

# Reconstructing see-saw models from low energy data

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Based on works in collaboration with Alberto Casas,  
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**Flavour in the era  
of the LHC  
CERN, Feb 2006**

# 1-Motivation

**Why reconstruct the see-saw?** Why determining  $Y_\nu$  and  $\mathcal{M}$ ?

★ The effective theory provides a good **description** of neutrino observations. But we would like to have a **deeper understanding** of neutrino observations (why tiny masses? why two large angles?)

★ The **clues to unravel the flavour puzzle** lie in the fundamental theory.

- Look for patterns. In the quark sector,  $m_u : m_c : m_t \sim \lambda^8 : \lambda^4 : 1$ ,  
 $m_d : m_s : m_b \sim \lambda^4 : \lambda^2 : 1$ ,  $\lambda$  being the Cabibbo angle.

- Look for similarities with the charged lepton sector

- Look for similarities with the quark sector (GUT symmetries?)

★ Determine the **scale of new physics** (masses of the right-handed neutrinos).

Implications for

- GUTs

- leptogenesis

## 2- Approaches to determine the see-saw parameters

★ **top-down:** Start with a concrete model (GUT, Froggatt-Nielsen, strings...) and compare the predictions with the experiments.

**Many different possibilities.**

Unfortunately, the simplest ideas do not seem to work... May be we are being misled by theoretical prejudices?

★ **bottom-up:** Exploit all the information available at low energies on the leptonic sector, in order to reconstruct the high-energy theory.

Completely phenomenological. Impossible to get misled by aesthetics, but

**very difficult in practice**

★ In the **Standard Model** it is **hopeless**

$\{Y_\nu, \mathcal{M}\}$  depend on **18** parameters

12 real  
6 phases

$\{\mathcal{M}_\nu\}$  depends on **9** parameters

6 real  
3 phases

**Part of the information is lost in the decoupling process**

★ In the **Minimal Supersymmetric Standard Model**, radiative corrections on slepton parameters provide additional information about the see-saw mechanism, through the combination  $Y_\nu^\dagger Y_\nu$

Davidson, A.I.  
Ellis, Hisano, Raidal, Shimizu

Assume universality at  $M_X$ :

$$m_L^2(M_X) = m^2 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \longrightarrow m_L^2(M_Z) \sim m^2 \begin{pmatrix} \times & \times & \times \\ \times & \times & \times \\ \times & \times & \times \end{pmatrix}$$

$$16\pi^2 \frac{dm_L^2}{dt} = \left( m_L^2 Y_e^\dagger Y_e + Y_e^\dagger Y_e m_L^2 \right) + \left( m_L^2 Y_\nu^\dagger Y_\nu + Y_\nu^\dagger Y_\nu m_L^2 \right) \\ + 2 \left( Y_e^\dagger m_e^2 Y_e + m_{H_1}^2 Y_e^\dagger Y_e + A_e^\dagger A_e \right) + 2 \left( Y_\nu^\dagger m_\nu^2 Y_\nu + m_{H_2}^2 Y_\nu^\dagger Y_\nu + A_\nu^\dagger A_\nu \right) \\ - \left( \frac{6}{5} g_1^2 |M_1|^2 + 6g_2^2 |M_2|^2 \right) \mathbf{I}_3 - \frac{3}{5} g_1^2 S \mathbf{I}_3$$

In the leading-log approximation:

$$(m_L^2)_{ij} \simeq \# \mathbf{I}_3 - \frac{1}{8\pi^2} (3m_0^2 + A_0^2) (\mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu)_{ij} \log \frac{M_X}{M}$$

★ Is it possible to reconstruct the see-saw parameters with the information from  $\mathcal{M}$  and  $Y_\nu^\dagger Y_\nu$ ? **YES!!**

In the basis where  $\mathcal{M} = D_{\mathcal{M}} = \text{diag}(M_1, M_2, M_3)$ , the Yukawa coupling reads  $\mathbf{Y}_\nu = V_R^\dagger \text{diag}(Y_1, Y_2, Y_3) V_L$ .

Then, the reconstruction follows in two steps: Davidson, A.I.

