

BSM theory overview: scalar extensions

Venus Keus

University of Helsinki & Helsinki Institute of Physics



Particle Physics Day, 20 October 2017

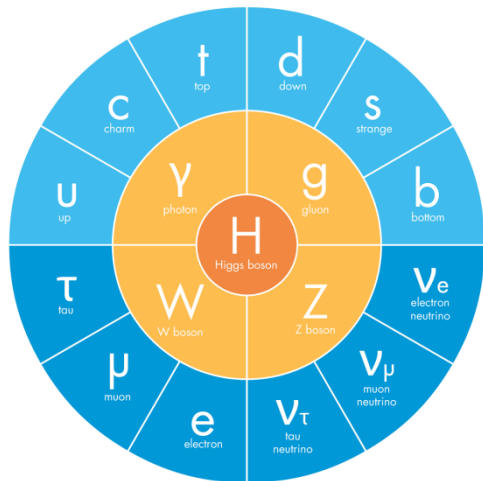
- 1 Introduction
- 2 Fermion mass hierarchy
- 3 Vacuum stability
- 4 Baryon asymmetry in the universe
- 5 Dark Matter
- 6 SM+singlet
- 7 SM+doublet
- 8 Summary

In praise of the Standard Model

Current formulation finalised
in the 70's predicted:

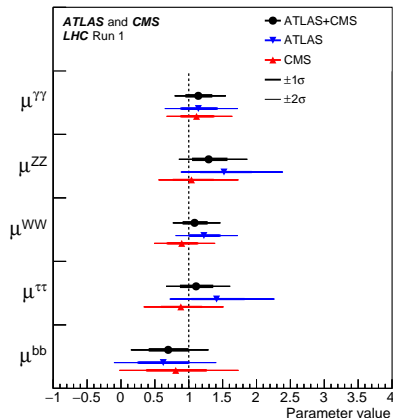
- the W & Z (1983)
- the top quark (1995)
- the tau neutrino (2000)
- “a” Higgs boson (2012)

FERMIONS (matter) | BOSONS (force carriers)
● Quarks ● Leptons | ● Gauge bosons ● Higgs boson



What's up at the LHC?

- Higgs looks SM-like
- No significant deviation from the SM
- No signs of new physics
- Is that all there is?



JHEP 08 (2016) 045

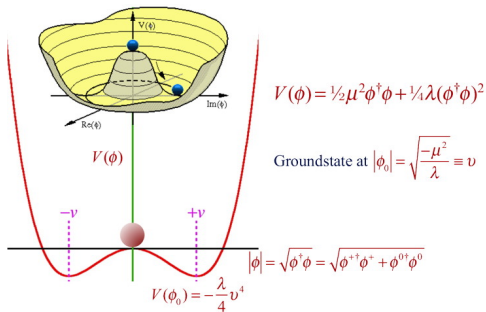
In criticism of the Standard Model

What is missing:

- An explanation for the Fermion mass hierarchy
- EW vacuum stability
- Baryon asymmetry in the universe
 - Strongly first order phase transition
 - Sufficient amount of CP-Violation
- No suitable candidate for Dark Matter

Let's focus on the scalar sector

$$\begin{aligned}
 \mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\
 & + i \bar{\Psi} \not{D} \Psi + \text{h.c.} \\
 & + \bar{\Psi}_i \gamma_{ij} \Psi_j \phi + \text{h.c.} \\
 & + \frac{1}{2} \partial_\mu \phi^2 - V(\phi)
 \end{aligned}$$

