

# Probing SUSY effects in $K_S^0 \rightarrow \mu^+ \mu^-$

Miriam Lucio

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V. Chobanova, G. D'Ambrosio, T. Kitahara, M. Lucio Martinez,  
D. Martinez Santos, I. Suarez Fernandez, K. Yamamoto

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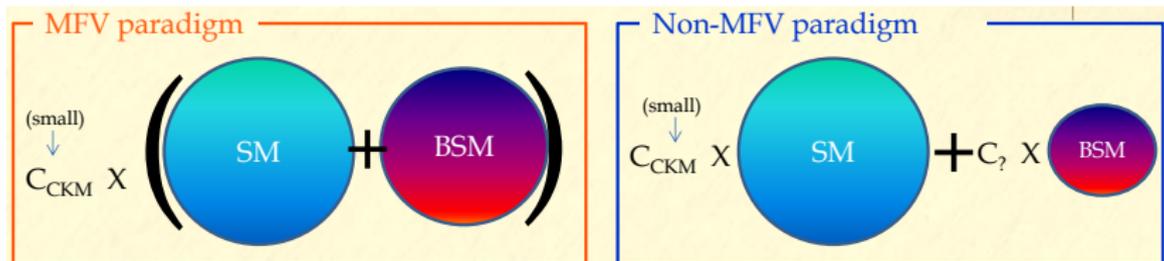
# Motivation

- $P \rightarrow l^+ l^-$  decays very sensitive to the Higgs sector of MSSM ( $\propto \tan^6 \beta / M_A^4$ )
- If BSM Physics is at energies  $> \mathcal{O}(\text{TeV})$ , the only possibility to see it in quark flavour physics is if it has new sources of Flavour Violation: **non-MFV**, not originating from the CKM Matrix

	$< 1 \text{ TeV}$	few TeV	$> \text{few TeV}$
Direct searches	Reachable @ 8 TeV	Reachable @ 14 TeV	Not reachable
Quark flavour			
Anarchic structure	$> \mathcal{O}(1)$ effects	$\mathcal{O}(1)$ effects	$\mathcal{O}(1)$ effects
Small misalignment	$\mathcal{O}(1)$ effects	$\mathcal{O}(10\%)$ effects	$\mathcal{O}(1-10\%)$ effects
MFV	$\mathcal{O}(10\%)$ effects	$\mathcal{O}(1\%)$ effects	Invisible

# Motivation

- **Minimal Flavour Violation:**  $B_s^0 \rightarrow \mu^+ \mu^-$  is the dominant constraint
- **New sources for Flavour Violation:**  $B_{s(d)}^0 \rightarrow \mu^+ \mu^-$ ,  $K_{S(L)}^0 \rightarrow \mu^+ \mu^-$  carry **complementary** information



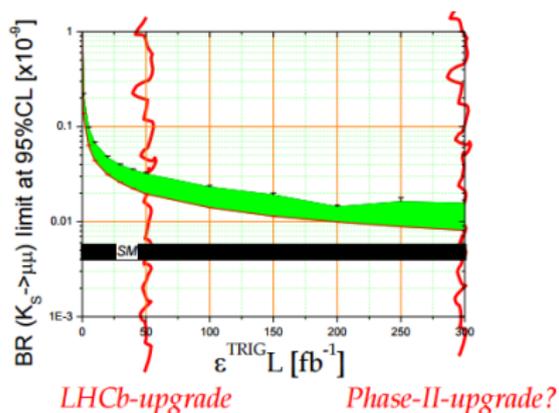
$C \sim 1 \Rightarrow$  bounds from kaon physics go up to a scale of  $10^5 \text{TeV}$

$$K_S^0 \rightarrow \mu^+ \mu^-$$

$$K_S^0 \rightarrow \mu^+ \mu^-$$

$$\mathcal{B}_{SM}(K_S^0 \rightarrow \mu^+ \mu^-) = (5.18 \pm 1.50_{LD} \pm 0.02_{SD}) \times 10^{-12}$$

[Nucl. Phys. B366 (1991) 189], [JHEP 01 (2004) 009]



- Strongly suppressed decay in the SM

LHCb Run1 data (EPJ C (678) 77)

$$\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) < 0.8(1.0) \times 10^{-9}$$

