoing.
- The **Detector system is ready** for the physics data taking.
- We are hoping to start soon after the completion of the beamline construction. No beam-time conflicts with T2K, KOTO or whatever the physics programs with the main ring of J-PARC.
- Meanwhile, we are keeping collaboration activities with detector improvements and by-product research.

# Mu2e and COMET: High-Level Differences

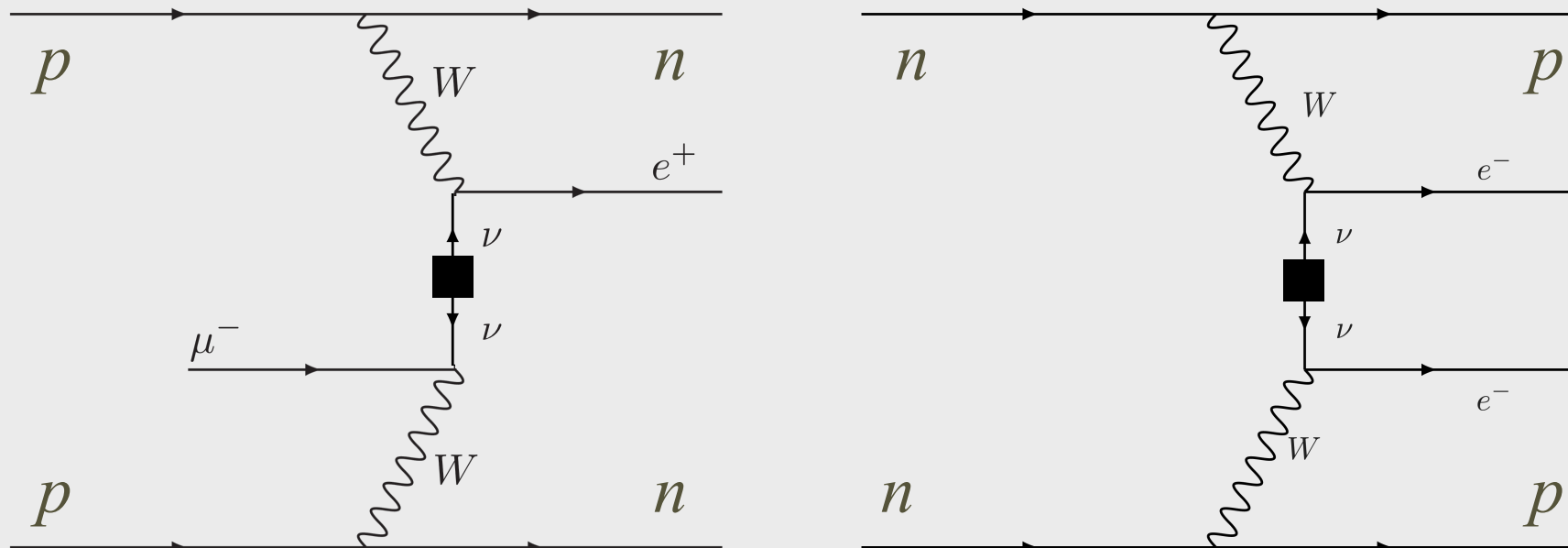
- COMET Staging has obvious advantages for learning as you go
- Charge Symmetry of Detector:
  - Mu2e: charge symmetric,  $e^+e^-$  the same
  - 2nd Bend in COMET Stage II momentum selects  $\sim 105$  MeV  $e^-$  only
  - Mu2e needs hole, COMET does not

# Effect of Charge-Symmetry Choice

- Backgrounds
  - many background sources of  $e^-$  (RPC!) produce equal  $e^+e^-$
  - Mu2e can measure them with  $e^+$  *in situ*
  - *look “inside the box”, no blinding needed since  $e^+$*
- Physics (next topic)
  - Mu2e can simultaneously measure  $\mu^-N \rightarrow e^+N'$ , different/unique physics
  - COMET would have to reverse field, not trivial: separate runs

$$\mu^- N \rightarrow e^+ N^*$$

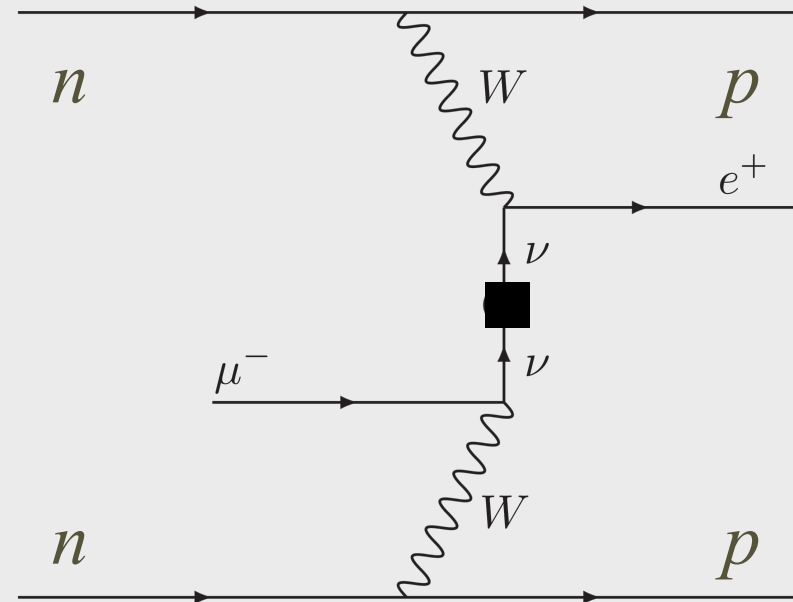
- Closely related to  $\theta\nu 2\beta$ 
  - or a leptoquark (not shown), ...



- <https://doi.org/10.1103/PhysRevD.95.115010>, <https://arxiv.org/pdf/1611.00032.pdf>, <https://arxiv.org/pdf/1705.07464.pdf>, Geib et al., Physics Letters B 764 (2017) 157–162 or <https://arxiv.org/abs/1609.09088> discusses relation to  $\theta\nu 2\beta$

# $\mu^- N \rightarrow e^+ N^*$ Changes Nucleus

- The final state nucleus is different by  $\Delta Z = 2$ 
  - different from muon-to-electron conversion, which is coherent and gets a  $Z^2$  enhancement



# One-Slide Summary

- Large Background from a physics process with the same  $\mu^-$  we use for  $\mu^- N \rightarrow e^- N$ , called radiative muon capture
  - Probably need to look for GS $\rightarrow$ GS Transitions between initial, final state nuclei. Numerically, transitions to ground state not well known
  - Probably Al is not the best, need higher  $Z$  material like Ti, but that is natural direction for Mu2e anyway
- These are not “show-stoppers”. Work underway.