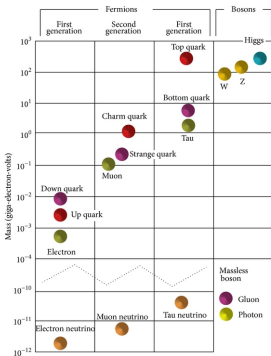


Fermion mass hierarchy

Fermion mass hierarchy

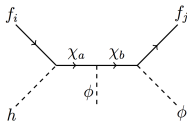
SM: no explanation for

- $m_t/m_e \approx 10^6$
- $m_t/m_\nu \approx 10^{11}$



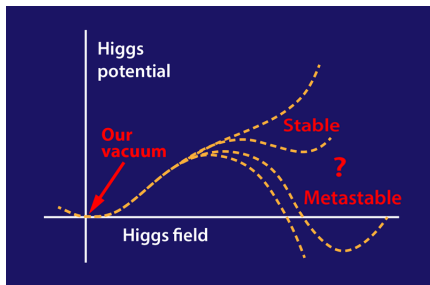
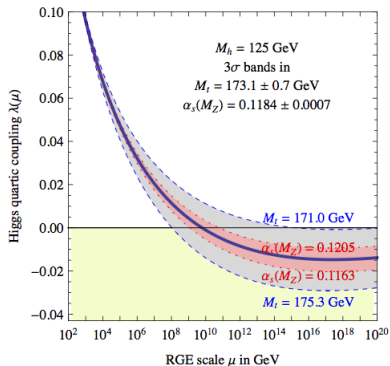
BSM: solutions

- SM + 2 scalar doublets
Weinberg's private Higgs model
- SM + singlet scalar + ...
Froggatt-Nielsen mechanism



EW vacuum stability

EW vacuum stability



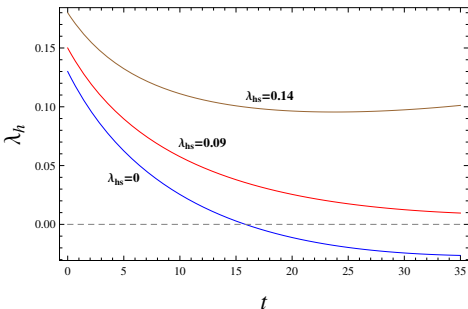
$$V = -\mu^2 \phi^\dagger \phi + \lambda (\phi^\dagger \phi)^2$$

[JHEP 1312, 089 (2013)]

Economic solution: extend the scalar sector

Simplest case: SM + real singlet scalar

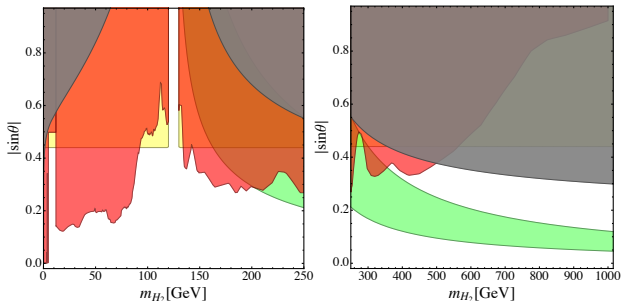
$$V = -\frac{\mu_h^2}{2}\phi^\dagger\phi - \frac{\mu_s^2}{2}S^2 + \frac{\lambda_h}{4}(\phi^\dagger\phi)^2 + \frac{\lambda_s}{4}S^4 + \frac{\lambda_{hs}}{4}(\phi^\dagger\phi)(S^2)$$



[Eur.Phys.J. C72 (2012) 2058]

How much mixing is allowed?

$$\text{If } \langle S \rangle \neq 0: \tan 2\theta = \lambda_{hs} v_h v_s / (\lambda_h v_h^2 - \lambda_s v_s^2)$$



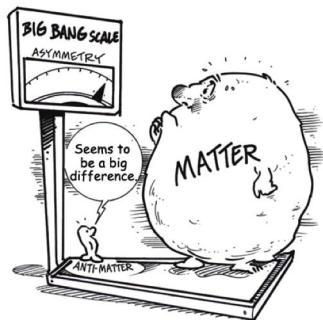
Excluded by direct searches (red), precision tests (gray), and H_1 couplings measurements (yellow).

