

# Heavy neutrinos meet the triple Higgs coupling

*Collider Physics and the Cosmos*, Galileo Galilei Institute, Florence

Julien Baglio | 10/10/2017

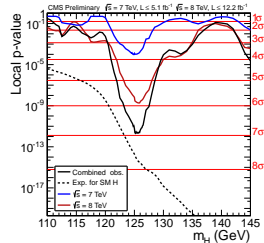
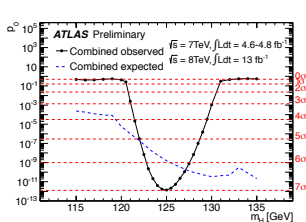
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- 1 Introduction
- 2 Neutrino effects on the triple Higgs coupling: 3+1 model
- 3 Neutrino effects on the triple Higgs coupling: Inverse seesaw
- 4 Outlook

# Once upon a time...

4/7/2012: CERN presents the discovery of a bosonic particle  
Its properties are compatible with those of the Higgs boson



$$M_H \simeq 125\text{ GeV}$$

[ATLAS, PLB 716 (2012) 1; CMS *ibid* 30]

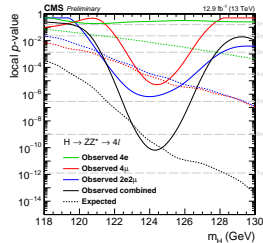
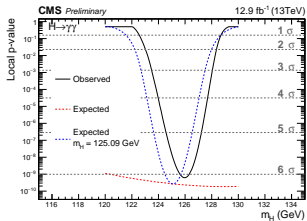
2013 analyses have confirmed the discovery of a Higgs boson: spin 0, couplings to fermions and bosons as a function of their masses  $\Rightarrow$  2013 Nobel Prize awarded to Englert and Higgs



**Key question: what is the exact nature of the observed Higgs boson? Standard Model (SM)-like or more exotics?**

# Once upon a time...

Observed again last year in 13 TeV data [CMS-PAS-HIG-16-020, CMS-PAS-HIG-16-033]



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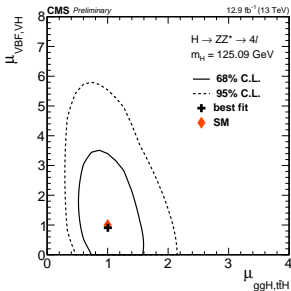
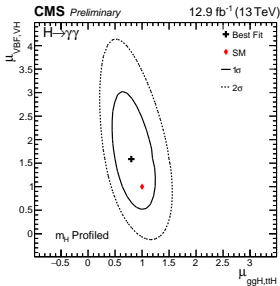
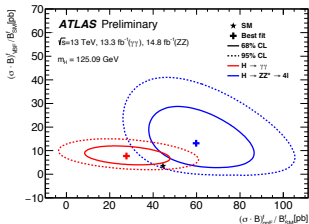
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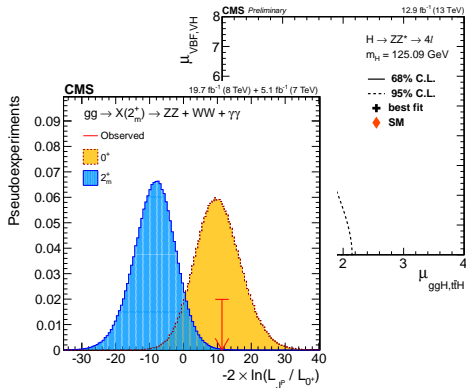
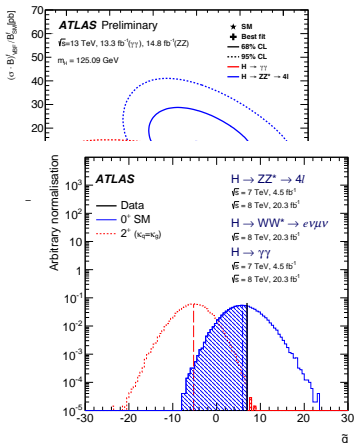


**Key question: what is the exact nature of the observed Higgs boson? Standard Model (SM)-like or more exotics?**

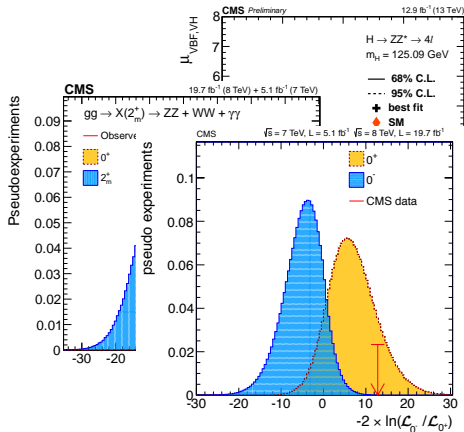
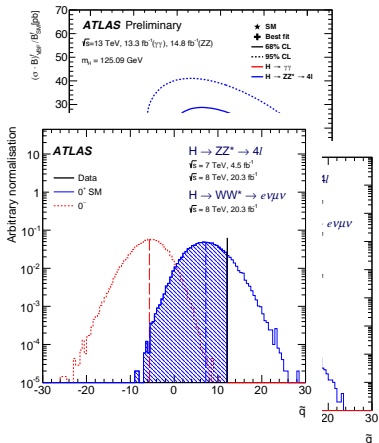
# Coining the scalar boson