90%, 95% CL

- Dispersive treatment of  $K_S^0 \rightarrow \gamma\gamma$  and  $K_S^0 \rightarrow \gamma l^+ l^-$  [EPJ C76 (2016) 604]:  $K_S^0 \rightarrow \gamma\gamma$ ,  $K_S^0 \rightarrow \gamma\mu^+\mu^-$ ,  $K_S^0 \rightarrow \mu^+\mu^-\mu^+\mu^-$ ,  $K_S^0 \rightarrow \mu^+\mu^-e^+e^-$  (measurable @ LHCb) are relevant to improve LD prediction

## Mass Insertion Approximation (MIA)

[Nucl. Phys. B830 (2010) 17–94], [arXiv:hep-ph/9511250]

- MI terms as phenomenological parameters at the SUSY scale,

$$\mu^{\text{SUSY}} = \sqrt{\tilde{m}_Q M_3}$$

$$\begin{aligned}(\delta_d^{LL})_{ij} &= \frac{[(\mathcal{M}_D^2)_{LL}]_{ij}}{\tilde{m}_Q^2} = \frac{(m_Q^2)_{ij}}{\tilde{m}_Q^2}, \\(\delta_d^{RR})_{ij} &= \frac{[(\mathcal{M}_D^2)_{RR}]_{ij}}{\tilde{m}_d^2} = \frac{(m_D^2)_{ij}}{\tilde{m}_d^2},\end{aligned}$$

- $\mathcal{M}_{D,U}^2$ : 6x6 squark mass matrices;  $\tilde{m}_Q^2$ ,  $\tilde{m}_d^2$ ,  $\tilde{m}_u^2$ : average of the Q, D, and U-squark masses

$\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)$ :

- Relation between  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)$  and  $\mathcal{B}(K_L^0 \rightarrow \mu^+ \mu^-)$  different for **left-handed** and **right-handed** new physics scenarios
  - 1 **left-handed**:  $C_P = C_S$  (P: pseudoscalar, S: scalar)
  - 2 **right-handed**:  $C_P = -C_S$  (P: pseudoscalar, S: scalar)
  - 3 **right-handed** + **left-handed**
- $\mathcal{B}(K_L^0 \rightarrow \mu^+ \mu^-)^{\text{EXP}}$  upper bound on  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)$  for 1 and 2, alleviated for 3

$\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)_{\text{eff}}$

- Includes and interference contribution with  $K_L^0 \rightarrow \mu^+ \mu^-$

**Direct CP asymmetry**

$$A_{CP}(K_S^0 \rightarrow \mu^+ \mu^-)_{D,D'} = \frac{\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)_{\text{eff}}(D) - \mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)_{\text{eff}}(D')}{\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)_{\text{eff}}(D) + \mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)_{\text{eff}}(D')}$$

D, D': initial  $K^0 - \bar{K}^0$  asymmetry, using  $K^-$  and  $K^+$  tagging

# Observables

- MSSM contribution to  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) \propto [(\delta_d^{LL(RR)})_{12} \mu \tan^3 \beta M_3 / M_A^2]^2$

Observable	Constraint
$\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)^{\text{EXP/SM}}$	unconstrained
$\mathcal{B}(K_L^0 \rightarrow \mu^+ \mu^-)^{\text{EXP/SM}}$ (*)	$1.00 \pm 0.12$ (+) [PRL 119 (2017) 201802], $0.84 \pm 0.16$ (-) [hep-ph/9411439]
$\Delta M_K^{\text{EXP/SM}}$	$1 \pm 1$
$\varepsilon_K^{\text{EXP/SM}}$	$1.05 \pm 0.10$ [1710.06614], [JHEP 04 (2018) 019]
$\Delta(\varepsilon'_K / \varepsilon_K)^{\text{EXP-SM}}$	$[15.5 \pm 2.3(\text{EXP}) \pm 5.07(\text{TH})] \times 10^{-4}$
$\mathcal{B}(B^+ \rightarrow \tau^+ \nu_\tau)^{\text{EXP/SM}}$	$0.91 \pm 0.22$ [PDG]
$\mathcal{B}(K^+ \rightarrow \mu^+ \nu_\mu)^{\text{EXP/SM}}$	$1.0004 \pm 0.0095$ [PDG]
$\Delta C_7$	$-0.02 \pm 0.02$ [JHEP 6 (2016) 92]
$\tan \beta: M_A$ plane	ATLAS limits for hMSSM scenario [1709.07242]
LSP	Lightest neutralino
$B_G$	$1 \pm 3(\text{TH})$ [NPB 566], [NPB 578]