- $\mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu = V_L^\dagger \text{diag}(Y_1^2, Y_2^2, Y_3^2) V_L$

From here we extract  $V_L$  and  $\text{diag}(Y_1, Y_2, Y_3)$

- $\mathcal{M}_\nu = \mathbf{Y}_\nu^T \mathcal{M}^{-1} \mathbf{Y}_\nu \langle H_u^0 \rangle^2 = V_L^\dagger D_Y V_R^* D_{\mathcal{M}}^{-1} V_R^\dagger D_Y V_L$

$$\frac{1}{\langle H_u^0 \rangle^2} D_Y^{-1} V_L \mathcal{M}_\nu V_L^\dagger D_Y^{-1} = V_R^* D_{\mathcal{M}}^{-1} V^\dagger$$

From here we extract  $V_R$  and  $D_{\mathcal{M}}$

**We have everything!!** However, in practice some of the low energy parameters could be very hard to measure (if not impossible)

neutrino mass matrix, $\mathcal{M}_\nu$			radiative effects, $P \equiv \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu$			
$m_1$	$m_1$	😊	$P_{11}$	mass splittings	largest	😊
$m_2$	$\Delta m_{atm}^2$	✓	$P_{22}$		smallest	😞
$m_3$	$\Delta m_{sol}^2$	✓	$P_{33}$	absolute scale		😞
$\theta_{12}$	✓		$ P_{12} $		$\mu \rightarrow e\gamma$	😊
$\theta_{13}$	😊		$ P_{13} $	rare decays	$\tau \rightarrow e\gamma$	😊
$\theta_{23}$	✓		$ P_{23} $		$\tau \rightarrow \mu\gamma$	😊
$\delta$		😊	$\arg P_{12}$	e-EDM	one of them	😊
$\phi$	$\nu 0\beta\beta$ phase	😞	$\arg P_{13}$	$\mu$ -EDM	the other two	😞
$\phi'$	orthogonal combination	😞 😞	$\arg P_{23}$	$\tau$ -EDM		

## An interesting implication of this procedure:

There is a one to one correspondence between low energy observables and see-saw parameters:  $\{\mathbf{Y}_\nu, \mathcal{M}\} \longleftrightarrow \{\mathcal{M}_\nu, \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu\}$

$\{\mathbf{Y}_\nu, \mathcal{M}\}$  depend on 18 parameters      12 real  
6 phases

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★ **Positive lesson:** We can use this trick as an alternative parametrization of the see-saw mechanism, using only low energy observables.

★ **Negative lesson:** For any set of observations, there is a see-saw scenario that accommodates them. **The see-saw cannot be ruled out!!**

# 3- A less radical (and more practical) approach

Both the top-down and the bottom-up approaches are interesting, but have limitations → **try a hybrid approach:**

## bottom-up approach with some well-motivated hypotheses about the high-energy theory

- GUT inspired  $Y_\nu$  symmetric  
 $Y_\nu$  eigenvalues as  $m_u, m_c, m_t$  Akhmedov, Frigerio, Smirnov

- texture zeros. Inspired by Gatto-Sartori-Tonin relation, A.I., Ross

$$\sin \theta_c \simeq \left| \sqrt{\frac{m_d}{m_s}} - e^{i\delta} \sqrt{\frac{m_u}{m_c}} \right|$$

- **two right-handed neutrino (2RHN) model**

Kleppe; Ma, Roy, Sarkar; Frampton, Glashow, Yanagida; Raidal, Strumia; Gonzalez-Felipe, Joaquim, Nobre; Raby Appelquist, Shrock; Endoh, Kaneko, Kang, Morozumi, Tanimoto Smirnov; King, Ross; Rodejohann; Dreiner, Murayama, Thormeier A.I., Ross; Altarelli, Feruglio, Masina...

Observations tell us that at least two new mass scales have to be introduced

$(\Delta m_{sol}^2, \Delta m_{atm}^2)$ . Two right-handed neutrinos can do the job.