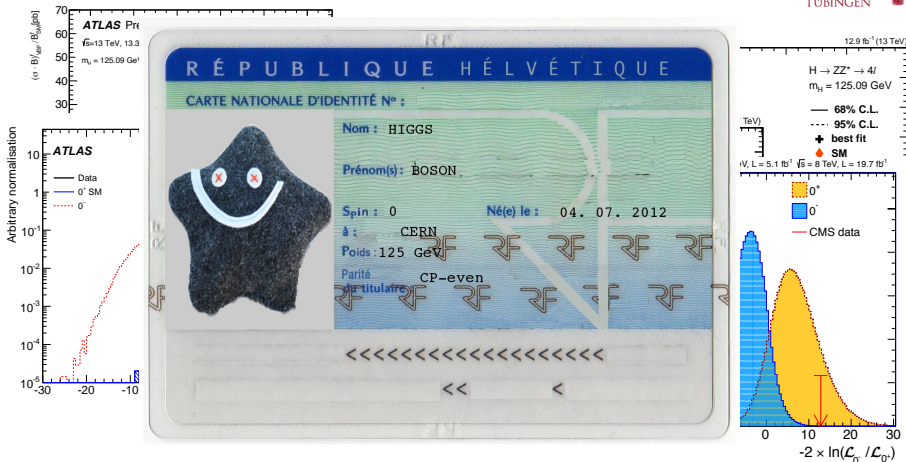
# Coining the scalar boson



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# Coining the scalar boson



**CP-even spin 0 hypothesis strongly preferred, no significant deviations from SM couplings: data up to now points toward an SM Higgs boson...**

[ATLAS Collaboration, EPIC 75 (2015) 476, ATLAS-CONF-2016-081; CMS Collaboration, PRD 89 (2014) 092007, PRD 92 (2015) 012004,

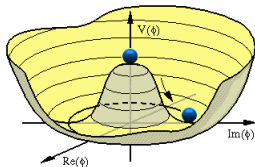
CMS-PAS-HIG-16-020, CMS-PAS-HIG-16-033]



# The SM ultimate test: probing the scalar potential

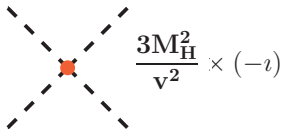
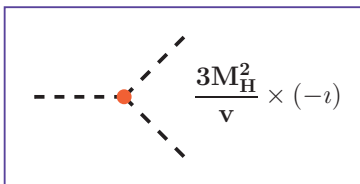
- From the scalar potential before EWSB ( $\phi$  as the Higgs field):

$$V(\phi) = -m^2|\phi|^2 + \lambda|\phi|^4$$



- To  $V(\phi)$  after EWSB, with  $M_H^2 = 2m^2$ ,  $v^2 = m^2/\lambda$ :

$$\phi = \begin{pmatrix} 0 \\ v + H(x) \\ \sqrt{2} \end{pmatrix} \Rightarrow V(H) = \frac{1}{2}M_H^2 H^2 + \frac{1}{2} \frac{M_H^2}{v} H^3 + \frac{1}{8} \frac{M_H^2}{v^2} H^4 + \text{constant}$$



- **Neutrino oscillations:** observed experimentally in 1998 [Super-Kamiokande, PRL 81 (1998) 1562]  
⇒ **neutrinos are massive!** ⇒ **new physics required** to account for their mass
- Different mixing pattern from CKM,  $\nu$  lightness; **Majorana or Dirac  $\nu$ ?**
- **No information through oscillations about:**
  - **Absolute mass scale:**  
cosmology  $\Sigma m_{\nu_i} < 0.23$  eV [Planck, A&A 594 (2016) A13]  
 $\beta$  decays  $m_{\nu_e} < 2.05$  eV [Mainz experiment, EPJC 40 (2005) 447; Troitsk, PRD 84 (2011) 112003]
  - **Neutrino nature (Dirac or Majorana):**  
Neutrinoless double  $\beta$  decays  $m_{2\beta} < 0.061 - 0.165$  eV [KamLAND-ZEN, PRL 117 (2016) 082503]

■ **Standard Model:**  $L = \begin{pmatrix} \nu_L \\ \ell_L \end{pmatrix}, \tilde{\phi} = \begin{pmatrix} H^{0*} \\ H^- \end{pmatrix}$