- (\*) due to unknown sign in long-distance S-wave contribution
- Consider degenerate Higgs masses:  $M_H \approx M_A$

- **Large  $\tan\beta$  regime**

$\varepsilon'_K/\varepsilon_K$

$\varepsilon_K$

- Hadronic matrix elements from lattice

e.g. [PRL 115 (2015) 212001]

$$\left(\frac{\varepsilon'_K}{\varepsilon_K}\right)^{\text{SUSY}} \simeq \left(\frac{\varepsilon'_K}{\varepsilon_K}\right)_{\text{box+pen}} + \left(\frac{\varepsilon'_K}{\varepsilon_K}\right)_{\text{chromo}}$$

[JHEP 12 (2016) 078]

- Controversial SM prediction for  $\varepsilon_K$
- $(\varepsilon_K^{\text{SM}})_{\text{SD}} \propto |V_{cb}|^4$   
[PRD 92 (2015) 034510]
- 4.1  $\sigma$  discrepancy between measured value using inclusive and exclusive decays  
[1710.06614], [HFLAV]
- Require  $\text{Re}(\varepsilon_K) > 0$   
[JHEP 12 (2006) 011]

# Parameter scan

- Ipanema- $\beta$  framework [arXiv:1706.01420], GPU GeForce GTX 1080
- Flat scans + scans based on genetic algorithms  
[IEEE Transactions on Evolutionary Computation 10 (2006) 646–657]

## Scenarios

**Scenario A** : universal gaugino masses + no DM relic density constraint

- Mostly Bino-like LSP, with some points with Higgsino LSP

**Scenario B** : universal gaugino mass + relic density function of the LSP mass:

$m_{\chi_1^0} \approx 1$  TeV e.g. [EPJ C78 (2018) 158] fulfills the measured value [Astronomy & Astrophysics 594 no. A1]

- motivated by scenarios with Higgsino LSP

**Scenario C** : relic density function of the LSP mass:  $m_{\chi_1^0} \approx 3$  TeV ([EPJ C77 (2017) 268], [PLB 646 (2007) 34-38])

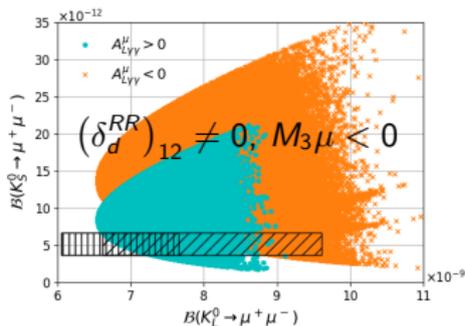
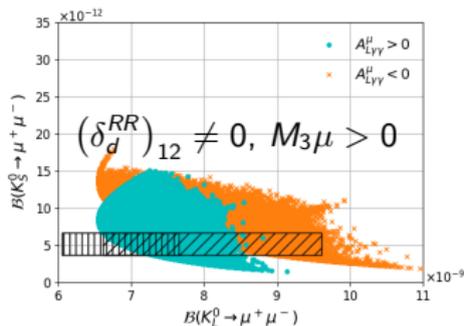
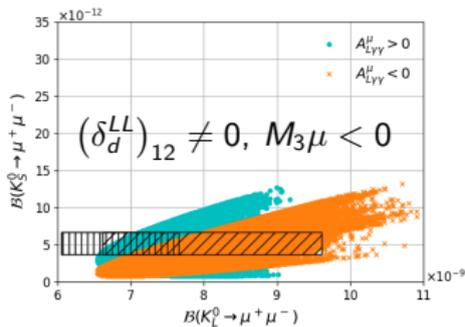
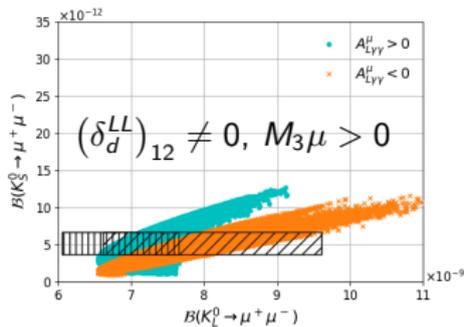
- motivated by scenarios with Wino LSP (mAMSB, pMSSM)
- nearly degenerate lightest neutralino and lightest chargino