**Further motivation:** there are some interesting situations where the 3RHN model can be well described by the 2RHN model. The mass matrix is:

$$\mathcal{M}_{ij} = \frac{y_{1i}y_{1j}}{M_1} + \frac{y_{2i}y_{2j}}{M_2} + \frac{y_{3i}y_{3j}}{M_3}, \text{ where } y_{ij} = (\mathbf{Y}_\nu)_{ij}.$$

Two RH neutrinos dominate when

- $\frac{y_{1i}y_{1j}}{M_1} \ll \frac{y_{2i}y_{2j}}{M_2}, \frac{y_{3i}y_{3j}}{M_3}$  It occurs when  $y_{1i} \ll y_{2i}, y_{3i}$ . In this case,

$$\mathcal{M}_{ij} \simeq \frac{y_{2i}y_{2j}}{M_2} + \frac{y_{3i}y_{3j}}{M_3} \longrightarrow \text{2RHN model!!}$$

$$(\mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu)_{ij} \simeq y_{2i}^* y_{2j} + y_{3i}^* y_{3j}$$

- $\frac{y_{2i}y_{2j}}{M_2} \ll \frac{y_{1i}y_{1j}}{M_1}, \frac{y_{3i}y_{3j}}{M_3}$  Not very interesting...

- $\frac{y_{3i}y_{3j}}{M_3} \ll \frac{y_{1i}y_{1j}}{M_1}, \frac{y_{2i}y_{2j}}{M_2}$  It occurs when  $M_3 \gg M_1, M_2$

But it could happen that  $(\mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu)_{ij} = y_{1i}^* y_{1j} + y_{2i}^* y_{2j} + y_{3i}^* y_{3j}$ . **Not 2RHN!**

In GMSB models, with messenger masses between  $M_2$  and  $M_3$ ,

$$\mathcal{M}_{ij} \simeq \frac{y_{1i}y_{1j}}{M_1} + \frac{y_{2i}y_{2j}}{M_2} \longrightarrow \text{2RHN model!!}$$

$$(\mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu)_{ij} = y_{1i}^* y_{1j} + y_{2i}^* y_{2j}$$



## ★ Predictions on the radiative effects

We want to incorporate in these predictions the information from the neutrino mass matrix. Working in the basis where  $\mathcal{M} = \text{diag}(M_1, M_2)$ , it can be checked that

$$\mathbf{Y}_\nu = D_{\sqrt{\mathcal{M}}} R D_{\sqrt{m}} U^\dagger / \langle H_u^0 \rangle$$

Casas, A.I.

is the most general Yukawa coupling that satisfies

$$\mathcal{M}_\nu = \mathbf{Y}_\nu^T \text{diag}(M_1^{-1}, M_2^{-1}) \mathbf{Y}_\nu \langle H_u^0 \rangle^2$$

Here,  $R$  is an orthogonal **complex** matrix

$$R = \begin{pmatrix} 0 & \cos z & \sin z \\ 0 & -\sin z & \cos z \end{pmatrix}$$

A.I., Ross

So, the independent parameters in  $\mathbf{Y}_\nu$  are  $M_1$ ,  $M_2$  and  $z$  (three real and one phase). On the other hand,  $Y_\nu^\dagger Y_\nu$  depends in general on 6 real parameters and three phases. **We can make predictions on three real parameters and two phases in  $Y_\nu^\dagger Y_\nu$ .**

To compute the predictions, we construct the matrix

$$U^\dagger \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu U = D_{\sqrt{m}} R^\dagger D_{\mathcal{M}} R D_{\sqrt{m}} / \langle H^0 \rangle^2$$

since  $m_1 = 0$ ,

$$(U^\dagger \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu U)_{1i} = 0 \quad \text{for } i = 1, 2, 3. \quad \text{A.I.}$$

