



### **Directional WIMP detection R&D**

Gianluca Cavoto Sapienza Univ Roma and INFN PIRE Workshop Feb 5th and 6th 2017

### Extending the reach





- If ton-scale dark matter detectors do not see anything, need to explore lower cross section (the solar-atm neutrino will be a background).
   Use <u>direction</u> of WIMPs to discriminate them
- If they do see something, study WIMP's wind
- Use smaller mass target (carbon, electrons) to go to sub-GeV masses



### TPC concept for nuclear recoil





A nuclear recoil!



example high energy F recoil in optical TPC (D. Loomba et al.)

DRIFT detector (low pressure TPC with anode wire readout)



- Various activities financed by INFN and EU (new detectors concepts, new target concepts, ...)
  - Time Projection Chamber readout with GEM
    - Study negative ion drift (new gas mixture based on SF6, charge readout - NITEC)
    - Optical readout of large GEM surfaces CYGNUS-RD (large surface detectors)
    - Carbon nanotubes as anisotropic target for WIMPs (demonstration of ion channeling) - DCANT

CYGNUS int'l proto-collaboration to create eventually a multi-ton multi-site observatory

Working on a proposal for measurement of neutrons flux at INFN Gran Sasso underground labs



### Sapienza and INFN



 Strong link between Sapienza Physics departments, INFN Roma division and INFN Frascati labs.





http://www.phys.uniroma1.it/fisica/en-welcome http://www.roma1.infn.it/en/index.html http://w3.lnf.infn.it/?lang=en

Access to Frascati Beam Test Facility (50-500 MeV electron or positron beam from DAFNE LINAC <u>http://www.lnf.infn.it/acceleratori/btf/</u>)





### People involved

- People involved:
  - Sapienza and INFN Roma: C.Antochi (st.), GC, D.Pinci, E.Di Marco, M.Marafini, A.D.Polosa, F.Renga, C.Voena
  - INFN LNF (Frascati): E.Baracchini, G.Mazzitelli, F.Murtas, A.Tomassini.
- Sapienza and INFN have a strong link for Ph.D. student support
- Student academic training provided by Sapienza Ph.D. school(s)
- Access to laboratories granted by INFN





**DCANT** INFN CSN5

2016-2017

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Section CYGNUS-RD INFN CSN5

2017-2018

Friple thin GEM amplification (10 x 10 cm<sup>2</sup>) + CMOS optical readout

Dark Matter is very high in INFN priority list Longer term support for investments might come from INFN if current R&D successful



- In the following slides presented at two recent conferences
  - **E.Baracchini** at CAASTRO-CoEPP Joint Workshop, (Univ. Melbourne, Australia, Jan 2016)
  - GC at IDM 2016 (Univ Sheffield, UK Jul 2017)



## Thin GEM Amplification



- 🗣 Particle conversion, charge amplification and signal induction zones are physically separated
- Large dynamic range: from 1 to 10 particle/ $cm^2$  /s
- <u>Gain up to >  $10^4$ </u>
- High stability/granularity



- external field drift gap Vdrift **GEM** field gas ampli fication collection gap Micro pattern gas detector
- Thin holes are etched in a metallised kapton foil and a potential is placed across it
- GEM Very large electric field around the holes (40 kV/cm) which creates a localised electron Pads avalanche



ion blocking

2.+3.GEM

for gas amplificatior

Two-Track-Resolution: ~mm<sup>3</sup> Q Signal distribution on the pads



## NITEC

### a Negative Ion Time Expansion Chamber for directional Dark Matter searches

10

## NITEC detector



**Carbon Nanotubes** (see later)



<u>New field (age</u>: rings support structure (in black in the picture) manufactured with 3D printer

## <u>GEMPix</u>

Developed by LNF in collaboration with CERN

TimePix

triple GEM



Triple GEM detector with HV filters and connector

Triple GEM

Quad Timepix ASIC

Quad Timepix ASIC board with naked devices (i.e. no silicon)

A dedicated, very

stable GEM HV up to

700 V per GEM fully

developed at LNF









pixel size 55 x 55 um Quad Timepix (512 x 512 pixels) = 4 Timepix chips 2.8 x 2.8 cm<sup>2</sup>

## TimePix



TimePix is a pixelated silicon detector developed by MediPix2 collaboration
 We use a 2x2 array for a total of 512x512 pixel of 55 um side WITHOUT silicon sensors
 Processing electronics, including preamplifiers, discriminator threshold and pseudo-random counter fit inside the footprint of the overlying semiconductor pixel.

Can be operated in counting TOA, TOA and TOT mode but also TOA/TOT MIXED mode



Timepix clock can run from <1 MHz up to 100 MHz</p>
Timepix counter depth is 11810, SUITED FOR BOTH ELECTRONS and NEGATIVE IONS DRIFT

## NITEC activities 2015-2016



## A NITEC event



## Measurement @ Beam Test Facility



E. Baracchini - TPC GEMs R&D for directional Dark Matter searches - CAASTRO-CoEPP Joint Workshop 2017, Melbourne

Time measurements (TOA)



Pure SF<sub>6</sub> at 75 Torr, 100 Torr, 150 Torr

GEM gain 1140 V 1240 V 1440 V



Apr 2016

He:CF<sub>4</sub>:SF<sub>6</sub> 60:40:120 Torr, 360:240:10 Torr

GEM gain 1460 V 1640 V

## We measured the Time Of Arrival for 5 different drift distances @ 250, 530, 640, 750 and 860 V/cm for each configuration

(less points in Apr 2016 data)

## SF<sub>6</sub> @ 150 Torr



Dec 2016



Global quantities analysis shown in this talk. On going work on single track analysis.