- No right-handed neutrino  $\nu_R \Rightarrow$  No Dirac mass term

$$\mathcal{L}_{\text{mass}} = -Y_\nu \bar{L} \tilde{\phi} \nu_R + \text{h.c.}$$

- No Higgs triplet  $T \Rightarrow$  No Majorana mass term

$$\mathcal{L}_{\text{mass}} = -\frac{1}{2} m \bar{L} T L^c + \text{h.c.}$$

■ **Necessary to go beyond the Standard Model for  $\nu$  mass**

- Radiative models?
- R-parity violation in supersymmetry?
- **Seesaw mechanisms?**  $\rightarrow \nu$  mass at tree-level  
 $\rightarrow$  **heavy sterile fermions**  
 $\Rightarrow$  **neutrino portal for Dark Matter?**

# Dirac neutrinos?

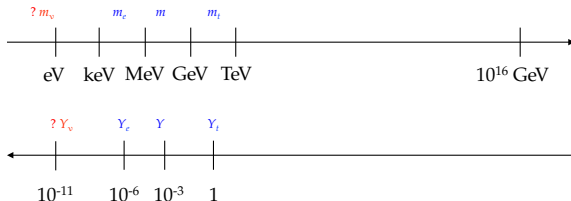
- Add **gauge singlet** (sterile), right-handed neutrinos  $\nu_R \Rightarrow \nu = \nu_L + \nu_R$

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -Y_\ell \bar{L} \phi \ell_R - Y_\nu \bar{L} \tilde{\phi} \nu_R + \text{h.c.}$$

$\Rightarrow$  After electroweak symmetry breaking:

$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -m_\ell \bar{\ell}_L \ell_R - m_D \bar{\nu}_L \nu_R + \text{h.c.}$$

$\Rightarrow$  **3** light active neutrinos:  $m_\nu \lesssim 1\text{eV} \Rightarrow Y^\nu \lesssim 10^{-11}$



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$$\mathcal{L}_{\text{mass}}^{\text{leptons}} = -Y_\ell \bar{L} \phi \ell_R - Y_\nu \bar{L} \tilde{\phi} \nu_R - \frac{1}{2} M_R \bar{\nu}_R \nu_R^c + \text{h.c.}$$

⇒ After electroweak symmetry breaking:

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3  $\nu_R$  ⇒ **6** mass eigenstates:  $\nu = \nu^c$

- $\nu_R$  gauge singlets  
⇒  $M_R$  not related to SM dynamics, not protected by symmetries  
⇒  $M_R$  between 0 and  $M_P$
- $M_R \bar{\nu}_R \nu_R^c$  violates lepton number conservation  $\Delta L = 2$

How to search for heavy neutrino with  $m_\nu > \mathcal{O}(1 \text{ TeV})$  ?

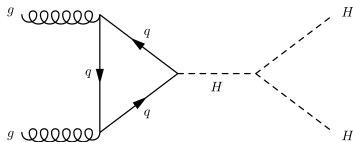
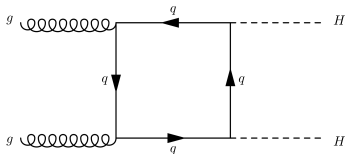
Use the Higgs sector to probe neutrino mass models

- TeV-scale neutrinos + Large Yukawa couplings  
⇒ Possibly **large deviations from SM properties** in the Higgs sector
- **$HH$  production:** one of the main motivation for high-luminosity LHC and future colliders  
⇒ need to study the impact of BSM on  $\lambda_{HHH} \Rightarrow$  **impact of heavy neutrino(s) on  $\lambda_{HHH}$ ?**
  - Sizeable SM 1-loop corrections ( $\mathcal{O}(10\%)$ )  $\Rightarrow$  Quantum corrections cannot be neglected
  - Sensitive to **diagonal** Yukawa couplings  $Y_\nu$

# Neutrino effects on the triple Higgs coupling The case of the 3+1 model

[J.B., Weiland, PRD 94 (2016) 013002]

## $\lambda_{HHH}$ extracted from $HH$ production

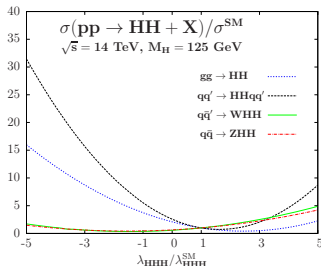


## Experimental prospects for the sensitivity to $\lambda_{HHH}$ :

- HL-LHC:  $\sim 50\%$  for ATLAS or CMS [CMS-PAS-FTR-15-002]  
 $\sim 35\%$  combined
- ILC: 27% at 500 GeV with  $4 \text{ ab}^{-1}$  [Fujii *et al.*, arXiv:1506.05992]  
10% at 1 TeV with  $5 \text{ ab}^{-1}$
- FCC-hh: 8% per experiment with  $3 \text{ ab}^{-1}$  using only  $b\bar{b}\gamma\gamma$  [Je, Ren, Yao, PRD 93 (2016) 015003]  
 $\sim 5\%$  combining all channels