[JHEP 05, 057 (2015)]

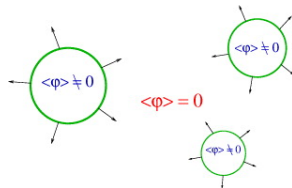
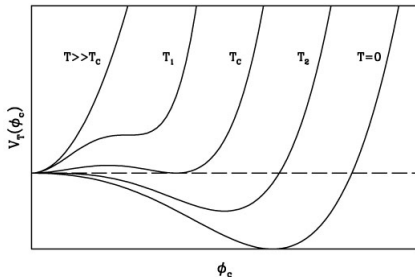
Baryon asymmetry in the Universe

Sakharov's conditions

- B-violation
- C & CP violation
- Departure from thermal equilibrium



Strongly 1st order phase transition



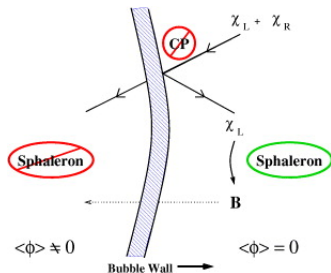
$$V_{\text{eff}}(\phi, T) = V_0(\phi) + V_1(\phi) + \Delta V_1^{(T)}(\phi, T)$$

SM scalar potential does not go through a first order phase transition.

⇒ Scalar extensions

[Phys. Rev. Lett. 77 (1996)]

C & CP violation



$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

$$V_{nb} \neq V_{nb}^*; V_{td} \neq V_{td}^* \Rightarrow \text{CPV}$$

Observation $\frac{N(B)}{N(\gamma)} \approx 10^{-9} \gg 10^{-20}$ provided by SM

New sources of CPV needed.

\Rightarrow Scalar extensions

Dark Matter

Dark Matter

How we know it exists:

- [Galaxy Clusters](#)
- Galactic Rotation Curves
- The CMB
- The Bullet Cluster
- Large-Scale Structure Formation

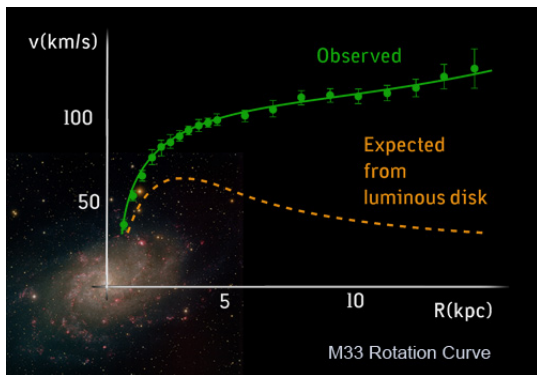


Fritz Zwicky in 1933 using the Virial theorem

Dark Matter

How we know it exists:

- Galaxy Clusters
- Galactic Rotation Curves
- The CMB
- The Bullet Cluster
- Large-Scale Structure Formation

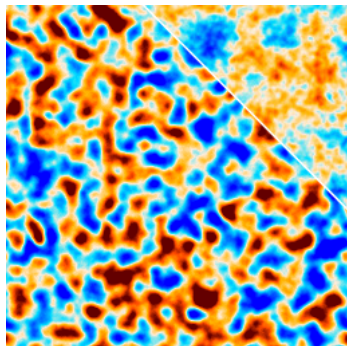


Vera Rubin & Kent Ford in 1960s

Dark Matter

How we know it exists:

- Galaxy Clusters
- Galactic Rotation Curves
- The CMB
- The Bullet Cluster
- Large-Scale Structure Formation



Planck CMB simulator

Dark Matter

How we know it exists:

- Galaxy Clusters
- Galactic Rotation Curves
- The CMB
- **The Bullet Cluster**
- Large-Scale Structure Formation

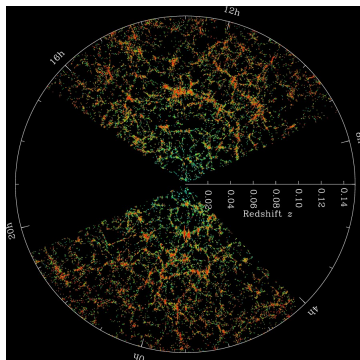


Two galaxies collide (2006)