## GS $\rightarrow$ GS in Mu2e

- mono-energetic positron at known energy If  $B$  is binding energy,  $R$  recoil (# for  $^{27}\text{Al}$  to  $^{27}\text{Na}$ )

$$\begin{aligned} E_{\mu^- e^+} &= m_\mu + M(A, Z) - [M(A, Z - 2) + 2m_e - B_\mu - R(A)] \\ &= 92.31 \text{ MeV} \end{aligned}$$

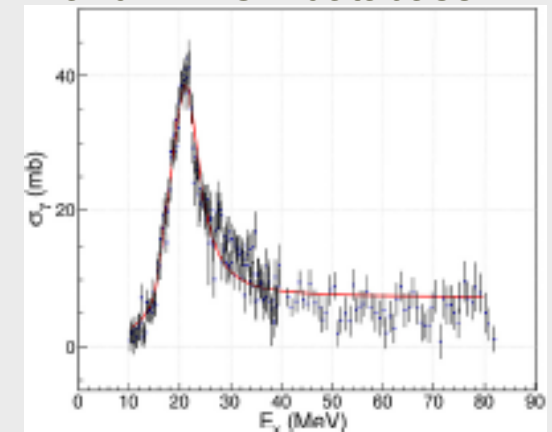
- Mu2e acceptance at 92 MeV/c is relatively low; may need to turn down field, implying special run

# Modeling $\mu^- N \rightarrow e^+ N^*$

Geib et al.

Physics Letters B 764 (2017) 157–162  
and EXFOR database

- Usually imagine two possibilities:
  - transition to ground state
    - narrow
  - transition to giant dipole resonance (GDR)
    - very broad, 20 MeV  $\rightarrow$  hard to do
  - and then since fraction to each not well known, quote separate limits



# Calculating Signal: GS Transition Rate

- Ground State transitions are mono-energetic so can pick out positron.
- Last estimate:
  - P.C. Divari et al. / Nuclear Physics A 703 (2002) 409–431
  - Estimate 41% to GS; much larger than closure approximation estimates of GS rate

# Background Processes

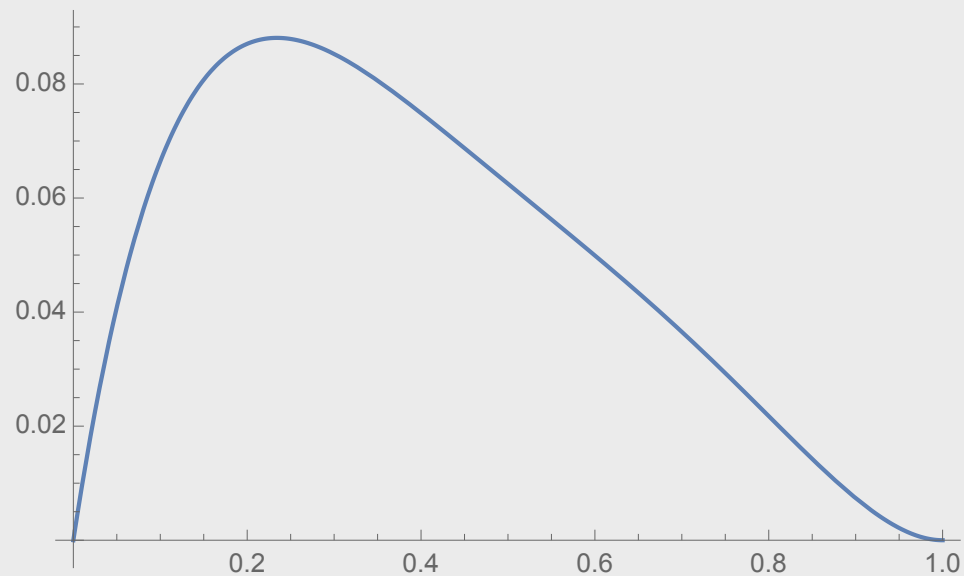
- Two that are relevant, and present different problems:
  - $\pi^- N \rightarrow \gamma N$  (Radiative Pion Capture = RPC)
  - $\mu^- N \rightarrow \gamma N$  (Radiative Muon Capture = RMC)
- The  $\gamma$  can convert either in the stopping target or undergo internal conversion (happen to be  $\sim$ the same in Mu2e and COMET by coincidence).
  - RPC suppressed already, but RMC can't be: from  $\mu^-$
  - background  $e^+$  for  $\mu^- N \rightarrow e^+ N^*$

# Calculating RMC Background: Closure Approximation

- Model transitions from an initial state to a single final state with a mean energy  $\tilde{E}$  which is now a parameter
- So we also have a parameter  $k_{max}$  in the closure approximation
  - the biggest it could be is the kinematic limit from mass, recoil, and binding energy  
 $\sim 101.865 \text{ MeV}/c^2$  in Al