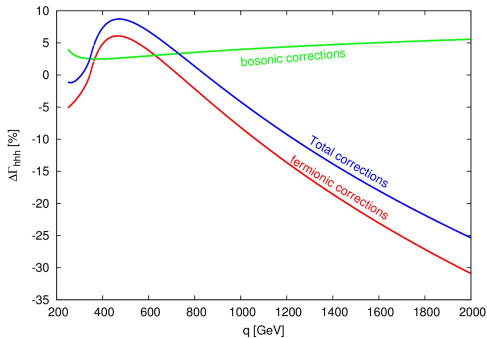
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[J.B *et al.*, JHEP 04 (2013) 151]

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 $\sim 5\%$  combining all channels



taken from [Arhrib *et al.*, JHEP 12 (2015) 007]

- tree-level:  $\lambda_{HHH}^0 = -\frac{3M_H^2}{v}$
- Dominant contribution from top-quark loops [Kanemura *et al.*, PRD 70 (2004) 115002]

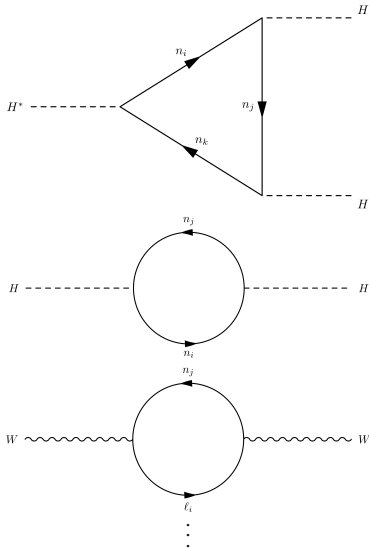
$$\lambda_{HHH}(q^2, m_H^2, m_H^2) = -\frac{3m_H^2}{v} \left[ 1 - \frac{1}{16\pi^2} \frac{16m_t^4}{v^2 m_H^2} \times \left\{ 1 + \mathcal{O}\left(\frac{m_H^2}{m_t^2}, \frac{q^2}{m_t^2}\right) \right\} \right]$$

- Opposite sign for the threshold ( $\sqrt{q^2} = 2m_t$ ) and  $m_t^4$  contributions

- A well-known class of renormalizable models for neutrino ( $\nu$ ) mass: **seesaw models**
  - To simplify the study of low-scale seesaw effects from a heavy  $\nu$ : Simplified models
  - Illustrate effects of new fermionic coupling through **neutrino portal**
- Simplified model with **3 light  $\nu$  ( $m_n = 1$  eV)** and **1 heavy sterile  $\nu$  ( $m_4$ )** parametrized by  $\nu$  masses and **active-sterile mixing  $B_{ij}$**

$$\begin{aligned}
 \mathcal{L} \ni & -\frac{g_2}{\sqrt{2}} \bar{\ell}_i \gamma^\mu W_\mu^- B_{ij} P_L n_j + \text{h.c.} \\
 & -\frac{g_2}{2M_W} \bar{n}_i (B^\dagger B)_{ij} H(m_{n_i} P_L + m_{n_j} P_R) n_j \\
 & -\frac{g_2}{2 \cos \theta_W} \bar{n}_i \gamma^\mu Z_\mu (B^\dagger B)_{ij} P_L n_j
 \end{aligned}
 \quad
 B_{3 \times 4} = \begin{pmatrix}
 B_{e1} & B_{e2} & B_{e3} & B_{e4} \\
 B_{\mu 1} & B_{\mu 2} & B_{\mu 3} & B_{\mu 4} \\
 B_{\tau 1} & B_{\tau 2} & B_{\tau 3} & B_{\tau 4}
 \end{pmatrix}$$

Active-sterile mixing matrix  $B$  constructed from the PMNS matrix



- Heavy  $\nu$  generates **new 1-loop diagrams and new counterterms**
- Counterterm to the triple Higgs coupling:

$$\begin{aligned}
 \frac{\delta\lambda_{HHH}}{\lambda_{HHH}^0} = & \frac{3}{2} \delta Z_H + \delta t_H \frac{e}{2M_W \sin\theta_W M_H^2} + \delta Z_e \\
 & + \frac{\delta M_H^2}{M_H^2} - \frac{1}{2} \frac{\delta M_W^2}{M_W^2} \\
 & + \frac{1}{2} \cot^2 \theta_W \left( \frac{\delta M_W^2}{M_W^2} - \frac{\delta M_Z^2}{M_Z^2} \right)
 \end{aligned}$$