# Parameter scan

Parameter	Scenario A	Scenario B	Scenario C
$\tilde{m}_Q$	[2, 10]	[2, 10]	[4, 10]
$\tilde{m}_Q^2 / \tilde{m}_d^2$	[0.25, 4]	[0.25, 4]	[0.25, 4]
$M_3$	[2, 10]	[4.5, 15]	[4, 15]
$\tan \beta$	[10, 50]	[10, 50]	[10, 50]
$M_A$	[1, 2]	[1, 2]	[1, 2]
$ \mu $	[1, 10]	1	[5, 20]
$M_1$	$\frac{\alpha_1(\mu^{SUSY})}{\alpha_3(\mu^{SUSY})} M_3$	$\frac{\alpha_1(\mu^{SUSY})}{\alpha_3(\mu^{SUSY})} M_3$	5
$M_2$	$\frac{\alpha_2(\mu^{SUSY})}{\alpha_3(\mu^{SUSY})} M_3$	$\frac{\alpha_2(\mu^{SUSY})}{\alpha_3(\mu^{SUSY})} M_3$	3
$B_G$	[-2, 4]	[-2, 4]	[-2, 4]
$\text{Re} \left[ (\delta_d^{LL(RR)})_{12} \right]$	[-0.2, 0.2]	[-0.2, 0.2]	[-0.2, 0.2]
$\text{Im} \left[ (\delta_d^{LL(RR)})_{12} \right]$	[-0.2, 0.2]	[-0.2, 0.2]	[-0.2, 0.2]

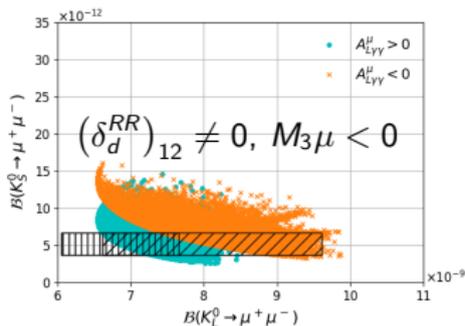
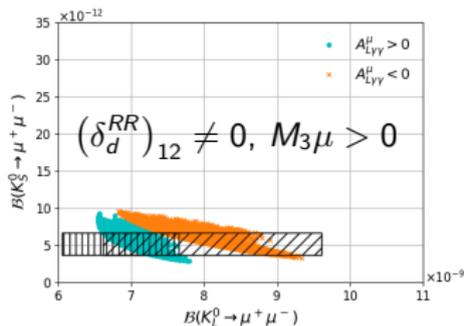
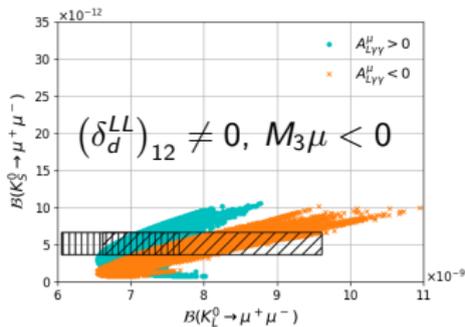
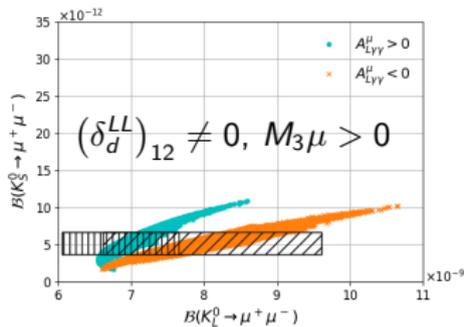
- Trilinear couplings and the MIs other than  $(\delta_d^{LL(RR)})_{12}$  and  $(\delta_u^{LL})_{12}$  set to **zero**
- $\tilde{m}_Q = \tilde{m}_u \neq \tilde{m}_d$