# RMC Spectrum

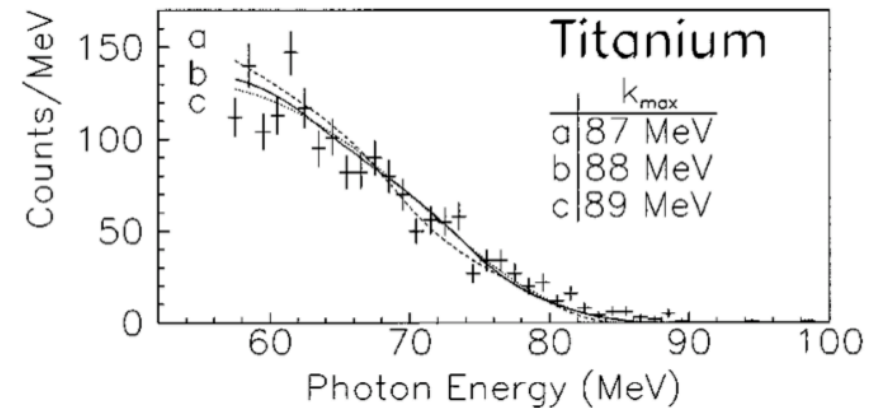
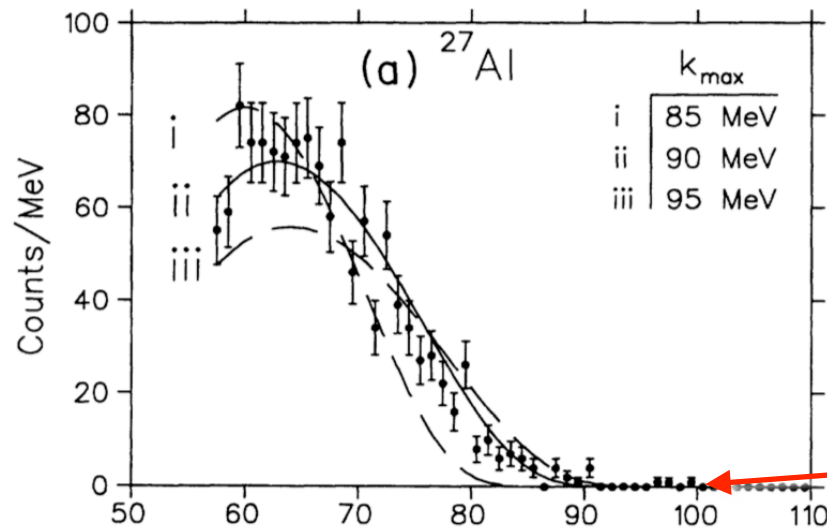
- Theory: use the closure approximation to derive  $N(x) \propto (1 - 2x + 2x^2) \times (1-x)^2$  where  $x = k/k_{max}$



- Data?

P. Christillin, Ref Th 2697, CERN

# RMC Data



- How reliable is the closure approximation?
- Is there a tail?
- Can we minimize these uncertainties?
  - have signal region near endpoint of RMC spectrum

# RMC vs Signal

- From Yeo et al. <https://arxiv.org/pdf/1705.07464.pdf>

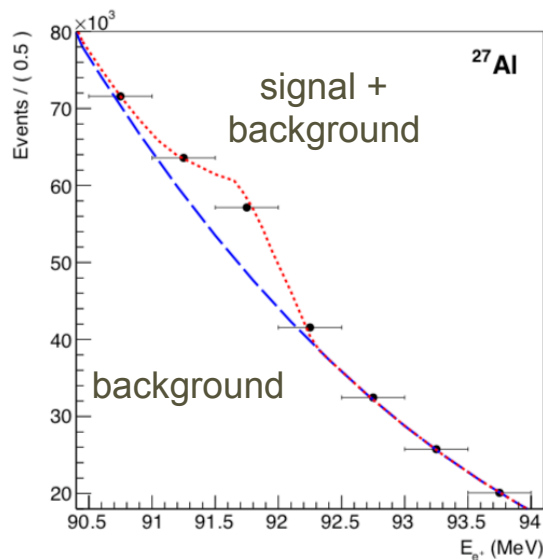
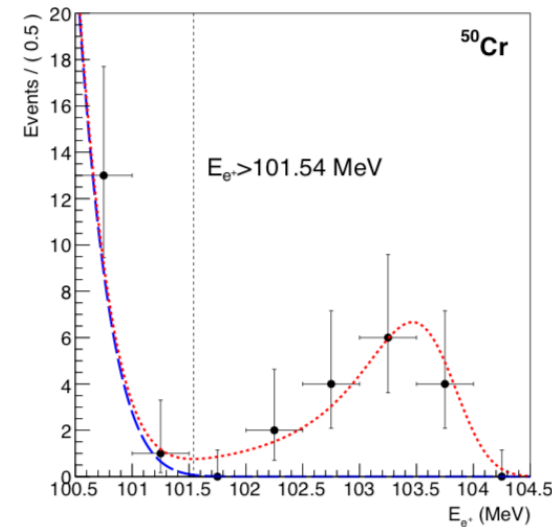
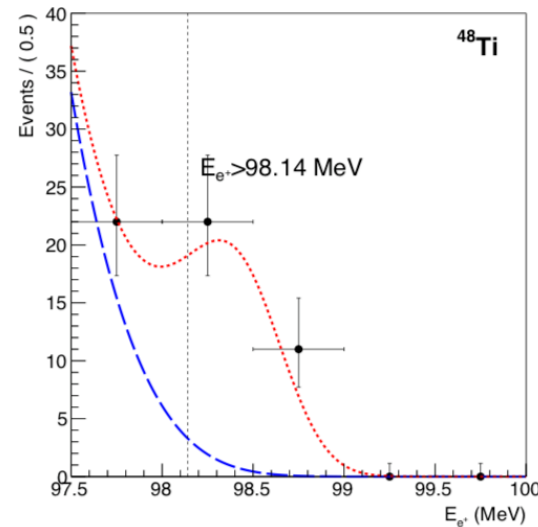


FIG. 1. Fitting result of the energy distribution of the  $\mu^- - e^+$  signal (short dashed red line) stacked on the on-shell RMC photon background (long dashed blue line) from  $^{27}\text{Al}$  muon-stopping target when  $Br(\mu^- - e^+) = 1.7 \times 10^{-12}$  and  $N_{\mu\text{-stop}} = 10^{18}$ . Black dots are pseudo data of positrons generated by the background and signal composite model.



kinematic  $k_{max}$  ;