Dark Matter

How we know it exists:

- Galaxy Clusters
- Galactic Rotation Curves
- The CMB
- The Bullet Cluster
- Large-Scale Structure Formation

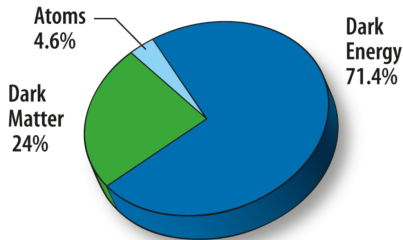


Sloan Digital Sky Survey (2013)

WIMP Dark Matter

Characteristics:

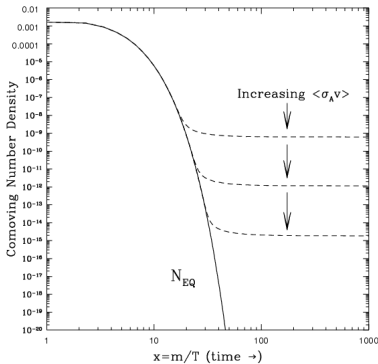
- Cold (non-relativistic at the onset of galaxy formation)
- Non-baryonic
- Neutral & weakly interacting
- Stable due to a discrete symmetry



$$\underbrace{\text{DM DM} \rightarrow \text{SM SM}}_{\text{pair annihilation}},$$

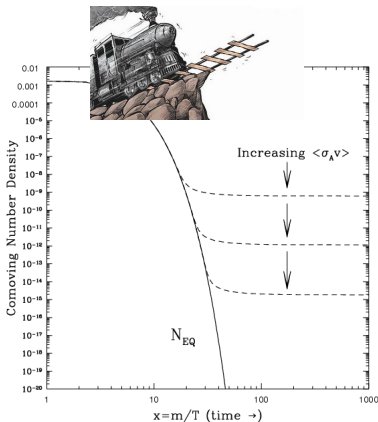
$$\underbrace{\text{DM} \not\rightarrow \text{SM}, \dots}_{\text{stable}}$$

WIMP Dark Matter freeze-out



Observed relic density: $\Omega_{DM} h^2 = 0.1199 \pm 0.0027$

WIMP Dark Matter freeze-out



Observed relic density: $\Omega_{DM} h^2 = 0.1199 \pm 0.0027$

BSMs to the rescue

Focusing on DM and CPV

SM + singlet scalar extension

DM, ϵ CPV

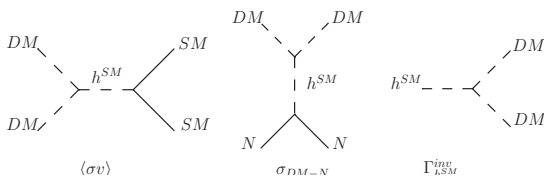
SM+RS: scalar DM

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{2}(\partial S)^2 + \frac{1}{2}m_{DM}^2 S^2 - \lambda_{DM} S^4 - \lambda_{hDM} \phi_{SM}^2 S^2$$

$$S \rightarrow -S, \quad \text{SM fields} \rightarrow \text{SM fields}$$

Higgs-portal interaction:

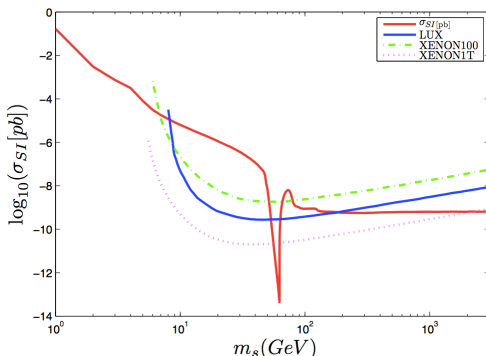
SM sector $\overset{\text{Higgs}}{\longleftrightarrow}$ DM sector



given by the same coupling

SM+RS: scalar DM

Relic density + direct detection constraints:

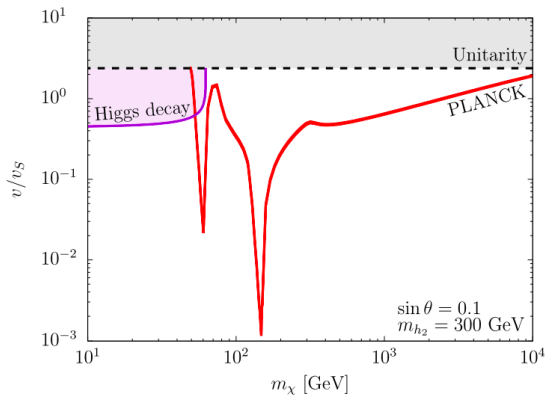
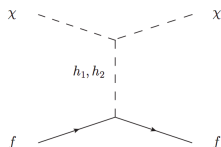


+ Higgs decays + SM vacuum stability + perturbativity constraints:

$$1.1 \text{ TeV} \leq m_{DM} \leq 2.0 \text{ TeV}$$

SM+CS: scalar DM

$S = (v_s + s + i\chi) \rightarrow Z_2$ -symmetry broken, CP conserves DM



[arXiv:1708.02253]

Further singlet extensions

SM+ 2 singlet scalars:

$S_1 \rightarrow -S_1$, $S_2 \rightarrow -S_2$, SM fields \rightarrow SM fields

- Introducing coannihilation channels
- DM: the lightest particle from the dark sector
- CPV: in the dark sector
- CPV not transferable to fermions \rightarrow EWBG

[Phys.Rev.D. 83 (2011)]

2 Higgs doublet model

DM, ~~CPV~~

~~DM~~, CPV

2HDM with CP-violation (DM)

The general scalar potential

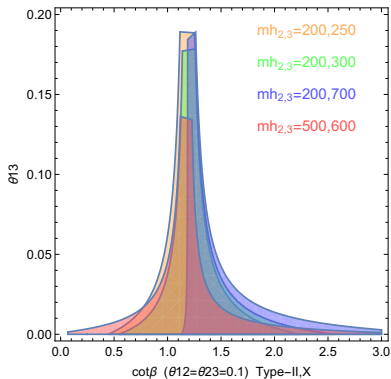
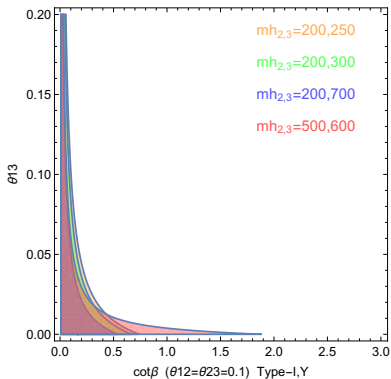
$$\begin{aligned}
 V = & -\mu_1^2(\phi_1^\dagger\phi_1) - \mu_2^2(\phi_2^\dagger\phi_2) - \left[\mu_3^2(\phi_1^\dagger\phi_2) + h.c. \right] \\
 & + \frac{1}{2}\lambda_1(\phi_1^\dagger\phi_1)^2 + \frac{1}{2}\lambda_2(\phi_2^\dagger\phi_2)^2 + \lambda_3(\phi_1^\dagger\phi_1)(\phi_2^\dagger\phi_2) + \lambda_4(\phi_1^\dagger\phi_2)(\phi_2^\dagger\phi_1) \\
 & + \left[\frac{1}{2}\lambda_5(\phi_1^\dagger\phi_2)^2 + \lambda_6(\phi_1^\dagger\phi_1)(\phi_1^\dagger\phi_2) + \lambda_7(\phi_2^\dagger\phi_2)(\phi_1^\dagger\phi_2) + h.c. \right].
 \end{aligned}$$

$$Z_2 \text{ symmetry} \Rightarrow \lambda_6 = \lambda_7 = 0$$

The doublets composition with $\tan\beta = v_2/v_1$

$$\phi_1 = \begin{pmatrix} \phi_1^+ \\ \frac{v_1 + h_1^0 + ia_1^0}{\sqrt{2}} \end{pmatrix}, \quad \phi_2 = \begin{pmatrix} \phi_2^+ \\ \frac{v_2 + h_2^0 + ia_2^0}{\sqrt{2}} \end{pmatrix}$$

How much CPV is allowed?