# Results: floating $\left(\delta_d^{LL(RR)}\right)_{12}$ separately



- A & C:** 95 % CL allowed regions for  $B(K_S^0 \rightarrow \mu^+ \mu^-)$ :  
 $[0.78, 14] \times 10^{-12}$  (LL),  $[1.5, 35] \times 10^{-12}$  (RR)

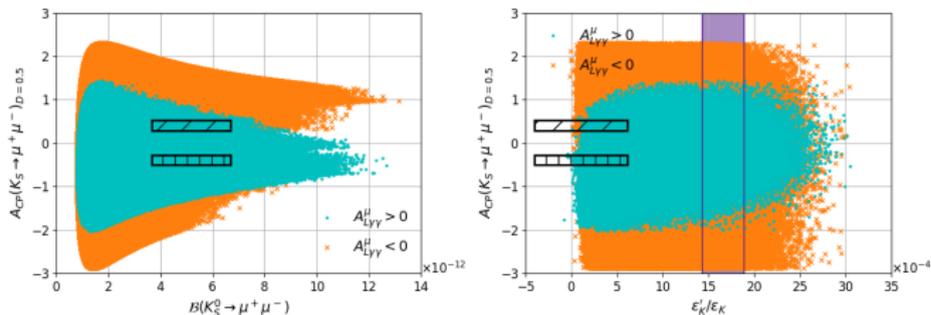
# Results: floating $\left(\delta_d^{LL(RR)}\right)_{12}$ separately



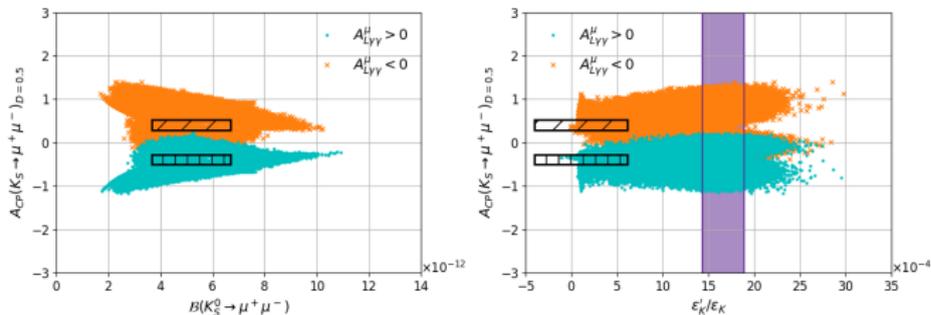
- **B**: smaller deviation of  $B(K_S^0 \rightarrow \mu^+ \mu^-)$  from the SM ( $C_{S,P} \propto \mu$ ,  $\mu$  is small relative to squark and gluino masses)

# Results: floating $\left(\delta_d^{LL(RR)}\right)_{12}$ separately

Scenario A,  $(\delta_d^{LL})_{12} \neq 0$  and  $(M_3 \cdot \mu) < 0$ ,  $D = -D' = 0.5$

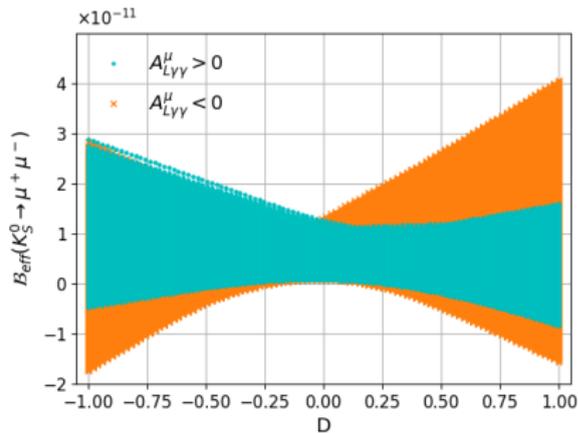
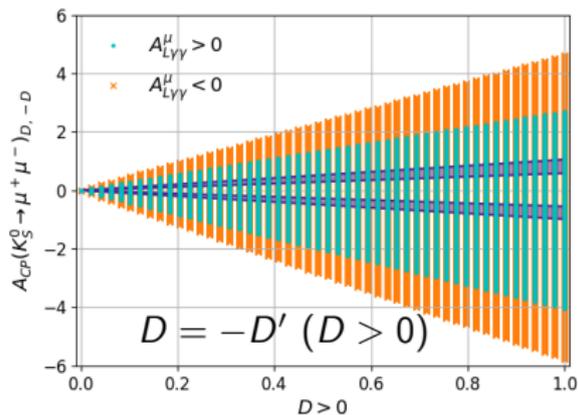