**five equations  $\rightarrow$  five relations/predictions.**

★ Still three independent real parameters and one phase. We choose as independent parameters  $P_{12}, |P_{13}|, |P_{23}|$  ( $P \equiv \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu$ ). The predictions are:

$$P_{11} = -\frac{P_{12}^* U_{21}^* + P_{13}^* U_{31}^*}{U_{11}^*}$$

$$P_{22} = -\frac{P_{12} U_{11}^* + P_{23}^* U_{31}^*}{U_{21}^*}$$

$$P_{33} = -\frac{P_{13} U_{11}^* + P_{23} U_{21}^*}{U_{31}^*}.$$

$$e^{i \arg P_{13}} = \frac{-i \operatorname{Im}(P_{12} U_{21} U_{11}^*) \pm \sqrt{|P_{13}|^2 |U_{11}|^2 |U_{31}|^2 - [\operatorname{Im}(P_{12} U_{21} U_{11}^*)]^2}}{|P_{13}| |U_{31}| U_{11}^*}$$

$$e^{i \arg P_{23}} = \frac{i \operatorname{Im}(P_{12} U_{21} U_{11}^*) \pm \sqrt{|P_{23}|^2 |U_{21}|^2 |U_{31}|^2 - [\operatorname{Im}(P_{12} U_{21} U_{11}^*)]^2}}{|P_{23}| |U_{31}| U_{21}^*}$$

# 4- Reconstructing the 2RHN model

The 2RHN model can be parametrized in terms of

$\{Y_\nu, \mathcal{M}\}$  depend on **11** parameters 8 real  
3 phases

$\left\{ \begin{array}{l} m_2, m_3 \\ \theta_{12}, \theta_{13}, \theta_{23}, \\ \delta, \phi \end{array} \right\}$   $|P_{12}|, |P_{13}|, |P_{23}|$   
 $\arg P_{12}$  depend on **11** parameters 8 real  
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It is possible to derive exact expressions for the high-energy parameters in terms of the low-energy parameters:

$$M_1 = \frac{1}{2} \left[ \sqrt{\left(\frac{Q_{33}}{m_3} + \frac{Q_{22}}{m_2}\right)^2 + \frac{(Q_{23} - Q_{23}^*)^2}{m_2 m_3}} - \sqrt{\left(\frac{Q_{33}}{m_3} - \frac{Q_{22}}{m_2}\right)^2 + \frac{(Q_{23} + Q_{23}^*)^2}{m_2 m_3}} \right] \langle H_u^0 \rangle^2$$

$$M_2 = \frac{1}{2} \left[ \sqrt{\left(\frac{Q_{33}}{m_3} + \frac{Q_{22}}{m_2}\right)^2 + \frac{(Q_{23} - Q_{23}^*)^2}{m_2 m_3}} + \sqrt{\left(\frac{Q_{33}}{m_3} - \frac{Q_{22}}{m_2}\right)^2 + \frac{(Q_{23} + Q_{23}^*)^2}{m_2 m_3}} \right] \langle H_u^0 \rangle^2$$

$$\cos 2z = \left( \frac{Q_{33}^2}{m_3^2} - \frac{Q_{22}^2}{m_2^2} + \frac{(Q_{23} + Q_{23}^*)(Q_{23} - Q_{23}^*)}{m_2 m_3} \right) \frac{\langle H_u^0 \rangle^4}{M_2^2 - M_1^2}$$

where  $Q = U^\dagger P U = U^\dagger \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu U$ .

Finally, the Yukawa coupling can be reconstructed  $\mathbf{Y}_\nu = D_{\sqrt{\mathcal{M}}} R D_{\sqrt{m}} U^\dagger / \langle H_u^0 \rangle$