## SF<sub>6</sub> @ 100 Torr TOA analysis

#### 750 V/cm



### He:CF<sub>4</sub>:SF<sub>6</sub> 360:240:10 Torr TOA analysis

#### 860 V/cm

(9200 -volume)\*0 5



1.4

1.2

0.8

0.6

200

**p**0

p1

E. Baracchini - TPC GEMs R&D for directional Dark Matter searches - CAASTRO-CoEPP Joint Workshop 2017, Melbourne

(9200.-volume)\*0.5

## Drift Velocity Measurements



## Mobility Measurements UPP



### NITEC gain (TOT) measurements in pure SF<sub>6</sub>



### NITEC gain measurement in pure SF<sub>6</sub>



E. Baracchini - NITEC: a Negative Ion Time Expansion Chamber for very rare events searches - IDM 2016, Sheffield



## DCANT Carbon Nanotube for for Dark Matter directional searches



### Outline

- WIMP directional searches
- Carbon Nanotubes (CNT)
  - Aligned CNT arrays as anisotropic target
  - Ion channeling inside a CNT
  - Ion trapping in the CNT array



### C ion channeling in CNT

 Use C ion in a Time Projection Chamber (*Triple-GEM TPC*) to demonstrate channeling is active

L.M. Capparelli, GC, D. Mazzilli, A.D. Polosa, *Phys.Dark Univ.* 9-10 (2015) 24-30, Corrigendum: Phys.Dark Univ. 11 (2016) 79-80 (<u>http://arxiv.org/abs/1412.8213</u>)

GC, E.N.M. Cirillo, F. Cocina, J. Ferretti, A.D. Polosa, **Eur.Phys.J. C** 76 (2016) no.6, 349 (http://arxiv.org/abs/1602.03216)





### Anisotropy in scattering rates

 Modelling based on elastic scattering of WIMP on *C* ions σ<sub>xp</sub> ~ 10<sup>-4</sup> pb





### Carbon nanotubes







#### M.G.Betti, C.Mariani Sapienza CNT with scattering electron microscope Univ. Roma



#### collaboration University of Mons, Belgium



#### length: 100 $\mu$ m (can be increased) ext. diameter: (20 ± 4) nm aspect ratio: $5x10^4$

#### commercial



#### length: 75 $\mu m$ ext. diameter: (13 ± 4) nm aspect ratio: 0.6 x10<sup>4</sup>





"target" mass on the CNT surface

### Scattering on a carbon nanotube

- CNT are "empty"
  - no electrons along the carbon ion path
- Large aspect ratio: ~10 nm diameter vs. ~100  $\mu m$  height











- 6+C ion scattered off E<sub>1</sub> = the CNT are  $= T \theta^{2} + U(R - x, \varphi)$ channeled in the CNT
- Little effect of electrons on **CNT** surface





~100 eV





#### Aligned and oriented CNT "brush"

- Recoiling C ions are emerging from CNTs with different rates depending on CNTs orientation.
- When C ions are not channeled they are absorbed within the brush
- Effect of rechanneling or inter-CNT trapping
   NOT included HERE





### Infinite horizon billiard





### Average **time** to **exit** the lattice from its **sides**

$$\langle \tau_{out} \rangle = \frac{L^2}{7} \tau_R$$
 with  $\tau_R = \frac{\pi (a^2 - \pi \rho^2)}{4v_1 \delta}$ 

Small  $\delta \rightarrow$  Machta-Zwanzig regime [1]J. Machta and R. Zwanzig, Phys. Rev. Lett. 50, 1959 (1983).

- CNT brush as an array of cylindrical obstacles (LxL): trapping within the inter-CNT space
- MC simulation of 2D motion
   (C ions below energy barrier)

Ratio of our simulation to a semi-analytical result







- C ion can leave the array at its top (*desired!*) or from the lateral sides (*avoid!*)
  - Fraction of particle leaving from the sides versus *side length L*

Lateral losses are negligible for realistic CNT brushes (L ~ 10<sup>5</sup>)







### C ion moving within the array

- Simulation including energy losses and scattering on the CNT walls (*C ions can penetrate CNT*)
  - Initial kinematics according to WIMP-C scattering



Trapping efficiency much larger than single CNT channeling





#### Detecting C ion escaping CNTs

- Use aligned CNT as target mass (~few g/cm3 density possible)
- Aligned CNTs as an anisotropic medium: scattered *C* ions are escaping from the top of the array when emitted almost parallel to CNT axes.
- Detect the channeled C ion in a very thin (low pressure) gas chamber
- Escaping C ion energy, C ion range in gas and direction measurements should be possible