Scenario B,  $(\delta_d^{LL})_{12} \neq 0$  and  $(M_3 \cdot \mu) > 0$ ,  $D = -D' = 0.5$



# Results: floating $\left(\delta_d^{LL(RR)}\right)_{12}$ separately

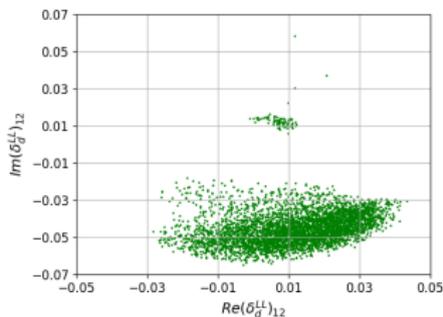
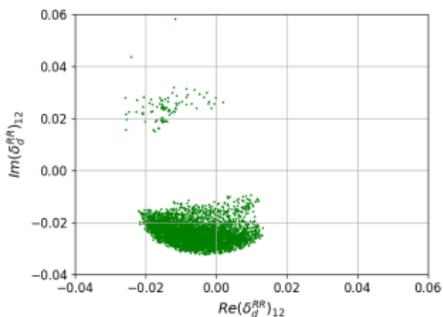
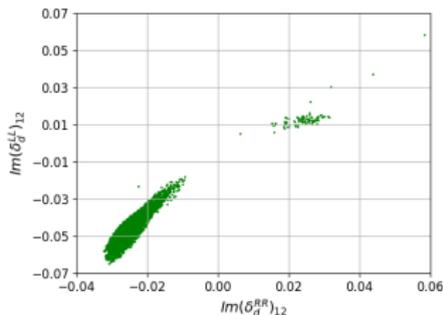
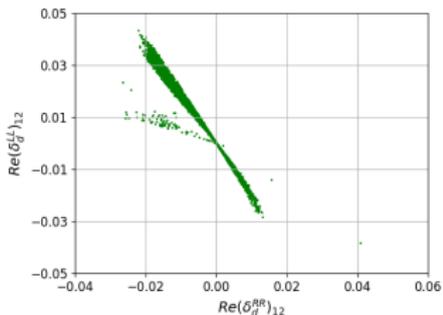
Scenario A,  $(\delta_d^{LL})_{12} \neq 0$  and  $(M_3 \cdot \mu) < 0$



- $CP$  asymmetries can be up to  $\approx 6$  (at  $D = 1$ )  $\Rightarrow 8 \times \text{SM}$
- Largest effects in left-handed scenarios

# Results: floating $LL$ and $RR$ MIs simultaneously

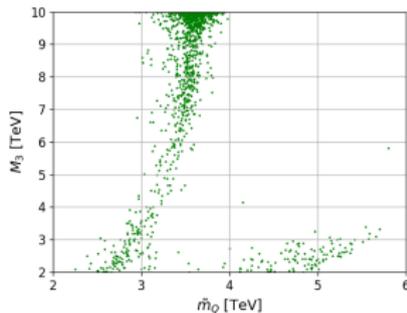
Scenario C  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) > 2 \times 10^{-10}$



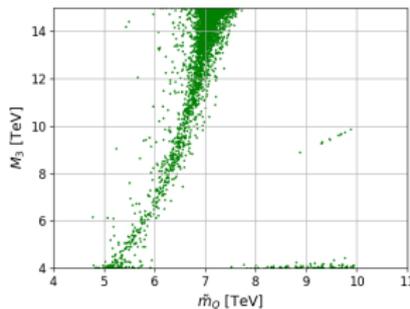
- Scenario A shows similar pattern

# Results: floating $LL$ and $RR$ MIs simultaneously

Scenario A



Scenario C



- $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) > 2 \times 10^{-10}$
- Region where  $|(\delta_d^{LL})_{12}| \approx 2|(\delta_d^{RR})_{12}| \sim 0.03$ ,  
 $\arg [(\delta_d^{LL})_{12}] \approx -\arg [(\delta_d^{RR})_{12}] + \pi$  particularly favorable
- Higher chances of finding regions with  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) > 10^{-10}$  using  $\epsilon_K^{\text{EXP/SM}} = 1.41 \pm 0.16(\text{TH})$  (from exclusive decays)