★ Is it feasible? In the 3RHN model

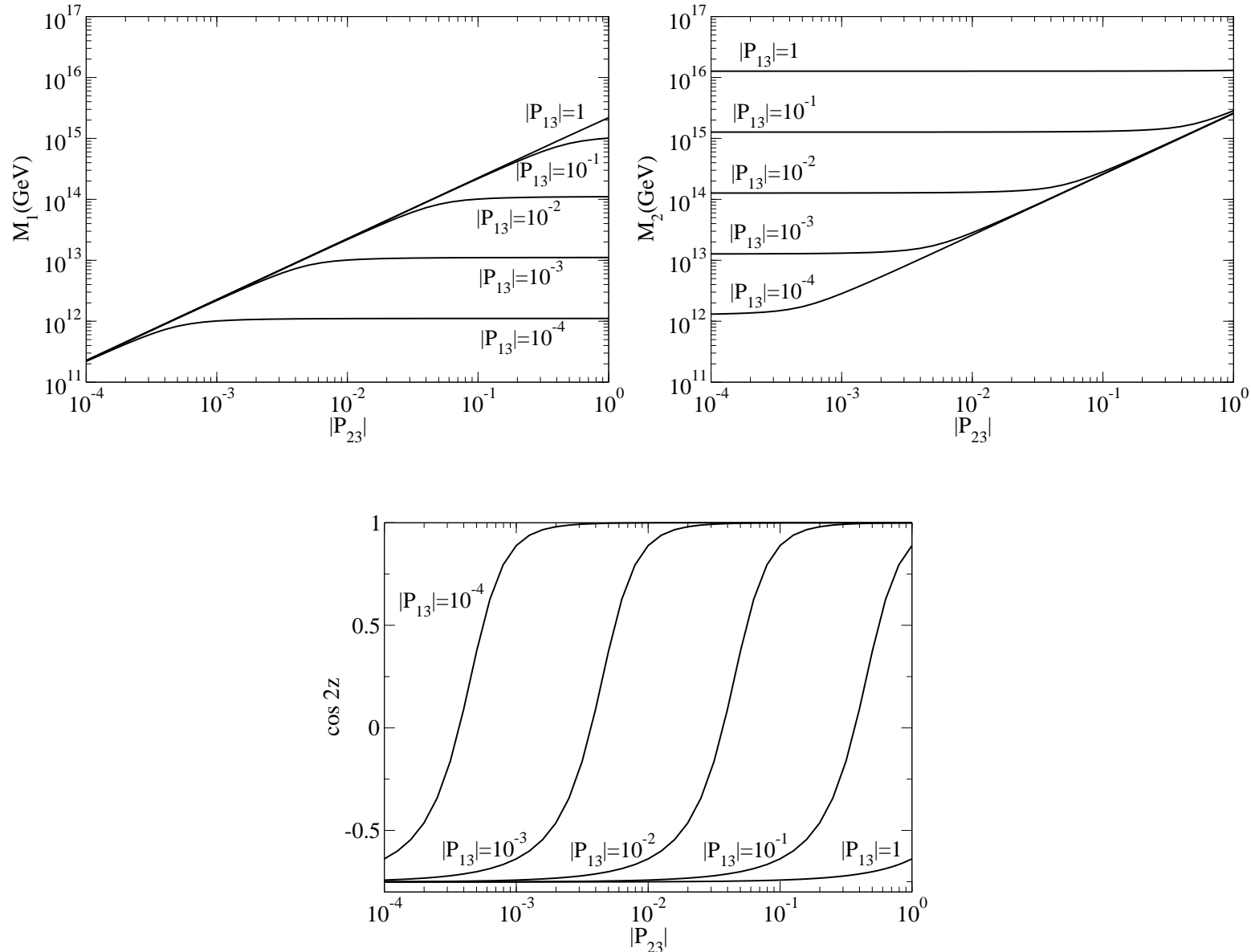
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★ Is it feasible? In the 2RHN model

neutrino mass matrix, $\mathcal{M}_\nu$	radiative effects, $P \equiv \mathbf{Y}_\nu^\dagger \mathbf{Y}_\nu$
<del><math>m_1</math></del> $m_1$ 😊	<del><math>P_{11}</math></del> mass splittings largest 😊
$m_2$ $\Delta m_{atm}^2$ ✓	<del><math>P_{22}</math></del> smallest 😞
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$\theta_{12}$ ✓	$ P_{12} $ $\mu \rightarrow e\gamma$ 😊
$\theta_{13}$ 😊	$ P_{13} $ rare decays $\tau \rightarrow e\gamma$ 😊
$\theta_{23}$ ✓	$ P_{23} $ $\tau \rightarrow \mu\gamma$ 😊
$\delta$ 😊	$\arg P_{12}$ e-EDM one of them 😊
$\phi$ $\nu 0\beta\beta$ phase 😞	<del><math>\arg P_{13}</math></del> $\mu$ -EDM the other two 😞
<del><math>\phi'</math></del> orthogonal combination 😞 😞	<del><math>\arg P_{23}</math></del> $\tau$ -EDM