Demonstrating that a 1-100 KeV *C* ion is effectively channeled in CNT and then detectable



### Scheme for detection of C ion







### Channeling of an ion





Demonstrate ~10-100 KeV C ions are trapped. Trapping has a larger effective θ<sub>c</sub> ~ 35 deg





- Use electron beam at LNF BTF to "extract" carbon ions from CNT
  - One carbon ion elastically scattered by a 500 MeV electron
    - PRO: trigger on scattered electron at well defined angle: beam clearly visible
    - CON: electron beam can induce a sizeable background into TPC





### Experiment at BTF: channeling





erc

### Experiment at BTF







## NITEC tests with carbon nanotube

#### **Carbon Nanotubes**



## Beam on the side of nanotubes at various heights to study modification of the drift field

## NITEC tests with carbon nanotube



## NITEC with carbon nanotubes





### We observe a consistent modification of the drift field due to the introduction of nanotubes structure AND support

- Support, kapton scotch and nanotube get polarized
- ANSYS simulation confirms observed results
- On going work to develop suitable support and substrate





## **CYGNUS-RD** Optical readout for a Negative Ion Time Projection Chamber

## CYGNUS-RD Detector







### CYGNUS-RD events (with electron drift)











## CYGNUS-RD potentialities "

#### May 2016, electron drift





~ 1000 photons/track mm, ~70 um track residuals, cluster structures visible

2) Ionization density (can be used to extrapolate track direction/sense)



## Light vs Distance Study

### How far can we go and still see light? i.e. how large area can we cover with one CMOS camera?



50 cm



## Light vs Lens Aperture



@ 20 cm distance 50 um pixels equivalent

## Light vs Distance



#### Both light vs lens aperture and light vs distance indicate isotropic light production

## Example of cosmic track @ 60 cm



## CYGNUS-RD @ BTF







## Conclusions & Outlooks

#### 🖗 NITEC

- Innovative SF<sub>6</sub> based negative ion gas mixtures tested and operated nearly atmospheric pressure (610 Torr) with triple thin GEMs at electron beam line
- Perform gain study with <sup>55</sup>Fe with the tested mixtures and explore new ones with higher He content (i.e. lower density)
- Perform neutron run at the ENEA facility (under discussion)

### **DCANT**

- Support and nanotubes observed to modify drift field
- On going study to develop suitable support and substrate

#### CYGNUS-RD

- Verified isotropic emission of photons by the GEMs, implying that collected light follows optical rules (one over distance squared)
- Cosmic tracks easy identified at 60 cm distance (i.e. 30 x 30 cm<sup>2</sup> area covered)
- New beam test beginning Feb with attempt of atmospheric negative ion operation following gas mixtures tested by NITEC
- Fest of PMT inside TPC volume to measure times (minority carriers)



# Backup

## Carbon Nanotubes





### GEMPix + NITPC: A Time Expansion Chamber

- At moderately high reduced fields, <u>anions drift at about 100 m/s, compared to about 104</u> <u>m/s for electron</u> in typical atmospheric pressure drift chamber conditions
- Excellent GEMPix time, energy and spatial resolutions
- Slow anions speed + typical separation of primary ionization clusters in gas + GEMPix performances = Time Expansion Chamber
  - Single ionization clusters drift slowly and could be individually observed with high precision: a relative time expansion between ionization process and signal readout has effectively been achieved
- Single ionization cluster observation can provide excellent dE/dx information, improved position resolution and possibility of superior energy resolution for low energy radiation

"The Time Expansion Chamber and single ionization measurement" (A.H.Walenta, IEEE TNS 26 73) "Suppressing drift chamber diffusion without magnetic field" (C.J.Martoff et al, NIM A 440)

### Also Graphene target !



#### http://arxiv.org/pdf/1606.08849.pdf

Directional Detection of Dark Matter with 2D Targets

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We propose two-dimensional materials as targets for direct detection of dark matter. Using graphene as an example, we focus on the case where dark matter scattering deposits sufficient energy on a valence-band electron to eject it from the target. We show that the sensitivity of graphene to dark matter of MeV to GeV mass can be comparable, for similar exposure and background levels, to that of semiconductor targets such as silicon and germanium. Moreover, a two-dimensional target is an excellent directional detector, as the ejected electron retains information about the angular dependence of the incident dark matter particle. This proposal can be implemented by the PTOLEMY experiment, presenting for the first time an opportunity for directional detection of sub-GeV dark matter.





### Functionalization





Figure 8. Different approaches to chemical modification of carbon nanotubes. (a) substitutional doped single-walled nanotubes (either during synthesis or by post-growth ion-implantation), (b,c) nanotube bundles intercalated with atoms or ions, (d,e) peapods: SWNTs filled with fullerenes (other endohedral fillings are possible), (f) fluorinated tubes, (g) covalently functionalised tubes and (h) functionalised nanotubes via *n*-stacking of the functionality and the tubes.

- CNT can be very efficiently **doped**
- Alkali metal can be bonded to CNT surface (Na,Cs,...) or F.
- WIMP can scatter on Na, Cs, ... and these ions can then be channeled