- Tools for the calculation:
  - FeynArts/FormCalc and LoopTools 2.12 for the 1-loop integrals
  - implementation of a Model File for the neutrino interactions

## Theoretical constraints

- Loose (tight) **perturbativity** bound:

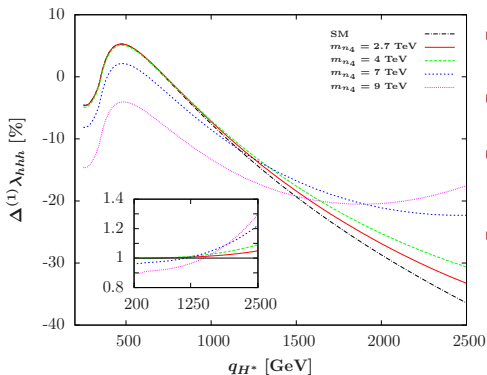
$$\left( \frac{\max |C_{i4}| g_2 m_{n_4}}{2M_W} \right)^3 < 16\pi (2\pi)$$

- Width** limit:  $\Gamma_{n_4} \leq 0.6 m_{n_4}$

## Experimental constraints

- PMNS matrix**: best fit of normal hierarchy with no CP-violation [Gonzalez-Garcia, Maltoni, Schwetz, JHEP 11 (2014) 052]
- Lepton flavor violating Higgs decays**  
 $\text{BR}(\mu \rightarrow e\gamma), \text{BR}(\tau \rightarrow e\gamma), \text{BR}(\tau \rightarrow \mu\gamma)$   
[MEG, EPJC 76 (2016) 434]
- Neutrinoless beta decay**: escaped as neutrino are Dirac
- Strongest experimental constraints on  $n_4$** :  
**Electroweak precision observables** [del Aguila, de Blas, Pérez-Victoria, PRD 78 (2008) 013010]

$$\begin{aligned} |B_{e4}| &\leq 0.041 \\ |B_{\mu 4}| &\leq 0.030 \\ |B_{\tau 4}| &\leq 0.087 \end{aligned}$$



- $\Delta^{(1)}\lambda_{HHH} = \frac{1}{\lambda^0} (\lambda_{HHH}^{1r} - \lambda^0)$

- $B_{\tau 4} = 0.087, B_{e4} = B_{\mu 4} = 0$

- Deviation of the BSM correction with respect to the SM correction in the insert

- $C_{44}m_{n_4} = m_t \Rightarrow m_{n_4} = 2.7 \text{ TeV}$

tight perturbativity bound:  $m_{n_4} = 7 \text{ TeV}$

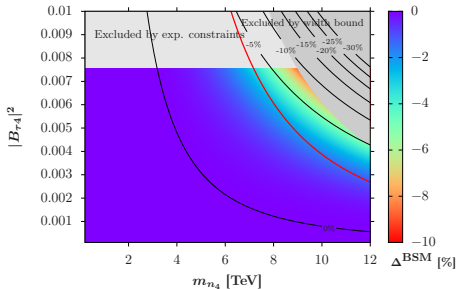
width bound:  $m_{n_4} = 9 \text{ TeV}$

- Largest positive correction at  $q_H^* \simeq 500 \text{ GeV}$ , heavy  $\nu$  decreases it

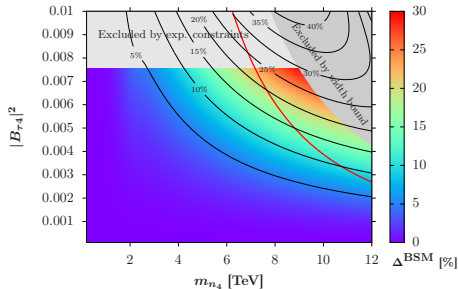
- Large negative correction at large  $q_H^*$ , heavy  $\nu$  increases it

# Contour maps in a 3+1 simplified model

$\Delta^{\text{BSM}}$  map with  $q_{H^\pm} = 500$  GeV



$\Delta^{\text{BSM}}$  map with  $q_{H^\pm} = 2500$  GeV



- $\Delta^{\text{BSM}} = \frac{1}{\lambda_{HHH}^{1r, \text{SM}}} \left( \lambda_{HHH}^{1r, \text{full}} - \lambda_{HHH}^{1r, \text{SM}} \right)$
- **Red line:** tight perturbativity bound
- Heavy  $\nu$  effects at the limit of HL-LHC sensitivity (35%)
- Heavy  $\nu$  effects clearly visible at the ILC (10%) and FCC-hh (5%)
- Similar plots for  $B_{e4}$  and  $B_{\mu 4}$

# Neutrino effects on the triple Higgs coupling The inverse seesaw

[J.B., Weiland, JHEP 1704 (2017) 038]