# Conclusions

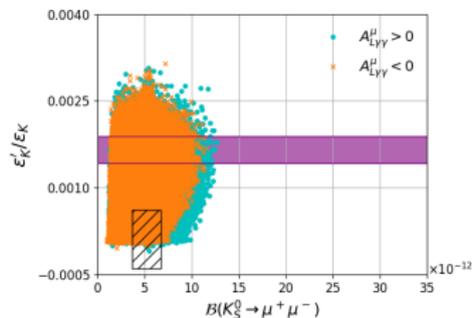
- $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)$  can surpass the SM contributions by a factor of 7 within the MSSM (**large  $\tan \beta$  regime**)
  - even for large SUSY masses
  - with no conflict with experimental (LHCb) data
- Fine-tuned regions can bring  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-)$  above the  $10^{-10}$  level
  - largest deviations from SM at  $|(\delta_d^{LL})_{12}| \approx 2|(\delta_d^{RR})_{12}| \sim 0.03$  and  $\arg [(\delta_d^{LL})_{12}] \approx -\arg [(\delta_d^{RR})_{12}] + \pi$  for large squark and gluino masses
- $CP$  asymmetry of  $K^0 \rightarrow \mu^+ \mu^-$  can be significantly modified by MSSM contributions
- If  $M_H \neq M_A$  (small  $\tan \beta$ )  $\Rightarrow$  looser constraints
  - regions  $\mathcal{B}(K_S^0 \rightarrow \mu^+ \mu^-) > 10^{-10}$  for mass differences of  $\mathcal{O}(33\%)$  or larger without fine-tuning the MIs

# Thanks for your attention!

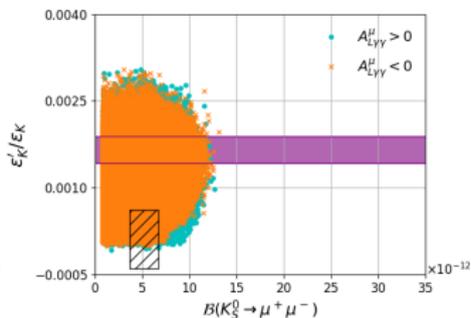
# BACKUP

# Scenario A

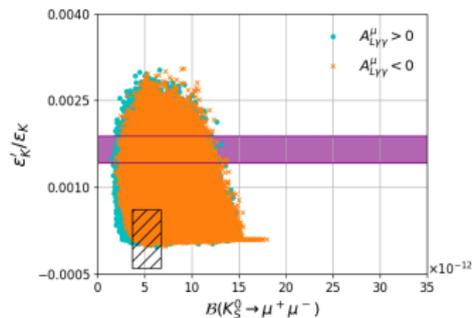
$$(\delta_d^{LL})_{12} \neq 0, M_3\mu > 0$$



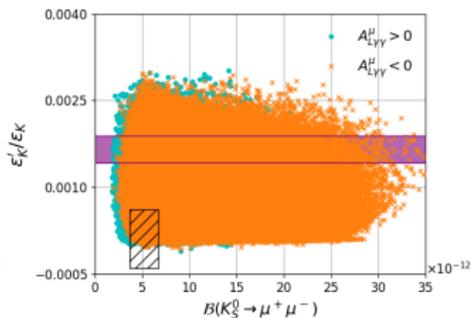
$$(\delta_d^{LL})_{12} \neq 0, M_3\mu < 0$$