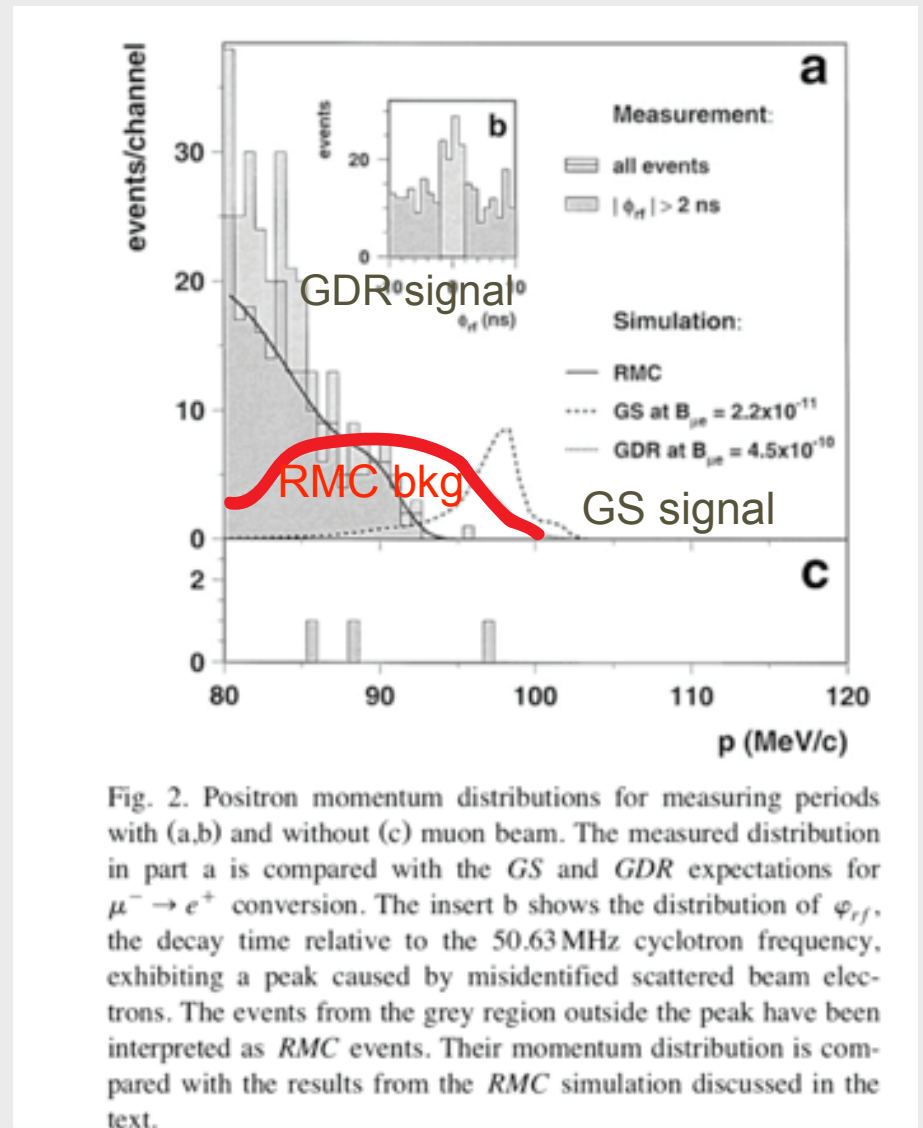
just to indicate how much better  
you do at higher Z

Mu2e is working on making these plots in our simulations



# $\mu^- N \rightarrow e^+ N^*$ on Ti: Real Data

- SINDRUM actually used much older data on Ca to transfer to Ti
- Can't do better without measuring RMC and RPC
- Mu2e Simulations underway



# Current Limits

- From SINDRUM-II, to ground state and GDR at 90% CL
  - other materials, just quoting Ti

$\mu^- + Ti \rightarrow e^+ Ca$	BR < $1.7 \times 10^{-12}$
$\mu^- + Ti \rightarrow e^+ Ca^*$	BR < $3.6 \times 10^{-11}$

# Which Material?

- Don't want  $k_{max, kinematic} > E_{e^+(GS)}$  to the extent possible

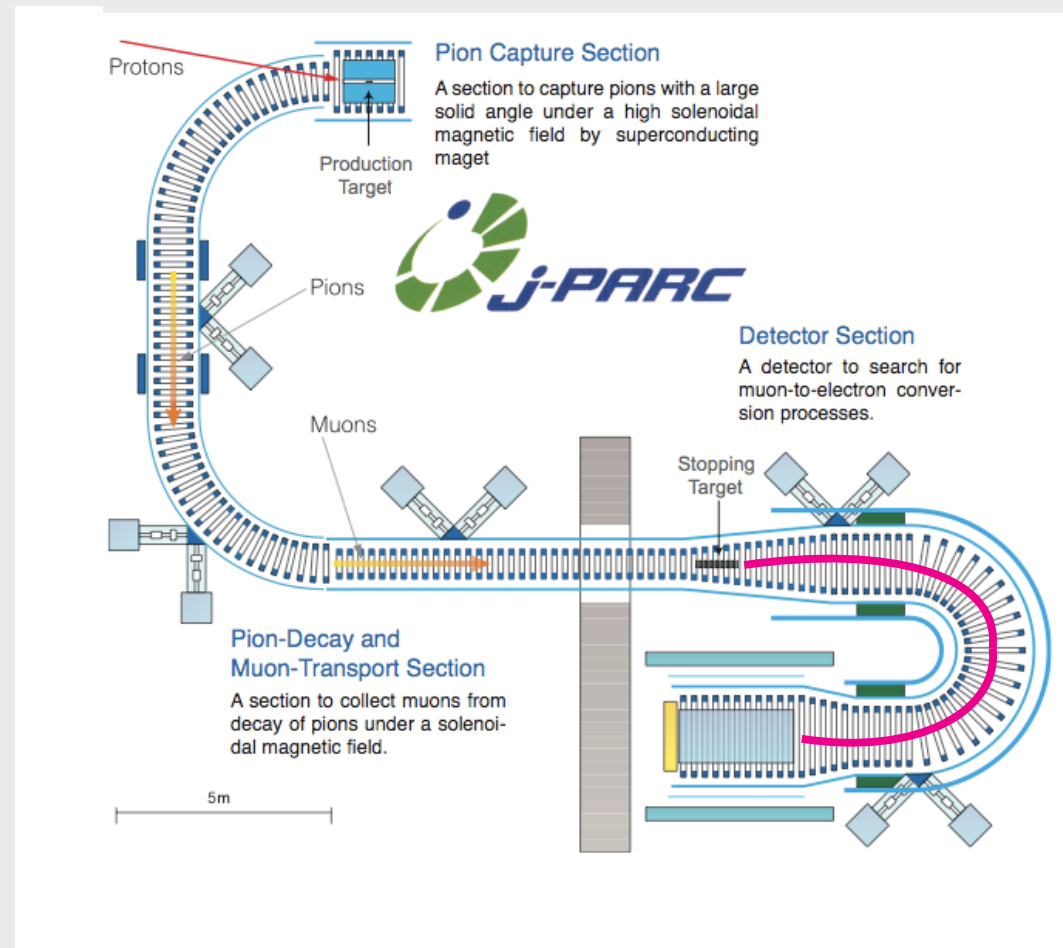
	rate > 57 MeV/ OMC	$k_{max}$	$k_{max, kinematic}$	conversion energy
$^{27}\text{Al}$	$1.43 \pm 0.13 \times 10^{-5}$	$90.0 \pm 2.0$	101.9	92.4
$^{28}\text{Si}$	$1.31 \pm 0.15 \times 10^{-5}$	$89.4 \pm 1.8$	99.8	97.4
$^{48}\text{Ti}$	$1.31 \pm 0.15 \times 10^{-5}$	$89.2 \pm 2.0$	99.7	98.9