- TeV-scale neutrino induces **sizeable corrections** to  $\lambda_{HHH}$ 
  - Decrease at  $q_H^* \simeq 500 \text{ GeV}$
  - Increase at large  $q_H^*$
- Effects could be used to **constrain the active-sterile mixing** at the ILC and FCC-hh
- What are the effects in an appealing low-scale seesaw model ?
  - ▶ Inverse seesaw  $\rightarrow$  Additional constraints need to be included

# The inverse seesaw (ISS) mechanism

- Lower seesaw scale from approximately conserved lepton number
- Add **fermionic gauge singlets**  $\nu_R$  ( $L = +1$ ) and  $X$  ( $L = -1$ )

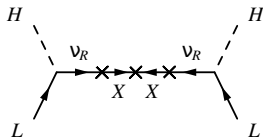
[Mohapatra, PRL 56 (1986) 561; Mohapatra, Valle, PRD 34 (1986) 1642; Bernabéu *et al.*, PLB 187 (1987) 303]

$$\mathcal{L}_{\text{ISS}} = -Y_\nu \bar{L} \tilde{\phi} \nu_R - M_R \bar{\nu}_R^c X - \frac{1}{2} \mu_X \bar{X}^c X + \text{h.c.}$$

with  $m_D = Y_\nu v / \sqrt{2}$ ,  $M^\nu = \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M_R \\ 0 & M_R^T & \mu_X \end{pmatrix}$

$$m_\nu \approx \frac{m_D^2}{M_R^2} \mu_X$$

$$m_{N_1, N_2} \approx \mp M_R + \frac{\mu_X}{2}$$



2 scales:  $\mu_X$  and  $M_R$

- **Decouple** neutrino mass generation from active-sterile mixing
- Inverse seesaw:  $Y_\nu \sim \mathcal{O}(1)$  and  $M_R \sim 1 \text{ TeV}$   
 $\Rightarrow$  **within reach of the LHC and low energy experiments**



- Accommodate low-energy neutrino data using **parametrization**

$$\nu Y_\nu^T = V^\dagger \text{diag}(\sqrt{M_1}, \sqrt{M_2}, \sqrt{M_3}) R \text{diag}(\sqrt{m_1}, \sqrt{m_2}, \sqrt{m_3}) U_{PMNS}^\dagger$$

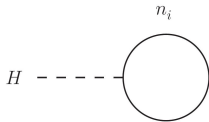
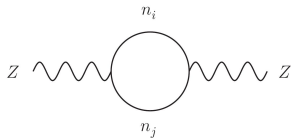
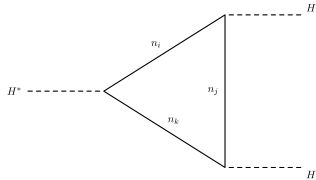
$$M = M_R \mu_X^{-1} M_R^T \text{ (Casas-Ibarra parametrization)}$$

or

$$\mu_X = M_R^T Y_\nu^{-1} U_{PMNS}^* m_\nu U_{PMNS}^\dagger Y_\nu^{T-1} M_R \nu^2 \quad \text{and beyond}$$

- Charged lepton flavor violation  
→ For example:  $\text{Br}(\mu \rightarrow e\gamma) < 4.2 \times 10^{-13}$  [MEG, EPJC 76 (2016) 434]
- Global fit to EWPO and lepton universality tests [Fernandez-Martinez *et al.*, JHEP 1608 (2016) 033]
- Electric dipole moment: **0** with **real** PMNS and mass matrices
- Invisible Higgs decays:  $M_R > m_H$ , **does not apply**
- Yukawa perturbativity:  $|\frac{Y_\nu^2}{4\pi}| < 1.5$

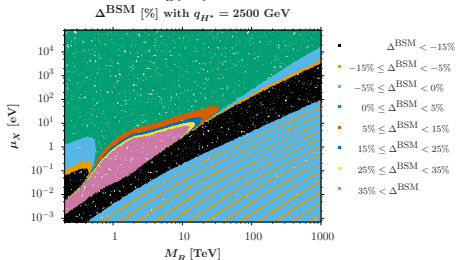
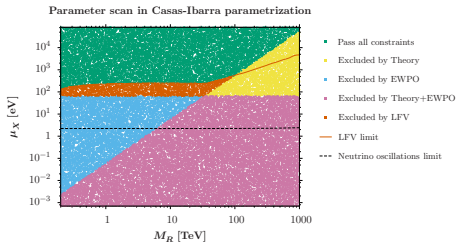
# Calculation in the ISS



- Similar diagrams to the 3+1 Dirac scenario but with Majorana neutrinos
- Evaluated with `FeynArts`, `FormCalc` and `LoopTools`
- More heavy neutrinos  
⇒ effects generically larger than 3+1 model