[work in progress]

The Inert Doublet Model (CPV)

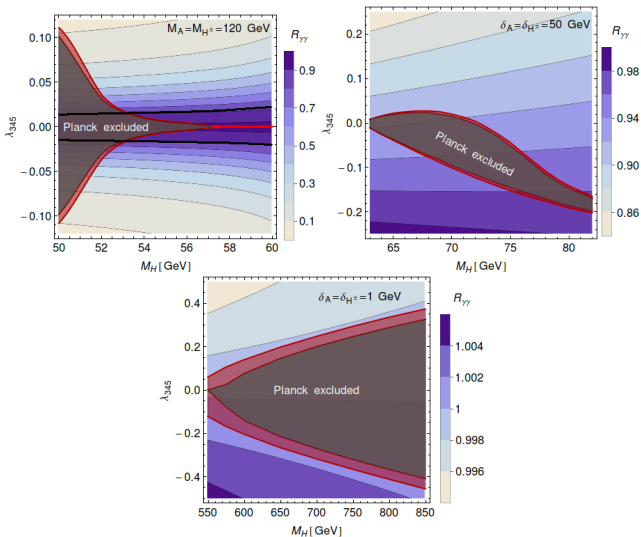
Scalar potential invariant under a Z_2 -transformation:

$$Z_2 : \quad \phi_1 \rightarrow \phi_1, \quad \phi_2 \rightarrow -\phi_2, \quad \text{SM fields} \rightarrow \text{SM fields}$$

$$\begin{aligned} V = & -\mu_1^2(\phi_1^\dagger\phi_1) - \mu_2^2(\phi_2^\dagger\phi_2) + \frac{1}{2}\lambda_1(\phi_1^\dagger\phi_1)^2 + \frac{1}{2}\lambda_2(\phi_2^\dagger\phi_2)^2 \\ & + \lambda_3(\phi_1^\dagger\phi_1)(\phi_2^\dagger\phi_2) + \lambda_4(\phi_1^\dagger\phi_2)(\phi_2^\dagger\phi_1) + \frac{1}{2}\lambda_5(\phi_1^\dagger\phi_2)^2 \end{aligned}$$

- All parameters are real \rightarrow no CP violation
- $\langle\phi_1\rangle = v$, $\langle\phi_2\rangle = 0$

All constraint taken into account



To explain:

DM, CPV, EWBG, Stability



beyond simple singlet/doublet extensions

Summary

Scalar extensions with or without a Z_2 symmetry:

- SM + scalar singlet(s)

- $\phi_{SM}, S \Rightarrow DM, CPV$
- $\phi_{SM}, S_1, S_2 \Rightarrow DM, CPV$

- 2HDM: SM + scalar doublet

- Type-I, Type-II, ...: $\phi_1, \phi_2 \Rightarrow CPV, DM$
- IDM - I(1+1)HDM: $\phi_1, \phi_2 \Rightarrow DM, CPV$

- 3HDM: SM + 2 scalar doublets

- Weinberg model: $\phi_1, \phi_2, \phi_3 \Rightarrow CPV, DM$
- I(1+2)HDM: $\phi_1, \phi_2, \phi_3 \Rightarrow DM, CPV$
- I(2+1)HDM: $\phi_1, \phi_2, \phi_3 \Rightarrow CPV, DM$

BACKUP SLIDES

Detailed summary

- Both DM and CPV from scalar sector → beyond 2HDM
- CP-Violation in I(1+2)HDM
 - IDM-like inert sector: CPC DM
 - CPV in the active sector: $\tilde{H}_1, \tilde{H}_2, \tilde{H}_3$
 - EWBG possible
- CP-Violation in I(2+1)HDM
 - SM-like active sector: $H_3 \equiv h^{SM}$
 - CPV in the inert sector: $H_{1,2}, A_{1,2} \rightarrow S_{1,2,3,4}$ CPV DM
 - EWBG possible