- Al not best choice? Ti (or Si with active target) might be much better; Ti also good for  $\mu^- N \rightarrow e^- N$  (later)

# COMET and $\mu^- N \rightarrow e^+ N^*$

- 105 MeV conversion electron from  $\mu^- N \rightarrow e^- N$
- $e^+$  follows opposite path
- have to reverse dipole portion of field in second section, so separate run is required

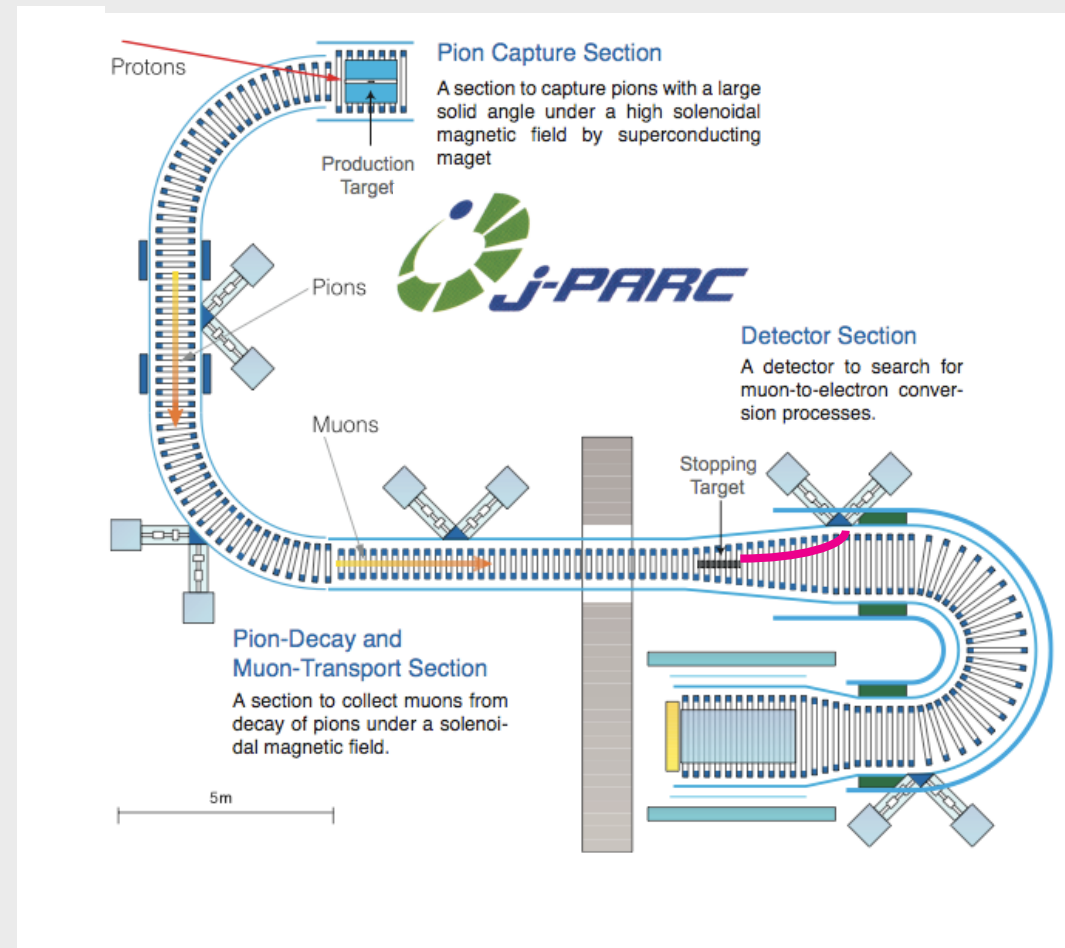
perhaps OK in Stage 1 where detector charge-symmetric



# COMET and $\mu^- N \rightarrow e^+ N^*$

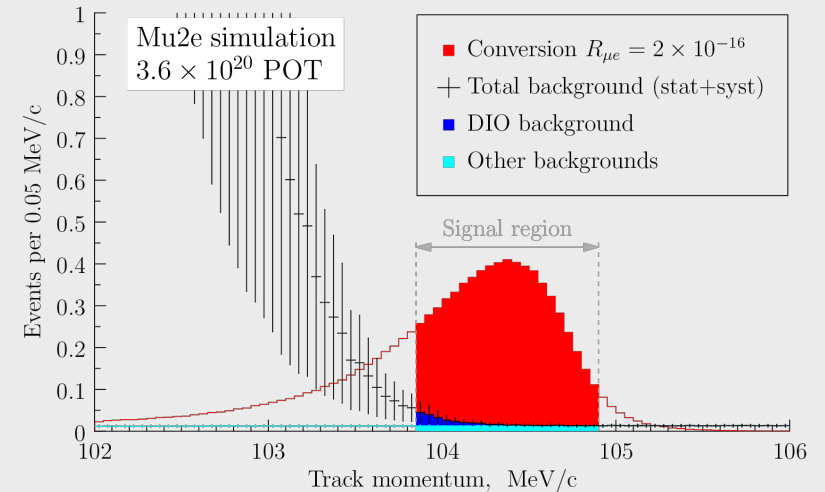
- 105 MeV conversion electron from  $\mu^- N \rightarrow e^- N$
- $e^+$  follows opposite path
- have to reverse dipole portion of field in second section, so separate run is required

perhaps OK in Stage 1 where detector charge-symmetric



# Mu2e Summary

- $\mu^- Al \rightarrow e^- Al$ 
  - SES  $\sim 3 \times 10^{-17}$
  - 90% CL  $\sim 8 \times 10^{-17}$
  - $5\sigma$  discovery  $\sim 2 \times 10^{-16}$
  - Run for 3 years; physics data starting 2022
- Upgrades in intensity and changes in material already under examination