Analytical formulae for both Dirac and Majorana fermions

# Results using the Casas-Ibarra parametrization



- Random scan: 180000 points with degenerate (diagonal)  $M_R$  and  $\mu_X, \theta_i$  angles of the matrix  $R$ ,

$$\begin{aligned}
 0 &\leq \theta_i \leq 2\pi, \quad (i = 1, 2, 3) \\
 0.2 \text{ TeV} &\leq M_R \leq 1000 \text{ TeV} \\
 7 \times 10^{-4} \text{ eV} &\leq \mu_X \leq 8.26 \times 10^4 \text{ eV}
 \end{aligned}$$

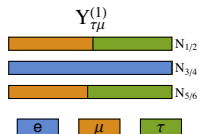
- $\Delta^{\text{BSM}} = \frac{1}{\lambda_{HHH}^{\text{SM}}} \left( \lambda_{HHH}^{\text{full}} - \lambda_{HHH}^{\text{SM}} \right)$
- Strongest constraints:
  - Lepton flavor violation, mainly  $\mu \rightarrow e\gamma$
  - Yukawa perturbativity (and neutrino width)
- Large effects necessarily excluded by LFV constraints ?

- How to evade the LFV constraints ?
- Approximate formulas for large  $Y_\nu$  [Arganda *et al.*, PRD 91 (2015) 015001]:

$$\text{Br}_{\mu \rightarrow e \gamma}^{\text{approx}} = 8 \times 10^{-17} \text{GeV}^{-4} \frac{m_\mu^5}{\Gamma_\mu} \left| \frac{v^2}{2M_R^2} (Y_\nu Y_\nu^\dagger)_{12} \right|^2$$

- Solution: Textures with  $(Y_\nu Y_\nu^\dagger)_{12} = 0$

$$Y_{\tau\mu}^{(1)} = |Y_\nu| \begin{pmatrix} 0 & 1 & -1 \\ 0.9 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix}$$

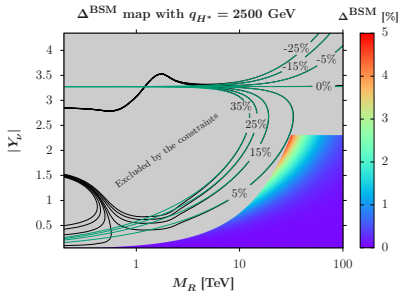
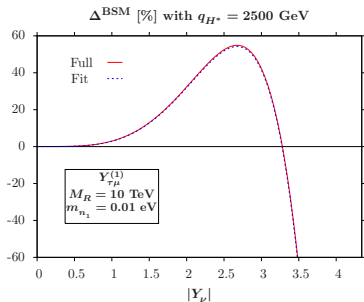


[taken from Arganda *et al.*, PLB 752 (2016)

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- Or even take  $Y_\nu$  diagonal

# Results for $Y_{\tau\mu}^{(1)}$



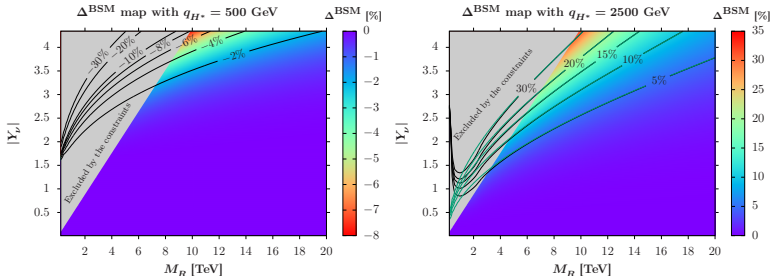
- $$\Delta^{\text{BSM}} = \frac{1}{\lambda_{HHH}^{\text{SM}}} \left( \lambda_{HHH}^{\text{full}} - \lambda_{HHH}^{\text{SM}} \right)$$

- Right: Full calculation in black, approximate formula in green
- Well described at  $M_R > 3$  TeV by approximate formula

$$\Delta_{\text{approx}}^{\text{BSM}} = \frac{(1 \text{ TeV})^2}{M_R^2} \left( 8.45 \text{Tr}(Y_\nu Y_\nu^\dagger Y_\nu Y_\nu^\dagger) - 0.145 \text{Tr}(Y_\nu Y_\nu^\dagger Y_\nu Y_\nu^\dagger Y_\nu Y_\nu^\dagger) \right)$$

- Can maximize  $\Delta^{\text{BSM}}$  by taking  $Y_\nu \propto I_3$

# Results in the ISS



- $\Delta^{\text{BSM}} = \frac{1}{\lambda_{HHH}^{1r, \text{SM}}} \left( \lambda_{HHH}^{1r, \text{full}} - \lambda_{HHH}^{1r, \text{SM}} \right)$
- Diagonal  $Y_\nu$  and hierarchical heavy neutrinos: full calculation in black, approximate formula in green; Maximize the correction
- Results agree with 3+1 Dirac analysis despite stronger constraints
- Heavy  $\nu$  effects at the limit of HL-LHC (35%) and ILC (10%) sensitivities clearly visible at the FCC-hh (5%)
- Complementary to existing observables  
→ Provide a new probe of the  $\mathcal{O}(10)$  TeV region of neutrino mass models