# An example of reconstruction: the case with $|P_{12}| \ll |P_{13}|, |P_{23}|$

★ Assume for simplicity that **all the parameters are real and  $\theta_{13} = 0$**





- $\boxed{|P_{13}| \ll |P_{23}|} \implies P \simeq |P_{23}| \begin{pmatrix} \lambda/\sqrt{6} & 0 & -\lambda \\ 0 & 1 & 1 \\ -\lambda & 1 & 1 \end{pmatrix}, \text{ with } \lambda = \frac{|P_{13}|}{|P_{23}|} \ll 1$

$$M_1 \simeq 2\sqrt{\frac{2}{3}} \frac{|P_{13}|}{m_2} \langle H_u^0 \rangle^2$$

$$M_2 \simeq \frac{2|P_{23}|}{m_3} \langle H_u^0 \rangle^2$$

$$\mathbf{Y}_\nu \simeq \sqrt{|P_{23}|} \begin{pmatrix} \sqrt{\frac{|P_{13}|}{\sqrt{6}|P_{23}|}} & \sqrt{\sqrt{\frac{3}{8}} \frac{|P_{13}|}{|P_{23}|}} & -\sqrt{\sqrt{\frac{3}{8}} \frac{|P_{13}|}{|P_{23}|}} \\ -\frac{|P_{13}|}{2|P_{23}|} & 1 & 1 \end{pmatrix}$$

- $|P_{23}| \ll |P_{13}| \implies P \simeq |P_{13}| \begin{pmatrix} 1/\sqrt{6} & 0 & -1 \\ 0 & \lambda & \lambda \\ -1 & \lambda & \sqrt{6} \end{pmatrix}$ , with  $\lambda = \frac{|P_{23}|}{|P_{13}|} \ll 1$

– When the light neutrinos have the same CP parities ( $\phi = 0$ )

$$M_1 \simeq \frac{8|P_{23}|}{3m_2+4m_3} \langle H_u^0 \rangle^2$$

$$M_2 \simeq \frac{(3m_2+4m_3)|P_{13}|}{\sqrt{6}m_2m_3} \langle H_u^0 \rangle^2$$

$$\mathbf{Y}_\nu \simeq \sqrt{\sqrt{6}|P_{13}|} \begin{pmatrix} \sqrt{\frac{\sqrt{6}|P_{23}|}{|P_{13}|} \frac{m_2}{3m_2+4m_3}} & \sqrt{\frac{|P_{23}|}{\sqrt{6}|P_{13}|}} & \sqrt{\frac{|P_{23}|}{\sqrt{6}|P_{13}|} \frac{-3m_2+4m_3}{3m_2+4m_3}} \\ -\frac{1}{\sqrt{6}} & \frac{\sqrt{6}m_2}{3m_2+4m_3} \frac{|P_{23}|}{|P_{13}|} & 1 \end{pmatrix}$$

– When they have opposite parities ( $\phi = \pi$ )