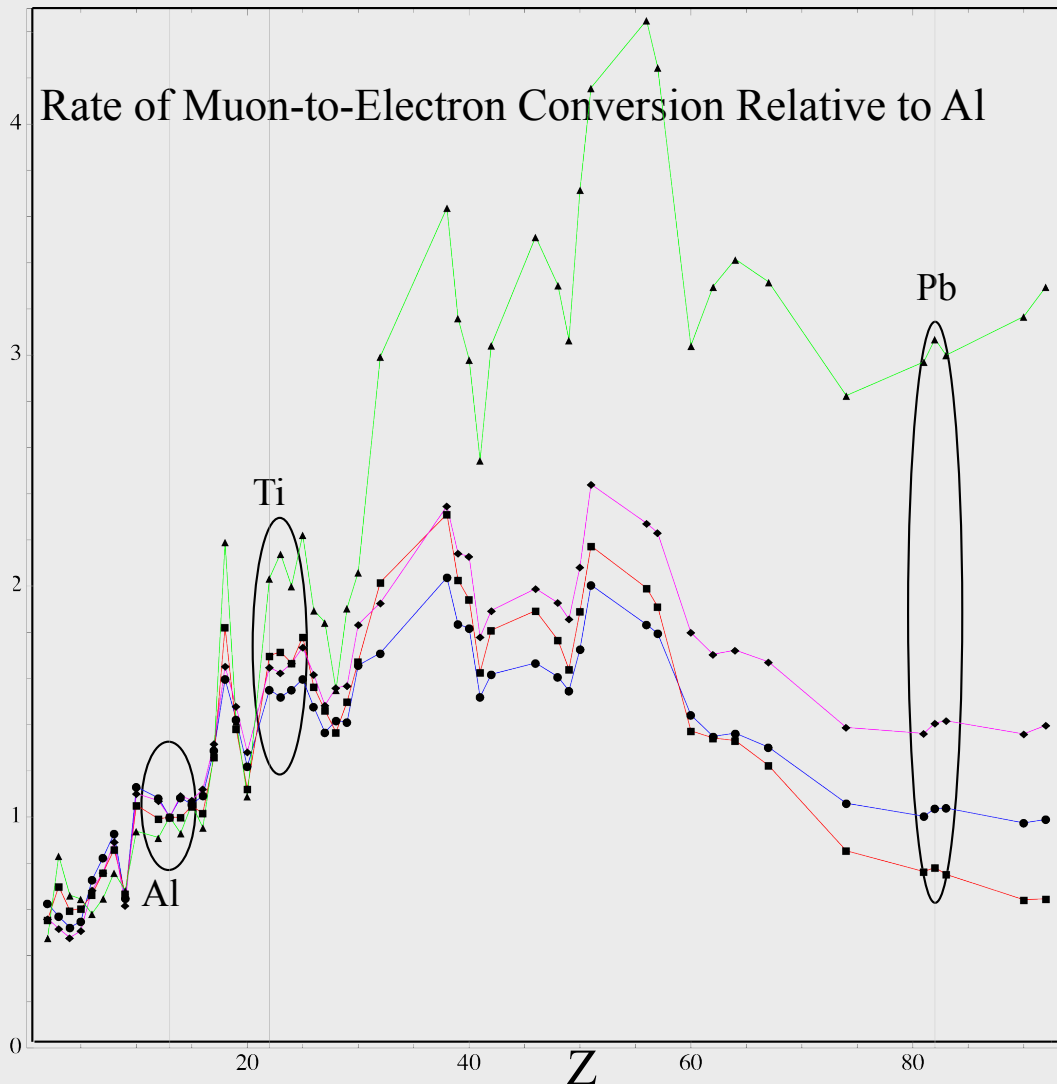


# Mu2e Upgrades

- Next Step in cLFV Program:



# Upgrades and Z-Dependence



- Different Operators have different Z-dependence
- Combine depending on the particular model

5% measurement on Al/Ti needed to see split

Lepton flavor violating mu - e conversion rate for various nuclei

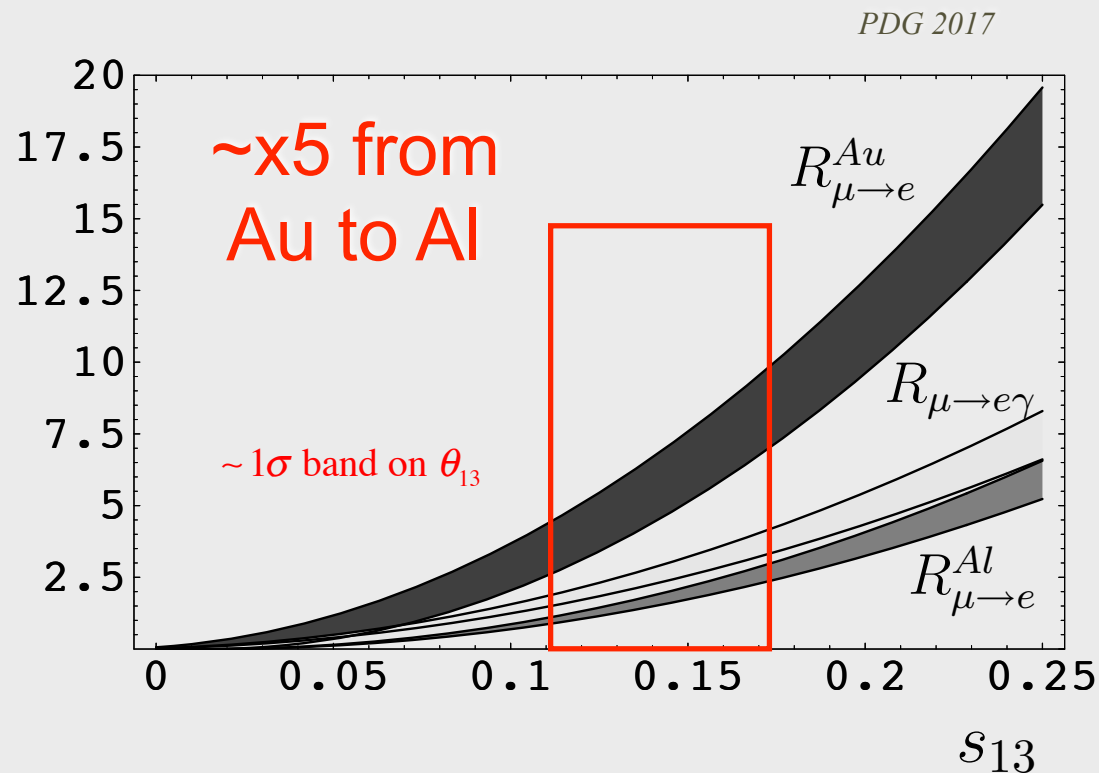
M. Koike et al., J.Phys. G29 (2003) 2051-2054