LHC bounds on CPV DM

Higgs invisible branching ratio and total decay

From ATLAS and CMS

$$\text{Br}(h \rightarrow \text{inv}) < 0.23 - 0.36$$

for $m_{i,j} < m_h/2$ if long lived

$$\text{BR}(h \rightarrow \text{inv}) = \frac{\sum_{i,j} \Gamma(h \rightarrow S_i S_j)}{\Gamma_h^{\text{SM}} + \sum_i \Gamma(h \rightarrow S_i S_j)}$$

The **total decay** signal strength

$$\mu_{\text{tot}} = \frac{\text{BR}(h \rightarrow \text{XX})}{\text{BR}(h_{\text{SM}} \rightarrow \text{XX})} = \frac{\Gamma_{\text{tot}}^{\text{SM}}(h)}{\Gamma_{\text{tot}}^{\text{SM}}(h) + \Gamma^{\text{inert}}(h)}$$

We use $\mu_{\text{tot}} = 1.17 \pm 0.17$ at 3σ level.

LHC signatures

$E_{miss}^T + e^+e^-$ cross section at the LHC

- Higgs-strahlung at **tree level**:

$$q\bar{q} \rightarrow Z \rightarrow S_1 S_{2,3,4} \rightarrow S_1 S_1 Z^* \rightarrow S_1 S_1 e^+ e^- \text{ with } \sigma \sim 10^{-2} \text{ pb}$$

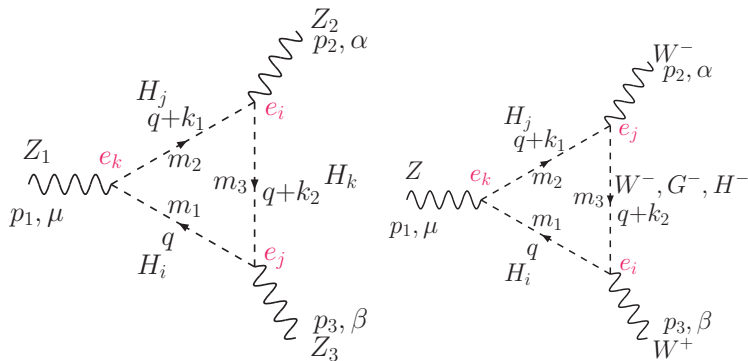
- Higgs-strahlung at **loop level**:

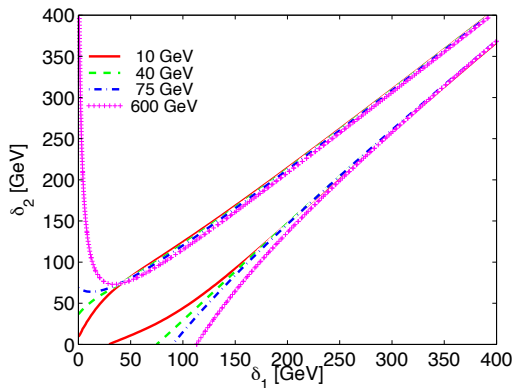
$$q\bar{q} \rightarrow Z \rightarrow S_1 S_{2,3,4} \rightarrow S_1 S_1 \gamma^* \rightarrow S_1 S_1 e^+ e^- \text{ with } \sigma \sim 10^{-3} \text{ pb}$$

- Gluon-fusion at **tree level**:

$$pp \rightarrow h \rightarrow S_1 S_{2,3,4} \rightarrow S_1 S_1 Z^* \rightarrow S_1 S_1 e^+ e^- \text{ with } \sigma \sim 10^{-5} \text{ pb}$$

ZZZ and ZWW vertices





$$\delta_1 = m_{H^\pm} - m_{DM} \text{ and } \delta_2 = m_A - m_{DM}$$