$$M_1 \simeq \frac{8|P_{23}|}{-3m_2+4m_3} \langle H_u^0 \rangle^2$$

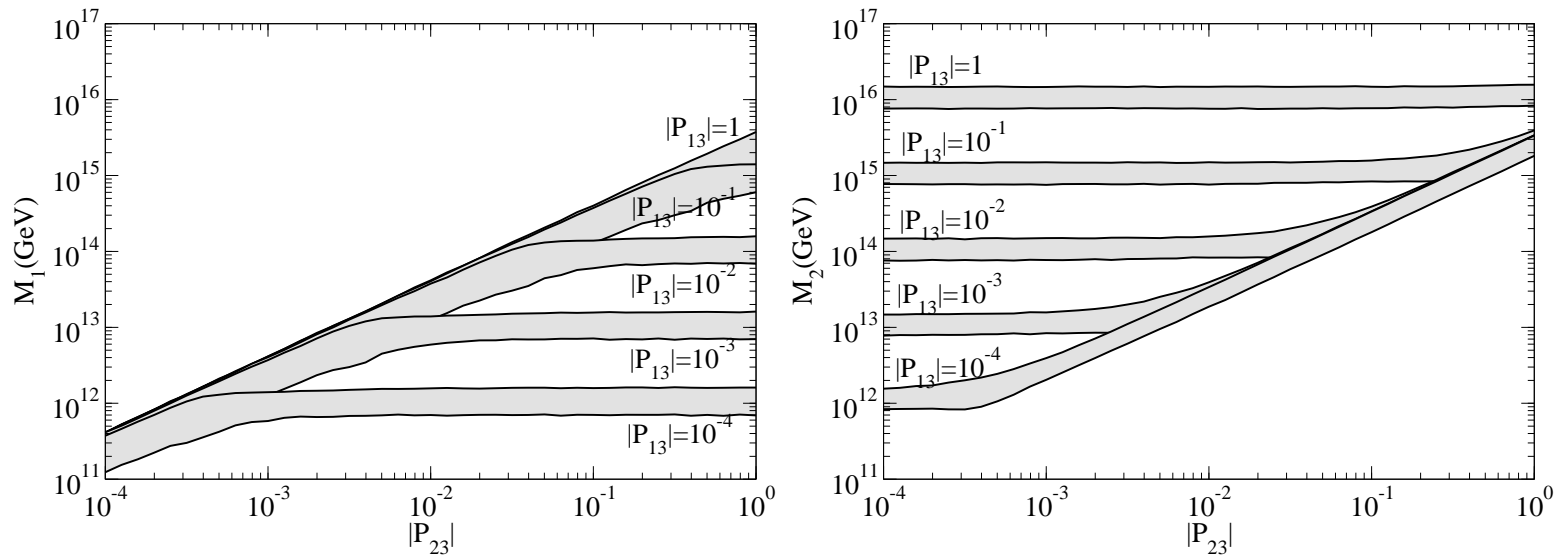
$$M_2 \simeq \frac{(-3m_2+4m_3)|P_{13}|}{\sqrt{6}m_2m_3} \langle H_u^0 \rangle^2$$

$$\mathbf{Y}_\nu \simeq \sqrt{\sqrt{6}|P_{13}|} \begin{pmatrix} -\sqrt{\frac{\sqrt{6}|P_{23}|}{|P_{13}|} \frac{m_2}{-3m_2+4m_3}} & \sqrt{\frac{|P_{23}|}{\sqrt{6}|P_{13}|}} & \sqrt{\frac{|P_{23}|}{\sqrt{6}|P_{13}|} \frac{3m_2+4m_3}{-3m_2+4m_3}} \\ -\frac{i}{\sqrt{6}} & \frac{-i\sqrt{6}m_2}{-3m_2+4m_3} \frac{|P_{23}|}{|P_{13}|} & i \end{pmatrix}$$

★ Effect of a non-vanishing  $\theta_{13}$ ,  $\delta$  and  $\phi$

The reconstruction of the Yukawa coupling requires to know  $\delta$  and  $\phi$ , but the order of magnitude of the right-handed masses could be estimated.

$\theta_{13} = 0.1$ ,  $\delta, \phi$  random



## 5- Conclusions

- ★ The reconstruction of the full see-saw model with three right-handed neutrinos is possible in theory, but extremely difficult in practice.
- ★ The model with two right-handed neutrinos is the **most simple see-saw scenario that can accommodate the observations**. Furthermore, there are interesting limits of the 3RHN model that resemble the 2RHN model.
- ★ In the 2RHN model, there are correlations among the elements of  $Y_\nu^\dagger Y_\nu \longrightarrow$  they could give rise to correlations among slepton parameters.
- ★ It could be possible to reconstruct the Yukawa coupling of the 2RHN model in terms of low energy observables (**but only after an enormous experimental effort!**). On the other hand, the order of magnitude of the right-handed masses could be estimated if SUSY and rare decays are observed.