DOI: [10.1088/0954-3899/29/8/401](https://doi.org/10.1088/0954-3899/29/8/401)



# Relationship to Neutrinos (aka neutral leptons)

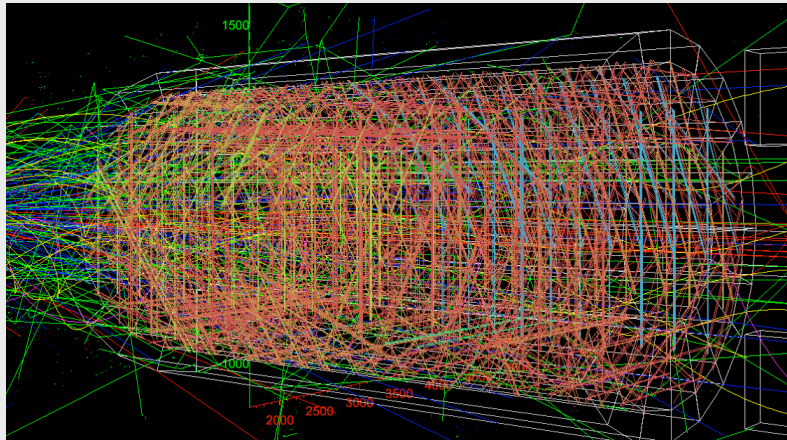
- want as high a  $Z$  as possible



V. Cirigliano, B. Grinstein, G. Isidori, M. Wise  
Nucl.Phys.B728:121-134,2005

# Choice of Z for Upgrade

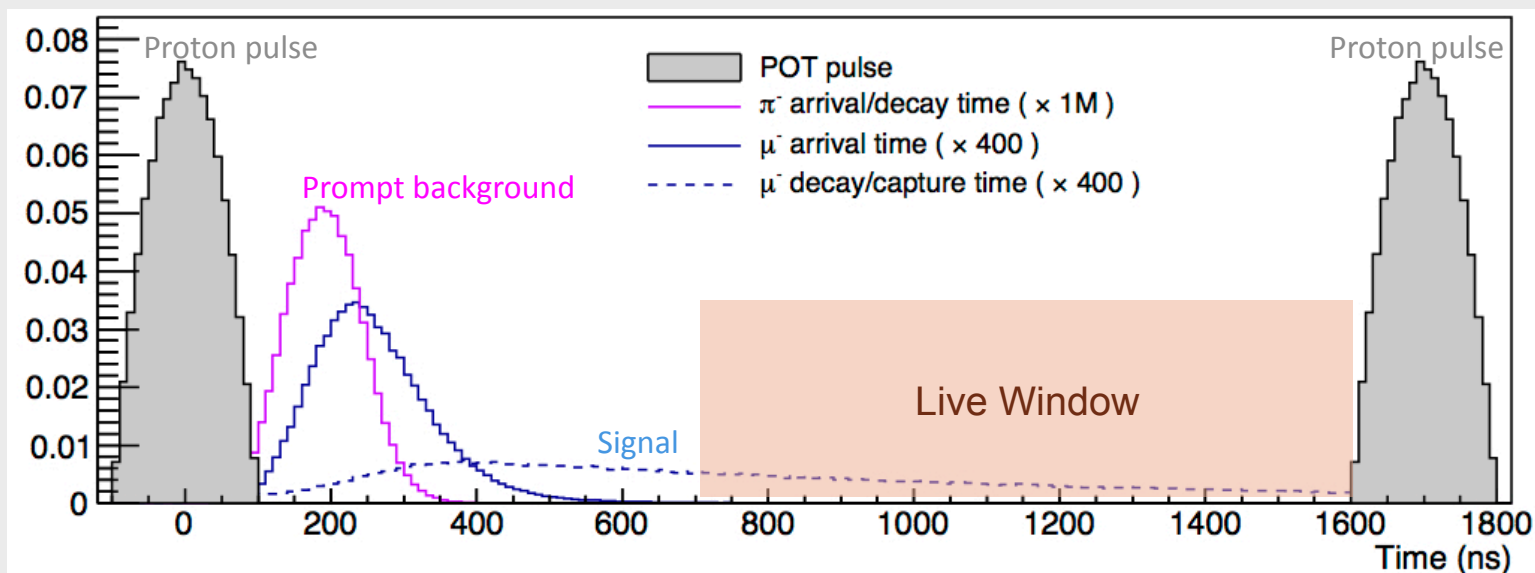
- What Sets Material Choice? Lifetime:



$$\tau_{\mu}(\text{Al}) = 864 \text{ ns}$$

$$\tau_{\mu}(\text{Ti}) = 338 \text{ ns}$$

$$\tau_{\mu}(\text{Au}) = 74 \text{ ns}$$



# Mu2e Upgrades

- Studies for x10 improvement with Ti look promising and will be continued; EOI written (1307.1168 and EOI at 1802.02599)
- We need detector and solenoid improvements
  - may need new production solenoid to handle lower energy beam and higher power.
- FNAL PIP-II natural for both pulsed and non-pulsed CLFV, could do  $\mu^- N \rightarrow e^\pm N$ ,  $\mu \rightarrow e \gamma$ ,  $\mu \rightarrow 3e$ ,  $\mu^- e^- \rightarrow e^- e^-$  at one facility

# Advertisement

- 3rd International Conference on Charged Lepton Flavor Violation
  - we cover everything in CLFV:  $\mu$ ,  $\tau$ ,  $B$ , talks split between theory and experiment
  - three days, plenary only (w/posters)
  - 17-19 June 2019 at Kyushu U.
- For examples:
  - <http://clfv2013.le.infn.it> (referenced by Emilie P.)
  - <https://indico.fnal.gov/event/11128/> (Julian H.)