$$(\delta_d^{RR})_{12} \neq 0, M_3\mu > 0$$

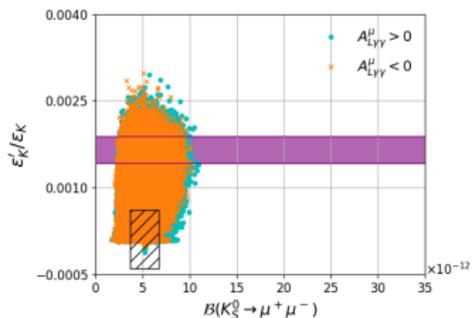


$$(\delta_d^{RR})_{12} \neq 0, M_3\mu < 0$$

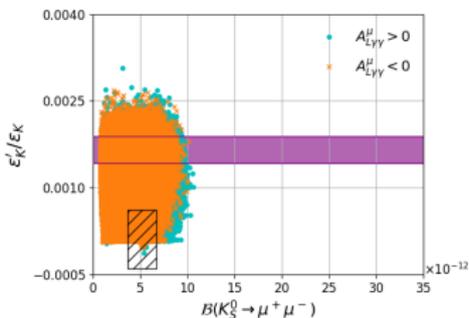


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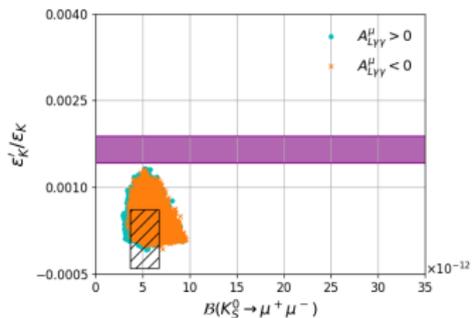
$$(\delta_d^{LL})_{12} \neq 0, M_3\mu > 0$$



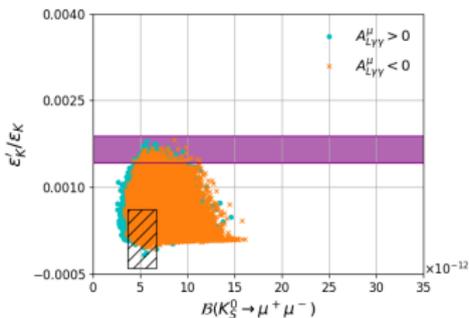
$$(\delta_d^{LL})_{12} \neq 0, M_3\mu < 0$$



$$(\delta_d^{RR})_{12} \neq 0, M_3\mu > 0$$

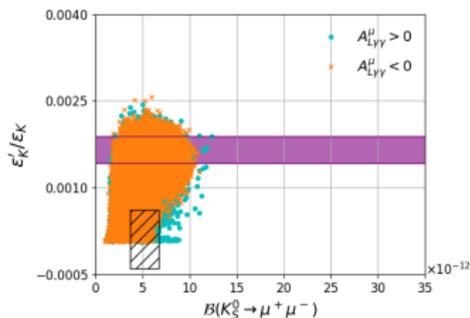


$$(\delta_d^{RR})_{12} \neq 0, M_3\mu < 0$$

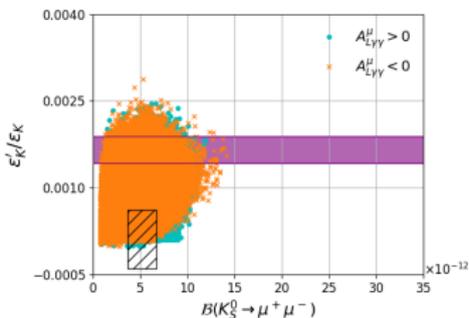


# Scenario C

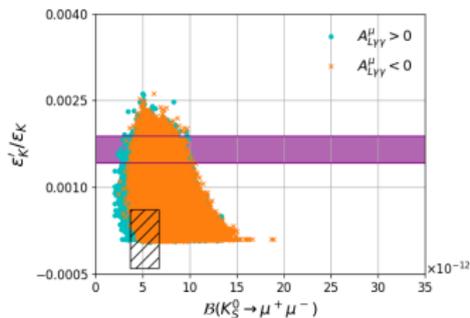
$$(\delta_d^{LL})_{12} \neq 0, M_3\mu > 0$$



$$(\delta_d^{LL})_{12} \neq 0, M_3\mu < 0$$



$$(\delta_d^{RR})_{12} \neq 0, M_3\mu > 0$$



$$(\delta_d^{RR})_{12} \neq 0, M_3\mu < 0$$