## HH production at hadron colliders: The Higgs frontier at the HL-LHC and at the future colliders

- **Major news since 2012:** a Higgs boson has been observed, now it is time to solve the next big question: **Is it standard or a first window on BSM physics?**
  - ⇒ Study of BSM effects on the triple Higgs coupling
- **Neutrino oscillations: New physics needed to generate  $\nu$  masses**
  - ⇒ low scale seesaw mechanism appealing model to generate tree-level  $\nu$  masses. Inverse seesaw:  $Y_\nu \sim \mathcal{O}(1)$  and  $M_R \sim \mathcal{O}(0.1 - 10)$  TeV
- **Neutrino effects on the triple Higgs coupling: Corrections as large as 30 %**
  - ⇒ measurable at future colliders
  - Probe a new part of the parameter space of the mass generation models
  - Larger effects expected with additional heavy neutrinos; generic effect, expected in all models with TeV fermions and large Higgs couplings
  - give new constraints on active-sterile neutrino mixing: impact on astrophysics, cosmology, neutrino physics
- **Future work:** corrections to  $HH$  production cross-section, UV completion to low-scale seesaw models

## Backup slides

# Renormalization procedure for the HHH coupling I



- No tadpole:  $t_H^{(1)} + \delta t_H = 0 \Rightarrow \delta t_H = -t_H^{(1)}$
- Counterterms:

$$M_H^2 \rightarrow M_H^2 + \delta M_H^2$$

$$M_W^2 \rightarrow M_W^2 + \delta M_W^2$$

$$M_Z^2 \rightarrow M_Z^2 + \delta M_Z^2$$

$$e \rightarrow (1 + \delta Z_e)e$$

$$H \rightarrow \sqrt{Z_H} = (1 + \frac{1}{2}\delta Z_H)H \quad (1)$$

- Full renormalized 1-loop triple Higgs coupling:  $\lambda_{HHH}^{1r} = \lambda^0 + \lambda_{HHH}^{(1)} + \delta\lambda_{HHH}$

$$\begin{aligned} \frac{\delta\lambda_{HHH}}{\lambda^0} &= \frac{3}{2}\delta Z_H + \delta t_H \frac{e}{2M_W \sin\theta_W M_H^2} + \delta Z_e + \frac{\delta M_H^2}{M_H^2} \\ &\quad - \frac{\delta M_W^2}{2M_W^2} + \frac{1}{2} \frac{\cos^2\theta_W}{\sin^2\theta_W} \left( \frac{\delta M_W^2}{M_W^2} - \frac{\delta M_Z^2}{M_Z^2} \right) \end{aligned}$$

- OS scheme

$$\begin{aligned}\delta M_W^2 &= \text{Re} \Sigma_{WW}^T(M_W^2) \\ \delta M_Z^2 &= \text{Re} \Sigma_{ZZ}^T(M_Z^2) \\ \delta M_H^2 &= \text{Re} \Sigma_{HH}(M_H^2)\end{aligned}\tag{2}$$

- Electric charge:

$$\delta Z_e = \frac{\sin \theta_W}{\cos \theta_W} \frac{\text{Re} \Sigma_{\gamma Z}^T(0)}{M_Z^2} - \frac{\text{Re} \Sigma_{\gamma\gamma}^T(M_Z^2)}{M_Z^2}$$

- Higgs field renormalization

$$\delta Z_H = -\text{Re} \left. \frac{\partial \Sigma_{HH}(k^2)}{\partial k^2} \right|_{k^2=M_H^2}$$

- Weaker constraints on diagonal couplings  
→ Large active-sterile mixing  $m_D M_R^{-1}$  for diagonal terms
- Previous parametrizations built on the 1st term in the  $m_D M_R^{-1}$  expansion →  
Parametrizations breaks down
- Solution: Build a parametrization including the next order terms
- The next-order  $\mu_X$ -parametrization is then

$$\begin{aligned}\mu_X \simeq & \left( \mathbf{1} - \frac{1}{2} M_R^{*-1} m_D^\dagger m_D M_R^{T-1} \right)^{-1} M_R^T m_D^{-1} U_{\text{PMNS}}^* m_\nu U_{\text{PMNS}}^\dagger m_D^{T-1} M_R \\ & \times \left( \mathbf{1} - \frac{1}{2} M_R^{-1} m_D^T m_D^* M_R^{\dagger-1} \right)^{-1}\end{aligned}$$