# Conclusions

- Mu2e will:
  - Reduce the limit for  $R_{\mu e}$  by more than four orders of magnitude, x10 in mass reach ( $R_{\mu e} < 8 \times 10^{-17}$  @ 90% C.L.)
  - Discover unambiguous proof of new physics or
  - Set powerful constraints on a wide variety of models
- Mu2e will therefore both complement LHC results and independently probe up to  $10^4$  TeV/ $c^2$
- COMET and DeeMe in progress
- Upgrades already under consideration with a new muon CLFV program being mapped out

**QUESTIONS**  
FOUND IN GOOGLE AUTOCOMPLETE

**WHY DO WHALES JUMP**  
WHY ARE WITCHES GREEN  
WHY ARE THERE MIRRORS ABOVE BEDS  
**WHY DO I SAY UH**  
WHY IS SEA SALT BETTER  
WHY ARE THERE TREES IN THE MIDDLE OF FIELDS  
WHY IS THERE NOT A POKEMON MMO  
WHY IS THERE LAUGHING IN TV SHOWS  
WHY ARE THERE DOORS ON THE FREEWAY  
WHY ARE THERE SO MANY SVOHOSTEXE RUNNING  
WHY AREN'T THERE ANY COUNTRIES IN ANTARCTICA  
WHY ARE THERE SCARY SOUNDS IN MINECRAFT  
WHY IS THERE KICKING IN MY STOMACH  
WHY ARE THERE TWO SLASHES AFTER HTTP  
WHY ARE THERE CELEBRITIES  
**WHY DO SNAKES EXIST**  
WHY DO OYSTERS HAVE PEARLS  
WHY ARE DUCKS CALLED DUCKS  
WHY DO THEY CALL IT THE CLAP  
WHY ARE KYLE AND CARTMAN FRIENDS  
WHY IS THERE AN ARROW ON AANG'S HEAD  
WHY ARE TEXT MESSAGES BLUE  
WHY ARE THERE MUSTACHES ON CLOTHES  
WHY ARE THERE MUSTACHES ON CARS  
WHY ARE THERE MUSTACHES EVERYWHERE  
WHY ARE THERE SO MANY BIRDS IN OHIO  
WHY IS THERE SO MUCH RAIN IN OHIO  
WHY IS OHIO WEATHER SO WEIRD  
**WHY ARE THERE MALE AND FEMALE BIKES**  
WHY ARE THERE BRIDESMAIDS  
WHY DO DYING PEOPLE REACH UP  
WHY AREN'T THERE WARDROBE PARTS  
WHY ARE OLD KINGDOMS DIFFERENT  
**WHY ARE THERE SQUIRRELS**  
WHY ARE THERE TINY SPIDERS IN MY HOUSE  
WHY DO SPIDERS COME INSIDE  
WHY ARE THERE HUGE SPIDERS IN MY HOUSE  
WHY ARE THERE LOTS OF SPIDERS IN MY HOUSE  
WHY ARE THERE SPIDERS IN MY ROOM  
WHY ARE THERE SO MANY SPIDERS IN MY ROOM  
**WHY DO SPIDER BITES ITCH**  
**WHY IS DYING SO SCARY**  
WHY IS THERE NO GPS IN LAPTOPS  
WHY DO KNEES CLICK  
WHY AREN'T THERE E GRADES  
WHY IS ISOLATION BAD  
WHY DO BOYS LIKE ME  
WHY DON'T BOYS LIKE ME  
WHY IS THERE ALWAYS A JAVA UPDATE  
WHY ARE THERE RED DOTS ON MY THIGHS  
**WHY IS LYING GOOD**  
WHY IS PROGRAMMING SO HARD  
WHY IS THERE A 0 OHM RESISTOR  
WHY DO AMERICANS HATE SOCCER  
WHY DO RHYMES SOUND GOOD  
**WHY DO TREES DIE**  
WHY IS THERE NO SOUND ON CHN  
WHY AREN'T POKEMON REAL  
WHY AREN'T BULLETS SHARP  
WHY DO DREAMS SEEM SO REAL

**WHY DO TESTICLES MOVE**  
WHY ARE THERE PSYCHICS  
WHY ARE HATS SO EXPENSIVE  
WHY IS THERE CRITIQUE IN MY SHIRTPOD  
WHY DO YOUR BOOBS HURT  
WHY DO TWINS HAVE DIFFERENT FINGERPRINTS  
WHY ARE AMERICANS AFRAID OF DRAGONS  
WHY ARE THERE SURVIVORS OF DANCES  
WHY IS THERE PALEOM  
WHY ARE THERE SO MANY CROWS IN ROCHESTER,  
WHY IS PSYCHIC WEAK TO BUG  
WHY DO CHILDREN GET CANCER  
WHY IS POSEIDON ANGRY WITH ODYSSEUS  
WHY IS THERE ICE IN SPACE  
**WHY ARE THERE ANTS IN MY LAPTOP**  
WHY IS THERE AN OWL IN MY BACKYARD  
WHY IS THERE AN OWL OUTSIDE MY WINDOW  
WHY IS THERE AN OWL ON THE DOLLAR BILL  
WHY DO OWLS ATTACK PEOPLE  
WHY ARE AK 47s SO EXPENSIVE  
WHY ARE THERE HELICOPTERS CIRCLING MY HOUSE  
WHY ARE THERE GODS  
WHY ARE THERE TWO SPOOKS  
WHY IS MT VESUVIUS THERE  
WHY DO THEY SAY T MINUS  
WHY ARE THERE OBEUSKS  
WHY ARE WRESTLERS ALWAYS WET  
WHY ARE OCEANS BECOMING MORE ACIDIC  
**WHY IS ARWEN DYING**  
WHY AREN'T MY QUAIL LAYING EGGS  
WHY AREN'T MY QUAIL EGGS HATCHING  
WHY AREN'T THERE ANY FOREIGN MILITARY BASES IN AMERICA

**WHY AREN'T ECONOMISTS RICH**  
WHY DO AMERICANS CALL IT SOCCER  
**WHY ARE MY EARS RINGING**  
WHY ARE THERE SO MANY AVENGERS  
WHY ARE THE AVENGERS FIGHTING THE X MEN  
WHY IS WOLVERINE NOT IN THE AVENGERS  
**WHY ARE THERE GHOSTS**  
WHY IS EARTH TILTED  
WHY IS SPACE BLACK  
WHY IS OUTER SPACE SO COLD  
WHY ARE THERE PYRAMIDS ON THE MOON  
WHY IS NASA SHUTTING DOWN

**WHY AREN'T MY ARMS GROWING**  
WHY ARE THERE WEEKS IN  
WHY DO I FEEL DIZZY

**WHY AREN'T THERE GUNS IN HARRY POTTER**  
WHY ARE ULTRASOUNDS IMPORTANT  
WHY ARE ULTRASOUND MACHINES EXPENSIVE  
WHY IS STEALING WRONG

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